

High Pressure Raman and Nd³⁺ Luminescence Spectroscopy of Bastnäsit-(REE)CO₃F

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Abstract

Bastnäsit-(Ce), a rare earth element (REE) bearing carbonate (Ce,La,Y,Nd,Pr)CO₃F, is one of the most common REE-bearing minerals, and has importance from both economic and geologic perspectives due to its large REE concentration. It also provides an example of the structural interplay between carbonate groups and fluorine ions, as well as the complex bonding properties of rare earth elements. We report Raman vibrational and Nd³⁺ luminescence (⁴F_{3/2}→⁴I_{9/2}, ⁴F_{3/2}→⁴I_{11/2} and ⁴F_{5/2+2}H_{9/2}→⁴I_{9/2}) spectra of natural bastnäsit-(Ce) to 50 GPa at 300 K. Two phase transitions are observed under compression. Bastnäsit-I remains the stable phase up to 25 GPa, where it undergoes a subtle phase transition to bastnäsit-II. This is likely produced by a change in symmetry of the carbonate ion. Bastnäsit-II transforms to bastnäsit-III at ~38 GPa, as demonstrated by changes in the luminescence spectra. This second transition is particularly evident within the ⁴F_{3/2}→⁴I_{9/2} luminescent transitions, and it appears that a new rare earth element site is generated at this phase change. This transition is also accompanied by modest changes in both the Raman spectra and two sets of luminescent transitions. Despite these transformations, the carbonate unit remains a stable, three-fold coordinate unit throughout this pressure range, with a possible increase in its distortion. Correspondingly, the rare-earth element site(s) appears to persist in quasi-9-fold coordination as well, implying that the general bonding

24 configuration in bastnäsite is at least metastable over a ~30% compression range. All pressure-
25 induced transitions are reversible, with some hysteresis, reverting to its ambient pressure phase
26 on decompression.

27 **Keywords:** Bastnäsite, high pressure, Raman spectroscopy, rare earth elements, neodymium
28 luminescence spectroscopy, carbon, deep Earth

29 **Introduction**

30 Bastnäsite-(Ce) is a rare earth element (REE) bearing fluorocarbonate
31 (Ce,La,Y,Nd,Pr)CO₃(F,OH) that is typically found associated with carbonatite deposits, and is
32 economically important due to its high content of REEs. Indeed, bastnäsite is one of the two
33 major mineral ores for REEs, with its chief competition being monazite. Geochemically, REEs
34 are typically incompatible in silicate minerals; their concentration in carbonatite melts (often
35 with bastnäsite as their crystallization product) is generally more than 3 orders of magnitude
36 higher than the bulk Earth content (Jones et al. 2013), and greater than 4 orders of magnitude
37 higher than within chondrites (Holtstam and Andersson 2007). Indeed, carbonate-bearing liquids
38 and solids likely play a primary role in transporting and retaining rare earth elements within the
39 Earth's mantle (e.g., Collerson et al. 2010). In addition, light REE have long been used to
40 determine the ages of continental rocks, particularly via the samarium-neodymium dating
41 method (McCulloch and Wasserburg 1978), and the evolution of the depleted mantle has been
42 probed using the distinct ¹⁴³Nd/¹⁴⁴Nd variation in the mantle compared to chondrites and crustal
43 rocks (DePaolo and Wasserburg 1976). Bastnäsite has significance beyond its REE content: as
44 carbonates are thought to be the dominant species containing carbon within oxidized regions of
45 the mantle (e.g., Brenker et al. 2007), the properties of bastnäsite can provide insight into the
46 bonding of carbon at depth. Specifically, bastnäsite, has importance for understanding the

47 potential behavior of REE at depth within carbon (and halogen) enriched zones of the mantle,
48 and within carbonatite melts (which are frequently halogen-enriched: e.g., Jago and Gittins 1991;
49 Williams and Knittle 2003). Additionally, the optical and magnetic properties of rare-earth
50 carbonate fluorides have attracted interest within the materials science community (e.g., Grice et
51 al. 2007).

52 Bastnäs site virtually never occurs as an endmember in terms of its REE content, as it
53 commonly contains a mixture of REEs (e.g. Ce, Nd, Eu, Y, etc.). Visible and infrared reflectance
54 spectroscopy have been used to probe the absorption bands of bastnäs site, which can constrain the
55 REE content (Turner et al. 2015); however, assignments to the absorptions and energy levels of
56 different REEs have generally not been made to the observed absorption peaks. In addition, the
57 rare earths in bastnäs site luminesce strongly; as discussed below, it appears that this fluorescence
58 has at times been associated with hydroxyl stretching vibrations in Raman spectroscopic
59 experiments (Frost and Dickfos 2007). In this study, we are able to assign the luminescence
60 peaks to transitions of the neodymium ion, and probe changes to the REE site with pressure by
61 using this Nd³⁺ luminescence.

62 Luminescence has been employed extensively as a probe for pressure-induced changes in
63 crystal structure or in electronic configuration (e.g., Dolan et al. 1986; Freire et al. 1994; Bray
64 2001; O'Bannon and Williams 2016). Neodymium luminescence has been investigated for use as
65 a pressure calibrant due to its high fluorescence intensity and insensitivity to changes in
66 temperature (e.g., Hua et al. 1996), with YAlO₃ being typically viewed as the most promising
67 Nd-bearing sensor due to its lack of phase change or amorphization up to 80 GPa (Hua and
68 Vohra 1997; Hernández-Rodríguez et al. 2018a, 2018b). Other high pressure studies using Nd³⁺
69 luminescence have investigated the pressure-shifts of different energy levels and their associated

70 crystal field parameters (e.g., Santiuste et al. 2017; Hernández-Rodríguez et al. 2018a). Thus,
71 Nd³⁺ luminescence has been investigated under pressure in a range of oxide crystals to derive
72 insights into the electronic properties of crystals used for solid state lasers under compression,
73 and at ambient pressure within minerals (e.g., Lenz et al. 2013). However, to our knowledge,
74 high-pressure Nd³⁺ luminescence spectroscopy has not been previously utilized to probe the
75 electronic and structural changes within a geologically relevant mineral. Here, we employ both
76 Raman and Nd³⁺ luminescence spectroscopy to constrain changes in the bonding environments
77 of both the carbonate and neodymium ions in bastnäsite; this is the first high pressure study of
78 bastnäsite using these complementary techniques.

79 **Experimental Methods**

80 Natural bastnäsite (Ce_{0.49}La_{0.25}Nd_{0.21}Pr_{0.04})CO₃(F_{0.95}(OH)_{0.05}) from Northern Pakistan was
81 used for these experiments. The chemical composition of the sample was confirmed with four
82 grains of bastnäsite using a ThermoScientific Apreo scanning electron microscope equipped with
83 an Oxford Instruments Ultim Max 100 mm energy dispersive X-ray spectrometer (0.8 nA and
84 15kV beam); SEM results are shown in Table S1 and Fig. S1 . Raman spectroscopy and single
85 crystal X-ray diffraction of this sample were in excellent agreement with previous determinations
86 (Ni et al. 1993; Frost and Dickfos 2007).

87 The ambient pressure crystal structure was measured at beamline 12.2.1 at the Advanced
88 Light Source. Diffraction images were collected at 100 K with a Bruker D8 diffractometer
89 equipped with a Photon II CPAD detector, and X-rays monochromated by silicon (111) to 17
90 keV (0.7288 Å). The single crystal was a clear block that was free of visible cracks and/or
91 imperfections, with an approximate size of 90 x 80 x 50 μm. The sample was mounted in
92 Paratone-N oil on a MiTeGen Micromount. Diffraction images were collected using the APEX3

93 program (Bruker 2016a). Images were integrated and cell refinements were computed using
94 SAINT V8.34A (Bruker 2013). The sample symmetry of the ambient structure was determined
95 using XPREP, and crystal absorption was corrected using the algorithms of SADABS-2016/2
96 (Bruker 2016b). The crystal structure from Ni et al. (1993) was used as an initial fit, and refined
97 using full-matrix least-squares on F^2 via SHELXL-2018/1 (Sheldrick 2015), using the software
98 package ShelXle Rev 806 (Hübschle et al. 2011).

99 Raman and luminescence spectra were collected using a Horiba LabRam HR Evolution
100 spectrometer equipped with either a 633 or 532 nm excitation laser, depending on the experiment
101 being conducted. The spectrometer focal length was 800 mm, and a grating of 1800 lines/mm
102 was used with a spectral resolution of $\sim 1 \text{ cm}^{-1}$. Spectra were collected up to a pressure of ~ 50
103 GPa at $\sim 300 \text{ K}$. Peak positions were calculated from fits that used a combination of Gaussian and
104 Lorentzian peak shapes with Horiba LabSpec6 software.

105 High pressures were generated using a Princeton type symmetric diamond anvil cell
106 equipped with type IIa diamonds with 250 μm culets. Rhenium was used as the gasket material.
107 Gaskets were preindented to $\sim 30 \mu\text{m}$ thickness, and the sample compartments were $\sim 125 \mu\text{m}$ in
108 diameter. Neon was used as a pressure medium, as it is closer to hydrostatic than most other
109 pressure media at high pressures (Klotz et al. 2009); however, supplementary data were collected
110 using a 4:1 methanol:ethanol pressure medium, as reported in Fig. S2. These are completely
111 consistent with the data reported here in a neon medium. . Ruby fluorescence from multiple ruby
112 grains within the sample compartment was used to both determine pressures, and pressure
113 variations within the chamber (Dewaele et al. 2008): our results are compatible with those of
114 Klotz et al. (2009) in there being an order of a few tenths of GPa of pressure variations at the
115 highest pressures of these measurements.

116

Results

117 The ambient pressure, low temperature crystal structure (Fig. 1) of our sample agrees
118 with the previously reported structure by Ni et al. (1993), rather than with the structure inferred
119 by Donnay and Donnay (1953). Bastnäsite is in the space group $P-62c$, with $a = 7.0838(2) \text{ \AA}$, $c =$
120 $9.7365(3) \text{ \AA}$, and 6 formula units in the unit cell. Its structure has alternating layers comprised of
121 $(\text{CO}_3)^{2-}$ units and REE ions juxtaposed with F ions in the c -axis direction (Fig. 1). The REE
122 atoms are in a 9-coordinate site (with a volume of 31.338 \AA^3), with a point symmetry of C_2 . The
123 atoms comprising the REE polyhedra are 3 fluorine atoms and 6 oxygen atoms. There is one
124 carbonate group within the unit cell; the carbon and one of its coordinating oxygens (O1) lie on a
125 mirror plane, and the other oxygens (O2) in the carbonate group are related by symmetry. The C-
126 O1 distance is $1.2901(45) \text{ \AA}$ and C-O2(x2) are $1.2845(27) \text{ \AA}$ and the carbon atom lies (within
127 error) in a plane with the three oxygen atoms ($0.0024(27) \text{ \AA}$ out of the plane generated by the
128 three oxygen atoms). There are two fluorine atom sites within the unit cell; F1 lies (roughly) in a
129 plane with the REE atoms that generate a plane at $x \sim 0.66$. F2 lies out of this plane, causing a
130 distortion of the REE polyhedra.

131 At ambient pressure and temperature, thirteen lattice modes and seven modes associated
132 with the carbonate ion are resolved in the Raman spectrum (Fig. 2a). There are a large number of
133 modes in the lattice region of the ambient spectrum, reflecting the relatively low symmetry of the
134 polyhedra of the REE ions, with bonding to both oxygen and fluorine. The modes associated
135 with the carbonate ions are easily assigned: the symmetric stretch (ν_1) is at $1,096 \text{ cm}^{-1}$, one mode
136 associated with the out-of-plane bend (ν_2) is at 870 cm^{-1} , an asymmetric stretch (ν_3) is at 1432
137 cm^{-1} , and the in-plane bend (ν_4) has components at $664, 686, 729, \text{ and } 737 \text{ cm}^{-1}$. The increased
138 number of carbonate modes compared to the isolated ion is due to the distortion of the carbonate

139 ion coupled with its low symmetry site within the crystal structure. Factor group analysis of the
140 optic modes of bastnäesite with the space group $P-62c$, yields:

$$141 \quad \Gamma_{\text{optic}} = 9A'_1 (\text{R}) + 7A''_1 (\text{Inactive}) + 11A'_2 (\text{Inactive}) + 8A''_2 (\text{IR}) + 19E' (\text{R,IR}) + 16 E''$$

142 (R) .

143 Thus, there are 44 Raman active vibrations and 27 infrared active vibrations; accordingly,
144 the numbers of anticipated Raman active modes and carbonate modes are substantially greater
145 than we observe, which is not unusual given that both weak modes and accidental degeneracies
146 are typically present in complex, low-symmetry molecular crystals. At high pressures, splittings
147 and mode discontinuities in the Raman data provide *prima facie* evidence for phase transitions
148 that occur under compression.

149 Bastnäesite does not naturally occur as the pure endmember with a given rare earth
150 element, so multi-element occupancy of the rare earth element site is expected. In our bastnäesite
151 sample, neodymium substitutes at the 24% level into the REE site. As such, we probe the Nd^{3+}
152 luminescence within the structural environment of the REE site (Fig. 2b), and monitor systematic
153 changes in city distortion via the luminescence spectroscopy. Within the bastnäesite crystal
154 structure, neodymium's electronic energy levels are split by three principal effects: (1) by
155 electrostatic interactions, (2) spin orbit coupling, and (3) the local electric field that generates 2
156 to 6 Stark components (depending on the energy level), as imposed by the crystal field
157 surrounding the ion (Fig. 3). The observed Nd^{3+} luminescence is entirely consistent with that of a
158 Nd^{3+} ion in a site with low symmetry (e.g., Table 4.3 of Kaminskii 1981). We observe
159 luminescence associated with energy transitions between these electronic states:

160 ${}^4F_{5/2+2}H_{9/2} \rightarrow {}^4I_{9/2}$ at ~790-822 nm, ${}^4F_{3/2} \rightarrow {}^4I_{9/2}$ at ~862-900 nm, and ${}^4F_{3/2} \rightarrow {}^4I_{11/2}$ at ~1046-1066

161 nm (Fig. 3). We are able to deconvolve the peaks and assign the energy levels of each Stark-split
162 level ($^4I_{9/2}$, $^4I_{11/2}$, $F_{3/2}$, $^4F_{5/2}$, $^2H_{9/2}$) of the spin orbit coupled energy levels (Fig. 2b).

163 **High Pressure Raman Results and Discussion**

164 **Bastnäsité-I: Raman spectroscopy under pressure**

165 We resolve mode shifts for both lattice and carbonate modes of the ambient pressure
166 phase of bastnäsité up to 25 GPa (Fig. 4). Mode assignments and their pressure shifts are
167 reported with associated mode Grüneisen parameters in Table 1, calculated using the bulk
168 modulus of 105 GPa from Rowland (2017). Below 25 GPa, the lattice modes all increase in
169 frequency with pressure, with no major changes in their number or peak shape morphology (Fig.
170 4a, d). Some apparent changes in intensity are observed in this pressure range, but these changes
171 are likely due to a combination of pressure-induced removal of accidental degeneracies and
172 possible effects of changes in crystal orientation between different spectra.

