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3 **Nature and timing of Sn mineralization in southern Hunan, South**
4 **China: Constraints from LA-ICP-MS cassiterite U-Pb geochronology**
5 **and trace element composition**

6 Tao Ren^a, Huan Li^{a*}, Thomas J. Algeo^{b, c, d}, Musa Bala Girei^e, Jinghua Wu^a, Biao Liu^a

7 ^a *Key Laboratory of Metallogenic Prediction of Nonferrous Metals and Geological*
8 *Environment Monitoring, Ministry of Education, School of Geosciences and Info-*
9 *Physics, Central South University, Changsha 410083, China*

10 ^b *State Key Laboratory of Geological Processes and Mineral Resources, School of*
11 *Earth Resources, China University of Geosciences, Wuhan 430074, China*

12 ^c *State Key Laboratory of Biogeology and Environmental Geology, School of Earth*
13 *Sciences, China University of Geosciences, Wuhan 430074, China*

14 ^d *Department of Geosciences, University of Cincinnati, Cincinnati, OH 42221-0013*
15 *USA*

16 ^e *Department of Geology, Bayero University Kano, Kano State, Nigeria*

17 *Corresponding author (H. Li): lihuan@csu.edu.cn

18 **ABSTRACT**

19 Accurately determining the timing and mechanism of metallogenesis of ore deposits
20 is essential for developing a robust genetic model for their exploration. In this paper,
21 we analyze the formation conditions of cassiterite in five major deposits of southern
22 Hunan Province, one of the most important tungsten-tin (W-Sn) provinces in South

23 China, using a combination of cathodoluminescence imaging, in situ U-Pb
24 geochronology, and trace-element concentration data. In situ cassiterite U-Pb
25 geochronology constrains the main period of Sn mineralization to between 155.4 and
26 142.0 Ma, demonstrating a temporal and genetic relationship to silicic intrusive
27 magmatism in the same area. Three stages of magmatic activity and metallogenic
28 evolution are recognized: (1) Early Paleozoic and Triassic: the initial enrichment stage
29 of tungsten and tin; (2) Jurassic: the metasomatic mineralization stage; and (3)
30 Cretaceous: the magmatic-hydrothermal superposition stage. The cassiterite in these
31 deposits takes four forms, i.e., quartz vein-type, greisen-skarn-type, greisen-type, and
32 granite-type, representing a progression characterized by the increasing content and
33 decreasing range of variation of high-field-strength elements (HFSEs), and reflecting
34 a general increase in the degree of evolution of the associated granites. Rare earth
35 element (REE) concentrations suggest that precipitation of cassiterite was insensitive
36 to the redox state of the fluid, and that precipitation of cassiterite in the southern
37 Hunan Sn deposits did not require a high- fO_2 environment. These findings provide
38 new insights into tin mineralization processes and exploration strategies.

39 **Keywords:** tin; tungsten; U-Pb dating; geochemistry; metallogenesis; Nanling

40 INTRODUCTION

41 Cassiterite (SnO₂), the most economically important tin (Sn)-bearing mineral, is
42 generally mined from primary magmatic-hydrothermal deposits that are spatially and
43 temporally associated with highly differentiated granites (Heinrich 1990; Cheng et al.
44 2019; Zhu et al. 2021). Accurately determining the timing and duration of
45 precipitation of cassiterite is essential to understanding Sn-ore mineralization
46 processes and thus generating genetic models that can aid in prospecting for rare large

47 tin deposits (Yuan et al. 2018). Given its tetragonal, rutile-type structure, cassiterite
48 typically has a high U-Pb ratio as well as a high closure temperature (Zhang et al.
49 2011, 2017; Neymark et al. 2018). These characteristics give it high resistance to
50 post-ore hydrothermal alteration (Plimer et al. 1991; Hu et al. 2021). Due to their
51 similarities to Sn⁴⁺ with respect to ionic charges, radii, and coordination numbers,
52 trace elements such as Hf, Zr, Sc, Ta, Nb, Ti, Fe, Mn, In, U and W are able to
53 substitute for Sn in cassiterite either directly or through a coupled substitution
54 mechanism. These elements can be used to trace the cassiterite growth environment as
55 well as the source of mineralizing fluids (Schmidt 2018; Cheng et al. 2019; Bennett et
56 al. 2020; Lehmann et al. 2021; Wu et al. 2021).

57 The southern Hunan Province, located within the western Nanling metallogenic
58 belt, is one of the most important W-Sn metallogenic regions in China (Mao et al.
59 2007; Hu et al. 2017; Jiang et al. 2020). This province consists of several major
60 Sn-polymetallic deposits including the Hehuaping, Bailashui, and Xitian deposits
61 (SnO₂ reserves of 140 metric kilotons [Kt], 420 Kt, and 178 Kt respectively; Yao et al.
62 2014; Wang et al. 2014; Li et al. 2019). Also located within this province are the
63 Dengfuxian (quartz vein-type), Dayishan (greisen-type), Shizhuyuan
64 (greisen-skarn-type), Xianghualing (granite-type), and Jiuyishan (greisen-type)
65 deposits that are the focus of this study. These deposits are spatially and temporally
66 related to highly differentiated granites of mid-Mesozoic (165-150 Ma) age that have
67 experienced pervasive hydrothermal alteration (Sun et al. 2018; Li et al. 2018a, b;
68 Yang et al. 2018; Xiong et al. 2020; Liao et al. 2021; Wu et al. 2021, Zhu et al. 2021).
69 Despite extensive geochronological research, the relative timings of magma
70 emplacement and Sn-polymetallic mineralization in this province are still poorly
71 known—partly because earlier studies used conventional radiometric dating systems

72 that are readily disturbed in mineralized granite systems (e.g., mica Ar-Ar) (Yin et al.
73 2002; Cai et al. 2012; Wu et al. 2018; Liu et al. 2019; Liao et al. 2021), and partly
74 because few studies analyzed cassiterite to establish constraints on the timing of tin
75 mineralization. Here, we address this issue by providing new U-Pb dates (i.e., for the
76 Dengfuxian Dayishan, and Xianghualing deposits) as well as the first U-Pb ages for
77 cassiterite (i.e., for the Shizhuyuan and Jiuyishan deposits) for these deposits.
78 Moreover, differences in initial fluid composition and fluid evolution between various
79 types of tin deposits are not clear owing to a paucity of geochemical studies targeting
80 the main tin-bearing phase (i.e., cassiterite).

81 In this contribution, we undertook a study of cassiterite in five major W-Sn
82 deposits of southern Hunan Province, using a combination of cathodoluminescence,
83 LA-ICP-MS U-Pb ages, and trace-element concentrations. Our objectives were: (1) to
84 better constrain the timing of Sn-polymetallic mineralization; (2) to determine genetic
85 links between Sn-polymetallic mineralization and the associated granitic rocks; and (3)
86 to constraint differences in geochemical characteristics and evolution of ore-forming
87 fluids among the different types of Sn deposits by analyzing cassiterite trace-element
88 compositions.

89 REGIONAL GEOLOGY

90 The South China Craton is composed of the Yangtze Block and the Cathaysia
91 Block, sutured along the Pingxiang-Jiangshan-Shaoxing Fault Zone (Fig. 1a and 1b;
92 Li and McCulloch 1996; Zhang et al. 2017). These two blocks were amalgamated
93 during the mid-Neoproterozoic Jiangnan Orogeny (Zhang et al. 2017; Song et al.
94 2020). The South China Craton experienced multiple tectonic events between the
95 Early Paleozoic and Mesozoic, triggering the emplacement of numerous magmatic
96 rocks that are associated with economically significant hydrothermal W-Sn ore

97 deposits ([Hua et al. 2003](#); [Hu et al. 2017](#); [Jiang et al. 2020](#)).

98 Southern Hunan Province is located in the western portion of the Nanling
99 metallogenic belt of South China, which is a world-class tungsten-tin mineral district.
100 This province consists of several W-Sn-Pb-Zn-Cu polymetallic deposits, including the
101 Dengfuxian, Dayishan, Shizhuyuan, Xianghualing, and Jiuyishan deposits ([Fig. 1c](#); [Li](#)
102 [et al. 2018b](#); [Wu et al. 2018, 2021](#); [Yuan et al. 2018](#); [Jiang et al. 2020](#); [Liao et al. 2021](#);
103 [Zhu et al. 2021](#)), most of which are distributed along major NW-SE- and
104 NE-SW-trending faults ([Li and Sasaki 2007](#)). Magmatic rocks in this area were
105 emplaced during three discrete stages, i.e., the Early Paleozoic (542-360 Ma), the
106 Triassic (251-200 Ma), and the Jurassic (195-146 Ma; [Hu et al. 2017](#)). Of these
107 magmatic episodes, the Jurassic granites are the most closely associated with
108 W-Sn-Pb-Zn-Cu polymetallic deposits and, hence, have attracted considerable
109 scientific and commercial interest ([Mao et al. 2007](#); [Hu et al. 2017](#); [Yan et al. 2018](#)).

110 **DEPOSIT GEOLOGY**

111 Southern Hunan Province contains several Sn-polymetallic deposits located in
112 proximity to the Chenzhou-Linwu Fault ([Fig. 1c](#)). These bodies are often spatially
113 associated with highly differentiated granites.

114 **Dengfuxian deposit**

115 The Dengfuxian (DFX) deposit, located in Chaling County, is a
116 W-Sn-Nb-Ta-Pb-Zn polymetallic deposit hosted by the Dengfuxian Granite Complex.
117 The ore occurs in a series of quartz veins that are controlled by a network of closely
118 spaced NEE-oriented fractures. The main ore minerals include cassiterite, wolframite,
119 molybdenite, arsenopyrite, pyrite, chalcopyrite, pyrrhotite, galena, sphalerite,
120 scheelite, and quartz with minor fluorite, sericite, and chlorite constituting the gangue

121 minerals (Xiong et al. 2020). The major wallrock alterations that accompanied
122 formation of these ore minerals include muscovitization, silicification, sericitization,
123 chloritization, and dolomitization.

124 The Dengfuxian Pluton is a large granitic complex that is spatially and
125 genetically associated with a Sn polymetallic deposit (Xiong et al. 2019, 2020). The
126 granitic rocks in this composite pluton cover a total area of about 171 km² and were
127 emplaced in three main stages (Fig. 2a). Stage I is represented by coarse-grained and
128 porphyritic biotite granites emplaced at ca. 230.0 ± 1.6 Ma (Huang et al. 2011). Stage
129 II was marked by emplacement of a medium- to fine-grained two-mica granite at
130 154.0 ± 2.2 Ma (Huang et al. 2013), and its associated mineralization has yielded a
131 molybdenite Re-Os isochron age of 150.5 ± 5.2 Ma (Cai et al. 2012). Stage III is
132 represented by a fine-grained muscovite granite emplaced at 145.2 ± 0.9 Ma (Xiong et
133 al. 2020). Sedimentary and metamorphic rocks spatially associated with the
134 Dengfuxian polymetallic deposit include limestone of the Lower Permian Longtan
135 Formation as well as metamorphosed sandstone, slate, and phyllite of middle
136 Cambrian age (Xiong et al. 2020).

137 **Dayishan deposit**

138 The Dayishan (DYS) deposit is located about 30 km northwest of Chenzhou city
139 in southern Hunan Province. This Sn deposit is spatially related to the Dayishan
140 Granitic Complex, which can be divided into four units, from oldest to youngest: the
141 Shuangfengan Unit (Triassic), Guankou Unit (Early Jurassic), Tangshipu Unit
142 (Middle Jurassic), and Nibatian Unit (Late Jurassic) (Sun et al. 2018, 2021; Zhang et
143 al. 2021). Granite related to tin mineralization was emplaced mainly in the Jurassic
144 (Sun et al. 2018). The granites of the Dayishan Granitic Complex were intruded into
145 carbonate strata of the Devonian to Carboniferous age, which overlie shale,

146 conglomerate, and sandstone beds of the Ordovician age. Major structures in the area
147 include NW-trending faults and NNE-trending anticlinal folds (Fig. 2b).

148 Well-studied ore bodies in the study area include the Tengshan'ao and
149 Maozaishan deposits. The mineralization is associated with greisen-quartz veins and
150 altered granites, both of which are associated with potash alteration, silicification, and
151 greisenization (Sun et al. 2018, 2021; Zhang et al. 2021). Cassiterite, the major ore
152 mineral in this deposit, is closely associated with arsenopyrite, chalcopyrite, pyrite,
153 sphalerite, and molybdenite. Common gangue minerals are quartz, potassium feldspar,
154 tourmaline, muscovite, topaz, albite, sericite, and chlorite (Sun et al. 2018, Zhang et al.
155 2021).

156 **Shizhuyuan deposit**

157 The Shizhuyuan (SZY) deposit is a multistage W-Sn-Mo-Bi deposit located 16
158 km southeast of Chenzhou city. This deposit is closely associated with the Qianlishan
159 Granite Complex, which covers a total area of ~11 km² and consists of three intrusive
160 stages: Stage I is a microfine-grained porphyritic biotite granite with an emplacement
161 age of 183.2-152.3 Ma (Liu et al. 1997; Chen et al. 2016; Liao et al. 2021); Stage II is
162 a fine-grained biotite granite with an intrusion age of 162.6-137.4 Ma (Liu et al. 1997;
163 Yin et al. 2002; Li et al. 2004; Chen et al. 2016; Liao et al. 2021); and Stage III is a
164 medium-grained equigranular zinnwaldite granite and granite porphyry, intruded at
165 154.4-144.4 Ma (Liu et al. 1997; Chen et al. 2016; Liao et al. 2021). Other units in the
166 area include Sinian metasedimentary rocks, Devonian carbonate, and clastic
167 sedimentary rocks. The Devonian rocks are the most significant ore-bearing host
168 rocks and can be subdivided into four formations (from bottom to top): (1) Tiaomajian
169 Formation, (2) Qiziqiao Formation, (3) Shetianqiao Formation, and (4) Xikuangshan
170 Formation.

171 The W-Sn-Mo-Bi mineralization is structurally controlled by NE-trending faults
172 and fractures (Fig. 2c). The styles of mineralization include greisen-quartz vein Sn-Cu
173 ore, disseminated W-Sn-Mo-Bi ore in skarn, greisen-skarn-type W-Sn-Mo-Bi ore, and
174 disseminated W-Sn-Mo-Bi-type ore in greisen at the roof of the porphyritic biotite
175 granite pluton (Liao et al. 2021). The major ore minerals are cassiterite, wolframite,
176 scheelite, molybdenite, bismuthinite, arsenopyrite, pyrrhotite, pyrite, magnetite,
177 sphalerite, galena, and chalcopyrite. The common gangue minerals are potassium
178 feldspar, albite, quartz, muscovite, topaz, and fluorite (Liao et al. 2021). The major
179 hydrothermal alterations that accompanied ore formation are skarn metasomatism,
180 tourmalinization, and greisenization.

