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Revision 1

Word Count: 10348

**Genesis of Mesozoic high-Mg dioritic rocks from the eastern North China  
Craton: Implications for the evolution of continental lithosphere**

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## ABSTRACT

17 Pre-Cenozoic High-Mg andesites (HMAs) are mostly present in continental interiors,  
18 but their genetic relationship with the continental lithosphere evolution remains unclear  
19 because of uncertainties of their mantle source, magmatic processes, and physicochemical  
20 conditions of formation. Early Cretaceous high-Mg dioritic rocks (HMDs, analogues of  
21 HMAs) of the Jinling complex in the Luxi area are typical intra-plate intrusions of the eastern  
22 North China Craton (NCC) and can be subdivided into two groups (Group-I and -II) on the  
23 basis of their petrographic and geochemical features. Group-I HMDs show low SiO<sub>2</sub> contents  
24 (52.47–56.10 wt%) and Sr/Y (34.5–39.6) and (La/Yb)<sub>N</sub> (10.3–13.6) ratios but high contents of  
25 MgO (7.86–9.13 wt%), Y (18.3–20.3 ppm), Yb (1.43–1.47 ppm), and compatible elements  
26 (Cr = 407–585 ppm; Ni = 117–216 ppm), classifying as sanukitic rocks. Group-II HMDs are  
27 characterized by high SiO<sub>2</sub> contents (63.81–64.87 wt%) and Sr/Y (47.1–63.4) and (La/Yb)<sub>N</sub>  
28 (16.1–17.5) ratios with low MgO (2.90–3.08 wt%), Y (0.88–1.04 ppm), Yb (0.88–1.04 ppm),  
29 and compatible elements (Cr = 201–213 ppm; Ni = 55–57 ppm) contents, belonging to  
30 adakitic rocks. Group-I and Group-II HMDs of the Jinling complex are closely related in  
31 spatial and temporal distribution, and all have enriched Sr-Nd isotopic compositions and  
32 arc-like trace elements patterns with abundant hydrous minerals. Therefore, the Jinling HMDs  
33 should share a common source of ancient sub-continental lithospheric mantle that had been  
34 metasomatized by aqueous fluids derived from the subducted Paleo-Pacific slab. The Jinling  
35 HMDs were not formed from interaction between slab-derived melts and mantle-wedge  
36 peridotites but were instead derived from partial melting of hydrous mantle peridotites in  
37 continental interior of the eastern NCC. The distinctly different petrography, geochemistry,  
38 and mineralogy of the two groups of rocks resulted mainly from differing magmatic processes  
39 at crustal depths. Thus, Pre-Cenozoic intra-plate HMAs/HMDs are genetically distinct from  
40 Cenozoic HMAs that were mostly present in arc settings and generally represent juvenile

41 crust growth. In a way, Archean tonalitic-trondhjemitic-granodioritic rocks (TTG) and  
42 sanukitoids, geochemically similar to HMAs/HMDs, could also be derived from interaction  
43 between slab-derived melts and mantle-wedge peridotites in arc settings or partial melting of  
44 hydrous mantle peridotites in continental interiors, and thus might not always be related with  
45 continental crustal growth and the onset of plate subduction.

46 **Key words:** High-Mg dioritic rocks; Magmatic processes; Fluid metasomatism;  
47 Sub-Continental lithospheric mantle; North China Craton.

48

## INTRODUCTION

49       The bulk composition of continental crust has been estimated to include 57–64 wt%  
50 SiO<sub>2</sub>, 4.4–6.7 wt% Na<sub>2</sub>O + K<sub>2</sub>O, and 3.2–4.7 wt% MgO and to have Mg# [= 100 ×  
51 Mg<sup>2+</sup>/(Mg<sup>2+</sup> + Fe<sup>total</sup>)] of 45–55 (e.g., [Rudnick 1995](#); [Rudnick and Gao 2014](#)). Cenozoic  
52 HMAs have similar compositional characteristics to the bulk crust (e.g., [Kelemen 1995](#)).  
53 Despite their small magmatic volumes in modern subduction zones, Cenozoic HMAs have  
54 attracted considerable research attention during the past three decades because they can  
55 provide insights into the geodynamics of continental growth and the onset of plate tectonics,  
56 for their compositional similarities to Archean TTG and sanukitoids (e.g., [Shirey and Hanson](#)  
57 [1984](#); [Kelemen 1995](#); [Rudnick 1995](#); [Tatsumi 2001, 2008](#); [Martin et al. 2005](#); [Wang et al.](#)  
58 [2020a](#); [Xu et al. 2020](#)). Cenozoic HMAs can be categorized into four sub-types on the basis of  
59 their petrographic and geochemical characteristics, i.e., adakitic, bajaitic, sanukitic, and  
60 boninitic HMAs (e.g., [Yogodzinski et al. 1995](#); [Kemei et al. 2004](#); [Tang and Wang 2010](#);  
61 [Wang et al. 2020a](#)). These sub-types are generated through different mechanisms and have  
62 distinct implications for slab–mantle interaction at modern convergent plate margins.  
63 Cenozoic HMAs occur mainly in oceanic subduction zones and subordinately in continental  
64 collision zones away from intracontinental settings (e.g., [Defant and Drummond 1990](#);  
65 [Yogodzinski et al. 1994, 1995](#); [Tatsumi 2001, 2008](#); [Wang et al. 2020a](#); [Xu et al. 2020](#)).  
66 However, Pre-Cenozoic HMAs, including Archean TTG and sanukitoids, have also been  
67 reported in continental interiors, such as the eastern NCC and the Central Asian orogenic belt  
68 (e.g., [Gao et al. 2004](#); [Yang et al., 2012a, 2012b](#); [Wang et al. 2020a](#)). Compared with the  
69 well-developed understanding of Cenozoic HMAs in arc settings, the nature and genesis of  
70 Pre-Cenozoic intra-plate HMAs remain unclear, including the mantle source, magmatic  
71 processes, and physicochemical conditions of formation.

72       Early Cretaceous HMDs are widely distributed in the eastern NCC (Fig. 1), which had

73 undergone a significant lithospheric thinning and destruction in Mesozoic with a thick and  
74 cold Paleozoic sub-continental lithospheric mantle (SCLM) replaced by a thin and hot  
75 Cenozoic SCLM (e.g., [Menzies et al., 1993](#); [Xu, 2001](#); [Gao et al. 2004](#); [Huang et al. 2012](#);  
76 [Yang et al. 2021](#)). These rocks generally have high contents of compatible elements and high  
77 Mg# values, and typically display arc-like trace-element characteristics, i.e., enrichment in  
78 light rare earth elements (REE) and large-ion lithophile elements (LILE) and depletion in  
79 high-field-strength elements (HFSE) and heavy REE, and enriched radiogenic isotope  
80 compositions (e.g., [Yang et al. 2012a, 2012b](#); [Jin et al. 2015](#); [Lan et al. 2019](#); [Sun et al. 2019](#);  
81 [Gao et al. 2021](#); [Zhang et al. 2021](#)). Although their whole-rock major and trace element  
82 features are similar to those of Cenozoic HMAs, the Early Cretaceous HMDs are commonly  
83 considered to have been generated in an intracontinental setting during the Mesozoic  
84 descratonization of the eastern NCC (e.g., [Gao et al. 2004](#); [Yang et al. 2012a, 2012b](#); [Jin et al.](#)  
85 [2015](#); [Lan et al. 2019](#); [Sun et al. 2019](#); [Gao et al. 2021](#); [Zhang et al. 2021](#); [Guo et al. 2022](#)).  
86 However, the origin of the Mesozoic intra-plate HMDs remains controversial, with four main  
87 models having been proposed: interaction between delaminated lower continental  
88 crust-derived melts and mantle peridotites (e.g., [Gao et al. 2004](#); [Yang et al. 2006](#); [Zhang et al.](#)  
89 [2010](#); [Jin et al. 2015](#)); partial melting of enriched lithospheric mantle metasomatized by felsic  
90 melts derived from the delaminated lower continental crust or the subducted continental crust  
91 (e.g., [Yang et al. 2012a, 2012b](#); [Lan et al. 2019](#); [Gao et al. 2021](#)); magma mixing between  
92 crust-derived felsic melts and mantle-derived mafic melts (e.g., [Chen et al. 2013](#)) and  
93 assimilation of mantle peridotite by monzodioritic magmas at crustal depths (e.g., [Qian and](#)  
94 [Hermann 2010](#)). Here, we present whole-rock element and isotopic data and in situ mineral  
95 (amphibole and plagioclase) compositions for the Jinling intrusions in the Luxi region, typical  
96 intra-plate HMDs of the eastern NCC, where the SCLM might have been affected by the  
97 subducted Yangtze continental crust in Triassic or the subducted Paleo-Pacific oceanic crust in

98 Jurassic-Cretaceous. The objective of the study was to investigate the magma sources,  
99 magmatic processes, and associated physicochemical conditions of the Jinling HMDs to gain  
100 insights into the petrogenesis and geodynamics of intra-plate HMAs and their role in the  
101 evolution of continental lithosphere.

102

### 103 GEOLOGICAL SETTING AND SAMPLE DESCRIPTIONS

104 The NCC is bounded by the early Paleozoic Qilianshan Orogen to the west, the  
105 Paleozoic Central Asian orogenic belt to the north, and the Qinling–Dabie–Sulu orogenic belt  
106 to the south and the east, respectively (Fig. 1a; e.g., [Zhao et al. 2005](#); [Zheng et al. 2013](#)). The  
107 NCC is subdivided into the Eastern and Western blocks by the Trans-North China Orogen  
108 (e.g., [Zhao et al. 2005](#)). The eastern NCC, lying to the east of the Daxinganling–Taihangshan  
109 Gravity Lineament (DTGL), underwent intensive tectono-magmatic activities during the  
110 Mesozoic and also pronounced lithospheric thinning (e.g., [Menzies et al. 1993](#); [Gao et al.](#)  
111 [2004](#); [Xu et al. 2004](#); [Zhu et al. 2011](#); [Huang et al. 2012](#); [Yang et al. 2021](#)). Shandong  
112 Province in the central part of the eastern NCC is separated by the Tan–Lu fault zone (TLFZ)  
113 into two parts (Fig. 1a and b), i.e., the Luxi block and the Jiaodong Peninsula (e.g., [Huang et](#)  
114 [al. 2012](#)).

115 The crystalline basement in the Luxi area is composed chiefly of Neoproterozoic Taishan  
116 Group TTG gneisses, which is unconformably overlain by Cambrian to Lower–Middle  
117 Ordovician clastic–carbonate successions. Late Carboniferous to Triassic marine–  
118 terrigenous-facies sedimentary rocks unconformably overlie Lower–Middle Ordovician  
119 limestones ([Liu et al. 1996](#)). Jurassic to Cretaceous terrestrial clastic rocks are unconformably  
120 overlain by Cenozoic strata composed predominantly of alluvial and lacustrine sediments  
121 ([Song 2008](#)). In addition to Precambrian magmatic rocks, voluminous Mesozoic intrusions

122 are widespread in the Luxi area and represent two stages of magmatism; i.e., Early to Middle  
123 Jurassic monzonitic–syenitic magmatism (ca. 160–155 Ma) and more extensive Early  
124 Cretaceous gabbrodioritic–dioritic–monzonitic magmatism (ca. 132–112 Ma; e.g., [Xu et al.](#)  
125 [2004](#); [Huang et al. 2012](#); [Zhong and Huang 2012](#); [Jin et al. 2015](#); [Gao et al. 2021](#); [Zhang et al.](#)  
126 [2021](#)).

127 The Jinling high-Mg dioritic complex is located in the Luxi area (Fig. 1b) and consists  
128 of a main body with several separate stocks that intruded Ordovician limestone and dolomite  
129 sequences of the Majiagou Formation (Fig. 1c; e.g., [Zhong and Huang 2012](#)). This complex is  
130 extensively covered by Quaternary deposits, meaning that it is difficult to observe the nature  
131 of contacts between different types of constituent rock. The Jinling HMDs have generally  
132 been subdivided into Group-I and Group-II HMDs in previous studies (e.g., [Yang et al. 2006](#),  
133 [2012a](#), [2012b](#); [Zhong and Huang 2012](#); [Jin et al. 2015](#); [Gao et al. 2021](#); [Zhang et al. 2021](#);  
134 [Guo et al. 2022](#)). Group-I HMDs consist of gabbroic diorite and hornblende diorite, whereas  
135 Group-II HMDs are monzonite. This subdivision is followed in this study.

136 Seven gabbroic diorite, one hornblende diorite, and two monzonite samples were  
137 collected from the Jinling complex for whole-rock geochemical and mineral electron  
138 microprobe analyses. The gabbroic diorites are fresh and show porphyritic texture (Fig. 2a–c),  
139 with phenocrysts of orthopyroxene (5–10 vol%), clinopyroxene (5–10 vol%), amphibole (~5  
140 vol%), and biotite (~5 vol%), and a matrix that is composed mainly of fine-grained  
141 plagioclase (30–35 vol%), amphibole (15–20 vol%), K-feldspar (~10 vol%), biotite (~5 vol%),  
142 and clinopyroxene (~5 vol%), with accessory minerals of magnetite, apatite, and zircon.  
143 Orthopyroxene and clinopyroxene phenocrysts are generally replaced by amphibole in rims  
144 (Fig. 2a–c), and amphibole phenocrysts show complex compositional zoning (Fig. 2d). The  
145 hornblende diorites also show porphyritic texture with amphibole phenocrysts of 20–25 vol%  
146 (Fig. 2d). The matrix consists primarily of fine-grained plagioclase (40–45 vol%), amphibole

147 (15–20 vol%), K-feldspar (~5 vol%), and biotite (~5 vol%), with accessory minerals of  
148 magnetite, apatite, and zircon. The monzonites show porphyritic texture (Fig. 2e and f), with  
149 phenocrysts of plagioclase (~20 vol%) and amphibole (15–20 vol%), and a matrix that  
150 consists predominantly of fine-grained plagioclase (~15 vol%), amphibole (5–10 vol%),  
151 K-feldspar (30–35 vol%), and anhedral quartz (<5 vol%). Accessory minerals include  
152 magnetite, apatite, titanite, and zircon. Amphibole and plagioclase phenocrysts show complex  
153 compositional zoning (Fig. 2e and f). In the gabbroic and hornblende diorites, plagioclases  
154 appear only in the matrix and lack zoned texture (Fig. 2a–d). Magnetites are usually present  
155 around and/or are included in rims of amphibole phenocrysts and matrix amphiboles.

156

## 157 ANALYTICAL METHODS

### 158 **Whole-rock major and trace elements**

159 Whole-rock major elements were analyzed using a Rigaku RIX 2000 X-ray  
160 fluorescence spectrometer (XRF) at Guangzhou Institute of Geochemistry, Chinese Academy  
161 of Sciences (GIG-CAS), Guangzhou, China. The analytical uncertainties are mostly less than  
162 2%. Whole-rock trace element concentrations were obtained by the Thermal X series 2  
163 inductively coupled plasma-mass spectrometry (ICP-MS) equipped with a Cetac ASX-560  
164 AutoSampler at the Tongwei Analytical Technology Company (TATC), Guizhou, China, and  
165 the ICP-MS procedure for trace element analysis followed the protocols of Eggins et al.  
166 (1997), with modifications described in Kamber et al. (2003) and Li et al. (2005). The  
167 analytical precisions are better than 5% for most trace elements, estimated from analytical  
168 results of the USGS Rock References W-2a and BHVO-2 in the same measurement session.

### 169 **Whole-rock Sr-Nd isotopes**

170 The separation and purification procedures for the whole-rock Sr–Nd isotopes were



171 performed at TATC. Chemical separation was performed by conventional ion-exchange  
172 techniques. The detailed chemical procedures are similar to those described in Pin et al.  
173 (1997), Deniel and Pin (2001) and Míková et al. (2007). Total procedure blanks are typically  
174 in the ranges of  $\leq 100$  pg for Sr and  $\leq 60$  pg for Nd.

175 The purified solution Sr and Nd isotope ratios were measured by a Neptune Plus  
176 MC-ICP-MS at GIG-CAS. The analytical precisions of isotopic ratio were reported as  $2\sigma$   
177 standard errors. Normalizing factors of  $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$  and  $^{146}\text{Nd}/^{144}\text{Nd} = 0.7219$  are  
178 used to correct the mass fractionations of Sr and Nd during the measurements, respectively.  
179 During the analytical sessions, the measured values for standards NBS987 Sr were  $^{87}\text{Sr}/^{86}\text{Sr} =$   
180  $0.710248 \pm 8$  ( $2\sigma$ ,  $n=8$ ) and those for JNdi-1 Nd were  $^{143}\text{Nd}/^{144}\text{Nd} = 0.512115 \pm 4$  ( $2\sigma$ ,  $n=9$ ).  
181 Two USGS reference materials W-2a and BHVO-2 were also processed for Sr–Nd isotopes to  
182 monitor the analytical accuracy and gave ratios of  $0.706957 \pm 10$  and  $0.703480 \pm 14$  for  $^{87}\text{Sr}/^{86}\text{Sr}$ ,  
183 respectively, and of  $0.512509 \pm 10$  and  $0.512989 \pm 18$  for  $^{143}\text{Nd}/^{144}\text{Nd}$ , respectively, which are in  
184 agreement with the recommended values by Fourny et al. (2016) within errors. The analytical  
185 procedures are principally similar to the description in Wei et al. (2002) and Liang et al.  
186 (2003).

### 187 **Back-scattered electron (BSE) images and Electron microprobe analyses**

188 BSE images of the amphibole and plagioclase crystals were captured using a Carl Zeiss  
189 SUPRA55SAPPHIRE Field Emission-Scanning Electron Microscope (SEM) at GIG-CAS. In  
190 situ major elemental analyses were obtained using a Cameca SXFive FE Electron Probe  
191 Microanalyzer (EPMA) at GIG-CAS. This EPMA is equipped with an electron optical column  
192 with field emission source, controlled by Cameca PeakSight software. An operating condition  
193 of 15 kV accelerating voltage and 20 nA beam current was used during the course of this  
194 study. A variable peak counting time (10–60 s) was designed based on the intensity of  
195 characteristic X-ray line and desired precision for the element. Calibration standards used for

196 feldspar analyses were albite (Na), almandine (Mg), sanidine (Si, K), hematite (Fe),  
197 plagioclase (Ca), and Celestite (Sr). Calibration standards for amphibole analyses were jadeite  
198 (Na, Al), diopside (Si, Mg, Ca), orthoclase (K), rutile (Ti), Cr<sub>2</sub>O<sub>3</sub> (Cr), hematite (Fe),  
199 rhodonite (Mn), topaz (F) and tugtupite (Cl) from SPI company. The PAP (Pouchou and  
200 Pichoir) procedure was used for matrix correction (Pouchou and Pichoir 1991). The detailed  
201 procedures are the same as those described in He et al. (2021).

202

203

## RESULTS

204 All the whole-rock and mineral geochemical data are listed in Supplementary Tables  
205 S1–3.

### 206 Whole-rock major and trace elements

207 Whole-rock major and trace elements compositions for the Jinling HMDs are presented  
208 in Supplementary Table S1. These compositions overlap with previously published data in  
209 diagrams of major and trace elements (Figs. 3–5). Group-I HMDs have low contents of SiO<sub>2</sub>  
210 (52.47–56.10 wt%), Al<sub>2</sub>O<sub>3</sub> (11.67–13.00 wt%) and K<sub>2</sub>O + Na<sub>2</sub>O (4.20–5.39 wt%), and they  
211 plot in the field of gabbro diorite in a total alkali versus silica (TAS) diagram (Table S1; Fig.  
212 3a). Group-I HMDs also have high contents of MgO (7.86–9.13 wt%), total (T)Fe<sub>2</sub>O<sub>3</sub> (8.63–  
213 10.97 wt%) and CaO (6.61–8.48 wt%) with high Mg# values (61.3–66.8). Compared with  
214 Group-I HMDs, Group-II HMDs have considerably higher contents of SiO<sub>2</sub> (63.81–64.87  
215 wt%), Al<sub>2</sub>O<sub>3</sub> (15.16–15.53 wt%) and K<sub>2</sub>O + Na<sub>2</sub>O (8.21–8.67 wt%), and they plot in the field  
216 of quartz monzonite in a TAS diagram (Table S1; Fig. 3a). Group-II HMDs also have much  
217 lower contents of MgO (2.90–3.08 wt%), TFe<sub>2</sub>O<sub>3</sub> (2.40–3.66 wt%), and CaO (3.44–4.40  
218 wt%). Despite their much lower MgO contents, Group-II HMDs show considerably variable  
219 Mg# values (61.1–71.8) that are comparable to those of Group-I HMDs. Both Group-I and

220 Group-II HMDs as classified as high-K calc-alkaline series, but Group-II HMDs have much  
221 higher  $K_2O$  contents than Group-I (Fig. 3b). In Harker diagrams of major element oxides or  
222 compatible elements versus MgO, Group-I and Group-II HMDs show roughly similar  
223 geochemical trends with a pronounced compositional gap in MgO contents (Fig. 4). For both  
224 Group-I and Group-II HMDs, the contents of CaO,  $P_2O_5$ ,  $TiO_2$ ,  $TFe_2O_3$ , and other compatible  
225 elements decrease with decreasing MgO content, whereas those of  $SiO_2$  and  $Al_2O_3$  increase.  
226 However, for most major and trace elements, the geochemical variations of Group-II HMDs  
227 are wider than those of Group-I HMDs.

