Revision 1

1	XANES Spectroscopy of Sulfides Stable under Reducing Conditions
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10

Abstract

11	X-ray absorption near-edge structure (XANES) spectroscopy is a powerful technique to
12	quantitatively investigate sulfur speciation in geologically complex materials such as minerals,
13	glasses, soils, organic compounds, industrial slags, and extraterrestrial materials. This technique
14	allows non-destructive investigation of the coordination chemistry and oxidation state of sulfur
15	species ranging from sulfide (2- oxidation state) to sulfate (6+ oxidation state). Each sulfur
16	species has a unique spectral shape with a characteristic K-edge representing the $s \rightarrow p$ and d
17	hybridization photoelectron transitions. As such, sulfur speciation is used to measure the
18	oxidation state of samples by comparing the overall XANES spectra to that of reference
19	compounds. Although many S XANES spectral standards exist for terrestrial applications under
20	oxidized conditions, new sulfide standards are needed to investigate reduced (oxygen fugacity,
21	fO ₂ , below IW) silicate systems relevant for studies of extraterrestrial materials and systems.
22	Sulfides found in certain meteorites (e.g., enstatite chondrites and aubrites) and predicted to exist
23	on Mercury, such as CaS (oldhamite), MgS (niningerite), and FeCr ₂ S ₄ (daubréelite), are stable at
24	fO ₂ below IW-3 but rapidly oxidize to sulfate and/or produce sulfurous gases under terrestrial
25	surface conditions. XANES spectra of these compounds collected to date have been of variable
26	quality, possibly due to the unstable nature of certain sulfides under typical (e.g., oxidizing)
27	laboratory conditions. A new set of compounds were prepared for this study and their XANES
28	spectra are analyzed for comparison with potential extraterrestrial analogs. S K-edge XANES
29	spectra were collected at Argonne National Lab for FeS (troilite), MnS (alabandite),
30	CaS(oldhamite), MgS (niningerite), Ni _{1-x} S, NiS ₂ , CaSO ₄ (anhydrite), MgSO ₄ , FeSO ₄ , Fe ₂ (SO ₄) ₃ ,
31	$FeCr_2S_4$ (daubréelite), Na_2S , Al_2S_3 , Ni_7S_6 , and Ni_3S_2 ; the latter five were analyzed for the first
32	time using XANES. These standards expand upon the existing S XANES endmember libraries at

a higher spectral resolution (0.25 eV steps) near the S K-edge. Processed spectra, those that have
been normalized and 'flattened', are compared to quantify uncertainties due to data processing
methods. Future investigations that require well-characterized sulfide standards such as the ones
presented here may have important implications for understanding sulfur speciation in reduced
silicate glasses and minerals with applications for the early Earth, Moon, Mercury, and enstatite
chondrites.

Keywords: XANES spectroscopy, chemical state of S, oxidation state of S, sulfides, sulfates,
endmember variability.

41

Introduction

Along with iron, sulfur is the most important heterovalent element in geologic systems. 42 43 Due to its range of charges from 2- to 6+, the behavior of sulfur is complex. Under different redox conditions, sulfur can bond with both more electropositive and more electronegative 44 45 elements in magmas, soils, industrial glasses and slags, and meteorites (Fleet 2005). This has significant effects on the partitioning behaviors of these elements between silicate melts (S²⁻ and 46 S^{6+}), liquid metals and sulfides (S^0 and S^{2-}), gases (H₂S, S₂, SO₂, and SO₃), and solids (S^{2-} , S^{1-} , 47 S^0 , S^{2+} , S^{4+} , and S^{6+}). Under reducing conditions inferred for enstatite chondrite and aubrite 48 meteorite parent bodies, as well as for lunar and Mercurian magmas, sulfur is stable as a number 49 of different sulfide compounds not typically found on Earth, and S is the most abundant volatile 50 51 element dissolved in silicate melts (>1 wt%). Sulfides stable below IW-3 include CaS (oldhamite), MgS (niningerite), MnS (alabandite), FeCr₂S₄ (daubréelite), NaCr₂S₂ 52 (caswellsilverite), and djerfisherite ($K_6Na(Fe^{2+},Cu,Ni)_{25}S_{26}Cl$), all of which rapidly oxidize to 53 54 sulfate and/or produce sulfurous gases in air under terrestrial surface conditions.

