## The effect of thermal decarbonation on stable isotope compositions of carbonates Z.D. SHARP,<sup>1,\*</sup> J.J. PAPIKE,<sup>1,2</sup> AND T. DURAKIEWICZ<sup>3,4</sup>

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## ABSTRACT

Oxygen and C isotope compositions of  $CO_2$  gas released by thermal decomposition of siderite, calcite, and dolomite were measured using a new "real-time" continuous-flow technique to determine whether fractionation associated with simple thermal decarbonation could explain the large isotopic variations and mineralogy such as those found in the ALH84001 meteorite.

Oxygen and C isotope fractionation between calcite or dolomite and evolved CO<sub>2</sub> gas during thermal decarbonation in a 3 bar He pressure environment is very small. The  $\delta^{13}$ C and  $\delta^{18}$ O values of evolved CO<sub>2</sub> gas are nearly identical to those of the carbonate, very different from the calculated equilibrium  $\Delta^{18}$ O<sub>calcite-CO<sub>2</sub></sub> value of -4 to -5‰ at 800-900 °C or from previous experimental results of decarbonation in vacuum. The kinetic  $\Delta^{18}$ O<sub>siderite-CO<sub>2</sub></sub> values are ~-2‰, whereas  $\Delta^{13}$ C<sub>siderite-CO<sub>2</sub></sub> values increase logarithmically with time, from ~1‰ for the earliest stages of decarbonation to >5‰ in the final stages. Incomplete siderite decomposition produces both magnetite ( $\delta^{18}$ O = 3.5‰ SMOW) and minor graphite. CO and O<sub>2</sub> were detected during the decarbonation process. The data can be explained by simultaneous oxidation and reduction by the reaction:

$$6FeCO_3 \rightarrow 2Fe_3O_4 + 2x CO + 4y CO_2 + (6 - 2x - 4y) C + (5 - x - 4y) O_2$$

where x and y are between 0 and 1. Siderite decomposition in the presence of  $H_2$  gas produces wüstite and Fe metal in place of oxidized Fe minerals.

The experiments in this study are not a perfect analog for possible decarbonation conditions that might have occurred to the carbonates in ALH84001. Nevertheless, the large  $\delta^{13}$ C and  $\delta^{18}$ O variations observed in ALH84001 (>10‰ for O) are significantly larger than those expected by thermal decarbonation, suggesting instead a low-temperature mechanism for their formation.