228 The two groups of HMDs show highly fractionated chondrite-normalized REE patterns  
229 with weak/negligible Eu anomalies (Fig. 5a) and are characterized by enrichment in LILE,  
230 negative Nb–Ta–Ti anomalies, and positive Pb anomalies in primitive mantle normalized  
231 multi-element diagrams (Fig. 5b). Although the two groups have similar LREE contents,  
232 Group-II HMDs have considerably lower contents of middle and heavy REE (Fig. 5).  
233 Furthermore, Group-I HMDs have substantially higher Yb (1.43–1.47 ppm) and Y (18.3–20.3  
234 ppm) contents but much lower  $(La/Yb)_N$  (10.3–13.6) and Sr/Y (34.5–39.6) ratios than  
235 Group-II HMDs [Yb = 0.88–1.04 ppm, Y = 10.5–12.3 ppm,  $(La/Yb)_N = 16.1–17.5$ , and Sr/Y =  
236 47.1–63.4; N denotes the normalization values relative to chondrite.]. In addition, Group-I  
237 HMDs exhibit slight negative Eu anomalies ( $Eu/Eu^* = 0.93–0.95$ ), and Group-II HMDs  
238 display slightly positive Eu anomalies ( $Eu/Eu^* = 1.03–1.17$ ; Fig. 5) owing to their much  
239 higher plagioclase contents (Fig. 2). Furthermore, Group-I HMDs have much higher contents  
240 of compatible elements (Cr = 407–585 ppm, Ni = 117–216 ppm) but lower contents of LILE  
241 (Rb = 28.8–48.1 ppm, Ba = 592–1030 ppm) compared with Group-II HMDs (Table S1; Figs.  
242 4g–i and 5b).

### 243 **Whole-rock Sr–Nd isotopes**

244 Whole-rock Sr–Nd isotopic compositions for the two groups of HMDs are listed in

245 Supplementary Table S1 and are plotted in Fig. 6c together with literature data. The studied  
246 samples commonly show enriched Sr–Nd isotopic compositions with distinct variations  
247 between the two groups. Group-I HMDs display variably low ( $^{87}\text{Sr}/^{86}\text{Sr}$ )<sub>i</sub> ratios (0.7048–  
248 0.7052) and small negative  $\epsilon_{\text{Nd}}(t)$  values (–6.61 to –3.75), with two-stage Nd model ages of  
249 1.47–1.23 Ga. In contrast, Group-II HMDs show slightly higher ( $^{87}\text{Sr}/^{86}\text{Sr}$ )<sub>i</sub> ratios (0.7054–  
250 0.7055) and more negative  $\epsilon_{\text{Nd}}(t)$  values (–9.29 to –8.60), corresponding to substantially older  
251 two-stage Nd model ages of 1.69–1.63 Ga (Table S1; Fig. 6c).

### 252 **BSE images and mineral geochemistry of amphibole and plagioclase**

253 The major element compositions of amphibole and plagioclase from the two groups of  
254 HMDs are presented in Supplementary Tables S2–3 and Figs. 7c and 8c–d. As shown in the  
255 BSE images (Fig. 7a and b), plagioclases from the two groups of HMDs show distinctive  
256 crystal morphology and internal texture. Plagioclase phenocrysts from Group-II HMDs show  
257 complex inner core (core I)–outer core (core II)–mantle–rim zoning with variable lightness in  
258 different domains (Fig. 7b). Euhedral plagioclases in the matrix and outer cores and rims of  
259 plagioclase phenocrysts have similar darkness, and they are much darker than the inner cores  
260 and mantles of plagioclase phenocrysts from Group-II HMDs. The outer cores of plagioclase  
261 phenocrysts from Group-II HMDs show relict textures of disequilibrium reaction (Fig. 7b).  
262 Different domains in the plagioclase phenocrysts have highly variable anorthite (An) contents.  
263 Inner cores have the highest An contents (34.6%–41.9%) and are andesine in composition. In  
264 contrast, the rims of plagioclase phenocrysts and matrix plagioclases have low An contents  
265 (8.8%–13.4% and 7.7%–13.5%, respectively) and are albite or oligoclase (Fig. 7c), suggesting  
266 normal zoning of plagioclase phenocrysts overall. However, the outer cores of plagioclase  
267 phenocrysts have much lower An contents (11.1%–19.4%; Fig. 7c) relative to mantles (An =  
268 22.0%–31.6%; Fig. 7c), indicating reverse zoning within plagioclase phenocrysts. The outer  
269 cores of plagioclase phenocrysts from Group-II HMDs are best interpreted to be relicts of

270 crustal contamination, as indicated by their relict texture and similar compositional  
271 characteristics to plagioclases from Archean TTGs (e.g., [Jahn et al. 1988](#)). Group-I HMDs  
272 lack plagioclase phenocrysts (Fig. 7a) and contain fine-grained euhedral plagioclase in the  
273 matrix (Fig. 7a). Matrix plagioclases in Group-I HMDs have An contents of 22.2%–30.0%  
274 and plot in the fields of oligoclase and andesine (Table S2; Fig. 7c), similar to the mantles of  
275 plagioclase phenocrysts from Group-II HMDs.

276 Both phenocryst and matrix amphiboles are found in the two groups of HMDs. The  
277 fine-grained amphiboles in the matrix are homogeneous, whereas amphibole phenocrysts  
278 from Group-I and Group-II HMDs show core–rim and core–mantle–rim compositional zoning  
279 patterns, respectively (Fig. 8a and b). The cores of amphibole phenocrysts from the Group-I  
280 and Group-II samples exhibit similar major element compositions, showing the lowest SiO<sub>2</sub>  
281 contents (42.18–46.09 wt% and 43.66–44.90 wt%, respectively) and the highest Al<sub>2</sub>O<sub>3</sub> (7.96–  
282 10.69 wt% and 9.26–11.50 wt%, respectively) and TiO<sub>2</sub> contents (1.75–3.51 wt% and 0.51–  
283 3.05 wt%, respectively) of all phenocryst domains, and most of them are pargasite (Fig. 8c).  
284 The rims of amphibole phenocrysts from the Group-I and Group-II HMDs are also  
285 compositionally similar to each other. These rims have higher SiO<sub>2</sub> contents (48.50–54.34 wt%  
286 and 51.11–53.39 wt%, respectively) and much lower Al<sub>2</sub>O<sub>3</sub> (1.91–6.37 wt% and 2.79–4.64  
287 wt%, respectively) and TiO<sub>2</sub> (0.58–1.38 wt% and 0.65–1.09 wt%, respectively) contents than  
288 the cores and are classified as magnesiohornblende (Fig. 8d). The mantles of amphibole  
289 phenocrysts from Group-II HMDs have contents of SiO<sub>2</sub> (44.55–46.97 wt%), Al<sub>2</sub>O<sub>3</sub> (8.18–  
290 9.36 wt%), and TiO<sub>2</sub> (1.09–2.67 wt%) that are intermediate between those of cores and rims,  
291 and most of their compositions fall in the field of edenite (Fig. 8c). Thus, amphibole  
292 phenocrysts from the two groups of HMDs show normal compositional zoning overall.  
293 Matrix amphiboles in Group-I and Group-II HMDs are similar to the rims of amphibole  
294 phenocrysts in terms of their internal textures and major element contents (Table S3; Fig. 8d),

295 indicating their formation under similar physicochemical conditions. However, subtle  
296 differences exist between Group-I and Group-II HMDs with respect to the internal textures  
297 and major element compositions of amphibole phenocrysts. For example, Group-II amphibole  
298 phenocrysts have mantles, whereas Group-I HMD amphibole phenocrysts do not have (Fig.  
299 8a), suggesting that Group-II HMDs may have undergone more complex magmatic processes  
300 in the crustal magma chamber. In addition, amphiboles of all phenocryst domains from  
301 Group-II HMDs have considerably higher Na<sub>2</sub>O contents than those from Group-I HMDs at  
302 given SiO<sub>2</sub> contents (Table S3).

303

## 304 **DISCUSSION**

### 305 **Classification of the Jinling high-Mg dioritic rocks**

306 High-Mg andesite, as a special type of andesite, is generally characterized by high SiO<sub>2</sub>  
307 (>52 wt%) and high Mg# values (>45) as well as low FeO<sup>T</sup>/MgO ratios (<1.5) (e.g.,  
308 [Yogodzinski et al. 1994, 1995](#); [Kelemen 1995](#); [Tatsumi 2001, 2008](#); [Tang and Wang 2010](#);  
309 [Wang et al. 2020a](#)). In general, Cenozoic HMAs occur mainly in arc settings regardless of the  
310 different sub-types (i.e., adakitic, bajaitic, sanukitic, and boninitic) (e.g., [Yogodzinski et al.](#)  
311 [1994, 1995](#); [Kemei et al. 2004](#); [Tang and Wang 2010](#); [Wang et al. 2020a](#)). Cenozoic adakitic  
312 HMAs show high Sr/Y and (La/Yb)<sub>N</sub> ratios and formed through minor interaction between  
313 slab-derived melts and mantle peridotite (e.g., [Kay 1978](#); [Defant and Drummond 1990](#);  
314 [Yogodzinski et al. 1995](#); [Kemei et al. 2004](#); [Tang and Wang 2010](#)). Cenozoic sanukitic HMAs  
315 show similar REE patterns to those of adakitic rocks but have much higher MgO contents  
316 (e.g., [Yogodzinski et al. 1994, 1995](#); [Tatsumi 2001, 2008](#); [Kemei et al. 2004](#); [Tang and Wang](#)  
317 [2010](#); [Wang et al. 2020a](#)). Bajaitic HMAs have approximately similar trace-element  
318 characteristics to those of adakitic rocks, except for their extreme enrichment in Sr and

319 depletion in Rb (e.g., [Rogers et al. 1985](#); [Kemei et al. 2004](#)). Boninitic HMAs are enriched in  
320 bronzite phenocrysts, glass matrix, and augite microlites but lacks plagioclase crystals, and  
321 they are characterized by  $\text{SiO}_2 > 52 \text{ wt\%}$ ,  $\text{MgO} > 8 \text{ wt\%}$ ,  $\text{TiO}_2 < 0.5 \text{ wt\%}$ , U-shaped REE  
322 patterns, and marked depletion in HFSEs and REEs (e.g., [Kemei et al. 2004](#); [Tang and Wang](#)  
323 [2010](#); [Wang et al. 2020a](#)). Although Pre-Cenozoic HMAs/HMDs occur widely in continental  
324 interiors worldwide, most of them have approximately similar major and trace elements  
325 compositions to those of Cenozoic adakitic and sanukitic HMAs (e.g., [Shirey and Hanson](#)  
326 [1984](#); [Martin et al. 2005](#); [Jin et al. 2015](#); [Lan et al. 2019](#); [Sun et al. 2019](#); [Gao et al. 2021](#);  
327 [Zhang et al. 2021](#)) with few resembling Cenozoic bajaitic and boninitic HMAs (e.g., [Polat et](#)  
328 [al. 2002](#)). Cenozoic adakitic and sanukitic HMAs show remarkable differences in petrography  
329 and whole-rock geochemistry (e.g., [Yogodzinski et al. 1994, 1995](#); [Kemei et al. 2004](#); [Tang](#)  
330 [and Wang 2010](#); [Wang et al. 2020a](#)). In general, adakitic HMAs contain more feldspathic  
331 minerals and less mafic minerals, have much higher  $\text{SiO}_2$  and LILE contents, higher Sr/Y and  
332  $(\text{La}/\text{Yb})_N$  ratios, and much lower MgO (and Mg# values), compatible-element (i.e., V, Cr, and  
333 Ni), Y, and Yb contents compared with sanukitic HMAs (e.g., [Yogodzinski et al. 1994, 1995](#);  
334 [Kemei et al. 2004](#); [Tang and Wang 2010](#)).

335 Group-I HMDs of the Jinling complex have low  $\text{SiO}_2$ , but high MgO,  $\text{TFe}_2\text{O}_3$ , and CaO  
336 contents with high Mg# values, which are similar to Cenozoic sanukitic HMAs (Fig. 9a and b;  
337 e.g., [Yogodzinski et al. 1994](#); [Tatsumi 2001, 2008](#); [Tang and Wang 2010](#); [Wang et al. 2020a](#)).  
338 Group-II HMDs also have high Mg# values but much higher contents of  $\text{SiO}_2$  and  $\text{K}_2\text{O} +$   
339  $\text{Na}_2\text{O}$ , and substantially lower contents of MgO,  $\text{TFe}_2\text{O}_3$  and CaO, and they are roughly  
340 comparable to the major-element compositions of Cenozoic adakitic HMAs worldwide (Fig.  
341 9a and b; e.g., [Yogodzinski et al. 1995](#); [Kemei et al. 2004](#); [Tang and Wang 2010](#); [Wang et al.](#)  
342 [2020a](#)). Furthermore, Group-I HMDs show substantially higher Yb (1.43–1.47 ppm) and Y  
343 (18.3–20.3 ppm) contents but lower  $(\text{La}/\text{Yb})_N$  (10.3–13.6) and Sr/Y (34.5–39.6) ratios than

344 those of Group-II HMDs ( $Yb = 0.88\text{--}1.04$  ppm,  $Y = 10.5\text{--}12.3$  ppm,  $(La/Yb)_N = 16.1\text{--}17.5$ ,  
345 and  $Sr/Y = 47.1\text{--}63.4$ ; Table S1; Fig. 9). In diagrams of Sr/Y versus Y (Fig. 9c and d), the  
346 majority of Group-II samples plot in the fields of adakite and adakitic HMAs, whereas  
347 Group-I samples plot in and/or near the fields of island-arc volcanic rocks and sanukitic  
348 HMAs. In addition, Group-I HMDs have much higher compatible-element contents ( $V =$   
349  $172\text{--}254$  ppm,  $Cr = 407\text{--}585$  ppm, and  $Ni = 117\text{--}216$  ppm) but lower LILE contents ( $Rb =$   
350  $28.8\text{--}48.1$  ppm and  $Ba = 592\text{--}1030$  ppm, and  $Pb = 8.58\text{--}22.0$  ppm) compared with Group-II  
351 HMDs (Table S1; Figs. 4g–i and 5). Therefore, the Group-I and Group-II HMDs of the Jinling  
352 complex are geochemically classified as sanukitic and adakitic HMDs, respectively.

### 353 **Generation of the Jinling high-Mg dioritic rocks: Magmatic processes and** 354 **physicochemical conditions**

355 **Magmatic processes.** The two groups of HMDs in the Jinling complex display different  
356 geochemical compositions (Table S1; Figs. 4 and 6), which suggests that they underwent  
357 different magmatic processes and/or differing degrees of magmatic evolution in the crustal  
358 magma chamber. Group-I HMDs show variable LREE contents (e.g., La) with uniform La/Sm  
359 ratios (Fig. 10a), indicating that their compositional variations were controlled mainly by  
360 fractional crystallization. Slight crustal contamination is also inferred to have been involved  
361 in the generation of Group-I HMDs. The chemical compositions of amphibole in igneous  
362 rocks have been widely used to determine magma sources (i.e., mantle or crust; e.g., [Jiang](#)  
363 [and An 1984](#); [Zhang et al. 2015](#); [Sun et al. 2019](#)). Although the absolute contents of major  
364 elements in amphibole to a certain extent depends on physical conditions of host melt ([Ridolfi](#)  
365 [et al. 2010](#)), there are good linear correlations between Si, Ti and Al in wide ranges of  
366 pressures (0–10 kbar) and temperatures (650–1075 °C; e.g., [Putirka 2016](#)). In particular, the  
367  $Si/(Si + Ti + Al)$  ratio of amphibole is a reliable indicator of the source of host magma, with  
368 values of  $\geq 0.775$  for crust-derived amphibole and  $\leq 0.765$  for mantle-derived amphibole (e.g.,



369 [Jiang and An 1984](#); [Li et al. 2021](#)). As shown in Fig. 11, the cores of amphibole phenocrysts  
370 from Group-I HMDs plot predominantly in the mantle source field, whereas the rims of  
371 amphibole phenocrysts and matrix amphiboles fall mainly in the crustal source and/or crust–  
372 mantle mixed source fields, suggesting that crustal involvement was mostly occurred in the  
373 latest stage of magma evolution. The cores of amphibole phenocrysts from Group-I HMDs  
374 have Si/(Si + Ti + Al) ratios of 0.74–0.81, whereas rims and matrix amphiboles show much  
375 higher Si/(Si + Ti + Al) ratios (0.85–0.95; Table S3). However, the effects of crustal  
376 contamination was negligible in the genesis of Group-I HMDs, as inferred from the narrow  
377 ranges of whole-rock elemental and isotopic compositions (Figs. 6c and 10a–b) and the  
378 absence of inherited Neoproterozoic zircons (e.g., [Yang et al. 2012b](#); [Zhong and Huang 2012](#)).

379 The MgO contents of the Group-I and Group-II samples are positively related to  
380 contents of CaO, TiO<sub>2</sub>, TFe<sub>2</sub>O<sub>3</sub>, and compatible elements and negatively related to SiO<sub>2</sub> and  
381 Al<sub>2</sub>O<sub>3</sub> contents (Fig. 4), indicating pronounced fractional crystallization of mafic minerals  
382 (i.e., olivine, orthopyroxene, and clinopyroxene) and Fe–Ti oxides. However, the Group-II  
383 samples show a weak positive trend in the diagram of La/Sm versus La (Fig. 10a), indicating  
384 that their compositional variations cannot be explained primarily by fractional crystallization  
385 and suggesting that crustal contamination also played an important role in generating  
386 Group-II HMDs. In diagrams of Si–Ti–Al, TiO<sub>2</sub> versus Al<sub>2</sub>O<sub>3</sub>, and (Na + K) versus Al<sup>IV</sup> (Fig.  
387 11), the cores of amphibole phenocrysts from Group-II HMDs plot mostly in the mantle  
388 source field, mantles fall in the crust–mantle mixing source field, and rims and matrix  
389 amphiboles plot predominantly in the crustal source field, suggesting substantial crustal  
390 contamination in different stage of magma evolution. In addition, the Si/(Si + Ti + Al) ratios  
391 of amphiboles gradually increase from cores (0.757–0.775) to mantles (0.784–0.811) to rims  
392 and matrix amphiboles (0.892–0.934; Table S3), consistent with an increasing influence of  
393 crustal contamination during magma evolution. This interpretation is supported by the relict

394 textures and low An contents of outer cores of plagioclase phenocryst from Group-II HMDs  
395 (Table S2; Fig. 7b and c). Crustal contamination can also account for the main patterns of  
396 variation in whole-rock geochemical and isotopic compositions of samples of Group-II  
397 HMDs (Figs. 6b–c and 10). Most Group-II HMD samples fall in the field of Taishan Group  
398 TTG gneisses (e.g., [Jahn et al. 1998](#); [Peng et al. 2013](#); [Chen et al. 2020](#)) in diagrams of La/Sm  
399 versus La and  $(\text{Hf/Sm})_N$  versus  $(\text{Ta/La})_N$  (Fig. 10a and b), consistent with the involvement of  
400 ancient crustal materials. In addition, samples from Group-II HMDs have higher  $(^{87}\text{Sr}/^{86}\text{Sr})_i$ ,  
401  $(\text{La/Yb})_N$ , and Sr/Y and lower  $\epsilon_{\text{Nd}}(t)$  values compared with Group-I HMDs (Fig. 10c and d),  
402 which suggests more substantial involvement of crustal materials during the generation of  
403 Group-II HMDs. The Taishan Group TTG gneisses, as the main components of crystalline  
404 basement in the Luxi area, are identified as the most likely candidate for the involvement of  
405 crust during magma evolution (Figs. 6b–c and 10; e.g., [Jahn et al. 1988](#); [Peng et al. 2013](#);  
406 [Chen et al. 2020](#)), as supported by the presence of inherited Neoproterozoic zircons in  
407 monzonites from the Jinling complex (e.g., [Jin et al. 2015](#); [Gao et al. 2021](#); [Zhang et al.](#)  
408 [2021](#)).

