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4 Magmatic degassing controlled the metal budget of the  
5 Axi epithermal gold deposit

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22 **ABSTRACT**

23 From integrated textural and compositional studies of auriferous and barren  
24 pyrite/marcasite in the epithermal Axi gold deposit, China, we have identified a  
25 relationship between multiple gold mineralizing events, mafic magma recharge and  
26 fluid-rock reactions. Three generations of pyrite (Py1–3) and four generations of  
27 marcasite (Mar1–4) record episodic gold mineralizing events, followed by  
28 silver-copper-lead-zinc-cadmium enrichment. The gold mineralizing events were  
29 recorded by high concentrations of sub-nanometer sized gold in Py1, Py3, and Mar3  
30 (max. = 147, 129 and 34 ppm, med. = 39, 34 and 12 ppm). Based on previous Re-Os  
31 ages determinations of pyrite and U-Pb zircon ages of the andesitic wallrock, these  
32 gold events slightly postdated pulsed mafic magma recharge, and represent the  
33 incursion of Au-As-S-rich magmatic volatiles into circulating meteoric water.  
34 Silver-Cu-Pb-Zn-Cd enrichment in Py2, Mar2 and Mar4 are consistent with quiescent  
35 degassing and gradual Ag-Cu-Pb-Zn-Cd enrichment in an evolved felsic magma.  
36 Barren Mar1 records the dominance of meteoric water, and a limited magmatic fluid  
37 contribution. High Co-Ni-V-Cr-Ti contents in porous cores of Py1 and Mar2 are  
38 attributed to wall rock alteration and dissolution-reprecipitation. The results provide  
39 convincing evidence that the metal budget (especially for Au, Ag, Cu, Pb, Zn, Sb) of  
40 the hydrothermal fluids and sulfides in epithermal systems is controlled by the influx  
41 of magmatic fluids and associated magma, whereas the enrichment of certain

42 fluid-immobile elements, such as Co, Ni, V, Cr and Ti, is caused in part by fluid-rock  
43 interaction.

44 **Keywords:** Mafic recharge; magma degassing; metal budget; fluid-rock interaction;  
45 epithermal deposit; Axi; Central Asian Orogenic Belt

46

## 47 INTRODUCTION

48 Andesites play an important role in the formation and evolution of continental  
49 crust at convergent margins (Rudnick, 1995). They can be produced through  
50 fractional crystallization of basalt and crustal assimilation, slab melting, hydrous  
51 melting of peridotite, or magma mixing between mafic and felsic magmas  
52 (Carmichael, 2002; Defant and Drummond, 1990; Laumonier et al., 2014; Lee and  
53 Bachmann, 2014; Reubi and Blundy, 2009). In the last of these mechanisms, the  
54 replenishment of a shallower, felsic and partially solidified magma reservoir by mafic  
55 magma yields evolved (andesitic to dacitic), volatile-rich magmas that are of  
56 sufficiently low density to erupt (Kent et al., 2010).

57 Epithermal deposits are commonly associated with volcanic rocks in magmatic  
58 arcs (Simmons et al., 2005). Notwithstanding the debate concerning the nature of the  
59 magmatic contribution to the epithermal mineralization, it has been proposed that the  
60 characteristics of epithermal ore-forming systems, as well as those of analogous  
61 geothermal systems, are mainly determined by the nature of the associated  
62 magmatism (Giggenbach, 1995; Sillitoe and Hedenquist, 2003; Simmons and Brown,

63 2007). Considering that mafic melt is intrinsically rich in gold and sulfur (Hattori and  
64 Keith, 2001; Nadeau et al., 2016), the question arises: Does mafic magma  
65 replenishment and the mixing of mafic magma with felsic magma at mid-crustal  
66 levels affect the upper-crustal epithermal mineralization?

67 We selected the Axi epithermal gold deposit (>70 t Au; Chen et al., 2012), NW  
68 China (Fig. 1), to provide the context in which to answer this question because it has  
69 been clearly shown that the host andesitic rocks were the products of episodic mafic  
70 magma injection and mixing with shallower felsic magmas (Zhang, 2020). Moreover,  
71 detailed geochronologic studies demonstrate that precipitation of auriferous pyrite (at  
72 ca. 355 Ma and 332 Ma, Re-Os isochron, Liu et al., 2020; Li et al., 2022) slightly  
73 postdated mafic magma injection (351–357 Ma and 340 Ma, zircon U-Pb age, Li et  
74 al., 2022), indicating a potential link between mafic magma replenishment and gold  
75 mineralization.

76 In this paper, we report the results of combined textural and *in situ* trace-element  
77 studies of auriferous and barren pyrite/marcasite. These results reveal how mafic  
78 magma intrusion, magma mixing, and fluid-rock interaction affect metal budgets in  
79 epithermal systems.

## 80 **THE AXI VOLCANICS AND GOLD MINERALIZATION**

81 The Axi district is located in the Chinese Western Tianshan in an area of  
82 long-lived, subduction-related arc magmatism (Yu et al., 2018; Ye et al., 2020). The  
83 deposit is hosted by volcanic-sedimentary rocks of the Late Paleozoic Dahalajunshan

84 Formation (Fig. 1). Volcanism began at ca. 410 Ma and focused in the vicinity of the  
85 Axi deposit from 361 to 340 Ma (Li et al., 2022 and references therein). The volcanics  
86 exposed in the deposit consist of crystal-rich (30–50%) andesitic to dacitic rocks in  
87 which plagioclase, clinopyroxene, amphibole and quartz phenocrysts are embedded in  
88 a microcrystalline groundmass of plagioclase, magnetite and ilmenite. A wealth of  
89 petrological and geochemical evidence shows that magma mixing played an important  
90 role in the petrogenesis of the Axi volcanics, including the development of  
91 disequilibrium textures (e.g., inverse zoning, patchy zoning, sieve textures) involving  
92 plagioclase phenocrysts, the resorption of quartz phenocrysts, the entrapment of  
93 quenched mafic and felsic silicate melt inclusions by quartz phenocrysts and the  
94 changes in bulk rock chemistry that led to simple linear trends on bivariate plots of  
95 major oxide components (Zhang, 2020).

96 Nearly 90% of the gold (ca. 60 tonnes, with an average grade of 5.6 g/t) were  
97 mined from the No. 1 orebody (Rui et al., 2002; Chen et al., 2012). As details of the  
98 alteration and mineralization have been reported by Zhai et al. (2009), An and Zhu  
99 (2018); Liu et al. (2018, 2020), and Zhang et al. (2018), only a brief summary of these  
100 features is presented below.

