

1 **Revision 1**

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3 **Viscosity of Earth's inner core constrained by Fe–Ni interdiffusion in Fe–Si**  
4 **alloy in an internal-resistive-heated diamond anvil cell**

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17

18 **ABSTRACT**

19 Diffusivity in iron (Fe) alloys at high pressures and temperatures imposes constraints on  
20 transport properties of the inner core, such as viscosity. Because silicon (Si) is among the  
21 most likely candidates for light elements in the inner core, the presence of Si must be  
22 considered when studying diffusivity in the Earth's inner core. In this study, we conducted  
23 diffusion experiments under pressure up to about 50 GPa using an internal-resistive-heated  
24 diamond anvil cell (DAC) that ensures stable and homogeneous heating compared with a  
25 conventional laser-heated DAC and thus allows us to conduct more reliable diffusion  
26 experiments under high pressure. We determined the coefficients of Fe–nickel (Ni)  
27 interdiffusion in the Fe–Si 2 wt.% alloy. The obtained diffusion coefficients follow a  
28 homologous temperature relationship derived from previous studies without considering Si.  
29 This indicates that the effect of Si on Fe–Ni interdiffusion is not significant. The upper limit  
30 of the viscosity of the inner core inferred from our results is low, indicating that the Lorentz  
31 force is a plausible mechanism to deform the inner core.

32 **Keywords:** Earth's inner core, diffusion, viscosity, iron, silicon, high pressure

33

## INTRODUCTION

34 Viscosity is a crucial physical property to understand dynamical processes in planetary  
35 interiors. The viscosity of the Earth's inner core constrains the modes and mechanisms of  
36 viscous flows responsible for the observed seismic anisotropy of the inner core (Lasbleis and  
37 Deguen 2015). Seismological observations of the inner core reveal its complicated  
38 anisotropic structures; the P wave travels ~3–4% faster along the polar direction compared to  
39 its equatorial direction (e.g., Poupinet et al. 1983). This anisotropy of the inner core can be  
40 attributed to lattice preferred orientation (LPO) induced by the viscous flow of inner core  
41 materials (Deguen 2012; Romanowicz and Wenk 2017). The Rayleigh–Bénard convection  
42 was first proposed as a possible mechanism of viscous flow in the inner core (Jeanloz and  
43 Wenk 1988). Yoshida et al. (1996) argued that the inner core grows preferentially to the  
44 equatorial direction due to the columnar convection of the outer core. Such heterogeneous  
45 inner core growth leads to isostatic disequilibrium, and the resultant differential stress gives  
46 rise to a viscous flow in the inner core. External forces related to the Earth's magnetic field,  
47 such as the Lorentz force, or force induced by heterogeneous joule heating, have been  
48 proposed as possible mechanisms for the viscous flow in the inner core (Karato 1999;  
49 Takehiro 2011). The strength and mode of viscous flows (i.e., the plausibility of the proposed  
50 mechanisms) heavily depend on the viscosity of the inner core; however, there is significant  
51 uncertainty in its estimates, which range from  $10^{10}$  Pa s to  $10^{22}$  Pa s (Buffett 1997; Davies et  
52 al. 2014; Frost et al. 2021; Jackson et al. 2000; Koot and Dumberry 2011; Reaman et al. 2011;  
53 Ritterbex and Tsuchiya 2020; Van Orman 2004; Yoshida et al. 1996). These studies estimated  
54 the viscosity of the inner core either from experiments, *ab initio* calculations, or geophysical  
55 observations.

56 From a mineral physics point of view, the approach often adopted is an estimate of the  
57 viscosity based on the diffusion coefficient of iron (Fe) under high pressure (Yunker and Van  
58 Orman 2007; Reaman et al. 2012; Ritterbex and Tsuchiya 2020). Terrestrial cores are thought  
59 to be composed of Fe alloyed with nickel (Ni) and some light elements. Solid Fe alloys  
60 composing the Earth's inner core assume a hexagonal closed packed (hcp) structure, while  
61 pressure and temperature conditions of the centers of smaller terrestrial planets, such as  
62 Mercury and Mars, favor face-centered cubic (fcc) structures of iron alloys as a dominant  
63 phase of their possible solid inner cores, depending on light element concentrations

64 (Komabayashi et al. 2019; Tsujino et al. 2013). Although the self-diffusivity of Fe under high  
65 pressure is a critical limiting factor of crystal plasticity of inner core materials, the self-  
66 diffusion coefficient of Fe under pressures relevant to the deep Earth has not been studied  
67 experimentally due to experimental difficulties. Instead, experimental attempts have been  
68 made to estimate the effects of pressure on Fe–Ni interdiffusion coefficients as an analogy of  
69 Fe self-diffusion coefficients (Reaman et al. 2012; Yunker and Van Orman 2007).

70 An open question unaddressed by previous studies is the effect of light elements on Fe–Ni  
71 interdiffusion coefficients, as the Earth’s inner core must contain ~1–3 wt.% light elements  
72 (Badro et al. 2007; Fei et al. 2016). In such alloys, the diffusivity of solvent atoms (Fe in this  
73 case) is subject to change by the presence of solute elements. One of the likely light elements  
74 to explain the density deficit and velocity reduction of the inner core with respect to pure Fe  
75 is silicon (Si) (Antonangeli et al. 2010; Mao et al. 2012; Sakairi et al. 2018). Therefore,  
76 the diffusivity in Fe–Si alloys under high pressure is of great importance for estimating  
77 diffusivity in the Earth’s inner core.

78 In this study, we conducted Ni diffusion experiments on an Fe–2 wt.% Si alloy to study the  
79 effect of Si on the Fe–Ni interdiffusion coefficient for pressures up to ~50 GPa. We used an  
80 internal-resistive-heated DAC (IRHDAC), where the sample is heated with smaller  
81 temperature gradients and lower temporal fluctuations than in the conventional laser-heated  
82 DAC (LHDAC). We compare our results with previous studies on the Fe–Ni interdiffusion  
83 coefficient (Reaman et al. 2012; Yunker and Van Orman 2007) and further discuss its  
84 implications on the diffusivity and viscosity of the Earth’s inner core.

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## EXPERIMENTAL METHODS

87 We prepared diffusion couples by coating a thin film of Ni (3N purity level) on foils of the  
88 Fe–2 wt.% Si alloy using a magnetron sputtering system. The thickness of the coated Ni layer  
89 is approximately 300 nm. The Ni coated Fe–Si foils were further fabricated by a focused ion  
90 beam (FIB) apparatus (FEI Versa 3D) to make diffusion couples whose size fits the sample  
91 chamber of a DAC. We place the prepared diffusion couple into an IRHDAC with single-  
92 crystal sapphire and alumina powder as pressure medium and thermal insulator against  
93 diamond anvils (Fig. 1). We used diamond anvils with a culet size of 300  $\mu\text{m}$  for pressure

94 generation. The surface of the preindented Re gasket was electrically insulated by cBN+TiO<sub>2</sub>  
95 powder and cement (Resbond<sup>TM</sup> 919), similarly to the literature (Inoue et al. 2020; Suehiro et  
96 al. 2019). Two Pt foils on the insulation layer act as electrical leads connecting the Ni+Fe–Si  
97 diffusion couple inside the sample chamber and Cu wires outside of the gasket, which creates  
98 a circuit for resistive heating of the diffusion couple itself.