409 **Physicochemical conditions.** Physicochemical conditions, such as melt water content  
410 ( $\text{H}_2\text{O}_{\text{melt}}$ ), temperature (T), pressure (P), and oxygen fugacity ( $f\text{O}_2$ ), commonly play an  
411 important role in controlling the paths of magmatic evolution and the petrographic and  
412 geochemical variations of genetically associated magmas ([Richards 2011](#)). By applying the  
413 thermobarometric formulations of [Ridolfi et al. \(2010\)](#), the major element compositions of  
414 amphibole can be used to calculate the  $\text{H}_2\text{O}_{\text{melt}}$ , T, and P conditions of the melts that formed  
415 amphibole-bearing calc-alkaline igneous rocks. The geochemical characteristics of  
416 amphiboles from the Jinling HMDs can be used to estimate these conditions. Jinling HMD  
417 amphiboles have Al# ( $= \text{Al}^{\text{VI}}/\text{Al}_T$ ) values of  $<0.21$ , indicative of a magmatic origin (Table S3;  
418 [Ridolfi et al. 2010](#)). The physicochemical parameters calculated from the compositions of

419 different domains of amphibole phenocrysts from Group-I HMDs are consistent with those  
420 from Group-II HMDs (Fig. 12), indicating a close genetic relationship between the two  
421 groups of rocks and suggesting that they formed under similar physicochemical conditions but  
422 underwent different magmatic processes. Cores of amphibole phenocrysts from Group-I and  
423 Group-II HMDs yield the highest crystallization temperatures (850–944 and 895–941 °C,  
424 respectively) and pressures (135–277 and 192–327 MPa, respectively) (Table S3; Fig. 12a).  
425 The rims of amphibole phenocrysts (as well as matrix amphiboles) from the two groups of  
426 HMDs show considerably lower crystallization temperatures (723–802 and 720–756 °C,  
427 respectively) and pressures (30–90 and 37–58 MPa, respectively) than the cores (Table S3;  
428 Fig. 12a).

429 The calculated crystallization pressures can be used to estimate the depths of magma  
430 chamber during magma evolution of the Jinling HMDs. Given an average density of the upper  
431 crust of  $\rho = 2.7 \text{ g/cm}^3$ , the estimated crystallization depths for the cores of amphibole  
432 phenocrysts from Group-I and Group-II HMDs are 5.1–10.5 and 7.2–12.3 km, respectively,  
433 whereas those for the rims of amphibole phenocrysts (including matrix amphiboles) are 1.1–  
434 3.4 and 1.4–2.2 km, respectively (Table S3). Estimated T–P conditions for the mantles of  
435 amphibole phenocrysts in Group-II HMDs are 857–895 °C and 143–195 MPa, respectively,  
436 corresponding to a depth of 5.4–7.4 km, intermediate between the estimated depths for cores  
437 and rims. These results suggest that the Jinling HMDs underwent multi-stage evolution in  
438 crustal magma chambers located at different depths.

439 The oxygen fugacity and water contents of magma are also key controls on the magma  
440 evolution paths and compositional variations of igneous rocks (e.g., [Ridolfi et al. 2008, 2010](#)).  
441 The oxygen fugacity of host magma can be estimated from amphibole compositions by  
442 applying the formula proposed by [Ridolfi et al. \(2010\)](#). The calculated  $f\text{O}_2$  of melts in  
443 equilibrium with cores of amphibole phenocrysts are relatively high (Fig. 12b;  $\Delta\text{NNO} + 0.5$  to

444  $\Delta\text{NNO} + 1.7$  for Group-I HMDs and  $\Delta\text{NNO} + 0.7$  to  $\Delta\text{NNO} + 1.7$  for Group-II HMDs), and  
445 amphibole phenocryst rims and matrix amphiboles yield even higher  $f\text{O}_2$  of equilibrium melts  
446 (Fig. 12b;  $\Delta\text{NNO} + 2.1$  to  $\Delta\text{NNO} + 3.0$  for Group-I HMDs and  $\Delta\text{NNO} + 2.3$  to  $\Delta\text{NNO} + 3.0$   
447 for Group-II HMDs). These results reveal that the Jinling HMDs were formed in a relatively  
448 oxidizing environment and that the oxygen fugacity changed during the operation of  
449 magmatic processes at different crustal depths. Thus, both fractional crystallization and  
450 crustal contamination processes played important roles in the formation of the two groups of  
451 HMDs. Furthermore, King et al. (2000) proposed that the  $\text{Fe}^{3+}/(\text{Fe}^{3+} + \text{Fe}^{2+})$  ratios of  
452 amphibole are a reliable indicator of (i.e., strongly positively related to) the oxygen fugacity  
453 of the host magma. The different domains of amphibole phenocrysts from the two groups of  
454 HMDs show different evolutionary trends in a diagram of  $\text{Fe}^{3+}/(\text{Fe}^{3+} + \text{Fe}^{2+})$  versus  $\Delta\text{NNO}$   
455 (Fig. 12d). The rims of amphibole phenocrysts and matrix amphiboles in the two groups of  
456 HMDs show substantially higher  $\text{Fe}^{3+}/(\text{Fe}^{3+} + \text{Fe}^{2+})$  ratios and  $\Delta\text{NNO}$  values relative to cores,  
457 with mantles having intermediate values between them (Fig. 12d).

458 Water contents of host magma can be calculated from amphibole compositions on the  
459 basis of the formula proposed by Ridolfi et al. (2010). The estimated water contents of melts  
460 in equilibrium with cores of amphibole phenocrysts from Group-I and Group-II HMDs are  
461 higher (Fig. 12c; 3.6–4.5 and 3.6–5.3 wt%, respectively) than those of melts in equilibrium  
462 with rims of amphibole phenocrysts and matrix amphiboles (Fig. 12c; 2.9–3.9 and 2.3–3.2  
463 wt%, respectively). These results imply that the parental magmas of the Jinling HMDs  
464 contained abundant water and that their mantle sources had been metasomatized by aqueous  
465 fluids. It is expected that the water content of a melt should increase with increasing  $\text{SiO}_2$   
466 content during magmatic evolution because  $\text{H}_2\text{O}$  behaves similarly to incompatible elements  
467 (e.g., Ridolfi et al. 2010). However, during magma ascent, the solubility of water in silicate  
468 melts decreases with decreasing pressure (Holtz et al. 1995). Fluid exsolution is a common

469 phenomenon during late-stage magmatic evolution, and the escape of fluids will reduce water  
470 contents of host magma. The evolved magmas in equilibrium with the rims of amphibole  
471 phenocrysts and matrix amphiboles in the two groups of HMDs have lower water contents  
472 compared with the near-primary magmas in equilibrium with cores (Fig. 12c), which suggests  
473 that magmatic fluids might have escaped from shallow chambers during late-stage magmatic  
474 evolution.

#### 475 **Magma sources of the Jinling high-Mg dioritic rocks**

476 Cenozoic HMAs in arc settings can be generated through different mechanisms and  
477 have been widely used to characterize slab–mantle interactions at convergent plate margins  
478 (e.g., [Kay 1978](#); [Rogers et al. 1985](#); [Martin 1986](#); [Defant and Drummond 1990](#); [Yogodzinski](#)  
479 [et al. 1994, 1995](#); [Tatsumi 2001, 2008](#); [Martin et al. 2005](#); [Wang et al. 2020a](#)). In general,  
480 Cenozoic HMAs in oceanic subduction zones show depleted or slightly enriched Sr-Nd-Pb-Hf  
481 isotopic compositions in all cases (e.g., [Defant and Drummond 1990](#); [Yogodzinski et al. 1994,](#)  
482 [1995](#); [Tatsumi 2001, 2008](#); [Wang et al. 2020a](#); [Xu et al. 2020](#)). In contrast, the Jinling HMDs  
483 display strongly enriched radiogenic isotopic compositions, suggesting that their mantle  
484 sources were distinct from those of Cenozoic HMAs in arc settings and that these rocks were  
485 derived from partial melting of ancient sub-continental lithospheric mantle rather than  
486 asthenospheric mantle (e.g., [Yang et al. 2006, 2012a, 2012b](#); [Zhong and Huang 2012](#); [Jin et al.](#)  
487 [2015](#); [Lan et al. 2019](#); [Gao et al. 2021](#); [Zhang et al. 2021](#)).

488 Despite the progress made by previous studies of the petrology and geochemistry of  
489 Mesozoic mafic–intermediate igneous rocks from the eastern NCC, uncertainty remains  
490 regarding the nature of their mantle sources, especially the nature and origin of metasomatic  
491 agents in the sub-continental lithospheric mantle of the eastern NCC (e.g., [Gao et al. 2004](#); [Xu](#)  
492 [et al. 2004](#); [Yang et al. 2006, 2012b](#); [Huang et al. 2012](#); [Zhong and Huang 2012](#); [Jin et al.](#)  
493 [2015](#); [Lan et al. 2019](#); [Sun et al. 2019](#); [Gao et al. 2021](#); [Zhang et al. 2021](#)). These mafic–

494 intermediate igneous rocks have been interpreted as being derived from partial melting of  
495 ancient sub-continental lithospheric mantle of the eastern NCC with minor but variable  
496 contributions from asthenospheric mantle (e.g., Xu 2001; Xu et al. 2004; Zhong and Huang  
497 2012), or partial melting of enriched lithospheric mantle that had been metasomatized by  
498 felsic melts originating from subducted or delaminated continental crust (e.g., Gao et al. 2004;  
499 Yang et al. 2006, 2012a, 2012b; Jin et al. 2015; Lan et al. 2019; Gao et al. 2021). A recent  
500 geochemical study has suggested that Mesozoic sub-continental lithospheric mantle of the  
501 eastern NCC that had been metasomatized by aqueous fluids derived from subducted  
502 Paleo-Pacific oceanic crust during the Early–Middle Jurassic effectively preserved its original  
503 radiogenic Sr-Nd-Hf isotopic compositions (e.g., Niu 2005; Wang et al. 2020b).

504 Group-I HMDs show enriched Sr-Nd isotopic compositions resembling those of ancient  
505 sub-continental lithospheric mantle of the eastern NCC (Fig. 6c; e.g., Zhang and Yang 2007;  
506 Yang et al. 2009), indicating that they might have been derived from partial melting of ancient  
507 sub-continental lithospheric mantle of the eastern NCC but did not require previous  
508 metasomatism by felsic melts derived from subducted or delaminated continental crust  
509 materials. In addition, Group-I HMDs fall in or near the field of subduction zone fluid-related  
510 metasomatism in a diagram of  $(\text{Hf}/\text{Sm})_N$  versus  $(\text{Ta}/\text{La})_N$  (Fig. 10b), which is consistent with  
511 the occurrence of abundant hydrous minerals (amphibole and biotite) and arc-like  
512 geochemical characteristics (i.e., enrichment in LILEs and LREEs and depletion in HFSEs  
513 and HREEs). In contrast, Group-II HMDs show slightly more enriched whole-rock Sr–Nd and  
514 zircon Lu-Hf isotopic compositions in comparison with Group-I HMDs (Fig. 6b and c),  
515 indicating the involvement of more ancient continental crustal components in the genesis of  
516 the Group-II HMDs. TTG gneisses of the Taishan Group are the favored candidate for the  
517 crustal materials involved during magma emplacement on account of their relatively high  
518  $(^{87}\text{Sr}/^{86}\text{Sr})_i$  and low  $\epsilon_{\text{Nd}}(t)$  values (e.g., Jahn et al. 1988; Peng et al. 2013; Chen et al. 2020).

519 Nevertheless, a possible role of asthenospheric mantle- and/or oceanic crust-derived melts in  
520 the generation of the Jinling HMDs cannot be completely precluded, given their distinctly  
521 higher  $\epsilon_{\text{Hf}}(t)$  values relative to ancient sub-continental lithospheric mantle in the eastern NCC  
522 (Fig. 6b; e.g., [Zhong and Huang 2012](#); [Zhang et al. 2021](#)).

### 523 **Genetic mechanisms of formation of the Jinling high-Mg dioritic rocks**

524 The genetic mechanisms and geodynamics of Pre-Cenozoic HMAs/HMDs in  
525 continental interiors are hotly debated, in contrast to the more established understanding of  
526 Cenozoic HMAs in arc settings. For instance, at least four genetic models have been proposed  
527 to explain the generation of Mesozoic HMAs/HMDs in the eastern NCC, i.e., partial melts of  
528 delaminated continental lower crust interacting with mantle peridotites (e.g., [Gao et al. 2004](#);  
529 [Yang et al. 2006, 2012a](#); [Jin et al. 2015](#)), partial melting of enriched lithospheric mantle  
530 metasomatized by subducted or delaminated continental crust (e.g., [Yang et al. 2012b](#); [Lan et](#)  
531 [al. 2019](#); [Gao et al. 2021](#)), magma mixing between crustal- and mantle-derived melts (e.g.,  
532 [Chen et al. 2013](#)) and assimilation of previously emplaced mantle peridotite by crust-derived  
533 melts at crustal depths (e.g., [Qian and Hermann 2010](#)).

534 The two groups of Jinling HMDs display similar crystallization ages (Fig. 6a),  
535 indicating a close petrogenetic relationship between them in the same tectonic setting.  
536 Group-I HMDs were likely derived from partial melting of ancient sub-continental  
537 lithosphere mantle that had been metasomatized by aqueous fluids from subducted oceanic  
538 crust but without the involvement of metasomatism or the interaction of felsic melts derived  
539 from subducted or delaminated continental crust. The cores of amphibole phenocryst from  
540 Group-I and Group-II HMDs plot predominantly in the mantle source field (Fig. 11), which  
541 further contradicts the model of partial melting of continental lower crust interacting with or  
542 assimilating mantle peridotites at mantle and/or crustal depths (e.g., [Gao et al. 2004](#); [Yang et](#)  
543 [al. 2006, 2012a](#); [Qian and Hermann 2010](#); [Jin et al. 2015](#)). Furthermore, Group-I HMDs show

544 a narrow range of Sr–Nd isotopic compositions that are similar to those of sub-continental  
545 lithospheric mantle of the eastern NCC (Fig. 6c), suggesting that continental crust-derived  
546 melts were only negligibly involved in the petrogenesis of these rocks. However, the  
547 sub-continental lithospheric mantle of the eastern NCC consists predominantly of harzburgite  
548 and dunite (e.g., [Menzies et al. 1993](#); [Niu 2005](#)), which are refractory and hard to melt under  
549 normal mantle P–T–H<sub>2</sub>O conditions. It is noted that melts in equilibrium with the cores of  
550 amphibole phenocrysts from the Jinling HMDs have high H<sub>2</sub>O contents (Fig. 12c), which  
551 suggests that the mantle source had been metasomatized by aqueous fluids that were probably  
552 derived from subducted Paleo-Pacific oceanic crust. In addition, experiments have shown that  
553 sanukitic HMAs may represent near-primary magmas in equilibrium with upper-mantle  
554 peridotites at T = 1050–1150 °C and P = 10–15 kbar and under H<sub>2</sub>O-rich conditions (e.g.,  
555 [Tatsumi and Ishizaka 1982](#); [Kelemen 1995](#); [Tatsumi 2008](#)), and partial melting of peridotites  
556 under H<sub>2</sub>O-rich conditions at uppermost-mantle pressures can produce high-Mg andesitic  
557 rather than basaltic melts (e.g., [Kelemen 1995](#); [Hirose 1997](#); [Tatsumi 2001, 2008](#)). Group-I  
558 HMDs of the Jinling complex show comparable compositions of major elements with the  
559 experimental HMAs melts (Fig. 9a and b; e.g., [Hirose 1997](#)). Collectively, the Group-I HMDs  
560 of the Jinling complex were derived from partial melting of sub-continental lithospheric  
561 mantle of the eastern NCC that had been metasomatized by aqueous fluids derived from the  
562 subducted Paleo-Pacific oceanic crust at relatively high-T, low-P, and H<sub>2</sub>O-rich conditions  
563 (Fig. 13).

564 In continental interiors, crustal materials might be involved in the formation of mafic–  
565 intermediate igneous rocks through source mixing or the process of assimilation and  
566 fractional crystallization. The temporal and spatial relationships between Group-I and  
567 Group-II HMDs suggest that they likely shared a common mantle source, which is further  
568 supported by their overall similar Sr–Nd–Pb–Hf isotopic compositions and similar



569 physicochemical conditions of primary melts (Fig. 6b and c; e.g., Yang et al. 2006, 2012a,  
570 2012b; Zhong and Huang 2012; Jin et al. 2015; Lan et al. 2019; Zhang et al. 2021). However,  
571 the differences in petrography, geochemistry, and radiogenic isotopes between the two groups  
572 of HMDs (Figs. 2–6) suggest that they likely underwent different magmatic processes during  
573 their evolution in crustal levels. The compositional variation of Group-I HMDs was controlled  
574 mainly by fractional crystallization, with only minor involvement of crustal materials that  
575 mostly occurred in the latest stage of magma evolution (Fig. 13), while Group-II HMDs  
576 underwent far more complex and intensive magmatic processes, with more intensive  
577 fractional crystallization of mafic minerals and more extensive involvement of crustal  
578 materials. Furthermore, TTG gneisses of the Taishan Group in the Luxi area have much  
579 higher Sr/Y, (La/Yb)<sub>N</sub>, and (<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> ratios and larger negative ε<sub>Nd</sub>(t) values than those of  
580 sub-continental lithospheric mantle of the eastern NCC (e.g., Jahn et al. 1988; Peng et al.  
581 2013; Chen et al. 2020) and are therefore the favored candidate for the contribution of crustal  
582 materials into Group-II HMDs during magma emplacement (Figs. 6c and 10c–d). Collectively,  
583 Group-II HMDs were also produced by partial melting of sub-continental lithospheric mantle  
584 of the eastern NCC that had been metasomatized by aqueous fluids from subducted  
585 Paleo-Pacific oceanic crust but underwent more extensive fractional crystallization of mafic  
586 minerals and greater involvement of crustal materials that were probably derived from TTG  
587 gneisses of the Taishan Group in the Luxi area (Fig. 13).

588

## 589 **IMPLICATIONS FOR THE EVOLUTION OF CONTINENTAL LITHOSPHERE**

### 590 **Petrogenesis of intra-plate HMAs**

591 The association of basalt–andesite–dacite–rhyolite is the most widely distributed  
592 Cenozoic arc igneous rock association and is generally considered to be derived from partial

593 melting of fluid-metasomatized peridotites in the mantle wedge and subsequent intra-crustal  
594 differentiation processes; i.e., crustal contamination and fractional crystallization (e.g., Wang  
595 et al. 2020a; Xu et al. 2020; Zheng et al. 2020). In addition to slab-derived fluids,  
596 metasomatism or interaction between slab-derived melts and mantle-wedge peridotites is also  
597 an important control on the source nature of Cenozoic arc magmatism (e.g., Wang et al. 2020a;  
598 Xu et al. 2020; Zheng et al. 2020). In subduction zones, basaltic oceanic crust, underlying  
599 peridotitic mantle, and overlying sediments can be effectively transported to sub-arc depths  
600 beneath arc volcanoes and are all potential source materials of arc igneous rocks. Cenozoic  
601 adakites, as a special type of arc magmatic rock, are generally produced by partial melting of  
602 metabasalts at high pressure within the stability fields of garnet and rutile but outside the  
603 stability field of plagioclase and are thus genetically associated with the subduction of young  
604 and warm oceanic lithosphere (e.g., Defant and Drummond 1990; Yogodzinski et al. 1995;  
605 Kemei et al. 2004; Martin et al. 2005; Tang and Wang 2010; Wang et al. 2020a). When  
606 adakitic melts are produced at the surface of a subducted slab, they may infiltrate and react  
607 with mantle-wedge peridotites during magma ascent. At low melt/rock ratios, adakitic melts  
608 would be completely consumed in metasomatic reactions with mantle wedge, producing  
609 mantle sources for sanukitic and/or bajaitic HMAs, whereas at high melt/rock ratios, they  
610 would become adakitic HMAs (e.g., Kay 1978; Yogodzinski et al. 1995; Rapp et al. 1999;  
611 Martin et al. 2005; Wang et al. 2020a; Xu et al. 2020). Accordingly, the study of the origin of  
612 Cenozoic HMAs in modern arc volcanoes is an important petrological topic with respect to  
613 young and warm oceanic subduction systems and has implications for the understanding of  
614 chemical geodynamics at convergent plate margins (e.g., Defant and Drummond 1990; Martin  
615 et al. 2005; Xu et al. 2020; Zheng et al. 2020).

