

1 **Revision 2**

2 **Mineralogy and bulk geochemistry of a fumarole at Hverir, Iceland: Analog for acid-**  
3 **sulfate leaching on Mars**

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## Abstract

25

26 Iceland's Námafjall geothermal area exhibits a range of alteration environments. Geochemical  
27 and mineralogical analyses of fumaroles and hot springs interacting with Holocene basaltic lavas  
28 at Hverir, and with Pleistocene hyaloclastites atop nearby Námaskarð, reveal different patterns of  
29 alteration depending on water-rock ratio, degree of oxidation, and substrate composition and age.  
30 The focus of this study is a transect of a Hverir fumarole that has formed a bulls-eye pattern of  
31 alteration of a Holocene basaltic lava flow. Surface samples and samples collected from shallow  
32 pits were analyzed by X-ray Diffraction (XRD), X-ray Fluorescence (XRF), and Scanning  
33 Electron Microscopy (SEM) to constrain changes in mineral assemblage and major elemental  
34 composition with both distance and depth. Elemental sulfur is concentrated near the vent, with  
35 leached deposits with amorphous silica and anatase nearby and kaolinite, hematite, and  
36 jarosite/alunite-group sulfate minerals farther out, with smectites and less altered material at the  
37 margins, though smaller-scale mineralogical diversity complicates this pattern.

38

39 Silica phases include amorphous silica (most samples), cristobalite (some samples, in the leached  
40 part of the apron), and quartz (minor constituent of few samples). The silica was concentrated  
41 through residual enrichment caused by leaching, and is accompanied by a significant enrichment  
42 in TiO<sub>2</sub> (in anatase). The presence of abundant cristobalite in a surface fumarole-altered  
43 Holocene basaltic lava flow most likely reflects cristobalite formed during the devitrification of  
44 volcanic glass or precipitation from fumarolic vapors, rather than high-temperature processes.  
45 Minor, localized quartz likely reflects diagenetic maturation of earlier-formed amorphous silica,  
46 under surface hydrothermal conditions. Natroalunite, natrojarosite, and jarosite are all present

47 and even exhibit compositional zonation within individual crystals, showing that under surface  
48 hydrothermal conditions, these minerals can form a significant solid solution.

49

50 The high iron content of the substrate basalt and the prevalence of Fe-sulfates and Fe-oxide  
51 spherules among the alteration products makes this geothermal area an especially useful analog  
52 for potential Martian hydrothermal environments. The residual enrichment of silica in the  
53 leached deposits of the Hverir fumarole apron could serve as an acid-sulfate leaching model in  
54 which amorphous silica forms without appreciable sulfur-bearing phases in many samples, a  
55 possible analog for silica-rich soils in the Columbia Hills on Mars. The co-existence of hematite  
56 spherules and jarosite-group minerals serves as an intriguing analog for a volcanic/hydrothermal  
57 model for hematite and jarosite occurrences at Meridiani Planum.

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## Introduction

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62 Volcanic hydrothermal deposits are key targets for astrobiological research, since they can  
63 provide warm, wet environments on or below otherwise inhospitably cold planetary surfaces, and  
64 since terrestrial examples provide a habitat for a variety of extremophile microorganisms. The  
65 characterization of likely hydrothermal deposits in the Columbia Hills explored by the Mars  
66 Exploration Rover (MER) Spirit (e.g., Yen et al., 2008), and the detection of likely hydrothermal  
67 deposits from orbit by the Mars Reconnaissance Orbiter (MRO, e.g. Skok et al., 2010), have  
68 made understanding the astrobiological potential of these environments a priority for NASA's  
69 Mars exploration program.

70

71 Iceland provides an excellent analog for potential Martian hydrothermal sites because of the  
72 wide variety of hydrothermal environments coupled with their interaction with high-iron basalts  
73 comparable to those found on Mars. This study describes an analog site in northeastern Iceland  
74 (Hverir), where acid-sulfate fumaroles interact with relatively young basaltic lavas, forming  
75 alteration and precipitated products representing a range of temperature and redox conditions.  
76 The redox gradients observed in these deposits could provide an energy source for iron or sulfur  
77 reducing or oxidizing microbes.

78

79 The objectives of this work were to (1) characterize the mineral assemblages and major element  
80 distribution for altered and precipitated samples at the surface and at shallow depth at increasing  
81 distances from an active fumarole at Hverir, (2) conceptually model the processes of leaching  
82 and mineral precipitation in terms of changing pH, oxidation, and temperature conditions, and  
83 (3) consider these trends in the context of Mars hydrothermal deposits.

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## Background

87

### Mars Hydrothermal Activity

89

90 A long, early history of basaltic volcanism, plus mineralogical and geomorphological  
91 evidence for surface water in the ancient rock record, suggest that volcanic hydrothermal activity  
92 likely occurred on early Mars. Hydrothermal environments on Mars could provide sources of

93 heat, energy, and water for life (Walter & Des Marais, 1993; Schulze-Markuch et al., 2007;  
94 Hynek et al., 2013, 2018), even when it is too dry or cold for life to persist nearby. A  
95 hydrothermal origin has been proposed for some Mars surface features, including the mineralogy  
96 and geochemistry of some outcrops and soils studied by Mars Exploration Rover (MER) Spirit in  
97 the Columbia Hills (e.g. Yen et al., 2008), and potential hydrothermal mineral assemblages  
98 identified from orbit (e.g. Ehlmann et al., 2009, 2011; Skok et al., 2010). The Observatoire pour  
99 la Minéralogie, l'Eau, les Glaces et l'Activité (OMEGA) instrument aboard Mars Express, the  
100 Compact Reconnaissance Imaging Spectrometer for Mars (CRISM) aboard the MRO, the MER  
101 missions, and the Mars Science Laboratory (MSL) Curiosity rover have identified globally  
102 distributed mono- and poly-hydrated Mg/Ca/Fe-sulfates within diverse geologic settings that  
103 include sedimentary and hydrothermal environments (Langevin et al., 2005; Yen et al., 2008;  
104 Nachon et al., 2014). MER Opportunity and Spirit, along with MSL Curiosity, detected iron  
105 oxides and oxyhydroxides that include hematite and goethite (Klingelhöfer et al., 2004; Morris et  
106 al., 2008; Fraeman et al., 2016). Chojnacki and Hynek (2008) attribute some of the widespread  
107 Valles Marineris sulfate deposits to high temperature basalt alteration, and widespread jarosite-  
108 bearing deposits at Syrtis Major could also have been formed hydrothermally (Ehlmann and  
109 Mustard, 2012; Ehlmann and Edwards, 2014). Potential products of Martian hydrothermal  
110 alteration have also been identified in Martian meteorites, particularly nakhlites (e.g. Bridges and  
111 Schwenzer, 2012). Schulze-Makuch et al. (2007) outline targets for Martian hydrothermal  
112 environments, and researchers have proposed hydrothermal activity as a possible explanation for  
113 certain mineralogical features of both Mawrth Vallis (Bishop et al., 2008) and Nili Fossae  
114 (Ehlmann et al., 2009; Viviano et al., 2013), perennial favorites for future landed missions. Gale  
115 crater, the MSL landing site currently under investigation by Curiosity, has abundant sulfates and

116 phyllosilicates (Milliken et al., 2010), likely formed in part by basalt alteration, and Mangold et  
117 al. (2012) have even suggested a hydrothermal origin for some of its deposits. This interpretation  
118 is bolstered by the discovery of the high-temperature SiO<sub>2</sub> polymorph mineral tridymite (Morris  
119 et al., 2016) and disordered adularia and specular hematite (Morris et al., 2020, Rampe et al.,  
120 2020) in the Gale crater sediments. Finally, Martian impact craters represent another class of  
121 features that likely hosted long-lived hydrothermal systems (e.g., Abramov and Kring, 2005;  
122 Osinski et al., 2013; Schwenzer and Kring, 2013).

123

124 In the Columbia Hills, the Spirit rover analyzed soils and outcrops and identified both  
125 mineralogical and geochemical patterns consistent with hydrothermal alteration of basalts. The  
126 Paso Robles class soils are sulfate-rich, with Ca, Mg, and ferric sulfates, along with amorphous  
127 silica and hematite; in the case of Paso Robles and Arad these have been interpreted as likely  
128 fumarolic deposits with lower water/rock ratios, while at the Tyrone locality conditions were  
129 likely wetter (Wang et al., 2008; Yen et al., 2008; Schmidt et al., 2008; 2009; Hausrath et al.,  
130 2013). Deposits rich in amorphous silica (Opal-A) in the Eastern Valley have been variously  
131 attributed to hydrothermal acid-sulfate leaching (e.g. Squyres et al., 2008; Morris et al., 2008) or  
132 sinter precipitation in a near-neutral hot spring environment (e.g. Ruff et al., 2011; Ruff and  
133 Farmer, 2016), though morphological and stratigraphic considerations favor the latter (Ruff et  
134 al., 2020). Reconciling these different hydrothermal interpretations is critical for assessing  
135 potential habitability.

136

137 Hematite (as spherules) and sulfates (including jarosite) are also abundant in the Burns  
138 Formation, investigated by the MER Opportunity rover at Meridiani Planum. While most often

139 interpreted as a sedimentary deposit (hematite precipitated as nodules in the subsurface by the  
140 action of groundwater, sulfates produced through evaporitic processes), a volcanic or  
141 hydrothermal origin has also been suggested (McCollom and Hynek, 2005, 2006, 2021). The  
142 presence of jarosite in particular, as identified based on Mössbauer spectroscopy, has been used  
143 to argue for overall acidic and oxidizing conditions, given the limited stability field of jarosite.  
144 Some studies have questioned this interpretation, based on the presence of jarosite in non-acidic  
145 environments on Earth (e.g. saline-alkaline paleolake Olduvai, Tanzania: McHenry et al., 2011)  
146 and concerns about the ability of Mössbauer to unambiguously distinguish between jarosite and  
147 iron-bearing natroalunite, which has a less restrictive stability field (McCollom et al., 2013b).

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## 149 **Iceland**

150

151 The Krafla central volcano in northern Iceland lies above the Icelandic mantle plume at  
152 the Mid-Atlantic Ridge within the Northern Volcanic Zone. Krafla volcano forms a low, broad  
153 shield (~25 km diameter); with an 8 by 10 km caldera bisected by a 100 km en echelon NNE-  
154 trending fissure swarm (Figure 1a) (Gudmundsson & Arnorsson, 2002, 2005). The Krafla central  
155 volcano exhibits a distinct dacitic welded-tuff caldera rim that formed ~110 ka ago (Björnsson et  
156 al., 1977). The caldera has been partially in-filled with predominantly basaltic eruptive products  
157 with lithologies dominated by hyaloclastite, subaerial lava flows, and fine-grained tuffs, while  
158 edifices of icelandite, dacite, and rhyolite have also been emplaced (Stefansson, 1981;  
159 Saemundsson, 1991). Pleistocene hyaloclastites formed during subglacial eruptions, while  
160 Holocene (3 ka – present) subaerial lavas were emplaced during six major eruptions with an  
161 interval period of 250-1000 years (Saemundsson, 1991). This study focuses on Holocene (post-

162 glacial) basalts and Pleistocene (subglacial) basaltic hyaloclastites currently undergoing intense  
163 surface hydrothermal alteration in the Námafjall area.

164

165 The Námafjall geothermal area lies at the southern edge of the Krafla volcano (Figure 1), where  
166 hydrothermal fluids reach the surface along faults. These hydrothermal fluids interact with  
167 basaltic hyaloclastites atop and along the margins of the Námaskarð hyaloclastite ridge, and with  
168 late Holocene basaltic lava flows at Hverir at its base. The Hverir geothermal field lies to the east  
169 of Námaskarð and is dominated by vigorous fumaroles along a fault, with abundant hot springs  
170 and mudpots on the downthrown side (Saemundsson et al., 2012). These hot springs and  
171 mudpots interact with surface materials derived from local drainages, including from the  
172 neighboring Námaskarð hill. Large circular fumarolic aprons (up to ~25 m diameter) that  
173 produce little to no volcanic vapors are scattered across the Late Holocene lava field nearby. One  
174 of these circular fumarolic aprons is the focus of the current study.

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## Methods

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### Site details

180

181 At the eastern edge of the Hverir hydrothermal area (Figure 1c), individual fumaroles emerge  
182 through Holocene lavas. This interaction leads to a concentric “bull’s-eye” fumarolic apron, with  
183 higher temperatures and more significant sulfur precipitation and alteration at the center (Figure  
184 2). At the largest of these fumaroles, we collected surface samples along a transect (targeting



185 obvious changes in color or appearance) in 2013 (samples I-V-13 5 through 14), and returned in  
186 2014 to excavate shallow pits to collect samples at depth (Figure 2) and in 2016 for more  
187 detailed observations. From these shallow pits, we collected both altered soil and pieces of  
188 altered basaltic lava (abbreviated PAB, samples I-V-14 6 through 20). Temperature, distance  
189 from fumarole center, and depth (if appropriate) are reported for each sample.

190

191 All samples analyzed in this study are summarized in Table 1. Precipitate samples were collected  
192 by carefully scraping surface materials over a small area into a plastic sample bag, and altered  
193 soils were scooped using a spatula. Pieces of altered basalt (PABs) were collected along with  
194 enclosing soil from the pits. Temperatures were measured for each sample using a thermometer.  
195 A less altered “substrate” sample was collected from the same lava flow, away from the fumarole  
196 field.

197

## 198 **Laboratory Analyses**

199

200 Samples were air dried and ground using an agate mortar and pestle without the addition of  
201 liquids. The substrate (basalt) sample was powdered first using a Rocklabs Shatterbox, then  
202 powdered more finely by hand. Samples were mounted as random powders for X-ray Diffraction  
203 (XRD) and analyzed at UW Milwaukee (UWM) using a Bruker D8 Focus XRD (Cu tube,  $0.02^\circ$   
204  $2\theta$  step size,  $2-60^\circ 2\theta$ , 1 s/step, scintillation detector: see McHenry et al., 2017). Phases were  
205 identified using Bruker’s EVA software and the International Centre for Diffraction Data Powder  
206 Diffraction Files (ICDD PDF) 2 database for comparison. Presence, absence, and qualitative  
207 relative abundance (trace vs. major) were estimated by comparing peak heights.

208

209 Samples that were determined not to be dominated by sulfur-bearing phases based on XRD were  
210 further prepared for X-ray Fluorescence (XRF) analysis. Powdered samples were dried overnight  
211 at 105°C. A split of each dried sample was analyzed for loss on ignition (LOI) using a muffle  
212 furnace. Another split (1.000 g) of the same dried powdered sample was then mixed with 10.000  
213 g of a 50/50 Lithium Metaborate/ Lithium Tetraborate flux with an integrated LiBr non-wetting  
214 agent and ~1 g of ammonium nitrate (oxidizer) in a platinum crucible and fused into a glass bead  
215 using a Claisse M4 fluxer. Beads were then analyzed for major and minor elements using a  
216 Bruker S4 Pioneer XRF, following methods of Byers et al. (2016). Concentrations were  
217 calculated using a calibration curve based on eleven USGS rock standards. Since sulfur is  
218 partially lost during the fusion process, select samples for which sufficient material was available  
219 were also prepared as pressed pellets. 7.5 grams of dried, powdered sample was combined with  
220 three GeoQuant wax binder pellets (0.94 g total) in a shatterbox, and then pressed at 25 tons for  
221 one minute in an Atlas T25 semiautomatic press. Pellets were then analyzed by XRF with  
222 concentrations determined using a calibration curve based on six USGS rock standards. Only  
223 sulfur is reported from the pressed pellets, as the fused bead calibration is better for all other  
224 elements. Methods, limits of detection, and analytical errors associated with both calibrations are  
225 described in detail in Byers et al., 2016).

226

227 Seven thin sections and small chips of fresh and altered basalt pieces were analyzed by Scanning  
228 Electron Microscopy (SEM). Thin sections and small chips of altered basalt were carbon coated  
229 using an Edwards Coating System E306A and placed into a geologic thin section holder, and  
230 small chips of altered basalt were mounted on an aluminum stub and embedded in epoxy to

231 expose a specific vesicle in-filling or exterior surface of interest. Samples were analyzed using a  
232 S-4800 Hitachi at UW Milwaukee, using a cold cathode field emitter in secondary electron or  
233 backscattered electron mode with an accelerating voltage of 15 kV, 10 uA emission current, and  
234 high probe current. Elemental compositions were determined qualitatively using an attached  
235 Energy-Dispersive X-Ray Spectrometer (EDS) and Bruker software.

