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3	Major and trace element composition of olivine from magnesian skarns
4	and silicate marbles
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16	Abstract
17	Olivine is a major rock-forming mineral in various magmatic and metamorphic rocks and upper
18	mantle. In this paper, we present first high-precision analyses of olivine from 15 samples of
19	magnesian skarns and silicate marbles (MSSM) from the collection of the Fersman
20	Mineralogical Museum (Moscow, Russia). Mg# (Mg/(Mg+Fe <sup>2+</sup> )*100, mol.%) of olivine from
21	the samples studied varies from 86 to nearly 100. The main distinctive features of the olivine are
22	anomalously low contents of Co (<51 $\mu$ g/g), Cr (< 5 $\mu$ g/g) and Ni (< 44 $\mu$ g/g) and high content
23	of B (23-856 $\mu$ g/g), which correlate with host-rock compositions. Phosphorus (5-377 $\mu$ g/g)
24	incorporation in olivine is charge-balanced by the incorporation of Li (0.15-61 $\mu g/g)$ and Na
25	(<14.3 $\mu$ g/g). Y and REE contents exhibit positive correlations with Na, which suggest that REE

26	incorporation into MSSM olivine could occur via charge-balanced coupled substitution with Na
27	at low temperature and low aSiO <sub>2</sub> conditions during MSSM formation. The documented
28	compositional features of olivine from magnesian skarns and silicate marbles can help
29	reconstruct the genesis of the host-rocks and identify xenocrysts of MSSM olivine in magmatic
30	rocks.
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32	Keywords: Olivine; LA-ICP-MS; magnesian skarn; silicate marble; contact metamorphism
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34	Introduction
35	Calc-silicate metamorphic and metasomatic rocks, mantle peridotites and mafic igneous rocks
36	are three major rock types of common occurrence of magnesian olivine (Mg#>80 after Plechov
37	et al. (2018)). Olivine-bearing calc-silicate metamorphic and metasomatic rocks are represented
38	by magnesian skarns and silicate marbles (hereafter referred to as MSSM).
39	MSSM have similar mineral assemblages, which include carbonates (calcite or dolomite),
40	olivine, phlogopite, diopside, spinel, periclase, clinohumite, and pargasite. Magnesian silicate
41	marbles are difficult to distinguish from magnesian skarns by mineralogy or chemistry. The only
42	major difference between them is their geological position. Magnesian skarns are the product of
43	contact metasomatism and always located in the aureoles of magmatic intrusives (Zharikov et al.,
44	2007), whereas olivine-bearing magnesian silicate marbles are not associated with magmatic
45	rocks. Olivine-bearing magnesian silicate marbles could form during high-grade metamorphism
46	or infiltration metasomatism of dolomites (Bucher and Grapes, 2011).
47	Olivine-bearing MSSM can form in wide range of pressures (from $\sim 1$ to $\sim 10$ kbar) and
48	temperatures (from ~550 to ~900 $^{\circ}$ C) as follows from thermodynamic estimates (Bucher and
49	Grapes, 2011; Pertsev, 1974; Zharikov, 1970). The majority of temperature and pressure
50	estimates for olivine growth in carbonate-silicate metamorphic or metasomatic rocks lies in the
51	ranges of 550-700°C and 0.5-5 kbar, for example: about 600°C and from 1 to 5 kbar for

Tazheran, Russia (Doroshkevich et al., 2017), 595 °C and 1.8±0.5 for Twin Lakes pendant,
California, USA (Ferry et al., 2011), 680 °C and 0.5 kbar for Beinn an Dubhaich aureole, Isle of
Skye, Scotland (Ferry et al., 2011), and 550 °C and 2.5 kbar for Gruvaisen, Sweden (Hellingwerf,
1984).

The concentration ranges of major, minor and trace elements in igneous and mantle olivines that 56 57 are formed at higher temperatures and wider range of pressures are increasingly wellcharacterized (e.g. Arai, 1994; Bussweiler et al., 2019; De Hoog et al., 2010; Foley et al., 2013; 58 Jaques and Foley, 2018; Neave et al., 2018; Rooney et al., 2020; Sobolev et al., 2007; Sobolev et 59 al., 2008; Su et al., 2019; etc.), whereas high-precision data on the composition of olivine from 60 MSSM are very scarce. These published data include several ion probe analyses of B-rich olivine 61 from the Tayozhnoye deposit (Grew et al., 1991), high-precision electron probe analyses of 62 olivine from Kuh-i-Lal deposit (Plechov et al., 2018), some routine (non-high-precision) 63 microprobe analyses of olivine from Tazheran deposit (Doroshkevich et al., 2017), Twin Lakes 64 pendant (California, USA) and Beinn an Dubhaich aureole (Isle of Skye, Scotland) (Ferry et al., 65 66 2011) and from primitive magmatic skarns found as xenoliths in magmatic rocks (Fulignati et al., 2005; Gilg et al., 2001; Wenzel et al., 2002). These published data are too scarce to place 67 constraints on the entire variations of olivine composition from MSSM. 68

In this work we systematically study the content of major, minor and trace elements in olivine from MSSM from different locations. The new data reveal distinctive compositional features of olivine from these rocks, which are manifested in strong enrichment in B combined with depletion in Ni, Co and Cr concentrations in comparison with typical igneous and mantle olivine.

