

1 **Revision 1; Word Count:** 6765 (excluding cover letter, abstract, reference list, tables and figures)

2 Low-temperature crystallography and vibrational 3 properties of rozenite ($\text{FeSO}_4 \cdot 4\text{H}_2\text{O}$), a candidate mineral 4 component of the polyhydrated sulfate deposits on Mars

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16 Abstract:

17 Rozenite ($\text{FeSO}_4 \cdot 4\text{H}_2\text{O}$) is a candidate mineral component of the polyhydrated sulfate
18 deposits on the surface and in the subsurface of Mars. In order to better understand its behavior
19 at temperature conditions prevailing on the martian surface and aid its identification in ongoing
20 and future Rover missions we have carried out a combined experimental and computational
21 study of the mineral's structure and properties. We collected neutron powder diffraction data at
22 temperatures ranging from 21 – 290 K, room temperature synchrotron X-ray data and Raman
23 spectra. Moreover, first-principles calculations of the vibrational properties of rozenite were
24 carried out to aid the interpretation of the Raman spectrum. We found, in contrast to a recent
25 Raman spectroscopic study, that there are no phase transitions between 21 and 290 K. We
26 confirm the heavy atom structure reported in the literature (space group $P2_1/n$) to be correct and
27 present, for the first time, an unconstrained determination of the hydrogen atom positions by
28 means of high-resolution neutron powder diffraction, and report the complete crystal structure at
29 290 K and 21 K. The anisotropy of the thermal expansion of the lattice vectors is $\alpha_a : \alpha_b : \alpha_c =$
30 $1.00 : 2.19 : 1.60$ at 285 K. Subsequent analysis of the thermal expansion tensor revealed highly
31 anisotropic behavior as reflected in negative thermal expansion approximately $\parallel \langle 101 \rangle$ and ratios
32 of the tensor eigenvalues of $\alpha_1 : \alpha_2 : \alpha_3 = -1 : 3.74 : 5.40$ at 285 K. Lastly, we demonstrated how
33 combining Raman spectroscopy and X-ray diffraction of the same sample material sealed inside
34 a capillary with complementary first-principles calculations yields accurate reference Raman
35 spectra. This workflow enables the construction of a reliable Raman spectroscopic database for
36 planetary exploration, which will be invaluable to shed light on the geological past as well as in
37 identifying resources for the future colonization of planetary bodies throughout the solar system.

38
39 **Keywords:** polyhydrated sulfates, Mars, rozenite, negative thermal expansion, Raman
40 spectroscopy, Density Functional Theory, Neutron Diffraction, Synchrotron Diffraction

41 Introduction

42 Ferrous iron sulfate minerals occur in various hydration states, $\text{FeSO}_4 \cdot n\text{H}_2\text{O}$ with $n = 1$,
43 4, 5, 6, 7, most commonly in the vicinity of sulfide ore deposits (Jambor et al. 2000). In such
44 environments, weathering of ore and waste materials releases oxidized iron and sulfur, as well as
45 a wide range of potentially toxic elements, into nearby surface waters (Hudson-Edwards et al.
46 1999; Nordstrom 2011), which subsequently precipitate as hydrous ferric and ferrous sulfates.
47 Widespread occurrences of minerals diagnostic to acid mine drainage environments have been
48 identified on the martian surface using both surface (Klingelhöfer et al. 2004) and orbit (Carter et
49 al. 2013) based measurements. These findings have been interpreted as evidence for the
50 existence of acidic aquatic environments early in martian history (Squyres et al. 2004). In detail,
51 it has been proposed that hydrated sulfate minerals on Mars are secondary minerals produced by
52 the evaporation of fluids involved in the aqueous alteration of martian basalt (Tosca et al. 2005).
53 In the Valles Marineris canyon system sulfate minerals occur in sandy stratified deposits,
54 typically exposed in canyon walls, and lie stratigraphically above widespread clay-mineral
55 deposits (Roach et al. 2010). This stratigraphic sequence records a transition from neutral-
56 alkaline (clay-forming) to acidic (iron sulfate-forming) aqueous environments indicating a
57 change in global climate from wetter to dryer conditions starting around 3.5 Gya (Bibring et al.
58 2006). Due to their apparent role as climatological archives a detailed mineralogical
59 characterization of these sulfate deposits is essential in order to decipher the nature and drivers of
60 changing environmental conditions during the planet's early history. However, attempts to assign
61 any single mineral species to the spectral data acquired for the polyhydrated sulfate deposits have
62 proven to be challenging. This is due to inherent difficulty in differentiating between the diffuse
63 near-IR reflectance spectra of different sulfate phases in the range $0.35 - 5.1 \mu\text{m}$ (corresponding
64 to the OMEGA spectrometer on-board the Mars Express orbiter (Langevin et al. 2006)) that is

65 typically measured from orbit, i.e., significant spectral similarities occur between sulfates of
66 different chemical composition and degrees of hydration (Bishop et al. 2009).

67 Numerous studies have been carried out in order to identify candidate minerals for the
68 polyhydrated sulfate phase. If these deposits indeed originate from the alteration of olivine in
69 martian basalt, the polyhydrated sulfate phase as its weathering product would likely be an iron-
70 or magnesium-bearing sulfate (Bibring et al. 2005) or a sulfate mineral of intermediate
71 composition. Among the iron sulfates, Carter et al. (2013) noted that rozenite matched the data
72 acquired by OMEGA and CRISM very well. Wang et al. (2016), however, observed by studying
73 the phase stability of melanterite ($\text{FeSO}_4 \cdot 7\text{H}_2\text{O}$), under variable temperature and relative
74 humidity conditions, that at 323 K rozenite merely occurs as a transient state and further
75 dehydrates to szomolnokite ($\text{FeSO}_4 \cdot \text{H}_2\text{O}$). Based on this finding they concluded that rozenite is
76 not stable under present day martian surface conditions. Nevertheless, we note that in the same
77 study Wang et al. (2016) reported that even after around 2030 hours (number obtained from
78 supporting information S4 in Wang et al. (2016)) of exposure to relatively dry air (i.e., 33 %
79 relative humidity) rozenite did not reveal any signs of dehydration at temperatures as high as 294
80 K, which is well above a maximum ground temperature of 280 K measured by the Curiosity
81 rover over the first 100 sols of data acquisition (Gómez-Elvira et al. 2014). Moreover,
82 rehydration of szomolnokite to rozenite and melanterite was observed at a temperature of 298 K
83 at a relative humidity of 65 % (Mitchell 1984). Hence, even if temperatures as high as 323 K
84 have prevailed at any point in martian history (leading to the dehydration of rozenite or
85 melanterite), the process may have been reversed under present day martian environmental
86 conditions. Thus, rozenite should still be regarded as a promising candidate mineral for the
87 polyhydrated sulfate phase.

88 In order to unambiguously confirm the absence or presence of rozenite in the martian
89 polyhydrated sulfate deposits, in-situ analytical techniques such as X-ray diffraction or Raman
90 spectroscopy are needed. Raman spectroscopy in particular has proven very effective in
91 discriminating between various sulfate mineral species (Košek et al. 2017), and thus appears well
92 suited to unravel the mineralogical phase composition of the polyhydrated sulfate phase. Raman
93 spectroscopic investigations of the martian soil have recently become possible. NASA's
94 Perseverance Rover, landing on the martian surface in spring 2021, as well as the ESA's
95 Rosalind Franklin Rover, which will likely arrive on Mars in 2023, both feature Raman
96 spectrometers (Rosalind Franklin: The Raman Laser Spectrometer (Rull et al. 2017);
97 Perseverance: SuperCam (Wiens et al. 2020) and SHERLOC (Bhartia et al. 2021)) as part of
98 their scientific payload.

99 Interestingly, Chio et al. (2007) observed an apparent splitting of Raman-active
100 vibrational modes of rozenite, which they suggested might be indicative of two structural phase
101 transitions ($T_{\text{crit}1} = 240 - 190$ K, $T_{\text{crit}2} = 140 - 90$ K). The first transition is well within the range
102 of temperatures relevant to the martian surface. Although Raman spectroscopy is a powerful tool
103 for the identification of polymorphic phase transitions, sharpening of closely-spaced vibrational
104 modes at low-temperatures might be erroneously interpreted as splitting due to a phase transition.

105 With rozenite being one of the most promising candidates for the polyhydrated sulfate
106 phase on the martian surface, there is an interest in revisiting these putative phase transitions by
107 means of neutron diffraction in order to determine the structural stability of rozenite at
108 temperatures relevant to the martian surface as well as to solve the crystal structure of any
109 proposed low-temperature polymorphs. For this reason, we have studied herein the structural
110 stability of $\text{FeSO}_4 \cdot 4\text{D}_2\text{O}$ at temperatures ranging from 290 K down to 21 K by means of high-

111 resolution time-of-flight (TOF) neutron diffraction. Furthermore, the samples used for the
112 neutron diffraction have been perdeuterated for the explicit reason of avoiding the large
113 incoherent scattering signal from ordinary ^1H , which produces a substantial background masking
114 weak Bragg peaks. Use of ^2D virtually eliminates this background and allows for rapid
115 acquisition of high-quality data.

116 Since subtle changes in relative humidity may cause hydrated sulfates to partially or
117 completely re- or dehydrate (Wang et al. 2016), it is entirely possible that the rozenite sample
118 studied by Chio et al. (2007) underwent a transformation in between the initial phase
119 identification by X-ray diffraction and the subsequent measurements of its vibrational properties.
120 In order to confirm that the Raman spectra reported by Chio et al. (2007) indeed corresponds to
121 rozenite, we have carried out Raman spectroscopy and synchrotron X-ray diffraction, which is
122 able to detect the smallest amounts (i.e., 0.1 wt% (Thompson et al. 2009)) of any contaminant
123 phases. Such measurements have been complemented with ab initio calculations in order to
124 further explore the vibrational properties of rozenite. This combined experimental and theoretical
125 approach allows us to cast light on the structural stability of rozenite at martian surface
126 temperatures and, therefore, will aid the identification of rozenite in future and ongoing Mars
127 missions.

