

1 **REVISION 1**

2 **Immiscible sulfide melts in primitive oceanic magmas: evidence and implications from**  
3 **picrite lavas (Eastern Kamchatka, Russia)**

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5 DMITRY P. SAVELYEV<sup>1</sup>, VADIM S. KAMENETSKY<sup>2,3,\*</sup>, LEONID V. DANYUSHEVSKY<sup>2</sup>, ROMAN E.  
6 BOTCHARNIKOV<sup>4</sup>, MAYA B. KAMENETSKY<sup>2</sup>, JUNG-WOO PARK<sup>5</sup>, MAXIM V. PORTNYAGIN<sup>6,7</sup>, PAUL  
7 OLIN<sup>2</sup>, STEPAN P. KRASHENINNIKOV<sup>7</sup>, FOLKMAR HAUFF<sup>6</sup>, AND MICHAEL E. ZELENSKI<sup>3</sup>

8 <sup>1</sup>Institute of Volcanology and Seismology FEB RAS, Petropavlovsk-Kamchatsky 683006,  
9 Russia

10 <sup>2</sup>Earth Sciences and CODES, University of Tasmania, Hobart, Tasmania 7001, Australia

11 <sup>3</sup>Institute of Experimental Mineralogy RAS, Chernogolovka 142432, Russia

12 <sup>4</sup>Institut für Mineralogie, Leibniz Universität, Hannover 30167, Germany

13 <sup>5</sup>School of Earth and Environmental Sciences & Research Institute of Oceanography, Seoul  
14 National University, Seoul 08826, South Korea

15 <sup>6</sup>GEOMAR Helmholtz Center for Ocean Research, Kiel 24148, Germany

16 <sup>7</sup>V.I. Vernadsky Institute of Geochemistry and Analytical Chemistry, Moscow 119991, Russia

17  
18 \* E-mail: [Dima.Kamenetsky@utas.edu.au](mailto:Dima.Kamenetsky@utas.edu.au)

19 **ABSTRACT**

20 Silicate-sulfide liquid immiscibility in mantle-derived magmas has important control on the  
21 budget of siderophile and chalcophile metals, and is considered to be instrumental in the origin  
22 orthomagmatic sulfide deposits. Data on primitive sulfide melts in natural samples, even those

23 representing most voluminous magmatism in oceanic rifts, are very scarce due to the small size  
24 and poor preservation of incipient sulfide melt globules. Here we present the first detailed report  
25 of the crystallized sulfides melts in the oceanic picrites of the (presumably) Cretaceous age  
26 Kamchatsky Mys ophiolite complex in Eastern Kamchatka (Far East Russia). Sulfide melts are  
27 present in three forms; 1) as inclusions in olivine (87.1–89.6 mol% Fo), 2) interstitial to the  
28 groundmass minerals (clinopyroxene, plagioclase and Ti-magnetite) of studied picrites, and 3) as  
29 daughter phases in silicate melt inclusions hosted by olivine and Cr-spinel phenocrysts. The  
30 sulfide melt inclusions in olivine and the groundmass of studied rocks are composed of several  
31 sulfide phases that correspond to the monosulfide (Fe–Ni; *Mss*) and intermediate (Fe–Cu–Ni;  
32 *Iss*) solid solutions. Several <0.5  $\mu\text{m}$  Pd–Sn, Pt– and Au–Ag phases are recorded within the  
33 matrix sulfides, commonly along phase boundaries and fractures. Major elements (S, Fe, Cu, Ni,  
34 Co), platinum group elements (PGE) and gold analyzed in the homogenized olivine-hosted  
35 sulfide melt inclusions, and phases identified in the matrix sulfides record the range of magmatic  
36 sulfide compositions. The most primitive sulfide liquids are notably enriched in Ni and Cu  
37 ((Ni+Cu)/Fe, at. > 0.5), continuously evolve with crystallization of (e.g., increasing Cu/Ni and  
38 Au/PGE) and demonstrate metal fractionation between *Mss* and *Iss*. Although the compositional  
39 systematics found in this study are consistent with those previously recorded, the compositions  
40 of individual sulfide phases are strongly affected by the noble metal (PGE, Au) ‘nuggets’ that  
41 exsolve at subsolidus temperatures and form during serpentinization of the rocks. We conclude  
42 that the budget of noble metals in the studied picrites is controlled by sulfides, but the  
43 abundances of Pt and Au are influenced by mobility in post-magmatic alteration. Our data can be  
44 also used for modelling sulfide saturation at crustal pressures and understanding behavior of the  
45 noble metals in primitive oceanic magmas.

46 **Keywords:** Silicate-sulfide immiscibility, mid-ocean ridge magmatism, picrite, olivine,  
47 sulfide melt, platinum-group elements

## 48 INTRODUCTION

49 Orthomagmatic models assume that many types of economic mineralization hosted in  
50 igneous rocks implicitly relate to magmas and magma-derived fluids (e.g., Burnham, 1998;  
51 Hedenquist and Lowenstern, 1994; Naldrett, 2004 and references therein). The origin of  
52 ‘mineralizing fluids’ is often linked to liquid immiscibility (unmixing) in parental silicate  
53 magmas (Roedder, 1992). The compositional divergence between unmixed phases can be  
54 extreme, and even though the physical amount of the new segregated phase may be small, it can  
55 extract, accumulate and transport large amounts of trace and rare elements (Roedder, 1992;

56 Veksler et al., 2012). Small amounts and transient, reactive qualities of immiscible phases render  
57 the understanding of their intrinsic compositions through the study of natural samples extremely  
58 difficult (e.g., Kamenetsky and Kamenetsky, 2010). The research problem persists with  
59 experimental and theoretical modelling being too simplistic to account for all physical and  
60 chemical variables in control of magmatic immiscibility.

61 Immiscibility between silicate and sulfide melts that occurs in a variety of mantle-derived  
62 magmas is believed to be a starting point in the formation of Cu–Ni sulfide deposits with  
63 platinum group elements (hereafter, PGE) (e.g., Naldrett, 2004). Unfortunately, the majority of  
64 continental rocks that host such deposits do not fit general criteria for primary/primitive melts  
65 (i.e., they feature <8 wt% MgO and the absence of high–forsteritic olivine phenocrysts). Even in  
66 the case of rare high–Mg rocks with primitive olivine in some large igneous provinces on  
67 continents (see review in Kamenetsky et al., 2017), a record of the early sulfide immiscibility is  
68 yet to be discovered. In contrast to continental magmas, sulfide globules in phenocrysts and  
69 glasses at mid-ocean ridge basalts (MORB) have been documented since the underwater samples  
70 became available in the 1960’s (e.g., Czamanske and Moore, 1977; Francis, 1990; Kanehira et  
71 al., 1973; MacLean, 1977; Mathez, 1976; Moore and Calc, 1971; Peach et al., 1990). These  
72 studies provided first insights into occurrence of silicate–sulfide unmixing in shallow plumbing  
73 systems during crystallization and eruption, but lacked specific details about the liquidus  
74 assemblage and compositions at the onset of immiscibility.

75 After a long period of quiescence the research on sulfide globules in MORB has been  
76 recently resurrected (e.g., Kamenetsky et al., 2013; Keith et al., 2017; Patten et al., 2012; Patten  
77 et al., 2013; Yang et al., 2014), however, scarcity of modern data still limits understanding of  
78 compositions of incipient sulfide melts, their compositional evolution with fractionation, and

79 metal partitioning between immiscible silicate and sulfide liquids. This paper is the first  
80 comprehensive report of several types of immiscible sulfide liquids in primitive oceanic  
81 magmas.

## 82 **GEOLOGICAL SETTING AND SAMPLES**

83 The Kamchatsky Mys (hereafter, “KM”) ophiolite complex in the easternmost Kamchatka  
84 Peninsula (Far East Russia) is renowned for its present-day tectonic position at the junction of  
85 the Kuril-Kamchatka arc, Aleutian arc and the inferred starting point of the Hawaiian-Emperor  
86 seamount chain (Fig. 1). Several studies have suggested that sedimentary rocks (e.g., shales,  
87 limestones, cherts), igneous rocks (e.g., lavas, hyaloclastites, dykes and gabbros) and mantle  
88 peridotites in the ophiolite assemblage are possibly different in age, and hence belong to  
89 different geodynamic settings (e.g., Zinkevich et al., 1985; Fedorchuk et al., 1989). The  
90 alternative explanation assumes mid-Cretaceous (Albian – Cenomanian) age (Palechek et al.,  
91 2010) for most rocks, and that the KM ophiolite represents magmatism and sedimentation in the  
92 former Kula–Pacific spreading center (Khotin and Shapiro, 2006; Portnyagin et al., 2005;  
93 Saveliev, 2003). The age and compositions of the ophiolite is consistent with the interpretation  
94 that it is the earliest expression of interaction of the Hawaiian “plume” and paleo-spreading  
95 center (Batanova et al., 2014; Portnyagin et al., 2009; Portnyagin et al., 2008; Portnyagin et al.,  
96 2005; Saveliev, 2003).

97 The volcanic rocks in the upper part of the KM ophiolite are basalts belonging to the  
98 tholeiitic, sub-alkaline and alkaline series (e.g., Fedorchuk, 1992; Portnyagin et al., 2008;  
99 Portnyagin et al., 2005; Saveliev, 2003). The enriched tholeiites are geochemically similar to  
100 those from the neighboring Meiji and Detroit Seamounts (Portnyagin et al., 2008), whereas ultra-  
101 depleted melt compositions contributed to N-MORB tholeiitic basalts (Portnyagin et al., 2009).

102           The olivine-phyric (picrite) rocks with magmatic sulfides reported here have been recently  
103 discovered in the serpentinite mélange that borders the ultramafic massif of Mt. Soldatskaya  
104 (Fig. 1; Savelyev, 2014), where ~1.5 m picrite boulders were found trailing in a small outcrop  
105 (~400 x 250 m) among debris of peridotites. The samples represent cumulates of euhedral  
106 olivine (40-50 %; Table 1) and rare plagioclase and Cr-spinel crystals set in a crypto- to  
107 holocrystalline (doleritic) matrix, composed of clinopyroxene– plagioclase– Ti-magnetite  
108 aggregates. In some varieties dendritic radial (sheaf-like) growth of groundmass minerals and  
109 altered interstitial glass are indicative of relatively fast cooling.

110           Olivine phenocrysts ranging in size up to 2-3 mm are moderately fractured and  
111 serpentinized along fractures and grain boundaries. Olivine is dominantly primitive ( $\text{Fo}_{87-90}$ ),  
112 does not show chemical zoning except for more Fe-rich thin (~200  $\mu\text{m}$ ) outer rims, and has trace  
113 element systematics typical for MORB olivine (i.e., 0.25–0.35 wt% NiO, 0.26–0.30wt% CaO,  
114 0.14–0.18 wt% MnO; Savelyev, 2014). Reddish-brown Cr-spinel, present as phenocrysts (up to  
115 0.4 mm) and inclusions (<0.15 mm) in olivine, is characteristically Mg- and Al-rich (68–76  
116 mol% Mg#, 19–40 mol% Cr#) and Ti-poor (0.25–0.43 wt%  $\text{TiO}_2$ ; Savelyev, 2014) and thus  
117 compositionally similar to MORB spinel. Both olivine and Cr-spinel phenocrysts contain  
118 crystallized inclusions of the silicate and sulfide melts.

119           The large abundance of olivine phenocrysts in the studied picrites is responsible for the  
120 high MgO contents (24–28 wt%) and low abundances of non-olivine major elements (8–9 wt%  
121  $\text{Al}_2\text{O}_3$ , 5–6 wt% CaO; Table 1) in the whole rock compositions. In addition to mineral  
122 compositions, the MORB-like geochemical characteristics of parental melts for the studied  
123 picrites are expressed in the contents and ratios of incompatible trace elements (e.g., La/Sm =  
124 1.5–2, Gd/Yb = 1.3–1.6, Nb/La = 0.8–1.4). The origin from a MORB-type depleted mantle is

125 further confirmed by unradiogenic  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.70243), radiogenic  $^{143}\text{Nd}/^{144}\text{Nd}$  (0.51315) and  
126 moderately radiogenic Pb ( $^{206}\text{Pb}/^{204}\text{Pb}=18.90$ ,  $^{207}\text{Pb}/^{204}\text{Pb}=15.50$ ,  $^{208}\text{Pb}/^{204}\text{Pb}=38.20$ ) isotope  
127 compositions (Appendix 1<sup>1</sup>).

128

## METHODS

### 129 Sample preparation

130 Sulfides were extracted from ~ 20 kg of rocks. The rocks were crushed and sieved (<0.2 mm, 0.2-0.5 mm and  
131 0.5-2 mm). A heavy fraction was then separated by panning in water. Olivine grains were hand-picked from the  
132 heavy fraction 0.5-2 mm and examined under a binocular microscope for sulfide inclusions. Rock fragments (0.2 to  
133 2 mm) from the heavy fraction were further separated into non-magnetic and magnetic fractions. The latter fraction  
134 yielded sulfide grains > 0.1 mm that were picked by hand, mounted in epoxy resin and polished. Sulfides were  
135 inspected in reflected light using the optical microscope, photographed, and selected for analytical work.

### 136 Electron microscope and electron microprobe

137 Backscattered electron (BSE) imaging and energy dispersive X-ray spectrometry (EDS) were performed on a  
138 Hitachi SU-70 Schottky field emission SEM fitted with an Oxford AZtec XMax 80 silicon drift detector EDS  
139 system at 15 kV at the University of Tasmania, and a VEGA 3 (Tescan, Czech Republic) scanning electron  
140 microscope with an energy-dispersive spectrometer X-Max 50 at the Institute of Volcanology and Seismology  
141 (Petrovavlovsk-Kamchatsky, Russia). Sulfur, Fe, Ni and Cu were calibrated on celestite, hematite, nickel silicide  
142 and cuprite (Astimex Standards Ltd, Toronto, Canada) using the K x-ray series for quantification. To correct for  
143 changes in the beam current between calibration and sample data acquisition, the cobalt Ka peak intensity was  
144 measured on pure cobalt metal at the start of each session.

145 Electron microprobe analysis (EMPA) using a Cameca SX100 electron microprobe (Institut für Mineralogy,  
146 Leibniz Universität Hannover, Germany) determined major element concentrations in sulfides. To minimize effects  
147 of small-scale heterogeneity caused by phase segregation (or exsolution) during cooling, sulfide samples were  
148 analyzed with the electron beam defocused to 20  $\mu\text{m}$ . Iron, S and Si were analyzed using 15 keV accelerating  
149 voltage, 15 nA beam current and 10 sec accumulation time. Nickel and Cu were analyzed using 15 keV, 100 nA  
150 beam current and 30 sec accumulation time. When possible, larger phases were analyzed in several locations to test  
151 for homogeneity. Calibration of the analyses was performed using a number of standards (Fe on  $\text{Fe}_2\text{O}_3$ , S on  $\text{FeS}_2$ ,  
152 Si on  $\text{CaSiO}_3$ , Ni on NiO, Cu on Cu metal).

