REVISION 1

Reduced charge transfer in mixed-spin ferropericlase inferred from its high-pressure refractive index

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13 ABSTRACT

14 Physical properties of mantle minerals are essential for comprehensive geodynamic 15 modelling. High-pressure experiments allow measurements of physical properties but fundamental insights into their evolution with pressure are often experimentally inaccessible. 16 17 Here we report the first in situ experimental determination of the optical refractive index, its 18 wavelength-dispersion, and optical absorption coefficient of ferropericlase up to ~140 GPa at 19 room temperature. All these properties change gradually in dominantly high-spin (below ~50 20 GPa) and low-spin (above ~80 GPa) ferropericlase. However, in the mixed-spin state (i.e., 21 significant presence of both high- and low-spin iron), the index dispersion and the absorption coefficient decrease by a factor of three and ~ 30 %, respectively. These anomalies suggest that 22 23 charge transport by small polaron is reduced in mixed-spin ferropericlase, providing fundamental 24 insights into the factor-of-three lower electrical conductivity of ferropericlase at ~50-70 GPa.

25

26 **KEYWORDS**

High-pressure, diamond anvil cell, refractive index, ferropericlase, MgO, DFT, spin
 transition, band gap, electrical conductivity.

29

30 **1. INTRODUCTION**

31 Physical properties of the Earth's mantle and core are at the center of our understanding of 32 planetary evolution. For example, if the electrical conductivity of the lowermost mantle is 33 sufficiently high, the conducting core and the mantle may exchange angular momentum 34 producing detectable intradecadal signals in the length of day (Duan and Huang, 2020; Holme 35 and de Viron, 2013) or imposing preferred paths of geomagnetic reversals along the Americas or 36 Eastern Asia (Buffett, 2015; Runcorn, 1992). Furthermore, the mantle acts as a filter on the 37 magnetic field of the core; thus, knowledge of the electrical conductivity of the mantle is needed 38 to decipher the dynamo history from the geomagnetic record (Alexandrescu et al., 1999; 39 Bloxham and Jackson, 1992; Constable, 2015). Ferropericlase (Fp), being the second most abundant mineral in the lower mantle and the dominant host of iron (Irifune et al., 2010; Piet et 40 41 al., 2016), likely governs the bulk electrical conductivity of that region. Accordingly, the 42 electrical conductivity of Fp has been the subject of many experimental and theoretical 43 investigations.

44 Earlier experimental measurements of electrical conductivity at pressures below ~30 GPa 45 have established that the conductivity of Fp is very sensitive to its overall iron content (Dobson et 46 al., 1997; Hansen and Cutler, 1966; Li and Jeanloz, 1990). Perhaps even more importantly, 47 pressure (P), temperature (T), and oxygen fugacity, all of which vary in the mantle with depth, 48 affect the conductivity of Fp (Dobson and Brodholt, 2000; Wood and Nell, 1991). Two different 49 charge transfer mechanisms have been identified in Fp. At $T \le -1000$ K, the activation energies and the dependence of electrical conductivity on the Fe^{3+}/Fe_{total} ratio indicate that the dominant 50 conduction mechanism is the electron hopping between Fe^{2+} and Fe^{3+} (small polaron) (Dobson et 51 al., 1997; Iyengar and Alcock, 1970). At mantle temperatures (T > -1000 K), however, the extant 52 53 experimental data are consistent with the mechanism that involves Fe-O charge transfer (large 54 polaron) (Dobson et al., 1997). Subsequent measurements to ~100 GPa found that the room-55 temperature conductivity of Fp increases by a factor of ~10 upon compression to 50 GPa, drops 56 by a factor ~3 at 50-70 GPa, and then either increases upon further compression (Lin et al., 2007) 57 or is almost insensitive to pressure (Ohta et al., 2007). The factor-of-three drop in roomtemperature electrical conductivity at 50-70 GPa is concomitant with the iron high-to-low spin 58 59 transition in Fp (Glazyrin et al., 2016; Lin et al., 2007) and has been attributed to the decreased mobility and/or density of charge carriers (small polaron) in low-spin Fp (Lin et al., 2007; Ohta et 60 al., 2007). Optical studies indirectly support this conclusion as the overall absorbance in the 61 visible range, which is a measure of high-frequency electronic conductivity, decreases with 62 pressure in low-spin Fp (Goncharov et al., 2006; Keppler et al., 2007; Schifferle and Lobanov, 63 64 2022). This decrease in absorbance, however, is somewhat questionable because of the unknown 65 sample thickness at high pressure, which is also expected to decrease with pressure in a strongly non-isotropic fashion (Lobanov and Geballe, 2022). Reliable in situ measurements of sample 66 67 thickness are thus needed to quantify the absorption coefficient of Fp at high pressure and to 68 resolve whether the electrical conductivity of Fp is linked to its optical properties.