101 The hydrothermally altered rock contains quartz, chalcedony, calcite, sericite and  
102 adularia (Bao et al., 2002; Zhai et al., 2009), an assemblage typical of low sulfidation  
103 epithermal deposits (White and Hedenqueist, 1990, 1995; Simmons et al., 2005). A  
104 zonal distribution of alteration is observed, from an inner silicification zone outwards

105 through adularia-sericite to propylitic alteration zones. Silicification is most closely  
106 associated with gold mineralization. The adjacent adularia-sericite alteration zone  
107 comprises sericite, quartz, adularia, plus minor illite and carbonate (Zhai et al., 2009;  
108 An and Zhu, 2018; Liu et al., 2020). The peripheral propylitic alteration is  
109 characterized by the mineral assemblage chlorite, carbonate, and epidote. Native gold,  
110 electrum, pyrite, marcasite, sphalerite, chalcopyrite, tetrahedrite and galena are the  
111 main ore minerals (Liu et al., 2018; Zhang et al., 2018).

## 112 **SAMPLING AND ANALYSIS TECHNIQUES**

### 113 **Sampling**

114 Over 500 samples from outcrop and eight drill holes (ZK2404, ZK2405,  
115 ZK4811, ZK3605, ZK3507, ZK4002, ZK4004, ZK4007) were examined  
116 petrographically in order to identify key minerals and textures. Twelve representative  
117 pyrite and marcasite samples from drill holes ZK2404 and ZK2405 were selected for  
118 further study.

### 119 **SEM-BSE imaging**

120 Thin sections were prepared and examined using transmitted and reflected light  
121 microscopy to characterize the morphology, textures and paragenesis of pyrite and  
122 marcasite. This was followed by SEM-BSE imaging using a FEI Quanta 650 FEG  
123 scanning electron microscope (SEM) at the Key Laboratory of Orogenic Belts and  
124 Crustal Evolution, Peking University. The SEM is equipped with an energy dispersive  
125 spectrometer (EDS), which allows combined backscattered electron (BSE) imaging of

126 mineral morphology and compositional analysis. The applied voltage was 10 to 20  
127 kV, with a spot-size of 4.0 to 5.0  $\mu\text{m}$ , and a dwell time of 20 to 30  $\mu\text{s}$ .

#### 128 **LA-ICPMS trace element analyses**

129 The trace element composition of the sulfide minerals was analyzed using  
130 laser-ablation inductively coupled plasma mass spectrometry (LA-ICPMS) at the  
131 Aarhus Geochemistry and Isotope Research platform (AGiR), Department of  
132 Geoscience, Aarhus University, Denmark. A 193 nm ArF Excimer laser from  
133 Resonetics attached to an Agilent 7900 ICPMS was used for this purpose. The  
134 ablation rate was 8 Hz and the laser energy was 70 mJ. Helium was used as the carrier  
135 gas, which was mixed with Ar before it entered the ICP. A total of 245 single-spot  
136 analyses were performed using a laser beam that varied from 6 to 60  $\mu\text{m}$ , depending  
137 on the size of the crystals. All the data were collected in time-resolved mode. The  
138 analysis time was 55 seconds, including an initial 25 seconds of background  
139 measurement and 30 seconds of ablation. Data for a set of 20 elements were reported  
140 for the spot analyses (Hg, Au, Sb, Ag, Mo, Se, As, Zn, Cu, Co, Mn, Cr, V, Cs, Tl, Pb,  
141 Sc, Ti, Ni and Al). Bismuth, Sn, In, Cd, Pt and U were also analyzed, but their  
142 concentrations were typically very low (<1 ppm) and in many cases below the limit of  
143 detection.

144 The NIST 610 synthetic glass and MASS 1 pressed sulfide reference materials  
145 were used as external calibration standards. The reproducibility and accuracy of  
146 MASS-1 for different analytical sessions are given in Table A1. The relative standard

147 deviation for the different elements is between 2 – 10% (Table A1). The standards  
148 were analyzed after roughly every 10th pyrite analysis. For the quantification of the  
149 data, the Fe concentration obtained by electron microprobe analysis (Zhang et al.,  
150 2018) was used as the internal standard. The data processing and quantification were  
151 done using Iolite software (Paton et al., 2011).

152 In order to develop a chemical profile across selected pyrite grains, two  
153 LA-ICPMS line scans were conducted. The sample was moved under the laser along a  
154 line, crossing the mineral grain at a scanning speed of 8  $\mu\text{m/s}$  with a spot size of 8  $\mu\text{m}$ .  
155 The elements mentioned above were chosen for these analyses. The profiles varied  
156 between 300 to 400  $\mu\text{m}$  in length.

## 157 **RESULTS**

### 158 **Types of pyrite/marcasite**

159 Previous studies have documented many features of the pyrite at Axi, including  
160 their textures, paragenetic relationships and composition (Zhang et al., 2018; Liu et  
161 al., 2015, 2018). These studies reveal that pyrite hosts nearly half of the total gold  
162 reserve, and that there are three generations of pyrite. In this study, we couple detailed  
163 SEM/EDS imaging of pyrite and marcasite with specific textures/minerals, including  
164 crustiform banding, hydrothermal breccias, adularia, platy calcite/quartz and  
165 replacement textures, and use them to refine previous textural and mineralogic  
166 interpretations.

167       The earliest pyrite, Py1 (50 – 1500  $\mu\text{m}$  in diameter, Fig. 2, 3) is associated with  
168 quartz and sericite. It comprises a corroded, porous core (Py1a), a compact,  
169 inclusion-free mantle (Py1b), an oscillatory-zoned rim (Py1c), and an arsenopyrite  
170 overgrowth (these features are not present in all the crystals). The anhedral Py1a  
171 replaced clinopyroxene phenocrysts as pseudomorphs (Fig. 3B) or is concentrated  
172 close to the fluid-rock reaction front (Fig. 3C). Subhedral to euhedral Py1b and Py1c  
173 are more common in quartz-dominated cements or veinlets and away from the  
174 fluid-rock reaction front (Fig. 3C). Anhedral Py2 (50 – 1000  $\mu\text{m}$  in diameter) coexists  
175 with chalcopryrite, sphalerite, tetrahedrite, quartz, calcite and chlorite (Fig. 2B, 3D).  
176 Tiny Py3 crystals (10 – 50  $\mu\text{m}$  in diameter) are euhedral to subhedral, and  
177 oscillatory-zoned (Fig. 2C). They are surrounded by chalcedony  $\pm$  calcite, and cement  
178 Py1 and/or Py2 clasts in breccias (Fig. 3F).

