1	<b>Revised Manuscript (Revision 2)</b>
2	Estimation of radiation damage in titanites using Raman spectroscopy
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8	Abstract
9	Recent studies have shown that $\alpha$ -damage in titanite influences He diffusivity and
10	thus the closure temperature of the (U-Th)/He system in titanite. We compare
11	different methods for measuring the $\alpha$ -dose in titanite by Raman spectroscopy.
12	293 Raman spectra of randomly oriented titanite fragments from the Archean
13	Karelian domain in eastern Finland along with some well-studied young titanites
14	and U-Pb standard reference materials were analyzed and related to the
15	concentration of $\alpha$ -emitting elements (U and Th) that generated damage in the
16	respective grains. Automated curve-fitting was performed by the IFORS software
17	and different curve-fitting protocols were tested and compared.

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18	The Raman bands at 424 cm <sup>-1</sup> and 465 cm <sup>-1</sup> show a good correlation of full width
19	at half maximum (FWHM) and position with the $\alpha$ -dose. However, these bands
20	are not always present because titanite is highly anisotropic implying that Raman
21	spectra are sensitive to orientation. The intensity-weighted mean FWHM (iw-
22	FWHM) of all Raman bands of a spectrum proves to be the most robust measure
23	of the $\alpha$ -dose. A simplified fitting approach considering 15 peaks is sufficient to
24	describe the accumulated $\alpha$ -dose. For $\alpha$ -doses below $5 \times 10^{16} \alpha/g$ the iw-FWHM
25	is independent of $\alpha$ -dose and ranges from 25 to 50 cm <sup>-1</sup> . Above this value the iw-
26	FWHM increases linearly with increasing $\alpha$ -dose up to $3 \times 10^{18} \alpha/g$ . The linear
27	correlation can be described as iw–FWHM[cm <sup>-1</sup> ] ~ $39(\pm 1.2)$ [cm <sup>-1</sup> ]+3.84(+0.61,-
28	$0.26$ ) × 10 <sup>-17</sup> [cm <sup>-1</sup> /( $\alpha$ /g)] × $\alpha$ -dose[ $\alpha$ /g]. The approach provides a pre–selection
29	method in order to optimize the range of $\alpha$ -doses of titanite crystals to be dated by
30	(U–Th)/He thermochronology.
31	keywords: titanite; (U–Th)/He; metamictization; radiation damage; $\alpha$ –dose;
32	Raman spectroscopy; thermochronology
33	Introduction
34	Low-T thermochronology is a widely-applied tool for deciphering thermal histories of cratonic
35	areas, especially when sedimentary successions constraining the timing of subsidence and
36	exhumation are missing. In contrast to zircon which has commonly effective uranium content
37	(eU=U[ppm]+0.235Th[ppm]) of $\geq$ 300 ppm (up to 5000 ppm), titanite has usually lower eU
38	contents in the range of 10-500 ppm and is thus less subjected to metamictization. Moreover,
39	titanite is common in intermediate to mafic rocks where zircon is typically absent.

40	The closure temperature (T <sub>c</sub> ) of the (U–Th)/He thermochronometers is influenced by the
41	composition and crystalline state of the dated crystals. Among all influencing parameters
42	metamictization is the dominant one (Flowers 2009; Orme et al. 2016; Johnson et al. 2017). The
43	influence of metamictization on the T <sub>c</sub> 's of the most widely used minerals, zircon and apatite, has
44	been intensively investigated (Shuster et al. 2006; Flowers et al. 2009; Gautheron et al. 2009;
45	Guenthner et al. 2013; Orme et al. 2016). In contrast, its influence on the titanite (U-Th)/He
46	thermochronometer (THe) has been reported only recently (Baughman et al. 2017; Guenthner et
47	al. 2017). These studies show that with increasing $\alpha$ -dose the T <sub>c</sub> of the THe system drops to a
48	significantly lower value. Unfortunately, several aspects of the THe system remain unclear
49	(Reiners and Farley 1999; Stockli and Farley 2004; Cherniak and Watson 2011). Titanite has
50	been discussed as host material for nuclear waste (Weber et al. 1998; Stefanovsky et al. 2004;
51	Lumpkin 2006) and various studies focus on the effects of radiation damage in this mineral
52	(Bismayer et al. 2010; Salje et al. 2011; Beirau et al. 2016).
53	The $\alpha$ -dose of zircon can be estimated from the width of its main Raman band at approximately
54	1000–1008 cm <sup>-1</sup> (Zhang et al. 2000; Nasdala et al. 2001; Palenik et al. 2003; Nasdala et al. 2004).
55	Therefore, Raman spectroscopy offers a quick, non-destructive method for selecting grains for
56	zircon (U–Th)/He (ZHe) thermochronology according to their $\alpha$ -dose (e.g. Ault et al. 2018). For
57	titanite, as for zircon, metamictization affects the Raman spectrum by broadening and shifting the
58	position of several bands (Salje et al. 1993; Meyer et al. 1996; Zhang and Salje 2003; Bismayer
59	et al. 2010; Beirau et al. 2012; Zhang et al. 2013). However, Raman spectroscopy of titanite is
60	more complicated when compared to zircon because of (i) strong anisotropy, (ii) higher amount
61	of Raman bands, and (iii) adjacent Raman bands often overlap in highly metamict titanite grains
62	(e.g. Beirau et al. 2012; Zhang et al. 2013). Moreover, titanite crystals usually have irregular

63 shapes. Therefore mineral fragments without crystal faces are typically used for THe dating, 64 which makes crystallographic orientation of the grains difficult. 65 So far, no advanced diffusion model exists for titanite that takes radiation damage into 66 consideration. As a consequence, THe ages measured on highly damaged grains are more 67 difficult to interpret than THe ages of slightly damaged grains. Therefore, the routine application 68 of THe thermochronology requires an easy-to-use, orientation independent method for damage 69 estimation and selection of the crystals to be dated. 70 This study compares different approaches considering single peaks as well as the entire Raman 71 spectrum for quantifying the  $\alpha$ -dose in titanite. We present coupled Raman, eU, and THe data of 72 titanite from the Archean Karelian domain in Fennoscandia. The area has been tectonically stable 73 since the Svecofennian orogeny (approx. 1.92–1.80 Ga) (Kähkönen 2005; Lahtinen et al. 2005; 74 Heller et al., in prep) and since then all samples have experienced the same thermal history. To 75 cover a broad range of  $\alpha$ -doses, measurements of younger titanite samples and titanite reference 76 materials with known age and simple cooling histories complement the dataset. The main goal of 77 this study is to establish a quick (and non-destructive) method for estimating the  $\alpha$ -dose from the 78 titanite Raman spectrum. We demonstrate that characterization of the entire Raman spectrum can 79 be used as a robust estimator independent of grain orientation and evaluation method.