Page 2 of 19

69 The thickness of (semi)transparent samples in diamond anvil cell (DAC) experiments can 70 be accurately measured if the refractive index of the sample is known. The refractive index and 71 its wavelength-dispersion also provide information on the electronic structure of materials. 72 Pressure-induced changes in the electronic conductivity (e.g. due to small polarons) may thus be accompanied by changes in the refractive index. To the best of our knowledge, the refractive 73 74 index of ferropericlase at elevated pressures has never been characterized. Recently, we 75 developed a method to measure the refractive index and its wavelength-dependence of fully 76 transparent solids compressed in a DAC to pressures greater than 100 GPa (Lobanov et al., 2022; 77 Schifferle et al., 2022). In the present work, we build up on that method to make it applicable for 78 semitransparent samples and report the refractive index of high spin (HS), mixed spin (MS), and 79 low spin (LS) ferropericlase. Our results show that the refractive index of Fp in the visible 80 spectral range is largely independent of pressure or iron spin state. The wavelength-dispersion of 81 the refractive index is also nearly invariant in the HS and LS Fp but is reduced abruptly by a 82 factor of three in the MS state. The absorption coefficient also decreases by ~35% at the onset of 83 the spin transition. Our observations allow an alternative interpretation of the previous high-84 pressure electrical conductivity data. We suggest that the mobility and/or density of small 85 polarons is reduced only in the MS regime, unlike previous studies that proposed reduced small 86 polaron transport in LS Fp.

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88 88 89 89 Samples and diamond anvil cell loading

90 All high-pressure experiments were performed using symmetrical DACs equipped with diamond pairs featuring beveled 300/100 µm or flat 300 µm culets. The beveled diamonds 91 92 feature 300 µm culets at 8° to a single 100 µm central flat. Rhenium gaskets indented to a thickness of 15-20 µm were laser-drilled to create cylindrical holes with diameters of ~45 µm 93 94 (when $300/100 \ \mu m$ culets were used) or ~140 μm (when 300 μm culets were used), which served 95 as sample chambers. For each loading we selected fragments of the samples with appropriate 96 dimensions to fill the DAC sample chamber, placed the samples without any pressuretransmitting medium, immediately sealed the DAC assembly and increased the pressure to ~20-97 98 30 GPa producing optically homogenous samples, which is required for refractive index measurements described below. We used synthetic single-crystalline samples of ferropericlase of 99 100 two different compositions: (Mg_{0.87}Fe_{0.13})O and (Mg_{0.76}Fe_{0.24})O, further referred to as Fp13 and Fp24 to indicate their iron content. Fp13 was produced by Fe diffusion into an MgO crystal in a 101 102 gas-mixing furnace at ambient pressure (Lobanov and Speziale, 2019). Fp24 was synthesized by Caterina Melai (Aprilis et al., 2020) in a 10/5 multi-anvil assembly at 15 GPa and 1800 °C using 103 starting material from (Longo et al., 2011). Electron energy loss spectroscopy yielded an 104 $Fe^{3+}/\Sigma Fe_{total}$ ratio of ~0.1 for Fp13 (Lobanov and Speziale, 2019), however, its Fe^{3+} content 105 might be considerably lower as discussed in (Schifferle and Lobanov, 2022). The Fe³⁺/∑Fe_{total} 106

107ratio of Fp24 is ~0.02, based on Mössbauer spectroscopy measurements (Aprilis et al., 2020).108Because the fragments of Fp13 and Fp24 were effectively crushed in the DAC sample chamber109upon sealing and pressure increase, they are likely present in the form of coarse powder at high110pressure (although this has not been verified by x-ray diffraction). The pressure at the center of111the sample chamber was gauged following the diamond Raman edge method with a relative112uncertainty of ~5% (Akahama and Kawamura, 2006).

113 **Refractive index measurements**

Because of significant light attenuation in Fp13 and Fp24 over thicknesses characteristic of samples in DAC, we needed to extend the reflectivity method used previously to study transparent samples (Lobanov et al., 2022; Schifferle et al., 2022; van Straaten and Silvera, 1988; Zha et al., 2007) to semi-transparent (absorbing) samples. The approach is based on the Fresnel equation for normal incidence, where the measured reflectivity of the diamond-sample interface $(R_{dia-smp})$ is related to the refractive index *n* of the sample (n_{smp}) and diamond (n_{dia}) , and the imaginary part of the refractive index of the sample (κ_{smp}) :

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$$R_{dia-smp} = \frac{(n_{smp} - n_{dia})^2 + \kappa_{smp}^2}{(n_{smp} + n_{dia})^2 + \kappa_{smp}^2}$$
(Eq. 1)

122 Previous reports on the absorption coefficient of ferropericlase indicate that for Fp13 and 123 Fp24 κ_{smp} is small (~0.01 for the studied pressure range, see Supplementary Data for further 124 details) and can thus be considered zero. Nonetheless, quantifying light attenuation in 125 ferropericlase due to absorption is necessary for the determination of $R_{dia-smp}$:

126
$$\frac{I_1 + I_2}{I_0} = R_{dia-smp} + T^2 (R_{dia-smp}^3 - 2R_{dia-smp}^2 + R_{dia-smp})$$
(Eq. 2),