179       In addition to pyrite, we identified four generations of marcasite (Figs. 2-4) at  
180 shallower depth (< 280 m) than pyrite (that is present to 720 m). The blade-shaped  
181 Mar1 (50 – 600  $\mu\text{m}$  in length) coexists with fine- to coarse-grained quartz, chalcedony  
182 and/or calcite. It may crosscut Py1 (Fig. 3A) and Py2 and hence crystallized later than  
183 them. This marcasite occurs in crustiform layers with a rhythmic alternation of quartz,  
184 chalcedony, Py3 and Mar1 (Fig. 4). Since Py3 occurs more commonly in the outer  
185 layers, and Mar1 in the inner layers, we propose that Mar1 crystallized  
186 contemporaneously with or slightly later than Py3. The second marcasite generation,  
187 Mar2, forms radial or fan-shaped aggregates (100 – 2000  $\mu\text{m}$  in length) and has an

188 inclusion-rich, porous core (Mar2a) that is surrounded by a homogeneous marcasite  
189 overgrowth (Mar2b) and arsenopyrite (Fig. 2D). It cements hydrothermal breccia  
190 fragments together with fine-grained quartz, chalcedony and minor calcite. This  
191 generation of marcasite was followed by Mar3 (100 – 1500  $\mu\text{m}$  in length) in calcite±  
192 chalcedony veins/cements. Commonly, Mar3 crosscuts anhedral Py2 (Fig. 3E), or  
193 cements breccia clasts containing earlier pyrite and marcasite generations. Locally,  
194 Mar3-calcite veins crosscut lattice structures and adularia, and calcite may fill the  
195 lattice interstices (Fig. 3I). The latest marcasite generation (Mar4) forms acicular  
196 crystals/aggregates (50 –1000  $\mu\text{m}$  in length) and is hosted by calcite veins, in which  
197 the calcite is euhedral to subhedral and occurs as rhombs with alternating clear and  
198 “dusted” zones due to sub-microscopic inclusions. These veins crosscut  
199 calcite-chalcedony-Mar3 veins (Fig. 3H), indicating that Mar4 was later than Mar3.  
200 On a modal basis, Py1, Py3, Mar2 and Mar3 are the most abundant pyrite and  
201 marcasite phases, followed by minor Py2, and traces of Mar4 and Mar1.

## 202 **The trace element record of episodic mineralization**

203 The trace element data for pyrite and marcasite are summarized in Table 1 and  
204 listed in Appendix A2. In the following text, the maximum (max.) values are reported  
205 for gold hosted by Py1b, Py3 and Mar3, and median (med.) values are reported for all  
206 other elements/samples (see also Fig. 5).

207 The compositional data reveal episodic gold mineralizing events. Each of these  
208 events was followed by a period of silver and base metal enrichment. The first gold

209 mineralizing event is recorded by Py1b that has the highest concentrations of Au (max.  
210 = 147 ppm, med. = 39 ppm) and As (67500 ppm). The precursor Py1a, by contrast, is  
211 characterized by higher Co (209 ppm), Ni (510 ppm), V (14 ppm), Cr (12 ppm) and  
212 Ti (3300 ppm) contents than Py1b (Co = 25 ppm, Ni = 35 ppm, V = 0.83 ppm, Cr =  
213 3.5 ppm, Ti = 12 ppm). The compositional variation of different parts of Py1 crystals  
214 is clearly illustrated by the laser ablation traverses across a Py1 grain (Fig. 6). The  
215 subsequent Py1c is relatively deficient in most trace elements in comparison to Py1a  
216 and Py1b, especially As (15000 ppm), Zn (1.1 ppm), Cu (96 ppm), Co (11 ppm), V  
217 (0.67 ppm), Cr (1.0 ppm) and Ti (7.7 ppm). Gold mineralization was followed by the  
218 introduction of Cu, Pb, Zn and Sb, which is recorded in Py2 (Cu = 928 ppm, Pb = 124  
219 ppm, Zn = 96 ppm, Sb = 884 ppm). This pyrite type has the highest concentration of  
220 Ag among all the samples (max. = 305 ppm, med. = 157 ppm) but is depleted in Au  
221 (6.0 ppm).

222 The second gold mineralizing event is documented by the composition of Py3  
223 (max. = 129 ppm, med. = 34 ppm Au). Later Mar1, however, is depleted in most trace  
224 elements. A second period of Ag-Zn enrichment is recorded by Mar2 (containing 243  
225 ppm Ag and 69 ppm Zn in the core, and 60 ppm Ag and 27 ppm Zn in the rim). Its  
226 porous core (Mar2a) has higher Co (298 ppm), Ni (697 ppm), V (9.7 ppm), Cr (8.0  
227 ppm) and Ti (32 ppm) contents than the compact Mar2b overgrowth (Co = 58 ppm,  
228 Ni = 195 ppm, V = 7.0 ppm, Cr = 3.8 ppm, Ti = 13 ppm), mimicking the  
229 compositional difference between Py1a and Py1b (Fig. 5).

230 A weak gold mineralizing event may be indicated by the composition of Mar3,  
231 although its Au content is much lower (max.= 34 ppm, med. = 12 ppm) than those of  
232 Py1b (39 ppm) and Py3 (34 ppm). It also has relatively high Ag (146 ppm), Pb (169  
233 ppm), Zn (29 ppm) and Sb (872 ppm) contents. The youngest marcasite variety,  
234 Mar4, is characterized by extremely low contents of Au (0.11 ppm), Ag (13 ppm) and  
235 As (3115 ppm), but moderately high Pb (1027 ppm), Zn (104 ppm) and Sb (417 ppm)  
236 contents.