173 The Raman active modes of the carbonate ion shift monotonically up to 25 GPa,
174 indicating that no phase change occurs up to this pressure. The symmetric stretching mode (ν_1)
175 shifts approximately linearly up to 25 GPa at a rate of $2.79 \text{ cm}^{-1}/\text{GPa}$. Interestingly, at ambient
176 pressure, this peak is more asymmetric than at 23 GPa. The carbonate out-of-plane bend (ν_2)
177 shifts negatively at $-0.29(8) \text{ cm}^{-1}/\text{GPa}$; such a modest negative shift is typical for carbonate
178 minerals (e.g., aragonite and dolomite: Kraft et al. 1991; Vennari and Williams 2018). This mode
179 grows in intensity on compression compared to the in-plane bend, and at 25 GPa, the mode has
180 twice its relative amplitude as at lower pressures. This shift in intensity is plausibly associated
181 with a Fermi resonance between higher-lying components of the ν_4 in plane bend and the ν_2 out-
182 of-plane bend, as these bands approach one another under pressure (their separation changes
183 from $\sim 130 \text{ cm}^{-1}$ to near 75 cm^{-1} at 25 GPa: Fig. 4b).

184 The four carbonate in-plane bends (ν_{4a} , ν_{4b} , ν_{4c} , ν_{4d}) all shift positively with pressure (Fig.
185 4b,d, Table 1), and their relative intensities also shift with pressure. The highest frequency band
186 (ν_{4d}) remains the most intense up to 25 GPa. The second highest frequency mode (ν_{4c}), which
187 initially is present as a shoulder, merges into the highest frequency band under pressure, and may
188 decrease in relative amplitude: we believe this decrease may be from a combination of a
189 compression-induced loss of Fermi resonance, coupled with a peak broadening of the most
190 intense peak. There appears to be a general increase in relative intensity of the third-highest
191 frequency in-plane bending mode with pressure (ν_{4b}), although the intensities of the two lowest
192 frequency in-plane bending modes (ν_{4a} , ν_{4b}) relative to one another appear largely unchanged
193 under compression. The pressure shifts of the two lower frequency bending modes (ν_{4a} , ν_{4b}) are
194 such that they diverge from the two highest frequency in-plane bending modes (ν_{4c} , ν_{4d}),
195 implying that the distortion of the carbonate group may increase under compression in bastnäsite.

196 **Bastnäsite-II: Raman spectroscopy under pressure**

197 The phase change that occurs at 25 GPa is primarily manifested by changes of the
198 carbonate vibrations, although there are some shifts in the lattice region, as well (Fig. 4). Most
199 lattice modes appear to persist through this phase change: however, three new low intensity
200 modes appear at 284, 430 and 480 cm^{-1} . Within the carbonate region, the symmetric stretch splits
201 (Fig. S4), with a second peak appearing on its lower frequency side; this splitting is first resolved
202 through deconvolution near 25 GPa (Fig. S4), but becomes progressively more pronounced as
203 pressure is increased. The two highest frequency in-plane bends (ν_{4c} , ν_{4d}), which below 25 GPa
204 progressively merge into a single peak, become a single, asymmetric peak. The other two in-
205 plane bend modes (ν_{4a} , ν_{4b}) and the out-of-plane bend mode each increase in intensity across the

206 transition (Fig. 4). Taken together, these changes are consistent with a change in symmetry of the
207 carbonate group, probably from *m* to lower symmetry.

208 The pressure shifts of the lattice modes in bastnäsite-II are listed in Table 2, and shown in
209 Fig. 4a,c. In addition to the modes that first are resolved at 25 GPa, there are two weak modes
210 that become unresolvable at this pressure (those with zero pressure frequencies of 140 and 185
211 cm^{-1}). Relative to the carbonate region, these are relatively subtle changes in the lattice region of
212 the Raman spectrum. Accordingly, the REE site within bastnäsite-II has likely only changed
213 modestly compared to bastnäsite-I; certainly, no shift in coordination of the REE site occurs. The
214 lowest frequency lattice mode splits at 30 GPa—we believe that this may be due to pressure-
215 induced elimination of an accidental degeneracy rather than the phase change itself, as no other
216 lattice modes split at the same pressure. The new intermediate frequency mode (at 284 cm^{-1} at
217 the transition) is only readily tracked until the transition to bastnäsite-III at 38 GPa. Overall, the
218 lattice modes undergo remarkably little alteration with pressure; the two higher frequency lattice
219 modes that appear at 25 GPa are resolvable to the highest pressure probed. This persistence of
220 many of the lattice modes suggests that the topological changes that occur at the two phase
221 transitions do not profoundly affect the lattice vibrations (which are likely more associated with
222 the REE-F layer within the structure and vibrations of the rare earth elements against the
223 carbonate framework). It is possible that a coexistence of multiple REE sites, with different
224 distortions but similar coordination, occurs in the pressure range from 25 GPa to 38 GPa.

225 The pressure shifts of the modes (ν_1 , ν_{1a}) associated with the symmetric stretch of the
226 carbonate ion decrease above 25 GPa, and, as pressure is increased, the two modes associated
227 with the symmetric stretch in the high-pressure phase separate in frequency. Notably, the lower
228 frequency in-plane bend (ν_{2b}) and the out-of-plane bending (ν_2) vibrations become sharper after

229 the transition at 25 GPa to bastnäsite-II up until 38 GPa: this may be a consequence of the
230 transition relieving internal stress within the crystal. Nevertheless, even though there is a change
231 in intensity and shape of a subset of the carbonate bending modes across the transition, there is
232 neither a change in the observed number of peaks, nor a discontinuous change in their
233 frequencies. Bastnäsite-II may have more distorted cation sites (due to the larger number of
234 lattice peaks observed), but does not have major changes associated with its bonding, nor an
235 apparent doubling of sites within the unit cell. The most marked effect of the transition is the
236 splitting of the symmetric stretching vibration (ν_1 , ν_{1a}) of the carbonate group, implying that the
237 local symmetry of the carbonate group may have shifted.

238 **Bastnäsite-III: Raman spectroscopy under pressure**

239 The transition from bastnäsite-II to -III is marked by further splitting of the carbonate
240 modes and modest changes in the lattice modes near 38 GPa. Again, additional modes split at
241 this transition, and there are significant changes in band intensities: for example, the most intense
242 lattice mode at $\sim 390\text{ cm}^{-1}$ at 38 GPa becomes markedly less intense compared to the other
243 modes. It is unclear whether this transition, which clearly initiates at 38 GPa, is complete near
244 this pressure, or whether it is kinetically impeded, and occurs gradually across a wide pressure
245 range. The net decline in intensity of a number of the vibrational modes of bastnäsite-II between
246 38 and 51 GPa, and apparent growth of new, or weak, bands is compatible with the latter
247 interpretation. If, however, this is a transition that is complete near 38 GPa, the new modes (at
248 120 and 271 cm^{-1}) that appear at the transition indicate that either the REE site is further
249 distorted, and/or there has been a doubling of the REE sites within the unit cell. We believe, in
250 conjunction with the luminescence data described below, that a new REE site is produced
251 associated with the transition to bastnäsite-III.

252 The changes of the carbonate vibrations at the transition to bastnäsite-III are more
253 dramatic than those associated with the bastnäsite-I to –II transition. The symmetric stretch,
254 which splits into two deconvolvable modes (ν_1 , ν_{1a}) above 25 GPa, requires a new component
255 (ν_{1b}) to fit its band shape above 38 GPa (Fig. 4c). The out-of-plane bend begins to become
256 asymmetric at 36 GPa, and splits from one band to three distinct modes (ν_2 , ν_{2a} , ν_{2b}) above 38
257 GPa. The highest frequency in-plane bending peak (ν_{4d}) sharpens across the transition to
258 bastnäsite-III; the two components (ν_{4c} , ν_{4d}) of this peak become difficult to deconvolve above
259 38 GPa. The next highest frequency in-plane bending peak (ν_{4b}) remains sharp and distinct
260 across the transition, but a change in its peak shape is apparent: the top of the peak remains
261 sharp, but the bottom broadens, and there may be additional, not-readily-resolved modes on its
262 high and low frequency sides. Accordingly, two of the three (and possibly all three) of the
263 different types of carbonate vibrations monitored have additional bands that appear following the
264 transition at 38 GPa.

265 The Raman spectrum of bastnäsite-III reflects both distorted carbonate ions and
266 potentially multiple REE sites that persist to the highest pressures probed (52 GPa). While
267 Raman spectroscopy is an excellent probe for characterizing changes in local bonding
268 environments, we supplement these results with Nd³⁺ luminescence results to characterize the
269 electronic transitions of Nd in the REE site in order to better constrain structural changes that
270 specifically involve the REE ion.

271 **High Pressure Nd³⁺ Luminescence Results and Discussion**

272 Figures 5-7 show both our luminescence results within different spectral regions, and the
273 ambient pressure energy level diagrams for the associated transitions. In each instance, we use
274 the peak nomenclature of Dieke (1969). The majority of the transitions have red shifts of their

275 respective luminescence with pressure – such an approach of long-lived luminescing electronic
276 states towards the ground state under compression is common, such as for Nd^{3+} in garnets (Hua
277 et al. 1996) and Cr^{3+} in ruby (Mao et al. 1986), although exceptions exist (e.g., Hua and Vohra
278 1997; Hernández-Rodríguez et al. 2018a). The most intense and highest energy bands associated
279 with the transition between the ${}^4\text{F}_{3/2}$ state and the ${}^4\text{I}_{9/2}$ ground state (which are, by convention,
280 named R1 and R2: Dieke 1969), do not redshift or blueshift notably; while they may have a
281 small and non-linear pressure dependence, they shift little with pressure up to 50 GPa (Fig. 5).
282 Thus, the pressure-induced shifts of the ${}^4\text{F}_{3/2}$ states are similar to that of the lowest energy ${}^4\text{I}_{9/2}$
283 state.

284 Overall, the luminescence changes subtly at the transition between bastnäsite-I and –II
285 compared to the changes that occur across the bastnäsite-II to –III transition; this is likely due to
286 the subtle nature of the first transition, and a more dramatic change in the REE site in the second
287 transition (possibly involving a doubling of the number of REE sites in the cell). We discuss
288 each of the major transition manifolds in sequence, since each show somewhat different behavior
289 under compression.

290 ${}^4\text{F}_{3/2} \rightarrow {}^4\text{I}_{9/2}$ luminescence under pressure

291 We observe all 10 peaks associated with the transitions between the Stark levels for the
292 ${}^4\text{F}_{3/2}$ and ${}^4\text{I}_{9/2}$ states. The R1 and R2 bands provide a direct measure of the excited state (${}^4\text{F}_{3/2}$)
293 energy above the ground state. Under pressure, up until ~ 15 GPa, we see a decrease in their peak
294 separation (Fig. 8). However, near 15 GPa this trend reverses, and the peaks begin to separate.
295 Because the Stark splitting is produced by the local electric field at the Nd ion, this turnover is
296 associated with an initial decrease in the electric field gradient, followed by an increase, or an
297 onset of a more anisotropic local bonding environment. Such a change may be indicative of a

298 change in local bonding, and hence a shift in compressional behavior within the REE site, but
299 does not necessarily indicate any change in crystal symmetry. The relative intensities of these
300 bands remain relatively unchanged to 25 GPa.

301 From 25 to 38 GPa, most of the luminescence peak positions continue to decrease in
302 energy at rates similar to those prior to the transition to bastnäsite-II; a representative spectrum at
303 31.2 GPa is peakfit in Fig. S5a. Interestingly, only one transition clearly increases in energy
304 above 25 GPa: R2→Z2(3). This anomalous shift indicates that the splitting of the $^4I_{9/2}$ state is
305 complex under compression, and that the electric field that is impacting the f-orbitals is
306 anisotropic in its shape.

307 Above 38 GPa, many of the pairs of luminescence bands induced by the two $^4F_{3/2}$ states
308 split from two to three or four peaks: this splitting strongly suggests that an additional rare earth
309 element site is present within the unit cell of bastnäsite-III, consistent with our Raman spectra,
310 with the new site having a modestly different set of Stark splittings. The number of Stark levels
311 is capped by the low symmetry of the REE site—so any decrease in symmetry of the single REE
312 site would not increase the number of peaks observed (Kaminskii 1981). Thus, the most
313 plausible mechanism for producing additional bands is via the onset of a new REE site within the
314 structure. Above 38 GPa, deconvolution of the luminescence bands and the presence of the new
315 bands provide primary evidence for a new electronic environment within the structure; a
316 representative deconvolved spectrum is shown in Fig. S6a. While we are able to determine the
317 new R1 and R2 energy levels of the new site, we are unable to identify the new site's Z3 level.
318 We believe this is due to the low intensity of the transitions from R1 and R2 to Z3 before the
319 phase transition; thus, after the transition it remains difficult to determine the full suite of peak
320 locations associated with the new site.

321 Interestingly, a new band begins to appear at the highest energies of this suite of
322 transitions (associated with an additional transition from R2→Z1) near 36 GPa—this is likely
323 either a precursor to, or the first indication of the phase change that we observe in the Raman
324 spectra at 38 GPa. The intensity of the peak associated with R1→Z2(4) becomes more intense in
325 bastnäsite-III. The change in intensity indicates that the excitation and/or emission cross sections
326 for those transitions have changed: since we have not measured excitation spectra, we cannot
327 distinguish which (or both) of these alternatives is the case.

328 $^4F_{3/2} \rightarrow ^4I_{11/2}$ luminescence under pressure

329 At room pressure and temperature using 532 nm excitation wavelength, we are able to
330 also observe all 12 peaks associated with the transitions from the $^4F_{3/2}$ to $^4I_{11/2}$ states (Figure 6).
331 At ambient pressure, the most intense peak is associated with the transition from R1→Y2 (14).
332 The relative intensities of the respective peaks associated with transitions between these states is
333 similar to that of Nd³⁺ in yttrium aluminum garnet, with the most intense being peak 14, with
334 generally comparable intensities of the nearby lower amplitude peaks (e.g., Pokhrel et al. 2012).
335 Figure 6 shows the changes produced by pressure, and Table 5 contains the pressure shifts of the
336 peaks. Interestingly, the highest wavelength peaks (lowest energy transitions) associated with R1
337 and R2→Y5 and Y6 are not observable with 633 nm excitation—they were only observable with
338 532 nm excitation (Figure S3). We attribute this to the existence of either rapid non-radiative
339 decay pathways associated with the absorptions that are pumped by the lower energy excitation,
340 or those lower energy transitions being accessible when the higher energy transitions are
341 overpopulated.

342 Under compression, all modes shifted to lower energy with increased pressure, and as
343 with the $^4I_{9/2}$ transitions, the modes with the lowest 1-bar energy exhibited the strongest negative

344 pressure dependence. The lowest energy mode (18) becomes unresolvable above 15 GPa.
345 Around this pressure, an additional peak 13* is observed at slightly lower energy than peak 13.
346 These changes in peak position and intensity may be associated with the change in the sign of the
347 trend of electric field near this pressure that is observed in the R2-R1 separation (Figure 8). At 25
348 GPa, peak 13 becomes unresolvable, and this shift may be associated with the transition from
349 bastnäsite-I to -II. Overall, across the bastnäsite-I to -II transition, there are no major changes in
350 either the number of luminescence peaks or their intensities, thus confirming our inference from
351 the Raman spectra that this transition likely involves (at most) a subtle change to the REE site.
352 Moreover, the pressures shifts of all bands are consistent in their migration to lower energy,
353 indicating that these transitions may be less sensitive to changes in the detailed electric field
354 gradient.

355 Above 25 GPa, the peaks associated with bastnäsite-II continue to broaden and decrease
356 in intensity. As pressure is increased, peak 14 decreases particularly in intensity relative to the
357 other bands in this multiplet. All peaks, except peak 16 (which becomes unresolvable) persist to
358 38 GPa; a representative spectrum at 31.2 GPa is peakfit in Fig. S5b.