55	Reduced silicate systems have applications for terrestrial mid-ocean ridge basalt
56	(MORB), meteorites such as enstatite chondrites and aubrites, and magmas on the Moon and
57	Mercury. Mercury is the most extreme example with surface sulfur detections of 1.5-4 wt%
58	(Nittler et al. 2011). Recent work on sulfide solubility in experimental Mercurian glasses (Namur
59	et al. 2016b) predicts MgS and CaS should be the dominant sulfides in extremely reduced silicate
60	melts, while our analyses of sulfide speciation have confirmed and quantified MgS and CaS in
61	reduced S-rich and Fe-poor silicate melts. Therefore, sulfur speciation is critical to understanding
62	the thermodynamics of these sulfur-rich silicate systems and has been shown to influence silicate
63	phase equilibria (Namur et al. 2016b, 2016a) and physical properties such as density,
64	polymerization, and viscosity (Holzheid and Grove 2002; Robert A. Fogel 2005; Namur et al.
65	2016b).
66	Sulfur speciation can be determined through a number of spectroscopic techniques
67	including x-ray emission and absorption spectroscopy, nuclear magnetic resonance (NMR),
68	Raman spectroscopy, and Infrared spectroscopy (Wilke et al. 2011). XANES is advantageous
69	because sulfur oxidation state and individual species can be determined quantitatively at
70	concentrations down to ~100 ppm with ~1 micrometer spot size. The XANES analyses can also
71	be compared with spectra estimated by ab initio methods (Fleet 2005).
72	X-ray Absorption Spectroscopy (XAS) works by absorbing x-ray photons into a core
73	orbital level followed by photoelectron emission at a characteristic energy. The resultant spectra
74	can be loosely divided into two main energy regions, X-ray Absorption Near Edge Structure
75	(XANES) and Extended Absorption Fine Structure (EXAFS) that are shown in an example
76	spectrum of MgS (niningerite) in Figure S1. XANES is the region within ~50 eV of the
77	absorption edge representing a Fermi level transition, which in this case is the primary S K-edge

corresponding to the photoelectron transition S $1s \rightarrow 3p$. Secondary absorption peaks shown in the near-edge structure may result from transitions to empty S 3d states or hybridization of S pstates with metal 3d states (i.e. Mg) (Fleet 2005). EXAFS is the oscillatory region ~30eV above the absorption edge that probes the local environment of atoms. For light elements such as S, the information EXAFS can provide is limited because the oscillatory structure decays rapidly to a line past the edge (e.g. Bunker & Stern, 1984; Fleet, 2005). For a more detailed overview of the theory and practice of S XANES see Fleet (2005) and references therein.

85 A number of sulfur species standards have been analyzed in previous studies, including FeS (troilite), MnS, CaS, MgS, Ni_{1-x}S, NiS₂, CaSO₄ (anhydrite), MgSO₄, FeSO₄, and Fe₂(SO₄)₃ 86 (e.g. Fleet, 2005; Langman et al., 2015). These data indicate that the S K-edge moves to higher 87 energies as the sulfide bonding environment shifts from metallic/covalent to ionic, and from 88 reduced (S^{2-}) to oxidized (S^{6+}) species as seen in Figure 1. Here, a new set of S XANES 89 90 standards for sulfide and sulfate compounds were synthesized. By protecting the samples from 91 air throughout synthesis, transport, and analysis, these standards are expected to be stoichiometrically stable during high-precision spectral measurements. S K-edge XANES 92 93 measurements were made at beamline 13-IDE at Argonne National Lab, which is tunable to 0.25 94 eV spectral resolution in the energy range of sulfur with a focused beam resolution of $\sim 1 \,\mu m$. New S K-edge XANES spectra are presented for FeS (troilite), MnS (alabandite), 95 96 CaS(oldhamite), MgS (niningerite), Ni_{1-x}S, NiS₂, CaSO₄ (anhydrite), MgSO₄, FeSO₄, Fe₂(SO₄)₃, FeCr₂S₄ (daubréelite), Na₂S, Al₂S₃, Ni₇S₆, and Ni₃S₂. Below, we also present the range in 97 processed XANES spectra that can result from different choices in data processing steps and 98 discuss how the resulting variations translate to uncertainties in the final processed spectra of the 99 100 S standards. This type of endmember spectral variability has been extensively studied for remote

