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## Carbonation and the Urey Reaction

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### Abstract

There are three major reservoirs for carbon in the Earth at the present time, the core, the mantle, and the continental crust. The carbon in the continental crust is mainly in carbonates (limestones, marbles, etc). In this paper we consider the origin of the carbonates. In 1952 Harold Urey proposed that calcium silicates produced by erosion reacted with atmospheric CO<sub>2</sub> to produce carbonates, this is now known as the Urey reaction. In this paper we first address how the Urey reaction could have scavenged a significant mass of crustal carbon from the early atmosphere. At the present time the Urey reaction controls the CO<sub>2</sub> concentration in the atmosphere. The CO<sub>2</sub> enters the atmosphere by volcanism and is lost to the continental crust through the Urey reaction. We address this process in some detail. We then consider the decay of the Paleocene-Eocene thermal maximum (PETM). We quantify how the Urey reaction removes an injection of CO<sub>2</sub> into the atmosphere. A typical decay time is 100,000 years but depends on the variable rate of the Urey reaction.

28

## Introduction

29 The continental crust is a major reservoir for carbon in the Earth. A major question in geology  
30 is the origin of the carbon, principally in calcium carbonates. The first successful attempt to explain the  
31 origin of calcium carbonates (limestones, marbles) in the continental crust was given by Urey (1952).

32 The basic equation he gave was of the form

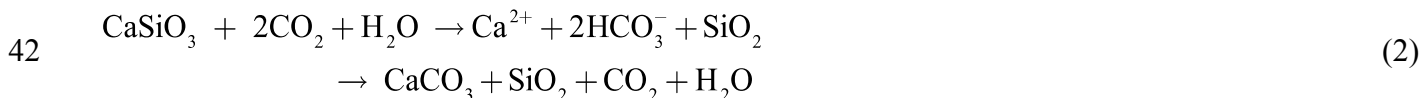
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36 He proposed that atmospheric CO<sub>2</sub> combines with a calcium silicate to generate a calcium carbonate  
37 plus silica. A direct quote from his paper states: “As carbon dioxide was formed it reacted with silicates  
38 to form limestone. Of course, the silicates may have been a variety of minerals, but the presence of CO<sub>2</sub>  
39 was always kept at a low level by this reaction or similar reactions just as it is now.”

40 In the current literature an expanded version of the Urey reaction is given (Blattler and Higgins  
41 2017). In order to include the role of acid rain the Urey reaction takes the form



43

44 The carbonation takes place when carbon dioxide (carbonic acid) in acid rain dissolves calcium silicate  
45 (wollastonite) sediments to give calcium, bicarbonate, and silica. The resulting calcium and bicarbonate  
46 ions flow in rivers to the oceans where either organic or inorganic precipitation produces the calcium  
47 carbonate.

48 The three large reservoirs for carbon in the Earth at the present time are the core, mantle, and  
49 continental crust. We assume that the core is an isolated reservoir and neglect its role. About one per

50 cent of the carbon in the Earth is in the continental crust. Wadepohl (1995) has given a comprehensive  
51 study of the composition of the continental crust with an emphasis on carbon. He gives an estimate for  
52 the total mass of carbon (c) in the continental crust (cc) at the present time (y) of  ${}^cM_{ccp} = 4.2 \times 10^7$  Gt.  
53 Hayes and Waldbauer (2006) have reviewed the literature on carbon in the continental crust and  
54 suggest that it may be as high as  ${}^cM_{ccp} = 10^8$  Gt. DePaolo (2015) gives a range of 6 to  $7 \times 10^7$  Gt. In  
55 this paper we take a representative value to be  ${}^cM_{ccp} = 5 \times 10^7$  Gt. The mass of carbon in the ocean is  
56 about a factor of  $10^3$  less than the mass of carbon in the continental crust (Houghton 2007).

57 Urey (1952, 1956) clearly recognized that the reaction he proposed would efficiently remove  
58  $\text{CO}_2$  from the Earth's atmosphere, but at that time little was known about the early atmosphere.  
59 Although the mass of carbon in the atmosphere today is small (850 Gt), the mass may have been much  
60 higher in the past. One of the major differences between Venus and the Earth is atmospheric  
61 composition. The atmospheric pressure on Venus is about a factor of 100 greater than the atmospheric  
62 pressure on Earth and is 96% carbon dioxide. The mass of carbon in the Venus atmosphere (a) at the  
63 present time (p) is  ${}^cM_{ap} = 1.28 \times 10^8$  Gt. Scaling the atmospheric carbon masses to the overall masses of  
64 Venus and the Earth gives an estimate of the mass of carbon (c) in the early atmosphere ( $t=0$ ) of the  
65 Earth. The estimated value is  ${}^cM_{a0} = 1.57 \times 10^8$  Gt (Kasting and Ackerman 1986).