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## Samples and methods

We studied a total of 15 samples of magnesian skarn (or silicate marbles in some cases) olivine from nine different locations in three countries (Table 1, Fig. 1) from the collection of the Fersman Mineralogical Museum of the Russian Academy of Sciences (Moscow, Russia). The

78 samples are representative of both magnesian skarns and silicate marbles. The samples of magnesian skarns are from three classical localities: Nikolai-Maximilian mine, South Urals, 79 Russia (e.g. Bocharnikova et al., 2011), Slyudvanka, Baikal area, Russia (e.g. Aleksandrov and 80 81 Senin, 2006) and Kukhilal, Tajikistan (Pautov et al., 2018). The sample from Belaya Vyemka, Baikal area, Russia, is a spinel-forsterite marble (Sal'nikova et al., 2007). Five samples are 82 selected from contact zones of carbonate rocks with alkaline intrusives (Tazheran, Russia, and 83 Hodzha-Achkan, Kyrgyzstan) (e.g. Doroshkevich et al., 2017; Pautov et al., 2013) and another 84 five samples are from skarns of phlogopite deposits in Yakutia, Russia, (Katalakh, Emeldhak and 85 Timpton) associated with granites (e.g. Mikhailov, 1997). 86

High-precision electron microprobe analysis (EPMA) of studied olivine samples was conducted 87 using a wavelength-dispersive X-ray (WDS) electron microprobe JEOL JXA-8230 (Lomonosov 88 Moscow State University) operated at an accelerating voltage of 20 kV and a beam current of 89 100 nA. San Carlos olivine USNM-111312/444 (Si, Mg, Fe), metal oxides (Ti, Mn, Cr, Al, Ni), 90 wollastonite (Ca) and Durango apatite (P) were used for calibration. The peak signal 91 92 accumulation time was 120 s for Cr, Ti, P, Mn and Ni, 180 s for Al, 40 s for Ca, respectively. Analysis at high beam current and long counting times significantly lowers detection limits of 93 trace element in olivine (Sobolev et al., 2007). Analytical precision, which was determined as 2 94 95 standard deviations of multiple repetitive analyses of San Carlos olivine (USNM-111312/444), was estimated to be 41 µg/g for Ni, 33 µg/g for Mn, 9 µg/g for Ca, 11 µg/g for P, 12 µg/g for Al, 96 19 µg/g for Cr and 8 µg/g for Ti. The results are presented in Supplementary Table 1. San Carlos 97 olivine (USNM-111312/444) was used as a secondary standard (Supplementary table 4). 98

99 Major and trace element concentrations of the same olivine samples were also analyzed by laser 100 ablation-inductively coupled plasma-mass spectrometry (LA-ICP-MS) using an Agilent 7900s 101 quadrupole mass-spectrometer coupled with the 193 nm Excimer Laser-Ablation system 102 GeoLas<sup>TM</sup> HD (Coherent) at the Institute of Geosciences, Kiel University. Instrumental 103 conditions, mass numbers analyzed, and details of data reduction are given in Supplementary

Table 5. In brief, the analyses were performed using two volume ablation cell (ETH Zürich, 104 Switzerland; Fricker et al., 2011) with a carrier gas flow of 0.7 L min<sup>-1</sup> He and addition of 14 mL 105 min<sup>-1</sup> H<sub>2</sub>. The carrier gas was mixed with Ar (~1 L min<sup>-1</sup>) prior to introduction to the ICP-MS. 106 107 The ablation was done with laser beam diameter ranging from 32 to 120  $\mu$ m, pulse frequency of 10 Hz and fluence of 5 J/cm<sup>2</sup>. Every analysis comprised 20 s background measurement (laser-108 109 off) and 40 s of signal measurement (laser-on). The initial data reduction was performed in the GLITTER software (Griffin et al., 2008). The intervals for integration of analytical signal were 110 manually defined to avoid possible contamination by inclusions in olivine, which are usually 111 detectable by strongly elevated concentrations of highly incompatible elements (K, Rb, U, Pb 112 113 etc.). The data were initially quantified using glass KL2-G (Jochum et al., 2006) obtained using laser spot of 120 µm and Si as reference element, assuming SiO<sub>2</sub>=40 wt% in olivine. A 114 secondary standard of olivine MongOL Sh11-2 (Batanova et al., 2019) was used for minor 115 matrix and spot size correction of concentrations of Si, Mg, Fe, Li, Na, Al, P, Ca, Sc, Ti, V, Cr, 116 Mn, Co, Ni, Cu, Zn, Ga, Sr, Y, Zr, Dy, Ho, Er, Tm, Yb and Lu (Supplementary Table 3). A 117 correction was applied for the <sup>29</sup>Si<sup>16</sup>O interference on <sup>45</sup>Sc by analysis of synthetic nominally Sc-118 free quartz analysed throughout the measurement session. The final step was proportional 119 adjustment of all concentrations to obtain the sum of SiO<sub>2</sub>, TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, FeO, MnO, MgO, CaO, 120 Na<sub>2</sub>O, K<sub>2</sub>O, NiO, Cr<sub>2</sub>O<sub>3</sub>, P<sub>2</sub>O<sub>5</sub>, ZnO, and V<sub>2</sub>O<sub>3</sub> equal to 100 wt%. The results are presented in 121 Supplementary Table 2. 122