128 Methods

129 Synthesis and phase analysis

130 $\text{FeSO}_4 \cdot 7\text{H}_2\text{O}$ (Sigma Aldrich ACS reagent grade) was dehydrated under vacuum for 48
131 hours at 473 K to prepare anhydrous FeSO_4 . A hot supersaturated solution of FeSO_4 in 0.5 M
132 D_2SO_4 (Sigma Aldrich > 99 atom % D) was then cooled to room temperature in order to
133 precipitate fine-grained $\text{FeSO}_4 \cdot 7\text{D}_2\text{O}$. The sealed container was stored for several years, during
134 which time diurnal and seasonal temperature changes resulted in the fine-grained material

135 altering into a fully dense coarse crystalline material. $\text{FeSO}_4 \cdot 7\text{D}_2\text{O}$ crystals were then ground to a
136 powder under helium in order to prevent exchange of ^2D with atmospheric ^1H . The powder was
137 loaded into rubber-sealed glass jars containing a saturated solution of MgCl_2 in D_2O (Sigma
138 Aldrich > 99 atom % D), which buffered the relative humidity at 33 % (Greenspan 1977), and
139 kept at 280 K for 3 days. The sample was then transferred, again under helium, into the sample
140 holder used for the neutron diffraction experiments. This comprised an aluminum frame
141 surrounding a cuboid sample cavity of dimensions $18 \times 23 \times 10$ mm ($w \times h \times d$), open at the
142 front and back. The rear opening was first covered with a vanadium foil window sealed by
143 indium wire; powder was then transferred into the sample holder and the front opening was
144 sealed with another vanadium foil window. Gadolinium and cadmium foils were used to mask
145 scattering from various aluminum and steel parts that might be exposed to the incident neutron
146 beam around the edges of the vanadium windows. A cartridge heater, used for temperature
147 control, and a RhFe thermometer were inserted into the aluminum frame of the sample container
148 either side of the sample.

149 The sample was mounted into a closed cycle refrigerator (CCR) at the High-Resolution Powder
150 Diffraction (HRPD) beamline, ISIS Neutron and Muon Spallation Facility, UK. This instrument
151 allows collection of neutron time-of-flight (TOF) data in various 100 ms-wide ‘windows’; those
152 used in this study were 30 – 130 ms and 100 – 200 ms. In HRPD’s highest resolution
153 backscattering detectors ($2\theta = 154 - 176^\circ$), these yield diffraction patterns covering d -spacings
154 from 0.65 – 2.60 Å and 2.20 to 3.90 Å, respectively. The latter is often best used for rapid phase
155 identification, even though the incident neutron flux is very low, since Bragg peaks are typically
156 fewer and better dispersed, whilst also being rather more intense than those at shorter d -spacings.

157 We could thus quickly confirm that the sample was mostly rozenite, but with a few weak
158 Bragg peaks identified as melanterite, indicating that dehydration of the starting material was
159 incomplete. Heating the sample to 305 K for 60 min and then 315 K for a further 20 minutes
160 resulted in the complete transformation of melanterite to rozenite. At 315 K, however, the high
161 temperatures also gave rise to slight further dehydration and the formation of szomolnokite.
162 Therefore, we lowered the temperature again to 305 K and kept the sample at this temperature
163 for a further 12 minutes to ensure the complete transformation from melanterite to rozenite in the
164 final sample material. The diffraction patterns acquired during dehydration of the sample are
165 provided in the supplementary data (Fig. s1). Rietveld refinement of the neutron diffraction
166 pattern acquired at 290 K suggests that the final sample material consists of 93.5(1) wt% rozenite
167 and 6.5(1) wt% szomolnokite.

168 Protiated $\text{FeSO}_4 \cdot 4\text{H}_2\text{O}$ for the synchrotron X-ray diffraction and Raman analysis was
169 synthesized using the dehydration method stated above. We placed the reagent $\text{FeSO}_4 \cdot 7\text{H}_2\text{O}$ in a
170 rubber sealed glass jar containing a saturated solution of MgCl_2 in H_2O for 3 days at around 290
171 K. Subsequently, the sample was loaded into a borosilicate glass capillary of 0.5 mm diameter.
172 The loaded specimen and the rest of the sample material were used for Raman analysis.

173

174 TOF Neutron diffraction and Rietveld refinement

175 High resolution TOF neutron diffraction data were collected at temperatures ranging
176 from 315 to 21 K upon cooling and warming. After the initial dehydration at 305 and 315 K, the
177 sample was cooled to 290 K. Datasets with long counting times of 3 h 17 min in the 30 – 130 ms
178 and 100 – 200 ms TOF window were then collected to allow crystal structure refinement. In
179 order to characterize the thermal expansion of rozenite, diffraction patterns with shorter counting

180 times of 50 min were collected on cooling in 10 K increments in the 100 – 200 ms TOF window
181 in the temperature range from 290 – 21 K. At the base temperature of 21 K, another dataset for
182 structural refinement was acquired in the 100 – 200 ms TOF window, again, for 3 h 17 min and
183 in 30 – 130 ms window for 4 h 6 min. Lastly, we collected diffraction patterns upon heating from
184 35 – 285 K in 10 K increments for 37 min each. To ensure good thermal equilibrium between the
185 heated aluminum frame of the sample holder and the powder sample itself, temperature changes
186 were done at 3 K min⁻¹ and a dwell time of 10 minutes after reaching a set-point was used prior
187 to the start of data collection.

188 The data were time-focused, normalized to the incident spectrum and corrected for
189 instrument efficiency using a V:Nb standard. Diffractometer calibration constants and
190 instrumental peak-profile coefficients were determined using NIST silicon SRM640e and CeO₂
191 standards.

192 All refinements were carried out using the Rietveld method (Albinati and Willis 1982;
193 van Laar and Schenk 2018) as implemented in the GSAS/EXPGUI software suite (Toby 2001).
194 The diffraction patterns collected in HRPD's backscattering detector bank at 290 K and 21 K are
195 displayed in Fig. 1 and the refinement parameters are given in Tab. 1 as well as in the CIF.
196 Clearly, the refined structural models are in excellent agreement with the observed intensity data.

197 Unit-cell parameters were refined from the shorter 100 – 200 ms TOF datasets using
198 initial least-squares cycles of Rietveld refinement, followed by a series of least-squares cycles
199 using the 'F(calc) weighted' method; this workflow typically results in the most precise lattice
200 parameters by virtue of fitting the intensities more accurately. For szomolnokite we used the
201 structure reported by Talla and Wildner (2019) for the refinements and varied the lattice
202 parameters and profile coefficients at each temperature. The small number of weak peaks

203 accessible in the 100 – 200 ms TOF window combined with the low symmetry of szomolnokite
204 precludes us from following the evolution of the lattice parameters accurately as a function of
205 temperature, which are therefore not reported in this study.
206

207
208 **Synchrotron X-ray diffraction and room temperature Raman**
209 **spectroscopy**

210 A synchrotron X-ray powder diffraction pattern of the protiated rozenite capillary sample
211 was collected at the I11 instrument, Diamond Light Source, UK using the Multi Analyzing
212 Crystal detectors (Thompson et al. 2009). The experiment was carried out at an ambient
213 temperature of 295 ± 0.5 K. The wavelength of $0.826574(9)$ Å and a zero-point error of
214 $0.000315(2)^\circ$ was determined from a *NIST SRM 640c* silicon standard.

215 Raman spectra of protiated rozenite were acquired using a *B&W Tek i-Raman Plus*
216 spectrometer equipped with a neodymium doped Yttrium-Aluminum-Garnet laser that was
217 frequency-doubled to 532 nm (maximum power of 30 mW as determined by the manufacturer).
218 We collected data over the entire spectral range accessible to the spectrometer (i.e., 65 – 4200
219 cm^{-1}) with a resolution of smaller than 3.5 cm^{-1} (as determined by the manufacturer at 614 nm)
220 on protiated rozenite powder samples both outside and inside of the very same borosilicate glass
221 capillary that was used for the synchrotron diffraction analysis. The spectra were collected for 53
222 and 50 seconds and averaged over three acquisitions for the sample outside and inside of the
223 capillary, respectively.

224
225 **Ab initio vibrational properties from density functional**
226 **theory (DFT):**

227 We performed Kohn-Sham DFT calculations (Hohenberg and Kohn 1964; Kohn and
228 Sham 1965) using ultrasoft pseudopotentials from the *GBRV* library (Garrity et al. 2014). Kinetic
229 energy cut-off values and *k*-point grid density were derived from convergence testing; values of
230 70 Ry and 840 Ry were adopted for the wave function and charge density cut-offs, respectively

231 and a Monkhorst–Pack (Monkhorst and Pack 1976) k -point grid of $2 \times 1 \times 2$ was applied to
232 sample the Brillouin zone.

233 The crystal structure as reported by Baur (1962) served as input geometry for an initial
234 relaxation using the PBE functional (Perdew et al. 1996) in conjunction with the D2 dispersion
235 correction (Grimme 2006). We have demonstrated in our previous work (Meusburger et al.
236 2021) that the PBE + D2 approach is very well suited to model geometries at temperatures close
237 to the ground state both for dispersion-dominated solids as well as hydrogen-bonded solids such
238 as the one under investigation.

239 The unit-cell and internal atomic coordinates were relaxed using the Broyden-Fletcher-
240 Goldfarb-Shanno algorithm (Pfrommer et al. 1997), with convergence thresholds 1×10^{-6} Ry,
241 1×10^{-5} Ry/Bohr, 5×10^{-1} kbar for the total energy, forces, and pressure, respectively. Moreover,
242 we applied a mixing factor of 0.3 and the local Thomas–Fermi charge mixing mode (Raczkowski
243 et al. 2001) in order to achieve convergence in the self-consistent field cycles.

244 The structure optimized at the DFT + D level served as basis for the self-consistent
245 computation of the Hubbard U by means of Density Functional Perturbation Theory (Timrov et
246 al. 2018). The formulation of the Hubbard model following Dudarev et al. (1998) as well as
247 nonorthogonalized atomic orbitals (Cococcioni and de Gironcoli 2005; Amadon et al. 2008) as
248 projectors for the strongly localized $3d$ states of Fe^{2+} were used for all DFT + U calculations. We
249 have tested multiple q -meshes and found that a grid of $2 \times 1 \times 2$ is necessary in order to derive
250 Hubbard U values converged to within 0.0007 eV. Subsequently, we used the U value converged
251 with respect to the q -mesh (i.e., 6.9806 eV), and again, computed the U value from this optimized
252 structure. This procedure was repeated three times until the U value (i.e., 6.0156 eV) was
253 identical with respect to the previous iteration.

254 The geometry obtained from this workflow was then used as input for our phonon
255 calculations at the Γ -point using Density Functional Perturbation Theory in order to compute the
256 Infrared (IR) and Raman spectra of rozenite. The formalism of phonon calculations at the DFT +
257 U level of theory as implemented in Quantum ESPRESSO's ph.x code is presented in detail in
258 Floris et al. (2011, 2020). The open-source Quantum ESPRESSO code suite (Giannozzi et al.
259 2009, 2017) was used for all ab initio calculations. All files necessary to reproduce our
260 calculations (i.e., input, output, and pseudopotential files) may be accessed on the
261 MaterialsCloudArchive (Meusburger et al. 2022).

