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2 **HIGHLIGHTS AND BREAKTHROUGHS**

3 **Titan mineralogy: A window on organic mineral evolution**

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8 Every planet and moon experiences mineral evolution—a change in the diversity and  
9 distribution of minerals through billions of years of physical, chemical, and (in the case of Earth)  
10 biological processes (Zhabin, 1981; Hazen et al., 2008; Hazen and Ferry, 2010). A question often  
11 pondered in this context asks if minerals exist somewhere in the cosmos that are not found on  
12 Earth. As Maynard-Casely et al. (2017) have elegantly shown, one needs look no farther than  
13 Saturn’s moon Titan for the answer—a resounding “Yes!” They paint a vivid picture of a frozen  
14 world perpetually at 92 K, where hydrocarbon rains splash on a landscape of organic-rich rocks  
15 containing a rich taxonomy of molecular crystals—ices of water and ammonia, cage-like  
16 clathrates hosting small guest molecules, varied organic hydrates, and novel “co-crystals” that  
17 incorporate two or more molecular species in stoichiometric ratios.

18 Multiple paragenetic processes likely enrich Titan’s postulated mineralogy. Changes of state  
19 from the orange organic-rich atmosphere to liquid and solid phases are important drivers of near-  
20 surface distillation and diversification. For example, condensation of tholins, complex molecules

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24 rich in C, N, and H produced by photochemistry in Titan’s atmosphere, are likely to result in a  
25 range of colorful crystals (Cable et al., 2011). Hydration reactions may be common, as water  
26 molecules from aqueous “cryo-volcanoes” react with organic species. Phase transitions featuring  
27 molecular orientational order-disorder reactions could lead to dramatic changes in density and  
28 hardness, as well as in electrical and elastic properties. Some organic molecular species would  
29 dissolve in Titan’s hydrocarbon seas, only to subsequently form exotic evaporite deposits  
30 (imagine fields of shiny faceted acetylene crystals sprouting along the margins of hydrocarbon  
31 lakes during a Titan “dry spell”).

32 Yet, in spite of the exotic carbon-rich chemistry and alien environmental conditions, Titan has  
33 revealed familiar geological features not unlike those of Earth: rivers and lakes, dunes and deltas,  
34 mountains and valleys, and those curious cryovolcanoes that erupt aqueous “magmas.” Maynard-  
35 Casely *et al.*’s contribution emphasizes the diverse physical and chemical properties of purported  
36 Titan minerals—their hardnesses, which may contribute to resistant landforms; their melting  
37 points, which result in varied fluid-rock interactions, both on the surface and in the warmer Titan  
38 interior; and their relative densities and elastic properties that play essential roles in creating and  
39 evolving Titan’s intriguing topography.

40 As Maynard-Casely *et al.* emphasize, the study of organic crystal chemistry and phase  
41 equilibria at modest pressures and cryogenic conditions—the experimental petrology of Titan—is  
42 in its infancy. Consequently, we can anticipate the discovery of many new plausible Titan  
43 minerals—as yet unknown carbon-bearing crystalline phases that must surely occur on countless  
44 cold, carbon-rich worlds throughout the cosmos.

45

46 *Earth's organic mineralogy in space and time*

47 Does Titan's intriguing organic mineralogy hold lessons for the story of Earth? At first blush  
48 the two worlds have little in common. Not only is Earth warmer and wetter, but the bulk  
49 compositions are strikingly different, as well. The crustal mineralogy of our planetary home is  
50 dominated by eight elements—oxygen, silicon, aluminum, iron, calcium, sodium, potassium, and  
51 magnesium comprise 98 percent of the mass? Carbon, by contrast, is a minor constituent,  
52 accounting for less than 0.1 weight percent of Earth's crust, and almost all of that carbon is  
53 locked into vast precipitated platforms of carbonate minerals (Reeder, 1983; Hazen et al.,  
54 2013a). Nevertheless, Maynard-Casely *et al.*'s study inspires a closer look at Earth's organic  
55 mineral evolution.

