

Revision 1

1

2

3 **Genetic implications, composition, and structure of trioctahedral micas in** 4 **xenoliths related to Plinian eruptions from the Somma-Vesuvius volcano (Italy)**

5

6 Giuseppina Balassone^{1,2,3*}, Emanuela Schingaro⁴, Maria Lacalamita⁴, Ernesto Mesto⁴,

7 Angela Mormone², Monica Piochi², Vincenza Guarino¹, Annamaria Pellino¹,

8 Loredana D'Orazio³

9

10 ¹ Dipartimento di Scienze della Terra, dell'Ambiente e delle Risorse, Università degli Studi di Napoli

11 Federico II, via Cintia 26, 80126 Napoli, Italy

12 ² Istituto Nazionale di Geofisica e Vulcanologia, Osservatorio Vesuviano, via Diocleziano 328, I-80124

13 Napoli, Italy

14 ³ Istituto per i Polimeri Compositi e Biomateriali IPCB, CNR, via Campi Flegrei, 34, 80078 Pozzuoli,

15 Napoli, Italy

16 ⁴ Dipartimento di Scienze della Terra e Geoambientali, Università degli Studi di Bari, via Orabona 4, I-

17 70125 Bari, Italy

18

19

20 * Corresponding author: balasson@unina.it

21

22 **RUNNING TITLE:** Genetic implications and crystal-chemistry of Vesuvius micas

23

24

ABSTRACT

25 The present work is part of a systematic mineralogical and petrographic characterization of mica-

26 bearing xenoliths from Somma-Vesuvius volcano (Roman Magmatic Province, southern Italy).

27 Skarns, composite skarns-marbles and cumulates from Pompeii Plinian eruption (AD 79), and
28 skarns and syenite from Avellino eruption (3945 ± 10 cal yr BP) were investigated to define the
29 crystal chemistry of the Somma-Vesuvius trioctahedral micas and to draw inferences on
30 petrogenetic processes they were subjected. Xenoliths were characterized by means of polarized
31 optical microscopy, scanning electron microscopy (SEM-EDS), X-ray powder diffraction (XRPD)
32 and bulk-rock geochemical analyses. Mica crystals were studied through electron microprobe
33 analysis (EMPA) and single-crystal X-ray diffraction (SCXRD).

34 Micas from skarns are variably associated with $Mg_{\pm}Ca$ silicates (clinopyroxene, vesuvianite,
35 humite, clinohumite, chondrodite, forsterite, garnet), other sporadic silicates (anorthite, sodalite,
36 titanite, britholite), apatite, calcite, various types of oxides, as well as rare sulfides and halides. In
37 composite skarn-marble rocks, the mineral assemblages show some differences compared to skarns,
38 as lack of clinopyroxene and the presence of dolomite. Cumulate samples consist of mica and
39 clinopyroxene, whereas syenite is mainly composed of mica, K-feldspar, feldspathoids and
40 clinopyroxene. Together with mica, apatite occurs in all the lithotypes.

41 Trace element arrays are scattered for skarn and composite skarn-marble samples. The REE
42 patterns have a general enrichment in light (La, Ce, Pr, Nd) and medium (Sm, Eu, Gd, Tb, Dy) rare
43 earth elements, in some case with slight positive Gd anomaly. Cumulate samples generally have
44 low amounts of Ba, Sr, Zr and Th, while syenite exhibits low concentrations of trace elements,
45 except for Rb, Cs and Tl.

46 Mica crystals occurring in the studied xenoliths are phlogopite with different Al and Mg content at
47 the octahedral site, a negligible tetraferriphlogopite component and variable dehydrogenation
48 degree. All samples belong to the *1M* polytype (*C2/m* and *C2* space group) and have a wide range
49 of unit cell parameters, especially of the *c* axis [$5.3055(1) \leq a \leq 5.3218(1) \text{ \AA}$, $9.1893(1) \leq b \leq$
50 $9.2188(4) \text{ \AA}$, $10.1803(2) \leq c \leq 10.2951(2) \text{ \AA}$]. The shortest *c* cell parameter pertains to de-
51 hydrogenated phlogopite from Avellino skarn whereas OH-rich phlogopite from Pompeii composite
52 skarns-marbles has a *c* cell parameter which approximates that of the endmember phlogopite.

53 Overall, it is observed that the crystal chemistry of the micas of the present study extend the known
54 range of the other Vesuvian micas from literature. The Ti-depletion and the wide degree of
55 dehydrogenation of phlogopites from skarns and composite skarns-marbles suggest that the studied
56 samples originated under variable pressure conditions. In addition, the presence of humite in the
57 mineral assemblage seems to indicate the occurrence of devolatilization reactions. The scarce mica
58 occurrence in cumulate and mainly in syenite, instead, may depend on pressure conditions in the
59 magma storage system exceeding the mica stability.

60

61

62 **Keywords:** phlogopite, crystal chemistry, ejecta, Vesuvius.

63

64

65

INTRODUCTION

66

67 Micas are common minerals in magmatic rocks from a wide range of physico-chemical
68 conditions and their importance as petrogenetic indicators is widely accepted (e.g., Lepore et al.,
69 2017, and references therein; Guarino and Brigatti, 2018; Lacalamita et al., 2020). In the peridotite
70 mantle, micas are stabilized at the expense of olivine and pyroxene in presence of hydrous silica-
71 rich metasomatizing agents (e.g., Sekine and Wyllie, 1982) and provide important information on
72 the early fractionation history of ultrapotassic rocks (Fritschle et al., 2013). According to various
73 authors (e.g., Fritschle et al., 2013; Lepore et al., 2017 and references therein), micas of the
74 phlogopite-annite series are particularly useful to define igneous processes leading to the wide
75 compositional spectrum of ultrapotassic rocks. Therein, several crystal chemical studies exist on
76 micas from different ultrapotassic rocks such as kamafugites (e.g., Lacalamita et al., 2012;
77 Schingaro et al., 2011; Scordari et al., 2012), lamproite-like rocks (e.g., Babushkina et al., 2000;
78 Brigatti et al. 1991; Brigatti and Poppi, 1993; Cruciani and Zanazzi, 1994; Schingaro et al., 2014),

79 leucitites (e.g., Brigatti et al., 2005; Laurora et al., 2009), Mica-Rich Enclaves (MRE; e.g., Laurora
80 et al., 2007), as well as on micas from alkaline-carbonatite complexes (Guarino and Brigatti, 2018).

81 The role of phlogopite in petrogenetic processes is strictly associated to the multiple
82 concurrent ionic substitutions occurring in the crystal structure: $M^{3+,4+}$ -Tschermak (e.g., Abrecht
83 and Hewitt, 1988; Dymek, 1983; Matarrese et al., 2008; Waters and Charnley, 2002), $M^{3+,4+} \leftrightarrow$
84 vacancy (Waters and Charnley, 2002) and $M^{3+,4+}$ -oxy (e.g., Ventruti et al., 2008 and references
85 therein). However, the definition of the substitutions in the crystal structure of phlogopite requires
86 the combination of structural analyses, along with a complete determination of the chemical
87 composition including the analysis of the H₂O content and of the iron speciation, often limiting the
88 application of phlogopite as petrogenetic indicator. In the literature only two complete crystal
89 chemical studies on mica from lamproite rocks (Lepore et al., 2017) and xenolithic ejecta
90 (Balassone et al., 2013) exist. The characterization of phlogopite from three main lamproite rocks -
91 Western Alps, Tuscany, and Northern Latium (northern-central Italy) as well as Corsica (France) -
92 was undertaken by Lepore et al. (2017), with the aim of shedding light on the possible different
93 origin of micas in relation with the different lamproite-like magmas (e.g., Conticelli et al., 2009).

94 Balassone et al. (2013) proposed the first crystal-chemical and petro-volcanological
95 investigations of trioctahedral micas from the Somma-Vesuvius volcano, one of the most famous
96 worldwide volcanoes located in Italian volcanic belt (at the southern end of the Roman Magmatic
97 Province; Conticelli et al., 2011, and references therein) and associated with the high potassic
98 magmatism developed in the Central Mediterranean area. These authors, by considering different
99 xenolithic ejecta (metamorphic/metasomatic skarns, pyrometamorphic/hydrothermally altered,
100 mafic cumulates) associated to historical lavas of the most recent volcanic activity, AD 1631 –
101 1944, found that the variability of the crystal chemical features of micas was consistent with the
102 remarkable variation of their host rocks and magma storage conditions.

103 The present work aims at continuing the previous study on the Somma-Vesuvius micas by
104 focusing on xenoliths from the older and larger eruptions, i.e., the so-called Pompeii and Avellino

105 Plinian eruptions (AD 79 and 3945 ka, respectively). This research wants to improve the knowledge
106 of the crystal chemistry of the Somma-Vesuvius trioctahedral micas and to draw inferences on
107 genetic processes they were subjected to, also offering insights to understanding other similar
108 volcanic systems.

109

110 **BACKGROUND INFORMATION**

111

112 **Geological and volcanological outline**

113

114 The Italian Peninsula is unique in the world for the high number, more than fifty, of recent to
115 active volcanoes characterized by extremely different petrochemical features and eruption styles. In
116 this area, the magmatism is alkaline-potassic and developed from Oligocene to present, with the last
117 eruption of leucite-bearing magmas occurring in AD 1944 at Vesuvius (Avanzinelli et al., 2017, and
118 references therein). Vesuvius, or also Somma-Vesuvius (Fig. 1), is a stratovolcano built-up on the
119 Mesozoic carbonate sedimentary sequence of the Campanian Plain (Brocchini et al., 2001). Somma
120 is the old edifice dissected by volcanic collapses consequent the emptying of magma chamber
121 during high-magnitude Plinian eruptions (e.g., Cioni et al., 1999); Vesuvius *s.s.* is the most recent
122 cone, discontinuously growth inside the caldera until the AD 1944. The volcano produced largely
123 variable magnitude eruptions from high-VEI Plinian and sub-Plinian to low-VEI lava effusion and
124 strombolian to vulcanian events; the last ones characterized the periods of semi-persistent open-
125 conduit - known as “inter-plinian” - volcanism (Santacroce, 1987). The classical and best-studied
126 “inter-plinian” volcanism started after the AD 1631 sub-plinian event and lasted until the most
127 recent eruption in AD 1944 (Arrighi et al., 2001).

128 Based on the rock-geochemistry of the erupted products (Ayuso et al., 1998; Piochi et al.,
129 2006a,b; Di Renzo et al., 2007; Santacroce et al., 2008), different volcanic periods were established:
130 the oldest one (>8-9 ka) produced slightly undersaturated rocks from K-basalt to K-trachyte, the

131 period between 8-9 ka and the AD 79 involved K-phonolite-tephrite to K-phonolite magma
132 compositions, and the period younger than AD 79 characterized for highly undersaturated magmas,
133 from leucititic tephrite to leucititic phonolite. Overall, a wide range of Sr and Nd isotope
134 compositions have been documented, from 0.7066 to 0.7081 and 0.5126 to 0.5124 respectively
135 (Ayuso et al., 1998; Piochi et al., 2006a, b; Di Renzo et al., 2007). The compositions mostly reflect
136 evolutionary processes that occurred during upward migration to shallow crustal levels in primary
137 magmas bearing contributions from different mantle reservoirs and slab-derived components. Some
138 authors ascribed the variations of Sr-Nd-Pb-O isotopic compositions in central Italian magmas
139 (including Somma-Vesuvius) to crustal assimilation of carbonates during storage in the upper few
140 kilometres of the crust (e.g., Piochi et al., 2006a; Di Renzo et al., 2007; Iacono-Marziano et al.,
141 2008, 2009; Scaillet et al., 2008; Dallai et al., 2011; Pichavant et al., 2014); other researchers favor
142 isotopic variability of associated mantle sources (e.g., Peccerillo, 2005; Moretti et al., 2013).
143 Refilling and mingling processes between different magmas within each volcanic period or eruption
144 event are evident from isotope variations (Cioni et al., 1999; Piochi et al., 2004; Pappalardo et al.,
145 2004 and references therein). Recent studies involved the petrological and geochemical evolution of
146 the chemically zoned juvenile samples come from Plinian and sub-Plinian eruptions (Melluso et al.,
147 2022). These authors point to the existence of a similar mineralogical paragenesis in the different
148 eruptions, but these minerals have different composition, suggesting the existence of independent,
149 zoned magma batches throughout the activity of the stratovolcano, which possibly started to
150 crystallize at similar depths.

151 However, the processes also involved carbonate contamination and entrapment of crystal
152 mush generated during previous magma storage in the crust (Jolis et al., 2013; Piochi et al., 2006a).

153 The “Avellino” (3945 ± 10 cal yr BP; Di Vito et al., 2019, and references therein) and
154 “Pompeii” (AD 79) are probably the most famous and widespread plinian eruptions of Somma-
155 Vesuvius, with the white and grey pumice deposits associated with sustained pyroclastic column
156 and variable collapsing styles (Gurioli et al., 2005; Sulpizio et al., 2010; Balcone-Boissard et al.,

157 2012 and references therein; Massaro et al., 2018; Doronzo et al., 2022). White pumices were
158 emitted prior to grey pumices, both phonolitic, but with the white type having a higher
159 differentiation degree (Cioni et al., 2008; Balcone-Boissard et al., 2012). According to the above
160 authors, the volume of fallout tephra is 1.5 km³ for Avellino and 3.3 km³ for Pompei. The Avellino
161 products contain sanidine and minor nepheline, clinopyroxene, biotite, amphibole, and oxides;
162 minerals of marialite-meionite (“scapolite”) have also found. The Pompeii tephra contains sanidine,
163 pyroxene, phlogopite, K-ferripargasitic amphibole, melanitic garnet and leucite in the groundmass.
164 Similarly to “Pomici di Base” (18.3 ka bp) and “Mercato” (8 ka bp), they determined caldera
165 collapses and dissection of the limestone basement (Cioni et al., 1999) in which their feeding
166 magma chambers were stored and evolved (Cioni et al., 1995; Cioni, 2000; Piochi et al., 2006a, b;
167 Fulignati et al., 2005; Signorelli et al., 1999; Balcone-Boissard et al., 2012; Vona et al., 2020).

168

169 **The Somma-Vesuvius ejecta**

170

171 A large variety of xenolith types is present in the Somma-Vesuvius eruptive products,
172 including weakly-to-highly metamorphosed carbonate, skarns, plutonic rocks (cumulates, alkali
173 syenites), lavas/pyroclasts (Jolis et al., 2015 and references therein), as well as hornfelses (Del
174 Moro et al., 2004) and sedimentary rocks (Barberi et al., 1981; Joron et al., 1987). Composite
175 xenoliths are also frequently found at Somma-Vesuvius, having sharp transitions from skarns to
176 marbles, hornfelses and igneous rocks; they document the close spatial association of these rock
177 types at depth and provide unequivocal evidence for magma–carbonate interaction during magma
178 differentiation (e.g., Gilg et al., 2001; Piochi et al., 2006a, b; Jolis et al., 2015).

179 The marbles derive from thermally affected limestone or dolostone of the Mesozoic basement
180 that has undergone recrystallization and, in part, devolatilization, but overall reflecting relatively
181 modest interaction with metasomatic fluids or silicate melts (Gilg et al., 2001; Jolis et al., 2015).
182 Carbonate clasts occur in most depositional sequences of the Somma-Vesuvius, in agreement with

183 entrapment pressure of inclusions of 3.5-5 km and deeper than 7 km (e.g., Marianelli et al., 1999;
184 Cioni, 2000; Fulignati et al., 2005) and Sr-isotope variation through magma storage depths (Piochi
185 et al., 2006a). Based on several authors (e.g., Dallai et al., 2011; Jolis et al., 2013; Piochi et al.,
186 2006a), Mesozoic carbonate rocks are the contaminant of magma in the roots of the volcanic
187 structure, mostly at shallow depth. Volcanic rock geochemistry registered carbonate assimilation
188 and is useful tool to study the consequent influence on the magma evolution, its volatile content and
189 degassing modes, the eruption dynamics. However, limestone xenoliths allow detecting the possible
190 physico-chemical changes produced by the magma and carbonate host interaction with implications
191 on rock mechanics, hydrothermal circulation and mineral/metal ore developing. Some authors
192 focused on mineralization and isotope deviance induced by thermo-metamorphism in carbonate
193 xenoliths (e.g., Jolis et al., 2015 and references therein). From petrographic and mineralogical point
194 of view, marble xenoliths mainly consist of calcite with a typical saccharoidal texture. Dolomite is
195 also detected as relic crystals in metadolostone ejecta, characterized by calcite and periclase
196 assemblage; brucite and hydromagnesite commonly occur as secondary retrograde rims around
197 periclase (Gilg et al., 2001). Accessory minerals include diopside, phlogopite, forsterite,
198 chondrodite, humite, clinohumite, norbergite, spinel, graphite, illite, and albite (Barberi and Leoni,
199 1980).

200 Hornfelses, rarely occurring at Somma-Vesuvius, consist of a fine-grained homogeneous
201 granoblastic calc-silicate assemblage of diopside, quartz, feldspars, and wollastonite (Barberi and
202 Leoni, 1980). Marbles and hornfelses are interpreted as thermally metamorphosed carbonate
203 (limestone or dolostone) or marly rocks that have suffered recrystallization and in part
204 devolatilization with modest metasomatic interaction with aqueous fluids or silicate melts (Gilg et
205 al., 2001; Jolis et al., 2015).

206 Skarns have calc-silicate compositions typically developing in contact-metamorphic aureoles
207 in response to interaction between limestone or dolostone with a silicate magma. The formation of
208 these rocks is known to progressively modify the composition of both the magma and the carbonate

209 wall-rock. Skarn ejecta are common at Vesuvius and display a wide range of mineral compositions
210 and textures also with zoning and metasomatic fronts (also few millimetres in thickness). Typical
211 skarn Ca- and/or Mg-silicate crystal phases are clinopyroxene, phlogopite, wollastonite, olivine,
212 clinohumite, vesuvianite, garnet, gehlenite, and anorthite; other skarn minerals can be calcite,
213 dolomite, meionite, leucite, fluorapatite, titanite, cuspidine, perovskite, baddeleyite, geikelinte,
214 zirconolite, calzirtite, etc. (Gilg et al., 2001; Pascal et al., 2011; Jolis et al., 2015 and references
215 therein). Different varieties of spinels have been also described, i.e., spinel, magnetite,
216 magnesioferrite and qandilite, the last two phases found in metasomatic forsterite-spinel-calcite
217 skarns (Pascal et al., 2011). Many studies characterized pressure and temperature of skarn
218 formation, the fluid nature and the isotope evolution related to the limestone/dolostone-skarn
219 transition, as reported by Jolis et al. (2015 and references therein). Skarn xenoliths in the Vesuvius
220 eruptive deposits can be linked to relatively shallow magma reservoirs (≤ 4 –10 km; Auger et al.,
221 2001; Civetta et al., 2004; Scaillet et al., 2008 and references therein) just developed within the
222 thick limestone and dolostone carbonate sequence which extend from approximately ≥ 2 km to at
223 least 8 km depth under the volcano (Zollo et al., 1996; Bruno et al., 1998; Auger et al., 2001;
224 Brocchini et al., 2001).

225 The igneous ejecta are volcanic to plutonic (or hypabyssal) rocks; the last one consists of
226 pyroxenites and alkali syenites. Pyroxenites are characterized by high proportions of clinopyroxene
227 with minor phlogopite and are likely of cumulate origin. The alkali syenites are fine to medium
228 grained and are dominated by subhedral to anhedral sanidine, garnet, and subordinate amphibole,
229 sodalite group minerals (Balassone et al., 2012, 2016), nepheline, and leucite etc. They have been
230 interpreted as part of a crystallizing magmatic mush.