### 153 Laser ablation ICPMS

154 Laser ablation inductively-coupled plasma mass spectrometry (LA-ICP-MS) analyses of sulfides were  
155 carried out at CODES Analytical Laboratories, University of Tasmania. The instrumentation was an ASI  
156 RESOLUTION-LR-S155 laser microprobe equipped with a Coherent Compex-Pro 193 nm Ar-F excimer laser,  
157 coupled to an Agilent 7700x quadrupole ICP-MS. A laser beam size of 29  $\mu\text{m}$  was used with a pulse rate of 5 Hz  
158 and a fluence of 2.7  $\text{J}/\text{cm}^2$  to drill holes into sulfide samples and standard materials. Ablation was performed in a He  
159 atmosphere flowing at ~0.35 L/min. The ablated aerosol was mixed with Ar (1.05 L/min) as a transport gas, before  
160 exiting the small volume cell. Tuning was performed to minimize oxide production (<0.15 % ThO/Th) and  
161 maximize sensitivity for the mid and high mass isotopes. The following isotopes measured were:  $^{34}\text{S}$ ,  $^{57}\text{Fe}$ ,  $^{59}\text{Co}$ ,  
162  $^{60}\text{Ni}$ ,  $^{63}\text{Cu}$ ,  $^{66}\text{Zn}$ ,  $^{99}\text{Ru}$ ,  $^{101}\text{Ru}$ ,  $^{103}\text{Rh}$ ,  $^{105}\text{Pd}$ ,  $^{106}\text{Pd}$ ,  $^{111}\text{Cd}$ ,  $^{185}\text{Re}$ ,  $^{189}\text{Os}$ ,  $^{193}\text{Ir}$ ,  $^{195}\text{Pt}$  and  $^{197}\text{Au}$ . Dwell time was 5 ms for S,  
163 10 ms for Fe, Co, Ni, Cu and Zn and 20 ms for all other elements, resulting in a total sweep time of 0.275 sec. Total  
164 acquisition time for each analysis was 90 seconds, consisting of a 30 second gas background and a 60 second  
165 ablation signal. Quantification was performed following the standard methods (Košler, 2001; Longrich et al.,  
166 1996). Calibration involved STDGL3 (Danyushevsky et al., 2011), NIS3 (Gilbert et al., 2013) and Po724-T  
167 (Sylvester et al., 2005) reference materials, and Fe as the internal standard element. Instrumental drift was corrected  
168 by regular (at least hourly) analyses of references materials across the analytical session. Due to the unknown bulk  
169 composition of the sulfide inclusions, quantification involved normalization to 100 wt% total. Corrections for base  
170 metal-argide interferences are required for the quantification of Ru, Rh and Pd concentrations by LA-ICP-MS  
171 techniques (Guillong et al., 2011; Sylvester, 2001). The analysis of magmatic sulfides has the potential to form

172 argide interferences of  $^{61}\text{Ni}^{40}\text{Ar}$  on  $^{101}\text{Ru}$ ,  $^{59}\text{Co}^{40}\text{Ar}$  on  $^{99}\text{Ru}$ ,  $^{63}\text{Cu}^{40}\text{Ar}$  on  $^{103}\text{Rh}$ ,  $^{65}\text{Cu}^{40}\text{Ar}$  on  $^{105}\text{Pd}$ , and  $^{66}\text{Zn}^{40}\text{Ar}$  on  
173  $^{106}\text{Pd}$ . For the quantification of Ru, Rh and Pd, the extent of base metal-argide production was determined by  
174 ablating pure Ni, Cu, Zn and Co metals and a correction factor applied to the results. All  $^{106}\text{Pd}$  results were corrected  
175 for isobaric interference from  $^{106}\text{Cd}$  which was monitored by recording signal on  $^{111}\text{Cd}$ .

## 176 Whole rock analyses

177 Two rock samples were analyzed for major, trace and platinum-group elements in three research institutes in  
178 South Korea. First, they were powdered in an agate mill at the Korea Polar Research Institute (KOPRI). A 0.95 g  
179 split of the powder was mixed with 9.5 g of Li-tetraborate and <0.03 g of Li-bromide. The mixture was fused in a  
180 platinum crucible in an open furnace at 1100°C for 15 min. The fused disks were analyzed on a S8 TIGER (Bruker,  
181 Germany) XRF at the National Center for Inter-university Research Facilities, Seoul National University (SNU).

182 Trace element contents were determined by LA-ICP-MS at the Korea Basic Science Institute. A 0.5 g split of  
183 the sample powder was mixed with 1.0 g of lithium tetraborate and 0.5 g of lithium metaborate. The mixture was  
184 fused at 1000°C for 30 minutes using a Katanax K2 Prime Electric Fusion Machine at KOPRI. The fused glass disks  
185 were analyzed by a LA-ICP-MS system, consisting of a Photon-Machines EXCITE193 excimer laser and a Thermo  
186 Scientific iCAP-Q ICP-MS. The analyses were performed at the laser pulse rate of 10 Hz, beam size of 135  $\mu\text{m}$ , and  
187 energy density of  $\sim 10 \text{ J/cm}^2$ . A NIST 612 glass standard was used as an external standard for calculating the trace  
188 element contents with Si contents obtained by XRF analysis. In order to check the reproducibility and precision of  
189 the analyses we performed multiple analyses of the reference material BHVO-2 (n=7) and BCR-2G (n=2). The  
190 reproducibility calculated using relative standard deviation ( $1\sigma$ ) was mostly less than 5 % for all elements.

191 Whole-rock PGE contents were measured using the Ni sulfide fire assay-isotope dilution method by Park et al.  
192 (2012). A 2.5 g split of the powder was thoroughly mixed with 0.25 g of Ni, 0.125 g of S and 2.5 g of sodium borax.  
193 An enriched PGE spike solution ( $^{99}\text{Ru}$ ,  $^{105}\text{Pd}$ ,  $^{191}\text{Ir}$  and  $^{195}\text{Pt}$ ) was added to the mixture. The spiked powder was dried  
194 at 100°C for 1 hour and fused in porcelain Coors<sup>TM</sup> crucibles in a preheated open furnace at 1150°C for 30 minutes.  
195 The reducing conditions were provided by flowing  $\text{N}_2$  gas into the furnace and adding  $\sim 0.1$  g of cereal flour into the  
196 second outer crucible. After quenching, Ni-sulfide beads were recovered and dissolved in 6M HCl. The solution was  
197 then filtered through a Millipore 0.45  $\mu\text{m}$  cellulose membrane filter to collect the PGE-bearing minerals. The filter  
198 paper containing the PGE minerals was digested in 4 ml of aqua regia, and after complete digestion, the solutions  
199 were dried down to <0.1 ml and diluted with 2%  $\text{HNO}_3$ . The solutions were analyzed for PGE isotopes on an Elan  
200 6100 quadrupole ICP-MS system at the Seoul National University. Potential molecular interferences on the analyzed  
201 isotopes by Ni, Cu, Zn, Co, Hf, Mo, Zr and Ta argides or oxides were monitored and the effects of the interferences  
202 were subtracted using measured oxide (0.2–1.9%) and argide production rates (0.002–0.01%). The interference  
203 correction is <0.2% for all PGE, except for Ru for which the correction is 3–4%. Concentrations of Ir, Ru, Pt, Pd and  
204 Re were determined by isotope dilution using the  $^{191}\text{Ir}/^{193}\text{Ir}$ ,  $^{99}\text{Ru}/^{101}\text{Ru}$ ,  $^{191}\text{Ir}/^{193}\text{Ir}$ ,  $^{195}\text{Pt}/^{194}\text{Pt}$ ,  $^{105}\text{Pd}/^{108}\text{Pd}$  and  
205  $^{185}\text{Re}/^{187}\text{Re}$  ratios, respectively. Concentrations of Rh were obtained using the method by Meisel et al. (2003) and  
206 Park et al. (2012), where count rates of  $^{103}\text{Rh}$  and  $^{106}\text{Pd}$  are used, assuming that the loss of Rh during the analytical  
207 procedure was similar to the loss of Pd. Average procedural blanks measured by sample-free analyses (n=4) were  
208  $22\pm 6$  pg Ir,  $36\pm 25$  pg Ru,  $9\pm 4$  pg Rh,  $268\pm 58$  pg Pt, and  $295\pm 87$  pg Pd. The method detection limits (MDL) were  
209 taken to be three standard deviations of the procedural blanks. MDLs for 2.5 g samples are 8 ppt for Ir, 30 ppt for  
210 Ru, 11 ppt for Rh, 69 ppt for Pt and 105 ppt for Pd. The accuracy and precision of the analyses were assessed by  
211 multiple analyses (n=4) of reference material TDB-1. The reproducibility is 4–11% for Ru, Pt and Pd, 17% for Ir and  
212 29% for Rh. The measured TDB-1 values were consistent with published values.

213

## RESULTS

### 214 Sulfide occurrences

215 Three types of sulfide appearance are recorded in the KM olivine-phyric rocks Figs. 2-4):

216 1) masses in the plagioclase-clinopyroxene-magnetite groundmass, 2) inclusions in olivine and

217 Cr-spinel and 3) daughter phases in silicate melt inclusions hosted by phenocrysts.

218 Sulfide in the groundmass (hereafter, “*matrix sulfide*”) is represented by relatively large  
219 grains (~ 0.1 mm, up to 0.6 mm) with highly variable shapes. Areal measurements of matrix  
220 sulfide grains in 18 thin sections (~ 4 cm<sup>2</sup> each) and S abundances in whole rocks (~ 0.05 wt.%,  
221 Table 1) indicate ~ 0.1 % sulfide is present in the studied rocks, however, the sizes and  
222 distributions of matrix sulfides are highly irregular. The shapes and textures of matrix sulfides  
223 reflect cooling and crystallization of both sulfide liquid and host silicate liquid, and an  
224 overwhelming number of matrix sulfides have had their surfaces “shaped” by other groundmass  
225 minerals (Fig. 2). The host silicate liquid was interstitial to earlier olivine and Cr-spinel  
226 phenocrysts, and crystallized well-formed, interlocking clinopyroxene, plagioclase and Ti-  
227 magnetite groundmass with irregularly shaped, partially flat-walled interstices into which matrix  
228 sulfides crystallized. Accordingly, matrix sulfides (once trapped liquid droplets) are rarely  
229 spherical (or circular to oval in cross-section), and only so if entrapped by plagioclase and Ti-  
230 magnetite (Figs. 2c and 2e). More often, matrix sulfides show curvature at contacts with olivine  
231 phenocrysts (Figs. 2b and 8a). Matrix sulfides are observed squished and even pierced through  
232 by neighboring silicate and oxide phases (Figs. 2d-2g) and wrapped around earlier formed Cr-  
233 spinel (Fig. 2a). The observed scalloped edges, lobate surfaces and tubular embayments of the  
234 matrix sulfides could result from interaction between a molten sulfide and crystallizing silicate-  
235 oxide groundmass. Most matrix sulfides bear signs of alteration (e.g., cracking, oxidation).

236 Sulfide inclusions in olivine and Cr-spinel phenocrysts represent a second type of sulfide  
237 melt in the samples (Figs. 3 and 7). The spherical and, rarely, cylindrical inclusions found in a  
238 very small number of olivine grains (~ 0.7%) are commonly singular (rarely 2-3 per olivine  
239 grain) and relatively small (< 100 μm). Phenocryst-hosted sulfides are undoubtedly entrapped  
240 sulfide melts (e.g., Ackermann et al., 1998; Czamanske and Moore, 1977; Francis, 1990;



241 Gurenko et al., 1987; Kamenetsky and Kamenetsky, 2010), and thus they are labelled “*sulfide*  
242 *melt inclusions*” in this study.

243         The third type of sulfide occurs as small globules (commonly < 1  $\mu\text{m}$ ) hosted in partially  
244 crystalline silicate melt inclusions in olivine and Cr-spinel (Fig. 4). They are interpreted as a  
245 daughter phase that formed during cooling in relation to the so-called Fe loss process, caused by  
246 post-entrapment crystallization and re-equilibration with the host mineral (Danyushevsky et al.,  
247 2002). A number of daughter sulfide globules in a given melt inclusion is hard to estimate due to  
248 opacity of the crystalline silicate and host Cr-spinel, but the number and size of individual  
249 globules are inversely correlated. A given section of any melt inclusion may contain several  
250 equally sized globules (0.3-0.4  $\mu\text{m}$ ; Fig. 4d) or a larger number of much smaller globules (<0.1  
251  $\mu\text{m}$ ). Co-entrapment of sulfide and silicate liquids could also be the case, where sulfides  
252 (typically > 5  $\mu\text{m}$ ) are volumetrically disproportional to enclosing silicate melt, and develop  
253 scalloped embayments and flattened edges at contacts with silicates (Figs. 4c and 4f). No  
254 correlation was found between the size and composition of sulfide globules and their position  
255 relative to the host inclusion walls, vapor bubble(s) or daughter pyroxene crystals (Figs. 4b and  
256 4e).

### 257 **Sulfide phase and major element compositions**

258         Matrix sulfides and sulfide melt inclusions are composed of at least four coarse phases that  
259 can be recognized in reflected light and backscattered electrons. A light-grey, euhedral Fe-oxide  
260 phase, most likely magnetite, is minor and can be absent in a given section, especially in sulfide  
261 melt inclusions. The phase boundaries between intimately intergrown sulfide phases are distinct,  
262 and can be both straight and curvilinear (Figs. 8, 12, 13). The proportion between sulfides

263 revealed in any given exposure is highly variable, which does not permit confident calculation of  
264 the bulk composition.

265 The matrix sulfides analyzed *in situ* by EMPA (Appendix 2<sup>1</sup>) belong to two main  
266 compositional populations that conform to their optical colors. One population is virtually Cu-  
267 free (< 2 wt% Cu) and characterized by variable Fe and Ni contents (Fig. 5a, Table 2, Appendix  
268 3<sup>1</sup>) and (Fe+Ni)/S ratios (at.) that vary between 1.0 and 1.14 and are negatively correlated with  
269 Fe. The second population has high Cu abundances (22-34 wt%), narrow range in Fe (30-37  
270 wt%), large range in Ni/Cu (0.04-0.50, at.; Fig. 5a, Table 2, Appendix 3<sup>1</sup>) and metal/S ratios,  
271 varying between 1.04 and 1.23 (at.) and are positively correlated with Fe. No stoichiometric  
272 minerals were confidently identified among the matrix sulfides, except Ni-poor pyrrhotite. Thus,  
273 we tentatively assigned widely used terms monosulfide solid solution (*Mss*) and intermediate  
274 solid solution (*Iss*) to Fe-Ni and Fe-Cu-Ni phases, respectively. The same two compositional  
275 populations are clearly recognizable in the majority of LA-ICP-MS analyses (Figs. 5b and 8,  
276 Appendix 3<sup>1</sup>).

