

1 **Revision 2**

2 **Can we use pyroxene weathering textures to interpret aqueous alteration conditions?**
3 **Yes and No.**

4 By: Charity M. Phillips-Lander^{1*}, Carey Legett IV¹, Andrew S. Elwood Madden¹, and
5 Megan E. Elwood Madden¹

6 ¹School of Geology and Geophysics, University of Oklahoma, 100 E. Boyd Street, Room
7 710, Norman, OK, 73071; *charity.m.lander@ou.edu

8 **Abstract**

9 Pyroxene minerals are a significant component of Shergottite-Nakhlite-Chassignite (SNC)
10 meteorites (e.g., Velbel, 2012) and detected across large areas of Mars' surface (e.g., Mustard,
11 2005). These minerals are associated with chloride, sulfate, and perchlorate salts that may
12 represent briny waters present in Mars' history. Previous textural analyses by Velbel and Losiak
13 (2010) comparing pyroxenes and amphiboles from a variety of natural weathering environments
14 showed no correlation between apparent apical angles (describing the morphology of denticular
15 weathering textures) and mineralogy or aqueous alteration history in relatively dilute solutions.
16 However, high salinity brines preferentially dissolve surface species, potentially leading to
17 different textures dependent on the brine chemistry. In this study, we performed controlled
18 pyroxene dissolution experiments in the laboratory on a well-characterized diopside to determine
19 if aqueous alteration in different high salinity brines, representative of potential weathering fluids
20 on Mars, produce unique textural signatures.

21 Following two months of dissolution in batch reactors, we observed denticles on etch pit
22 margins and pyroxene chip boundaries in all of the solutions investigated: ultrapure water (18
23 MΩ cm⁻¹; *a*H₂O=1); low-salinity solutions containing 0.35 M NaCl (*a*H₂O=0.99), 0.35 M
24 Na₂SO₄ (*a*H₂O=0.98) and 2 M NaClO₄ (*a*H₂O=0.9); and near-saturated brines containing 1.7 M

25 Na₂SO₄ (*a*H₂O=0.95), 3 M NaCl (*a*H₂O=0.75), and 4.5 M CaCl₂ (*a*H₂O=0.35). No systematic
26 change in denticle length or apical angle was observed between any of the solutions investigated,
27 even when altered in brines with significantly different salinity, activity of water, and anion
28 composition. Based on these and previous results from natural systems, apical angle
29 measurements are not a useful proxy for determining the extent or nature of aqueous alteration.
30 However, since dendrites form relatively slowly during weathering at circum-neutral pH, denticle
31 length may be a useful proxy for chemical weathering duration. All of the experimental solutions
32 produced median denticle lengths ≤ 1 μm , likely due to the brief weathering experiments.
33 However, perchlorate brines produced a significantly wider range of denticle lengths than those
34 observed in all the other experimental solutions tested. Since perchlorate is likely a common
35 constituent in martian soils (Glotch et al., 2016), denticle length measurements should be used
36 cautiously as proxies for extent of aqueous alteration on Mars, particularly in samples which also
37 contain perchlorate.

38 **Key Words**

39 weathering texture, pyroxene, dendrites, apical angles, Mars, perchlorate, brines

40 **Introduction**

41 Interpreting the history of aqueous alteration on Mars relies on chemical, mineralogical, and
42 textural clues recorded in rocks and minerals. Nano- to microscale textural features on primary
43 rock-forming minerals frequently contribute to interpretations of mineral-fluid interactions in
44 terrestrial samples (e.g., Hochella and Banfield, 1995). However, micro/nano-scale textural
45 analyses of minerals from Mars remain mostly limited to observations of SNC meteorites.
46 Pyroxene minerals are a significant component of SNC meteorites (McSween, 1994; Trieman,
47 2005; Velbel and Losiak, 2010; Velbel, 2012), indicative of their abundance in the martian crust

48 (Bandfield et al., 2000). They typically display indicators of weathering formed through aqueous
49 alteration on Mars (Trieman, 2005; Velbel, 2012; Ling and Wang, 2015), including the
50 formation of secondary clay minerals and Fe-oxides (Trieman et al., 1993; Ling and Wang,
51 2015) and dissolution textures on primary mineral phases. For example, jarosite, hematite,
52 cristobalite, and iddingsite are observed in veins in the Miller Range meteorite MIL 03346 (Ling
53 and Wang, 2015). In addition, denticles formed during pyroxene dissolution have been observed
54 in the Allen Hills meteorite ALH 84001 (Thomas-Keptra et al., 2009; Velbel, 2012). Pre-
55 terrestrial alteration in these meteorites has been established using cross-cutting relationships
56 including impact-related features overprinting alteration textures and minerals, as well as
57 weathering features predating the fusion crust (Gooding et al., 1988; Velbel, 2011).

