

1 **Textural and chemical evolution of magnetite from the Paleozoic**
2 **Shuanglong Fe-Cu deposit: implications for tracing ore-forming**
3 **fluids**

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ABSTRACT

19 The Aqishan-Yamansu belt in Eastern Tianshan (NW China) hosts several
20 important Fe and Fe-Cu deposits, the origin of which is the subject of considerable
21 debate. The coexistence of various types of ore-forming fluids makes it difficult to
22 distinguish the genesis of the Fe-Cu deposits. We present detailed textural and
23 compositional data on magnetite from the Paleozoic Shuanglong Fe-Cu deposit to
24 constrain the formation of iron oxides and the evolution of the ore-forming fluids, to
25 define the genesis of the Fe-Cu ores.

26 Based on the mineral assemblages and crosscutting relationships of veins, two
27 mineralization stages were established, including the early Fe mineralization and late
28 Cu mineralization stage. Three types of magnetite, i.e., platy (MA), massive (MB),
29 and granular (MC) magnetite occur in the Fe mineralization. In backscattered electron
30 (BSE) images, an early hematite and transformational mushketovite phase (MA-I)
31 characterized by abundant porosity and inclusions, and two later generations
32 including an early dark (MA-II, MB-I, and MC-I) and later light magnetite (MA-III,
33 MB-II, and MC-II) were identified, which display oscillatory zoning. The MA-I has
34 extremely high W contents and mostly displays as micro- and invisible scheelite
35 inclusions, which were probably caused by the W expulsion during
36 mushketovitization. The texture and composition of magnetite suggest that the later
37 light magnetite formed via dissolution and reprecipitation of the precursor dark
38 magnetite, and the temperature and oxygen fugacity of fluids decreased over time.

39 Our study also shows the MB-II magnetite and coexisting chlorite display
40 synchronous oscillatory zoning, with the calculated temperature from 444 to 212 °C.
41 Such variations could indicate the incursion of external low temperature fluids with
42 high salinity, which can dissolve the primary dark magnetite. This study provides a
43 good example for using magnetite to trace the complex evolution and multiple sources
44 of ore-forming fluids.

45 Key words: Magnetite, Texture and chemistry, ore-forming fluid, Fe-Cu deposit,
46 Eastern Tianshan

47

INTRODUCTION

48 Magnetite is a common mineral in many types of ore deposits, including
49 Kiruna-type, BIF (banded iron formation), magmatic Fe-Ti oxide, Fe-skarn, IOCG
50 and porphyry deposits, and also in many igneous, metamorphic and sedimentary rocks
51 (Dare et al., 2014; Dupuis and Beaudoin, 2011; Hu et al., 2015; Hu et al., 2017;
52 Huang and Beaudoin, 2019; Nadoll et al., 2014; Nadoll et al., 2012; Wu et al., 2019;
53 Zhao et al., 2018a). It can accommodate a variety of trace elements into its inverse
54 spinel structure by substitution with Fe^{3+} and Fe^{2+} in tetrahedral or octahedral sites
55 (Nadoll et al., 2014), and its textures and chemical compositions vary in response to
56 different mineralizing systems and physicochemical conditions (Dare et al., 2014;
57 Dupuis and Beaudoin, 2011; Huang et al., 2019a; Nadoll et al., 2012). Previous
58 studies have developed a series of discriminant diagrams to fingerprint various deposit
59 types or ore-forming processes based on the compositional variety in magnetite
60 (Dupuis and Beaudoin 2011; Nadoll et al., 2012). However, recently studies have also
61 shown that the textural and chemical composition of magnetite can be significantly
62 modified or reequilibrated by hydrothermal fluid and supergene processes (Broughm et
63 al., 2017; Bain et al., 2020; Hu et al., 2015; Hu et al., 2014; Huang and Beaudoin, 2019;
64 Huang et al., 2018; Wen et al., 2017; Zhang et al., 2020b), for instance, mixing with
65 high-salinity, chlorine-rich fluids has been suggested as an important mechanism for
66 inducing coupled dissolution-reprecipitation (DR) reactions in magnetite from
67 hydrothermal deposits such as skarn deposits and iron oxide-copper-gold deposits (Hu

68 et al., 2015; Hu et al., 2014; Huang and Beaudoin, 2019; Liang et al., 2020; Yin et al.,
69 2017).

70 The Aqishan-Yamansu belt in the Chinese eastern Tianshan hosts many important
71 Fe-Cu deposits, including Heijianshan, Duotoushan, Shuanglong and Shaquanzi (Fig.
72 1c; Zhao et al., 2017). These Fe-Cu deposits have been described as skarns (Mao et al.,
73 2005), but the absence of some typical features of typical skarn deposit such as the
74 close relationships of orebodies with carbonate host rocks, and garnet-pyroxene skarn
75 alteration, indicating that they are not typical skarn deposits. Recent studies reveal
76 that these are possibly IOCG-like deposits according to the presence of hematite and
77 calcite-rich Cu mineralization as well as the less amounts of pyrite, and the fluid
78 inclusions and sulfur isotopes studies show involvement of seawater or basinal brine
79 (Jiang et al., 2018; Zhao et al., 2017; Zhang et al., 2018b). Thus, investigation of the
80 fluid evolution is very important for its ore genesis. As magnetite is a common mineral
81 in these Fe-Cu deposits, the study of its texture and composition can provide important
82 information about the ore-forming fluid and can be used to investigate the source of the
83 mineralization. The Shuanglong system is a typical Fe-Cu deposit in the
84 Aqishan-Yamansu belt, located in the east of the Bailingshan complex (Fig. 1c). The
85 existence of abundant magnetite in the Shuanglong deposit provides a good
86 opportunity to investigate the sources and evolution of ore-forming fluids during Fe
87 and Cu mineralization based on the textural and chemical characteristics of magnetite.

88 We present detailed textural and compositional data, using backscattered electron

89 (BSE) imaging combined with electron probe microanalyzer (EPMA) and laser
90 ablation-inductively coupled plasma-mass spectrometry (LA-ICP-MS) analyses of
91 magnetite from the Shuanglong Fe-Cu deposit, to characterize the factors controlling
92 the formation of the various iron oxides and the evolution of the corresponding
93 ore-forming fluids, to constrain the genesis of the regional Fe-Cu mineralization.

94 **GEOLOGIC SETTING**

95 **Regional geology**

96 The Central Asian Orogenic Belt (CAOB; Fig. 1a) is a complex collage of
97 ophiolite suites, magmatic arc remnants, accretionary wedges, and continental
98 fragments (Sengor, 1996; Sengor et al., 1993; Windley et al., 2007), which formed
99 during the evolution of the Paleo-Asian Ocean (PAO) through a long period from the
100 Neoproterozoic to Neopaleozoic (Han and Zhao, 2018; Sengor et al., 1993; Windley et
101 al., 2007; Xiao et al., 2013). The Eastern Tianshan, situated to the north of the Tarim
102 Basin in North Xinjiang (Fig. 1b), can be divided into the Dananhu-Tousuquan Island
103 Arc Belt, Kangguer Shear Zone, and Aqishan-Yamansu Belt (from north to south) by
104 the E-W-trending Kangguer and Yamansu faults (Fig. 1c).

105 The Dananhu-Tousuquan Island Arc Belt is situated to the north of the Kangguer
106 shear zone and contains Devonian to Carboniferous volcanic and plutonic rocks, and
107 hosts several porphyry copper deposits, including the Tuwu-Yandong deposit (Fig. 1c;
108 Mao et al., 2005). The Kangguer Shear Zone is an important regional structure that

109 separates the Junggar Block from the Tarim Block (Chen et al., 2019), and contains
110 ductility deformed Carboniferous volcanoclastic rocks that have been metamorphosed
111 to greenschist facies (Xiao et al., 2004) and Permian mafic-ultramafic intrusions (Qin
112 et al., 2011). The shear zone hosts Au deposits in the west (e.g., Shiyingtang and
113 Kangguer; Xiao et al., 2004) and Cu–Ni deposits in the east (e.g., Huangshan and
114 Huangshandong; Fig. 1c; Qin et al., 2011).

115 The Aqishan-Yamansu Belt (Fig. 1c) contains mainly Carboniferous volcanic,
116 volcanoclastic, and clastic rocks overlain by Permian clastic and volcanic rocks with
117 local carbonate interbeds (Mao et al., 2005). The Carboniferous rocks consist of the
118 Early Carboniferous Yamansu and Late Carboniferous Tugutubulak formations (Zhao
119 et al., 2018b). Late Carboniferous to Early Triassic felsic intrusions intruded the
120 Carboniferous Aqishan-Yamansu strata (Zhao et al., 2018b; Zhou et al., 2010), and are
121 associated with Fe- and Fe-Cu deposits, including Hongyuntan, Bailingshan,
122 Chilongfeng, Heijianshan, Yamansu, and Shaquanzi (Huang et al., 2018; Jiang et al.,
123 2018; Zhao et al., 2017). Previous studies have proposed that the Aqishan-Yamansu
124 belt represents an arc-related basin formed by southward subduction of the Kangguer
125 oceanic plate beneath the Central Tianshan block during the Early Carboniferous
126 (350–330 Ma), and the basin was then inverted during the Late Carboniferous (320–
127 305 Ma; Zhang et al., 2020b).

128 **Ore deposit geology**

129 The Shuanglong Fe-Cu deposit is located ~260 km south of Hami city, NW
130 China. The Shuanglong deposit is hosted in the Upper Carboniferous Tugutubulak
131 Formation which can be subdivided into three members: the lower member consisting
132 of basalt, volcanic breccia and volcanic tuff which is volcanic explosive facies, the
133 middle member which is volcanic overflowing facies of volcanic tuff and rhyolitic
134 ignimbrite, and the upper member which is volcano-sedimentary facies, comprising
135 mainly andesitic tuff with carbonate interlayers. These volcanic rocks and
136 volcanoclastic rocks all belong to the calc alkali series. Previous studies have shown
137 that the volcanic tuff in the Shuanglong deposit has LA-ICP-MS zircon U-Pb ages of
138 319.1 ± 1.9 Ma to 305.4 ± 1.6 Ma (Zhang et al., 2020b). The Tugutubulak Formation
139 at Shuanglong was intruded by the Bailingshan intrusive complex, including diorite
140 (317.1 ± 3.3 Ma; Zhang et al., 2020), monzogranite (311.0 ± 1.4 Ma; Zhang et al.,
141 2020b), granodiorite (307.5 ± 1.7 Ma; Zhao et al., 2019), and minor granite and
142 granite porphyry (306.9 ± 1.7 Ma; Fig. 2; Zhang et al., 2020b). Multiple faults have
143 been recognized in the Shuanglong ore district, and the ore bodies are associated with
144 the E- trending Shuanglong fault (Fig. 2).

145 Two ore zones have been defined in the Shuanglong district. The main Fe-Cu
146 orebody (L3) occurs in the northern zone, extending for approximately 120 m in
147 length with a width ranging from 9 to 27m (Fig. 2). The orebody was controlled by
148 the Shuanglong faults, and its dip direction is consistent with the fault dip and infilled

149 in the fault. The host rocks of the orebody are Carboniferous Tugutubulak Formation
150 volcanoclastic rocks and diorite intrusions. The ore minerals consist primarily of
151 magnetite, hematite, and sulfides including pyrite, pyrrhotite, and chalcopyrite. The
152 wall rock alteration includes garnet (andradite), K-feldspar, epidote, amphibole
153 (actinolite), quartz, calcite, and chlorite (pynochlorite) alteration. Three
154 hydrothermal alteration and mineralization stages can be identified at Shuanglong,
155 namely (I) Skarn alteration, (II) Fe mineralization, and (III) Cu mineralization (Fig.
156 3).

157 Stage I skarn alteration contains abundant garnet occurring as euhedral to
158 subhedral fine-grains in the magnetite ore and wall rock and is commonly replaced or
159 cut by the Stage II K-feldspar-magnetite veins. The stage II Fe mineralization mainly
160 contains platy mushketovite that coexists with amphibole (Fig. 4a–b), massive
161 magnetite that occurs with amphibole, chlorite, quartz and pyrite (Fig. 4c–d), and
162 disseminated magnetite with K-feldspar, or epidote-amphibole in wall rocks that has
163 undergone early K-feldspar alteration and secondary epidote-amphibole-(calcite)
164 alteration (Fig. 4e–f). In addition, residual hematite can be found in the cores of the
165 mushketovite (Fig. 4b), indicating that hematite stage occurred before the magnetite.
166 Stage III Cu mineralization is dominated by veins or aggregates of
167 quartz-hematite-chalcopyrite (Fig. 4g–h), and epidote-calcite-hematite -chalcopyrite,
168 with locally minor chlorite.