99 After compression to the pressure of interest, electricity was supplied to the diffusion  
100 couple from a DC electric power supply (TAKASAGO ZX-400LA) through the prepared  
101 circuit. The temperature of diffusion couple during joule heating is determined from the  
102 thermal radiation spectrum to fit Planck's law. The thermal radiation emitted from the heated  
103 sample is collected through the refractive lens and reflected by dichroic and silver mirrors,  
104 and finally was recorded with a Princeton grating spectrometer combined with a CCD array  
105 detector. The system optical response was calibrated with a certified tungsten lamp. For the  
106 data reduction, we used the LightField software package. Target temperatures were reached  
107 within 10 seconds and maintained during diffusion experiments (Table 1). The temperature  
108 error from fitting of a spectrum is typically less than 0.3%. 1-dimensional temperature  
109 distributions of the specimens were obtained (Fig. 2) and errors of our temperature  
110 measurements are estimated from standard deviation of the temporal and radial temperature  
111 fluctuation of the specimens. After diffusion experiments, the specimens were quenched  
112 rapidly by shutting off the electric power supply. Pressures inside the sample chamber were  
113 determined by the Raman shift of a diamond anvil (Akahama and Kawamura 2006), and  
114 thermal pressure during heating was corrected using an estimation from Sinmyo et al. (2019).  
115 The overall uncertainty in pressure may be less than 5%.

116 After the diffusion experiments, samples were recovered to ambient condition, and thin  
117 cross-sections across the heated region of the diffusion couple were prepared by FIB for  
118 transmission electron microscope (TEM) analysis. We further conducted microstructural  
119 observations and chemical analysis during sample preparation using the scanning ion  
120 microscope (SIM) and energy dispersive X-ray spectroscopy (EDS) mounted on the FIB  
121 apparatus. After sample preparation, microstructural observations and chemical analysis were  
122 conducted using TEM (JEOL JEM-2800). The acceleration voltage was 200 kV, and the  
123 probe size was about 1.0 nm.

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125

## RESULTS AND DISCUSSION

126 We performed four separate diffusion experiments, in which pressure–temperature  
127 conditions were in the stability field of the fcc phase of Fe–2 wt.% Si alloy according to  
128 Komabayashi et al. (2019) (Table 1). Temperature variations along the radial direction of the  
129 heated area were at most ~50 K (Fig. 2). Moreover, the temperature difference between both  
130 sides of the specimen and temperature fluctuation during the diffusion experiment was at  
131 most ~100 K. The temperature homogeneity and heating stability of our method were  
132 significantly improved compared to LHDAC (Reaman et al. 2012), and thus desirable for  
133 precise diffusion experiments.

134 The SIM observation during FIB cross-sectioning allows us to observe grain boundaries in  
135 the specimen after the diffusion experiment (Fig. 3a). The SIM images show the boundary  
136 between Fe–Si alloy and thin Ni layer and it can be observed that thickness of the Ni layer is  
137 reduced from initial ~300 nm to 100–150 nm after high-pressure experiment (Fig. 3a). The  
138 sample texture shown in Fig. 3a can be attributed to the phase transformation of Fe alloys  
139 from fcc to hcp after the temperature quench and the martensitic phase transformation from  
140 hcp to bcc during decompression. Large domains having a similar orientation of twin  
141 boundaries can be considered as former fcc grains during diffusion experiments. Domains of  
142 former fcc grains have sizes of several micrometers at the central part of the specimen,  
143 whereas grains have finer sub-micrometer sizes at the sample corners in contact with Pt  
144 electrodes. This grain size distribution in the specimen suggests grain growth of Fe alloys  
145 during diffusion experiments. Element mapping of the sample cross-sections obtained by  
146 EDS indicates the existence of high-diffusivity paths in the specimens, and the paths match  
147 boundaries of large domains observed in the SIM image, which is evidence of grain boundary  
148 diffusion (Fig. 3b). Contamination of specimens from Pt electrodes would be negligible as  
149 temperature of areas having contacts with Pt electrodes are relatively low. We also confirmed  
150 that there is no significant contamination of specimens from Pt electrodes after heating for 3  
151 hours under 1900 K and 40 GPa.

152 We further conducted TEM-EDS elements mapping analyses for all specimens and  
153 confirmed that the grain boundary diffusion is active in our experiments (Fig. 4). In some  
154 regions, Fe diffusion into grain boundaries of the Ni layer is also observed (Fig. 4c). Since  
155 the focus of this study is to determine lattice diffusion coefficient of Fe–Ni interdiffusion, the

156 TEM-EDS point analyses were performed in several areas within the quenched sample along  
157 lines perpendicular to the diffusion boundary, avoiding areas of significant grain boundary  
158 diffusion, based on Ni element mapping results (FIGURE 4b). Furthermore, diffusion profiles  
159 with evidence of grain boundary diffusion were excluded from the obtained line profile  
160 results. Fig. 5 shows two representative diffusion profiles obtained from separate  
161 experiments. The measurement error of Ni concentration is evaluated by the variation of the  
162 value of the diffusion profile obtained within the same sample, and it was found that the Ni  
163 concentration has a maximum error of only 2%. Although Run #2 experienced a shorter  
164 heating duration compared with Run #3, higher temperature and lower-pressure condition of  
165 Run #2 during the diffusion experiment makes its diffusion distance longer than that of Run  
166 #3. This reduction in the diffusion distance with increasing pressure and decreasing  
167 temperature is consistent with the known pressure and temperature dependence of diffusion  
168 (Mehrer 2007). The TEM-EDS analyses confirmed that the quenched samples had a Si  
169 content of 2 wt.%, the same composition as the starting material.

170 The diffusion of particles inside an anisotropic medium can be described by following  
171 Fick's second law:

$$172 \quad \frac{\partial C}{\partial t} = \nabla \cdot (D \nabla C), \quad (1)$$

173 where  $C$  is concentration of diffusion particles,  $D$  is diffusion coefficient, and  $t$  is time during  
174 which the diffusion has been taking place. Due to the limited sample sizes in this study, it is  
175 challenging to evaluate concentration dependence on diffusion coefficient. If we ignore the  
176 concentration effect on diffusivity, we can express Eq. (1) as Eq. (2):

$$177 \quad \frac{\partial C}{\partial t} = D \Delta C. \quad (2)$$

178 Because we have a thin layer of Ni (having thickness of 100-150 nm) as tracers in our  
179 experiments, it is appropriate to apply the 'thin film geometry' (Crank 1975). Applying the  
180 thin film geometry to Eq. (2) yields the following 'thin film solution':

$$181 \quad C(x, t) = \frac{M}{\sqrt{\pi D t}} \exp\left(-\frac{x^2}{4 D t}\right), \quad (3)$$

182 where  $M$  is the initial amount of the tracer, and  $x$  is diffusion distance.