236

237

## Results

238

### 239 XRD results

240

241 XRD results (Table 1) reveal diverse mineral assemblages that vary depending on the specific  
242 environment. Sulfur-bearing minerals are abundant and include a wide range of sulfates (Ca (e.g.  
243 gypsum),  $\text{Fe}^{3+}$  (e.g. rhomboclase  $((\text{H}_5\text{O}_2)^+\text{Fe}^{3+}(\text{SO}_4)_2 \cdot 2(\text{H}_2\text{O}))$ , ferricopiapite  
244  $((\text{Fe}^{3+}_{2/3}\text{Fe}^{3+}_4(\text{SO}_4)_6(\text{OH})_2 \cdot 20(\text{H}_2\text{O}))$ , and mixed cation (e.g. natroalunite  $(\text{NaAl}_3(\text{SO}_4)_2(\text{OH})_6)$ ,  
245 natrojarosite  $(\text{NaFe}^{3+}(\text{SO}_4)_2(\text{OH})_6)$ , jarosite  $(\text{KFe}^{3+}(\text{SO}_4)_2(\text{OH})_6)$ , and minamiite  
246  $((\text{Na,Ca,K})\text{Al}_3(\text{SO}_4)_2(\text{OH})_6)$  varieties), and elemental sulfur. Other phases include silica phases  
247 (amorphous silica, opal A, quartz, and cristobalite), oxides and hydroxides (anatase, hematite,  
248 goethite), clay minerals (smectite, kaolinite), and residual igneous phases (plagioclase, augite).  
249 As the clay size fraction was not separated for analysis, our characterization of the clay phases is  
250 limited to distinguishing between “smectite” (clays with first-order basal spacing around 14  
251 angstroms, based on XRD) and kaolinite.

252

253 Figures 3 and 4 show representative XRD patterns that highlight the diversity in alteration  
254 mineralogy in the Hverir fumarole transect. Fumarolic alteration of the basaltic lava substrate  
255 and precipitation from fumarolic vapors, acting in conjunction with meteoric water, produced the  
256 following mineral assemblages: (i) elemental sulfur, at the fumarole vent and in lenses at greater  
257 distances, (ii) SiO<sub>2</sub>-rich soils that include amorphous silica + anatase ± cristobalite ± quartz,  
258 sampled within ~2 meters of the vent, (iii) purple sediments that include natroalunite or jarosite ±  
259 hematite ± amorphous silica ± kaolinite, and (iv) red sediments that include hematite ± goethite ±  
260 jarosite. Other sulfates observed in samples collected near the Hverir fumarole include gypsum  
261 and minamiite (rarely observed, but both abundant in two samples each), and ferricopiapite and  
262 rhomboclase (only observed in one sample each). Along the margins of the acid-sulfate fumarole  
263 are clay-rich sediments that include smectite ± kaolinite ± anatase ± hematite and/or goethite.

264

## 265 **XRF results**

266

267 Altered rocks and soils in the vicinity of the fumaroles show a range of compositions (OM  
268 Tables 2-3), reflecting their formation environments and mineral assemblages. Heavily leached  
269 soils show elevated SiO<sub>2</sub> and TiO<sub>2</sub> concentrations (up to 93 wt% and 23 wt%, respectively),  
270 while red soils are elevated in Fe<sub>2</sub>O<sub>3</sub> (up to 33 wt%). The less altered basaltic lava substrate had  
271 49 wt% SiO<sub>2</sub> and relatively high iron (15.8 wt% Fe<sub>2</sub>O<sub>3</sub>T). TiO<sub>2</sub> for the substrate sample was 1.82  
272 wt%, much lower than the TiO<sub>2</sub> observed in most of the altered samples. Sulfur was only  
273 analyzed for a subset of samples, but where analyzed shows enrichment in altered samples (0.2-  
274 5.4 wt% SO<sub>4</sub>) compared to the substrate (0.11 wt% SO<sub>4</sub>). Most cations (Na, K, Ca, Mg, Mn) are  
275 lower in the altered rocks and soils than in the substrate. These trends are shown in a spider plot,

276 in which the composition of altered basalt samples (PABs) are plotted against the less altered  
277 substrate (Figure 5). Iron (reported at  $\text{Fe}_2\text{O}_3\text{T}$ ) is variable throughout the study area, as reflected  
278 by the colors observed during sampling.

279

280 Samples from the Hverir transect were subdivided into sediment, rock, and precipitate samples,  
281 and pieces of altered basalt (designated PABs) that were collected along with surrounding  
282 sediments. The composition of the PABs reflects the alteration of the basaltic lava in contact  
283 with the fumarolic environment, and are plotted (in a spider plot) against the composition of the  
284 basaltic lava in Figure 5. Differences in composition on a small spatial scale are evident for  
285 individual samples that were subdivided (after collection) based on color and texture differences  
286 observed within individual collected samples, in addition to expected differences between PABs  
287 and their enclosing sediments. For example, samples IV-13-10 and IV-13-14 PAB were  
288 separated into sub-samples IV-13-10W (much higher in  $\text{SiO}_2$ ) and IV-13-10MIX (containing  
289 elemental sulfur, and higher  $\text{Fe}_2\text{O}_3$ ), and IV-13-14 PAB, where the pinker portion of the altered  
290 rock (IV-13-14 PAB W+P) is higher in  $\text{Fe}_2\text{O}_3\text{T}$ , at 25.16 wt%, and contains abundant hematite,  
291 compared to IV-13-14 PAB MIX, which has 9.00 wt%  $\text{Fe}_2\text{O}_3\text{T}$ , less than the substrate basalt. In  
292 Figure 6, samples collected at the same distance from the fumarole are assigned the same symbol  
293 (most come from different depths within the same pit, representing different colored layers or  
294 lenses). The spread of compositions from samples from the same sampling locations  
295 demonstrates small-scale variability that is largely independent of temperature and distance from  
296 the fumarole.

297

298 **SEM/EDS results**

299

300 SEM/EDS analysis provided a detailed geochemical and mineralogical analysis for pieces of  
301 altered basalt (PABs) from the Hverir fumarole transect for four distinct alteration groups: (i)  
302 PABs in silica-rich soils (IV-13-7), (ii) PABs in purple sediments (IV-13-14 PAB), (iii) PABs in  
303 red sediments (IV-13-13), and (iv) a slightly altered basalt along the distal margin (IV-13-11).  
304 Examination of the interior of altered basalt shows varying degrees of alteration of primary  
305 igneous textures and the precipitation of secondary minerals.

306

307 Pieces of altered basalt within silica-rich soils and colored sediments showed extensive alteration  
308 and had few if any remaining primary phenocrysts, which were completely or partially replaced  
309 by amorphous silica. In sample IV-13-7, a PAB from a red-orange soil, a small section of  
310 primary minerals remained and exhibited extensively etched surfaces along augite and  
311 plagioclase crystals (Figure 7a). Qualitative EDS results showed relict plagioclase laths with Al  
312 and Si, while mobile elements Ca and Na were depleted. Titanomagnetite retained its skeletal  
313 and cruciform texture, but with a notable depletion in Fe (referred to as ‘ghost’ texture by Urcia  
314 et al., 2010). In one instance in this sample, FeOx spherules are observed near the cruciform  
315 texture of titanomagnetites (Figure 7b).

316

317 Basaltic glass is completely absent in the silica-rich soils, whereas abundant sideromelane with  
318 etched surfaces partially replaced by amorphous silica and possibly clays and iron oxides/  
319 hydroxides is observed in the colored sediments (Sample IV-13-13, Figure 8a,b). In Figure 8b,  
320 the dark layer near the rim of the sideromelane fragments likely represents palagonite, i.e. poorly  
321 crystalline smectite-like material, based on its texture. In the less altered deposits at the margin

322 of activity (e.g. IV-13-11), basaltic glass is still preserved with minimal surface etching (Figure  
323 8e). Secondary minerals observed by SEM for more altered samples were commonly associated  
324 with amorphous silica (Figure 9a,b for samples IV-13-13 and IV-13-14 PAB). The most  
325 frequently observed secondary minerals were members of the alunite-jarosite group and  
326 hematite. A microscopic needle-like barium and sulfur bearing mineral, presumably barite, was  
327 sparsely distributed in amorphous silica veins and surface coatings in the same sample (IV-13-  
328 13, Figure 8c). This sample also contains gypsum in amorphous silica veins (Figure 8d). XRD  
329 patterns for sample IV-13-13 did not show evidence for barite or gypsum, likely due to their  
330 minor abundance.

331

332

## Discussion

333

### Distribution of phases, based on environment

335

336 The Hverir fumarole transect revealed different mineral assemblages, following a coarse-scale  
337 concentric pattern from most altered (center) to least altered (margin), with local small-scale  
338 variation. The sample collected at the central fumarole vent consisted of almost pure elemental  
339 sulfur, surrounded by leached deposits dominated by amorphous silica with anatase (enriched in  
340  $\text{SiO}_2$  and  $\text{TiO}_2$  and depleted in mobile elements). Elevated  $\text{Fe}_2\text{O}_3$ , associated with abundant  
341 hematite and goethite, is observed in some samples, primarily at greater distance from the vent  
342 (e.g. 33.46 wt% in IV-13-13, at 550 cm), though lenses of colored sediment with higher hematite  
343 and/or goethite are observed closer in (e.g. minor hematite in IV-13-7, at 120 cm). Figure 6 plots  
344 wt%  $\text{Fe}_2\text{O}_3$  against other oxides, showing how it is depleted in  $\text{SiO}_2$ -rich samples and enriched

345 in P<sub>2</sub>O<sub>5</sub>-rich samples. Trends of wt% Fe<sub>2</sub>O<sub>3</sub>T vs. MgO, K<sub>2</sub>O, and Al<sub>2</sub>O<sub>3</sub> vary depending on the  
346 samples, but overall show increasing concentrations with increasing Fe<sub>2</sub>O<sub>3</sub>T. Clay minerals  
347 (including kaolinite and smectites) are rare in the vicinity of this fumarole, occurring in only the  
348 most distal pit (IV-14-17 and 18, at 800 cm), and are associated with elevated Al<sub>2</sub>O<sub>3</sub>. Sulfate  
349 minerals occur in about 1/3 of the samples from the fumarole transect, and include the Ca-sulfate  
350 mineral gypsum, members of the alunite-jarosite series (natroalunite, natrojarosite, and jarosite),  
351 Fe<sup>3+</sup> sulfate minerals ferricopiapite and rhomboclase, and the mixed-cation, Al-rich sulfate  
352 mineral minamiite. Compared to the least altered sample of the basaltic substrate, all samples  
353 from the transect analyzed for sulfur have significantly elevated SO<sub>4</sub>, even those for which  
354 sulfur-bearing phases (sulfur, sulfates) were not identified by XRD. This overall trend shows  
355 more reducing conditions at the fumarole vent (with elemental sulfur), while sulfates and oxides  
356 further from the vent reveal more oxidizing conditions. The amorphous silica and anatase rich  
357 zone surrounding the central vent reveals significant leaching and removal of most mobile  
358 elements. No systematic differences were observed between samples collected at depth and those  
359 collected near the surface, suggesting no significant change in redox conditions with shallow  
360 depth in the fumarole apron.

361

## 362 **Element Mobility**

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364 The mobility of elements during alteration of basaltic lava by the Hverir fumarole can be traced  
365 by comparing the bulk compositions of pieces of altered basalt collected from shallow depths  
366 within the fumarole transect (the 2013 “PAB” samples) against the composition of unaltered  
367 basalt from the same lava flow (Fig 5). The sample farthest from the vent is the closest in



368 composition to the unaltered basalt, as expected. Sulfur is concentrated in all altered samples for  
369 which sulfur was analyzed, added to the system by the fumarolic vapors. Also enriched in all but  
370 the least altered sample are Si, Ti, and Zr, consistent with residual enrichment after acid-sulfate  
371 leaching or mobile components. Al is depleted in all but the least altered sample, consistent with  
372 its higher mobility in acidic systems. Except for the least altered samples near the margin, all  
373 samples are depleted in most mobile elements, particularly Na, Ca, Mn, and Mg. K follows this  
374 trend for some samples but retains its original concentration in samples that contain jarosite. Fe  
375 varies from highly depleted to moderately enriched- a trend shown even more clearly in Fig 6. It  
376 is most enriched in those samples that contain hematite as measured by XRD.

377

378 These trends show the strong influence of the acid-sulfate leaching environment. Extensive  
379 leaching of mobile components (that are not included in the sulfur-bearing phases being  
380 precipitated) leads to the residual enrichment of the least mobile elements- in this case Si, Ti, and  
381 Zr. The Ti enrichment is exceptional- altered soil sample IV-13-8 has 23.16 wt% TiO<sub>2</sub> (though  
382 the associated piece of altered basalt, IV-13-8PAB, has only 4.65 wt%), compared to 1.82 wt%  
383 in the substrate basalt. Anatase is recognized in the XRD patterns of most samples, in some cases  
384 (e.g. IV-13-6, 9, and 12) as the only crystalline phase in samples otherwise dominated by  
385 amorphous silica.

386

### 387 **Alunite-Jarosite Group**

388

389 Members of the alunite-jarosite group are the most abundant sulfate phases observed in the  
390 Hverir fumarole transect. They were identified in XRD patterns and by SEM (for samples IV-13-

391 3, IV-13-13, and IV-13-14 PAB), where they were observed to have distinct pseudo-cubic  
392 crystals, ranging from <5 to 20  $\mu\text{m}$  in diameter (Figure 9). These minerals line interior walls of  
393 vesicles associated with amorphous silica and spheroidal FeOx. In other instances, dense  
394 aggregates partially fill vesicle interiors (e.g. IV-13-14 PAB).

395

396 The alunite-jarosite group minerals have an idealized chemical formula [ $\mathbf{AB}_3(\text{SO}_4)_2(\text{OH})_6$ ] where  
397 the A site is occupied by monovalent species  $\text{K}^+$ ,  $\text{Na}^+$ ,  $\text{H}_3\text{O}^+$ , while the B site is occupied by  
398 trivalent species  $\text{Al}^{+3}$  and  $\text{Fe}^{+3}$ . When  $\text{Fe} > \text{Al}$  the mineral is classified as jarosite, while those  
399 minerals with  $\text{Al} > \text{Fe}$  are in the alunite family. EDS provides reliable measurements for the  
400 alunite-jarosite compositions for the relative proportions of Fe and Al, but underestimates the  
401 amount of Na and K in the sample (see McCollom et al., 2013a). The qualitative chemical  
402 composition at Námafjall ranges from natroalunite to Al-bearing jarosite, and often shows  
403 enrichment of Fe and K in the crystal rims (Figure 10). Significant compositional zoning of the  
404 A site was also observed within individual crystals. Polished grain mounts showed similar  
405 pseudo-cubic morphology despite differences in chemical composition.

406

407 The alunite-jarosite group is known to range in composition as a result of significant solid  
408 solution mixing in the A site, while substitution of the B site is less common. Published data  
409 reports compositions that lie close to Fe and Al end-members, which likely reflects a miscibility  
410 gap between alunite and jarosite (Stoffregen et al., 2000; Papike et al., 2006). Possible  
411 explanations for the miscibility gap include differences in hydrolysis constants for dissolved Fe  
412 and Al species or oxidation state during deposition (Stoffregen et al., 2000; Papike et al., 2006,  
413 2007). However, a few studies have reported intermediate compositions between natroalunite

414 and Al-bearing jarosite in zoned pseudocubic crystals that were found in the early stages of acid-  
415 sulfate alteration of basalt in natural systems (Morris et al., 2005; McCollom et al., 2013a; Black  
416 et al., in revision).

417

418 Compositional zonation of the A site in Hverir samples likely reflects local and temporal  
419 variability in ambient fluid compositions during mineral precipitation (e.g. Juliani et al., 2005;  
420 Papike et al., 2006). Although local basalts lack significant K-bearing minerals, the abundance of  
421 jarosite can be attributed to the release of K from altered glass (e.g. IV-13-14 PAB). Na-jarosite  
422 was associated with orange precipitates in fumarolic settings. The low water/rock ratio near the  
423 fumarole (as opposed to hot springs) can allow Na and other soluble elements to be more readily  
424 retained in the products of alteration.

425

426 The substitution of Fe and Al in the B site is likely attributed to local variation in the chemical  
427 environment caused by the decomposition of Al-bearing (plagioclase) and Fe-bearing (olivine,  
428 augite, primary oxide) minerals. This is supported by the wide compositional variation within an  
429 individual sample. On the other hand, the enrichment of Fe in the crystal rims may reflect a  
430 transition to more oxidizing conditions, which makes more ferric Fe available for precipitation in  
431 the minerals (Papike et al., 2006). Regardless, EDS measurements record significant Fe-for-Al  
432 substitution within the crystal structure in alunite-jarosite group minerals during acid-sulfate  
433 alteration at Hverir, which corresponds well to alunite-jarosite observed in synthetic and natural  
434 basaltic systems (Morris et al., 2005; McCollom et al., 2013a,b; Black et al., in revision).

435

436 Members of the alunite-jarosite group were often associated with spheroidal hematite. This  
437 observation suggests that the alunite-jarosite group may be unstable over time, eventually  
438 breaking down and forming more stable spheroidal hematite (Desborough et al., 2010;  
439 McCollom et al., 2013b).