169

SAMPLING AND ANALYTICAL METHODS

170 Representative samples of three types of magnetite from Stage II Fe
171 mineralization were selected for this study. The magnetite samples include platy
172 magnetite (MA Mag - SL16-6, SL16-8, SL17-157) and massive magnetite (MB Mag -
173 SL17-116, SL17-118) collected from the ore bodies, as well as disseminated
174 magnetite (MC Mag - SL17-88, SL17-136, SL17-137) from the wall rocks. The
175 chlorite and pyrite which coexist with magnetite were also selected for analysis in this
176 study, including chlorite in platy magnetite sample (SL17-157) and massive magnetite
177 sample (SL17-116, SL17-118), and pyrite in platy magnetite sample (SL17-157),
178 massive magnetite sample (SL17-116) and disseminated magnetite sample (SL17-88).
179 All samples were prepared as standard polished thin sections and subsequently
180 analyzed for their textural characteristics by Scanning Electron Microscope (SEM) in
181 backscattered electron (BSE) mode, and magnetite, chlorite and pyrite major
182 chemistry by Electron Probe Micro-Analyzer (EPMA), and magnetite trace element
183 geochemistry by Laser Ablation Inductively Coupled Plasma Mass Spectrometry
184 (LA-ICP-MS).

185 The representative sample sections were carbon coating and then investigated
186 with a SIGMA SEM in BSE mode, at Guangzhou Institute of Geochemistry, Chinese
187 Academy of Sciences (GIGCAS). The major element geochemistry of magnetite,
188 chlorite, and pyrite was analyzed using a JEOL JXA-8230 electron microprobe at the
189 Shandong Analysis Center of China Metallurgical Geology Bureau. The operating

190 conditions were 15 kV accelerating voltage, 20 nA beam current, 1 μm beam diameter,
191 10 s counting time and ZAF correction procedure for the data reduction. Zoning in
192 magnetite grain was mapped using EPMA for Fe, Si, Ti, Al, Mn, and V, and zoning in
193 chlorite grain was mapped for Si, Fe, Al and Mg. The operating conditions were 20
194 kV voltage, 300 nA probe current, 0.5 to 4 μm beam size, and a dwell time of 100 to
195 200 ms for each point were used for mapping.

196 Magnetite trace element geochemistry was analyzed by LA-ICP-MS at the In
197 Situ Mineral Geochemistry Lab, Ore Deposit and Exploration Centre (ODEC), Hefei
198 University of Technology, Hefei, Anhui Province, China. The analyses were carried
199 out on an Agilent 7900 Quadrupole ICP-MS coupled to a Photon Machines Analyte
200 HE 193-nm ArF Excimer Laser Ablation system. Argon was used as the make-up gas
201 and mixed with the carrier gas via a T-connector before entering the ICP (Ning et al.,
202 2017; Wang et al., 2017). Each analysis was performed with a uniform spot size
203 diameter of 30mm at 8 Hz with an energy of $\sim 4 \text{ J/cm}^2$ for 40s after measuring the gas
204 blank for 20s. Standard reference materials NIST 610, NIST 612, BCR-2G, and
205 MASS-1 were used as external standards to plot calibration curves using preferred
206 element concentrations for the USGS reference glasses from the GeoReM database
207 (<http://georem.mpch-mainz.gwdg.de/>). Off-line data processing was undertaken using
208 the ICPMS Data Cal software package (Liu et al., 2008), and ^{57}Fe was used as internal
209 standardization.

210

RESULTS

211 **Petrography and texture of magnetite and chlorite**

212 As mentioned above, the magnetite in the Shuanglong deposit is composed of
213 platy (MA), massive (MB) and granular (MC) types from the Stage II Fe
214 mineralization stage. The MA magnetite commonly coexists with amphibole and
215 shows three different zones in BSE images (central light, median dark, and outer light;
216 Fig. 5a-b), indicating that it may have evolved through three generations. The light
217 magnetite (MA-I) in the core is characterized by residual hematite, abundant porosity
218 and inclusions that commonly consist of tiny W-bearing minerals such as scheelite
219 (Fig. 5b) and is commonly overgrown by the later dark MA-II magnetite that grows
220 parallel to the contact interface of MA-I magnetite (Fig. 5b). Median dark magnetite
221 (MA-II) is characterized by abundant silicate inclusions (i.e., amphibole) and weak to
222 obvious oscillatory zoning including dark and light bands (MA-IIID and MA-IIIL; Fig.
223 5b). The outermost light magnetite (MA-III) lacks porosity and inclusions, and
224 several dark bands occur within its rim (MA-III; Fig. 5b).

225 The MB magnetite occurs as aggregates of euhedral grains, coexisting with
226 amphibole, minor chalcopyrite and chlorite (Fig. 4c-d). It shows two generations
227 including primary dark magnetite which is characterized by well-developed
228 oscillatory zoning, and secondary light magnetite (Fig. 5d). In the primary dark
229 magnetite with oscillatory zoning, the dark magnetite (MB-ID) has been overgrown

230 by later light magnetite (MB-II), and abundant porosity in the light magnetite (Fig.
231 5d). Complete MB-I magnetite grains have been replaced by secondary MB-II light
232 magnetite, with abundant porosity and inclusions such as chlorite and chalcopyrite,
233 distributed in the secondary MB-II magnetite at the interface between MB-I and
234 MB-II magnetite (Fig. 5d and f). In addition, some MB-II also overgrown with
235 primary dark magnetite grain and display oscillatory zoning with dark bands (Fig. 5d).

236 The MB-II magnetite mainly coexists with chlorite, as shown in the
237 micro-inclusions and cracks (Fig. 6a). The chlorite occurs as pseudo-uniaxial plates
238 arranged in rosette shaped aggregates (Fig. 6a and b). SEM images show that rosette
239 chlorite displays oscillatory zoning pattern and mainly include two zones (Fig. 6a and
240 b), the zone 1 is composed of stacks of 20-80 μm thick layers, suggesting that the
241 primary chlorite grew in an open space. Chlorite in zone 2 overgrows along the
242 crystal in zone 1 and some also exists as euhedral to subhedral crystals. Chlorite in
243 both zones show alternated dark and light bands, and the zone 2 chlorite coexists with
244 outermost layer of MB-II magnetite (Fig. 6a).

245 The MC magnetite commonly occurs as euhedral or subhedral granular
246 magnetite in the wall rock, accompanied by two stages of alteration an early
247 K-feldspar (-calcite) phase and a late epidote-amphibole phase, where the K-feldspar
248 is replaced by later epidote and calcite (Fig. 5g). These two phases of alteration may
249 also be reflected in the textures of the magnetite, as the primary dark magnetite (MC-I)
250 has been replaced by the secondary light magnetite (MC-II) which coexists with

251 epidote (Fig. 5g-h), and the MC-I magnetite probably formed with the early
252 K-feldspar based on the sequence of alteration identified in wall rock.

253 Therefore, based on the petrography and texture of the different types of
254 magnetite, they can be divided into three phases an early MA-I mushketovite, a
255 second phase including MA-II, MB-I, and MC-I magnetite which occur as “dark
256 bands” in BSE images; and the third phase of magnetite including MA-III, MB-II, and
257 MC-II which occur as “light bands” in BSE images.

258 **EPMA**

259 The EPMA elemental mapping shows that the three generations in MA magnetite
260 including the first generation of MA-I magnetite transformed by precursor hematite,
261 and the later phrases (MA-II and MA-III) are overgrown with it. And the dark
262 magnetite (MA-II) contains higher content of Si, Al, but lower Fe content than the
263 light magnetite (MA-I and MA-III; Fig. 7a). In the MB magnetite, the dark magnetite
264 (MB-I) contains higher content of Si, Al, but lower Fe and V contents than light
265 magnetite (MB-II; Fig. 7b). For the dark and light bands in the oscillatory zoned MB-I
266 magnetite displays similar variations trend as the dark MB-ID magnetite contains
267 higher Si, Al, and lower Fe contents than that of light MB-IL magnetite (Fig. 7b). The
268 EPMA elemental mapping shows that the dark zone in MC magnetite (MC-I) contains
269 higher contents of Si and Al, but lower Fe contents than the light magnetite (MC-II;
270 Fig. 7c).

271 A total of 201 EPMA spot analyses were completed on the magnetite, with 30
272 spots on MA-I, 29 on MA-IID, 21 on MA-IIL, 36 on MA-III, 27 on MB-ID, 11 on
273 MB-IL, 17 on MB-II, 13 on MC-I and 18 on MC-II. The full results of chemical
274 contents of the Shuanglong magnetite are presented in in Table 1. Figure 8a shows
275 comparative box and whisker diagrams of major elements in the various magnetite
276 generations.

277 The mushketovite (MA) contains three generations of magnetite, and the MA-II
278 magnetite with oscillatory zoning has dark and light zones (MA-IID and MA-IIL),
279 thus corresponding to four compositional groups (Fig. 7a). MA-I magnetite has the
280 highest average Fe (72.9 wt%), the lowest Si (0.03 wt%), Al (0.02 wt%), Ca (0.03
281 wt%), Mn (0.08 wt%) and Ti (0.03 wt%), and the contents of Mg, V, Cr and Ni are
282 mostly below the detection limit; MA-IID magnetite has the lowest average Fe (68.5
283 wt%), the highest average Si (1.4 wt%), Al (0.43 wt%), Ca (0.68 wt%), Mg (0.08
284 wt%), Mn (0.13 wt%), Ti (0.03 wt%); MA-IIL magnetite contains moderate average
285 Fe (67.0 wt%), and moderately high average Si (0.94 wt%), Al (0.22 wt%), Ca (0.42
286 wt%), Mg (0.06 wt%); MA-III magnetite has similar composition with MA-I,
287 containing (average) contents of Fe (72.2 wt%), Si (0.19 wt%), Al (0.09 wt%), Ca
288 (0.25 wt%), Mn (0.08 wt%), Ti (0.04 wt%), and Mg (0.02 wt%), and with the contents
289 of Cr and Ni mostly below the detection limit.

290 The MB magnetite contains two generations, and the primary generation (MB-I)
291 contains obvious oscillatory zoning including dark and light zones (MB-ID and

292 MB-IL). MB-ID magnetite has the lowest average Fe (68.8 wt%), the highest average
293 Si (1.07 wt%), Al (0.52 wt%), Ca (0.56 wt%), Mg (0.07 wt%), and Ti (0.11 wt%);
294 MB-IL magnetite contains moderately low average Fe (70.0 wt%), and moderately
295 high average Si (0.73 wt%), Al (0.33 wt%), Ca (0.40 wt%), Mg (0.05 wt%), and Ti
296 (0.06 wt%); MB-II magnetite has the highest average Fe (72.1 wt%), the lowest Si
297 (0.11 wt%), Al (0.03 wt%), Ca (0.04 wt%), and Ti (0.05 wt%). All the MB magnetite
298 subtypes contain similar Mn (0.072, 0.071, 0.058 wt%, respectively) and V (0.024,
299 0.024, 0.03 wt%, respectively). The concentrations of Ni and Cr are mostly below the
300 detection limit. For the MC magnetite, MC-I magnetite has higher average Si (0.69
301 wt%), Al (0.16 wt%), Ca (0.36 wt%), Mg (0.04 wt%), Mn (0.14 wt%), Ti (0.04 wt%),
302 V (0.06 wt%), and lower average Fe (70.0 wt%) compared to MC-II magnetite which
303 has the lower average Si (0.18 wt%), Al (0.02 wt%), Ca (0.06 wt%), Mg (0.01 wt%),
304 Mn (0.1 wt%), Ti (0.03 wt%), V (0.04 wt%) and higher average Fe (69.2 wt%); and
305 the concentrations in Ni and Cr of both subtypes are mostly below the detection limit.

306 The tungsten inclusions in MA-I magnetite have also been analyzed by EPMA,
307 however, due to its extremely small size, the data seems to be a mixture with
308 magnetite. Even so, the analyses reveal high Ca (5.4-6.4 wt%) and W (19.3-28.6 wt%)
309 contents, with significant Mo (0.6-0.9 wt%; Appendix Table 1).

310 The EPMA elemental mapping shows that the alternated dark and light bands in
311 chlorite have compositional variations, the dark bands contain higher content of Si,
312 Mg, but lower Fe and Al content than the light bands (Fig. 6). The EPMA spot

313 analysis shows that the oscillatory contrasts are mainly caused by FeO and MgO
314 content variation. The average FeO and MgO content in zone 1 have variations from
315 dark (21.7 wt%, 17.8 wt%; Appendix Table 1) to light bands (25.9 wt%, 14.7 wt%),
316 and ranges from dark (19.5 wt%, 18.8 wt%) to light bands (24.8 wt%, 15.9 wt%) in
317 zone 2. SiO₂ and Al₂O₃ contents show smaller variations (less than 2 wt%). The
318 EPMA spot analysis shows that the pyrite which coexists with magnetite has variable
319 and high Co and Ni contents (0.19-1.54 wt%, 0.02-0.25 wt%).

320 **LA-ICP-MS**

321 Due to the small-scale oscillatory zoning in MA-II and MB-I magnetite, light
322 and dark domains of magnetite could not be separately measured by LA-ICP-MS, and
323 thus, analytical results of MA-II and MB-I magnetite represent the average
324 composition of light and dark domains in oscillatory zoning. Full LA-ICP-MS
325 analytical results and the corresponding detection limits are provided in Appendix
326 Table 1 and illustrated in Figure 8b. Most of the data for Cu and rare earth elements
327 (REE) are below detection limits. In general, the dark magnetite including MA-II,
328 MB-I, and MC-I magnetite have higher Mg, Al, Si, Ca, Ti, Zn, Ga, and Sr
329 concentrations than those in the light magnetite including MA-III, MB-II, and MC-I.
330 The mushketovite including MA-I and MA-II magnetite have relatively high W
331 (69-3311 ppm; 1-271 ppm), Sn (1.2-31.7 ppm; 1.1-14.6 ppm), and U (2.7-78.6 ppm;
332 0.2-1.9 ppm) concentrations than other magnetite subtypes, although the content of

333 these elements in other magnetite types is all below 1 ppm or below the detection
334 limit (Appendix Table 1).

