

Revision 1

Precipitates of α -cristobalite and silicate glass in UHP clinopyroxene from a
Bohemian Massif eclogite

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21 **ABSTRACT**

22 Low pressure metastable nano-scale crystals of α -cristobalite have been observed
23 epitaxially-exsolved in cores of UHP clinopyroxene from the Bohemian Massif, Czech Republic.
24 SAED patterns and HRTEM images detail the close structural relationship between host
25 clinopyroxene and α -cristobalite precipitate: $[001]_{Di} \parallel [010]_{\alpha}, (010)_{Di} \sim \parallel (101)_{\alpha}$. TEM results
26 indicate that α -cristobalite exsolved from host clinopyroxene. Non-crystalline Al-bearing silicate
27 phases, also exsolved from UHP clinopyroxene, possesses Al/Si ratios close to eutectic
28 compositions in the system $\text{NaAlSi}_3\text{O}_8\text{-SiO}_2\text{-H}_2\text{O}$ system. Presence of glass exsolution suggests
29 a high temperature formation environment and presence of water. The α -cristobalite formed in a
30 localized low-pressure, micro-environment formed through exsolution of vacancies and excess
31 silica from the host pyroxene lattice. This micro-environment may be a result of negative
32 density changes due to excess lower density silica exsolving from higher density pyroxene
33 during an exsolution process that involved no localized volume change. Interface-controlled
34 exsolution via lattice matching at the diopside/cristobalite interface, and stability changes and
35 melting point depression due to nano-scale size effects contributed to formation and persistence
36 of this metastable phase. Amphibole in association with α -cristobalite and some non-crystalline
37 silicate phases may be a clue to localized water quantities; silica exsolution with amphibole may
38 have formed below the eutectic temperature and at a later stage than non-crystalline silicate
39 phases without amphibole. Silica rods in Nové Dvory clinopyroxenes were previously thought
40 to be quartz; however, our investigation reveals a variety of low pressure, high temperature,
41 and/or metastable phases greatly affected by the presence of vacancy and OH in clinopyroxenes.
42 The results will help us better understand OH in the UHP pyroxene and even water release in the
43 mantle.

44

45 INTRODUCTION

46 Discoveries of coesite (Smith 1984) and microdiamond in the Kokchetav Massif
47 (Sobolev and Shatsky 1990) have clearly established that rocks of continental origin were
48 subjected to ultrahigh-pressure metamorphism (UHPM), which required their subduction to a
49 depth of not less than 80-120 km (Chopin 1984; Smith 1984; Schreyer et al. 1987; Sobolev and
50 Shatsky 1990; Xu et al. 1992; Coleman and Wang 1995; Dobrzhinetskaya et al. 1995; Wain
51 1997; Ye et al. 2000; Van Roermund et al. 2002; Carswell et al. 2006; Dobrzhinetskaya et al.
52 2006). It is in these deeply-subducted terranes that micro- and nano-scale minerals within stable
53 host minerals such as garnet and pyroxene are often preserved, even though prograde textures
54 and mineral parageneses are typically obliterated during exhumation from great depth. Micro-
55 and nano-scale minerals of these orogenic terranes are of particular interest to investigators
56 because they are windows to the special conditions required to develop UHPM terranes when no
57 other evidence persists, and aids in elucidating the geodynamic processes endured by them.

58 Crystallographically-oriented silica rods exsolved in clinopyroxene, many coexisting
59 with amphibole, are well-documented from ultrahigh-pressure (UHP) terranes around the world:
60 (Bakun-Czubarow 1992; Bi, et al. 2018; Smith 1984; Gayk et al. 1995; Wain 1997; Katayama et
61 al. 2000; Schmadicke and Muller 2000; Tsai and Liou 2000; Dobrzhinetskaya et al. 2002; Zhu
62 and Ogasawara 2002; Klemd 2003; Song et al. 2003; Janak et al. 2004; Sajeev et al. 2010;
63 Dokukina and Konilov 2011; Konilov et al. 2011; Song et al. 2018). Although there is at present
64 evidence of HP (and not UHP) clinopyroxenes possessing these crystallographically-oriented
65 silica precipitates in the Blue Ridge Mountains, USA, the Kontum Massif, central Vietnam, and
66 the Greek Rhodope (Page et al. 2005; Anderson and Moechner 2007; Nakano et al. 2007a;
67 Nakano et al. 2007b; Proyer et al. 2009; Faryad and Fisera 2015; Li et al. 2018), in the dearth of

68 other UHP evidence the siliceous rods are often used as indicators of rocks experiencing UHP
69 conditions. The breakdown textures of clinopyroxene and its chemical variations may help us to
70 investigate pre- and post-peak rock evolution of the host rock.

71 Reports of silica exsolution in clinopyroxene have previously been identified by micro-
72 and macroscale techniques as either coesite or quartz (Bakun-Czubarow 1992). For instance,
73 Zhang et al. (2005) demonstrated that silica exsolution took place in the coesite stability field and
74 that the coesite easily transforms to quartz during retrograde metamorphism. More recently,
75 High-resolution transmission electron microscopy (HRTEM) investigation of
76 crystallographically-oriented silica in clinopyroxenes of the Kokchetav Massif discovered
77 anomalous low pressure silica polymorphs (the first confirmed natural occurrence of keatite) in
78 UHP pyroxene (Hill et al. 2013), but no coesite or quartz.

79 This study characterized crystallographically-oriented silica in clinopyroxenes from an
80 eclogite of well-established UHP origin from the Bohemian Massif to confirm mineral phase
81 identification of silica exsolution and to understand the processes of this occurrence more
82 thoroughly. As in the Hill et al. (2013) study, this HRTEM investigation has found neither of
83 the expected higher density silica polymorphs coesite or quartz. Instead, low-density and low-
84 pressure phases α -cristobalite and Na-Al silicate glass are observed exsolved from UHP
85 diopside. We explore possible mechanisms of formation for the metastable, low P , low T α -
86 cristobalite and Al-bearing silica glass.

87 The terms exsolution, precipitate, and inclusion are used interchangeably by researchers
88 to describe the phenomenon of crystallographically-oriented silica within pyroxene. We opt for
89 “exsolution” in this manuscript to imply that the siliceous rods are likely formed from
90 breakdown of the clinopyroxenes during exhumation. Further information about this is found in

91 the Discussion section. We have, at times still used the word precipitate as a descriptor when the
92 use of the term “exsolvent” is somewhat cumbersome.

93

94 **GEOLOGIC BACKGROUND AND SAMPLE DESCRIPTION**

95 Orogenic peridotite bodies are components of metamorphic terranes in most major
96 mountain systems. Some peridotite lenses contain (or contained) garnet-bearing assemblages,
97 particularly where they occur within eclogite-facies metamorphic crustal rocks that recrystallized
98 at high (HP) or ultrahigh (UHP) pressures. They are rich in petrological, mineralogical and
99 geochemical information about the interaction of the mantle with the crust during mountain
100 building (Brueckner and Medaris 1998, 2000; Rampone and Morten 2001;
101 Scambelluri et al. 2008; Spengler et al. 2009). This study focuses on clinopyroxenes of the Nové
102 Dvory eclogite from the Gföhl nappe (340 Ma (van Breemen et al. 1982)), a nappe within the
103 Bohemian Massif in the Czech Republic. Clinopyroxenes in this eclogite exhibit
104 crystallographically-oriented silica in their cores. It is also within the UHP range. Calculated *P-T*
105 estimates of the Gföhl n ppe mantle garnet peridotite (which hosts the thin layers of the Nové
106 Dvory eclogite) have been calculated at 875–1150°C, 3.3–6 GPa (Medaris et al. 2006). In
107 addition to calculated UHP pressures, microdiamonds have been found as inclusions in garnet,
108 kyanite, and zircon in the Eclogite-Gneiss Unit (Massonne 1998). Coesite has also been found in
109 the Eclogite-Gneiss Unit (Massonne 2001).

110 The Nové Dvory eclogite is fine- to medium-grained, massive, and generally
111 equigranular. It is composed primarily of Mg-Fe-garnet and Ca-rich clinopyroxene (between
112 omphacite and diopside) (Table 1). A transmitted light micrograph of the rock (eclogite) (Fig.
113 1a) demonstrates the predominance of garnet and pyroxene. Thin rims of Na-rich clinopyroxene

114 are observed in electron backscatter images surrounding pyroxenes (arrows, Fig. 1a). Observed
115 accessory minerals are rutile, apatite, zircon, iron sulfides, and iron oxides (possible retrograde
116 phases ilmenite and Mg-bearing ilmenite). Some later-stage retrograde reactions (pyroxene-
117 plagioclase symplectites and amphiboles) are observed at grain boundaries (Fig. 1a), otherwise,
118 the rock is relatively fresh displaying only minor fracturing and alteration.

119 A plane polarized light image with a clinopyroxene grain oriented at some angle to the *c*-
120 axis (Fig. 1b) displays siliceous phases concentrated in the core. In this orientation silica appears
121 as elongated dark euhedral lines. In contrast, when clinopyroxenes are oriented with the *c*-axis
122 parallel to the viewing direction, they appear as dark euhedral dots. This shape change with
123 differing orientations of pyroxene demonstrates that the precipitates are crystallographically-
124 oriented within the pyroxene lattice. A pyroxene grain boundary at the upper left in Fig. 1b
125 (black circle) demonstrates that silica rods are few near grain boundaries.

126

127 **EXPERIMENTAL METHODS**

128 High-resolution TEM (HRTEM) images and selected-area electron diffraction patterns
129 (SAED) were obtained utilizing an FEI Titan aberration-corrected S/TEM at the University of
130 Wisconsin-Madison at 200 kV accelerating voltage. In addition to clinopyroxene (CPX)
131 chemical compositions from the electron microprobe, compositions of silica precipitates were
132 also analyzed by means of X-ray energy-dispersive spectroscopy (EDS) with small spot sizes in
133 HRTEM mode using an EDAX EDS system on the FEI Titan S/TEM.

134 Major element compositions of bulk CPX crystals were determined via wavelength
135 dispersive spectrometry performed with the five spectrometer Cameca SX51 electron microprobe
136 at the University of Wisconsin-Madison. Analytical operating conditions were 15 kV

137 accelerating voltage, Faraday cup current of 20 nA, and 20 s counting times per element. Beam
138 spot size was 1 μm . Samples and standards were coated with ~ 20 nm carbon. Microprobe
139 analytical standards were: hematite for Fe, augite for Ca, olivine for Mg, rutile for Ti, chromite
140 for Cr, Mn-olivine for Mn, and hornblende for Si, Na, Al, and K.