277 The phases in the sulfide melt inclusions are coarse and compositionally similar to those in  
278 matrix sulfides (Figs. 3b and 3c). In order to estimate bulk compositions, six sulfide melt  
279 inclusions in olivine phenocrysts (87.1-89.6 mol% Fo) were homogenized in a vertical furnace at  
280 the V.I. Vernadsky Institute of Geochemistry (Moscow, Russia). They were heated to 1355 °C  
281 for 5 min at 1 atm of CO<sub>2</sub>:H<sub>2</sub> gas mixture corresponding to the QFM oxygen buffer and then  
282 quenched in water. The resulting textures are fine-grained with different sulfide phases  
283 intermixed on a much smaller scale than in nature (Fig. 7). The heated inclusions show a good  
284 match between the compositions at exposed surfaces and bulk compositions analyzed by SEM-  
285 EDS and LA-ICP-MS methods, respectively (Fig. 6a). The major element compositions of all

286 six homogenized inclusions are very similar, but resemble neither *Mss* nor *Iss*; instead, they  
287 likely represent a mixture of these two end-member sulfide phases in roughly equal proportions  
288 (Fig. 5). The sulfide melt inclusions demonstrate conspicuous changes with fractionation of the  
289 silicate melt, namely positive correlation between Ni, Ni/Fe and Ni/Cu and Fo content of the host  
290 olivine (Fig. 6b, Table 2). One of the homogenized sulfide inclusions is cylindrical and  
291 distinctively larger than other five spherical globules (Fig. 7c), but not different in major element  
292 composition (Table 2, #3).

293 Fifty-nine daughter sulfide globules (0.5-1  $\mu\text{m}$ ) in 35 silicate melt inclusions hosted by Cr-  
294 spinel (Figs. 4d and 4e) were studied by SEM-EDS. Twenty-nine inclusions contained only one  
295 exposed sulfide globule, whereas from 3 to 8 sulfides globules were exposed and analyzed in  
296 other inclusions. Most compositions are Cu-free or Cu-poor (< 5 at%), and belong to the *Mss*  
297 endmember with variable Ni/Fe (0-0.93, average  $0.19 \pm 0.22$ , at.), whereas Cu-bearing sulfides  
298 (> 10 at%) are rare (Fig. 5c). The compositions of daughter sulfides in a given melt inclusion are  
299 highly variable (Fig. 5c) and show no correspondence with the size and position of globules.

### 300 **Sulfide trace and minor element compositions**

301 Thirty-nine matrix sulfides and six sulfide melt inclusions were analyzed using the LA-  
302 ICP-MS method (79 spots) to determine concentrations and distribution of noble metals (PGE,  
303 Au and Re) and trace metals (Zn and Cd). The selection of grains of matrix sulfides was based  
304 on occurrence of both *Mss* and *Iss* phases on polished surface and their size large enough (> 30-  
305 40  $\mu\text{m}$ ) for the laser ablation analysis. Major elements of the sulfide matrix (Fe, Ni, Co, Cu and  
306 S) were also measured together with trace elements, and used for data reduction and to  
307 monitoring the phase compositions. The behavior of the elements during analysis was registered  
308 with time-resolved laser ablation spectra (Figs. 7 and 8), used to determine if sulfides are

309 homogeneous “down hole” (i.e. constant element ratios in the ablated volume), and if they  
310 contain noble-metal phases (hereafter, *nuggets*) large enough to be recorded. The majority of  
311 single point analyses demonstrate homogeneous major element compositions, however several  
312 analytical spots returned “mixed” compositions (i.e. *Mss* + *Iss*), encountered within the analyzed  
313 volume (Fig. 5b; also see Fig. 8A “mt3-5b”). Importantly, all sulfide melt inclusions appeared  
314 homogeneous below their exposed surface (Fig. 7) and show good correspondence between the  
315 surface SEM-EDS and the volume LA-ICP-MS (Fig. 6a).

316 All elements analyzed in the sulfide melt inclusions are above the detection limit (Table 2)  
317 and correspond well (except Pt) to the average composition of the analyzed matrix sulfides (Figs.  
318 9, 10 and 11b). The time-resolved laser ablation spectra of the sulfide melt inclusions are smooth  
319 and bear no evidence of nuggets (Fig. 7). Six sulfide melt inclusions show contents of PGE and  
320 Au that are unexpectedly variable (Figs. 10, 11a and 14, Table 2), given their similar major  
321 element compositions and limited compositional range of their host olivine. It is worth  
322 mentioning that one unusually PGE-enriched composition belongs to the inclusion that is largest  
323 and, unlike other inclusions, cylindrical (#3 in Table 2 and Fig. 7c). Even if this anomalous  
324 composition is excluded, the PGE contents in five smaller and spherical sulfide globules are still  
325 highly variable (by a factor of 2.5 for Rh to 13.8 for Ir) and positively correlated with Fo content  
326 of the host olivine (Figs. 10 and 11a, Table 2).

327 The chalcophile and noble metal compositions of the KM matrix sulfides encompass the  
328 range of the sulfide melt inclusions, but are even more fluctuating for both *Mss* and *Iss* phases  
329 (Figs. 9 and 10, Appendix 3<sup>1</sup>). The chalcophile elements Zn and Cd concentrate in the Cu-rich  
330 sulfide (*Iss*, Figs. 9a and 9b), whereas Ni-rich sulfide (*Mss*) is more enriched in Re (Fig. 10b).  
331 Cobalt, which has chalcophile and siderophile affinities, partitions strongly to *Mss*, however, it

332 correlates positively with Ni in both *Mss* and *Iss* phases of the matrix sulfides (Fig. 9c). Some  
333 analyzed PGE (e.g., Os, Ir and Pt) appear to be below detection limits (~ 0.005-0.01 ppm) in a  
334 large number of analyses (Appendix 3<sup>1</sup>). Based on this detection limits, the overall compositional  
335 variation is at least two orders of magnitude (Fig. 10). The iridium-group PGE (IPGE) correlate  
336 exceptionally well with each other in both *Mss* and *Iss* of the matrix sulfides and sulfide melt  
337 inclusions, but are 3-4 times more abundant in *Mss*.(Fig. 10). In contrast, the platinum-group  
338 PGE (PPGE) and Au show different systematics to each other and IPGE. For example, the  
339 average abundances of Rh are similar in *Mss*, *Iss* and sulfide melt inclusions (~ 0.2 ppm),  
340 covariate with IPGE only in *Mss* and melt inclusions, but remains relatively constant in *Iss* (Figs.  
341 10d and 10e, Table 2). Palladium and gold show weak and strong preference, respectively, for  
342 *Iss*, however, they are not correlated with each other and other PGE in either of studied phases  
343 (Fig. 10f, Table 2). As mentioned above, Pt is often undetectable in the matrix sulfides,  
344 especially *Iss* (Fig. 10h), which is in contrast to its high and variable abundances in the sulfide  
345 melt inclusions (0.44-7.03 ppm, Table 2). Moreover, Pt contents in the matrix sulfides do not  
346 correlate with other PGE and demonstrate a pronounced negative anomaly relative to normalized  
347 (primitive mantle (PM) using values from McDonough and Sun (1995) concentrations of  
348 neighboring Rh and Pd (Fig. 11b).

349 Compositions of the sulfide melt inclusions and the average matrix sulfide from the KM  
350 picrite are similar in terms of major elements (Fig. 5b), IPGE and Re (Figs. 9–11). The average  
351 matrix sulfide is notably more enriched in Zn, Rh and Pd compared to the average sulfide melt  
352 inclusion, but is relatively depleted in Co, Cd, Au and especially Pt (Figs. 9–11). For comparing  
353 PGE abundances in sulfides and whole rock analyses we have assumed that the matrix sulfide is  
354 largely responsible for the overall PGE budget of the rocks. The amount of sulfide present, based

355 on counting statistics and sulfur abundances in whole rocks, is  $\sim 0.1\%$ . Using this value, the  
356 measured PGE contents are multiplied by  $10^3$  and show good agreement with the sulfide PGE  
357 pattern (Fig. 11b), therefore, reinforcing our original assumption.

358 The time-resolved laser ablation spectra recorded the presence of dominantly Pd- or/and  
359 Au-bearing nuggets of variable size in 57% of analytical spots (Fig. 8). These nuggets,  
360 previously observed by SEM-EDS on polished surfaces of matrix sulfides, are described in more  
361 detail below.

### 362 **Noble-metal nuggets in matrix sulfides**

363 Noble metal nuggets are common in all matrix sulfide grains, with occasionally 10-15  
364 particles observed on a given surface ( $\sim 500 \mu\text{m}^2$ ; Fig. 12). They usually occur on phase  
365 boundaries and in cracks, and range significantly in size ( $< 0.05$  to  $1 \mu\text{m}$ ) even in the same field  
366 of view (Fig. 12). The small size of nuggets hampers accurate determination of compositions by  
367 energy-dispersive spectrometry because the analytical volume (at an accelerating voltage of 15-  
368 17 keV, specimen density  $\sim 4.5 \text{ g/cc}$ ) is around  $1\text{--}1.2 \mu\text{m}$  across, exceeding the size of most  
369 nuggets. Semi-quantitative compositions of nuggets  $\geq 0.3 \mu\text{m}$  can be estimated by subtracting  
370 matrix elements (Fe, Ni, Cu, and S) from the analyses. In most cases the neighboring nuggets  
371 show diverse compositions and morphology (Fig. 12). There is no correlation between the  
372 compositions of the sulfide matrix (i.e.  $M_{ss}$  or  $I_{ss}$ ) and included nuggets. Palladium is by far the  
373 most abundant element occurring with Sn as a Pd-Sn alloy with variable Pd/Sn (1.7-2.5), but  
374  $\text{Pd}_2\text{Sn}$  (paolovite) is the most common. Alloys of Au and Ag mixed in different proportions (0.5-  
375 2.7 at.), as well as pure Au, is the second dominant phase. Typically both types of alloys (Pd-Sn  
376 and Au-Ag) occur in close vicinity to each other (Fig. 12), which is also seen in the time-

377 resolved laser ablation spectra (Fig. 8). Sometimes Ag-free gold contains minor Pd (3-14 %),  
378 whereas low quantity of Zn and Pt maybe present in Au-Ag alloys.

379 In addition to common Pd-Sn and Au-Ag nuggets in matrix sulfides we record phases that  
380 contain Pt-As ± Ni-As ± Au ± Ag ± Te, Bi. They are larger (up to 2-3 μm) and occur along and at  
381 intersection of cracks, at the contacts of sulfides with the alteration silicates (chlorite, serpentine)  
382 and within the latter (Fig. 13). Typically they appear as a chain-like “decoration” of the sulfide  
383 grain boundaries and even take the form of micro-veinlets (Fig. 13). Sperrylite (PtAs<sub>2</sub>), Au-Ag  
384 alloys, maucherite (Ni<sub>11</sub>As<sub>8</sub>) and Pt<sub>3</sub>(As,Te,Bi)<sub>7</sub> dominate the population.

## 385 **DISCUSSION**

### 386 **Early sulfide saturation in oceanic magmas**

387 Petrography and chemical data suggest that the MORB-type magmas from Kamchatsky  
388 Mys achieved the sulfide saturation. The common occurrence of sulfides in the pyroxene-  
389 plagioclase-Ti-magnetite groundmass and their irregular, often deformed shapes (Fig. 2) imply  
390 that these sulfides were in a liquid or plastic state at the time of eruption and were affected by  
391 groundmass minerals growing around and through them. A close similarity between the average  
392 composition of matrix sulfides and sulfide globules included in olivine phenocrysts from same  
393 samples (Figs. 5b, 9 and 10) suggests that studied sulfides belong to the same population of  
394 immiscible melts.

395 Our data are consistent with the records of omnipresent sulfide globules in the Atlantic  
396 and Pacific basalts and glasses (e.g., Czamanske and Moore, 1977; Kanehira et al., 1973;  
397 MacLean, 1977; Mathez, 1976; Moore and Calc, 1971) and reconfirm sulfide saturation of  
398 common MORB melts during crystallization (e.g., Haughton et al., 1974; Mathez, 1976).  
399 Moreover, sulfide melt inclusions in primitive olivine in this (Fig. 3, 7; Table 2) and other

400 studies (e.g., Kamenetsky and Kamenetsky, 2010) should be considered as the earliest  
401 manifestation of the silicate-sulfide immiscibility in oceanic magmas. Such records are  
402 exceptionally rare, but provide the strongest support for sulfide immiscibility in the crust without  
403 the requirement of magma contamination by sulfur-enriched crustal material and/or extensive  
404 fractionation, advocated in some studies (see review in Naldrett, 2004).

#### 405 **Composition of incipient sulfide melts**

406 The homogenized sulfide melt inclusions in primitive olivine (Fig. 7) provide a snapshot  
407 of natural sulfide liquids at the onset of immiscibility. Once entrapped in olivine the sulfide melt  
408 inclusions are effectively insulated from further interactions with their parental silicate melts,  
409 and thus they are considered to maintain their original composition in terms of major and trace  
410 elements, including noble metals.

411 In contrast to constraints from theoretical and experimental modelling (e.g., Naldrett,  
412 2004; Zhang, 2015), the olivine-hosted sulfide globules in this and other studies (e.g.,  
413 Ackermann et al., 2007; Czamanske and Moore, 1977; Francis, 1990; Zelenski et al., 2017) are  
414 not pure FeS, but contain substantial amounts of other metals (Fig. 5;  $(\text{Ni}+\text{Cu})/\text{Fe} = 0.47\text{-}0.59$ ).  
415 Similar Ni- and Cu-rich compositions of primitive sulfide melts are reported for globules in  
416 MORB glasses (e.g., Kamenetsky et al., 2013; Keith et al., 2017; Patten et al., 2012; Peach et al.,  
417 1990; Yang et al., 2014). In particular, there is a good match between the major element  
418 compositions of KM sulfide melt inclusions and high Ni/Cu sulfide globules in the FAMOUS  
419 glass ALV 526-1 (Fig. 5a). Noteworthy, a significant range in the individual compositions of 22  
420 sulfide globules in a single glass ALV 526-1 ( $\text{Ni}/\text{Cu} = 0.6\text{-}1.7$  (at) as in Fig. 5a; 1220-1870 ppm  
421 Co, 2-23 ppm Au, 35-1472 ppb Ir; Peach et al., 1990) clearly demonstrates that sulfide melts can  
422 be compositionally heterogeneous on a very local scale (i.e., within the same silicate liquid).



423 Likewise, small-scale disequilibrium is demonstrated by variable compositions of daughter  
424 sulfide globules in single melt inclusions in our study (Figs. 4 and 5c). On the other hand, “*a*  
425 *state close to chemical equilibrium with the host magma*“ was proposed for individual sulfide  
426 globules in seven MORB glasses (Patten et al., 2013), based on corresponding compositions in  
427 terms of Ni, Cu and Co abundances. The results by Patten et al. (2013) and compositional  
428 relationships between incipient sulfide melts and host olivine (Figs. 6b and 11a) show that the  
429 evolution of MORB sulfide liquids is characterized by decreasing Ni/Cu. This mirrors the effects  
430 of olivine crystallization (i.e., Ni intake) in parental magmas (e.g., Kamenetsky and Eggins,  
431 2012). On the other hand, compositions of individual sulfide globules hosted by the same glass  
432 (Fig. 5a; Peach et al., 1990) strongly imply their derivation from compositionally different  
433 batches of magma and incomplete re-equilibration with the melt that transported them to the  
434 surface (e.g., Yang et al., 2014). Additional insights into equilibration of incipient sulfide melts  
435 can be provided by noble metal compositions of sulfide melt inclusions and matrix sulfides.