335 The statistical techniques (factor analysis) were used in this study, to examine
336 underlying relations among elements content obtained with LA-ICP-MS. We used a
337 principal component extraction method with a Kaiser Varimax rotation (Kaiser, 1958)
338 and a Bartlett factor score (Bartlett, 1937). Factors score coefficients for each analysis
339 are derived from the factor loadings. The factor score represents the significance of a
340 given factor for an analysis. The factor score line outside -0.5 to 0.5 indicate
341 significant relations. The factor analysis results provide three multiple elements
342 factors: factors 1 comprising Si, Al, Ca, Mg, Sr, Mn and Co, factors 2 comprising Ti,
343 Ga and Zn, and factors 3 comprising V, Ni, Cr and Sc (Fig. 9).

344 DISCUSSION

345 **Controlling factors for magnetite composition**

346 Magnetite has a general stoichiometry XY_2O_4 , where X is divalent cations such
347 as Mg, Fe, Mn and Ni, and Y represents trivalent or tetravalent cations such as Si, Al,
348 Fe and V. For the EPMA results, positive correlations ($R^2 \geq 0.7$) are exhibited among Si,
349 Al, Ca, Mg and Mn, and Ti positively correlates with Al ($R^2=0.6$), whereas Ti and V
350 do not correlate with any of the other trace elements. The factor analysis of
351 LA-ICP-MS results also reveal three factors (Fig. 9), suggesting that some trace
352 elements have underlying correlation, e.g., Si, Ca, Al, Mg, Mn, Sr, Sn and Co; Ti, Zn

353 and Ga; V, Ni, Sc and Cr. These correlations of trace elements could probably be
354 attributed to several reasons, such as the inclusions (amphibole) and partition
355 coefficient. Fe^{3+} is negatively correlated with Si and Al (Fig. 10a), indicating that
356 these elements were incorporated into the structural sites of magnetite by substitution
357 of Si^{4+} and Al^{3+} for Fe^{3+} . Si also positively correlates with Ca and Mg, indicating
358 these divalent elements play as a valence state balance related to the substitution. Fe^{2+}
359 is negatively correlated with Ca and Mg (Fig. 10b), indicating Mg^{2+} and Ca^{2+}
360 substitute Fe^{2+} . Micro-inclusions or nano-inclusions in magnetite can influence the
361 magnetite geochemistry detected. For MA-I magnetite, the few anomalously high W
362 and Ca signal may point to the presence of invisible W-bearing micro-inclusions (Fig.
363 11a and b), probably be scheelite (CaWO_4), and few visible scheelite inclusions were
364 also found in MA-I magnetite (Fig. 5b). Moreover, the positive correlate between W
365 and Mn indicate the invisible wolframite ($(\text{Fe}, \text{Mn}) \text{WO}_4$) inclusions may occur (Fig.
366 12h). However, some spots show extremely high W content with stable signal,
367 indicating W may also incorporate into the lattice of magnetite (Fig. 11c). For other
368 magnetite, especially the trace element rich zone, our EPMA spot analyses can avoid
369 the visible inclusions, however, most LA-ICP-MS analyses could not avoid
370 nano-scale inclusions. For example, the positive correlate between Sr and Al, Ca may
371 be caused by the silicate inclusions due to that Sr is highly incompatible for magnetite
372 (Fig. 12f and g, Nielsen et al., 1994).

373 The composition of magnetite is mainly controlled by some factors including: (1)

374 co-precipitating minerals (mainly Fe-sulfides), (2) fluid geochemistry and rock-water
375 reaction, and (3) physicochemical conditions including temperature and oxygen
376 fugacity (Whalen and Chappel, 1988; Frost and Lindsley, 1991; Ghiorso and Sack,
377 1991). The presence of other minerals within the mineral assemblage has significant
378 effect on magnetite geochemistry. Minerals coprecipitated with magnetite can
379 preferentially partition some certain elements, for instance, chalcophile (e.g., Cu, Co,
380 Pb) and siderophile elements (e.g., Ni) preferentially partition into sulfides minerals
381 (Cygan and Candela 1995; Fleet et al., 1996; Simon et al. 2008), whereas lithophile
382 elements (e.g., Mg, Al, Si, Ti) partition into silicates minerals (Frost 1991; Toplis and
383 Corgne 2002). The MA and MB magnetite produced from infilling coexist with
384 similar minerals, primary dark magnetite with mainly amphibole and secondary light
385 with mainly chlorite and few pyrite, whereas the MC magnetite was produced by
386 K-Fe and subsequent Ca-Fe alteration. The MB and MC have similar low Cu, Pb, Co
387 and Ni contents (Fig. 8b), however the rare coexisting pyrite has extremely high Co
388 and Ni contents (0.19-1.54 wt%, 0.02-0.25 wt%), indicating that coexisting sulfides
389 effectively scavenge certain elements. Although some difference in lithophile
390 elements (Fig. 8a, MA has higher Si, Al, Mg, Ca, Sr contents), it is difficult to know
391 what caused it due to the influence of potential nano silicate inclusions, and such
392 systematic difference was probably controlled by the fluid geochemistry or degree of
393 rock-water reaction. Extensive fluid-host rock interaction commonly results in
394 hydrothermal magnetite enriching in Si, Al, Mg and Mn (Carew 2004; Nadoll et al.

395 2014; Deditius et al. 2018). The MC magnetite has low Si, Al, Mg, and Mn contents
396 (Fig. 8a), reflecting a low degree of fluid-rock interaction, and the MC has high Mn
397 content probably due to the recharge of Mn during K-Fe alteration (Corriveau et al.
398 2016). In addition, MA magnetite samples have significant W, Sn, Pb, and U contents
399 which are basically below the detection line in MB and MC (Fig. 8b), and these
400 elements are actually concentrated in MA-I (core of MA). Considering that MA-I was
401 transformed from hematite, it may represent the compositional information of the
402 original hematite.

403 Little experimental result is currently available on the partitioning efficiency of
404 trace elements between magnetite and hydrothermal fluids (Chou and Eugster, 1977;
405 Ilton and Eugster, 1989; Simon et al., 2004). It is generally considered that some
406 factors such as temperature and oxygen fugacity control partitioning of trace elements
407 in hydrothermal-derived magnetite like that in igneous system. For example, Al and Ti
408 concentrations in igneous-derived magnetite are high positively correlative with
409 temperature (Nielsen et al., 1994; Toplis and Carroll, 1995), and these elements also
410 decrease from high-temperature porphyry and skarn magnetite to low-temperature
411 banded iron formation (BIF) (Nadoll et al., 2014). For the oscillatory zoning in MA-II
412 and MB-I magnetite (Fig. 8a), it is clear that the light bands (MB-IL) have lower Ti
413 contents than the dark bands (MB-ID, D and L mean dark and light bands), however,
414 no obvious differences in Ti content were observed between dark (MA-IID) and light
415 (MA-IIL) bands in MA-II magnetite (Fig. 8a), likely due to the very low abundances.

416 Combined with the decrease of trace element such as Al and Mg being related to
417 temperature in MA-II and MB-I magnetite, it is likely that the oscillatory zoning in
418 MA-II and MB-I magnetite was mainly controlled by changes in temperature. The
419 above factor analysis result also shows that Ti, Ga and Zn have correlations (Fig. 9,
420 12a and 12b). Nadoll et al. (2014) proposed that Ga concentration follows the trend
421 that highest concentrations are in high temperature igneous magnetite and the lowest
422 in BIF magnetite. Previous study also showed that Zn is more effectively fractionated
423 into hydrothermal fluids than Fe with decreasing temperature (Ilton and Eugster 1989;
424 Wu et al., 2019). Therefore, Ga and Zn in Shuanglong magnetite were also controlled
425 by temperature. Nadoll et al. (2014) proposed the Ti + V vs. Al + Mn plot can reflect
426 the variation in temperature, with high-temperature magnetite plotting at high values
427 of Ti + V and Al + Mn values. The Shuanglong magnetite displays a descending
428 temperature trend suggesting they formed from a fluid under middle to
429 high-temperature conditions > 200 °C (Fig. 13; Deditius et al., 2018; Nadoll et al.,
430 2014). Some elements, such as V and Cr, can occur in various valence states (e.g., V³⁺,
431 V⁴⁺, V⁵⁺), and therefore, their behavior is strongly linked to fO_2 (Nadoll et al., 2014;
432 Toplis and Carroll, 1995; Toplis and Corgne, 2002). For example, Toplis and Corgne
433 (2002) proposed vanadium partitioning is strongly dependent on oxygen fugacity,
434 decreasing by approximately one order of magnitude with increasing fO_2 from
435 NNO-0.7 to NNO+2.6 for magnetite at a constant temperature (1068 °C), and this
436 behaviour is inferred to be dominated by partitioning of V³⁺ due to the relative

437 proportions of V^{n+} (Toplis and Corgne, 2002). However, it is less well constrained at
438 cooler temperature hydrothermal fluids (Nadoll et al., 2014). The MA magnetite has
439 low V contents, reflecting a more oxidized fluid. For the dark zone and light zone in
440 MA and MB magnetite, the contents of V were slightly increased, indicating the
441 decreasing of oxygen fugacity in fluids which is consistent with the coexisting pyrite
442 in light magnetite. The Ni content also shows strong positive correlation with V (Fig.
443 12c), probably also related with the oxygen fugacity. Moreover, the EPMA data show
444 that the V content of MC-II decreases obviously compared to MC-I (Fig. 8a),
445 revealing an opposite trend.

446 **Tungsten remobilization during mushketovitization**

447 Previous studies have shown that platy magnetite (also call mushketovite) is
448 common in IOCG deposits, such as the Heijianshan deposit in Eastern Tianshan and
449 the Candelaria deposit in Chile (Marschik, 2001; Zhao et al., 2018a). The
450 mushketovite is interpreted to result from transformation of magnetite after hematite
451 during reducing fluid reactions, a process called mushketovitization (Hu et al., 2020;
452 Mucke and Cabral 2005; Ohmoto, 2003). At Shuanglong, abundant platy magnetite
453 (MA Mag) coexists with actinolite, pyrite and chlorite. In fact, only the core of MA
454 magnetite is the result during the reducing reactions according to the presence of
455 residual hematite and abundant micro porosity probably formed by the volume
456 shrinkage during the reactions (Hu et al., 2020; Mucke and Cabral 2005).

457 Significantly, abundant scheelite inclusions occur in the MA-I magnetite, and the
458 previous discussion revealed that the high W contents of MA-I were also attributed to
459 invisible scheelite inclusions and lattice-tungsten. Measured partition coefficients for
460 igneous magnetite indicate that W is highly incompatible (Klemme et al., 2006),
461 however, some authors have reported the incorporation of W into magnetite at
462 significant levels (up to dozens of ppm; Candela, 1997; Carew, 2004; Huang et al.,
463 2019). Thus, such a high contents in mushketovite could have been inherited from the
464 precursor hematite. Indeed, W-U-Sn-Mo enriched hematite is common in many
465 hydrothermal deposits, such as the Olympic Dam Cu-U-Au-Ag deposit (Ciobanu et
466 al., 2013; Verdugo-Ihl et al., 2017), and these element were incorporated into the
467 structural sites of magnetite by: $2\text{Fe}^{3+} \leftrightarrow \text{Me}^{6+} + \square$, in which the Me could be U, W
468 and Pb (Ciobanu et al., 2013). Verdugo-Ihl et al. (2020) also proposed significant W
469 can be released from hematite due to coupled dissolution–reprecipitation during
470 interaction with fluids, and was confirmed at nanoscale. Therefore, when
471 mushketovitization of hematite occurred, the reduction in volume results in abundant
472 porosity and cracks which provide pathways for fluid percolation and the release of W.
473 In addition, the estimate values of $f\text{O}_2$ for many scheelite or wolframite deposits
474 mostly lie between quartz-fayalite-magnetite (QFM) and Ni-NiO (NNO) buffers
475 (Wood, 2000 and reference therein), and no tungsten inclusions have been found in
476 the hematite, indicating that these inclusions are not coexisting with the previous
477 hematite. Therefore, these scheelite inclusions might be released from the previous

478 hematite and precipitated in open spaces due to local supersaturation.