440

#### 441 **Spheroidal Hematite**

442

443 Abundant spheroidal Fe-oxides were observed in sulfate-rich fumarolic deposits (Figures 7, 9,  
444 and 11). In samples IV-13-13 and 14 PAB, two FeOx populations were observed with similar  
445 spheroidal shape: (i) micrometer-sized (typically ~1-3 $\mu$ m diameter) spherules observed forming  
446 dense clusters around basaltic glass fragments and relict plagioclase and (ii) larger (~5 to 20  $\mu$ m  
447 diameter) individual spheroids that often form botryoidal clusters (Figure 11). FeOx spherules  
448 lining vesicles or exterior surfaces (see Figure 11a,b) are interpreted as hematite, which is  
449 abundant in the XRD patterns, though goethite is also identified in lesser abundances in sample  
450 IV-13-13. Additional spherules within the interior of the basalt were observed (e.g. Figure 11c).  
451 All examined FeOx spherules contain impurities of Al, Si, S, and P that may serve as a  
452 nucleation site for mineral formation (Golden et al., 2008). Spheroidal FeOx associated with the  
453 alunite-jarosite group and amorphous silica lining vesicle walls and exterior surfaces are a  
454 commonly found mineral assemblage in volcanic acid-sulfate environments (Morris et al., 2005;  
455 Bishop et al., 2007; McCollom et al., 2013a). Additionally, these FeOx spherules exhibit a radial  
456 growth pattern consisting of rods or fibers (see Figure 11d). Morphology, elemental composition,  
457 and association with Fe-bearing sulfates support the interpretation that spheroidal hematite in the  
458 fumarolic deposits is related to aqueous precipitation during acid-sulfate alteration of basalt.

459

460 The occurrence of spherules at Hverir provides direct evidence for the formation of spheroidal  
461 hematite during acid-sulfate alteration, but it is unknown whether hematite formed from the  
462 dissolution of basaltic precursors or through a multi-step process. The latter involves a simplified  
463 two-step process: (i) the formation of Fe-bearing sulfates (e.g. jarosite) and (ii) the dissolution of  
464 sulfates to form spheroidal hematite through forced hydrolysis. Morris et al. (2005) noted the  
465 apparent absence of hematite in altered basaltic tephra at Mauna Kea where hydrothermal  
466 solutions did not form sulfate minerals. This observation implied that the formation of jarosite  
467 and Fe-bearing natroalunite under acid-sulfate conditions was a necessary precursor for the  
468 formation of hematite spherules.

469

470 Published laboratory experiments support the multi-step hypothesis. Several studies have shown  
471 that hematite spherules can directly precipitate from aqueous solutions during forced hydrolysis  
472 at ~100°C (e.g. Kandori et al., 2000). Spheroidal growth is affected by a variety of factors that  
473 include: anions (SO<sub>4</sub>, Cl), super-saturation, extent of hydrolysis, viscosity of the medium, and  
474 impurities (Golden et al., 2010). Golden et al. (2008) synthesized both jarosite and spheroidal  
475 hematite from Mg-Al-Fe-SO<sub>4</sub>-Cl solutions in the following reaction sequence: (i) initial  
476 precipitation of metastable hydronium jarosite at pH 1.2-1.5, (ii) jarosite dissolution and  
477 precipitation of an intermediate ferric hydroxyl species through forced hydrolysis that  
478 polymerizes into hematite, and (iii) precipitation of hydronium alunite upon depletion of  
479 hydronium jarosite. Hydrothermal acidic conditions and high Fe<sup>3+</sup> concentrations promote the  
480 formation of hematite over goethite (Cornell and Schwertmann, 2003). However, goethite is  
481 predicted to form from the breakdown of jarosite under reducing conditions (King and

482 McSween, 2005), which is supported by observations in acid-mine drainage (AMD)  
483 environments (Fernández-Remolar et al., 2005). Elwood Madden et al. (2012) describe jarosite  
484 dissolution experiments under a range of different pH and temperature conditions, and found that  
485 iron (hydro)oxides formed quickly in experiments with pH >3.5. While various phases formed as  
486 initial reaction products (e.g. hematite, maghemite, schwertmannite, and ferrihydrite), reactions  
487 did lead to hematite over time. They observed goethite in only one reaction (at pH 4.4) and only  
488 as a minor phase, and postulate that goethite may not form as readily under water-limited  
489 Martian conditions. Both hematite and goethite were observed in the fumarole transect at Hverir,  
490 though hematite was more common and usually more abundant.

491

## 492 **Phyllosilicates**

493

494 Kaolin and smectite group minerals (as detectable using only bulk powder XRD) are the main  
495 phyllosilicates in the Hverir fumarole transect. Kaolinite is only observed in the most distal  
496 samples (IV-14-17 and 18, at 800 cm) and smectite is only identified in one sediment sample  
497 (IV-13-11, at 600 cm), alongside plagioclase and augite, reflecting a low degree of alteration.  
498 Increased distance from the pathway of hydrothermal fluids will raise the pH and neutralize the  
499 acidity within the system and, thereby, precipitate out minerals containing more mobile elements  
500 (e.g. Ca, Mg). The composition of smectite depends on a combination of several factors  
501 including the pH of the fluid, element mobility, and the pathway of the fluids. It is important to  
502 remember that the genesis of poorly crystalline smectite clays comes from the hydration of  
503 sideromelane fragments, i.e. palagonite. Through continued alteration, palagonite transforms into

504 smectite, hence, smectite can be an abundant mineral phase in areas of low surface activity  
505 further from the source of heat and fumarolic vapors.

506

507 Thermodynamic models indicate that most phyllosilicate formation is favored by moderate to  
508 alkaline pH conditions (e.g. Velde, 1995). However, kaolinite can precipitate at low pH (Fialips  
509 et al., 2000). The behavior of phyllosilicate stability in low pH systems with high ionic strength  
510 is not well understood. For example, Mg-smectites (e.g. saponite) are more susceptible to  
511 alteration in acid waters than Al-phyllosilicates. Even so, Story et al. (2010) showed that these  
512 phyllosilicates can persist in sediments saturated in acid-saline lakes in Western Australia. In  
513 addition, Peretyazhko et al. (2014) synthetically produced saponite and nontronite from the  
514 alteration of basaltic glass under mildly acidic conditions (pH 4). Also, Hynek et al. (2013)  
515 documented kaolinite in acid-sulfate alteration of basalts in fumaroles and hot springs with pH as  
516 low as 2 and montmorillonite at pH 4 at Cerro Negro volcano, Nicaragua. The presence of  
517 smectite at the outskirts of a fumarole apron, associated with both primary igneous phases and  
518 Fe-sulfate minerals, provides further evidence that smectite can form and persist under acid  
519 sulfate fumarolic conditions, in an area of low surface activity.

520

## 521 **Silica and anatase**

522

523 Most samples from the Hverir fumarole transect display a significant amorphous silica “hump”  
524 in their XRD patterns. Some samples (all from Pit 1, 200 cm from the Hverir fumarole: samples  
525 IV-14-6 and IV-14-8 through 12) contain cristobalite as well. Opal A was only observed in  
526 sample IV-14-6, in a white, fibrous coating of an altered basaltic fragment near the surface of the

527 same pit. Quartz was observed in two samples collected at or near the surface- one sandy  
528 sediment (sample IV-13-09) and one PAB (sample IV-13-14 PAB).

529

530 The abundant amorphous silica in the leached apron surrounding the fumarole, often associated  
531 with abundant anatase (and elevated concentrations of both SiO<sub>2</sub> and TiO<sub>2</sub> measured by XRF), is  
532 consistent with acid-sulfate leaching of the surrounding rocks and soils at low to moderate  
533 temperatures. Other elements are leached away, leaving a residual concentration of these less  
534 mobile elements. SiO<sub>2</sub> is initially housed in amorphous silica, while TiO<sub>2</sub> is housed within  
535 anatase.

536

537 Since the substrate is Holocene basaltic lava, quartz and cristobalite are not expected as primary  
538 volcanic minerals. Their presence within pieces of altered basalt rules out detrital contamination  
539 as a likely source (e.g., wind-blown from degrading rhyolites in the Krafla area), though this  
540 remains a possible source for the quartz observed in the sandy sediment sample (IV-13-09).

541 While cristobalite is most often considered a high-temperature silica polymorph (if quenched  
542 rapidly from high temperatures, it can exist metastably at the surface), it can also form from the  
543 devitrification of volcanic glass (Heaney, 1994). Opal-CT, which consists of disordered mixtures  
544 of cristobalite and tridymite, can form authigenically within sediments (Heaney, 1994), though  
545 its XRD pattern reflects this disorder, with broader peaks than the sharp cristobalite peaks  
546 observed in this study (e.g. Graetsch, 1994). Cristobalite is also described in steam vent samples  
547 from Sulfur Banks at Kilauea, Hawaii, where Morris et al. (2000) attribute its formation to  
548 precipitation from dissolved SiO<sub>2</sub> under hydrothermal conditions. Quartz and cristobalite both  
549 generally require higher temperatures to form, above the range typically associated with surface



550 hydrothermal alteration, or longer periods of time, to reach a higher level of diagenetic  
551 maturation. Estimates for the time it takes for opal-A formed through hydrothermal sinter  
552 precipitation or acid-sulfate fumarole leaching to mature into quartz under terrestrial surface  
553 conditions vary from thousands or tens of thousands of years (Herdianita et al., 2000; Rodgers et  
554 al., 2002, 2004) to as little as hundreds of years (Taupo volcanic area: Lynne et al., 2007) or even  
555 months (Lynne et al., 2006), depending on the specific conditions. These shorter episodes of  
556 maturation are assisted by higher temperatures, such as those in persistent acid-sulfate fumarole  
557 environments.

558

559 Anatase is a major mineral constituent on some of the acid leached samples, and in some cases is  
560 the only crystalline phase (associated with amorphous silica).  $\text{TiO}_2$  concentrations reach as high  
561 as 23.16 wt% in one leached soil sample (IV-13-8), and a PAB collected at the same level (IV-  
562 13-8-PAB) had 4.65 wt%  $\text{TiO}_2$ , significantly concentrated over the 1.82 wt% observed in the less  
563 altered substrate. This overwhelming pattern of  $\text{TiO}_2$  enrichment reveals significant acid leaching  
564 and the removal of more mobile constituents, even silica, and demonstrates clearly that the  
565 associated amorphous silica is associated with residual enrichment due to leaching, rather than  
566 due to sinter precipitation (as might occur under more neutral conditions). With over 12x the  
567 amount of  $\text{TiO}_2$  in the leached sample with the highest concentration compared to the less altered  
568 basalt,  $\text{TiO}_2$  is clearly being enriched relative to  $\text{SiO}_2$ , and their residual enrichment thus does not  
569 appear to be coupled. This contrasts with the observations of Morris et al. (2000) for Kilauea,  
570 where  $\text{TiO}_2$  is enriched at a more moderate level during fumarolic alteration of basalt, though  
571 most Hverir samples from the current study show comparable levels of enrichment.

572

## 573 **Overall Fluid-Rock Interaction and Secondary Mineralogy at the Surface**

574

575 The alteration halos or aprons around fumaroles exhibit distinct mineralogical and geochemical  
576 trends as the result of element mobility and the formation and precipitation of secondary  
577 minerals. The dominant processes acting within a volcanic geothermal system include: (i) the  
578 supply of acid and H<sub>2</sub>S, (ii) the extent of the reaction and pH of the fluids, and (iii) the redox  
579 conditions (Markússon and Stefánsson, 2011).

580

581 Acid-sulfate waters are produced upon the boiling of aquifer fluids, resulting in phase  
582 segregation of the vapor/steam from the boiling water. Rising volcanic vapors enriched in H<sub>2</sub>S  
583 and CO<sub>2</sub> condense into oxygenated ground- and surface-waters where the H<sub>2</sub>S is oxidized into  
584 sulfuric acid and the CO<sub>2</sub> is mostly degassed (Kaasalainen & Stefánsson, 2012). The results are  
585 acidic waters with pH <4 and reducing conditions that alter basalt in gas- (low water-rock ratios)  
586 and fluid- (high water-rock ratios) dominated settings. The SEM analyses of altered basalts in the  
587 current study show relict plagioclase and augite, largely replaced by an amorphous  
588 aluminosilicate (Figure 7). Glass is not observed in highly altered samples, but can be seen  
589 undergoing alteration in less altered samples (Figure 8).

590

591 In gas-dominated settings, such as the Hverir fumarole, secondary minerals can precipitate  
592 directly from the volcanic vapors or during alteration of primary igneous phases. Sulfur-bearing  
593 phases at the Hverir fumarole are limited to elemental sulfur at the source and sulfate minerals  
594 further out, indicating more oxidizing conditions. No sulfide minerals were observed in the XRD  
595 patterns for any samples from the Hverir fumarole, in contrast to what is observed in hot spring

596 and mud pot contexts to the west within the Hverir hydrothermal area (Mínguez et al., 2011) or  
597 on the slopes or summit of neighboring Námaskarð hill (el-Maarry et al., 2017). The fumarole  
598 environment thus appears to be significantly more oxidizing than the higher water/rock ratio hot  
599 spring and mud pot environments of the same field.

600

### 601 **Fumarolic Alteration Model**

602

603 A conceptual model for fumarolic alteration illustrates what happens when volcanic vapors  
604 interact with Holocene basalt (Figure 12). Fumarolic alteration of the high-Fe basaltic substrate  
605 (sample IV-13-15) exhibited isochemical trends, where mobile cations were extremely depleted  
606 near the discharge source ( $X_4$  – Figure 13) but retained in the alteration apron ( $X_1$  – Figure 13).  
607 Here, the low-water rock environment effectively mobilized major rock-forming and trace  
608 elements. Sulfur exhibits a slight enrichment in the silica-rich soils, but most is mobilized into  
609 the purple and red sediments, likely through lateral transport. Dissolved metals were transported  
610 in solution and precipitated as a variety of sulfate and oxide phases such as barite, gypsum,  
611 alunite-jarosite group minerals, and hematite along exterior surfaces and vesicle walls. In the  
612 purple and red sediments, Fe and Al mobility was lower because of a more oxidizing  
613 environment, which enriched the samples in immobile ferric Fe. Both Fe and Al were  
614 incorporated into  $Fe^{3+}$ -bearing sulfates and oxides (e.g.  $X_2$ ,  $X_3$  – Figure 13). These phases were  
615 important secondary minerals in medium to low activity areas. The least altered basalts show  
616 minimal leaching along the margins of surface activity, where only minor depletion of Mg, Ca,  
617 and Na were observed.

618

619 The low levels of volcanic gas emissions at Hverir fumaroles may explain the lack of surface  
620 efflorescence (e.g. Fe<sup>2+</sup>-sulfates), but sulfate precipitates were found in shallow depth profiles.  
621 Here, members of the alunite-jarosite group are the dominant sulfates. The dominance of Al-  
622 sulfates over other sulfate phases corresponds well to fumarolic gas emissions, where Al<sup>3+</sup> >>  
623 Ca<sup>2+</sup> + Mg<sup>2+</sup> (e.g. Hynek et al., 2013). Fe-bearing natroalunite, spheroidal hematite, and  
624 amorphous silica were commonly associated alteration minerals that formed along vesicle walls  
625 and exterior surfaces. These secondary phases have been observed in other basalt-hosted  
626 volcanic environments at Haleakala and Cerro Negro volcanoes (Bishop et al., 2007; McCollom  
627 et al., 2013b), and thus likely represent a key mineralogical assemblage of fumarolic acid-sulfate  
628 alteration.

629

### 630 **Relevance to Martian Hydrothermal Systems**

631

632 Extrapolation of the mineralogical and chemical trends associated with the Námafjall geothermal  
633 system to analogous environments on Mars requires the consideration of several factors. First,  
634 the basaltic composition in Iceland differs from that of Mars. Icelandic basalts have similar,  
635 albeit slightly lower, Fe, higher Al, and lower Mg than Martian basalts (Table 4). This is  
636 significant because substrate largely controls the aqueous geochemistry, which affects the  
637 secondary mineral phases present (e.g. Tosca et al., 2004). As a result, the basalt composition  
638 will control in part the relative abundance of secondary minerals. Second, Icelandic geothermal  
639 areas occur in open-system environments, where soluble salts and Mg, Ca, Na and other cations  
640 are leached out of the system by meteoric water. Today, Martian atmospheric conditions differ  
641 significantly from modern terrestrial conditions; however our understanding of atmospheric

642 conditions during the Hesperian-Noachian period is limited. The abundance of ferric minerals  
643 and phyllosilicates in Noachian terrains may indicate at least locally oxidizing conditions and/or  
644 abundant surface water that altered basaltic substrate. The gas chemistry is wholly unconstrained,  
645 but given the high levels of sulfur in Martian basalts, volcanism likely produced abundant  
646 volatiles in the form of SO<sub>2</sub> and/or H<sub>2</sub>S gas (Gaillard et al., 2009).

647

648 **Home Plate, Gusev crater.** Gertrude Weise class soils were exposed by Spirit's right front  
649 wheel in the Eastern Valley between Home Plate and the Mitcheltree/Low Ridge complex. The  
650 light-toned soils are enriched in silica (as much as 91 wt% SiO<sub>2</sub>), Ti, Cr, and Zn, while most  
651 other major elements show a downward trend in abundance with increasing element atomic  
652 number relative to typical Martian soils (Squyres et al., 2008). The hydrated silica phase is  
653 identified (using measurements of Spirit's mini-TES instrument, compared against laboratory  
654 spectra of different silica polymorphs and mixtures) as opal-A and no high-temperature silica  
655 phases (e.g. cristobalite or quartz) were observed, indicating a lack of diagenetic maturation  
656 (Ruff et al., 2011). The exact depositional mechanism, fluid-rock ratios, and environmental  
657 parameters for the silica-rich soils in the Eastern Valley are not well defined.