231 As already stated, the mica-bearing samples considered in this study come from Avellino and
232 Pompeii deposits (Table 1); these eruptions have brought a large amount of ejecta, especially of
233 skarn type (Gilg et al., 2001; Jolis et al., 2015), to the surface (Joron et al., 1987). The mica
234 varieties commonly observed at Somma-Vesuvius are phlogopite and “biotite” (Russo and Punzo,

235 2004). As reported by Gilg et al. (2001) for micas from skarn and composite ejecta, Mg-rich and Ti-
236 poor phlogopite, typically occur in zoned skarn ejecta associated to diopsidic/hedenbergitic to (less
237 common) “fassaitic” clinopyroxene, F-bearing vesuvianite, wollastonite, gehlenite, meionite,
238 forsterite, clinohumite, anorthite, and Mg-poor calcite, with many accessory minerals (i.e., apatite,
239 spinel, magnetite, perovskite, baddeleyite, and various undefined REE-, U-, Th-, Zr-, and Ti-rich
240 minerals).

241

242

243 **SAMPLES AND EXPERIMENTAL METHODS**

244

245 The analyzed micas occur in xenoliths collected from outcrops close to Somma Vesuviana,
246 Ercolano and Terzigno towns (Fig. 1), i.e., in skarns, cumulates, syenites, and composite (marble-
247 skarn) lithotypes (see Table 1 and Fig. 2a). After preliminary mesoscopic observations, a
248 combination of polarized optical microscopy, X-ray powder diffraction (XRPD) and scanning
249 electron microscopy equipped with an energy-dispersive spectrometer (SEM-EDS) was used to
250 characterize the mica-bearing xenoliths from the textural, mineralogical, and petrographic point of
251 view.

252 Bulk rock geochemistry was carried out at two laboratories, i.e. Activation Laboratories LTD
253 (Ancaster, Canada) and Bureau Laboratories Ltd. (Vancouver, Canada). At Activation Laboratories
254 (sample set PS40, PS41, PS44, PS50, PS54, PS55, PS57, PS58, PS60, PS80, PS81, PS82, PS84,
255 PSK-R), major elements were analyzed via ICP-OES (inductively coupled plasma-optical emission
256 spectrometry) using a Thermo Jarrell-Ash ENVIRO II ICP. Trace elements were determined by
257 ICP-MS with a Perkin Elmer SCIEX ELAN 6000. Uncertainty is less than 3% for major oxides,
258 less than 15% for Co, Y, Zr, and Tb, and less than 5% for all other trace elements (see
259 www.actlabs.com). At Bureau Laboratories (sample set PS62, PS65, PS68, PS70, PS73, PS77),
260 major elements were analyzed by X-ray fluorescence (XRF) and inductively coupled plasma

261 emission spectrometry (ICP-ES), using $\text{LiBO}_2/\text{Li}_2\text{B}_4\text{O}_7$ fusion. Minor and trace elements were
262 determined by inductively coupled plasma-mass spectrometry (ICP-MS), using a four-acid (HNO_3 -
263 HClO_4 - HF - HCl) digestion. Loss on ignition (LOI) was calculated by weight loss after ignition at
264 1000°C . The uncertainty is less than 3% for major/minor oxides, less than 5-10% for trace elements
265 (see <http://acmelab.com>). In both cases, loss on ignition (LOI) was calculated by weight loss after
266 ignition at 1000°C .

267 XRPD was conducted on powdered rocks using a Seifert-GE diffractometer ID 3003
268 (DiSTAR, University of Naples, Italy). Intensity profiles were collected in the 2θ range of 3 - 80°
269 using Ni-filtered $\text{CuK}\alpha$ radiation at 40 kV and 30 mA, with a step size 0.02° , at a scanning time of
270 10 s/step. The diffraction patterns were processed using the RayfleX software package. A minor set
271 of measurements were performed by using a X'Pert Powder diffractometer by PANalytical (Istituto
272 Nazionale di Geofisica e Vulcanologia, Osservatorio Vesuviano, Napoli, Italy), equipped with a
273 high speed PIXcel detector, Ni-filtered, $\text{CuK}\alpha$ radiation (40 kV and 40 mA in the 3 - $70^\circ 2\theta$ range,
274 0.02° steps at 8 s/step) and a PANalytical B.V. software HIGHScore Plus version 4.9 for the
275 elaboration of the diffraction profiles (Mormone et al., 2014).

276 Electron dispersive spectroscopy (EDS) microanalyses and backscattered electron images
277 (BSE) of the whole mineral assemblages in thin sections were obtained with an INCA X-stream
278 pulse processor and the 4.08 version Inca software (Oxford Instruments detector), interfaced with
279 the JEOL JSM 5310 (DiSTAR, Naples, Italy). Observations were made in backscattered electron
280 mode (BSE), at accelerating voltage 15 kV, 50–100 μA filament current and variable spot size and
281 a working distance of 20 mm. The following reference standards were used: albite (Si, Al, Na),
282 orthoclase (K), wollastonite (Ca), diopside (Mg), almandine (Fe), rutile (Ti), barite (Ba),
283 strontianite (Sr), metallic chromium (Cr), rhodonite (Mn), pyrite (S), sphalerite (Zn), galena (Pb),
284 fluorite (F), apatite (P), sylvite (Cl), Smithsonian phosphates (La, Ce, Nd, Sm, Y), gallium arsenide
285 (As), and metallic vanadium (V). Analytical errors are 1% relative for major elements and 3%
286 relative for minor elements.

287 In order to determine the chemical composition by electron microprobe analysis (EPMA) of
288 mica crystals, they were selected under the binocular microscope from the grains of crushed rocks,
289 and then prepared as polished opaque mounts (by means of Buehler EpoKwich™ epoxy resin, as
290 well as Buehler polishing cloths and diamond suspensions). Chemical analyses were carried out
291 using a Cameca SX50 WDS instrument (CNR-IGAG, Rome) under the following experimental
292 conditions: 15 nA sample current, 10 kV accelerating voltage, 10 μm beam diameter; no Na loss
293 was detected during the analyses. From each sample, we selected three to five mica grains, and
294 performed five to eight microanalyses, reporting the averages in Table 3. The standards were:
295 wollastonite (Si, Ca), titanite (Ti), corundum (Al), magnetite (Fe), periclase (Mg), jadeite (Na), and
296 orthoclase (K). Data were corrected using the PAP program (Pouchou and Pichoir, 1991). Relative
297 uncertainty on the analytical measurements was 1% for major elements, 4% for minor elements,
298 and 10% for F.

299 Scanning electron microscopy (SEM) examination of morphological features of selected mica
300 crystals were carried out by means of both the JEOL JSM 5310 instrument, and a small subset of
301 samples was also observed bFEI Quanta 200 SEM (IPCB-CNR Pozzuoli, Italy).

302 Single crystal X-ray diffraction (SCXRD) data were collected on mica crystals from selected
303 rocks (PS68, PS77, PS50, PS62, PS73) both from Avellino and Pompeii eruptions by means of a
304 Bruker AXS X8 APEXII automated diffractometer (Dipartimento di Scienze della Terra e
305 Geoambientali, University of Bari, Italy), equipped with a CCD detector and graphite-
306 monochromatized Mo K α radiation ($\lambda = 0.71073 \text{ \AA}$) operating at 50 kV, 30 mA and 40 mm crystal-
307 to-detector distance. The whole Ewald sphere ($\pm h, \pm k, \pm l$) was recorded up to $\theta \sim 45^\circ$ with a scan
308 width $1.0^\circ/\text{frame}$ and an exposure time 20 s/frame . The APEX program suite allowed to optimize
309 the collection strategy (Bruker, 2010); the SAINT package (Bruker, 2007) was used for the
310 integration of the intensities of reflections and the correction of the Lorentz and polarization effects;
311 the SADABS software (Bruker, 2009) was employed for the empirical absorption correction;

312 XPREP (Sheldrick, 2008) was used for the subsequent analysis of the intensity data and the
313 assignment of the space group.

314 In the case of PS50 sample, CELL_NOW program (Sheldrick, 2003a) identified two twinned
315 individuals rotated of 180° around a twin axis parallel to [310]; the TWINABS software provided
316 for a semi-empirical absorption correction (Sheldrick, 2003b).

317 Structure refinements were carried out in $C2/m$ space group by means of the program
318 CRYSTALS (Betteridge et al., 2003) considering the reflections with $I > 3\sigma(I)$. Overall scale factor,
319 atomic positions, cation occupancies and anisotropic atomic displacement parameters were refined,
320 starting from the coordinates of micas belonging to the phlogopite-fluorophlogopite series in
321 Lacalamita et al. (2020). The occupancy of the tetrahedral site was constrained to 1; the Mg vs. Fe
322 occupancies in the octahedral site were varied with full occupancy constraint for all samples except
323 for PS50 where only Mg was refined both in $M1$ and $M2$; a restraint (1.00 ± 0.08) on the occupancy
324 of the K atoms at the interlayer site was used to allow the occupancy of this site to assume values
325 higher or lower than 1. The analysis of the difference-Fourier maps for all samples except for PS77
326 evidenced the presence of a residual electron density peak $\sim 1 \text{ e}^-/\text{\AA}^3$ which may be considered as
327 the position of the hydrogen atom, at $\sim 0.9 \text{ \AA}$ from the O4 oxygen. It was included in the refinement
328 using the strategy of refinement described elsewhere (Lacalamita et al., 2020).

329 For PS77 sample, at the end of the refinement in $C2/m$ space group, an R -value of 4% was
330 obtained; analysis of the difference Fourier synthesis showed two peaks ~ 3.5 and $1.5 \text{ e}^-/\text{\AA}^3$ at 0.57,
331 0, 0.22 and 0, 0.16, 0, respectively. These peaks are hints of $\pm b/3$ shifts of the octahedral sheet,
332 which leads to desymmetrization of the whole structure (Schingaro et al., 2001). Therefore, a
333 reduction of symmetry from $C2/m$ to $C2$ was attempted considering extra (disordered) atoms with
334 respect to the ordered part of the structure. The latter was refined using proper constraints to impose
335 the $C2/m$ symmetry. The final refinement converged to R 2.9% and no significant residual electron
336 density was observed. The same disorder affected the PS73 although at minor extent (residual peaks
337 $\leq 1.3 \text{ e}^-/\text{\AA}^3$). Since the electron density of the disorder atoms was $\leq 0.2 \text{ e}^-$, the disorder modelling

338 was not included in the final refinement of this sample. Crystallographic interchange format (CIF)
339 files are also provided as supplementary materials.

340

341

342 **RESULTS AND DISCUSSION**

343

344 **The xenoliths**

345

346 *Petrographic description*

347 The Pompeii xenolithic samples (# 12 samples) are represented by three lithologies, i.e., skarns,
348 composite skarns-marbles and cumulates, whereas the Avellino samples (# 8 samples) mostly
349 consist of skarns and one syenite (Fig. 2b). SEM-BSE micrographs of selected samples in thin
350 sections are shown in Fig. 3.

351 The skarn samples are characterized by a texture variable from granoblastic with equigranular
352 crystals of phlogopite, clinopyroxene and calcite often randomly oriented (decussate texture), to
353 poikiloblastic with small inclusions of various minerals mainly in larger megacrysts (Fig. 3a-h).
354 These samples are mostly composed of Ca±Mg silicates, among which the most common are
355 clinopyroxene, vesuvianite, humite, clinohumite, sodalite, forsteritic olivine, leucite, and garnet,
356 detected in variable amounts (Table 1), locally accompanied by calcite and a wide variety of other
357 µm-sized mineral phases, difficult to identify with just chemical analyses and optical observations.
358 An association with phlogopite-clinopyroxene±spinel±calcite±apatite - recognized in many skarns
359 of Somma-Vesuvius - was observed both in Pompeii and Avellino samples (Table 1 and Fig. 3a-h).
360 Sodalite and anorthite occur in two Avellino skarns (Table 1 and Fig. 3b). Humite occurs in
361 Avellino (PS55, PS77, PSK-R) and Pompeii (PS70, PS73) skarns, chondrodite in Avellino skarn
362 PSK-R and clinohumite in the Avellino skarn PS41 (i.e., Fig. 3c,f). Titanite was found only in
363 Avellino PS84 skarn, where rare mixed Ca-Ti and Zr-Ti oxides (as perovskite, pyrochlore

364 supergroup minerals, and zirconolite) also occur (Fig. 3g). Magnetite is present in Pompeii skarn
365 PS58, PS68, and PS70 (Table 1). Baddeleyite (Fig. 3a,d,h) is quite common and observed in five
366 skarn samples in tiny and disseminated crystals (Avellino: PS41, PS60, PS82, and PSK-R; Pompeii:
367 PS84); thorianite also occurs in Pompeii skarn PS58 (Fig. 3c) in tiny individuals. Carbonates are
368 represented by rare and small calcite crystals, observed in five samples both from Pompeii and
369 Avellino eruptions (Table 1). Tiny pyrite inclusions were observed in Avellino skarns (PS60 and
370 PSK-R) (Fig. 3d). Other trace minerals are cuspidine, as well as likely bismoclite (Table 1).

371 In composite Pompeii skarns-marbles (Table 1), the contacts between the two lithologies can
372 be variously sharp, with the carbonate components relatively more (PS44, PS54, PS62, PS65) or
373 less (PS40, PS50) prevailing over skarn-related assemblages (i.e., sample PS44 in Fig. 2). Calcite
374 displays a granoblastic saccharoidal texture, occasionally very friable, and often with quite fine (<
375 500 μm) crystals (Fig. 3i,l,m); dolomite is recorded in samples PS44, PS50 and PS54 (Fig. 3k-m).
376 A poikiloblastic texture is also observed (partly as in Fig. 3l). The silicate minerals include, together
377 with mica, humite, clinohumite, chondrodite, forsteritic olivine, vesuvianite, titanite and meionite
378 (Fig. 3i,j,l,m); Accessory spinel, magnetite, apatite, galena, pyrite, thorianite, baddeleyite, and
379 stolzite are locally detected (Table 1 and Fig. 3i). An undetermined Si-Zn-Fe-Pb phase was
380 observed in sample PS44 (Fig. 3j); this could correspond to a mixture of a silicate-rich phases
381 alternating with sulfide-rich phases forming a colloform texture, but no reliable chemical analyses
382 were carried out.

383 Syenite PS57, with holocrystalline texture, is composed of K-feldspar, feldspathoids (leucite,
384 nepheline, and sodalite), phlogopite, clinopyroxene, with small amounts of apatite and britholite
385 (Fig. 3n).

386 The two cumulate samples (Table 1) are basically formed by clinopyroxene and phlogopite
387 (Fig. 3o, p); small inclusions and/or vuggy linings of apatite, sodalite, baddeleyite, bismoclite, and
388 vesuvianite are locally found, together with trace amounts of late fluorite, mimetite, and sylvite
389 (Table 1; Fig. 3,o,p).

390

391 ***Mineral chemistry***

392 The main mineral compositions detected in the investigated mica-bearing lithics are
393 illustrated in Tables S1-S4. Clinopyroxenes, recorded in skarns, syenite and cumulates of both
394 Pompeii and Avellino ejecta (Table S1), have zoned diopside/hedenbergite compositions (i.e., Fig.
395 4b, h), or also subsilicic ferroan aluminian diopside (“fassaitic”) with significant Ca-Al-Tschermak
396 exchange component (i.e., skarn PS60 and cumulate PS80; table S1). As reported by Gilg et al.
397 (2001), the Vesuvian “fassaitic” skarn pyroxenes generally have higher Ti and lower Mn contents
398 than Al-poor skarn diopsides.

399 Olivine was detected in skarns PS58, PS73, and PSK-R and in composite rock PS62 (Table
400 1); it has a forsteritic composition ($Fo_{88.99}$) comparable to forsterites from metasomatic cumulates
401 and skarns, but Mg-richer than magmatic olivines (Fo_{67-87}) (Gilg et al., 2001, and references therein;
402 Fulignati et al., 2005). Among various silicates detected in the analyzed xenoliths, Table S1 also has
403 representative chemical analyses of anorthite, feldspathoids (sodalite, leucite, and nepheline),
404 garnets (grandite), and humite. Anorthite, nepheline, spinel, and garnets have been already reported
405 in skarns from AD 79 (Fulignati et al., 2005). However, apart from sodalite and humite, in our
406 sample set, these silicates mostly occur in the Avellino ejecta (Table 1).

407 F-rich apatite, with low Cl amounts, was furthermore detected in ejecta from both eruptions
408 (samples PS58, PS60, PS80, and PS84), whereas mimetite was found only in sample PS81 as last
409 phase to crystallize (Table S2). In particular, the humite and clinohumite in carbonate-rich rocks is
410 known in the literature (e.g., Rice, 1980; Young and Morrison, 1992), although in our samples not
411 often associated with olivine and pyroxene.

412 Accessory baddeleyite, zirconolite, and pyrochlore supergroup minerals, testifying variable
413 leaching/mobilization processes of HFSE elements (such as Ti, Zr, Nb) during metasomatism which
414 concentrated these elements in the skarn as oxides (Pascal et al., 2009), are presented in Table S3,

415 together with magnetite, spinel, thorianite, and perovskite; sulfides are also sporadically found and
416 represented by galena and pyrite (Tables 1 and S3).

417 Calcite can be found both in skarns and skarns-marbles (Tables 1 and S4); calcite exhibits the
418 following oxide content: 0.09-3.43 wt% of MgO, 0-1.20 wt% FeO, 0-5.58 wt% MnO, and up to
419 1.13 wt% SrO and 1.21 wt% BaO. Dolomite, occurring only in three composite samples, has up to
420 1.40 wt% FeO, 0.04 wt% MnO and 0.18 wt% BaO.

421 The mineralogical assemblage of the Pompeii and Avellino skarn and composite xenoliths
422 investigated in this study shows some differences from that reported in Balassone et al. (2013).
423 Indeed, the xenoliths investigated here do not contain britholite and glass, instead detected in the
424 Ca-Mg-silicate xenoliths from the AD 1631 and AD 1944 eruptions from Balassone et al. (2013);
425 also cuspidine was found only in trace amounts in the studied samples.

426

427 *Xenoliths geochemistry*

428 The chemical composition of the studied xenoliths has a wide compositional range in terms of
429 major elements contents, as SiO₂, CaO, and MgO (Fig. 4; Table 2), closely related to their complex
430 mineralogical assemblages. Skarns and skarn-marble xenoliths from the Pompeii eruption are
431 characterized by a broad compositional spectrum, with SiO₂, CaO, and MgO in the ranges 6.4-48.7
432 wt%, 4.2-38.1 wt% and 3.7-30.9 wt%, respectively. The highest CaO concentration is detected in
433 rock samples with a composite nature, mainly reflecting the prevailing presence of calcite (and
434 minor dolomite) in the marble components. The Avellino skarns have a SiO₂ narrower range of
435 variation (30.1-44.6 wt.%) compared to the Pompeii xenoliths (Fig. 4a), also due to the absence of
436 composite rocks in that analysed sample set, and CaO and MgO ranges of 3.7-23.4 wt.% and 8.9-
437 31.2, respectively (Fig. 4b). The more frequent occurrence of pyroxene can justify the coupled
438 higher SiO₂ and lower CaO content in the Avellino with respect to the Pompeii skarns (Table 1). In
439 CaO vs. SiO₂ and MgO vs. CaO diagrams (Fig. 4a, b) Pompeii and Avellino skarns fall
440 approximately close to the compositional field of AD 1631 skarns of Balassone et al. (2013, see

441 Fig. 4); some samples also plot in or very close to the AD 79 and AD 472 skarn field of Jolis et al.
442 (2015), while the Pompeii composite ejecta PS44, PS54, and PS62 are close to the marble
443 compositions of these authors.

444 The Avellino syenite sample (PS57) has the highest SiO₂ values (49.5 wt%) of the analyzed
445 sample batch, and CaO and MgO values of 6.5 wt% and 4.6 wt%, respectively (Fig. 4a, b); it falls
446 in the Somma-Vesuvius syenites-foiolites-essexites field from literature (see Fig. 4 in Balassone et
447 al., 2013, and references therein), as well as in the igneous rocks compositional field of Jolis et al.
448 (2015).