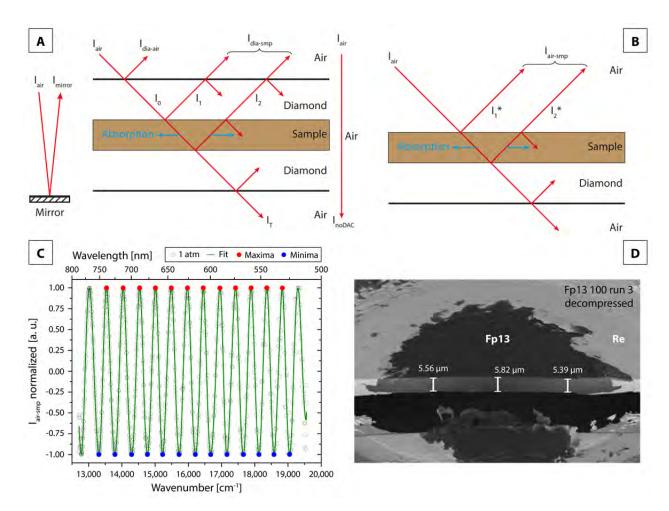
where I_0 , I_1 , and I_2 are individual reflections defined graphically in **Figure 1 A**. *T* is sample transmission (T = 0 for opaque and T = 1 for transparent samples). We independently express *T* through I_T/I_{noDAC} :

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$$\frac{I_T}{I_{noDAC}} = T(R_{dia-smp}^2 - 2R_{dia-smp} + 1)(1 - R_{dia-air})^2$$
(Eq. 3)

131 where I_T and I_{noDAC} are the signals measured through the sample and without the DAC and $R_{dia-air}$ 132 is the reflectivity of the diamond-air interface. The derivations of Eq. (2) and (3) are provided in 133 the Supplementary Data. All measured signals were averaged over the 550-650 nm spectral 134 range, where the reflectivity of the reference mirror is well-characterized (~0.985-0.995), 135 allowing for precise measurements of the probe intensity impinging on the upstream diamond-air interface. Eq. (2) and (3) contain only two unknowns ($R_{dia-smp}$ and T) and are solved 136 simultaneously. Two of the three sets of roots included imaginary $R_{dia-smp}$ and T and were 137 138 discarded. Using $R_{dia-smp}$ from the real set of roots, we solve Eq. (1) for n_{smp} , the refractive index of ferropericlase at 600 nm (because of the averaging of $\frac{I_1+I_2}{I_2}$ and T over 550-650 nm). In solving 139

Page 4 of 19

140 Eq. (1), we assume a pressure-independent refractive index of diamond $(n_{dia} = 2.418 \text{ (Hynes,})$ 141 2016)). The validity of this assumption and possible effects of a pressure dependence of n_{dia} have 142 been discussed in detail by Schifferle et al. (2022). Their results on MgO support the hypothesis 143 n_{dia} = constant. However, we provide all necessary data for the evaluation of n_{smp} in 144 **Supplementary Table 1** which can be reanalyzed when more detailed information on the pressure-dependence of n_{dia} is available. The solutions to Eq. (1-3) have been found with Python 145 146 SciPy and SymPy libraries (Meurer et al., 2017; Walt et al., 2011). We independently tested the 147 reflectivity method for semitransparent samples developed in this work by reproducing the high-148 pressure evolution of the refractive index of an Fe-bearing basaltic glass (i.e. absorbing in the 149 visible) reported by Kuryaeva and Kirkinskii (1997). We empirically estimate the overall 150 uncertainty in the measured n_{smp} at ~1%. The effect of uniaxial stress (due to the lack of pressure-151 transmitting medium) on the measured index is likely smaller than the overall experimental 152 uncertainty because the pressure dependence of the refractive index of MgO (and ferropericlase) 153 is small. That is, moderate pressure gradients across the sample chamber (e.g., of 5 %) would 154 cause index gradient with the total difference at the sample center and periphery of $\ll 1\%$. This 155 is further supported by first-principles simulations: e.g., the pressure-dependence of the refractive 156 index of MgO measured in the DAC with the same method as used in the present work is in 157 excellent agreement with DFT computations (Schifferle et al., 2022).



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159 Figure 1: A: Reflectivity measurements in a partially absorbing sample at high-pressure conditions. The 160 perpendicularly-incident probe laser is partially reflected at the air-mirror, diamond-air and diamond-sample 161 interfaces. Oblique probe incidence is depicted for clarity. B: Reflectivity measurements for the decompressed 162 sample in an opened diamond anvil cell. C: A normalized intensity spectrum of Fp13 (run 3) decompressed to 1 atm. 163 The local extrema (blue and red circles) served to find the optical path $(n_{1atm}d_{1atm})$, where d_{1atm} is sample 164 thickness. D: A secondary electron microscope image of a focused ion beam cross-section of the of Fp13 (run 3) 165 decompressed to 1 atm used to directly image the sample thickness. The light grey part is the Re gasket and dark 166 grey is Fp13.

167 We also measured the refractive index of Fp13 decompressed to 1 atm after one of the 168 DAC runs. To this end, we removed the upper diamond anvil to ensure no pressure was applied to the sample and measured the reflectivity of the air-sample interface where $I_{air-smp}$ is 169 composed of I_1^* and I_2^* , which are individual reflections from the air-sample and sample-170 171 diamond interfaces (Figure 1 B). The normalized reflectance spectrum was analyzed for the 172 average spectral separation of the interference fringes to obtain the optical path (OP =173 $n_{1atm}d_{1atm}$ (Figure 1 C). A radial cross-section through the retrieved sample (Figure 1, D) cut by a focused ion beam allowed to directly measure the thickness of decompressed ferropericlase 174

175 at the center of the DAC sample cavity (d_{1atm}) and to obtain the refractive index of Fp13 as 176 $n_{1 atm} = \frac{OP}{d_{1atm}}$.