359 Above 38 GPa, the relative intensities of individual peaks continue to change. As
360 pressure is increased, the lower energy transitions become less intense relative to the higher
361 energy transitions; specifically, peak 14, which at ambient pressure is the most intense, becomes
362 the third most intense, following peaks 12 and 13. These intensity changes may be due to a
363 splitting of peak 14. At 38 GPa, a lower energy peak branches off of peak 14. This splitting,
364 coupled with the overall drop in intensity of the multiplet of peaks, further indicates that there is
365 likely an additional, distinct REE site in the unit cell above 38 GPa; the deconvolution of a
366 representative spectrum at 50.6 GPa is shown in Fig. S6b.

367 **${}^4F_{5/2}+{}^2H_{9/2}\rightarrow{}^4I_{9/2}$ luminescence under pressure**

368 At ambient pressure and temperature, we observe seven peaks associated with the
369 transition from ${}^4F_{5/2}+{}^4H_{9/2}\rightarrow{}^4I_{9/2}$. There is a well documented overlap between the energy levels
370 of ${}^4F_{5/2}$ and ${}^4H_{9/2}$ (5 in ${}^4H_{9/2}$ and 3 in ${}^4F_{5/2}$), and identification of the individual Stark levels of the
371 ${}^4F_{5/2}+{}^4H_{9/2}$ transitions to the ${}^4I_{9/2}$ state is helped by the fact that these transitions happen solely to
372 the lowest energy state in ${}^4I_{9/2}$ (e.g., Muñoz Santiuste et al. 2017). Additionally, we are able to
373 separate and identify the majority of the individual Stark levels due to differences in intensity,
374 line shape and line width in both sets of transitions (Henderson et al. 1967). Once compressed
375 above ~ 1 GPa, a fifth band associated with the transition from ${}^4H_{9/2}\rightarrow{}^4I_{9/2}$ is observed (Fig. 7).
376 As an aside, we note that the appearance of this sequence of luminescence bands is essentially
377 indistinguishable in both amplitude and absolute position (laser frequency minus reported Raman
378 frequency) from bands that have been attributed to Raman-active hydroxyl stretching vibrations
379 observed under 633 nm excitation in bastnäsite from Pakistan (Frost and Dickfos 2007).

380 Around 2 GPa, the pressure-induced removal of accidental degeneracies in the higher
381 energy transitions becomes apparent, and the appearance of new bands continues as pressure is
382 increased in bastnäsite-I. For example, the intensity of peak 25 decreases, likely also from the
383 loss of degeneracy. As with most other emission bands in this system, the nine bands associated
384 with these transitions all shift to lower energy with pressure up to 25 GPa (Fig. 7).

385 Above 25 GPa, two new bands emerge, and the highest energy mode 23 undergoes an
386 anomalous shift to higher energy with pressure from 25 to 38 GPa. This shift is consistent with
387 the changes in the highest-lying bands associated with the ${}^4F_{3/2}\rightarrow{}^4I_{9/2}$ transitions. All other peaks
388 shift to lower energy with pressure until 38 GPa; a representative deconvolved spectrum at 31.2
389 GPa is shown in Fig. S5c.

390 Bastnäsite-III has broader and lower intensity luminescence peaks for this set of
391 transitions (Fig. 7). This drop in intensity and increase in breadth is consistent with there being
392 an additional REE site within the unit cell of this phase, but resolution of a separate full manifold
393 of bands associated with the new site is not feasible. In addition to the intensity changes, there
394 are discontinuous changes in some of the bands. For example, the lowest energy peak 30
395 becomes unresolvable above the transition, peak 25 shifts discontinuously to a higher energy,
396 and a new band, (peak 26a**), branches out from between peaks 26a and 26a* (Fig. 7, Table 6);
397 the deconvolution of the highest pressure spectrum is shown in Fig. S6c. As with the other
398 luminescent sets of transitions, the higher pressure phase change involves more marked changes
399 in the spectra than the lower pressure transition from bastnäsite-I to -II.

400 **R1-R2 separation**

401 The energy differences between the R1 and R2 bands are perhaps most readily associated
402 with structural changes in the REE site under pressure (Figure 8). While the absolute shift of
403 each energy level with pressure is quite small (Figure 5), the separation in energy between these
404 two bands does change. An increase in energy separation is characteristic of a REE site that is
405 decreasing in symmetry/increasing in distortion, and a decrease indicates the converse: a
406 transition towards higher symmetry/lower distortion (e.g., Hua et al. 1996). From 0 to 15 GPa,
407 the difference between these two modes initially decreases indicating a REE site that is
408 becoming more symmetric or, in other words, R1 and R2 are becoming closer to degenerate (Fig.
409 8). At 15 GPa, the difference between these two energy levels increases with pressure until 25
410 GPa, thus indicating that the REE site is becoming more distorted. Above 25 GPa, the change in
411 energy level separation is negligible until 38 GPa, where the onset of a new REE site occurs. The
412 energy level splitting of the new and old REE site differ dramatically. The most striking change

413 is the energy difference between R1 and R2: at 50 GPa, the difference between R1 and R2 in the
414 original REE site is $\sim 42 \text{ cm}^{-1}$, and in the new site, the difference is $\sim 96 \text{ cm}^{-1}$. While this
415 difference is large, an energy difference between R1 and R2 of above 100 cm^{-1} is common
416 within ambient pressure, highly distorted Nd^{3+} sites. For example, the R2-R1 separation is 123
417 cm^{-1} in CAZGAR ($\text{CaZn}_2\text{Y}_2\text{Ge}_3\text{O}_{12}$; Sardar and Yow 1998) and 116 cm^{-1} in $\text{CaY}_2\text{MgGe}_3\text{O}_{12}$
418 (Sharp et al. 1974). In addition, if site distortion increases under pressure, increased separation is
419 expected, so this doubling of the R2-R1 separation in the new site to nearly 100 cm^{-1} is not
420 outside the bounds of sites observed at ambient pressure in other phases.

421 Separation of energy levels with pressure

422 Over most of the pressure range (15-50 GPa) of this experiment, the luminescence bands
423 within each set of transitions separate under compression, indicating a progressive increase in the
424 crystal field splitting, which is likely associated with a pressure-induced increase in the distortion
425 of the REE site. The change in the separation energy of ${}^4\text{I}_{9/2}$ (Z1-Z5) is $6.03 \text{ cm}^{-1}/\text{GPa}$, ${}^4\text{I}_{11/2}$
426 (Y1-Y4), is $3.61 \text{ cm}^{-1}/\text{GPa}$, ${}^4\text{F}_{3/2}$ (R1-R2) is $0.32 \text{ cm}^{-1}/\text{GPa}$, ${}^4\text{F}_{5/2}$ is $2.81 \text{ cm}^{-1}/\text{GPa}$, and ${}^2\text{H}_{9/2}$ is
427 $3.63 \text{ cm}^{-1}/\text{GPa}$. Figure 9 displays the overall splitting of the spin orbit coupling energy levels,
428 with the barycenters (the arithmetic means) shown as the energy levels (for all but the ground
429 state). Similar rates of changes in the separation of the energy levels ${}^4\text{I}_{9/2}$, ${}^4\text{I}_{11/2}$, and ${}^4\text{F}_{3/2}$ have
430 been observed previously in $\text{YAlO}_3:\text{Nd}$ at high pressures (Hernández-Rodríguez et al. 2018a).
431 Several distinct and separable effects are present in Figure 9. First, the net separation between
432 the Stark levels of the ${}^2\text{H}$, ${}^4\text{F}$ and ${}^4\text{I}$ levels decreases under compression. As shown in Figure 3,
433 this indicates that the splitting of the f-levels induced by Coulombic splitting is reduced under
434 compression. This implies that the interaction between f-orbitals is lessened, indicating that
435 covalency of the REE site in bastnäsite increases under compression (this is a manifestation of an

436 increased nephelauxetic effect). The shift in Coulombic splitting can be semi-quantitatively
437 assessed from Figure 9: the separations of the averages of the H, F, and I levels decrease by ~2%
438 for both the H to I level differences and the F to I level differences between 0 and 40 GPa. The
439 relative tradeoffs between a pressure-induced contraction in the size of the f-orbitals relative to
440 decreased electron localization in these levels induced by increased covalency in generating this
441 change in splitting by Coulombic repulsion are unclear. If, as an endmember upper bound on the
442 maximum shift in electron occupancy, the entire shift in Coulombic splitting is attributed to a
443 change in occupancy, then a decrease of ~1% in electron occupancy in the f-orbitals is indicated
444 by the results of Figure 9: a relatively small, but spectroscopically significant, shift relative to the
445 magnitude of the change in volume of the system (which is of order ~19% at 40 GPa using
446 Rowland's (2017) bulk modulus value). In passing, we note that this semi-quantitative estimate,
447 derived from luminescence spectra, could be rigorously tested via single-crystal diffraction
448 constraints on the pressure-dependence of electron density distributions.

449 In terms of the spin-orbit splittings (Figure 3), the separation between the spin-orbit split
450 levels decreases for the ⁴F levels, but may increase modestly for the ⁴I levels (Figure 9). A
451 decreased electron density induced by enhanced covalency would be generally anticipated to
452 decrease spin-orbit coupling, as is clearly the case for the ⁴F level. We have no simple
453 explanation for the enhanced spin-orbit coupling of the ⁴I level. Lastly, the splittings of the Stark
454 levels are expected to be enhanced by both increased distortion and enhanced crystal field
455 strength (e.g., Tröster 2003). This is clearly observed over the pressure range of our experiments,
456 with the spread between levels (the shaded regions in Figure 9) clearly expanding under
457 pressure. Indeed, the maximum value of Stark splitting has been shown to scale with the crystal
458 field strength at the site (Auzel and Malta 1983).

459 **Anomalous pressure shifts in Nd³⁺ luminescence**

460 Overall, the general expectation is that luminescence lines will shift to lower energy with
461 pressure; as pressure is increased, the ligands bound to the Nd³⁺ ion are brought closer to the
462 Nd³⁺ ion, and all electronic levels move to higher energy, but the overarching trend is that higher
463 lying levels tend to shift less quickly to higher energies than lower levels. Thus, the effect of the
464 change in spatial distance between the REE and ligands causes the energy difference between the
465 ground state and the excited states to become lessened with compaction. Previously, anomalous
466 behavior has been observed in the perovskite structured Nd³⁺ doped YAlO₃: blueshifts of a
467 subset of peaks associated with the transition from ⁴F_{3/2} to ⁴I_{9/2} have been observed (Barnett et al.
468 1973; Hua and Vohra 1997; Hernández-Rodríguez et al. 2018a). These anomalous shifts have
469 been attributed to a combination of increasing crystal field strength (and hence splitting) and
470 hybridization of the f-orbitals (Hernández-Rodríguez et al. 2018a). The vast majority of the
471 luminescence bands shift to lower energy with pressure in bastnäsite. However, two
472 luminescence bands have pressure shifts that involve an increase in energy in bastnäsite-II
473 (²H_{9/2} → ⁴I_{9/2}, peak 23; and ⁴F_{3/2} → ⁴I_{9/2}, peak 3), and two increases in bastnäsite-III: the new mode
474 associated with R2 peak 1 and peak 6 in ⁴F_{3/2} → ⁴I_{9/2}. Notably, these bands are among the higher
475 energy lines associated with each of their respective transitions; hence, the role of enhanced
476 Stark splitting counterbalances the general decrease in energy of the barycenter of their
477 respective transitions. Thus, these atypical shifts can be attributed to complex crystal field
478 interactions between the Nd³⁺ ion and O/F ligands induced by an enhanced site distortion, and
479 hence anisotropic compression within the REE site.

480 **Implications**

481 The bonding of both the carbonate units and neodymium ions under pressure is
482 constrained in this critically important ore mineral. Notably, bastnäsite undergoes two apparently
483 distortional phase transitions under compression: the first, near 25 GPa, appears to involve
484 primarily a shift in the symmetry of the carbonate ion, while the second, which initiates at 38
485 GPa, generates a second REE site in this material. From an overarching perspective, the overall
486 layered bonding of the bastnäsite structure, with REE-F layers and carbonate layers, appears
487 remarkably stable under compression: no indication is seen within this pressure range of any
488 weakening of C-O bonds that would indicate an incipient increase in coordination, or onset of
489 dimerization associated with the carbonate group (e.g., Vennari et al. 2018). The general
490 structural stability of this phase indicates that REE retention in the Earth's mantle may continue
491 to involve bonding within phases containing both halogens and carbonate ions. Hence, REE
492 element cycling, which has been used to constrain mantle differentiation, may critically depend
493 on the presence of trace carbonate- and/or halogen-bearing phases at depth. Thus, the affinity of
494 REE for carbonatite-associated lithologies in the near surface may persist to depth within the
495 planet.

496 From a high-pressure science perspective, bastnäsite (which is quite stable under
497 compression at 300 K) could represent a candidate for a high pressure luminescent calibrant in
498 the near IR region from 0 to ~38 GPa: the transition from R1 to Y2 (Band 14), which is an
499 intense mode in the NIR region does not split under pressure and shifts robustly and
500 continuously through the transition from bastnäsite-I and -II. Furthermore, the splitting between
501 two of the other intense peaks (R1 and R2) may provide a sensitive probe of the degree of
502 distortion of REE sites within a range of REE-bearing oxides and halides under compression.

503

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622

623

624

Tables

625

Table 1. Raman modes as a function of pressure and calculated mode Grüneisen

626

parameters of bastnäsite-I up to 25 GPa. *Bulk modulus of 105 GPa used from the La/F

627

bastnäsite endmember (Rowland 2017).