479 **Origin of synchronous oscillatory zoning in magnetite-chlorite**

480 **aggregates**

481 Oscillatory zoning in MA and MB magnetite is very common in Shuanglong
482 deposit and can be divided into two phases in MA-II, MB-I, and MA-III, MB-II
483 respectively. They are composed of alternately dark and light bands under BSE
484 imaging. The coexisting mineral chlorite occurring as rosette-shape can be divided
485 into two oscillatory zones. Similar oscillatory zoning in magnetite is widely observed
486 in Fe skarn deposit (Dare et al., 2014; Huang et al., 2018), IOCG deposit (Huang et al.,
487 2019) and IOA deposit (Knipping et al., 2015b; Deditius et al., 2018). A common
488 explanation for the oscillatory zoning in magnetite is that the changes in fluid
489 compositions and/or physicochemical parameters (such as temperature and oxygen
490 fugacity) during the crystal growth, that could also periodically change the
491 partitioning behavior of trace elements into magnetite (Dare et al., 2015; Knipping et
492 al., 2015b; Sievwright et al., 2017; Huang et al., 2019). The oscillatory zoning is
493 widely distributed in MA and MB magnetite, suggesting that oscillatory zoning is not
494 due to ultra-local fluid disequilibrium. The sharp compositional boundaries between
495 the dark and light zones within single magnetite grains imply preservation of the
496 original oscillatory zoning, which requires the growth rate of the crystal to be higher
497 than intracrystalline diffusion (Watson and Liang, 1995). EPMA point analyses

498 indicate the trace element-rich zones (MA-IID, MB-ID, and dark bands in MA-III and
499 MB-II) involve all elements except V, and Ti for MA (Fig. 14), implying a relatively
500 high temperature and stable oxygen fugacity. Therefore, oscillatory zoning in
501 magnetite from MA and MB was likely probably caused by variation in temperature.

502 The rosette-like chlorite with similar oscillatory zoning was also observed from a
503 green-schist facies fault zone in Spain (Trincal et al., 2015). Previous studies showed
504 that chlorite chemistry can be a good indicator of the temperature of crystallization
505 (Cathelineau and Nieva 1985; Inoue et al. 2009; Bourdelle et al. 2013; Lanari et al.
506 2014). In this study, we use thermodynamic thermometer calibration of Lanari et al.
507 (2014) based on the following equilibrium: 2 clinocllore + 3 sudoite = 4 amesite + 4
508 H₂O + 7 quartz. The ideal contributions of amesite, clinocllore, and sudoite are
509 calculated using a site mixing model for chlorite end-members, and the ln(K) can be
510 calculated using the following equation: $\ln(K) = \ln(a^4Ame / a^2Clna^3Sud)$. Chlorite
511 temperatures were then calculated using: $T_{Chl} (^\circ C) = 172341 / [-R\ln(K) + 315.149] -$
512 273.15. The average calculated temperature of the dark and light band was 243 °C and
513 444 °C in zone 1, and 430 °C and 212 °C in zone 2, which were consistent with the
514 estimated temperature range of MA-III and MB-II magnetite (Fig. 13).

515 Texture features of MB magnetite reveal that the primary MB-I underwent a
516 process of dissolution and precipitation according to the sharp boundaries with
517 secondary magnetite, and the chlorite inclusions near the boundaries reveal that it
518 occurred in a lower temperature than primary magnetite (Fig. 6a). Moreover, some

519 MB-II magnetite also overgrows with primary MB-I grains (Fig. 5d), and these
520 overgrowing MB-II magnetite usually have oscillatory zoning (Fig. 5d). It is
521 reasonable to speculate that MB-II magnetite precipitated in two processes, including
522 the early dissolution of massive MB-I magnetite, and subsequent precipitation of
523 MB-II by two forms. For the coexisting chlorite, the zone 2 was synchronously
524 formed with the MB-II magnetite overgrowing with the MB-I grains, and the zone 1
525 most likely formed during the early dissolution of MB-I magnetite. Previous studies
526 have shown that increasing temperature, increasing chloride concentration, or
527 decreasing pH are the most important controls on the dissolution of magnetite because
528 these can enhance the solubility and the undersaturation of iron (Crerar et al., 1978;
529 Hemley and Hunt, 1992). Experimental studies also revealed that additions of HCl
530 could strongly enhance the solubility of magnetite in hydrothermal fluid (Holser and
531 Schner, 1961; Chou and Eugster, 1977). External fluids such as basin brines or
532 formation waters that dissolved evaporites, which will increase the salinity and Cl
533 contents in ore-forming fluids, promoting the dissolution and reprecipitation process,
534 have been identified as important components in many Fe skarn and IOCG deposits
535 (Broughm et al., 2017; Bain et al., 2020; Hu et al., 2014; Huang and Beaudoin, 2019).
536 Therefore, the fluid composition especially the Cl content in equilibrium with MB-I
537 magnetite could have changed, causing the dissolution of MB-I magnetite. Combined
538 with the temperature variation in zone 1 of chlorite, it may indicate the incursion of
539 low-temperature external fluid which caused the change of fluid composition. The

540 composition of MB-II magnetite and chlorite, such as the higher Ti content in dark
541 bands in MB-II magnetite (Fig. 14), and the decreasing of temperature from dark band
542 to light band in zone 2 of chlorite, all indicate that the decreasing of temperature in
543 fluid may result in the precipitation of MB-II magnetite.

544 **IMPLICATIONS**

545 At Shuanglong, the MA-I magnetite represents a transformational mushketovite
546 phase characterized by abundant porosity and extremely high W content, which
547 mainly display as micro- and invisible scheelite inclusions probably caused by the W
548 expulsion during mushketovitization. The sharp contacts between primary dark
549 magnetite and later light magnetite, suggest that they may have formed via coupled
550 dissolution and reprecipitation. Moreover, the synchronous oscillatory zoning in light
551 magnetite and coexisting chlorite was related to the variation of fluids temperature,
552 and probably caused by the incursion of external low temperature and high salinity
553 fluids which can dissolve the primary dark magnetite.

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841 **Figure captions**

842 **Figure 1.** (a) Simplified tectonic map of the Central Asian orogenic belt (CAOB), modified after
843 Şengör (1996). (b) Tectonic map of north Xinjiang, modified after Chen et al. (2012). (c)
844 Geological map of the Eastern Tianshan and distribution of the ore deposits, modified after Zhao
845 et al. (2019).

846 **Figure 2.** Geological map of the Shuanglong Fe-Cu deposit (modified after Xinjiang Uygur
847 Autonomous Region Geological Survey (2003)).

848 **Figure 3.** Alteration and mineralization paragenesis of the Shuanglong Fe–Cu deposit.

849 **Figure 4.** Photographs showing representative alteration/mineralization paragenesis in the
850 Shuanglong Fe-Cu deposit. (a) platy MA magnetite. (b) photomicrographs of platy MA magnetite.
851 (c) massive magnetite coexisting with amphibole, chlorite and chalcopyrite. (d) photomicrographs
852 of massive MB magnetite. (e) MC magnetite coexisting with K-feldspar, epidote, and amphibole
853 in host rock of the Tugutubulak Formation. (f) photomicrograph of granular MC magnetite
854 coexisting with K-feldspar and epidote. (g-h) quartz-hematite-chalcopyrite vein cutting platy MA

855 magnetite. Abbreviations: *Mag* = magnetite, *Kfs* = potassium feldspar, *Cal* = calcite, *Ep* = epidote,
856 *Amp* = amphibole, *Qtz* = quartz, *Hem* = hematite, *Ccp* = chalcopyrite, *Chl* = chlorite.

857 **Figure 5.** Photomicrographs (**a, c, e, and g**) and BSE images (**b, d, f, and h**) of the Shuanglong
858 magnetite. (**a**) Platy magnetite (MA). (**b**) BSE images of platy magnetite, which shows three
859 different zones. MA-I magnetite is light with abundant porosity and inclusions. MA-II magnetite
860 is dark and shows chemical oscillatory zoning. And MA-III magnetite is light and lack of porosity
861 and inclusions. (**c**) massive magnetite (MB). (**d**) Two generations of magnetite in massive
862 magnetite (MB). The first generation of magnetite (MB-I) shows oscillatory zoning composed of
863 dark-gray (MB-ID) and light-gray (MB-IL) zones, and MB-ID magnetite was replaced by MB-IL
864 magnetite which contains abundant porosity and inclusions. The second generation of magnetite
865 (MB-II) crosscuts MB-I magnetite and contains porosity and inclusions locally. (**e**) Massive
866 magnetite (MB) and later mineral assemblage of chalcopyrite-sphalerite-chlorite. (**f**) BSE image of
867 MB magnetite shows that the secondary magnetite (MB-II) contains chalcopyrite inclusions
868 crosscuts the first generation of magnetite (MB-I). (**g**) Granular magnetite (MC) coexisting with
869 K-feldspar and epidote, and K-feldspar was replaced by late epidote and calcite. (**h**) BSE image of
870 MC magnetite, which shows two different zones. MC-I magnetite is dark and replaced by MC-II
871 light magnetite. Abbreviations: *Mag* = magnetite, *Kfs* = potassium feldspar, *Ep* = epidote, *Amp* =
872 amphibole, *Qtz* = quartz, *Hem* = hematite, *Chl* = chlorite, *Sch* = scheelite, *Ccp* = chalcopyrite, *Sp*
873 = sphalerite, *Cal* = calcite.

874 **Figure 6.** BSE images (**a and b**) of the MB magnetite and coexisting chlorite, and EMPA mapping
875 of selected elements in chlorite. Abbreviations: *Hem* = hematite, *Chl* = chlorite.

- 876 **Figure 7.** EMPA mapping of selected elements in magnetites from the Shuanglong deposit.
- 877 **Figure 8.** Box diagram of major (a) and trace (b) elements concentrations for the Shuanglong
878 magnetite.
- 879 **Figure 9.** Rotated component plot for magnetite showing factor analysis results from trace
880 elements with significant content.
- 881 **Figure 10.** Plots displaying negative correlations between (a) Si and Al vs. calculated Fe^{3+} in atom
882 per formula unit (a.p.f.u.) and (b) Mg and Ca vs. Fe^{2+} (a.p.f.u.).
- 883 **Figure 11.** Time-resolved analytical signal of LA-ICP-MS analysis for MA-I magnetite. CPS =
884 counts per second.
- 885 **Figure 12.** Binary diagrams of (a) Ga vs. Ti, (b) Zn vs. Ti, (c) Ni vs. V, (d) Cr vs. V, (e) Sc vs. V,
886 (f) Sr vs. Al, (g) Sr vs. Ca and (h) W vs. Mn for the Shuanglong magnetite.
- 887 **Figure 13.** Plots of (Al + Mn) vs. (Ti + V) diagram, modified after Nadoll et al. (2014) and
888 Deditius et al. (2018).
- 889 **Figure 14.** EPMA transects across the oscillatory zoning in dark and light magnetite.
890

Table 1. Mean, minimum, and maximum contents (wt%) of magnetite from the Shuanglong deposit by electron microprobe analyses

Magnetite type		Fe	Si	Al	Ca	Mg	Mn	Ti	V	Cr	Ni
MA-I (n=30)	Mean	72.868	0.025	0.018	0.033	0.008	0.084	0.031	0.018	0.015	b.d.1
	Min	70.019	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1
	Max	73.711	0.097	0.037	0.090	0.011	0.154	0.037	0.020	0.023	b.d.1
MA-IIID (n=29)	Mean	68.488	1.400	0.428	0.675	0.081	0.130	0.034	0.024	0.017	0.026
	Min	67.238	1.014	0.209	0.417	0.061	0.026	b.d.1	b.d.1	b.d.1	b.d.1
	Max	69.885	1.646	0.597	0.855	0.125	0.197	0.052	0.044	0.027	0.032
MA-IIL (n=21)	Mean	69.965	0.942	0.217	0.418	0.060	0.117	0.040	0.027	0.037	0.025
	Min	67.444	0.089	0.025	0.044	0.004	0.029	b.d.1	b.d.1	b.d.1	b.d.1
	Max	71.229	1.305	0.409	0.666	0.103	0.177	0.073	0.033	0.097	0.025
MA-III (n=36)	Mean	72.171	0.192	0.087	0.253	0.022	0.083	0.036	0.022	0.025	b.d.1
	Min	69.058	0.014	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1
	Max	73.217	1.070	0.414	0.398	0.054	0.147	0.046	0.040	0.032	b.d.1
MB-ID (n=27)	Mean	68.763	1.072	0.518	0.560	0.068	0.072	0.112	0.024	0.023	b.d.1
	Min	67.133	0.850	0.343	0.404	0.034	b.d.1	0.061	0.015	b.d.1	b.d.1
	Max	69.990	1.447	0.677	0.748	0.159	0.132	0.187	0.034	0.040	b.d.1
MB-IL (n=11)	Mean	69.971	0.732	0.333	0.399	0.049	0.071	0.060	0.024	0.017	b.d.1
	Min	69.373	0.620	0.248	0.340	0.028	b.d.1	0.027	b.d.1	b.d.1	b.d.1
	Max	70.976	0.837	0.372	0.447	0.094	0.114	0.100	0.030	0.020	b.d.1
MB-II (n=17)	Mean	72.101	0.113	0.030	0.040	0.010	0.058	0.049	0.030	0.024	b.d.1
	Min	71.138	0.043	0.006	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1	b.d.1
	Max	72.790	0.326	0.076	0.090	0.013	0.097	0.064	0.055	0.031	b.d.1
MC-I (n=13)	Mean	70.018	0.691	0.175	0.361	0.039	0.140	0.041	0.061	0.023	0.037
	Min	69.165	0.364	0.066	0.151	0.023	0.092	0.002	0.004	b.d.1	b.d.1
	Max	71.273	0.858	0.268	0.539	0.055	0.170	0.069	0.088	0.037	0.037
MC-II (n=18)	Mean	71.595	0.181	0.018	0.063	0.008	0.095	0.031	0.041	0.028	0.036
	Min	71.005	0.098	0.005	0.023	b.d.1	0.055	b.d.1	0.017	b.d.1	b.d.1
	Max	72.092	0.261	0.041	0.234	0.013	0.132	0.039	0.073	0.072	0.037