658

659 Eastern Valley soils and nodular outcrops are enriched in silica (~64-92 wt%) and titanium  
660 (0.42-1.74 wt%) and were initially interpreted as leached acid-sulfate fumarole deposits due to  
661 their proximity to the Paso Robles (e.g. Tyrone) class soils (Ming et al., 2006). The sulfate-rich  
662 soils and Fe<sup>3+</sup>/Fe total indicate low pH and highly oxidizing conditions (Lane et al., 2008). Acid-  
663 sulfate alteration could explain the leaching of most minerals from Gusev basalts and the relative  
664 enrichment of the most insoluble components (e.g. SiO<sub>2</sub>, TiO<sub>2</sub>). However, the solubility of silica

665 is relatively independent of pH under acidic to neutral conditions (Alexander et al., 1954;  
666 Krauskopf, 1956) and, therefore, the composition of the Si-rich deposits does not constrain the  
667 pH of the fluids. Ruff et al. (2011) argued that the silica enrichment is the result of precipitation  
668 of silica-enriched fluids in a hot spring setting rather than acid leaching. They favored near-  
669 neutral hot spring sinter deposition, noting that the lack of sulfur enrichment was inconsistent  
670 with acid-sulfate leaching. Furthermore, the high water-rock ratios associated with hot spring  
671 settings would be necessary to transport excess sulfur out of the system. Ruff et al. (2019) have  
672 also made a strong case for a hot spring (rather than pervasive fumarole) origin for the silica-rich  
673 deposits based on morphological and stratigraphic evidence.

674

675 Námafjall has areas of enriched silica- (~75-92 wt%) and titanium- (~3-23 wt%) soils, as the  
676 result of intense fumarolic acid-sulfate leaching. Samples of these deposits were dominated by  
677 amorphous silica and anatase. As anatase was not observed in XRD patterns for the less altered  
678 substrate, and since many of these silica-rich samples contain significantly higher  $\text{TiO}_2$ , this  
679 anatase is interpreted as a secondary, residual phase that formed from the concentration of  $\text{TiO}_2$   
680 following leaching of other cations. Near-neutral sinter deposits were not observed at these sites  
681 and all hydrothermal fluids measured during our field studies in Námafjall were acidic. Thus a  
682 direct comparison between these two formation mechanisms cannot be assessed in this study,  
683 though the high concentrations of  $\text{TiO}_2$  observed are likely unique to acid-sulfate leaching, as  
684  $\text{TiO}_2$  is not easily mobilized in aqueous solutions (e.g., Gray and Coolbaugh 1994) and thus not  
685 enriched in amorphous silica deposits formed through precipitation.

686

687 Alternatively, low to moderate sulfur contents were observed in thick Si-rich soils (<1 m depth)  
688 in the Hverir fumarolic apron, never exceeding ~1 wt% SO<sub>4</sub> in soil samples for which it was  
689 analyzed, which is below all measured values for Eastern Valley materials on Mars, though  
690 altered basalt pieces (PAB samples) from Hverir often had higher SO<sub>4</sub> (see OM Table 2). At  
691 Hverir, amorphous silica and anatase were the dominant minerals near the vent (beyond the  
692 elemental sulfur halo), though in one sampling pit at 200 cm distance from the vent, cristobalite  
693 was also abundant, potentially indicating a higher degree of diagenetic maturity. Cristobalite was  
694 not found in surface samples or in samples from any other pits at Hverir, suggesting a localized  
695 occurrence. In sum, results show that silica deposits related to fumarolic alteration in the  
696 Námafjall area exhibit highly variable sulfur enrichment, ranging from <2 to 65x greater than the  
697 parent rock.

698

699 On Mars, Eastern Valley deposits did not exhibit other secondary minerals or other high-  
700 temperature silica phases (such as cristobalite), as interpreted from mini-TES data. Ruff et al.  
701 (2011) concluded that the lack of other mineral phases, especially sulfates, indicated a near-  
702 neutral hot spring setting. However, our study (and Hynek et al., 2013) demonstrates that  
703 fumarolic acid-leaching can remove most of the excess sulfur from the system and produce a  
704 deposit dominated by amorphous silica (and anatase) with only minor sulfates, which are not  
705 observed in most samples from the Hverir transect. These thick silica-rich soils generated by  
706 acid-sulfate fumaroles may be an effective analog for Eastern Valley soils (e.g. Lefty Ganote and  
707 Kenosha Comets). Therefore, the pH of the fluids and formation mechanism (near-neutral sinter  
708 deposition versus acid-sulfate leaching) that formed the silica deposits in the Eastern Valley may  
709 not be easily constrained by SO<sub>3</sub> concentrations and the presence or lack of associated sulfate

710 minerals, as this could be the result of either process. The morphologic and stratigraphic context  
711 cited by Ruff et al. (2019), and the lack of the extreme TiO<sub>2</sub> enrichments observed at Hverir in  
712 the Eastern Valley silica-rich deposits, are stronger justifications for a non-fumarole origin.  
713 However, Yen et al. (2008) acknowledge that Kenosha Comets has TiO<sub>2</sub> concentrations 50%  
714 higher than those observed in average basaltic soils and provisionally attribute their formation to  
715 a high-temperature acid bleaching process.

716

717 **Jarosite and Spheroidal Hematite at Meridiani Planum.** Jarosite ( $\text{KFe}^{3+}_3(\text{SO}_4)_2(\text{OH})_6$ ) is an  
718 important environmental indicator mineral, which has been detected by orbiters (e.g. Milliken et  
719 al., 2008; Farrand et al., 2009; Roach et al., 2010; Thollot et al., 2012; Bishop et al., 2021) and  
720 the Opportunity rover (e.g. Klingelhofer et al., 2004) in diverse geologic settings on Mars. In  
721 terrestrial environments, jarosite is typically found in low-pH volcanic sulfur-rich fumarolic  
722 deposits, acid-mine drainage, and saline-acid lakes (e.g. Elwood Madden et al., 2004; Tosca et  
723 al., 2005; Papike et al., 2006; Story et al., 2010). The stability of jarosite in terrestrial settings has  
724 placed constraints on the environmental conditions present in the sulfate-bearing rocks in the  
725 Burns Formation at Meridiani Planum that consists of ephemeral mildly to strongly acidic waters  
726 with a pH <4.5 and oxidizing conditions, since Fe must be present in ferric form (e.g. Burns,  
727 1987; Tosca et al., 2005).

728

729 The presence of jarosite in the Meridiani outcrops would also place constraints on the conditions  
730 after deposition. Elwood-Madden et al. (2004) concluded that low water-rock ratios are needed  
731 to preserve the jarosite and Ca-sulfate weathering products, because progress towards  
732 equilibrium causes early-formed minerals to become unstable. Therefore, the presence of jarosite



733 in Meridiani outcrops indicates that basaltic weathering on Mars likely did not go to completion.  
734 Limited water near the surface/subsurface would explain the presence of soluble sulfate-salts.  
735 Perhaps, after deposition from an ephemeral hydrothermal or non-thermal aqueous solution, the  
736 solution evaporated and stayed dry until the present day.

737

738 In terrestrial systems, jarosite is a relatively unstable mineral and breaks down to form Fe<sup>3+</sup>  
739 oxides and hydroxides; hematite and/or goethite, depending on the environmental conditions. At  
740 Meridiani Planum, one proposed pathway to form hematite spherules involves the incongruent  
741 dissolution of a jarosite precursor; both jarosite and hematite have been proposed to be products  
742 of acid-sulfate alteration of basalt (Morris et al., 2005; Golden et al., 2008; Elwood Madden et  
743 al., 2012; McCollom et al., 2013a). Námafjall FeOx spherules are closely associated with jarosite  
744 and Fe-bearing natroalunite, providing a natural analog to further support this scenario.

745 McCollom et al.'s (2013a) work further confirms that hematite spherules can be formed from Fe-  
746 bearing natroalunite in addition to jarosite by the same process.

747

748 Golden et al. (2008) presented the following model, which is also supported by our studies of  
749 acid-sulfate alteration in basalt-hosted volcanic environments at Hverir, and studies elsewhere  
750 (e.g. Morris et al., 2005; McCollom et al., 2013b). Acid-sulfate alteration of basaltic materials at  
751 Meridiani Planum and Námafjall formed Fe- (e.g. jarosite or Fe-natroalunite), Ca- and Mg-  
752 sulfates. The source of sulfur at Meridiani is uncertain and a recent review of the formation  
753 models, considering mass balance, indicated that none of the proposed models can account for  
754 the S enrichments (Hynek et al., 2019). The MER team infers that evaporative conditions of  
755 acidic groundwater would enrich S concentrations to form sulfates (Grotzinger et al., 2005).

756 Alternatively, S enrichment could be associated with the condensation of fumarolic emissions  
757 and/or the exposure and oxidation of iron-sulfides in the presence of water (Knauth et al., 2005;  
758 McCollom & Hynek, 2005), processes both present at Námafjall. Acid-sulfate solutions would  
759 have quickly dissolved the primary igneous phases, a scenario supported by only residual  
760 presence of igneous Fe-bearing phases and plagioclase, and released Fe into solution to form  
761 jarosite and hematite (Glotch et al., 2006). Low water-rock ratios are inferred because the bulk  
762 chemical composition is basaltic when calculated as S-free (McCollom and Hynek, 2005; 2006;  
763 Morris et al., 2006; McCollom, 2018), indicating that mobile elements were not removed as  
764 would be expected at higher water-rock ratios.

765

766 More detailed studies are required to determine the characteristics of hematite at Hverir, but it  
767 seems likely that spheroidal hematite there may have similar properties to those synthesized by  
768 Golden et al. (2008), McCollom et al. (2013b), and Marcucci and Hynek (2014). Hverir  
769 spherules are also similar to spherules in other volcanic settings (e.g. Morris et al., 2005;  
770 McCollom et al., 2013a); though they are substantially smaller than those at Meridiani (which  
771 are ~40 to 400x larger than terrestrial hydrothermal spherules). This difference in size may be the  
772 result of high-temperature (>100°C) conditions, which would improve the kinetics of spheroidal  
773 hematite formation, whereas the Hverir spherules are found in samples at temperatures <70°C.  
774 Longer duration of aqueous activity is also inferred at Meridiani Planum, while the altered  
775 Hverir substrate was deposited ~6100 years ago and is therefore a geologically young feature.  
776 Additionally, the greater availability of Fe from Martian parent rocks compared to terrestrial  
777 basalts could also potentially influence the size and abundance of iron oxides. Prolonged acid-

778 sulfate alteration on Mars could help form more extensive hematite deposits and possibly alunite  
779 within the altered siliciclastic materials.

780 Textural considerations, in addition to spherule size, limit the direct application of the Hverir  
781 hematite spherules to those at Meridiani. The hematite spherules in the Hverir fumarole transect  
782 was largely observed within vesicles or on interior surfaces within altered basaltic lava, which is  
783 notably different from the siliciclastic environment in which the Meridiani hematite spherules are  
784 found (e.g. McLennan et al., 2005). Hematite spherules from volcanic environments where  
785 tephra (rather than lava) is the substrate, such as at Mauna Kea (Morris et al., 2005) would thus  
786 be a better textural analog, though those spherules are also much smaller than the Meridiani  
787 hematite. Morris et al. (2005) also note that not all of the hematite at Meridiani is in the form of  
788 large spherules, and that spherules as small as the ones found in the Mauna Kea study (or the  
789 current study) would not be imageable by the MI instrument on Opportunity. Temperature is  
790 another consideration; if elevated temperatures are required to form gray hematite such as that at  
791 Meridiani (e.g. Golden et al., 2008), hydrothermal processes may have been required (e.g.  
792 McCollom and Hynek, 2005, 2006), though others have suggested mechanisms that could allow  
793 for coarsely crystalline hematite to form at lower temperatures (e.g. Glotch et al., 2004, 2006;  
794 Madden et al., 2010). The hematite at Hverir may have formed under either ambient or elevated  
795 temperatures, as it is observed both in the hotter sediments closer to the fumarole and at the  
796 margin.

797

798 The presence of jarosite at Meridiani was largely determined by Mössbauer spectra of outcrops.  
799 A range of synthetic Fe-bearing natroalunite compositions produced by McCollom et al. (2013a)  
800 yield spectra that closely resemble the Mössbauer spectral signature for jarosite, suggesting that

801 Fe<sup>3+</sup>-bearing natroalunite should be considered as an alternative to a pure Fe<sup>3+</sup>-sulfate component  
802 in the Meridiani outcrop. The presence of Fe<sup>3+</sup>-bearing natroalunite in the Burns Formation would  
803 indicate variable Fe-Al substitution in the B site. Christensen et al. (2004) calculated that  
804 samples for which jarosite was identified at Meridiani had insufficient Fe and too much sulfur  
805 for all of the sulfate component to be jarosite. The APXS-measured Al<sub>2</sub>O<sub>3</sub> concentrations of  
806 “jarosite”-bearing samples at Meridiani are not consistently higher than those of other samples  
807 (typically 6-8 wt. %: Rieder et al., 2004), though differences in the calculated abundances of  
808 feldspars (the main host of Al in the assemblage) makes it difficult to assess the amount of Al  
809 available for natroalunite.

810

811 If natroalunite, rather than jarosite, were the source of the “jarosite” signal in the Opportunity  
812 Mössbauer data, then the environmental conditions during the formation of the Burns Formation  
813 would be less well constrained. Fe-bearing natroalunites have stability ranges much higher than  
814 jarosite, up to pH 7.5 and under less oxidizing conditions (e.g. Brown, 1971). Natroalunite is also  
815 more stable than jarosite, which could affect the conditions required for its long-term  
816 preservation on Mars (e.g. McCollom et al., 2013).

817

818 In terrestrial settings, intermediate compositions (between natroalunite and jarosite) reflect early  
819 stages of acid-sulfate alteration of basaltic rocks in hydrothermal volcanic environments  
820 (McCollom et al., 2013b). This observation may offer further support for the volcanic origin for  
821 the inferred jarosite and hematite deposits at Meridiani (e.g. McCollom & Hynek, 2005). Fe-  
822 bearing members of the alunite family would also considerably expand the range of possible  
823 environmental conditions present at the time of formation. Alunite is thermodynamically stable

824 up to pH 7.5 and under less oxidizing conditions, and can precipitate from  $\text{Fe}^{2+}$  solutions (Burns,  
825 1987). If further evidence determines the Martian ‘blueberries’ formed under hydrothermal  
826 conditions, then this would suggest significantly different paleoenvironmental conditions at  
827 Meridiani than models that propose low-temperature interaction of sediments with groundwater  
828 (e.g. McLennan et al., 2005).

829

830

### Implications

831

832 The alteration mineral assemblages and patterns of element enrichment and depletion observed  
833 in this Hverir fumarole are relevant to the interpretation of potentially hydrothermal deposits on  
834 Mars, including silica-rich soils investigated by the Mars Exploration Rover (MER) Spirit in the  
835 Columbia Hills and jarosite and hematite occurrences at Meridiani Planum, investigated by MER  
836 Opportunity. Several key observations:

837

838 1) The co-occurrence of jarosite with smectitic clays show that these smectites formed under  
839 acid-sulfate conditions in the vicinity of these fumaroles. Clay formation is often  
840 attributed to neutral to alkaline fluids in models for early Mars (unlike sulfate-rich  
841 deposits, attributed to later acidic conditions, e.g. Bibring et al., 2005). Clay formation  
842 and/or stability under more acidic conditions is consistent with recent geochemical  
843 modeling by Peretyazhko et al. (2018) and observations from Australian acid lakes (Story  
844 et al., 2010) and other fumarolic environments (Hynek et al., 2013).

845

- 846 2) Amorphous silica rich deposits formed through acid-sulfate leaching in a fumarolic  
847 environment, yet many samples lack sulfate minerals. The lack of accessory sulfate  
848 minerals in hydrothermally-formed silica-rich deposits near Home Plate in the Columbia  
849 Hills has been interpreted to favor precipitation of silica sinters from near-neutral fluids,  
850 rather than acid-sulfate leaching (e.g. Ruff et al., 2011), though other lines of evidence,  
851 including morphology and stratigraphy, are now interpreted to favor a sinter model (e.g.  
852 Ruff et al., 2019).
- 853  
854 3) The presence of cristobalite in fumarole-altered basalt confirms that this phase is not  
855 limited to volcanic or high-temperature hydrothermal environments and can precipitate in  
856 fumarolic environments under milder hydrothermal conditions through direct  
857 precipitation or the breakdown of basaltic glass.
- 858
- 859 4) The extreme concentration of  $\text{TiO}_2$  (12x) in some fumarole-leached samples shows that  
860  $\text{TiO}_2$  enrichment can outpace  $\text{SiO}_2$  enrichment in an acid-leaching environment and it  
861 may not be the most appropriate “immobile” element to use in determining the degree of  
862 leaching. While not enriched to the same level, the silica-rich soil Kenosha Comets from  
863 the Eastern Valley (Gusev crater, Mars) also has  $\text{TiO}_2$  concentrations 50% higher than the  
864 likely precursor basalt (Yen et al., 2008).
- 865
- 866 5) These fumarole alteration deposits contain naturally occurring Fe-bearing natroalunite.  
867 Only limited solid solution between natroalunite and jarosite has been previously  
868 recognized (e.g. Papike et al., 2006). This is important because Fe-bearing natroalunite  
869 could have similar Mössbauer spectroscopy signatures to jarosite (e.g. McCollom et al.,

870 2013a,b), which could call into question the identification of jarosite at Meridiani Planum  
871 on Mars (e.g. Klingelhöfer et al., 2004). Natroalunite is less restricted in its pH range, and  
872 its presence (rather than jarosite) could allow a wider range of conditions that could have  
873 formed these sulfates on Mars.