141 Clinopyroxene lattice parameters were refined from X-ray diffraction (XRD) patterns by
142 means of least squares analysis using MDI, Inc. Jade 9.0 software. XRD patterns were collected
143 with a Rigaku Rapid II X-ray diffraction system with a 2-D image plate (Mo $K\alpha$ radiation), with
144 an accelerating voltage of 50kV, a current of 50 mA, and a 0.1mm diameter collimator.
145 Diffraction data were collected on an image-plate detector. Two-dimensional images were then
146 integrated to produce conventional 2-theta -intensity patterns using Rigaku's 2DP software.

147 Polished petrographic thin sections (30 μm thick) of the previously described CPX grains
148 were examined to find pyroxenes oriented parallel to the c -axis. Clinopyroxene grains were cut
149 out and thinned to ~ 20 μm by hand with silicon carbide grinding disks. Grains were then
150 mounted on molybdenum TEM grids. These foils were thinned to electron transparency via Ar^+
151 ion milling in a Fischione 1010 ion mill.

152

153 **RESULTS**

154 *Transmission electron microscope results*

155 A bright-field (BF) TEM image details a nano-crystal of α -cristobalite (Fig. 2).
156 Dominant pyroxene crystallographic planes observed in the $[001]_{\text{Di}}$ orientation are labeled $\{010\}$
157 and $\{110\}$. These planes govern α -cristobalite crystal morphology, which appears in almost all
158 cases as a euhedral 'negative' pyroxene shape. The SAED pattern (Fig. 2b) was calibrated with
159 clinopyroxene lattice parameters calculated from collected X-ray diffraction data from a single
160 crystal of clinopyroxene, then refined via least squares method. Those parameters are: $a = 9.630$

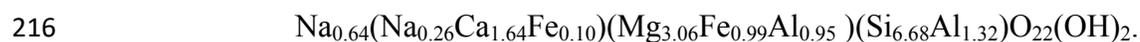
161 Å, $b = 8.818$ Å, $c = 5.244$ Å, $\alpha = \gamma = 90^\circ$, $\beta = 106.2^\circ$. The SAED pattern confirms host diopside,
162 [001], is epitaxially-oriented with α -cristobalite, [010]. Cristobalite hkl 's are denoted with white
163 numbers and white arrows; diopside is labeled with black numbers and arrows in a white
164 rectangle. Simulated diffraction patterns (DP) with indexing of both phases observed in the
165 collected SAED pattern. Only α -cristobalite and diopside are shown for clarification in Fig. 2d,
166 e, and f, respectively. Spots which appear in the α -cristobalite SAED pattern (001, 002, 003) are
167 labeled as X's in the simulated DP (Fig. 2e, f). These are typically at extinction due to screw
168 axis (4_1) but appear due to multiple diffraction effects.

169 Composition is straightforward. The silica precipitate EDS spectrum displays only silicon
170 and oxygen peaks (Fig. 2c). Both the TEM images and the SAED pattern reveal homogeneous
171 pyroxene structure with $C2/c$ symmetry. Ordered omphacite lamellae ($P2/n$ symmetry) known
172 to occur in clinopyroxene structures were not observed. In addition, nano-scale amphibole
173 occurs at the α -cristobalite interface; the crystals display excellent pyroxene/amphibole lattice
174 matching (Fig. 2g). In this HRTEM lattice image of all three phases, the cristobalite can be seen
175 in the midst of transforming to an amorphous phase as a result of damage from the electron
176 beam. This happened more quickly at the amphibole/pyroxene interfaces, where no cristobalite
177 lattice can be observed. In Fig. 2a, a planar boundary marked $(101)\alpha$ is observed in the
178 cristobalite and explained in Fig. 3, where black arrows denote the location of the boundary in
179 (a-c). Boxes denote areas of Fast Fourier Transform (FFT) patterns (Fig. 3c) on either side of the
180 boundary in the HRTEM lattice image (Fig. 3a) and demonstrate that the planar boundary is the
181 cristobalite twin plane (101). The amphibole and host diopside are labeled in (a) for reference.
182 A schematic representation of the α -cristobalite lattice with twinned unit cells outlined in black
183 (b) shows the near-perfect lattice fit on either side of the composition plane.

184 A second category of siliceous precipitates are non-crystalline Al-bearing silicate phases
185 enclosed by the euhedral negative pyroxene shape. They are quite common; a select four are
186 shown in Fig. 4. Dominant euhedral CPX planes are sharp and clear; however, SAED patterns
187 from precipitates (e.g., Fig. 4b) generally show diffuse rings indicative of amorphous phases. A
188 ring with weak intensity and some possible diffraction spots are attributed to crystallites of an
189 unidentified mineral that was quickly destroyed by the electron beam. There is no other
190 evidence of this mineral. The only chemical constituents observed via EDS (Fig. 4c) are Si, O,
191 and Al. Al/Si ratios in the exsolved silica in Fig. 4 (as shown by EDS) range from 0.1 to 0.2.
192 Preferential loss of alkali elements and low-atomic-number elements, such as Na and K, and
193 during analytical electron microscopy gives rise to characteristically low count rates for those
194 elements. In addition, with a focused probe for X-ray analyses such as used in this investigation,
195 the damage rate to the sample is exacerbated by a high current density and results in the loss of
196 these elements during X-ray collection (Ma et al. 1998). No amphiboles (or other H₂O-rich
197 phases) are observed in the non-crystalline silica in Fig. 4, unlike the cristobalite or the third
198 category of siliceous precipitates, which is non-crystalline, but also possesses amphibole (Fig. 5).

199 This group of siliceous phases (Fig. 5(a)) possess very similar features to the non-
200 crystalline siliceous phases in Fig. 4, with one main difference: the presence of amphibole at the
201 pyroxene/silica interface. Fig. 5 shows one of the siliceous, Al-bearing glass phases with
202 amphibole. The EDS spectrum (c) displays an Al peak as in the non-crystalline phase in Fig. 4.
203 Euhedral negative pyroxene morphology of the silica is still observed here, and, although the
204 SAED pattern (Fig. 5d) shows this image is oriented at $[T01]_{Di}$, at an angle to CPX c-axis, the
205 interface between diopside and the amphibole precipitate can still be seen to be coherent. Broad
206 diffuse rings are observed in the silica SAED pattern (Fig. 5b), establishing that the silicate is

207 amorphous. White arrows in this SAED pattern point to a distinct faint ring representative of
208 multiple small crystals. The extremely weak diffraction ring likely stems from the minerals
209 indicated by black arrows in the white precipitate center (Fig. 5a). Identity of these minerals is
210 unknown as they quickly disappeared under the energy of the electron beam. Inspection of the
211 CPX/amphibole interface (Fig. 5e) illustrates the excellent lattice matching between the single
212 and double chain silicates, and FFT patterns of both minerals further detail their crystallographic
213 relationship: $[T01]_{Di} \parallel [100]_{Amp}$. A chemical formula calculated from the EDS spectrum using
214 Cliff-Lorimer-based k -factors (Fig. 5f). This formula confirms a sodic-calcic amphibole (OH is
215 assumed, not measured by EDS):



217 It is also of note that multiple crystals of phyllosilicates were observed in close proximity
218 to some siliceous precipitates (Fig. 6a, b) in this investigation. In these precipitates, the shape of
219 the non-pyroxene crystals is modified from the typical euhedral ‘negative’ pyroxene shape, with
220 the sheet silicates fully covering the siliceous phase (Fig. 6a), but not the pyroxene. Dominant
221 pyroxene planes, specifically $\{010\}$ and $\{110\}$ are highly degraded and the familiar euhedral
222 negative pyroxene shape is unrecognizable. The composition (EDS spectrum taken from the
223 area in (b)) is that of a talc-like phyllosilicate with $\sim 9 \text{ \AA}$ (001) lattice fringes (Fig. 6c), and the
224 SAED pattern (e) confirms pyroxene plus diffraction rings from multiple talc-like crystals.

225

226 *Electron microprobe results*

227 Table 1 presents recalculated electron microprobe (EMPA) data from a representative
228 clinopyroxene grain in the Bohemian Massif eclogite. All clinopyroxenes analyzed via EMPA
229 exhibit silica precipitates in their cores. Standard deviation and standard error of these data are

230 reported (SDEV= Standard deviation, SERR=Standard error, and n=17). Oxygen was
231 calculated, not measured, and all calculations are normalized to six O atoms.

232 Almost all CPX analysis locations show sufficient Al to fill ^{IV}Al, the rest fills ^{VI}Al. The
233 clinopyroxene is highly calcic (~Di₆₈, between diopside and omphacite), with very little to no K.
234 UHP (non-stoichiometric) clinopyroxenes are characterized by significant cation deficiencies,
235 with cations equaling <4 total cations/6 oxygen; this has been well-established in the literature
236 (Smyth 1980; McCormick 1986; Proyer et al. 2004; Day and Mulcahy 2007; Konzett et al.
237 2008b; Zhao et al. 2011). Here, we broadly constrain possible vacancy content in clinopyroxene
238 grains by assuming cation deficiencies exist in the structure as vacancies (cations + vacancies =
239 four). Iron was measured via EMPA, and structural formula calculations involving only Fe²⁺
240 resulted in cations sum to ~ 4 or even slightly larger than 4, permitting no possible calculation of
241 potential vacancy content. In natural UHP clinopyroxenes, it is highly unlikely that all iron is Fe²⁺,
242 and some usefulness is found in estimating some amount of Fe³⁺ (Deer et al 1966) to calculate
243 formulae. In general, Fe³⁺ is not measured, it is estimated by charge balance; in fact, these
244 authors found no instances of measured Fe³⁺ data for the Nové Dvory eclogitic clinopyroxenes.
245 A comparable UHP clinopyroxene bulk composition was found in another UHP regime—the
246 Norwegian UHP clinopyroxenes of the Hareitland eclogite possess extremely similar bulk
247 clinopyroxene compositions. Calculated Fe³⁺ data from Mysen (1972) show 1.1-1.7 wt. % Fe³⁺
248 in the Hareitland eclogitic clinopyroxenes. Utilizing this as a proxy and making the assumption
249 that 2% of the total iron was Fe³⁺ in the UHP Nové Dvory clinopyroxenes, we tabulate both
250 measured Fe²⁺ and assumed 2% Fe³⁺ amounts in Table 1. The host pyroxene is close to
251 stoichiometric.