Eight elements in olivine were measured both by EPMA and LA-ICP-MS methods: Si, Fe, Mg, Ti, Al, Mn, Ca, P. Comparison of the data obtained with ca. 5  $\mu$ m (EPMA) and 32-120  $\mu$ m (LA-ICP-MS) spots is instructive about the existence of a small scale compositional heterogeneity of the samples. Statistical analysis of this comparison (Ludwig, 1980; York et al., 2004) shows that the linear regressions for Mg#, Ti, Mn and P do not significantly differ from the line of equal concentrations (1:1 line). MSWD is less than 1.8 for Ti and P, which means that the two analytical methods give statistically indistinguishable results for these elements with probability

of 95% (Wendt, Carl, 1991; Fig. 2). High MSWD for Mg# and Mn can be explained by some 130 compositional heterogeneity of the samples, exceeding the analytical precision. There is minor 131 systematic difference between the EPMA and LA-ICP-MS data for Ca, which exceeds 2 132 133 standard deviations and is probably linked to O<sup>-</sup>, OH<sup>-</sup> and N<sup>-</sup> based interferences from Mg and Si on minor isotope of <sup>43</sup>Ca used for LA-ICP-MS analysis (e.g. Batanova et al., 2019). Therefore, 134 135 EPMA data for Ca were preferred in this study. Significantly higher concentrations of Al obtained by EPMA (Fig. 2) on some samples can result from contamination of olivine surface by 136 Al<sub>2</sub>O<sub>3</sub> used for polishing (e.g. Batanova et al., 2015). LA-ICP-MS analyses appear to be free of 137 this problem and were used in our study. 138

Results

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#### 140 Petrography

MSSM samples studied are represented by olivine-calcite-spinel (Fig. 3a), olivine-calcite-141 phlogopite (Fig. 3b) and olivine-calcite-spinel-phlogopite (Fig. 3c) mineral associations. 142 Diopside and amphiboles are also occur in some samples as rock-forming minerals. Most 143 144 samples are unaltered. The primary rock textures are granoblastic and grano-porphyroblastic. 145 Proportions of rock-forming minerals in studied samples vary widely. In general, calcite makes up more than 50 vol.% in most of them, olivine – up to 50 vol.%, while the contents of spinel 146 147 and phlogopite are rarely exceed 20 vol.% of the rocks. Olivine in these samples has idiomorphic shape and size ranging from 0.1 to 2 mm. Some samples (74607, 67723, 67792, and 77329) are 148 strongly serpentinized (up to 90 % of olivine). Olivine in these samples occurs as relics in the 149 serpentine matrix (Fig. 3d). Many olivines studied contain crystal inclusions of the other rock-150 151 forming minerals and also accessory minerals such as apatite, pyrrhotite, magnetite and ilmenite. 152 Some olivine grains contain abundant fluid inclusions.

#### 153 Olivine Mg#

Mg# of studied olivine samples varies from 86.3 to 99.8, assuming all measured Fe as  $Fe^{2+}$ . Only four studied samples have Mg# < 95, which overlaps with Mg# of typical olivine from mantle

xenoliths and magmatic rocks (e.g. De Hoog et al., 2010; Sobolev et al., 2007; Su et al., 2019). 156 157 The other samples are more Mg-rich (Fig. 4). The high-Mg composition of the studied olivine samples is in a good agreement with the previous descriptions of olivine from magnesian skarns 158 159 (e.g. Plechov et al., 2018; Zharikov, 1970). It should be noted that extremely high-Mg olivine (Mg#>96) can also occur in magmatic rocks (Blondes et al., 2012; Bussweiler et al., 2015; 160 161 Plechov et al., 2018). However, these olivines are usually interpreted to result post-magmatic processes like oxidation and low-temperature re-equilibration. Some of these previously reported 162 ultra Mg-rich olivines can also be xenocrysts from skarns. 163

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# 165 Minor elements (Al, Ca, Mn, Ni, Co, and Cr)

166 As minor elements we here consider elements which occur in olivine at concentrations < 10,000

 $\mu g/g$  and are typically measured by EPMA in other studies (e.g. Sobolev et al., 2007), i.e. Al, Ca,

168 Mn, Ni, Co and Cr. Some of these elements (e.g., Co, Cr and Al) have concentrations below the

detection limit of EPMA in the studied olivine, thus they were analyzed by LA-ICP-MS.
Contents of Al, Ca and Mn in the studied olivine samples vary widely: Al – from 0.5 to 108

171  $\mu g/g$ ; Ca – from 40 to 1880  $\mu g/g$ ; Mn – from 140 to 4140  $\mu g/g$  (Fig. 4). The contents of Ca and 172 Al do not show any correlation with the olivine Mg# and largely overlap with the typical values

for mantle olivine (Fig. 4). Mn content in studied olivine reveals clear negative correlation with