616 Besides Cenozoic HMAs, numerous Pre-Cenozoic HMAs/HMDs are also reported in  
617 continental interiors, and their petrogenesis is keenly debated. Contrary to Cenozoic HMAs,

618 Pre-Cenozoic intra-plate HMAs/HMDs generally have strongly enriched radiogenic  
619 Sr-Nd-Pb-Hf isotopic compositions, indicating that they were produced mainly by partial  
620 melting of enriched lithospheric mantle rather than asthenospheric mantle (e.g., [Gao et al.](#)  
621 [2004](#); [Xu et al. 2004](#); [Yang et al. 2006, 2012a, 2012b](#); [Zhong and Huang 2012](#); [Jin et al. 2015](#);  
622 [Lan et al. 2019](#); [Sun et al. 2019](#); [Zhang et al. 2021](#)). In this study, the Group-I and Group-II  
623 HMDs of the Jinling complex correspond geochemically to sanukitic and adakitic HMAs,  
624 respectively. However, both the Group-I and Group-II HMDs were derived from partial  
625 melting of sub-continental lithospheric mantle of the eastern NCC metasomatized by aqueous  
626 fluids from the subducted Paleo-Pacific slab and underwent variable and extensive fractional  
627 crystallization and incorporation of crustal materials. Thus, the Jinling HMDs are not related  
628 to the subduction of young and warm oceanic crust nor to the growth of continental crust at  
629 convergent plate margins.

630         Although the Jinling HMDs are commonly considered to have been emplaced into the  
631 upper continental crust during the Mesozoic lithospheric destruction of the eastern NCC (e.g.,  
632 [Yang et al. 2012a, 2012b](#); [Zhong and Huang 2012](#); [Jin et al. 2015](#); [Lan et al. 2019](#); [Guo et al.](#)  
633 [2022](#)), their radiogenic isotopic compositions resemble those of sub-continental lithospheric  
634 mantle of the eastern NCC (Fig. 6c; e.g., [Zhang and Yang 2007](#); [Yang et al. 2009](#)). Since the  
635 Early-Middle Jurassic, the Paleo-Pacific plate has been subducted westward under the eastern  
636 Asian continental margin (e.g., [Maruyama et al. 1997](#); [Zheng et al. 2013](#)). The subducted  
637 Paleo-Pacific oceanic crust would have undergone metamorphic dehydration and/or partial  
638 melting below sub-arc depths, and the resultant aqueous fluids and hydrous melts would have  
639 further modified sub-continental lithospheric mantle of the eastern NCC, which would not  
640 only have formed metasomatized mantle sources with high oxygen fugacities and water  
641 contents but also significantly changed the rheological properties of the cratonic lithospheric  
642 mantle. Subsequently, the fluid-metasomatized lithospheric mantle of the eastern NCC would

643 have been heated and melted to form mafic-intermediate magmas as the asthenosphere  
644 upwelled in late Mesozoic (Fig. 13).

#### 645 **Implications for the evolution of continental lithosphere**

646 Cenozoic adakitic HMAs are formed by slab melting and are the products of the most  
647 common parental magmas in modern arc volcanoes. Subduction of young and warm oceanic  
648 lithosphere is fundamental to the slab-melting genesis of Cenozoic HMAs in modern  
649 subduction systems (e.g., [Kay 1978](#); [Rogers et al. 1985](#); [Defant and Drummond 1990](#);  
650 [Yogodzinski et al. 1994, 1995](#); [Tatsumi 2001, 2008](#); [Martin et al. 2005](#); [Wang et al. 2020a](#)).  
651 On average, subducting oceanic lithosphere was much younger and hotter during the Archean  
652 than for modern Earth, and slab melting should therefore have been a common phenomenon  
653 during the Archean (e.g., [Martin 1986](#)). Neoproterozoic sanukitoids of the southwestern Superior  
654 Province, Canada are a typical example of Pre-Cenozoic HMAs and are geochemically  
655 analogous to Cenozoic HMAs in the Japanese Setouchi belt, and it has been proposed that  
656 they are derived from partial melting of mantle peridotites that had been metasomatized by  
657 aqueous fluids or hydrous melts of subducting oceanic slab ([Shirey and Hanson 1984](#)).  
658 Furthermore, Neoproterozoic sanukitoids and low-silica adakitic rocks are widely considered to  
659 develop in oceanic subduction environments and might be diagnostic petrological records for  
660 the onset of plate subduction and Archean continental crustal growth. For example, numerous  
661 studies have argued that oceanic subduction was occurring during the Archean (e.g., [Shirey](#)  
662 [and Hanson 1984](#); [Polat et al. 2002](#); [Martin et al. 2005](#); [Hastie et al. 2015](#)) on the basis of  
663 geochemical similarities between Archean sanukitoids and Cenozoic HMAs. The key question  
664 associated with the timing of onset of plate subduction is whether Archean sanukitoids  
665 developed exclusively in oceanic subduction systems.

666 In oceanic subduction zones, the subduction of oceanic plate not only forms the  
667 mantle-wedge structure but also generates arc volcanic rocks at convergent plate margins (e.g.,

668 [Zheng et al. 2020](#)). Partial melting of subducting oceanic slabs and subsequent interaction  
669 between slab-derived melts and mantle-wedge peridotites can generate adakitic rocks and  
670 mantle sources for sanukitic and/or bajaitic HMAs (e.g., [Rogers et al. 1985](#); [Defant and](#)  
671 [Drummond 1990](#); [Yogodzinski et al. 1995](#); [Kemei et al. 2004](#); [Martin et al. 2005](#); [Wang et al.](#)  
672 [2020a](#); [Xu et al. 2020](#)). As mentioned above, Group-I HMDs of the Jinling complex belong to  
673 sanukitic rocks, and their primary magmas were derived from partial melting of  
674 sub-continental lithospheric mantle; while Group-II HMDs of the Jinling complex as adakitic  
675 rocks were formed through crustal contamination and fractional crystallization of the primary  
676 magmas of Group-I HMDs rather than by partial melting of the subducted oceanic slab. Thus,  
677 there is no requirement for a relationship between slab-derived melts and the formation of  
678 Pre-Cenozoic HMAs. If so, Neoproterozoic sanukitoids cannot be simply used to infer the  
679 operation of oceanic subduction. Furthermore, the Archean mantle was 200–300 °C hotter  
680 than modern mantle, and its derivative komatiitic magmas commonly contained several  
681 percent water (e.g., [Martin 1986](#); [Grove et al. 2004](#); [Sobolev et al. 2016](#)). Therefore, although  
682 we cannot completely rule out the possibility of slab dehydration in the Archean, Neoproterozoic  
683 TTG and sanukitoids could have been derived from partial melting of hydrous mantle  
684 peridotites at high temperatures in within-plate settings and might have experienced different  
685 magmatic processes in crustal chambers. Indeed, Pre-Cenozoic intra-plate HMAs and their  
686 intrusive equivalents appear to have been more common during the Archean than on modern  
687 Earth (e.g., [Shirey and Hanson 1984](#); [Martin 1986](#); [Polat et al. 2002](#); [Martin et al. 2005](#)).

688       However, Archean TTG, the primary rock type in Archean continental crust ([Condie](#)  
689 [2005](#)), is compositionally similar to Cenozoic adakites (e.g., [Defant and Drummond 1990](#);  
690 [Martin et al. 2005](#)). Some studies have compared the two rock types to establish the  
691 mechanisms of Archean continental crustal growth ([Condie 2005](#); [Martin et al. 2005](#)).  
692 Archean TTGs were produced primarily through partial melting of a basaltic source under

693 eclogite-facies conditions, which could have occurred either in subduction zones or at the  
694 base of thickened continental crust (e.g., [Defant and Drummond 1990](#); [Kay and Kay 1991](#);  
695 [Martin et al. 2005](#); [Wang et al. 2020a](#)). In fact, the composition of Archean TTGs varied over  
696 time. In general, >3.5 Ga TTGs have lower Mg, Cr, Ni, and Sr content than <3.0 Ga TTGs  
697 that have high Mg, Cr, Ni, and Sr contents ([Martin et al. 2005](#)), which is considered to be  
698 related to a change in oceanic lithosphere subduction; i.e., >3.5 Ga TTGs lack a mantle-wedge  
699 compositional contribution, whereas <3.0 Ga TTGs record such a contribution ([Martin et al.](#)  
700 [2005](#)). The Group-II HMDs of the Jinling complex have high contents of compatible elements  
701 (e.g., Ni = 55–57 ppm; Cr = 201–213 ppm) and MgO (2.90–3.08 wt%) with high Mg# (61.1–  
702 71.8) and slightly positive Sr anomalies (Table S1; Figs. 4a and 5b), which show similar  
703 geochemical characteristics to the <3.0 Ga TTGs (MgO = ~5 wt%; Mg# = ~65; Cr = ~200  
704 ppm; Ni = ~70 ppm; slightly positive Sr anomalies; e.g., [Martin and Moyen 2002](#); [Martin et al.](#)  
705 [2005](#)). The petrogenesis of the Group-II HMDs implies that subduction and melting of  
706 oceanic slabs were not required to generate Pre-Cenozoic adakitic rocks. Accordingly, the  
707 geodynamics of widespread Archean TTGs in ancient cratons should be reconsidered when  
708 investigating the mechanisms of Archean continental crustal growth and for comparisons with  
709 Cenozoic adakites in modern arc settings.

710

711

## CONCLUSIONS

712 The Early Cretaceous Jinling complex in the Luxi area of the eastern NCC can be  
713 subdivided into two groups of HMDs. Group-I HMDs, comprising gabbroic diorites and  
714 hornblende diorites, have geochemical features similar to those of Cenozoic sanukitic HMAs,  
715 whereas Group-II HMDs are monzonites and are geochemically classified as adakitic HMAs.

716 Group-I and Group-II HMDs of the Jinling complex share a common source of ancient

717 sub-continental lithospheric mantle of the eastern NCC that had been metasomatized by fluids  
718 from subducted Paleo-Pacific slab. However, the two groups of HMDs show highly distinct  
719 petrographic and whole-rock geochemical characteristics as a result of different magmatic  
720 processes that occurred in the crustal magma chambers.

721 The two groups of HMDs of the Jinling complex were not formed by interaction  
722 between slab-derived melts and mantle-wedge peridotites but were instead derived from  
723 partial melting of hydrous mantle peridotites in continental interior of the eastern NCC.

724 Pre-Cenozoic intra-plate HMAs/HMDs may have had a different role in the evolution of  
725 continental lithosphere compared with Cenozoic HMAs in arc settings. Neoproterozoic  
726 sanukitoids and TTGs were produced by either interaction between slab-derived melts and  
727 mantle-wedge peridotites or partial melting of hydrous mantle peridotites.

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729

#### ACKNOWLEDGMENTS

730 We appreciate L. Zhang for the assistance with whole-rock Sr-Nd isotope analysis. We  
731 are grateful to Prof. Chao Zhang and an anonymous reviewer, for their careful reviews and  
732 constructive comments, which considerably improved the manuscript. This work was  
733 financially supported by the National Natural Science Foundation of China (NSFC projects  
734 42021002, 41625007) and the Key Special Project for Introduced Talents Team of Southern  
735 Marine Science and Engineering Guangdong Laboratory (Guangzhou) (GML2019ZD0202).  
736 This is contribution No. IS-XXXX from GIGCAS.

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982

## FIGURE CAPTIONS

983 **Figure 1.** (a) Geological sketch of the North China Craton (NCC) showing the distribution of  
984 major tectonic units and the location of Shandong Province (SDP). (b) Geological  
985 map of SDP and the location of the study area (modified from [Huang et al. 2012](#)). (c)  
986 Geological map of the Jinling high-Mg dioritic complex showing the distribution of  
987 the two groups of high-Mg dioritic rocks (modified from [Zhong and Huang 2012](#)).  
988 Abbreviations: DGTL, Daxinganling–Taihangshan Gravity Lineament; TLFZ, Tan–  
989 Lu fault zone; YTC, Yangtze Craton; CYB, Cathaysia Block; XMOB, Xing–Meng  
990 Orogenic Belt; SOB, Sulu Orogenic Belt; DOB, Dabie Orogenic Belt.

991 **Figure 2.** Petrographic characteristics of samples from the Jinling high-Mg dioritic complex.  
992 (a–c) Gabbroic diorites containing orthopyroxene and clinopyroxene phenocrysts  
993 rimmed by amphiboles. (d) Hornblende diorites containing core–rim-zoned  
994 amphibole phenocrysts with magnetite inclusions in the rim. Both amphiboles and  
995 plagioclases are common minerals in the matrix of gabbroic diorites and hornblende  
996 diorites. (e–f) Monzonites containing core–mantle–rim-zoned amphibole and  
997 plagioclase phenocrysts. Mineral abbreviations: Opx, orthopyroxene; Cpx,  
998 clinopyroxene; Amp, amphibole; Bi, biotite; Pl, plagioclase; Mt, magnetite.

999 **Figure 3.** Diagrams of (a) total alkalis versus SiO<sub>2</sub> and (b) K<sub>2</sub>O versus SiO<sub>2</sub> for the Jinling  
1000 high-Mg dioritic rocks from the Luxi area. Literature data are from [Yang et al. \(2006,](#)  
1001 [2012b\)](#), [Jin et al. \(2015\)](#), [Lan et al. \(2019\)](#), [Gao et al. \(2021\)](#), and [Zhang et al. \(2021\)](#).  
1002 Abbreviations: PG, Peridotgabbro; FG, Foid gabbro; FMD, Foid monzodiorite; FMS,  
1003 Foid monzosyenite; MG, Monzogabbro; MD, Monzodiorite; GD, Gabbroic diorite;  
1004 QM, Quartz monzonite.

1005 **Figure 4.** (a–i) Harker diagrams of major and trace elements for Jinling high-Mg dioritic  
1006 rocks from the Luxi area. Literature data sources are the same as for Fig. 3. The

1007 fields of > 3.5 Ga TTG and < 3.0 Ga TTG are after Martin et al. (2005).

1008 **Figure 5.** (a) Chondrite-normalized REE diagram and (b) primitive-mantle-normalized  
1009 multi-element variation diagram for the Jinling high-Mg dioritic rocks from the Luxi  
1010 area. Normalization values are from McDonough and Sun (1995). Literature data  
1011 sources are the same as for Fig. 3.

1012 **Figure 6.** Histograms of (a) zircon U–Pb ages and (b) zircon  $\varepsilon_{\text{Hf}}(t)$  values and (c) diagram of  
1013 whole-rock  $\varepsilon_{\text{Nd}}(t)$  versus ( $^{87}\text{Sr}/^{86}\text{Sr}$ )<sub>i</sub> for the Jinling high-Mg dioritic rocks from the  
1014 Luxi area. The field for sub-continental lithospheric mantle (SCLM) of the North  
1015 China Craton is after Zhang and Yang (2007) and Yang et al. (2009); data for the  
1016 Taishan Group TTG gneisses are from Jahn et al. (1988), Peng et al. (2013), and  
1017 Chen et al. (2020). Literature data for the Jinling high-Mg dioritic rocks are from  
1018 Yang et al. (2006, 2012b), Zhong and Huang (2012), Jin et al. (2015), Lan et al.  
1019 (2019), Gao et al. (2021), and Zhang et al. (2021).

1020 **Figure 7.** (a–b) Back-scattered electron images of plagioclase in (a) Group-I and (b) Group-II  
1021 high-Mg dioritic rocks. (c) An–Ab–Or ternary diagram for plagioclase from the  
1022 Jinling high-Mg dioritic rocks of the Luxi area. Red circles indicate EPMA analytical  
1023 spots, and numbers adjacent to circles are An contents. Mineral abbreviations are the  
1024 same as for Fig. 2.

1025 **Figure 8.** (a–b) Back-scattered electron images of amphibole phenocrysts in (a) Group-I  
1026 (adapted from Guo et al. 2022) and (b) Group-II high-Mg dioritic rocks. (c–d)  
1027 Classification of amphiboles from the Jinling high-Mg dioritic rocks of the Luxi area.  
1028 Red circles indicate EPMA analytical spots, and numbers adjacent to circles are  
1029  $\text{Al}_2\text{O}_3$  contents. Major-element data for amphiboles from Group-I high-Mg dioritic  
1030 rocks are from Guo et al. (2022). Mineral abbreviations are the same as for Fig. 2.

1031 **Figure 9.** Diagrams of (a) Mg# versus SiO<sub>2</sub> (modified after [Rapp et al. 1999](#)), (b) TiO<sub>2</sub> versus  
1032 MgO/(MgO + FeO<sup>T</sup>) (modified after [Kemei et al. 2004](#)), (c) Sr/Y versus Y (modified  
1033 after [Defant and Drummond 1990](#)), and (d) Sr/Y versus Y (modified after [Kemei et al.](#)  
1034 [2004](#)) for the Jinling high-Mg dioritic rocks from the Luxi area. The experimental  
1035 melt compositions are from [Hirose \(1997\)](#). The plot regions of Sanukitic, Adakitic,  
1036 Bajaitic and Boninitic HMAs are from [Kemei et al. \(2004\)](#), and other literature data  
1037 sources are the same as for Fig. 3.

1038 **Figure 10.** Diagrams of (a) La/Sm versus La, (b) (Hf/Sm)<sub>N</sub> versus (Ta/La)<sub>N</sub> (modified after  
1039 [LaFlèche et al. 1998](#)), (c) whole-rock (<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> versus (La/Yb)<sub>N</sub>, and (d)  
1040 whole-rock ε<sub>Nd</sub>(t) versus Sr/Y for the Jinling high-Mg dioritic rocks from the Luxi  
1041 area. Data for the Taishan Group TTG gneisses are from [Jahn et al. \(1988\)](#), [Peng et al.](#)  
1042 [\(2013\)](#), and [Chen et al. \(2020\)](#). Other literature data sources are the same as for Fig. 3.  
1043 Abbreviations: DM, depleted mantle; N-MORB, normal mid-oceanic ridge basalt;  
1044 OIB, oceanic island basalt.

1045 **Figure 11.** Diagrams of (a) Si–Ti–Al ternary (modified after [Jiang and An 1984](#)), (b) TiO<sub>2</sub>  
1046 versus Al<sub>2</sub>O<sub>3</sub> (modified after [Sun et al. 2019](#)), and (c) (Na + K) versus Al<sup>IV</sup> (modified  
1047 after [Jiang and An 1984](#)) for amphiboles from the Jinling high-Mg dioritic rocks of  
1048 the Luxi area. Major-element data for amphiboles from the Group-I high-Mg dioritic  
1049 rocks are from [Guo et al. \(2022\)](#).

1050 **Figure 12.** Diagrams of (a) P versus T, (b) log(fO<sub>2</sub>) versus T, (c) T versus H<sub>2</sub>O<sub>melt</sub>, and (d)  
1051 Fe<sup>3+</sup>/(Fe<sup>3+</sup> + Fe<sup>2+</sup>) versus ΔNNO for amphiboles from the Jinling high-Mg dioritic  
1052 rocks of the Luxi area (modified after [Ridolfi et al. 2010](#)). Fields for the cores of  
1053 amphibole phenocrysts (I) and rims of amphibole phenocrysts and matrix amphiboles  
1054 (II) of high-Mg dioritic rocks from the Han–Xing district are from [Zhang et al.](#)  
1055 [\(2015\)](#).