183 We apply the thin film solution to the diffusion profiles obtained in this study (Fig. 5).

184 Multiple diffusion profiles obtained at several different regions were used to estimate the  
185 diffusion coefficients. We calculated the averages of diffusion coefficients estimated from  
186 diffusion profiles and used their standard deviation as the estimated error (Table 1). To  
187 validate the use of the thin film solution, we performed finite element method (FEM) analysis  
188 on diffusion in Run #3 by using commercial software COMSOL Multiphysics (COMSOL  
189 Inc.) (Fig. 6a). We assume diffusion of tracers having thickness of 120 nm for 400 seconds  
190 with diffusion coefficient of  $4.08 \times 10^{-17} \text{ m}^2 \text{ s}^{-1}$  estimated from the thin film solution. The result  
191 of finite element analysis reasonably reproduces the experimentally observed diffusion  
192 profile and support validity of using the thin film solution to estimate diffusion coefficients in  
193 this study (Fig. 6b).

194 Because the kinetics of diffusion are dictated by the thermal vibration of atoms, the  
195 diffusion coefficient can frequently be expressed by the following Arrhenius formula:

$$196 \quad D = D_0 \exp\left(-\frac{E_A + PV_A}{RT}\right), \quad (4)$$

197 where  $E_A$  is the activation energy,  $P$  is the pressure,  $V_A$  is the activation volume,  $R$  is the gas  
198 constant,  $T$  is the temperature, and  $D_0$  is a pre-exponential factor. According to Eq. (4),  
199 activation volume  $V_A$  can be derived from the slope of the logarithm of diffusion coefficients  
200 against pressure. We plotted the diffusion coefficients at similar temperature conditions  
201 against pressure and derived the activation volume using Eq. (4) (Fig. 7). Present diffusion  
202 coefficients are consistent with the linear fit line suggested by Yunker and Van Orman (2007)  
203 and relatively low compared to other fits, including the result of Reaman et al. (2012).  
204 While Yunker and Van Orman (2007) suggests the activation volume of  $3.1 \pm 0.7 \text{ cm}^3 \text{ mol}^{-1}$   
205 under a pressure of  $\sim 23 \text{ GPa}$ , Reaman et al. (2012) suggests a decrease in the activation  
206 volume to  $2.62 \text{ cm}^3 \text{ mol}^{-1}$  at  $65 \text{ GPa}$ . The decrease in the activation volume with increasing  
207 pressure is anticipated with the homologous temperature relationship and our current  
208 understanding of the high-pressure melting curve of Fe. However, the results of Reaman et al.  
209 (2012) might contain systematic errors in the temperature measurement, as they used  
210 LHDAC showing large temperature gradients and possible unstable heating. In contrast, a  
211 multi-anvil apparatus and a piston-cylinder apparatus employed by Yunker and Van Orman  
212 (2007) exhibit a rather homogeneous temperature distribution, high stability, and accuracy in  
213 the temperature measurement compared to LHDAC. The activation volume estimated solely

214 from our results is  $3.49 \pm 0.3 \text{ cm}^3 \text{ mol}^{-1}$ , which is consistent with the experimental result of  
215 Yunker and Van Orman (2007), and equal to their theoretical prediction based on the  
216 homologous scaling. Our results suggest that the diffusion process and lattice geometry of the  
217 specimens in this study approximately identical to those obtained in previous studies on Fe–  
218 Ni interdiffusion, and the effects of 2 wt.% of Si on Fe–Ni interdiffusion is negligible.

219 Previous studies suggested that the ‘homologous temperature relationship’ worked under  
220 high-pressure conditions:

$$221 \quad D = D_0 \exp\left(-\frac{AT_m}{RT}\right), \quad (5)$$

222 where  $A$  is an empirical constant, and  $T_m$  is the melting temperature. We plotted diffusion  
223 coefficients against the homologous temperature estimated from the melting temperature of  
224 Fe reported by Sinmyo et al. (2019) (Fig. 8). We use the melting temperature of pure Fe, as  
225 the incorporation of 2 wt.% of Si into Fe does not decrease the melting temperature  
226 significantly (e.g., Fischer et al. 2013; Komabayashi 2020). The plots of our data and Yunker  
227 and Van Orman (2007) follow homologous temperature relationships suggested in previous  
228 studies under lower pressure regimes (Goldstein et al. 1965; Ritterbex and Tsuchiya 2020).  
229 The results of Reaman et al. (2012) also seem to be consistent with other studies, although the  
230 results of Reaman et al. (2012) show slight deviations from the trend, probably due to  
231 relatively large temperature uncertainties originated from LHDAC. The slope represents  
232 empirical constant  $A/R$  in Eq. (5), which depends on the crystal structure and bonding  
233 properties.  $A/R$  derived from this study is 20.37, while Yunker and Van Orman (2007) and  
234 Reaman et al. (2012) report 20.4 and 19.3, respectively. Considering the temperature and  
235 pressure conditions of diffusion experiments in this study and that of Yunker and Van Orman  
236 (2007), the specimens in both studies were assumed to be the fcc structure. Further, we can  
237 assume that the bonding properties of Fe–2 wt.% Si alloy are not significantly different from  
238 that of pure Fe, as alloying of 2 wt.% of Si does not change the melting temperature  
239 significantly (e.g., Fischer et al. 2013; Komabayashi 2020). Therefore, we conclude that Ni  
240 diffusion in Fe–2 wt.% Si would show a similar slope to the result of the previous study on  
241 Fe–Ni interdiffusion (Yunker and Van Orman 2007). Our estimation of the diffusion  
242 coefficient in the inner core is  $3.29(\pm 1.55) \times 10^{-14} \text{ m}^2 \text{ s}^{-1}$ , assuming  $T_m/T$  of the Earth’s inner  
243 core is 1.05 (Reaman et al. 2012). As shown by Yunker and Van Orman (2007), there are



244 cases where concentration dependence exists in Fe–Ni interdiffusion. However, since the  
245 results of our FEM analysis without considering the concentration dependence are in good  
246 agreement with the experimental data (Fig. 6), the effect of the concentration dependence of  
247 Fe–Ni interdiffusion would be small under our experimental conditions. The concentration  
248 dependence of Fe–Ni interdiffusion decreases with increasing temperature up to about 2000  
249 K at the equivalent pressures (Yunker and Van Orman 2007). Therefore, the concentration  
250 dependence of the diffusion coefficient is not expected to have a significant effect under high  
251 temperature conditions such as those in the Earth's inner core.

252 The Fe–Si alloy containing up to 7 wt.% Si has a hcp structure under pressure and  
253 temperature conditions relevant to the inner core (Tateno et al. 2015). Because 7 wt.% Si is  
254 more than sufficient to explain the density deficit of the inner core, and the Si composition in  
255 the inner core is thought to be ~1–2 wt.% (Antonangeli et al. 2010; Badro et al. 2007; Sakairi  
256 et al. 2018), the inner core is still expected to be composed of Fe alloys having a hcp  
257 structure, even when taking into account the presence of Si. Even though we conducted  
258 experiments under pressure and temperature conditions of the fcc stable region according to  
259 previous studies (Fischer et al. 2013; Komabayashi et al. 2020), our result can be applied to  
260 the Earth's inner core, as the diffusivity in the fcc crystal would be similar to that in hcp  
261 crystal, because they both assume the closest packed structure. Indeed, numerous hcp and fcc  
262 metals show comparable diffusion coefficients at the same homologous temperatures (Brown  
263 and Ashby 1980). The inner core also likely contains some amount of light elements other  
264 than Si such as S, O, C, and H (Hirose et al. 2021). Effects of the presence of these multiple  
265 light elements on diffusivity in the inner core should be further studied.