436 It is expected that PGE abundances will decrease with fractionation of olivine and Cr-  
437 spinel and concomitant separation of immiscible sulfide liquids. The olivine-hosted sulfide melts  
438 in our study, excluding one anomalously enriched composition (#3 in Table 2, Fig. 7c),  
439 demonstrate a four-fold to seven-fold decrease in PPGE and IPGE, respectively, and co-variation  
440 of PGE with olivine Fo and sulfide Ni/Cu (Figs. 11a and 14). However, anticipated smooth and  
441 (sub) parallel trends for normalized compositions are not observed for these early sulfide melts  
442 (Fig. 11a). Moreover, the correlations between PGE observed in our large dataset (Fig. 10) are  
443 not mimicked by dependence of PGE ratios (except Pd/Os, Rh/Ru and Rh/Os) on the  
444 compositions of sulfide matrix (Ni/Cu) and host olivine (Fig. 14, Table 2). Unlike total PGE  
445 abundances, Au contents seem to be unrelated to olivine and sulfide fractionation, however,

446 Au/PGE ratios are remarkably consistent (Fig. 14, also see Mungall and Brennan, 2014). It is  
447 noteworthy that each trend of the fractionation-driven increase in Au/PGE ratios, except Au/Os,  
448 has at least one outlier represented by sulfide melt inclusions #2 (Au/Ru), #3 (Au/Pd), and #4  
449 (Au/Pt, Au/Ir) in Table 2. Five analyzed sulfide melts have moderately to high fractionated PGE  
450 except the inclusion #5 (c.f., PPGE/IPGE = 3.0–6.4 and 0.90; Table 2, Fig. 14).

451 Although the overall compatible behavior of PGE in magmatic fractionation is  
452 confirmed, we are concerned with 1) the large magnitude of PGE variation observed in a small  
453 number of studied primitive sulfide melts in olivine of restricted composition  $FO_{87.1-89.6}$ ; and 2)  
454 the lack of correlation between the degree of fractionation and PGE ratios (except very few  
455 shown on Fig. 14). Additionally, at least four out of six studied sulfide melt inclusions have  
456 abundances of certain noble metals that do not fit anticipated fractionation trends and are at odds  
457 with other PGE. The above concerns are highly pertinent to the crystallization of noble metal  
458 phases in primitive magmas (e.g., Kamenetsky et al., 2015; Zelenski et al., 2016; Zelenski et al.,  
459 2017 and references therein). Such phases may increase the availability of PGE+Au to  
460 concurrent sulfide melts through direct entrapment or have the opposite effect by reducing the  
461 noble metal budget of the silicate melt undergoing sulfide immiscibility.

#### 462 **Effects of crystallization and alteration of sulfide melts**

463 The shapes, textures, and compositions of matrix sulfide in the KM olivine-phyric rocks  
464 record a prolonged and complex process starting with solidification of the sulfide liquid as both  
465 Fe-Ni and Fe-Cu-Ni solid solutions, followed by solid-state recrystallization and transformation,  
466 and ultimately modifications by alteration and weathering. The shapes of the matrix sulfides  
467 (Fig. 2) conform to an essentially liquid state at the time of olivine accumulation. Bulk metal  
468 contents of the matrix sulfides do not seem to influence their shape and texture. Although a few

469 compositions might represent Ni-bearing pyrrhotite, other in-situ analyses do not correspond to  
470 stoichiometric minerals (Fig. 5a). The two principal components of the matrix sulfides represent  
471 *Mss* and *Iss* that are recorded in MORB sulfide globules worldwide (e.g., Keith et al., 2017;  
472 Patten et al., 2012; Patten et al., 2013; Yang et al., 2014 and references therein). Our  
473 observations are not conclusive about the order of crystallization/solidification of *Mss*, *Iss* and  
474 small magnetite grains in the KM matrix sulfides.

475 Only estimates of phase compositions are possible due to the heterogeneity of matrix  
476 sulfides. As they represent former sulfide melts cogenetic with primitive olivine, which makes  
477 the bulk of picrite rocks, their compositions are envisaged as also relatively primitive. This is  
478 supported by the close similarity between the average matrix sulfide (*Mss* and *Iss* in 39 grains)  
479 and sulfide melt inclusions in terms of major and chalcophile trace elements and IPGE (Figs. 9,  
480 10 and 11b). The average matrix sulfide is slightly enriched in Pd compared to most sulfide melt  
481 inclusions, but depleted in Pt (c.f., Pt/Pd 0.11 vs 0.7-3.1) and Au (c.f., Au/Pd 0.6 vs 1.0-7.1).  
482 This relative depletion of matrix sulfides in Pt and Au could be caused by pervasive low-  
483 temperature alteration of KM rocks. Preferential low-temperature extraction of Pt and Au from  
484 matrix sulfides, mobility and subsequent precipitation is confirmed by occurrence of minerals  
485 and alloys of Pt and Au in the vicinity of sulfides (Fig. 13). Occurrences of Pt and Ni arsenides  
486 and Au-Ag alloys in chain-like aggregates along fractures in sulfides, at sulfide-silicate  
487 interfaces, and in hydrous silicates (Fig. 13) support their non-magmatic origin. Low-temperature  
488 crystallization of PGE-bearing phases in the rocks containing magmatic sulfides and Cr-spinel is  
489 not unusual (e.g., Barkov and Fleet, 2004).

490 The effects of alteration on *Mss* and *Iss* are expected to be selective and different for  
491 certain elements. Therefore, the predicted partitioning of IPGE into Fe-Ni and PPGE into Fe-Cu

492 sulfide phases (e.g., Ballhaus et al., 2001) cannot be fully evaluated in our study. However, if  
493 disturbance of magmatic abundances of IPGE by alteration was insignificant, these elements do  
494 show preference for *Mss*, but are highly variable in both *Mss* and *Iss* (Fig. 10). Average contents  
495 of Rh are similar in both phases, however, in contrast to nearly constant Rh in *Iss*, Rh in *Mss*  
496 varies over two orders of magnitude and correlates with IPGE (Fig. 10). Chalcophile elements  
497 Zn and Cd show strong preference for *Iss*, whereas Co and Re are concentrated in *Mss*. The  
498 anticipated partitioning of Pd into *Iss* is not confirmed in our study, as both sulfide phases in  
499 matrix sulfides can be extremely depleted or enriched in Pd (0.01 to ~ 30 ppm; Fig. 10, Table 2,  
500 Appendix 2<sup>1</sup>). Furthermore, the presence of abundant micro- to nanoscale Pd-Sn alloys in both  
501 *Mss* and *Iss* (Figs. 8 and 12) makes understanding of Pd partitioning even more difficult.

502 Finally, the PGE systematics in the KM picrites is comparable to those of sulfide melt  
503 inclusions and matrix sulfides, except Pt (Fig. 11b). The comparison is based on the petrographic  
504 estimate of 0.1% sulfide and demonstrates that the whole rock PGE budget is dominated by PGE  
505 in matrix sulfides. The deficit of Pt in matrix sulfides compared to whole rock and olivine-hosted  
506 inclusions (Fig. 13) can be explained by formation of hydrothermal Pt-rich phases during  
507 serpentinization.

## 508 **IMPLICATIONS**

509 Our study of the MORB-type picrite rocks from Kamchatka demonstrate that magmatic  
510 sulfides in volcanic rocks and associated silicate and oxide minerals and their silicate melt  
511 inclusions is a natural laboratory for understanding silicate-sulfide immiscibility and related  
512 partitioning of noble and chalcophile metals. Our data and observations imply that magmatic  
513 sulfide liquids undergoing cooling and solidification separate into Fe-Ni and Fe-Cu-Ni phases  
514 that have significant compositional ranges in terms of metal/metal, metal/sulfur and PGE ratios.

515 These two principal sulfide phases not only share the trace and noble metals dissolved in the  
516 parental sulfide liquid, but are also responsible for in-situ fractionation of these elements. The  
517 established partitioning of IPGE and PPGE+Au into *Mss* and *Iss*, respectively, should be further  
518 studied using matrix sulfides from modern MORB. We draw attention to redistribution of noble  
519 metals in magmatic sulfides during their crystallization (i.e. formation of “nuggets”) and low-  
520 temperature alteration of their host rocks. Our results suggest that the latter process affected the  
521 matrix sulfides and was responsible for leaching some noble metals and their deposition as  
522 hydrothermal phases of Pt, Au and Ag. Thus any ‘orthomagmatic’ sulfide deposit, composed of  
523 stoichiometric phases (pyrrhotite, pentlandite and chalcopyrite) and having local enrichments  
524 and depletions in certain noble metals should not be considered ‘primary magmatic’. Instead,  
525 further research into the origin of accumulations of Fe-Ni-Cu sulfides with high PGE contents in  
526 magmatic rocks may pay more attention to post-solidus modifications of original sulfide liquids  
527 including hydrothermal mobility and redistribution of the noble metals.

528

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687

688 **FIGURE 1.** Schematic geological map of Kamchatsky Mys (modified after Savelyev, 2014)  
689 showing location of studied rocks (star). (1-4) – sedimentary rocks of the Pliocene-Quaternary  
690 (1), Miocene (2), Turonian-Campanian (3) and Albian–Cenomanian (4) ages. Early Cretaceous  
691 ophiolite is represented by ultramafic rocks (5), gabbro and dolerites (6) and serpentinites  
692 mélange (7). Geological boundaries are shown as faults (8) and thrusts (9).

693 **FIGURE 2.** Sulfides in the groundmass of Kamchatsky Mys picrites. Shapes of matrix sulfides (S)  
694 vary from round to semi-spherical to cusped and angular, depending on surrounding grains of  
695 olivine (ol), Cr-spinel (sp), Ti-magnetite (mt), plagioclase (pl) and clinopyroxene (cpx). All  
696 images are taken in backscattered electrons, except transmitted light in Fig. 2c.

697 **FIGURE 3.** Sulfides globules (S) in olivine (ol) phenocrysts. Globules, interpreted as sulfide melt  
698 inclusions, are composed of Fe-rich, Ni poor phase, resembling pyrrhotite (*Fe-Mss*), Fe-Ni (*Mss*)  
699 and Fe-Cu-Ni phase (*Iss*). Sulfide globules associate with inclusions of Cr-spinel (sp) and silicate  
700 melt (MI). Hydrated silicate minerals (sr) replace olivine along numerous fractures.  
701 Fragmentation of olivine by fractures and related alteration significantly reduce the number of  
702 sulfide globules available for homogenization and subsequent analyses. Images are taken in plain  
703 transmitted light (Fig. 3a) and backscattered electrons (Figs. 3b and 3c).

704 **FIGURE 4.** Daughter sulfides globules (S) in partially crystallized silicate melt inclusions hosted  
705 by phenocrysts of olivine (ol, Figs. 4a–4c) and Cr-spinel (sp, Figs. 4d–4f). Apart from small  
706 daughter sulfides (Figs. 4a, b and e), the melt inclusions contain daughter clinopyroxene (cpx),  
707 vapor bubbles (b), residual silicate glass (G) and accidentally trapped sulfide melts. The latter are  
708 distinguished by larger size and non-spherical shapes (Figs. 4c and 4f). Images are taken in plain  
709 transmitted light (Fig. 4a) and backscattered electrons (Figs. 4b–4f).

710 **FIGURE 5.** Compositional variability of sulfide globules (at%) in the Kamchatsky Mys picrites.  
711 Sulfide melt inclusions in olivine SMI (open circles, Table 2) are compared to matrix sulfides  
712 (Fe-Ni (*Mss*) and Fe-Cu-Ni (*Iss*) phases), analyzed by the electron microprobe (diamonds, Fig.  
713 5a, Appendix 2<sup>1</sup>) and laser ablation ICPMS (triangles, Fig. 5b), daughter sulfide globules in Cr-  
714 spinel-hosted melt inclusions (Fig. 5c) and 22 individual sulfide globules in MORB glass ALV  
715 526-1 (closed circles on Fig. 5a after Peach et al., 1990). Fig. 5b also shows the average matrix  
716 sulfide (av.MS, black circle) as analyzed by laser ablation ICPMS, and fields of *Mss* and *Iss*  
717 compositions, denoted by the electron microprobe data from Fig. 5a. Solitary daughter sulfides in  
718 melt inclusions are shown by crosses on Fig. 5c; all other symbols represent multiple (from 3 to  
719 8) globules in single inclusions.

720 **FIGURE 6.** The compositions of homogenized sulfide melt inclusions, analyzed by SEM-EDS at  
721 exposed surfaces and by LA-ICPMS in volume (a) and their relationships with the host olivine  
722 forsterite content (b).

723 **FIGURE 7.** Time-resolved laser ablation signals (y axis, counts per second (cps)) recorded for  
724 selected elements (masses) in homogenized sulfide melt inclusions in olivine. The  
725 microphotographs show the appearance of sulfides in reflected light and their corresponding  
726 numbers, as per Table 2.

727 **FIGURE 8.** Reflected light microphotographs showing representative matrix sulfides composed  
728 of *Mss* and *Iss* and time-resolved laser ablation signals (y axis, counts per second (cps)) recorded  
729 for selected elements (masses) in these phases (see Appendix 3<sup>1</sup>). Note distinct Au- and Pd-rich  
730 phases in sulfides (see text and Fig. 12 for more details).

731 **FIGURE 9.** Relationships between Co and chalcophile elements (Ni, Cd, and Zn) in the matrix  
732 sulfides and sulfide melt inclusions. Note that Co and Ni may have siderophile and chalcophile  
733 affinities.

734 **FIGURE 10.** Relationships between the noble metal (PGE, Au) contents in the matrix sulfides and  
735 sulfide melt inclusions.

736 **FIGURE 11.** Primitive mantle (after McDonough and Sun, 1995) – normalized concentrations of  
737 noble metals in (a) individual sulfide melt inclusions and (b) two Kamchatsky Mys picrites,  
738 compared to average sulfide melt inclusion (av.SMI) and average matrix sulfide (av.MS)  
739 compositions. For comparison, the PGE abundances in the rocks are increased by a factor of  $10^3$ ,  
740 based on estimated amounts of matrix sulfide (~ 0.1%, see text for details). Analytical precision  
741 ( $2\sigma$ ) for individual analyses is shown as bars on Fig. 10a.

742 **FIGURE 12.** Backscattered electron images showing noble metal “nuggets” (circled) of  
743 dominantly Pd-Sn and Au-Ag compositions in matrix sulfides.

744 **FIGURE 13.** Backscattered electron images showing phases containing PtAs<sub>2</sub>, Au–Ag and Ni  
745 arsenides in fractures within matrix sulfides (Figs. 12a–12c), and in surrounding alteration-  
746 related hydrous silicates (Figs. 12b–12d).

747 **FIGURE 14.** Relationships between the noble metal ratios in sulfide melt inclusions (open circles)  
748 and their Ni/Cu (at.) and host olivine composition (Fo). The average matrix sulfide is shown with  
749 closed circle.