58 In addition to alteration textures, salts are observed in SNC meteorites (Wentworth and
59 Gooding, 1994; Velbel, 2012). Perchlorate salts were observed in Mars meteorite EETA 9001
60 (Kounaves et al., 2014). Sulfates, including gypsum, anhydrite, and epsomite, as well as
61 chlorides, including halite, are common secondary minerals in SNC meteorites (Trieman, 1985;
62 Gooding et al., 1988; Chatzitheodoridis and Turner, 1990; Gooding et al., 1991; Trieman et al.,
63 1993; Wentworth and Gooding, 1994, 2005; Bridges and Grady, 1999, 2000; Bridges et al.,
64 2001; Lee et al., 2013; Ling and Wang, 2015). These salts occur as coatings on fracture surfaces
65 and as discontinuous veins in primary minerals, including pyroxene, in Chassigny (Wentworth
66 and Gooding, 1994) and MIL 03346 (Ling and Wang, 2015). Sulfate salts and jarosite also occur
67 along grain boundaries in MIL 03346 (Ling and Wang, 2015) and sulfate salts (gypsum,
68 basanite, epsomite) and halite are intergrown with smectite in Nakhla (Gooding et al.,
69 1991). Significant S and Cl are adsorbed to secondary clays and Fe-oxides in Lafayette (Trieman
70 et al., 1993). The presence of these salts associated with secondary minerals in SNC meteorites

71 provides evidence of episodic weathering in saline waters on or near Mars' surface (Trieman et
72 al., 1993; McSween, 1994).

73 During chemical weathering, surface textures form on silicate minerals as a result of mineral
74 dissolution (Berner et al., 1980; Velbel and Losiak, 2010). In most cases, the rate of silicate
75 dissolution is controlled by chemical reactions at the mineral/solution interface (Berner, 1978),
76 producing dissolution textures related to both mineral chemistry/structure and solution chemistry
77 (Hochella and Banfield, 1995). Laboratory experiments (e.g., Berner et al., 1980) and field
78 studies (e.g., Berner et al., 1980; Berner and Schott, 1982; Cremeens et al., 1992; Velbel and
79 Barker, 2008; Phillips-Lander et al., 2014) of pyroxene weathering demonstrated the progressive
80 development of distinctive dissolution textures, including lenticular etch pits lined with
81 "hacksaw" or "sawtoothed" terminations called denticles. These surface weathering textures can
82 be modified during transport and diagenesis; however, they may also be preserved through
83 weathering, transportation, deposition, and cementation (Stieglitz and Rothwell 1978; Velbel,
84 2007; Velbel et al. 2007; Velbel and Losiak, 2010). Preservation of surface weathering textures
85 may aid in the interpretation of aqueous alteration histories (Berner and Schott, 1982; Velbel,
86 2007; Andó et al. 2012).

87 Field-studies initially focused on etch pit sizes, morphologies, and distributions to understand
88 the factors that influence mineral dissolution (Berner et al., 1980; Berner and Schott, 1982;
89 Brantley et al., 1986; Cremeens et al., 1992; Brantley et al., 1993; MacInnis and Brantley, 1993;
90 White and Brantley, 2003). Berner et al. (1980) demonstrated surface dissolution textures are
91 primarily crystallographically controlled because textures observed from field studies were
92 replicated in laboratory weathering experiments conducted under a wide range of pH and
93 temperature conditions. However, individual etch pits observed in field-based studies are

94 generally larger than those observed in laboratory studies (Berner and Schott, 1982). Also,
95 denticles are commonly observed in field-weathered chain silicates (e.g., Berner et al., 1980;
96 Berner and Schott, 1982; Cremeens et al., 1992; Velbel and Barker, 2008; Phillips-Lander et al.,
97 2014) but were only observed in laboratory-based dissolution experiments at elevated
98 temperature (50-90 °C) and low pH (≤ 4) (Brantley and Chen, 1995; Chen and Brantley, 1998).
99 Denticles likely only form in those experiments where alteration was accelerated due to
100 temperature and pH effects since experiments are typically brief in duration (30 days). At neutral
101 pH and low temperatures experiments with durations of ≤ 30 days, the extent of weathering may
102 not have been sufficient to produce denticles due to slow pyroxene dissolution rates (Velbel,
103 2007). Therefore, previous short-term laboratory dissolution experiments suggest denticle
104 formation correlates with extent of chemical weathering.