252

253 **DISCUSSION**

254 The occurrence of α -cristobalite and Na-Al-bearing silicate glass in UHP pyroxene raises
255 several questions. First, did the cristobalite form as β -cristobalite (as commonly happens) and
256 transform to α -cristobalite? Secondly, did the cristobalite form within its own range of
257 metastability (~ 200 - 275 °C (Hatch and Ghose 1991)) or at UHP pyroxene stability? Thirdly, we
258 interpret both the silica glass and α -cristobalite, which exhibit similar morphology, as primary
259 exsolution phases—what mechanism may have formed these low-density phases, and is the same
260 mechanism responsible for both?

261 Crystallographically-oriented silica in clinopyroxene from high- and ultrahigh-pressure
262 pyroxenes, such as that investigated here, is widely interpreted to be the result of a complex
263 combination of rock bulk composition, P , and T with a generally-accepted exsolution mechanism
264 of excess silica from host clinopyroxene during decompression via the reaction:



266 ($\text{Ca}_{0.5}\square_{0.5}\text{AlSi}_2\text{O}_6 = \text{CaEs}$ (CaEs) vacancy solid solution in pyroxene, and $\square = \text{vacancy}$)
267 (Smyth 1980; Liou et al. 1998; Katayama et al. 2000; Dobrzhinetskaya et al. 2002; Klemd 2003;
268 Zhang et al. 2005; Konzett et al. 2008a; Zhao et al. 2011). Substitution of CaEs component into
269 pyroxenes results in non-stoichiometry, and pyroxene may then dissolve excess silica charge
270 balanced by vacancies in M sites (Wood and Henderson 1978). Day and Mulcahy (2007)
271 document clear evidence that vacancy consumption in non-stoichiometric clinopyroxene can
272 release free silica. Vacancies are not the only way to result in pure silica from pyroxenes. Smith
273 (2006) demonstrates that it is also theoretically possible to exsolve pure SiO_2 from non-
274 stoichiometric pyroxene other than by the CaEs formula mechanism and leave behind

275 stoichiometric compositions. However, the non-stoichiometric character of HP/UHP
276 clinopyroxene is well-established (Smyth 1980; McCormick 1986; Proyer et al. 2004; Day and
277 Mulcahy 2007; Konzett et al. 2008b; Zhao et al. 2011), and we can conclude with some
278 confidence that vacancies were most likely present in the UHP pyroxenes of the Bohemian
279 Massif eclogite prior to decompression. These mechanisms produce silica irrespective of silica
280 polymorph, and exsolution may not produce and maintain α -cristobalite and Na-Al-silicate glass
281 phases to the exclusion of other silica polymorphs. Exsolution of excess silica from non-
282 stoichiometric pyroxenes is merely the beginning of the answer, and formation of these unusual
283 phases is likely an intricate combination of many processes.

284 In addition to excess silica, evidence of hydroxyl (water) observed in this investigation
285 may play a part in the vacancy content and the particular polymorph exsolved in the CPX
286 structure. Experimental studies have documented the incorporation of varying amounts of
287 hydroxyl in synthetic clinopyroxene (Skogby and Rossman 1989; Skogby 1994), as well as
288 natural clinopyroxenes (Skogby et al. 1990; Smyth et al. 1991; Bell et al. 1995; Koch-Muller et
289 al. 2004; Katayama et al. 2006). It is reasonable to assume that OH was present in Nové Dvory
290 clinopyroxenes. Observed micro- and nano-scale amphiboles are evidence of at least a sufficient
291 amount of OH to transform clinopyroxene to amphibole, and to produce a siliceous melt during
292 expulsion of excess silica and OH. Smyth et al. (1991) found a positive correlation between
293 hydroxyl content and cation vacancy concentration in pyroxene from mantle-derived eclogite in
294 the Roberts Victor kimberlite pipes. Katayama and Nakashima (2003) reported water contents as
295 high as 3000 ppm in omphacite within a diamond-grade eclogite from the Kokchetav massif,
296 Kazakhstan, although they assert this could be overestimated as much as 30%. Much lower OH

297 quantities are also known—31-60 ppm in diamond-bearing eclogite xenoliths of the Mir
298 kimberlite pipe (Koch-Müller et al. 2004).

299

300 *Cristobalite and silicate glass*

301 High- (or β -) cristobalite is the stable low pressure polymorph of SiO_2 between
302 1470°C and its anhydrous melting point at $\sim 1728^\circ\text{C}$, but exists in a metastable form down to
303 $\sim 275^\circ\text{C}$. There, it undergoes a rapid and reversible displacive transition to the metastable phase
304 α - (low-) cristobalite (Hatch and Ghose 1991). Beta-cristobalite is cubic (denoted below as
305 subscript 'c') ($Fd\bar{3}m$); α -cristobalite is tetragonal ($P4_12_12$).

306 Metastable silica phases are not uncommon in nature, and cristobalite can be found in a
307 variety of volcanic and metamorphic environments. Low- (or, α -) cristobalite is often found in
308 volcanic ejecta, lunar rocks and meteorites due to quenching (Christie et al. 1971), and
309 metastable cristobalite can also crystallize authigenically within sedimentary deposits as opal-CT
310 (Jones and Segnit 1971). Planar boundaries and microstructures within cristobalite are physical
311 evidence of the $\beta \rightarrow \alpha$ transition. Microstructures in α -cristobalite following the $\beta \rightarrow \alpha$ transition
312 are well documented by Hatch and Ghose (1991), and via electron microscopy (Lally et al.
313 1978). None of these $\beta \rightarrow \alpha$ transformation microstructures described above are observed in Fig.
314 2, indicating that this silica polymorph may have formed as α -cristobalite, not the expected
315 higher density polymorphs coesite or quartz previous researchers have reported in UHP
316 clinopyroxenes. Although none of the above transformation microstructures were observed in
317 the Nové Dvory clinopyroxene, a single planar boundary along (101) in the α -cristobalite crystal
318 is observed in Fig. 2. Planar boundaries along (101)_t have been observed in previous studies
319 (Withers et al. 1989), but the exact nature of these boundaries and their relationship with β -

320 cristobalite is still unclear. Diopside surfaces likely served as sites for heterogeneous nucleation
321 and epitaxial growth of cristobalite. Although there is evidence for the (101) boundary to be
322 interpreted as a growth twin, it is possible this may be an inherited twin (not a transformation
323 twin) from the β -phase (Lemmens et al. 2000). Thus, we cannot rule out the possibility that the
324 cristobalite transformed from the high temperature β - to α - cristobalite. It may be that the crystal
325 is simply too small to display all of the transformation microstructures typical of this
326 transformation.

327 Silicate glass can be found in many environments. Nanometer-size P/K-rich silica glass
328 inclusions are found in microdiamond from gneisses of the Kokchetav and Erzgebirge massifs
329 (Hwang et al. 2006). These are interpreted as former melt inclusions that reflect the chemical
330 character of their host rocks and may have been generated through partial melting of gneissic
331 rocks and/or dehydration/decarbonation processes in marble (Hwang et al. 2006). Pertinent to
332 this investigation, highly silicic glass inclusions are found commonly present in Mg-rich
333 minerals (olivine and pyroxene) of xenolithic terrestrial upper mantle rocks and have many
334 similarities with glass inclusions from mafic minerals of achondritic meteorites (Varela et al.
335 1998).

336

337 *Mechanisms of formation—silicate glass*

338 As shown above, the non-stoichiometric character of UHP clinopyroxene has been well-
339 established. If exsolution of excess silica is intimately linked to the non-stoichiometry of UHP
340 clinopyroxene, it is possible that both non-crystalline Al-bearing silicate phases and α -
341 cristobalite nucleated and exsolved from Ca-Es-bearing clinopyroxene. Simultaneous expulsion
342 of OH held within the clinopyroxene structure at high temperatures most likely caused the partial

343 melting of silica that created Na-Al-silicate glass phases, and growth of amphiboles associated
344 with α -cristobalite. It is at this point in the story that formation mechanisms of α -cristobalite and
345 silica glass diverge slightly.

346 We propose that the glass precipitates in this study (Figs. 4, 5) are likely formed through
347 partial melting of the silica and albite components, as silica, a jadeite component, and OH are
348 expelled from the pyroxene lattice. Although silica-rich, aluminous alkali-rich glasses have been
349 found in Mg-rich minerals of both terrestrial upper mantle rocks and achondrites, mechanisms of
350 formation in the two cases are different. Varela et al. (1997) observed that although many
351 processes were likely to have contributed to their formation— low-degree partial melting,
352 precipitation from silicate-loaded fluids, fractional crystallization and liquid immiscibility, or
353 residual glass from a parent magma—glass inclusions in olivine and pyroxene of upper mantle
354 peridotites were more likely related to a very low degree of partial melting. Precipitation from
355 silicate-loaded fluids can be ruled out in the formation of glass in the Nové Dvory eclogite and
356 garnet pyroxenite, as it is more often associated with fractures and cross-cutting grain boundaries
357 that suggest migration of fluids through the grain, not isolated precipitation in cores of grains.
358 And, as can be seen in Fig. 1, the cores of clinopyroxenes in the rock are quite fresh, without
359 obvious subsequent alteration. Fractional crystallization and liquid immiscibility may be ruled
360 out on the basis that this phenomenon is typically extremely FeO-rich and, with these
361 compositions, is generally found in lunar environments. Table 1 compositions show these
362 pyroxenes possess some FeO, but may not be enriched enough for that. Residual glass from a
363 parent magma can also be ruled out. Varela et al. (1997) state that silica-rich glass inclusions in
364 upper mantle xenoliths are incompatible with this genesis, as residual melts that form glass likely
365 represent independent components trapped during crystallization of primary magma components.

