

## **Sound wave velocities of Fe<sub>5</sub>Si at high-pressure and high-temperature conditions: Implications to lunar and planetary cores**

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### **ABSTRACT**

Elastic properties of Fe alloys are critical in constraining the compositions of planetary bodies by comparing to the planetary observations. The sound wave velocities and density of an Fe<sub>5</sub>Si (9 wt% Si) alloy in body-centered cubic (bcc) structure were measured by combining an ultrasonic technique with synchrotron X-ray radiography at pressure ( $P$ ) and temperature ( $T$ ) conditions of 2.6–7.5 GPa and 300–1173 K, respectively. At room temperature, it is observed that adding Si to bcc-Fe increases the compressional wave velocity ( $v_p$ ) but decreases the shear wave velocity ( $v_s$ ). At high temperatures, we observed a pronounced effect of pressure on the  $v_s$ - $T$  relations in the Fe<sub>5</sub>Si alloy. The  $v_p$ -density ( $\rho$ ) relationship of the Fe<sub>5</sub>Si alloy is found to follow the Birch's law in the  $P$ - $T$  range of this study, whereas the  $v_s$ - $\rho$  relation exhibits complex behavior. Implications of these results to the lunar core and the Mercurian core are discussed. Our results imply that adding Si to a pure Fe lunar core would be invisible in terms of  $v_p$ , but exhibit a decreased  $v_s$ . Including Si in a sulfur-rich lunar core would display an increased  $v_p$  and a decreased  $\rho$ . Our density and sound wave velocity model provide lower and upper limit for a Si-bearing lunar core if 1–3 wt% Si content of enstatite chondrite is taken as compositional analog. A Si-rich (>9 wt%) Mercurian core model is derived to satisfy newly observed moment of inertia values by Messenger spacecraft.

**Keywords:** Fe-Si alloy, elastic wave velocity, lunar and planetary cores, high pressure, high temperature