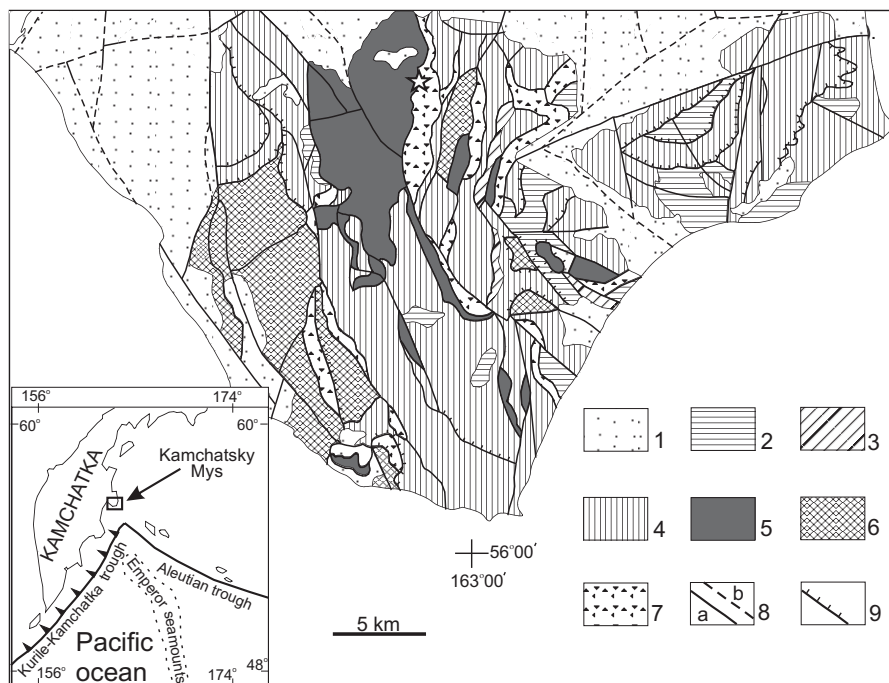


Fig. 1

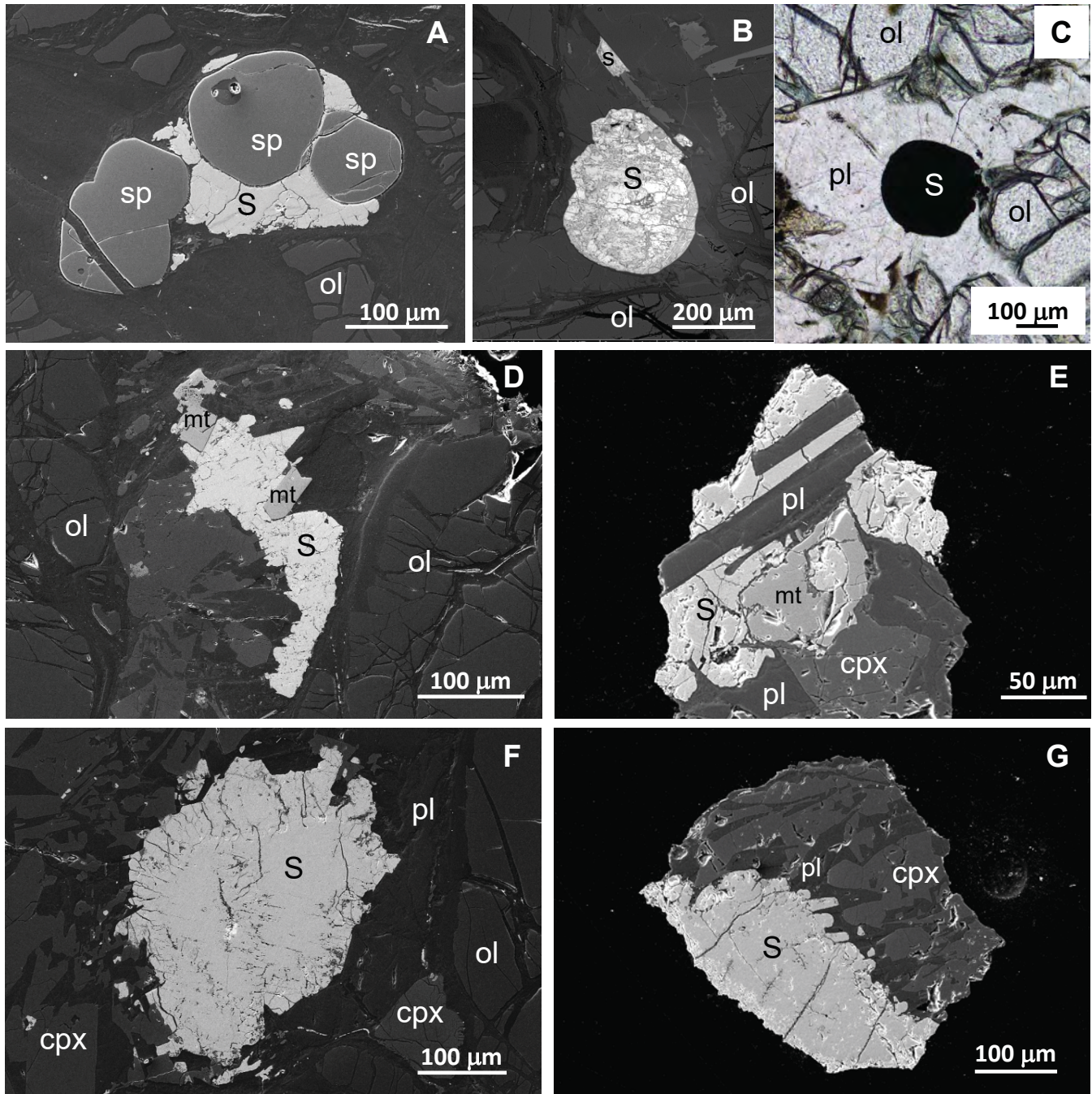


Fig. 2.

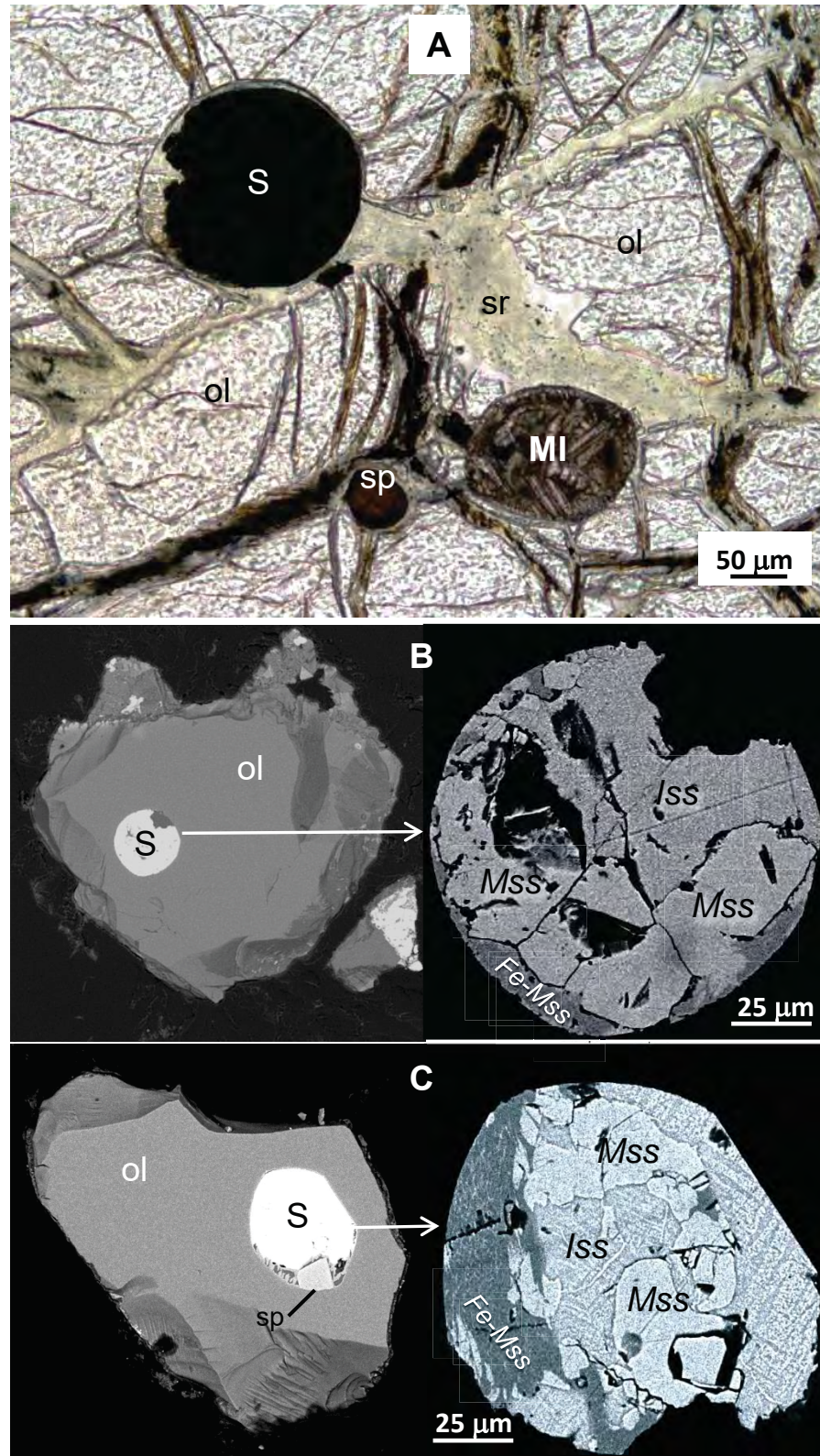


Fig. 3.

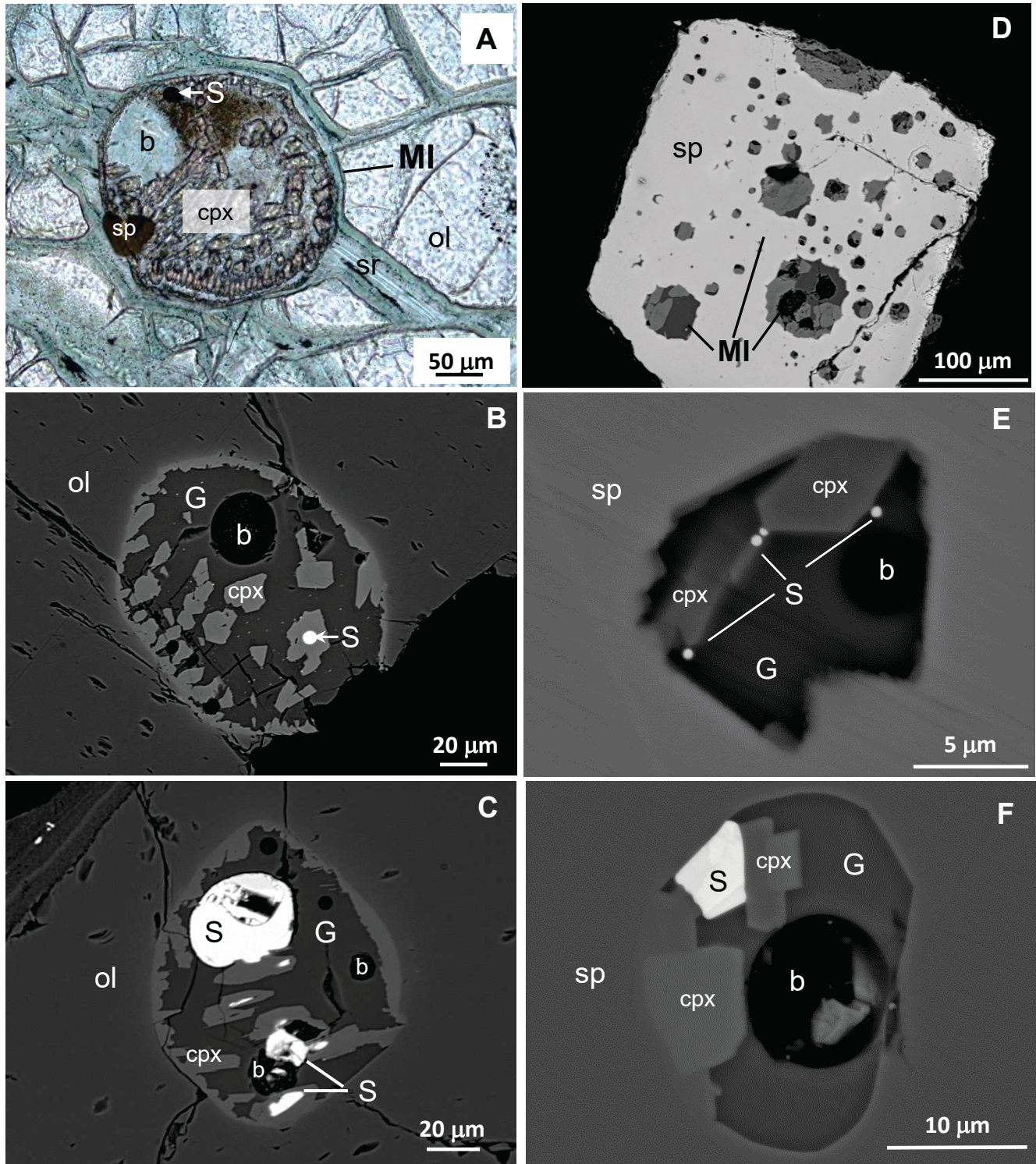


Fig. 4.

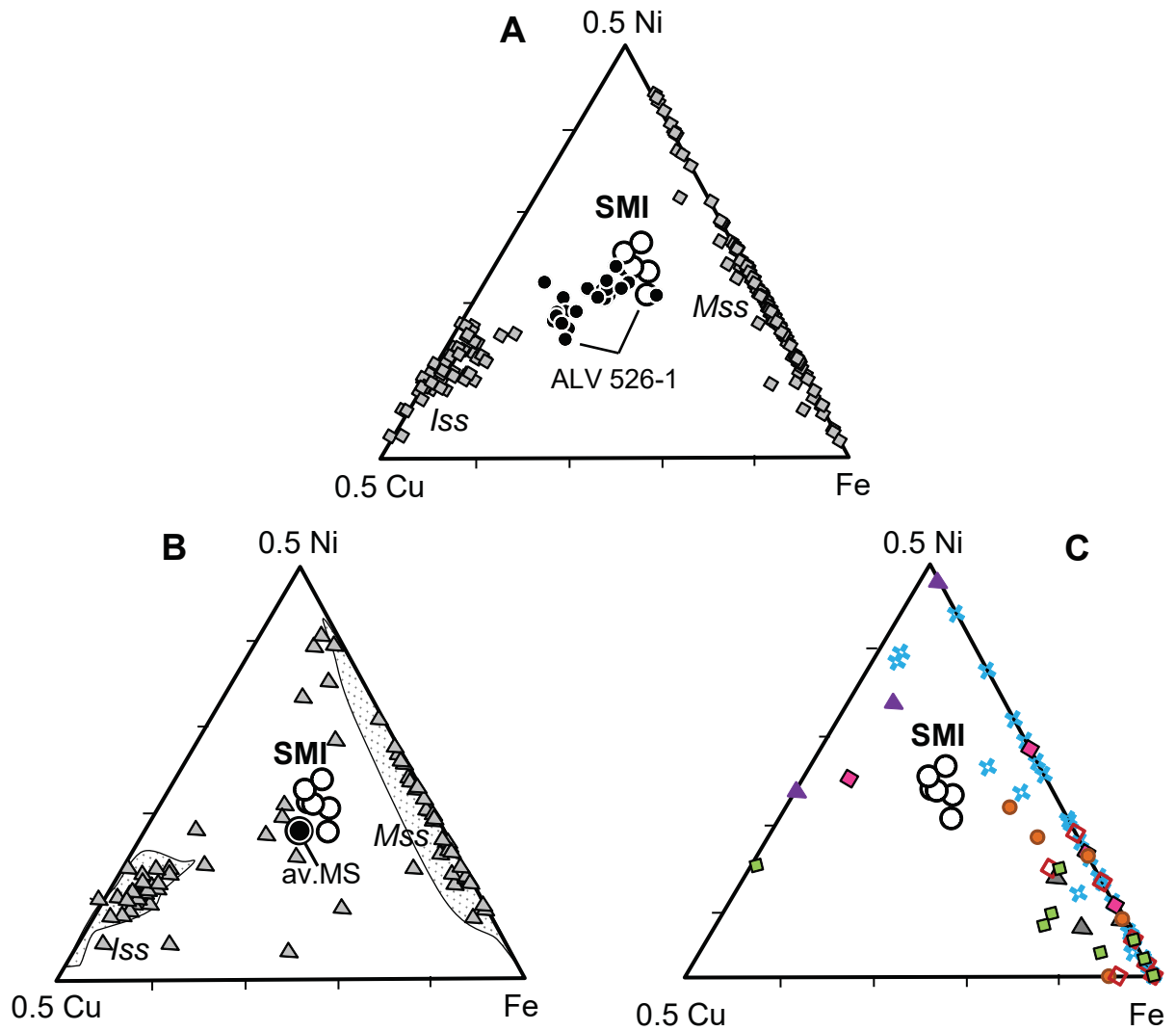


Fig. 5.