1056 **Figure 13.** Schematic model for the petrogenetic relationship between the Group-I and  
1057 Group-II high-Mg dioritic rocks of the Jinling complex in the Luxi area (modified  
1058 after Guo et al. 2022). Abbreviations: SCLM, sub-continental lithospheric mantle of  
1059 the North China Craton; Amp, amphibole; Pl, plagioclase.  
1060

Figure 1

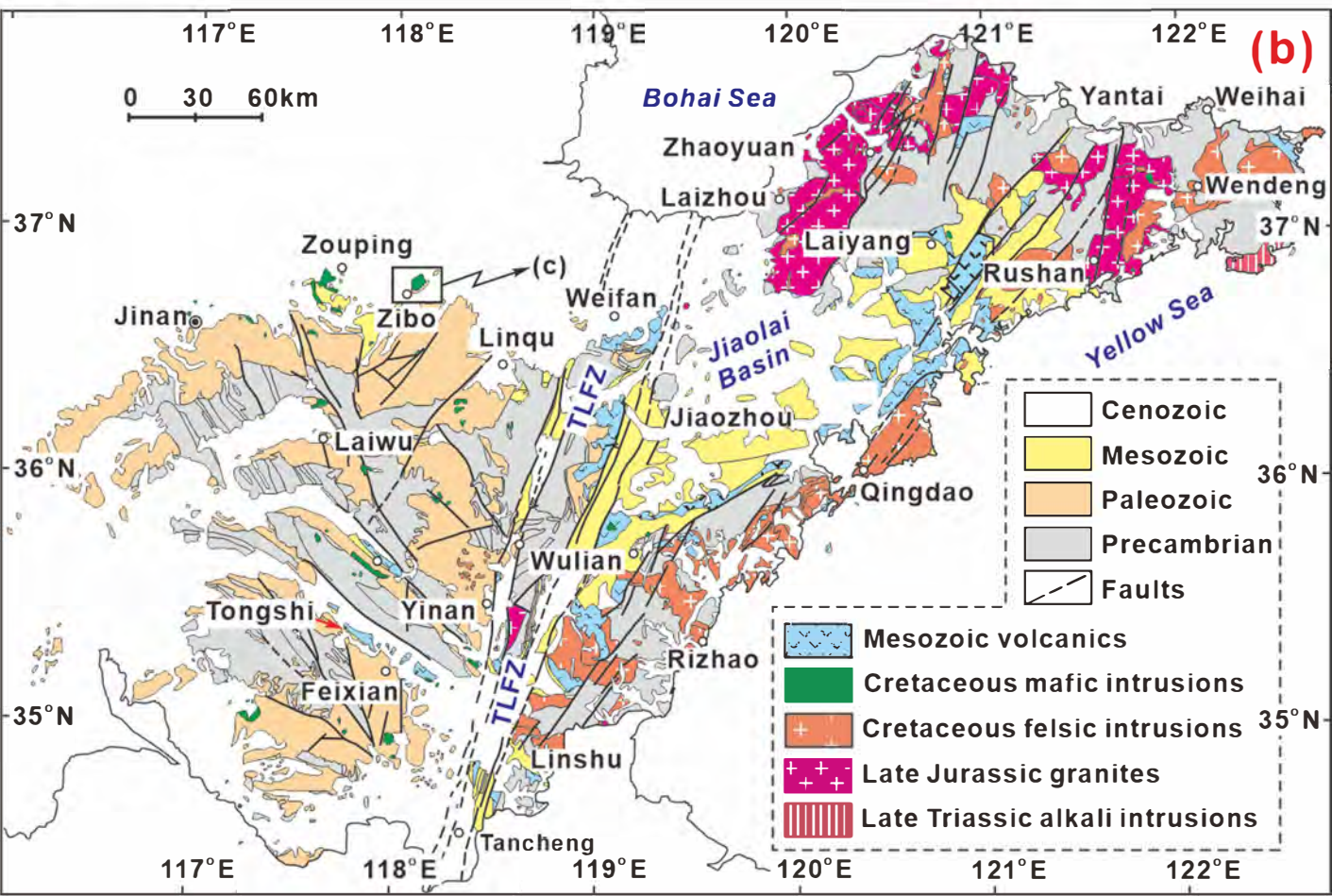
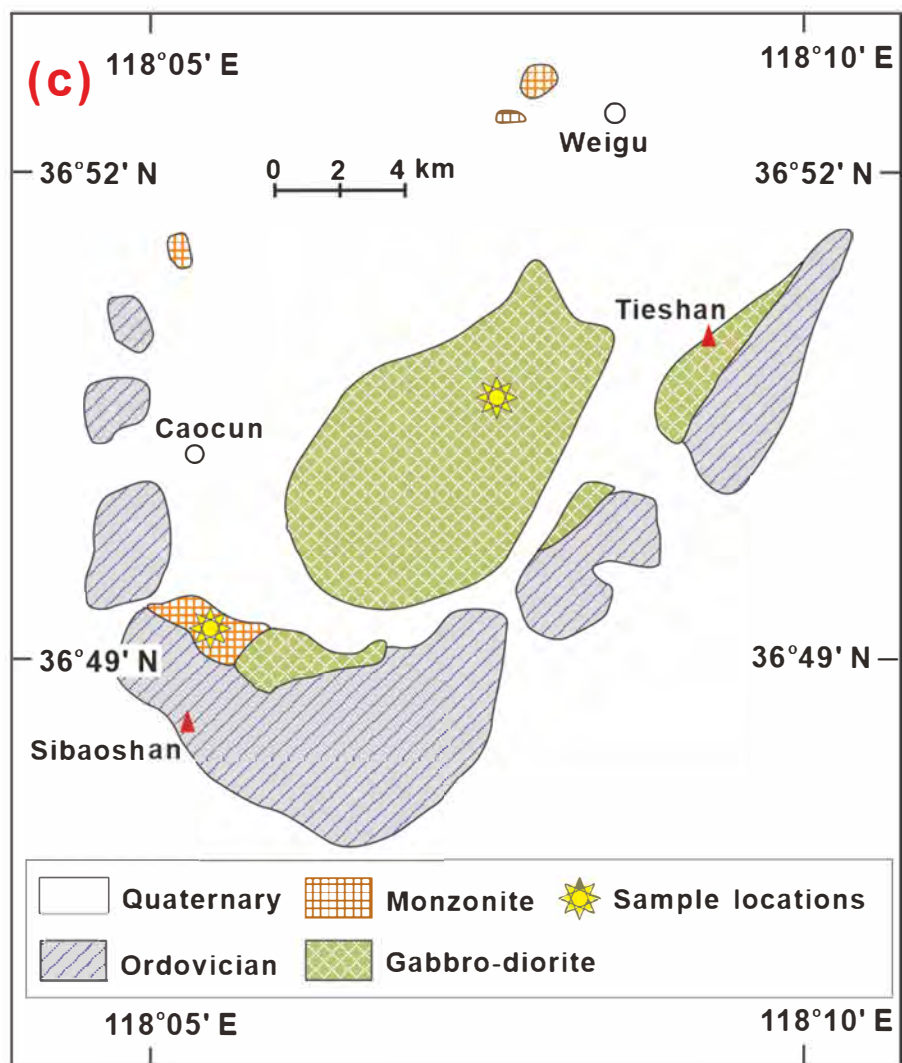
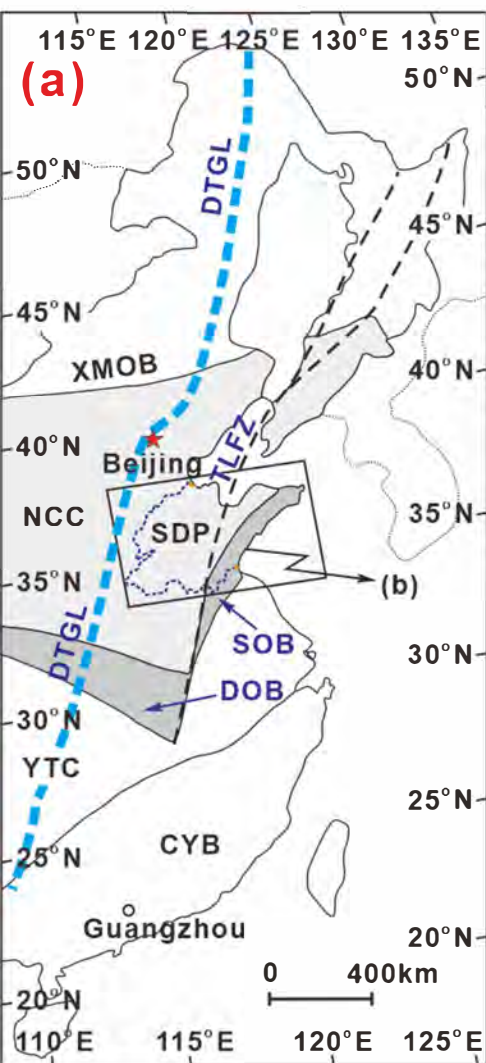


Figure 2

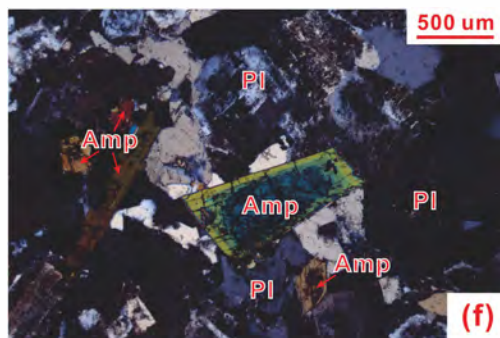
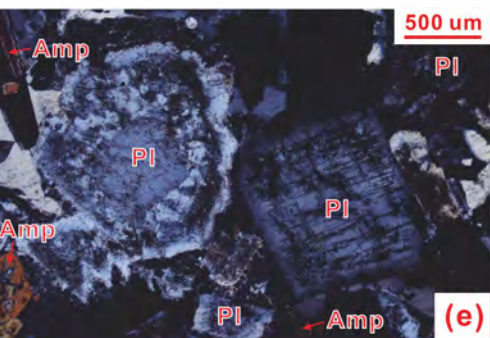
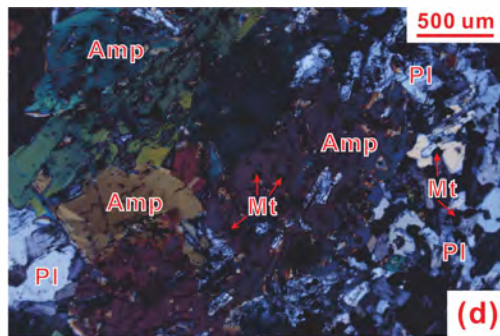
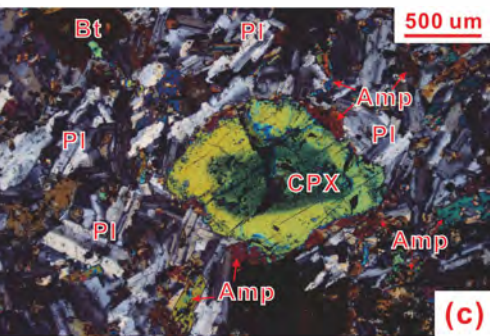
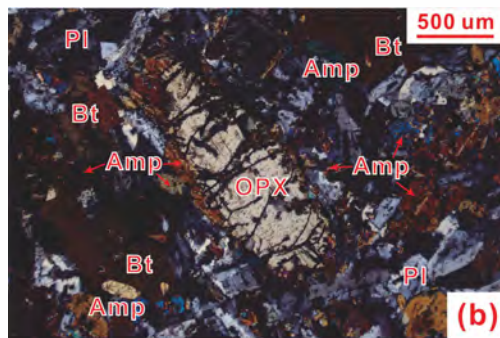
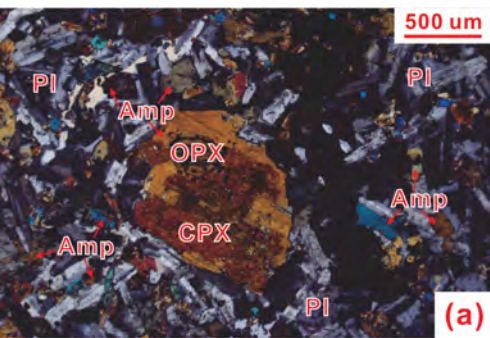




Figure 3

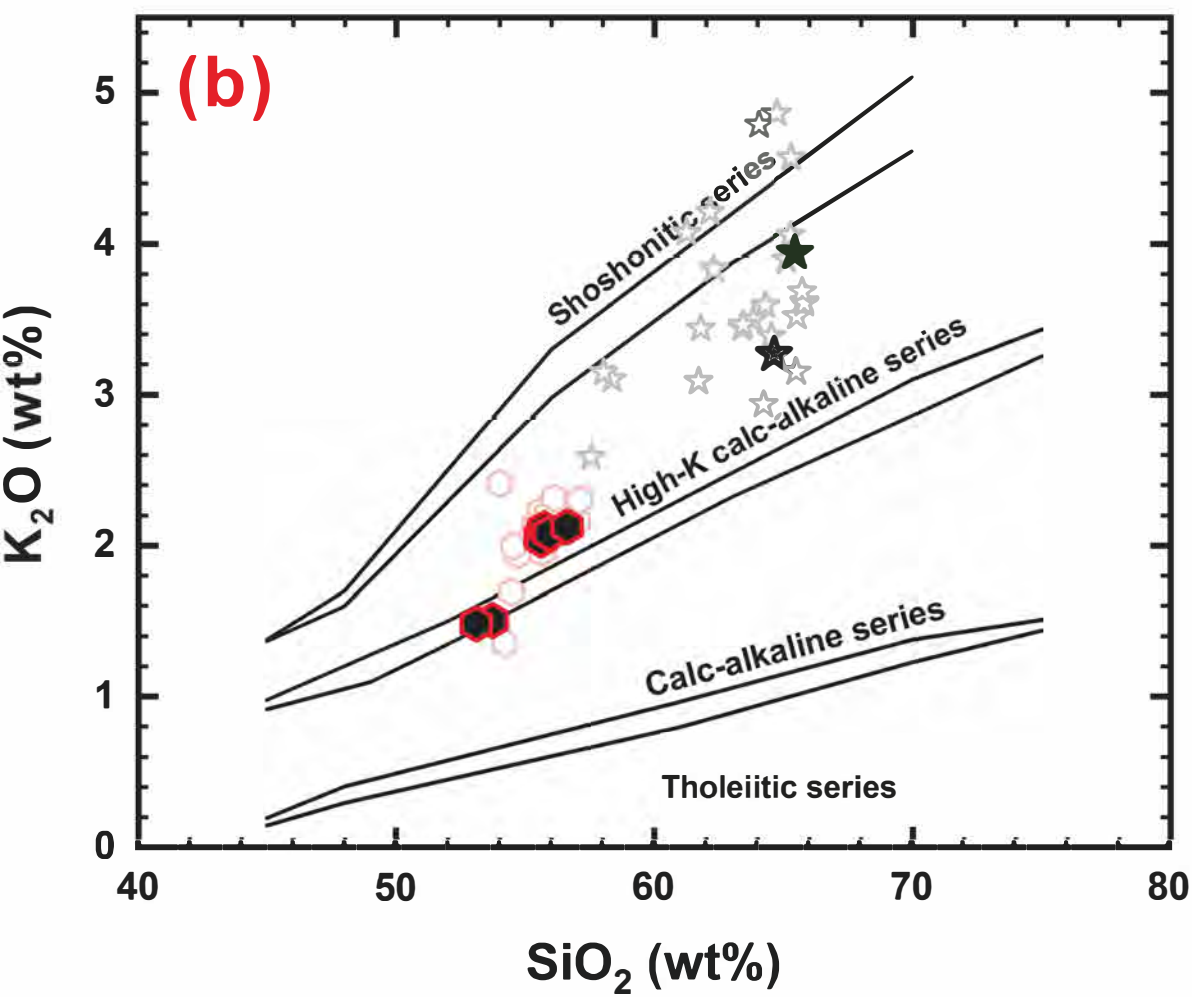
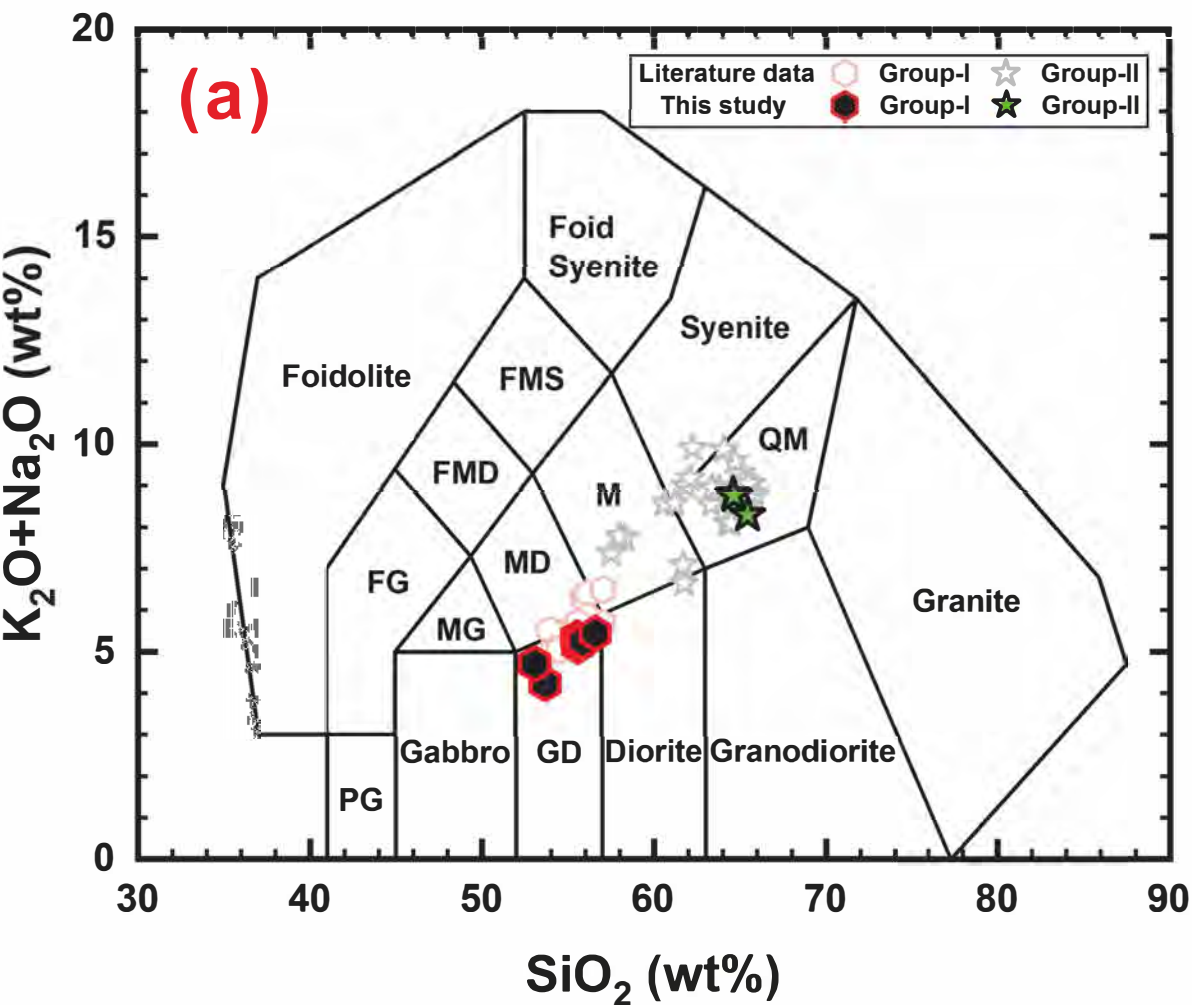