262 Results and discussion

263 Complete crystal structure and hydrogen bonding of $\text{FeSO}_4 \cdot 4\text{D}_2\text{O}$ at 290 264 K

265 The crystal structure of rozenite was first determined by Baur (1962) and consists of
266 isolated cyclic $[\text{Fe}(\text{H}_2\text{O})_4\text{SO}_4]_2$ units (Fig. 2a) interconnected by a hydrogen bond network (Fig.
267 2b). $[\text{Fe}(\text{H}_2\text{O})_4\text{SO}_4]_2$ may be further divided into $\text{Fe}(\text{H}_2\text{O})_4\text{O}_2$ octahedral units that engage in
268 corner-sharing of O atoms with the SO_4 tetrahedral units (Fig. 2a).

269 Considering that the crystal structure of rozenite has not been revisited since its initial
270 determination 60 years ago by means of single crystal X-ray diffraction, the results presented in
271 this study improve the literature data substantially. In contrast to X-ray diffraction, which
272 severely underestimates O – H bond lengths (Baur 1972), neutron diffraction is particularly well
273 suited to accurately locate the hydrogen atom positions, allowing us to resolve long-standing
274 ambiguities in the hydrogen bonding network of rozenite. In the following section the geometry
275 of the octahedral and tetrahedral units will be discussed and compared to the structure reported
276 by Baur (1962) (Tab. 2).

278 Overall, the heavy atom (i.e., Fe, S, O) structure refined from the 290 K dataset is in
279 excellent agreement with the structure reported by Baur (1962). This is reflected by the Mean
280 Unsigned Differences (MUD)¹ of 0.011 Å (Fe – O bonds), 0.605 ° (O – Fe – O angles), 0.016 Å
281 (S – O bonds) and (O – S – O angles) 0.76 ° being either well below or very close to the
282 estimated standard deviations reported by Baur (1962) on the respective quantities.

283 Moving on to the hydrogen bond network it is noteworthy that Baur (1962) determined
284 the heavy atom structure of rozenite and did not refine the position of the hydrogen atoms.
285 Instead, he fixed the hydrogen atoms to the positions determined for the isotopic material
286 MgSO₄·4H₂O (starkeyite) during the refinement process. As noted above, Baur (1962)
287 determined the hydrogen atom positions by means of X-ray diffraction, which typically yields
288 underestimated bond lengths as well as inaccurate H – O – H angles (Baur 1972). Baur (1962)
289 counteracted this shortcoming by setting multiple constraints upon the hydrogen bond
290 geometries. Therefore, we present for the first time the complete crystal structure of rozenite
291 with all atomic positions derived from an unconstrained refinement. We assess the differences in
292 the hydrogen bonding network and compare our results to the Baur (1962) study (Tab. 3).

293 The O ••• O contacts are in excellent agreement (MUD = 0.0168 Å, largest individual
294 difference (LID) = 0.043 Å), which was expected since the distance of donor acceptor oxygens
295 of the hydrogen bonding network is solely determined by the heavy atom positions. Despite the
296 fact that X-ray diffraction (XRD) underestimates O – H bond lengths, the covalent O – H bond
297 distances in our neutron diffraction and the Baur (1962) XRD study agree very well (MUD =
298 0.020 Å, LID = 0.063 Å). The reason for this is likely that the bond-length constraints for O – H

¹ The MUD is defined as $\frac{\sum_i^n |x_B - x_T|}{n}$ with x_B and x_T being the values as observed by Baur and in this study, respectively for the quantities of interest (i.e., bond-length and angle).

299 bonds as applied by Baur (1962) are very close to the values we observed in rozenite. Although
300 the constraints used by Baur (1962) yielded H – O distances in close agreement with our
301 refinement, the constrained model fails for the geometry of the H – O – H molecules, as reflected
302 by a large MUD of 1.05 ° and LID of 1.80 °. Moreover, the H ••• O contacts revealed a LID of
303 0.209 Å. This failure to accurately describe the hydrogen bonding geometry is most pronounced
304 in the O_w – H ••• O angles which exhibit a MUD of 9.51 ° and LID of up to 30.80 °, reinforcing
305 the importance of neutron diffraction data in order to derive accurate geometries for the complete
306 crystal structure.

307 The hydrogen bonding system in rozenite-type compounds has been subject of intensive
308 discussions (Baur 1962, 1964, 2002; Kellersohn 1992; Held and Bohaty 2002; Anderson et al.
309 2012). Based on a long donor acceptor distance of 3.02 Å Baur (1962) suggested that no
310 intermolecular hydrogen bonding takes place for the Ow2 – H2b ••• O2 contact (Fig. 3a). More
311 recent studies on the rozenite-type compounds ZnSO₄·4D₂O (Anderson et al. 2012),
312 MnSO₄·4D₂O (Held and Bohaty 2002; Anderson et al. 2012) and CoSO₄·4D₂O (Kellersohn
313 1992) interpret H2b to partake in a three-centered interaction (i.e., a bifurcated H-bond, Ow2 –
314 H2b :: O2/O2') with long donor-acceptor distances of 3.02 Å and 3.26 Å (Fig. 3b). Kellersohn
315 (1992) investigated the Ow2 hydrogen bond system based on bond-valance considerations and
316 noted that exclusion of the two long hydrogen bonds results in the O2 atom exhibiting a
317 deficiency of 0.22 valence units. Our own bond valance calculations (Brown and Altermatt 1985;
318 Brese and O'Keeffe 1991; Alig et al. 1994) for rozenite (supplementary information; Tab. s1)
319 yield a deficiency of 0.10 valence units on the O2 atom, reducing to 0.054 if the intermolecular
320 contacts are included in the calculations. Clearly, the long intermolecular hydrogen bonds play a
321 vital role in achieving charge neutrality for the O2 atom and thus should not be neglected when

322 describing the hydrogen bond network. Supporting this evidence in favor of a three-centered
323 hydrogen bond in rozenite-type compounds, Anderson et al. (2012) suggested, based on an
324 exhaustive comparison with numerous hydrogen bonding systems, that not only the H2b, but
325 also the H3a atom is involved in a three-centered hydrogen bond (Ow3 – H3a ::: O4/Ow4) (Fig.
326 3c). According to our bond valance calculations the H3a ••• Ow4 contact exhibits a valency of
327 0.017, making it a weak but non-negligible hydrogen bond. Furthermore, when assessing the
328 geometry of Ow3 – H3a ••• Ow4 contact (i.e., H3a ••• Ow4 = 2.425(5) Å; \angle (Ow3 - H3a •••
329 Ow4) = 124.0(3) ° at 290 K) it is evident that this bond falls well within even the conservatively
330 defined limits of hydrogen bonding (i.e., H ••• O < 3 Å; \angle (Ow - H ••• O) = 110 ° - 180 °;
331 (Steiner 2002)). Therefore, based on the large range of evidence in favor of three-centered
332 hydrogen bonding involving the H3a and H2b atoms, we have adopted the hydrogen bonding
333 system proposed by Anderson et al. (2012) (Fig. 3c) for our study.
334

335 Low temperature crystallography, thermal expansion, and absence of 336 phase transition

337 The powder diffraction data revealed no splitting of Bragg peaks nor appearance of new
338 peaks that might be indicative of a phase transition. Consequently, the refined lattice parameters
339 also varied smoothly (Fig. 4, Tab. s3, Tab. s4). It is noteworthy that the β angle decreased
340 systematically on cooling such that, at ~ 137 K, the unit-cell became metrically orthorhombic.
341 However, β continued to decrease $< 90^\circ$ below 137 K and it is important to emphasize that the
342 Laue symmetry of the diffraction pattern retains its monoclinic character throughout the whole
343 temperature range under investigation. Thus, rozenite does not undergo any structural phase
344 transition down to at least 21 K.

345 There are various approaches to quantify the temperature dependency of the lattice
346 parameters of crystalline solids, with varying degrees of usefulness and capability for accurate
347 extrapolation. These range from polynomial fits through to parameterizations based on the
348 classical Einstein model of the internal energy and the more accurate Debye-type model, which
349 is used here.

350 The Debye model is derived from the Grüneisen relation of the thermoelastic properties.

$$351 \quad \gamma = \frac{\alpha_V V K_0}{C_V} \quad (1)$$

352 where γ is the Grüneisen ratio, α_V is the volume thermal expansion coefficient, K_0 is the
353 isothermal bulk modulus, C_V is the isochoric specific heat capacity, and V the unit-cell volume.
354 Equation (1) is integrated with respect to the temperature, whereby γ and K_0 are assumed to be
355 independent of the temperature. This yields the first order approximation to the Grüneisen zero
356 pressure equation of state which expresses the thermal expansion of the crystal as a function of
357 its internal energy $U(T)$ and isothermal bulk modulus K_0

358
$$V(T) = V_0 \left(1 + \frac{U(T)}{Q} \right) \quad (2)$$

359 where $Q = (V_0 K_0 / \gamma)$.

360 $U(T)$ may be derived by the Debye approximation

361
$$U(T) = 9Nk_B T \left(\frac{T}{\theta_D} \right)^3 \int_0^{\theta_D/T} \frac{x^3}{e^x - 1} dx \quad (3)$$

362 where N is the number of atoms per formula unit, k_B is Boltzmann's constant, θ_D is the Debye
363 temperature, and $x = \hbar\omega/k_B T$. Note that the zero-point energy of $9Nk_B\theta_D/8$ is included in
364 Equation (2) via the term V_0 .

365 It was apparent from the residuals obtained from fitting this model to $V(T)$ that the first-
366 order Grüneisen model failed to provide an accurate description of the thermal expansion (see
367 supplementary Fig. s2). Consequently, we fitted a second-order approximation to the Grüneisen
368 zero pressure equation of state (Equation (4)), which introduces the first derivative of the bulk
369 modulus K_0' with respect to pressure via the parameter b

370
$$V(T) = V_0 \left(1 + \frac{U(T)}{Q - bU(T)} \right) \quad (4)$$

371 where $b = \frac{1}{2} (K_0' - 1)$.

372 The sum of squared residuals is thus reduced from 0.32 \AA^6 for the first-order fit to 0.014
373 \AA^6 for the second order model.