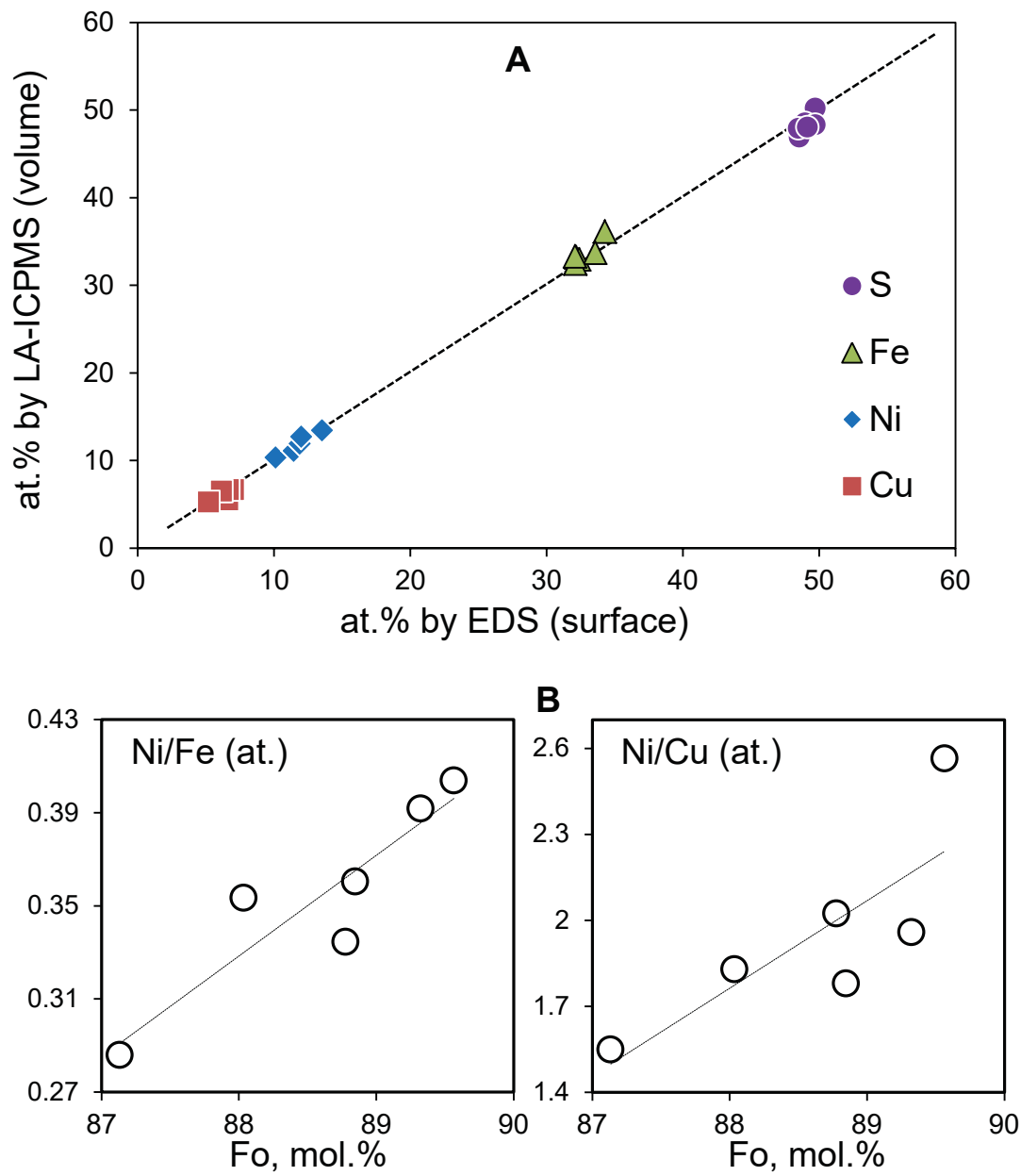


Fig. 6.

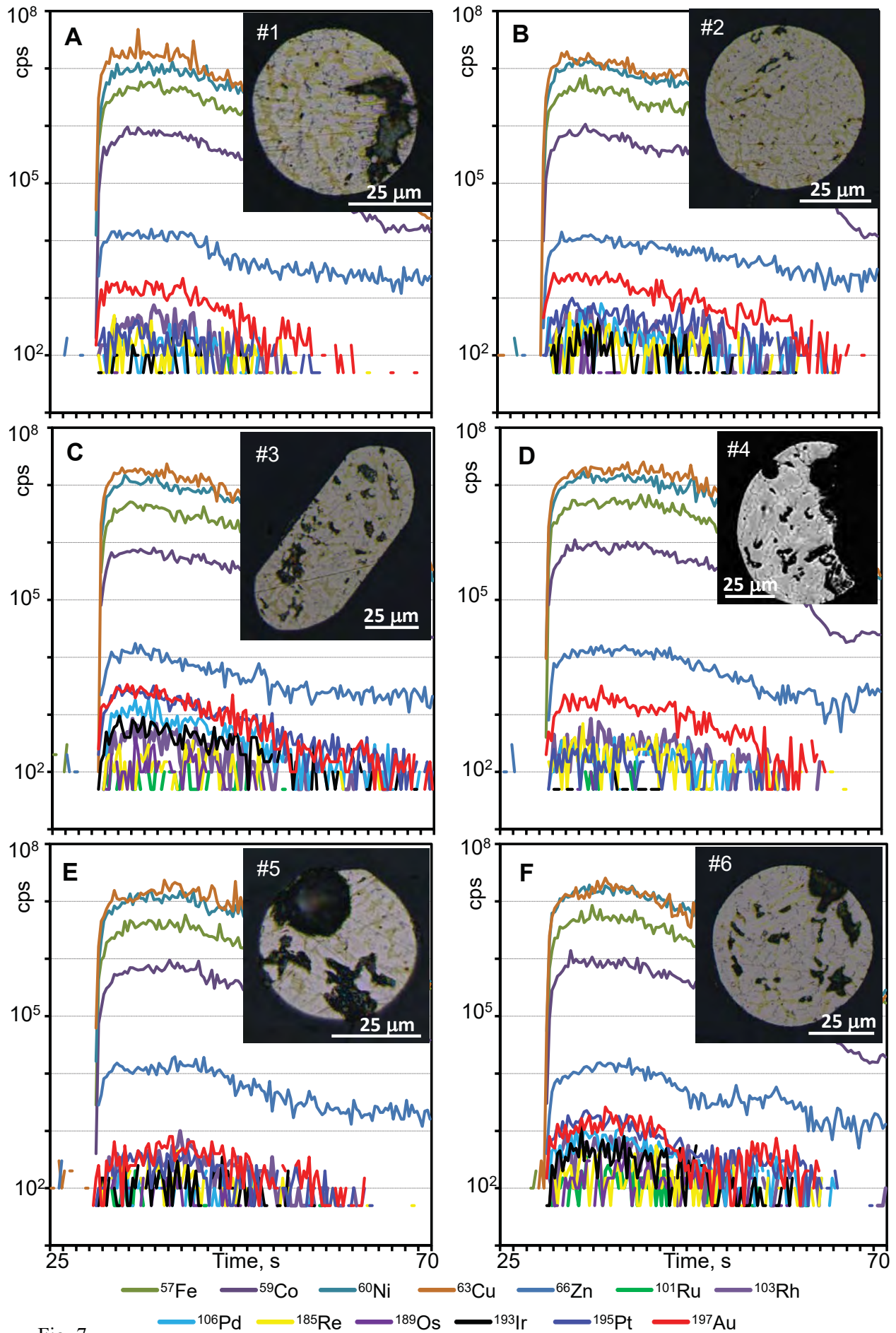


Fig. 7.

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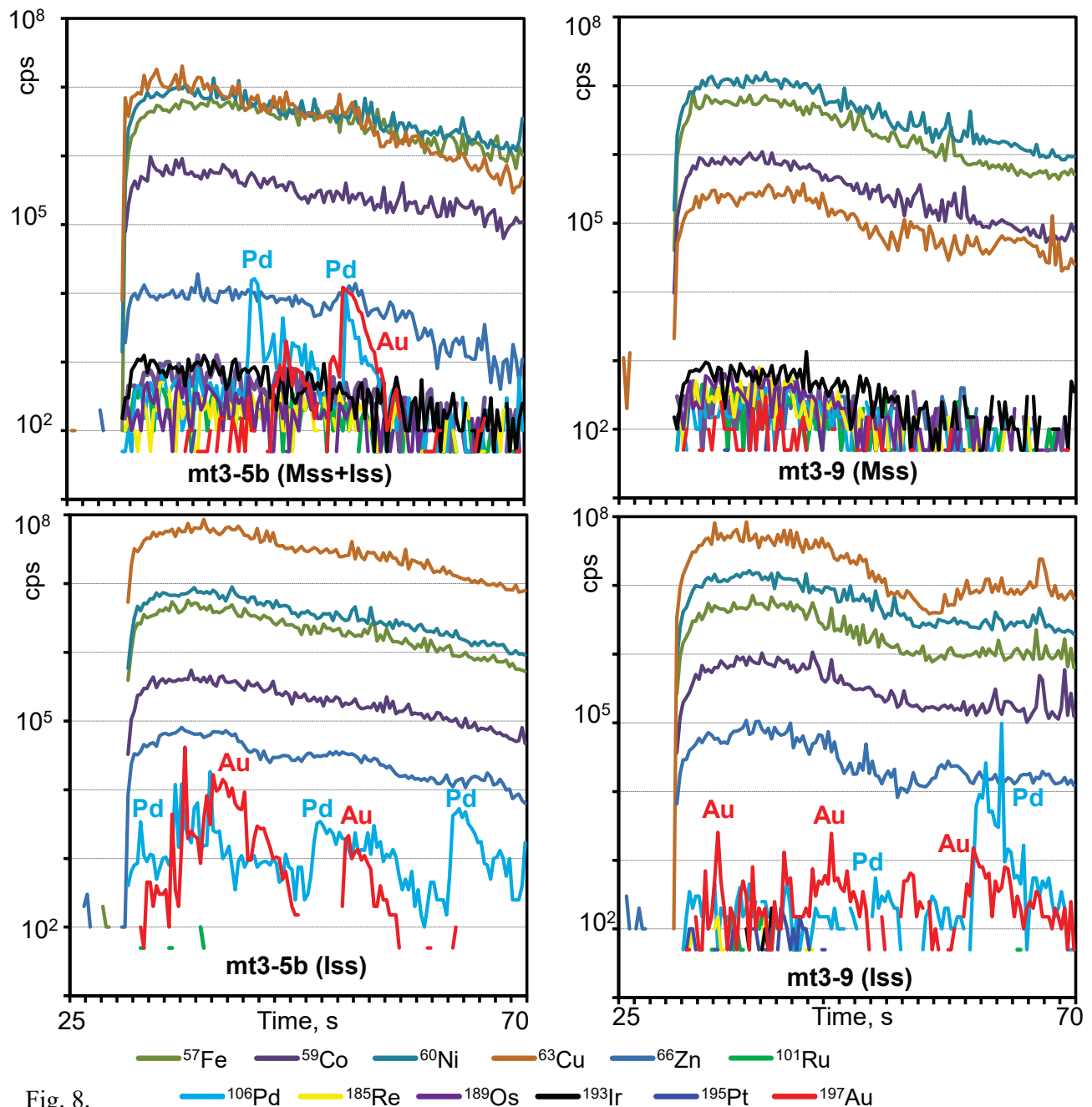
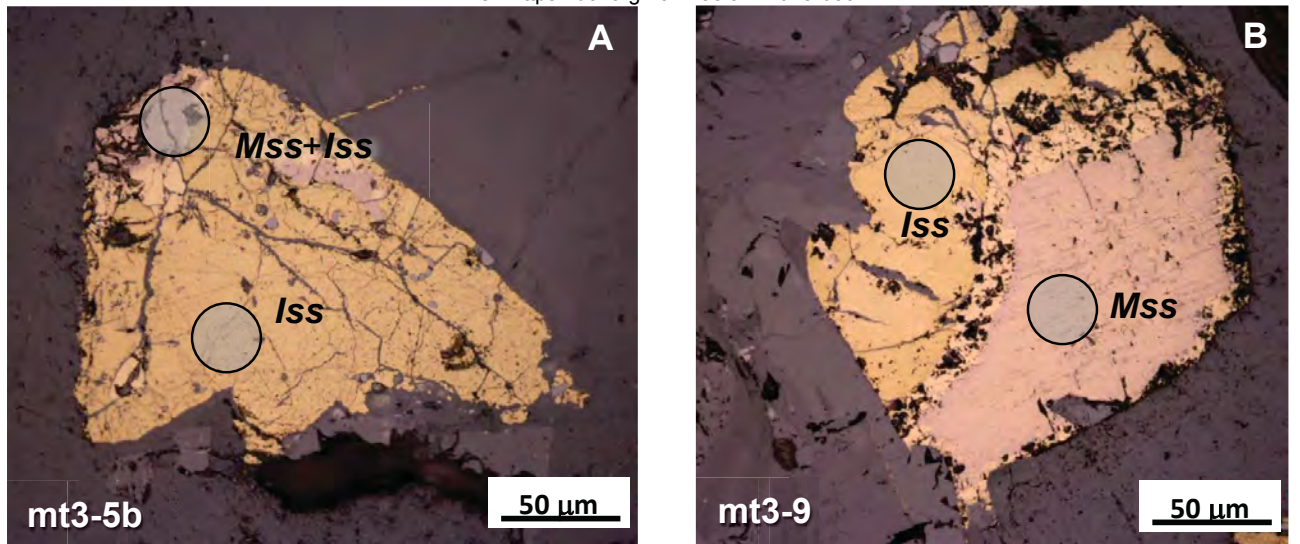


Fig. 8.