Figure 4

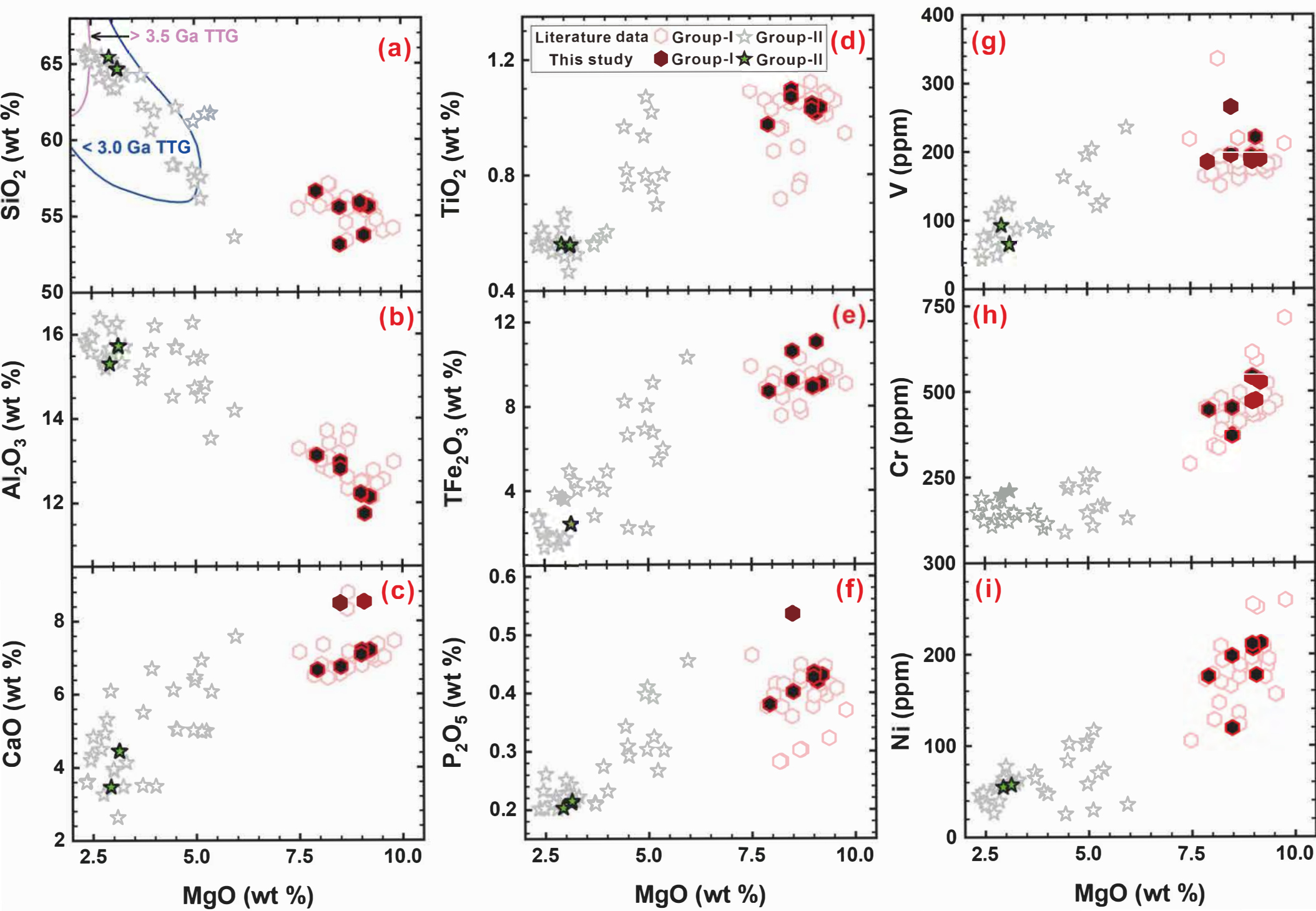


Figure 5

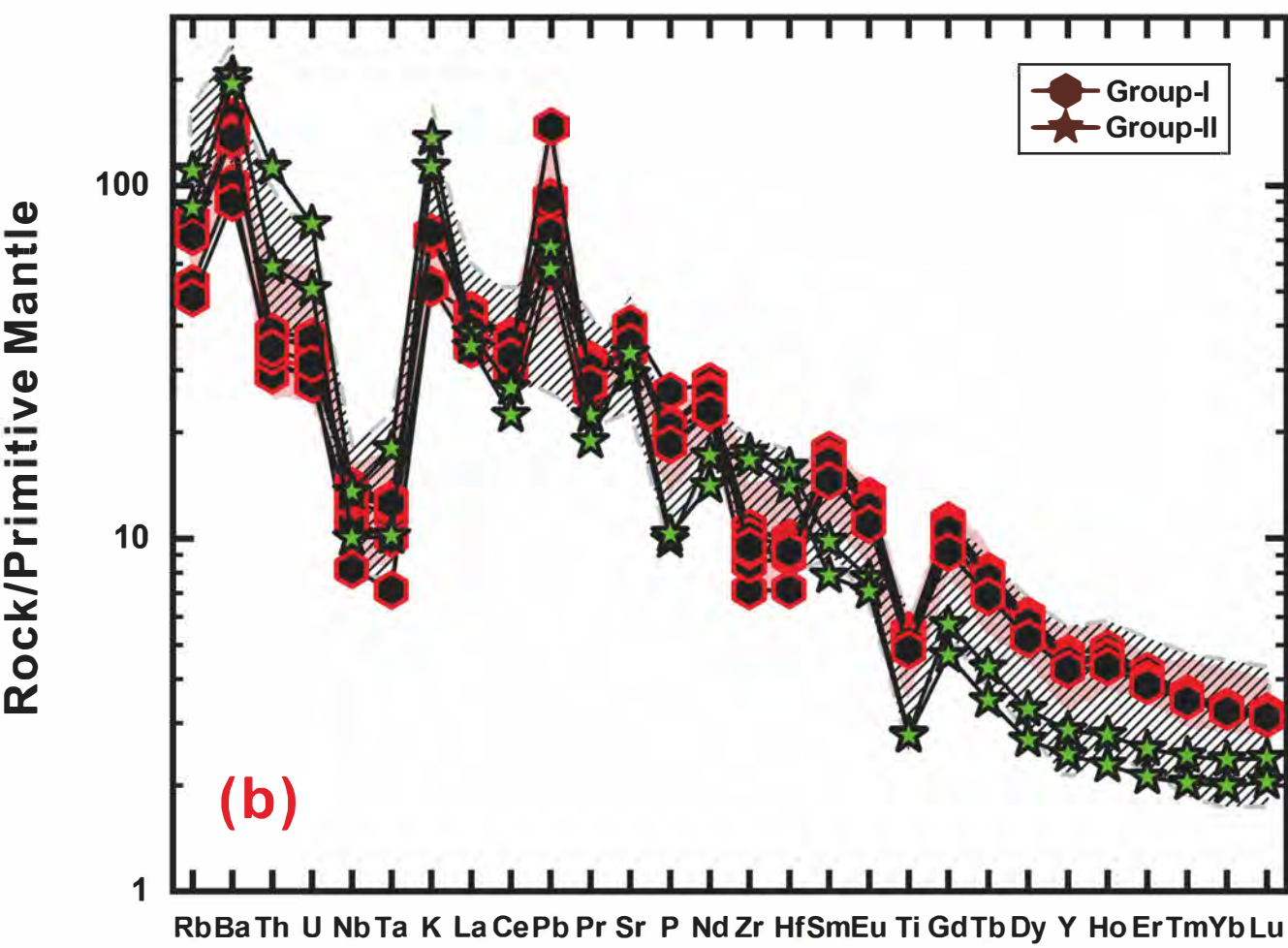
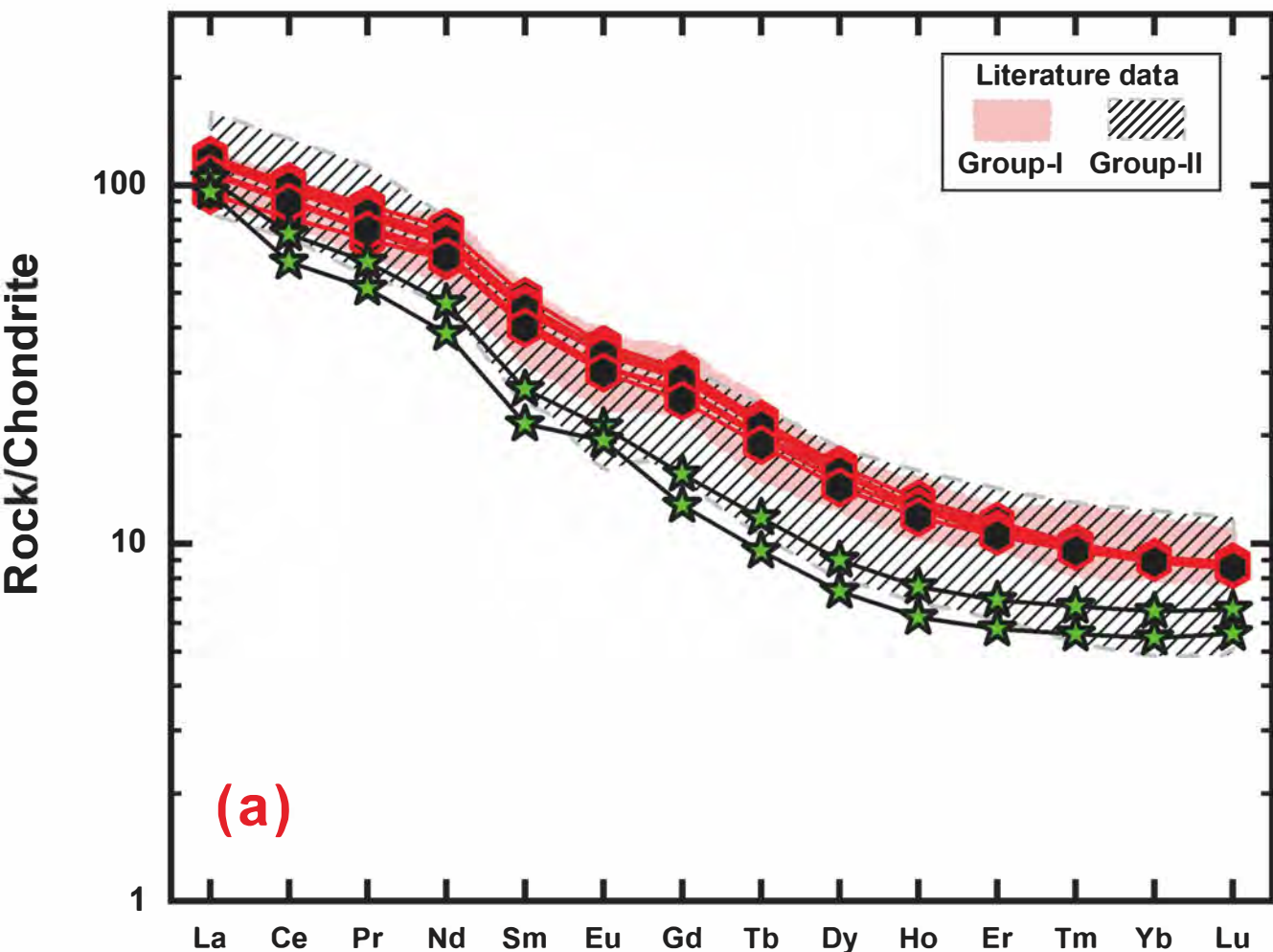


Figure 6

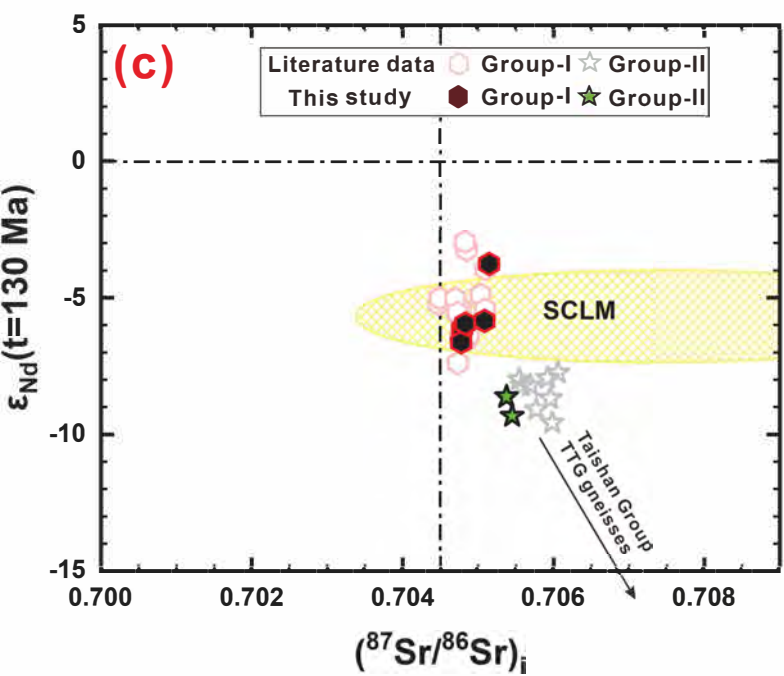
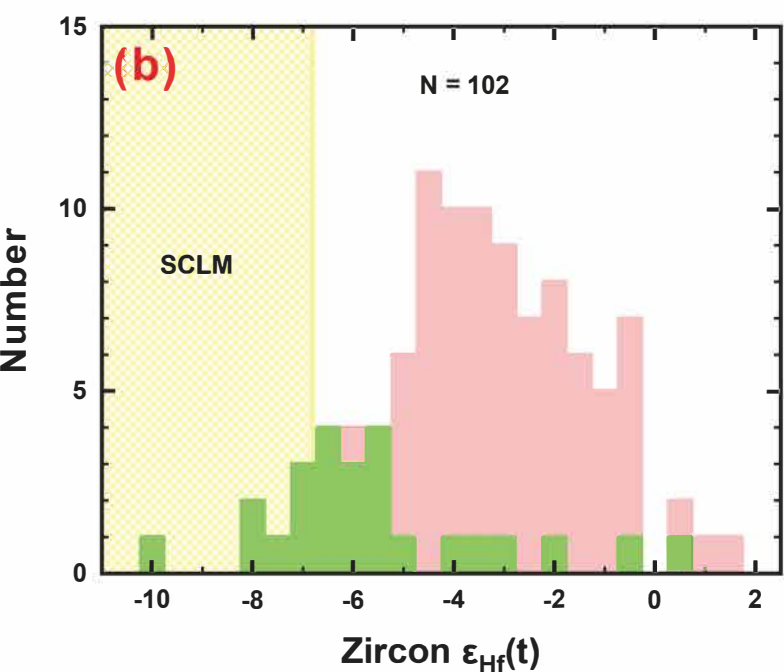
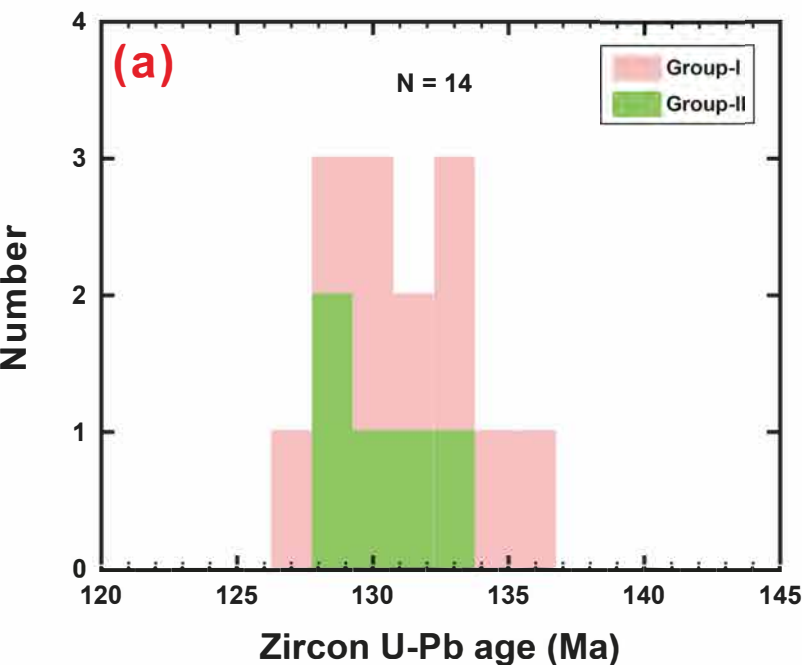


Figure 7

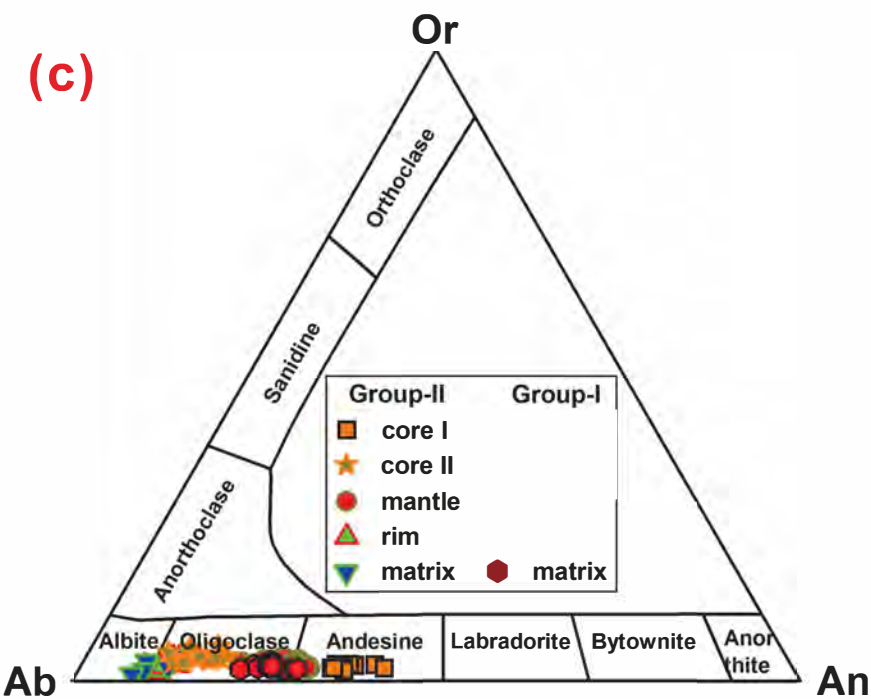
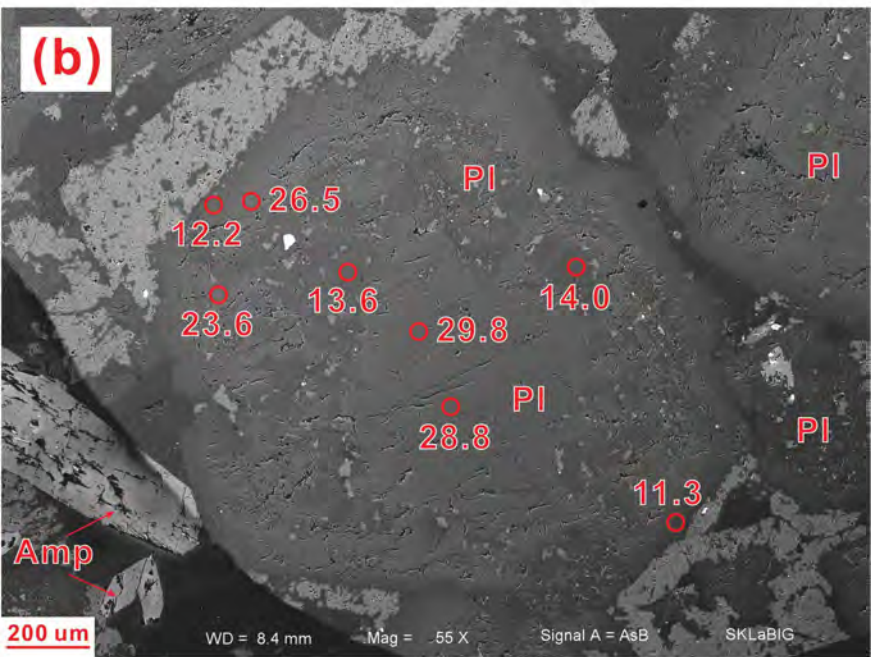
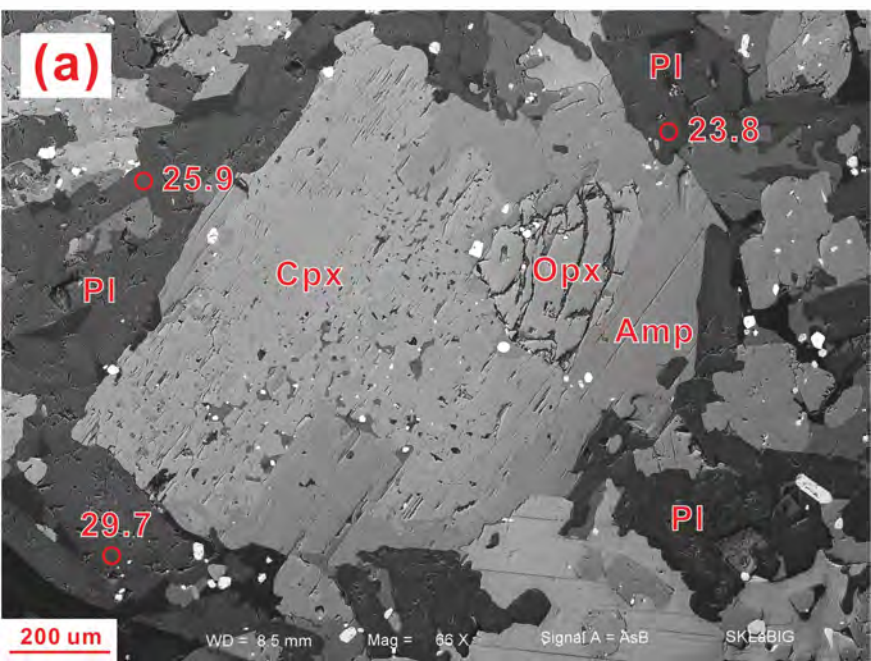