266

## 267 **IMPLICATION: ESTIMATION OF UPPER LIMIT OF EARTH'S INNER CORE** 268 **VISCOSITY**

269 The seismic anisotropy of the Earth's inner core can be interpreted as a result of viscous  
270 flow that induces LPO of the inner core material (Deguen 2012; Romanowicz and Wenk  
271 2017). Various models have been suggested for explaining the viscous flow in the Earth's  
272 inner core (Buffett and Bloxham 2000; Jeanloz and Wenk 1988; Karato 1999; Takehiro 2011;  
273 Yoshida et al. 1996), and the plausibility of each model depends on the viscosity of the

274 Earth's inner core, which can be evaluated from mechanical properties and diffusion  
275 coefficient of the inner core material. Herein, we use the diffusion coefficient estimated in  
276 this study to place a constraint on the upper limit of the viscosity of the Earth's inner core. If  
277 the seismic anisotropy can be attributed to the LPO of Fe alloys induced by deformation, the  
278 active creep mechanism in the inner core would be dislocation-controlled creep, rather than  
279 diffusion creep, as the latter may not develop LPO or rather destroy pre-existing LPO  
280 (Boullier and Gueguen 1975). Moreover, all previous studies on the grain size of the inner  
281 core support a large grain size, which makes diffusional creep of the inner core infeasible.  
282 Considering low seismic attenuation of the inner core, its grain size might be within a range  
283 of  $10^2$ – $10^3$  m (Cormier and Li 2002). Grain growth experiments on hcp Fe under high  
284 pressure suggest the inner core grain size ranging from several hundred meters to several  
285 kilometers (Yamazaki et al. 2017). Irrespective of the above-mentioned estimation adopted,  
286 diffusional creep will lead to a notably high viscosity of the inner core ranging from  $\sim 10^{24}$  to  
287  $10^{27}$  Pa s (Ritterbex and Tsuchiya 2020). This estimation is higher than the upper end of the  
288 range of previously estimated viscosity (Buffett 1997; Davies et al. 2014; Frost et al. 2021;  
289 Jackson et al. 2000; Koot and Dumberry 2011; Reaman et al. 2011; Ritterbex and Tsuchiya  
290 2020; Van Orman 2004; Yoshida et al. 1996), which in turn makes dislocation-based creep  
291 the dominant creep mechanism.

292 Based on the high homologous temperature and low-stress condition of the inner core,  
293 Harper-Dorn creep has been suggested as a controlling creep mechanism in the inner core  
294 (Van Orman 2004). Harper-Dorn creep is a dislocation-based mechanism that has been  
295 reported in experimental studies under low deviatoric stress and high homologous  
296 temperature conditions (Harper and Dorn 1957), similar to those in the inner core. However,  
297 this has been observed only in high purity metals, while the inner core includes around 5% Ni  
298 and a certain number of light elements. Further, its underlying physical mechanisms and even  
299 its existence remain controversial (Blum et al. 2002).

300 Among dislocation-based creep mechanisms, the most widely accepted one is the p  
301 ower law creep:

$$302 \quad \dot{\epsilon} = \alpha \frac{D G b}{kT} \left(\frac{\sigma}{G}\right)^n, \quad (6)$$

303 where  $\dot{\epsilon}$  is the strain rate,  $\alpha$  is a constant,  $D$  is the diffusion coefficient,  $G$  is the shear

304 modulus,  $b$  is the burgers vector,  $k$  is the Boltzmann constant,  $\sigma$  is the differential stress, and  
305  $n$  is the stress exponent. Viscosity ( $\eta$ ) is defined as an equation:

$$306 \quad \eta = \frac{\sigma}{2\dot{\epsilon}}. \quad (7)$$

307 With a  $T_m/T$  lower than 2.5, the creep mechanism suggested by Weertman (1957) may be  
308 active, which assume dislocation glide produces strain and dislocation climb is the rate  
309 limiting process. Weertman's creep can be expressed by  $n = 3$  and stress-dependent  
310 parameter  $\alpha$ , depending on the lattice plane distance  $d$  between slip planes (Ritterbex and  
311 Tsuchiya 2020). We employ the basic assumptions of Ritterbex and Tsuchiya (2020) and  
312 calculate the upper limit of viscosity in the Earth's inner core. For the diffusion coefficient,  
313 we used  $3.29(\pm 1.55) \times 10^{-14} \text{ m}^2 \text{ s}^{-1}$  derived from the present results (Fig. 8). For lattice  
314 parameters  $a$  and  $c$  of hcp Fe at the inner core conditions, we use experimental results by  
315 Tateno et al. (2010). We employ the shear modulus of 212 GPa (Vočadlo et al. 2009) and  
316 assume that a basal slip of hcp Fe is active in the inner core conditions (hence,  $b = a$  and  $d =$   
317  $c/2$ ).

318 To estimate the maximum viscosity of the inner core, we first define the minimum strain  
319 rate in the inner core that develops its deformation-induced anisotropy. Nishihara et al. (2018)  
320 conducted shear deformation experiments on hcp Fe, which revealed that the texture  
321 development was initiated at a strain of 0.5. If we assume that the viscous flow has been  
322 active since the beginning of the inner core formation, the strain of 0.5 divided by the inner  
323 core age would yield the minimum strain rate  $\dot{\epsilon}$  for explaining deformation-induced  
324 anisotropy in the inner core. Recent estimates on mantle global circulation heat flow suggest  
325 the inner core age ranging from  $\sim 400$  to  $\sim 1100$  Ma (Olson et al. 2015). The anomalously low  
326 Ediacaran paleomagnetic field intensity suggests the onset of inner core growth at around  
327  $\sim 500$  Ma (Bono et al. 2019). Core thermal conductivity estimated from high-pressure  
328 experiments also can constrain inner core age to be  $\sim 700$  Ma (Ohta et al. 2016). Here, we  
329 employ the inner core age of 1100 Ma for conservative estimation, thus the minimum strain  
330 rate of the inner core would be  $1.44 \times 10^{-17} \text{ s}^{-1}$ . We then substitute the estimated minimum  
331 strain rate and relevant physical parameters into Eq. (6) to derive minimum stress conditions  
332  $\sigma$  in the inner core. The minimum stress condition of the inner core would be around 0.12 Pa  
333 when using the diffusion coefficient estimated in this study. The resultant maximum viscosity

334 of the inner core is  $4.23(\pm 0.68) \times 10^{15}$  Pa s (Eq. 7). This upper limit of the Earth's inner core  
335 viscosity is consistent with a geophysical observation suggesting a viscosity below  $3 \times 10^{16}$  Pa  
336 s (Buffett 1997). Considering that typical estimations on the mantle viscosity are in the range  
337 of  $10^{21}$  to  $10^{24}$  Pa s, the inner core with a viscosity below  $3 \times 10^{16}$  Pa s would easily  
338 accommodate viscous flow. It should be emphasized that the  $3 \times 10^{16}$  Pa s is a result assuming  
339 the inner core strain of 0.5, which is a minimum strain for showing deformation induced LPO  
340 and if the inner core's viscosity is larger than  $3 \times 10^{16}$  Pa s, seismic anisotropy of the inner  
341 core should be explained by mechanisms other than LPO such as shape preferred orientation  
342 (for review, see Deguen 2012). The upper bound of viscosity estimated in this study  
343 constrains geodynamical mechanisms responsible for the inner core anisotropy. For instance,  
344 Lasbleis and Deguen (2015) built regime diagrams of dominant viscous flow in the inner  
345 core, and under a given thermal instability regime in the diagram, the dominant viscous flow  
346 mechanism is dependent on the viscosity of the inner core. The inner core with a viscosity  
347 below  $3 \times 10^{16}$  Pa s suggests that it is likely to be deformed by plume convection under an  
348 unstable stratification regime. If the inner core is under a stable stratification regime and the  
349 viscosity of the inner core is below  $10^{12}$  Pa s, the azimuthal Lorentz force resulting from the  
350 Earth's magnetic field could represent the dominant mechanism. In any case, the Lorentz  
351 forces would be significant driving forces of viscous flows, as their effects increase with  
352 decreasing viscosity under both the stable and unstable stratification regime (Lasbleis and  
353 Deguen 2015).