105 The correlation between denticle length and extent of chemical weathering has been further
106 demonstrated in field studies examining amphiboles from a variety of glacial and periglacial
107 systems (e.g., Hall and Michaud, 1988; Argast, 1991; Hall and Horn, 1993; Mikesell et al., 2004;
108 Schatzel et al., 2006). Amphibole and pyroxene group minerals exhibit similar denticular
109 dissolution textures in response to similar weathering conditions (Velbel, 2007). Denticle lengths
110 decrease logarithmically with increasing depth in weathered soil profiles, which correlates with
111 sediment age and degree of soil development (Hall and Michaud, 1988; Argast, 1991; Hall and
112 Horn, 1993; Mikesell et al., 2004; Schatzel et al., 2006). However, these measurements only
113 provide a relative indicator of weathering, as sediment size, mean annual precipitation, reactive
114 fluid chemistry, and organic acids influence surface texture development in pyroxenes and
115 amphiboles (Hall and Horn, 1993; Schatzel et al., 2006). In addition to denticle lengths, the
116 presence of microdenticles, denticles ≤ 1 μm in length observed on the surfaces of larger

117 denticles, have been previously interpreted as forming in low temperature solutions
118 undersaturated with respect to pyroxene and amphibole (Velbel, 2011) or acidic solution
119 conditions (Brantley et al., 1993). Therefore, microdenticles on pyroxenes in ALH 84001 may
120 have formed during weathering in similar solutions (Velbel, 2012).

121 Based on these previous terrestrial studies, Velbel and Losiak (2010) examined whether SEM
122 analysis, including quantitative measurements of denticle apparent apical angle (i.e. the angle
123 subtended by a pair of denticle surfaces) could be used to determine the mineralogy of weathered
124 grains and/or the weathering conditions that produced the observed morphologies. They sought
125 to determine whether surface weathering textures on individual grains, like grain surface textures
126 observed using atomic force microscopy (AFM) at the Phoenix landing site (Hecht et al., 2009),
127 could be used to understand pyroxene weathering textures on another planetary body where both
128 the mineralogy and aqueous alteration history may be unknown. Their results indicated denticle
129 apical angles varied significantly, but apparently randomly, between mineral compositions (i.e.
130 augite v. diopside), weathering conditions, and even different silicates (amphibole v. pyroxene)
131 from the same terrestrially weathered rock unit, yielding no unique determining characteristics
132 based on quantitative analysis of weathering textures. However, this survey of weathering
133 textures was subject to the complications of studying natural pyroxenes from several localities
134 that were influenced by several variables (mineralogy, solution chemistry, contact time with
135 reactive solutions, temperature, saturation state, etc.) that may have affected denticle formation.

136 In order to reduce potential variability in textural weathering observations and measurements
137 and extend the range of textural observations to brines, we conducted a series of experiments in
138 which we reacted a diopside of known composition with different Mars-representative
139 weathering solutions. A range of pyroxene compositions are common on the surface of Mars,

140 including diopside (e.g., Mustard, 2005). Previous studies of mineral weathering in Mars-
141 representative solutions have demonstrated solution chemistry (Elwood Madden et al., 2012;
142 Phillips-Lander et al., 2016; Parnell et al., 2016) and activity of water (Pritchett et al., 2012;
143 Legett et al., 2014; Olson et al., 2015; Miller et al., 2016; Parnell et al., 2016; Steiner et al.,
144 2016) influence mineral dissolution rates. We sought to determine if pyroxene dissolution in
145 brine solutions with different activities of water and solution chemistries would result in
146 quantifiable differences in denticle length and/or apparent apical angles. In other words, we seek
147 to determine whether diopside weathering textures vary in a systematic way during low-
148 temperature aqueous alteration. The results of this study will inform interpretations of pyroxene
149 weathering textures in meteorites and future studies of samples returned from Mars.

150 **Methods**

151 **Pyroxene Characterization**

152 We obtained a sample of diopside ($\text{Ca}_{0.96}\text{Na}_{0.04}\text{Fe}_{0.32}\text{Mg}_{0.65}\text{Al}_{0.04}\text{Si}_{1.9}\text{O}_6$), a Ca-rich
153 pyroxene, from Wards Scientific (#466474). Bulk geochemistry was determined using energy-
154 dispersive X-ray spectroscopy (EDXA) in a Cameca SX-50 electron microprobe (EMPA)
155 operated at 20 kV and 10 nA, analyzing 160 x 160 μm areas (Supplemental Table 1).