### 66 **Carbon from the atmosphere to the continental crust**

67 One hypothesis for the origin of the carbon in the continental crust is that it was extracted  
68 directly from the atmosphere relatively early in Earth's history. The estimated mass of carbon (c) in the  
69 early atmosphere (a) given above,  ${}^cM_{a0} = 1.57 \times 10^8$  Gt, is substantially larger than the total estimated  
70 carbon in the continental crust given above,  ${}^cM_{ccp} = 5 \times 10^7$  Gt. The hypothesis of direct extraction  
71 from the atmosphere has been discussed in some detail by Kramers (2002) and by Lowe and Tice  
72 (2004).

73 The basic hypothesis is that the mass flux of carbon from the atmosphere to the continental  
 74 crust,  ${}^cJ_{a-cc}$ , is controlled by the availability of calcium silicates. In order for the Urey reaction to extract  
 75  $\text{CO}_2$  from the atmosphere the early Earth must have had continental crust in order to generate surface  
 76 deposits of calcium silicate. In addition, the Earth must have had oceans in order for the acid rain to  
 77 catalyze the Urey reaction between atmospheric  $\text{CO}_2$  and the service deposits of calcium silicates. Little  
 78 data are available for timing the initiation of the extraction of  $\text{CO}_2$  from the atmosphere. We will  
 79 assume that the process begins at a time  $t_0$  after the early bombardment and the solidification of the  
 80 magma ocean at about 4.4 Ga. We further assume that the Urey reaction extracted carbon from the  
 81 atmosphere at a constant rate  ${}^cJ_{a-cc}$  until the concentration of  $\text{CO}_2$  in the atmosphere was reduced to a  
 82 very low level. During the time,  $t_0 < t < t_0 + \tau_{a-cc}$ , the Urey reaction extracts atmospheric carbon to the  
 83 continental crust. We will specify the mass of carbon extracted from the atmosphere and obtain

$$84 \quad {}^cJ_{a-cc} = \frac{{}^cM_{a0}}{\tau_{a-cc}} \quad (3)$$

85  
 86 We assume that the mass of carbon in the atmosphere  ${}^cM_a$  decreases linearly in time from  ${}^cM_{a0}$  to zero  
 87 during the time period  $\tau_{a-cc}$  and the mass of carbon in the continental crust increases linearly in time.

88 Assuming all the carbon in the continental crust  ${}^cM_{ccp}$  was extracted from the atmosphere the  
 89 dependence on time is given by

$$90 \quad {}^cM_{cc} = 0 \quad 0 \leq t \leq t_0$$

$$91$$

$$92 \quad {}^cM_{cc} = {}^cM_{ccp} [(t - t_0)/(\tau_{a-cc})] \quad t_0 \leq t \leq t_0 + \tau_{a-cc} \quad (4)$$

$$93$$

$$94 \quad {}^cM_{cc} = {}^cM_{ccp} \quad t_0 + \tau_{a-cc} \leq t \leq t_p$$

$$95$$

96 Taking  ${}^cM_{ccp} = 5 \times 10^7$  Gt,  $t_0 = 1$  Gyr, and  $\tau_{a-cc} = 1$  Gyr, the dependence of  ${}^cM_{cc}$  on  $t$  is given in Fig. 1.

97 The required flux of carbon from the atmosphere to the continental crust is  ${}^cJ_{a-cc} = 50 \text{ Mtyr}^{-1}$ . It must be

98 emphasized that the value of  $\tau_{a-cc}$  is uncertain and the flux  ${}^cJ_{a-cc}$  is expected to have considerable  
99 variability in time. However it is quite clear that the extraction of carbon from the atmosphere to the  
100 continental crust would have been carried out early in Earth's history.

101

### 102 **Removal of the volcanic addition of carbon to the atmosphere**

103 When excess carbon in the atmosphere has been depleted by the Urey reaction an approximate  
104 steady-state balance is established between the volcanic input of carbon into the atmosphere and the  
105 extraction by the Urey reaction. We approximate this balance by the relation

$$106 \quad {}^cJ_{a-cc} = \frac{{}^cM_a}{\tau_u} \quad (5)$$

107 where  ${}^cJ_{a-cc}$  is the rate of volcanic input of carbon into the atmosphere. We assume that this extraction  
108 rate is constant and is proportional to the mass of carbon in the atmosphere  ${}^cM_a$ . The characteristic time  
109  $\tau_u$  takes account of the rate at which acid rain can interact with calcium silicate sediments, and although  
110 we assume that  $\tau_u$  is constant it clearly can be a function of time.