177 **Dispersion analysis**

The interference fringe spectra can be analyzed for the wavelength-dispersion of the
refractive index following (Schifferle et al., 2022) (for more details, see their section 2.3).
Briefly, for each extremum the interference condition is:

181
$$\lambda k(\lambda) = 2dn_{smp}(\lambda)$$
 (Eq. 4),

182 where $k(\lambda)$, is the interference order number of the minimum/maximum (half-integer step size) observed at the wavelength λ , d is sample thickness at high pressure, and $n_{smp}(\lambda)$ is the 183 refractive index at the extremum. Using Eq. 4 for each pair of extrema we estimated sample 184 thickness assuming n_{smp} is wavelength independent, which is accurate within ~1 % for Fe-poor 185 Fp because the index dispersion of MgO across the studied spectral range is small (~1 %) 186 187 (Stephens and Malitson, 1952). We will provide evidence below that the index dispersion of 188 Fp13 is indeed comparable to MgO. The sample thickness averaged over all used extrema pairs is 189 then used to assign the interference order number to the extremum closest to 600 nm (e.g., k_{600nm} 190 = 70 at 16.9 GPa, k_{600nm} = 41.5 at 109.1 GPa) and subsequently to all the other observed extrema. 191 Finally, the refractive index at each λ is calculated using Eq. 4 for all the observed extrema. The 192 propagation of random and systematic relative errors in the measured index dispersion yields 193 overall errors of ~0.9% at ~23 GPa and ~2.6% at ~109 GPa. The increase in this error is likely 194 due to diamond cupping at high pressure (Schifferle et al., 2022).

195 Computation of refractive index and index dispersion for MgO and ferropericlase

196 Atomistic model calculations were carried out within the framework of DFT (Hohenberg and 197 Kohn, 1964) and the pseudopotential method using the CASTEP simulation package (Clark et al., 198 2005). Ultrasoft pseudopotentials were generated "on the fly" using the parameters provided with 199 the CASTEP distribution. These pseudopotentials have been extensively tested for accuracy and 200 transferability (Perdew et al., 1996). The pseudopotentials were employed in conjunction with 201 plane waves up to a kinetic energy cutoff of 630 eV (Lejaeghere et al., 2016). Spin-polarised 202 calculations were carried out with the PBE exchange-correlation functional and a Hubbard U of 203 2.5 eV for the Fe-d-states. As spin-polarized DFT-GGA-PBE + U calculations of low symmetry 204 structures (such as MgO with Fe defects) are computationally expensive, the calculations here 205 were limited to 64-atom supercells. In such a cell, a single Fe-atom is about 8.5 Å away from the next Fe atom, and hence interactions between Fe atoms are small. Increasing the concentration of 206 207 Fe atoms, e.g. by placing a second Fe Atom in this simulation cell, would then have to consider 208 the numerous relative arrangements, possible clustering and Fe-Fe interactions. Using a larger 209 simulation cell or studying Fe-Fe interactions in the 8.5 Å cell were beyond the scope of the 210 present study. Monkhorst-Pack grids (Monkhorst and Pack, 1976) were used for Brillouin zone

integrations with a distance of < 0.029 Å⁻¹ between grid points. Convergence criteria included an energy change of $< 5 \times 10^{-6}$ eV/atom, a maximal force of < 0.01 eV/Å, and a maximal deviation of the stress tensor < 0.02 GPa from the imposed stress tensor. All calculations were carried out in the athermal limit, *i.e.*, the influences of temperature and zero-point motion were not taken into account.

216

217 **3. RESULTS AND DISCUSSION**

218 The refractive indices of Fp13 and Fp24 measured in this work are shown in Figure 2. At 219 all pressures, the index of Fp24 at 600 nm is ~1-2 % higher than that of Fp13, consistent with the 220 results of Henning et al. (1995) who measured the optical properties of $Fe_xMg_{1-x}O(x = 0.4-1)$ at 1 221 atm. The indices of both Fp13 and Fp24 increase with pressure with no sharp discontinuity over 222 the expected spin transition range of 40-80 GPa (Glazyrin et al., 2016; Kantor et al., 2009; Lin et 223 al., 2006). In one of the runs (Fp13, run3), the index measured at 600 nm appears ~1% too low 224 than what may be expected from its value independently measured at 1 atm as well as expected from the results of Henning et al. (1995). It is plausible that submicron impurities are present in 225 that loading which contributed an up-shift to the measured $\frac{I_1+I_2}{I_0}$ through light back scattering, 226 increasing the apparent $R_{dia-smp}$ and thus decreasing n_{smp} . This is indirectly supported by our data 227 on Fp24, which remained fully opaque upon compression, in which case the evaluation of $R_{dia-smp}$ 228 is more accurate because only the upstream diamond-sample reflection (I_l) contributes to $\frac{I_1+I_2}{I_2}$ at 229 T = 0 (see Eq. 2). At P > 60 GPa, the increase in refractive indices of Fp13 (+0.0014 (±1.31×10⁻³)) 230 per 10 GPa) and Fp24 (+0.0021 (±7.75×10⁻⁴) per 10 GPa) is small, but resolvable within the 231 232 uncertainty of our determination; the refractive index of mantle-like Fp is not independent of 233 pressure (dashed lines in Figure 2). The extrapolation of the high-pressure index of Fp24 at 600 234 nm to 1 atm is in excellent agreement with the expectation from the literature (Henning et al., 235 1995).