366 Although Na is not observed now as it disappears quickly under the electron beam and can be
367 difficult to capture, the potential presence of Na allows for us to consider the Al-bearing glass as
368 a mixture of albite and silica components in a silica-albite eutectic system. The calculated range
369 of Al/Si ratios (0.1-0.2) from the glasses is close to the eutectic composition in the NaAlSi₃O₈-
370 SiO₂-H₂O (Ab + Qz + H₂O) system—at 0.1 GPa and 800 °C (Wen and Nekvasil 1994). The
371 eutectic temperature at low pressures is similar for hydrous versus anhydrous conditions—at 1
372 bar, the temperatures of the eutectics are shown to be 1007 °C and 1009 °C, respectively. In a
373 H₂O-saturated system, which is not the case here, an increase in pressure causes the temperature
374 of the eutectic to drop—at 7 kbar, the eutectic is 682 °C. In an anhydrous system, a pressure
375 increase to 7 kbar causes the eutectic temperature to increase dramatically to 1132 °C, and at 15
376 kbar (1.5 GPa), the eutectic moves to 1228 °C. Calculated *P-T* estimates of the UHP mantle
377 garnet peridotite which hosts the thin layers of the Nové Dvory eclogite are 875–1150 °C, 33–60
378 kbar (Medaris et al. 2006). At our calculated Al/Si ratio range of 0.1-0.2, the temperature of the
379 eutectic in an anhydrous system which is similar to a hydrous system, as mentioned above, (at 5
380 kbar) is ~1100°C, as shown by Wen and Nekvasil (1994). This temperature is possible for glass
381 formation. Figure 8 is a diagram illustrating the pressure difference between the precipitates
382 (curve c) and the host minerals (curve b). Melt inclusions will be over-pressured (curve a) with
383 respect the host mineral and precipitates will be under-pressured with respect to the host mineral.

384 It is unlikely that the glass-forming components (jadeite component, silica, vacancy, and
385 OH) were expelled from the clinopyroxene lattice at peak metamorphic *P-T*. Rather, they may
386 have been expelled at varying high temperatures post-metamorphic peak during decompression,
387 when these elements formed a melt at the eutectic. In fact, modeling of thermal boundary
388 conditions of exhumation rates in UHP rocks has shown that in the absence of lateral heating,

389 rocks exhuming from great depth always cool or decompress isothermally, and the main cooling
390 period at shallower crustal levels follows the main period of exhumation (Grasemann et al.
391 1998). We may use the modeling results of Grasemann et al (1998) to propose that a glass
392 formed at the eutectic in the Ab + Qz + (low) H₂O system with an Al/Si ratio of 0.1-0.2 may
393 have formed at many different depths along the exhumation pathway. To narrow the
394 temperature range in which the glass formed, we must also consider contributing factors that
395 likely lowered the temperature at which the melt can form. See the section entitled
396 “*Contributing factors that aid in α -cristobalite and silica glass formation*” for a discussion of
397 these factors.

398

399 ***Mechanisms of formation— α -cristobalite***

400 The key to α -cristobalite formation may be the stability of the CPX lattice during subduction
401 and exhumation. Metastable cristobalite has previously been found in high pressure garnet
402 granulite from Gore Mountain, New York. This cristobalite within garnet likely originated as
403 small trapped samples of hydrous Na-Al-siliceous melt where diffusive loss of water from the
404 inclusions under isothermal, constant volume conditions may have resulted in a large enough
405 internal pressure decrease to promote the metastable crystallization of cristobalite (Darling et al.
406 1997). In the Nové Dvory eclogite as in the Gore Mountain garnet, a crystal may form under
407 constant volume conditions due to the high tensile strength and low thermal
408 expansivity/compressibility of diopside compared with α -cristobalite (Berman 1988). No
409 evidence of volume change is observed in the Nové Dvory clinopyroxenes in optical (Fig. 1) or
410 HRTEM (Fig. 2) micrographs as a result of silica polymorphic transformations or changes in
411 density from host clinopyroxene (~ 3.3 g/cm³) to cristobalite (2.32 g/cm³). Darling et al. (1997)

412 attribute formation of low pressure cristobalite to an open system in which mass is removed from
413 the system, causing lowered pressure. We hypothesize a locally-closed system at the point of
414 silica exsolution in the case of the α -cristobalite crystal in Fig. 2. Exsolution of vacancies
415 (assumed density of $\sim 0.0 \text{ g/cm}^3$) that coalesce may cause a void and aid in this localized density
416 decrease. The density of the surrounding pyroxene ($\sim 3.3 \text{ g/cm}^3$) minus precipitated α -cristobalite
417 (2.32 g/cm^3) plus a potential coalesced vacancy void (0.0 g/cm^3) is a negative change of ~ 1
418 g/cm^3 . An arrow in Fig. 2 points to a potential void at the interface of silica and pyroxene. We
419 interpret this as an area of expelled, coalesced vacancies. If we interpret Fig. 2 in such a way as
420 to assume overall volume is preserved in clinopyroxene during silica exsolution, and assume
421 simultaneous vacancy expulsion and coalescence, a low pressure micro- or nano-environment
422 may be generated at any depth in response to density changes, perhaps via a reaction similar to
423 this:



425 Vacancies are point defects in the crystal structure that create non-stoichiometry and are
426 also notoriously difficult to image with the TEM. We can assume the unstable vacancies have
427 been expelled from the destabilized pyroxene lattice and cannot be imaged in situ. Fortunately,
428 cases exist in which point defects form clusters and have been imaged readily with TEM
429 methods (Veblen and Cowley 1994). We believe this may be the mechanism which aids in
430 forming a low pressure micro-environment. As in the case of the Gore Mountain garnet, this
431 interpretation requires that the host mineral (clinopyroxene) maintain a pressure differential
432 between the small inclusions and the outside of the clinopyroxene. This pressure differential, in
433 addition to the high tensile strength and low thermal expansivity/compressibility of the
434 clinopyroxene (Berman 1988) likely serves to preserve micro- and nano-scale stable and

435 metastable minerals within the clinopyroxene during exhumation. Results of recent synthetic
436 fluid inclusion studies on quartz demonstrate that high pressure differences can be maintained
437 over small distances at high temperatures in cases where the internal inclusion pressure is lower
438 than external confining pressure (Sternner and Bodnar 1989; Vityk et al. 1995). In this way,
439 metastable phases such as α -cristobalite and the siliceous non-crystalline phases may persevere
440 to the surface from any depth.

441 This mechanism has been used by these authors (Hill et al, 2013) to explain another
442 instance of silica exsolution from UHP diopside in the Kokchetav Massif. In this diopside, low-
443 pressure keatite with accompanying amphibole was identified, also with no evidence of volume
444 change within the clinopyroxene crystal. Formation of α -cristobalite is favored over keatite in
445 the Nové Dvory eclogite as a result of higher temperatures than Kokchetav Massif garnet
446 pyroxenite.

447

448 *Contributing factors that aid in α -cristobalite and silicate glass formation*

449 Although it appears likely that both α -cristobalite and silica glass initiated formation in a
450 similar manner, the crystalline nature of α -cristobalite requires a different formation environment
451 than glass. Already discussed is the low-pressure micro-environment that may have been
452 produced by vacancy expulsion/coalescence and subsequent density decrease under constant
453 volume conditions, thus creating a suitable P and T for α -cristobalite. A number of other factors
454 may have contributed to the formation of metastable α -cristobalite.

455 Favorable interface energy due to excellent lattice matching and nano-scale size effects
456 may have played integral parts in lowering the nucleation and crystallization energy.

457 Generation/exsolution of metastable phases takes place when the kinetically-preferred path

458 differs from the preferred thermodynamic path (Baricco et al. 2004). In this way, exsolution of a
459 metastable phase may be preferred over a more stable phase. Mineral nano-particles commonly
460 behave differently as a function of their size within the nano-scale range. Mineral nano-particles
461 also behave differently than larger micro- and macroscopic crystals of the same structure and
462 composition. At the nano-scale, three factors compete to stabilize a given polymorph: enthalpy
463 of polymorphic transition, surface enthalpy, and enthalpy of hydration (Barnard and Xu 2008;
464 Hochella et al. 2008). Navrotsky et al. (2004) reported that competition between surface
465 enthalpy and the energetics of phase transformation leads to the general conclusion that
466 polymorphs that are metastable as micron-sized or larger crystals can often be
467 thermodynamically stabilized at the nano-scale. Nanocrystals of protoenetsite within plagioclase
468 are quenchable and stable at low-temperature (Xu et al. 2017a). This phenomenon has been
469 observed in TiO₂ polymorphs, aluminum and iron oxides, and oxyhydroxides (Navrotsky 2003;
470 Navrotsky et al. 2004; Chen and Fu 2006; Barnard and Xu 2008; Navrotsky et al. 2009; Lee and
471 Xu 2016; Lee et al. 2016). Thus, we infer metastable nano-scale α -cristobalite may be stabilized
472 in part due to its small size.

473 Additionally, the size at which a nano-scale mineral forms/melts is temperature
474 dependent. Extensive research has shown that decreasing particle size also decreases the
475 temperature at which the particle forms (Barnard and Xu 2008; Barnard et al. 2009). This
476 melting/crystallization point depression occurs for nearly all free nanoparticles, e.g. Au, TiO₂, Pt
477 (Barnard et al. 2009; Barnard et al. 2011), and for some nano-particles embedded in a matrix.
478 Melting point depression is also observed as a function of dimension (Zhang et al. 2000).

479 The melting point depression can be dramatic. As the particle shrinks to the nanoscale,
480 the interfacial energy (γ) penalty becomes significant when compared to total bulk energy (ΔH).

481 This, in turn, destabilizes the particle with respect to the melt (DeHoff 2006). Assuming an $l \times l$
482 $\times 3l$ rectangle particle shape as observed in TEM (Fig. 4), this destabilizing effect for a pure
483 silica glass results in nearly a 300°C melting temperature depression for an $l = 50$ nm particle as
484 the melting temperature drops from 1723°C to 1430°C. The same calculations for $l = 100$ and
485 200 nm yield a 140°C and 70°C depression, respectively:

$$486 \quad \lambda_m = \frac{3\gamma M}{\rho_s \Delta H} \quad (1)$$

$$487 \quad T_m(\ell) = T_m(\infty) \left[1 - \lambda_m \frac{1}{\ell} \right] \quad (2)$$

488 where, l (length), γ (interfacial energy) = 4.4 J/m², M (molar mass) = 60.08 g/mol, ρ_s (density) = 2.203E6
489 g/m³, $T_m(\infty) = 1996$ K (1723°C), ΔH (bulk enthalpy of fusion) = 50.55 kJ/mol

490

491 In combination with nano-scale size effects, mechanisms of exsolution are closely related
492 to the types of interfaces that develop between two phases. The geometry of the substrate in
493 epitaxial mineral growth is known to cause growth of phases not stable at existing P - T - X
494 conditions. The morphology of silica glass and α -cristobalite is certainly dominated by the
495 clinopyroxene lattice (Figs. 2, 4, 5), and a high degree of lattice match between the two likely
496 serves to stabilize the nano-scale α -cristobalite crystal. The structural relationship of α -
497 cristobalite with diopside derived from the SAED pattern (Fig. 2a) is schematically-represented
498 in Fig. 7. In all models $[001]_{\text{CPX}} \parallel [010]_{\alpha\text{-cristobalite}}$. Here we show $(010)_{\text{CPX}} \sim \parallel (-101)\alpha$ (Fig. 7a)
499 and $(\bar{1}\bar{1}0)_{\text{CPX}} \sim \parallel (100)\alpha$ (Fig. 7b). These models illustrate that diopside and silica lattices join
500 with little to no bending of the silica tetrahedra. Diopside (010) is parallel to α -cristobalite ($\bar{1}01$),
501 and Fig. 7d illustrates the close lattice match with unit cell dimensions: d_{100} of diopside is 9.3 Å,
502 while α -cristobalite d_{101} 4.04 Å. This is a mere 5% difference over just two unit cells of diopside

503 and three unit cells of α -cristobalite, where $5d_{101\alpha}$ ($5 \times 4.04 \text{ \AA}$) is approximately equal to $2d_{100\text{CPX}}$
504 ($2 \times 9.3 \text{ \AA}$). In addition, very minor mismatch between α -cristobalite and diopside likely
505 minimizes the nucleation energy which aids a favorable configuration of silica tetrahedra.