174 Mg#, which is also a typical feature for olivine of magmatic and metamorphic genesis (e.g. De

Hoog et al., 2010; Sobolev et al., 2007; Su et al., 2019). Some of studied MSSM olivines with

176 Mg# of ~88 contain twice as much Mn as the most Mn-rich igneous olivine (Fig. 4). Average

177 Mn/Fe\*100 of MSSM olivine is  $\sim$ 5.3, which is significantly higher than typical values for

igneous and metamorphic olivine (~0.5-2) (De Hoog et al., 2010; Sobolev et al., 2007; Su et al.,

179 2019). However, Mn/Fe\*100 ranges from 1.4 to 20 for different studied samples (Fig. 4) making

180 it difficult to discriminate olivine formed in different environments solely on the base of this

181 parameter.

182	Relative variations of Ni (0.2 - 44 $\mu$ g/g) and Co (0.3 - 51 $\mu$ g/g) contents in studied olivine are
183	even larger than for Al, Ca and Mn. The contents of both these elements correlate negatively
184	with Mg# (Fig. 4). However, even the highest Co and Ni contents in olivine from MSSM do not
185	reach the lowest content of these elements in high-Mg# olivine of mantle and magmatic origin
186	(e.g. De Hoog et al., 2010; Sobolev et al., 2007; Su et al., 2019). Variations of Cr content in the
187	studied olivine samples are the smallest – from 0.5 to 4.9 $\mu$ g/g. The contents do not correlate
188	with olivine Mg# and are lower than in high-Mg# olivine of mantle and magmatic genesis (Fig.
189	4).

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#### Trace element contents in olivine 191

192 Significant amounts of Li, B, Na, Sc, V, Cu, Zn, Sr, Y, Zr, Dy, Ho, Er, Tm, Yb and Lu were 193 detected in the studied olivine samples by LA-ICP-MS (Supplementary Table 2).

Li content varies from 0.15 to 61  $\mu$ g/g and shows a positive correlation with the content of P, 194 which varies from 5 to 377  $\mu$ g/g (Fig. 5a). 195

196 Y (<1.3 μg/g), Dy (<0.04 μg/g), Ho (<0.03 μg/g), Er (<0.22 μg/g), Tm (<0.08 μg/g), Yb (<1.1

 $\mu g/g$ ) and Lu (<0.31  $\mu g/g$ ) reveal clear positive correlations with Na content (from 0.15 to 14 197 198  $\mu g/g$ ) (Fig. 6).

199 Ti content varies widely from 3.7 to 159  $\mu$ g/g and does not correlate with Mg# or the contents of other elements. On average, Ti in studied MSSM olivine has nearly the same content as olivine 200 from magmatic rocks and mantle xenoliths (Fig. 7a). Boron content in studied MSSM olivine 201 ranges from 23 to 856 µg/g and similarly to Ti does not correlate with Mg# or with contents of 202 203 other elements. This B content is 2-3 orders of magnitude higher than in olivine from mantle 204 xenoliths (Fig. 7b).

Zn in studied MSSM olivine samples has a similar distribution pattern as Mn. Zinc contents for 205

206 most of studied olivine samples exhibit a clear negative correlation with Mg# and increases from

207 5-10  $\mu$ g/g at Mg#99.8 to ~380  $\mu$ g/g at Mg#86-88, which is higher than in olivine with similar

208	Mg# from igneous and mantle rocks (De Hoog et al., 2010; Kamenetsky et al., 2018; Nekrylov et
209	al., 2018; Neumann et al., 1999; Savelyev et al., 2018; Su et al., 2019) (Fig. 7c). The exceptions
210	are olivine samples PB-15/3, 75567 and 67792, which do not show any dependence of Zn
211	content on Mg#, and all have Zn content below 20 µg/g. Fe/Zn ratio in studied olivine varies
212	from ~300 to ~3000, which covers the range of Fe/Zn in olivine from igneous and mantle rocks
213	(mostly 1000-2000) (Fig. 7c). The measured contents of Sc (0.4-30 $\mu$ g/g), V (0.03-3.3 $\mu$ g/g), Cu
214	(up to 0.3 $\mu$ g/g) and Zr (0.05-17.7 $\mu$ g/g) do not show any clear distribution patterns.
215	
216	Discussion
217	Comparison of olivine from MSSM, magmatic rocks and mantle peridotites
218	Magmatic rocks often contain olivine xenocrysts of questionable origin (e.g. Brett et al., 2009;
219	Chayka et al., 2020; Plechov et al., 2017 etc.). Reliable discrimination of the olivine provenance
220	should be based on high quality reference chemical data. Such data exist now for olivine from
221	magmatic silicate rocks (e.g. Foley et al., 2013; Jaques and Foley, 2018; Neave et al., 2018;
222	Sobolev et al., 2007 etc.) and from mantle peridotites (e.g. Bussweiler et al., 2019; De Hoog et
223	al., 2010; Foley et al., 2013; Su et al., 2019). The composition of olivine from high-Mg
224	metamorphic rocks is very poorly known that hampers its identification using chemical criteria.
225	Studied olivine from MSSM has some distinctive features, which can be used to discriminate it
226	from magmatic and mantle olivine. The most prominent feature is low content of Co ( $<51 \mu g/g$ ),
227	Cr (< 5 $\mu$ g/g) and Ni (< 44 $\mu$ g/g) in comparison with olivine of the same Mg# (> 85) from mafic
228	magmatic rocks and mantle xenoliths (Fig. 4). This compositional feature is likely related to low
229	Co, Cr and Ni content in sedimentary dolomites (e.g. Carmichael et al., 2008; Ephraim, 2012;
230	Weber, 1964), which are common protoliths for MSSM.
231	Another distinctive feature of olivine from MSSM is high content of B (23-856 $\mu$ g/g). Boron
232	content in olivine of other origins is poorly known. The available data indicate concentrations of