## 237 **INTERPRETATION AND DISCUSSION**

238 In a previous study, we showed that the Axi andesites accumulated incrementally  
239 through injections of mafic magma into a shallow, felsic magma reservoir (Li et al.,  
240 2022). We documented the injection events by the textural features of zircon in which  
241 an anhedral, antecrystic core is resorbed and embayed by a thin, irregular zone (<5  
242  $\mu\text{m}$ ) prior to overgrowth by an oscillatory-zoned autocrystic rim (Li et al., 2022).  
243 Using the ages of the core and rim of the zircon, the main period of injections is  
244 interpreted to have been between 351 and 357 Ma, contemporaneous with the  
245 crystallization of Py1b ( $357 \pm 18$  Ma,  $353 \pm 6$  Ma, Re-Os isochron, Liu et al., 2020;  
246 Li et al., 2022). This chronology, in combination with our new sulfide trace element  
247 data, provides compelling evidence for a link between magma injection and gold  
248 mineralization. We propose that fluid exsolution associated with mafic magma  
249 injections and interspersed periods of magma quiescence exerted a first-order control  
250 on the metal budget of the ores (Fig. 7). We further propose that enrichment of certain

251 fluid-immobile elements, such as Co, Ni, V, Cr and Ti, was caused, in part, by  
252 fluid-rock interaction. A detailed discussion is provided below.

253 The earliest pyrite, anhedral Py1a, is interpreted to have formed as a result of  
254 fluid-mediated sulfidation, or coupled dissolution-reprecipitation reactions (Putnis,  
255 2009), based on the observation that it replaced clinopyroxene phenocrysts as  
256 pseudomorphs, or is concentrated close to the sharp fluid-rock reaction front (Fig.  
257 3C). Intense alteration led to the breakdown of clinopyroxene, amphibole, plagioclase,  
258 magnetite-ilmenite and other constituents of the andesite host. Consequently,  
259 elements that were once compatible in these phases (such Co, Ni, V, Cr and Ti) were  
260 released into the circulating fluids (Mustard et al., 2006; Jenner et al., 2010), and were  
261 scavenged by the alteration products. Thus, it was the alteration of the parent mineral,  
262 coupled in space and time with the precipitation of daughter minerals and  
263 co-generation of porosity, that led to the development of porous, Co-Ni-V-Cr-Ti-rich  
264 Py1a.

265 The gold mineralization event recorded by Py1b is interpreted to have been  
266 associated with mafic magma intrusion into the shallow, felsic magma reservoir  
267 referred to above. Although the sulfur and gold contents of igneous rocks in the Axi  
268 area have not been analyzed, mafic magma is generally considered to have high sulfur  
269 (more than several thousand ppm; Nadeau et al., 2016) and Au contents (Keith et al.,  
270 1997; Hattori and Keith, 2001; Halter et al., 2005; Stern et al., 2007; Zajacz et al.,  
271 2012; Guo and Audetat, 2017), and thus, is likely to have been the source of sulfur

272 and gold. Injections of mafic magma would have induced a sudden over-saturation in  
273 volatiles (Kent et al., 2010; Nadeau et al., 2016), with sulfur partitioning  
274 preferentially into the exsolving supercritical fluid (a  $D_S^{\text{fluid/melt}}$  value of 20 has been  
275 proposed based on hydrothermal S partitioning experiments, Webster and  
276 Botcharnikov, 2011). Elements such as Au, As and Hg are readily transported by  
277 aqueous vapor, as shown by the compositions of volcanic gases and their sublimates  
278 (Williams-Jones and Heinrich, 2005). This reflects the high volatility of Hg and As  
279 (Williams-Jones et al., 2002); and Au prefers reduced sulfur, which also partitions  
280 preferentially into the vapor (Hurtig and Williams-Jones, 2014; Williams-Jones and  
281 Migdisov, 2014). In contrast, elements such as Pb and Zn are likely to have low  
282 volatility because of the very strong bonds that they form with chloride ions in  
283 aqueous liquids. On rising above the depth of exsolution, the supercritical fluid, which  
284 would have been liquid-like because of the high temperature of the mafic magma,  
285 would have condensed a small proportion of brine and Au-As-S would have been  
286 transported upwards dominantly by the more buoyant vapor. Incursion of this  
287 Au-As-S-rich vapor thus would have greatly increased the gold content of the  
288 circulating meteoric fluids and decreased their pH (Zhai et al. 2009), leading to gold  
289 mineralization and associated quartz-sericite-pyrite (Py1b) alteration. Later Py1c is  
290 interpreted to have precipitated from fluids with lesser inputs of magmatic  
291 components, and thus less Au and As.

292 The Ag and base metal-rich Py2 postdated Py1 (Liu et al., 2020; Zhang et al.,  
293 2018) and the related mafic magma injection. It possibly accompanied a period of  
294 magma quiescence, during which there was a gradual enrichment of Ag, Cu, Pb, Zn  
295 and Sb in the evolved magma and exsolved fluids (Nadeau et al., 2016). In the  
296 absence of mafic magma injection, the supercritical fluid would have been liquid-like  
297 because of its lower temperature, and would have separated to produce a small  
298 proportion of vapor leaving a brine more enriched in Cl (Giggenbach, 1997; Zajacz et  
299 al., 2008). This, in turn, would have enhanced the transportation of Cl-complexed  
300 elements (such as Ag, Cu, Pb, Zn and Sb, Wood and Samson, 1998; Seward et al.,  
301 2014) into shallow meteoric water. As a representative of the second gold  
302 mineralizing event, Py3 incorporated comparable amounts of Au and As to Py1b,  
303 indicating a similar source, i.e., mafic magma. The concentrations of Ag, Cu, Pb, Zn  
304 and Sb in Py3, however, are slightly higher than those of Py1b. Invoking a  
305 contribution from the felsic magma would reconcile this issue, by providing an  
306 additional metal reservoir for Ag, Cu, Pb, Zn and Sb. The deposition of Py3 is  
307 attributed to fluid boiling, as shown by: (1) the development of hydrothermal breccias  
308 (with Py3 ± chalcedony as the cement); and (2) the occurrence of adularia and bladed  
309 calcite (pseudomorphed by quartz) (Fig. 3I). Boiling not only caused precipitation of  
310 gold, bladed calcite and adularia (Simmons et al., 2005), but fractured the surrounding  
311 rocks, accelerating mixing with meteoric fluids. Marcasite has been shown to form  
312 from aqueous solutions with a pH < 5, at temperatures of 80 – 240 °C, in the presence

313 of H<sub>2</sub>S<sub>2</sub> (aq) (Murowchick, 1992), whereas Py3 is interpreted to have deposited at 325  
314 – 385 °C, as determined by the arsenopyrite geothermometer (Zhang et al., 2018).  
315 The occurrence of Py3 and Mar1 in alternating crustiform layers (Fig. 4) may have  
316 been the result of fluid mixing, with the high temperature, Au-rich Py3 recording a  
317 greater magmatic fluid contribution and the low temperature, trace-element-poor  
318 Mar1 recording the predominance of meteoric water and a minor magmatic fluid  
319 component.