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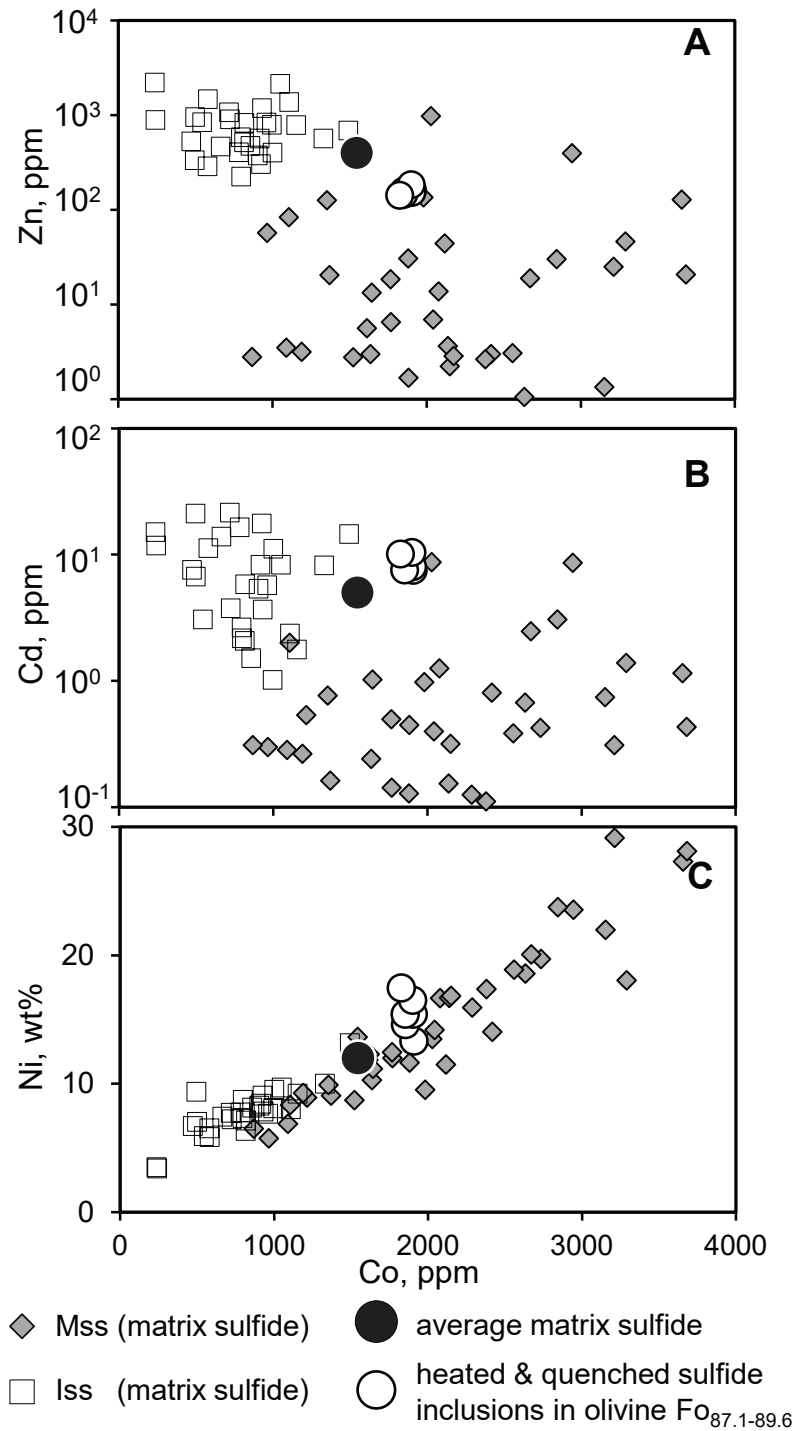


Fig. 9.

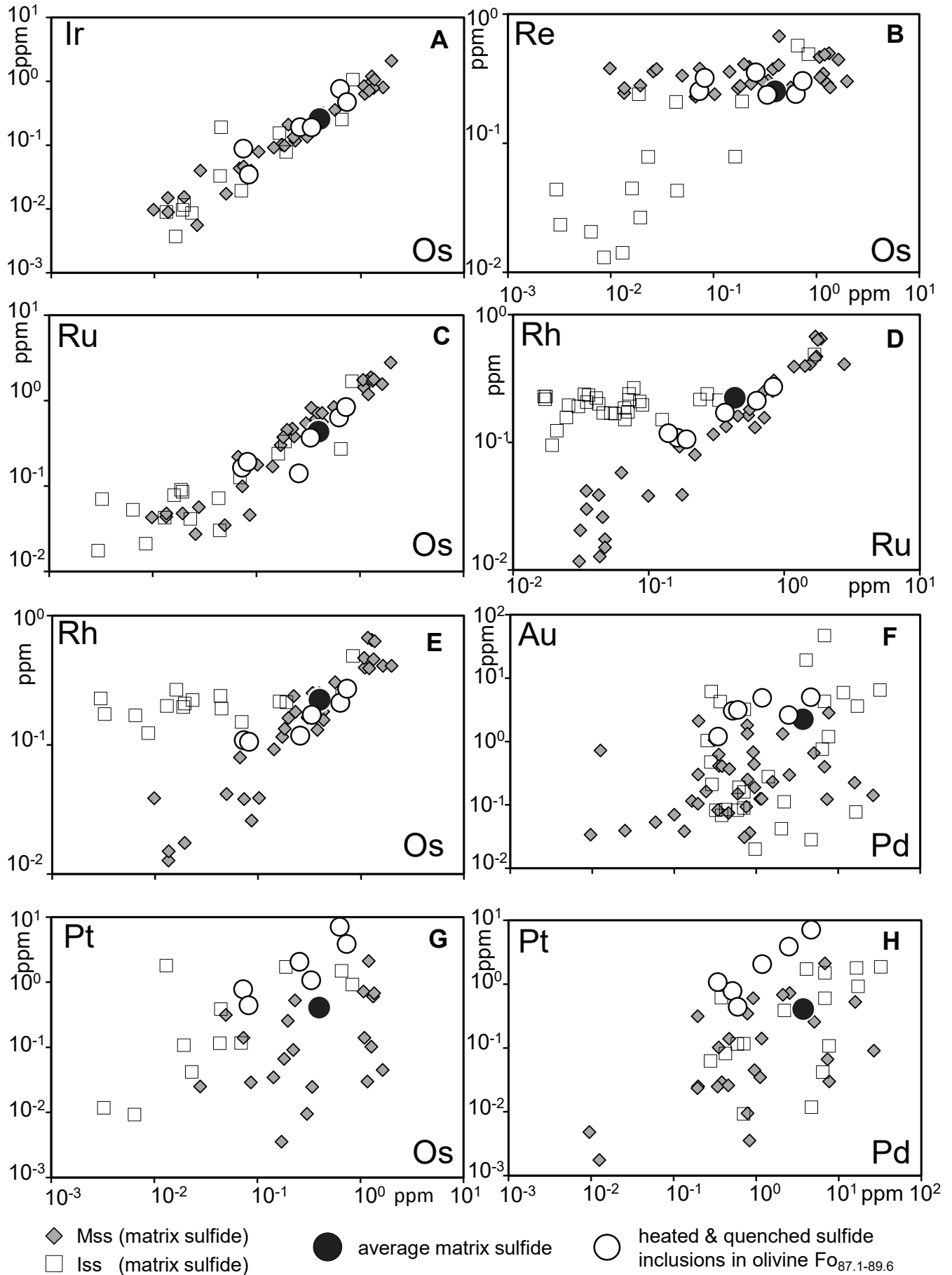


Fig. 10.

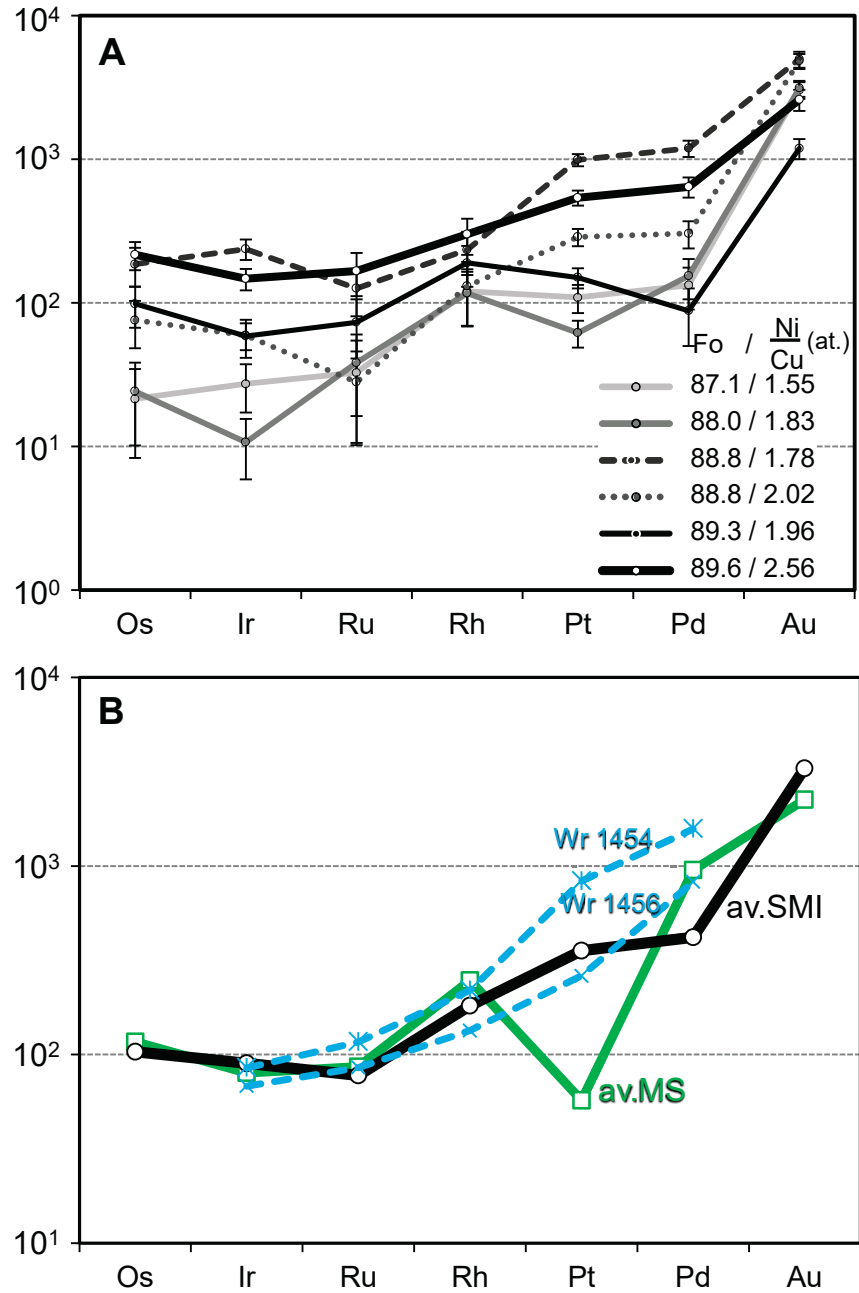


Fig. 11.

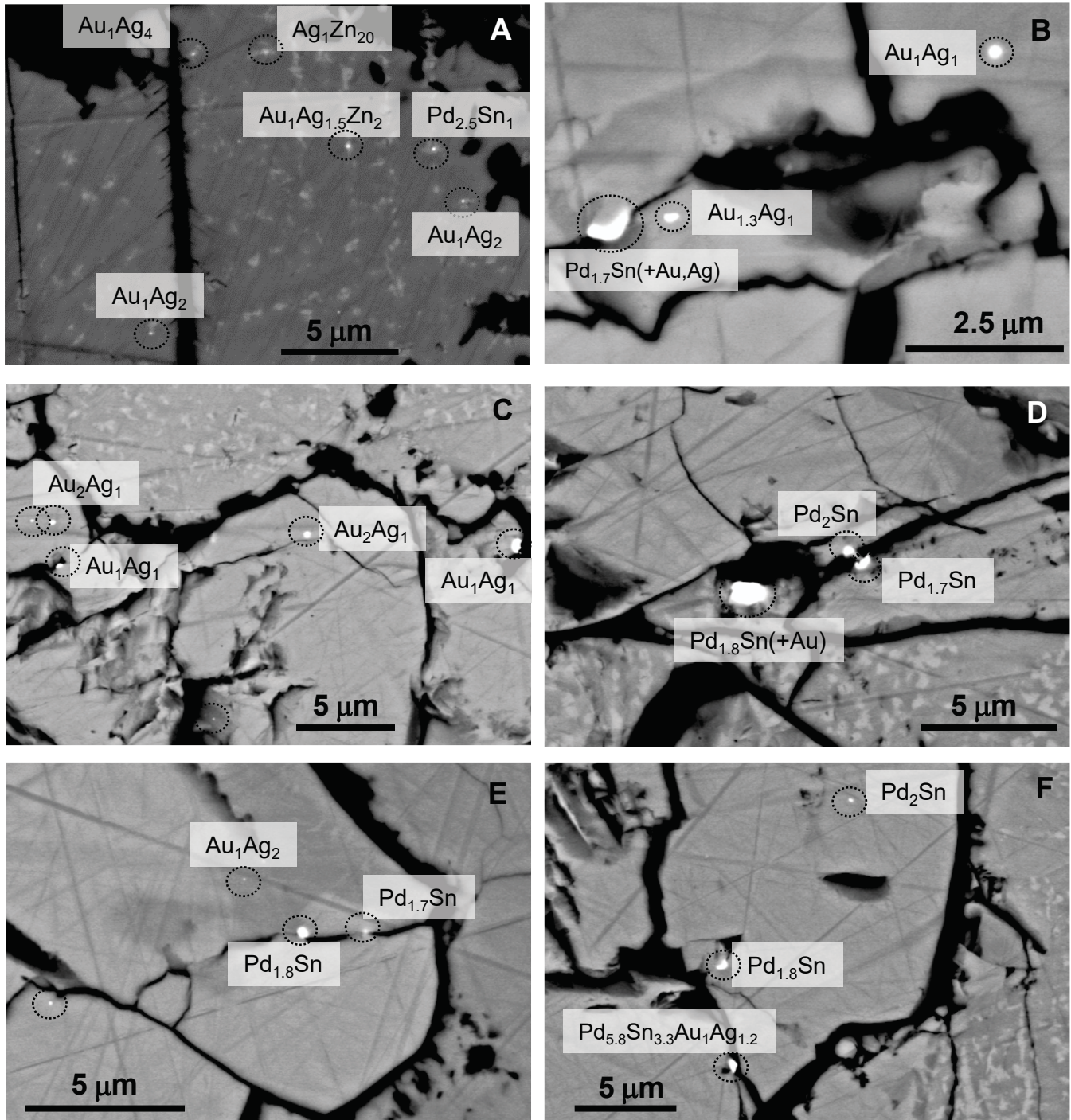


Fig. 12.