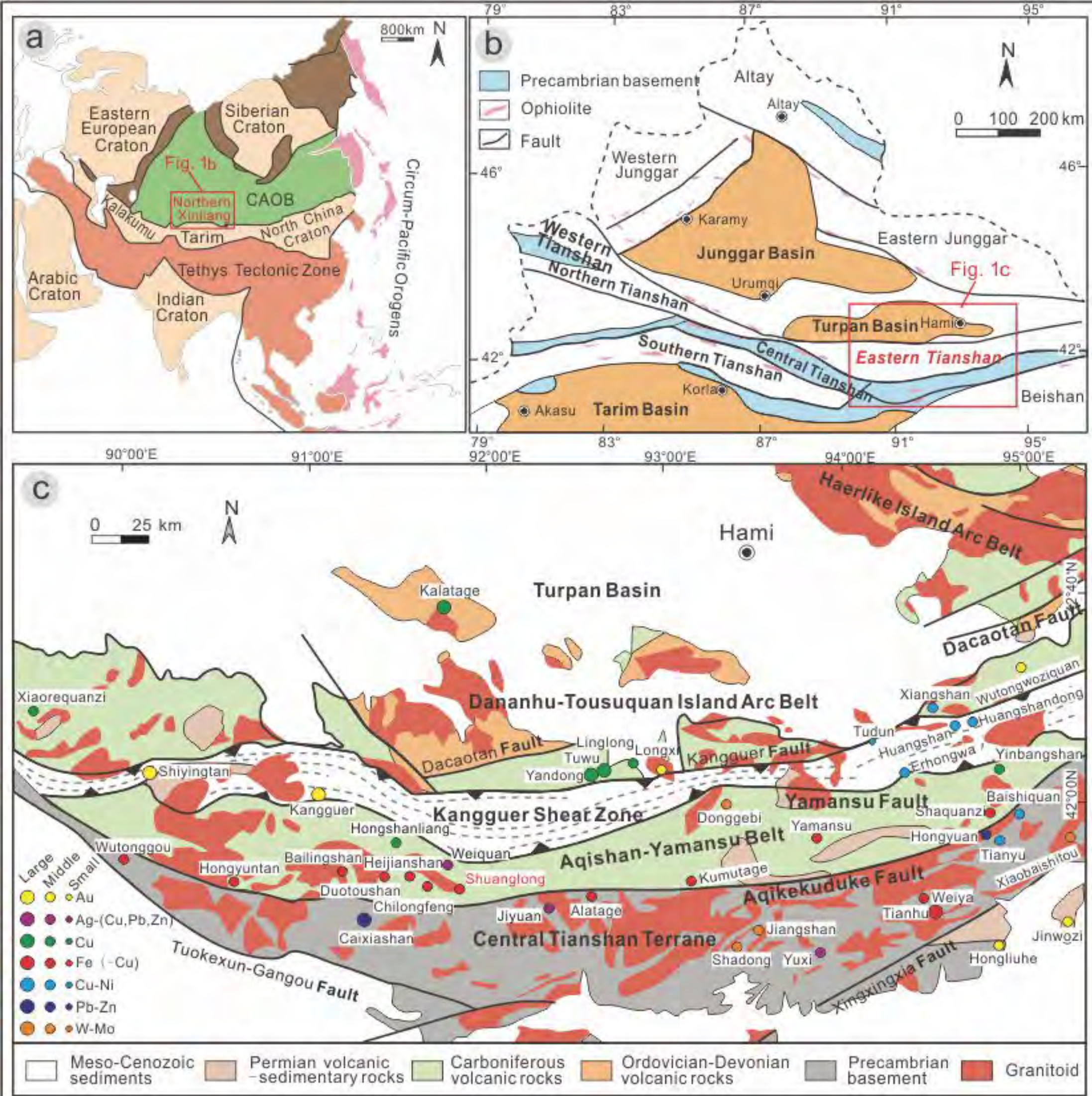


Fig 1

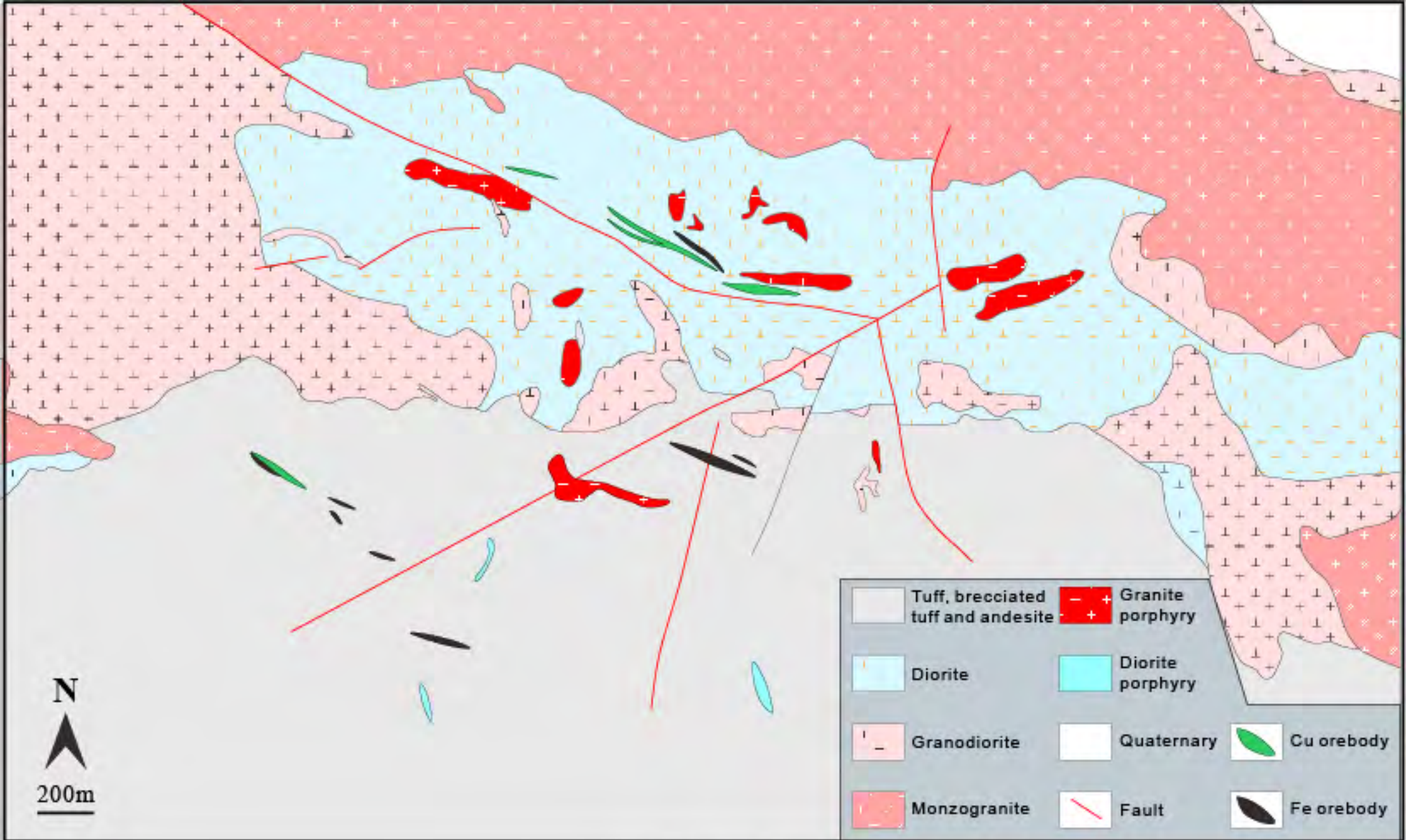


Fig 2

Minerals	Stage I skarn alternation	Stage II Fe mineralization	Stage III Cu mineralization
Garnet	Abundant		
Hematite		Local	Local
Epidote		Abundant	Abundant
K-feldspar		Local	
Amphibole		Abundant	
Magnetite		Abundant	
Pyrite		Abundant	Abundant
Calcite		Trace	Abundant
Apitite		Trace	
Quartz		Local	Local
Chlorite		Local	Abundant
Chalcopyrite		Trace	Abundant
Gypsum			Local
Galena			Trace
Bornite			Trace
Sphalerite			Trace