449 As regards the two Pompeii cumulate samples (PS80 and PS81), they have SiO₂, CaO, and
450 MgO in the ranges 41.5-44.6 wt%, 4.1-12.3 wt% and 14.2-16.6 wt%, respectively, which are close
451 to data from literature (Fig. 4a, b), also considering the igneous rocks compositional field of Jolis et
452 al. (2015).

453 In the main variation diagrams of selected major elements against SiO₂ (Fig. S1 a-f); a semi-
454 linear trend for K₂O, Na₂O, Al₂O₃, Fe₂O₃, and TiO₂ can be defined for the skarn and skarn-marble
455 samples, with significantly higher values of K₂O for Pompeii sample PS73 and of Fe₂O₃ for
456 Pompeii samples PS58 and PS73. The MgO concentrations in the skarn suite are more scattered, as
457 similarly observed by Jolis et al. (2015). A negative trend can be also observed in Fig. S1 g-j, where
458 K₂O, Na₂O, Al₂O₃, Fe₂O₃ are plotted against CaO concentrations. In general, some samples with
459 composite nature (i.e., PS44, PS62, and PS62) are featured by low amounts in major elements
460 oxides as K₂O, Na₂O, Al₂O₃, Fe₂O₃, and TiO₂, as expected for marble-related rocks. By comparing
461 the investigated samples and those of Balassone et al. (2013), the skarn samples have similar values
462 (Fig. S1). As already pointed out by Jolis et al. (2015), Somma-Vesuvius skarns typically have a
463 large compositional range in terms of major elements, owing to the occurrence of variable silicates-
464 oxides-carbonates assemblages (with different amounts of clinopyroxene, phlogopite, olivine,
465 humites, spinel, calcite, etc.). In particular, the investigated Pompeii and Avellino skarn xenoliths
466 and Pompeii composite skarn-marble plot between the marble xenoliths and the igneous

467 compositions in Jolis et al. (2015) (Fig. 4), depicting a continuous array; a similar trend can be
468 pointed out in various Harker's diagrams (Fig. S1). The lack of systematic trend for MgO against
469 SiO₂ suggests a combination of limestone and dolostone protoliths (Jolis et al., 2015).

470 Cumulate sample PS80 is quite close to the cumulate samples of Balassone et al. (2013) in
471 terms of K₂O, Na₂O, Al₂O₃, Fe₂O₃ vs. SiO₂ and CaO correlations, whereas sample PS81 plots
472 behaves differently (Fig. S1); it is worth noting that none of the Pompeii cumulates from this study
473 display the high TiO₂ amount of up to 1.9 wt% (Fig. S1 f) in the AD 1944 eruption cumulates
474 recorded by Balassone et al. (2013).

475 Syenite sample PS57 is characterized by high values of Al₂O₃ (20.1 wt%) and alkalis (9.76
476 wt% K₂O and 4.58 wt% Na₂O), as reported in Fig. S1a,b,c. In any case, cumulate and syenite ejecta
477 display a narrow range of major oxide composition that plots in the intermediate position of the
478 trend depicted by the entire sample set (Fig. 4; Fig. S1).

479 Whole-rock trace element concentrations are reported in Table 2 and a selection of elements
480 is plotted against CaO wt% in Fig. S2. Trace element arrays are scattered for the skarn and
481 composite skarn-marble samples, similarly to literature data. Broadly negative trends can be
482 observed for Ba (with higher values for Pompeii composite sample PS40 and Avellino skarns PS60
483 and PS84), Zr (except to skarns PS55, Avellino, and PS82, Pompeii), Rb, Nb, V, Cs, and Tl when
484 plotted against CaO (Fig. S2 a-h). The higher Sr concentrations are detected in some skarn-marble
485 samples (PS44, PS62, and PS65, Pompeii), usually associated with the occurrence of calcite in the
486 mineral assemblage (Fig. S2b). The Th values are commonly lower than 6 ppm, while it is higher in
487 the skarn samples PS70 and PS73 (Avellino) and PS68 and PS77 (Pompeii) (Fig. S2i), without
488 correlation with Th phases (Table 1). The concentrations of Cr, Ni, and Co are commonly low in
489 the skarn samples (Table 1). Among base metals, Cu is very low, excepting for some Avellino
490 samples (i.e., up to 40 ppm in PS55); Zn concentrations from 110 ppm to 660 ppm in Avellino
491 skarn PS58 (Table 1), and Pb in the range <5-30 ppm up to 240 ppm in Pompeii skarn-marble PS44
492 are measured.

493 Cumulate samples PS80 and PS81 have low amounts of Ba, Sr, Zr, and Th, while sample
494 PS81 has higher concentrations of Rb, V, Nb, Cs, and Tl compared to both all other xenoliths (Fig.
495 S1c, d, f, g, h) and cumulate nodule in Balassone et al. (2013).

496 The Pompeii syenite PS57 generally have low concentrations of trace elements, except for
497 Rb, Cs, and Tl amounts, which are closer to the cumulate samples.

498 The chondrite-normalized REE patterns shown in Fig. 5a display an enrichment in LREE
499 (light rare earth elements: La, Ce, Pr, Nd) and MREE (medium rare earth elements: Sm, Eu, Gd,
500 Tb, Dy) relative to HREE (heavy rare earth elements: Ho, Er, Tm, Yb, Lu), with a nearly flat
501 pattern for HREE. These patterns follow the trend those of recent activity ejecta of Balassone et al.
502 (2013, see Fig. 5a), with a wider variation in the LREE values; moreover, the positive Gd anomaly
503 found in the cumulate sample of these authors was not detected in the Pompeii cumulates from the
504 present study. Moreover, the composite skarn-marbles are more depleted in REE respect to most of
505 the samples.

506 In the multi-elemental patterns (Fig. 5b), some elements such as Ba and Ta have a deep
507 trough for all samples; a similar behavior is observed in most of samples for Th, Sr, P, and Ti, with
508 some exceptions represented by skarns and/or composite rocks. A strongly Pb enrichment is
509 observed in composite sample PS44 (Pompeii), whereas Pompeii skarn PS82 is characterized by
510 higher concentrations of U and REE. Compared to the data of Balassone et al. (2013), a similar
511 trace elements trend can be pointed out; higher concentrations of Ba, P, Eu, Tb, Y, Tm, and Yb of
512 the AD 1944 cumulate of the cited authors are observed in comparison with the Pompeii cumulates
513 of this work.

514

515 **Micas**

516

517 Mica crystals are commonly light green in color in skarn ejecta and composite rocks, whereas
518 they have variable color from dark green to brown in the cumulates and syenite. Via SEM

519 observation, the micas habit is mainly euhedral, and clusters of hexagonal platy crystals are
520 common. Micas can be found as laminae reaching up to 1 mm in length in cumulate and syenite
521 samples, whilst smaller size of tens or few hundreds of μm can be detected in the marbles and
522 skarns (Fig. 6).

523

524 ***Chemical composition***

525 In the classification diagram *feal* against *mgli* (Fig. 7a), the Pompeii and Avellino micas fall within
526 the compositional field of phlogopite¹. They also define a linear trend between ^[VI]Al-rich and Mg-
527 poor “end-member”, represented by PS68 (Avellino skarn), and a group of micas characterized by
528 low ^[VI]Al and high Mg contents, mainly belonging to Pompeii composite xenoliths. In the central
529 part of this array, PS57 (Avellino syenite) and PS80 and PS81 (Pompeii cumulates) crystals are
530 found. A similar trend was also observed for Vesuvian micas from Balassone et al. (2013), with
531 ^[VI]Al-rich and Mg-poor biotite from cumulate (AD 1944) on one side and micas from skarn (AD
532 1631) on the opposite site. The Si vs. Al content in tetrahedral site is nearly close to 4 apfu (atoms
533 per formula unit; Fig. 7b), suggesting negligible ^[IV]Fe³⁺ occupancy in this site (^[IV]Fe³⁺ = 0-0.07
534 apfu); also, the general decrease of the total content of Al with increasing Si (Table 3 and Fig. 7b)
535 is compatible with an Al³⁺-Tschermak substitution.

536 Using the tetrahedral diagram reporting the main octahedral Ti and ^[VI]Al cations compared
537 with F and Mg# (Fig. 8a), it is possible to observe that the skarn-related Avellino phlogopites can
538 reach lower Mg#/2 values (up to 0.78, sample PS68) and higher ^[VI]Al (up to 0.25 apfu, sample
539 PS55) and Ti (up to 0.15 apfu, sample PS68) values compared to skarn- and composite rocks-
540 deriving Pompeii phlogopite (see also Fig. S3a, b). The fluorine amounts vary in the interval 0.09-
541 0.71 apfu for the Avellino phlogopite from skarn and in a slightly higher range (0-0.79 apfu) for
542 micas of Pompeii metamorphic-metasomatic samples (Fig. 8a and Fig. S3c). Regarding the Pompeii
543 cumulates, micas PS80 and PS81 have the highest ^[VI]Al (up to 0.27 apfu) and F (up to 0.85 apfu)

¹ Possible lithium amounts, not analyzed in this study, may be considered negligible for the classification in the diagram of Fig. 7a.

544 values, respectively, of the whole analyzed samples (Table 3). Mica from Avellino syenite lies
545 approximately between the fields of skarn/composite rocks dataset from the two eruptions (Fig. 8a
546 and Fig. S3). Micras from Avellino skarns and Pompeii cumulates mainly have higher iron contents
547 compared to the other samples (Fig. 8b), as also resulting from bulk-rock geochemistry (Fig. S1).
548 Moreover, most of the investigated micras have a negligible “brittle mica component”, due to a Ba
549 amount lower than 0.005 apfu, except for micras from Avellino skarns PS60, PS84, and PSK-R
550 which display Ba amounts in the interval 0.02-0.03 apfu (Fig. S4; Table 3).

551 In Fig. 8, the comparison between micras from the present study and those from recent activity
552 of Vesuvius (Balassone et al., 2013) emphasizes that micras from AD 1631 skarns appear slightly
553 closer than those from Pompeii eruption (Fig. S3 and S4); micras from
554 pyrometamorphic/hydrothermally altered ejecta of AD 1872 are chemically rather different (as
555 exemplified by higher F concentrations), and also mica from cumulate of AD 1944 eruption do not
556 fit with samples from Pompeii cumulates from this study.

557 The substitution of Ba for K is mostly accounted for the coupled substitution $^{[XII]}Ba + ^{[IV]}Al$
558 $\leftrightarrow Si + ^{[XII]}K$ (Fig. 9), with a linear array and end-members represented by PSK-R (Avellino skarn)
559 and PS44 (Pompeii composite nodule) micras. A similar trend can be observed for Vesuvius micras
560 from the most recent volcanic activity (Balassone et al., 2013). It is worth noting that in few
561 Avellino and Pompeii micras small amounts of Sr (0.02-0.03 apfu) were detected (see samples PS41,
562 PS60, PS84, and PSK-R in Table 3).

563 Plotted in the Al_2O_3 vs. TiO_2 and FeO_t (wt%) diagrams (Fig. 10a, b), the chemical feature of
564 the studied micras from skarns and composite rocks fits most of those from the Vesuvian skarns
565 (mainly Avellino, Pompeii, and AD 472 Pollena eruptions) of Gilg et al. (2001) in terms of Al_2O_3
566 and TiO_2 , whereas have different values compared to micras in various xenoliths (lavas,
567 metamorphic nodules, and composite rocks) from AD 1872 ejecta (Petti et al., 2010). The
568 concentrations of FeO_t (wt%) and to some extent of the Mg# values of micras from skarn of Avellino
569 and Pompeii (Fig. 10c) eruptions, variously different from the other studied micras (see also Fig.

570 8b), likewise differ from the other literature data. The FeO_t content is higher in Avellino than in
571 Pompeii micas.

572

573 *Crystallographic features*

574 Mica crystals representative of the two clusters in the skarn and composite skarn-marble lithotypes
575 identified based on the classification plot in Fig. 7 were selected for a deeper investigation by
576 means of SCXRD. Unfortunately, it was not possible to obtain good-quality crystals for SCXRD
577 investigations both from the syenite and cumulates.

578 Crystals from samples PS68 and PS77 (Avellino skarns) which exhibit slightly low Mg# as
579 well as crystal from PS50, PS62, and PS73 (Pompeii composite skarns-marbles and skarns) with
580 high Mg# were deeper analyzed. The results of the structure refinements converged to good values
581 of the agreement indexes, $R_1 \leq 3.29\%$ and $wR_2 \leq 4.98\%$. All samples are 1M polytypes and
582 crystallize in the *C2/m* space group, except for the PS77 which was refined in *C2* space group
583 (Table 4). The PS50 crystal resulted to be twinned with 0.67(2):0.271(9) refined proportion
584 between the two individuals. The samples have a great variability of the lattice constants, especially
585 of the *c* axis [$5.3055(1) \leq a \leq 5.3218(1) \text{ \AA}$, $9.1893(1) \leq b \leq 9.2188(4) \text{ \AA}$, $10.1803(2) \leq c \leq$
586 $10.2951(2) \text{ \AA}$]. The PS68 crystal has the shortest *c* cell parameters [$10.1803(2) \text{ \AA}$] which is typical
587 of de-hydrogenated phlogopite (Lacalamita et al., 2011; Schingaro et al., 2011; Lacalamita et al.,
588 2017). On the contrary PS50 and PS62 crystals from Pompeii have a *c* cell parameter [$10.2932(4)$
589 and $10.2951(2) \text{ \AA}$, respectively], which approximates that of the endmember phlogopite
590 (Redhammer and Roth, 2002), in agreement with the trend in Fig. 7. Regarding the bond lengths
591 (Table 5), almost all the studied samples are geometrically ‘homo-octahedral’ (i.e., $\langle M1-O \rangle$ equals
592 the $\langle M2-O \rangle$ distance within three standard deviations, Weiss et al., 1992). Among the individual
593 distances, the PS68 crystal has the M2-O4 bond length remarkably shorter with respect to the other
594 two $\langle M2-O3 \rangle$ distances (Table 5) which is one of the clues of the possible occurrence of oxy-type
595 substitutions. Consistently, it also has the smallest cell volume (Table 4).

596

597 ***Crystal-chemistry***

598 The structural formulae of the selected mica crystals are listed in Table 6. They were obtained by
599 combining the results from the microprobe analysis (Table 3) and structure refinements (Tables 5
600 and 7), with estimated H₂O content (Ventruti et al., 2008) following the approach reported
601 elsewhere (see for instance Schingaro et al., 2012; Lacalamita et al., 2017).

602 The formulae are balanced based on the main substitutions: M³⁺-Tschermak [^{VI}M²⁺ + ^{IV}Si⁴⁺
603 ↔ ^{VI}(Al³⁺, Fe³⁺) + ^{IV}Al³⁺]; Ti-oxy [^{VI}M²⁺ + 2(OH)⁻ → ^{VI}Ti⁴⁺ + 2(O)²⁻ + H₂↑]; M³⁺-oxy [^{VI}M²⁺ +
604 (OH)⁻ ↔ ^{VI}M³⁺ + O²⁻ + ½ H₂]; OH⁻ → F⁻. The tetraferriphlogopite [^{IV}Fe³⁺ ↔ ^{IV}Al] substitution also
605 characterizes the PS73 mica.

606 Fig. 11 exhibits the variation of the *c* cell parameter versus the Mg number [Mg/(Mg+Fe_{tot})].
607 The highest values of the Mg number pertain to the OH-rich phlogopite (PS50 and PS62 Pompeii
608 samples), as well as fluorophlogopite from the AD 1872 eruption of Vesuvius. The latter exhibits *c*
609 shortening typical of dehydrogenation (see below). For the remaining crystals, as well as for other
610 phlogopite and biotite from Vesuvius xenoliths (Balassone et al., 2013), the *c* parameter decreases
611 with the increase of the Fe content, that, alone or with other metals, may be involved in oxy-type
612 substitution in micas.

613 In the case of the more dehydrated (see section above) PS68 crystal, the entry of Ti and M³⁺
614 cations in the octahedral site by oxy-mechanism involving the deprotonation process at the anionic
615 O4 site, strongly affects the values of geometrical parameters (short *c* cell parameter, M2-O4 and
616 K-O4 distances; high values of M2 bond lengths distortion, BLD_{M2}; strong displacement, shift_{M2}, of
617 the M2 octahedral cation from the geometric center of the octahedron, etc., see Tables 4 and 5)
618 which are consistent with those found for Ti-oxy phlogopite (Cruciani and Zanazzi, 1994). On the
619 contrary, the very low extent of oxy-mechanisms and OH⁻ → F⁻ replacement in PS50 and PS62
620 end-member crystals explains the high values of the *c* cell parameter and the low values of the
621 BLD_{M2} and shif_{M2} parameters (Table 5). For these crystals, the high charge cations are equally

622 distributed between the M1 and M2 octahedral sites as testified by the very close values of the
623 $\langle M1-O \rangle$ and $\langle M2-O \rangle$ distances and volumes of the octahedra (Table 5). These crystals also have
624 the highest value ($\sim 11^\circ$ vs 9° , Table 5) of the tetrahedral rotation angle, α , with respect to the other
625 studied crystals. This parameter describes the in-plane rotation of adjacent tetrahedra in opposite
626 directions respect with c^* . The large α parameter of PS50 and PS62 crystals indicate a good
627 congruence between the tetrahedral and octahedral sheet.

628 The PS77 and PS73 samples exhibit the same degree of dehydrogenation ($O^{2-}+F^-$ content \sim
629 0.8 apfu) resulting in a c cell parameter (~ 10.23 Å, Table 4) intermediate between those of
630 dehydrogenated (PS68 sample) and OH-rich phlogopite (PS50 and PS62 samples). The high charge
631 cation at the octahedral sites is essentially involved in M^{3+} -Tschermak substitutions (Table 6). In
632 the case of the PS73 crystal, the best structural formula was obtained by considering the total iron
633 as Fe_2O_3 which was distributed between both the octahedral and tetrahedral sites.

634 In the $Shift_{M2}$ parameter against Ti and F (Fig. S5a) four points fall close to the AD 1631
635 literature data and define a slightly flat pattern, whereas a particularly high Ti and $Shift_{M2}$ values for
636 the mica from Avellino skarn PS68 can be observed. Also in the $Shift_{M2}$ vs. F (apfu) diagram (Fig.
637 S5b) sample PS68 plots as an outlier, due to its high value of $Shift_{M2}$, while the other samples
638 behave as in Fig. S5a. Fig. S5c, d shows the in-plane rotation angle, i.e., α parameter ($^\circ$), as a
639 function of the Ti and F contents; in the α vs. Ti diagram (Fig. S5c) micas from the Pompeii skarn-
640 marble samples PS50 and PS62 and from the skarn PS77, as well as the Avellino skarn PS73, are
641 closer to the literature AD 1631 samples, whereas mica from the Avellino skarn PS68 plots very
642 close to micas from the AD 1944. In the α vs. F plot of Fig. S5d a broad negative trend is observed,
643 with the investigated micas close to the AD 1631 and AD 1944 data.

644

645 *Crystallization conditions*

646

647 Considering the mineralogical assemblages observed in most xenoliths, indicative of crystallization
648 conditions during their formation, and following Guarino and Brigatti (2018, and references
649 therein), we try to estimate the crystallization temperature at which the micas are formed. Using the
650 geothermometer of Luhr et al. (1984) [$T \text{ (K)} = 838 / (1.0337 - \text{Ti/Fe}^{2+})$] based on the ratio between
651 octahedral Ti and Fe^{2+} , we have obtained similar ranges of temperatures, i.e., 541 - 583 °C and 545
652 - 596°C for the Pompeii and Avellino xenoliths (Table 3). A higher temperature (824 °C) was
653 obtained for mica in skarn PS68, due to its high Ti/Fe^{2+} ratio. For the Pompeii composite sample
654 PS62, the Ti/Fe^{2+} value of 0.54 made the above formula inapplicable.

