Wassonite: A new titanium monosulfide mineral in the Yamato 691 enstatite chondrite

KEIKO NAKAMURA-MESSENGER,^{1,2,*} SIMON J. CLEMETT,^{2,3} ALAN E. RUBIN,⁴ BYEON-GAK CHOI,⁵ SHOULIANG ZHANG,^{2,6,7} ZIA RAHMAN,^{1,2} KATSUNARI OIKAWA,⁸ AND LINDSAY P. KELLER²

¹ESCG/Jacobs Technology, Texas 77058, U.S.A.

²Robert M Walker Laboratory for Space Science, Astromaterials Research and Exploration Science Directorate/NASA Johnson Space Center, Houston, Texas 77058, U.S.A.

³ESCG/ERC Inc., Texas 77058, U.S.A.

⁴Institute of Geophysics and Planetary Physics, University of California, Los Angeles, California 90095-1567, U.S.A. ⁵Earth Science Education, Seoul National University, Seoul 151-748, South Korea

⁶Lunar and Planetary Institute, Houston, Texas 77058, U.S.A.

⁷Texas Material Institute, University of Texas, Austin, Texas 78712, U.S.A.

⁸Department of Metallurgy, Graduate School of Engineering, Tohoku University, Sendai, 980-8579, Japan

ABSTRACT

Wassonite, ideally stoichiometric TiS, is a titanium monosulfide not previously observed in nature, that was discovered within the Yamato 691 EH3 enstatite chondrite. Twelve Ti-S phase grains were identified in a rare barred olivine (BO) chondrule; three of the grains were extracted by the focused ion beam technique. Because of the submicrometer size of the wassonite grains, it was not possible to determine conventional macroscopic properties. However, the chemical composition and crystal structure were well constrained by extensive quantitative energy-dispersive X-ray analysis and electron diffraction using transmission electron microscopy (TEM). The crystal system for wassonite is rhombohedral ($a = 3.42 \pm 0.07$, $c = 26.50 \pm 0.53$ Å) with space group: $R\overline{3}m$, cell volume: 268.4 ± 0.53 Å³, Z = 9, density (calculated): 4.452 g/cm³, empirical formula: (Ti_{0.93},Fe_{0.06},Cr_{0.01})S. The wassonite grains crystallized from the chondrule melt that was itself formed in the solar nebula, not on the parent asteroid. The other crystalline phases in the BO chondrule include forsterite, enstatite, troilite, metallic Fe-Ni, and osbornite (as well as the new Ti-S-bearing minerals and schollhornite) are highly reduced and indicate formation at low-oxygen fugacities.

Keywords: New mineral, wassonite, TiS, Antarctic meteorite, TEM, electron diffraction