874

875 6) Spheroidal hematite lining vesicle walls likely formed from the dissolution of Fe-bearing  
876 natroalunite and/or jarosite in Hverir fumaroles. This multistep process is similar to  
877 experiments conducted by Golden et al. (2008) and McCollom et al. (2013b), suggesting  
878 a pathway of formation for hematite spherules on Mars. If Fe-bearing natroalunite is the  
879  $\text{Fe}^{3+}$ -sulfate component in the Burns Formation, this may be a key signature of  
880 hydrothermal acid-sulfate alteration, though the larger size and textural context of the  
881 Meridiani “blueberries” needs to be considered.

882

883 Together, these observations inform our interpretations of the environments under which fluid-  
884 rock interactions occurred on the surface of ancient Mars. Acid-sulfate fumarolic alteration (and  
885 associate leaching) is a plausible mechanism for the formation of some Martian deposits,  
886 including some silica-rich soils in the Columbia Hills (explored by the Spirit rover) and jarosite  
887 and hematite (explored by Opportunity). Smectitic clays can also form in these environments and  
888 co-exist with minerals typically associated with lower pH conditions (e.g. jarosite), and their  
889 presence thus does not preclude acid-sulfate hydrothermal processes.

890

891

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892

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898 History.  
899



## Figures

900

901

902 Figure 1: Maps of (a) Iceland and the Krafla area, adapted from Gudmundsson and Arnorsson  
903 (2005), and (b and c) Námafjall geothermal field, with sampling sites from this study. Geologic  
904 map in (b) adapted from Saemundsson et al. (2012), satellite imagery in (c) from Google Earth.

905

906 Figure 2: Plan map and field photos for the sampled Hverir fumarole. (a) Map of surface sample  
907 locations and pits relative to the center of the fumarole apron. (b) field photograph of the  
908 fumarolic apron, with locations of shallow excavated pits. (c) close-ups of the shallow pits, with  
909 sample IDs.

910

911 Figure 3: Photos of the Hverir fumarole transect (including positions of four pits) and XRD  
912 patterns for samples from Pit 1 (200 cm from fumarole). The tan soil from the surface (IV-13-  
913 8R, 68.5°C) contains abundant amorphous silica and anatase, and minor quartz. IV-14-12, an  
914 orange soil sample collected at 18 cm depth (69.6°C), contains amorphous silica, cristobalite,  
915 goethite, and anatase. Sample IV-14-6, a white coating on a piece of altered basalt collected from  
916 16 cm depth (73.2°C), contains amorphous silica, cristobalite, and anatase. Sample IV-14-9,  
917 purple sediment collected at 38 cm depth (71.3°C), contains amorphous silica, cristobalite,  
918 anatase, natroalunite, and hematite. Its elevated background in XRD is attributed to iron  
919 fluorescence, as this sample contains 15.40 wt%  $\text{Fe}_2\text{O}_3\text{T}$ , compared to 0.9-2.63 wt% of other  
920 samples reported from this pit.

921

922 Figure 4: Photos of the Hverir fumarole transect (including positions of surface samples) and  
923 XRD patterns for surface samples along it. Sample IV-13-5 was collected at the fumarole  
924 (97.8°C) and consists entirely of elemental sulfur. Sample IV-13-6 is surface sediment collected  
925 at 30 cm from the source (66.9°C), and consists entirely of amorphous silica and anatase,  
926 representing a leached deposit. Sample IV-13-14 PAB W+P consists of a white and pink altered  
927 rock collected at the surface 500 cm from the source, and contains amorphous silica, anatase,  
928 quartz, hematite, and minamiite. Sample IV-13-10 White consists of white sediment collected  
929 550 cm from the source at 5 cm depth (17.2°C) and contains amorphous silica, anatase, gypsum,  
930 and rhomboclase. Sample IV-14-18 consists of red mud collected 800 cm from the source in a  
931 shallow pit (Pit 4) at 15 cm depth, and contains hematite, natroalunite, and kaolinite. Its high  
932 background in XRD is attributable to iron fluorescence, given the significant hematite  
933 occurrence in this sample.

934

935 Figure 5: Spider plot of altered basalt (PAB) samples from the Hverir fumarole transect  
936 normalized against the least altered basaltic substrate sample (IV-13-15). Missing data points  
937 indicate concentrations below the limits of detection. All samples indicate enrichment in sulfur  
938 compared to the fresh basalt, consistent with interactions with fumarolic sulfur, and all but the  
939 most distant from the center show elevated  $\text{TiO}_2$ ,  $\text{SiO}_2$ , and Zr, consistent with residual  
940 enrichment through acid leaching.  $\text{MnO}$ ,  $\text{CaO}$ ,  $\text{Na}_2\text{O}$ , and  $\text{K}_2\text{O}$  are all significantly depleted in  
941 these leached samples. Sample IV-13-7-PAB is anomalous, as it is close to the fumarole center  
942 yet shows a less leached element abundance pattern similar to more distal samples.

943

944 Figure 6: Plots of XRF-derived bulk compositions for samples from the Hverir fumarole transect.  
945 All plots vs. Fe<sub>2</sub>O<sub>3</sub>T (wt%). The same symbol is used for all samples collected at a specific  
946 distance from the fumarole vent at the center of the feature. SiO<sub>2</sub> is highly concentrated (up to 92  
947 wt%) in many samples, Fe<sub>2</sub>O<sub>3</sub>T is more enriched in lower SiO<sub>2</sub> samples. Al<sub>2</sub>O<sub>3</sub>, K<sub>2</sub>O, and P<sub>2</sub>O<sub>5</sub>  
948 abundances are typically higher for the higher Fe<sub>2</sub>O<sub>3</sub>T samples, while TiO<sub>2</sub> concentrations are  
949 highest in a few samples with low Fe<sub>2</sub>O<sub>3</sub>T. The variation exhibited by samples from the same  
950 distance show the degree of small-scale compositional heterogeneity.

951

952 Figure 7: SEM BSE images of a thin section of a heavily altered piece of basalt within red-  
953 orange soils (sample IV-13-7). (a) Small section of primary minerals in a highly leached basalt;  
954 Fe-S mineral is tentatively identified as pyrite. Altered clinopyroxene (augite) and plagioclase  
955 leached of mobile elements (e.g. Ca, Na) and replaced by amorphous alumino-silicate. (b)  
956 Altered titanomagnetite crystals (sample IV-13-7). Botryoidal growth of Fe-oxide spheroids  
957 forming in close association with titanomagnetite.

958

959 Figure 8: SEM image of altered basaltic glass fragments from sample IV-13-13 (a-d) and IV-13-  
960 11 (e). (a,b) Sideromelane fragments altered to Fe-hydr(oxides) and clays, in thin section. SEM  
961 BSE image of secondary minerals identified in amorphous silica veins. (c) SEM BSE image of  
962 needle-like barite crystals (labeled “ba”) and (d) SEM SE image of sample chip, showing  
963 authigenic gypsum (labeled “gy”) exhibits bladed star-shaped crystals on an amorphous silica  
964 spheroid from sample IV-13-13. (e) SEM BSE image of IV-13-11, showing less altered glass (gl)  
965 and basaltic lava fragments (bas) from the margin of activity (600 cm). sm = smectite.

966

967 Figure 9: Examples of the pseudocubic habit of the alunite-jarosite group from altered basalts  
968 from the Hverir geothermal field. Members of this group shown are natroalunite determined by  
969 EDS, except in (b), where jarosite was observed. (a) Interior vesicle wall with natroalunite,  
970 hematite, and amorphous silica (IV-13-14 PAB). (b) Adjacent vesicle walls with jarosite (light)  
971 forming on top of natroalunite (dark) with hematite and amorphous silica intermixed (IV-13-14  
972 PAB). (c) Natroalunite cubes with dissolved interiors. (d) Natroalunite cubes with K-rich rims  
973 and Al-rich interiors (IV-13-14 PAB). (e) Pseudocubic natroalunite closely associated with small  
974 hematite spherules (IV-13-13). (f) Higher magnification image of (e), showing textural  
975 relationship between natroalunite and hematite (IV-13-3). NaL = natrojarosite, Si = amorphous  
976 silica, He = hematite, Jar = jarosite.

977

978 Figure 10: Natroalunite exhibits compositional zoning (IV-13-13). Alunite-jarosite group  
979 minerals have the chemical formula  $[AB_3(SO_4)_2(OH)_6]$  (A site occupied by monovalent species  
980  $K^+$ ,  $Na^+$ ,  $H_3O^+$ , while the B site is occupied by trivalent species  $Al^{+3}$  and  $Fe^{+3}$ ). Zoned  
981 natroalunite (orange line = transect) shows Al-rich and Fe-depleted cores with elevated Fe and K  
982 along the rims. K is also slightly enriched in the core. Amorphous Silica (Si), Natroalunite (NaL).

983

984 Figure 11: Two distinct hematite spherule size populations: >3 and 5-20 micrometer range, with  
985 minor Al, Si, and S concentrations. (a) Hematite spheroids formed along the wall of vesicles and  
986 relict plagioclase crystals (IV-13-13). Inset shows an expanded view of spheroids. (b) Botryoidal  
987 morphology along vesicle walls with natroalunite cubes present (IV-13-14 PAB) and (c) Al-rich  
988 centers in Fe-oxide spherules may represent Al cores (Al) of remnant natroalunite (IV-13-13).

989 (d) Broken spherule shows fibrous/radial morphology of hematite. Hematite was confirmed by  
990 XRD analysis.

991

992 Figure 12: Conceptual model for the Hverir fumarolic apron. Alteration apron is up to 8 meters  
993 in diameter. The vertical shaded bar represents the oxidation front (between elemental sulfur and  
994 sulfate), as determined by surface mineralogy. The sulfur-rich layer found at depth may signify  
995 diffuse sulfur-bearing vapors that did not penetrate the surface. The maroon shaded area  
996 represents the changing color of the sediments, as amorphous silica dominated leached deposits  
997 give way to redder, more hematite-rich deposits. X1-X4 correspond to isocon plots in Figure 13.

998

999 Figure 13: Isocon plots for four samples from the Hverir fumarole, plotted relative to the less  
1000 altered basaltic lava substrate (IV-13-15). Sample locations are marked as X1, X2, X3, and X4  
1001 on Figure 12. (X1) Along the distal margin (600 cm away, IV-13-11), a slightly altered basalt  
1002 has minor depletions in Ca, Mg, Na, and Mn. (X2) In the red sediments (550 cm away, IV-13-  
1003 13), Fe, K, and P are enriched and Al is slightly depleted, while, Ca, Mg, and Na are  
1004 substantially depleted. (X3) Purple rocks and sediments (500 cm away, IV-13-14 PAB) exhibit  
1005 pronounced depletions in most major elements, except for enrichments in Fe and K. (X4)  
1006 Heavily altered basalts within the silica-rich soils (280 cm away, IV-13-12) are highly depleted  
1007 of major cations with a residual enrichment of Si, Ti, and Zr. S is enriched in all samples.

1008

1009

## Tables

1010

1011 Table 1: Hverir fumarole transect sample descriptions, and XRD-based phase identifications for  
1012 non-sulfates and sulfates.

1013

1014 Table 2 (Online Materials): XRF (major elements).

1015

1016 Table 3 (Online Materials): XRF (trace elements)

1017

1018 Table 4: Bulk chemistry, Iceland basalts vs. Mars

1019

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Table 1: Hverir fumarole transect sample details, and XRD-determined relative abundances of sulfur-bearing and other phases.

	Type	Distance (cm)	Depth (cm)	T (°C)	Description	Non-sulfur bearing		Sulfur-bearing	
						Abundant	Minor	Abundant	Minor
IV-13-05	precip.	0	0	97.8	elemental sulfur at fumarole vent			S	
IV-13-06	sed	30	0	66.9	Pale orange, sandy	Si, At			
IV-13-07	sed	120	5	51.5	Reddish-orange sed under top crust	Si, At	Hm		Na
IV-13-08	sed	200	0	68.5	Tan with dark orange, sandy	Si, At	Q		
IV-13-09	sed	280	5	14.4	Tan, pale orange sed.	Si, At			
IV-13-10mix	sed	550	5	17.2	Mixed sample- dark brown sandy top, sulfur and multi-colored minerals below	Si	At, Hm, Gt	S	
IV-13-10W					White component of IV-13-10	Si, At		Rh, Gyp	
IV-13-11	sed	600	0	16.3	Dark brown fine sand with some small white crystals	Si, Pl	Sm, Aug		Fc
IV-13-12	precip.	280	0	14.4	Surface white above sample 9	Si, At			
IV-13-13	sed	550	0	17.2	Surface dark red/black above sample 10	Si, Hm	Gt		
IV-13-13PAB	rock					Si, Hm	Gt	J	
IV-13-14MIX	rock	500	0		Purple altered basalt near surface	Si	At, Hm	Mi	J
IV-13-14-W+P	rock					Si, Hm	Q, At	Mi	
IV-13-15	rock				Unaltered substrate basalt	Pl, Aug			
Hverir fumarole pits									
Pit 1									
IV-14-06	precip.	200	16	73.2	white, fibrous crystals coating lava rocks	Si, OA, Ct	At		
IV-14-07	sed	200	26	76.2	red layer with pieces of altered basalt		Si, Hm, At		
IV-14-08	sed	200	38	76.8	Massive yellow material	Si, Ct	At		
IV-14-09	sed	200	38	71.3	Purple lens	Si, Ct, Hm	At	Na	
IV-14-10	sed	200	68	69.6	White, gummy material below sample 9.	Si, Ct	At		Na
IV-14-11	rock	200	0	47.7	Chunks of altered basalt	Si, Ct	At, Hm		
IV-14-12	sed	200	18	69.6	Orange, muddy soft material	Ct, Gt	Si, At, Hm		
Pit 2									
IV-14-13	precip.	400	60	38.1	Yellowish altered soil (uniform with depth)	Si, At			
Pit 3									
IV-14-14	sed	600	30	20.6	Orange lenses (mined from yellow matrix)		Si, At, Gt	Nj	Na
IV-14-15	sed	600	50	40.5	Yellow matrix material.	Si	At		
IV-14-16 PAB	rock	600	40		Altered basalt pieces		Hm	Gyp	Na, Nj
Pit 4									
IV-14-17	sed	800	10	19.5	Orange mud layer	Ka			
IV-14-18	sed	800	15	25.1	Red mud layer	Ka	Hm	Na	
IV-14-19	sed	800	65	39.3	Light yellow/white/purple material	Si, At		S	
Pit 5									
IV-14-20	sed	0	40	97.0	White and beige mud (under sulfur crystals)	Si, At	Gt		

Phases: At= anatase, Aug = augite, Ct = cristobalite, Fc = ferricopiapite, Gt = goethite, Gyp = gypsum, Hm = hematite, J = jarosite, Ka = kaolinite, Mi = minamiite, Na = natroalunite, Nj = natrojarosite, OA = opal-A, Pl = plagioclase, Q = quartz, Rh = rhomboclase, S = sulfur, Si = amorphous silica, Sm = smectite

Table 4: Major element compositions for Námafjall and other select basalts, compared to Mars

Oxide (wt. %)	> 6100 ka						
	1724-29 Eruption	Eruption IV-13-15	Hyaloclastite IV-14-36	Mauna Kea <sup>a</sup>	Cerro Negro <sup>b</sup>	OCBP-1 <sup>c</sup>	Adirondack <sup>d</sup>
SiO <sub>2</sub>	48.80	48.77	49.27	48.70	48.40	48.19	45.30
Al <sub>2</sub> O <sub>3</sub>	12.70	13.13	13.43	1.40	17.70	8.06	10.40
Fe <sub>2</sub> O <sub>3</sub> T	15.80	15.42	14.09	12.00	9.90	20.23	21.10
MnO	0.18	0.23	0.16	0.21	0.20	0.43	0.42
MgO	5.70	6.13	5.24	3.90	6.70	10.40	11.90
CaO	10.10	10.59	10.60	6.60	11.90	6.46	7.80
Na <sub>2</sub> O	2.30	2.14	2.20	4.30	2.00	2.19	2.10
K <sub>2</sub> O	0.34	0.29	0.27	1.90	0.40	0.75	0.03
TiO <sub>2</sub>	2.00	1.82	2.19	2.80	0.70	1.01	0.49
P <sub>2</sub> O <sub>5</sub>	0.22	0.18	0.23	0.85	0.09	1.37	0.54
SO <sub>4</sub>	0.12	0.11	0.18	0.09	0.09	-	-
LOI	0.25	-0.62	0.62	1.60	-	-	-
Total	98.15	98.79	98.39	99.83	99.90	97.38	99.81

Bulk compositions of Icelandic basalts determined by XRF, this study.

<sup>a</sup>Mauna Kea (Morris et al., 2005), <sup>b</sup>Cerro Negro (Hynek et al., 2013).

<sup>c</sup>OCBP1 (Meridiani) (Morris et al., 2006), <sup>d</sup>Adirondack (McSween et al, 2006).

Total Fe reported as Fe<sub>2</sub>O<sub>3</sub>. "-" = not reported. LOI = loss on ignition, where applicable.