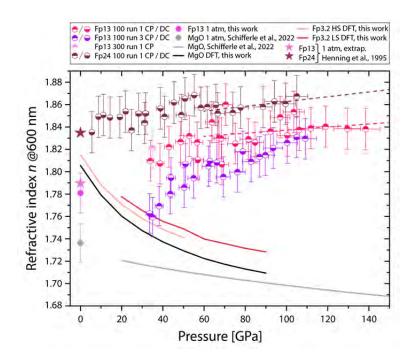




Figure 2 The refractive index of Fp13 and Fp24 at 600 nm measured in this work, compared to DFT calculations of Fp3.2 (high- and low-spin) and MgO, as well as experimental data on MgO from Schifferle et al. (2022). Stars represent extrapolations based on the indices of iron-rich Fp reported by Henning et al. (1995) (at 1 atm). Red (Fp13) and magenta (Fp24) dashed lines are visual guides based on a linear dn/dP fit for to our data at P > 60 GPa. Abbreviations: CP = compression, DC = decompression.

The refractive index of ferropericlase is considerably higher than that of MgO (Figure 2) 242 because of the higher polarizability of Fe^{2+} compared to Mg^{2+} at 1 atm (^[6]Fe²⁺ 2.040 Å³, ^[6]Mg²⁺ 243 0.651 Å³ (Shannon and Fischer, 2016)). This results in a compositional dependence of the 244 245 refractive index of ferropericlase (higher Fe content leads to higher index, Supplementary Figure 246 **S1**). Interestingly, while the pressure-derivatives of the indices of Fp13 and Fp24 are positive, 247 that of MgO is negative (Balzaretti and Da Jornada, 1990; Fratanduono et al., 2013; Oganov et 248 al., 2003; Schifferle et al., 2022). The increase in refractive index with pressure entails that the Lorenz-Lorentz factor $\left(\frac{n^2-1}{n^2+2} \propto \frac{\rho}{\alpha_{LL}}\right)$ also increases with pressure (ρ is density, α_{LL} is Lorenz-249 250 Lorentz polarizability). Accordingly, we attribute the positive pressure-dependence of the 251 refractive index of ferropericlase to its anomalous polarizability (compared to MgO; Supplementary Figure S2), because the compressibility of MgO, Fp13, and Fp24 are similar. 252 However, our DFT computations indicate that at low Fe²⁺ concentrations (Fp3.2) the pressure-253 254 derivative of the refractive index is similar to that of MgO (see black and grey lines in Figure 2). 255 We note that Fe-Fe interactions, which are very weak or absent in our DFT computations as the 256 Fe-Fe-distances are large (1 Fe, 31 Mg atoms in the supercell), are likely present in Fp13 and 257 Fp24 because their Fe content exceeds the percolation limit (12% (Lorenz and Ziff, 1998)). Above the percolation limit, an interconnected network of overlapping t_{2g} orbitals between 258 259 adjacent edge-sharing FeO₆ octahedra is present. We propose that pressure-induced electron 260 delocalization along the interatomic t_{2g} - t_{2g} joints (Diamond et al., 2022) maintains higher

Page 9 of 19

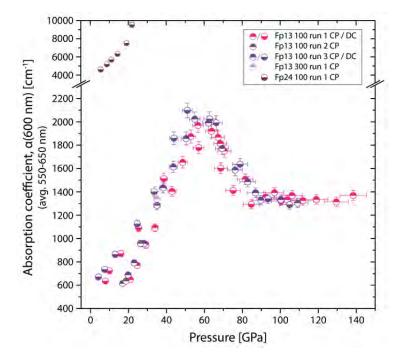
polarizability in Fp and is thus the reason for the anomalous positive pressure-dependence of the
 refractive indices of HS Fp13 and Fp24. Such electron delocalization is complete in the LS state;
 hence, the refractive index of LS Fp is almost pressure independent (*i.e.*, MgO-like).

264 We used the high-pressure refractive index and the interferometric data to obtain d, the 265 thickness of ferropericlase at high pressure. To avoid fluctuations in d due to the apparently 266 random error in the index, n is derived from a linear P vs. n fit (see Supplementary Data). Please 267 note, that possible pressure differences across the sample were not considered because pressure 268 differences at the sample center and periphery were not resolvable within the uncertainty of the 269 pressure determination method (5% relative). Such moderate pressure gradients across the sample 270 chamber would cause index gradient with the total difference at the sample center and periphery 271 of << 1%. This implies that the differences in optical path measured at the center and periphery 272 of the sample chamber (which may be as high as 2-3 %) are primarily due to differences in 273 sample thickness. Next, we quantify the absorption coefficient of ferropericlase as: $\alpha(600 nm) = A \cdot \ln(10)/d$, where $A = -log_{10}T$ is the measured optical absorbance of the 274 275 sample (T is transmission). Please note that even at low pressure (< 30 GPa) transmission data 276 and the interference fringe pattern quality were sufficient to obtain accurate thicknesses (the 277 variation in optical path inferred for different fringes is less than 1%) and absorption coefficients 278 (3% error). Due to the high absorption coefficient of Fp24, transmission could only be detected in 279 very thin samples (< 6 µm), achieved only on decompression as samples in DACs continue to 280 thin upon pressure release (Lobanov and Geballe, 2022). Here we extend this seemingly general 281 behavior to ferropericlase (Supplementary Figure S3), and show that thickness estimations based 282 on the assumption of isotropic compression severely overestimate the real sample thickness by up 283 to ~50%, similar to the conclusion of (Lobanov and Geballe, 2022). A similar gasket thinning on 284 decompression has also been noted previously by Dewaele et al. (2003) with He as a sample, 285 albeit the gasket thickness and the maximum pressure were very different (~48 μ m at ~20 GPa). 286 These suggest that sample size and hardness likely play a role in the gasket thinning upon 287 decompression.