181 **Xianghualing deposit**

182 The Xianghualing (XHL) deposit is located in northern Linwu County, Hunan
183 Province, proximally to the Laiziling plutons. The Laiziling Pluton covers a total area
184 of 2.2 km² and has yielded a zircon U-Pb age of 156.4 ± 1.4 Ma (Yang et al. 2018). It
185 can be divided into four vertical zones, from top to bottom, the pegmatite, greisen,
186 albite granite, and albite microcline granite zones (Wu et al. 2018). Minor rocks in
187 this area include shallow-marine sand and clay deposits of Cambrian age,
188 shallow-marine carbonates of Devonian to Carboniferous age, and sandstones and
189 shales ranging from the Permian to Cretaceous in age. NE-trending faults and
190 fractures are the main ore-bearing structures (Fig. 2d).

191 The style of mineralization is skarn type, and hydrothermal alteration processes
192 such as chloritization accompanied the formation of major ore minerals including
193 cassiterite, pyrrhotite, pyrite, arsenopyrite, sphalerite, magnetite, and galena. Common
194 gangue minerals in the deposit are actinolite, tremolite, diopside, garnet, and
195 wollastonite (Wu et al. 2018).

196 **Jiuyishan deposit**

197 The Jiuyishan (JYS) deposit, located in southwestern Hunan Province, is one of
198 the most important W-Sn deposits in the Nanling region of China. This deposit is
199 closely related to the Jiuyishan Granitic Complex, which consists of the Xuehuading,
200 Pangxiemu, and Jinjiling plutons, covering 130 km², 350 km², and 49 km²,
201 respectively (Fig. 2e; Su 2017). Although the Xuehuading Pluton dates to the Early
202 Paleozoic (432.0 ± 21.0 Ma; Fu et al. 2004), the other granites were emplaced in the
203 Jurassic (156-145 Ma), with cross-cutting relationships demonstrating Pangxiemu to
204 be younger than Jinjiling (Liu et al. 2019; Li et al. 2021). This area, which also
205 exposes carbonates of Devonian to Triassic age, is dominated by NNE- and
206 N-trending fractures, accompanied by NE-, NW-, and E-trending faults.

207 The main W-Sn deposits in the Jiuyishan area, e.g., Da'ao, Shazichong, and
208 Xiangyuan, are concentrated in the western part of the Jinjiling Pluton and exhibit
209 diverse mineralization types including disseminated greisen-type, altered granite-type,
210 and greisen-quartz vein-type (Fu et al. 2007). Greisenization and silicification were
211 the major hydrothermal alteration processes that accompanied formation of the main
212 ore minerals, which include cassiterite, wolframite, pyrite, chalcopyrite, sphalerite,
213 and molybdenite. Common gangue minerals are potassium feldspar, plagioclase,
214 topaz, albite, quartz, and muscovite (Fu et al. 2007).

215 **SAMPLING AND ANALYTICAL METHODS**

216 Representative samples of Sn ore were collected from the five deposits. Sample
217 DFX is from a sulfide-bearing W-Sn quartz vein in muscovite granite of the No. 13
218 tunnel of the Dengfuxian tungsten deposit (Fig. 3a-c), and sample DYS is from the
219 greisen ore body (ore vein width 1 m) of the Xilingxi granite in Dayishan (Fig. 3d-f).
220 Sample SZY is from the greisen-skarn vein (not intensely skarn mineralized) of the

221 No. 490 tunnel of the Shizhuyuan deposit (Fig. 3g-i), sample XHL is from the
222 skarn-related Sn- and As-bearing granite in exploration line No. 47 of Tunnel 272 of
223 the Xinfeng Ore Block (Fig. 3j-l), and sample JYS is from the greisen-quartz vein of
224 the Xiangyuan mine (Fig. 3m-o).

225 After observation by optical microscopy (Fig. 4), each sample was crushed and
226 different phases were separated for further analysis. The analytical techniques used in
227 this study, i.e., cathodoluminescence (CL) imaging and mapping, geochronology and
228 trace elements analyses, and statistical methods (PCA and OPSL-DA) are detailed in
229 the Supplementary Material 1 ([10.6084/m9.figshare.22187755](https://doi.org/10.6084/m9.figshare.22187755)) and have been made
230 publicly available in Figshare (<https://figshare.com/account/home>, 2023/02). SnO₂
231 reserve data for the tin deposits involved in this study were obtained from the
232 National Geological Archive of China (<https://www.ngac.cn/>, 2023/02).

233 RESULTS

234 Cassiterite petrography

235 All five of the studied tin deposits contain only a single generation of cassiterite,
236 each showing similar textural characteristics. Sample DFX grains are
237 quasi-automorphic, grayish-yellowish in color, ~400 μm in diameter, and lightly
238 altered (Fig. 4a-b). Sample DYS grains are yellow-brown, up to 1 mm in diameter
239 (the largest among the five deposits), and strongly altered in greisen veins (Fig. 4c-d).
240 Sample SZY grains are irregular in shape, yellow-brown, ~200 μm in size, and
241 strongly altered (Fig. 4e-f). Sample XHL grains are elongated (~300 μm in length)
242 and usually broken, gray-yellow (Fig. 4g-h). Sample JYS grains are
243 quasi-automorphic, brown, and ~200 μm in size, and strongly altered (Fig. 4i-j).

244 **Cassiterite U-Pb geochronology**

245 Analytical results of cassiterite U-Pb dating are presented in Supplementary
246 Material 2 ([10.6084/m9.figshare.22187752](https://doi.org/10.6084/m9.figshare.22187752)). Analyzed grains range in size from 70 to
247 300 μm and in color from dark grey to black, most having euhedral to subhedral
248 shapes and displaying evident oscillatory zonation or patchiness in CL images (Fig.
249 5a-e). They frequently contain hydrothermal alteration veins or veinlet cracks. These
250 characteristics can provide a reference for the location of laser ablation spots, and
251 avoid the ablation of grains with exceptionally low or high U or Pb contents (which
252 are usually pure black or off-white) prone to yielding poor intercept ages. Because the
253 128 cassiterites tested have relatively low U (0.1-53 ppm) and Th (0-6.5 ppm)
254 contents, crystallization ages and initial $^{207}\text{Pb}/^{206}\text{Pb}$ are calculated from a
255 Tera-Wasserburg concordia plot. Cassiterite analyses from the five deposits form
256 Tera-Wasserburg isochrons defining lower-intercept dates of 142.0 ± 17.1 Ma
257 (MSWD = 1.40; $n = 25$, initial $^{207}\text{Pb}/^{206}\text{Pb} = 0.884 \pm 0.020$) for Dengfuxian (Fig. 6a),
258 154.7 ± 3.0 Ma (MSWD = 0.96; $n = 40$, $^{207}\text{Pb}/^{206}\text{Pb} = 0.852 \pm 0.013$) for Dayishan
259 (Fig. 6b), 154.9 ± 2.1 Ma (MSWD = 1.50; $n = 21$, $^{207}\text{Pb}/^{206}\text{Pb} = 0.851 \pm 0.081$) for
260 Shizhuyuan (Fig. 6c), 155.4 ± 4.8 Ma (MSWD = 0.43; $n = 20$, $^{207}\text{Pb}/^{206}\text{Pb} = 0.833 \pm$
261 0.018) for Xianghualing (Fig. 6d), and 152.0 ± 5.8 Ma (MSWD = 1.50; $n = 23$,
262 $^{207}\text{Pb}/^{206}\text{Pb} = 0.828 \pm 0.007$) for Jiuyishan (Fig. 6e). Overall, cassiterites from the five
263 deposits yielded lower intercept ages that range from 155.4 ± 4.8 Ma to 142.0 ± 17.1
264 Ma, indicating a Late Jurassic to earliest Cretaceous age of formation.

265 **Trace element compositions**

266 The trace element compositions of cassiterite grains from the five studied W-Sn
267 deposits are given in Supplementary Material 2 ([10.6084/m9.figshare.22187752](https://doi.org/10.6084/m9.figshare.22187752)). To
268 show more clearly the trace element characteristics of cassiterite in each type of

269 deposit, we present here a pooled dataset (note: all ranges represent 16th-84th
270 percentiles in order to avoid the influence of outliers).

271 Cassiterite in the quartz vein-type deposit (DFX) has the highest concentrations
272 of Fe (median = 1892 ppm, 243-5091 ppm), Mn (median = 8.5 ppm, 1.1-171 ppm)
273 and W (median = 4224 ppm, 51.6-18799 ppm), and the lowest concentrations of Zr
274 (median = 8.3 ppm, 3.8-14.0 ppm), U (median = 0.8 ppm, 0.3-1.3 ppm), Nb (median
275 = 20.3 ppm, 7.3-278 ppm), Ta (median = 2.2 ppm, 0.2-4.4 ppm) and Hf (median = 0.5
276 ppm, 0.1-1.2 ppm).

277 Cassiterite in the greisen-skarn-type (SZY) deposit has the highest
278 concentrations of Ti (median = 1067 ppm, 352-4701 ppm) and U (median = 34.1 ppm,
279 25.8-47.3 ppm), but rather low concentrations of Nb (median = 302 ppm, 163-549
280 ppm), Ta (median = 83.0 ppm, 44.9-122 ppm), Zr (median = 203 ppm, 127-347 ppm)
281 and Hf (median = 15.9 ppm, 6.2-26.4 ppm).

282 Cassiterites in both the granite-type (XHL) and greisen-type deposits (DYS and
283 JYS) contain relatively high concentrations of Nb (XHL, median = 567 ppm,
284 101-1367 ppm; DYS, median = 2772 ppm, 1582-5513 ppm; JYS, median = 1867 ppm,
285 315-9789 ppm), Ta (XHL, median = 680 ppm, 92-1763 ppm; DYS, median = 292
286 ppm, 189-1019 ppm; JYS, median = 994 ppm, 15.5-3788 ppm), Zr (XHL, median =
287 145 ppm, 57.2-145 ppm; DYS, median = 151 ppm, 101-283 ppm; JYS, median = 118
288 ppm, 37.1-498 ppm), Hf (XHL, median = 43.0 ppm, 13.9-48.4 ppm; DYS, median =
289 19.8 ppm, 12.7-45.5 ppm; JYS, median = 10.0 ppm, 2.1-74.4 ppm), and Ti (XHL,
290 median = 2023 ppm, 742-2576 ppm; DYS, median = 97.0 ppm, 47.3-163 ppm; JYS,
291 median = 1139 ppm, 376-1675 ppm). The total REE (Σ REE) contents of cassiterites
292 from all five deposits are quite low and not significantly different (DFX, 0-0.40 ppm;
293 DYS, 0-0.40 ppm; SZY, 0-1.00 ppm; XHL, 0-0.40 ppm; JYS, 0-0.43 ppm), and REEs

294 of importance (i.e., Eu and Ce) will be considered in the Discussion section below.

295 **DISCUSSION**

296 **Timing of the Sn-polymetallic mineralization**

297 Only small amounts of U and Th usually partition into cassiterite, which
298 precludes considerable radiation-related damage of the mineral ([Lehmann et al. 2021](#)).
299 Thus, in situ U-Pb dating of cassiterite provides a robust constraint on the timing of
300 evolved magmatic or magmatic-hydrothermal systems responsible for rare-metal
301 mineralization ([Zhang et al. 2017](#)). In this section, we discuss our newly obtained
302 cassiterite U-Pb crystallization ages in the context of the magmatic and metallogenic
303 evolution of the five studied deposits of southern Hunan Province ([Fig. 7](#)). Overall,
304 these ages offer novel insights into the timing of magmatic-hydrothermal processes
305 responsible for the production of the largest Sn reserves in China.

306 The main episode of tin mineralization at Dengfuxian is considered to be
307 associated with the formation of Stage II granites ([Cai et al. 2012](#); [Huang et al. 2013](#)).
308 However, recent studies have found that the muscovite granite (i.e., Stage III granite),
309 which was the product of residual magma fractional crystallization, also has tin
310 mineralization potential (141-137 Ma; [Xiong et al. 2020](#)). Our cassiterite grains are
311 spatially correlated with the Stage III granite, and the U-Pb dating results (142.0 ±
312 17.1 Ma) also support this relationship despite significant age uncertainty. The main
313 reason for the large age uncertainty is the very low U and Pb contents of these
314 cassiterite grains, which were probably precipitated late in the mineralization history
315 of the deposit. This observation implies that Early Cretaceous granites (~145-130 Ma)
316 in this region have a strong potential for Sn-polymetallic mineralization.

317 There is as yet no consensus on the emplacement age of the numerous granite

318 bodies in the Dayishan area (Zeng 2013). Geochronological studies of
319 granite-associated deposits can help to resolve this issue. Several radiometric studies
320 have yielded Middle Jurassic dates, e.g., 151.1 ± 1.5 Ma for a muscovite ^{40}Ar - ^{39}Ar
321 plateau age of the Tengshan'ao deposit (Zhang et al. 2021), 157.9 ± 7.7 Ma for a
322 molybdenite Re-Os isochron age of the Maozaishan deposit (Sun et al. 2018), and
323 156.5 ± 2.8 Ma for a cassiterite U-Pb age of the Maozaishan deposit (Sun et al. 2018).
324 The geochronological results for these nearby deposits overlap, within analytical
325 uncertainty, with our cassiterite U-Pb date of 154.7 ± 3.0 Ma, thus suggesting that tin
326 mineralization in the Dayishan area was related to the intrusion of granite plutons and
327 occurred during the Late Jurassic.

328 Cassiterite grains from the Shizhuyuan deposit yield a concordia intercept age of
329 154.9 ± 2.1 Ma, which is more precise than existing zircon U-Pb ages of ca. 160-150
330 Ma reported for the porphyritic biotite granite hosting this deposit (Chen et al. 2016).
331 This date overlaps with a molybdenite Re-Os isochron age of 151.0 ± 3.5 Ma reported
332 for W-Sn-Mo-Bi mineralization (Li and Mao 1996), which also supports late-stage
333 formation of greisen-skarn-type cassiterite in the SZY deposit, during which granitic
334 fluids provided the heat source and metals (Liao et al. 2021).

335 The age of Sn mineralization of the Xianghualing deposit is constrained by in
336 situ U-Pb dating of cassiterite, which yielded a Tera-Wasserburg lower intercept age
337 of 155.4 ± 4.8 Ma (MSWD = 0.43), consistent with a zircon U-Pb age of 156.4 ± 1.4
338 Ma reported for granite of the Laiziling Pluton (Yang et al. 2018). This finding
339 suggests that Sn mineralization was coeval with granite emplacement.