111 A comprehensive model for the variability of atmospheric CO<sub>2</sub> over Phanerozoic times has been  
112 given by Berner and Kothavala (2001). This model, GEOCARB 3, is complex and involves both  
113 organic and inorganic processes. Transport of carbon between the atmosphere, oceans, and continental  
114 crust is quantified on the million year time scale. The balance is dominated by the exchange of carbon  
115 between carbonates in the continental crust and carbon in the surficial reservoirs (oceans and  
116 atmosphere) and organic carbon (Berner and Calderia, 1997). When erosion is high, the Urey reaction  
117 extracts CO<sub>2</sub> from the atmosphere adding carbonates to the continental crust. High erosion rates are  
118 associated with low sea level and large continental areas. When erosion is low, the Urey reaction  
119 operates in the opposite direction (from right to left in Eq.(1)) with carbonates decomposing to give

120 CO<sub>2</sub>. An example of this metamorphic process is the subduction of carbonate sediments and the  
121 generation and return to the atmosphere of CO<sub>2</sub> in subduction zone volcanics (Frezzotti et al. 2011).

122 The present mass of carbon in the atmosphere is 860 Gt (400 ppmv CO<sub>2</sub>), but this is not a quasi-  
123 equilibrium value because of the anthropogenic addition at high fluxes (3.5 Gtyr<sup>-1</sup>). We will take the  
124 1900 value of 650 Gt (300 ppmv CO<sub>2</sub>) as the present equilibrium value. This is a typical value for the  
125 current glacial epoch (0 to 50 Ma). Values given by the GEOCARB 3 Model are generally consistent  
126 with observations (Royer 2014). Between 50 and 250 Ma the average values were about 3000 Gt.  
127 During the major glacial epoch between 250 and 350 Ma low observed values near 650 Gt are found.  
128 Between 350 and 550 Ma values were considerably higher, typically near 10,000 Gt. This variability  
129 reflects variations in both of the variables in Eq. (3), the volcanic flux  ${}^cJ_{a-cc}$  into the atmosphere and the  
130 characteristic time  $\tau_u$ .

### 131 **Carbon from the mantle to the continental crust**

132 The second hypothesis for the origin of the carbon in the continental crust is that it comes from  
133 the mantle. If the volcanic flux of carbon out of the mantle at ocean ridges and hot spots exceeds the  
134 carbon lost to the mantle at subduction zones, the difference will be added to the continental crust.  
135 Some of the volcanic carbon input will enter the atmosphere and will be transferred to the continental  
136 crust through the Urey reaction. However, some will enter the oceans and will be converted directly to  
137 carbonates without entering the atmosphere.

138 Rates of carbon loss from the mantle by volcanism and lost by subduction will certainly vary  
139 over geologic time, but the variations are uncertain. Again, we assume that the plate tectonic processes  
140 required for carbon transfer began at a time  $t_0$  after the solidification of the magma ocean at about 4.5  
141 Ga. We further assume that the transfer of carbon out of the mantle has been at a constant rate  ${}^cJ_{m-cc}$   
142 until the present time  $t_p$ . Assuming all the carbon in the continental crust has been extracted from the  
143 mantle, the dependence on time is given by

144

$$\begin{aligned} 145 \quad & {}^cM_{cc} = 0 & 0 \leq t \leq t_0 \\ 146 \quad & {}^cM_{cc} = {}^cM_{ccp} [(t - t_0)/(t_p - t_0)] & t_0 \leq t \leq t_p + \tau_{a-cc} \end{aligned} \quad (6)$$

147

148 The mass of carbon in the continental crust increases linearly in time over the period  $t_0$  to  $t_p$ . The  
149 required flux of carbon from the mantle to the continental crust is given by

150

$$151 \quad {}^cJ_{m-cc} = \frac{{}^cM_{ccp}}{t_p - t_0} \quad (7)$$

152

153 Taking  ${}^cM_{ccp} = 5 \times 10^7$  Gt,  $t_0 = 1$  Gyr, and  $t_p = 4.4$  Gyr the dependence of  ${}^cM_{cc}$  on  $t$  is given in Fig 1. The  
154 required flux of carbon from the mantle to the continental crust is  ${}^cJ_{m-cc} = 14.7$  Mtyr<sup>-1</sup>.