506 The presence of amphibole in association with α -cristobalite and some glass precipitates
507 may be a clue to localized water quantities. Silica glass exsolution with amphibole may have
508 formed below the eutectic temperature and at a later stage than glass precipitates without
509 amphibole. Nano-crystals of α -cristobalite plus amphibole likely possessed less localized water
510 than the glass precipitates and may have formed after glass precipitates with and without
511 amphibole. The amphibole in Fig. 3 appears to have been encompassed by the exsolving silica
512 rather than transforming from pyroxene along the edge. A possible explanation is that the
513 diopside expelled only a small amount of water at this locality and only a nano-scale portion of
514 the pyroxene rehydrated to amphibole.

515 In this paper, we propose that the unique properties of HP/UHP pyroxenes discussed
516 above (excess silica balanced by structural vacancies, OH held within the lattice and expelled
517 during precipitate formation) drive exsolution from the clinopyroxene structure and form silicate
518 minerals and melts. In addition, we interpret the silicate glass Al/Si ratios of 0.1-0.2 to correlate
519 well with Al-bearing silicate glasses forming from a siliceous melt at high temperature on the
520 eutectic in the Ab-Qz-H₂O system. a slightly later stage and lower temperature, the metastable
521 low density silica polymorph α -cristobalite nucleated on diopside planes that served as sites for
522 solid state heterogeneous nucleation and epitaxial growth. Assuming that the pyroxene (a stable
523 crystal) has not experienced volume change at the point of low density silica exsolution, a
524 metastable phase may form and persist within the host due to an internal pressure decrease.

525 Thus, a low pressure micro-environment may be created at the point of low density cristobalite
526 formation.

527

528 *Later stage alteration of silica precipitates*

529 Some silica precipitates in this investigation are observed to have altered to talc-like
530 phyllosilicates (Fig. 6). This appears similar to exsolved silica altered to high-silica sheet
531 silicates from UHP clinopyroxene of the Kokchetav Massif (Hill et al. 2013). Here, as in the
532 Kokchetav Massif, the dominant planes of the euhedral clinopyroxene have degraded and are
533 nearly unrecognizable (Fig. 6a). Multiple crystals of sheet silicates in random orientation show a
534 composition that is unexplainably high in silica (Fig. 6c), where no silica precipitate is observed.
535 Due to the similarities of the two cases, we continue to interpret the sheet silicates as a later stage
536 alteration of silica by replacement reaction due to fluid migration, similar to the results of
537 Dobrzhinetskaya et al. (2009), localized on silica during rock retrogression. Refer to Hill et al.
538 (2013) for a more detailed explanation.

539

540 **IMPLICATIONS**

541 Nano-scale investigation of formation processes of precipitates of silicate glass and
542 cristobalite / amphibole intergrowths in UHP pyroxene may be important to help us better
543 understand OH (or water) in eclogitic host pyroxene during subduction, and its ultimate release
544 in the mantle. As geologists have little access to minerals from UHP terranes to study, the
545 findings here have great significance for investigation of rocks that weren't commonly believed
546 to exist a mere 20-30 years ago. This study revealed vacancies and OH in high-pressure
547 pyroxene that formed during subduction. When the host rock is exhumed, the incorporated OH

548 migrates out from the UHP pyroxene structure together with vacancies and excess silica in
549 multiple forms of glass and cristobalite / amphibole precipitates. If the pyroxene-bearing slab
550 subducts further into the mantle, the pyroxene may release the incorporated OH into mantle
551 during its phase transformation to garnet and/or perovskite structures. The subsequent effects on
552 mantle rocks are worthy of investigation.

553 Characterization of nano-scale phases inside UHP rocks is also significant for
554 understanding the larger scale geodynamics of collisional terranes, particularly when unexpected
555 (metastable) phases make an appearance. Geologic processes occurring deep in the Earth and
556 under UHP conditions are difficult to recognize due to subsequent retrogressive reactions and
557 fluid and deformation events experienced by rocks during exhumation. Only by identifying
558 minerals and their mechanisms of exsolution at the nano-scale to track the path of these rocks
559 can we understand in detail the processes operating deep in the Earth. Silica precipitates in Nové
560 Dvory clinopyroxenes were previously thought to be quartz; however, our investigation reveals a
561 variety of low pressure, high temperature, and/or metastable phases greatly affected by the
562 presence of water in clinopyroxenes. The high tensile strength and low thermal expansivity /
563 compressibility of diopside serves to preserve phases typically unlikely to make it to the surface
564 without recrystallizing. When compared to observed silica phases, this may mean that the
565 exsolved phases have little to no effect on the rheology of the exhuming rock but could affect the
566 interpretation of the *P-T* path. Some low-density precipitates containing Na and K may be
567 formed through similar precipitation processes, instead of melt inclusions. *P-T* evolution for the
568 precipitates and inclusions are very different. The precipitates are under-pressured with respect
569 to the host mineral. However, the melt inclusions are over-pressured.

570

571

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- 873

FIGURE CAPTIONS

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876 Figure 1. Plane polarized light photomicrograph of Bohemian Massif eclogite (a) and
877 clinopyroxene (b). The rock (a) is dominated by garnets (black) and clinopyroxenes (most other
878 grains in the image). Arrows point to decompression garnet/CPX and CPX/CPX symplectites;
879 the symplectites are generally thin and the CPX grains are quite fresh, with little alteration. The
880 lack of alteration can be seen in the cores of the clinopyroxenes (b), which are filled with
881 siliceous precipitates. In this orientation viewed at an angle to the *c*-axis, the precipitates appear
882 primarily as rods and some dots (likely rods on end). Precipitates are concentrated in the grain
883 core, with few precipitates in the rims near grain boundaries. A black circle denotes one of the
884 CPX grain boundaries at the top left of (b), where there is a much lower concentration of
885 precipitates. Note that no optical evidence of volume change (i.e., cracking due to less dense
886 phase transformation) surrounds the precipitates.

887 Figure 2. (a) Bright-field TEM image of α -cristobalite nano-crystal. Dominant euohedral
888 diopside planes and α -cristobalite (101) α twin plane are labeled; nano-scale amphibole at
889 cristobalite/diopside interface indicated with white arrow at top left of precipitate. White arrow
890 at bottom of image highlights an area of great strain contrast, and may be a void. (b) SAED
891 pattern details epitaxial orientation of diopside (black arrows/numbers on white box) with α -
892 cristobalite (white arrows/numbers). (c) EDS spectrum from just the α -cristobalite nano-crystal.
893 Simulated diffraction patterns (DP) with indexing of (d) both phases observed in the SAED
894 pattern, only α -cristobalite (e), and only diopside (f). Diffraction spots from α -cristobalite spots
895 which appear in the SAED (001, 002, 003), labeled as X's in the simulated diffraction pattern,
896 are typically at extinction but appear due to multiple diffraction. (g) HRTEM image of nano-
897 scale amphibole at interface of diopside and α -cristobalite.

898 Figure 3. (a) HRTEM image showing cristobalite domains in a twin relationship. Twin plane
899 (101) corresponds to (111) in β -cristobalite. White squares are locations of Fast Fourier
900 Transform (FFT) patterns in (c), which detail the crystallography on either side of the (101) twin
901 plane. (b) Schematic representation of the α -cristobalite lattice with twinned unit cells in black;
902 black arrows in (a, b, c) denote the location of the twin plane.

903

904 Figure 4. (a) Bright field TEM image of non-crystalline siliceous precipitate exhibiting
905 “negative” pyroxene morphology. Dominant pyroxene planes evident in crystalline silica
906 precipitates are also evident here. (b) Diffraction pattern is from the glass precipitate in (a).
907 Ring with weak intensity, and possible diffraction spots are attributed to crystallites of an
908 unidentified mineral that was quickly destroyed by the electron beam; (c) EDS spectrum from
909 precipitate in (a). All non-crystalline precipitates possess Al in addition to Si, just as in this
910 spectrum. (d-f) More examples of Al-bearing non-crystalline silica precipitates with negative
911 pyroxene morphology.

912 Figure 5. (a) Non-crystalline siliceous precipitate in negative pyroxene shape surrounded by
913 diopside (dark area). (b) SAED pattern shows evidence of a faint ring (white arrows point to the
914 ring) likely contributed by the small crystals indicated by black arrows in the non-crystalline
915 precipitate. (c) EDS spectrum shows there is some Al in the precipitate. (d) Diopside SAED
916 pattern. (e) HRTEM image and (g, h) FFT patterns of diopside and Na-Ca-amphibole across the
917 interface, respectively, detailing crystallographic orientation and excellent lattice fit of the
918 single/double chains. (f) EDS spectrum from amphibole exhibits Na and Al.

919 Figure 6. (a) Low magnification TEM image and (b) HRTEM image of multiple sheet silicate
920 crystals in diopside. Note euhedral [001] pyroxene shape of former silica precipitate is degraded,
921 with few recognizable pyroxene planes evident. (c) EDS spectrum taken from the area in (b)
922 showing Mg, Fe, and Al peaks, in addition to a large silica peak. Carbon peak is from carbon
923 coating on sample. (d) SAED pattern from diopside. (e) SAED pattern from diopside and
924 multiple mica crystals (f) SAED pattern from talc-like nano-crystals showing (001) reflections.