655 The formula proposed by Uchida et al. (2007) for pressure determination cannot be applied because
656 the studied micas have the total aluminum less than 1.5 apfu.

657 We tested other thermobarometers based on different mineral pairs, as apatite-phlogopite
658 (Ludington, 1978; Seifert et al., 2000), garnet-clinopyroxene (Duchêne and Albarède, 1999) and
659 garnet-biotite (Reche and Martinez 1996; Wu, 2015), but unreliable results were obtained because
660 the minerals/rocks are commonly not in equilibrium.

661

662

663

IMPLICATIONS

664

665 The mineralogical features and genetic inferences of the studied mica-bearing ejecta may
666 provide valuable information for reconstructing the environmental conditions at Somma-Vesuvius
667 and in similar geological systems.

668 Firstly, the crystal chemical analysis of micas from different xenoliths related to Avellino and
669 Pompeii Plinian eruptions of Somma-Vesuvius volcano evidenced that they belong to the
670 phlogopite-annite join. They are affected by different cations substitutions and dehydrogenation
671 degrees, that largely extend the known range of the other Vesuvian micas (as in Balassone et al.,
672 2013), providing some genetic indications, that can be of certain utility also for the understanding

673 the quiescent volcano dynamics. Their host rocks have a wide composition reflecting the abundance
674 of carbonate or silicate phases (i.e., CaO richer Pompeii skarns-to-marbles or SiO₂ enriched
675 Avellino ejecta, respectively). The main mineralogical and petrological aspects can be summarized
676 as follows:

677

678 **Micas from skarns and composite skarn-marble rocks.** The studied Avellino and Pompeii
679 phlogopites from skarns and composite skarn-marbles are generally Ti-depleted, as also found in
680 most of phlogopites from Avellino, Pompeii, and Pollena of Gilg et al. (2001), and slightly more Ti-
681 depleted than the AD 1631 micas of Balassone et al. (2013). Concurrently, they have variable
682 amounts of Al and F, in keeping with the compositional ranges reported by the cited literature.
683 However, the micas set of the present work (i.e., both samples of metamorphic/metasomatic and of
684 igneous origin) never reach the high F concentrations detected in micas in AD 1872 xenoliths
685 (Balassone et al., 2013), which correspond to fluorophlogopite composition and testify for a
686 strongly dehydrated condition. Similarly, the studied phlogopites display chemical characteristics
687 significantly different from the micas batch related to the AD 1872 nodules of Petti et al. (2010).
688 These AD 1872 xenoliths are pyrometamorphic/hydrothermally altered nodules explaining the very
689 different nature and crystallization conditions of micas.

690 In terms of Mg# and Fe, the micas of this study are more variable compared to the previously
691 cited literature data; in particular, three samples from skarn of Avellino eruption, i.e., PS55, PS68,
692 and PS77, have higher Fe contents and lower Mg# values. Interestingly, mica PS68 is also distinctly
693 highest in Ti concentration of the whole sample set, accompanied by a lower *c* parameter;
694 concurrently, the host rock is Fe-rich and have quite high MgO and TiO₂ values, in relation with
695 silicate-oxide (mica+clinopyroxene+magnetite+vesuvianite) mineral assemblage. According to
696 Laurora et al. (2007), micas with the highest Ti contents can be ideally in equilibrium with a more-
697 evolved melt enriched in Fe and Ti.

698 The chemical and structural data indicate that micas PS73 (Pompeii) and PS77 (Avellino)
699 samples exhibit the same degree of dehydrogenation, resulting in a *c* cell parameter intermediate
700 between those of dehydrogenated (PS68 sample; Avellino) and OH-rich phlogopite (P50 and PS62
701 Pompeii samples). It appears that the phlogopite dehydrogenation variability is higher in ejecta
702 from highest VEI Pompeii eruptions, whereas it is more clustered in those from the AD 1631 and
703 1944 eruptive events. Highly dehydrated mica structure coupled with a fluorophlogopite
704 composition found in AD 1872 ejecta (Balassone et al., 2013), characterized by lower P and higher
705 T environment, have never been found in the present study. Following their reasoning, more
706 variable pressure conditions should be consistent with the environment of fluid-carbonate wall rock
707 interaction developed around the Pompeii and Avellino magma chambers.

708 Concerning the temperatures, fluid inclusion studies of Gilg et al. (2001) give a range of 800-
709 890 °C for most skarn-derived minerals (vesuvianite, calcite) indicating initial temperatures related
710 to the entrapment of these fluid inclusions during crystal growth, similarly was obtained for mica in
711 sample PS68 (824 °C). Differently the generally low temperatures estimated for micas in the skarns
712 and composite skarn-marbles (541-596 °C) well represent the minimum crystallization temperatures
713 related to contact metamorphism.

714 This can be supported by the presence of humite, suggesting that dehydrated micas can occur
715 under high CO₂/H₂O (and F) conditions during their growth by contact metamorphism with with
716 carbonatic limestones. As proposed by Young and Morrison (1992), the humite features of our
717 samples would result by devolatilization reactions driven by addition of aqueous fluid and
718 extraction of heat in endoskarn around plutons at less than 600°C. The presence of humite species
719 in the ejecta paragenesis deserves further studies in order to evaluate stable isotope shift in skarn's
720 calcite correlated with its growth proposed by the Authors.

721 As concerns the Ba content, it is worth noting that micas from Avellino skarn PS60, PS84,
722 and PSK-R have higher Ba amounts compared to the other samples; at the same time, the related
723 host rocks PS60 and PS84 also has the highest Ba concentrations, even though this correlation does

724 not occur for PSK-R sample. Finally, the higher Sr concentrations detected in three of the Pompeii
725 skarn-marble nodules are not reflected in micas composition.

726 Unlike Pompeii, Avellino magmas were usually leucite-free. Leucite in Avellino skarns could
727 suggest that the K enrichment already occurred in the Vesuvius magmatism before the AD 79,
728 either as a primary magma feature or as an effect of carbonate assimilation (Piochi et al., 2006a;
729 Mollo et al., 2010).

730 All the skarns and skarn-marbles contain several F-bearing minerals supporting the role of
731 “fluorine-bearing aqueous-carbonic fluids” in the magma-host rock processes suggested in previous
732 studies (Gilg et al., 2001 and references therein).

733

734 **Micas from cumulates and syenite.** Micas related to cumulates (Pompeii eruption) are
735 compositionally different from cumulate sample from the AD 1944 eruption of Balassone et al.
736 (2013); they are distinctly more depleted in Fe, Ti and Ba and enriched in F and Mg, whereas the Al
737 concentration is variable.

738 Mica from syenite (Avellino eruption) appears to record similar low to medium
739 concentrations of selected elements detected in the source rock, as Ti, Ba, Mg, and Fe.

740 Syenite and one cumulate xenolith have the highest content of Cs and Rb, as expected from
741 the major role of the magmatic processes. The low content of crystallizing mica mainly in syenite
742 can determine the observed enrichment of Cs and Rb. This can be associated with pressure
743 conditions exceeding the mica stability, i.e., depth shallower than 4 km (Fabbrizio and Carroll,
744 2008; Scaillet et al., 2001). Similar minimum crystallization temperatures related to contact
745 metamorphism are also obtained for micas in syenite (554 °C) and and other skarn to cumulate
746 xenoliths (558-583 °C).

747 Remarkably, sodalite is common in Avellino and not in Pompeii magmas (Piochi et al., 2006a
748 and references therein). Its occurrence in the Pompeii cumulate should be associated with an
749 Avellino’s remnant thrown out by the AD 79 eruption.

750

751 Finally, this study highlights the importance of micas crystal chemical features in the genetic
752 contexts.

753 As stated by Lepore et al. (2017), commonly petrologists use only microprobe data and
754 petrographic details to classify mica and the formula is derived by placing the obtained chemical
755 data in worksheets. Often, to get charge neutrality, fanciful cation distributions are hypothesized
756 (e.g., Mg in the tetrahedral site), neglecting crucial information like the role of the light elements
757 (including a possible dehydrogenation) and a good estimation of the $\text{Fe}^{2+}/\text{Fe}^{3+}$ ratio. Indeed, both
758 the estimation of the $\text{Fe}^{2+}/\text{Fe}^{3+}$ ratio and the quantification of light elements (H and Li) are also key
759 variables to obtain insights on the relationship between the studied phase and the crystallization
760 conditions.

761 In the case of absence of direct estimation of the above variables the crystal chemical details
762 may be of help: the hydrogen content can be estimated from the c parameter and the $\text{Fe}^{2+}/\text{Fe}^{3+}$ ratio
763 the one providing the best agreement between the EPMA-derived and the structure refinement-
764 derived mean atomic numbers as well as between calculated and observed bond distances (Ottolini et
765 al., 2012). Structural evidence (i.e., geometrical distortions) of substitution mechanisms can also be
766 employed. This alternative approach is indeed widely recognized and has been pursued in the
767 present work.

768

769

770 **ACKNOWLEDGEMENTS**

771

772 This work is dedicated to the memory of Prof. Enrico Franco, who introduced GB to the study
773 of the Somma-Vesuvius minerals and donated her his personal collection of Vesuvian specimens.
774 We are grateful to Warren Huff for helpful suggestions and for the editorial handling of the
775 manuscript. The revision of Alex Asavin and an anonymous reviewer provided us useful insights

776 that greatly enhanced the quality of the paper. Roberto de Gennaro (DiSTAR, Naples) and Marcello
777 Serracino (IGAG CNR, Rome) are thanked for some SEM-EDS and WDS analyses, respectively.
778 This study was supported by the research funds ‘Ricerca Dipartimentale 2016-2017’ from the
779 Dipartimento di Scienze della Terra, dell' Ambiente e delle Risorse of the Università degli Studi di
780 Napoli Federico II (Italy) granted to GB. VULCAMED and EPOS Research Infrastructure through
781 the contribution of the Italian Ministry of University and Research (MUR) provided financially
782 supports to Istituto Nazionale di Geofisica e Vulcanologia for the X-ray diffraction laboratory at the
783 Osservatorio Vesuviano.

784

785

786 REFERENCES CITED LIST

787

788 Abrecht, J., and Hewitt, D.A., (1988) Experimental evidence on the substitution of Ti in biotite.
789 American Mineralogist, 73, 1275-1284.

790 Arrighi, S., Principe, C., and Rosi, M. (2001) Violent strombolian and subplinian eruptions at
791 Vesuvius during post-1631 activity. Bulletin of Volcanology, 63, 126-150.

792 Auger, E., Gasparini, P., Virieux, J., and Zollo, A. (2001) Seismic evidence of an extended
793 magmatic sill under Mt. Vesuvius. Science, 294, 1510-1512, doi:10.1126/science.1064893.

794 Avanzinelli, R., Cioni, R., Conticelli, S., Giordano, G., Isaia, R., Mattei, M., Melluso, L., and
795 Sulpizio R. (2017) The Vesuvius and the other Volcanoes of Central Italy. Geological Field
796 Trips 9(1.1), [http://www.isprambiente](http://www.isprambiente.gov.it/it/pubblicazioni/periodici-tecnici/geological-field-trips), gov.it/it/pubblicazioni/periodici-tecnici/geological-
797 field-trips [accessed on 5 January 2021].

798 Ayuso, R.A., Vivo, B.D., Rolandi, G., and Paone, A. (1998) Geochemical and isotopic (Nd–Pb–Sr–
799 O) variations bearing on the genesis of volcanic rocks from Vesuvius, Italy. Journal of
800 Volcanology and Geothermal Research, 82, 53-78.

- 801 Babushkina, M.S., Lepekhina, E.N., Nikitina, L.P., Ovchinnikov, N.O., and Lokhov, K.I. (2000)
802 Crystal structure defects of micas from lamproites (Mossbauer and infrared spectroscopy
803 data). *Doklady Akademii Nauk*, 371, 797-801 (in Russian).
- 804 Balassone, G., Bellatreccia, F., Mormone, A., Biagioni, C., Pasero, M., Petti, C., Mondillo, N., and
805 Fameli, G. (2012) Sodalite-group minerals from the Somma-Vesuvius volcanic complex,
806 Italy: a case study of K-feldspar-rich xenoliths. *Mineralogical Magazine*, 76, 191-212.
- 807 Balassone, G., Scordari, F., Lacalamita, M., Schingaro, E., Mormone, A., Piochi, M., Petti C., and
808 Mondillo N. (2013) Trioctahedral micas in xenolithic ejecta from recent volcanism of the
809 Somma-Vesuvius (Italy): Crystal chemistry and genetic inferences. *Lithos*, 160-161, 84-97.
- 810 Balassone, G., Bellatreccia, F., Ottolini, L., Mormone, A., Petti, C., Ghiara, M.R., Altomare, A.,
811 Saviano, M., Rizzi, R., and D’Orazio L. (2016) Sodalite-group minerals from Somma-
812 Vesuvius volcano (Naples, Italy): A combined EPMA, SIMS and FTIR crystal chemical
813 study. *The Canadian Mineralogist*, 54, 583-604, doi.10.3749/canmin.1500083.
- 814 Balcone-Boissard, H., Boudon, G., Ucciani, G., Villemant, B., Cioni, R., Civetta, L., and Orsi, G.
815 (2012) Magma degassing and eruption dynamics of the Avellino pumice Plinian eruption of
816 Somma–Vesuvius (Italy). Comparison with the Pompeii eruption. *Earth and Planetary
817 Science Letters*, 331-332, 257-268.
- 818 Barberi, F., and Leoni, L. (1980) Metamorphic carbonate ejecta from Vesuvius Plinian eruptions:
819 evidence of the occurrence of shallow magma chambers. *Bulletin of Volcanology*, 43, 107-
820 120.
- 821 Barberi, F., Bizouard, H., Clocchiatti, R., Metrich, N., Santacroce, R., and Sbrana, A. (1981). The
822 Somma-Vesuvius magma chamber: a petrological and volcanological approach. *Bulletin
823 Volcanologique*, 44, 295-315.
- 824 Betteridge, P.W., Carruthers, J.R., Cooper, R.I., Prout, K., and Watkin, D.J. (2003) Crystals
825 version: software for guided crystal structure analysis. *Journal of Applied Crystallography*,
826 36, 1487.

- 827 Brigatti, M.F., Galli, E., and Poppi, L. (1991) Effect of Ti substitution in biotite-1M crystal
828 chemistry. American Mineralogist, 76, 1174-1183.
- 829 Brigatti, M.F., and Poppi, L. (1993) Crystal chemistry of Ba-rich trioctahedral micas-1M. European
830 Journal of Mineralogy, 5, 857-871.
- 831 Brigatti, M.F., Caprilli, E., Funicello, R., Giordano, G., Mottana, A., and Poppi, L. (2005) Crystal
832 chemistry of ferroan phlogopites from the Albano maar lake zone (Colli Albani volcano,
833 central Italy). European Journal of Mineralogy, 17, 611-621.
- 834 Brocchini, D., Principe, C., Castradori, D., Laurenzi, M.A., and Gorla, L. (2001) Quaternary
835 evolution of the southern sector of the Campanian Plain and early Somma-Vesuvius activity:
836 insights from Trecase 1 well. Mineralogy and Petrology, 73, 67-91.
- 837 Bruno, P.P.G., Cippitelli, G., and Rapolla, A. (1998) Seismic study of the Mesozoic carbonate
838 basement around Mt. Somma-Vesuvius, Italy. Journal of Volcanology and Geothermal
839 Research, 84, 311-322.
- 840 Bruker (2007) SAINT, Bruker AXS Inc., Madison, Wisconsin, USA.
- 841 Bruker (2009) SADABS, Bruker AXS Inc., Madison, Wisconsin, USA.
- 842 Bruker (2010) APEX2 v.2010.7-0, Bruker AXS Inc., Madison, Wisconsin, USA.
- 843 Cioni, R., Civetta, L., Marianelli, P., Metrich, N., Santacroce, R., and Sbrana, A. (1995)
844 Compositional layering and syn-eruptive mixing of a periodically refilled shallow magma
845 chamber: the AD 79 Plinian eruption of Vesuvius. Journal of Petrology, 36, 739-776.
- 846 Cioni, R., Santacroce, R., and Sbrana, A. (1999) Pyroclastic deposits as a guide for reconstructing
847 the multi-stage evolution of the Somma-Vesuvius Caldera. Bulletin of Volcanology, 60, 207-
848 222.
- 849 Cioni, R. (2000) Volatile content and degassing processes in the AD 79 magma chamber at
850 Vesuvius (Italy). Contribution to Mineralogy and Petrology, 140, 40-54.

- 851 Cioni, R., Bertagnini, A., Santacroce, R., and Andronico, D. (2008) Explosive activity and eruption
852 scenarios at Somma-Vesuvius (Italy): Towards a new classification scheme. *Journal of*
853 *Volcanology and Geothermal Research*, 178, 331-346.
- 854 Civetta, L., D'Antonio, M., de Lorenzo, S., Di Renzo, V., and Gasparini, P. (2004) Thermal and
855 geochemical constraints on the 'deep' magmatic structure of Mt. Vesuvius. *Journal of*
856 *Volcanology and Geothermal Research*, 133, 1-12.
- 857 Conticelli, S., Guarnieri, L., Farinelli, A., Mattei, M., Avanzinelli, R., Bianchini, G., Boari, E.,
858 Tommasini, S., Tiepolo, M., Prelević, D., and Venturelli, G. (2009) Trace elements and Sr–
859 Nd–Pb isotopes of K-rich, shoshonitic, and calc-alkaline magmatism of the Western
860 Mediterranean Region: genesis of ultrapotassic to calc-alkaline magmatic associations in a
861 post-collisional geodynamic setting. *Lithos*, 107, 68-92.
- 862 Conticelli, S., Laurenzi, M.A., Giordano, G., Mattei, M., Avanzinelli, R., Melluso, L., Tommasini,
863 S., Boari, E., Cifelli, F., and Perini, G. (2011) Leucite-bearing (kamafugitic/leucititic) and -
864 free (lamproitic) ultrapotassic volcanic rocks and associated shoshonites in the Italian
865 Peninsula: constraints on petrogenesis and geodynamics. In: "The Geology of Italy: tectonics
866 and life along plate margins", Beltrando, M., Peccerillo, A., Mattei, M., Conticelli, S.,
867 Doglioni, C. (Eds.), *Journal of the Virtual Explorer* 36, paper 20, 1-95,
868 <https://dx.doi.org/10.3809/jvirtex.2009.00251>.
- 869 Cruciani, G., and Zanazzi, P.F. (1994) Cation partitioning and substitution mechanisms in 1 M
870 phlogopite: a crystal chemical study. *American Mineralogist*, 79, 289-301.
- 871 Cucciniello, C. (2016) Tetra-plot: a Microsoft Excel spreadsheet to perform tetrahedral diagrams.
872 *Periodico di Mineralogia*, 85, 115-119.
- 873 Dallai, L., Cioni, R., Boschi, C., and D'Oriano, C. (2011) Carbonate-derived CO₂ purging magma at
874 depth: Influence on the eruptive activity of Somma-Vesuvius, Italy. *Earth and Planetary*
875 *Science Letters*, 310, 84-95.