ν_0 (cm ⁻¹)	Assignment	dv/dP (cm ⁻¹ /GPa)	Grüneisen parameter*
112	Lattice mode	0.59(9)	0.53
129	Lattice mode	1.95(4)	1.59
162	Lattice mode	2.53(16)	1.64
188	Lattice mode	2.74(14)	1.54
238	Lattice mode	3.63(28)	1.60
258	Lattice mode	4.19(16)	1.70
272	Lattice mode	4.36(18)	1.68
303	Lattice mode	4.33(14)	1.50
350	Lattice mode	3.90 (24)	1.17
396	Lattice mode	4.35(14)	1.15
-	(CO ₃) ²⁻ in-plane bend (ν_{4a})	0.89(10)	-
684	(CO ₃) ²⁻ in-plane bend (ν_{4b})	1.28(6)	0.20
726	(CO ₃) ²⁻ in-plane bend (ν_{4c})	1.51(8)	0.22
734	(CO ₃) ²⁻ in-plane bend (ν_{4d})	1.60(6)	0.23
864	(CO ₃) ²⁻ out-of-plane bend (ν_2)	-	-
868	(CO ₃) ²⁻ out-of-plane bend (ν_2)	-0.29(8)	-0.04
1094	(CO ₃) ²⁻ symmetric stretch (ν_1)	2.79(13)	0.30
1434	(CO ₃) ²⁻ asymmetric stretch (ν_3)	-	-

628

629

630 Table 2. Raman modes as a function of pressure of bastnäsite-II from 25 to 38 GPa. * ν_0 is at 25
 631 GPa.

ν_0 (cm ⁻¹)*	Assignment	d ν /dP (cm ⁻¹ /GPa)
128	Lattice mode	0.52(3)
175	Lattice mode	1.66(4)
228	Lattice mode	1.77(6)
284	Lattice mode	1.76(8)
322	Lattice mode	2.20(11)
362	Lattice mode	2.61(9)
403	Lattice mode	2.92(11)
436	Lattice mode	2.82(9)
444	Lattice mode	2.88(8)
473	Lattice mode	3.35(15)
488	Lattice mode	2.92(8)
506	Lattice mode	3.06(1)
687	(CO ₃) ²⁻ in-plane bend (ν_{4a})	0.87(6)
717	(CO ₃) ²⁻ in-plane bend (ν_{4b})	0.98(2)
765	(CO ₃) ²⁻ in-plane bend (ν_{4c})	1.39(3)
775	(CO ₃) ²⁻ in-plane bend (ν_{4d})	1.09(7)
858	(CO ₃) ²⁻ out-of-plane bend (ν_2)	-0.49(2)
1158	(CO ₃) ²⁻ symmetric stretch (ν_{1a})	1.66(7)
1164	(CO ₃) ²⁻ symmetric stretch (ν_1)	2.09(7)

632

633 Table 3. Raman modes as a function of pressure of bastnäsite-III from 38 to 50 GPa. * ν_0 is at 38
 634 GPa.

ν_0 (cm ⁻¹)*	Assignment	dv/dP (cm ⁻¹ /GPa)
119	Lattice mode	0.007(116)
127	Lattice mode	0.32(8)
139	Lattice mode	0.40(4)
203	Lattice mode	1.39(8)
258	Lattice mode	1.59(5)
271	Lattice mode	1.84(12)
367	Lattice mode	2.50(17)
404	Lattice mode	1.74(4)
425	Lattice mode	2.75(17)
488	Lattice mode	2.07(29)
537	Lattice mode	2.65(6)
558	Lattice mode	2.43(5)
703	(CO ₃) ²⁻ in-plane bend (ν_{4a})	0.78(4)
733	(CO ₃) ²⁻ in-plane bend (ν_{4b})	0.85(4)
794	(CO ₃) ²⁻ in-plane bend (ν_{4d})	1.23(5)
848	(CO ₃) ²⁻ out-of-plane bend (ν_{2a})	-0.68(9)
852	(CO ₃) ²⁻ out-of-plane bend (ν_2)	-0.31(4)
864	(CO ₃) ²⁻ out-of-plane bend (ν_{2b})	-0.007(162)
1186	(CO ₃) ²⁻ symmetric stretch (ν_{1a})	1.54(5)
1192	(CO ₃) ²⁻ symmetric stretch (ν_{1b})	1.79(8)
1197	(CO ₃) ²⁻ symmetric stretch (ν_1)	1.81(7)

635

636 Table 4. ${}^4F_{3/2} \rightarrow {}^4I_{9/2}$ luminescence as a function of pressure in bastnäsite.

Assignment	λ_0 (nm)	ν_0 (cm ⁻¹)	Bastnäsite-I to 25 GPa (cm ⁻¹ /GPa)	Bastnäsite-II 25 to 40 GPa (cm ⁻¹ /GPa)	Bastnäsite-II 40 to 50 GPa (cm ⁻¹ /GPa)
1*					1.05(33)
1	862.91	11,588.7	0.42(2)	-0.27(11)	-0.12(7)
2	864.96	11,561.2	0.20(9)	-0.31(16)	-0.83(11)
3	868.47	11,514.5	-1.84(24)	1.43(48)	-1.13(13)
4	870.63	11,485.9	-1.67(10)	-0.48(2)	-0.87(29)
4*					-1.76(37)
5	874.51	11,435.0	-3.51(26)	-1.14(94)	-1.37(88)
6	876.80	11,405.1	-3.35(13)	-1.61(15)	0.16(62)
7*					-2.76(36)
7	887.95	11,261.9	-5.99(23)	-3.57(14)	-4.28(7)
8	889.68	11,240.0	-6.40(26)	-4.63(15)	-4.08(24)
9	897.84	11,137.8	-6.73(19)	-5.26(10)	-5.62(25)
10	900.23	11,108.3	-6.95(8)	-4.93(40)	-4.75(134)
10*					-5.16(2)

637
638

639 Table 5. ${}^4F_{3/2} \rightarrow {}^4I_{11/2}$ luminescence as a function of pressure in bastnäsite.

Assignment	λ_0 (nm)	ν_0 (cm ⁻¹)	Bastnäsite-I to 25 GPa (cm ⁻¹ /GPa)	Bastnäsite-II 25 to 40 GPa (cm ⁻¹ /GPa)	Bastnäsite-II 40 to 50 GPa (cm ⁻¹ /GPa)
11	1046.83	9,552.7	-1.91(11)	-1.26(40)	-2.35(51)
12	1049.86	9,525.1	-2.21(4)	-1.59(5)	-2.19(10)
13*			-2.47(12)	-2.99(14)	-2.33(22)
13	1051.0	9,514.7	-4.00(27)		
14	1054.01	9,487.6	-4.42(14)	-3.08(13)	-2.78(33)
14*					-3.54(50)
15	1056.82	9,462.3	-4.77(7)	-4.57(27)	-5.84(37)
16	1060.09	9,433.2	-4.91(13)		
17	1062.86	9,408.6	-5.26(4)	-6.00(39)	-5.35(86)
18	1065.77	9,382.9	-6.35(27)		

640
641

642 Table 6. ${}^4F_{5/2}+{}^2H_{9/2} \rightarrow {}^4I_{9/2}$ luminescence as a function of pressure in bastnäsite.

Assignment	λ_0 (nm)	ν_0 (cm ⁻¹)	Bastnäsite-I to 25 GPa (cm ⁻¹ /GPa)	Bastnäsite-II 25 to 40 GPa (cm ⁻¹ /GPa)	Bastnäsite-II 40 to 50 GPa (cm ⁻¹ /GPa)
23 (${}^2H_{9/2}$)	791.46	12,634.9	-0.37(8)	0.57(9)	-0.49(18)
24 (${}^2H_{9/2}$)			-0.04(7)	-0.73(7)	-1.30(13)
24* (${}^2H_{9/2}$)			-1.52(33)	-2.01(12)	-1.90(21)
25* (${}^2H_{9/2}$)					-2.00(24)
25 (${}^2H_{9/2}$)	793.65	12,600.0	-1.94(6)	-2.51(11)	
26a (${}^2H_{9/2}$)			-2.30(25)	-0.80(19)	-0.60(23)
26a* (${}^2H_{9/2}$)				-1.83(15)	-2.04(33)
26a**					-2.00(37)
26* (${}^2H_{9/2}$)				-3.72(29)	-0.73(70)
26 (${}^2H_{9/2}$)	798.04	12,530.7	-5.05(47)	-3.20(11)	-3.08(75)
27 (${}^2H_{9/2}$)	803.34	12,448.0	-3.82(18)	-3.19(37)	2.79(28)
28 (${}^4F_{5/2}$)	807.94	12,377.2	-4.41(68)	-3.93(54)	-3.16(76)
29 (${}^4F_{5/2}$)	814.90	12,271.4	-5.98(32)	-4.36(21)	-0.93(43)
30 (${}^4F_{5/2}$)	822.71	12,155.0	-7.22(14)	-4.99(22)	

643

644

645

Figure Captions

646 Figure 1. The ambient crystal structure of bastnäsite with the unit cell delineated by black lines;
647 as per the labels, dark green spheres are fluorine, lime green is the REE, black is carbon, and red
648 is oxygen.

649 Figure 2. Ambient pressure and temperature (a) Raman spectrum and (b) luminescence spectrum
650 of bastnäsite; black is the measured spectrum, gray are the deconvolved peaks, red is the
651 calculated intensities, and blue is the background. The relative intensities for the
652 $^4F_{5/2+2}H_{5/2} \rightarrow ^4I_{9/2}$ are increased by 3x, and the $4F_{3/2} \rightarrow ^4I_{9/2}$ section was collected for 6x longer than
653 the other two luminescence regions.

654 Figure 3. (left) Schematic representation of the f-electron energy levels of Nd^{3+} in bastnäsite
655 showing the effects of Coulombic repulsion, spin-orbit coupling, and crystal-field interactions.
656 (right) Partial energy level diagram of Nd^{3+} in bastnäsite showing the major spin-orbit coupling
657 induced energy levels.

658 Figure 4. (a and b) Representative Raman spectra of the lattice and carbonate spectral regions
659 under compression, arrows indicate new peaks in Raman spectrum; (c) deconvolution of the
660 symmetric stretch of the Raman spectrum of bastnäsite at 50.6 GPa, the highest pressure probed;
661 (d and e) peak positions of observed Raman modes as a function of pressure at room temperature
662 under compression.

663 Figure 5. (a) Luminescence spectra of the $^4F_{3/2} \rightarrow ^4I_{9/2}$ transition in Nd^{3+} as a function of pressure.
664 Partial energy level diagram is shown on the right. (b) Peak positions of observed luminescence
665 modes with pressure.

666 Figure 6. (a) Luminescence spectra of the ${}^4F_{3/2} \rightarrow {}^4I_{11/2}$ transition in Nd^{3+} as a function of pressure.
667 Luminescence spectra are produced by 633 nm excitation. (b) Peak positions of the observed
668 luminescence bands with pressure.

669 Figure 7. Luminescence spectra of the ${}^4F_{5/2} + {}^2H_{9/2} \rightarrow {}^4I_{9/2}$ transition in Nd^{3+} as a function of
670 pressure. Luminescence spectra are produced by 633 nm excitation. (b) Peak positions of
671 observed luminescence bands with pressure.

672 Figure 8. R-line separation of Nd^{3+} (R_1 and R_2) in the original REE site (closed circles), and
673 associated with the new high-pressure site in bastnäsite-III (closed triangles), as a function of
674 pressure.

675 Figure 9. Energy levels of Nd^{3+} with pressure; shaded regions designate the observed spread in
676 energy levels, and lines represent the barycenters of the Stark levels, with the exception of the
677 ${}^4I_{9/2}$ level, where the line represents the ground state of the system.

Figure 1

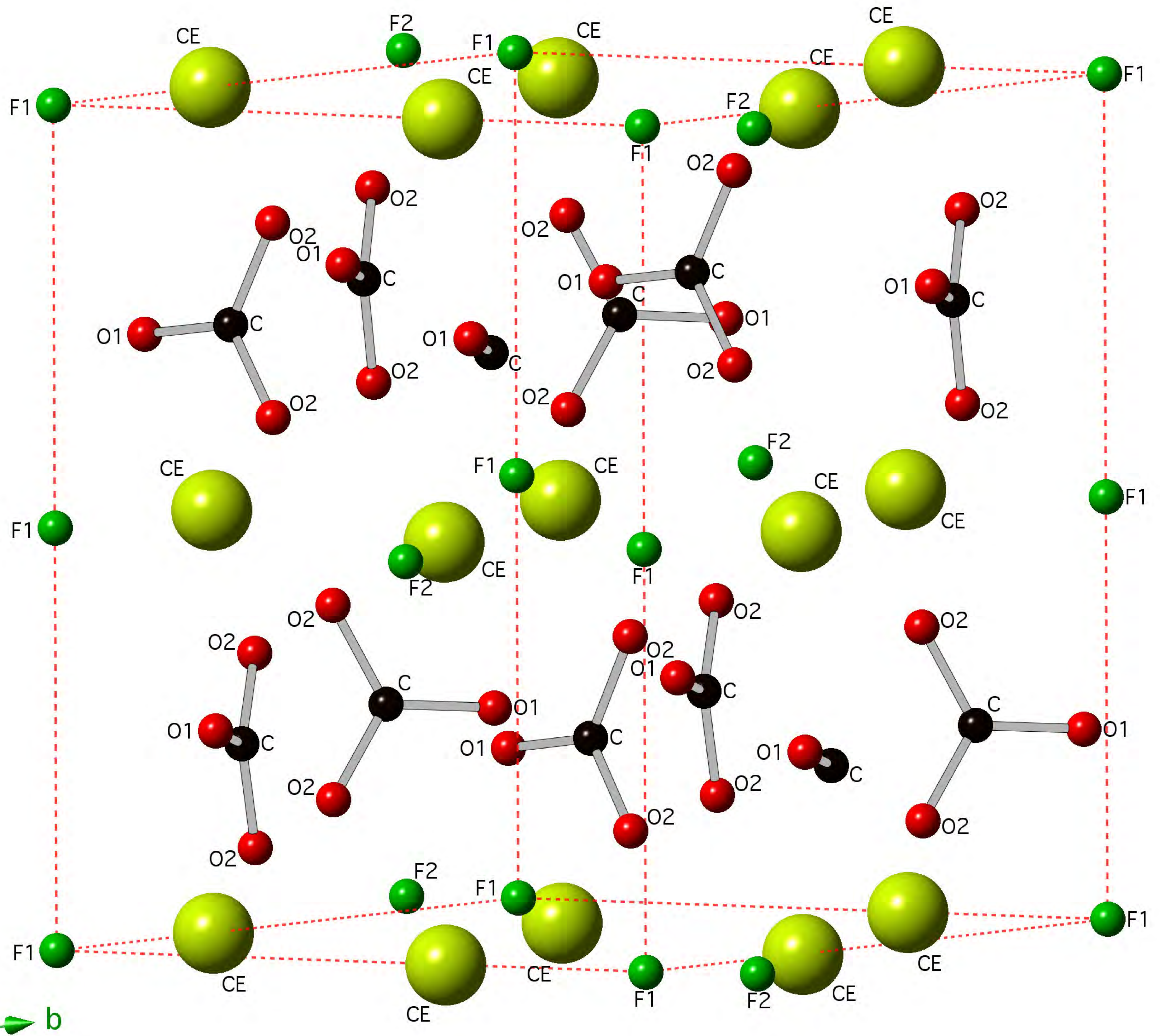
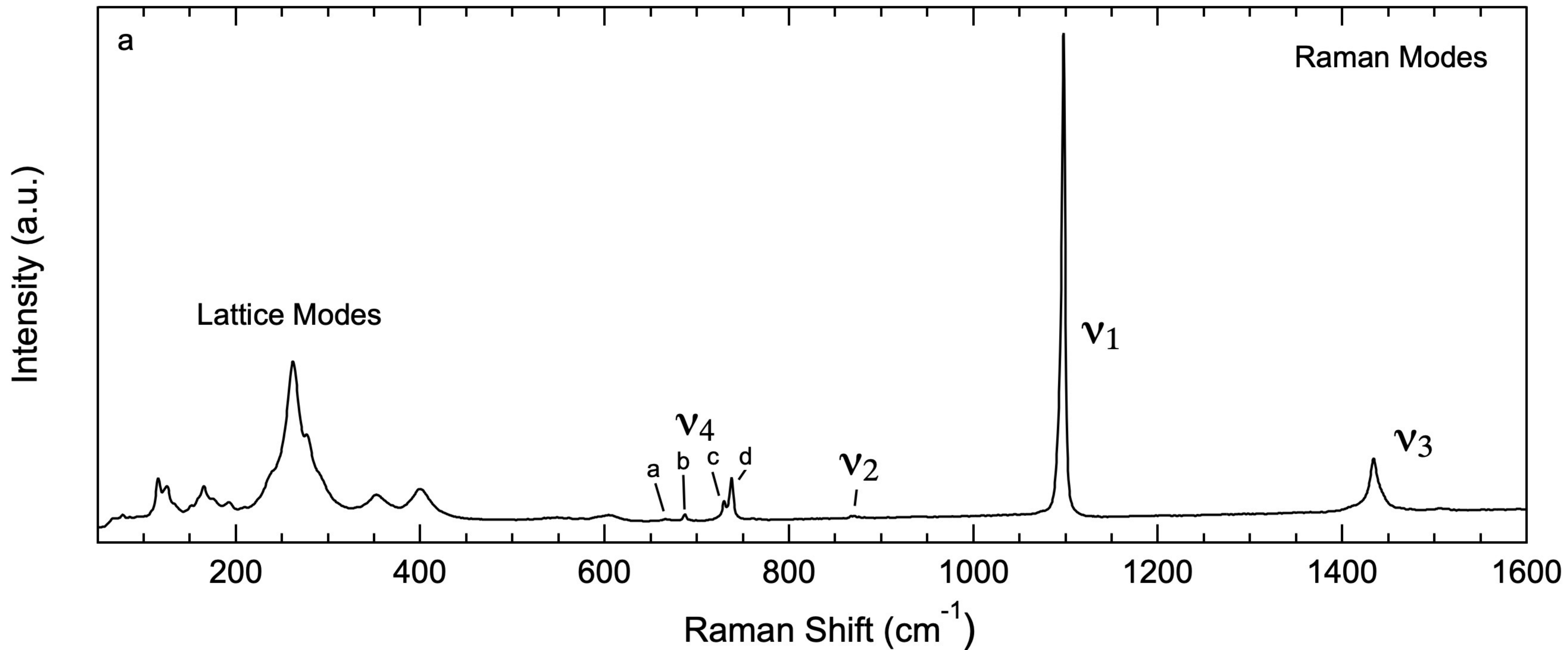