288 For Fp13 we find $\alpha(600 \text{ nm})$ to increase from 20 GPa up to ~57 GPa by a factor of ~3, where the maximum is reached at ~2000 cm⁻¹ (Figure 3). For the pressure range > 57 to ~85 GPa 289 we find a sharp decrease in $\alpha(600 \text{ nm})$ by ~30%. At P > 85 GPa, the absorption coefficient 290 reaches a plateau at ~1300 cm⁻¹. The pressure range of the negative $d\alpha/dP$ suggests that the 291 decrease of α is related to the spin transition in Fp13 (Schifferle and Lobanov, 2022). In HS Fp, 292 the transition energy of the single multiplicity allowed crystal field (d-d) band $({}^{5}T_{2g} \rightarrow {}^{5}E_{g})$ is 293 almost pressure independent (Keppler et al., 2007; Schifferle and Lobanov, 2022); thus, this band 294 does not contribute to the increase in the absorption coefficient up to 60 GPa. One of the d-d 295 296 bands of LS Fp appears close to 600 nm (16,667 cm⁻¹) at ~60 GPa (Schifferle and Lobanov, 297 2022) and would only increase the absorption coefficient, in contrast to what is observed. At the 298 same time, the absorption edge shows a qualitative red-shift (in HS) and blue-shift (in LS) with 299 pressure (Goncharov et al., 2006; Keppler et al., 2007; Schifferle and Lobanov, 2022). We

Page 10 of 19

300 conclude, therefore, that the trend in $\alpha(600 \text{ nm})$ is mostly related to the intensity of the UV-301 absorption edge in ferropericlase with a minor contribution of the crystal field bands.



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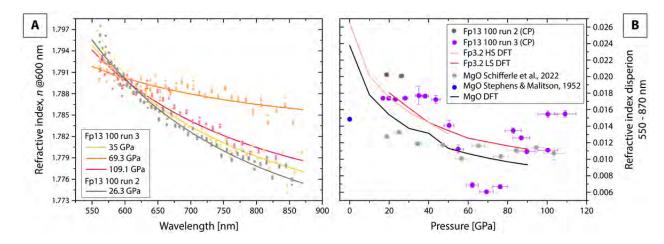
Figure 3 Absorption coefficients of Fp13 and Fp24 at 600 nm based on direct measurements of the optical path $(n_{smp}d)$. Thicknesses used for the calculation of the absorption coefficient are obtained by dividing the optical path by n_{smp} at 600 nm and neglecting the index dispersion in the analyzed spectral range (~500-800 nm) which is ~1% or less (see discussion in the main text). The uncertainty in *d* is that of refractive index and is ~1%. Please note the break in y-axis. Abbreviations: CP = compression, DC = decompression.

308 The variation of $\alpha(600 \text{ nm})$ with pressure is reversible upon decompression down to ~40 309 GPa. Our $\alpha(600 \text{ nm})$ allows to derive the wavelength-dependent absorption coefficients (and thus 310 the imaginary part of the refractive index) from previously published absorbance data for the same Fp13 sample (Schifferle and Lobanov, 2022) (Supplementary Figure S4). Compared to 311 312 previous reports on the absorption coefficient of ferropericlase where the sample thickness at 313 high pressure was estimated using its equation of state (Lobanov et al., 2021), we show that α is 314 lower by up to ~50%. This is because Lobanov et al. (2021) measured the thickness of their Fp13 315 sample after decompression and assumed the sample was thinner at high pressure. As show by 316 Lobanov and Geballe (2022) samples in DACs continue to thin upon decompression; and their 317 data can be used to revise the Fp13 absorption coefficients reported in Lobanov et al. (2021). An 318 approximate correction suggests that the radiative thermal conductivity of ferropericlase at the 319 base of the mantle has been underestimated by roughly a factor of two in that previous study.