Figure 8

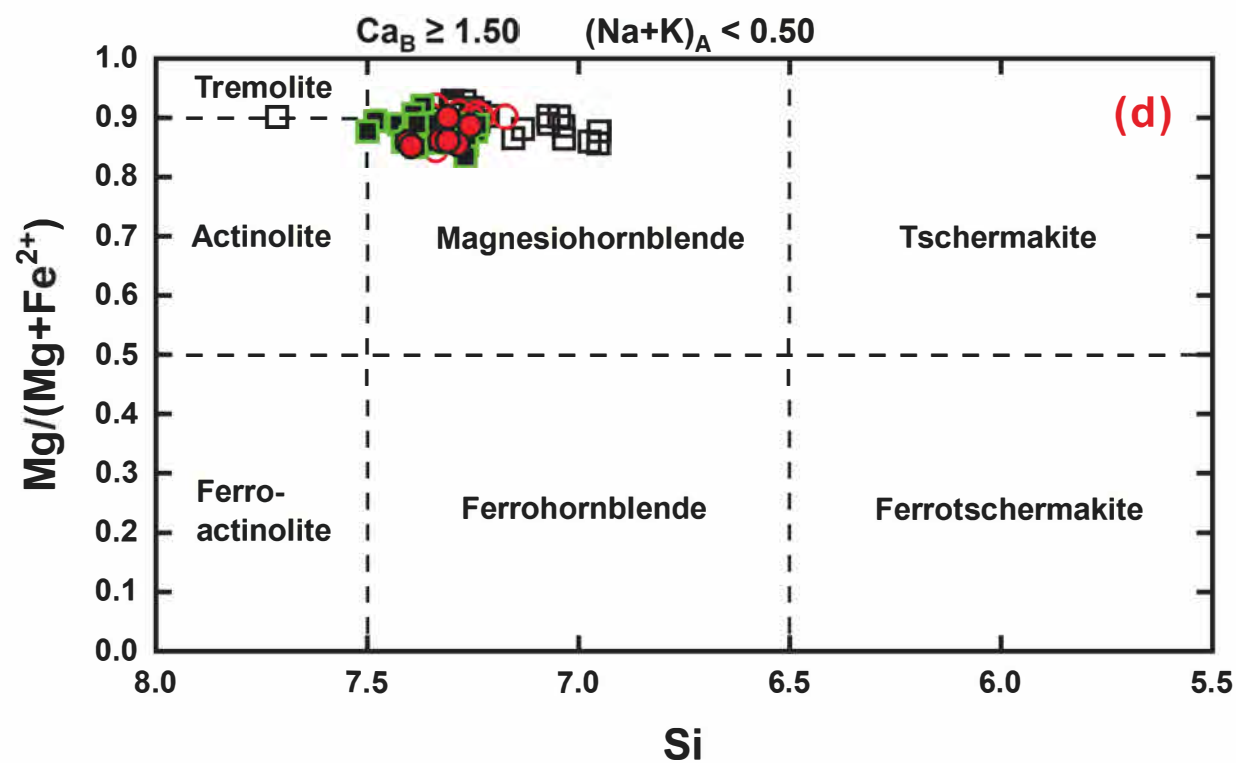
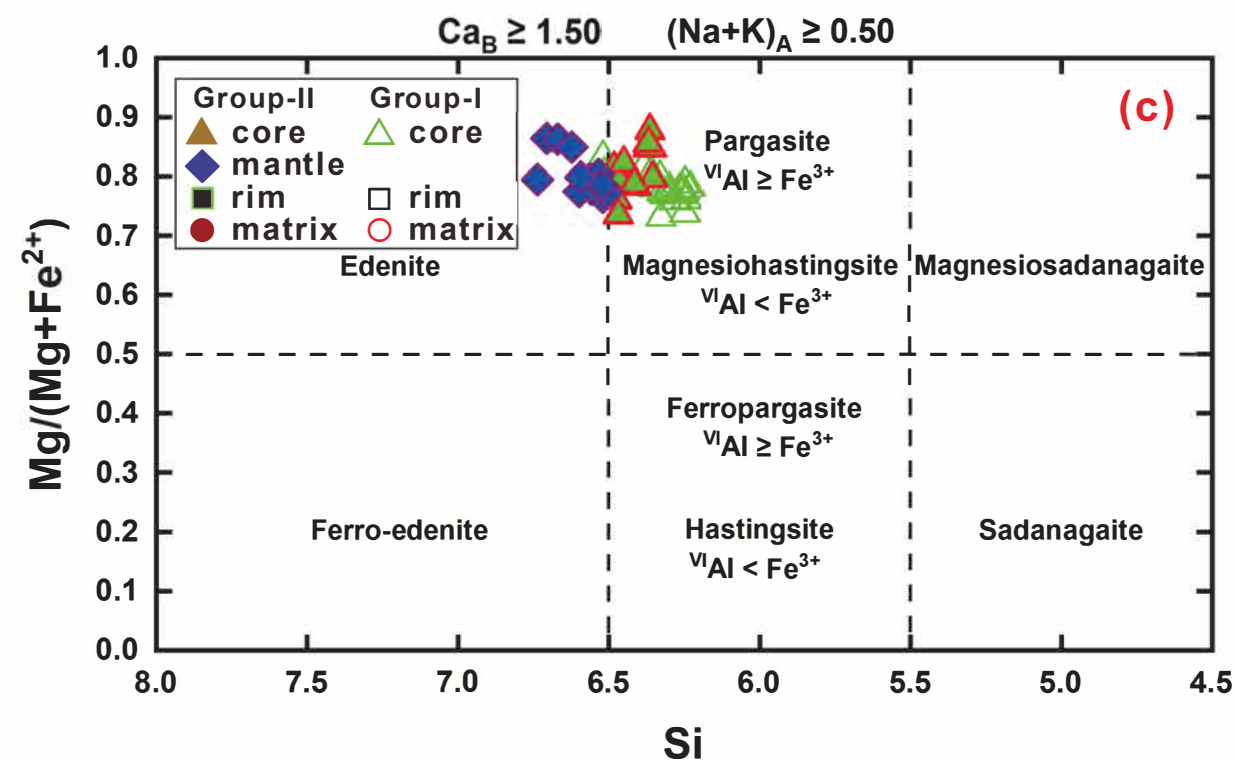
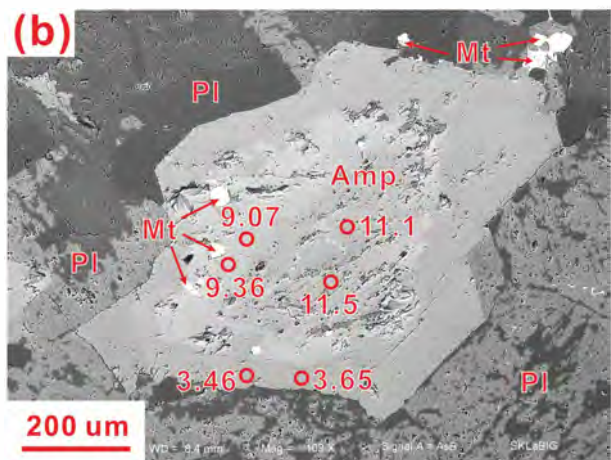
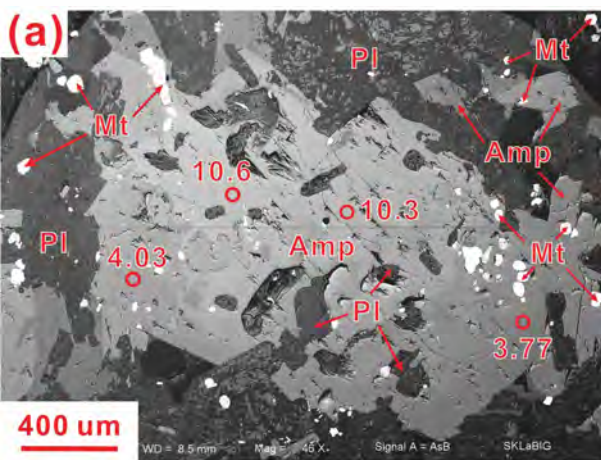


Figure 9

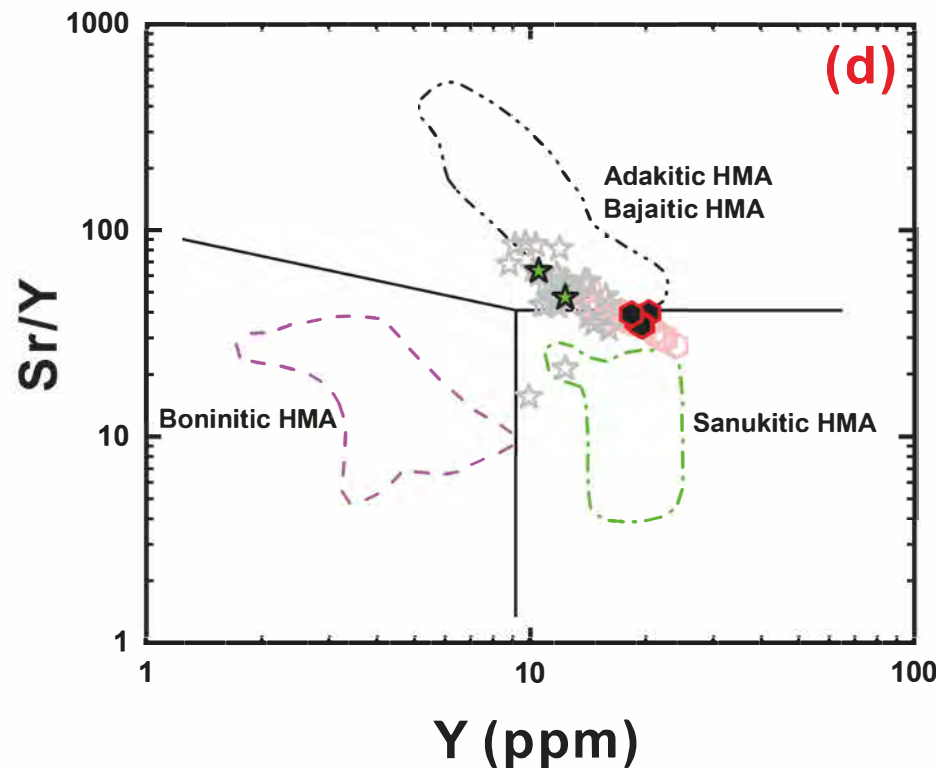
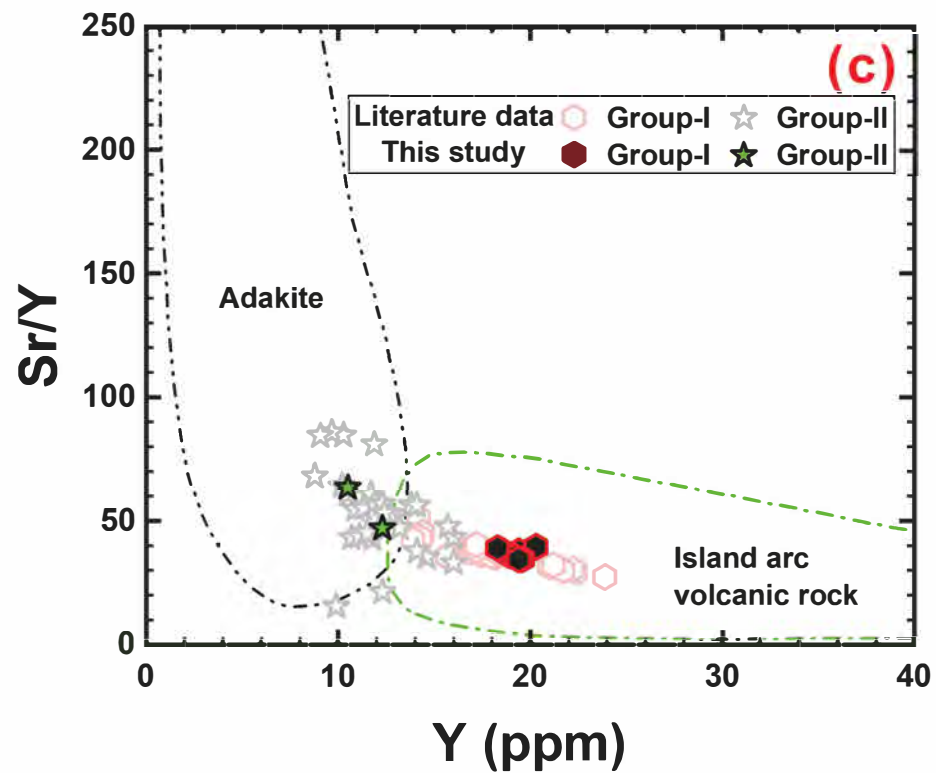
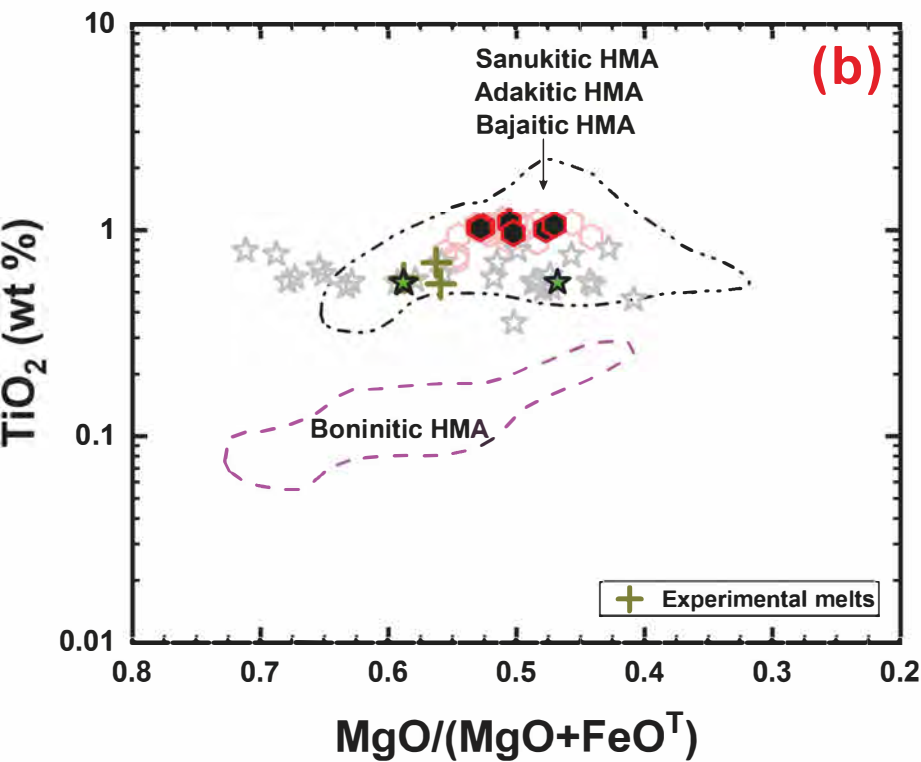
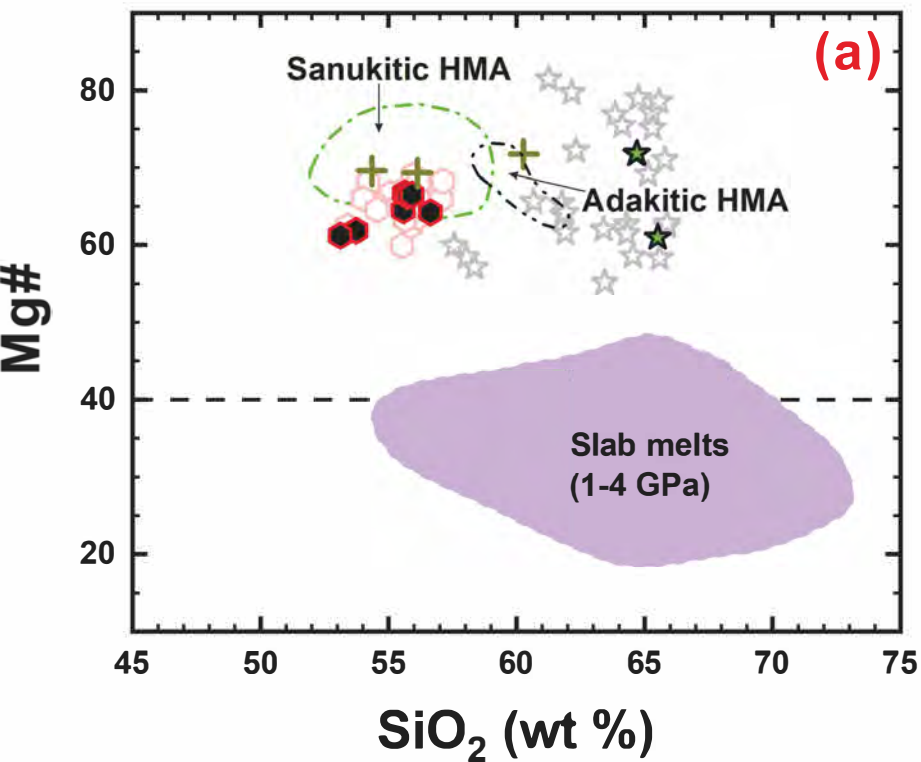