Abundant
 Local
 Trace

Fig 3

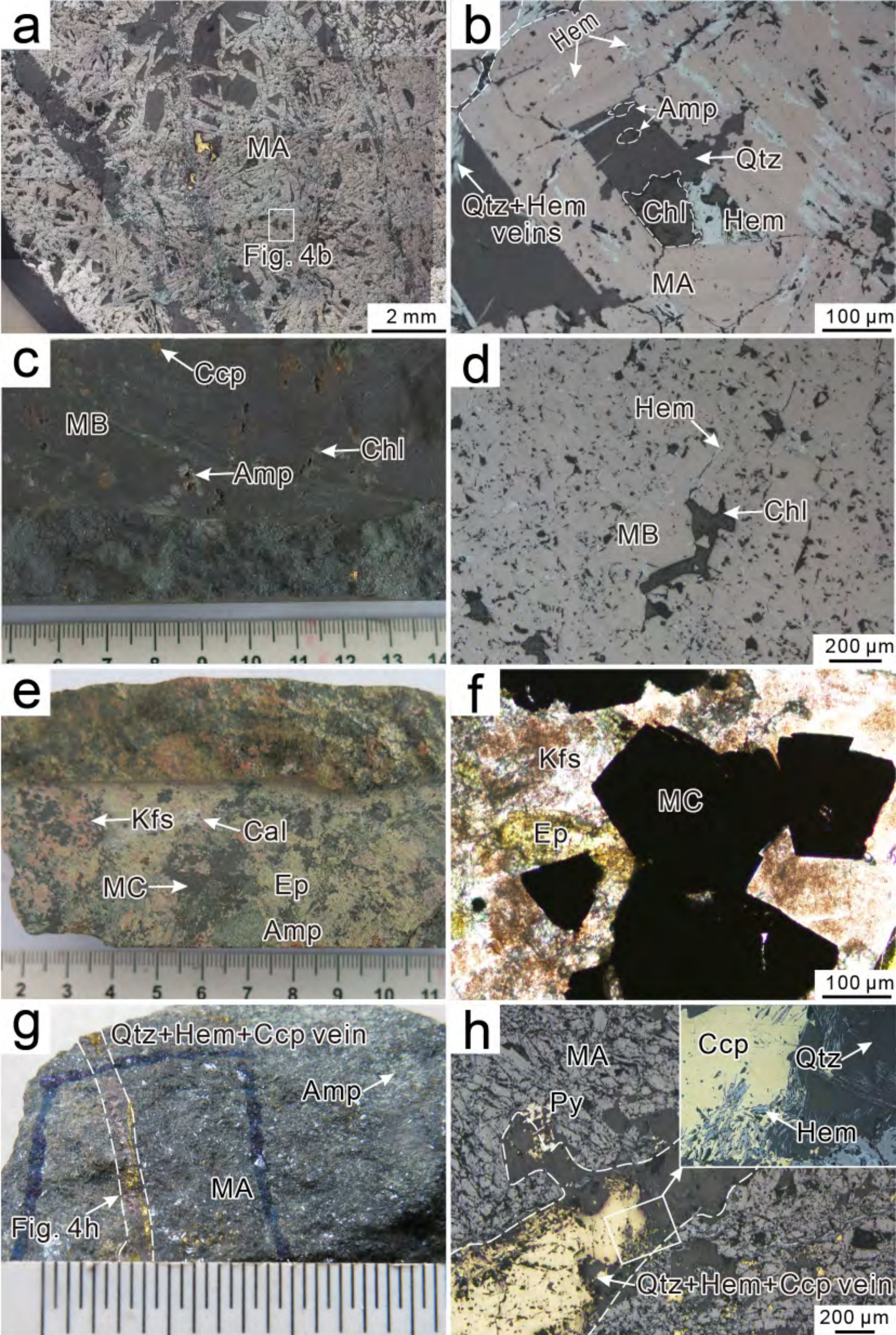


Fig 4

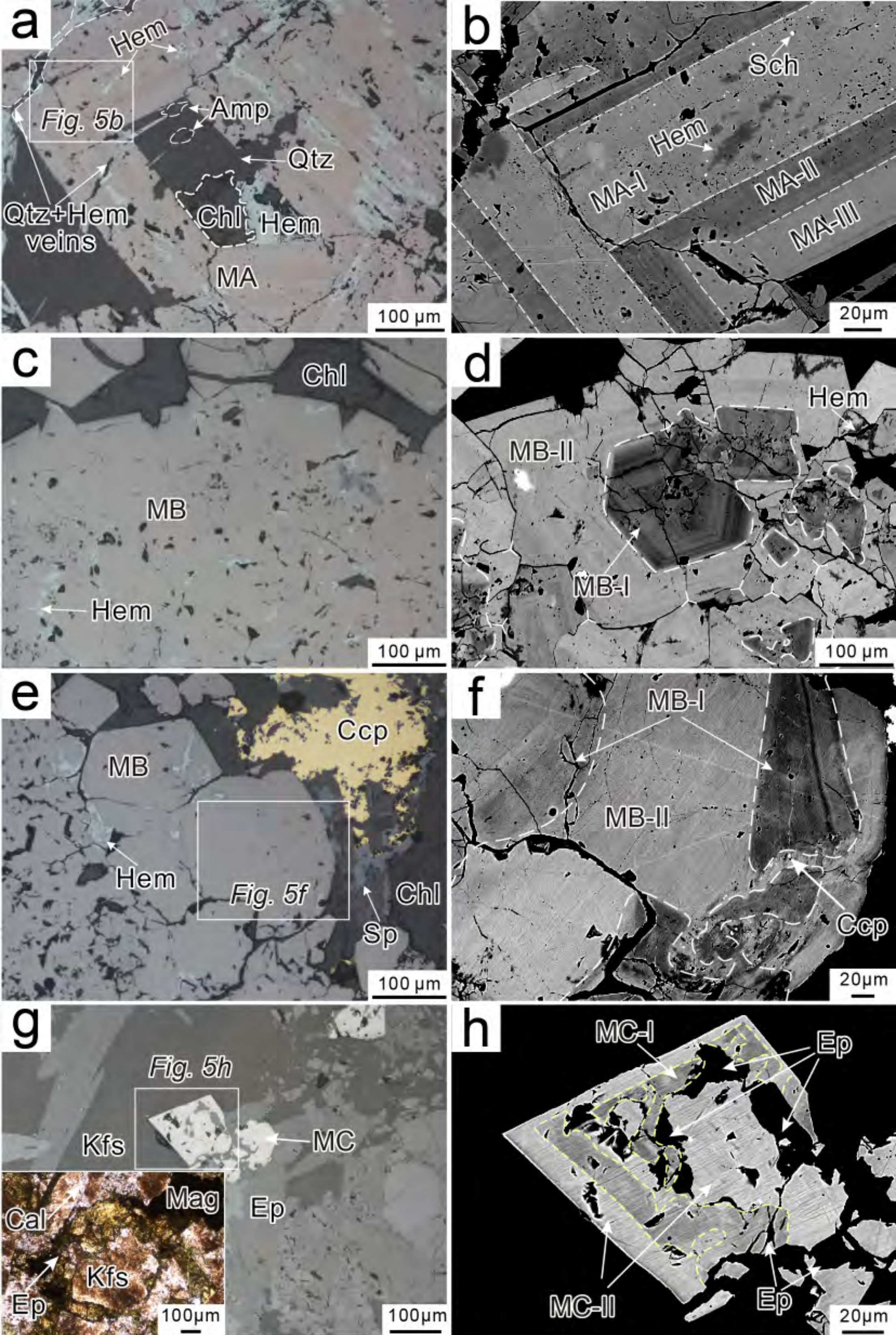


Fig 5

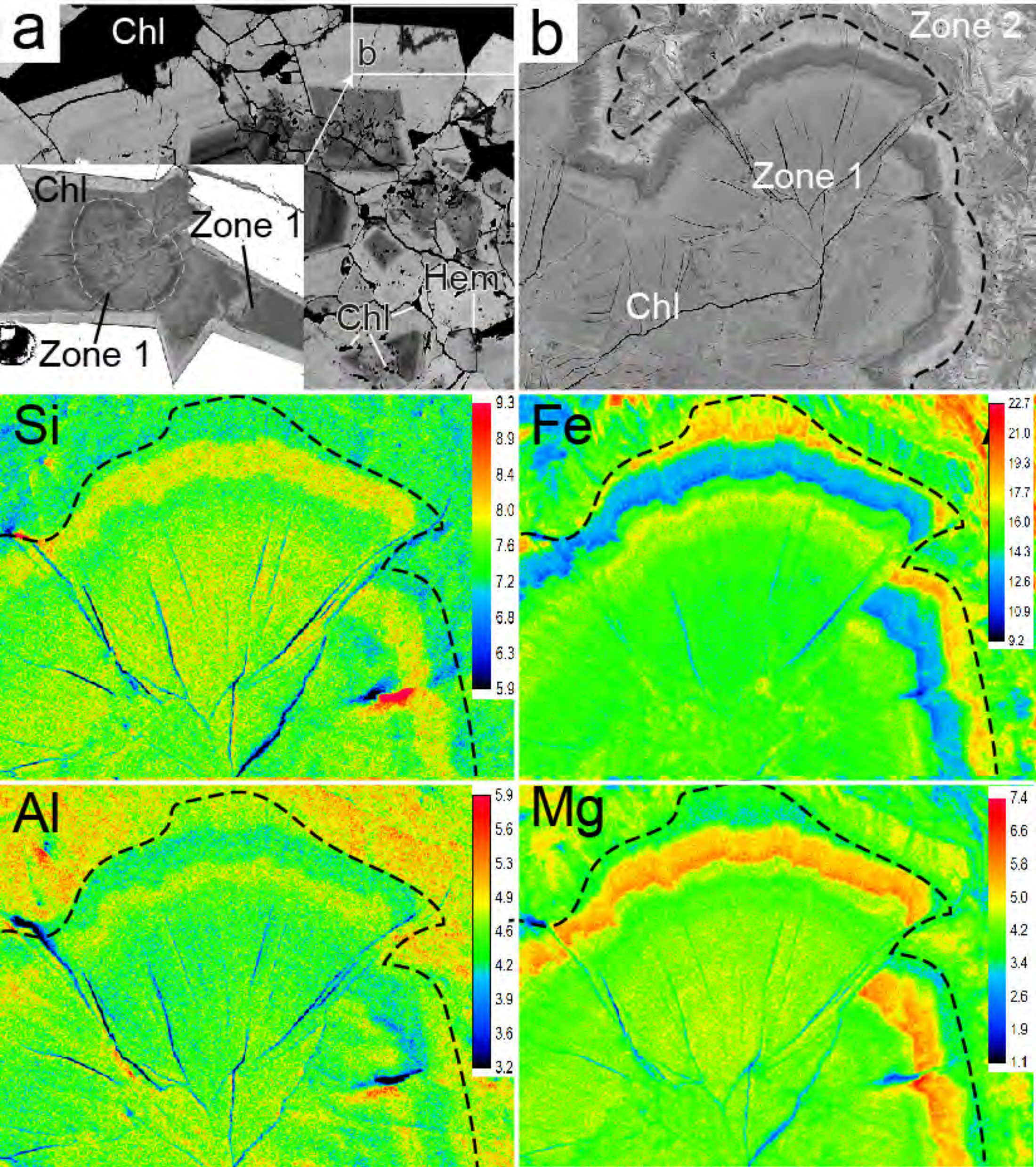


Fig 6

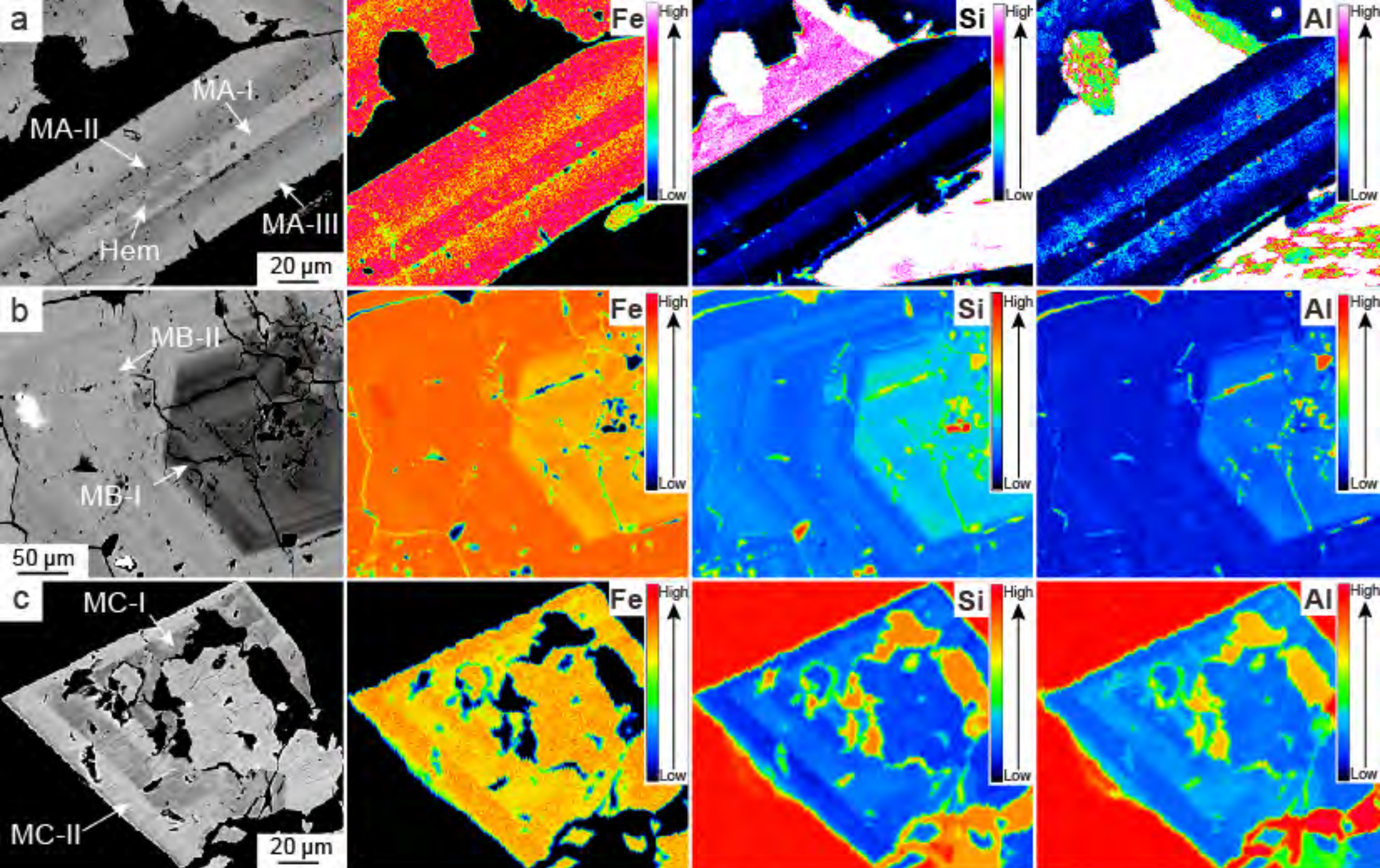