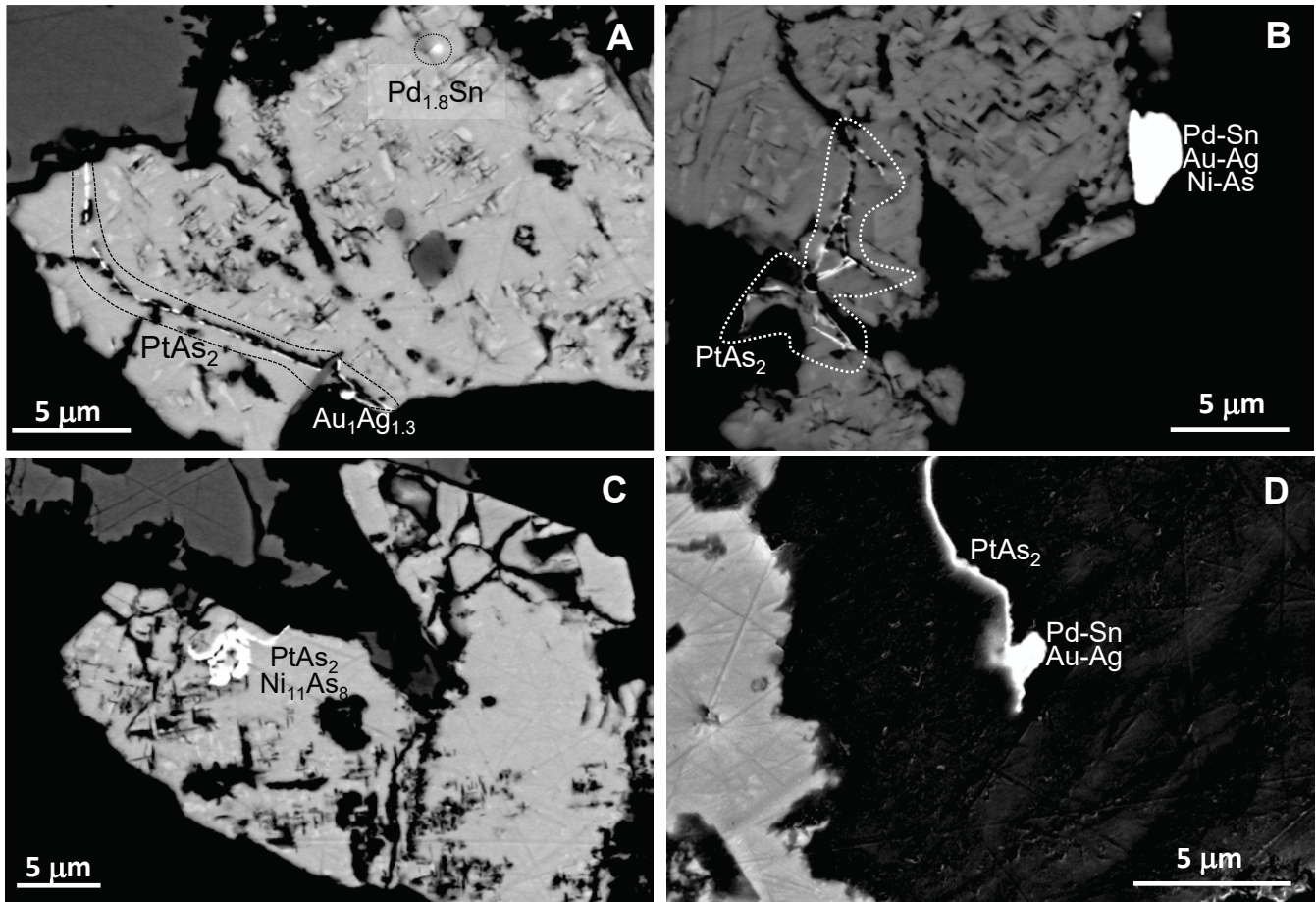
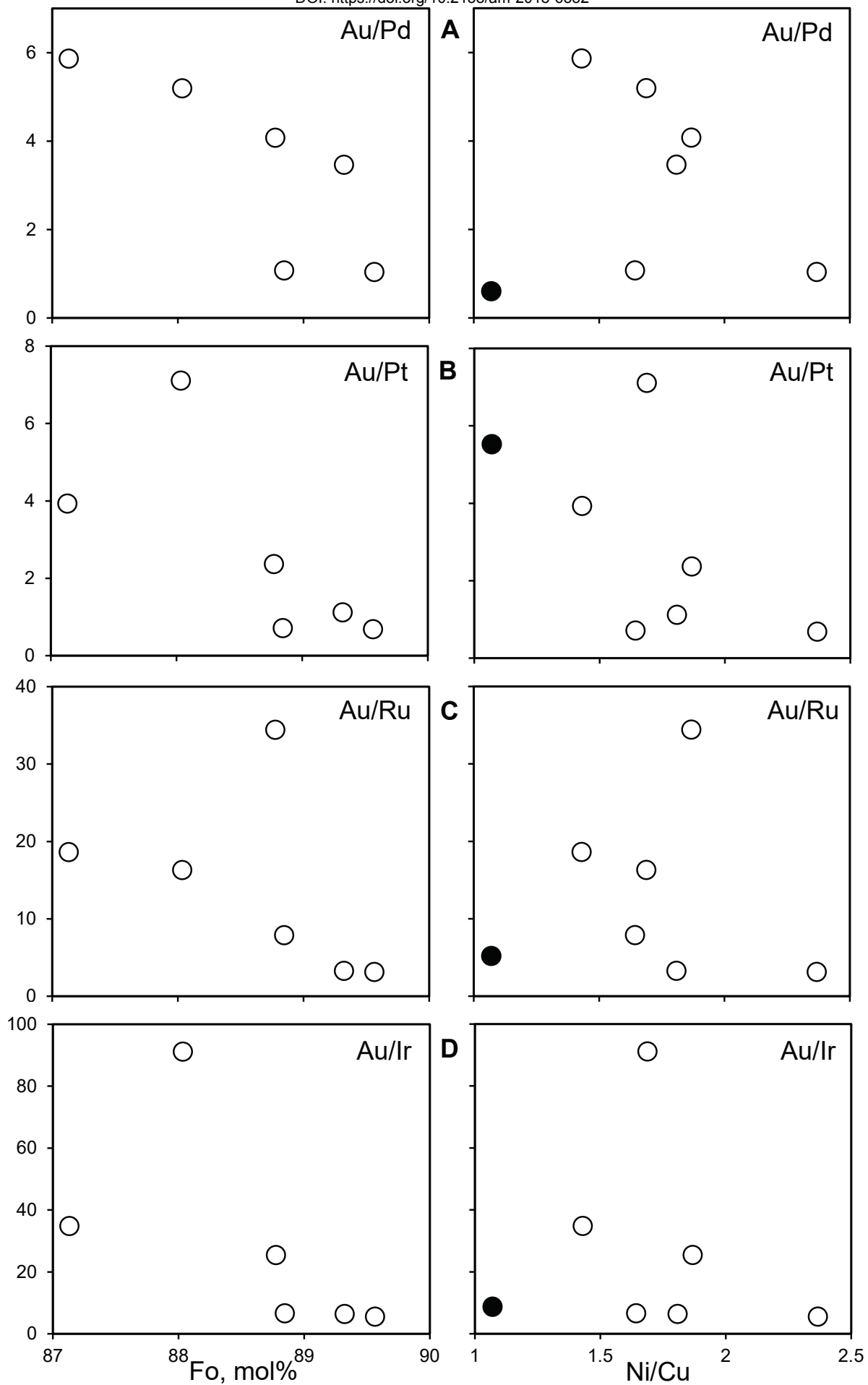


Fig. 13.





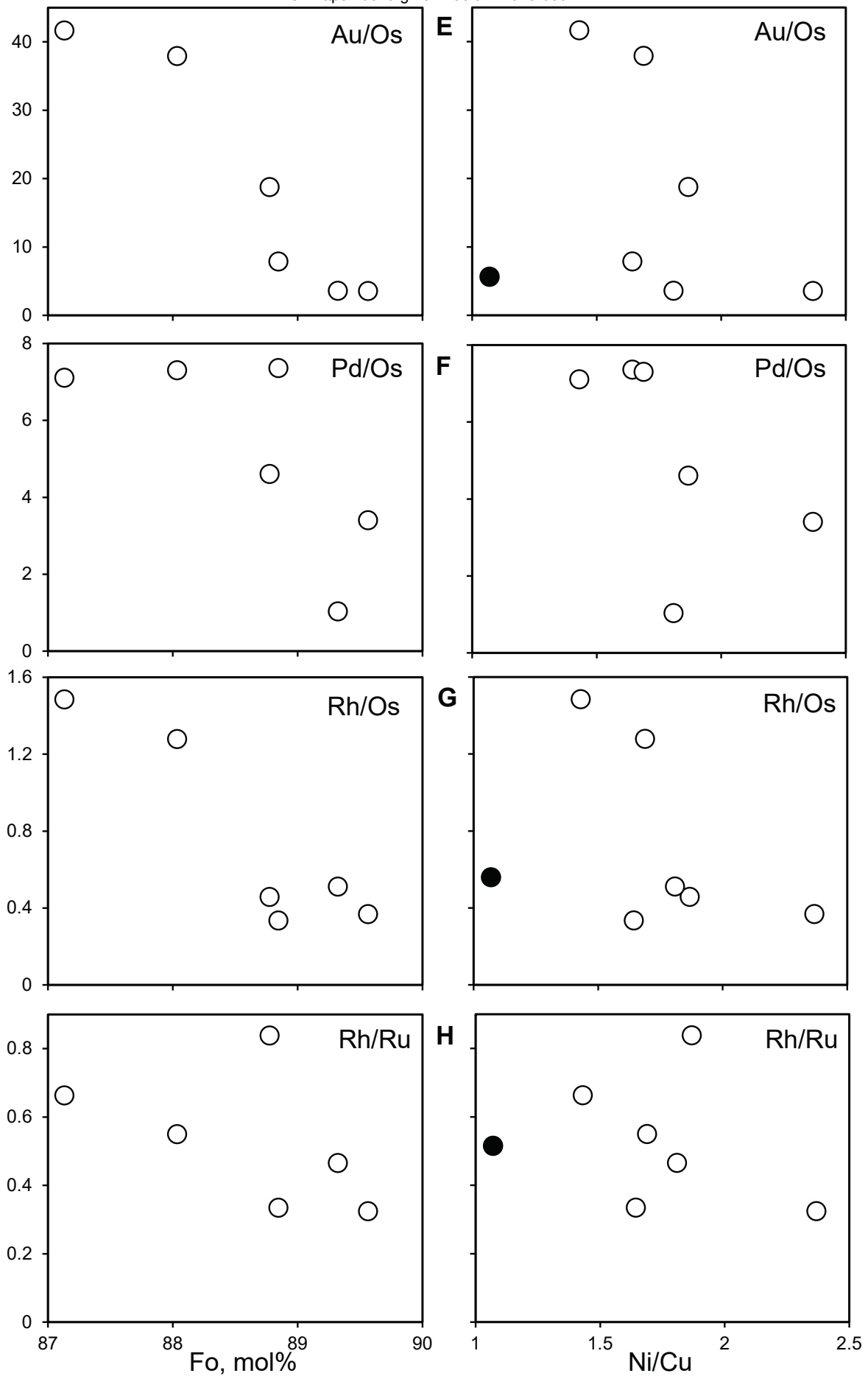


Fig. 14.

Table 1. Compositions of sulfide-bearing olivine-phyric rocks from the Kamchatsky Mys, eastern Kamchatka

Sample	D1454-1	D1456-1		1454-1	1456-1		1454-1	1456-1
wt. %			ppm			ppb		
SiO <sub>2</sub>	43.02	45.21	Sc	21.8	23.4	Ir	0.27	0.22
TiO <sub>2</sub>	0.40	0.45	V	110	121	σ	0.01	0.01
Al <sub>2</sub> O <sub>3</sub>	7.89	8.84	Cr	1372	1185	Ru	0.58	0.43
Fe <sub>2</sub> O <sub>3</sub>	11.05	9.96	Co	102	102	σ	0.01	0.01
MnO	0.16	0.15	Ni	1252	1249	Rh	0.20	0.12
MgO	28.05	25.77	Cu	66.4	70.9	σ	0.01	0.00
CaO	5.47	5.54	Zn	54.8	62.1	Pt	5.91	1.85
Na <sub>2</sub> O	0.85	0.88	Y	7.2	8.0	σ	0.14	0.04
K <sub>2</sub> O	0.04	0.04	Zr	21.4	24.1	Pd	6.13	3.21
P <sub>2</sub> O <sub>5</sub>	0.04	0.04	Nb	1.8	2.0	σ	0.20	0.15
S	0.04	0.06	La	1.4	1.6			
LOI	3.00	3.06	Gd	1.08	1.24			
Total	100.01	100.00	Yb	0.81	0.90			

Table 2. Compositions of homogenized sulfide melt inclusions and average *Mss* and *Iss* phases matrix sulfides analyzed by LA-ICPMS

	1	2	3	4	5	6	7	8
S, wt. %	33.0	36.0	34.4	33.8	34.2	34.0	35.21	34.31
Fe, wt. %	44.2	41.5	40.7	41.5	40.0	41.1	48.52	32.48
Ni, wt. %	13.3	14.6	15.4	15.4	16.5	17.4	15.32	7.75
Cu, wt. %	9.3	7.8	9.4	9.1	9.1	7.4	0.75	25.38
(Fe+Ni+Cu)/S, at.	1.13	0.99	1.06	1.09	1.07	1.08	1.04	1.04
Co, ppm	1911	1854	1906	1853	1901	1827	2081	807
Zn, ppm	159	156	151	144	182	141	61.87	789
Cd, ppm	7.44	9.25	7.94	7.49	10.33	10.03	1.10	9.37
Re, ppm	0.25	0.35	0.24	0.32	0.24	0.30	0.34	0.12
Os, ppm	0.07	0.26	0.63	0.08	0.33	0.74	0.50	0.13
Ir, ppm	0.09	0.19	0.76	0.03	0.19	0.47	0.32	0.10
Ru, ppm	0.16	0.14	0.63	0.19	0.37	0.83	0.60	0.13
Rh, ppm	0.11	0.12	0.21	0.11	0.17	0.27	0.21	0.21
Pt, ppm	0.77	2.04	7.03	0.44	1.06	3.82	0.24	0.62
Pd, ppm	0.52	1.19	4.64	0.60	0.34	2.50	2.25	4.40
Au, ppm	3.03	4.83	4.96	3.12	1.19	2.60	0.46	3.87
Host olivine (Fo, mol. %)	87.1	88.8	88.8	88.0	89.3	89.6		

Analyses of homogenized melt inclusions in olivine, labelled from 1 to 6, correspond to images and time-resolved laser ablation spectra on Figs. 7a–7f. Compositions 7 and 8 represent average *Mss* (n=40) and *Iss* (n=29), respectively, in matrix sulfides.