354 Frost et al. (2021) recently refined the seismic dataset of the inner core and suggested an  
355 inner core viscosity larger than  $10^{18}$  Pa s, which was restricted by their geodynamic model for  
356 simulating observed seismic anisotropy of the inner core. In contrast to Frost et al. (2021), we  
357 estimated the upper limit of the viscosity more directly from constraints provided by mineral  
358 physics. We suggest that further studies are needed to establish plausible geodynamic models  
359 capable of explaining both the low viscosity of the inner core and its characteristic seismic  
360 features, such as increasing seismic anisotropy with depth or hemispherical dichotomy (Frost  
361 et al. 2021; Tanaka and Hamaguchi 1997).

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## ACKNOWLEDGEMENTS

364 This study was supported by the Japan Society for the Promotion of Science (JSPS)  
365 KAKENHI (grant number 19H00716).

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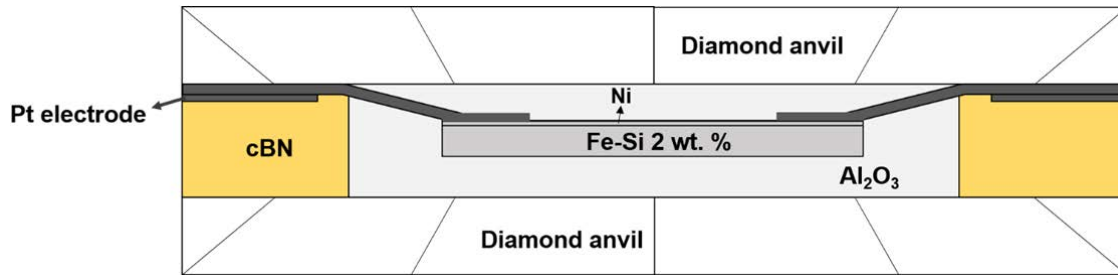
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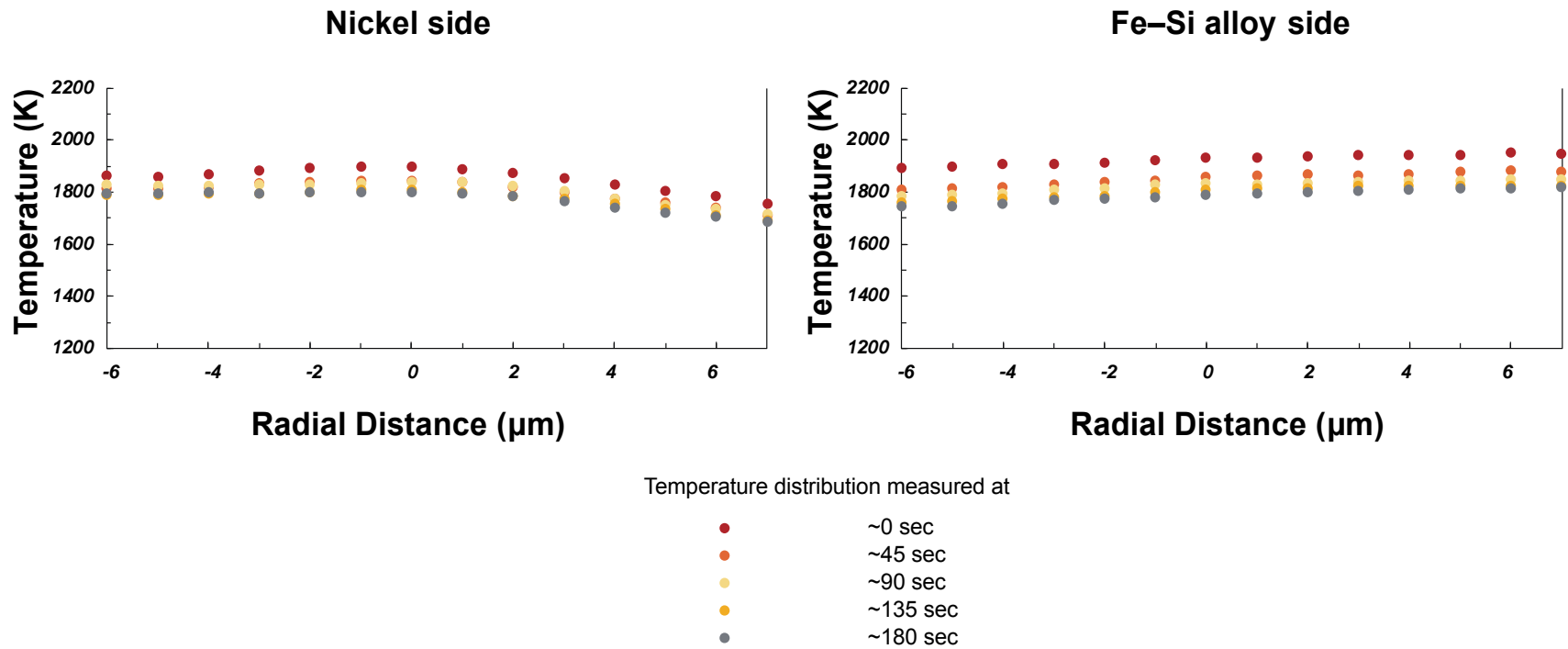
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496 **FIGURE 1.** Schematic illustration of sample configuration of an internal-resistive-heated  
497 DAC.