156 **Dissolution Experiments and Textural Analysis**

157 After EMPA analysis, we coarsely crushed the diopside to ~2-4 mm chips. Resulting mineral
158 chips were sonicated, rinsed, and air-dried. We conducted batch dissolution experiments at 22 °C
159 and 1 atm using diopside chips added to solution at a ratio of 1 g pyroxene L^{-1} solution. Pyroxene
160 was reacted with solutions with varying chemistry and salinity: 18 $\text{M}\Omega\text{ cm}^{-1}$ ultrapure water
161 (UPW) that has an activity of water ($a_{\text{H}_2\text{O}}=1$); low-salinity solutions containing 0.35 M NaCl
162 ($a_{\text{H}_2\text{O}}=0.99$), 0.35 M Na_2SO_4 ($a_{\text{H}_2\text{O}}=0.98$), and 2 M NaClO_4 ($a_{\text{H}_2\text{O}}=0.9$); and near-saturated

163 brines containing 1.7 M Na₂SO₄ (*a*H₂O=0.95), 3 M NaCl (*a*H₂O=0.75), and 4.5 M CaCl₂
164 (*a*H₂O=0.35). We mixed the samples at 120 rpm on a shaker table over 60 days. These solutions
165 were chosen to examine whether denticle length or apparent apical angle vary systematically
166 with either changes in solution chemistry between dilute (0.35 M) and near-saturated (1.7 M
167 Na₂SO₄ or 3 M NaCl) brines or decreasing activity of water.

168 At the end of the experiment, we filtered the contents of the reactor to retrieve the diopside
169 chips, then rinsed the reacted diopside chips three times in 18 MΩ cm⁻¹ water and air-dried the
170 samples. Chips from each experimental condition, including unreacted controls, were then placed
171 on Al pin stubs in random orientations with carbon tape and iridium coated for imaging analysis
172 on the Zeiss NEON 40EsB field-emission scanning electron microscope (SEM) at the University
173 of Oklahoma's Sam Noble Microscopy Laboratory. Images were collected at 15 kV and ~50 pA
174 beam current with a vacuum of ~4x10⁻⁶ torr using both in-lens and Everhart-Thornley secondary
175 electron detectors.

176 We took SEM images at scales ranging from 100 nm to 10 μm in order to evaluate mineral
177 weathering textures, including microdenticles. SEM images were imported into ImageJ
178 (Schneider et al., 2012), an open source imaging software program, which has the capability to
179 measure both denticle length and apparent denticle apical angles formed during dissolution. We
180 measured denticle lengths from the initial parting on the mineral surface to the tip of the denticle
181 (Figure 1A). Denticle lengths may be considered minimum values, due to the possibility that in
182 some cases the base of a denticle is obscured by an overlapping denticle. We estimated our
183 measurement precision based on measuring the same denticle three times to be ±0.13 μm. We
184 also measured apparent apical angles using the angle measurement tool, which drops anchor
185 points on each side of the denticle and at the tip. The precision of our apparent apical angle

186 measurements (1.3°) was determined by measuring the same apical angle three times for a single
187 denticle. The number of images with etch pits and denticles was controlled by the degree of
188 dissolution of the mineral surface. However, we aimed to quantify denticles from a minimum of
189 five images per solution condition, including all visually distinguishable denticles in each image.
190 Statistical analysis of denticle measurements was performed by constructing box and whisker
191 plots. The number of denticle measurements and the boundary of each interquartile range is
192 presented in Table 1. We compared denticle lengths and apparent apical angles observed on
193 diopside surfaces in order to determine how these data vary with changes in fluid chemistry and
194 activity of water, which have been previously demonstrated to influence mineral dissolution
195 rates.

196 Results

197 Denticles are present along etch pit margins on diopside surfaces weathered under all
198 experimental conditions, with the exception of unreacted grains (Figure 1). Median denticle
199 lengths are approximately $\leq 1 \mu\text{m}$; however, denticle length did not vary systematically with
200 decreasing activity of water (Table 1; Figure 2A). Denticle lengths in UPW ($a_{\text{H}_2\text{O}}=1$), 1.7 M
201 Na_2SO_4 ($a_{\text{H}_2\text{O}}=0.95$), and 2 M NaClO_4 ($a_{\text{H}_2\text{O}}=0.93$) are all very similar, with median lengths of
202 0.67-1.0 μm . However, the range (0.18-5.93 μm) of denticle lengths in 2 M NaClO_4 is
203 significantly larger than the range of denticle lengths observed in all other experimental samples
204 (~ 0.1 -2.94 μm ; Table 1; Figure 2A). Slightly shorter median denticle lengths were observed in
205 0.35 M NaCl ($a_{\text{H}_2\text{O}}=0.99$) and 4.5 M CaCl_2 ($a_{\text{H}_2\text{O}}=0.35$) with denticle lengths of 0.50 and 0.59
206 μm respectively. The shortest denticles were observed in 0.35 M Na_2SO_4 ($a_{\text{H}_2\text{O}}=0.98$) and 3 M
207 NaCl ($a_{\text{H}_2\text{O}}=0.75$) with median lengths of 0.14 and 0.36 μm respectively. Microdenticles

208 (Velbel, 2011) are observed in Figure 1 in (C) 0.35 M Na₂SO₄, (D) 2 M NaClO₄, and (E) 3 M
209 NaCl.