 $233 < 0.5 \ \mu g/g$  in olivine from mantle xenoliths (with one anomalous grain containing 67  $\mu g/g$ ) (e.g.

Kaliwoda et al., 2008; Kent and Rossman, 2002), <21.5 µg/g in olivine from highly 234 metasomatized mantle veins in sub-arc mantle (Bouilhol et al., 2009), <8 µg/g in subcontinental 235 lithospheric mantle (Chayka et al., 2020), <23 µg/g in olivine from high-pressure serpentinites 236 237 (De Hoog et al., 2014),  $<11 \mu g/g$  in olivine from kimberlites and lamprophyres (Nosova et al., 2017; Nosova et al., 2018), <165  $\mu$ g/g in lamproites (4 ± 3.5  $\mu$ g/g without one anomalous 238 olivine crystal) (Chayka et al., 2020), ~4 µg/g in olivine from Stapafell tholeiite (Jeffries et al., 239 1995), ~12  $\mu$ g/g in olivine from Hawaiian ankaramite (Jeffries et al., 1995) and <1  $\mu$ g/g in arc-240 related magmatic olivine from South Sulawesi (Elburg et al., 2006). All of these data suggest 241 that the observed concentrations of B content in MSSM olivine are anomalously high in 242 243 comparison with olivine from other environments. Our results are in a good agreement with previously reported data on B enrichment in olivine from magnesian skarns (Grew et al., 1991; 244 Sykes et al., 1994). This compositional feature of olivine from MSSM is also likely linked to the 245 compositions of their protoliths – dolomites, which are strongly enriched in boron (Harder, 1970; 246 Weber, 1964) relative to mafic magmatic and mantle rocks. Such enrichment is caused by the 247 248 lithophile behavior of boron and its high solubility in aqueous fluids and enrichment in seawater 249 (e.g. Grew, 2015), which makes it abundant in marine sediments.

The new data on the composition of MSSM olivine can be used now to place constraints on the 250 251 origin of some previously enigmatic olivine xenocrysts of magmatic rocks. For example, we can now discuss the origin of extremely high-Mg olivine (Mg#>96) from lavas of Pian di Celle 252 volcano, Italy (Plechov et al., 2017). In our previous study we argued that this olivine is unlikely 253 of magnesian skarn origin because of low Mn and Ca content, not correlating with the 254 255 composition of host rocks (Plechov et al., 2018). This interpretation requires revision in light of 256 our new data. It is evident from our new data that low Mn and Ca contents of extremely high-Mg olivine described in lavas of Pian di Celle volcano are in fact typical for olivine with Mg# > 98257 258 from MSSM. This implies that typical metasomatic MSSM olivine can occur as xenocrysts in

volcanic rocks as well as olivine from primitive magmatic skarns (Fulignati et al., 2005; Gilg et

al., 2001; Wenzel et al., 2002).

Wang et al. (2020) have recently described an olivine-rich rock, which they called "jianite". This rock of uncertain origin was found among various Proterozoic felsic volcanic rocks and marbles of Mayihe formation (Ji'an County, Jilin Province, Northeast China) (Wang et al, 2020). LA-ICP-MS analyses of the olivine from "jianite" revealed extremely high content of B (1763  $\pm$  23  $\mu$ g/g) and low contents of Cr (0.92  $\pm$  0.08  $\mu$ g/g), Co (0.58  $\pm$  0.03  $\mu$ g/g) and Ni (0.91  $\pm$  0.06  $\mu$ g/g) (Wang et al., 2020). Our new data strongly suggest a MSSM-related origin of this olivine. Therefore, we suggest that jianites represent altered olivine-rich magnesian skarns.