155 We next consider the estimate for the present loss of carbon from the mantle to the atmosphere.

156 Dasgupta and Hirschmann (2010) have summarized the available data on the loss of carbon from the

157 mantle to the surface reservoirs and give values in the range  ${}^cJ_{m-s} = 36 \pm 24$  Mtyr<sup>-1</sup>. Just as carbon is

158 lost from the mantle by volcanism, carbon is returned to the mantle by subduction. A detailed study of

159 carbon fluxes at subduction zones has been given by Kelemen and Manning (2015). These authors

160 suggest that the downward flux of carbon at global subduction zones is  $53 \pm 13$  Mtyr<sup>-1</sup>. However a

161 substantial fraction of this carbon never makes it to the mantle due to subduction zone volcanism. They

162 suggested that  $24 \pm 24$  Mtyr<sup>-1</sup> reach the mantle. Clearly it is quite possible that all the carbon in the

163 continental crust could have come from the mantle. This conclusion was also given by Hayes and

164 Waldbauer (2006).

165 In Fig. 1 we give examples of the two limiting cases, the carbon in continental crust comes

166 entirely from the atmosphere and the carbon comes entirely from the mantle. In the first case the

167 addition is early in time and in the second case it is more uniform in time. Observations of the mass of  
168 carbonates in the continental crust as a function of age could distinguish between the two cases, but the  
169 data are sparse. Observations of the mass of carbon in the atmosphere as a function of time could also  
170 be a constraint. An example given by Rye et al. (1995) utilizing studies of paleosols concluded that the  
171 mass of carbon in the atmosphere at 2.2 to 2.75 Ga was less than  $10^5$  Gt. The conclusion is that the  
172 extraction of a significant mass of carbon from the atmosphere to the continental crust was completed  
173 by  $t = 2$  Gyr. However how large this mass was is uncertain.

### 174 **PETM**

175 The decay of the Paleocene-Eocene thermal maximum (PETM) can be used to quantitatively  
176 constrain the role of the Urey reaction. The PETM was a period of elevated global temperatures (4 to 5  
177 degrees C) and high atmospheric  $\text{CO}_2$  beginning at 56.3 Ma, the onset lasted less than 10 Kyr and the  
178 subsequent decay lasted about 100 Kyr (McInerney and Wing 2011). Storey et al. (2007) have made a  
179 strong case for associating the PETM with flood volcanism resulting from the opening of the north  
180 Atlantic.

181 Isotope studies have quantitatively documented the PETM. These studies have been reviewed  
182 by Gutjahr et al. (2017). These authors also provided estimates for the carbon content of the  
183 atmosphere during the PETM. They suggest that the background carbon mass in the atmosphere before  
184 and after the PETM was  ${}^cM_{ab} = 1400$  Gt and the peak mass of carbon was  ${}^cM_{a0} = 3050$  Gt.

185 We now carry out an analysis of the decay of the PETM due to the loss of  $\text{CO}_2$  from the  
186 atmosphere by the Urey reaction. We extend the balance given in Eq. (5) to include the transient  
187 removal of carbon from the atmosphere and write

188

$$189 \quad \frac{d {}^c M_a}{dt} = {}^c J_{(a-cc)b} - \frac{{}^c M_a}{\tau_u} \quad (8)$$



190

191 From Eq. (5) the background mass of carbon in the atmosphere is given by

192

$$193 \quad {}^cM_{ab} = \tau_u {}^cJ_{(a-cc)b} \quad (9)$$

194

195 We prescribe an initial mass of carbon in the atmosphere at  $t = 0$ ,  ${}^cM_{a0}$  and solve Eq. (7) taking  $\tau_u$  to be

196 constant with the result

197

$$198 \quad {}^cM_a = ({}^cM_{a0} - {}^cM_{ab})e^{-t/\tau_u} + {}^cM_{ab} \quad (10)$$

199

200 The excess mass of carbon in the atmosphere  ${}^cM_{a0} - {}^cM_{ab}$  decays exponentially as the Urey reaction

201 extracts carbon from the atmosphere.

202 We next obtain the dependence of atmosphere carbon mass on time during PETM based on the

203 model dependence given in Eq. (10). Taking the values  ${}^cM_{ab} = 1400$  Gt and  ${}^cM_{a0} = 3050$  Gt with  $\tau_u$

204  $= 100$  kyr the model results are given in Figure 2.