80

### Background

## 81 Crystal structure and metamictization of titanite

82 Titanite, CaTi(SiO<sub>4</sub>)(O, OH, F), also called sphene, is a monoclinic nesosilicate. Zig-zag chains

83 of corner–sharing TiO<sub>6</sub> octahedra are cross–linked by isolated SiO<sub>4</sub>–tetrahedra, building a

84 framework. This (TiOSiO<sub>4</sub>)<sub>2</sub>-framework contains large cavities which enclose sevenfold-

85 coordinated Ca atoms (Speer and Gibbs 1976). Through lattice substitutions, titanite can 86 incorporate significant amounts of rare earth elements (REE), U, Th, Sr, Y, Mn and Pb at the 87 sevenfold Ca sites and at the octahedral Ti sites (Higgins and Ribbe 1976; Tiepolo et al. 2002; 88 Xu et al. 2015) and may also concentrate significant amounts of HFSEs, such as Nb, Ta, Zr and 89 W, on the octahedral Ti site (Groat et al. 1985; Lucassen et al. 2011). 90 The consecutive radioactive decay of U and Th atoms induces metamictization in titanite crystals 91 (Zhang and Salje 2003; Beirau et al. 2010; Bismayer et al. 2010; Beirau et al. 2012; Zhang et al. 92 2013; Beirau et al. 2014, 2016). In metamict titanites both crystalline and amorphous domains 93 coexist (Hawthorne et al. 1991; Lumpkin et al. 1991). Hawthorne et al. (1991) and Farges (1997) 94 suggest that metamictization leads to disordering around the Ti positions and a partial reduction 95 of the Ti coordination down to five and / or four. This further disturbs the Ti–O–Ti intra–chain 96 linkages (Beirau et al. 2012), triggering an anisotropic response of the titanite lattice. This means 97 that the plane containing Si-O-Ti-O bond closed contours is less susceptible to a self-98 accumulation of defects as compared to the perpendicular planes, which contain chains of Ti-O 99 bonds (Beirau et al. 2012).

## 100 Raman spectroscopy of titanite

101 Several studies have applied Raman spectroscopy to titanite (Salje et al. 1993; Meyer et al. 1996;

102 Zhang and Salje 2003; Bismayer et al. 2010; Beirau et al. 2012; Zhang et al. 2013). Titanite is an

anisotropic phase, exhibiting Raman spectra strongly dependent on the crystallographic

104 orientation. Up to 40 different Raman bands have been reported for natural titanite samples

105 (Zhang et al. 2013), whereas other authors (Salje et al. 1993) reported only up to 30 bands.

106 Increasing degree of metamictization (Fig. 1) changes the Raman spectrum by broadening bands,

107 shifting their frequency, decreasing the overall intensity and appearance of additional bands

- 108 (Zhang et al. 2013). As result of metamictization, weakening of the bands at 163 cm<sup>-1</sup>, 252 cm<sup>-1</sup>,
- $109 \quad 466 \text{ cm}^{-1}, 544 \text{ cm}^{-1}, 605 \text{ cm}^{-1}$  (see Fig. 1), appearance of new bands at 574 cm<sup>-1</sup> and 643 cm<sup>-1</sup>,
- 110 shifting of bands from 993 cm<sup>-1</sup> to 978 cm<sup>-1</sup>, 1017 cm<sup>-1</sup> to 1008 cm<sup>-1</sup> and 643 cm<sup>-1</sup> to 675 cm<sup>-1</sup>
- and appearance of a shoulder at 650 cm<sup>-1</sup> have been documented by (Zhang et al. 2013).
- 112 For estimating the degree of metamictization, Zhang et al. (2013) proposed to use the 605 cm<sup>-1</sup>
- 113 peak as it gets broader, shifts towards 612 cm<sup>-1</sup> and becomes increasingly asymmetric. Beirau et
- al. (2012) suggested estimating the degree of metamictization from the full width at half
- 115 maximum (FWHM) of the 465 cm<sup>-1</sup> band which arises from the  $SiO_4$ -bending mode (Heyns et
- al. 2000) and is narrow at low damage doses and broader at high doses. Unfortunately, both
- approaches can be realized only at given, limited crystallographic orientation and are therefore
- 118 not useful as universal method for spectra taken from fragments laying in different orientations.

## 119 Importance of radiation damage in (U–Th)/He thermochronology

- 120 The radioactive decay of U, Th, and Sm produces  $\alpha$ -particles (<sup>4</sup>He nuclei). Measuring the amount
- 121 of both <sup>4</sup>He and the parent isotopes makes it possible to calculate the duration of He
- accumulation (Farley 2002; Zeitler 2015).
- 123 He has a high diffusivity in solid phases, thus the preserved amount of He in a crystal depends on
- 124 diffusion that took place in the lifetime of the dated phase. He diffusion kinetics in crystals has
- been intensely studied; it is controlled by temperature, mineral type, grain size (diffusion domain
- size) and cooling rate (Zeitler et al. 1987; Wolf et al. 1996). Furthermore, mineral composition
- 127 and radiation damage are significant (Wolf et al. 1996; Nasdala et al. 2004; Reiners 2005).

128	Reiners and Farley (1999) have reported grain–size and cooling rate dependence of $T_c$ in the THe
129	system at 165–225°C. However, recent studies by Baughman et al. (2017) have shown that this is
130	not valid for highly metamict grains. These authors report T <sub>c</sub> 's of 150–210°C at low damage
131	levels and a rapid drop above a radiation damage threshold of ${\sim}5\times10^{17}\alpha/g$ (for minimal
132	estimates of the temperature where damage accumulation starts, the threshold is $\sim 3.5 \times 10^{17} \alpha/g$ ).
133	Despite the lower threshold of titanite compared to zircon (~ $1.5 \times 10^{18} \alpha/g$ ; Guenthner et al. 2013;
134	Baughman et al. 2017), titanite is likely to better document the higher temperature portion of a
135	protracted time-temperature path than zircon because of its usually order-of-magnitude lower
136	actinide contents (Baughman et al. 2017).
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149	The Ilomantsi greenstone belt was affected by two metamorphic events. The Neoarchean
150	metamorphism reached amphibolite facies peak conditions (550±50°C, 3–5 kbar; O'Brien et al.
151	1993). An ~2.70 Ga minimum (e.g. cooling) age for this metamorphism was deduced from 2.708
152	and 2.696 Ga U-Pb ages of titanite and monazite from granodiorite and leucogranite intrusions,
153	respectively (Vaasjoki et al. 1993). However, Hölttä et al. (2016) proposed 2.66 Ga and 2.62 Ga
154	age for the Neoarchean metamorphism based on dating of monazite growth. During the Lapland-
155	Savo orogenic phase of the Paleoproterozoic Svecofennian orogeny, thick, east-verging nappes
156	were emplaced on the Archean basement of eastern Finland and temperatures in the Ilomantsi
157	greenstone belt reached 400-500°C as indicated by partial reset of hornblende K-Ar ages
158	(Kontinen et al. 1992) and Pb isotope exchange studies (Halla and Heilimo 2009). Biotite K-Ar,
159	as well as biotite and muscovite Ar–Ar ages suggest cooling below ~350–300°C at ~1.8 Ga
160	(Kontinen et al. 1992; Molnár et al. 2016).
161	The later orogenies and accretionary events that occurred in Fennoscandia, such as the Gothian
162	(1.64–1.52 Ga) and the Telemarkian accretion (1.52–1.48 Ga), the Hallandian and Danopolonian
163	orogenies (1.47–1.42 Ga) and the Sveconorwegian orogeny (1.14–0.90 Ga) (Bingen et al. 2008)
164	affected mainly the southwestern margin of Fennoscandia and were of little and poorly
165	understood impact for eastern Finland. The lack of sedimentary record makes the reconstruction
166	of the post-Svecofennian thermal evolution difficult (Kohonen 2005).
167	A kaoline deposit at Virtasalmi, ~200 km WSW of the Ilomantsi greenstone belt, indicates
168	surface exposure of the Svecofennian crystalline basement during the Mesoproterozoic at ~1.18
169	Ga (Sarapää 1996). Late Neoproterozoic, in Fennoscandia referred to as Vendian, and Cambrian
170	sediments are common further south in Estonia. The extent of this unconformity and sediment

171 coverage towards north (e.g. Finland) is not clear and has been a matter of discussion (Puura et172 al. 1996; Larson et al. 1999; Sliaupa et al. 2006).