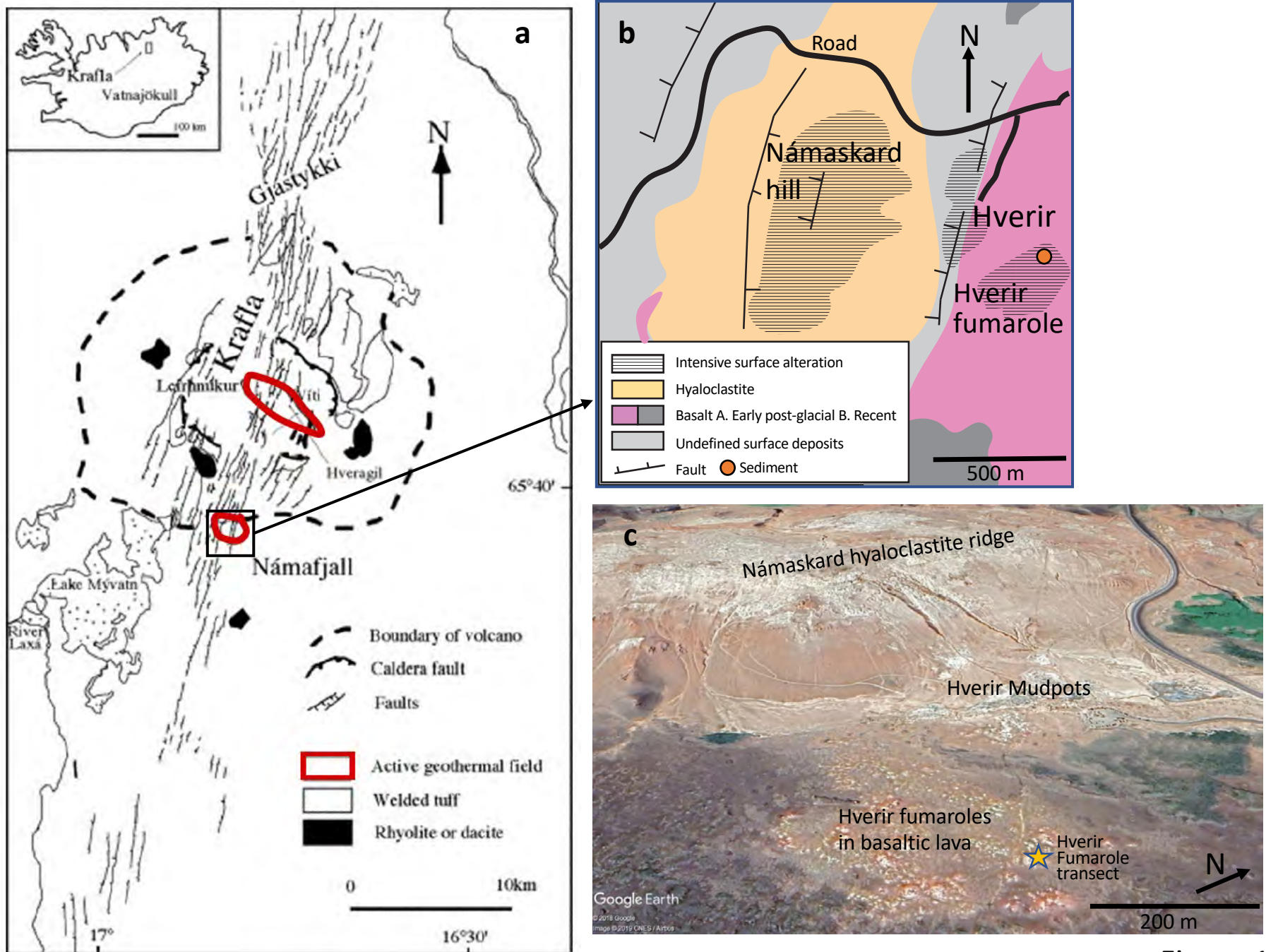


Figure 1

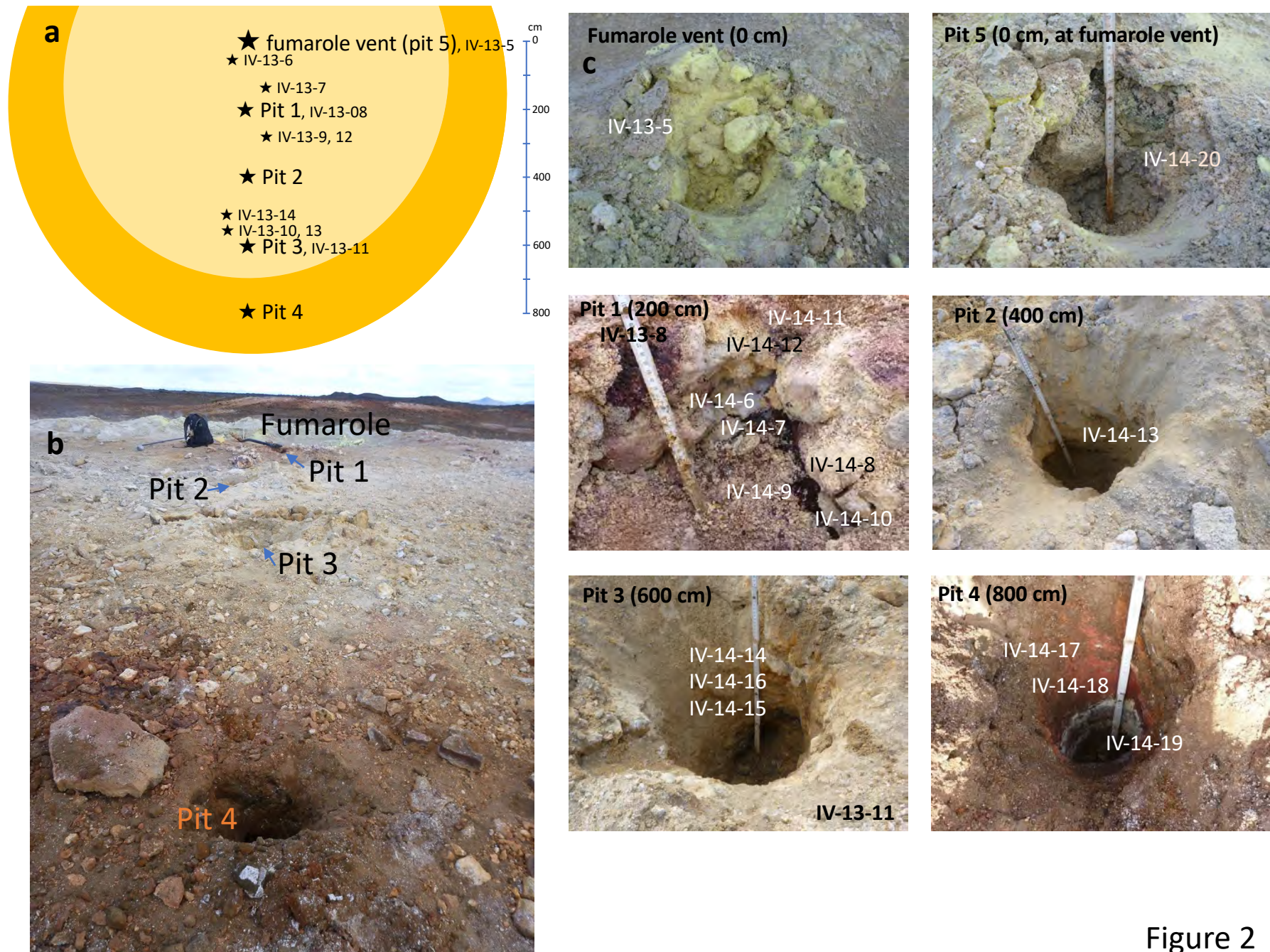
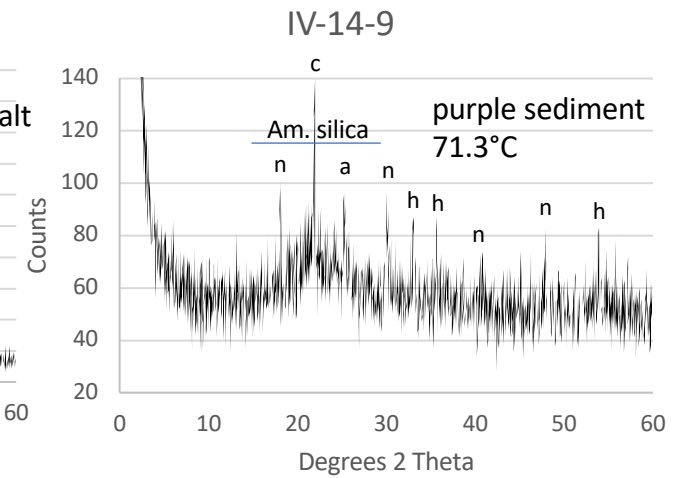
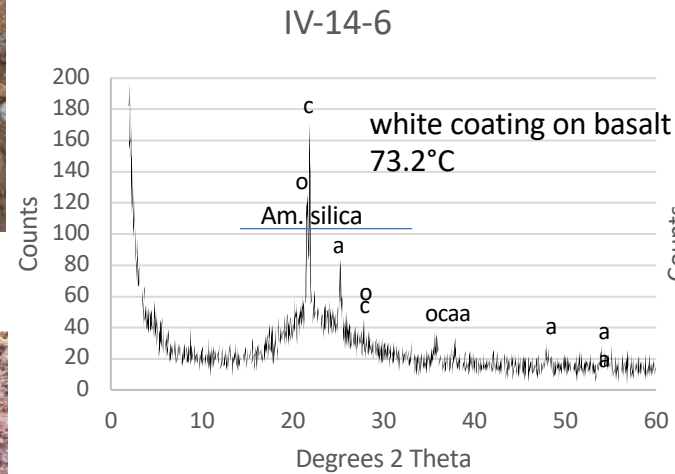
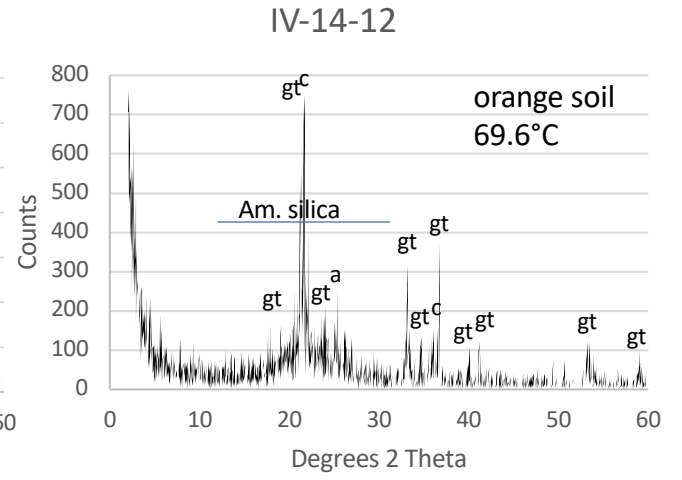
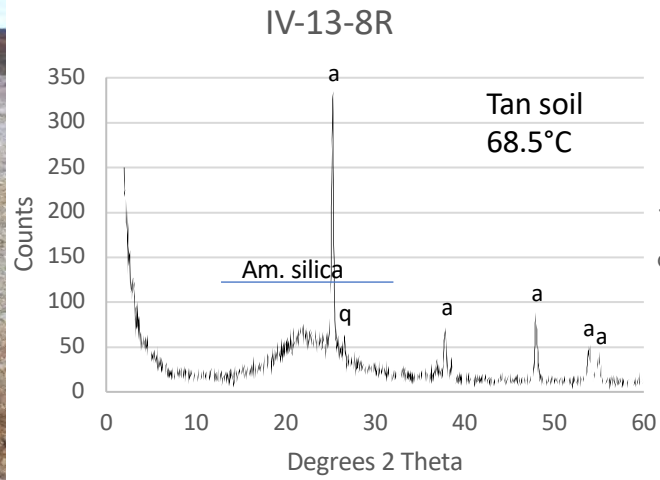
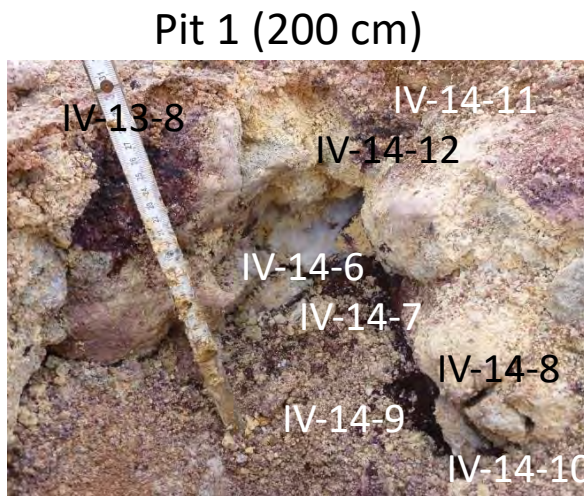
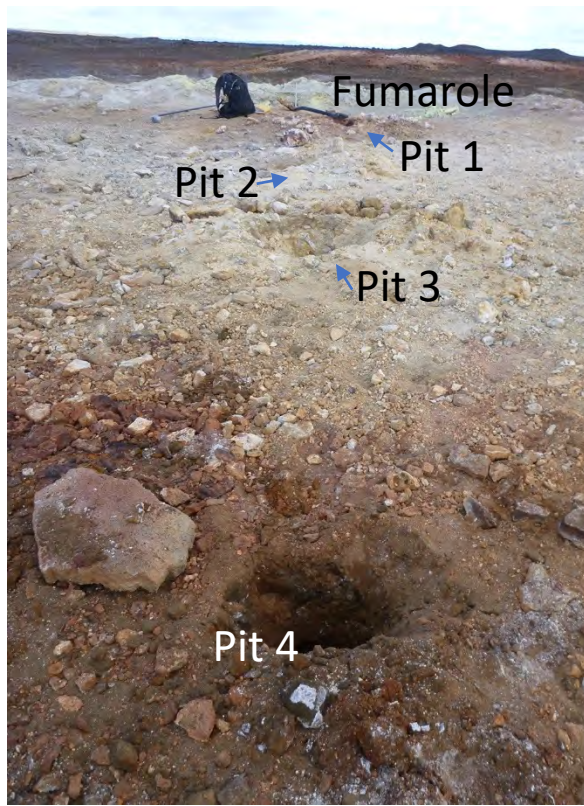


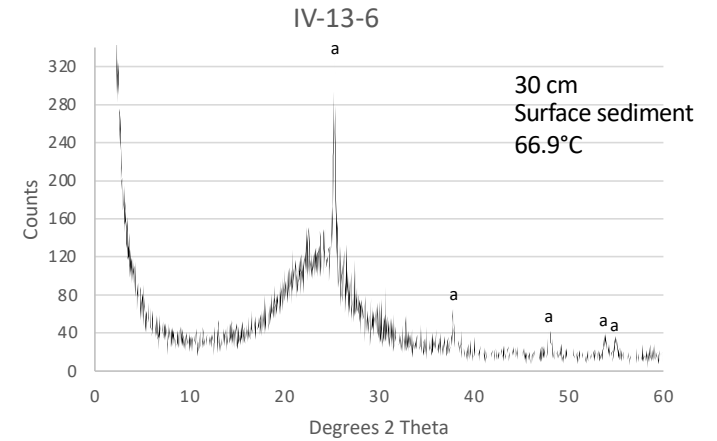
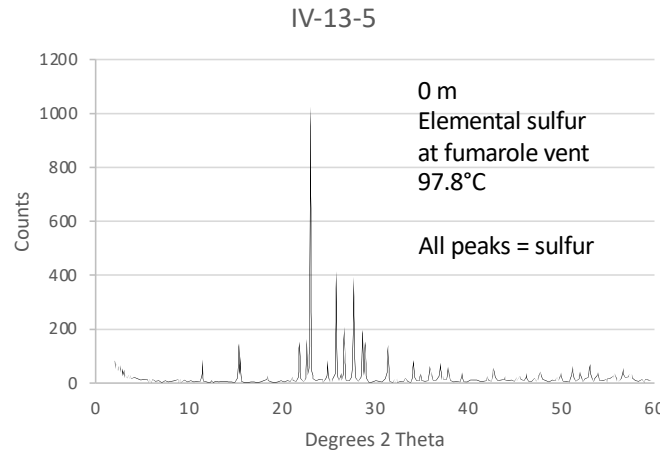
Figure 2



a = anatase  
 c = cristobalite  
 o = opal A  
 q = quartz  
 gt = goethite  
 n = natroalunite  
 h = hematite

Figure 3





a = anatase      q = quartz      g = gypsum  
 h = hematite    k = kaolinite    r = rhomboclase  
 mi = minamiite   n = natroalunite