210 Median apparent apical angle measurements in UPW, 0.35 M Na₂SO₄, 2 M NaClO₄, and 4.5
211 M CaCl₂ range from 18.1 to 23.5° (Table 1; Figure 2B). However, apparent apical angles were
212 significantly smaller in 0.35 M NaCl (9.5°) and 1.7 M Na₂SO₄ (11.1°). The range of apical angles
213 (up to 40°) overlaps for experimental solutions with activities of water >0.9 (Figure 2B).
214 Apparent apical angles are larger (42.4°) in 3 M NaCl. The range of apical angles for diopside
215 chips weathered in near-saturated NaCl and CaCl₂ brines are greater than the range for more
216 dilute solutions, if outlier data points are not considered (Figure 2B).

217 Discussion

218 Activity of water and weathering textures

219 Velbel and Losiak (2010) surveyed denticle apical angles of both amphiboles and pyroxenes
220 with varying weathering histories to determine whether the mineralogy and alteration history of
221 an unknown grain, like one imaged on Mars' surface by the Phoenix lander (Hecht et al., 2009),
222 could be obtained. Their study demonstrated apparent apical angles were not diagnostic of grain
223 composition (Velbel and Losiak, 2010). Our research extends Velbel and Losiak's (2010) work
224 and demonstrates apparent apical angle also does not change systematically with activity of
225 water. Median apparent apical angles, with the exception of those measured on diopside grains
226 reacted with 3 M NaCl, overlap and have similar average values (Figure 2B).

227 Measured median denticle lengths in our study are also similar, with <1 μm difference in
228 average values between experimental trials (Figure 2A), indicating activity of water also does not
229 systematically influence denticle lengths. For example, our 0.35 M Na₂SO₄ (*a*H₂O=0.98) and 3
230 M NaCl (*a*H₂O=0.75) had the smallest median denticle lengths, but vary significantly in activity

231 of water (Figure 2A). In addition, while median denticle length decreased slightly with
232 decreasing activity of water in NaCl brines, the opposite trend was observed in NaSO₄ brines.
233 Therefore, our data indicate even when the parent material is well-constrained, neither denticle
234 length nor apparent apical angle are indicative of the activity of water of the weathering solution,
235 as similar denticle lengths and apparent apical angles can be formed under very different aqueous
236 alteration conditions.

237 **Anion chemistry and weathering textures**

238 In order to further clarify and distinguish the potential effects of different anions on denticle
239 formation, length, and apparent apical angles, we compared results of sodium chloride and
240 sodium sulfate experiments at different salt concentrations: low salinity (0.35 M) and near-
241 saturated brines (3 M NaCl and 1.7 M Na₂SO₄). We observed clear differences in the trends
242 between these two sets of solutions with different anion chemistries; however, overall quartile
243 ranges for the quantitative measurements overlap. In NaCl brines, median denticle lengths
244 decrease slightly (from 0.50 to 0.36 μm) and apical angles increase significantly (9.5 to 42.4°)
245 with increasing chloride salt concentration. However, the opposite trend is observed with
246 increasing concentration in Na₂SO₄ brines, where median denticle lengths increase significantly
247 (from 0.14 to 0.67 μm) and apical angles decrease slightly (from 23.5 to 11.1°). The contrasting
248 trends in measurements between the two solution sets suggests the dissolution mechanism may
249 be different between sodium chloride and sodium sulfate solutions at circum-neutral pH and/or
250 there is no clear trend in denticle length and apparent apical angle measurements with changing
251 anion chemistry. Given the non-systematic relationship between solution chemistry and denticle
252 length, as well as the general overlap in measurements between all the experiments, neither
253 apical angle nor denticle length are useful indicators of anion chemistry, particularly in more

254 complicated natural solutions that likely contain a mixture of different ions.