- 876 Di Renzo, V., Di Vito, M.A., Arienzo, I., Civetta, L., D'Antonio, M., Giordano, F., Orsi, G., and
877 Tonarini, S. (2007) Magmatic history of Somma-Vesuvius on the basis of new geochemical
878 and isotopic data from a deep borehole (Camaldoli della Torre). *Journal of Petrology*, 48,
879 753-784.
- 880 Del Moro, A., Fulignati, P., Marianelli, P., and Sbrana, A. (2001) Magma contamination by direct
881 wall rock interaction, constraints from xenoliths from the walls of a carbonate-hosted magma
882 chamber (Vesuvius 1944 eruption). *Journal of Volcanology and Geothermal Research*, 112,
883 15-24.
- 884 Di Vito, M.A., Talamo, P., de Vita, S., Rucco, I., Zanchetta, G., and Cesarano, M. (2019) Dynamics
885 and effects of the Vesuvius Pomici di Avellino Plinian eruption and related phenomena on the
886 Bronze Age landscape of Campania region (Southern Italy). *Quaternary International*, 499,
887 231-244.
- 888 Doronzo, D.M., Di Vito, M.A., Arienzo, I., Bini, M., Calusi, B., Cerminara, M., Corradini, S., de
889 Vita, S., Giaccio, B., Gurioli, L., Mannella, G., Ricciardi, G.P., Rucco, I., Sparice, D.,
890 Todesco, M., Trasatti, E., and Zanchetta, G. (2022) The 79 CE eruption of Vesuvius: A lesson
891 from the past and the need of a multidisciplinary approach for developments in volcanology,
892 *Earth-Science Reviews*, 231, 104072, doi.org/10.1016/j.earscirev.2022.104072.
- 893 Duchêne, S., and Albarède, F. (1999) Simulated garnet-clinopyroxene geothermometry of eclogites.
894 *Contributions to Mineralogy and Petrology*, 135(1), 75.
- 895 Dymek, R.F. (1983) Titanium, aluminum and interlayer cation substitutions in biotite from high
896 grade gneisses, West Greenland. *American Mineralogist*, 68, 880-899.
- 897 Fabrizio, A., and Carroll, M.R. (2008) Experimental constraints on the differentiation process and
898 pre-eruptive conditions in the magmatic system of Phlegraean Fields (Naples, Italy). *Journal*
899 *of Volcanology and Geothermal Research*, 171, 88-102.

- 900 Fritschle, T., Prelević, D., Foley, S.F., and Jacob, D.E. (2013) Petrological characterization of the
901 mantle source of Mediterranean lamproites: Indications from major and trace elements of
902 phlogopite. *Chemical Geology*, 353, 267-279.
- 903 Fulignati, P., Panichi, C., Sbrana, A., Caliro, S., Gioncada, A., and Del Moro, A. (2005) Skarn
904 formation at the walls of the 79AD magma chamber of Vesuvius (Italy): Mineralogical and
905 isotopic constraints. *Neues Jahrbuch für Mineralogie, Abhandlungen* 181/1.
- 906 Gilg, H.A., Lima, A., Somma, R., Belkin, H.E., De Vivo, B., and Ayuso, R.A. (2001) Isotope
907 geochemistry and fluid inclusion study of skarns from Vesuvius. *Mineralogy and Petrology*,
908 73, 145-176.
- 909 Guarino, V., and Brigatti, M.F. (2018) Crystallization conditions of micas in oxidized igneous
910 systems. *American Mineralogist*, 103, 1999-2010, doi: <https://doi.org/10.2138/am-2018-6567>.
- 911 Gurioli, L., Houghton, B.F., Cashman, K.V., and Cioni, R. (2005) Complex changes in eruption
912 dynamics during the 79 AD eruption of Vesuvius. *Bulletin of Volcanology*, 67, 144-159,
913 doi:10.1007/s00445-004-0368-4.
- 914 Iacono-Marziano, G., Gaillard, F., and Pichavant, M. (2008) Limestone assimilation by basaltic
915 magmas: an experimental re-assessment and application to Italian volcanoes. *Contributions to*
916 *Mineralogy and Petrology*, 155, 719-738, doi: 10.1007/s00410-007-0267-8.
- 917 Iacono-Marziano, G., Gaillard, F., Scaillet, B., Pichavant, M., and Chiodini, G. (2009) Role of non-
918 mantle CO₂ in the dynamics of volcano degassing: The Mount Vesuvius example. *Geology*,
919 37, 319-322.
- 920 Jolis, E.M., Freda, C., Troll, V.R., Deegan, F.M., Blythe, L.S., McLeod, C.L., and Davidson, J.P.,
921 (2013) Experimental simulation of magma–carbonate interaction beneath Mt. Vesuvius, Italy.
922 *Contrib. Mineralogy and Petrology*, 166, 1335-1353, [http://dx.doi.org/10.1007/s00410-013-](http://dx.doi.org/10.1007/s00410-013-0931-0)
923 0931-0.

- 924 Jolis, E.M., Troll, V.R., Harris, C., Freda, C., Gaeta, M., Orsi, G., and Siebe, C. (2015) Skarn
925 xenolith record crustal CO₂ liberation during Pompeii and Pollena eruptions, Vesuvius
926 volcanic system, central Italy. *Chemical Geology*, 415, 17-36.
- 927 Joron, J.L., Métrich, N., Rosi, M., Santacroce, R., and Sbrana, A. (1987) Chemistry and
928 petrography. In: Santacroce, R. (Ed.), *Somma-Vesuvius*. Quaderni de “La Ricerca
929 Scientifica” CNR: Roma, 8, 105-171.
- 930 Lacalamita, M., Schingaro, E., Scordari, F., Ventruti, G., Fabbrizio, A., and Pedrazzi, G. (2011)
931 Substitution mechanisms and implications for the estimate of water fugacity for Ti-rich
932 phlogopite from Mt. Vulture (Potenza, Italy). *American Mineralogist*, 96, 1381-1391.
- 933 Lacalamita, M., Mesto, E., Scordari, F., and Schingaro, E. (2012) Chemical and structural study of
934 1M- and 2M1-phlogopites coexisting in the same Kasenyi kamafugitic rock (SW Uganda).
935 *Physics and Chemistry of Minerals*, 39, 601-611.
- 936 Lacalamita, M., Balassone, G., Schingaro, E., Mesto, E., Mormone, A., Piochi, M., Ventruti, G.,
937 and Joachimski, M. (2017) Fluorophlogopite-bearing and carbonate metamorphosed xenoliths
938 from the Campanian Ignimbrite (Fiano, southern Italy): crystal chemical, geochemical and
939 volcanological insights. *Mineralogical Magazine*, 81, 1165-1189.
- 940 Lacalamita, M., Schingaro, E., Mesto, E., Zaccarini, E., and Biagioni, C. (2020) Crystal-chemistry
941 of micas belonging to the yangzhumingite-fluorophlogopite and phlogopite-fluorophlogopite
942 series from the Apuan Alps (northern Tuscany, Italy). *Physics and Chemistry of Minerals*, 47,
943 54, 8-10.
- 944 Laurora, A., Brigatti, M.F., Mottana, A., Malferrari, D., and Caprilli, E. (2007) Crystal chemistry of
945 trioctahedral micas in alkaline and subalkaline volcanic rocks: A case study from Mt. Sassetto
946 (Tolfa district, Latium, central Italy). *American Mineralogist*, 92, 468-480.
- 947 Laurora, A., Malferrari, D., Brigatti, M.F., Mottana, A., Caprilli, E., Giordano, G., and Funicello,
948 R. (2009) Crystal chemistry of trioctahedral micas in the top sequences of the Colli Albani
949 volcano, Roman Region, central Italy. *Lithos*, 113, 507-520.

- 950 Lepore, G.O., Bindi, L., Pedrazzi, G., Conticelli, S., and Bonazzi, P. (2017) Structural and chemical
951 variations in phlogopite from lamproitic rocks of the Central Mediterranean region. *Lithos*,
952 286-287, 191-205.
- 953 Ludington, S. (1978) The biotite-apatite geothermometer revisited. *American Mineralogist*, 63(5-6),
954 551-553.
- 955 Luhr, J.F., Carmichael, I.S.E., and Varekamp, J.C. (1984) The 1982 eruptions of El Chichón
956 Volcano, Chiapas, Mexico: Mineralogy and petrology of the anhydrite bearing pumices.
957 *Journal of Volcanology and Geothermal Research*, 23(1-2), 39-68.
- 958 Lyubetskaya, T., and Korenaga, J. (2007) Chemical composition of Earth's primitive mantle and its
959 variance: 1. Methods and results. *Journal of Geophysical Research*, 112, 1-21, B03211.
- 960 McDonough, W.F., and Sun, S.S. (1995) The composition of the Earth. *Chemical Geology*, 120,
961 223-253.
- 962 Marianelli, P., Metrich, N., and Sbrana, A. (1999) Shallow and deep reservoirs involved in magma
963 supply of the 1944 eruption of Vesuvius. *Bulletin of Volcanology*, 61, 48-63.
- 964 Massaro, S., Costa, A., and Sulpizio, R. (2018) Evolution of the magma feeding system during a
965 Plinian eruption: the case of Pomici di Avellino eruption of Somma-Vesuvius, Italy. *Earth
966 and Planetary Science Letters*, 482, 545-555.
- 967 Matarrese, S., Schingaro, E., Scordari, F., Stoppa, F., Rosatelli, G., Pedrazzi, G., and Ottolini, L.
968 (2008) Crystal chemistry of phlogopite from Vulture-S.Michele Subsynthem volcanic rocks
969 (Mt. Vulture, Italy) and volcanological implications. *American Mineralogist*, 93, 426-437.
- 970 McDonough, W., and Sun, S.S. (1995) The composition of the Earth. *Chemical Geology*, 67, 1050-
971 1056.
- 972 Melluso, L., Scarpati, C., Zanetti, A., Sparice, D., and de' Gennaro, R. (2022) The petrogenesis of
973 chemically zoned, phonolitic, Plinian and sub-Plinian eruptions of Somma-Vesuvius, Italy:
974 role of accessory phase removal, independently filled magma reservoirs with time, and

- 975 transition from slightly to highly silica undersaturated magmatic series in an ultrapotassic
976 stratovolcano. *Lithos*, 430, 106854.
- 977 Mollo, S., Gaeta, M., Freda, C., Di Rocco, T., Misiti, V., and Scarlato, P. (2010) Carbonate
978 assimilation in magmas: a reappraisal based on experimental petrology. *Lithos*, 114(3-4), 503-
979 514.
- 980 Mormone, A., Piochi, M., Balassone, G., Carlino, S., Somma, R., Troise, C., and De Natale, G.
981 (2014) Mineralogical and petrological investigations of rocks cored from depths higher than
982 440 m during the CFDDP drilling activities at the Campi Flegrei caldera (southern Italy). In
983 EGU General Assembly Conference Abstracts (p. 14198).
- 984 Moretti, R., Arienzo, I., Civetta, L., Orsi, G., and Papale, P. (2013) Multiple magma degassing
985 sources at an explosive volcano. *Earth and Planetary Science Letters*, 367, 95-104.
- 986 Ottolini, L.P., Schingaro, E., and Scordari, F. (2012) Ceramics: contribution of secondary ion mass
987 spectrometry (SIMS) to the study of the crystal chemistry of mica minerals. *Mass
988 Spectrometry Handbook*, Mike S. Lee Editor, Wiley, pp.1017-1059.
- 989 Pappalardo, L., Piochi, M., and Mastrolorenzo, G. (2004) The 3550 year BP-1944 A.D. magma-
990 plumbing system of Somma-Vesuvius: constraints on its behaviour and present state through
991 a review of Sr-Nd isotope data. *Annals Geophysics*, 47, 1471-1483.
- 992 Pascal, M.-L., Di Muro, A., Fonteilles, M., and Principe, C. (2009) Zirconolite and calzirtite in
993 banded forsterite-spinel-calcite skarn ejecta from the 1631 eruption of Vesuvius: inferences
994 for magma-wallrock interactions. *Mineralogical Magazine*, 73(2), 333–356.
- 995 Pascal, M.L., Fonteilles, M., Boudouma, O., and Principe, C. (2011) Qandilite from Vesuvius skarn
996 ejecta: Conditions of formation and miscibility gap in the ternary spinel–qandilite–
997 magnesioferrite. *The Canadian Mineralogist*, 49, 459-485.
- 998 Peccerillo, A. (2005) Plio-Quaternary Volcanism in Italy. *Petrology, Geochemistry, Geodynamics*.
999 Springer Berlin Heidelberg, New York, pp. 365.

- 1000 Petti, C., Balassone, G., Bernardi, A.M., Mormone, A., Rossi, M., and Ghiara, M.R. (2010) Ejected
1001 rocks from the 1872 eruption of Vesuvius, Italy: a petrographic and mineralogical overview.
1002 *Periodico di Mineralogia*, 79, 1-25.
- 1003 Pichavant, M., Scaillet, B., Pommier, A., Iacono-Marziano, G., and Cioni, R. (2014) Nature and
1004 evolution of primitive Vesuvius magmas: an experimental study. *Journal of Petrology*, 55,
1005 2281-2310.
- 1006 Piochi, M., Pappalardo, L., and De Astis, G. (2004) Geochemical and isotopical variations within
1007 the Campanian Comagmatic Province: implications on magma source composition. *Annals of*
1008 *Geophysics* 44, 1485-1499.
- 1009 Piochi, M., Ayuso, R.A., De Vivo, B., and Somma, R. (2006a) Crustal contamination and crystal
1010 entrapment during polybaric magma evolution at Mt. Somma-Vesuvius volcano, Italy.
1011 Geochemical and Sr isotope evidence. *Lithos*, 86, 303-329.
- 1012 Piochi, M., De Vivo, B., Ayuso, and R.A. (2006b) The magma feeding system of Somma-Vesuvius
1013 (Italy) strato-volcano: new inferences from a review of geochemical and Sr, Nd, Pb and O
1014 isotope data. *Developments in Volcanology*, 9, 181-202.
- 1015 Pouchou, J.L., and Pichoir, F. (1991) Quantitative analysis of homogeneous or stratified
1016 microvolumes applying the model "PAP". In *Electron probe quantitation*, Springer, Boston,
1017 MA, pp. 31-75.
- 1018 Reche, J., and Martinez, F. J. (1996) GPT: an Excel spreadsheet for thermobarometric calculations
1019 in metapelitic rocks. *Computers & Geosciences*, 22(7), 775-784.
- 1020 Redhammer, G.J., and Roth, G. (2002) Single-crystal structure refinements and crystal chemistry of
1021 synthetic trioctahedral micas $KM_3(Al^{3+}, Si^{4+})_4O_{10}(OH)_2$, where M= Ni^{2+} , Mg^{2+} , Co^{2+} , Fe^{2+} , or
1022 Al^{3+} . *American Mineralogist*, 87, 1464-1476.
- 1023 Rice, J.M. (1980) Phase equilibria involving humite minerals in impure dolomitic limestones, Part
1024 I. Calculated stability of clinohumite. *Contributions to Mineralogy and Petrology* 71, 219-
1025 235.

- 1026 Russo, M., and Punzo, I. (2004) I minerali del Somma-Vesuvio. Associazione Mineralogica
1027 Italiana, Cremona, pp. 317.
- 1028 Santacroce, R. (1987) Somma–Vesuvius. Quaderni della Ricerca Scientifica CNR 114, pp. 1-230.
- 1029 Santacroce, R., Cioni, R., Marianelli, P., Sbrana, A., Sulpizio, R., Zanchetta, G., Donahue, D.J.,
1030 Joron, J.-L., 2008. Age and whole rock-glass composition of proximal pyroclastics from the
1031 major explosive eruptions of Somma–Vesuvius: a review as a tool for distal
1032 tephrostratigraphy. *Journal of Volcanology and Geothermal Research*, 177, 1-18.
- 1033 Scaillet, B., Pichavant, M., Cioni, R., Sbrana, A., and Marianelli, P. (2001) Experimental
1034 constraints on pre-eruption conditions of plinian events at Vesuvius. In *Geophysical Research*
1035 *Abstracts*. EGU.
- 1036 Scaillet, B., Pichavant, M., and Cioni, R. (2008) Upward migration of Vesuvius magma chamber
1037 over the past 20.000 years. *Nature*, 455, 216–219.
- 1038 Schingaro, E., Kullerud, K., Lacalamita, M., Mesto, E., Scordari, F., Zozulya, D., Erambert, M., and
1039 Ravna, E.J. (2014) Yangzhumingite and phlogopite from the Kvaløya lamproite (North
1040 Norway): Structure, composition and origin. *Lithos*, 210-211, 1-13.
- 1041 Schingaro, E., Lacalamita, M., Scordari, F., Brigatti, M.F., and Pedrazzi, G. (2011) Crystal
1042 chemistry of Ti-rich fluorophlogopite from Presidente Olegario, Alto Paranaíba igneous
1043 province, Brazil. *American Mineralogist*, 96, 732-743.
- 1044 Schingaro, E., Scordari, F., and Ventruti, G. (2001) trioctahedral micas-1M from Mt. Vulture
1045 (Italy): Structural disorder and crystal chemistry. *European Journal of Mineralogy* 13, 1057-
1046 1069.
- 1047 Scordari, F., Schingaro, E., Lacalamita, and M., Mesto, E. (2012). Crystal chemistry of trioctahedral
1048 micas-2M1 from Bunyaruguru (SW Uganda) kamafugite. *American Mineralogist*, 97, 430-
1049 439.

- 1050 Seifert, W., Kämpf, H., and Wasternack, J. (2000) Compositional variation in apatite, phlogopite
1051 and other accessory minerals of the ultramafic Delitzsch complex, Germany: implication for
1052 cooling history of carbonatites. *Lithos*, 53(2), 81–100.
- 1053 Sekine, T., Wyllie, P.J., 1982. Synthetic systems for modeling hybridization between hydrous
1054 siliceous magmas and peridotite in subduction zones. *The Journal of Geology*, 90, 734-741.
- 1055 Sheldrick, G.M., 2003a. CELL_NOW, Program for unit cell determination. University of
1056 Göttingen, Germany.
- 1057 Sheldrick, G.M. (2003b) SADABS, Program for empirical absorption correction of area detector
1058 data. University of Göttingen, Germany.
- 1059 Sheldrick, G.M. (2008) A short history of SHELX. *Acta Crystallographica Section A: Foundations*
1060 *of Crystallography* 64(1), 112-122.
- 1061 Signorelli, S., Vaggelli, G., and Romano, C. (1999) Pre-eruptive volatile (H₂O, F, Cl and S)
1062 contents of phonolitic magmas feeding the 3550-year old Avellino eruption from Vesuvius;
1063 southern Italy. *Journal of Volcanology and Geothermal Research*, 93, 237-256.
- 1064 Sulpizio, R., Bonasia, R., Dellino, P., Mele, D., Di Vito, M.A., and La Volpe, L. (2010) The
1065 Avellino eruption of Somma–Vesuvius (3.9 ka BP). part II: sedimentology and physical
1066 volcanology of pyroclastic density current deposits. *Bulletin of Volcanology*, 72, 559-577,
1067 doi:10.1007/s00445-009-0340-4.
- 1068 Uchida, E., Endo, S., and Makino, M. (2007) Relationship between solidification depth of granitic
1069 rocks and formation of hydrothermal ore deposits. *Resource Geology* 57, 47–56.
- 1070 Vona, A., Romano, C., Giordano, G., and Sulpizio, R. (2020) Linking magma texture, rheology and
1071 eruptive style during the 472AD Pollena Subplinian eruption (Somma-Vesuvius). *Lithos*,
1072 370-371, 105658.
- 1073 Ventruti, G., Zema, M., Scordari, F., and Pedrazzi, G. (2008) Thermal behaviour of a Ti-rich
1074 phlogopite from Mt. Vulture (Potenza, Italy): An in situ X-ray single-crystal diffraction study.
1075 *American Mineralogist* 93, 632-643.