Fig7

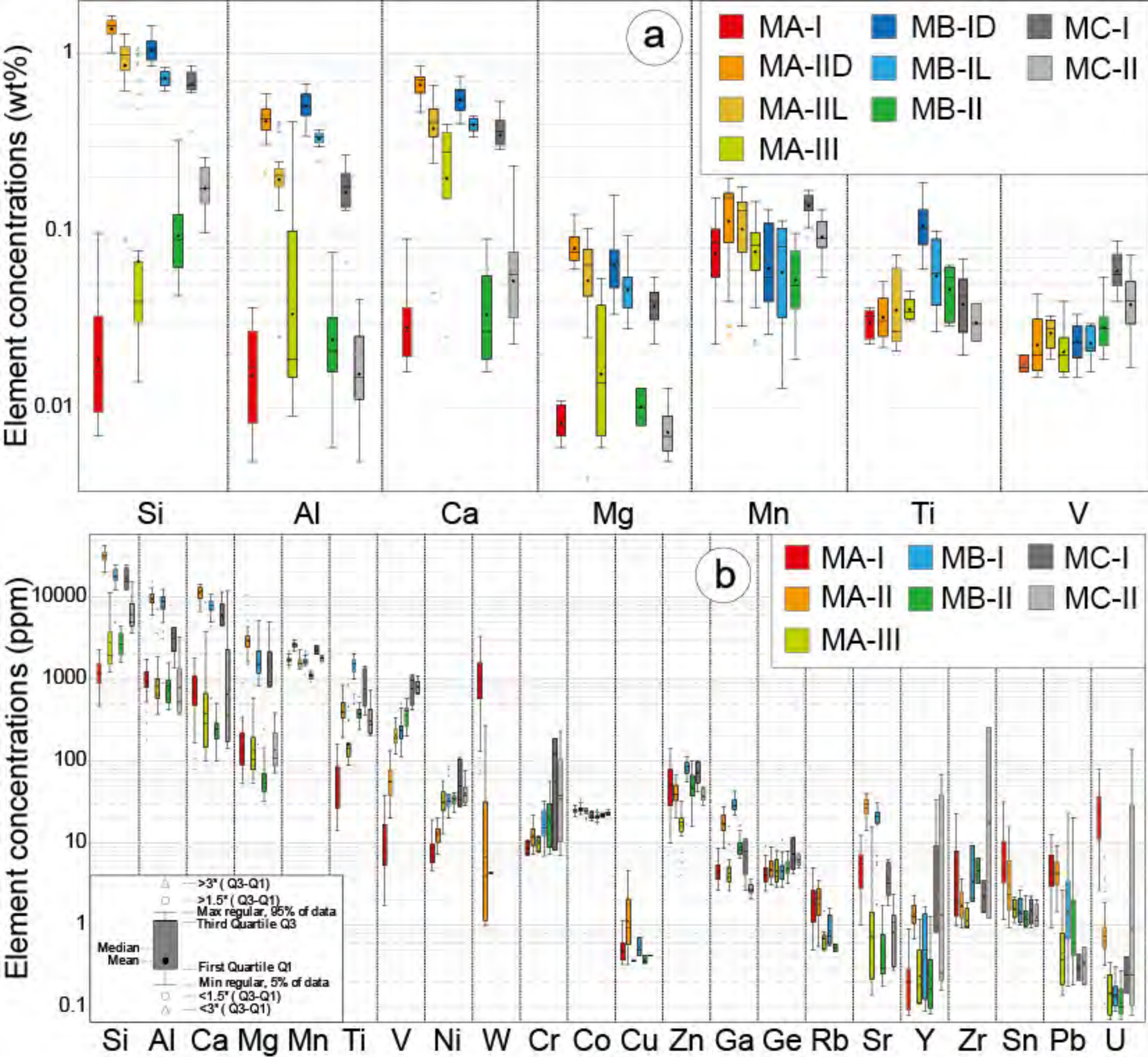


Fig 8

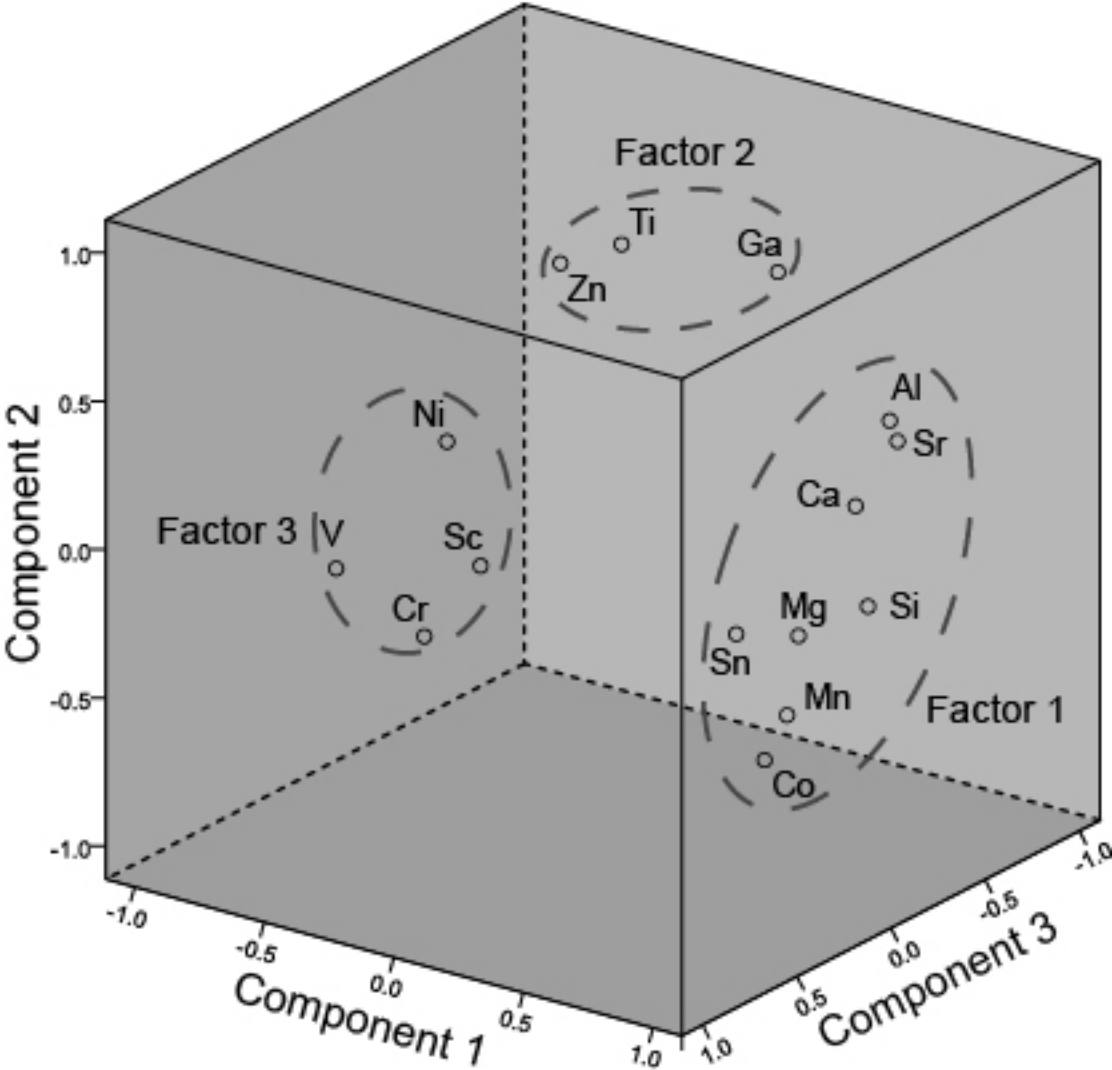


Fig 9

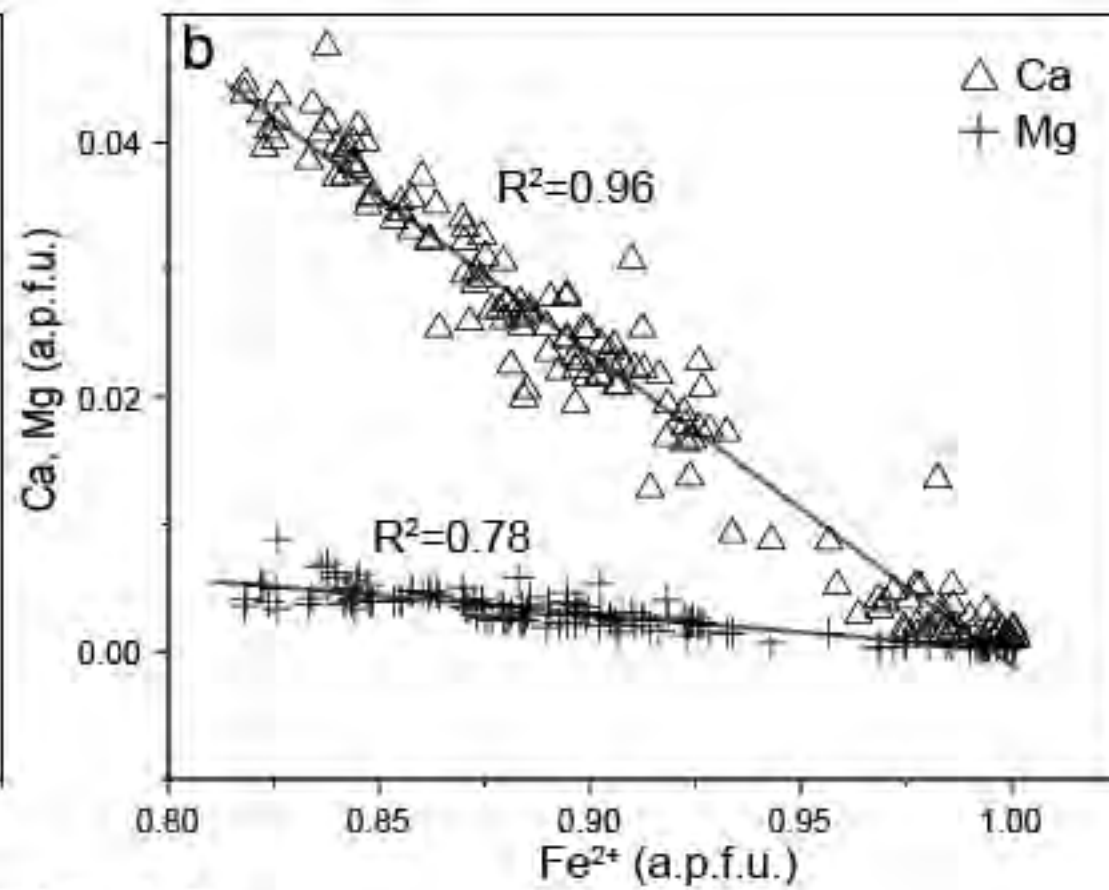
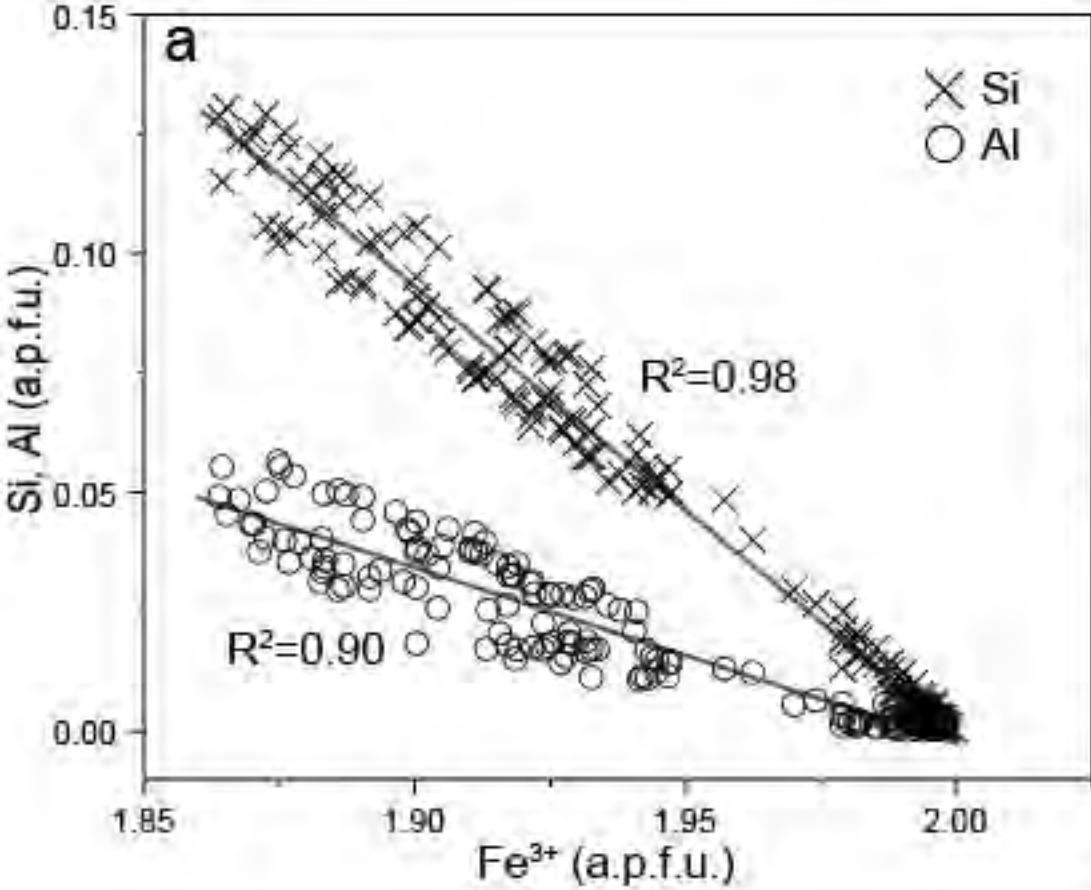