Figure 2a



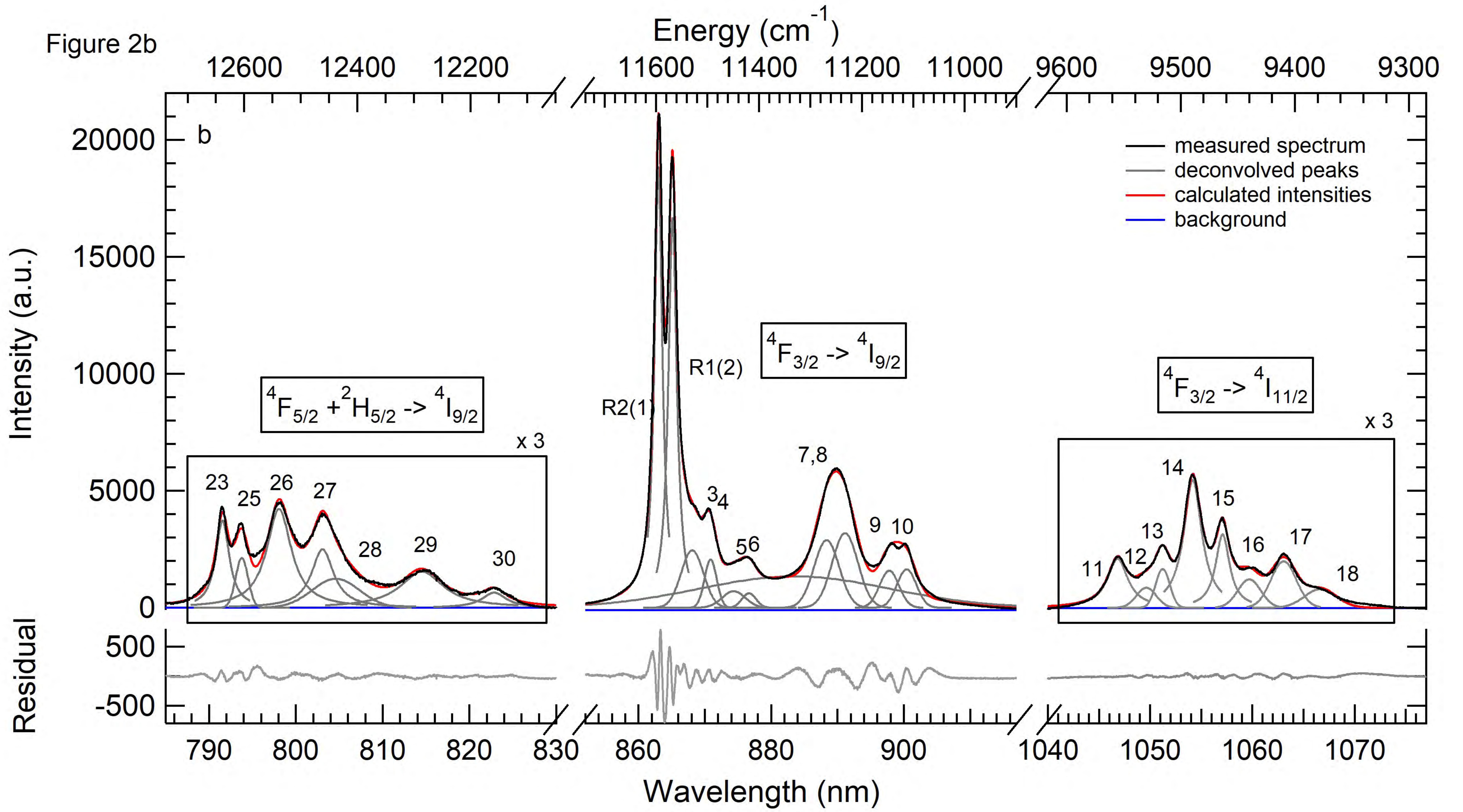


Figure 3

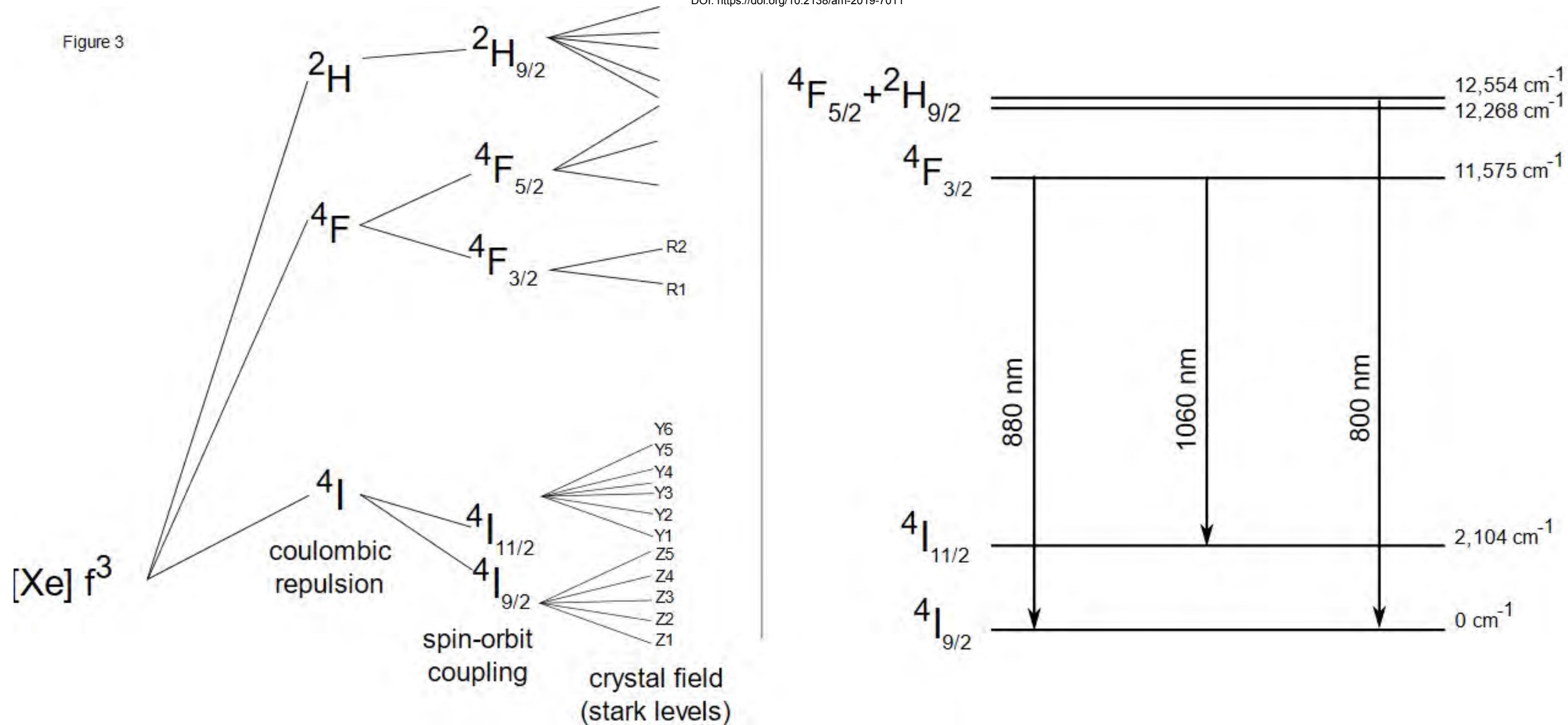


Figure 4a

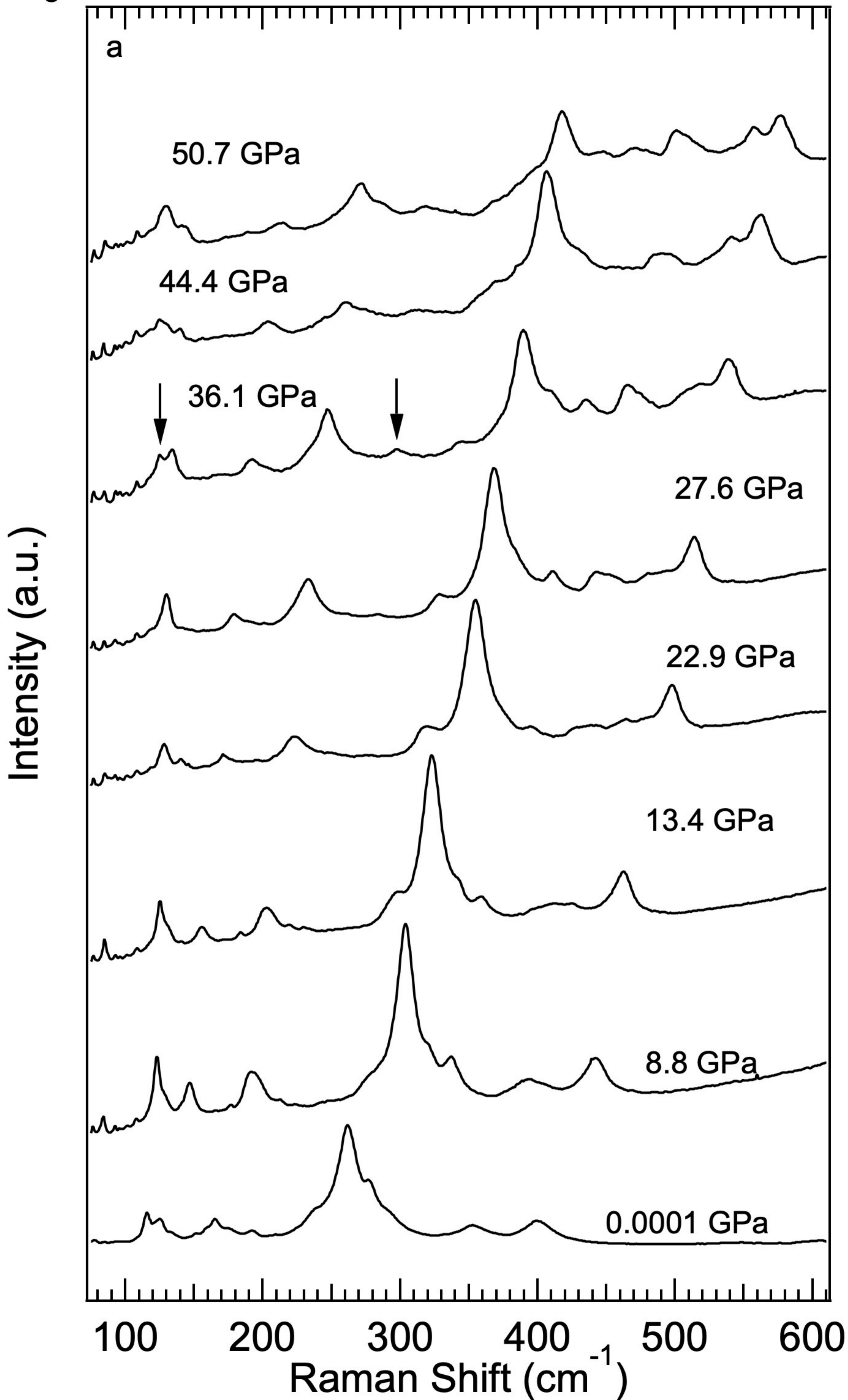
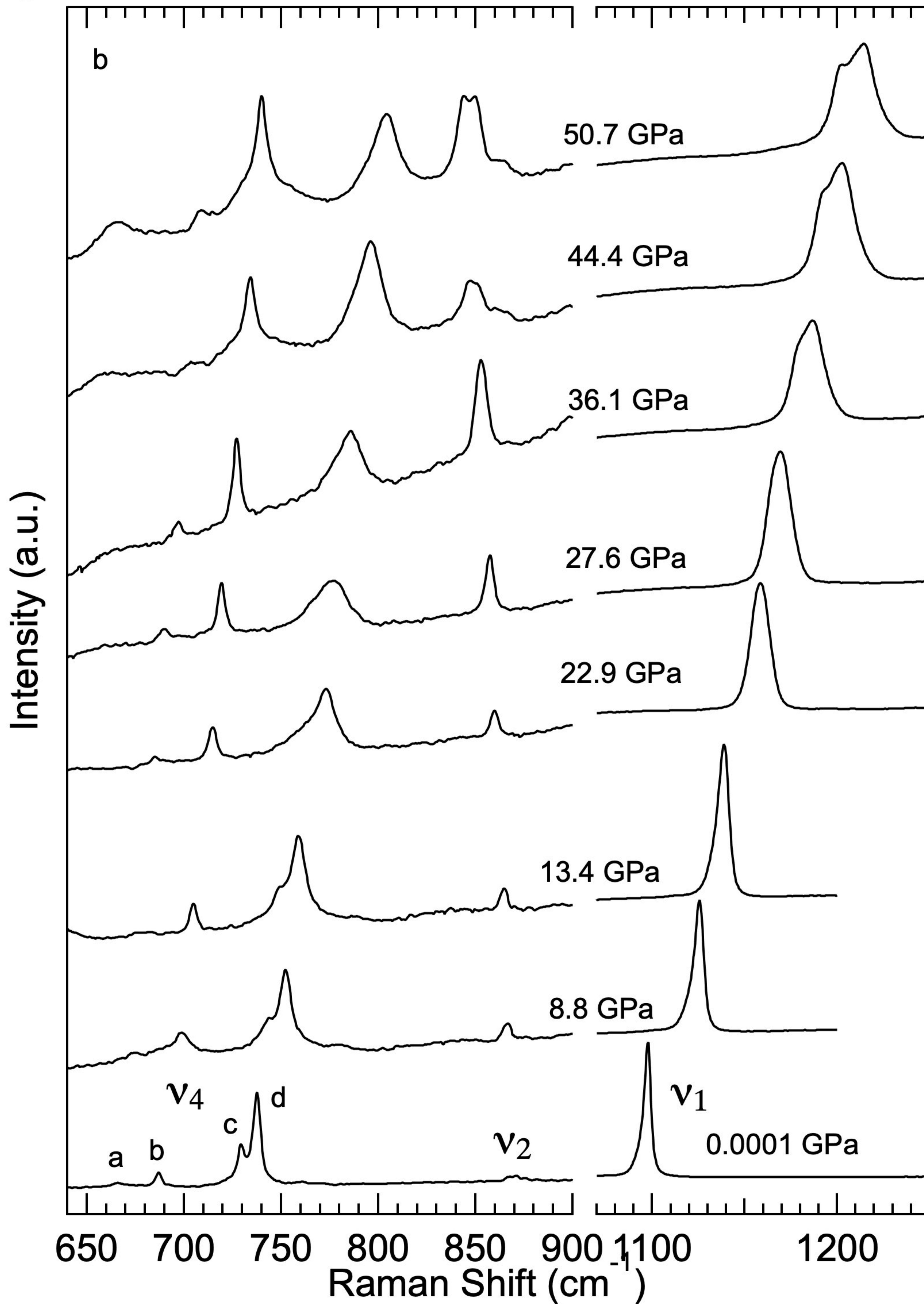