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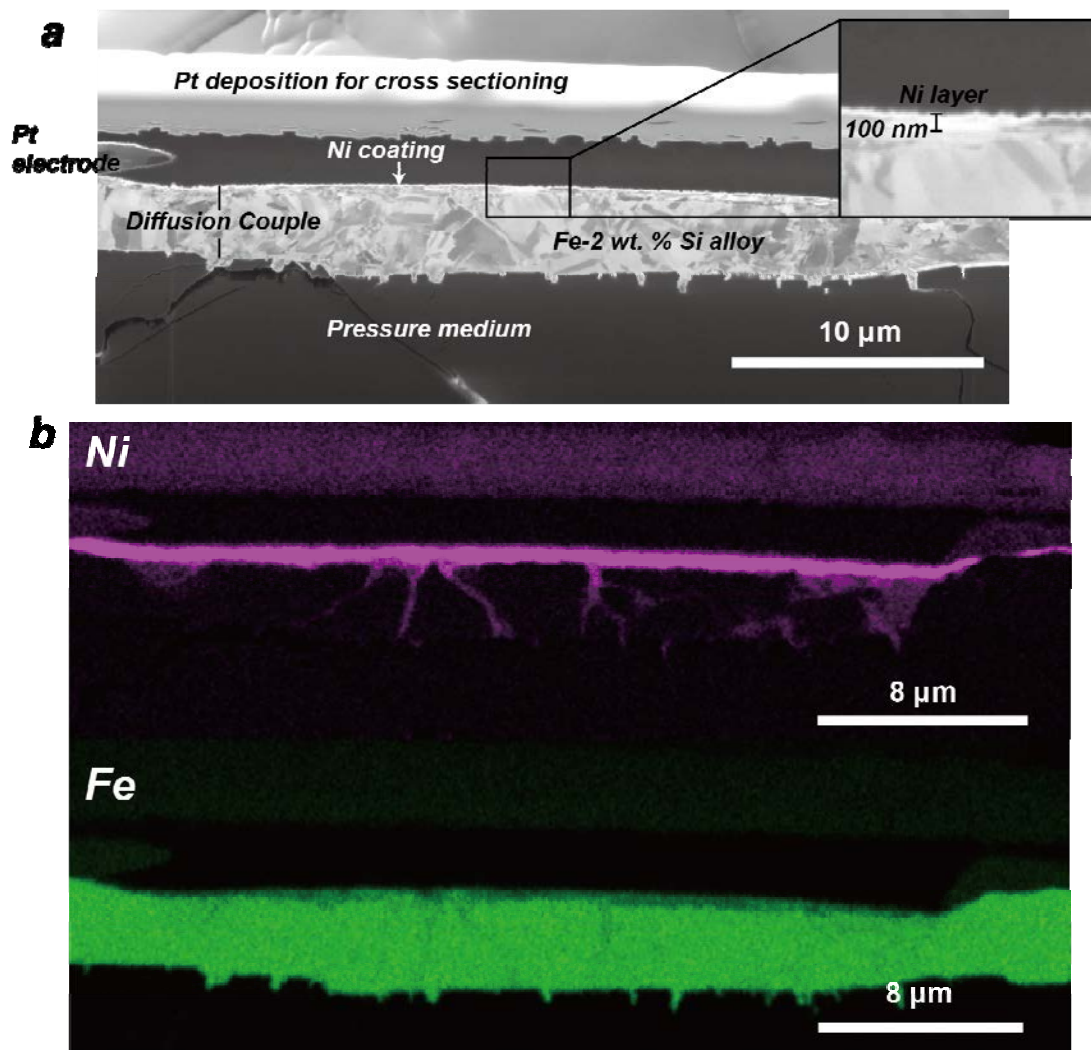




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500 **FIGURE 2.** Temperature profiles of a representative specimen during a diffusion experiment (Run #1). The lateral extent of each temperature  
501 point is 0.5 µm to the left and right.

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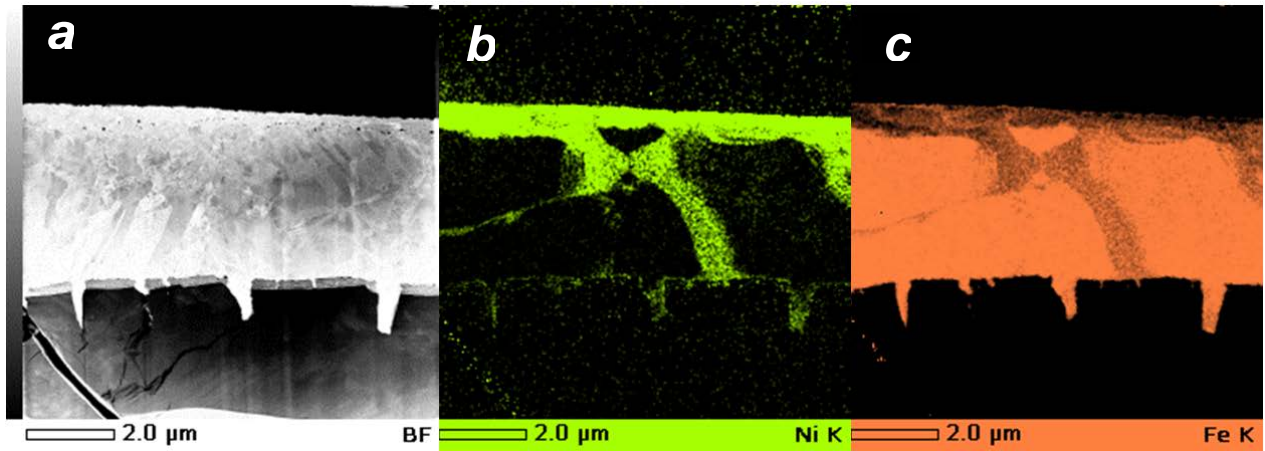


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505 **FIGURE 3.** (a) SIM image, and (b) SEM-EDS mapping images of sample cross section (run  
506 #3). The inset (a) shows enlarged view of boundary between Ni layer and Fe-2 wt.% Si alloy.