320 The marcasite variety, Mar2a, shares textural (porous) and compositional  
321 (Co-Ni-V-Cr-Ti-rich) characteristics similar to Py1a. The compact overgrowth  
322 contains proportions of Ag, Pb, Zn and Sb overlapping with or slightly lower than  
323 those of Py2, suggesting a similar metal control by quiescent degassing. Thus, the  
324 conditions of formation of Mar2a and Mar2b were comparable to those of Py1a and  
325 Py1b, with the exception that the invading fluid was Pb-Zn-Sb-rich instead of  
326 Au-As-S-rich. Crystallization of Mar2 was followed by that of Mar3, which contains  
327 slightly less Au and As than Py1b and Py3 and may record a small injection of mafic  
328 magma that was not identified by zircon age dating. The elevated Ag, Sb and Zn  
329 contents suggest a hybrid input involving a contribution from the evolved felsic  
330 magma. In the final variety of marcasite, Mar4, the fingerprint of the mafic magma is  
331 absent, and only that of the felsic magma (high Pb, Zn and Sb) is recorded.

### 332 **CONCLUSION AND IMPLICATION**

333           The present work provides convincing evidence for a linkage between mafic  
334 recharge, andesite formation and epithermal gold mineralization. It suggests that the  
335 bulk metal budget (especially for Au, Ag, Cu, Pb, Zn, Sb) of an epithermal system is  
336 controlled primarily by the composition of the fluids, which, in turn, is controlled by  
337 the nature of the magma in a shallow magma chamber (Keith et al., 2018; Martin et  
338 al., 2020). The episodic gold mineralizing events are attributed to repeated mafic  
339 magma injections into a shallow, felsic magma chamber and associated magma  
340 degassing. Silver-Cu-Pb-Zn-Sb enrichment, by contrast, is related to quiescent  
341 degassing of the felsic magma. It also acknowledges that fluid-rock interaction would  
342 have involved coupled dissolution and reprecipitation and facilitated immobile  
343 element enrichment such as Co, Ni, V, Cr and Ti. Considering that mafic magma  
344 replenishment and its mixing with felsic magma play an important role in the  
345 formation of andesite (Anderson, 1976; Eichelberger, 1978; Reubi et al., 2009; Kent  
346 et al., 2010), the metal budget reported for Axi may also apply to similar cases in  
347 convergent margins.

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540

#### 541 **FIGURE CAPTIONS**

542 FIGURE 1. A simplified geological map of the Axi epithermal deposit (modified after  
543 [Zhang et al., 2018](#)).

544

545 FIGURE 2. Representative SEM images, showing textures of different types of pyrite  
546 and marcasite.

547 A: Py1 characterized by a porous core (Py1a), a compact mantle (Py1b) and an  
548 oscillatory-zoned rim (Py1c). Near it is bladed Mar1 with an arsenopyrite overgrowth.

549 Also shown is a laser ablation traverse across the crystal; B: Py2 accompanied by  
550 coeval chalcopyrite and tetrahedrite; C: tiny, oscillatory-zoned Py3; D. Mar2 with a

551 laminated relict “core” (Mar2a), clear “rim” (Mar2b) and an arsenopyrite overgrowth;

552 E. fan-shaped Mar3 (with an arsenopyrite overgrowth) coexisting with chalcedony

553 and calcite; F. Homogeneous Mar4 characterized by the absence of an arsenopyrite

554 overgrowth and coexistence with zoned calcite.

555 Abbreviations: Asp-arsenopyrite; Cc-calcite; Ccp-chalcopyrite; Mar-marcasite;

556 Py-pyrite; Qz-quartz; Td-tetrahedrite.

557

558 FIGURE 3. Photomicrographs showing paragenetic relationship among different  
559 types of pyrite and marcasite.

560 A: Py1 crosscut by Mar1; B: pseudomorphic replacement of clinopyroxene by Py1; C:  
561 a breccia clast containing Py1 cemented by Py1+quartz; Py1 crystals at the reaction  
562 front or in the clast are morphologically irregular (Py1a), whereas those away from  
563 the front or in the cement are subhedral to euhedral Py1b or Py1c; D: Py1 included in  
564 or surrounded by Py2; E: Py2 crosscut by Mar3; F: Py1-bearing clast cemented by  
565 Py3+chalcedony; G: a clast containing Py1, Py3 and Mar1 cemented by Mar2; H: a  
566 calcite-Mar4 vein crosscutting a calcite-chalcedony-Mar3 vein; I: a lattice texture  
567 composed of platy quartz (pseudomorphs after calcite) and adularia, which was  
568 crosscut or infilled by later calcite. The yellow lines illustrate the inner textural  
569 boundary (A), and the white lines outline the phenocryst (B), breccia (C, F, G), or  
570 veinlet (H).

571 Mineral abbreviation: Adl-adularia; Apy-arsenopyrite; Cc-calcite; Cln-chalcedony;  
572 Mar-marcasite; Py-pyrite; Qz-quartz.

573

574 FIGURE 4. Crustiform banding characterized by the rhythmic alternation of quartz,  
575 chalcedony, Py3 and Mar1.

576 The crustiform bands comprise eight thin layers (numbered 1 to 8); Mar 1 either  
577 forms thin laminae (layer 7 in Fig. C)) alternating with aggregates of Py3 and

578 chalcedony (layer 1), or coexist with Py3 and arsenopyrite in a single band, indicating  
579 nearly contemporaneous precipitation of Mar1 and Py3.

580 Mineral abbreviation: Apy-arsenopyrite; Mar-marcasite; Py-pyrite.

581

582 FIGURE 5. Boxplots of selected trace element concentrations in pyrite and marcasite,  
583 illustrating the episodic nature of the gold mineralizing events (Au1, Au2, Au3) and  
584 Ag-Pb-Zn-Sb enrichment (BM1, BM2, BM3).

585

586 FIGURE 6. Two laser ablation traverses across a Py1 grain illustrating the elevated  
587 Au and As contents of Py1b. There are numerous inclusions (e.g., rutile and galena) in  
588 the core (Py1a).

589

590 FIGURE 7. A schematic model showing the relationship between mafic magma  
591 replenishment, magmatic degassing, sulfide precipitation and mineralization.