- 1076 Young, E.D., and Morrison, J. (1992) Relations among net-transfer reaction progress, ^{18}O - ^{13}C
1077 depletion, and fluid infiltration in a clinohumite-bearing marble. *Contribution to Mineralogy*
1078 and *Petrology*, 111, 391-408.
- 1079 Warr, L.N. (2021) IMA–CNMNC approved mineral symbols. *Mineralogical Magazine* 85, 291-
1080 320, doi:10.1180/mgm.2021.43.
- 1081 Waters, D.J., and Charnley, N.R. (2002) Local equilibrium in polymetamorphic gneiss and the
1082 titanium substitution in biotite. *American Mineralogist*, 87, 383-396.
- 1083 Weiss, Z., Rieder, M., and Chmielová, M. (1992) Deformation of coordination polyhedra and their
1084 sheets in phyllosilicates. *European Journal of Mineralogy*, 4, 665-682.
- 1085 Wu, C. M. (2015) Revised empirical garnet–biotite–muscovite–plagioclase geobarometer in
1086 metapelites. *Journal of Metamorphic Geology*, 33(2), 167-176.
- 1087 Zollo, A., Gasparini, P., Virieux, J., Le Meur, H., Natale, G. de, Biella, G., Boschi, E., Capuano, P.,
1088 Franco, R. de, dell’Aversana, P., Matteis, R. de, Guerra, I., Iannaccone, G., Mirabile, L., and
1089 Vilaro, G. (1996) Seismic evidence for a low-velocity zone in the upper crust beneath Mount
1090 Vesuvius. *Science*, 274, 592-594

1091

1092

1093 **LIST OF FIGURE CAPTIONS**

1094

1095 Figure 1. Geological sketch map of Somma-Vesuvius (modified after Peccerillo, 2005) with
1096 samples provenance.

1097

1098 Figure 2. (a) Selected xenoliths from Somma-Vesuvius (see Table 1): PS44 (composite skarn-
1099 marble) and PS80 (cumulate) are from Pompeii eruption; PS57 (syenite) and PS77 (skarn) are
1100 from Avellino eruption (bar scale equal to 5 cm). (b) Distribution of the samples set according

1101 to the lithotype and the occurrence.

1102

1103 Figure 3. Selected SEM micrographs (backscattered electrons, BSE mode) of mica-bearing
1104 xenoliths in thin sections. (a-h) skarn xenoliths; (i-m) skarn-marble xenoliths; (n) syenite
1105 skarns; (o,p) cumulate xenoliths. A thorough description of the mineral associations are given
1106 in the text. Mineral abbreviations (after Warr et al., 2021) are as in Table 1.

1107

1108 Figure 4. Selected major elements variations diagrams for the investigated Somma-Vesuvius
1109 xenoliths compared with literature data, (a) CaO against SiO₂ and (b) MgO against CaO
1110 (wt%).

1111

1112 Figure 5. Geochemical characteristics of the investigated mica-bearing xenoliths. (a) Chondrite-
1113 normalized REE and (b) primitive mantle-normalized minor to trace element patterns
1114 (normalization values from McDonough and Sun, 1995, and Lyubetskaya and Korenaga,
1115 2007, respectively). Data from Vesuvius xenoliths (recent activity) of Balassone et al. (2013)
1116 are also reported.

1117

1118 Figure 6. SEM micrographs of micas from Somma-Vesuvius samples. (a) PS54; (b) PS57; (c) and
1119 (d) PS80, (Sdl, sodalite).

1120

1121 Figure 7. (a) Octahedral sheet composition of the studied micas in the *feal-mgli* plot, where *feal* =
1122 $^{[IV]}(\text{Fe}_{\text{tot}} + \text{Mn} + \text{Ti} - \text{Al})$ and *mgli* = $^{[VI]}(\text{Mg} + \text{Li})$ of Tischendorf et al. (2007). (b) Plot of
1123 tetrahedral Si vs. Al; the correlation between Si and total Al is reported in the inset. Symbols
1124 used for micas are the same as in Fig. 4.

1125

1126 Figure 8. Discrimination diagrams for the studied micas, $\text{Ti}^{-\text{VI}}\text{Al}^{-\text{F}^-}\text{-Mg}\#2$ (a) and $\text{Ba-Fe}_{\text{tot}}\text{-F}^{-}\text{-Mg}\#$

1127 (b. The tetrahedral diagram has been made using the Tetra-Plot spreadsheet (Cucciniello,
1128 2016). Symbols as in Fig. 4.

1129 Figure 9. Plot of the $^{[XII]}Ba + ^{[IV]}Al \leftrightarrow Si + ^{[XII]}K$ substitution mechanism for the studied micas
1130 (symbols as in Fig. 4).

1131 Figure 10. Plot of Al_2O_3 vs. TiO_2 (a) and FeO_t (b), as well as Mg # vs. TiO_2 (c) for micas of the
1132 present study, compared to selected Somma-Vesuvius micas composition from literature.
1133 Symbols explanation as in Fig. 4; black dotted line field = micas from AD 1872 eruption
1134 ejecta (Petti et al., 2010); grey solid line field = micas from skarn xenoliths mainly from
1135 Avellino, Pompeii and AD 472 Pollena eruptions (Gilg et al., 2001).

1136

1137 Figure 11. Plot of the c parameter vs. Mg number $[Mg/(Mg+Fe_{tot})]$ for the investigated trioctahedral
1138 micas from the Avellino and Pompeii xenoliths (symbols as in Fig. 4). For comparison, values
1139 of phlogopite, fluorophlogopite and biotite of Vesuvius xenoliths from the AD 1631, AD
1140 1872 and AD 1944 eruptions (Balassone et al., 2013) are reported.

1141

1142

1143 LIST OF TABLE CAPTIONS

1144

1145 Table 1. Labels, occurrence, host rock and mineral assemblages of the Somma-Vesuvius samples.

1146 Sample provenances (see Fig. 1) are Somma Vesuviana and Ercolano (Avellino eruption) and
1147 Terzigno (Pompei eruption).

1148

1149 Table 2. Whole rock composition of the investigated mica-bearing xenoliths (major and minor
1150 elements in wt. % and trace elements in ppm).

1151

1152 Table 3. Chemical composition (EPMA, mean values on multiple point analyses expressed in wt%
1153 of oxides) of the investigated micas. Estimated values of formation temperature for selected
1154 samples are also given (see text for further explanation).

1155

1156 Table 4. Experimental details and relevant crystallographic data of the investigated micas.

1157

1158 Table 5. Selected bond distances (Å) and distortional parameters (calculated as reported in
1159 Lacalamita et al., 2011) of the studied mica crystals.

1160

1161 Table 6. Structural formulas in atoms per formula unit (apfu) of the selected micas for the
1162 crystallographic and crystal-chemical characterization.

1163

1164 Table 7. Mean atomic numbers (electrons, e^-) and average distances (Å) of cation sites derived by
1165 structure refinement (X-ref), and chemical data (EMPA) of the investigated mica crystals.
1166 Average error for mean atomic numbers is $\pm 0.5 e^-$.

1167

1168

1169 **SUPPLEMENTAL MATERIALS**

1170

1171 Supplemental Figure S1. Selected oxides of major and minor elements against SiO₂ and CaO (wt%)
1172 of Somma-Vesuvius xenoliths. Symbols as in Fig. 4.

1173

1174 Supplemental Figure S2. Selected trace elements plots against CaO (wt%) of Somma-Vesuvius
1175 xenoliths. Symbols as in Fig. 4.

1176

1177 Supplemental Figure S3. Plot of octahedral Al (a), Ti (b), and F (c) contents against Mg# for the
1178 studied micas (symbols as in Fig. 4).

1179

1180 Supplemental Figure S4. Plots of F vs. Ba for the studied micas (symbols as in Fig. 4).

1181

1182 Supplemental Figure S5. Plots of Shift_{M2} (Å) parameter vs. Ti (a) and F⁻ (b) contents (apfu) and of α
1183 (°) parameter vs. Ti (c) and F⁻ (d) contents for the investigated trioctahedral micas from the
1184 Avellino and Pompeii xenoliths (see Tables 3 and 5). For comparison, values of phlogopite,
1185 fluorophlogopite, and biotite of Vesuvius xenoliths from the AD 1631, AD 1872 and AD
1186 1944 eruptions (Balassone et al., 2013) are reported (symbols as in Fig. 4).

1187

1188

1189 Supplemental Table S1. Chemical composition (wt% oxide) of miscellaneous silicates in the
1190 studied ejecta (mineral symbols as in Table 1; AV = Avellino eruption, PP = Pompeii
1191 eruption).

1192

1193 Supplemental Table S2. Chemical composition of phosphates and arsenates (wt% oxide) in the
1194 studied samples (mineral symbols as in Table 1; AV = Avellino eruption, PP = Pompeii
1195 eruption).

1196

1197 Supplemental Table S3. Chemical composition of oxides (wt% oxide) and sulfides (wt% elements)
1198 detected in the studied samples (mineral symbols as in Table 1; AV = Avellino eruption, PP =
1199 Pompeii eruption).

1200

1201 Supplemental Table S4. Chemical composition of carbonates (wt% oxide) in the studied samples

1202 (mineral symbols as in Table 1; AV = Avellino eruption, PP = Pompeii eruption).

1203

1204 **CIF file:** CIF FILES SV PHL.

1205

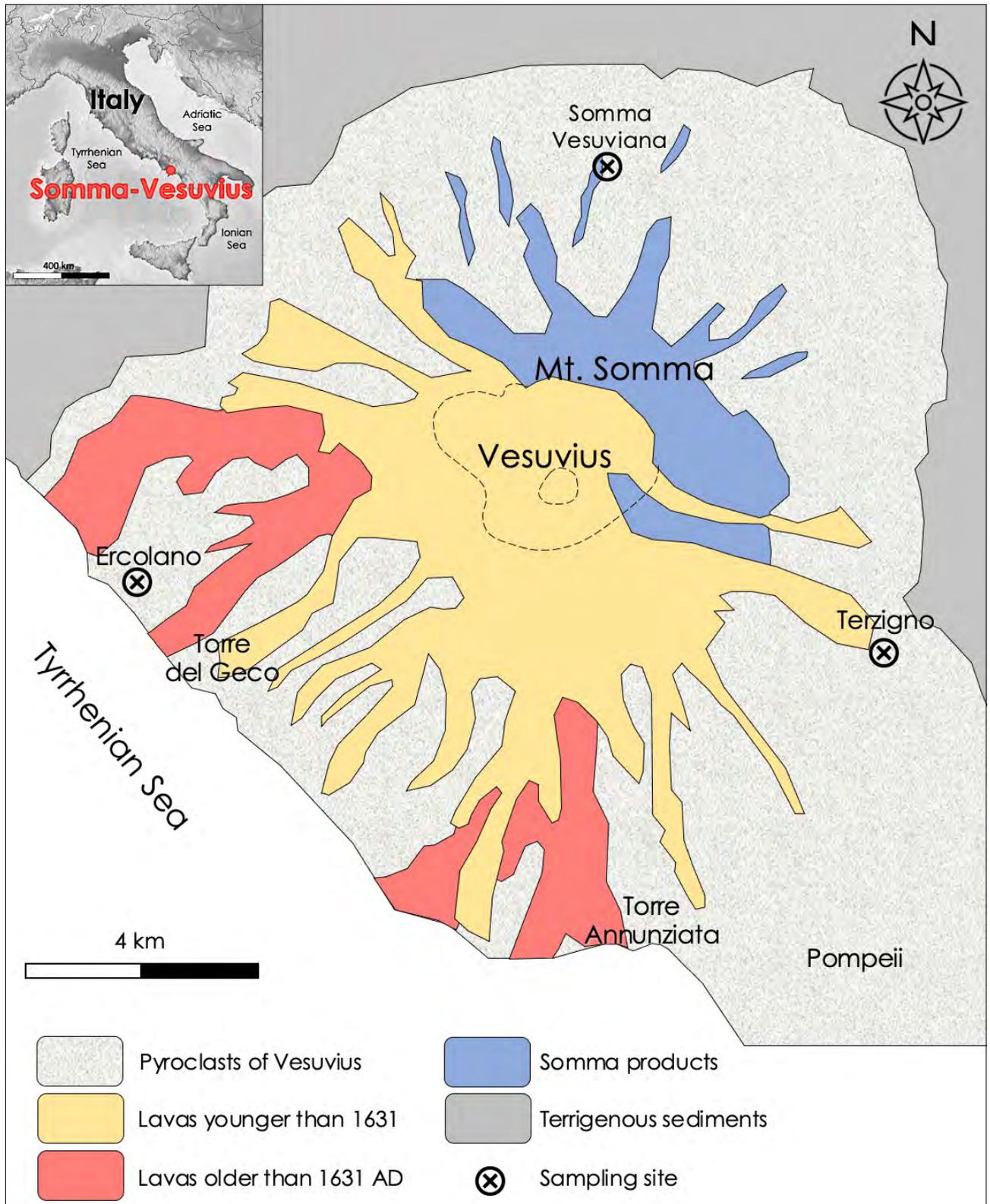


Figure 1

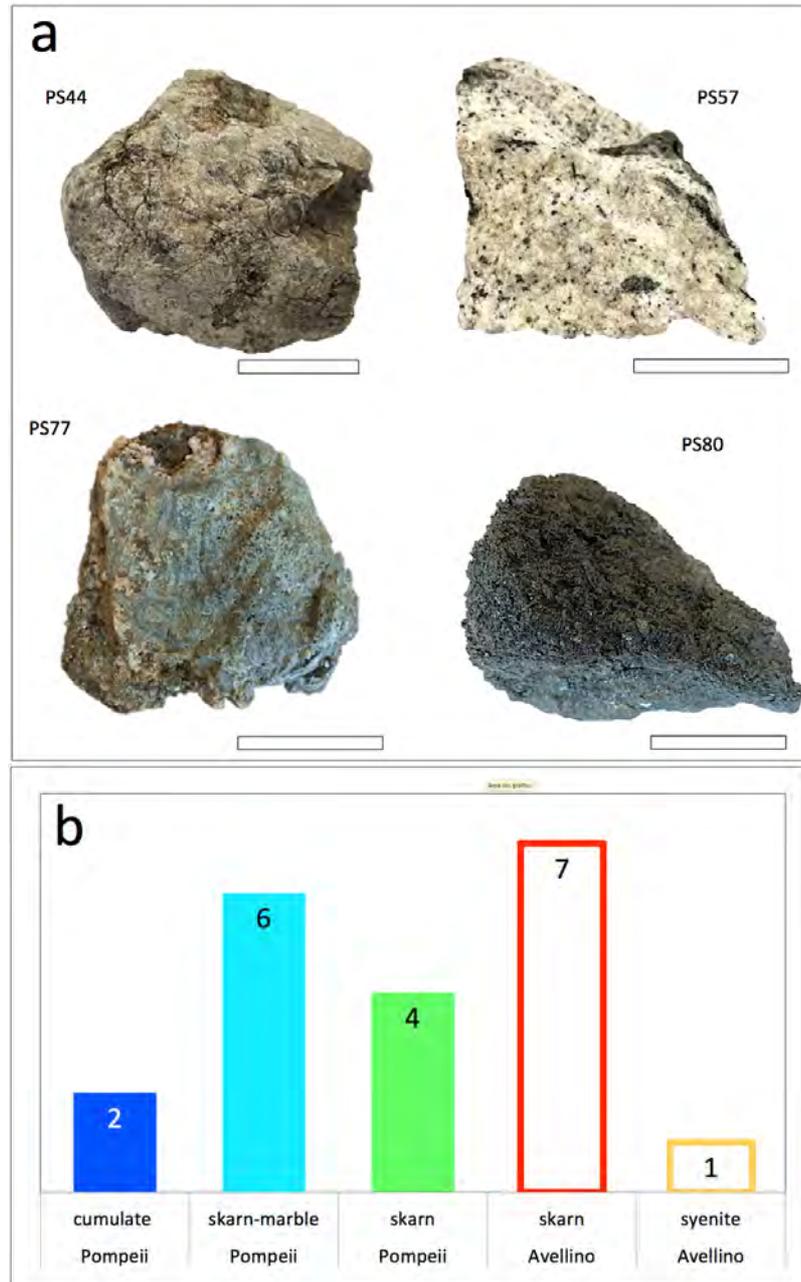


Figure 2

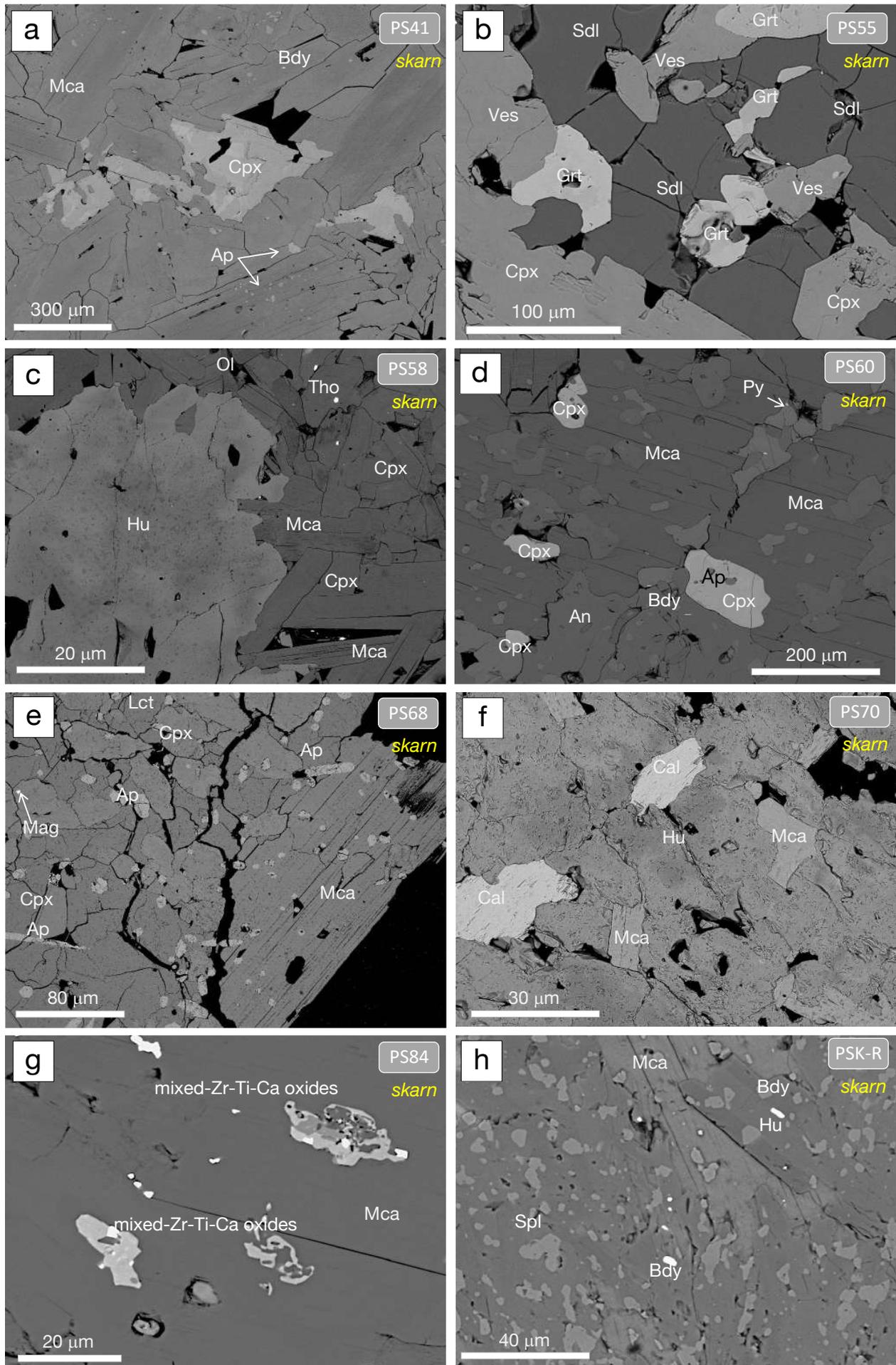


Fig. 3 (continued)