Figure 4b



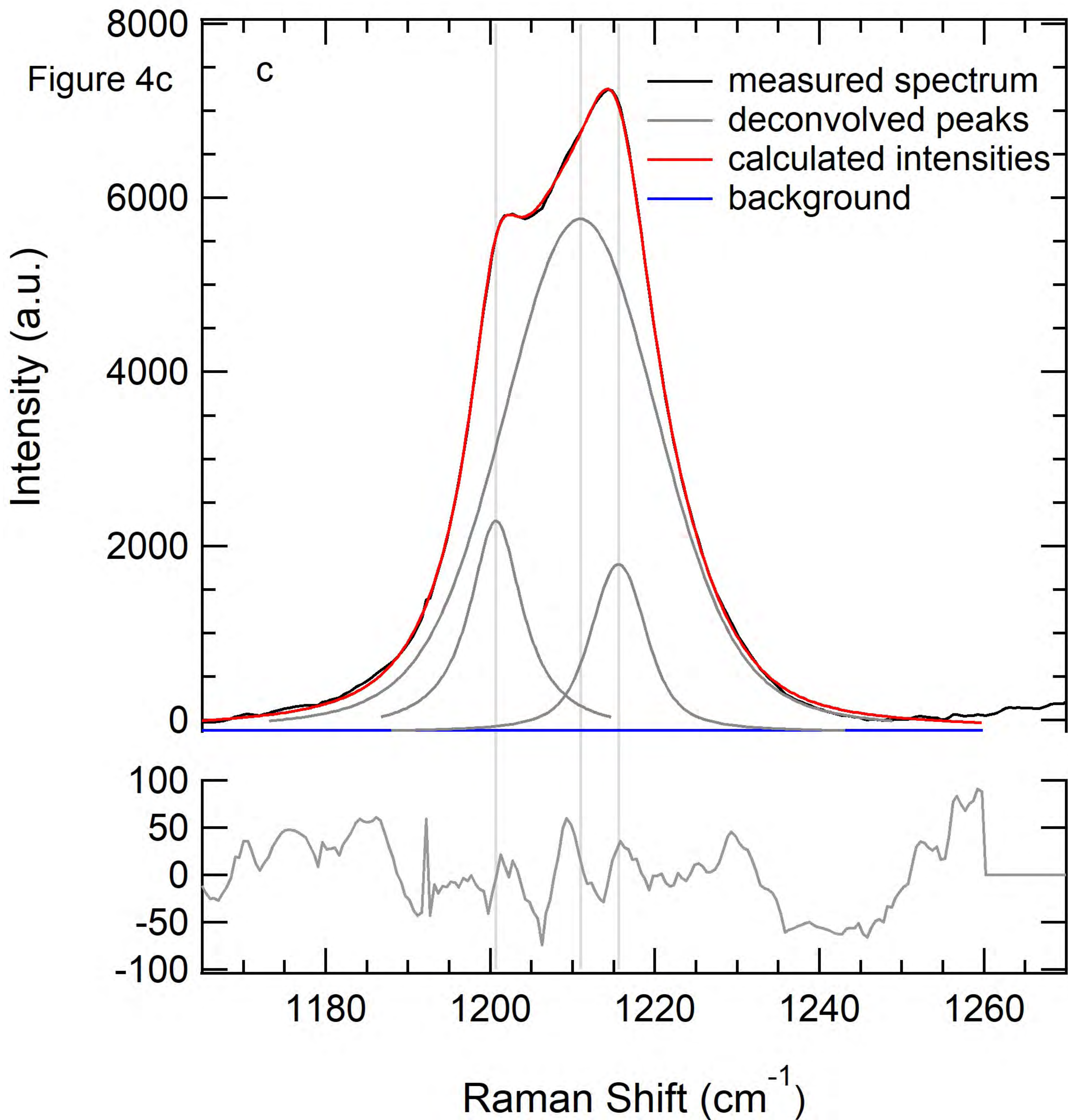


Figure 4d

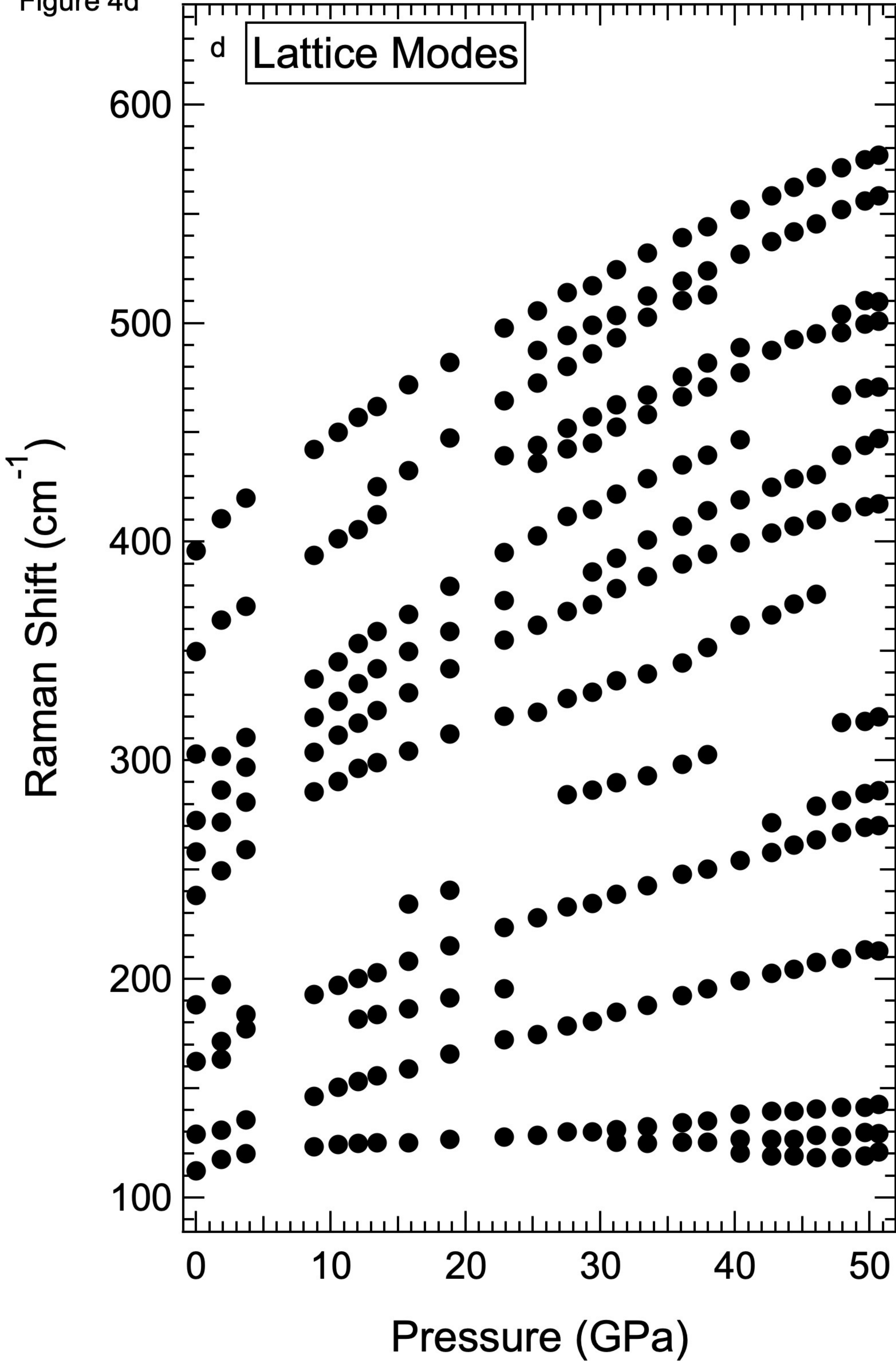


Figure 4e

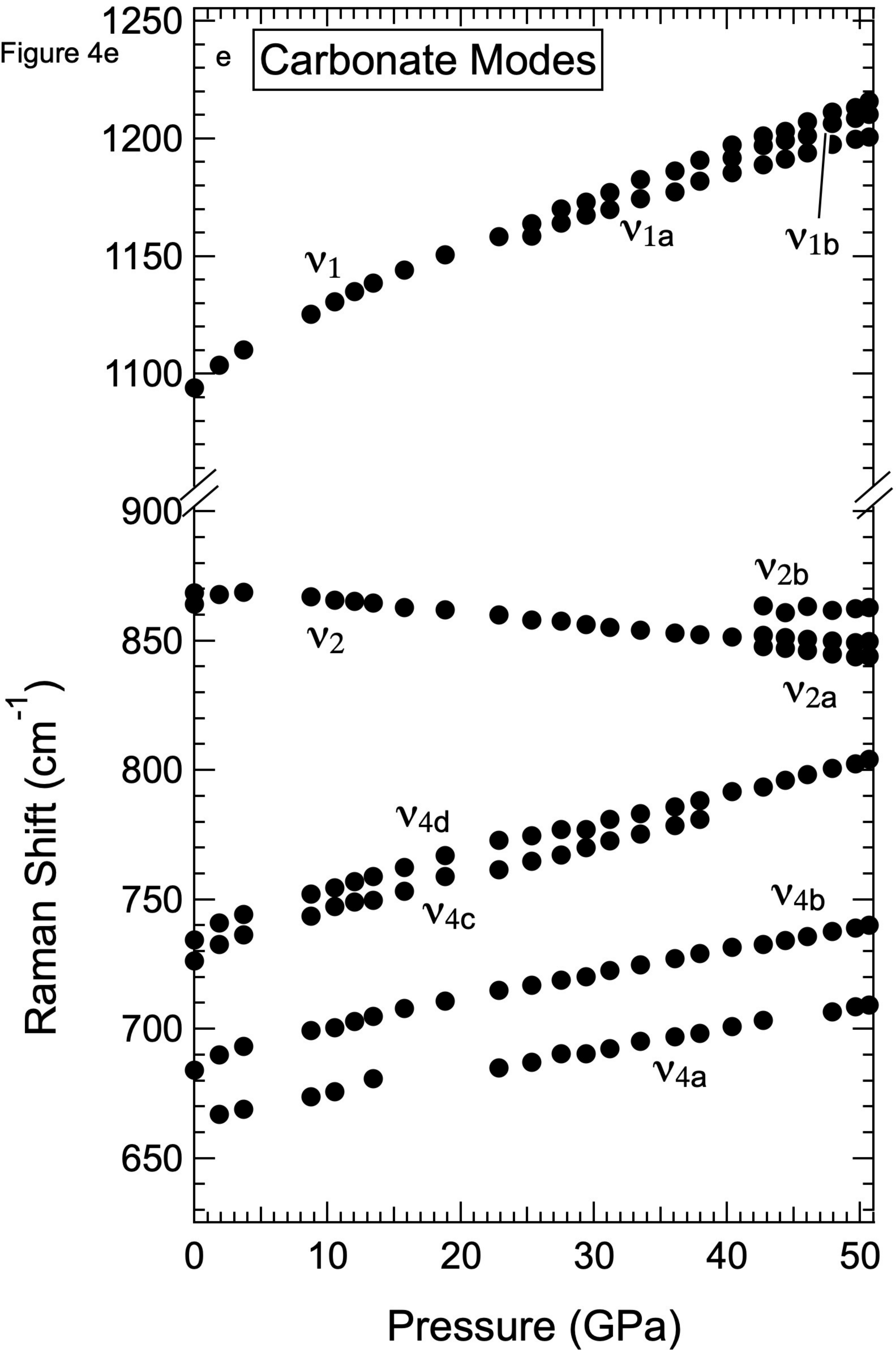


Figure 5ab

Energy (cm⁻¹)