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# 269 Occurrence of low-Ca olivine and monticellite in MSSM

270 Olivine from MSSM might be expected to have high Ca content, because CaO is a major component for these rocks, which mainly contain more than 50 vol.% of calcite. However, 271 olivine studied in our work has rather low Ca content (40-1880  $\mu$ g/g), which is similar to Ca 272 273 content in olivine from mantle peridotites and significantly lower than Ca content in olivine from 274 magmatic rocks (Fig. 4). It is however important to note that MSSM do not necessarily contain low-Ca olivine; some of them host pure monticellite (MgCaSiO<sub>4</sub>). Both low-Ca olivine and 275 276 monticellite occur in common MSSM association with calcite, spinel, phlogopite and diopside, sometimes in the same samples (Tracy et al., 1978, Wenzel et al., 2002). Thus, there is no 277 obvious host-rock compositional control on the crystallization of monticellite or olivine in 278 MSSM; both minerals appear to crystallize at relatively high CaO activity in fluid in contact 279 zones of magmatic intrusions and carbonates. 280

Stability of monticellite or olivine can be described via the following reaction:  $2MgCaSiO_4$ + $2CO_2 = Mg_2SiO_4 + 2CaCO_3 + SiO_2$ . According to this reaction, there can be also compositional control on the stability of these olivine polymorphs in MSSM. Crystallization of low-Ca olivine in association with calcite should be favored at high  $fCO_2$  and low  $aSiO_2$ . In opposite,

monticellite should be more stable at lower  $fCO_2$  and in more SiO<sub>2</sub> rich systems, for example,

during low-pressure (<1.5 kbar) contact metamorphism of clay-rich limestones.

In addition to system composition, temperature has important control on the incorporation of Ca 287 288 in olivine, promoting crystallization of more Ca-rich olivine with increasing temperature (e.g. Adams, Bishop, 1982, Shejwalkar, Coogan, 2013). It is thus likely that the occurrence of 289 290 monticellite or olivine is also controlled by the temperature of skarn formation so that low-Ca olivine forms at low temperature (generally <700 °C) and is replaced by monticellite at higher 291 292 temperature (e.g. Fig. 112 in Zharikov, 1970). Indeed some empirical data support this view as monticellite is usually found closer to the contact of scarns with intrusive body (Sinyakov, 293 294 1961).

Although more strict thermodynamic analysis is required to more precisely determine the 295 conditions of olivine crystallization in MSSM, our data allowed us to place semi-quantitative 296 constraints on these conditions. We propose that the paradoxically low Ca content in MSSM 297 olivine may reflect specific conditions of the olivine origin, which is crystallization from a CO<sub>2</sub>-298 299 rich fluid at low-temperature and low aSiO<sub>2</sub>. The olivine is replaced by monticellite at higher 300 temperature, lower  $fCO_2$  and at higher  $aSiO_2$ . This conclusion is in well agreement with the previously published data and thermodynamic modelling on the conditions of MSSM formation 301 302 (e.g. Bucher and Grapes, 2011; Pertsev, 1974; Zharikov et al., 2007).

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### 304 Incorporation of trace elements in MSSM olivine

The observed positive correlation of P and Li (Fig. 5a) is in a good agreement with the chargebalanced coupled substitution  $^{IV}Si^{4+} + ^{VI}(Fe, Mg)^{2+} = ^{IV}P^{5+} + ^{VI}Li^+$  (Buseck and Clark, 1984; Woodland et al., 2004). Sodium content reveals a weaker positive correlation with the content of P (Fig. 5b) than that of Li. However, a similar reaction could be the main mechanism of Na incorporation into the olivine structure (Mallmann et al., 2009). Indeed, molar Li + Na content nearly balance the P content for most of the studied olivine samples from MSSM (Fig. 5c).

Relatively high MSWD value for molar P and Li+Na 1:1 line suggest that there are most likely 311 involved other mechanisms of P incorporation into the olivine structure along with this one. 312 Excess of P relative to Li and Na observed in some samples (Fig. 5c) can be balanced by the 313 octahedral vacancies via reaction  $2^{IV}Si^{4+} + {}^{VI}(FeMg)^{2+} = 2^{IV}P^{5+} + {}^{VI}(vac)^0$  or trivalent cations 314 such as Cr, Al, Fe<sup>3+</sup> or even REE via reaction  $^{IV}Si^{4+} + {}^{VI}(FeMg)^{2+} = {}^{IV}P^{5+} + {}^{VI}(Cr, Al, Fe, Sc, I)$ 315 REE)<sup>3+</sup> (e.g. Milman-Barris et al., 2008; Shchipalkina et al., 2019; Shea et al., 2019). However, 316 we found no significant correlation of the P excess with the amount of trivalent cations in MSSM 317 olivine. Therefore, we propose that the former reaction involving octahedral vacancies is the 318 most likely mechanism for the incorporation of the P excess in the MSSM olivine. 319

320 The prevailing mechanisms of REE and Y incorporation in olivine structure is a matter of longyear discussion. In a recent study, Burnham and O'Neill (2020) concluded that these elements 321 replace  $Mg^{2+}$  on octahedral sites, but with charge-balance achieved by two different 322 mechanisms: (1) cation vacancies (2  $REE^{3+}$  + vacancy = 3  $Mg^{2+}$ ); and (2) substitution of Al for 323 Si  $(REE^{3+} + Al^{3+} = Mg^{2+} + Si^{4+})$ . These results confirm conclusions of previous studies (e.g. 324 Colson et al. 1989; Beattie, 1994; Di Stefano et al., 2019; Spandler and O'Neill, 2010), but show 325 that REE can enter olivine simultaneously using different mechanisms, depending on SiO<sub>2</sub> and 326 Al<sub>2</sub>O<sub>3</sub> activity in the system. 327