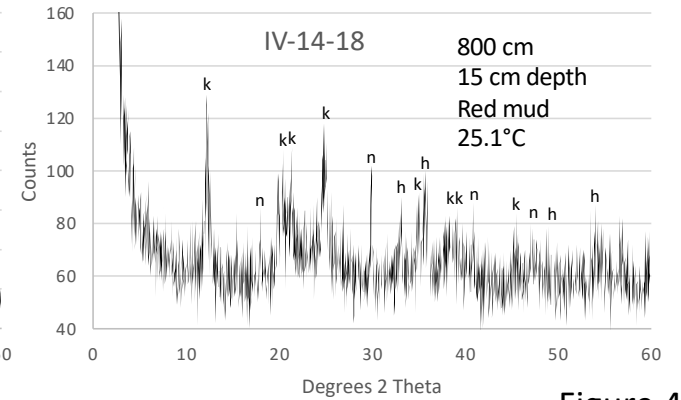
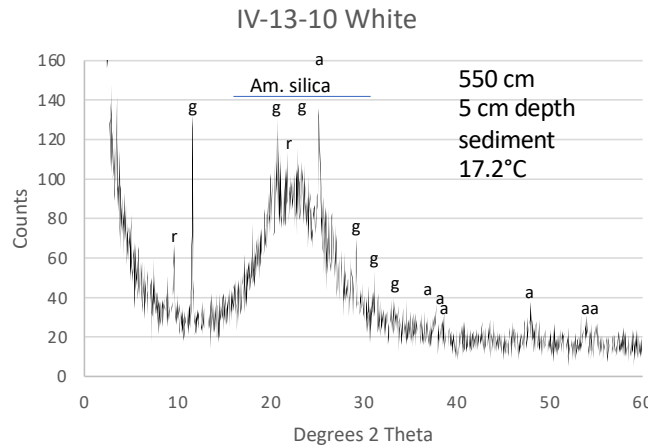
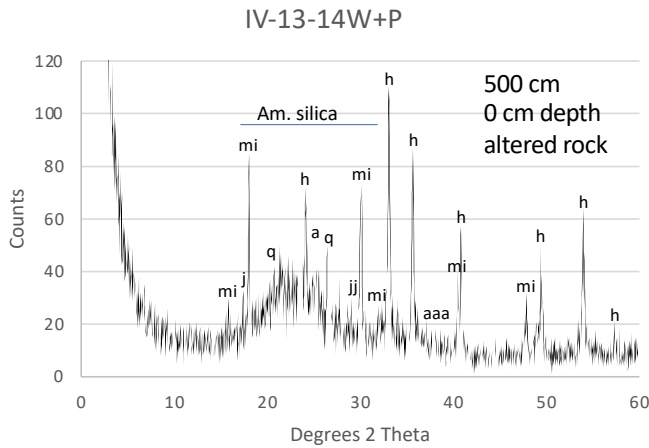
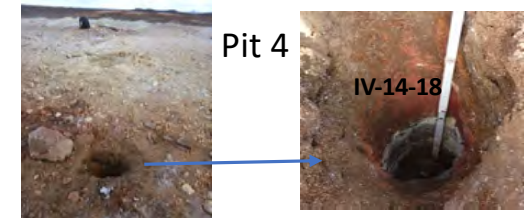


Figure 4

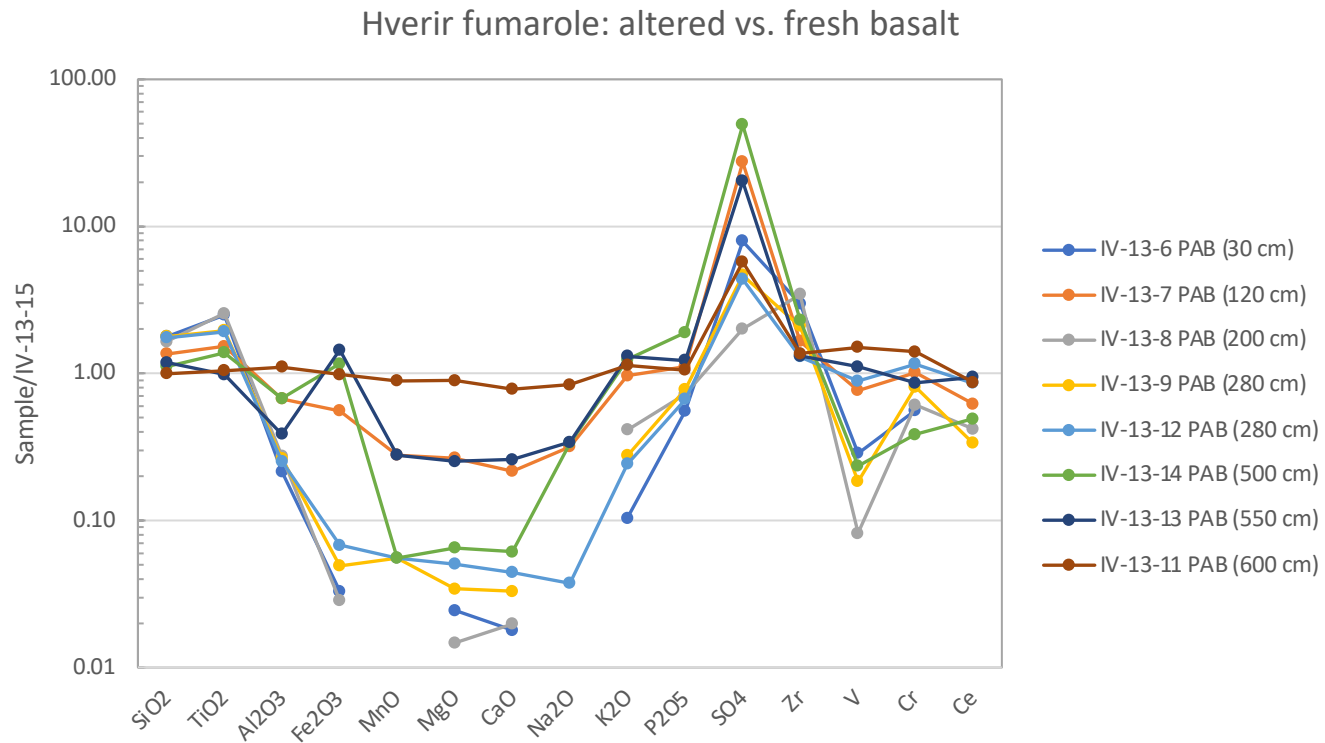


Figure 5

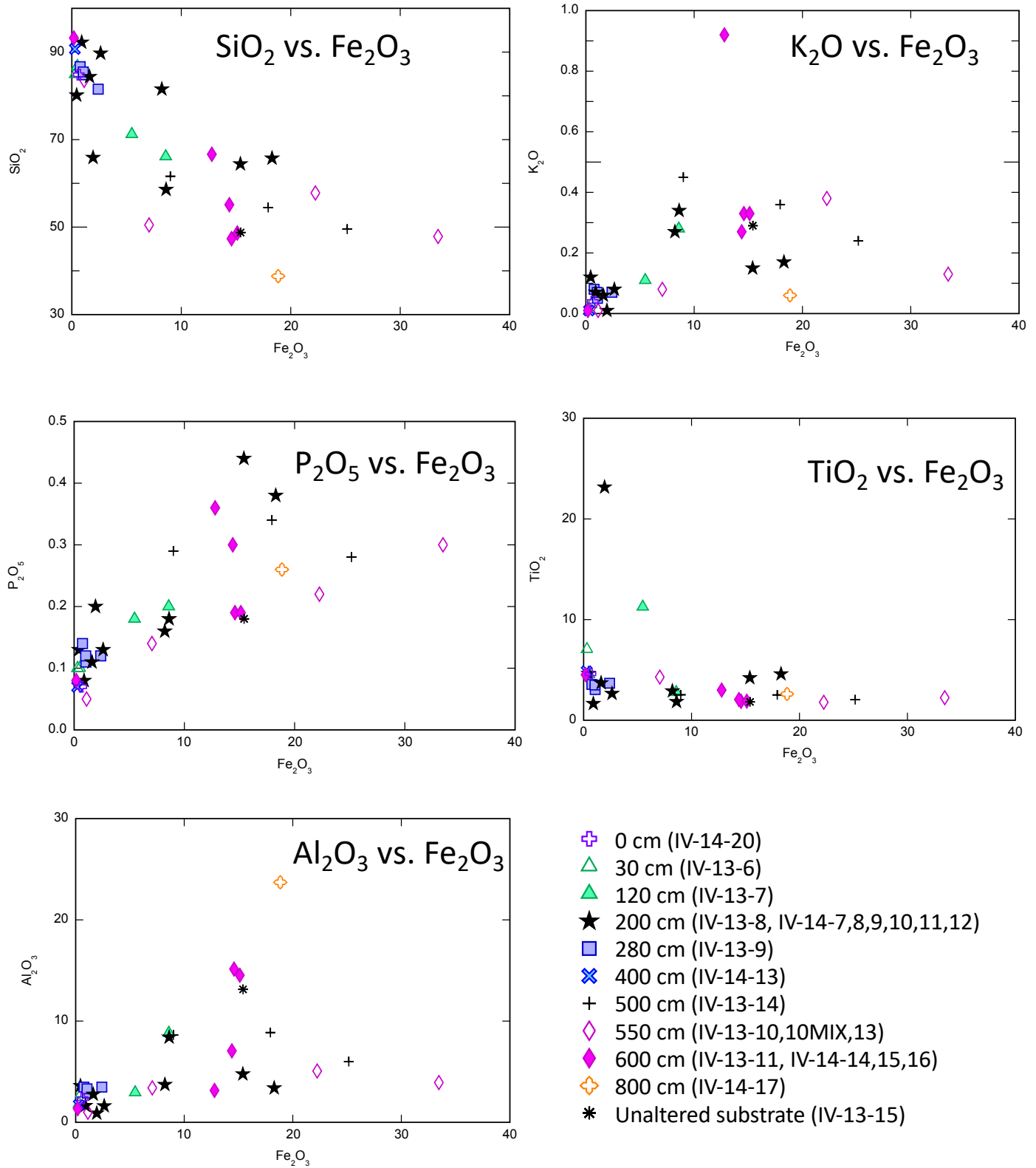


Figure 6

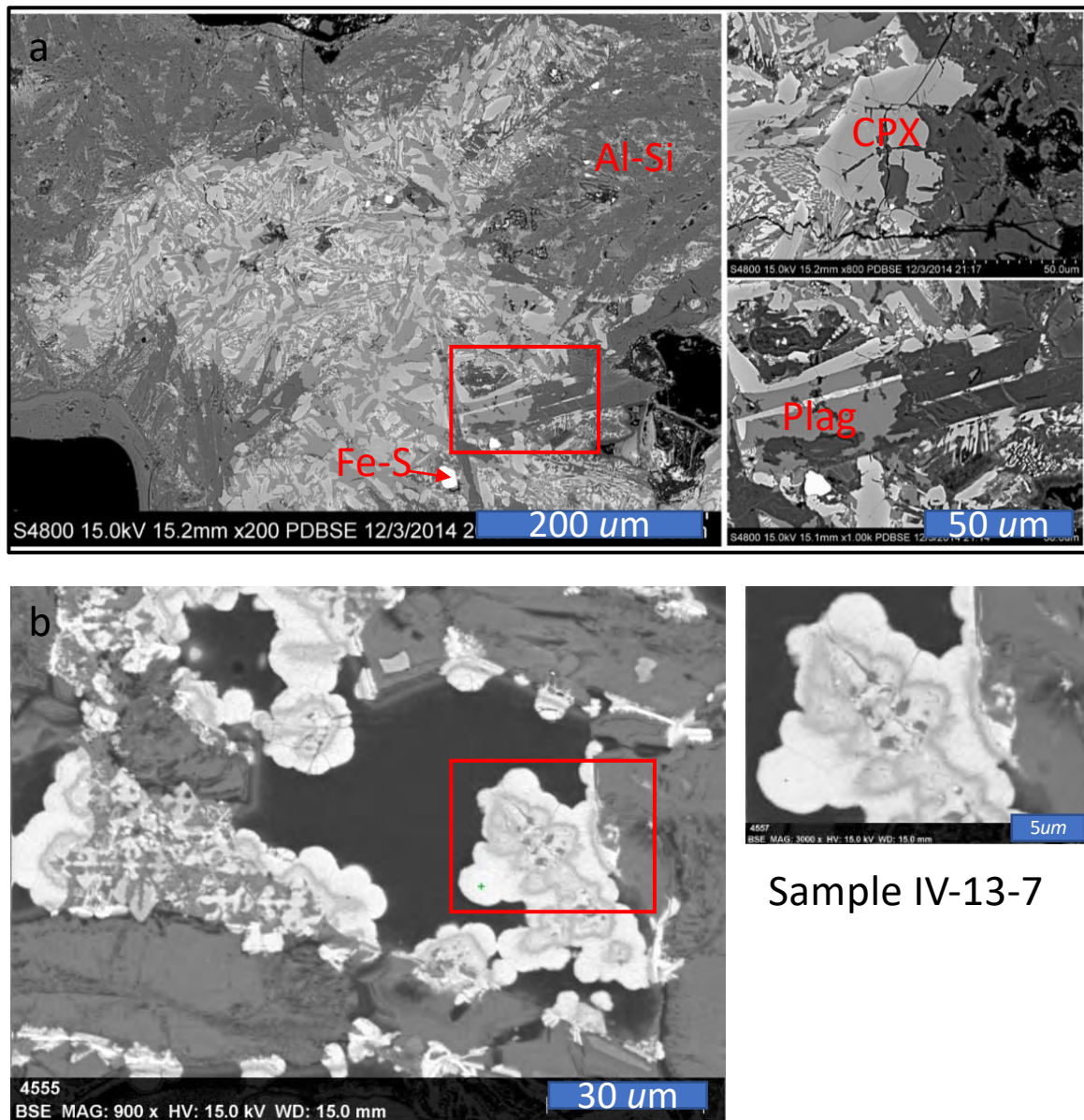


Figure 7

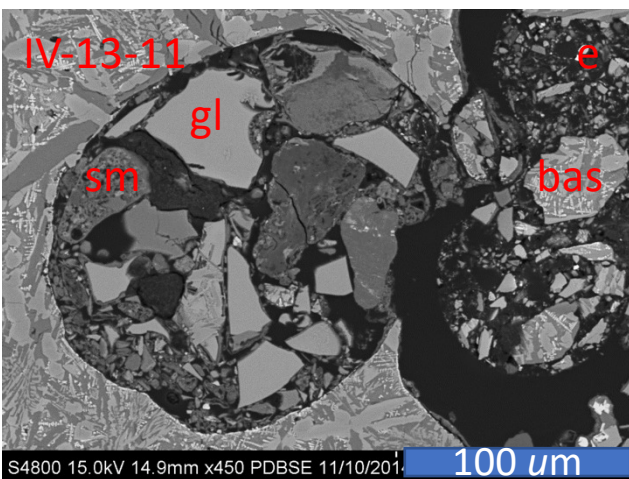
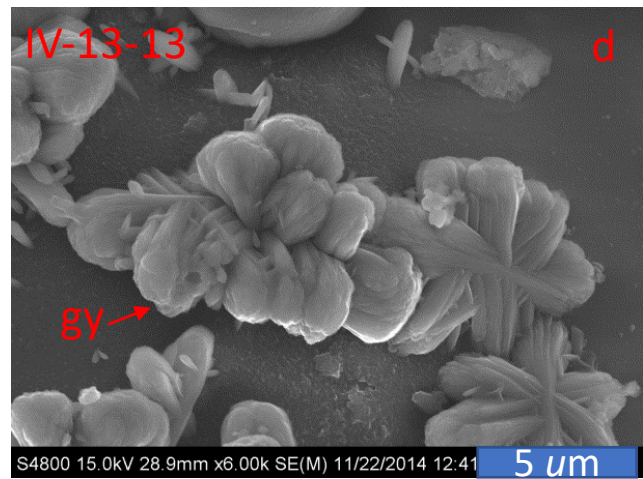
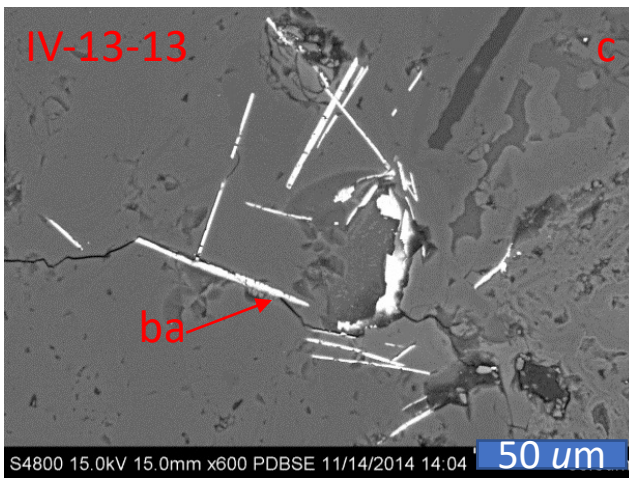
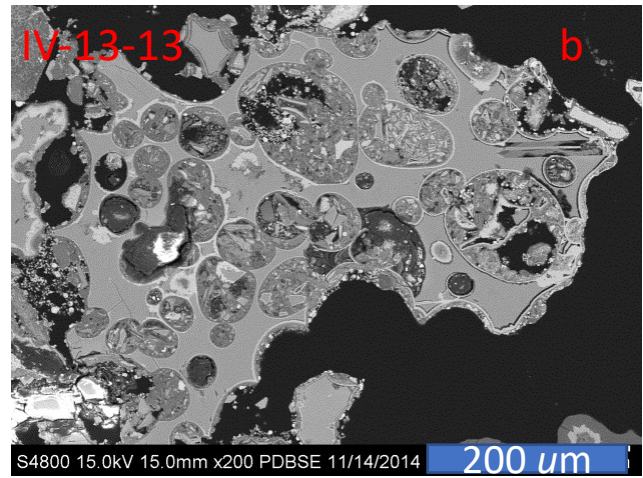
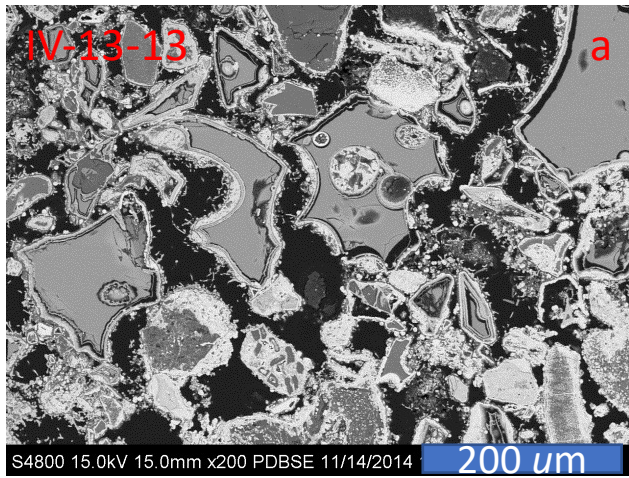