Concomitantly to the directly measured decrease in the absorption coefficient by ~ 30 % at 57-85 GPa, we observe a sharp decrease in the absolute wavelength dispersion of the refractive index of Fp13 (**Figure 4**). Below ~ 40 GPa, the wavelength-dispersion of the refractive index of Fp13 is essentially pressure independent. At P = 43.6-69.3 GPa, however, the index dispersion is

Page 11 of 19

324 dramatically decreased by a factor of ~ 3 from ~ 0.0174 ($\pm \sim 1$ %) to 0.006 ($\pm \sim 2$ %) (orange line, 325 Figure 4, A). Yet, at P > 69.3 GPa, the absolute index dispersion is almost restored at ~0.016 (±1 %) (Figure 4, B). The index dispersion is ~50 % stronger in HS and LS Fp13 than in MgO, 326 327 which is qualitatively consistent with our DFT computations that produced ~20 % higher 328 dispersion in the case of Fp3.2 than in the case of pure MgO. Likewise, our computations support 329 the notion that the index dispersion in HS and LS Fp3.2 are similar. In contrast to pure HS and 330 LS ferropericlase probed in experiments and computations, the measured index dispersion of MS 331 Fp13 is anomalously low: ~50 % lower than that of MgO at 60-80 GPa. We note, that the 332 dispersion data is analyzed assuming a constant refractive index at 600 nm of 1.79 (based on the 333 expectation for Fp13 at 1 atm (Henning et al., 1995) to better account for deviations due to 334 random errors in the refractive index. This, however, does not significantly affect the inferred 335 index dispersion, because possible differences in refractive index at 600 nm of 2-3 % result 336 mostly in a vertical shift of the index dispersion curve and not its slope (Figure 4 A).



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Figure 4 (A) Dispersion of the refractive index of Fp13 (compression) at selected pressures (assuming a pressureindependent refractive index *n*(600 nm) of 1.79) and (B) as the difference between refractive indices at 550 and 870 nm compared to DFT calculations for Fp3.2 and MgO as well as previously published experimental data (Schifferle et al., 2022; Stephens and Malitson, 1952).

We analyzed the dispersion by the single-effective-oscillator model of Wemple and DiDomenico (1971). In this model, the wavelength dependence of the refractive index is related to two fitting parameters E_0 and E_d , which describe the single oscillator energy and dispersion energy (in eV), respectively. These quantities are related to the band gap width and bonding environment:

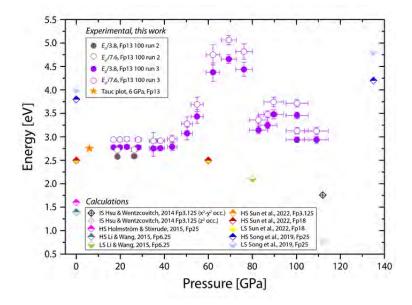
347
$$n(\lambda) = \sqrt{\frac{E_d \cdot E_0}{E_0^2 - \left(\frac{h \cdot c}{\lambda}\right)^2} + 1}$$
(Eq. 4)

Here, *h* is the Planck's constant in eV·s and *c* the speed of light in m/s and λ the wavelength in m. By fitting Eq. (4) to the measured index dispersion at 550-870 nm, we find E_d and E_0 . An indetail description of the procedure can be found in Schifferle et al. (2022).

Page 12 of 19

351 Wemple and DiDomenico (1971) proposed that E_0 is related to the average optical band 352 gap, and that the ratio $E_0/1.5$ approximates independently measured band gaps. The denominator 353 in this ratio, however, is material dependent. To estimate it for Fp13 we performed a Tauc plot 354 analysis of the low-pressure optical absorption data of Fp13 reported in Schifferle and Lobanov 355 (2022), which yields the minimum band gap energy of 2.75 eV. We thus divide all fitted E_0 by 356 3.8 to match 2.75 eV at low pressure (*i.e.* at $P < \sim 40$ GPa, where the wavelength-dispersion is 357 essentially pressure-independent). Although the values of $E_0/3.8$ at P < 40 GPa are broadly 358 consistent with previous DFT calculations (Holmstrom and Stixrude, 2015; Hsu and 359 Wentzcovitch, 2014; Li and Wang, 2015; Song et al., 2019; Sun et al., 2022), our results likely 360 underestimate the band gap as broad-range UV spectra are not available. Despite these 361 uncertainties, we can derive semi-quantitative trends in band gap energy.

362 In Figure 5 we show $E_0/3.8$ and $E_d/7.6$ (for illustration only, *i.e.*, to scale E_d close to $E_0/3.8$) of Fp13 from the Wemple and DiDomenico (1971) single-effective oscillator analysis and 363 compare it with existing computations of the optical band gap in ferropericlase. Regardless of the 364 chosen scaling factor for E_0 , our experimental data (Figure 5) suggests a constant band gap 365 366 energy up to ~40 GPa. In the MS state ($P \approx 50-80$ GPa), we observe an increase in $E_0/3.8$ with a 367 maximum at ~ 69.3 GPa of ~ 4.7 eV. This peak in the band gap probably represents 50% HS and 368 50% LS, consistent with the LS fraction inferred by Mössbauer spectroscopy (Kantor et al., 369 2009). At P > 100 GPa, the band gap energy is comparable to that at P < 40 GPa. These semi-370 quantitative insights into the band gap energy are consistent with DFT computations (Song et al., 371 2019; Sun et al., 2022), which suggest a pressure-independent band gap energy for high- and 372 low-spin states. The possible intermediate spin (IS) state, although energetically disfavored (Hsu 373 and Wentzcovitch, 2014), also does not deviate from the general trend of a near-constant band 374 gap in DFT calculations. However, to the best of our knowledge, there are no DFT computations 375 of the MS state available, which we probed optically here for the first time. E_d shows a trend qualitatively similar to that of E_0 . Because E_d represents the distribution of charge around the 376 anion (Wemple and DiDomenico, 1971), its increase and decrease over the spin transition might 377 378 reflect the decrease/increase in *p-d* orbital overlap. For example, charge donation from oxygen to 379 the empty e_g orbitals in the LS state has been inferred from the decrease in the Fe-O bond covalency (Schifferle and Lobanov, 2022). A similar ligand-to-Fe donation has been proposed as 380 a consequence of the Fe²⁺ HS to LS transition in S- and Sn-thiospinels (Womes and Jumas, 381 382 2013).