101	sensing applications of reflectance spectroscopy over the past several decades (Tompkins et al.
102	1997; Settle 2006; Somers et al. 2011), and it is similarly important to understand how
103	uncertainties associated with processing of XANES data propagate through spectral unmixing of
104	such data for multi-component mixtures (e.g., materials with multiple S-bearing phases).
105	Methods
106	Experimental Methods
107	Pure sulfides were synthesized by reaction of element powders or purchased from Alfa
108	Aesar©. MgS (niningerite) was synthesized from Mg metal powder and excess S powder in a
109	sealed evacuated silica tube following the method of Osborne & Fleet (1984). Mg was reacted
110	with S at 600 °C for 1 day and 700 °C for 2 days. The product was then crushed and reloaded
111	with excess S in a 15 cm long silica tube with the reagents located in the hotspot of the
112	horizontal furnace at 700 °C for 1 day and 900 °C for 1 day. Crushing and reloading decreases
113	the effect of armoring that limits reaction progress. FeS (troilite) was synthesized from Fe metal
114	powder and S powder at stochiometric proportions in a sealed evacuated silica tube held at 800
115	$^{\circ}$ C for 48 hours. NiS ₂ (vaesite), Ni _{1-x} S, Ni ₇ S ₆ , and Ni ₃ S ₂ (heazlewoodite) was synthesized from
116	Ni metal powder or Ni metal rod and S powder at stochiometric proportions in a sealed
117	evacuated silica tube similarly held at 800 °C for 48 hours. FeCr ₂ S ₄ (daubréelite) was
118	synthesized from reagent grade powders in experimental Mercurian samples at $Ts > 1250$ °C and
119	1GPa for 24 hours in a piston cylinder apparatus at Brown University. The Mercurian sample
120	composition was a modified CH chondrite composition ALH85085 (Weisberg et al. 1988). CaS
121	(oldhamite), Na ₂ S, MnS (alabandite), and Al ₂ S ₃ of >99.9% purity were purchased from Alfa
122	Aesar©.

123	Fe, $FeCr_2S_4$, Ni_{1-x} , Ni_7S_6 , and Ni_3S_2 were mounted in epoxy and polished under water
124	using steps of 600 and 1200 grit sandpaper, and 0.5 and 0.03 micrometer Al_2O_3 grit. All other
125	sulfides were kept in powder form. All sulfides were stored and transported in a nitrogen filled
126	desiccator to minimize potential reaction with oxygen and water.

127 Analytical Methods

- 128 Experimental sulfide compositions were confirmed using the Cameca SX100 electron
- 129 microprobe (EMPA) or by powder x-ray diffraction (XRD) using a Bruker D2 Phaser
- instrument. For EMPA analyses, metals were used as Ti, Mn, Ni, Cr, and Fe standards, while
- diopside was used as the Ca and Si standard. Forsterite was used as the Mg standard. Pyrite was
- used as the S standard. Synthesized sulfides were analyzed using a 15 keV accelerating voltage,
- a 10 nA beam current, and a 1 µm spot size. Based on repeated analyses, purity is >95% for MgS

134 (~5% MgO) and FeCr₂S₄ (<5% Mn) and >99% for FeS, NiS₂, Ni_{1-x}S, Ni₇S₆, and Ni₃S₂ with

- relative errors less than 0.5%. MgS (niningerite) was left in powdered form to prevent oxidation
- during any subsequent sample preparation and thus analyzed using XRD. No sulfate peaks were
- 137 detected in the XRD or XANES data during repeated analyses.