Alternatively it was suggested on the basis of thermodynamic calculations that REE substitution 328 on the octahedral site can be charge-balanced by monovalent cations via reaction  $2^{VI}$  (Fe, Mg)<sup>2+</sup> = 329 <sup>VI</sup>(H, Li, Na)<sup>+</sup> + <sup>VI</sup>(Y, REE)<sup>3+</sup> (e.g., Colson et al, 1989). The coupled substitution of 2Mg with 330 REE and proton in olivine (REESiO<sub>3</sub>(OH)) has been shown by Berry et al. (2007). Coupling 331 with Li was not considered as important mechanism because of low Li abundance in most 332 333 olivine-bearing rocks. The mechanism involving Na has been proposed as potentially significant (e.g., Burnham, O'Neill, 2020), but it was not studied experimentally in sufficient details thus 334 335 far.

Our data show general positive correlation of Y+REE and Na contents in MSSM olivine (Fig. 6). 336 This correlation may suggest that Na-rich fluids, from which the olivine crystallized, are also 337 enriched in REE+Y, and thus this correlation reflects the compositional peculiarity of fluid 338 339 composition. On the other hand, the correlation may also suggests the coupled incorporation of REE+Y and Na in MSSM olivine. Because only  $\sim 2$  % of all Na atoms in olivine are required to 340 charge-balance REE+Y content in olivine (Fig. 6), this mechanism does not contradict the 341 previous conclusion about the coupled P, Li and Na incorporation into the MSSM olivine 342 structure. 343

Although the exact mechanisms of REE+Y incorporation in MSSM olivine are uncertain, we propose that the coupled substitution of REE+Y with monovalent cations, particularly with Na and H, as well as coupled substitution with Al in tetrahedral site maybe important in low temperature hydrothermal environment, where olivine crystallize from  $CO_2$ -H<sub>2</sub>O fluid at low SiO<sub>2</sub> activity. In contrast, the mechanism involving cation vacancies is likely more important at high SiO<sub>2</sub> activity in magmatic systems.

- 350
- 351

# Conclusions

We present results of a systematic study of major, minor and trace elements contents in olivine from MSSM. A total of 15 olivine samples originate from nine different locations in three countries and are part of the collection of the Fersman Mineralogical Museum RAS (Russia). Our data allowed us to place constraints on the compositional variations of this olivine, highlight its distinctive features and provide new insights on the incorporation of trace elements in the olivine structure. Our main conclusions are the following:

Mg# of MSSM olivine varies from ~ 86 to nearly 100. The olivine has anomalously low
 contents of Ni, Co and Cr and high content of B inherited from sedimentary carbonate
 rocks, protoliths for MSSM. These features can be used to discriminate MSSM olivine
 from magmatic and mantle olivine.

362 2) Low-Ca olivine in MSSM may reflect crystallization from  $CO_2$  –rich fluid at relatively 363 low temperature (<700 °C) at low  $aSiO_2$ . At higher temperature and/or lower  $fCO_2$  and 364 higher  $aSiO_2$ , the olivine is expected to be replaced by monticellite via reaction with 365 calcite and SiO<sub>2</sub> in fluid.

3) Strong correlation of the total amount of Li and Na with the content of P suggests a coupled incorporation of these elements in the MSSM olivine via heterovalent substitution at the octahedral and tetrahedral sites:  ${}^{IV}Si^{4+} + {}^{VI}(Fe, Mg)^{2+} = {}^{IV}P^{5+} + {}^{VI}(Li,$ Na)<sup>+</sup>.

4) Correlation of REE+Y and Na contents suggests that REE+Y could be incorporated into the MSSM olivine structure via heterovalent substitution at the octahedral site:  $2^{VI}$ (Fe, Mg)<sup>2+</sup> = <sup>VI</sup>Na<sup>+</sup> + <sup>VI</sup>(Y, REE)<sup>3+</sup>. In contrast to high temperature magmatic systems, this mechanism maybe more important for olivine crystallizing at low temperature in low aSiO<sub>2</sub> environments such as in magnesian scarns.

Although our results provide a significant piece of information about the occurrence and composition of olivine in MSSM, we have to admit that the comprehensive understanding of the conditions of olivine crystallization in these rocks is currently missing. In particular, the conditions of crystallization of low-Ca olivine or monticellite in MSSM should be investigated in more details.

Besides MSSM, there is another class of rocks, where compositional variations of olivine were poorly characterized so far. These rocks are carbonatites, which could host olivine similar to that from MSSM by some parameters, for example, by extremely high Mg# approaching 100 (e.g. Guzmics et al., 2011). Although one can expect an elevated content of Ca in olivine from magmatic carbonatites (e.g. Di Stefano et al., 2018), there are reports of both relatively high (Guzmics et al., 2011) and very low Ca contents (Xie et al., 2019). The low Ca olivines from carbonatites are currently undistinguishable from olivine of MSSM origin. This could indicate a

similar metasomatic origin, or could be due to a lack of precise compositional data. Thus, further

388 systematic study of olivine from carbonatites is required to fill this gap in our knowledge.