173 Thermochronologic data from eastern Finland comprise of a wide range of Neoproterozoic and

174 early Paleozoic apatite fission track (AFT) and apatite (U–Th)/He (AHe) ages (Lehtovaara 1976;

175 Lorencak 2003; Murrell 2003; Murrell and Andriessen 2004; Kohn et al. 2009). In combination

176 with new ZHe, THe and AHe data (Heller et al. in preparation) they indicate that the Ilomantsi

177 greenstone belt experienced slow cooling since the Svecofennian orogeny and that temperatures

178 did not exceed 100°C for the last 1.0 Ga.

179 For calculation of radiation damage densities the duration of  $\alpha$ -damage accumulation has to be

180 estimated. However, neither the thermal history of the samples in the time interval of 1.8-1.5 Ga

181 (see below), nor the temperature range of  $\alpha$ -damage annealing are well defined. While it is a

182 common praxis to estimate that annealing of  $\alpha$ -damage occurs at similar temperatures as

annealing of fission tracks (Baughman et al. 2017 for titanite and Pidgeon 2014 for zircon),

184 experiments by Gleadow (1978) indicate that  $\alpha$ -damage annealing of titanites occurs mainly in

185 the lower part of the titanite fission track partial annealing zone (TPAZ). For the TPAZ different

estimates exist (197°C: Naeser and Forbes 1976; 260±20°C Gleadow and Lovering 1978; 250 ±

187 50°C: Gleadow and Brooks 1979;  $240 \pm 40$ °C: Harrison et al. 1979;  $275 \pm 25$ °C: Fitzgerald and

188 Gleadow 1988;  $275 \pm 25$ C°: Kohn et al. 1993; 265–310°C: Coyle and Wagner 1998). Studies

189 including thermal annealing experiments (Bismayer et al. 2010; Beirau et al. 2012; Salje et al.

190 2012) indicate that  $\alpha$ -damage annealing is a complex process occurring over a temperature range;

- 191 while recovery of point defect is possible at lower temperatures (T<300°C), relevant
- 192 recrystallization of amorphous areas happens only at T>320°C. However, note that temperature

- 193 estimates from thermal annealing experiments are often too high as they don't include geological194 timescales.
- 195 In our case the onset of  $\alpha$ -damage accumulation can be bracketed to 1.5–1.8 Ga. The upper limit
- 196 is given by the ~1.8 Ga biotite and muscovite K–Ar and Ar–Ar ages (Kontinen et al. 1992;
- 197 Molnár et al. 2016). Note that Kohn et al. (1993) report that K-Ar ages are always greater or
- 198 equal to titanite fission track ages. The lower bracket is given by the oldest THe Ages(~1.5 Ga)
- 199 of our samples, which in our case of typical grains sizes and slow cooling correspond to
- 200 temperatures of  $\sim 180^{\circ}$ C.
- 201 The time interval from 1.5–1.8 Ga thus corresponds to a temperature drop from ~300°C to
- $\sim 180^{\circ}$ C, the onset of  $\alpha$ -damage accumulation lies somewhere in between. Assuming that post-
- 203 metamorphic cooling was faster in the beginning and then slowed down we decided to calculate
- with 1.7 Gy as damage accumulation duration. However, due to the uncertainties involved, we
- also present the results for 1.8, 1.6 and 1.5 Gy accumulation duration and discuss the respectiveerror below.
- 207 The sample set with relatively high damage densities from the Ilomantsi belt was complemented
- 208 by some well–studied titanites (partly reference materials for U-Pb geochronology) of
- 209 considerably lower  $\alpha$ -doses (Table 2).
- 210

#### Methods

## 211 Sample preparation

The mineral separation has been performed at the Geological Survey of Finland. The samples
were crushed, sieved and titanite fractions were separated by density and magnetic methods. For

single grain analysis, inclusion–free titanite fragments from the 63–200 µm sieve fractions were handpicked. 2 to 5 aliquots of 19 different samples from Finland were chosen. For the polished mineral grain mounts 9 Finnish samples as well as 6 low–damage samples were mounted in epoxy resin discs, ground and diamond–polished.

## 218 Raman spectroscopy

219 Raman measurements were done on (i) unmounted crystal fragments and (ii) polished grain

220 mounts. The measurements were performed with a Horiba XPloRa Plus system equipped with a

532 nm excitation laser (25 mW maximum output power) and an Olympus BX41 microscope,

which is coupled to a 200 mm focal length spectrograph with a four-grating turret. The system

has a Multi-Pinned-Phase (MPP) open-electrode CCD detector with a precision of  $\pm 1$  pixel.

Using 532nm the spectral resolution is 1.4 cm<sup>-1</sup> with the 2400 grooves/mm grating and 2.5 cm<sup>-1</sup>

with the 1800 grooves/mm grating. For the unmounted crystal fragments, the 1800 grooves/mm

grating was used and the Raman spectra were collected for  $2 \times 10$  seconds at 50% laser power

227 with 20x (0.4 NA) or 50x LWD (0.5 NA) objective in 1–3 spots per crystal fragment. The grain

228 mounts were measured with the 2400 grooves/mm grating at 10% laser power with 50x LWD

229 objective. Measurement times were  $3 \times 5$  to  $3 \times 20$  seconds depending on signal strength.

The positional drift of the spectrometer was checked regularly by measurement of a Si standard and is negligible for this dataset. In transparent crystals the penetration depth is about ~15  $\mu$ m for the 50x LWD objective and ~20  $\mu$ m for the 20x objective. The theoretical diameter of the Raman spots are ~1.3  $\mu$ m (50x LWD) and ~1.6  $\mu$ m (20x).

Spectra were recorded in the range 100–1800 cm<sup>-1</sup>. For detection of possible photoluminescence
bands additional Raman spectra were taken on some selected grains with a Horiba Jobin–Yvon

Labram HR800UV spectrometer equipped with a 633 nm excitation laser. Titanite spectra with a
superposed epoxy signal were not used for further evaluation.

238 Curve-fitting of the Raman spectra

239 The curve-fitting and spectral evaluation were performed with the IFORS software (Lunsdo f

and Lünsdor 2016). The signal part of the spectrum was modeled by pseudo-Voigt functions

241 (pV-functions). In each iteration, a new function was added to the model, and the sum of the

squared residuals was reduced by randomly changing a function parameter (e.g. center position,

height, width, or shape-factor) by a random value. Two main approaches were applied: (i) using

244 IFORS for background correction and curve-fitting, and (ii) manual background correction using

the Fityk software (Wojdyr 2010) followed by automated curve–fitting with IFORS.