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Figure 5 Results of the Wemple and DiDomenico (1971) fit ($E_0/3.8$ and $E_d/7.6$) to the measured index dispersion of Fp13 (circles). Orange star represents the band gap estimation based on Tauc plot analysis (Tauc, 1968) of data from Schifferle and Lobanov (2022). In addition, we show calculations from literature for the band gap in different ferropericlase compositions and spin configurations (diamonds). Abbreviations: high-spin (HS), intermediate-spin (IS) and low-spin (LS). Please note, datapoints from Sun et al. (2022) for the three spin configurations are directly overlapping at 0 GPa and 2.5 eV (HS Fp 3.125, HS Fp18 and LS Fp18) as well as at 60 GPa and 2.5 eV (HS Fp18, LS Fp18).

391 The increase in $E_0/3.8$ by ~1.95 eV and the reduction of the absorption coefficient by ~35% over 392 the spin transition range offers a qualitative understanding of the previously reported drop of 393 electrical conductivity by a factor of up to three in ferropericlase at ~50-70 GPa (Lin et al., 2007; 394 Ohta et al., 2007). In contrast to these previous studies, where a lower small polaron mobility 395 and/or density in LS ferropericlase was proposed as an explanation (Lin et al., 2007; Ohta et al., 396 2007), our data suggests that lower electron mobility is characteristic of the MS state while that in 397 HS and LS is similar. We propose that qualitatively this can be understood in terms of small 398 polaron transport through the crystal lattice, which requires electron exchange between adjacent Fe sites: $Fe^{2+} \xrightarrow{e^-} Fe^{2+} = Fe^{3+} + Fe^+$. In particular, the absorption coefficient of Fp may be 399 400 strongly enhanced due to magnetic coupling between adjacent iron sites, which offers an efficient means of relaxing the spectroscopic selection rules, as has been shown for $Fe^{3+}-O_6$ octahedra 401 (Sherman and Waite, 1985). Considering that magnetic collapse is a defining characteristic of LS 402 Fe²⁺ (Cohen et al., 1997), the magnetic relaxation of spectroscopic selection rules is only possible 403 404 in HS-HS pairs. The collapse of magnetic moments of individual Fe atoms in the MS and LS 405 states is thus a viable explanation for the $\sim 30\%$ decrease of the absorption coefficients at P > 57 406 GPa. This mechanism, however, does not explain the relatively high refractive index dispersion 407 of the LS Fp.

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Page 14 of 19

409 **4. IMPLICATIONS**

410 A strong reduction of the electrical conductivity of MS ferropericlase opens a new 411 scenario for the overall conductivity of the lowermost mantle. Ab initio computations indicate 412 that ferropericlase is predominantly in the mixed spin state at depths greater than ~1900 km and 413 down to the core-mantle boundary at ~2900 km because of the high temperatures of the lower 414 mantle, which broaden the mixed spin pressure range (Holmstrom and Stixrude, 2015). If the 415 drop in electrical conductivity at ~50-70 GPa is due to the lower small polaron mobility/density 416 in LS Fp as proposed by Lin et al. (2007) and Ohta et al. (2007), then one expects only a 417 moderate drop in mantle conductivity if any at depths >1900 km, because LS Fp is never 418 dominant even at the base of the mantle. In contrast, if the drop in Fp conductivity is indeed 419 characteristic of the MS state (as inferred in this work), then one expects a significant reduction 420 of mantle electrical conductivity over the lowermost 1000 km of the mantle. The magnetic data 421 recorded by SWARM, which is a low-orbit satellite mission dedicated to the study of Earth's 422 magnetic field, now allows 3D mapping the conductivity of the Earth's mantle down to ~2000 423 km depth (Kuvshinov et al., 2021; Velimsky and Knopp, 2021). Deeper conductivity models will 424 be enabled by the continuing geomagnetic observations and may thus be sensitive to the changes 425 in conductivity due to the spin crossover in ferropericlase. Yet another alternative is that small 426 polaron is a relatively unimportant mechanism in the lower mantle because of the crossover to the 427 large polaron mechanism at $T > \sim 1000$ K, as suggested by earlier studies at P < ~ 30 GPa (Dobson and Brodholt, 2000; Dobson et al., 1997). A transition to a band conduction mechanism in Fp at T 428 429 $> \sim 2000$ K has been proposed by theoretical computations (Holmstrom et al., 2018), consistent 430 with the results of optical experiments at high P-T that indicate a crossover to opaque Fp at T >431 2500 K (Lobanov et al., 2021). Direct measurements of Fp electrical conductivity at realistic 432 lower mantle P-T conditions are thus needed to improve our understanding of lower mantle 433 conductivity.

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608