340 Finally, our U-Pb cassiterite date for the Jiuyishan deposit (152.0 ± 5.8 Ma)
341 agrees, within analytical uncertainty, with the Re-Os molybdenite date of the
342 greisen-quartz vein of the Da' ao tin mine (151.3 ± 2.4 Ma, Fu et al. 2007) and the

343 U-Pb age of zircons from the Jinjiiling granite with which it is spatially associated
344 (156.4 ± 0.7 Ma, [Li et al. 2021](#)).

345 Overall, cassiterite grains from the five studied tin deposits yield a relatively
346 narrow range of Late Jurassic to earliest Cretaceous ages (155.4 to 142.0 Ma),
347 indicating that the deposits are temporally linked to the multistage emplacement of
348 granitic plutons in the study area ([Xiong et al. 2020](#)). In a geodynamic context, the
349 Middle Jurassic to Early Cretaceous (180-125 Ma) magmatism and tin-tungsten
350 mineralization events in the South Ridge area may have been related to the westward
351 subduction of Pacific Ocean lithosphere beneath the Eurasian continent, which
352 triggered asthenospheric upwelling owing to lithospheric thinning in an extensional
353 tectonic setting ([Mao et al. 2007](#)). The new cassiterite U-Pb ages determined here
354 overlap with the main mineralization interval (165-150 Ma) of the world-class W-Sn
355 metallogenic province in the Nanling region ([Mao et al. 2007](#)). However, the
356 timescale of W-Sn mineralization in southern Hunan Province extends into the earliest
357 Cretaceous, possibly due to the continued enrichment of mineralizing elements in
358 residual magmas ([Xiong et al. 2019, 2020](#)). This inference indicates that Early
359 Cretaceous granites in the study area should be regarded as targets for W-Sn
360 exploration.

361 **Cassiterite compositions trace the nature of ore-forming fluids**

362 **Origin of Sn.** Cassiterite typically contains a wide range of trace elements such
363 as Ti, Nb, Ta, Fe, Mn, and W, which can either substitute for Sn^{4+} in the mineral
364 lattice or be present in the form of exsolved mineral inclusions ([Taylor 1979](#)). These
365 trace elements can be used to track the nature of mineralizing fluids as well as the
366 growth environment of the mineral ([Tindle and Breaks 1998; Cheng et al. 2019; Wu et](#)
367 [al. 2021](#)). In particular, Nb, Ta, Fe, W, and Mn in cassiterite are often used to

368 determine the type of tin deposits and the origin of tin-bearing fluids (Tindle and
369 Breaks 1998). A Nb+Ta versus Fe+Mn discriminant diagram shows that all five
370 deposits are typical magmatic-hydrothermal tin deposits related to granite (Fig. 8).
371 Considering the possible positive correlation between iron and tungsten content in
372 cassiterite (Yu and Jiang 2001) and that sample DFX is a quartz vein-type sample
373 dominated by W in a W-Sn deposit (n.b. the only one out of the five studied deposits),
374 the contents of W and Fe in this sample are high (Fig. 9d). Moreover, its low Nb and
375 Ta contents are consistent with its origin as a hydrothermal vein-type cassiterite
376 (Möller et al. 1988; Fig. 9a).

377 **Metallogenesis temperature.** Previous studies of tin deposits have shown that
378 the concentrations of HFSEs such as Nb, Ta, Zr, and Hf in cassiterite reflect the
379 temperature of the mineralizing fluid. Specifically, cassiterite formed at high
380 temperatures is more enriched in Nb, Ta, Zr, and Hf than that formed at low
381 temperatures (Cheng et al. 2019; Gemmrich et al. 2021; Hu et al. 2021; Wu et al.
382 2021). In addition to temperature, the early-precipitated minerals play a crucial role in
383 controlling the chemical composition of later-formed minerals. Experimental studies
384 of the partitioning of HFSEs in zircons have shown that Hf in the residual melt or hot
385 liquid decreases substantially at lower temperatures (Wang et al. 2010). Furthermore,
386 because HFSE-enriched accessory minerals are more enriched in Ta relative to Nb in
387 felsic melts ($D_{\text{Nb}}/D_{\text{Ta}} = 0.3\text{-}0.4$ for sphene, $0.6\text{-}0.7$ for rutile, $0.7\text{-}0.8$ for ilmenite, and
388 0.8 for titanomagnetite in high-silica rhyolites; Green and Pearson, 1987), their
389 precipitation leads to greater enrichment of Nb relative to Ta and Zr relative to Hf in
390 the residual melt and associated fluid phase. Therefore, variation in the elemental
391 composition of cassiterite is a function of temperature and mineral fractionation (Wu
392 et al. 2021). During high-temperature mineralization events (e.g., greisen and granite),

393 hydrothermal fluids become enriched in Nb, Ta, Zr, and Hf, which can then be readily
394 incorporated into the cassiterite lattice by isomorphic substitution. As the temperature
395 of the ore-forming fluid decreases, cassiterites precipitated at this stage (e.g., skarn
396 and veins) show a sharp decrease in HFSEs, and the Zr/Hf and Nb/Ta ratios increase
397 with the precipitation of incompatible element-enriched minerals (Fig. 9a-c). The high
398 contents of Nb, Ta, Zr, and Hf in cassiterite grains from greisen, greisen-quartz veins,
399 and altered granites (e.g., samples DYS, JYS, and XHL) indicate that they were
400 formed from higher-temperature fluids than cassiterite grains from greisen-skarn and
401 quartz veins (e.g., samples SZY and DFX), which have lower Nb, Ta, Zr, and Hf
402 contents (Fig. 9a-c). Thus, Nb, Ta, Zr, and Hf contents are indicative of mineralization
403 temperatures (Fig. 9a-c), which is consistent with the results of other
404 geothermometers that have been applied to these deposits (Xiong et al. 2019; Wang et
405 al. 2020; Zhao et al. 2022a, b).

406 **Variable chemistry of ore-forming fluids.** Differences in chemical composition
407 and physio-chemical conditions between ore-forming fluids of tin mineralization and
408 the precursor magma are reflected in the HFSE content of cassiterite, its precipitation
409 temperature, and the type of tin mineralization (Cheng et al. 2019; Wu et al. 2021). In
410 a geochemical system that is characterized by charge- and radius-controlled (ChaRaC)
411 behavior, some trace-element pairs with similar ionic radii and valence states (e.g.,
412 Nb-Ta and Zr-Hf) exhibit coherent behavior and maintain chondritic to
413 near-chondritic ratios (Bau 1996; Rudnick et al. 1993). Moreover, the interelement
414 ratios measured in magmatic systems can be influenced by chemical reactions and
415 fluid mixing, and non-ChaRaC behavior reflects a specific magmatic-hydrothermal
416 system that is highly evolved and enriched in H₂O, Li, B, F, P, and/or Cl (Bau 1996).
417 Cassiterites from the five studied W-Sn deposits have a wide range of Nb/Ta ratios

418 (DFX 14.16-68.5; DYS 6.64-15.4; XHL 3.79-45.0; SZY 5.43-54.0; JYS 7.36-48.0)
419 and Zr/Hf ratios (DFX 1.04-297; DYS 0.85-14.8; XHL 0.33-81.6; SZY 0.52-14.6;
420 JYS 0.81-718), implying derivation from a highly evolved melt. The wide range of
421 Nb/Ta and Zr/Hf ratios of cassiterite in these deposits (Fig. 9a-c) can be explained by
422 the influence of multiple processes, as described in previous studies. For example,
423 quartz vein-type cassiterite (DFX) can be produced by fluid immiscibility (Xiong et al.
424 2019), greisen-skarn-type cassiterite (SZY) by mixing ore-forming magmatic fluid
425 with meteoric water (Zhu et al. 2015), and greisen-type cassiterite (DYS and JYS) by
426 fluid-rock interactions (Korges et al. 2018; Schmidt et al. 2020; Liu et al. 2021; Zhao
427 et al. 2022a, b). In Nb-Ta-Zr-Hf discriminant plots (Fig. 9a-c), granite-type cassiterite
428 (XHL) exhibits a limited range of Nb/Ta and Zr/Hf ratios. In contrast, the greisen-type
429 cassiterite is more enriched in magmatic volatiles such as F and B, with the
430 hydrothermal fluid source of DYS exhibiting a narrower range of Nb/Ta and Zr/Hf
431 ratios than that of JYS (Rubin et al. 1993; Cheng et al. 2019; Zhao et al. 2022a, b).

432 Titanium (Ti) and uranium (U) are both incompatible elements that are difficult
433 to incorporate into the mineral lattice of rock formations, and they tend to be
434 relatively enriched in residual magmas and hydrothermal fluids (Han et al. 2003). The
435 Ti and U concentrations in cassiterite are thought to be related to the degree of
436 evolution of the parent magma, but the mechanism of their uptake during mineral
437 precipitation is not well understood (Hu et al. 2021). To gain a deeper understanding
438 of the influence of associated granites on the Ti and U contents of cassiterite, we
439 conducted a comparative Ti-U analysis of the cassiterite and associated granites in the
440 five studied tin deposits. Interestingly, our findings suggest that the Ti and U contents
441 of cassiterite are positively influenced by their associated granites (Fig. 10). Although
442 both are incompatible elements, Ti is relatively less incompatible than U and more

443 easily enters the cassiterite lattice. Additionally, the ionic radius of Ti (0.605\AA) is
444 slightly smaller than that of Sn (0.69\AA), allowing it to readily replace Sn in cassiterite
445 (Cheng et al., 2019), leading to lower TiO_2 contents in highly evolved tin-bearing
446 granites. Combined with the fact that sample XHL is associated with evolved granites,
447 this may be the reason for the low Ti content of the DYS sample, whereas the Ti
448 content of cassiterite in the low- TiO_2 Laiziling granite is not low (Fig. 10a).
449 Furthermore, the consistent geochemical behavior of the cassiterite and associated
450 granites of these deposits in terms of U content confirms that the granite associated
451 with quartz vein-type tin ore exhibits the lowest degree of evolution among the
452 associated granites (Fig. 10b). In summary, Ti and U concentrations reflect the degree
453 of evolution of granitic parent magmas for tin mineralization, with an increasing
454 evolutionary degree from quartz vein-type to greisen-type to granitic-type cassiterite.

455 **Redox conditions of ore-forming processes.** Patterns of REE distribution in
456 cassiterite have been used to constrain the source(s) of ore-forming fluids, the
457 physicochemical characteristics of the fluids, and the dynamics of hydrothermal
458 systems (Brugger et al. 2000; Wu et al. 2021). Four of the five of the studied deposits
459 (excepting XHL) contain minerals coprecipitated with cassiterite that can affect the
460 REE patterns of the latter, i.e., wolframite and/or scheelite at DFX, SZY, and JYS, and
461 tourmaline at DYS. For this reason, the following discussion will focus on the REE
462 behavior of cassiterite, wolframite, scheelite, tourmaline, and hydrothermal fluids.

463 The direction (i.e., positive or negative) and size of the Eu anomaly in
464 hydrothermal minerals are chiefly controlled by three factors: (1) the redox state of
465 the fluid, (2) the partitioning coefficients of Eu^{2+} and Eu^{3+} between the host mineral
466 and the fluid, and (3) the Eu content of the fluid (Shannon 1976). In oxidizing systems,
467 given the significantly smaller ionic radius of Sn^{4+} (0.69\AA) relative to Eu^{2+} (1.25\AA)

468 and Eu^{3+} (1.07 Å), valence changes have a limited effect on the Eu anomaly of
469 cassiterite. In this case, cassiterite directly inherits the Eu anomaly of the ore-forming
470 fluid because of the similar geochemical behavior of Eu^{3+} to other REEs (Wu et al.
471 2021). On the contrary, when the ore-forming fluid is reducing, Eu^{2+} readily replaces
472 Ca^{2+} (1.12 Å) because of their equivalent electrovalences and similar ionic radii. In
473 addition, Eu^{2+} tends to partition more readily into wolframite, scheelite and
474 tourmaline in a reducing environment (Sverjensky 1984; Brugger et al. 2000; Zhu et
475 al. 2014), resulting in a negative Eu anomaly in cassiterite. In this study, samples DFX,
476 DYS, SZY, and JYS show significant negative Eu anomalies (Fig. 11), reflecting the
477 reducing condition of the ore-forming fluids, whereas sample XHL from the Laiziling
478 Granite has no Eu anomaly, indicating that its crystallization environment was
479 probably less reducing than those of the other four deposits.

480 Little is known about the factors controlling the Ce content of cassiterite.
481 Precipitation of cassiterite and scheelite does not require high- $f\text{O}_2$ conditions (Schmidt,
482 2018). The geochemical behavior of Ce^{3+} under reducing conditions is similar to that
483 of other REEs, and the Ce anomalies in hydrothermal minerals may reflect the REE
484 patterns of their precipitation environments unless Ce-rich minerals (e.g., monazite)
485 are coprecipitated. Conversely, in an oxidizing magmatic-hydrothermal system, Ce^{4+}
486 (0.97 Å) becomes more incompatible than Ce^{3+} (1.14 Å) due to the larger difference
487 in its ionic radius from Ca^{2+} (1.12 Å), resulting in negative Ce anomalies in cassiterite,
488 wolframite, and scheelite following precipitation of Ce-rich minerals like monazite
489 (Wu et al. 2021). Based on the lack of significant Ce anomalies in cassiterite (Fig. 11),
490 we infer that tin mineralization in southern Hunan Province occurred in a reducing
491 environment.

492 To summarize, the characteristic behavior of trace elements in cassiterite

493 indicates that the crystallization of that changes in physicochemical conditions (e.g.,
494 temperature) during highly evolved magmatic hydrothermal processes are important
495 factors driving the precipitation of Sn-enriched minerals, but the redox conditions of
496 the precipitated liquid are irrelevant. Based on our findings, we suggest that
497 progressively more evolved granites are responsible for quartz vein-type, greisen-type,
498 and granite-type cassiterite respectively. Through this magmatic sequence, the HFSE
499 contents of cassiterite gradually increased and their ranges of variation gradually
500 decreased, implying a relationship between these two patterns that need further study.