246 Reasonable fitting parameters were found by testing different evaluation constraints and visually

checking the quality of the fit. The tested parameter values as well as the best configuration setup

can be found in Table 3 and Appendix 1.

249 During the curve–fitting, the spectrum is first scaled to 100 arbitrary intensity units and, in case

of automated background correction, the estimated baseline is subtracted in a second step. Thus,

the comparison of integrated intensities (by summing the areas of all pV-functions) and function

areas is only possible in evaluations without automated background correction.

253 The fitted band widths (FWHM = full width at half maximum) of the titanite spectra were

corrected for the apparatus function after Irmer (1985) and Nasdala et al. (2001).

### 255 Determination of the concentration of α-emitting elements and the (U-Th)/He ages

### 256 Unmounted titanite crystal fragments

257 For U and Th concentration measurements and (U-Th)/He age determination clear, inclusion-258 free fragments without external crystal surfaces were hand-picked using a stereo- and a 259 polarizing microscopes. These fragments were photographed and wrapped in platinum capsules 260 with 1x1 mm size. The analysis involved three separate steps: 1) He extraction and measurement 261 2) chemical digestion and 3) measurement of the parent isotopes. 262 For the He measurement the crystals were degassed in a high vacuum extraction line using an 263 infrared laser. A Ti–Zr getter at 450°C was used to purify the gas. The inert gases were measured 264 by a Hiden triple-filter quadrupole mass spectrometer with a positive ion-counting detector. For 265 every sample a re-extraction was done to check the complete degassing of the crystal fragment. 266 For determination of the mass of U. Th and Sm the degassed titanite fragments were dissolved in 267 a mixture of concentrated ultrapure HCl (500 µl) and HF (100 µl) (Reiners and Farley 1999) in pressurized teflon vials during one day at 220°C, and spiked with calibrated <sup>230</sup>Th and <sup>233</sup>U 268 269 solutions. After cooling, the samples were dried, converted to nitrate by re-dissolving in 0.25-270 1 ml 65% HNO<sub>3</sub> and after a further drying step dissolved in a mixture of 4% HNO<sub>3</sub> with 0.05% 271 HF. 272 U, Th, Sm and other REE measurements were performed using a Perkin Elmer Elan DRC II

assembled with an APEX MicroFlow nebulizer and a Thermo iCAP Q ICP–MS instrument. The

actinides were determined by isotope dilution technique, while for Ca, Ti and the other trace

- elements the external calibration method was applied.
- 276 Mounted titanite fragments
- The trace element content of titanite grains mounted in polished epoxy resin discs was analyzed
  by Laser–Ablation ICP–MS. A Resonetics excimer laser was used, coupled to a Thermo

Element 2 sector field mass spectrometer. Measurements were conducted at the same spots as the
Raman measurements and performed with 10 µm laser beam diameter and short ablation times.
For conversion of the cps data to concentrations the NIST610 and TNT1500 glass standards were
used (Klemme et al. 2008).

283

### Results

## 284 Raman spectra

285 82 Raman spectra of 63 unmounted crystal fragments from 19 Finnish samples were acquired, as

well as 152 spectra of 9 titanite mounts of the Finnish samples and 59 spectra of 6 mounted low-

287 damage samples, totaling 293 titanite Raman spectra.

288 The Raman spectra show a broad variation in band widths, absolute and relative intensities of the

bands, as well as shifts in band positions. Figure 1 shows two very different spectra obtained on

290 grains with different  $\alpha$ -doses. At the acquisition of the spectra the crystals were randomly

291 oriented, thus the entire range of anisotropy effects is represented (see e.g. Beirau et al. 2012;

292 Zhang et al. 2013).

- 293 Due to the edge–filter, which blocks the signal below  $\sim$ 70 cm<sup>-1</sup>, most spectra start with signal and
- 294 many have high Raman intensities at  $100-200 \text{ cm}^{-1}$ . Therefore, the background signal is

detectable only at higher wavenumbers (>1000 cm<sup>-1</sup>, Figure 1).

- For an excitation wavelength of 532 nm the major photoluminescence bands (Sm, Nd) occur at
- higher wavenumbers than the titanite Raman bands (Lenz et al. 2015). The comparison of the
- spectra with spectra acquired with 633 nm excitation shows no differences, which indicates that
- 299 no REE-related photoluminescence signals overlie the Raman signal. The only

- 300 photoluminescence that theoretically occurs in the Raman wavenumber range is an Er-generated
- 301 peak at 564 nm (i.e.  $=1060 \text{ cm}^{-1}$ ), but this has not been detected.

## 302 Effective uranium (eU) contents and α-doses

303 The eU contents of the studied titanite crystal fragments range from 6 ppm to 935 ppm. The  $\alpha$ -

304 doses were calculated as:

$$D_{\alpha} = 8 \cdot \frac{C_U \cdot N_A \cdot 0.9928}{M_{238} \cdot 10^6} \cdot (exp^{\lambda 238} \cdot t - 1)$$
  
+  $7 \cdot \frac{C_U \cdot N_A \cdot 0.0072}{M_{235} \cdot 10^6} \cdot (exp^{\lambda 235} \cdot t - 1)$   
+  $6 \cdot \frac{C_{Th} \cdot N_A}{M_{232} \cdot 10^6} \cdot (exp^{\lambda 232} \cdot t - 1)$   
(Eq. 1)

305

where C = actinide concentrations (in ppm),  $N_A = \text{Avogadro's number } (6.022 \times 10^{23} \text{ atoms/mol})$ , M = the molecular masses,  $\lambda = \text{the decay constants for the different isotopes, and } t = \text{the onset of}$ damage accumulation (t is 1.7 Ga for the Finnish samples as discussed above or the respective ages of the low damage titanites, see Table 2). The  $\alpha$ -doses of the Finnish samples range from  $4 \times 10^{16}$  to  $2.9 \times 10^{18} \alpha/\text{g}$  and those of the additional samples with lower  $\alpha$ -dose from  $1.3 \times 10^{13}$ to  $1.8 \times 10^{17} \alpha/\text{g}$  (Appendix 2). Altogether, the  $\alpha$ -dose in the samples ranges over more than five orders of magnitude.

## 313 Quality of the curve-fitting of the Raman spectra

- 314 In order to establish reasonable curve–fitting parameters and comparison of their effects on
- further calculations, we applied a systematic sensitivity study and several curve–fitting protocols