138	S K-edge XANES spectra were collected inside a helium environment using an
139	undulator-based, hard x-ray microprobe at beamline 13IDE of GSECARS, Argonne National
140	Laboratory. The spot size for XANES spectroscopy was 1-2 micrometers, with a nominal
141	penetration depth of 2-3 micrometers for S. XANES spectra were collected from 2442 to 2542
142	eV with a step size of 0.25 eV near the S K-edge and 2 eV in the other energy regions (1 s
143	acquisition time per step).

144 Data Processing Methods

Processing of XANES spectra included converting raw data to values of absorption 145 146 coefficient (μ (E)). S K-edge XANES spectra were calibrated with the white line maxima energy for FeS (troilite) at 2470.30 eV, CaS at 2474.30 eV, MgS at 2473.60 eV, CaSO₄ (anhydrite) at 147 148 2481.90 eV, and MgSO₄ at 2481.70 eV. Dead-time corrections were completed at APS. XANES 149 spectra were deglitched in Athena to remove anomalous data points associated with Bragg spots 150 and irregular x-ray input energies (Ravel and Newville 2005). Additionally, the spectra for Al₂S₃ 151 were smoothed to reduce the noise in the data using a 3-point boxcar average. These steps account for analytical, instrument, and facility errors and uncertainty. Two steps, normalization 152 153 and flattening, can be applied to XANES spectra. Although these steps are common in spectral processing, they require input from the user and can thus be a source of uncertainty in the 154 155 absolute values of the final spectra. We evaluate and discuss below how these user inputs may 156 affect the final processed spectra of our S-bearing samples.

157 Spectra were normalized using an edge step normalization algorithm that fits the energy 158 region before the K-edge and the energy region after the edge using separate linear regressions 159 (Ravel & Newville, 2005), as shown in **Figure S1**. The raw data is subtracted by the pre-edge

160	curve and then divided by the edge-step, where the edge step is the difference between the post-
161	edge curve and pre-edge curve at E0. E0 is a parameter that represents the k=0 continuum level
162	at which electrons have just enough energy to propagate through material. Normalization of
163	XANES spectra allows data to be directly compared regardless of the details of the experiment
164	by regularizing the effects of sample preparation, sample thickness, absorber concentration, and
165	detector and amplifier settings (Ravel & Newville, 2005). Normalization operates to remove the
166	pre-edge background by setting the absorption intensity values in the pre-edge region to 0 and to
167	scale data to a per-atom basis by scaling the edge jump to 1. Example raw and normalized
168	spectra for MgS (niningerite) and CaS (oldhamite) are shown in Figure 2.
169	Flattening is an optional processing step after normalization that scales the absorption
170	values in the EXAFS region (oscillatory part of the data after the absorption edge) to be near the
171	y=1 line, thus minimizing the spectral slope at post-edge energies. To 'flatten' a XANES
172	spectrum, the normalized spectrum is subtracted by the difference between the normalized post-
173	edge curve and normalized edge-step at E0, where E0 is calculated as the first inflection point on
174	the absorption edge.

175 XANES normalization optimization allows one to quantify the uncertainty introduced in the processed spectra when choosing pre-edge and post-edge normalization bounds. Once 176 quantified, the uncertainty that results from this processing step may be incorporated into 177 178 weighted linear combination fits of spectra of multicomponent samples, and this may be particularly important if spectra of potential endmember components exhibit K-edge positions 179 that are close to one another. During normalization and flattening there are 4 free parameters 180 (referred to here as PER1, PER2, NR1, and NR2), excluding E0. PER1 and PER2 represent the 181 energy positions that define the bounds of the spectral data that are used in the linear regression 182