Figure 10

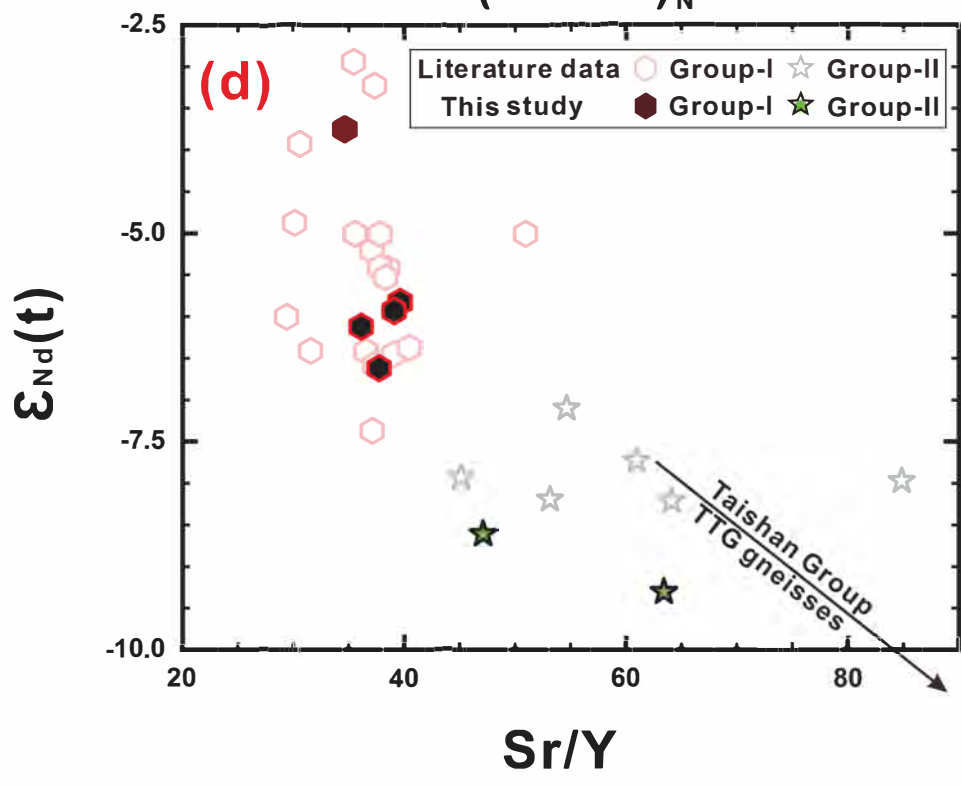
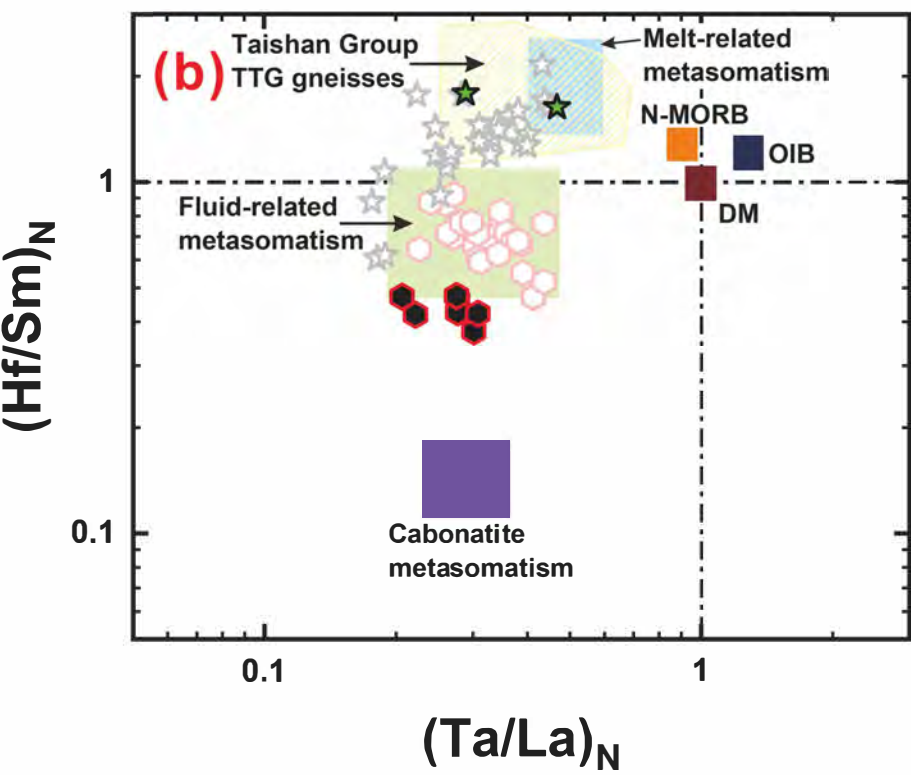
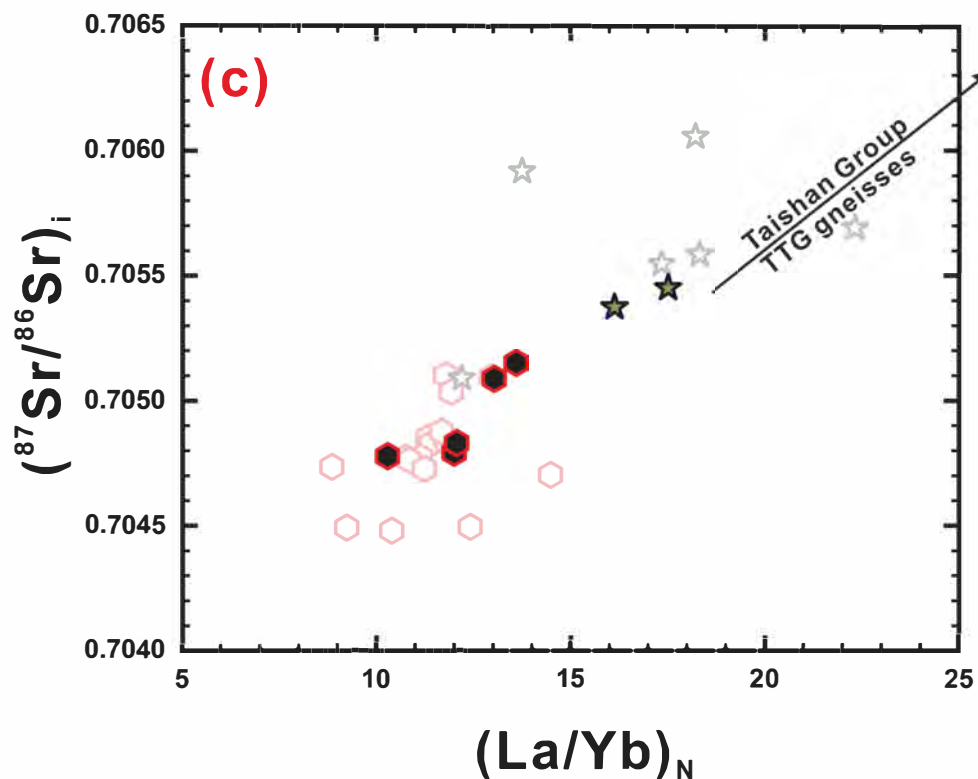
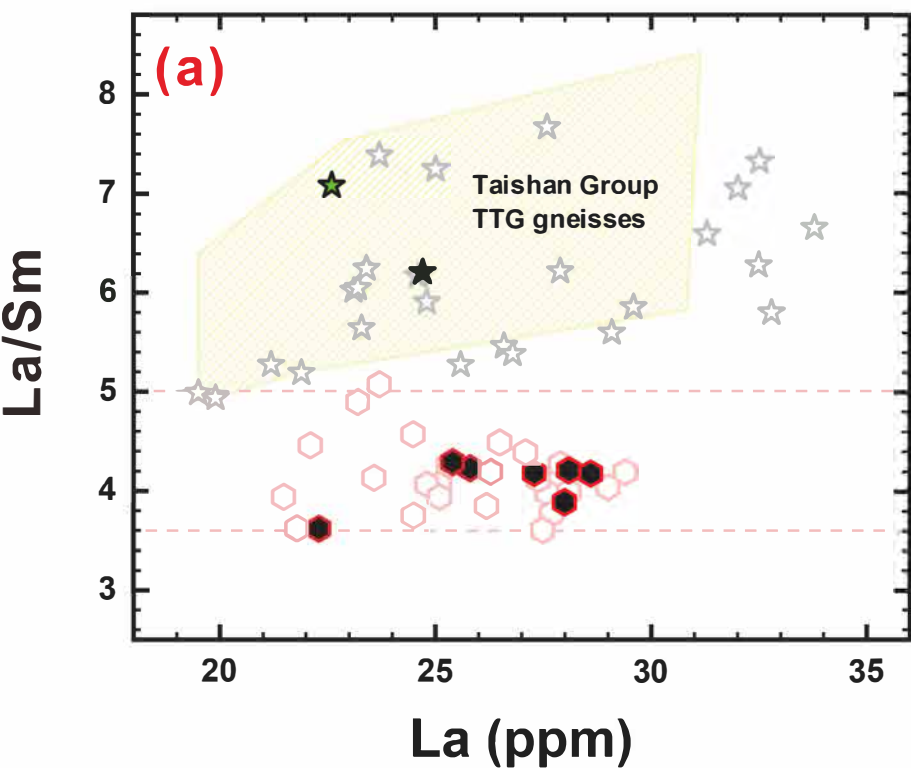




Figure 11

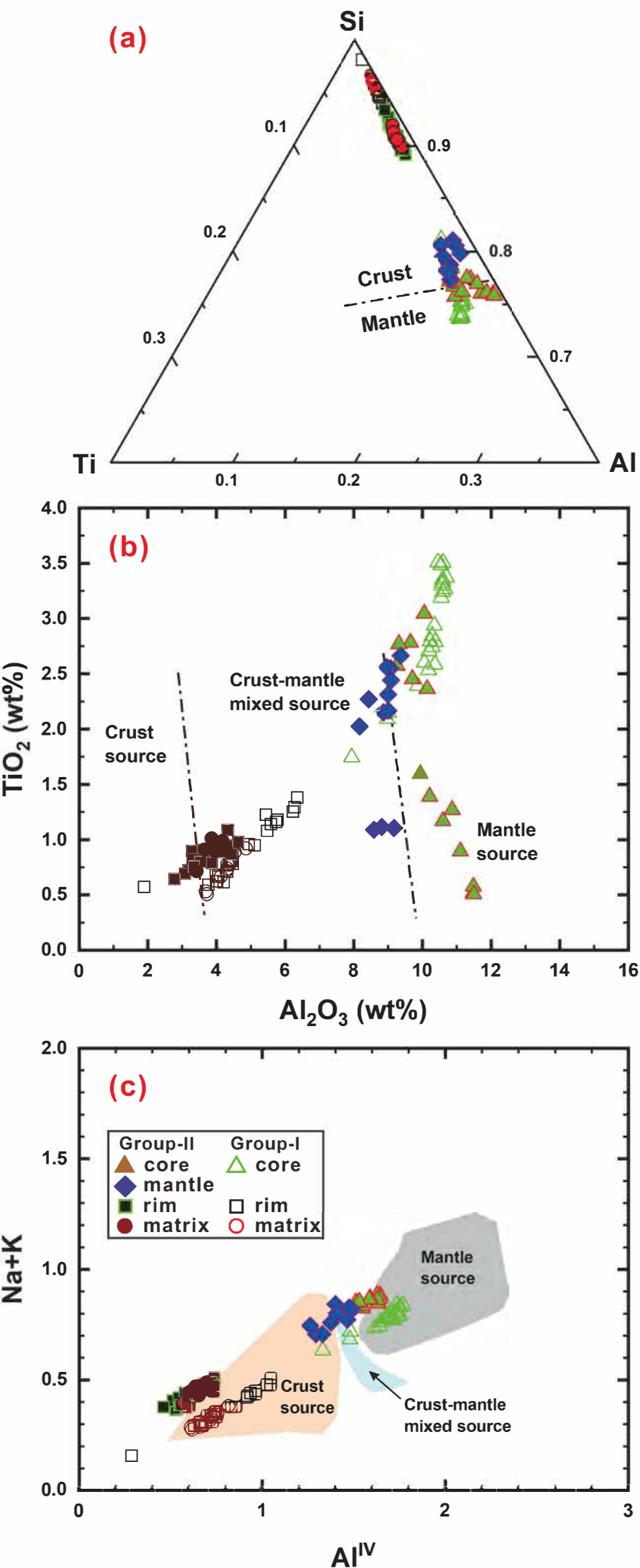


Figure 12

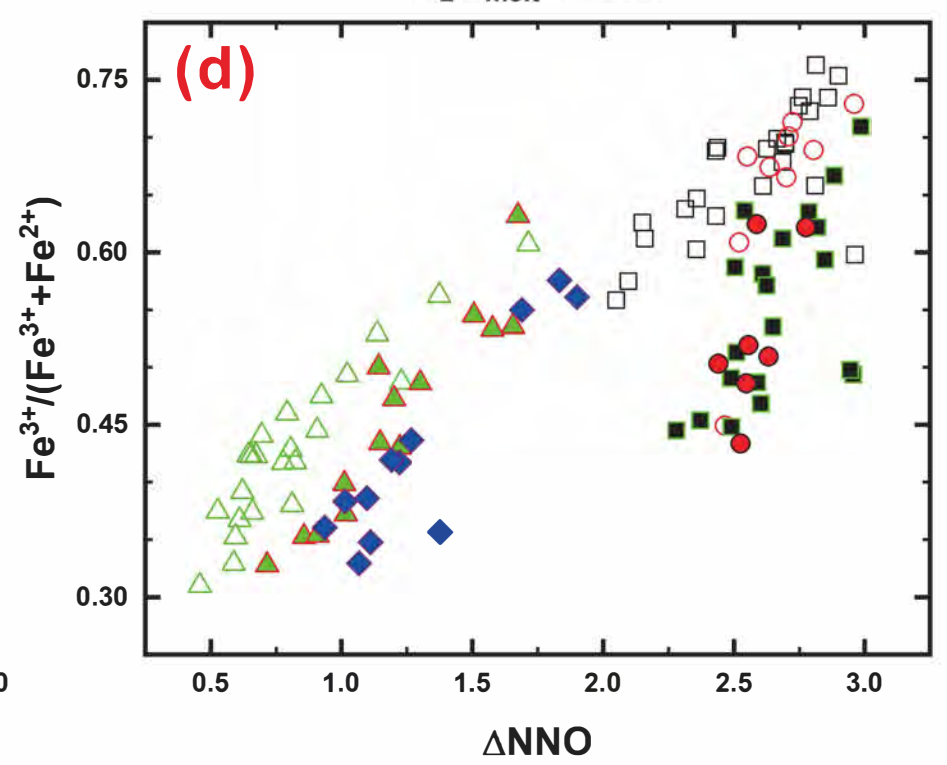
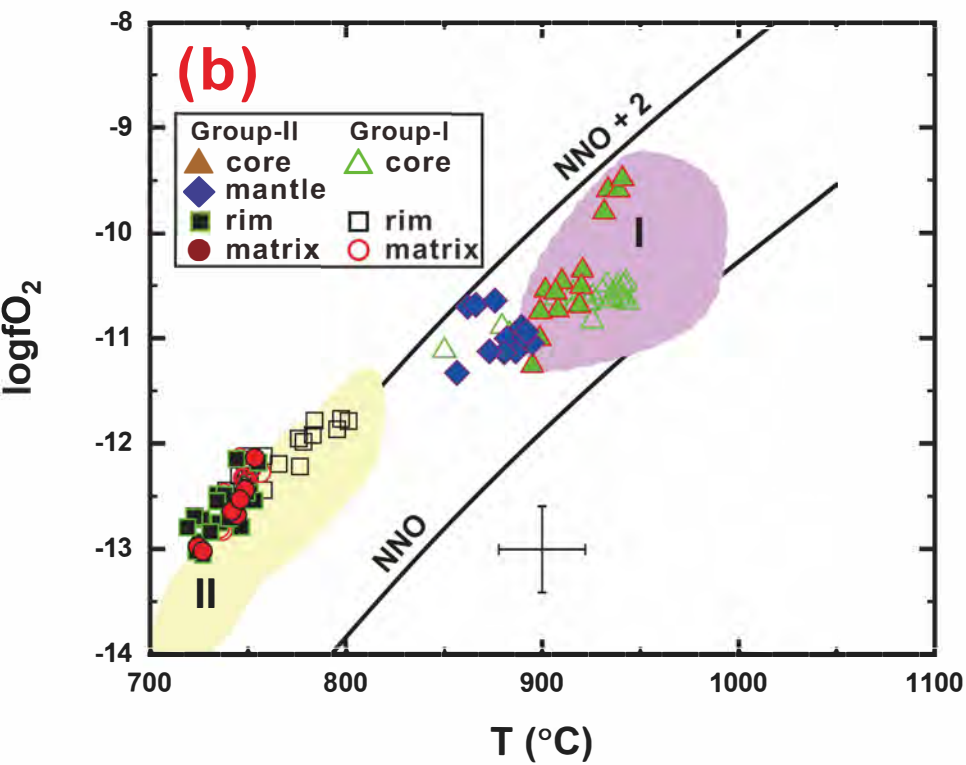
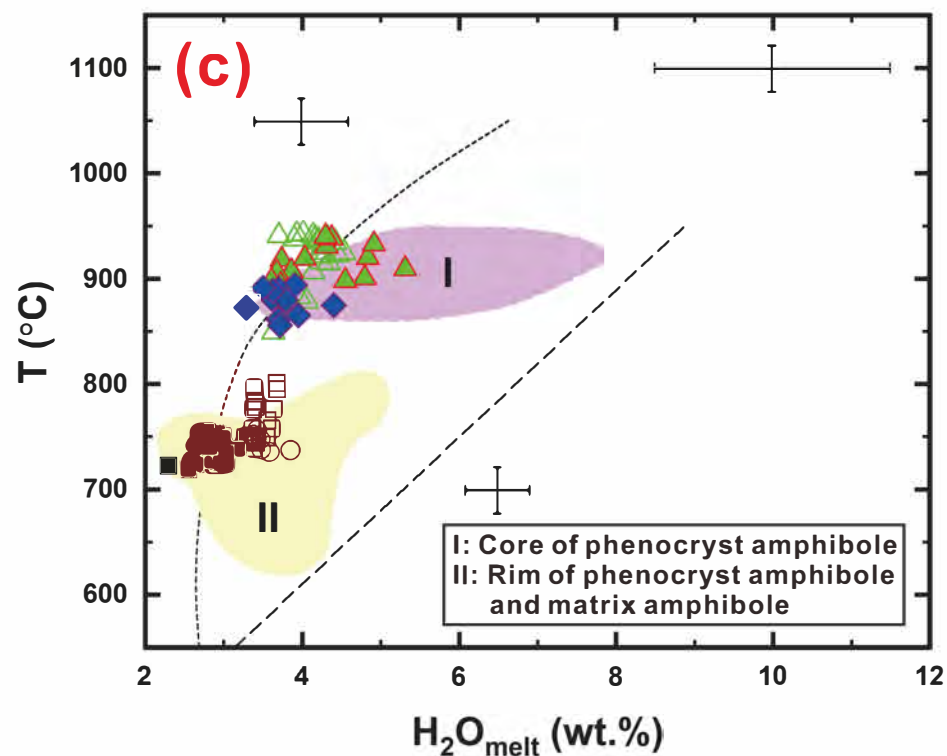
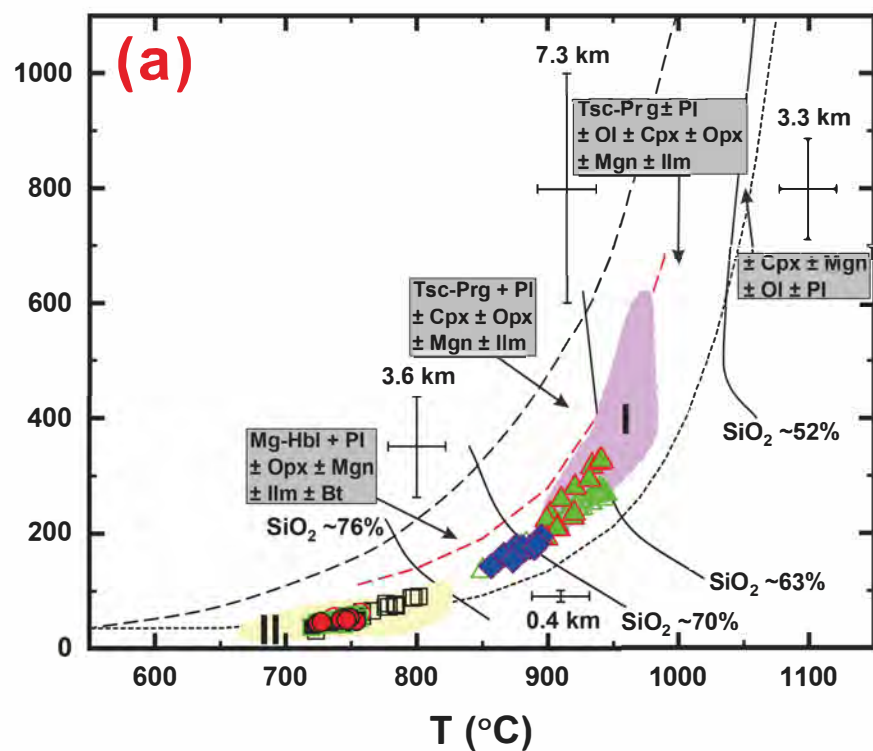


Figure 13