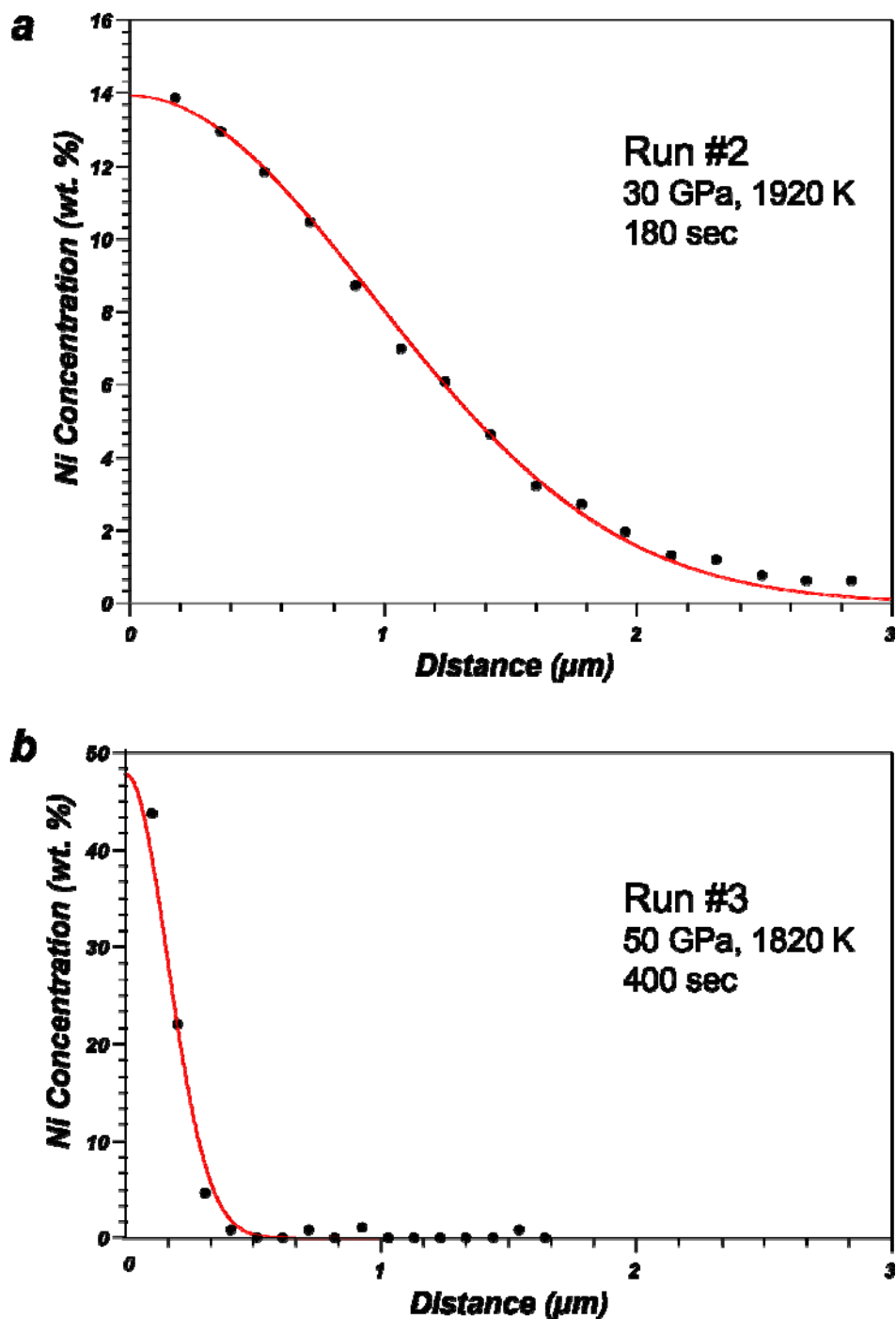
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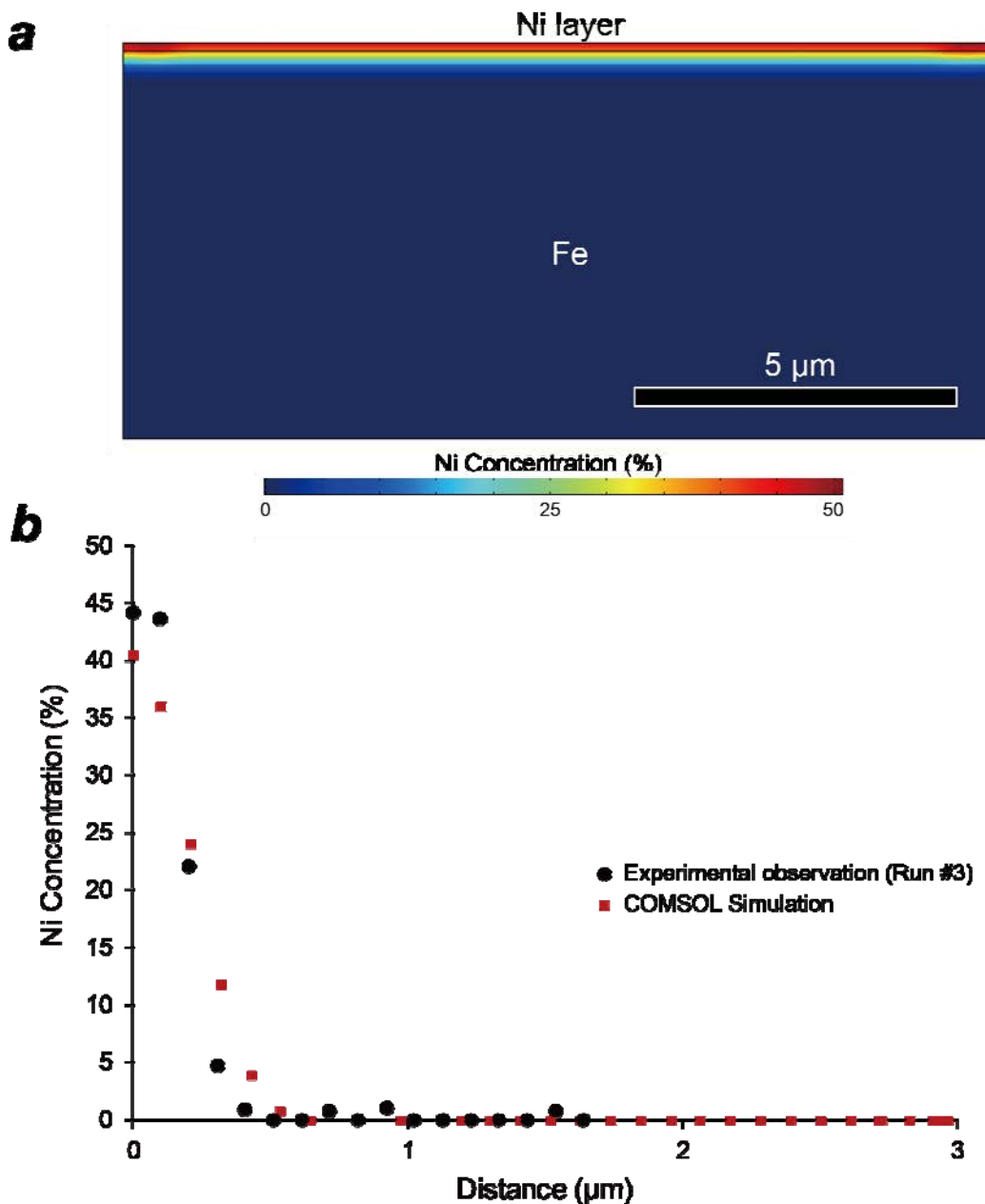
509 **FIGURE 4.** (a) TEM bright field image (run #3). (b), (c) Corresponding TEM-EDS mapping  
510 images of Ni and Fe.