183	of the pre-edge region (Figure S1). Similarly, NR1 and NR2 define the energy bounds of the
184	spectral data that is fit in the linear regression of the post-edge region (Figure S1). During
185	spectral normalization and flattening, the bounds of the pre-edge were varied for PER1 and
186	PER2 between the lowest-energy value of the spectrum plus 5 eV and E0-14:-4 in steps of 1 eV
187	respectively. The bounds of the post-edge, NR1 and NR2, were allowed to vary between
188	E0+5:30 and the highest-energy value of the spectrum minus 10 eV in steps of 1 eV respectively.
189	This resulted in 29766 unique combinations of input parameters and thus 29766 processed
190	XANES spectra per sample.
191	Results
192	Normalization Optimization
193	Normalization and flattening of XANES spectra are deemed to be satisfactory when they
194	remove the pre-edge background by setting the pre-edge region to 0 and scale the edge jump to
195	1, allowing direct comparison of spectra. As shown in Figure 3, there are a number of possible
196	solutions for the linear regressions to the pre- and post-edge regions depending on the range of
197	range of data chosen for the fit (i.e., positions of PER1, PER2, NR1, and NR2). However, it is
198	logical to take the furthest points (PER1 and NR2) along with middle points that allow the fit to
199	bisect data spread within the spectral region of interest. In the pre-edge case, the middle point
200	would preferably be before any pre-edge or K-edge absorption feature. In the post-edge case, the
201	middle point would be any point that allows the EXAFS region to be bisected by the fit (which
202	may be many as the EXAFS region is characterized by its oscillatory nature). Good
203	normalization has a convergent solution with minimal error and effect on spectral shape as
204	determined by slope and RMSE of the pre-edge and edge-step regions. Good normalizations

were determined by setting the tolerance of the edge-step to between 0.95 and 1.05, the pre-edge

RMSE below 0.3, and the pre-edge slope desired to be y=0. The top ten and best normalizations were chosen by ranking the processed spectra based on root mean square error of the pre-edge slope, pre-edge RMSE, and edge-step RMSE. The full range, reduced range (top 10), and best normalized and flattened spectra are shown in **Figure 3a and 4b**.

210

Spectra of New Sulfur Standards

211 New S K-edge XANES standards were collected for a number of S species including FeS, MgS, CaS, MnS, Ni_{1-x}S, NiS₂, CaSO₄ (anhydrite), MgSO₄, FeSO₄, and Fe₂(SO4)₃, FeCr₂S₄ 212 213 (daubréelite), Na₂S, and Al₂S₃, Ni₇S₆, and Ni₃S₂. All S species XANES spectra are shown as normalized spectra in Figure 4, Figure 5, and Figure 6 with flattened spectra in Figures S2 and 214 215 Figure S3. Unlike sulfite and sulfate species, sulfides with different cations display characteristic 216 XANES spectra K-edge energies and shapes (Table 1). XANES spectra of MnS, $Ni_{1-x}S$, NiS_2 217 and CaSO₄ are comparable to previous measurements (e.g. Fleet, 2005) except FeS (troilite), 218 CaS (oldhamite), and MgS (niningerite) all have lower K-edge energies. The primary K-edge of FeS, CaS, and MgS are measured at 2469.3(1) eV, 2471.2(1) eV, and 2472.0(1) eV compared to 219 2470.0 eV, 2474.0 eV, and 2475.1 eV respectively (Fleet 2005). Additionally, we find MgS has 220 221 3 resolvable K-edge peaks instead of just the 2 present in the data of Fleet (2005), highlighting 222 the advantage of higher spectral resolution data (Figure 1 compared to Figure 4). Spectra of FeS (troilite) and Na₂S exhibit similar K-edge energy of 2469.3 eV indicating the Na-S bond is very 223 224 metallic. A weak sulfate peak at 2482 eV is observed in the CaS and Na₂S spectra, likely due to their rapid reactivity in air to form trace amounts of sulfate (Figure 4). The XANES spectrum of 225 FeCr₂S₄ (daubréelite) is very similar to that of pure FeS (troilite) (Figure 4), which suggests the 226 structure and bonding environments of Fe and Cr are very similar. All nickel sulfides exhibit 227

similar spectra with no systematic trend in K-edge position or peak strength due to differing Ni/S
ratios (Figure 5).