255 While perchlorate has been typically considered an “indifferent” electrolyte for investigations
256 of mineral reactivity (e.g., Brown and Parks, 2001), we observed a much larger range of denticle
257 lengths in 2 M NaClO₄ brine compared to our NaCl and Na₂SO₄ experiments. Sanemasa and
258 Katsura (1973) stated that solutions containing up to 2 N HClO₄ and HCl had little effect on
259 synthetic diopside dissolution rates and extent of cation release compared to H₂SO₄; however,
260 they presented no data supporting their findings. In contrast, Sidhu et al. (1981) demonstrated a
261 range of perchlorate concentrations up to 2 M had no influence on the dissolution rate of iron
262 oxide minerals, while Cl⁻ ions produced faster dissolution rates. Our results show a much wider
263 range in denticle lengths produced in the perchlorate dissolution experiments compared to all the
264 other solution chemistries. This suggests perchlorate ions may interact differently with the
265 diopside surface, producing a wider variety of denticle lengths. We hypothesize that this may be
266 a result of redox reactions occurring at specific surface sites driven by sorption of highly
267 oxidized perchlorate anions. Although redox reactions between perchlorate and ferrous iron-
268 bearing minerals have not yet, to our knowledge, been demonstrated, it is well-established that
269 redox reactions influence the dissolution of iron-bearing pyroxene (e.g., Siever and Woodford,
270 1979; Schott and Berner, 1983; White et al., 1985; Burns, 1993; Hoch et al. 1996). These redox-
271 driven dissolution reactions may accelerate chemical weathering in specific locations on the
272 diopside surface leading to formation of much longer dendrites in some areas of the sample. The
273 patchy development of protective iron (hydr)oxide and clay surface layers may influence
274 diffusion of reactants, thus leading to variability in dissolution textures across the mineral surface
275 (e.g., Siever and Woodford, 1979; Benzerara et al., 2005; Behrends et al., 2015). However, other
276 studies suggest that protective surface layers may not ultimately control the overall dissolution

277 rate observed in natural weathering environments, depending on the ratio of molar volumes
278 between dissolving phases and secondary reaction products (Velbel, 1993) and the porosity of
279 the precipitates (e.g., Hodson, 2003; Lee et al., 2008). Therefore, individual measurements of
280 denticle lengths on weathered pyroxene minerals surfaces in systems exposed to perchlorate may
281 over-estimate the extent of aqueous alteration. Since perchlorate is commonly observed in Mars
282 surface samples (0.4-0.6 wt.%; Hecht et al., 2009; Navarro-Gonzalez et al., 2010; Glavin et al.,
283 2013), broader assessments of pyroxene weathering in the presence of perchlorate are needed to
284 provide context for any future denticle length measurements of pyroxene on/from Mars.

285 **Denticle length as a quantitative assessment of weathering extent**

286 Early laboratory studies conducted at either low pH (<4) and/or high temperature (50-90 °C)
287 formed etch pits and denticle features similar to those observed in natural weathered soil profiles
288 (Brantley and Chen, 1995; Chen and Brantley, 1998). In contrast, 30-day dissolution experiments
289 at circumneutral pH and 22 °C did not result in denticle formation (Werner et al., 1995), perhaps
290 due to the relatively short reaction times and slow reaction rates. However, etch pits and
291 denticles similar to those typically observed in natural systems formed in all of our 60 day
292 experiments, suggesting the duration of aqueous alteration may be the major control on denticle
293 formation. We observed microdenticles that have been previously linked to reactive solutions
294 that are acidic and/or undersaturated with respect to pyroxene (Velbel, 2011; Velbel, 2012) in
295 several of our experiments; however, we find no systematic variation with activity of water.

296 Previous studies have correlated pyroxene denticle length and extent of chemical weathering
297 in soil profiles, where denticle lengths decrease with increasing distance below the depth of first
298 chain silicate occurrence, reaching a minimum either where chemical weathering conditions
299 change (Argast, 1991) and/or as fresh bedrock surfaces are approached (Mikesell et al., 2004;

300 Schaetzl et al., 2006). For example, in amphibole-bearing soils developed on glacially-influenced
301 sediments, denticle lengths $\geq 2 \mu\text{m}$ were common near the surface, but shorter denticles ($< 1 \mu\text{m}$)
302 were observed at greater depths approaching incipiently altered bedrock (Hall and Michaud,
303 1988; Hall and Horn, 1993; Mikesell et al., 2004; Schaetzl et al., 2006). Therefore, quantitative
304 measures of median denticle lengths may be useful in constraining the duration of chemical
305 weathering, with denticles $\leq 1 \mu\text{m}$ indicating relatively brief periods of aqueous alteration.
306 Indeed, the relatively rapid formation of denticles observed in this study suggests even
307 geologically brief aqueous alteration events can result in significant changes to pyroxene surface
308 textures. In contrast, longer denticles would be expected to be observed on pyroxene and
309 amphiboles that experienced longer weathering durations.

310 **Implications**

311 Our results demonstrate quantitative measurements of apical angles and denticle lengths do
312 not correlate with changing solution chemistry or activity of water in chloride, perchlorate, and
313 sulfate brines. Therefore, it is unlikely quantitative denticle measurements can be used to
314 constrain past aqueous chemistry on Mars. As this supports Velbel and Losiak's (2010) findings
315 in dilute solutions, we recommend abandoning quantitative measures of apparent apical angles as
316 a tool to determine aqueous alteration histories of pyroxene grains.