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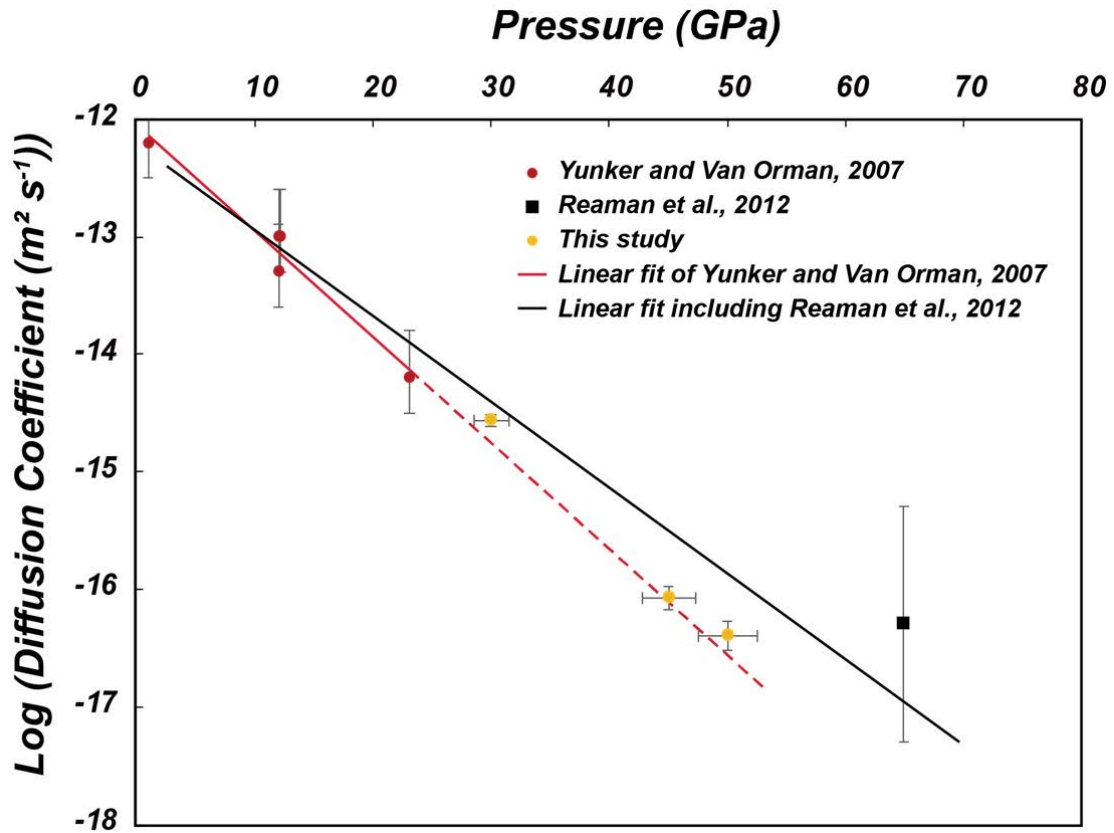
513 **FIGURE 5.** Representative diffusion profiles of Ni obtained by TEM-EDS analyses after  
514 diffusion experiments at temperature and pressure conditions of (a) ~1920 K, ~30 GPa, and  
515 (b) ~1820 K, ~50 GPa. The measurement error for Ni concentration is less than 2%. The  
516 distance uncertainty is about 1.0 nm, which corresponds to the diameter of the EDS electron  
517 probe. Red curves indicate best fitting curves of the thin film solution to the measured  
518 diffusion profiles.



519

520 **FIGURE 6.** (a) A result of COMSOL simulation assuming diffusion tracer thickness of 120  
521 nm and diffusion coefficient of  $4.08 \times 10^{-17} \text{ m}^2 \text{ s}^{-1}$  obtained by fitting the diffusion profile of  
522 Run #3 to the thin film solution. (b) Comparison of diffusion profiles from an actual  
523 experimental observation (Run #3) and the COMSOL simulation. The measurement error for  
524 Ni concentration is less than 2%. The distance uncertainty is about 1.0 nm, which  
525 corresponds to the diameter of the EDS electron probe.

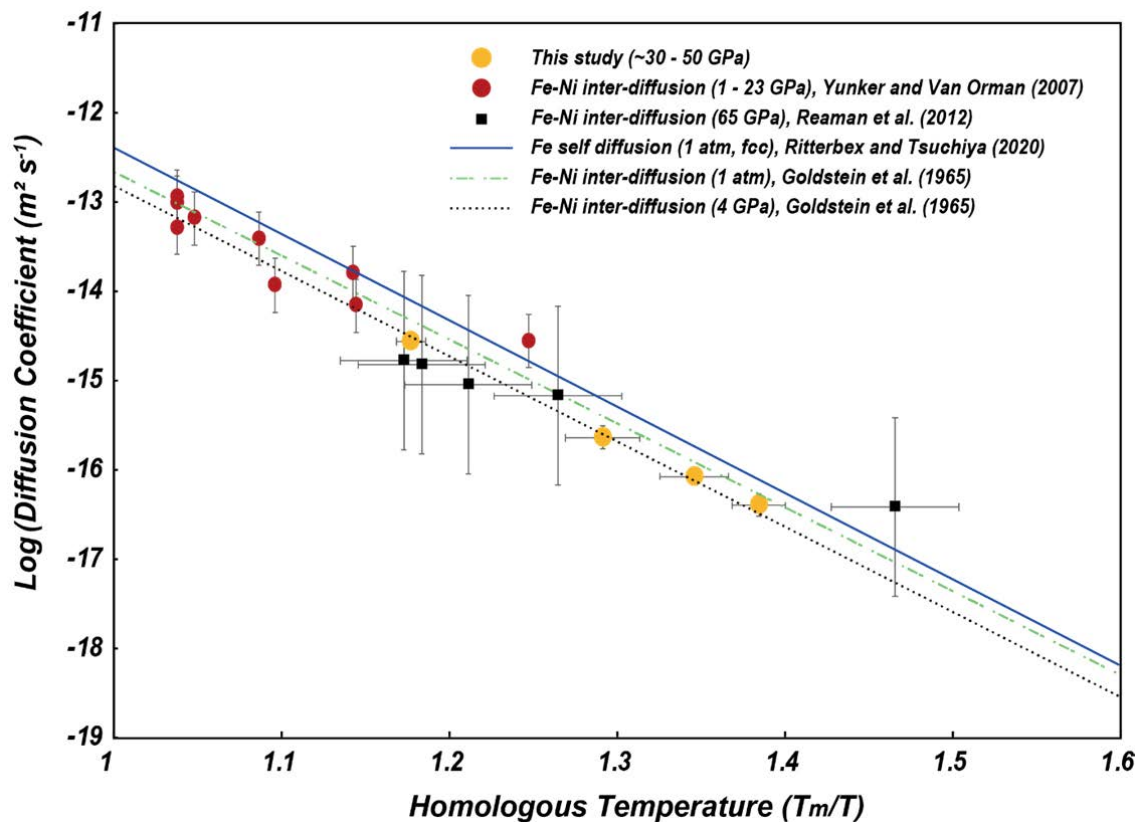
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528 **FIGURE 7.** Diffusion coefficients against pressure under similar temperature conditions.  
529 Temperature conditions are in a range from ~1800 to ~1900 K. Red dots and a black dot  
530 indicate Fe–Ni interdiffusion coefficients from Yunker and Van Orman (2007) and Reaman et  
531 al. (2012), respectively, without considering silicon.

532



533

534 **FIGURE 8.** Diffusion coefficients in this study versus homologous temperature and previous  
535 results of Fe–Ni interdiffusion experiments under various pressure ranges (1 atm to 65 GPa,  
536 Goldstein et al. 1965; Yunker and Van Orman 2007; Reaman et al. 2012) and Fe self-  
537 diffusion coefficients in fcc Fe under atmospheric pressure estimated from density functional  
538 theory calculations (Ritterbex and Tsuchiya 2020).

539 **TABLE 1.** Summary of diffusion experiments conducted in this study

Run #	Pressure (GPa)	Temperature (K)	Diffusion time (s)	Diffusion coefficient ( $\text{m}^2 \text{s}^{-1}$ )
#1	45 ( $\pm 2$ )	1820 ( $\pm 50$ )	180	$8.52 \times 10^{-17}$ ( $\pm 1.14 \times 10^{-17}$ )
#2	30 ( $\pm 2$ )	1920 ( $\pm 20$ )	180	$2.74 \times 10^{-15}$ ( $\pm 1.85 \times 10^{-16}$ )
#3	50 ( $\pm 3$ )	1820 ( $\pm 40$ )	400	$4.08 \times 10^{-17}$ ( $\pm 8.18 \times 10^{-18}$ )
#4	31 ( $\pm 2$ )	1750 ( $\pm 50$ )	250	$2.37 \times 10^{-16}$ ( $\pm 7.12 \times 10^{-17}$ )

540