11600 11400 11200 11000 10800

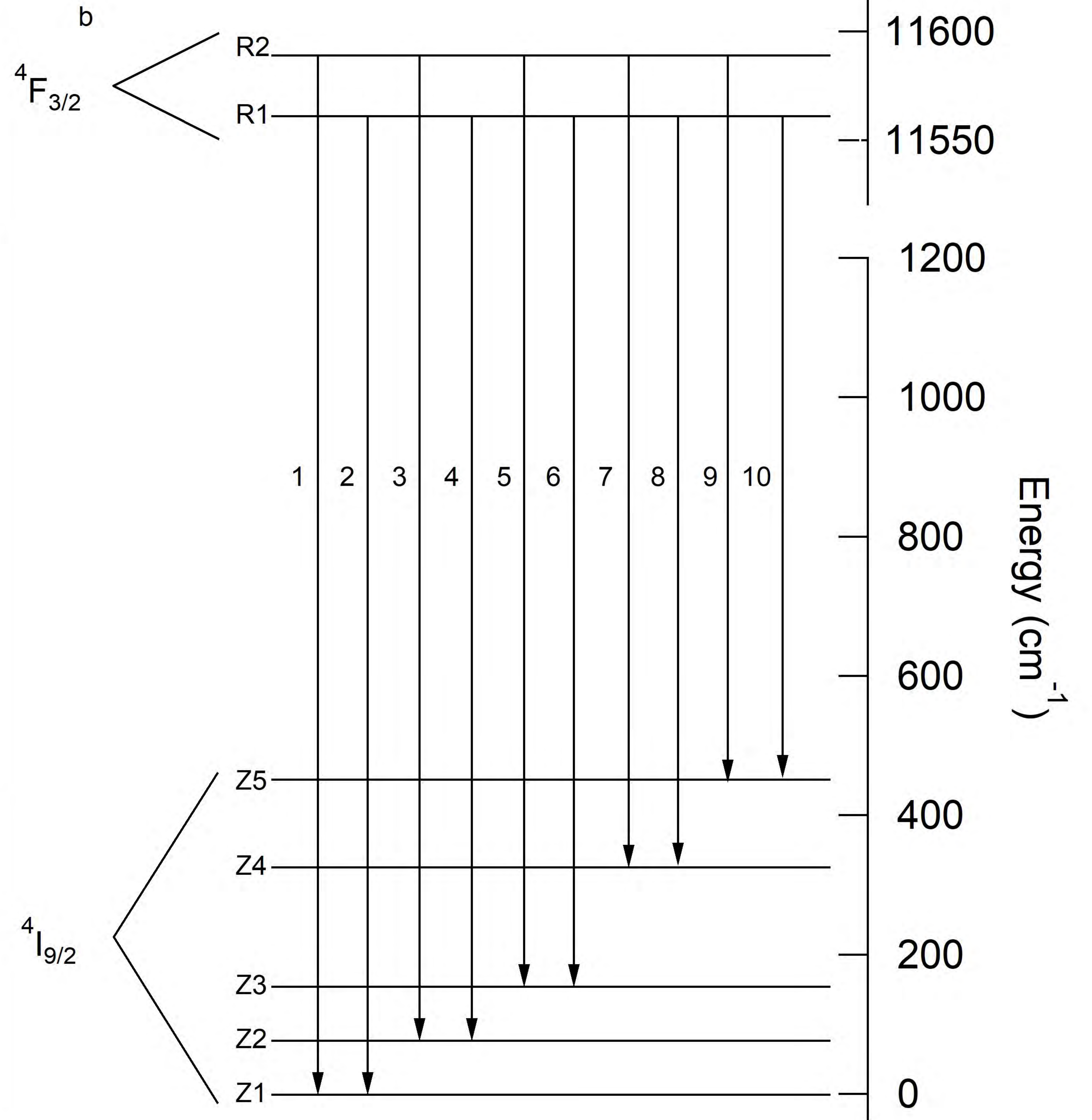
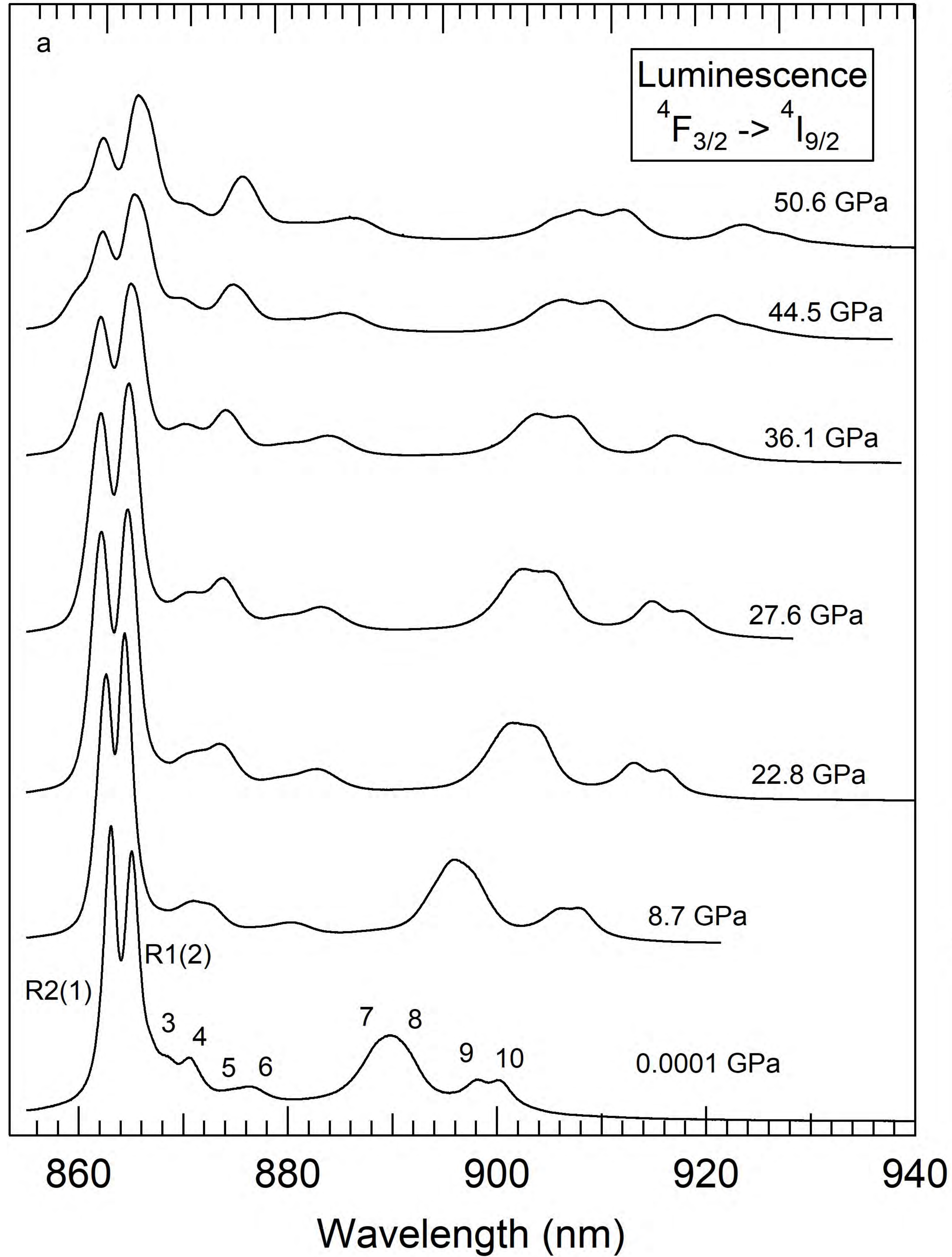


Figure 5c

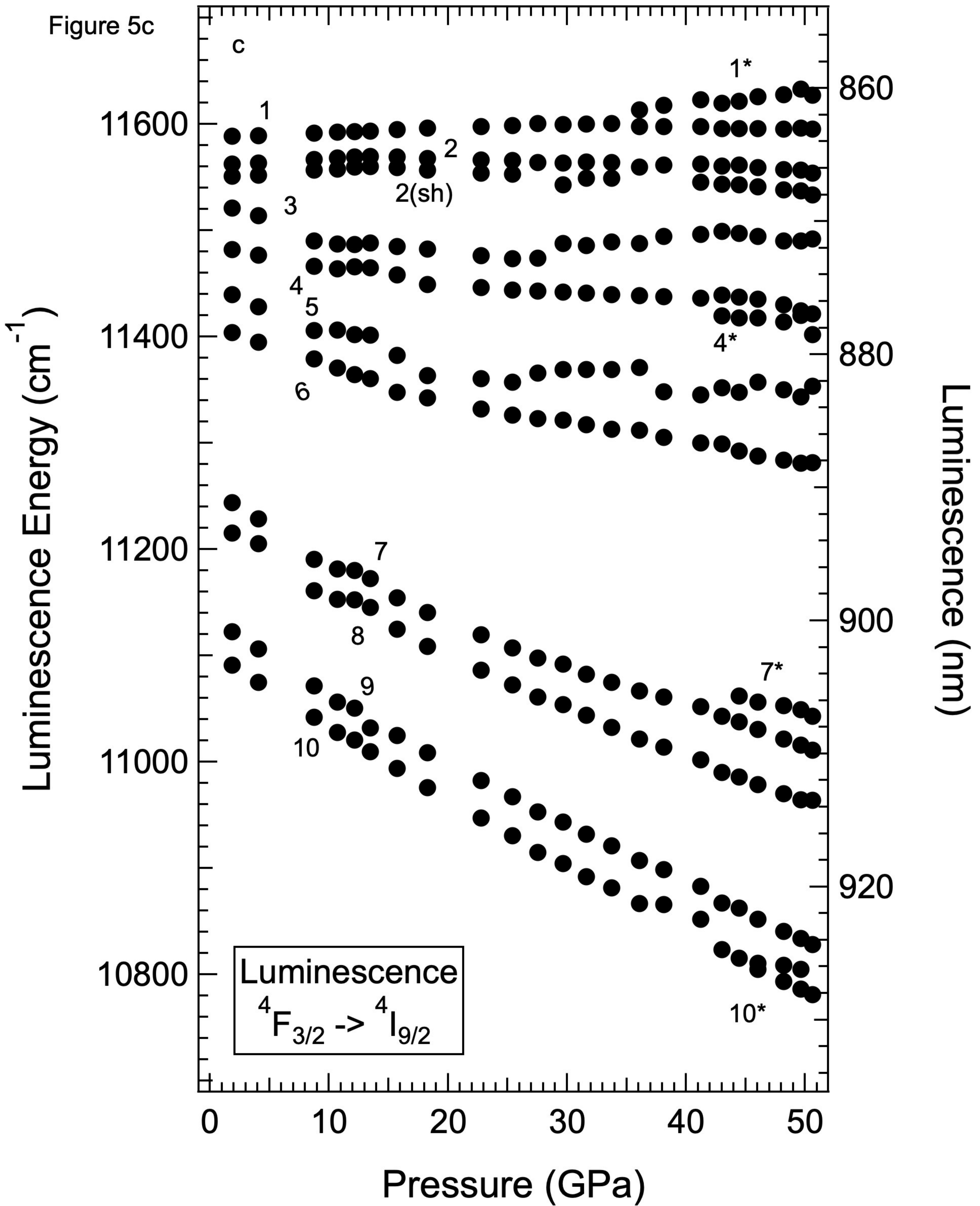


Figure 6ab

Energy (cm^{-1})

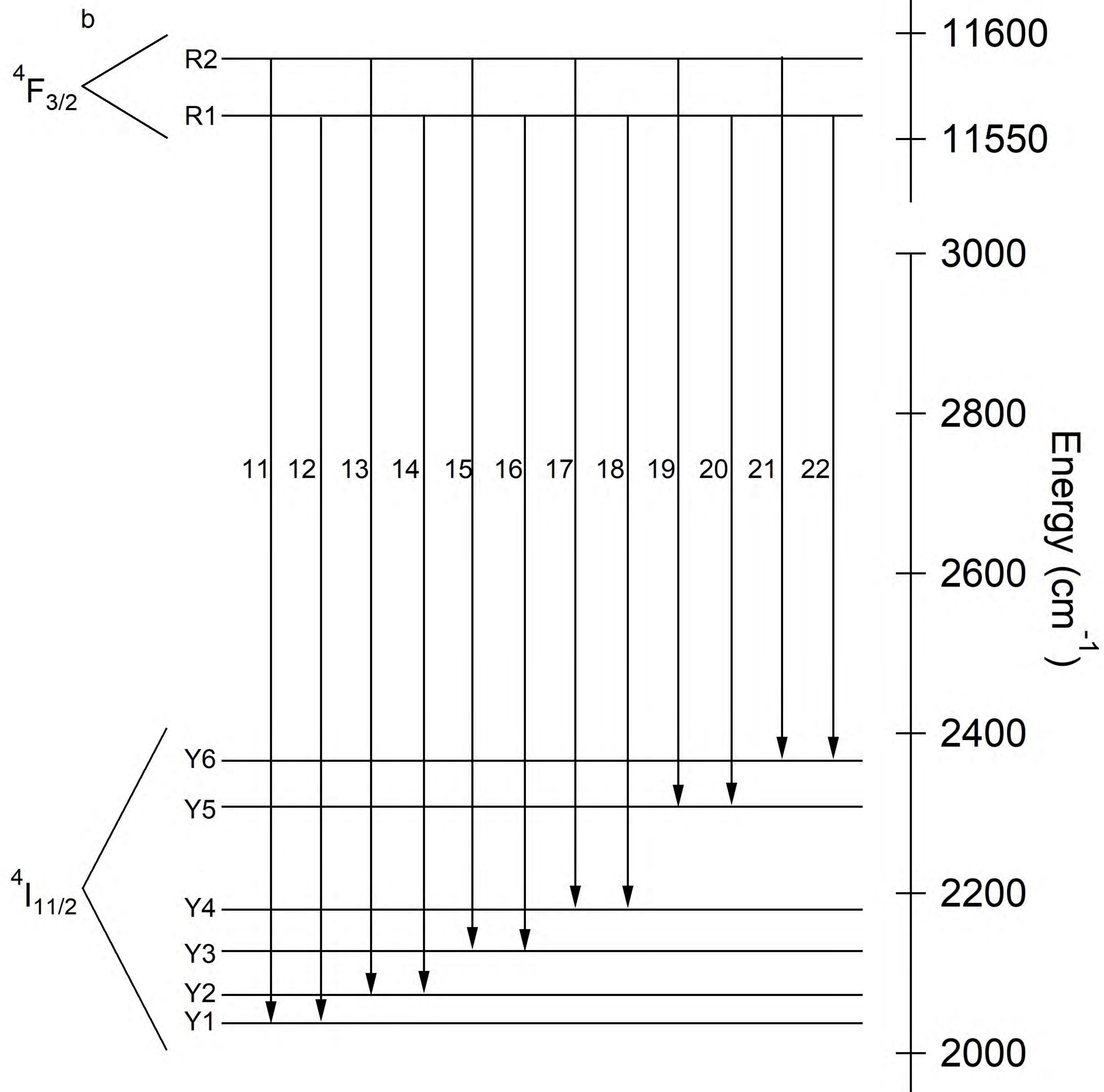
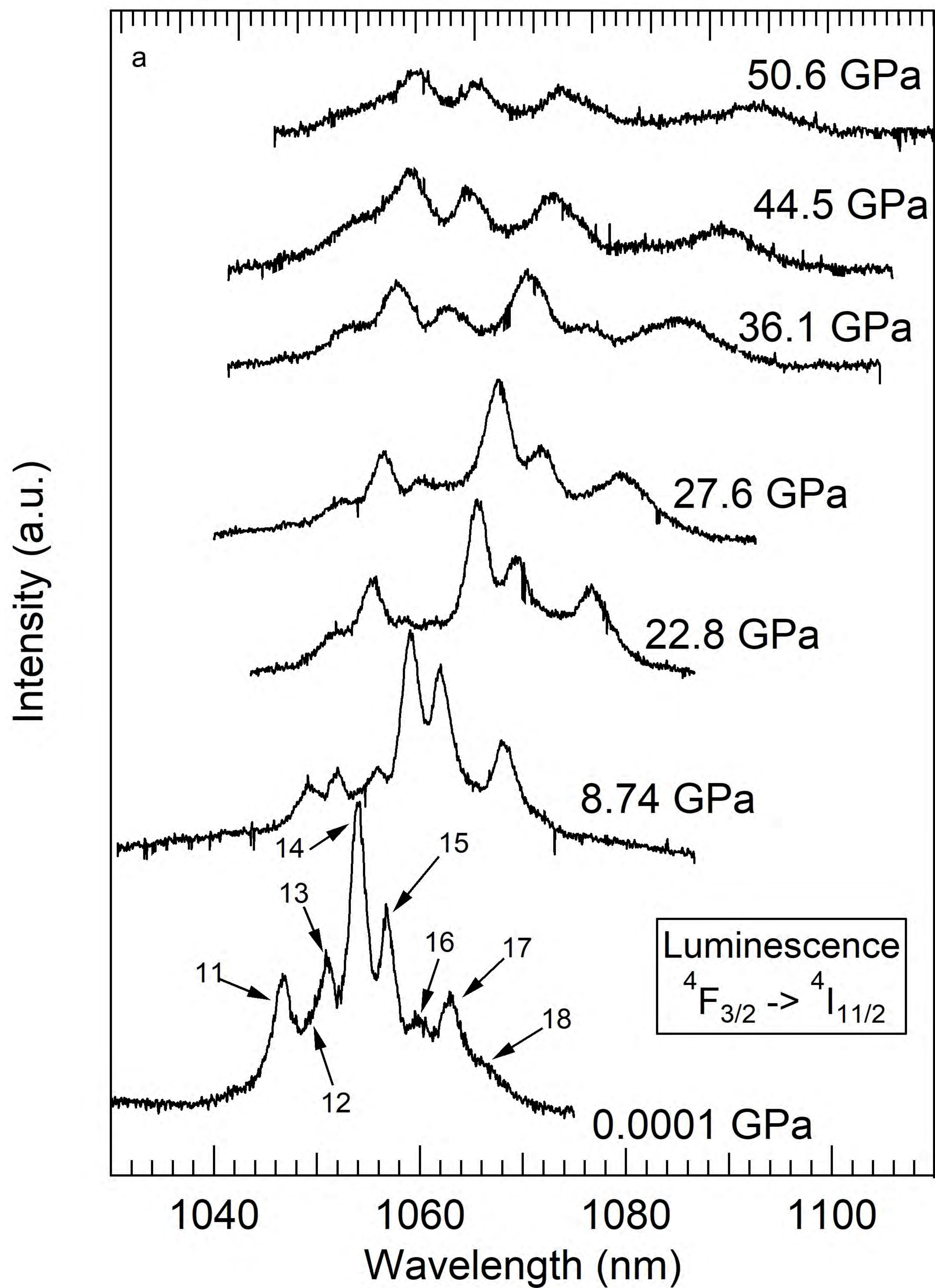