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# 647 **Figure captions:**

Fig. 1. Schematic map with marked sampling localities. Numbers (N), sample names andlocations correspond to those in Table 1. The map is drawn using https://mapchart.net website.

Fig. 2. Comparison of EPMA and LA-ICP-MS data on the olivine composition from the same samples. Sample names correspond to those in Table 1 and on Figure 1. The error bars correspond to two standard deviations. Uncertainties of primary standards are not included, because they contribute insignificantly to the total uncertainty (Batanova et al., 2015, 2019). Linear regressions are calculated following the method of York et al. (2004), uncertainties of the regressions are calculated following Ludwig (1980).

- 656 Fig. 3. Petrographic features of studied MSSM samples. a Ol-Cal-Sp association of the PB-15-
- 657 3 sample in plane polarized (left) and crossed polarized (right) light; b Ol-Cal-Phl association
- of the WS-15-4b sample in in plane polarized (left) and crossed polarized (right) light; c back-
- scattering electrons image of Ol-Cal-Sp-Phl association of the 50363 sample with spots of LA-
- 660 ICP-MS analyses; d back-scattering electrons image of highly serpentinized olivine of the
- 661 74186 sample with spots of LA-ICP-MS analyses. Sp spinel, Cal calcite, Phl phlogopite,
- 662 Srp serpentine, Mt magnetite, Ilm ilmenite.
- Fig. 4. Minor elements (Mn, Ni, Al, Ca, Cr and Co) contents in studied skarn and silicate marble
- olivine samples (LA-ICP-MS data, Supplementary Table 2) in comparison with magmatic (De
- Hoog et al., 2010; Sobolev et al., 2007) and mantle olivine (De Hoog et al., 2010; Su et al.,
- 666 2019). Sample names correspond to those in Table 1 and on Figure 1.
- Fig. 5. P, Li and Na content in studied olivine samples (a, b) and comparison of Li + Na with P
  concentrations calculated as atoms per formula unit (apfu) (c). Sample names correspond to
  those in Table 1 and on Figure 1. Linear regression for the comparison of Li + Na with P (in
- apfu) is calculated following the method of York et al. (2004), uncertainties of the regression are
- 671 calculated following Ludwig (1980).

- Fig. 6. Dependence of measured Y, Dy, Yb and total REE contents in studied olivine samples
- from the content of Na. Sample names correspond to those in Table 1 and on Figure 1.
- Fig. 7. Ti (a), B (b) and Zn (c) contents versus Mg# in studied olivine samples. Sample names
- 675 correspond to those in Table 1 and on Figure 1. Ti and Zn contents in mantle olivine are from De
- Hoog et al. (2010) and Su et al. (2019), B on the basis of Kent and Rossman (2002) and
- 677 Chayka et al. (2020) data; Ti and Zn contents in magmatic olivine are shown for olivine from
- 678 Baffin Island picrites (De Hoog et al., 2010), basalts from Canary Islands (Neumann et al.,
- 1999), MORBs (Savelyev et al., 2018), arc basalts (Kamenetsky et al., 2018) and back-arc basalt
- 680 (Nekrylov et al., 2018). B contents in magmatic olivine are shown only for kimberlites,
- lamproites and lamprophyres (Chayka et al., 2020; Nosova et al., 2017; Nosova et al., 2018) due
- to the lack of data on other magmatic suites.

	Sample		
N	#	Location	Short sample description
	WS-15-		
1	4b	Belaya Vyemka (Russia)	Phl-Ol skarn
2	PB-15-1	Tazheran Massif (Russia)	Ol-Sp skarn
3	PB-15-3	Tazheran Massif, Perovskite mine (Russia)	Ol-Sp skarn
4	74607	Tazheran Massif (Russia)	Fo with Di
5	75567	Tazheran Massif (Russia)	Fo
6	62949	Kukhilal (Tajikistan)	Ol Mg-skarn
7	82188	Kukhilal (Tajikistan)	Ol Mg-skarn
8	67723	Katalakh, gallery #2 (Russia)	Ol Mg-skarn
		Emeldzhak ore deposit, Taborny site, Gorely mine	
9	67792	(Russia)	Ol Mg-skarn
10	74186	Emeldzhak ore deposit (Russia)	Amp-Ol Mg-skarn
11	68009	Timpton river, Leglier mine (Russia)	Amp-Ol Mg-skarn
12	77329	Timpton river, Kurung-Honku (Russia)	Phl-Ol Mg-skarn
13	32900	Nikolai-Maximilian mine (Russia)	Fo with Cal, Prv and Di
14	50363	Hodzha-Achkan alkaline massif, Tilbe pass (Kyrgyzstan)	Fo
15	81924	Slyudyanka, Uluntui pad' (Russia)	Fo

Table 1: List of studied MSSM samples from Fersman Mineralogical Museum RAS.



me	Object
	Belaya Vyemka (Russia)
4607	Tazheran (Russia)
188	Kukhilal (Tajikistan)
	Katalakh (Russia)
186	Emeldzhak (Russia)
329	Timpton (Russia)
71	Nikolai-Maximilian (Russia)
	Hodzha-Achkan (Kyrgyzstan)
	Slyudyanka (Russia)





# Figure 3