317 In contrast, quantitative comparisons of denticle lengths may provide insight into the
318 duration of aqueous alteration. Small scale ($< 1 \mu\text{m}$) denticles are commonly observed in martian
319 meteorites, including ALH84001 (Thomas-Keptra et al., 2009; Velbel, 2011) and MIL 03346
320 (Velbel, 2012). These meteorites also contain veins with sulfate and chloride minerals, which
321 may indicate martian meteorites experienced aqueous alteration in saline conditions, similar to
322 those examined in this study, over relatively brief periods of geologic time.

323 However, perchlorate salts may complicate simple applications of denticle length
324 measurements to interpret extent of weathering since a much wider range of measured denticle
325 lengths (but similar average value $\leq 1 \mu\text{m}$) were produced in perchlorate brines compared to the
326 other solutions. Perchlorate brines may have been stable on the surface of Mars throughout its
327 history, and may be active on Mars' surface today (Chevrier et al., 2009; Ojha et al., 2015).
328 Perchlorate salts have been detected in martian meteorite EETA 79001 at concentrations of
329 0.6 ± 0.1 ppm (Kounaves et al., 2014). Perchlorate was also detected by the Phoenix lander at 0.5-
330 0.7 wt.% (Hecht et al., 2009), suggesting it may be a common constituent in martian sediments.
331 Therefore, care must be taken when using quantitative denticle lengths as proxies for the extent
332 of aqueous alteration on Mars, including in returned samples, in order to ensure accurate
333 interpretations of chemical weathering duration.

334 Laboratory studies using well-mixed reactors typically represent the 'maximum' weathering
335 intensity possible; the ultimate lifetime and persistence of mineral grains in natural systems may
336 be much longer than that predicted from short-term weathering rates (Velbel, 2007; Olsen and
337 Rimstidt, 2007; Pritchett et al., 2012). For example, the relatively pristine clinopyroxene
338 textures compared with more extensive olivine alteration observed in nakhlites may be due to
339 fluid saturation with respect to pyroxene but remaining undersaturated with respect to olivine
340 (Velbel, 2016). On the other hand, quantitative weathering rates from olivine laboratory
341 experiments accurately predicted the time required to form etch pits in different laboratory
342 experiments (Velbel, 2014), indicating that laboratory experiments do constrain the timing of
343 mineral surface textural evolution.

344 Given these caveats, we recommend using qualitative comparisons of denticle lengths as
345 proxies for relative extent of weathering on Mars only in areas where the mineralogy, including

346 salts associated with pyroxene grains, are well-constrained. We further recommend incorporating
347 additional data, including qualitative weathering assessments that examine etch pit density, and
348 secondary mineral coverage, including clays formed on the mineral surface (Phillips-Lander et
349 al., 2017), to increase confidence in interpretations of weathering duration on Mars.

350 **Acknowledgements**

351 This study was supported by a NASA Mars Fundamental Research Grant #NNX13AG75G.
352 We gratefully acknowledge the assistance of Preston Larson of OU's Sam Noble Microscopy
353 Laboratory and George Morgan of OU's Electron Microprobe Laboratory for their analytical
354 assistance. We are grateful for three reviews, including those of Prof. Velbel and Prof. Andò that
355 led to substantial improvements.

356 **References**

- 357 Andò, S., Garzanti, E., Padoan, M., and Limonta, M. (2012) Corrosion of heavy minerals
358 during weathering and diagenesis: A catalog for optical analysis. *Sedimentary Geology*, 280,
359 165–178.
- 360 Argast, S. (1991) Chlorite vermiculitization and pyroxene etching in an aeolian periglacial
361 sand dune, Allen County, Indiana. *Clays and Clay Minerals*, 39(6), 622-633.
- 362 Bandfield, J.L., Hamilton, V.E., and Christensen, P.R. (2000) A global view of Martian
363 surface compositions from MGS-TES. *Science*, 287, 1626-1630.
- 364 Behrens, R., Bouchez, J., Schuessler, J.A., Dultz, S., Hewawasam, T., and von
365 Blanckenburg, F. (2015) Mineralogical transformations set slow weathering rates in low-porosity
366 metamorphic bedrock on mountain slopes in a tropical climate. *Chemical Geology*, 411, 283–
367 298.
- 368 Benzerara, K., Yoon, T., Menguy, N., Tyliszczak, T., and Brown, G. (2005) Nanoscale
369 environments associated with bioweathering of a Mg-Fe-pyroxene. *Proceedings of the National*