374 The value of K_0/γ obtained from the second-order model is $84.5(4)$ GPa, increased from
375 $K_0 = 58.4(7)$ GPa in the poorer first-order fit. We note that Meusburger et al. (2019) obtained a
376 bulk modulus of $45.2(2)$ GPa for szomolnokite ($\text{FeSO}_4 \cdot \text{H}_2\text{O}$). It is highly unlikely that rozenite,
377 featuring isolated cyclic $[\text{Fe}(\text{H}_2\text{O})_4\text{SO}_4]_2$ units loosely interconnected by intermolecular
378 hydrogen bonds, exceeds the stiffness of szomolnokite featuring a rigid three dimensional

379 framework structure. Fortes et al. (2006) report a bulk modulus of 21.5(1) GPa for epsomite
380 ($\text{MgSO}_4 \cdot 7\text{H}_2\text{O}$), a mineral featuring isolated $\text{Mg}(\text{H}_2\text{O})_6$ and SO_4 units loosely interconnected by
381 hydrogen bonds. We should expect, therefore that rozenite will exhibit a compressibility
382 intermediate between these hydrates, implying $\gamma = 1.9 - 4.0$. We further note that an extremely
383 large value for K' of 41.1 as obtained from the second-order Debye fit clearly is unphysical but
384 typical of the case where the complexity of the underlying phonon spectrum is represented by a
385 simple model characterized with a single Debye frequency cut-off. The parameters derived from
386 the second order single Debye fits are given in [Tab. 4](#) and from the first order single Debye
387 model in the supplementary material ([Tab. s2](#)). In order to be dimensionally correct, the
388 individual lattice parameters were fitted as a^3 , b^3 and c^3 .

389 For a more detailed assessment of the thermal expansion of rozenite the orientation and
390 magnitude of the principal axes ([Fig. 5a](#)), and components ([Fig. 5b](#)) of the thermal expansion
391 tensor were calculated using the Win_strain software (Angel 2011). Strain tensors were
392 computed incrementally between consecutive data points using the finite Lagrangian strain
393 definition. The reference temperatures correspond to the average of the temperatures of the two
394 data points used to calculate the strains. The thermal expansion tensor is constrained to have the
395 α_2 principal axis parallel to [010], and α_1 and α_3 lying in the (010) plane. The orientation of the
396 crystallographic axes relative to the cartesian reference system as specified by the Institute of
397 Radio Engineers (i.e., X || a^* ; Z || c) was used for all calculations.

398 Cross-sections through the representation surface of the thermal expansion tensor,
399 evaluated at 285 K ([Fig. 6](#)), allow us to understand the relationship to both the monoclinic lattice
400 vectors and the underlying structural elements. The expansion parallel with the a and c directions
401 of the crystal are similar. The thermal expansion || b , however, is substantially larger (i.e., $\alpha_a : \alpha_b :$

402 $\alpha_c = 1.00 : 2.19 : 1.60$ at 285 K). This behavior may be attributed to a large structural ‘void’ and
403 the absence of hydrogen bonding between the rigid cyclic dimer units $\parallel b$ enabling the structure
404 to undergo greater expansion in this direction upon heating (Fig. 6a).

405 Mapping of the thermal expansion onto the crystallographic reference frame conceals the
406 true extent of the anisotropy, manifested in the eigenvalues and eigenvectors of the expansion
407 tensor, i.e., the principal thermoelastic strains, $\alpha_1 : \alpha_2 : \alpha_3 = -1.00 : 3.74 : 5.40$ at 285 K. The
408 principal axis α_1 , which is approximately oriented parallel to the $\langle 101 \rangle$ direction ($a \angle 40.6(3)^\circ$ b
409 $\angle 90^\circ$ $c \angle 49.8(3)^\circ$), exhibits negative thermal expansion over the entire temperature range
410 under investigation. Contrary to this behavior, α_3 , oriented approximately $\parallel \langle 10\bar{1} \rangle$ ($a \angle 49.4(3)^\circ$
411 $b \angle 90^\circ$ $c \angle 139.8(3)^\circ$) is the true direction of maximum thermal expansion. The stark contrast
412 between the negative thermal expansion approximately $\parallel \langle 101 \rangle$ and the direction of maximum
413 expansion approximately $\parallel \langle 10\bar{1} \rangle$ originates from the evolution of β with temperature. Shrinkage
414 of β , the angle spanning the $\langle 100 \rangle$ and $\langle 001 \rangle$ directions, is mirrored by a simultaneous increase
415 of the complementary angle spanning the $\langle 100 \rangle$ and $\langle 00\bar{1} \rangle$ directions (Fig. 6b). From a
416 structural perspective an angle β close to 90° results in adjacent cyclic dimers being almost
417 perfectly stacked on top of each other both in the a and c direction, despite this is not being a
418 requirement imposed by symmetry. The evolution of β may then be explained as being due to a
419 subtle displacement of the cyclic dimers relative to each other as a function of temperature (Fig.
420 6b). Furthermore, the distances between the central atoms of the diagonally opposing
421 $\text{FeO}_2(\text{H}_2\text{O})_4$ and SO_4 units (Fig. 6b), are increasing upon cooling from $9.91385(3) \text{ \AA}$ (290 K) to
422 $9.92414(3) \text{ \AA}$ (21 K) in the $\langle 101 \rangle$ direction (i.e., approximately $\parallel \alpha_1$) and shrinking from
423 $9.98531(3) \text{ \AA}$ (290 K) to $9.90118(3) \text{ \AA}$ (21 K) in the $\langle 10\bar{1} \rangle$ (i.e., approximately $\parallel \alpha_3$). Lastly, the

424 relatively smaller thermal expansion in the a than the c direction (Fig. 6c) may be attributed to a
425 denser population of the dimer units in this direction.

426 Comparing these findings to other hydrated sulfate minerals it is evident that both large
427 thermoelastic anisotropy and negative linear thermal expansion in particular is a rather common
428 phenomenon in such materials. Negative linear expansion has been observed in epsomite
429 ($\text{MgSO}_4 \cdot 7\text{H}_2\text{O}$) (Fortes et al. 2006), meridianiite ($\text{MgSO}_4 \cdot 11\text{H}_2\text{O}$) (Fortes et al. 2008), and most
430 recently several members of the kieserite group ($\text{M}^{2+}\text{SO}_4 \cdot \text{H}_2\text{O}$ with $\text{M}^{2+} = \text{Mg, Fe, Co, Ni}$)
431 (Wildner et al. 2022). The negative thermal expansion found in our study is smaller than in
432 $\text{MgSO}_4 \cdot 7\text{H}_2\text{O}$, $\text{MgSO}_4 \cdot \text{H}_2\text{O}$, $\text{FeSO}_4 \cdot \text{H}_2\text{O}$, $\text{CoSO}_4 \cdot \text{H}_2\text{O}$ and of approximately the same magnitude
433 as in $\text{NiSO}_4 \cdot \text{H}_2\text{O}$ and $\text{MgSO}_4 \cdot 11\text{H}_2\text{O}$ (Tab. 5). The volume thermal expansion of rozenite is
434 intermediate between values found in the higher hydrates (e.g., $\text{MgSO}_4 \cdot 7\text{H}_2\text{O}$ and
435 $\text{MgSO}_4 \cdot 11\text{H}_2\text{O}$) and in the lower hydrates (e.g., $\text{FeSO}_4 \cdot \text{H}_2\text{O}$ and $\text{MgSO}_4 \cdot \text{H}_2\text{O}$) (Tab. 5).

436 These trends in the volume thermal expansion reflect the changing degree of
437 polymerization of the $\text{M}^{2+}\text{O}_x(\text{H}_2\text{O})_{6-x}$ and SO_4 units as a function of hydration state, n , changing
438 from isolated polyhedra ($n = 6$ to 11) through cyclic dimers and chains ($n = 3$ to 5) to rigid 3D
439 corner-sharing frameworks ($n = 1$).

440 Vibrational properties of rozenite and absence of phase transition

441 Hydrated sulfate minerals may de- or rehydrate under changing relative humidity
442 conditions. Thus, it is possible that a sample or a fraction might transform after its identity and
443 phase purity is confirmed by X-ray diffraction and any subsequent measurements may not reflect
444 the original state. Fig. 7(a, b) displays a comparison of the Raman spectra as observed by Chio et
445 al. (2007) and in this study, both inside and outside the glass capillary. Although the background
446 is increased for the measurement performed through the borosilicate glass, a comparison with

447 data collected outside of the capillary demonstrates that no additional sharp bands that might be
448 mistaken for Raman-active vibrational modes appear in the spectrum (Fig. 7a).

449 Synchrotron X-ray diffraction analysis performed on the capillary sample confirms the
450 phase purity of the material (Fig. 7c). As for the data collected by Chio et al. (2007), these
451 authors noted that a shoulder at 1018 cm^{-1} observed in their Raman spectra likely originate from
452 szomolnokite impurities. A comparison of the Raman spectra collected in this study in and
453 outside of the glass capillary shows the absence of the 1018 cm^{-1} band for phase pure rozenite
454 (Fig. 7b), thus confirming this hypothesis. For this reason, acquiring Raman spectra and
455 performing X-ray diffraction on the same capillary, appears to be a suitable workflow in order to
456 produce reliable fingerprint spectra for salt hydrate phases whose hydration state is susceptible to
457 changes in relative humidity.

458 The neutron diffraction results clearly demonstrate the absence of any structural phase
459 transition down to at least 21 K. This raises the question of how the vibrational mode splitting
460 observed by Chio et al. (2007) may be explained. To understand this problem, we have computed
461 the spatial frequencies of the Raman-active vibrational modes of rozenite from first principles
462 (Fig. 8).

463 DFT predicts a total of 16 Raman-active vibrational modes in the region of the H_2O
464 stretching modes (Fig. 8a). Therefore, assuming the occurrence of vibrational mode splitting and
465 consequently a structural phase transition is not necessary in order to explain the 10 bands
466 evidenced by Chio et al. (2007) at temperatures lower than 90 K. We further note that our
467 computation suggests that several of the water stretching modes are overlapping; thus merely 10
468 instead of the total of 16 Raman-active modes are resolvable even at temperatures as low as 8 K.

469 The same holds true for the putative splitting of a mode located in the spectral region of
470 the external vibrations (Fig. 8b) where our ab initio calculations also predict two Raman-active
471 vibrational modes as closely spaced as 2.9 cm^{-1} . Therefore, we conclude that the putative
472 splitting of vibrational modes is a misinterpretation of the sharpening of these modes upon
473 cooling.