316	were tested (Table 3 and Appendix 1). For the completely automated fittings (including
317	automated background correction), reasonable results (setting "wf_sm") were achieved when
318	using a "window_filter" parameter with 0.35 window size for baseline calculation, setting the
319	"noise" parameter to 2.5 or 3, allowing 40 pV-functions and smoothing the spectra with a
320	"smoothing size" parameter of 9 (for all other parameter see Appendix 1). Lower "noise"
321	parameter values lead to a solution with many small pV-functions while higher "noise"
322	parameter values or severe restrictions on the maximum number of allowed pV-functions either
323	lead to incomplete curve-fits (i.e. some bands were not described by pV-function) or imprecise
324	fitting of broad bands. In most cases, three repeated curve-fits yielded consistent, well
325	reproducible results. Only a very minor proportion of the spectra showed poor reproducibility.
326	A problem during curve-fitting of metamict titanite spectra is the inherently unknown
327	background at low wavenumbers. Assuming a flat baseline, even at low wavenumbers seems to
328	be reasonable, however, dispersed light at low wavenumbers might cause some increase in the
329	background intensity. This is why we tested both the automated and the manual background
330	correction. The automated curve-fitting often estimates a baseline which is elevated at low
331	wavenumbers (Fig. 1). Too small "window size" parameter values (e.g. 0.25) produced
332	unreliable, bended baselines. For manual background correction a flat baseline (i.e. not elevated
333	at lower wavenumbers) was assumed.
334	The curve-fitting sometimes produces unapparent solutions; the most common type is when a
335	range in the spectrum is modeled by fitting of one flat, wide pV-function and three small narrow
336	pV-functions instead of a triplet with three medium sized functions. The only way to avoid the
337	dubious flat wide pV-functions is by limiting the "maximum peak width" parameter. Setting the
338	"maximum peak width" parameter to 80 or 100 cm <sup>-1</sup> (given as HWHM = half width at half

maximum) caused fewer artifacts than allowing it to become as high as  $150 \text{ cm}^{-1}$ . However, harsh restriction of the pV-function width (e.g. values below  $80 \text{ cm}^{-1}$ ) is critical because in our case the widths are the most important result of the curve-fitting procedure.

342

## Possibilities of α-dose estimation from the Raman spectra

## 343 Single band testing

Beirau et al. (2012) proposed the use of the FWHM of the 465 cm<sup>-1</sup> band, which is assigned to 344 345 the SiO<sub>4</sub>-bending mode (Heyns et al. 2000; Heyns and Harden 2013), as an estimator for the 346 degree of metamictization because it broadens with increasing  $\alpha$ -dose. Figures 3A and 3B show the FWHM of this band (defined by the center values between 458  $\text{cm}^{-1}$  and 468  $\text{cm}^{-1}$ ) versus the 347 348  $\alpha$ -dose of each spot (Appendix 3). The colors of the data points indicate the band position. In very low damage samples, the band has a width of  $\sim 20$  cm<sup>-1</sup> and centers at  $\sim 468$  cm<sup>-1</sup>. Above 349  $\sim 5 \times 10^{16} \alpha/g$  the band starts getting broader and shifts towards lower wavenumbers. However, for 350 351 highly damaged crystals the band widths are very scattered (Fig. 3A) and an accurate estimation 352 becomes difficult. Another problem with the use of this band is its absence in some spectra. 353 The band at 424 cm<sup>-1</sup> (415 cm<sup>-1</sup> to 425 cm<sup>-1</sup>, Figs. 3C and 3D), which is also assigned to the SiO<sub>4</sub>– 354 bending mode (Heyns et al. 2000; Heyns and Harden 2013), shows a very similar behavior. Both width and position correlate with the  $\alpha$ -dose (Appendix 4). Starting from ~5 × 10<sup>16</sup>  $\alpha$ /g, band 355 broadening from  $\sim 18$  cm<sup>-1</sup> to >50 cm<sup>-1</sup> and positional shift from 424 cm<sup>-1</sup> to 416 cm<sup>-1</sup> are 356 observed. The 424 cm<sup>-1</sup> band also does not appear in all spectra. 357 The bands at 163 cm<sup>-1</sup>, 252 cm<sup>-1</sup>, 544 cm<sup>-1</sup> and 605 cm<sup>-1</sup> for which Zhang et al. (2013) report 358 band broadening, as well as the band at 535 cm<sup>-1</sup> were likewise investigated. However, we could 359 not determine significant correlations. For the bands at 163 cm<sup>-1</sup>, 252 cm<sup>-1</sup> and 535 cm<sup>-1</sup> the band 360

17

361	widths are scattered and show poor positive correlation with the $\alpha$ -dose while the positions
362	appear to be uncorrelated. For the band at 544 cm <sup>-1</sup> neither the width nor the position shows any
363	correlation with the $\alpha$ -dose.
364	According to Zhang et al. (2013) the band at 605 cm <sup>-1</sup> broadens, shifts towards 612 cm <sup>-1</sup> and
365	becomes more asymmetric with increasing metamictization. However, Beirau et al. (2012) show,
366	that depending on the orientation the band either shifts towards higher wavenumbers or shifts
367	towards lower wavenumbers. Verifying this behavior is hindered by the asymmetric character of
368	the band. Most of the curve-fits contain at least one pV-function in the range of 590–620 cm <sup>-1</sup> ,
369	some contain 2, rarely 3 pV-functions could be fitted. No correlations with $\alpha$ -dose were
370	observed.

## 371 Characterization of the entire Raman spectrum

372 Metamictization influences some bands more than others but generally it broadens all Raman 373 bands (Zhang et al. 2013). In order to relate this general broadening with the corresponding  $\alpha$ -374 doses we looked for parameters which describe the modification of the spectra with a single 375 value. We calculated medians, unweighted and weighted means of all band widths for every 376 spectrum. The use of the mean pV-function areas was also considered in some test calculations, 377 however, it turns out that the comparison of the function areas is biased in all spectra using 378 automated background correction, because the absolute areas are affected by the baseline height. 379 The use of mean width values instead of single band widths has distinct advantages in the case of 380 metamict titanite where the fitting is hampered by overlapping and asymmetric bands. 381 For the loose grains where two or three spectra were recorded, means and medians combining the 382 two or three spectra were calculated. FWHM means were weighted either by (i) scaled pV-

383	function intensity (iw-FWHM) or by (ii) scaled pV-function area (aw-FWHM). To test the
384	reproducibility all spectra were curve-fitted three times. Means and medians were calculated for
385	each curve-fit as well as for the average spectrum joining all three evaluations. Usually, the
386	numeric difference between the three repetitions is less than 10%, but can be as high as 20% for
387	some spectra (due to overlapping broad bands). Therefore, the use of repeated spectral
388	evaluations and means over all evaluations is highly recommended. The (weighted) means and
389	medians were compared for all tested curve-fitting protocols. All data can be found in Appendix
390	2.
391	All means and medians show good correlation with the $\alpha$ -dose (nearly all p-values $<<2.2 \times 10^{-10}$
392	<sup>16</sup> ). However, unweighted means and medians are much more sensitive to the curve–fitting
393	parameters and spectral properties (e.g. signal-to-noise ratio) than weighted means. Examples
394	are presented in Appendix 5.
395	The use of weighted mean FWHM has the advantage that it is less sensitive to the total number
396	of used pV-functions and the larger pV-functions have stronger influence than small peaks.
397	Moreover, the weighted mean FWHMs, especially iw-FWHMs, are much more robust because
398	different curve-fitting protocols yield nearly the same results (see Fig. 4). The iw-FWHM yield a
399	better result than the aw-FWHM. The dependency of some bands on the orientation of the crystal
400	seems to have only minor impact on the iw-FWHM.
401	
401	I ne most important sensitivity test: influence of the maximum number of pv-functions on

402 the curve–fitting results

403 We demonstrated that the iw-FWHM procedure leads to more consistent results than the area-

404 weighted procedure. Thus, a systematic sensitivity test series was performed to study the impact