592 Fig. A shows intense magmatic degassing associated with the intrusion of mafic  
593 magma into a shallow, felsic magma reservoir, where the degassed, supercritical  
594 volatile is enriched in Au, As and S. Fig. B illustrates a period of magma quiescence,  
595 during which the magma and exsolved supercritical fluids are gradually enriched in  
596 Ag, Cu, Pb, Zn and Sb. During intrusion of the mafic magma, the supercritical fluid is  
597 vapor-like and condenses brine, with S, Au and As transported dominantly by the  
598 vapor before injection into circulating meteoric water, whereas during magma

599 quiescence the temperature is lower, the supercritical fluid is liquid-like and Cl, Cu,  
600 Ag, Pb, Zn are transported by the brine.

601 Fig. C illustrates the association between magma degassing and precipitation of FeS<sub>2</sub>.  
602 The formation of porous, Co-Ni-V-Cr-Ti-rich Py1a and Mar2a was caused by coupled  
603 dissolution-reprecipitation of the host andesite (here represented by the clinopyroxene,  
604 i.e. Cpx). The incursion of Au-As-S-rich volatile induced precipitation of Py1b, Py3  
605 and Mar3, while the Ag-Cu-Pb-Zn-Cd-rich volatile favored formation of Py2, Mar2b  
606 and Mar4. Interlayered Mar1 with Py3 in crustiform banding resulted from mixing  
607 with (or diluted by) meteoric water.

608

#### 609 **TABLE CAPTIONS**

610 Table 1 Summary of compositions of pyrite and marcasite from the Axi gold deposit  
611 (ppm)

612

#### 613 **Appendix**

614 Appendix A1. Composition and reproducibility of the standard MASS-1.

615

616 Appendix A2. Compositions of selected pyrite and marcasite crystals from the Axi  
617 gold deposit (ppm).

**Table 1.** Summary of compositions of pyrite and marcasite from the Axi gold deposit (ppm)

No.	Hg	Au	Sb	Ag	Mo	Se	As	Zn	Cu	Co	Mn	Cr	V	Cs	Tl	Pb	Sc	Ti	Ni	Al
<b><i>Py1a</i></b>																				
Min.	0.08	1.3	47	6.3	0.04	4.4	9400	0.60	258	6.8	5.4	0.93	1.5	0.83	7.3	108	2.9	306	63	117
Max.	6.2	129	1160	95	27	17000	66300	37	2070	690	336	32	820	50	244	2550	50	38700	1400	17900
Med.	1.5	23	168	27	0.77	70	44000	9.6	459	209	28	12	14	10	43	381	11	3300	510	3860
Std.	2.2	35	276	22	7.6	4198	15257	9.3	370	159	83	8.0	168	11	67	627	13	7279	376	4537
<b><i>Py1b</i></b>																				
Min.	0.32	1.0	44	6.2	0.26	15	18900	0.20	56	0.61	0.80	0.24	0.37	0.32	0.16	2.3	0.09	8.3	3.9	2.5
Max.	4.4	147	2070	70	14	44000	168000	23	790	326	1600	34	1630	88	570	1090	36	18300	500	11200
Med.	1.2	39	107	23	1.5	195	67500	1.8	265	25	7.3	3.5	0.83	1.3	1.8	25	0.22	12	35	175
Std.	1.6	58	682	20	4.5	19149	48377	7.5	244	101	521	10	511	30	182	335	13	5745	148	4491
<b><i>Py1c</i></b>																				
Min.	0.03	1.0	11	1.5	0.02	1.10	3010	0.25	0.75	0.05	0.19	0.25	0.04	0.01	0.47	0.06	0.01	4.1	0.05	1.0
Max.	2.8	137	702	253	276	104	84300	2330	1710	570	202	266	164	5.7	125	3190	37	27000	936	32000
Med.	0.20	13	119	27	2.7	11	15000	1.1	96	11.0	4.9	1.0	0.67	1.0	8.0	30	0.16	7.7	40	153
Std.	0.57	35	138	44	43	18	17816	356	416	100	31	45	26	1.3	22	506	7.1	4453	206	5241
<b><i>Py2</i></b>																				
Min.	0.10	1.0	156	53	0.02	0.3	1890	20	296	0.74	0.42	0.18	0.04	0.02	6.2	6.3	0.01	4.2	5.3	0.40
Max.	0.45	15	2700	305	1.4	27	9600	1620	10200	221	11	6.8	6.3	2.1	39	521	0.17	270	242	1160
Med.	0.25	6.0	884	157	0.28	5.2	4610	96	928	38	2.4	0.64	0.59	0.53	18	124	0.07	5.3	66	341
Std.	0.08	3.3	596	86	0.38	6.1	1582	402	2388	51	2.4	1.5	1.6	0.59	9.8	117	0.05	77	69	326
<b><i>Py3</i></b>																				
Min.	0.20	0.17	10	0.04	0.02	6.0	5340	1.5	5.2	0.21	1.0	0.80	0.30	0.64	0.40	1.3	0.10	1.30	7.3	4.5
Max.	6.2	129	1590	330	10	650	142000	73	729	715	590	205	232	15	134	3300	75	95000	1660	62000