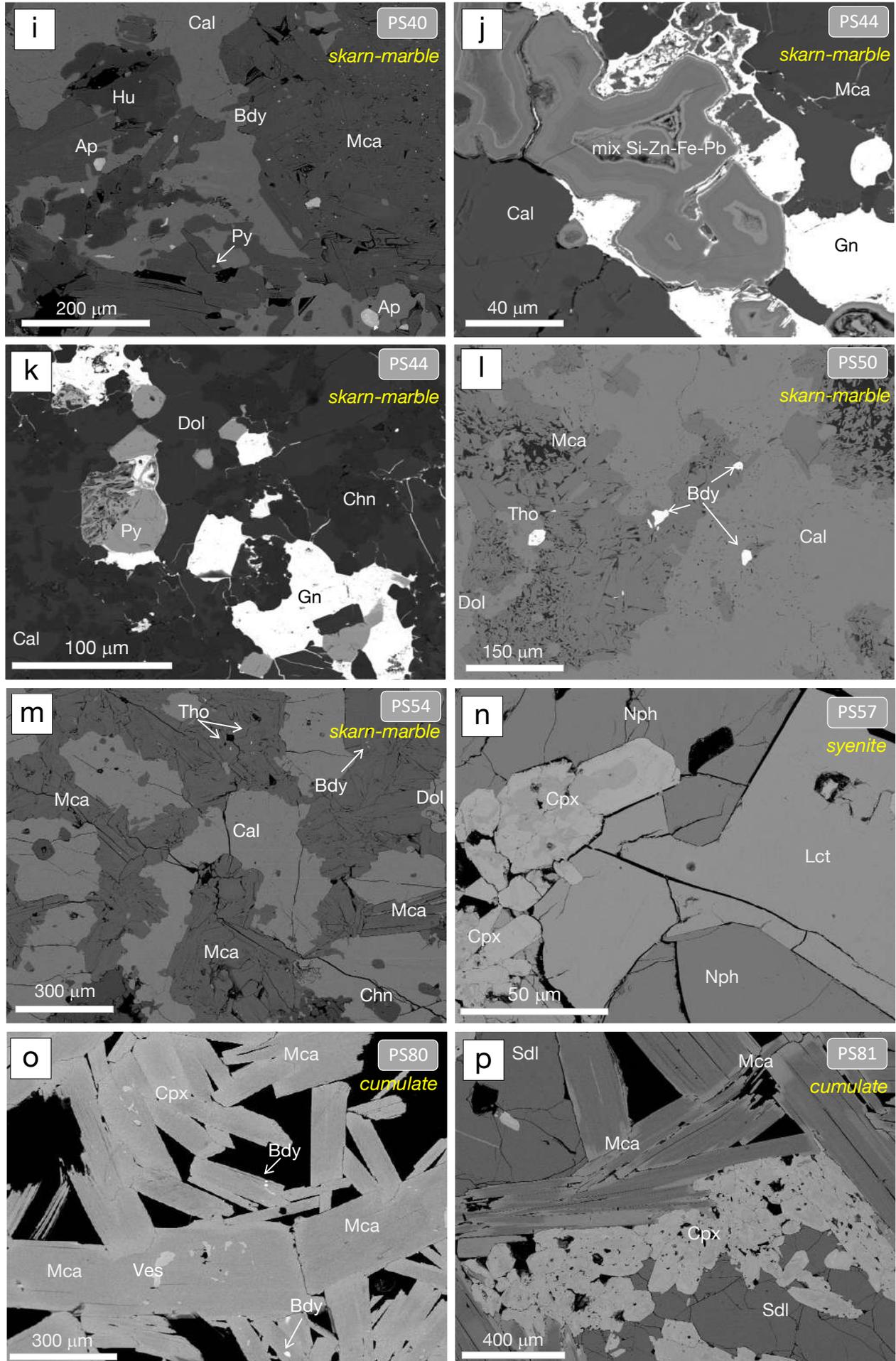


Fig. 3 (end)