501 **Substitution mechanism of trace elements in cassiterite**

502 We evaluated the trace element assemblages of cassiterite in the five studied tin
503 deposits using two chemometric methods (PCA and OPLS-DA). The PCA results
504 reveal coupling between trace elements in cassiterite, with the first two components
505 accounting for 62.0% of total variance (36.4% for PC1, and 25.6% for PC2),
506 demonstrating the feasibility of dimensionality reduction. The main elemental
507 loadings are Nb, Ta, Zr, and Hf on PC1, and Fe, Mn, and W on PC2 (Fig. 12). In
508 Figure 12, vector angles smaller than 90° indicate a positive correlation between
509 elemental pairs, as for Fe + Mn vs. Nb + Ta ($r = +35.20^\circ$; Fig. 8a), W vs. Fe ($r =$
510 $+67.54^\circ$; Fig. 9d), and Nb + Ta vs. Fe ($r = +22.56^\circ$; Fig. 9f). Therefore, the typical
511 trace element substitution mechanisms of $(\text{Fe, Mn})^{2+} + 2(\text{Nb, Ta})^{5+} = 3\text{Sn}^{4+}$, $\text{W}^{6+} +$
512 $2\text{Fe}^{3+} = 3\text{Sn}^{4+}$, and $\text{Fe}^{3+} + (\text{Nb, Ta})^{5+} = 2\text{Sn}^{4+}$ (Černý and Ercit 1985; Černý et al. 1985;
513 Cohen et al. 1985; Möller et al. 1988) can be confirmed in the cassiterite samples of
514 DYS, SZY, XHL, and JYS. However, the mechanism of trace element substitution in
515 the lattice of cassiterite from the tungsten-dominated quartz vein-type cassiterite is
516 dominantly $\text{W}^{6+} + 2\text{Fe}^{3+} = 3\text{Sn}^{4+}$.

517 The PCA plot reveals the general similarities of the trace element assemblages in

518 the cassiterite of all five tin deposits. To better investigate differences in their trace
519 element assemblages, we made use of OPLS-DA. In plots of this type, R^2X_{cum} shows
520 differences in variables between groups, and R^2Y_{cum} indicates differences in variables
521 within groups. The five studied tin deposits exhibit large differences in the trace
522 element assemblages of cassiterite (for each deposit, $R^2X_{cum} > 0.5$; Fig. 13). This
523 observation indicates that the nature and content of trace elements in cassiterite are
524 significantly different depending on mineralization type and location.

525 **Insights for W-Sn exploration**

526 The in situ cassiterite U-Pb ages of the present study offer insights regarding
527 genetic links between Sn-polymetallic ores and their spatially related granites in
528 southern Hunan Province, with useful implications for exploration of economically
529 significant deposits. Critical elements (W, Sn, Cu, Pb, Zn) in granites of the Nanling
530 region show a gradual increase from Early Paleozoic (542-360 Ma) to Triassic
531 (251-200 Ma), Jurassic (195-146 Ma), and Cretaceous (145-80 Ma) plutons, with
532 tungsten and tin being especially enriched in Early Cretaceous granites (Liu et al.
533 2022b). This trend conforms broadly to the size of mining operations in southern
534 Hunan Province, with Early Paleozoic deposits being non-productive, Triassic
535 deposits being rather small (e.g., Longshang deposit, 16 Kt SnO₂; Liu et al. 2022), and
536 Jurassic deposits being quite large (e.g., Hehuaping, 130 Kt SnO₂; Yao et al. 2014).
537 However, the comparatively small size of mines extracting Cretaceous-age ores (e.g.,
538 the Jiepailing deposit, 92.1±0.7 Ma, 48 Kt SnO₂; Yuan et al. 2015) is not
539 commensurate with their high concentrations of mineralized elements (W, Sn, Cu, Pb,
540 Zn), and Early Cretaceous plutons warrant closer inspection in the future as possible
541 mining targets.

542 Based on previous studies of the southern Hunan tin province and the present

543 study of its geochronology, the magmatic emplacement and metallogenic evolution of
544 this province can be summarized as occurring in three main stages (Li et al. 2018a;
545 Luo et al. 2022; Liu et al. 2022b): (1) Stage I (Early Paleozoic and Triassic), during
546 which magma was formed by melting of lower crust and enrichment of tungsten and
547 tin occurred through crustal-mantle interactions; (2) Stage II (Jurassic), during which
548 magma was emplaced at high crustal structural levels where it reacted with wall rocks
549 and extracted metals to further enrich tin-tungsten ores; and (3) Stage III (Cretaceous),
550 during which magmatic-hydrothermal reactions were superimposed on earlier
551 mineralization events. Small dykes and deeper intrusions formed during the Early
552 Cretaceous, with associated hydrothermal fluids depositing additional tin and tungsten
553 (Fig. 14a). Therefore, ore exploration should focus on the highly evolved granites
554 from the late stages of granite complex formation.

555 Previous and present studies have inferred that the contents of Nb, Ta, Zr, and Hf
556 and the ratios of Nb/Ta and Zr/Hf gradually decrease from granite-type to
557 greisen-type and then to quartz vein-type cassiterites (Cheng et al. 2019; Hu et al.
558 2021; Wu et al. 2021). However, previous studies have not used the chemical
559 behavior of these trace elements to analyze the distance between deposits and
560 magmatic systems associated with intrusive rocks. In a systematic study of tin
561 deposits in the Bolivian tin metallogenic belt, Gemmrich et al. (2021) concluded that
562 variations in trace element content of ores emplaced at different levels within the
563 xenothermal and epithermal environments recorded genetic compositional trends.
564 Deposits that formed proximally to related intrusive complexes are typically enriched
565 in Nb and Ta relative to epithermal and shallow xenothermal deposits. The present
566 study shows that similar patterns of enrichment of Nb, Ta, and other HFSEs are
567 present in the tin deposits of southern Hunan Province, which are more closely

568 associated with granitic bodies than other types of tin mineralization ([Sun et al. 2018](#);
569 [Zhang et al. 2021](#); [Zhao et al. 2022a, b](#)). The sum of these elements provides a solid
570 ground for the establishment of a comprehensive, idealized tin metallogenic model for
571 the southern Hunan tin province and similar provinces worldwide ([Fig. 14b](#)).

572 **IMPLICATIONS**

573 A combination of cassiterite cathodoluminescence, U-Pb geochronology, and
574 trace-element data offers important insights into the nature of ore-forming fluids and
575 the extended formation history of Sn-polymetallic deposits in the southern Hunan tin
576 province, as follows:

577 1. LA-ICP-MS in situ cassiterite U-Pb dating of five major deposits (i.e.,
578 Dengfuxian, Dayishan, Shizhuyuan, Xianghualing, and Jiuyishan) yielded ages
579 ranging from 155.4 to 142.0 Ma, corresponding to the Late Jurassic to earliest
580 Cretaceous. These ages indicate that these deposits are temporally and genetically
581 linked to their spatially associated granitic rocks and expand the known age range of
582 Sn-polymetallic mineralization in southern Hunan Province.

583 2. The characteristic behavior of trace elements in cassiterite records changes in
584 physicochemical conditions (e.g., temperature) of the source magma/hydrothermal
585 fluids, which were important factors driving the precipitation of Sn-enriched minerals
586 and were not related to fluid redox conditions. HFSE content tends to increase and its
587 range of variation becomes more limited from quartz vein-type to skarn-type to
588 granite-type cassiterites. Accordingly, the degree of magmatic evolution of related
589 granites increases.

590 3. The magmatic emplacement and metallogenic evolution of the southern
591 Hunan tin province can be divided into three main stages: (1) Stage I (Early Paleozoic

592 and Triassic): initial enrichment stage of tungsten and tin; (2) Stage II (Jurassic):
593 metasomatic mineralization stage; and (3) Stage III (Cretaceous):
594 magmatic-hydrothermal superposition stage. This study shows that the highly
595 differentiated Cretaceous-age granites and their peri-plutonic areas are important
596 targets for Sn ore exploration in the Nanling region.

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906

907 **Figure captions**

908

909 **Fig. 1.** (a) Sketch map showing the main terranes of China (modified after [Wu et al.](#)
910 [2018](#)); (b) Geological map of South China Craton (modified after [Yuan et al. 2018](#)); (c)
911 Geologic map of the Nanling Range showing the age distribution of granites and
912 spatially associated tungsten/tin deposits (modified after [Mao et al. 2007](#); [Yuan et al.](#)
913 [2018, 2019](#)).

914

915 **Fig. 2.** Simplified geological maps of study areas of five tin deposits: (a) Dengfuxian
916 (after internal data from 214 Geological Team of Hunan Province, 2010, and [Xiong et](#)
917 [al. 2020b](#)); (b) Xianghualing (after [Wu et al. 2018](#)); (c) Dayishan (after [Sun et al. 2018](#)
918 and [Zeng 2013](#)); (d) Shizhuyuan (after [Liao et al. 2021](#)); (e) Jiuyishan (after [Li et al.](#)
919 [2019, 2021](#)).

920

921 **Fig. 3.** Field pictures and hand specimen photomicrographs from five tin deposits. (a)
922 Sulfide-bearing wolframite quartz veins at DFX; (b) and (c) hand specimens of
923 sulfide-bearing wolframite quartz vein; (d) greisen-quartz veins at DYS; (e) and (f)
924 hand specimens of greisen-quartz veins; (g) skarn-greisen network veins at SZY; (h)
925 and (i) hand specimens of skarn-greisen network veins; (j) biotite granite at XHL; (k)
926 and (l) hand specimens of granite; (m) to (o) greisen veins at JYS. Mineral
927 abbreviations: Py = pyrite, Cst = cassiterite, Grt = garnet, Bt = biotite.

928

929 **Fig. 4.** Transmitted light photomicrographs of cassiterite grains from five tin deposits,
930 (a, c, e, g, i plane-polarized transmitted light; b, d, f, h, j cross-polarized transmitted
931 light). DFX (a-b): Semi-idiomorphic cassiterite crystallized in quartz vein; DYS (c-d):

932 Metasomatic cassiterite in greisen vein; SZY (e-f): Cassiterite crystallized in greisen
933 skarn vein with obvious metasomatism; XHL (g-h): Semi-idiomorphic cassiterite
934 crystallized in granite; JYS (i-j): Semi-idiomorphic cassiterite crystallized in greisen.
935 Mineral abbreviations: Qtz = quartz, Cst = cassiterite, Ms = muscovite, Bt = biotite.

936

937 **Fig. 5.** Cathodoluminescence (CL) images of cassiterites from five tin deposits: (a)
938 DFX; (b) DYS; (c) SZY; (d) XHL; and (e) JYS.

939

940 **Fig. 6.** Tera–Wasserburg concordia diagrams for cassiterite grains from five tin
941 deposits: (a) DFX, (b) DYS, (c) SZY, (d) XHL, and (e) JYS.

942

943 **Fig. 7.** Geochronologic chart of Sn-ore mineralization and related granites in southern
944 Hunan Province, including the new U-Pb cassiterite data of the present study.

945

946 **Fig. 8.** Plots of selected trace elements in cassiterite grains from five tin deposits: (a)
947 Nb+Ta vs Fe+Mn in cassiterite (modified from [Tindle and Breaks 1998](#)); (b) Fe vs W
948 in cassiterite (analyzed by LA–ICP–MS). The granite-related tin deposits field is
949 based on data from [Hennigh and Hutchinson \(1999\)](#); [Wang et al. \(2014\)](#); and [Pavlova](#)
950 [et al. \(2015\)](#), and the VMS/SEDEX tin deposits field is from [Hennigh and Hutchinson](#)
951 [\(1999\)](#). The dashed line represents the approximate detection limit for W by PIXE
952 probe ([Hennigh and Hutchinson 1999](#)).

953

954 **Fig. 9.** Log-scale scatter plots of selected trace elements for cassiterite grains from
955 five tin deposits: (a) Nb vs Ta, (b) Zr vs Hf, (c) Zr/Hf vs Nb/Ta; (d) Fe vs W, (e) Zr vs
956 Ti, and (f) Fe vs Nb+Ta. Sources: [Guo et al. 2018](#); [Chen et al. 2019](#); [Cheng et al. 2019](#);

957 [Hu et al. 2021](#); [Wu et al. 2021](#).

958

959 **Fig. 10.** Log-scale scatter plots of U and Ti content for cassiterite grains and
960 associated granite from five tin deposits: (a) Ti content, (b) U content. Sources: [Li et](#)
961 [al. 2019](#); [Sun et al. 2021](#); [Liao et al. 2021](#); [Yang et al. 2018](#); [Liu et al. 2019](#); [Li et al.](#)
962 [2021](#).

963

964 **Fig. 11.** Chondrite-normalized REE patterns for southern Hunan cassiterite-hosted tin
965 deposits (normalization after [Sun and McDonough 1989](#)). (a) DFX, (b) DYS, (c) SZY,
966 (d) XHL, (e) JYS, and (f) all samples. Data for quartz vein-type cassiterite in the
967 Tongshanling deposit ([Wu et al. 2021](#)) is included for reference.

968

969 **Fig. 12.** Score graph and loading graph of principal component analysis (PCA) of
970 cassiterite from southern Hunan Province.

971

972 **Fig. 13.** Score plot of the orthogonal projection to latent structures-discriminant
973 analysis (OPLS-DA) models between the five tin deposits. (a) DFX vs. DYS, (b) DFX
974 vs. SZY, (c) DFX vs. XHL, (d) DFX vs. JYS, (e) SZY vs. XHL, (f) SZY vs. JYS, (g)
975 DYS vs. JYS, (h) DYS vs. SZY, and (i) DYS vs. XHL.

976

977 **Fig. 14.** Diagenesis and mineralization model diagram of Sn polymetallic deposits in
978 southern Hunan Province. (a) Schematic diagram of the multistage
979 magmatic-hydrothermal evolution developed from previously reported data ([Li et al.](#)
980 [2018a](#); [Luo et al. 2022](#); [Liu et al. 2022b](#)) and the results of this study, and (b) ideal tin
981 metallogenic model diagram (inspired by [Gemmrich et al. 2021](#)).