Med.	0.60	34	189	53	1.1	40	37100	9.7	292	106	27	15	11	6.2	34	306	3.9	1905	238	4175
Std.	1.8	37	375	59	3.4	166	31994	14	213	162	105	44	52	4.8	30	752	16	21683	482	12864
<b>Mar1</b>																				
Min.	0.03	4.1	44	20	8.2	1.9	3820	0.38	4.5	0.22	0.70	0.35	0.11	0.01	5.2	2.7	0.01	4.1	3.3	28
Max.	1.8	24	338	117	368	14	12700	3.7	186	41	6.5	14	3.1	1.7	27	374	1.8	14	527	1770
Med.	0.45	8.3	206	66	37	9.7	9555	0.64	18	6.4	1.6	0.89	0.91	0.67	15	13	0.66	5.4	168	336
Std.	0.54	5.7	82	27	114	3.3	2632	0.94	64	12	1.6	3.9	0.98	0.45	6.4	116	0.60	2.7	162	579
<b>Mar2a</b>																				
Min.	0.52	2.7	135	89	1.2	0.8	2120	2.1	35	11	0.42	0.47	1.0	0.23	3.4	6.0	1.0	4.9	186	5.7
Max.	12	16	1169	350	83	3200	29500	219	225	841	5090	215	39	9.0	461	840	38	180	2440	8000
Med.	4.4	5.3	581	243	3.7	4.5	6380	69	139	298	1190	8.0	9.7	0.68	71	206	12	32	697	205
Std.	3.2	2.7	261	72	22	1049	8566	61	42	182	1374	67	8.3	3.2	135	235	12	56	701	1914
<b>Mar2b</b>																				
Min.	0.36	0.18	48	9.9	0.19	2.8	15	0.43	11	0.91	2.0	0.96	0.20	0.11	5.6	0.22	0.03	4.1	2.2	3.6
Max.	5.3	19	1430	253	65	1300	52600	532	980	313	19400	66	189	5.7	402	361	133	160	2530	18800
Med.	1.5	4.0	190	60	6.0	8.0	3235	27	71	58	43	3.8	7.0	1.3	62	38	2.2	13	195	410
Std.	1.4	5.4	272	54	16	503	10791	143	195	79	3906	17	43	1.7	91	91	32	50	533	3710
<b>Mar3</b>																				
Min.	1.3	4.4	348	77	0.40	1.1	8300	0.90	95	150	4.0	1.7	0.50	0.08	16	47	0.13	7.2	359	48
Max.	10	34	2710	369	5.2	180	94300	1510	357	2870	14500	40	59	9.1	92	2370	92	150	8430	41900
Med.	2.6	12	872	146	1.4	20	25400	29	138	290	1030	16	7.7	2.9	56	169	9.7	30	1460	5700
Std.	3.0	7.2	610	63	1.1	44	22709	366	66	634	4139	9.8	14	2.6	21	554	25	37	1747	10101
<b>Mar4</b>																				
Min.	8.4	0.07	325	8.5	1.4	1.6	2500	52	563	401	189	1.3	0.86	0.08	17	619	0.23	5.7	158	32
Max.	15	0.19	797	15	4.9	12	4450	341	2900	955	3560	189	44	3.5	40	1252	29	830	407	15600
Med.	13	0.11	417	13	2.6	5.8	3115	104	779	706	1056	8.2	15	0.68	29	1027	6.7	7.2	258	114

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Std.	2.1	0.03	126	2.2	0.87	3.1	566	67	546	152	990	49	14	1.0	6.4	190	10	199	64	3737
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Fig. 1