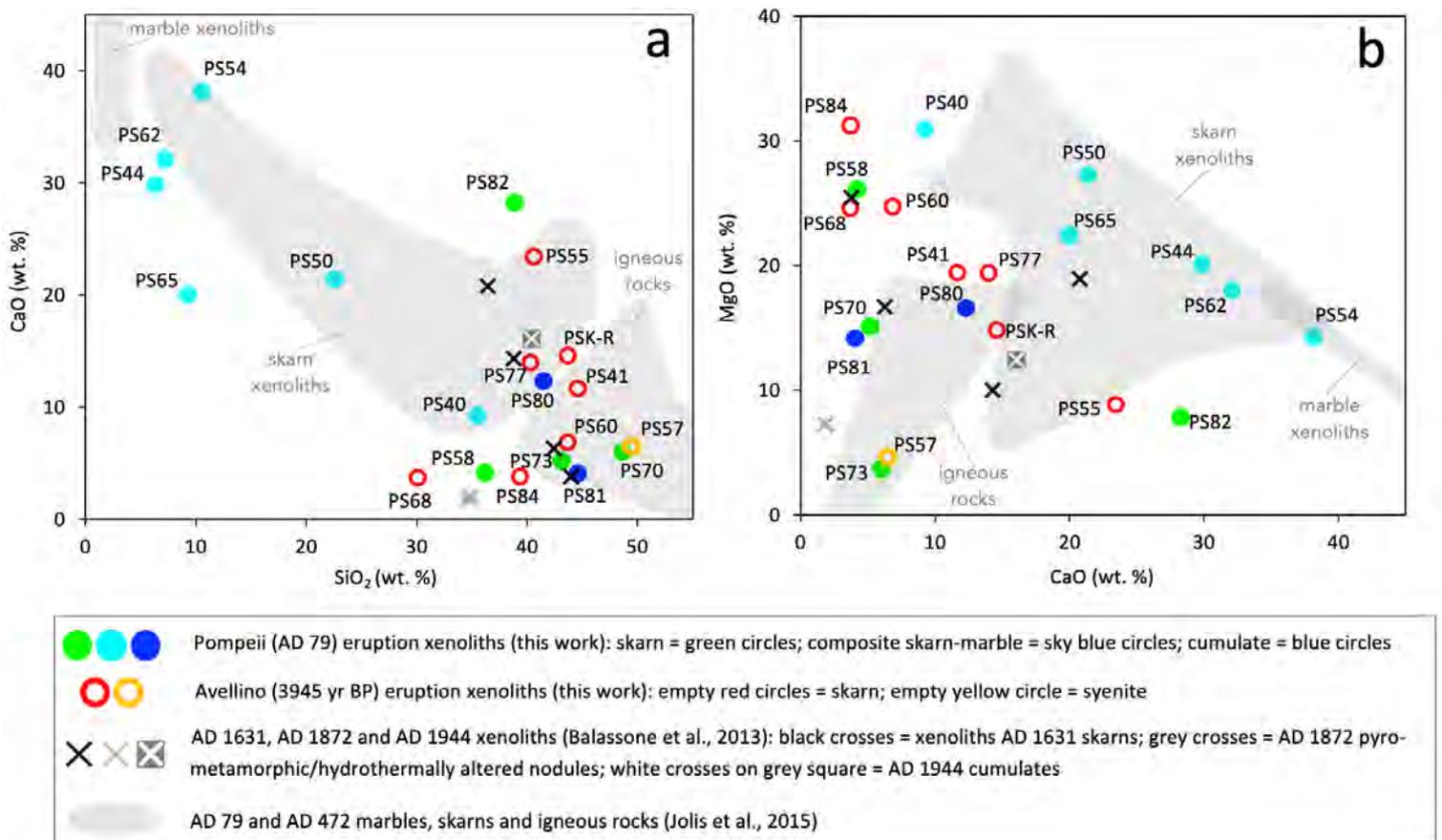


Figure 4

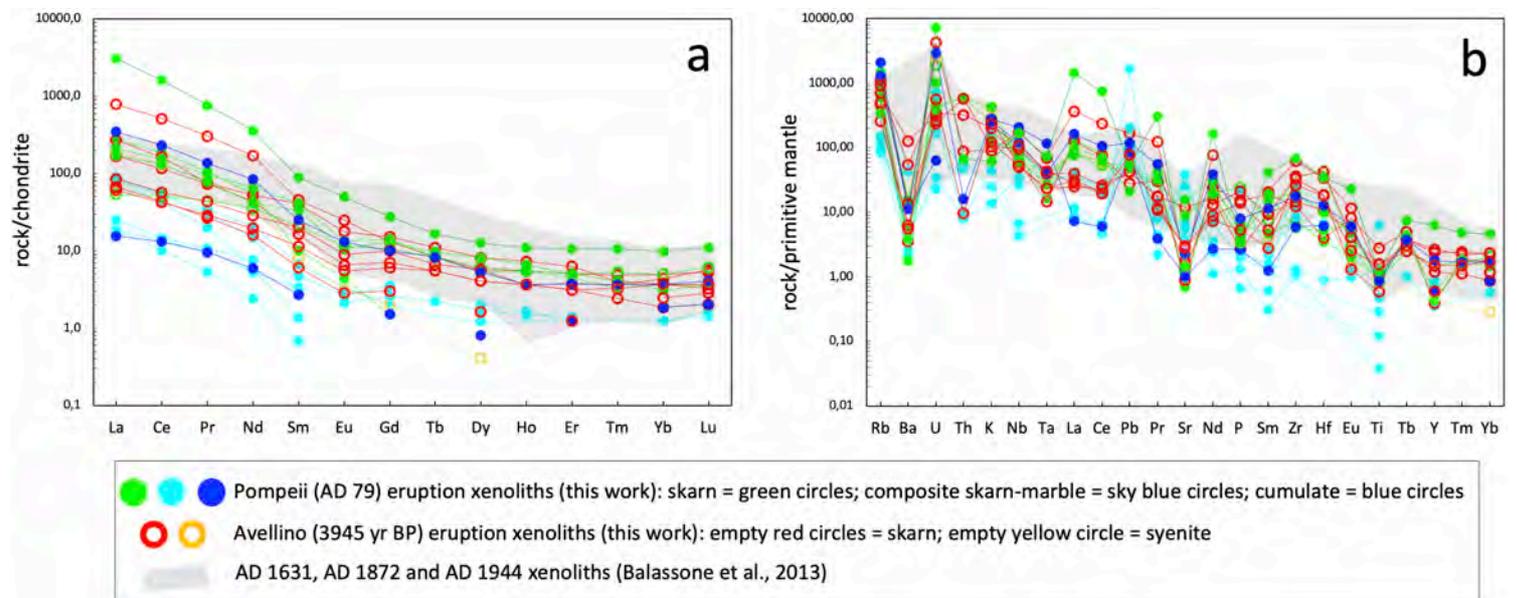


Figure 5

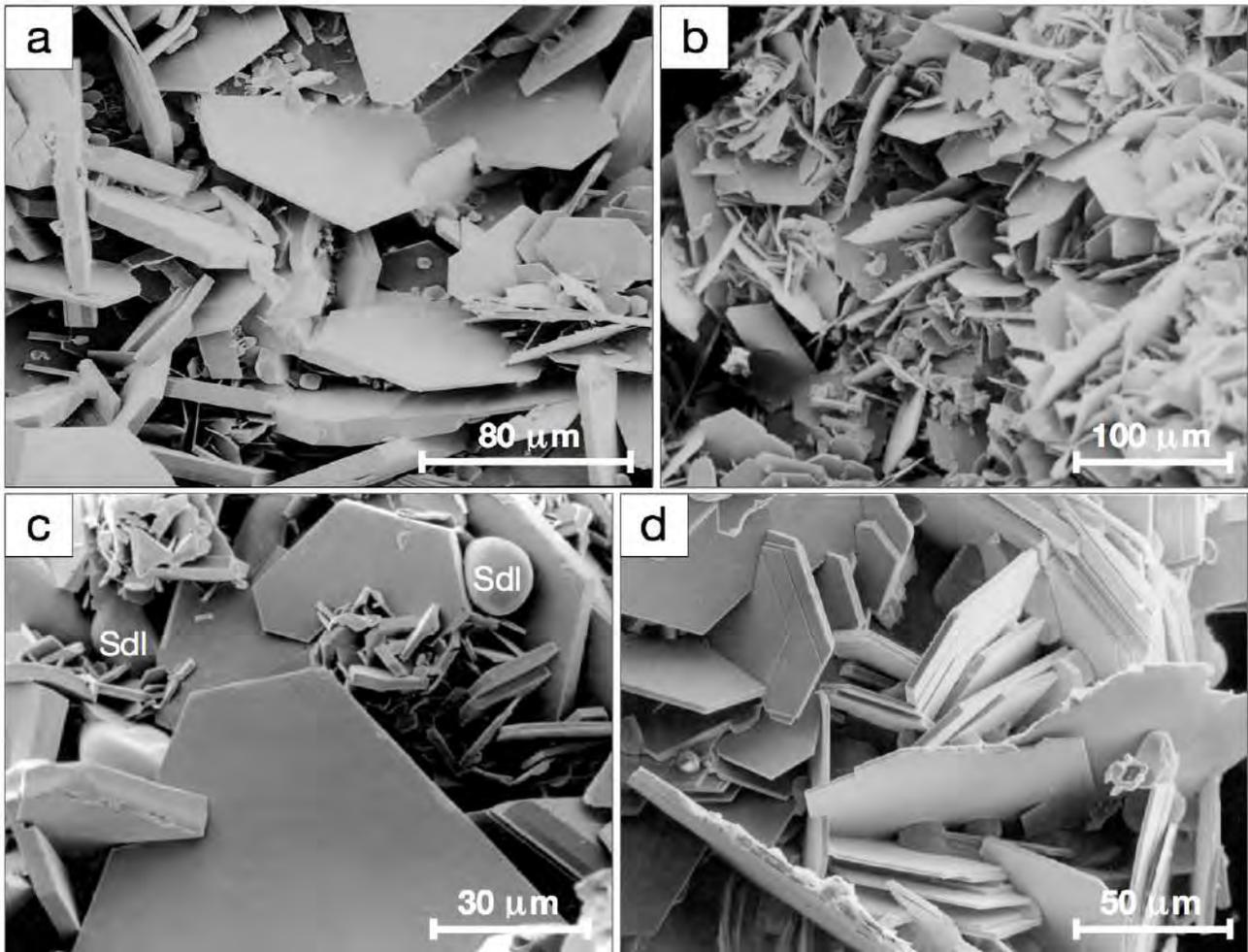


Figure 6

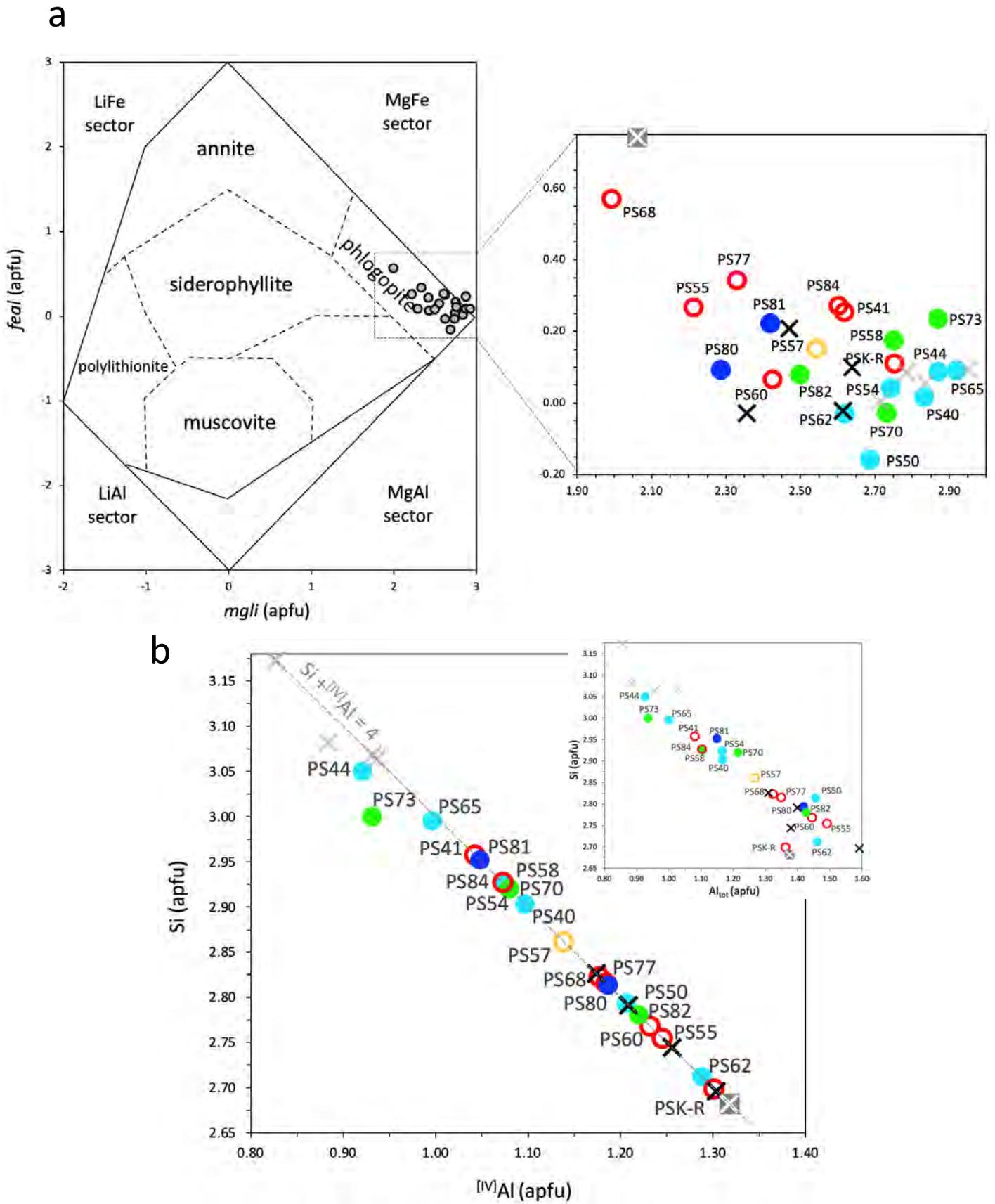


Figure 7

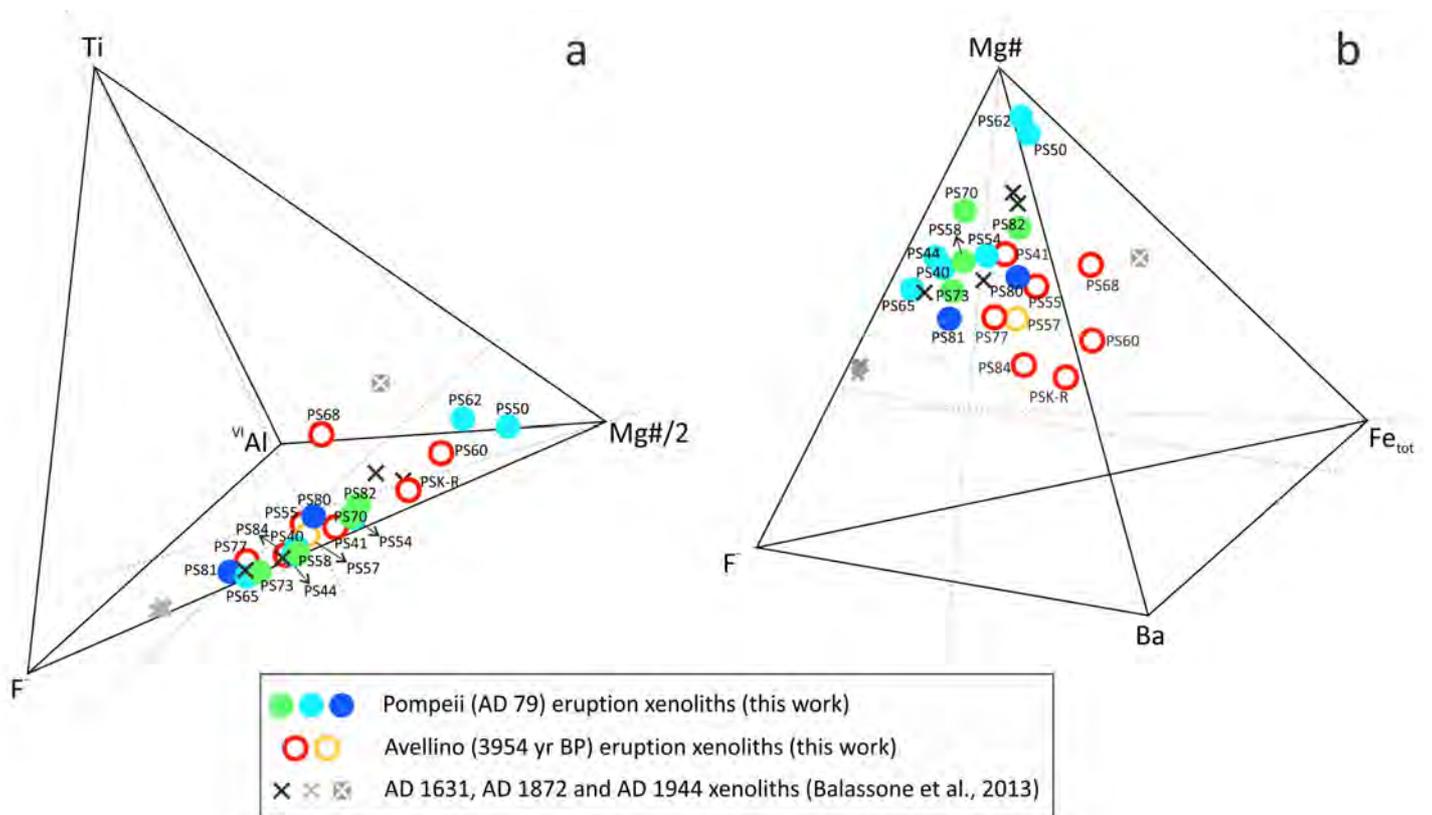


Figure 8

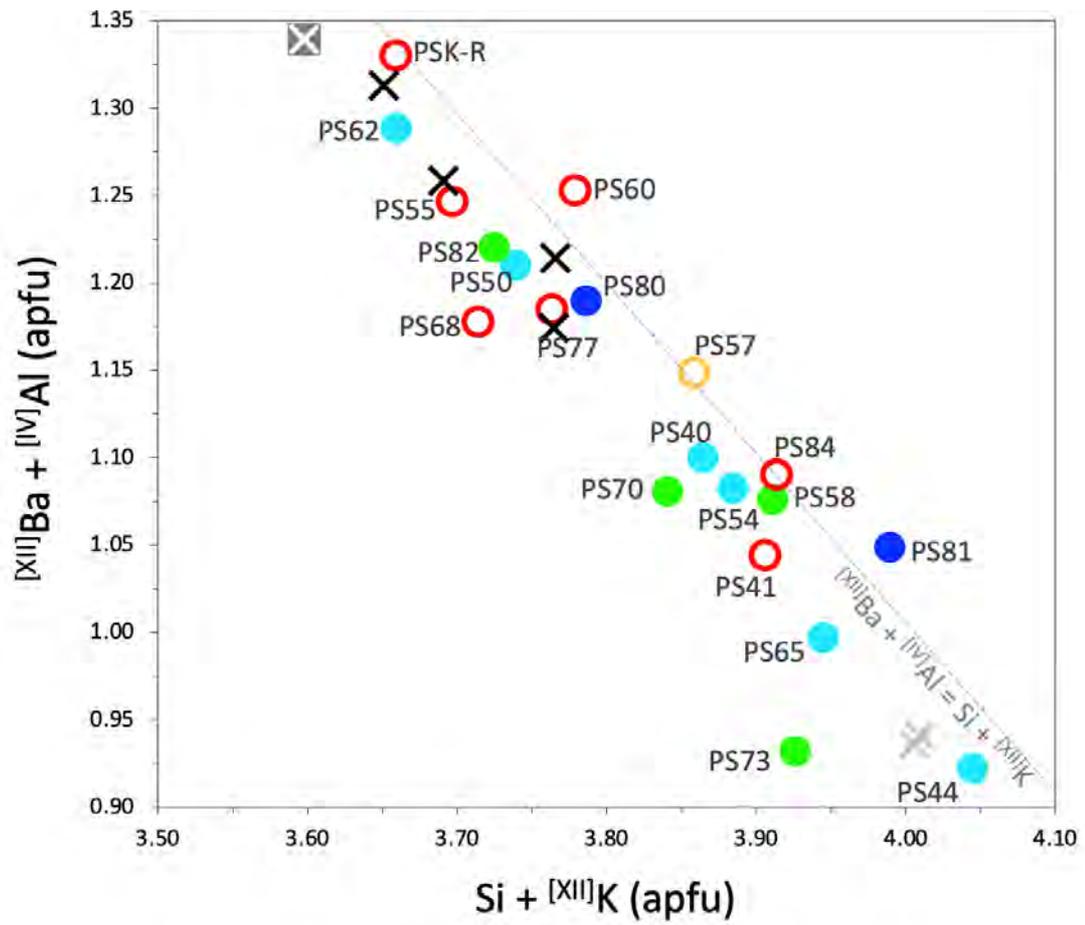


Figure 9

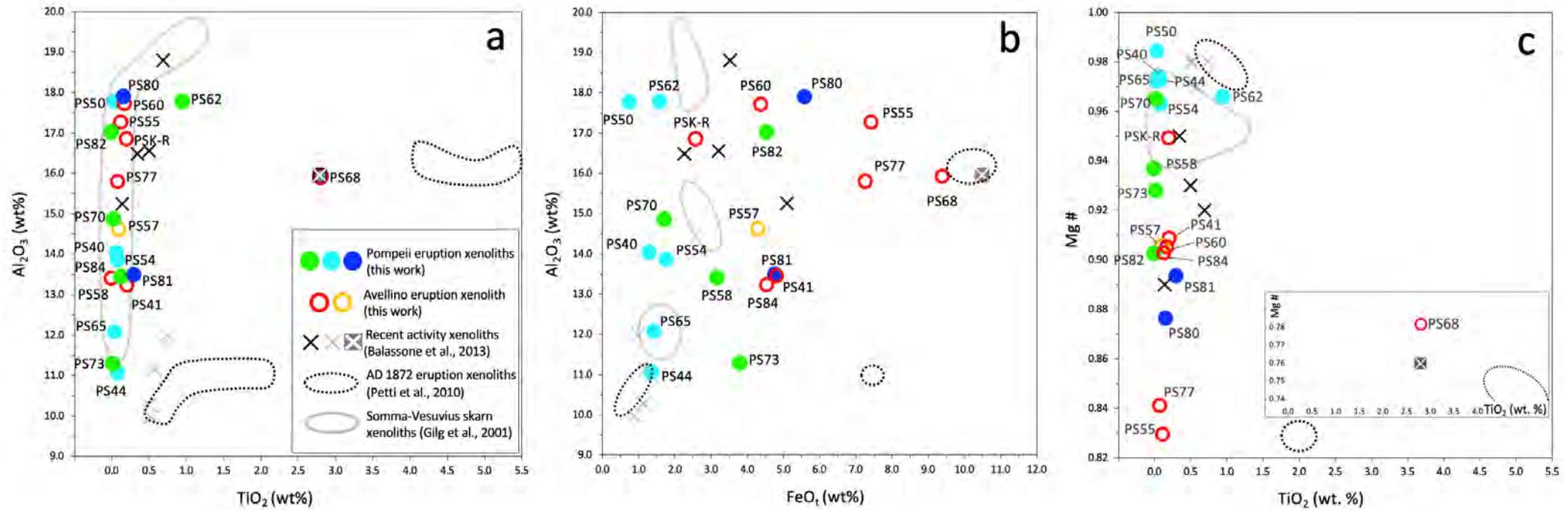


Figure 10

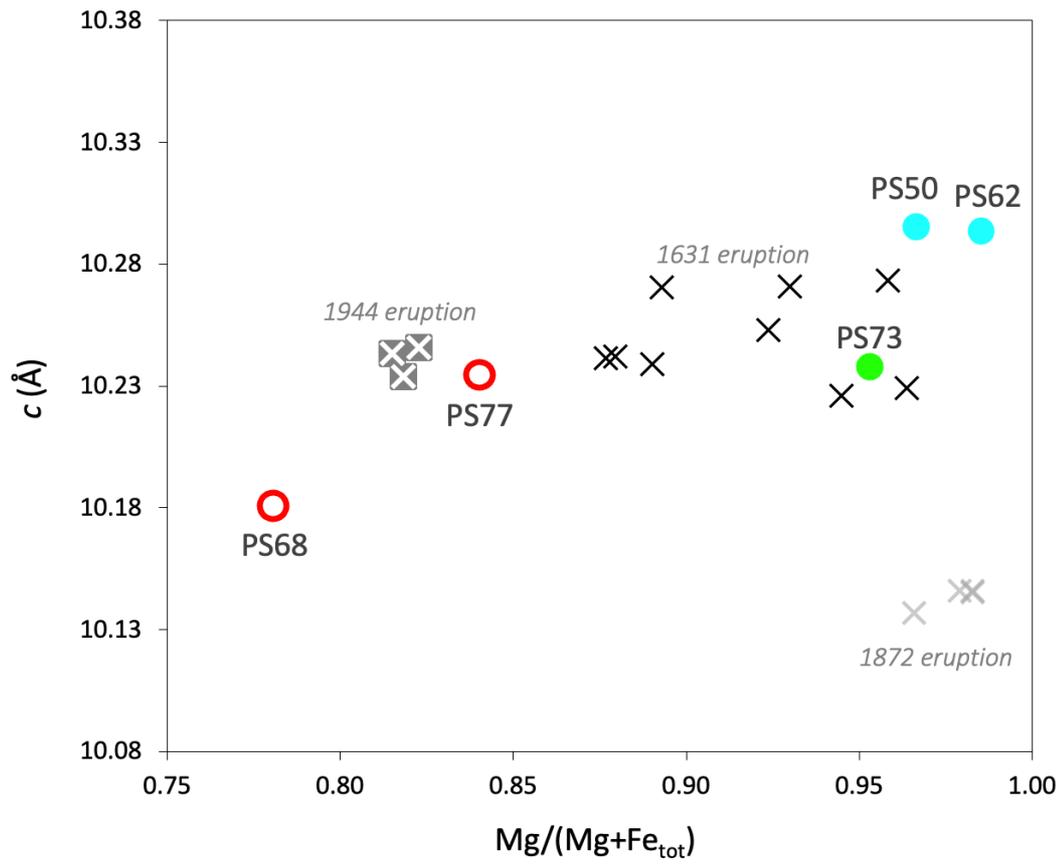


Figure 11

Table 4

Experimental details and relevant crystallographic data of the studied crystals.

	Avellino xenoliths		Pompeii xenoliths		
	PS68	PS77	PS50	PS62	PS73
<i>Crystallographic data</i>					
Space group	<i>C2/m</i>	<i>C2</i>	<i>C2/m</i>	<i>C2/m</i>	<i>C2/m</i>
<i>a</i> (Å)	5.3169(1)	5.3218(1)	5.3072(2)	5.3055(1)	5.3188(1)
<i>b</i> (Å)	9.2041(2)	9.2188(4)	9.1945(3)	9.1893(1)	9.2104(1)
<i>c</i> (Å)	10.1803(2)	10.2343(2)	10.2932(4)	10.2951(2)	10.2376(1)
β (°)	99.983(2)	100.011(2)	99.942(2)	99.902(1)	99.978(1)
Cell volume (Å ³)	490.65(2)	494.46(2)	494.74(3)	494.45(2)	493.94(1)
Z	4	4	4	4	4
<i>Experimental details</i>					
Reflections collected	5033	6066	3816	6048	6117
Reflections unique	1256	1829	1292	1805	1787
R _{merging} [R _(int)] (%)	2.79	2.76	2.26	1.79	1.86
Reflections used (<i>I</i> > 3σ(<i>I</i>))	953	1333	1124	1585	1432
No. of refined parameters	56	71	58	56	56
Goof ^a	1.087	0.983	0.965	0.947	1.045
R ₁ ^b (%)	2.66	2.85	3.27	1.98	2.63
wR ₂ ^c (%)	3.23	4.23	4.96	2.39	3.50
$\Delta\rho_{\min}/\Delta\rho_{\max}$ (e ⁻ /Å ³)	-0.41/0.71	-0.51/1.01	-0.53/0.96	-0.44/0.49	-0.39/1.38
a: Goodness-of-fit = $[\sum[w(F_o^2 - F_c^2)^2]/(N-p)]^{1/2}$, where <i>N</i> and <i>p</i> are the number of reflections and parameters, respectively.					
b: $R_1 = \sum[F_o - F_c]/\sum F_o $.					
c: $wR_2 = [\sum[w(F_o - F_c)^2]/\sum[w(F_o)^2]]^{1/2}$; <i>w</i> = Chebyshev optimized weights.					

Table 5

Selected bond distances (Å) and distortional parameters (calculated as reported in Lacalamita et al. 2011) of the studied crystals.

	Avellino xenoliths		Pompeii xenoliths		
	PS68	PS77	PS50	PS62	PS73
T-O1	1.656(2)	1.660(2)	1.664(2)	1.6660(5)	1.6572(8)
T-O1'	1.656(2)	1.661(2)	1.666(2)	1.6669(5)	1.6577(8)
T-O2	1.656(1)	1.6594(9)	1.666(1)	1.6668(3)	1.6584(5)
T-O3	1.670(2)	1.661(1)	1.660(2)	1.6681(5)	1.6504(7)
<T-O>	1.660(4)	1.660(3)	1.664(4)	1.6670(9)	1.656(1)
M1-O4(x2)	2.031(2)	2.048(2)	2.043(3)	2.0420(6)	2.0454(8)
M1-O3(x4)	2.086(1)	2.083(1)	2.079(2)	2.0737(5)	2.0853(6)
<M1-O>	2.068(2)	2.071(2)	2.067(4)	2.0631(8)	2.072(1)
M2-O4(x2)	1.987(2)	2.057(1)	2.057(2)	2.0495(5)	2.0558(6)
M2-O3(x2)	2.081(1)	2.071(1)	2.063(2)	2.0587(5)	2.0741(7)
M2-O3'(x2)	2.112(2)	2.071(1)	2.077(2)	2.0721(5)	2.0848(6)
<M2-O>	2.060(3)	2.066(2)	2.065(3)	2.0601(9)	2.072(1)
<M-O>	2.063(4)	2.068(3)	2.066(5)	2.061(1)	2.072(1)
K-O1(x4)	2.946(2)	2.962(2)	2.929(2)	2.9121(5)	2.9834(9)
K-O1'(x4)	3.346(2)	3.368(2)	3.428(2)	3.4566(6)	3.3380(9)
K-O2(x2)	2.948(3)	2.961(2)	2.926(3)	2.9111(7)	2.981(1)
K-O2'(x2)	3.362(3)	3.368(2)	3.429(3)	3.4594(9)	3.338(1)
<K-O> _{inner}	2.947	2.962(3)	2.928	2.912	2.9826
<K-O> _{outer}	3.351	3.368(3)	3.428	3.458	3.338
<K-O>	3.149		3.178	3.185	3.160
T _D -O1		1.65(4)			
T _D -O11		1.61(4)			
T _D -O2		1.72(5)			
T _D -O3		1.66(1)			
<T _D -O>		1.67(7)			
K _D -O1(x2)		2.31(1)			
K _D -O1'(x2)		2.307(6)			
K _D -O2(x2)		2.32(2)			
K _D -O2'(x2)		3.02(5)			
K _D -O11(x2)		2.96(4)			
K _D -O11'(x2)		3.36(4)			
<K _D -O> _{inner}					
<K _D -O> _{outer}					
α (°)	8.86	8.94	11.06	11.53	8.80
Volume _T (Å ³)	2.344	2.347	2.363	2.367	2.340
BLD _T	0.302	0.032	0.105	0.190	0.349
Volume _{M1} (Å)	11.556	11.638	11.581	11.491	11.665
BLD _{M(1)}	1.172	0.733	0.774	0.670	0.857
Volume _{M2} (Å)	11.433	11.608	11.552	11.446	11.659
BLD _{M(2)}	2.392	0.409	0.343	0.393	0.519
Shift _{M2} (Å)	0.062	-0.019	-0.019	-0.017	-0.016
Δ _{K-O} (Å)	0.404	0.406	0.500	0.520	0.400

Errors on distortion parameters, estimated by varying the refined positional parameters within one standard deviation are in the following ranges: < 0.5% for volumes and shifts; 0.1-13% for α and Δ_{K-O}.

Table 6

Structural formulas in atoms per formula unit (apfu) of the selected micas for the crystallographic and crystal-chemical characterization.

Sample		
Avellino xenoliths	PS68	$(K_{0.89}Na_{0.04})_{\Sigma=0.93}(Mg_{2.03}Fe^{2+}_{0.09}Fe^{3+}_{0.48}Mn_{0.01}Al_{0.23}Ti_{0.15})_{\Sigma=2.99}(Si_{2.88}Al_{1.12})_{\Sigma=4.00}O_{10.83}F_{0.25}Cl_{0.01}OH_{0.91}$
	PS77	$(K_{0.94}Na_{0.06})_{\Sigma=1.00}(Mg_{2.32}Fe^{2+}_{0.37}Fe^{3+}_{0.07}Mn_{0.06}Al_{0.16})_{\Sigma=2.98}(Si_{2.80}Al_{1.20})_{\Sigma=4.00}O_{10.03}F_{0.71}OH_{1.26}$
Pompeii xenoliths	PS50	$(K_{0.95}Na_{0.08}Ca_{0.02})_{\Sigma=1.05}(Mg_{2.70}Fe^{3+}_{0.04}Al_{0.24}Mn_{0.01})_{\Sigma=2.99}(Si_{2.81}Al_{1.19})_{\Sigma=4.00}O_{10.16}OH_{1.84}$
	PS62	$(K_{0.95}Na_{0.02})_{\Sigma=0.97}(Mg_{2.62}Fe^{3+}_{0.09}Al_{0.18}Cr_{0.05}Ti_{0.05})_{\Sigma=2.99}(Si_{2.71}Al_{1.29})_{\Sigma=4.00}O_{10.10}F_{0.04}OH_{1.86}$
	PS73	$(K_{0.92}Na_{0.06})_{\Sigma=0.98}(Mg_{2.85}Fe^{3+}_{0.14}Mn_{0.01})_{\Sigma=3.00}(Si_{2.98}Al_{0.93}Fe^{3+}_{0.09})_{\Sigma=4.00}O_{10.10}F_{0.69}OH_{1.21}$

Table 7

Mean atomic numbers (electrons, e^-) and average distances (Å) of cation sites derived by structure refinement (X-ref), and chemical data (EMPA). Average error for mean atomic numbers is $\pm 0.5 e^-$.

	Avellino xenoliths		Pompeii xenoliths		
	PS68	PS77	PS50	PS62	PS73
T e^- X-ref	14.00	14.24	14.00	14.00	14.00
T e^- EMPA	13.72	13.70	13.70	13.68	14.04
e^- (M1+2M2) X-ref	45.34	41.47	36.94	39.01	38.07
e^- (M1+2M2) EMPA	45.72	42.86	36.91	38.42	38.09
K e^- X-ref	17.92	17.61	16.89	18.89	17.23
K e^- EMPA	17.35	18.52	19.33	18.27	18.14
<T-O> X-ref	1.660	1.660	1.664	1.667	1.656
<T-O> EMPA	1.656	1.659	1.659	1.662	1.655
<M-O> X-ref	2.063	2.068	2.066	2.061	2.072
<M-O> EMPA	2.054	2.075	2.068	2.067	2.079