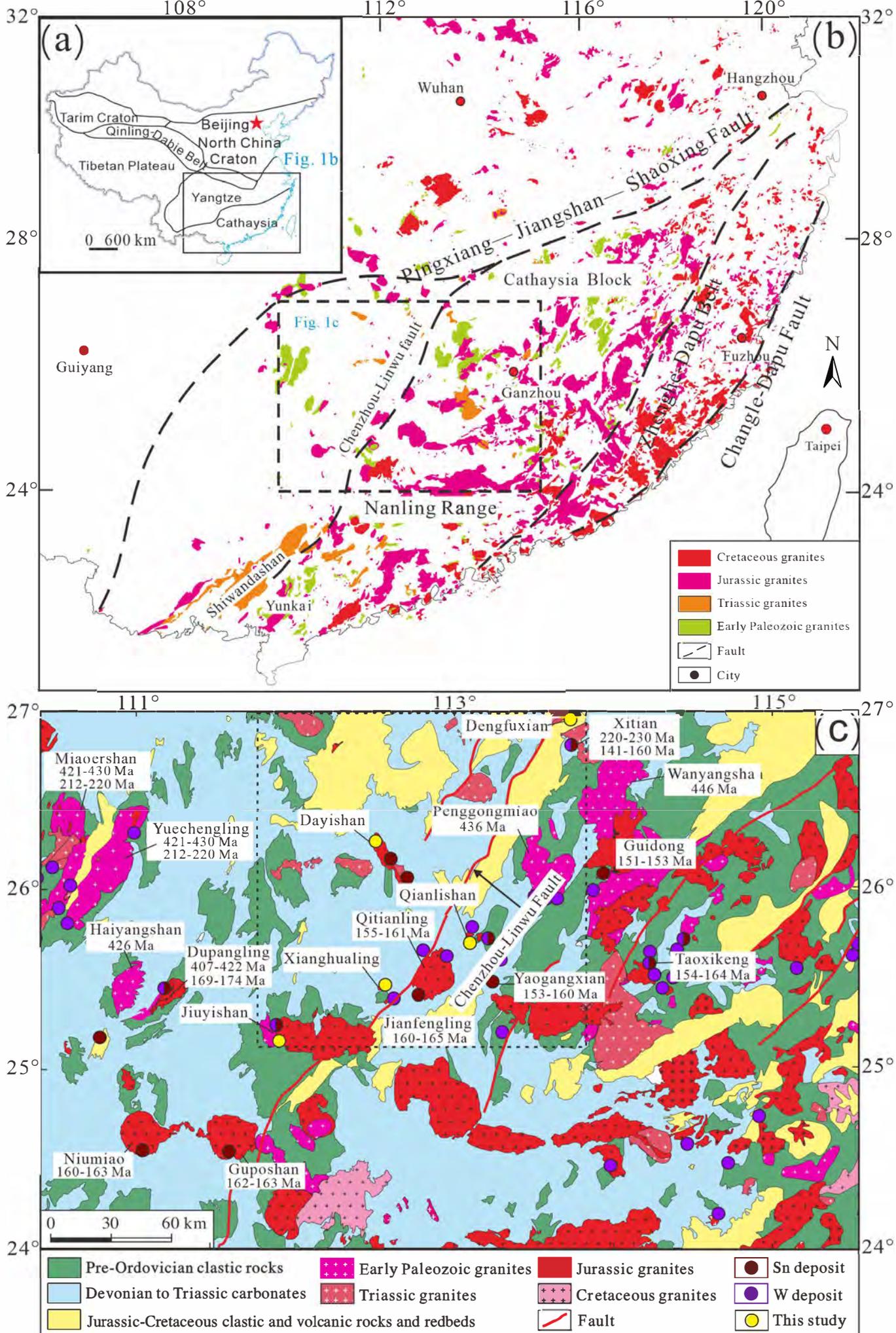


Figure 1

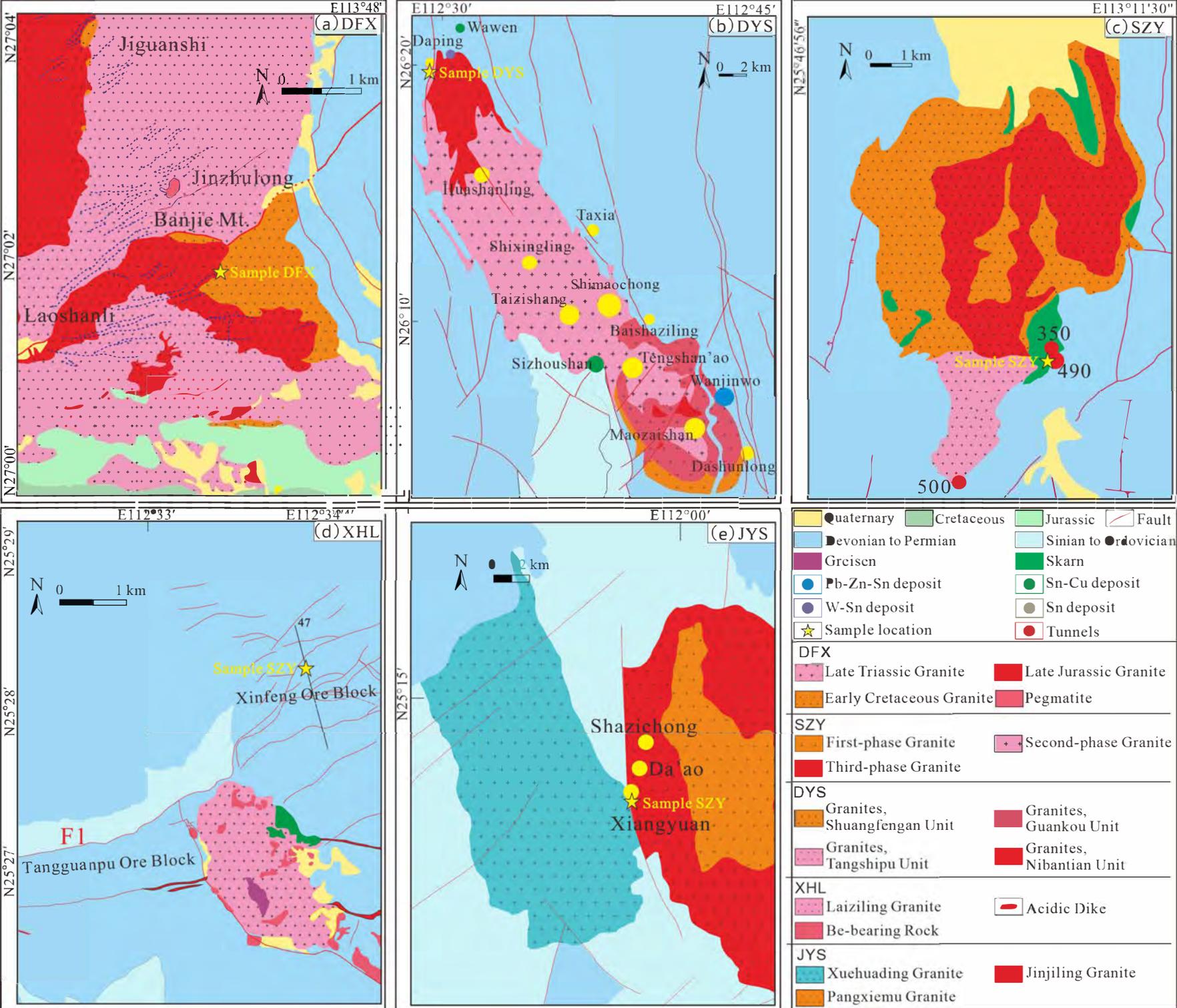


Figure 2

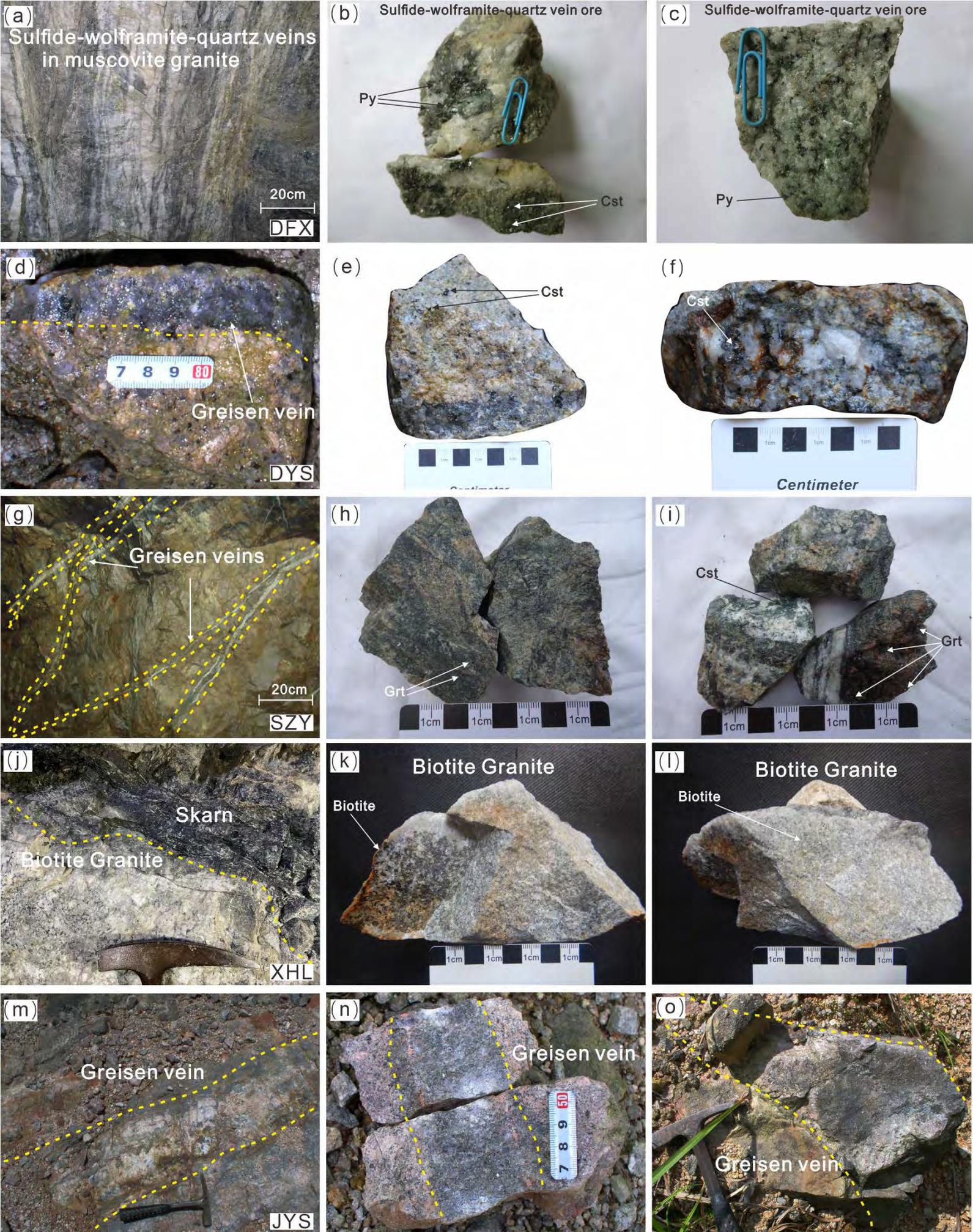


Figure 3

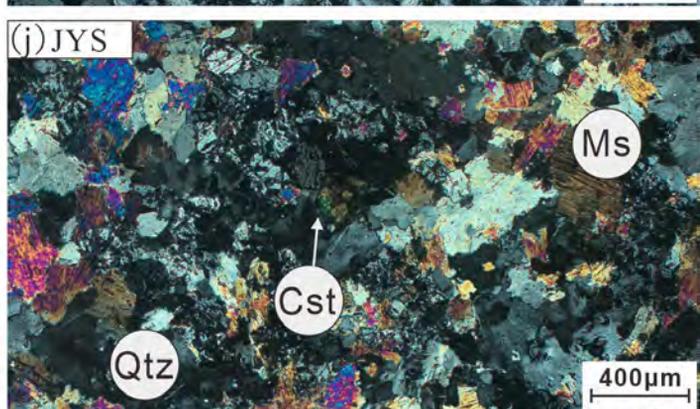
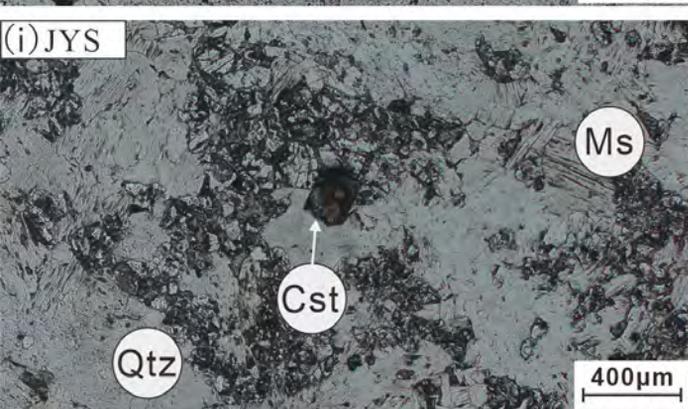
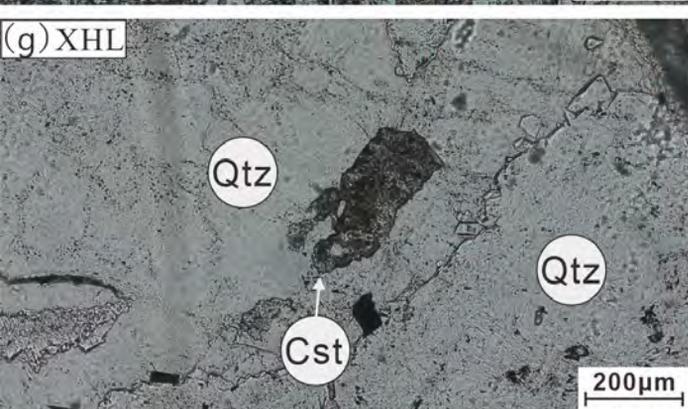
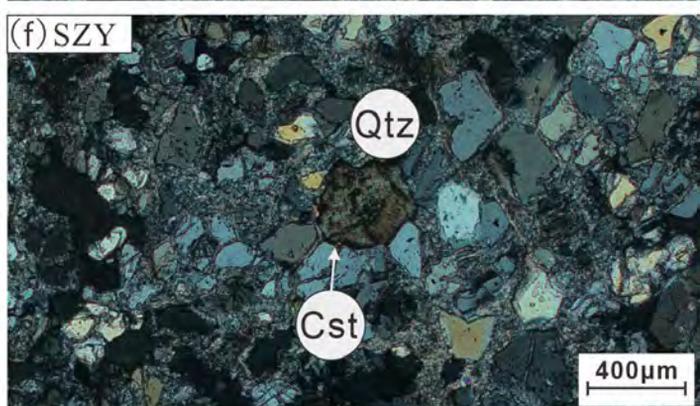
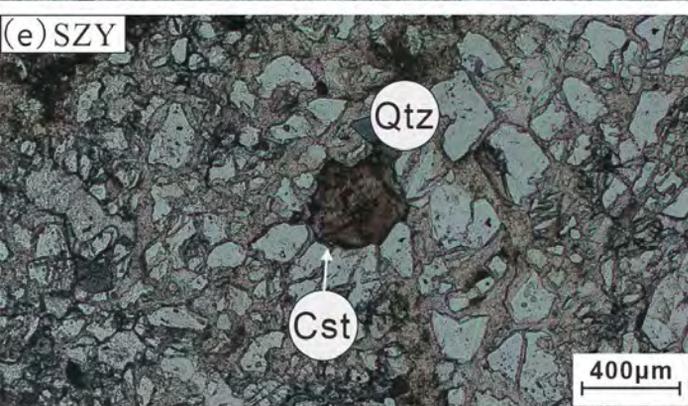
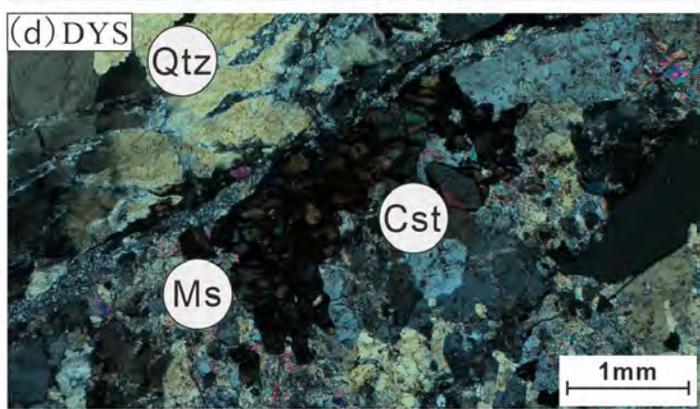
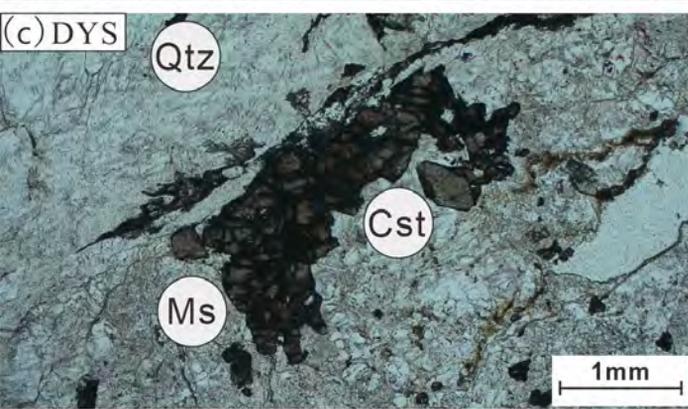
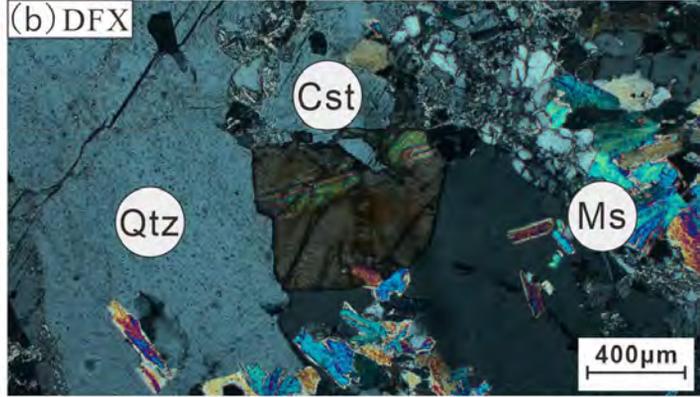
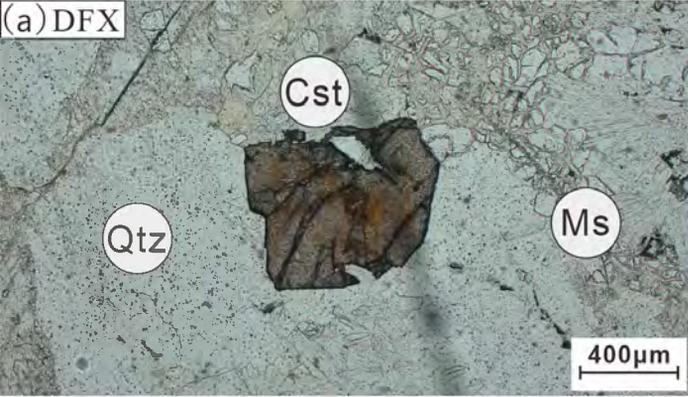