405	of the maximum number of pV-functions on the results of the curve-fitting. The repeated
406	evaluations allowed for 10, 15, 20, 30 or 40 pV-functions. Allowing 15 up to 40 pV-functions
407	we recorded always the same trend (see Figs. 4 and 5). However, with only 10 pV-functions the
408	spectra were often poorly fitted and especially in the well-crystalline samples not all bands were
409	described by pV-functions resulting in higher iw-FWHM values. Thus, curve-fitting with a
410	maximum of 15 pV-functions is sufficient to characterize the $\alpha$ -dose in titanites by the iw-
411	FWHM method, even if these 15 pV-functions might not represent the spectra in all details.
412	Discussion
413	Our results suggest that the intensity weighted mean FWHM is a reliable parameter to describe
414	the $\alpha$ -dose in titanite. The restriction to maximal 15 pV-functions led to consistent results.
415	Below an $\alpha$ -dose of 5 × 10 <sup>16</sup> $\alpha$ /g the iw-FWHM is independent of $\alpha$ -dose and scatters between
416	25 cm <sup>-1</sup> and 50 cm <sup>-1</sup> . Above this value, the iw–FWHM increases linearly with increasing $\alpha$ –dose
417	up to $3 \times 10^{18} \alpha/g$ .
418	Figure 5A compares the linear regression parameters (slope, y-axis intercept and R <sup>2</sup> ) of all
419	calculated values for all tested curve-fitting protocols. All linear regression data can be found in
420	Appendix 6. The iw-FWHM linear regressions yield the tightest cluster of results associated with
421	high coefficients of determination (Fig. 5A). All the other methods yield more scattered results
422	and especially the median linear regressions have often very low coefficients of determination.
423	This implies that regardless of the curve-fitting protocol used for spectral evaluation, the iw-
424	FWHM method produced nearly the same linear regression, which makes this method the most
425	robust of all tested procedures.

426 Thus, we recommend to use the iw–FWHM as radiation damage estimator for titanite. For the 427 highest density point in Figure 5A the formula for the linear increase of the iw–FWHM with growing  $\alpha$ -dose is iw-FWHM[cm<sup>-1</sup>]~39.2[cm<sup>-1</sup>]+3.82 × 10<sup>-17</sup>[cm<sup>-1</sup>/( $\alpha$ /g)] ×  $\alpha$ -dose[ $\alpha$ /g]. This is 428 429 similar to the formula for the most satisfying protocol "wf sm" (see details in Appendix 1), 430 which is iw-FWHM[cm<sup>-1</sup>]~38.8(±1.2)[cm<sup>-1</sup>]+3.86(±0.13) × 10<sup>-17</sup>[cm<sup>-1</sup>/( $\alpha/g$ )] ×  $\alpha$ -dose[ $\alpha/g$ ]. Combining both formulae we get iw-FWHM[cm<sup>-1</sup>]~39.0( $\pm 1.2$ )[cm<sup>-1</sup>]+3.84( $\pm 0.13$ ) × 10<sup>-17</sup>[cm<sup>-1</sup>] 431 432  $\frac{1}{(\alpha/g)} \propto \alpha$ -dose[ $\alpha/g$ ]. The constants of the regression equation are determined by ca. 3% relative 433 uncertainties. 434 The other, even more important bias that controls the accuracy of the constants comes from the 435 assumed duration of the damage accumulation. As discussed above, the temperature range where 436 radiation damage accumulation starts for the Finnish samples is not known precisely, neither is

the annealing behavior in titanite. It is not clear how the density of accumulated damage

438 influences the annealing kinetics, and the effect of the variable chemical composition of titanite is

439 not known. Additionally the thermal history of the Finnish samples bears some uncertainty in the

440 interval of 1.8–1.5 Ga. Figure 5B visualizes the effect of this uncertainty on the linear

441 regressions. While the y-intercepts remain unchanged, shorter damage accumulation durations

lead to steeper slopes as the absolute damage gets smaller. This uncertainty is higher than the

443 error derived from the spread of the data. Including this uncertainty into the error of the linear

regression leads to a corrected calibration line: iw-FWHM[ $cm^{-1}$ ]~39.0(±1.2)[ $cm^{-1}$ ]+3.84(+0.61, -

445 
$$0.26$$
 × 10<sup>-1</sup>/[cm<sup>-1</sup>/( $\alpha$ /g)] ×  $\alpha$ -dose[ $\alpha$ /g]

446 The onset of the significant change of the titanite Raman peak width is around 
$$5 \times 10^{16} \alpha/g$$
.

- 447 Baughman et al. (2017) studied the He diffusivity and the THe closure temperature in relation to
- 448 the  $\alpha$ -dose. Their data indicate a threshold, where T<sub>c</sub> strongly decreases at around ~5 × 10<sup>17</sup>  $\alpha$ /g

449	(Fig. 4). It is remarkable that our procedure based on the characterization of the entire Raman
450	spectrum is more sensitive and detects the change of the lattice at one order of magnitude lower
451	level. This indicates the need of further He diffusion experiments in order to quantify the
452	diffusion parameters for titanite in the relatively low damage range.

453

### Comparison with (U-Th)/He results

- 454 For some of the Finnish samples (U–Th)/He ages were determined. Figure 6A shows these ages
- 455 compared to their  $\alpha$ -dose and eU content. The data show a negative correlation and the covered
- 456 eU range is relatively small. Compared to our data, the data presented by Baughman et al. (2017)
- 457 shows a better negative correlation and includes higher eU contents.
- 458 Figure 6B shows the THe age versus the iw–FWHM (fitting with 15 pV–functions). As expected,
- the data show a negative correlation, however, the spread is high. Possible reasons for the
- 460 dispersion include the moderate correlation of THe age and  $\alpha$ -dose, especially at low  $\alpha$ -doses,
- 461 influence of REE on THe ages, alpha ejection correction (Farley et al. 1996; Stockli and Farley
- 462 2004), zoning and inclusions (the details are discussed in Heller et al. in prep).
- 463 Estimation of the α-dose cannot substitute the THe age determination, however, it enables a pre-
- selection of the grains most appropriate for the target of a specific study.
- 465

## Implications

- 466 We present a robust Raman spectroscopy based method for estimating the  $\alpha$ -dose in titanite by
- 467 Raman spectroscopy. The titanite (U-Th)/He thermochronometer is relevant for the
- 468 reconstruction of thermal histories in the upper crust, because titanite occurs in more mafic
- 469 lithologies, in contrast to zircon, and its typically lower actinide content allows for obtaining

470	reliable thermochronologic information in cratonic areas that experienced cooling in Precambrian
471	times. The method can be used in two ways for the routine work in (U-Th)/He
472	thermochronology: (i) It can provide an easy-to-apply and quick tool for identifying highly
473	damaged titanite grains that have extremely low closure temperature due to the high $\alpha$ -dose. In
474	such crystals the recoil and fission tracks are already in contact and provide a kind of
475	"percolation" leading to very quick helium diffusion (Baughman et al. 2017). Rejecting these
476	titanite grains can optimize the dating procedure by analyzing only the better crystalline grains
477	for which the closure temperature is considerable above the ambient temperature. (ii) The other
478	approach is to concentrate on the wide range of the $\alpha$ -doses and aim at involving grains that
479	represent both extremely low and high $\alpha$ -doses in order to perform a batch modelling (e.g. Orme
480	et al. 2016; Johnson et al. 2017; Hueck et al. 2018). In this way the thermal history may be
481	described with better confidence and even the lower temperature ranges may be better
482	constrained by the titanite grains with very low closure temperature. The method proposed here
483	provides a pre–selection method in order to optimize the range of $\alpha$ –doses of the dated grains.
40.4	