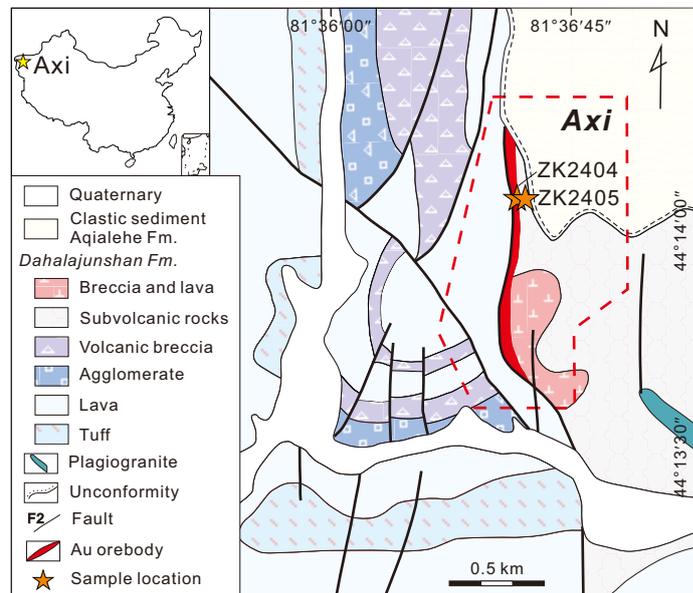


Fig. 2

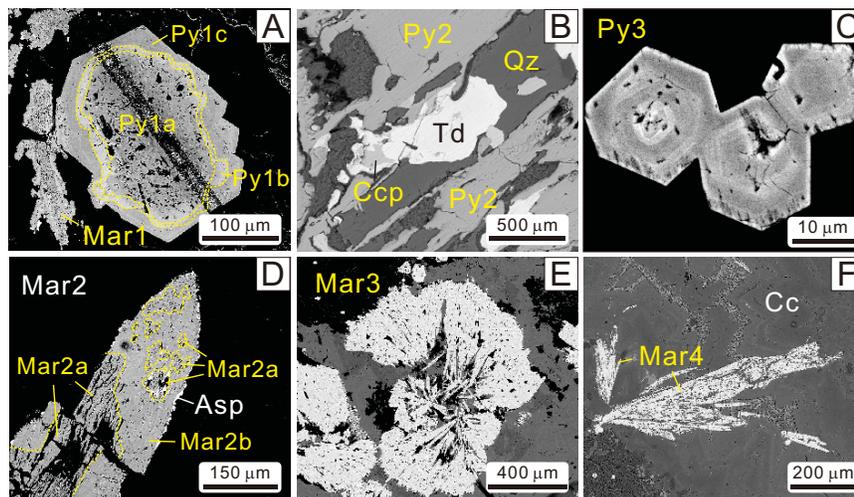


Fig. 3

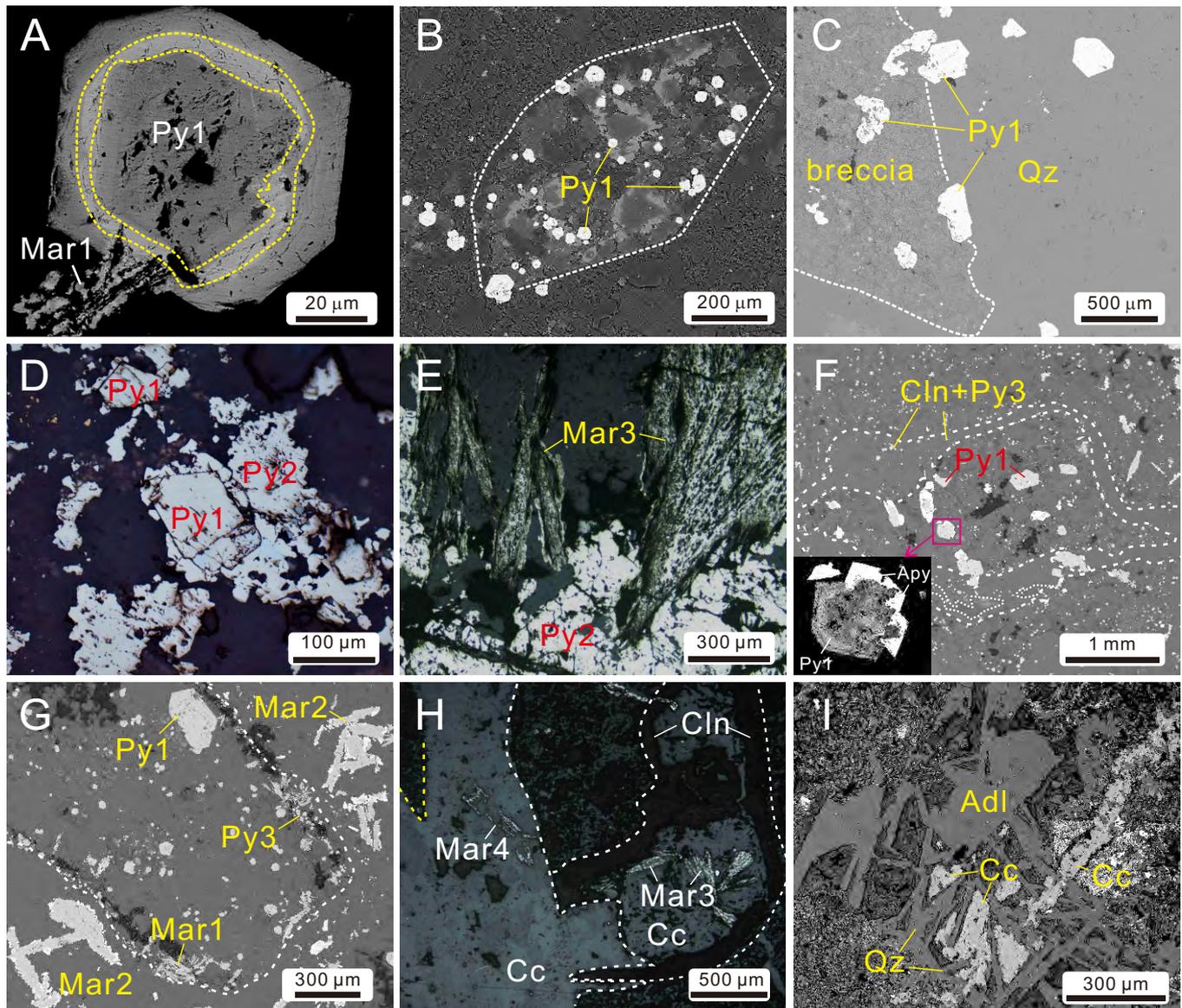


Fig. 4

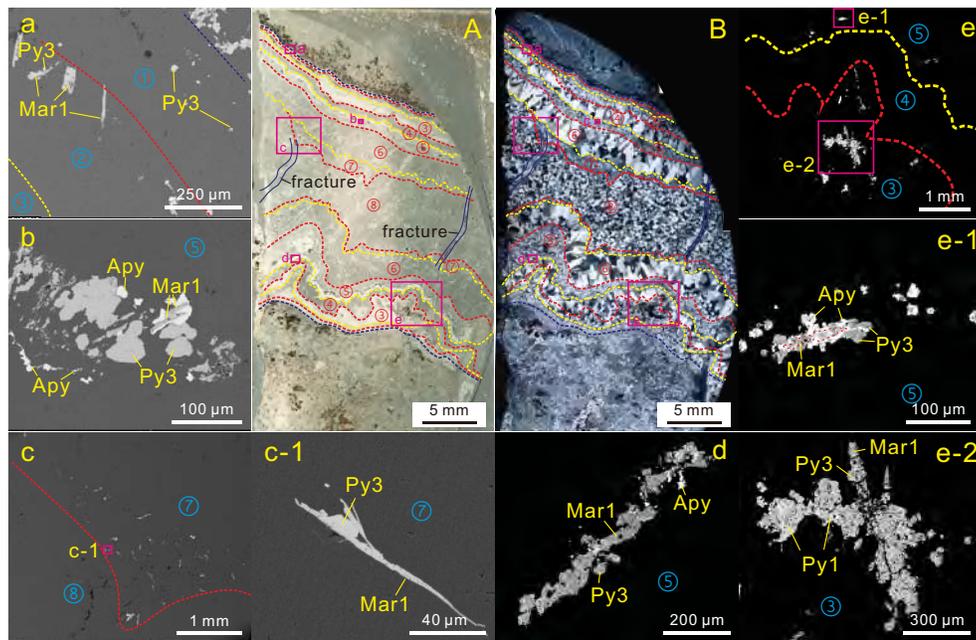


Fig.5

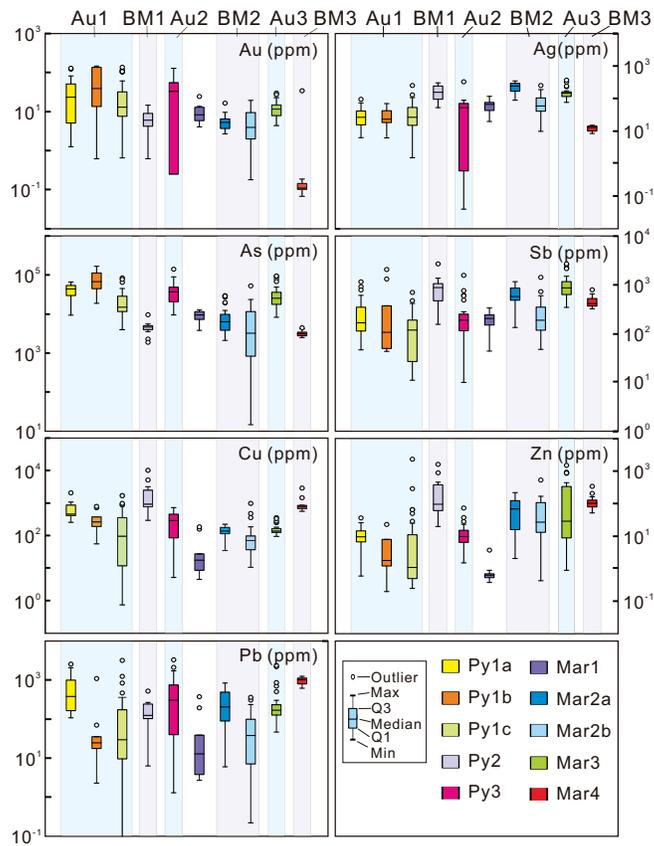


Fig.6