Fig 10

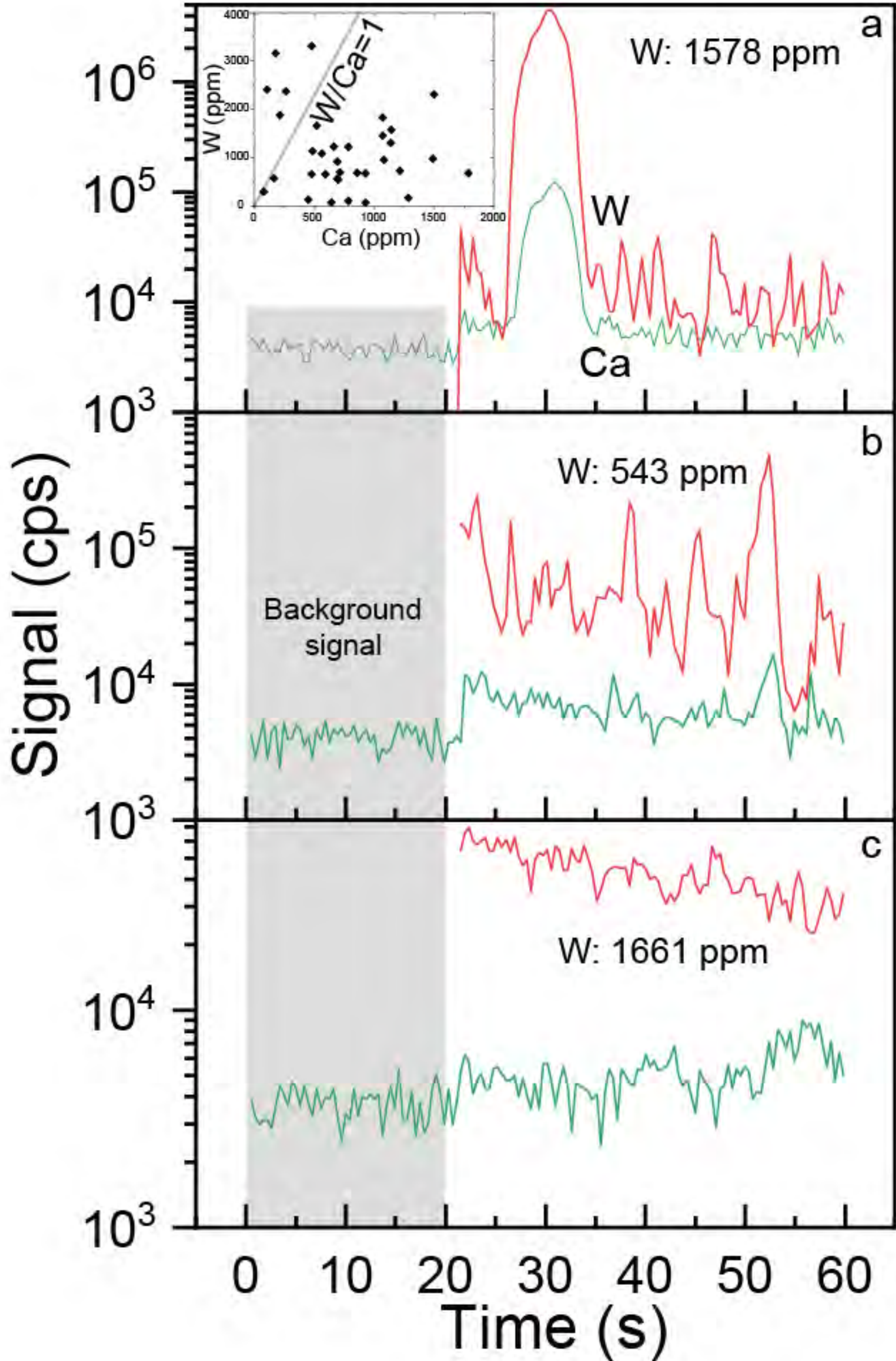


Fig 11

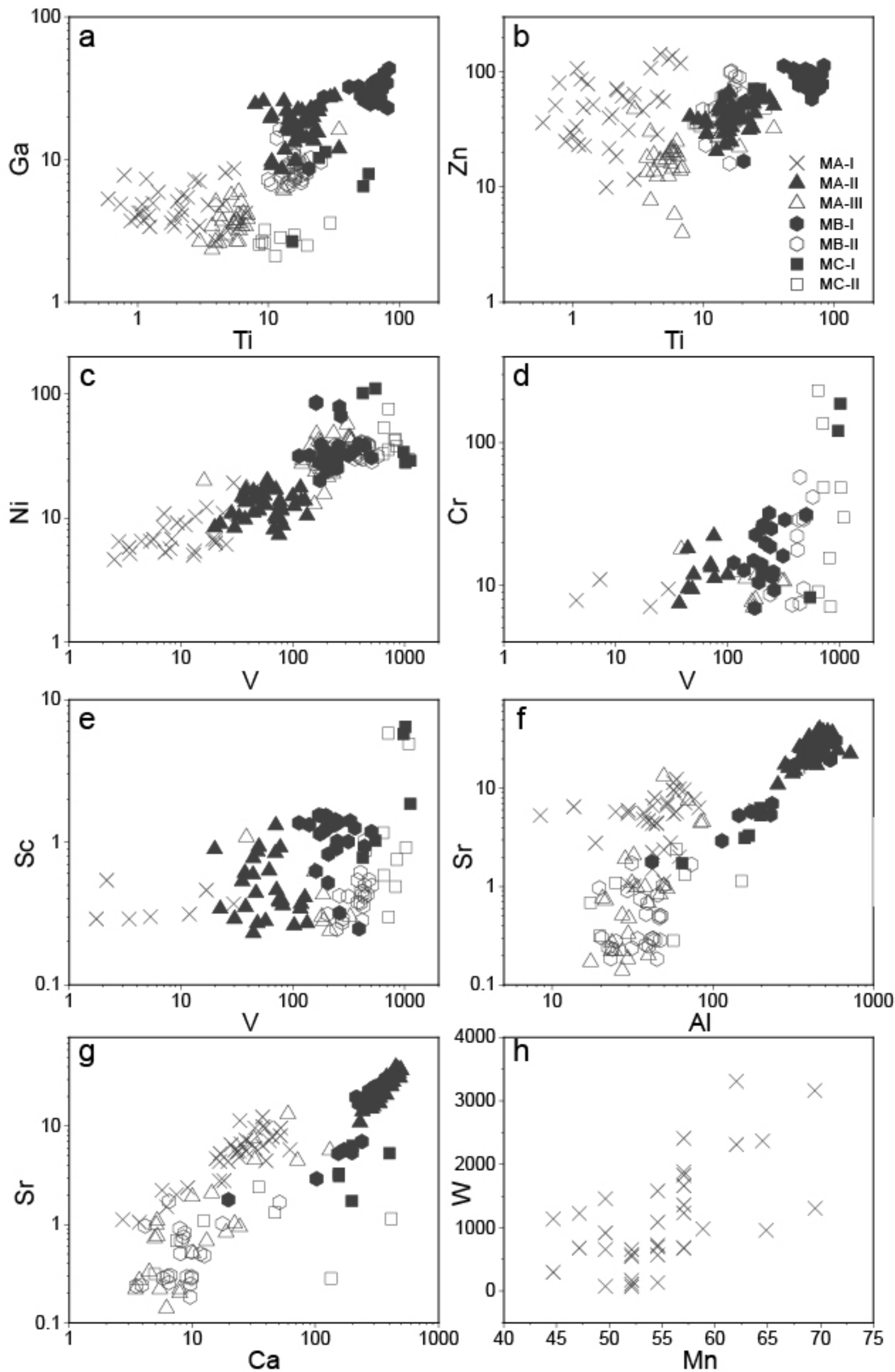


Fig 12

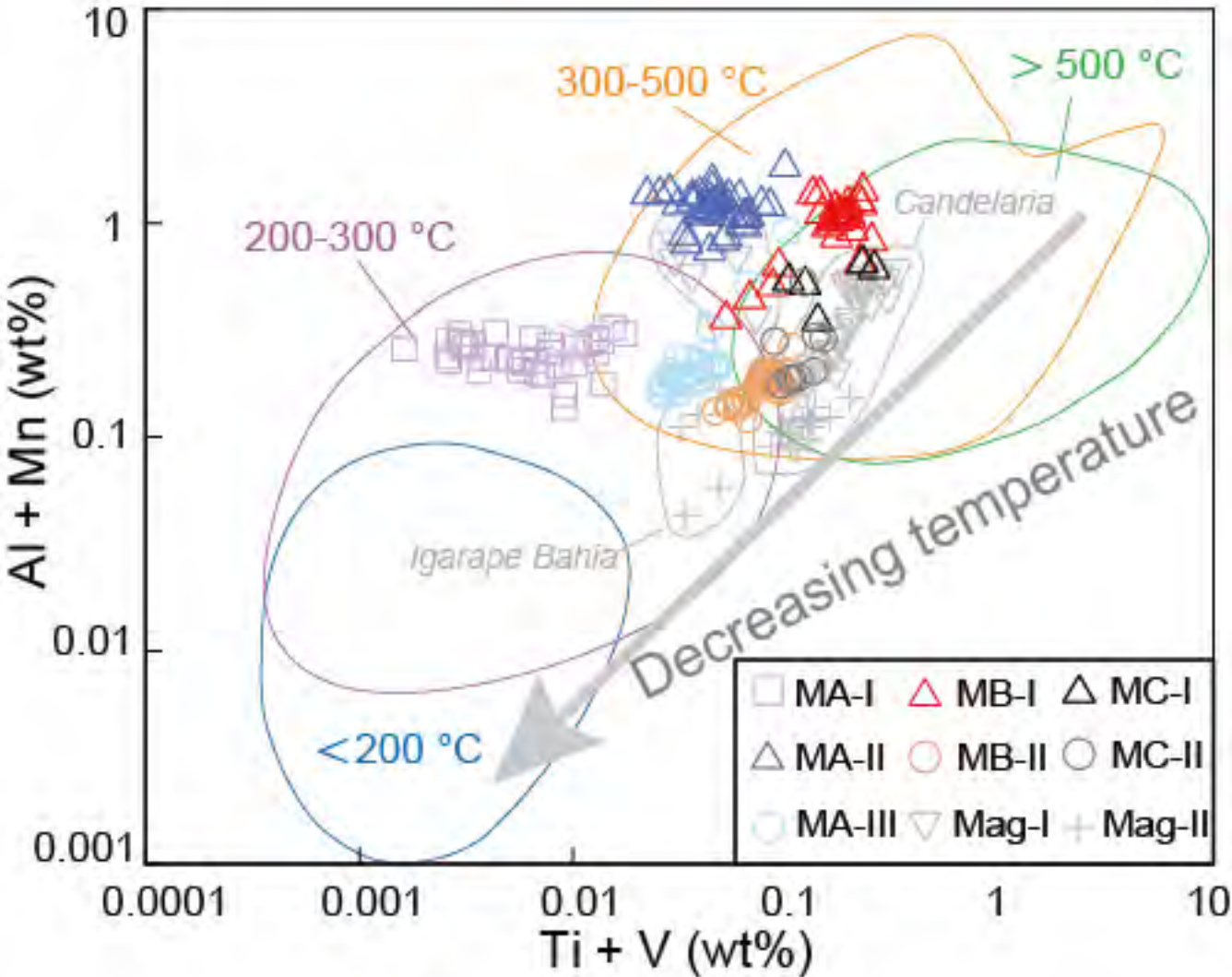


Fig 13

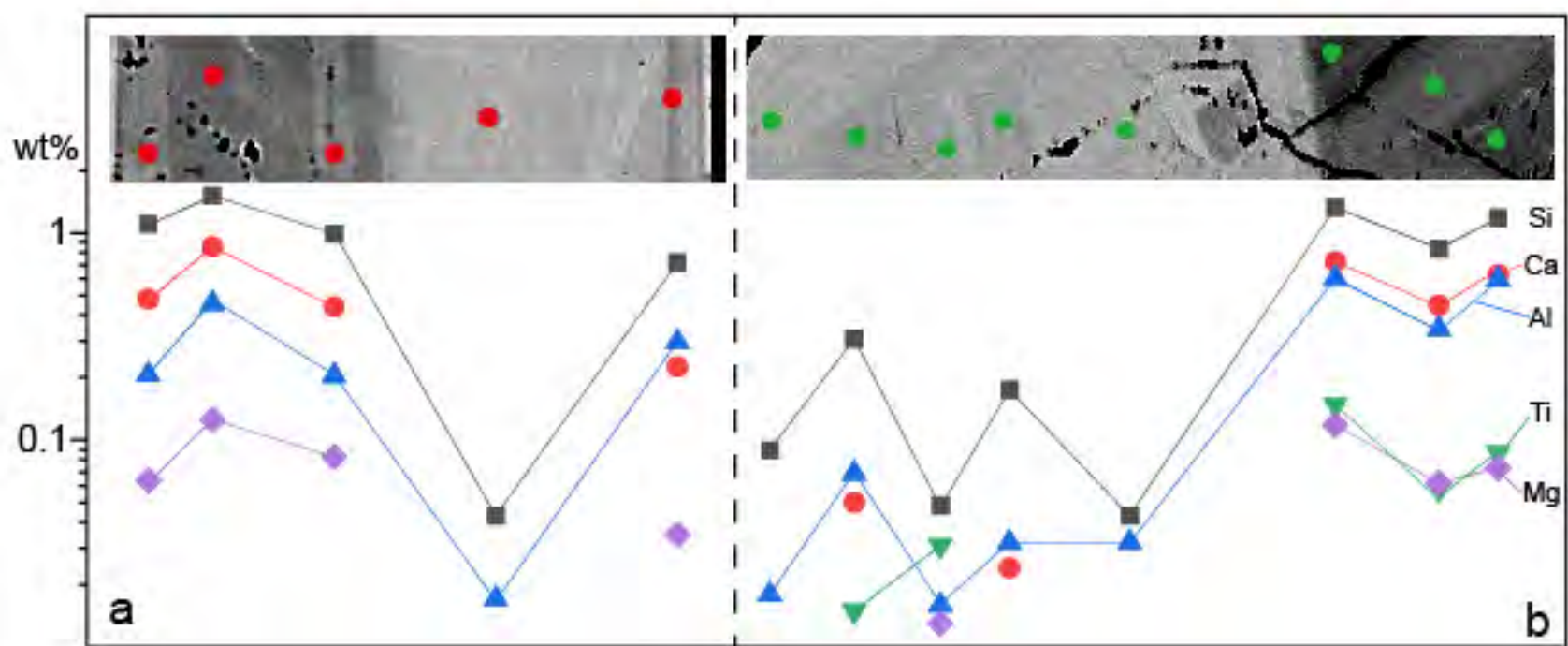


Fig 14