925 Figure 7. (a-d) Schematic representation of the structural relationship between diopside and α -
926 cristobalite based upon SAED pattern. (a-c) are same orientation. (a, b) illustrate that diopside
927 and cristobalite lattices join at the interface with little to no bending or rotation of silica
928 tetrahedra. (a) $(010)_{Di} \parallel (-101)_{\alpha}$. (b) $(-1-10)_{Di} \parallel (100)_{\alpha}$ (c-d) close lattice fit at the interface in
929 (a, b) is shown utilizing unit cells of both crystals, where $(100)_{Di}$ is approximately $\parallel (101)_{\alpha}$. (c)
930 $5d_{101\alpha} = 20.2 \text{ \AA} \sim \parallel 2(a \cdot \sin\beta) d_{100\text{CPX}} = 18.6 \text{ \AA}$. (d) this diagram is the orientation in (c) rotated
931 90 degrees to look at CPX [010], and illustrates the structural match between $5d_{101\alpha} = 20.2 \text{ \AA} \sim$

932 $\parallel 2(a \cdot \sin\beta) d_{100 \text{ CPX}} = 18.6 \text{ \AA}$. The slight mismatch may cause the observed strain at the
933 interface between diopside and cristobalite (Fig. 2a).

934

935 Figure 8: A diagram illustrating the pressure difference between the precipitates (curve c) and the
936 host minerals (curve b). Melt inclusions will be over-pressured (curve a) with respect the host
937 mineral. Modified from Burnham and Davis (1974).

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TABLE CAPTION

942 Table 1. Recalculated EMPA clinopyroxene analyses from Bohemian Massif CPX grain,
943 calculated with an assumed (not measured) Fe content of 2 wt. % Fe_2O_3 ; $\text{FeO} = \text{Fe} - 2 \text{ wt } \%$
944 Fe_2O_3 . Measurements made with $1 \mu\text{m}$ beam. Oxygen calculated, all calculations based on 6 O.
945 Potential calculated vacancy component = ([]).