205 We now return to Eq. (9). This result relates the background atmospheric carbon mass  ${}^cM_{ab}$  to

206 the background rate of volcanic input of CO<sub>2</sub> carbon into the atmosphere  ${}^cJ_{(a-cc)b}$  and the Urey reaction

207 rate  $\tau_u$ . During the PETM we have taken the background carbon mass  ${}^cM_{ab} = 1400$  Gt. Taking  $\tau_u = 100$

208 kyr we find from Eq. (9) that  ${}^cJ_{(a-cc)b} = 14$  Mtyr<sup>-1</sup>. This is an independent determination of the volcanic

209 flux of carbon into the atmosphere at that time. As discussed above we take the present equilibrium

210 mass of carbon in the atmosphere to be  ${}^cM_{ab} = 6500$  Gt. Assuming that  ${}^cJ_{(a-cc)b} = 14$  Mtyr<sup>-1</sup> we require

211 from Eq. (9) that  $\tau_u = 50$  kyr. This is our estimated relaxation time for a carbon excursion today.

212

## Discussion

213  
214 Urey (1952) proposed the Urey reaction, Eq. (1), to explain the origin of carbonates in the  
215 continental crust. He argued that the reaction would essentially remove all CO<sub>2</sub> from the atmosphere. It  
216 is now accepted that in analogy to Venus, there may have been a large mass of carbon in the Earth's  
217 early atmosphere, as much as 10<sup>8</sup> Gt. However, only a fraction of this may have survived the moon  
218 forming impact. We give a very simplified model for the extraction of carbon from the atmosphere to  
219 the continental crust, taking the extraction rate  ${}^cJ_{a-cc}$  to be constant. There are basically no constraints  
220 on the variation of this rate with time. If a significant fraction of the carbon in the continental crust was  
221 extracted from the atmosphere, it is likely that it occurred early in Earth's history as illustrated in Fig. 1.

222 The Urey reaction also controls the equilibrium mass of carbon in the atmosphere after the  
223 removal of any large initial concentration. The input of carbon to the atmosphere is from volcanism and  
224 we show that the equilibrium mass of carbon in the atmosphere  ${}^cM_{ab}$  is proportional to the rate of  
225 volcanic injection  ${}^cJ_{a-cc}$  divided by a characteristic Urey time  $\tau_u$ . We quantify the value of  $\tau_u$  by  
226 studying the observed relaxation of the Paleocene-Eocene thermal maximum, which occurred at 56 Ma,  
227 and the relaxation time is about  $\tau_u = 10^5$  yrs.

228 We also give a simplified model for the extraction of carbon from the atmosphere to the  
229 continental crust. If the volcanic loss of carbon from the mantle by volcanism exceeds the return of  
230 carbon by subduction, the difference is added to the continental crust. If the volcanic carbon enters the  
231 oceans organic precipitation creates carbonates. If the volcanic carbon enters the atmosphere it enters  
232 the continental crust by the Urey reaction. Current estimates of carbon fluxes from and to the mantle  
233 are sufficient to have produced all the carbon in the continental crust. At the present time it is not  
234 possible to quantify the relative importance of carbon addition to the continental crust from the early  
235 atmosphere and the mantle.

236

237

### Implications

238 We have addressed two major questions concerning carbon in the atmosphere in this paper. The  
239 first is the origin of the carbon in the continental crust. We conclude that it is possible the carbon could  
240 have been extracted either from the early atmosphere or from the mantle over a longer period of time.  
241 Studies of the concentration of carbon in the atmosphere and continental crust over geologic time are  
242 required and should receive a high priority.

243 The second question we have addressed is the relaxation of injections of carbon into the  
244 atmosphere back to equilibrium values. We quantify this by studying the Paleocene-Eocene thermal  
245 maximum (PETM). This has obvious implications for the recovery from the process of anthropogenic  
246 injection of carbon into the atmosphere. We find the relaxation time to be about 50,000 yrs.

247

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### 296 **Figure Captions**

297 Figure 1. Dependence of the mass of carbon in the continental crust  ${}^cM_{cc}$  on time. Two limiting models  
298 are given for adding the present mass  ${}^cM_{ccp} = 5 \times 10^7$  Gt. 1) Addition from the atmosphere beginning at  
299  $t_0 = 1$  Gyr. All atmospheric carbon is transferred in  $\tau_{a-cc} = 1$  Gyr at a constant flux  ${}^cJ_{a-cc} = 50$  Mtyr<sup>-1</sup>. 2).  
300 Addition from the mantle beginning at  $t_0 = 1$  Gyr. Carbon is added at a constant flux  ${}^cJ_{m-cc} = 14.7$   
301 Mtyr<sup>-1</sup> to the present.

302

303

304 Figure 2: Dependence of the atmosphere carbon mass values  ${}^cM_a$  for the PETM anomaly on time  $t_{PETM}$   
305 relative to the onset of the anomaly. The values are from our relaxation model given in Eq. (11) with  
306  ${}^cM_{ab} = 1400$  Gt and  ${}^cM_{a0} = 3050$  Gt and  $\tau_u = 100$  kyr.

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