Figure 8

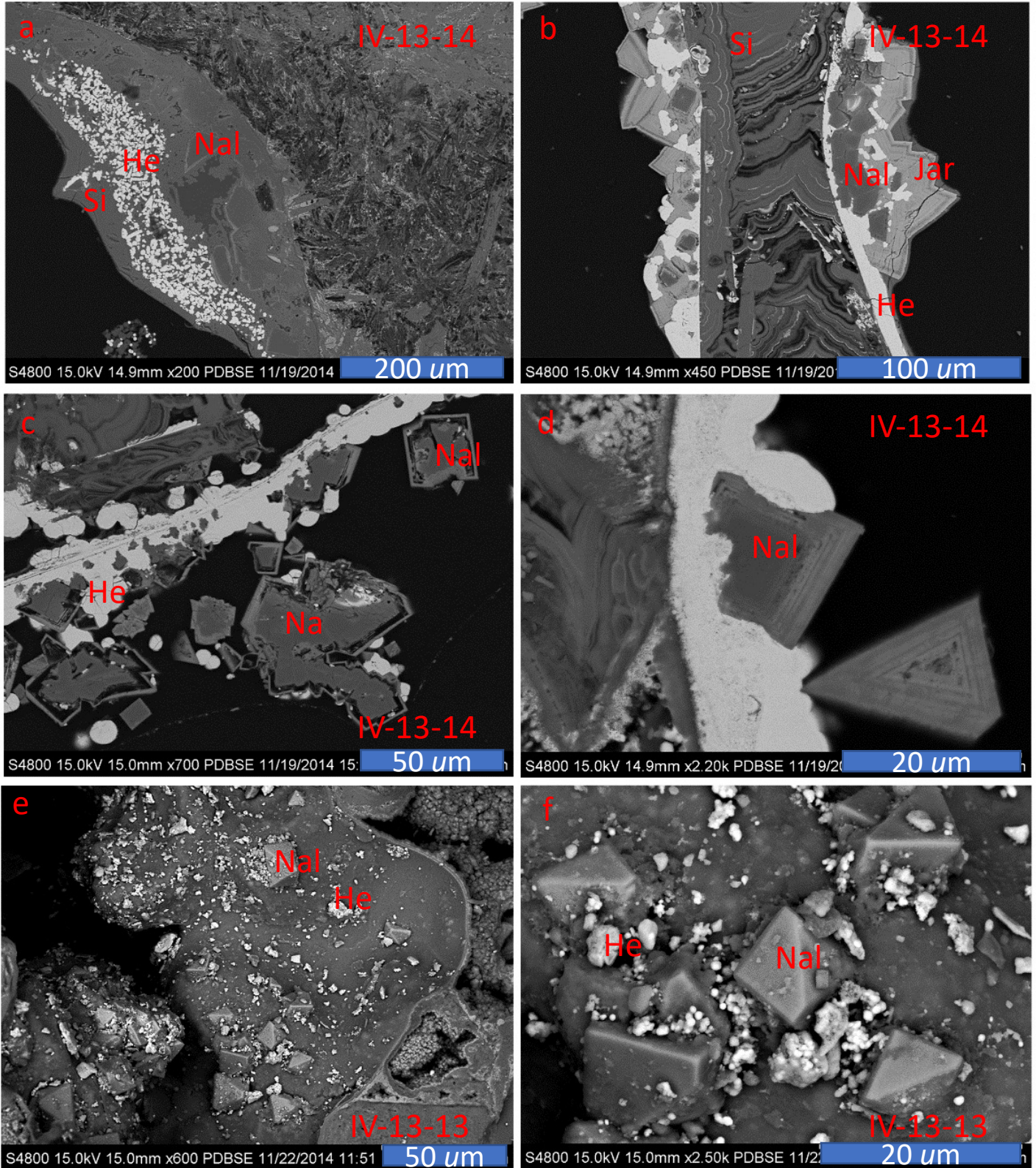


Figure 9

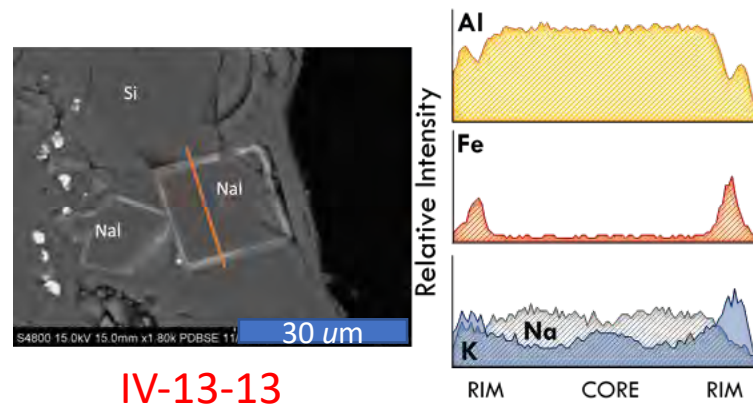


Figure 10

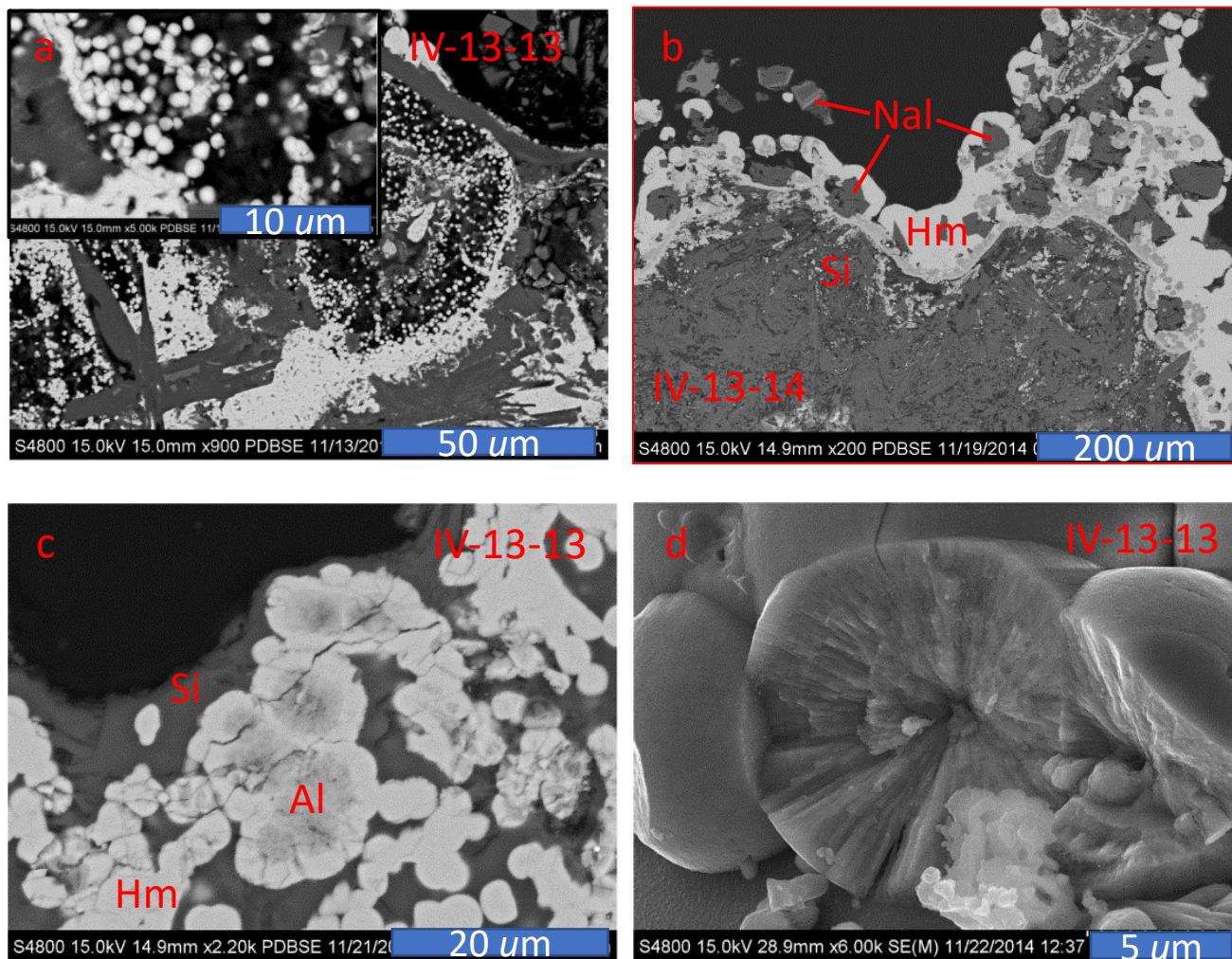


Figure 11



## Fumarole transect model

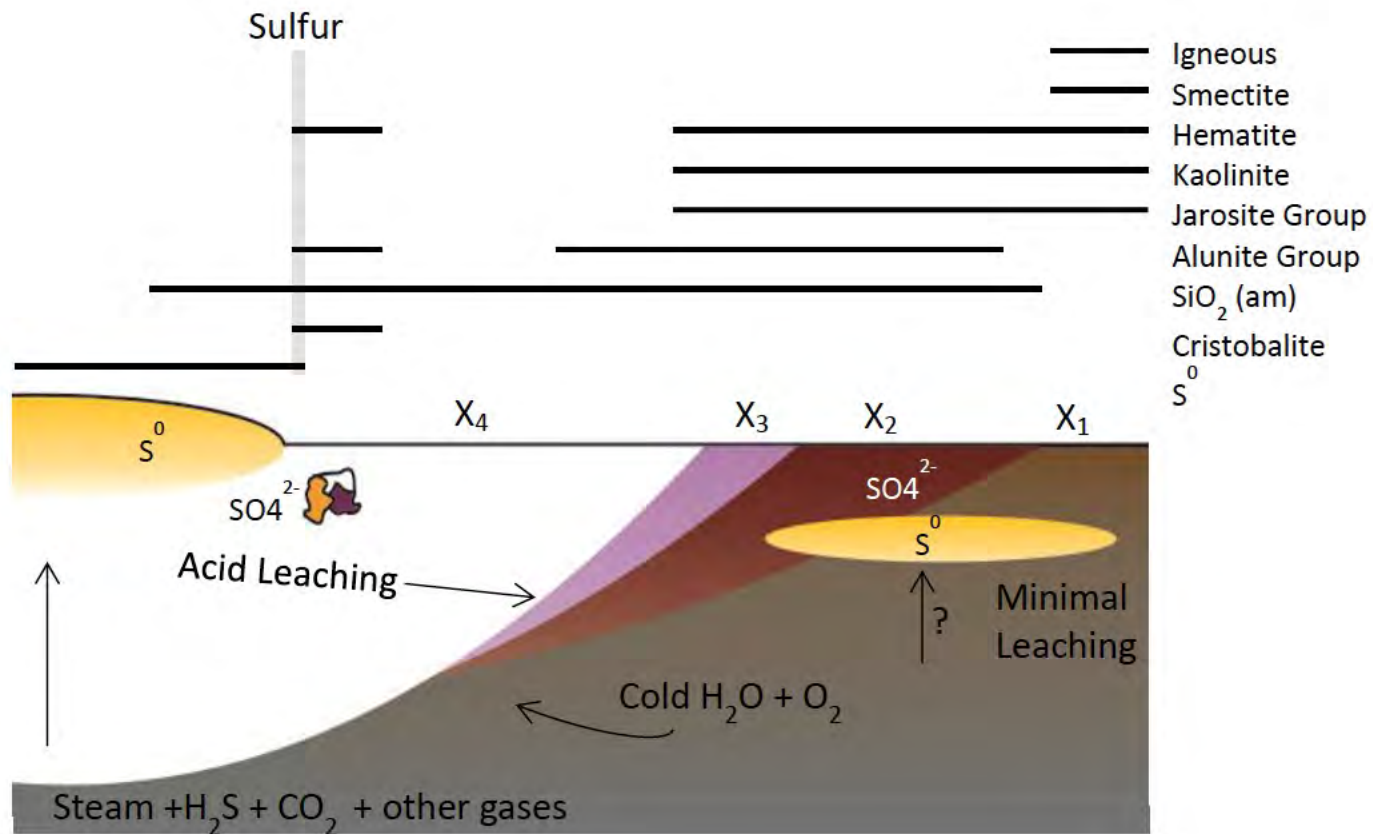


Figure 12

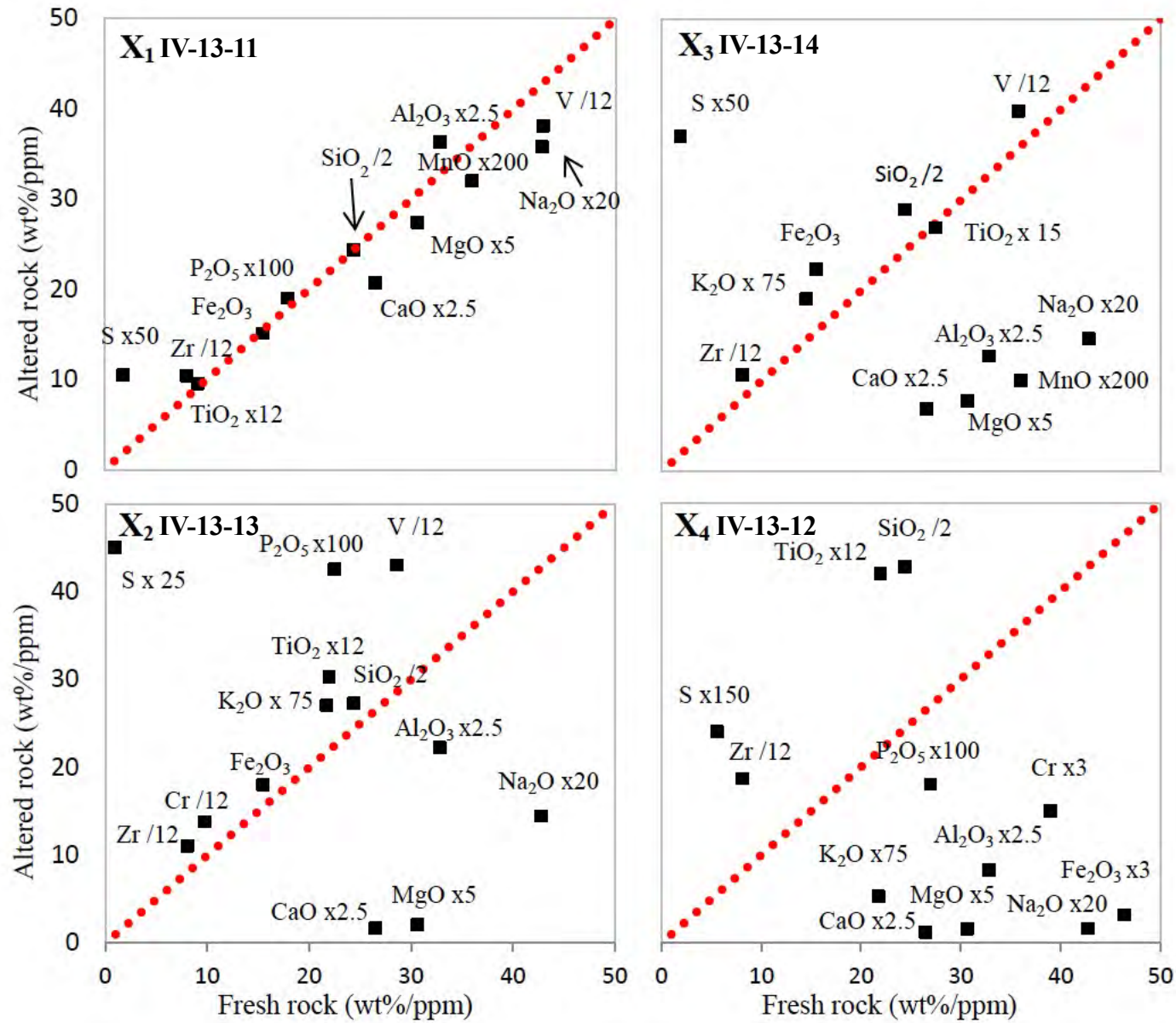


Figure 13