56 More than 50 “organic minerals” have been identified on Earth (Perry et al., 2007; Downs,  
57 2016; Hazen et al., 2013a), all of which are indirectly derived from biology, including crystals  
58 formed by the subaerial alteration of excrement (especially in cave guano deposits), from  
59 decaying vegetation and animal matter, or derived from coal and oil shale (e.g., Nasdala and  
60 Pekov, 1993; Perry et al. 2007; Hummer et al., 2017), most notably as crystals that condense  
61 from a hot vapor phase near burning coal mines (Oftedal, 1922; Rost, 1942; Belakovskiy, 1998;  
62 Chesnokov et al., 2008; Witzke et al., 2015; Pekov et al., 2016). Occurrences of these few dozen  
63 organic minerals are rare and localized; in several cases species are known from just a few  
64 microscopic crystals (mindat.org). Benner et al. (2010; personal communications), furthermore,  
65 emphasize that the most common of these compounds are highly oxidized oxalates and  
66 carboxylates—phases that consequently possess meager nutritional value on our oxygen-  
67 saturated world. Most crystals of more reduced organic species represent food and are thus likely  
68 to be consumed almost as soon as they form. The few notable edible exceptions, such as the

69 aluminum benzenhexacarboxylate mineral, mellite, may owe their survival to their low  
70 solubilities.

71 However, early in Earth's history at a time when carbonaceous chondrites and icy comets  
72 laced with organics rained with regularity onto the sterile surface, in the prebiotic age before the  
73 opportunistic predation of any stray food-like molecule by cellular life, there may have existed a  
74 profusion of now extinct organic mineral species (Benner et al., 2010). An inventory of such  
75 primordial organic minerals—phases overlooked in prior surveys of carbon mineral evolution  
76 (Hazen et al., 2013b) and a preliminary list of potential Hadean mineral species (Hazen, 2013)—  
77 is fertile ground for speculation.

78 Several intriguing crystalline organic compounds not found on Earth today have been invoked  
79 in origins-of-life scenarios, both on Earth (Benner et al., 2010) and early Mars (Benner et al.,  
80 2000). Prior to the origins of life, these bio-building blocks may have accumulated in  
81 concentrated crystalline form until they reached useful amounts. Benner et al. (2010) focus on  
82 reactive carbohydrates and related nitrogen-bearing compounds, while Ritson and Sutherland  
83 (2013) point to the possible role of crystalline ferrocyanide in the assembly of RNA. The  
84 Sutherland group has also studied the facile crystallization of 2-aminooxazole, a plausible  
85 prebiotic mineral that could have been an intermediate compound on the pathway to RNA  
86 synthesis (Powner et al., 2009; Szostak, 2017).

87 In addition, early Earth may have boasted a rich variety of crystalline or otherwise self-  
88 organized forms of lipids, aromatic compounds, and amino acids, some of which may have  
89 precipitated as chirally pure “left-“ and “right-handed” crystalline forms (Springsteen and Joyce,  
90 2004; Hein et al., 2011)—a key and as yet poorly understood step from geochemistry to  
91 biochemistry (Lahav, 1999; Hazen, 2005). Organic minerals may have abounded in zones where

92 crystals grew from evaporating organic-rich fluids or condensed from a hot gas phase. This  
93 geological selection and concentration of life's essential ingredients through crystallization of  
94 early Earth's diverse organic milieu is a problem ripe for continued discovery.

95

### 96 *Organic mineralogy of the Anthropocene Epoch*

97 Earth's organic mineral evolution is far from being over. The modern age of human  
98 exploitation of carbon-rich coal, oil, tar sands, and other hydrocarbon-rich "fossil" fuels—a time  
99 dubbed by some advocates the "Anthropocene Epoch"—has seen an explosion in the diversity  
100 and distribution of both inorganic and organic solids at or near Earth's increasingly engineered  
101 surface (Zalasiewicz et al., 2013; Hazen et al., 2017). Our geological era is marked by the  
102 synthesis of millions of diverse organic compounds, some of which will persist in crystalline or  
103 polymeric forms for eons. However we choose to characterize those diverse products of human  
104 ingenuity—minerals or mineraloids or "mineral-like" phases—solid organic compounds have  
105 once again, and for the first time since the prebiotic Hadean Eon, risen to a place of prominence  
106 in the chemistry of Earth's near-surface environment.

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