474 Both our DFT calculations as well as group theoretical analysis within the C_{2h} point
475 group using the SAM tool (Kroumova et al. 2003) as implemented in the Bilbao Crystallographic
476 Server (Aroyo et al. 2006) predict a total of 108 Raman active modes ($54A_g + 54B_g$). Since first
477 principle calculations at the Γ -point also yield the frequency of the IR-active modes as well as of
478 the acoustic phonon modes we further note that rozenite exhibits 105 IR active optical modes
479 ($53A_u + 52B_u$) as well as 3 ($A_u + 2B_u$) acoustic phonon modes, totaling to 216 phonon modes of
480 which 213 may be assigned to the optical branch and 3 to the acoustic one. The frequency and
481 symmetry for each of the Raman (Tab. s5) and IR-active (Tab. s6) vibrational modes computed
482 from first principles are provided in the supplementary material. Ruggiero et al. (2015) have also
483 modelled the vibrational properties of rozenite from first principles and curiously reported a total
484 of 213 IR-active vibrational modes (Table S3 in supplementary material of Ruggiero et al.
485 (2015)), at odd with both our first principles calculations as well as our group theoretical
486 analysis. Based on the very low intensity reported for several modes i.e., of the 32 modes in the
487 O – H stretching region 16 exhibit an intensity of less than 5% of the strongest IR active O – H
488 stretching mode, and a total number of 213 IR-active modes which corresponds to the number of
489 optical modes obtained in our study, we believe that Ruggiero et al. (2015) computed the Γ -point
490 frequencies of the complete set of optical vibrational modes but wrongly assigned them all to be
491 IR-active. Comparing the vibrational frequencies of the O – H stretching modes obtained by

492 Ruggiero et al. (2015) (i.e., 3451.2 – 3807.3 cm⁻¹) and the ones computed in our study (3278.7 –
493 3635.3 cm⁻¹) to the experimental values (3252 – 3582 cm⁻¹), it is evident that our calculations
494 predict the vibrational mode frequencies much more accurately. Ruggiero et al. (2015) computed
495 the vibrational frequencies at a higher level of theory and substantially higher computational cost
496 using the B3LYP hybrid functional, yet again proving the effectiveness of the self-consistent
497 DFT + U approach (Kulik et al. 2006; Timrov et al. 2021) in accurately predicting the material
498 properties of transition metal compounds at a fraction of the computational cost.

499 Implications

500 We have demonstrated, by combining neutron diffraction with ab initio calculation within
501 the framework of DFT, that the known monoclinic phase of rozenite is structurally stable over
502 the entire temperature range it might be exposed to on the martian surface and does not undergo
503 any structural phase transition down to temperatures of at least 21 K. Rozenite is amongst the
504 most promising candidates for the polyhydrated sulfate phase on the martian surface.
505 Constraining the low temperature stability and vibrational fingerprint of candidate phases such as
506 rozenite is of vital importance to understand both the climatological history as well as for the
507 future colonization of Mars. Characterization of the stratigraphic sequence (i.e., sulfate vs clay
508 mineral layers) of the polyhydrated sulfate deposits may enable us to decipher the nature, drivers
509 and onset of changing environmental conditions during the planet's early history.

510 As for the future colonization of Mars, hydrated sulfate minerals may be a valuable
511 resource in the equatorial region where water ice is unstable (Feldman et al. 2004). Therefore, it
512 is not only important to merely delineate between various mineral families, but a detailed
513 mineralogical phase identification, including the mineral's hydration state, is crucial. For
514 instance, 321 mL of water may be extracted from one kg rozenite, whereas complete dehydration

515 of the same amount of the monohydrate yields 106 mL. Moreover, rozenite may be dehydrated
516 to szomolnokite at 315 K, producing 241 mL H₂O/kg, whereas complete desiccation of
517 szomolnokite to anhydrous FeSO₄ takes place at much higher temperature (approximately 500
518 K) (Kanari et al. 2018). In contrast, the dehydration of rozenite to szomolnokite is more efficient
519 process, thus making it a potentially valuable resource for the extraction of water on Mars.

520 Raman spectroscopy has proven to be a very effective tool in discriminating between the
521 various hydrated mineral species (Košek et al. 2017), but such efficacy relies on an accurate
522 spectral library of the most promising candidate phases. As part of the scientific payload of the
523 Perseverance rover the first two Raman spectrometer were successfully deployed on the martian
524 surface (Bhartia et al. 2021; Wiens et al. 2020). The European Space Agency's Rosalind
525 Franklin Rover (Rull et al. 2017), which is expected to land on the martian surface in 2023, will
526 also rely on Raman spectroscopy for mineralogical phase identification. Accurate vibrational
527 fingerprint data is becoming increasingly important in planetary exploration. In our study, we
528 have demonstrated that the challenges associated with the sensitivity of the sample material to
529 changing relative humidity conditions may be circumvented by analyzing the sample inside a
530 glass capillary and performing synchrotron X-ray diffraction on the very same capillary. Lastly,
531 we want to highlight the role DFT may play in vibrational mode assignment and interpretation,
532 in particular at low-temperatures where sharpening of vibrational modes might be misinterpreted
533 as mode splitting. This will become even more important in the future as computational
534 resources become cheaper and more readily available, and likewise quantum chemical codes
535 become increasingly user friendly and optimized for high-throughput calculations. Combining
536 such calculations with experimentally determined reference spectra will enable the construction
537 of a reliable Raman spectroscopic database for planetary exploration, which will be invaluable to

538 shed light on the geological past as well as in identifying resources for the future colonization of
539 planetary bodies in the solar system.

540 Acknowledgment

541 We want to thank STFC's Scientific Computing Department for the provision of
542 computing resources on the SCARF cluster and in particular Jon Roddom for compiling the
543 Quantum ESPRESSO code. Moreover, we want to thank two anonymous reviewers for their
544 thorough reading and constructive feedback. We further acknowledge the STFC for access to
545 beamtime at the ISIS Neutron and Muon Source (RB2010354) and the Diamond Light Source
546 (CY26409). JMM acknowledges funding from an ISIS Facility Development and Utilisation
547 Studentship (50 %) and the University of Exeter (50 %).

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Tables

Crystal data

Chemical Formula	FeSO ₄ •4D ₂ O	FeSO ₄ •4D ₂ O
Space group	<i>P</i> 2 ₁ / <i>n</i>	<i>P</i> 2 ₁ / <i>n</i>
Temperature	290 K	21 K
<i>a</i> , <i>b</i> , <i>c</i> , β	5.966031(12) Å, 13.609756(31) Å, 7.962529(14) Å, 90.4288(2)°	5.942863(15) Å, 13.521390(40) Å, 7.933688(20) Å, 89.8617(2)°
<i>V</i>	646.509(2)	637.516(2)
<i>Z</i>	4	4

Refinement

R-factors	R _p = 0.0237,	R _p = 0.0316, R _w p
goodness of fit	R _w p = 0.0179, χ ² = 2.190	= 0.0257, χ ² = 5.472
Number of refined parameters	192	182*

Tab. 1. Selected details of the crystal structure refinement, for a full description we refer to the CIF. *Number of refined parameters is lower for the measurement at 21 K since the background coefficients had to be fixed for the 30 – 130 ms TOF window.

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Source	Baur	This study	Diff*	This study	Diff**
T (K)	RT	290		21	
Fe-O1	2.120(9)	2.121(3)	-0.001	2.130(3)	-0.009
Fe-O2	2.120(9)	2.130(3)	-0.010	2.127(3)	0.003
Fe-Ow1	2.099(11)	2.096(3)	0.003	2.108(3)	-0.011
Fe-Ow2	2.129(11)	2.139(4)	-0.010	2.145(3)	-0.006
Fe-Ow3	2.127(11)	2.114(4)	0.013	2.113(3)	-0.001
Fe-Ow4	2.126(11)	2.097(3)	0.029	2.094(3)	0.003
MSiD ¹			0.004		-0.003
MUD			0.011		0.006
O1-Fe-Ow1	86.2(2.4)	87.44(11)	-1.24	87.44(10)	0.00
O1-Fe-Ow2	95.7(2.4)	95.51(12)	0.19	95.50(11)	0.01
O1-Fe-Ow3	87.0(2.4)	87.13(12)	-0.13	86.88(10)	0.25
O1-Fe-O2	91.9(2.4)	91.00(10)	0.90	90.81(10)	0.19
Ow4-Fe-Ow1	85.5(2.4)	84.73(12)	0.77	84.36(11)	0.37
Ow4-Fe-Ow2	92.5(2.4)	92.31(13)	0.19	92.69(12)	-0.38
Ow4-Fe-Ow3	92.5(2.4)	92.64(13)	-0.14	92.86(11)	-0.22
Ow4-Fe-O2	88.8(2.4)	89.17(12)	-0.37	89.48(11)	-0.31
Ow1-Fe-Ow3	90.7(2.4)	91.97(13)	-1.27	92.47(11)	-0.50
Ow2-Fe-O2	94.0(2.4)	93.10(12)	0.90	93.33(11)	-0.23
Ow2-Fe-Ow3	86.8(2.4)	87.35(13)	-0.55	86.56(11)	0.79
MSiD			-0.068		-0.003
MUD			0.605		0.295
S-O1	1.512(8)	1.487(5)	0.025	1.478(5)	0.009
S-O2	1.492(8)	1.482(5)	0.010	1.485(5)	-0.003
S-O3	1.488(8)	1.468(5)	0.002	1.484(5)	-0.016
S-O4	1.473(8)	1.481(5)	-0.008	1.481(6)	0.000
MSiD			0.012		-0.003
MUD			0.016		0.007
O1-S-O2	109.4(1.1)	108.2(3)	1.2	109.7(4)	-1.5
O1-S-O3	108.3(1.1)	110.4(3)	-2.1	109.9(3)	0.5
O1-S-O4	108.3(1.1)	108.7(3)	-0.4	109.0(3)	-0.3
O2-S-O3	109.7(1.1)	109.3(3)	0.4	108.3(3)	1.0
O2-S-O4	111.8(1.1)	111.2(3)	0.6	111.0(3)	0.2
O3-S-O4	109.3(1.1)	109.0(3)	0.3	108.9(4)	0.1
MSiD			-0.24		0.3
MUD			0.76		0.42

Tab. 2. Bond lengths and angles for the octahedral and tetrahedral units as determined in this study at 290 K and 21 K and compared to the values reported by Baur (1962). * refers to the difference between Baur and this