230

Discussion

231 S XANES K-edge Location

The location of the S K-edge and overall shape of the XANES spectra depends on the 232 233 cation, bond character, and bonding environment. Sulfide species with low-lying empty 3d orbitals (transition metal sulfides including Fe, Ni, and Cr sulfides) have an initial absorption 234 peak due to hybridization of the S antibonding and metal 3d states and a second broad absorption 235 236 peak corresponding to transitions to S p-like states hybridized with the metal 4sp orbitals (Farrell and Fleet 2001; Farrell et al. 2002b; Kravtsova et al. 2004; Soldatov et al. 2004; Fleet 2005) as 237 238 shown in Figures 4 and 5. Sulfide species with bonds that are more covalent/metallic are 239 expected to have lower K-edges than those with ionic bonds (Fleet 2005). S K-edge energy 240 increases as the ionic character of the bond increases (Pauling 1960) showing a linear trend for 241 monosulfides with a cation charge of 2+ in Figure 7. Nickel disulfide has a higher K-edge energy due to higher mean oxidation state of the sulfur atom relative to the nickel monosulfide 242 243 similar to FeS₂ compared with FeS (Head et al. 2018). However, bond length and structural 244 effects also influence the K-edge energy influencing the nickel sulfide trends for monoclincic $(Ni_{1-x}S)$, orthorhombic (Ni_7S_6) , isometric (NiS_2) , and trigonal (Ni_3S_2) nickel sufides (Fleet 1972). 245 246 As the oxidation state of S increases (2- for sulfides to 6+ for sulfates), the energy of the S Kedge increases as well (Li et al. 1995; Fleet 2005) (Table 1). Ultimately, the K-edge energy of 247 the S species correlates linearly with an increase in the direct energy-band gap (Li et al. 1994). 248

249 XANES Normalization

250	The chief aim of normalization, whether or not a spectrum is subsequently flattened, is to
251	regularize data to allow for direct comparison of different samples measured under different
252	conditions. The three main methods of normalization are functional normalization, edge-step
253	normalization, and flattening. Edge-step normalization is used here to avoid issues of inverted
254	regions of a functional normalized spectrum due to division by negative background absorption
255	possibly arising from detector settings (Ravel and Newville 2005). Unfortunately edge-step
256	normalization can introduce a small amount of attenuation that is linear with energy and
257	quadratic in wavenumber, but does represent a much less severe issue than functional
258	normalization because it does not invert portions of a spectrum (Ravel and Newville 2005).
259	As expected, normalized spectra exhibit higher variance in the post-edge EXAFS region
260	(standard error ≈ 0.005 -0.01 for normalized and standard error ≈ 0.001 -0.01 for flattened),
261	whereas the flattened spectra exhibit higher variance in the pre-edge XANES region (standard
262	error ≈ 0.00001 -0.001 for normalized and standard error ≈ 0.01 -0.05 for flattened) as shown in
263	Figure 3c and 3d with standard errors reported in supplementary tables. Variance is still highest
264	at the K-edge and near-edge structure (standard error ≈ 0.01 -0.02 for sulfides and standard error
265	\approx 0.01-0.05 for sulfates). Although flattened spectra can be beneficial for displaying the data and
266	linear combination fitting (Webb 2005), it does inherently introduce processing error into the
267	pre-edge region. Thus, we recommend using edge-step normalized XANES spectra for XANES
268	spectroscopy deconvolution studies, especially if pre-edge features are important in the spectral
269	fits. Nevertheless, K-edge position is insensitive to the choice of the processing parameters; the
270	K-edge is found to vary by 0.1 eV for all sulfides and less than 0.3 eV for all sulfates as seen in
271	Table 1.