Figure 4

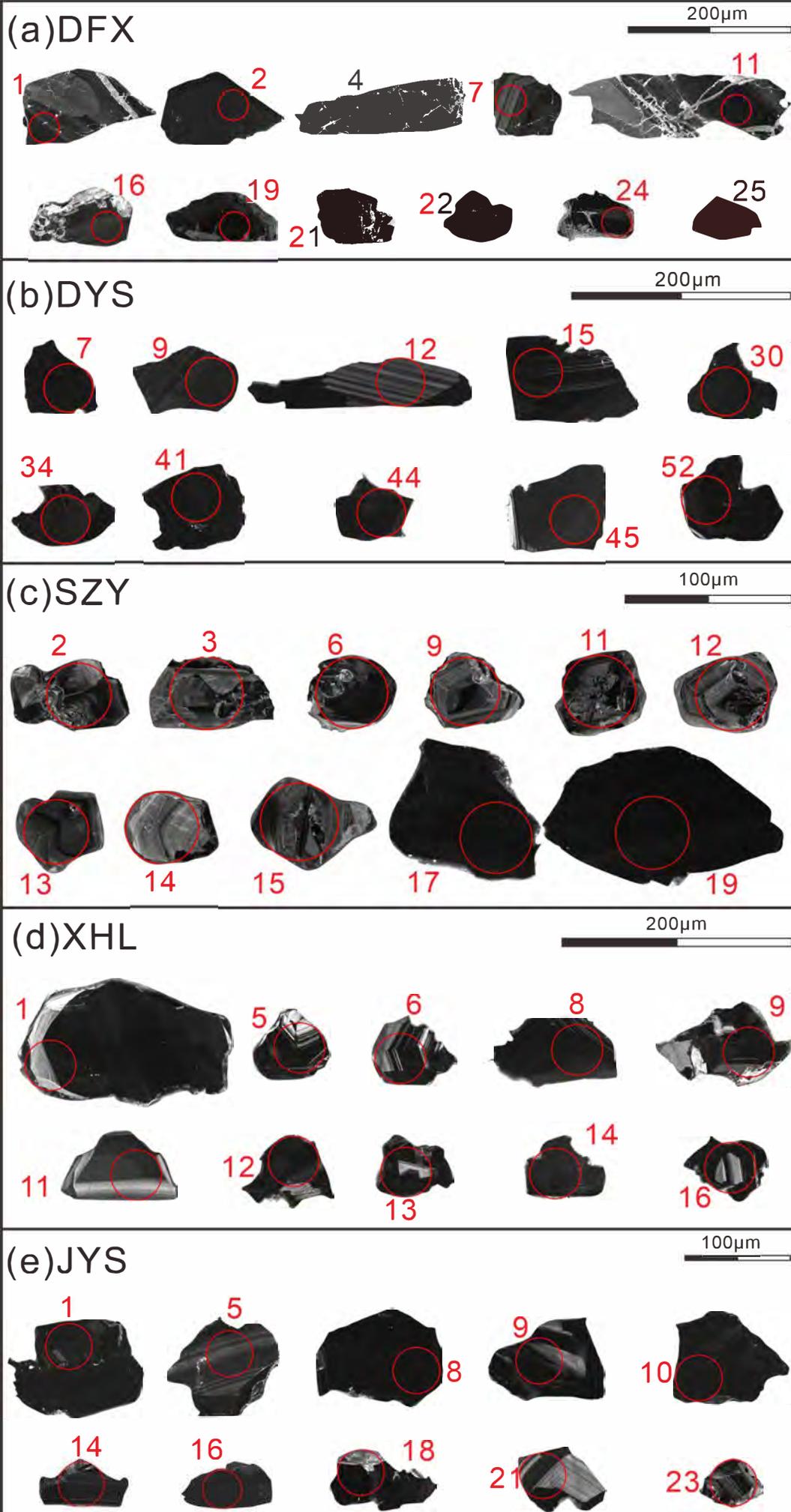


Figure 5

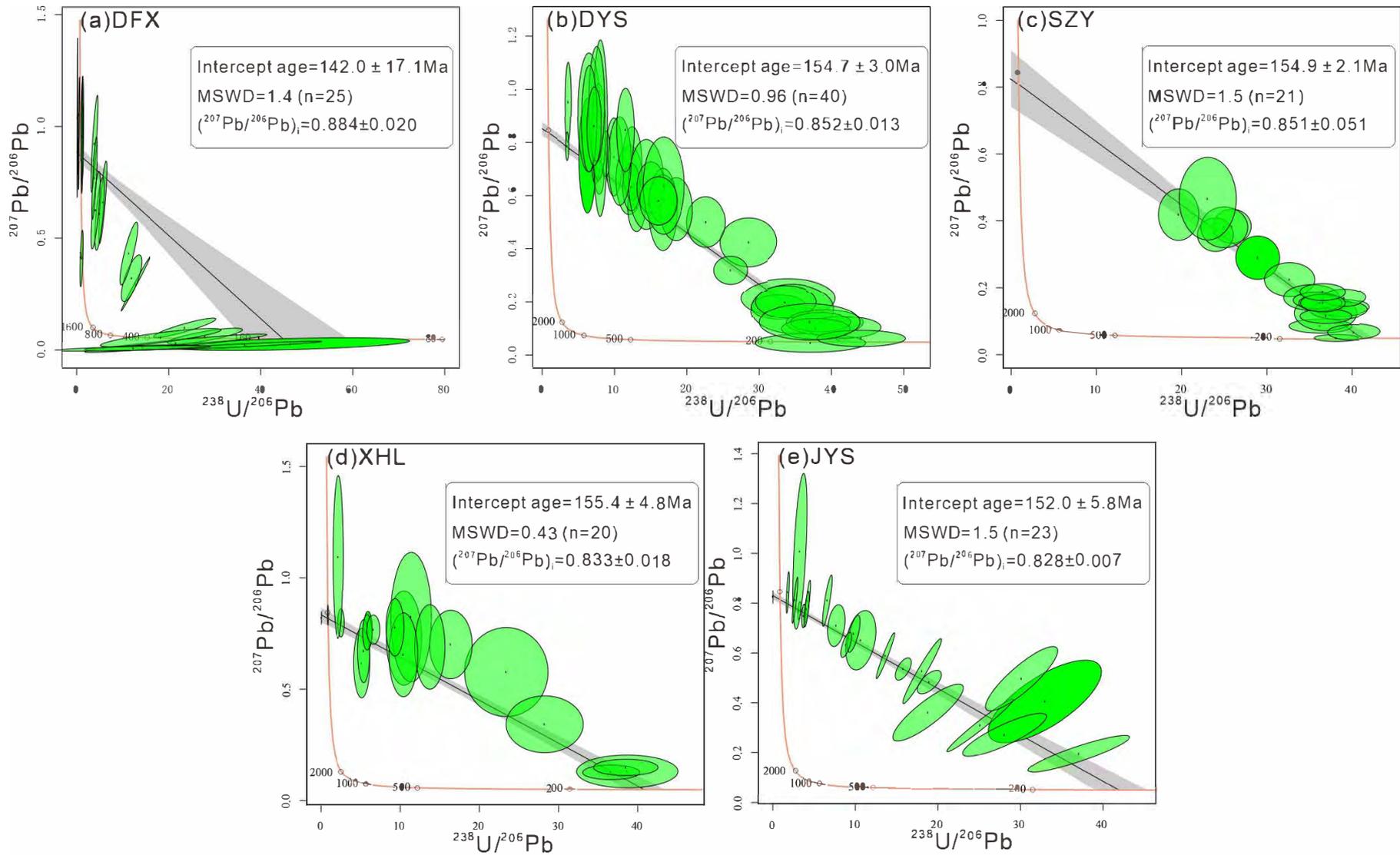
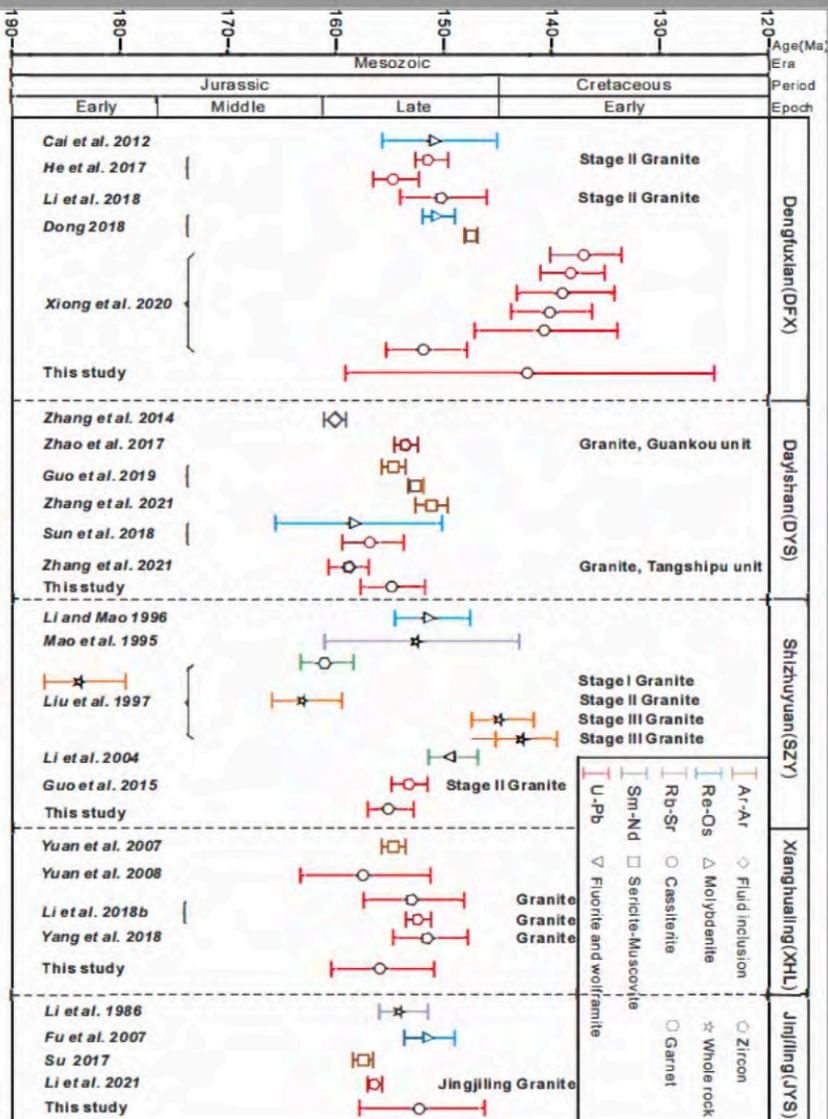