Figure 6c

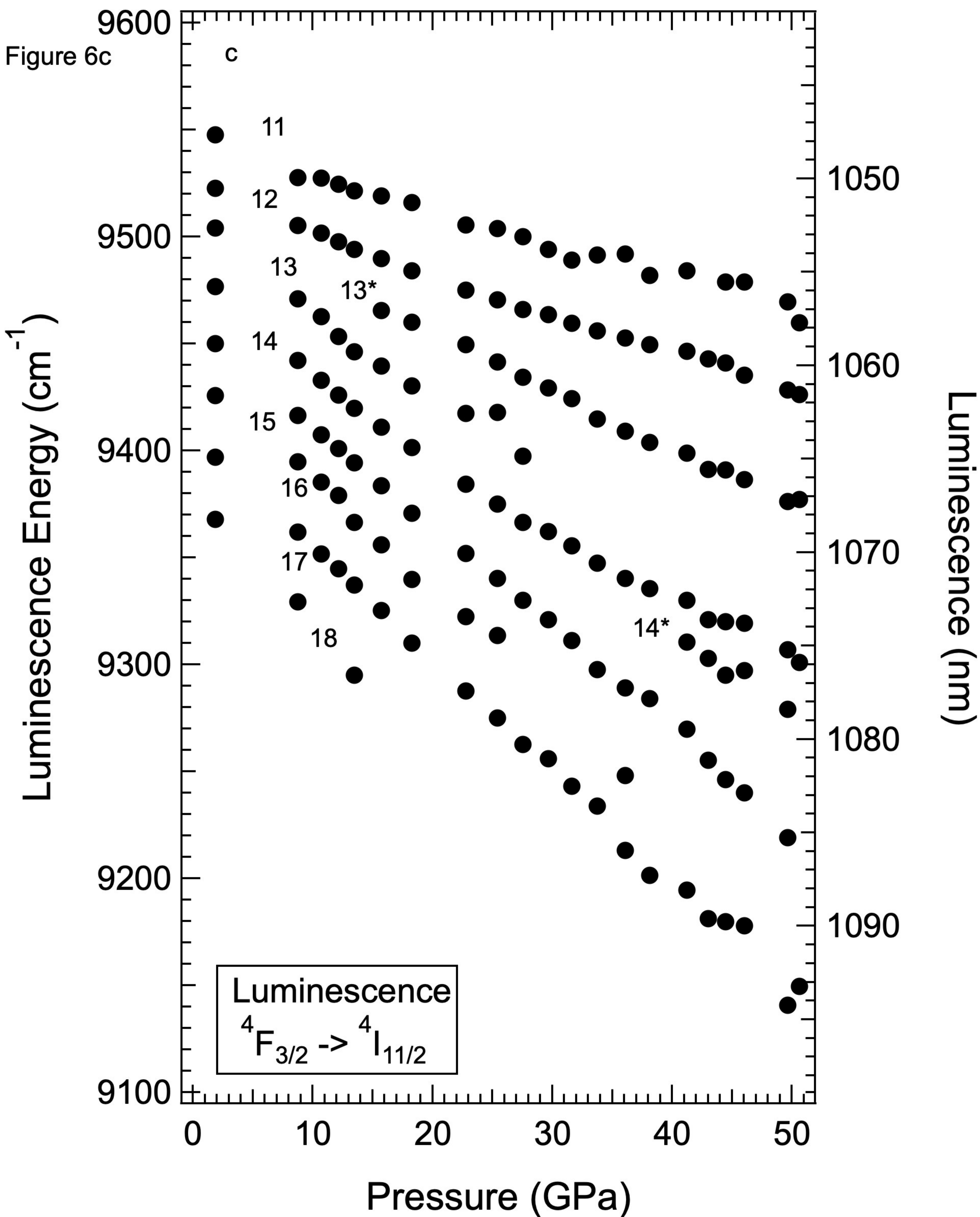


Figure 7ab

Energy (cm⁻¹)

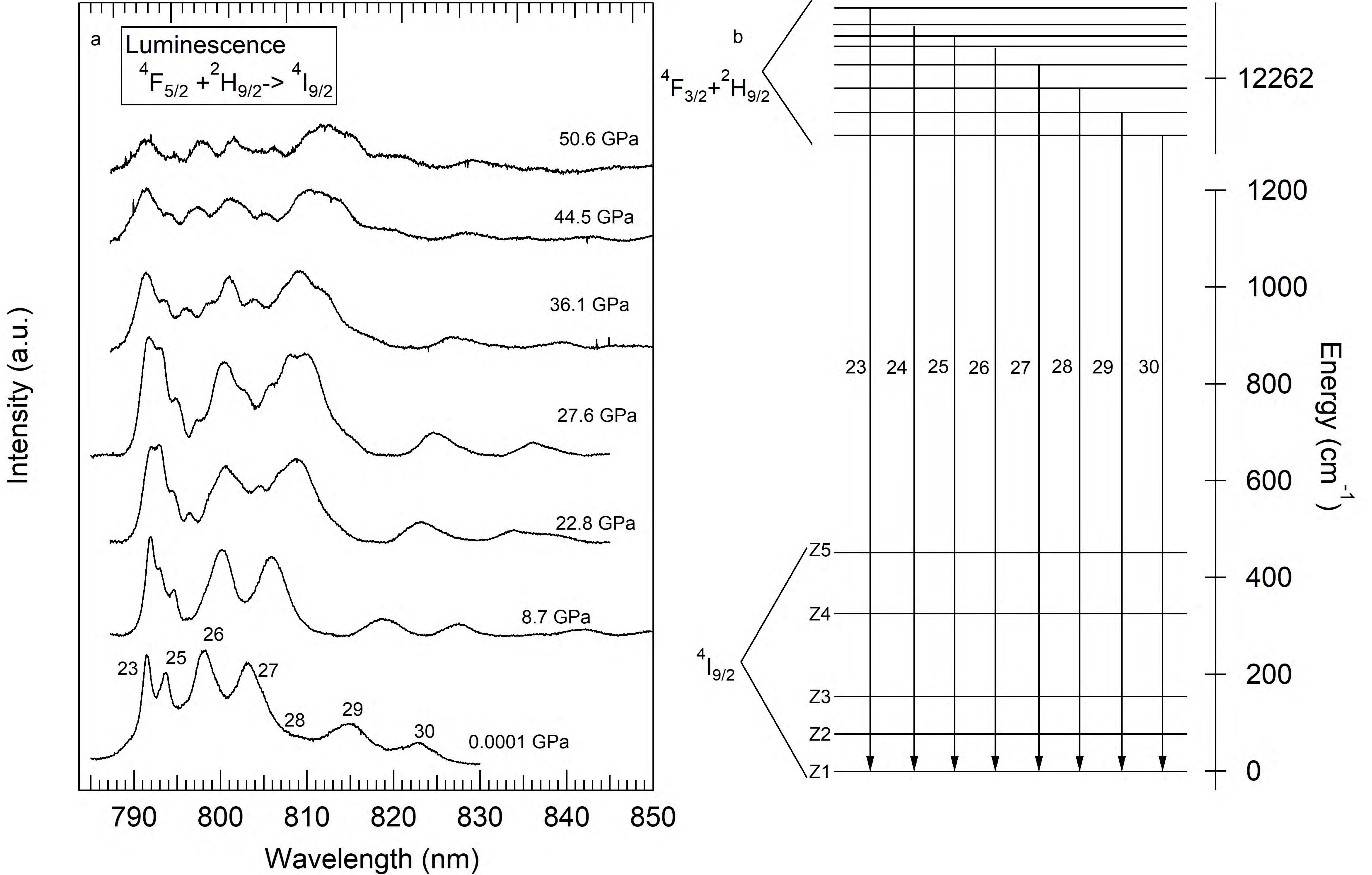


Figure 7c

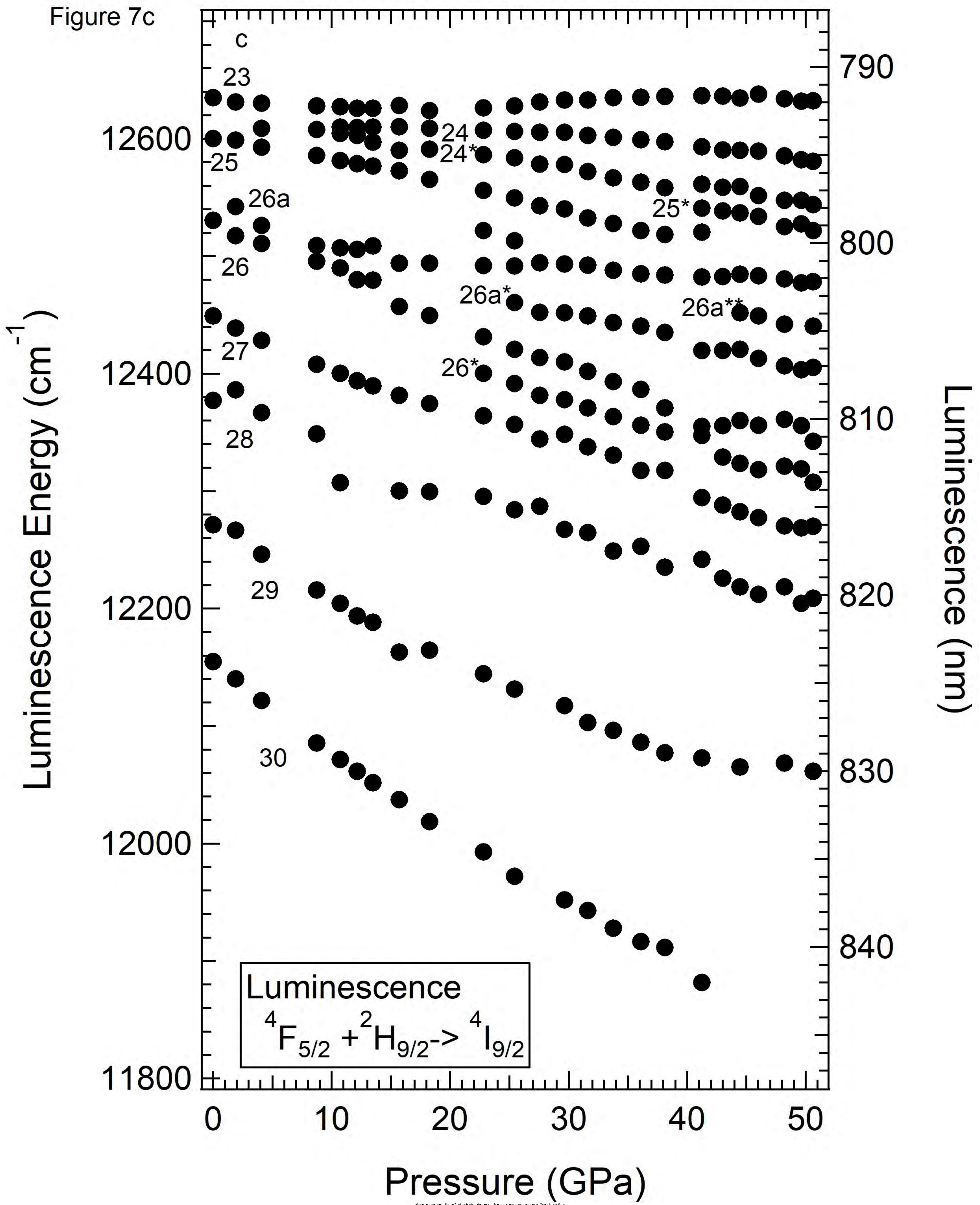


Figure 8

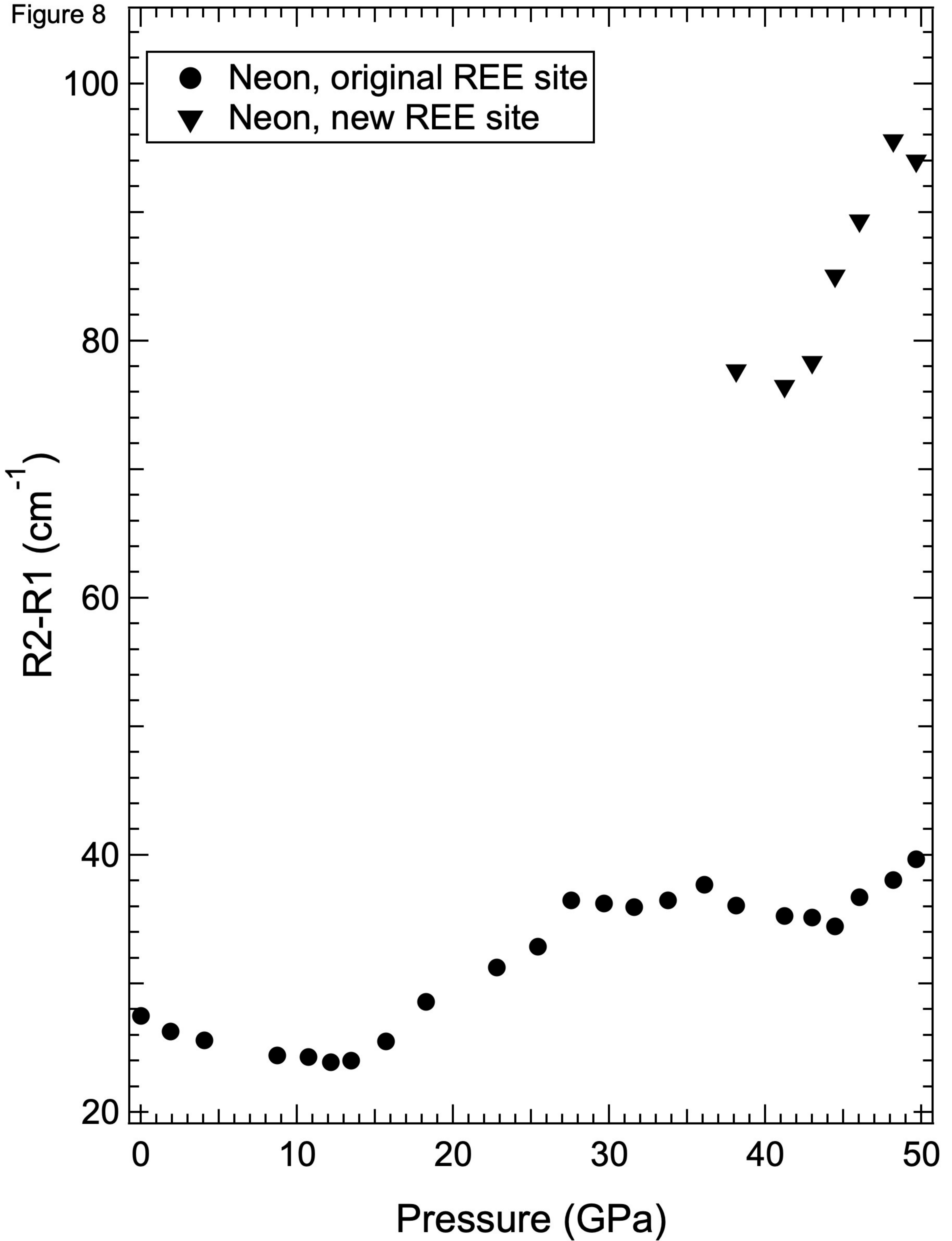


Figure 9