946

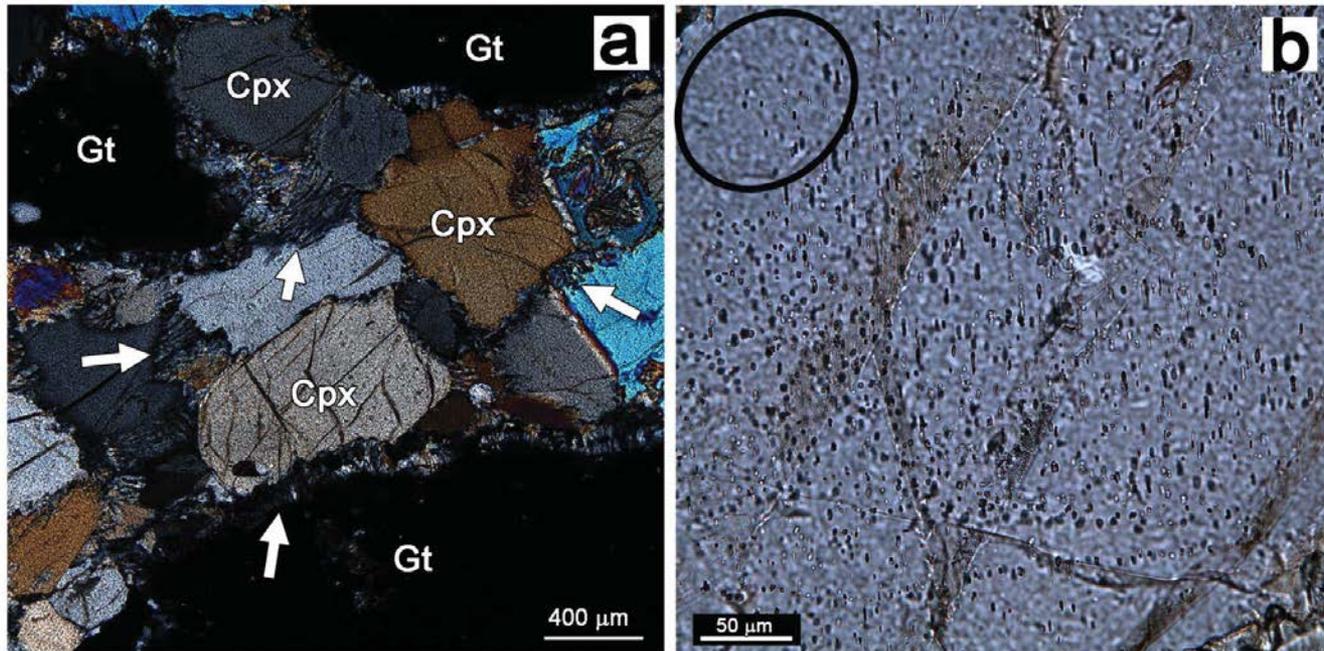


Fig. 1

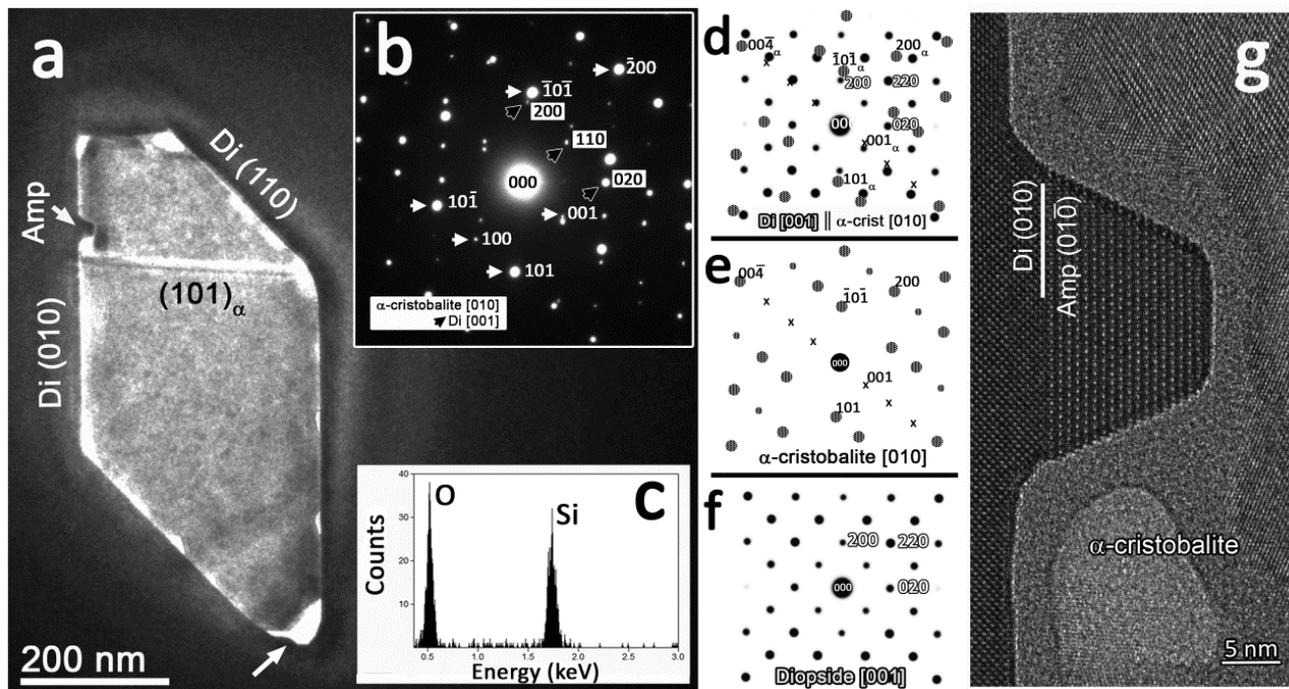


Fig. 2

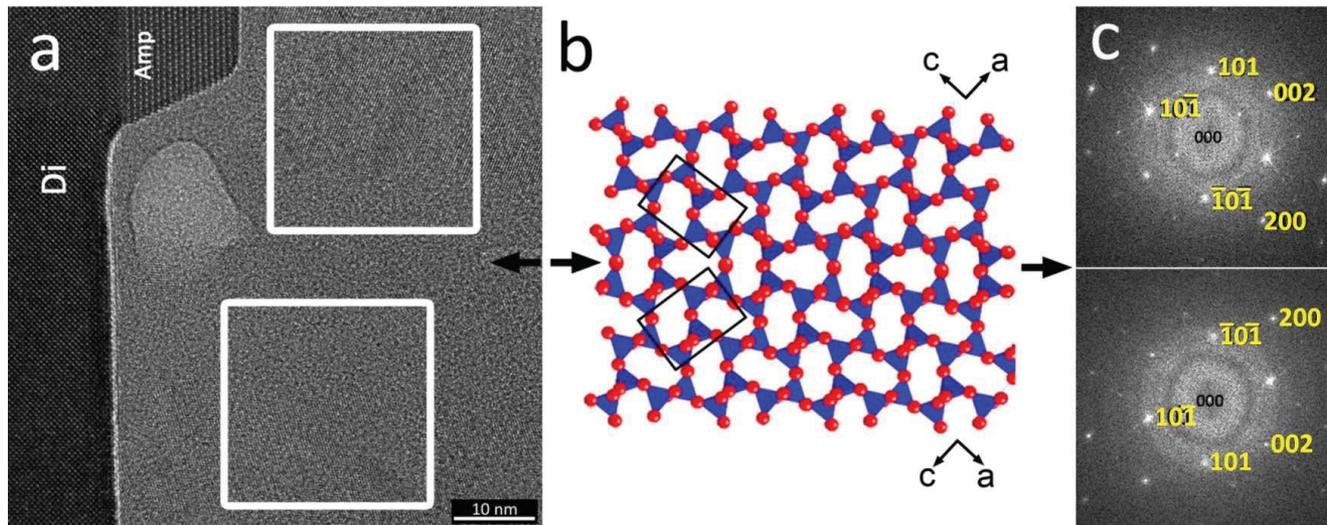


Fig. 3

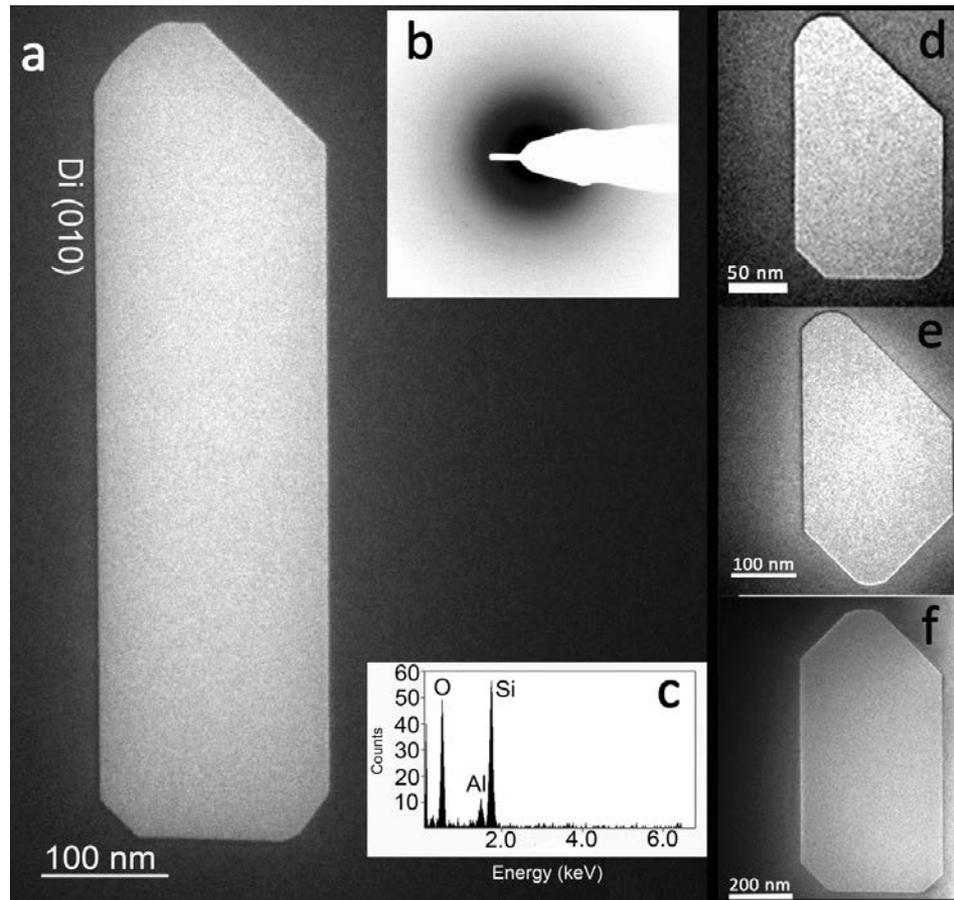


Fig. 4

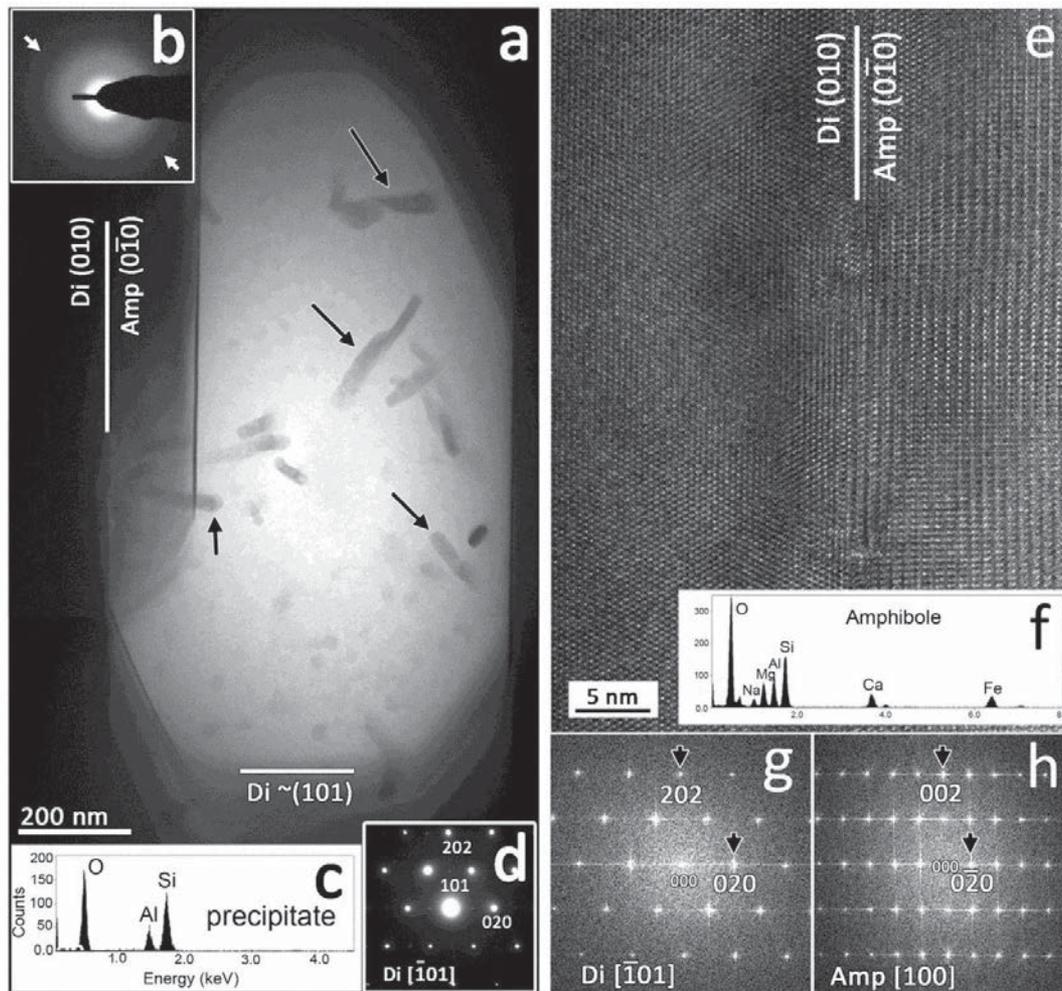


Fig. 5

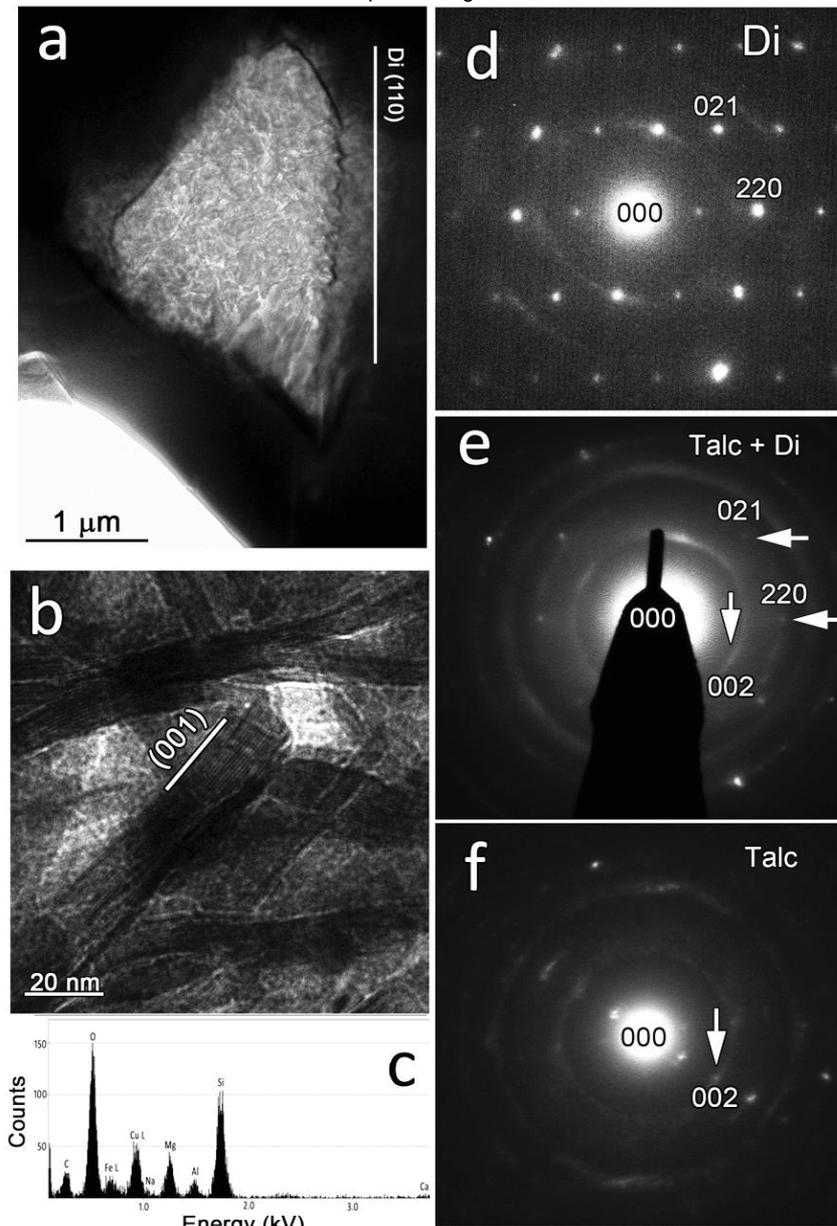


Fig. 6

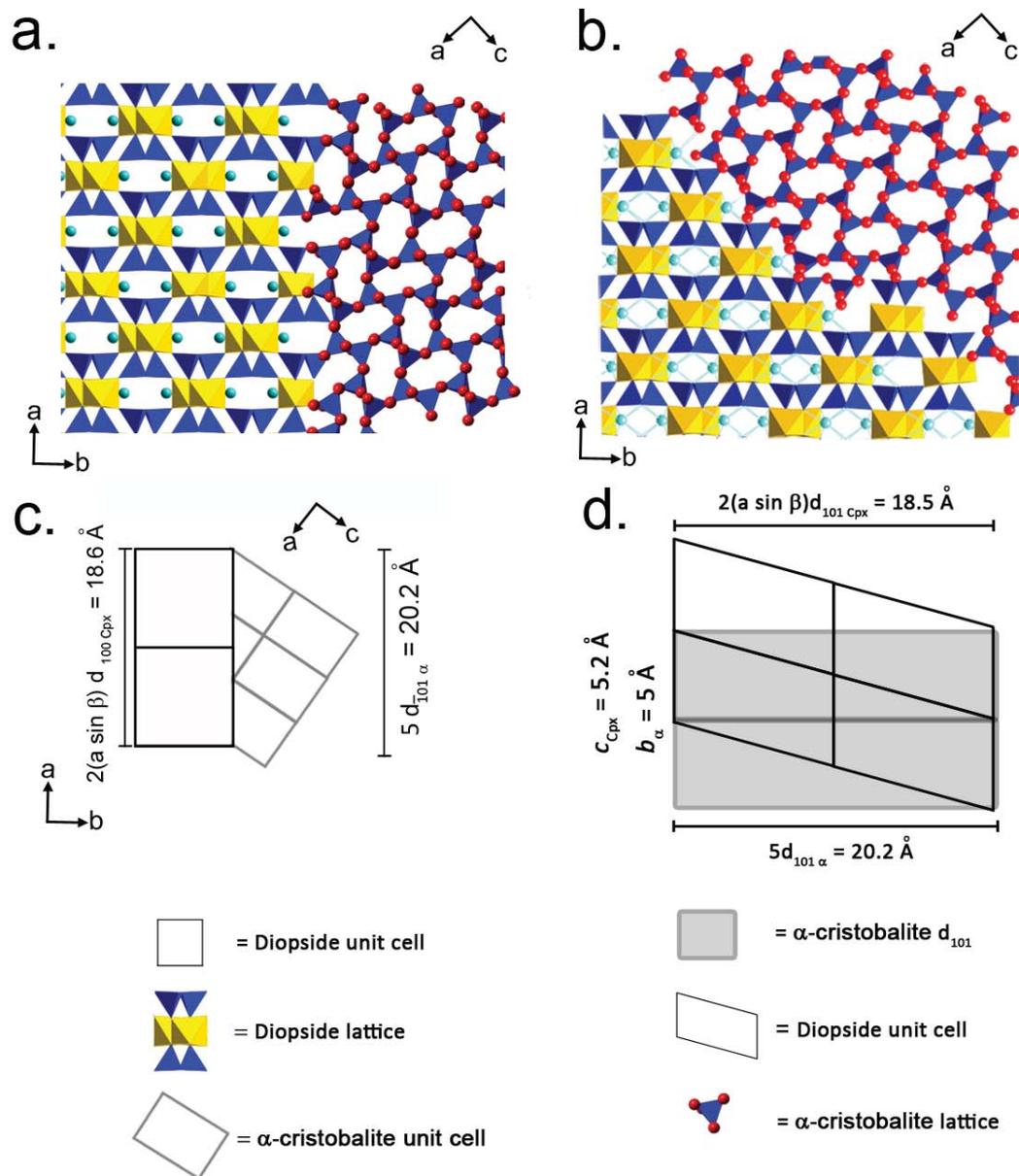


Fig. 7

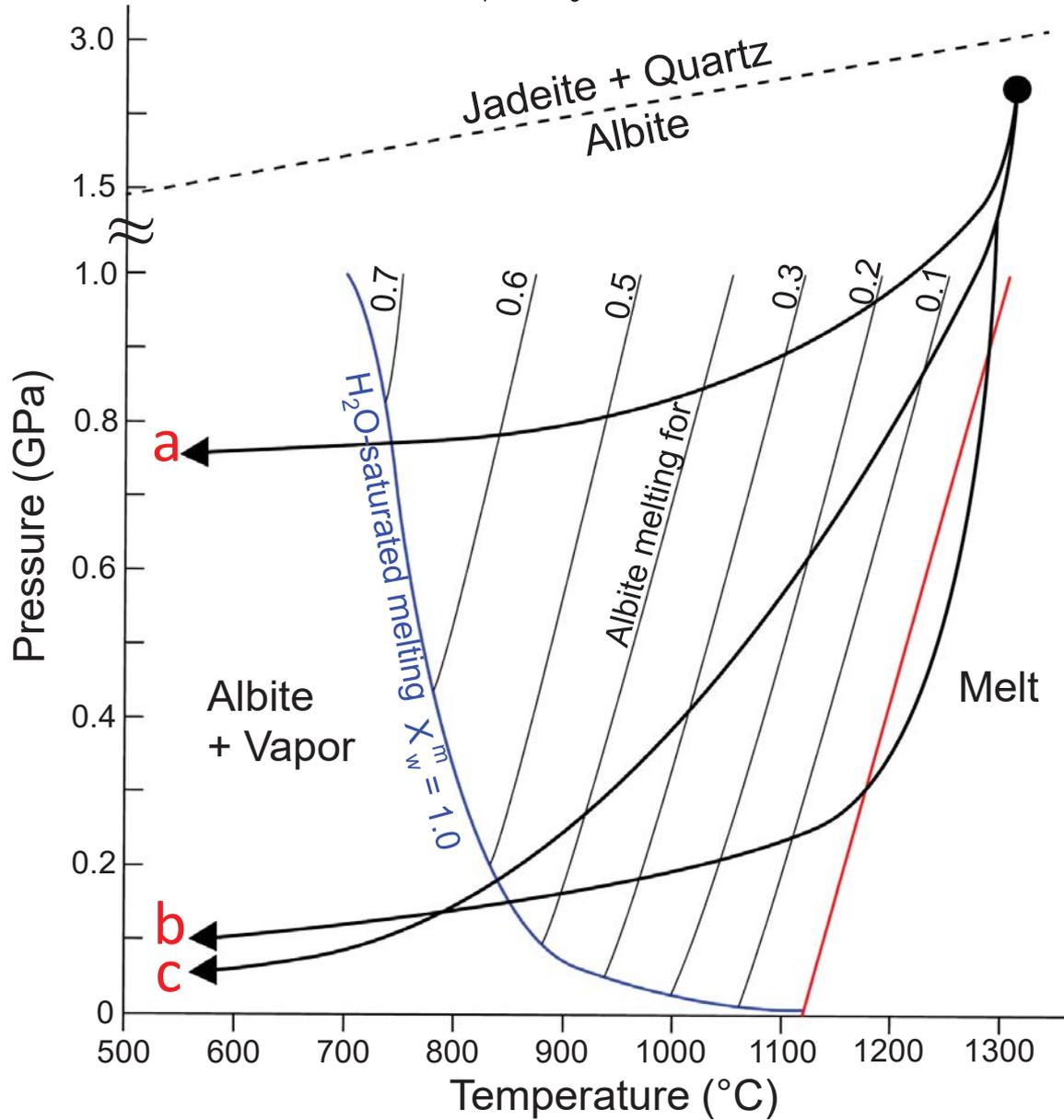


Fig. 8

Table 1. Recalculated EMPA clinopyroxene analyses from Bohemian Massif CPX grain, calculated with an assumed (not measured) Fe content of 2 wt. % Fe₂O₃; FeO = Fe – 2 wt % Fe₂O₃. Measurements made with 1µm beam. Oxygen calculated, all calculations based on 6 O. Potential calculated vacancy component = ([]).

Wt. % Oxide	Pt. 1	Pt. 2	Pt. 3	Pt. 4	Pt. 5	Pt. 6	Pt. 7	Pt. 8	Pt. 9	Pt. 10	AVG	STDEV	SERR
SiO ₂	52.88	53.11	53.50	53.32	53.48	53.73	53.24	53.12	52.93	53.11	53.24	0.26	0.08
TiO ₂	0.20	0.30	0.32	0.34	0.25	0.27	0.35	0.36	0.38	0.26	0.30	0.06	0.02
Al ₂ O ₃	8.39	7.67	7.38	7.63	7.62	7.63	8.05	7.99	8.26	8.34	7.90	0.36	0.12
Cr ₂ O ₃	0.06	0.08	bd	bd	bd	0.09	bd	0.07	0.05	bd	0.07	0.04	0.01
FeO	3.18	3.06	3.24	3.01	3.38	3.11	3.10	3.23	3.45	3.11	3.19	0.14	0.04
Fe ₂ O ₃ *	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	2.00	n/a	n/a	n/a
MnO	0.07	0.01	0.01	bd	0.02	0.05	0.09	0.11	0.09	0.06	0.06	0.04	0.01
MgO	11.30	11.66	11.63	11.71	11.87	11.63	11.75	11.40	11.17	11.19	11.53	0.24	0.08