Figure 6

Figure 7



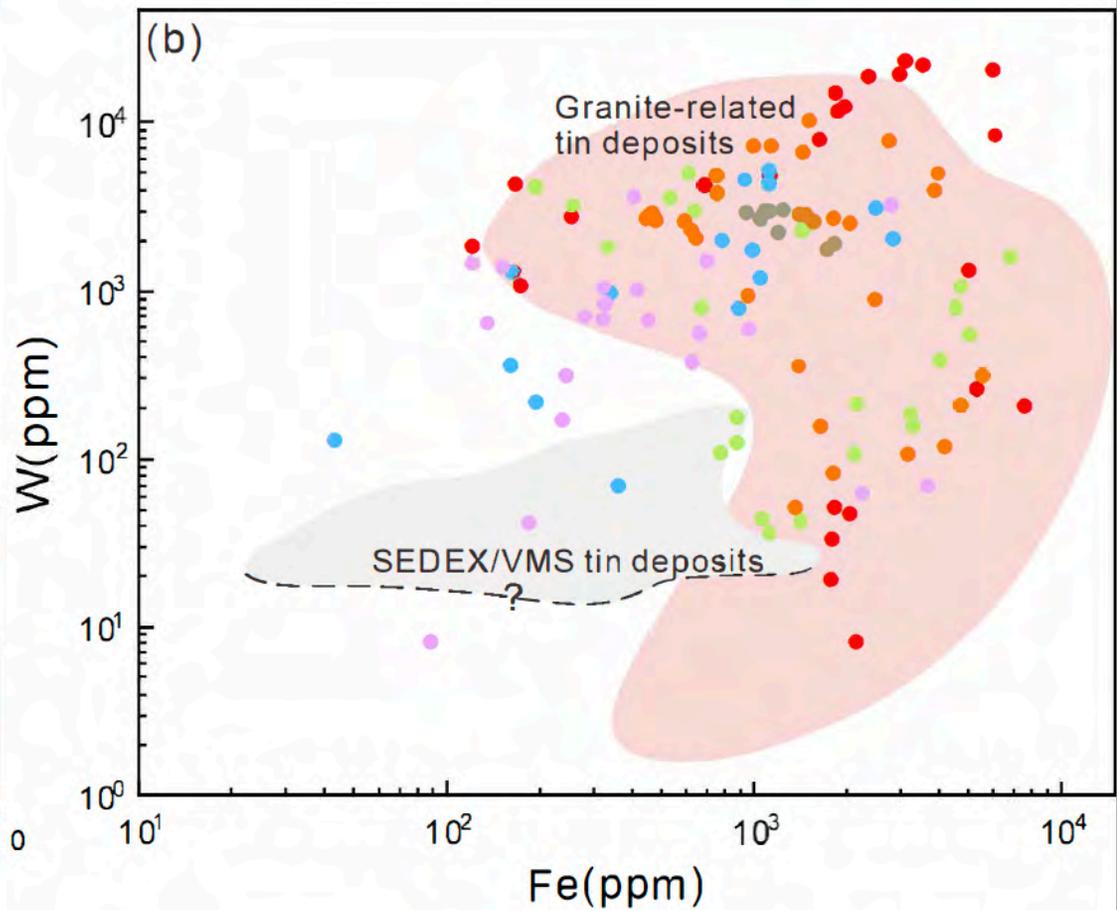
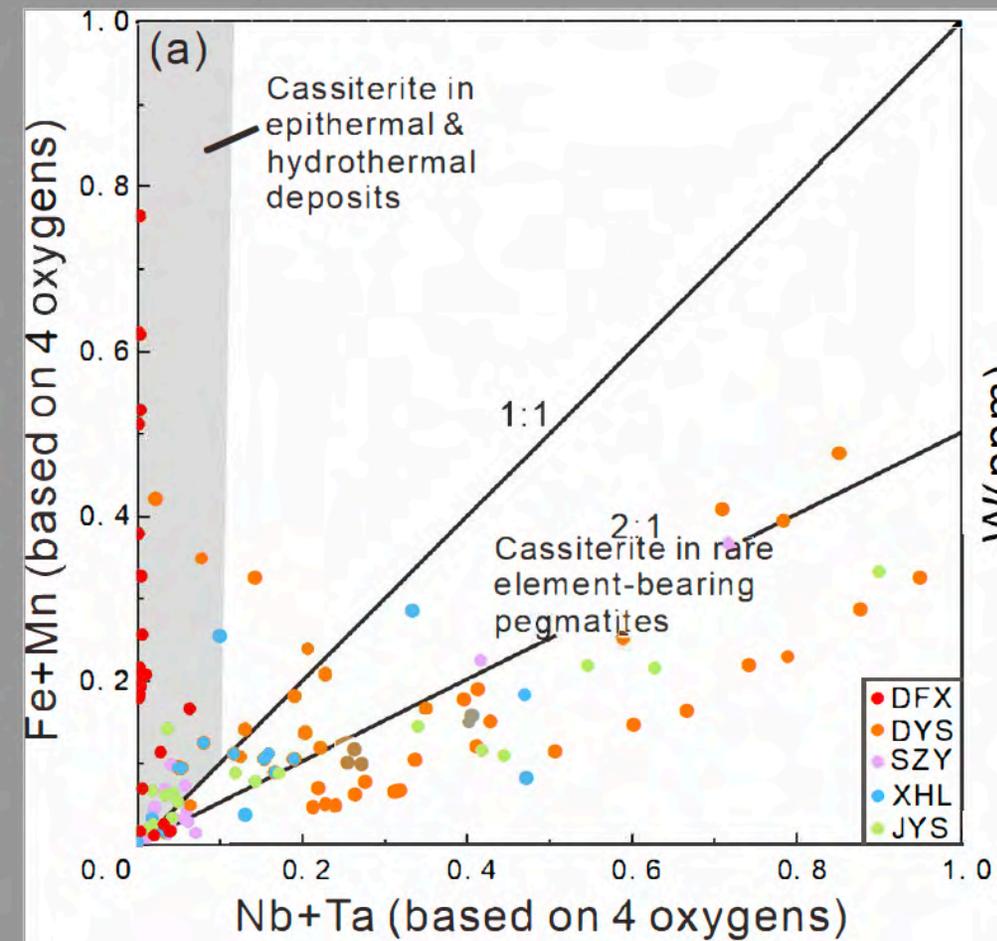