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747 study's 290 K structure. ** refers to the difference of this study's 290 K and 21 K structures, thus elucidating
 748 the influence of temperature on the respective quantity. ¹The Mean signed difference is defined as $\frac{\sum_i^n x_B - x_T}{n}$
 749 with x_B and x_T being the values as observed by Baur and in this study, respectively for the quantities of interest
 750 (i.e., bond-length and angle).
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Source	Baur	This study	Diff*	This study	Diff**
T (K)	RT	290		21	
Ow1-H1a	1.010(8)	0.947(5)	0.063	0.965(5)	-0.018
Ow1-H1b	0.962(9)	0.962(4)	0.000	0.972(4)	-0.01
Ow2-H2a	0.933(9)	0.963(5)	-0.03	0.959(4)	0.004
Ow2-H2b	0.958(9)	0.944(5)	0.014	0.975(4)	-0.031
Ow3-H3a	0.955(8)	0.950(5)	0.005	0.949(4)	0.001
Ow3-H3b	0.975(9)	0.943(4)	0.032	0.962(4)	-0.019
Ow4-H4a	0.960(8)	0.953(5)	0.007	0.968(4)	-0.015
Ow4-H4b	0.964(9)	0.973(4)	-0.009	0.960(4)	0.013
MSiD			0.010		-0.009
MUD			0.020		0.014
H1a-Ow1-H1b	105.2(8)	106.0(4)	-0.80	107.8(4)	-1.80
H2a-Ow2-H2b	104.0(8)	105.5(5)	-1.50	107.1(4)	-1.6
H3a-Ow3-H3b	110.2(8)	108.4(4)	1.80	108.1(4)	0.3
H4a-Ow4-H4b	105.1(8)	105.0(4)	0.10	106.9(4)	-1.90
MSiD			-0.10		-1.25
MUD			1.05		1.40
Ow1-O3	2.855(10)	2.898(4)	-0.043	2.848(4)	0.050
Ow1-O3'	2.790(10)	2.764(4)	0.026	2.744(4)	0.020
Ow2-O4	2.867(10)	2.858(4)	0.009	2.834(4)	0.024
Ow2-O2	3.023(10)	3.015(4)	0.008	2.986(4)	0.029
Ow2-O2'	3.267(10)	3.274(4)	-0.007	3.250(4)	0.024
Ow3-O4	2.845(10)	2.836(4)	0.009	2.794(4)	0.042
Ow3-Ow4	3.025(12)	3.059(5)	-0.034	3.029(4)	0.030
Ow3-O1	2.805(10)	2.816(4)	-0.011	2.795(4)	0.021
Ow4-O4	2.837(10)	2.848(4)	-0.011	2.832(4)	0.016
Ow4-O3	2.723(10)	2.733(4)	-0.010	2.718(4)	0.015
MSiD			-0.0064		0.027
MUD			0.0168		0.027
H1a-O3	1.865(6)	1.957(4)	-0.092	1.888(4)	0.069
H1b-O3'	1.834(6)	1.808(4)	0.026	1.776(4)	0.032
H2a-O4	2.201(6)	1.992(5)	0.209	1.967(4)	0.025
H2b-O2	2.417(6)	2.330(5)	0.087	2.249(4)	0.081
H2b-O2'	2.593(6)	2.551(4)	0.042	2.525(4)	0.026
H3a-O4	2.113(6)	2.015(4)	0.098	1.974(4)	0.041
H3a-Ow4	2.345(9)	2.425(5)	-0.080	2.399(4)	0.026
H3b-O1	1.833(6)	1.874(4)	-0.041	1.834(4)	0.040
H4a-O4	1.969(6)	1.948(4)	0.021	1.918(4)	0.093
H4b-O3	1.899(6)	1.764(4)	0.135	1.764(4)	0.000
MSiD			0.041		0.037
MUD			0.083		0.037
O3-Ow1-O3'	104.8(4)	104.90(11)	-0.10	104.46(10)	0.44
O4-Ow2-O2	144.5(4)	143.43(14)	1.07	142.81(12)	0.62
O4-Ow2-O2'	92.6(3)	93.12(11)	-0.52	93.75(9)	-0.63
O4-Ow3-O1	137.2(4)	136.23(13)	0.97	135.54(11)	0.69
Ow4-Ow3-O1	70.8(3)	70.28(10)	0.52	69.45(8)	0.83
O4-Ow4-O3	117.3(4)	116.68(12)	0.62	118.30(11)	-1.62
MSiD			0.43		0.055
MUD			0.63		0.805
Ow1-H1a-O3	165.8(5)	172.0(4)	-6.2	173.5(4)	-1.60
Ow1-H1b-O3'	171.8(6)	172.5(4)	-0.7	173.1(4)	-0.60
Ow2-H2a-O4	125.7(6)	148.5(4)	-22.8	149.2(3)	-0.70
Ow2-H2b-O2	122.6(6)	129.0(4)	-6.4	131.5(3)	-2.50
Ow2-H2b-O2'	129.6(6)	133.6(4)	-4.0	131.0(3)	2.60
Ow3-H3a-O4	132.3(6)	143.6(3)	-11.3	143.4(3)	0.20
Ow3-H3a-Ow4	127.7(6)	124.0(3)	3.7	123.6(3)	0.40

Ow3-H3b-O1	173.7(5)	175.4(4)	-1.7	176.0(3)	-0.60
Ow4-H4a-O4	149.2(6)	156.7(4)	-7.5	156.4(3)	0.30
Ow4-H4b-O3	141.8(5)	172.6(4)	-30.8	172.2(3)	0.40
MSiD			-8.77		-0.20
MUD			9.51		0.98

758 Tab. 3. Geometry of the hydrogen bonds as determined in this study at 290 K and 21 K, and compared to the
 759 values reported by Baur (1962). *refers to the difference between Baur and this study's 290 K structure. **
 760 refers to the difference of this study's 290 K and 21 K structures, thus elucidating the influence of temperature
 761 on the respective quantity.
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	a^3	b^3	c^3	β	V
θ_D (K)	240(6)	303(2)	382(4)	381(13)	278(2)
X_0 (cm ³ mol ⁻¹)	31.5994(6)	372.182(2)	75.1823(4)	13.5290(4)	95.9829(4)
X_0 (Å, Å ³)	5.9428(1)	13.5214(8)	7.9337(5)	934(19)	637.533(3)
Q (x10 ⁴ J cm ⁻³)	850(15)	519(3)	1047(12)	1263(26)	811.(3)
K_0/γ (GPa)	269(5)	13.94(7)	139(2)	1617(17)	84.5(4)
K'	12.6	19.92	90.76	146.55	41.11

764 Tab. 4. Parameters derived from fitting a second order single Debye model upon the lattice parameters of
 765 rozenite.
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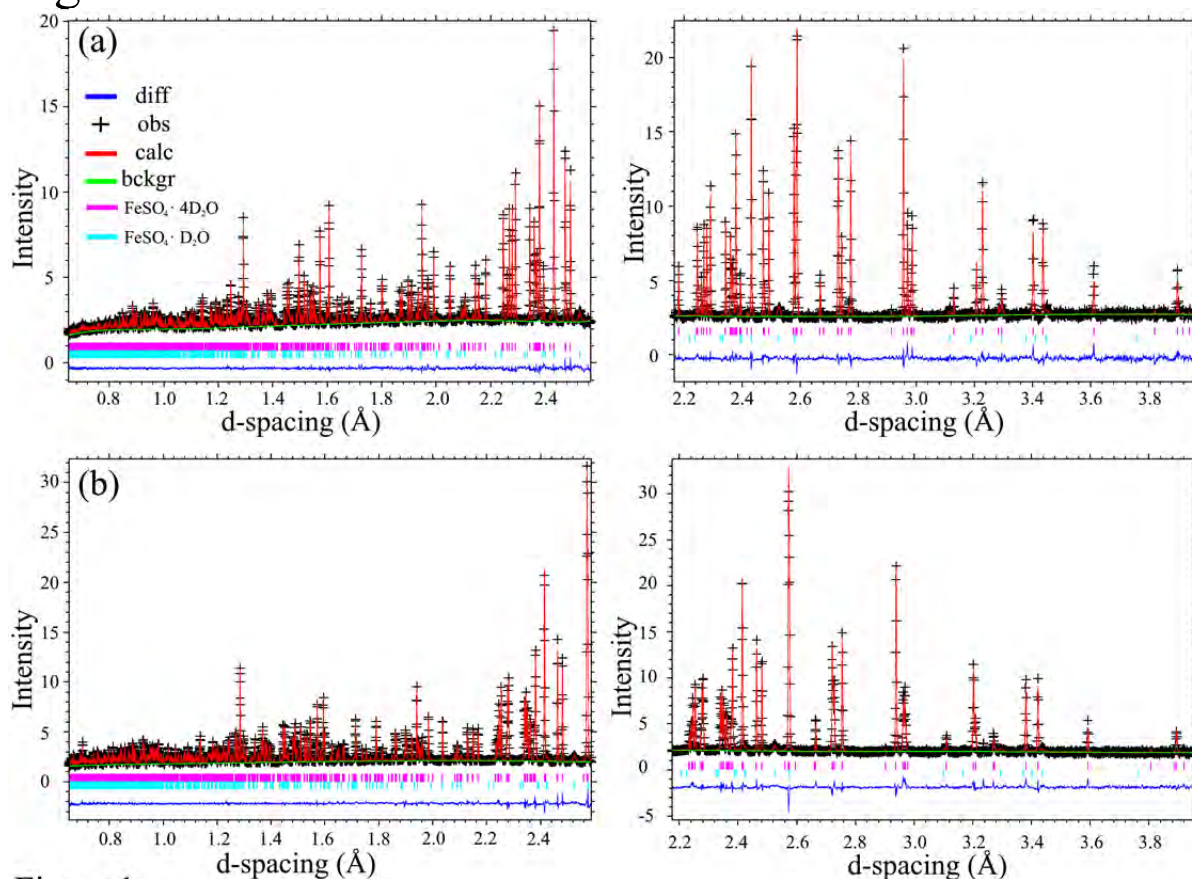
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Compound	NTE (K ⁻¹)	T _{ref}	VTE expansion (K ⁻¹)	T _{ref} (K)	Source
MgSO ₄ ·11H ₂ O	-1.2(5) x 10 ⁻⁵	50	6.8(4) x 10 ⁻⁵	240	Fortes et al. (2008)
MgSO ₄ ·7H ₂ O	-2.0(2) x 10 ⁻⁵	125	10.8(3) x 10 ⁻⁵	290	Fortes et al. (2006)
FeSO ₄ ·4H ₂ O	-1.0(2) x 10 ⁻⁵	285	8.2(5) x 10 ⁻⁵	285	This study
MgSO ₄ ·H ₂ O	-4.4(3) x 10 ⁻⁵	293	3.4(7) x 10 ⁻⁵	293	Wildner et al. (2022)
FeSO ₄ ·H ₂ O	-1.7(2) x 10 ⁻⁵	293	4.7(5) x 10 ⁻⁵	293	Wildner et al. (2022)
CoSO ₄ ·H ₂ O	-2.7(2) x 10 ⁻⁵	293	3.3(3) x 10 ⁻⁵	293	Wildner et al. (2022)
NiSO ₄ ·H ₂ O	-0.8(5) x 10 ⁻⁵	293	3.6(4) x 10 ⁻⁵	293	Wildner et al. (2022)