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## 765 Appendices

- 766 Data Repository for "Estimation of radiation damage in titanites using Raman spectroscopy"
- The material in this data repository containing 5 Tables, one figure and one zip-file includesinformation on:
- 36

- 769 Appendix 1: Used IFORS configuration files (14) and one overview table. (Appendix-
- 770 3\_IFORS\_config\_files.zip)
- 771 Appendix 2: Alpha-emitting element contents of the samples and compilation of mean peak
- values generated by the evaluation of the entire spectra. Table shows all results of the 14 different
- settings applied (see explanation in text and Appendix 1). (Appendix-
- 2\_Compil\_damage\_and\_Raman\_entire\_evaluation\_all\_settings.xls)
- Appendix 3: Peak width and position values of the band 465 cm-1 using setting wf\_sm (see
- details in Appendix 1). (Appendix-3\_Single-band-evaluation\_465.xls)
- Appendix 4: Peak width and position values of the band 424 cm-1 using setting wf\_sm (see
- details in Appendix 1). (Appendix-4\_Single-band-evaluation\_424.xls)
- Appendix 5: Plot of (unweighted) mean FWHMs versus  $\alpha$ -dose for 3 different curve-fitting
- 780 protocols (Appendix-5\_meanfwhm-alphadose.pdf).
- 781 Appendix 6: Regression line parameters for all curve–fitting protocols for different damage
- accumulation durations (plots presented in Figure 5). (Appendix-
- 783 6\_Regression\_line\_parameters.xls)
- 784 Appendix 7: (U–Th)/He age data (Appendix-7\_THe\_ages.xls)
- 785 Figure captions
- Figure 1: Two Raman spectra of titanite crystals with different  $\alpha$ -doses (a:  $1.3 \times 10^{13} \alpha/g$  and b:
- $1.25 \times 10^{18} \text{ a/g}$ ). The spectra were evaluated with the software IFORS (Lünsdo f and Lünsdor
- 2016). The black lines are the spectrum data, which are mostly covered by the dashed green lines
- representing fitted data, blue dashed lines are the calculated background correction and thin red

lines are the respective pseudo–Voigt functions. The black numbers give the positions of the
important bands in cm<sup>-1</sup>.

792	Figure 2: (a): Geological overview of Fennoscandia, the black rectangle refers to the area in (b)
793	that shows the entire Ilomantsi-Kostomuksha greenstone belt. (c): Geological map of the
794	Ilomantsi greenstone belt (purple) with the sampling locations (orange circles) in the Karelian
795	domain of Fennoscandia (modified after Koistinen et al. 2001). The Ilomansti greenstone belt
796	consist of the Hattu schist belt (purple polygon) and the Kovero belt (dashed purple polygon).
797	Figure 3: $\alpha$ -dose of the analyzed titanite grains versus the FWHMs of the 465 cm <sup>-1</sup> band (a & b)
798	and the 424 cm <sup>-1</sup> band (c & d). The linear and logarithmic plots show the same data, but the semi-
799	logarithmic presentation emphasizes better the onset of the significant changes in the crystal
800	lattice - which has prominent importance at helium thermochronology (see in Discussion). The
801	color of the points indicates the position of the center of the pV-functions. Only pV-functions
802	with an area greater than 1% of the total area and a $FWHM < 75 cm^{-1}$ were considered. Black
803	continuous lines in the linear presentations (a) and (c) are linear regressions shown with 0.95
804	confidence intervals (grey belts). Their formulae are FWHM[cm <sup>-1</sup> ]~19.5(±0.5)[cm <sup>-1</sup> ]
805	<sup>1</sup> ]+1.42( $\pm 0.06$ ) × 10 <sup>-17</sup> [cm <sup>-1</sup> /( $\alpha/g$ )] × $\alpha$ -dose[ $\alpha/g$ ], R <sup>2</sup> =0.52 for the 465 cm <sup>-1</sup> band (a) and
806	$FWHM[cm^{-1}] \sim 15.6(\pm 0.3)[cm^{-1}] + 1.30(\pm 0.03) \times 10^{-17}[cm^{-1}/(\alpha/g)] \times \alpha - dose[\alpha/g], R^2 = 0.69 \text{ for the}$
807	424 cm-1 band (c). The dashed fit in (a) and (b) with FWHM[cm <sup>-1</sup> ]~16.9( $\pm 0.3$ )[cm <sup>-1</sup> ]
808	$^{1}$ ]+1.83(±0.07) × 10 <sup>-17</sup> [cm <sup>-1</sup> /( $\alpha$ /g)] × $\alpha$ -dose[ $\alpha$ /g], R <sup>2</sup> =0.61 was calculated for all samples with
809	FWHM<40cm <sup>-1</sup> and damage density $< 1 \times 10^{18} \alpha/g$ and is more adequate for the low–damage
810	range. The grey dashed vertical lines in (b) and (d) correspond to $5 \times 10^{16} \alpha/g$ and $5 \times 10^{17} \alpha/g$ .
811	For data see Appendices 5 and 6.

812 Figure 4: Intensity weighted mean FWHMs (iw-FWHM) versus the  $\alpha$ -dose in linear (a) and

- semi-logarithmic (b) scale for two curve-fitting protocols, wf\_sm allowing up to 40 pV-
- functions and *wf\_sm\_max15* allowing up to 15 pV-functions. Black continuous lines correspond
- to the linear regression calculated over both datasets, shown in (a) with 0.95 confidence interval
- 816 (grey belt). The formula is iw-FWHM[cm<sup>-1</sup>]~39.3( $\pm 0.8$ )[cm<sup>-1</sup>]+3.89( $\pm 0.09$ ) × 10<sup>-17</sup>[cm<sup>-1</sup>]
- 817  $\frac{1}{(\alpha/g)} \times \alpha$ -dose[ $\alpha/g$ ], R<sup>2</sup>=0.75. The dashed lines indicate the possible threshold values of
- 818  $5 \times 10^{16} \alpha/g$  (this study) and  $5 \times 10^{17} \alpha/g$  [2]. Both setups show the same trend which is flat below
- 819  $5 \times 10^{16} \alpha/g$  and linear above this value. For fitting details see Appendix 1.