CaO	17.21	17.61	17.54	17.67	17.65	17.37	17.57	17.48	17.24	17.41	17.47	0.16	0.05
Na ₂ O	4.03	3.83	3.79	3.76	3.82	3.85	3.89	3.83	3.88	3.98	3.87	0.09	0.03
K ₂ O	0.00	bd	0.00	bd	0.02	0.02	0.01	bd	0.02	0.00	0.01	0.01	0.01
Total	99.32	99.34	99.41	99.43	100.10	99.74	100.06	99.60	99.47	99.46			
Si	1.944	1.955	1.970	1.960	1.959	1.970	1.944	1.953	1.949	1.952			
Ti	0.006	0.008	0.009	0.009	0.007	0.007	0.010	0.010	0.011	0.007			
^{IV} Al	0.056	0.045	0.030	0.040	0.041	0.030	0.056	0.047	0.051	0.048			
^{VI} Al	0.308	0.288	0.290	0.291	0.288	0.300	0.290	0.299	0.307	0.313			
Cr	0.002	0.002	bd	bd	bd	0.003	bd	0.002	0.001	bd			
Fe ²⁺	0.076	0.082	0.100	0.093	0.080	0.095	0.072	0.099	0.106	0.091			
Fe ³⁺	0.022	0.012	0.000	0.000	0.012	0.000	0.023	0.000	0.000	0.004			
Mn	0.002	0.000	0.000	bd	0.000	0.002	0.003	0.003	0.003	0.002			
Mg	0.619	0.640	0.638	0.642	0.648	0.636	0.640	0.625	0.613	0.613			
Ca	0.678	0.694	0.692	0.696	0.693	0.682	0.687	0.689	0.680	0.686			
Na	0.287	0.273	0.271	0.268	0.271	0.274	0.275	0.273	0.277	0.284			
K	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.000			
[] = CaEs	0.000	0.002	0.000	0.001	0.001	0.001	0.000	0.000	0.001	0.000			
Cations per 6 O	4.000	4.000	4.000	4.000	4.000	4.000	4.000	4.000	4.000	4.000			

Notes for Table 1:

bd = below detection

[] = vacancy component

* 2 wt. % Fe₂O₃ chosen for calculations, not measured

Cations per 6 O

Oxygen calculated

Statistics:

n=19

AVG= average

STDV= standard deviation

SERR= standard error