768 Tab.5. Comparison of the negative (NTE) and volume (VTE) thermal expansion of various M²⁺SO₄·nH₂O.
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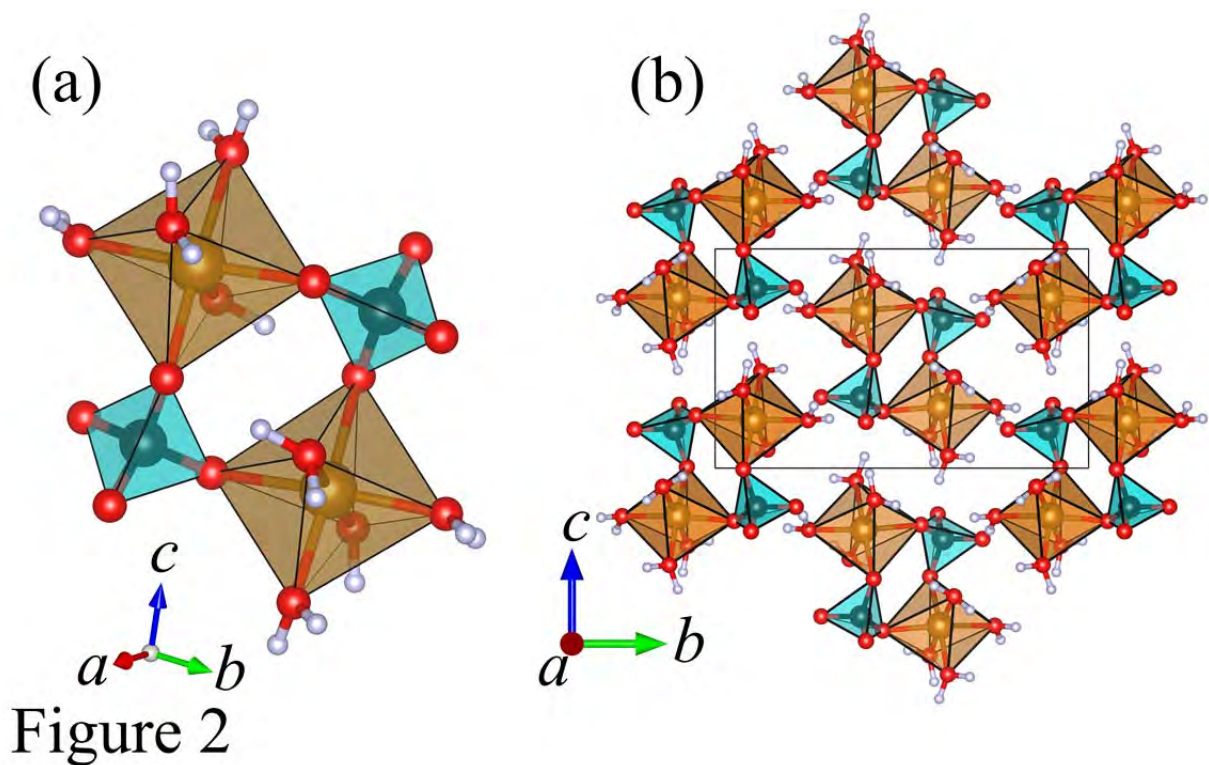
Figures



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Figure 1

Fig. 1. Neutron diffraction patterns acquired at (a) 290 K and (b) 21 K in the backscattering detector bank collected in the 30 – 130 (left) and 100 – 200 ms (right) TOF window. The observed data plotted as crosses, the red line represents the fitted model, and the blue line the difference profile. The tick marks corresponding to each of the Bragg peaks of FeSO₄·4D₂O and FeSO₄·D₂O are displayed in magenta and cyan, respectively.



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Fig. 2. Crystal structure of rozenite drawn using the VESTA software (Momma and Izumi 2011). (a) Detailed view of the $[\text{Fe}(\text{H}_2\text{O})_4\text{SO}_4]_2$ units. Note that the non-hydrated oxygens of the $\text{Fe}(\text{H}_2\text{O})_4\text{O}_2$ units form bridges to the SO_4 tetrahedra (orange and cyan, respectively). (b) The $[\text{Fe}(\text{H}_2\text{O})_4\text{SO}_4]_2$ isolated units are linked via a complex network of intermolecular hydrogen bonds.

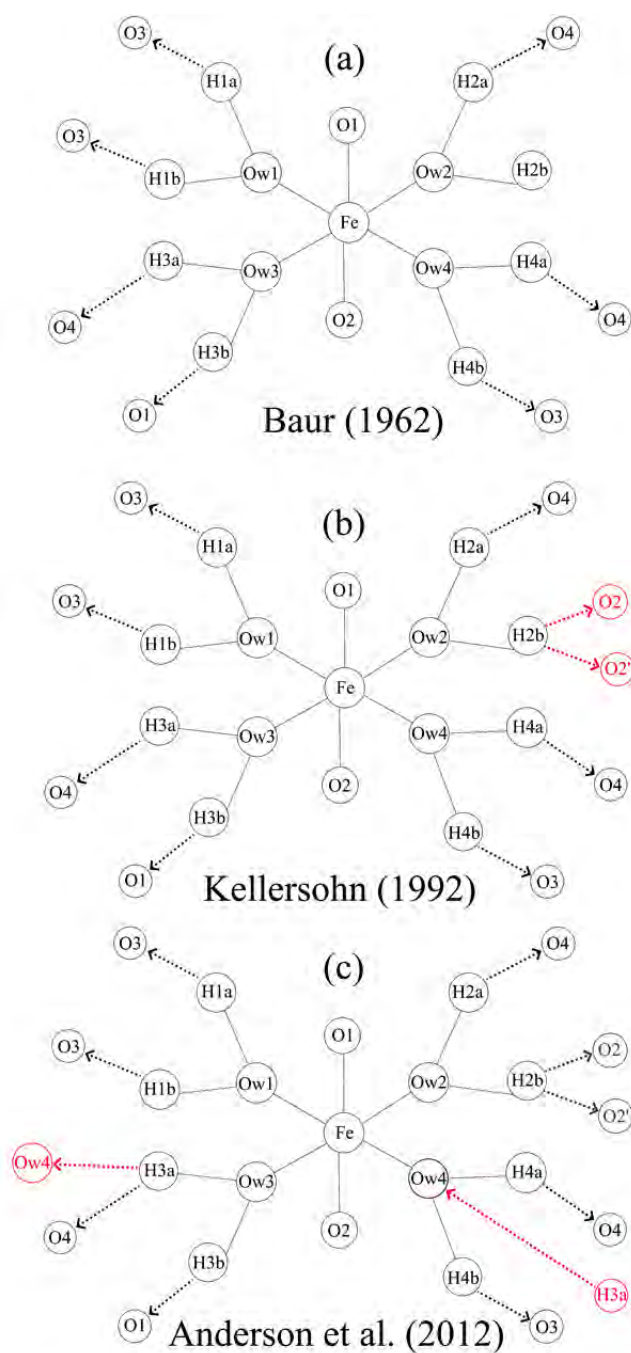


Figure 3

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Fig. 3 Interpretation of the hydrogen bond network by (a) Baur (1962), (b) Kellersohn (1992) and (c) Anderson et al. (2012). The changes suggested by these authors with respect to the previous interpretation are marked in red.

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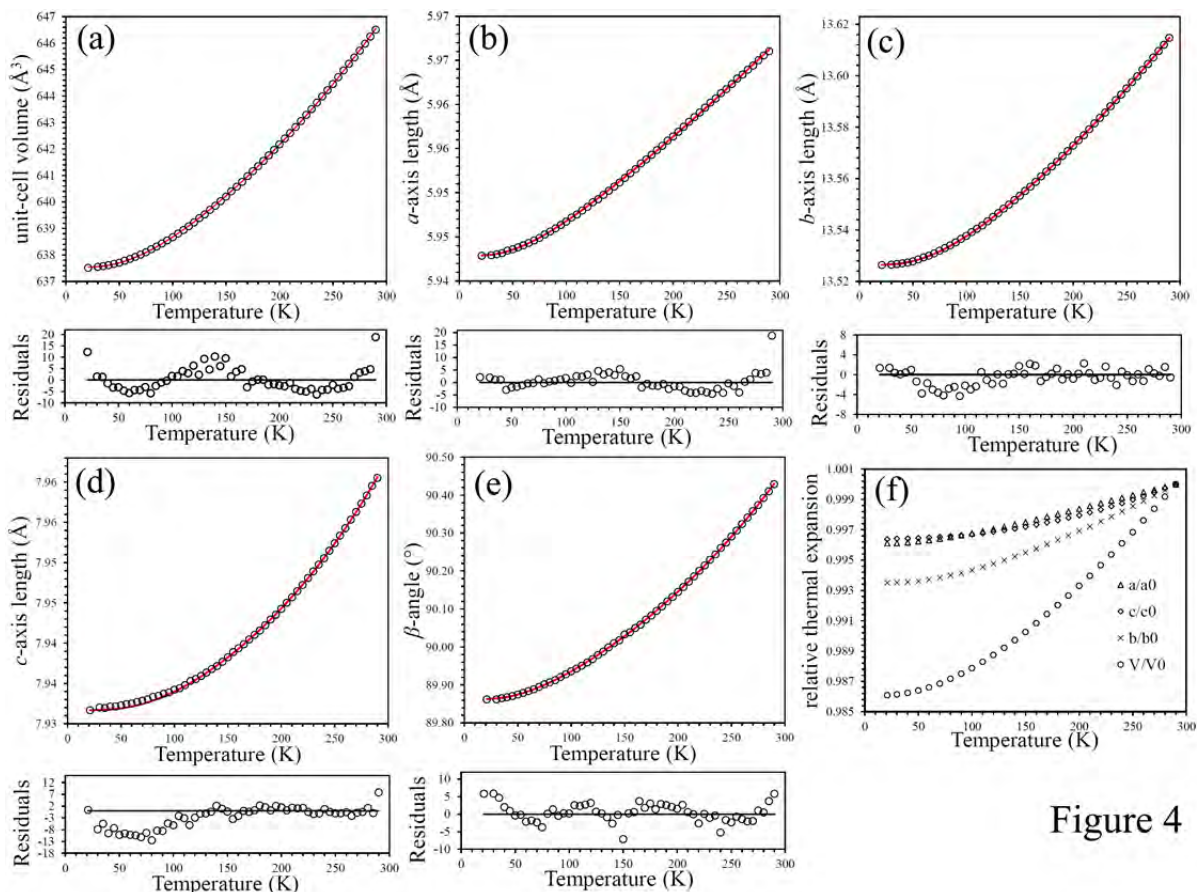


Figure 4

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792 Fig. 4. (a-e) Second-order Debye model fit (red line) to the lattice parameters (black open circle) of rozenite in
793 at temperatures ranging from 290 to 21 K. Error bars are smaller than the symbol size. Residuals are defined
794 as the difference between observed and fitted values divided by the experimental estimated standard deviation
795 determined for each data point. (f) Relative thermal expansion of each of the lattice parameters as a function of
796 temperature. Note the crossover in the evolution of the *a* and *c* axes ($T \sim 100$ K).