Figure 5: (a) Summary of the results of the different fitting methods. The axes represent the linear regression parameters of the FWHM vs.  $\alpha$ -dose correlations. Symbol shapes refer to calculation

method; big symbols are curve–fitting protocols with automated background correction, small

symbols are protocols with manual background correction. The color coding gives the coefficient

of determination of the linear regressions, the blue lines contour the density distribution. All

825 linear regression parameters for each setting can be found in Appendix 5. The generally higher

values for weighted means compared to unweighted means and medians are due to the intensity

and area weighting. The tightest distribution was generated by the intensity weighted mean

- 828 FWHMs (iw-FHWM), which yielded very similar results for all different settings, show
- 829 consistently high R<sup>2</sup> and are thus considered the most robust calculation method. (b) shows the
- 830 linear regression parameters for only two of the curve–fit protocols, *wf\_sm* (open symbols)
- allowing up to 40 pV-functions and *wf\_sm\_max15* (filled symbols) allowing up to 15 pV-
- 832 functions. Linear regressions were calculated for different α–damage accumulation durations
- 833 which are displayed in different colors. For the intensity weighted mean FWHM the two

- 834 protocols give very similar results. Lowering the  $\alpha$ -damage accumulation duration by 0.1 Gyr
- 835 increases the slope values by  $\sim 7-8\%$ .
- Figure 6: Titanite (U–Th)/He ages in relation to their  $\alpha$ -dose, eU contents and the intensity
- 837 weighted mean FWHM (iw-FHWM) of the corresponding Raman spectra. (a) Titanite (U-
- 838 Th)/He (THe) ages versus their  $\alpha$ -dose, color indicates the eU contents. (b) Relation of the THe
- ages vs. iw-FHWM. The color indicates the  $\alpha$ -dose. The observed dispersion is mainly due to the
- rather poor THe age vs. eU correlation of our dataset.

Table 1: Locality, lithology and emplacement age of the studied titanite-bearing formations from Finland.

					Zircon U-	Titanite U-	
					Pb Age	Pb Age	
Sample)	YKJ-north <sup>1</sup> Y	'KJ-east <sup>1</sup>	Location	Rock type	[Ma] (±2s)	[Ma] (±2s)	Reference
				Granodiorite			Vaasjoki et al. 1993;
A0050	6958398 3	3702479	llomantsi Kk	(sanukitoid)	2728 (7)	2707 (10)	Heilimo et al., 2011
							Sorjonen-Ward &
							Claoue-Long, 1993;
A0284	6973076 3	3710317	Lehtovaara Ilomantsi	Granodiorite	2752 (4)	2705 (10)	Vaasjoki et al 1993
							Vaasjoki et al., 1993;
A0301	6968216 3	3697164	Vehkavaara Ilomantsi	Porphyry (felsic, dyke)	2755 (4)	2720 (10)	Huhma et al., 2012a
A0338	6968216 3	3697164	Vehkavaara Ilomantsi	Porphyry (felsic, dyke)		2747 (15)	Vaasjoki et al., 1993
A0339	6983951 3	3706826	Silvevaara Ilomantsi	Granodiorite (porphyrit	ic)	2658 (10)	Vaasjoki et al., 1993
A0357	6945820 3	3530110	Paukarlahti	Gneiss (granitic, dome	)	1824 (10)	Gaal, 1980
A0804	6982475 3	3586352	Luikonlahti Kaavi	Granite			
A0951	6946074 3	3696122	Paavonvaara Ilomants	Granite			
				Diorite (quartz-,			Vaasjoki et al., 1993;
A1078	6953174 3	3715702	Oinassalmi Ilomantsi	sanukitoid)	2744 (3)		Heilimo et al., 2011
A1095	6974229 3	3713990	Kivisuo Ilomantsi	Porphyry (Qtz-Fldsp)	2756 (6)		Vaasjoki et al., 1993
							Halla, 2002; Mikkola
A1340	7003399 3	3669913	Jaakonvaara Lieksa	Diorite	2702 (5)		et al., 2013
A1614	7003710 3	3542680	Pajulahti Nilsiä	Syenite			
A1626	6944376 3	3674468	Rasisuo	Gabbro	2756 (4)		Huhma et al., 2012a
A1640	6928741 3	3675065	Huhtilampi	Granite			
A1641	7009265 3	3711422	Kierrosaho	Granite (microcline)			
A1762	7027497 3	3661337	Emonvaara	Tonalite		2732 (15)²	Huhma et al., 2012b
A1963	6959880 3	3708970	Kuikanniemi Ilomantsi	Granodiorite			
A1964	7021049 3	3685422	Kitsi	Granite			
A2196	6992639 3	3718302	Viluvaara	Granodiorite	2751 (4)	2737 (15)	Käpyaho et al., 2016

<sup>1</sup>: Finnish national grid. <sup>2</sup>:WR Sm-Nd age

## Sheet1

 Table 2: Sources and major parameters of the selected titanite reference

 samples having low radiation damage densities.

Sample	Locality	Source	U-Pb Age [Ma]
ECS	Ecstall pluton, British Columbia	Butler et al. (2002)	91.5
94T	U-Pb reference material; Univ. Arizor	from George Gehrels	51.5
ADA-15	Adamello massif, S. Alps	Schoene et al. (2012)	41.2 <sup>1</sup>
LIM	Kaiserstuhl ash, Rhine valley	Kraml et al. (2006)	16.5
MK-221b	Csomád volcano, Carpathians	Molnár (2018)	0.14
MK-5	Csomád volcano, Carpathians	Harangi et al. (2015)	0.03

<sup>1</sup>: 20 Ma considered as cooling age and used for  $\alpha$ -dose calulation (Heberer et al. 2017).

# IFORS\_tested\_parameters

Table 3: Table of (preliminary) tested values for IFORS automated curve–fitting procedure. For detailed description of the parameters see Lünsdorf and Lünsdorf (2016), for used curve–fitting protocols see Appendix 1.

Parameter	tested values	function of parameter
		Specifies the start wavenumber (Raman shift) of
start_wn	100	the spectrum
		Specifies the stop wavenumber (Raman shift) of
stop_wn	1300, 1800	the spectrum
		Sets the number repetitions / determines how often
repetitions	1, 3	the spectrum is repeatedly evaluated
		If set to 'yes', smoothing is enabled and the
		spectrum will be convoluted with a Hanning-
smoothing	yes, no	Window filter
<b>`</b>		Controls the window size of the Hanning-Window
smoothing size	9	used for smoothing
		Controls the minimum half width at half maximum
		(HWHM) of the pseudo-Voigt functions during
min width	1, 2	curve-fitting
	,	Controls the maximum half width at half maximum
		(HWHM) of the pseudo-Voiat functions during
max width	50, 80, 100, 150	curve-fitting
		Controls the minimum distance between two
min distance	2 5 8 10	pseudo-Voiat functions during curve-fitting
	_, 0, 0, 10	Style of baseline approximation. If set to
		'window filter' the baseline is approximated by a
		convolution of the baseline-data with a moving
		Hanning-window If set to 'off' no baseline will be
haseline mode	window filter off	used during curve-fitting
		Sets the size of the Hanning-window used during
		baseline approximation when baseline mode is set
window size	0 25 0 35	to 'window filter'
window_size	0.20, 0.00	This parameter is a multiple of the standard
noise intensity	1_3.5	deviation of the estimated noise distribution
noise_intensity	1-0.0	Sets the maximum number of pseudo-Voigt
		functions allowed during curve-fitting e.g. if this
max noak num	-10 15 20 30 35 4	D number is reached the program stops
	510, 15, 20, 50, 55, 40	During curve fitting, this parameter controls the
		width of the normal distribution from which the
		value is abasen that is added or subtracted from a
alaha	0.09	value is chosen that is added of subfracted from a
арпа	0.90	Controle the duration of the mutation loop mare
		directly logitical for former threshold and
		the interval (2) a signa threshold are
		irom the interval 2 > sigma_threshold > 0'. Values
		close to 2 result in short and values close to 0 in
sigma_threshold	0.01, 0.1, 0.5, 1	long mutation-loop duration





Figure 2



Figure3



Figure4



Figure5