Figure 8

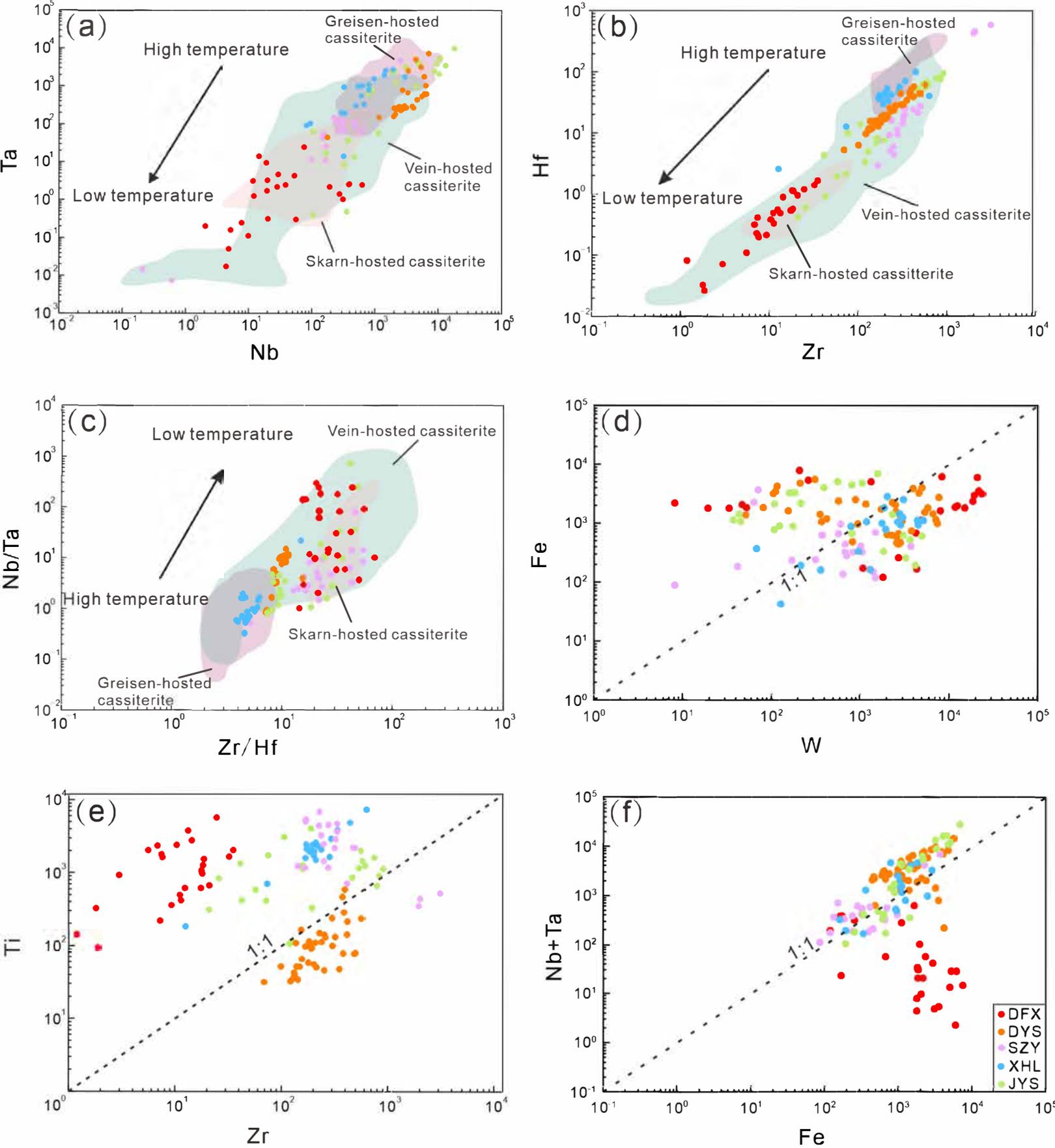


Figure 9

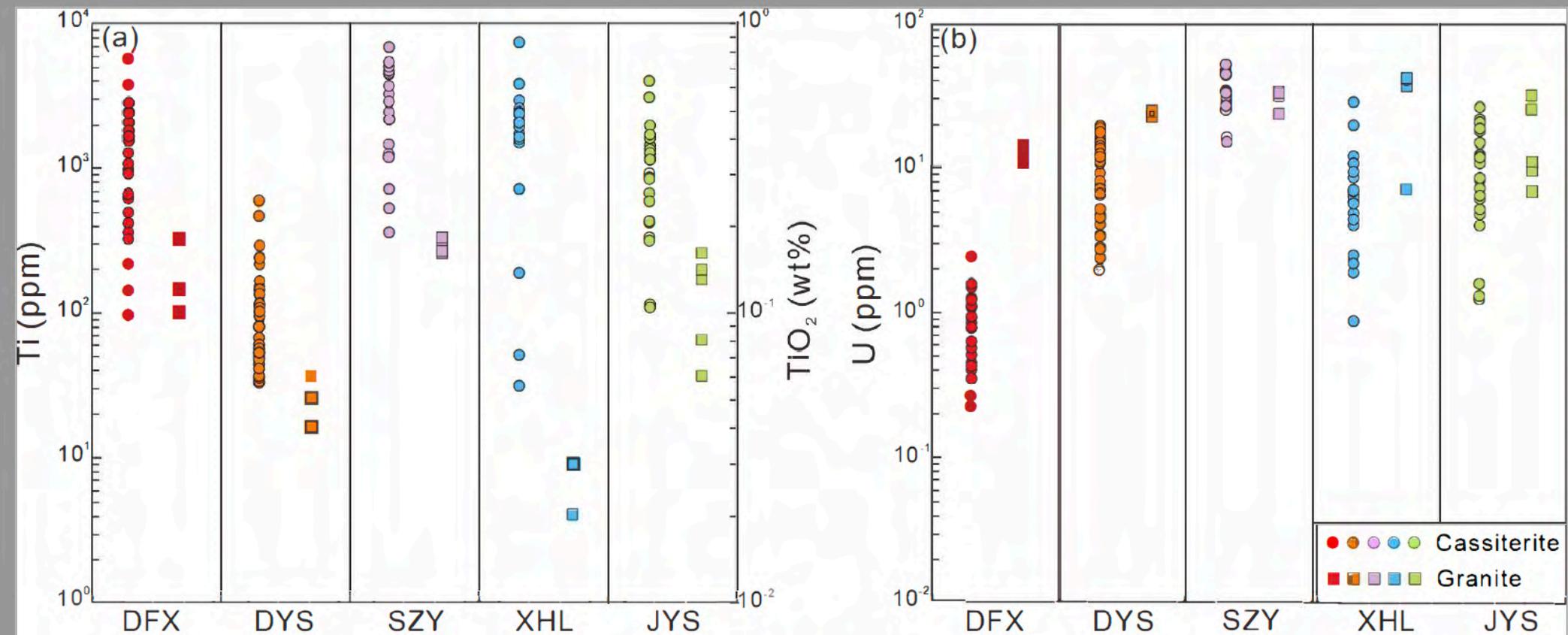


Figure 10

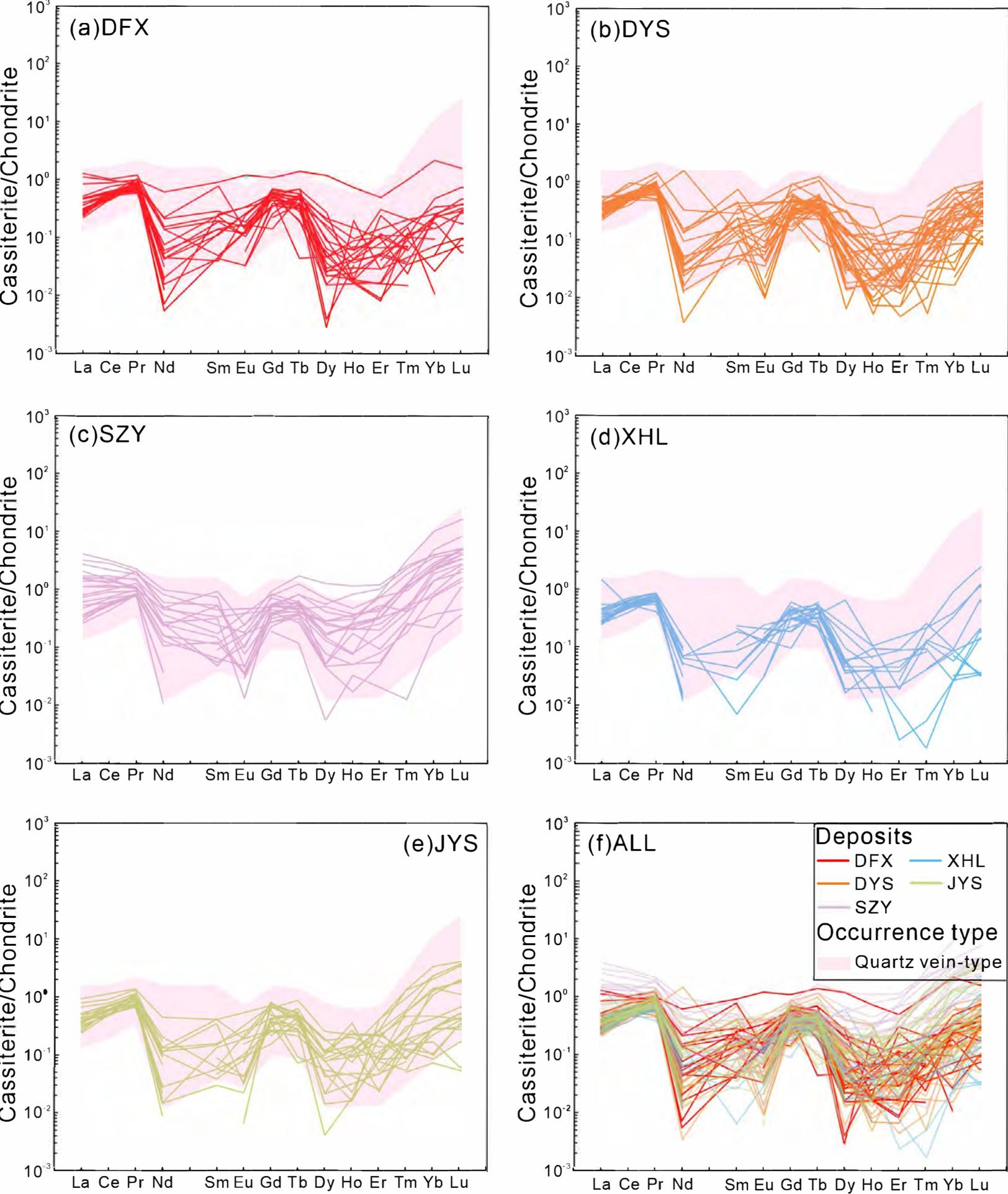
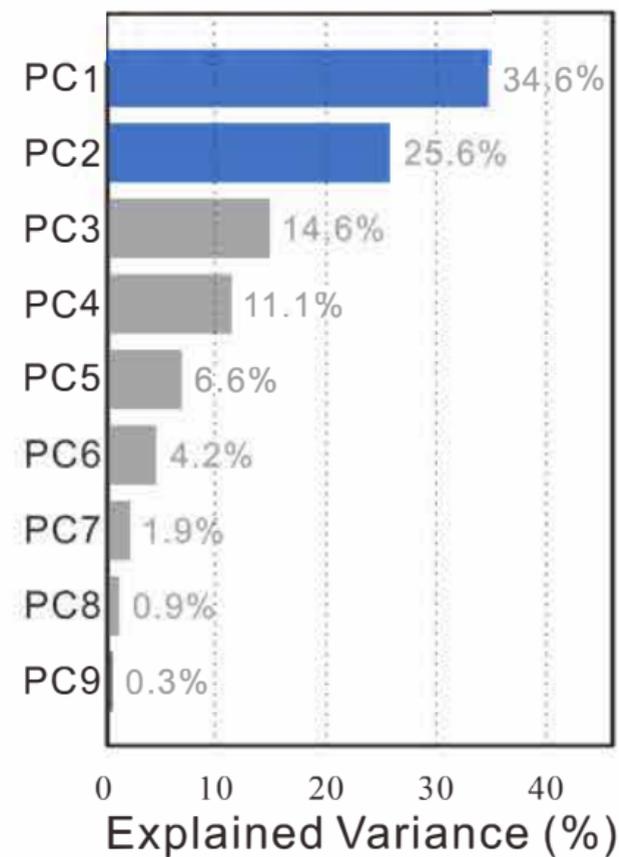
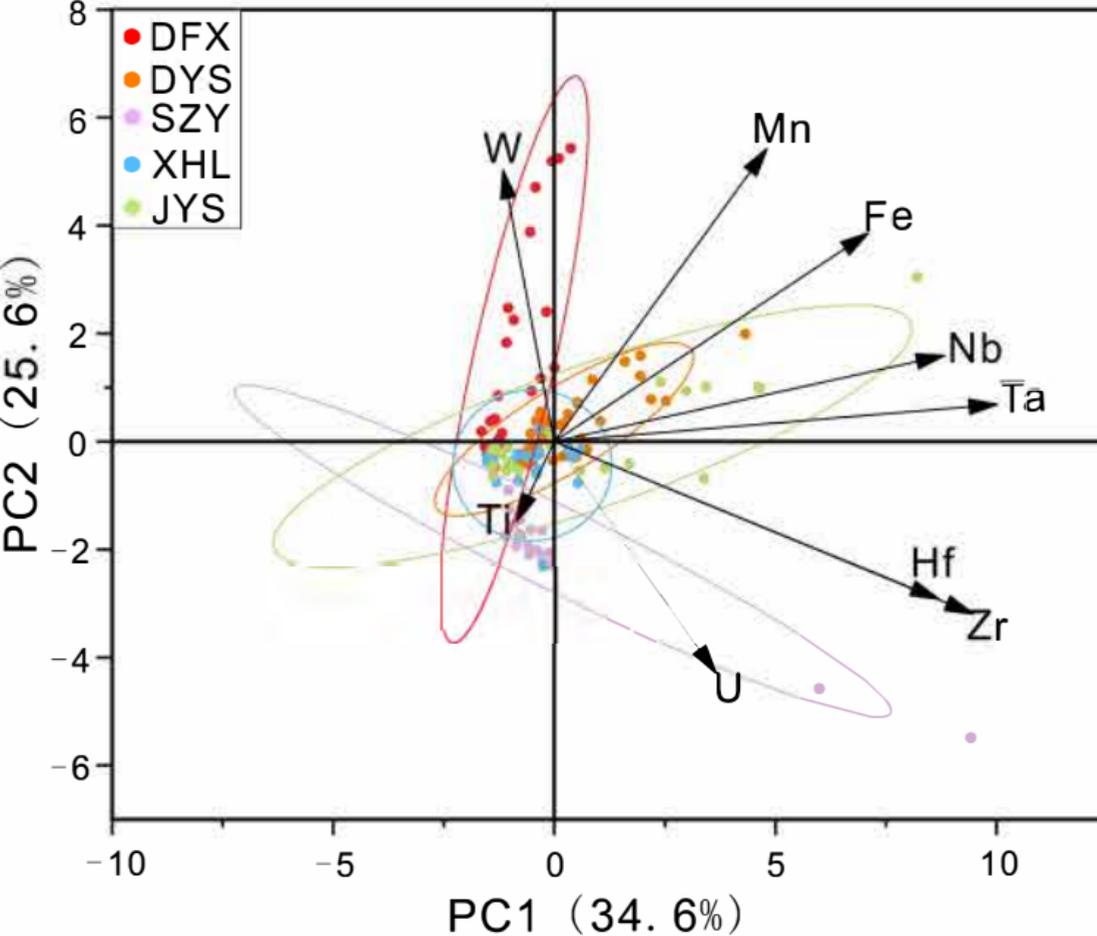


Figure 11



Loadings	PC1	PC2	PC3	PC4
LN(Ti)	-0.038	-0.137	0.603	0.589
LN(Mn)	0.227	0.518	0.342	-0.015
LN(Fe)	0.338	0.369	-0.043	0.076
LN(Zr)	0.457	-0.299	0.124	-0.306
LN(Nb)	0.425	0.166	-0.263	0.335
LN(Hf)	0.419	-0.276	0.111	-0.443
LN(Ta)	0.481	0.073	-0.170	0.285
LN(W)	-0.061	0.472	0.455	-0.385
LN(U)	0.181	-0.395	0.430	0.122

Figure 12

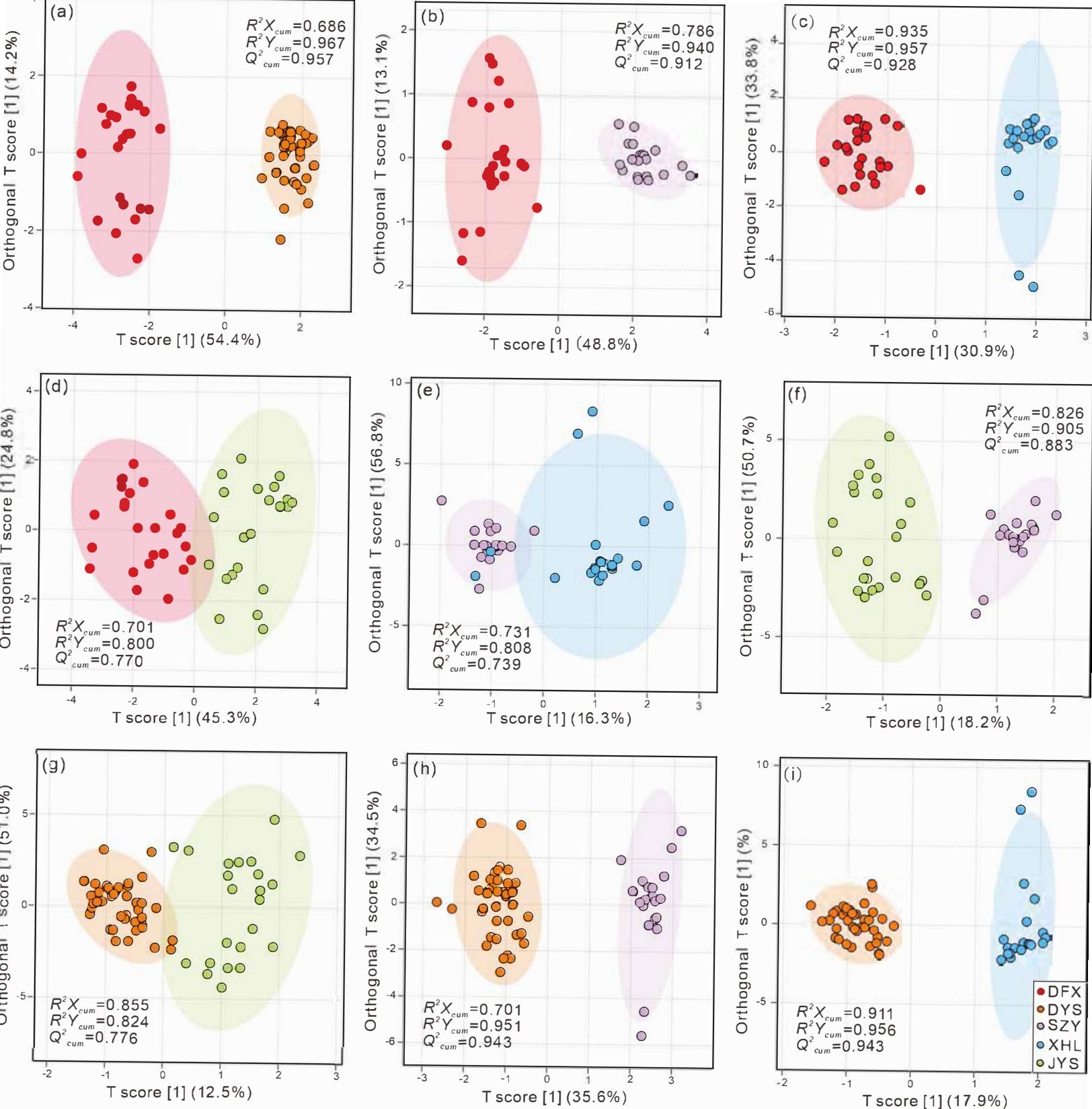


Figure 13

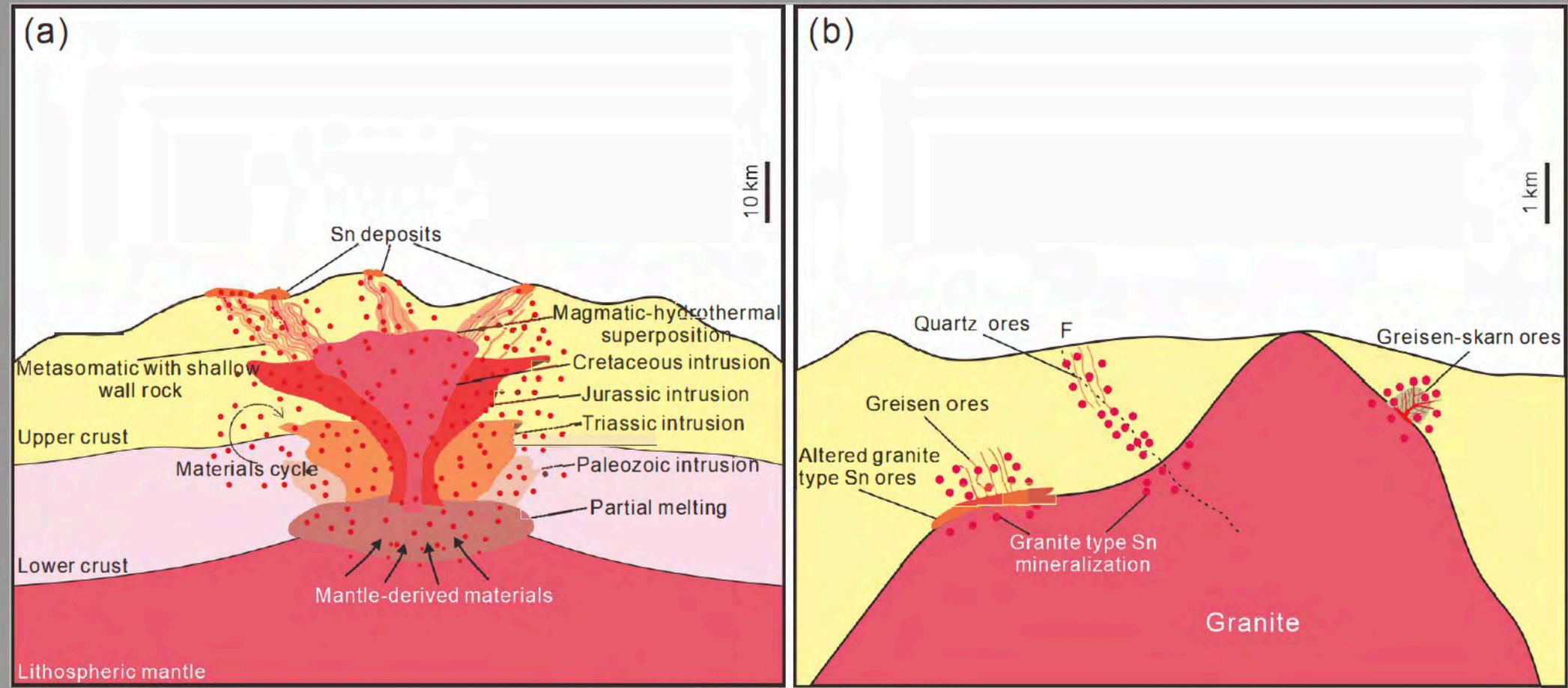


Figure 14