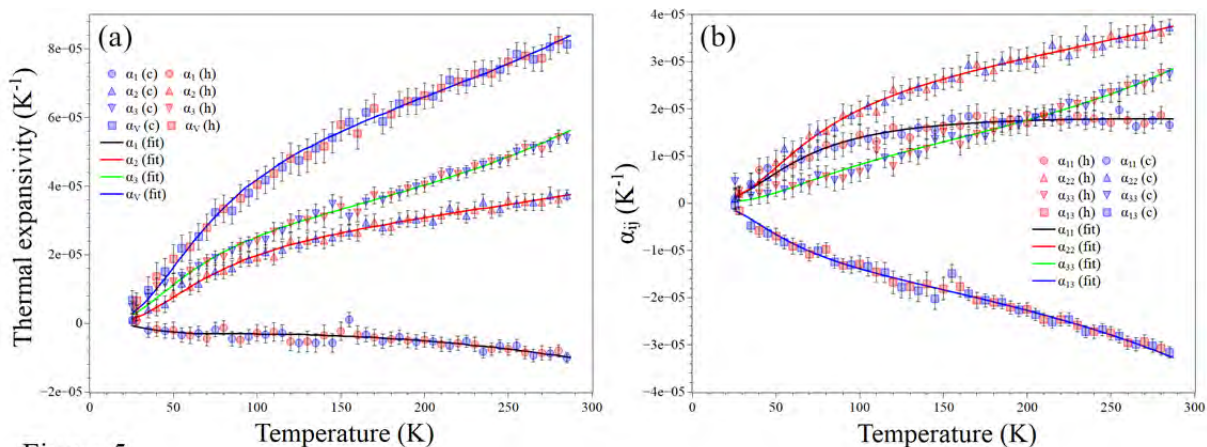


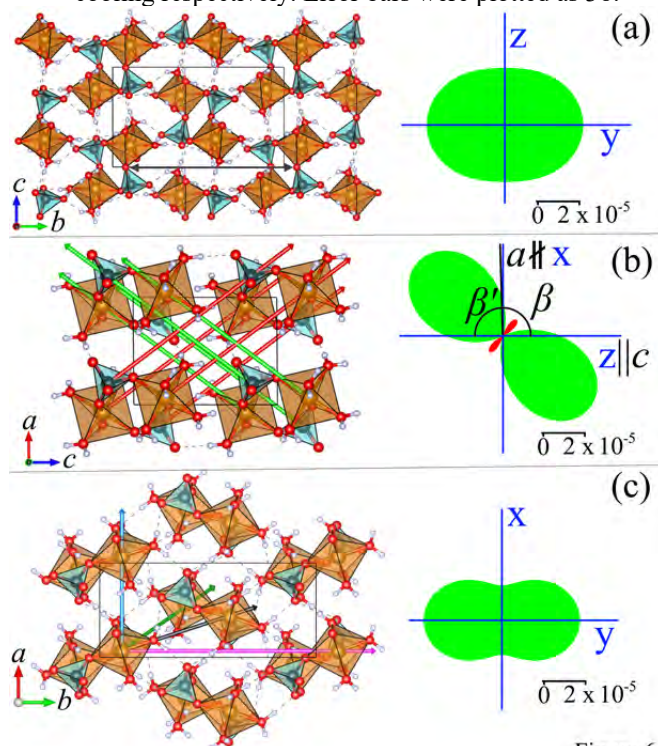
Figure 5

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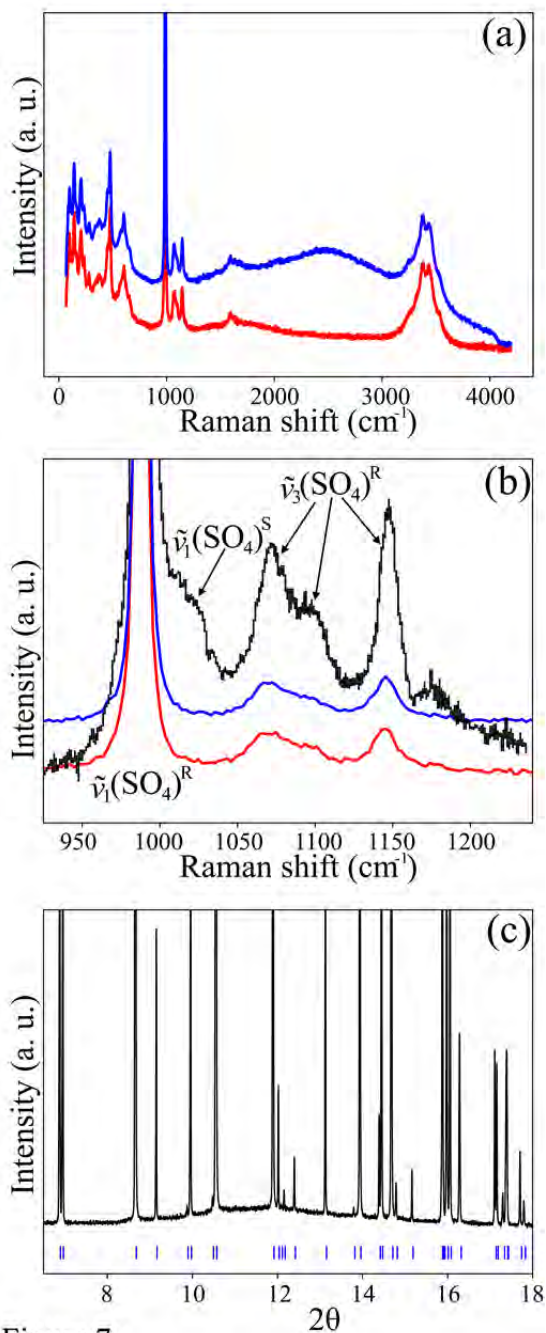
Fig. 5. The evolution of the thermal expansion tensor's the principal axes (a), and components (b) is continuous over the entire temperature range under investigation. Solid lines represent the 2nd order Debye model fit; red and blue symbols are calculated from the lattice parameter data acquired upon heating and cooling respectively. Error bars were plotted as 3 σ .



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Figure 6

803 Fig. 6. View of the crystal structure (left) and cross-sections of the thermal expansion tensor (green represents
804 positive, red negative values) at 285 K (right). (a) view || *a*: Large 'void' (black arrow: 11.932(5) Å at 290 K and
805 11.878(5) Å at 21 K) || *b* as well as absence of hydrogen bonding in this direction allows the structure to experience
806 substantial expansion and contraction in this direction as compared to the *c*-direction. (b) view || *b*: Red arrows point
807 || <101> (i.e., the direction of negative thermal expansion, whereas green arrows are oriented <10 $\bar{1}$ > (i.e., the
808 direction of maximum thermal expansion). The cyclic dimer units are stacked on top of each other and oriented by
809 the angle of β (~90 °). Furthermore, the distance between diagonally opposing Fe and S atoms (connected by the red
810 and green arrows) is increasing upon cooling shrinking in the direction of maximum thermal expansion. (c) view || *c*:
811 arrows point towards to the central atoms of the FeO₂(H₂O)₄ units of neighboring FeO₂(H₂O)₄ units.



812

Figure 7

813 Fig. 7. (a) Raman spectra acquired in this study outside (red) and inside (blue) of a borosilicate glass capillary.
814 (b) Selected spectral range for comparison between our data (red and blue curves) and Chio et al. (2007)
815 (black curve). (c) The phase purity of our sample has been confirmed by means of synchrotron X-ray
816 diffraction, suggesting that the shoulder at 1018 cm⁻¹ in the Chio et al. (2007) spectrum (black) stems from a
817 contamination of the sample with szomolnokite. $\tilde{\nu}_1(\text{SO}_4)^{\text{R}}$ and $\tilde{\nu}_1(\text{SO}_4)^{\text{S}}$ refer to the sulfate stretching modes
818 assigned to rozenite and szomolnokite, respectively. The Raman spectrum reported by Chio et al. (2007) was
819 digitized using the Webplotdigitizer tool (Rohatgi 2021). The intensity in all graphs was normalized with
820 respect to the maximum intensity.

821

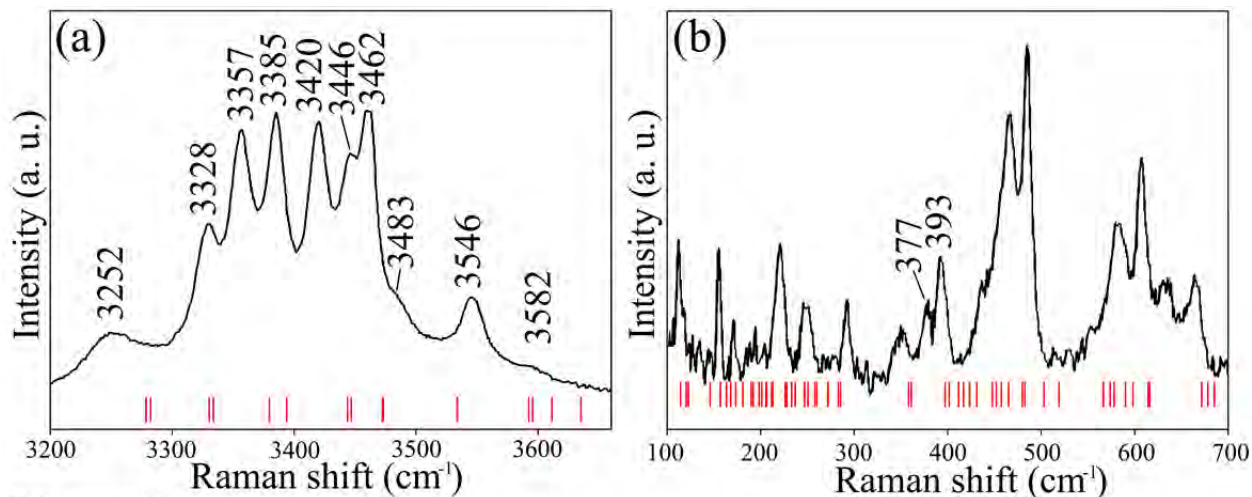


Figure 8

822 Fig. 8. Spectral region of the (a) water stretching vibration and (b) external modes of the Raman spectrum
823 acquired by Chio et al. (2007) at 8 K. Red markers indicate the Raman-active vibrational mode positions as
824 predicted by DFT. The Raman spectrum reported by Chio et al. (2007) was digitized using the
825 Webplotdigitizer tool (Rohatgi 2021).
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