272

Implications

273	X-ray absorption near-edge structure (XANES) spectroscopy is an excellent technique to
274	non-destructively investigate sulfur speciation regardless of sample type for both geologic and
275	industrial applications. S K-edge XANES spectra are presented for FeS, FeCr ₂ S ₄ (daubréelite),
276	MgS, CaS, MnS, Na ₂ S, Ni ₇ S ₆ , Ni ₃ S ₂ , Ni _{1-x} S, NiS ₂ , Al ₂ S ₃ , CaSO ₄ (anhydrite), MgSO ₄ , FeSO ₄ , and
277	Fe ₂ (SO4) ₃ . These new endmember spectra will be especially useful in deconvolving reduced
278	silicate glasses that may be dominated by (Mg,Ca)S (Robert A Fogel 2005; Stockstill-Cahill et
279	al. 2012; Namur et al. 2016b) and other sulfides stable under reducing conditions rather than
280	FeS. While this represents an important expansion of the existing S XANES endmember library,
281	more analyses of S-bearing phases are needed. It is of interest to measure alkali sulfides such as
282	caswellsilverite (NaCrS ₂) and djerfisherite (K ₆ Na(Fe ²⁺ ,Cu,Ni) ₂₅ S ₂₆ Cl) that are common in
283	reduced enstatite chondrites but difficult to synthesize in pure forms. Fortunately, XANES
284	analyses of sulfide solid solutions have shown that the S K-edge peaks vary with local cation
285	environment (or average cation environment) in structurally similar sulfides to their constituent
286	parts as long as the proportion of Fe remains the same (Farrell and Fleet 2001; Farrell et al.
287	2002a). S K-edge XANES spectra for solid solutions of (Mn,Fe)S and (Mg,Fe)S have shown that
288	the pre-edge peak area does not increase proportional to increasing Fe and progressive
289	participation of 3 <i>d</i> orbitals in metal-S bonding (Farrell et al. 2002a). Also, because the structure
290	is factored into the multiple scattering and resultant spectra, one might think that synthesizing
291	and measuring structurally similar endmembers (glass vs. crystalline) would improve
292	deconvolution fits of spectra of complex mixtures. In fact, some electron microprobe standards
293	have been developed for both crystalline and glassy materials. In the future, MgS dissolved in
294	enstatite or forsterite glass could improve deconvolution of silicate glass XANES spectra. In
295	addition, questions remain as to whether there is a normalization algorithm that minimizes

296	uncertainties due to spectral processing methods across the entire spectrum. The spectral
297	uncertainty measured for these new sulfur standards and presented here will also be useful in
298	determining how endmember variability affects quantitative spectral unmixing of XANES
299	spectra of materials containing multiple S-bearing phases. These new sulfide standards should
300	improve future studies of sulfur speciation in reduced silicate glasses and minerals in both
301	industrial and geologic contexts, with applications for the early Earth, Moon, Mercury, and
302	enstatite chondrites.
303	Acknowledgements
304	We acknowledge the experimental and analytical facilities at Brown University and
305	GeoSoilEnviroCars (Sector 13), Advanced Photon Source, Argonne National LaboratoryThis
306	research used resources of the Advanced Photon Source, a U.S. Department of Energy (DOE)
307	Office of Science User Facility operated for the DOE Office of Science by Argonne National
308	Laboratory under Contract No. DE-AC02-06CH11357. We acknowledge the funding of NASA
309	grant No. NNX15AH63G and NASA Earth and Space Sciences fellowship No.
310	80NSSC18K1245.
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388	
389	List of figure captions
390	Figure 1. S K-edge energy depends on the coordination chemistry and oxidation state of S for
391	the compound. The S K-edge moves to higher energies as the bonding environment shifts from
392	metallic/covalent to ionic and from reduced (S^{2-}) to oxidized (S^{6+}) species. The dashed line
393	indicates E0 of the compound.
394	
395	Figure 2. S XANES spectra of MgS (niningerite) and CaS (oldhamite) showing a) raw data and
396	b) normalized data. Raw XANES spectra may differ in background (pre-edge level), absorption
397	(near-edge peaks) that is nonlinearly proportional to concentration, and energy-dependent decay
398	(post-edge). Normalized XANES spectra of different samples with different concentrations and
399	sample form can be directly compared and used in linear combination fitting of mixed spectra.
400	
401	Figure 3. Normalization optimization of CaS (oldhamite) S K-edge XANES spectra. a)
402	Normalized and b) flattened spectra showing the full range (n=29766) in red as well as the best
403	according to the optimization program in black. Variance of the full range for c) normalized and
404	d) flattened data generally follow the shape of the data but have error introduced in either the c)
405	post-edge or d) pre-edge regions. Normalized data is better to compare the XANES region, while
406	flattened data is better to compare the EXAFS region.

408	Figure 4. Normalized Sulfur K-edge XANES spectra for reference compounds and minerals.		
409	Ni _{1-x} S, FeS (troilite), MnS (alabandite), CaS (oldhamite), MgS (niningerite), and CaSO ₄		
410	(anhydrite) spectra are comparable to previous studies (e.g. Fleet, 2005), however MgS has three		
411	near-edge peaks as opposed to two seen in Fleet (2005). FeCr ₂ S ₄ , Na ₂ S, and Al ₂ S ₃ were		
412	measured for the first time. The dashed line indicates E0 of the compound. The spectra were		
413	collected in total fluorescence yield.		
414			
415	Figure 5. Normalized Sulfur K-edge XANES spectra for nickel sulfides. $Ni_{1-x}S$ is comparable to		
416	Fleet (2005). The S K-edge moves to higher energies in progression of $Ni_{1-x}S$, Ni_7S_6 , Ni_3S_2 , to		
417	NiS ₂ . The dashed line indicates E0 of the compound. The spectra were collected in total		
418	fluorescence yield.		
419			
420	Figure 6. Normalized Sulfur K-edge XANES spectra for sulfates. Unlike sulfides, sulfates do		
421	not have significant variation in the S K-edge energy with bonding cation (CaSO ₄ , MgSO ₄ , and		
422	FeSO ₄) or oxidation state (FeSO ₄ and Fe ₂ (SO ₄) ₃ . There does appear to be a small peak at \sim 2478		
423	for FeSO ₄ that has been attributed to some degree of photo-reduction in previous analyses (Nash		
424	et al. 2019). The dashed line indicates E0 of the compound. The spectra were collected in total		
425	fluorescence yield.		
426			
427	Figure 7. Position of S K-Edge for reference compounds show the general relationship of		
428	increasing S K-Edge energy with increasingly ionic (as opposed to covalent/metallic) bond		
429	behavior calculated as percent ionic character (Pauling, 1960). Monosulfides with 2+ cation		

430	charge (filled circles: Ni _{1-x} S, FeS, MnS, MgS, CaS) show a linear trend. Monosulfides with
431	different cation charges (open circles: Ni ₇ S ₆ , Ni ₃ S ₂ , NiS ₂ , Al ₂ S ₃ , Na ₂ S) are also shown.
432	Daubréelite (cross: FeCr ₂ S ₄) percent ionic character was calculated using a Cr-S bond. The
433	marker size corresponds to the K-edge uncertainty of 1 eV for all sulfides.

- 434
- 435

Tables

436 Table 1. Position of S K-edge peaks in some sulfur species

⁴³⁷

Mineral or compound	K-edge energy (eV)	Mineral or compound	K-edge energy (eV)
FeCr ₂ S ₄ (daubréelite)	2469.0(1)	Al_2S_3	2471.9(1)
FeS (troilite)	2469.4(1)	Ni _{1-x} S	2469.0(1)
Na ₂ S	2469.2(1)	Ni ₇ S ₆	2469.4(1)
MnS (alabandite)	2470.6(1)	NiS ₂ (vaesite)	2470.3(1)
	2471.8(1)	Ni ₃ S ₂ (heazlewoodite)	2470.2(1)
	2473.2(1)	CaSO ₄ (anhydrite)	2480.5(1)
CaS (oldhamite)	2471.5(1)	$MgSO_4$	2480.7(3)
	2474.1(1)	$Fe_2(SO4)_3$	2480.9(2)
MgS (niningerite)	2472.2(1)	FeSO ₄	2481.3(1)
	2473.3(1)		
	2475.0(1)		

438

439

Figures













Energy (eV)

444