- 370 Academy of Sciences of the United States of America, 102, 979–982.
- 371 Berner, R.A. (1978) Rate control of mineral dissolution under earth surface conditions.
372 American Journal of Science, 278(9), 1235-1252.
- 373 Berner, R.A. and Schott, J. (1982) Mechanism of pyroxene and amphibole weathering; II,
374 Observations of soil grains. American Journal of Science, 282(8), 1214-1231.
- 375 Berner, R.A., Sjöberg, E.L., Velbel, M.A. and Krom, M.D. (1980) Dissolution of pyroxenes
376 and amphiboles during weathering. Science, 207(4436), 1205-1206.
- 377 Brantley, S.L., Blai, A.C., Cremeens, D.L., MacInnis, I. and Darmody, R.G. (1993) Natural
378 etching rates of feldspar and hornblende. Aquatic Sciences, 55(4), 262-272.
- 379 Brantley, S.L. and Chen, Y. (1995) Chemical weathering rates of pyroxenes and
380 amphiboles. Reviews in Mineralogy and Geochemistry, 31(1), 119-172.
- 381 Brantley, S.L., Crane, S.R., Crerar, D.A., Hellmann, R. and Stallard, R. (1986) Dissolution at
382 dislocation etch pits in quartz. Geochimica et Cosmochimica Acta, 50(10), 2349-2361.
- 383 Bridges, J.C., Catling, D.C., Saxton, J.M., Swindle, T.D., Lyon, I.C. and Grady, M.M. (2001)
384 Alteration assemblages in Martian meteorites: Implications for near-surface processes. In R.
385 Kallenbach, J. Geiss, and W.K. Hartmann (Eds.) Chronology and evolution of Mars, p. 365-392.
386 Springer, Netherlands.
- 387 Bridges, J.C. and Grady, M.M. (2000) Evaporite mineral assemblages in the nakhlite
388 (martian) meteorites. Earth and Planetary Science Letters, 176(3), 267-279.
- 389 Bridges, J.C. and Grady, M.M. (1999) A halite-siderite-anhydrite-chlorapatite assemblage in
390 Nakhla: Mineralogical evidence for evaporites on Mars. Meteoritics and Planetary Science, 34,
391 407-415.
- 392 Brown Jr., G.E. and Parks, G.A. (2001) Sorption of trace elements on mineral surfaces:

393 modern perspectives from spectroscopic studies, and comments on sorption in the marine
394 environment. *International Geology Review*, 43(11), 963-1073.

395 Burns, R.G. (1993) Rates and mechanisms of chemical weathering of ferromagnesian silicate
396 minerals on Mars. *Geochimica et Cosmochimica Acta*, 57, 4555–4574.

397 Chatzitheodoridis, E. and Turner, G. (1990) Secondary minerals in the Nakhla meteorite.
398 *Meteoritics*, 25, 354-354.

399 Chen, Y. and Brantley, S.L. (1998) Diopside and anthophyllite dissolution at 25 and 90 C
400 and acid pH. *Chemical Geology*, 147(3), 233-248.

401 Chevrier, V.F., Hanley, J. and Altheide, T.S. (2009) Stability of perchlorate hydrates and
402 their liquid solutions at the Phoenix landing site, Mars. *Geophysical Research Letters*, 36(10),
403 L10202.

404 Cremeens, D.L., Darmody, R.G. and Norton, L.D. (1992) Etch-pit size and shape distribution
405 on orthoclase and pyriboles in a loess catena. *Geochimica et Cosmochimica Acta*, 56(9), 3423-
406 3434.

407 Elwood Madden, M.E., Madden, A.S., Rimstidt, J.D., Zahrai, S., Kendall, M.R. and Miller,
408 M.A. (2012) Jarosite dissolution rates and nanoscale mineralogy. *Geochimica et Cosmochimica*
409 *Acta*, 91, 306-321.

410 Glavin, D.P., Freissinet, C., Miller, K.E., Eigenbrode, J.L., Brunner, A.E., Buch, A., Sutter,
411 B., Archer, P.D., Atreya, S.K., Brinckerhoff, W.B. and Cabane, M. (2013) Evidence for
412 perchlorates and the origin of chlorinated hydrocarbons detected by SAM at the Rocknest
413 aeolian deposit in Gale Crater. *Journal of Geophysical Research: Planets*, 118(10), 1955-1973.

414 Glotch, T., Bandfield, J., Wolff, M., Arnold, J., and Che, C. (2016) Constraints on the
415 composition and particle size of chloride salt-bearing deposits on Mars. *Journal of Geophysical*

416 Research: Planets, 121, 454–471.

417 Gooding, J.L., Wentworth, S.J. and Zolensky, M.E. (1991) Aqueous alteration of the Nakhla
418 meteorite. *Meteoritics and Planetary Science*, 26(2), 135-143.

419 Gooding, J.L., Wentworth, S.J. and Zolensky, M.E. (1988) Calcium carbonate and sulfate of
420 possible extraterrestrial origin in the EETA 79001 meteorite. *Geochimica et Cosmochimica*
421 *Acta*, 52(4), 909-915.

422 Hall, R.D. and Horn, L.L. (1993) Rates of hornblende etching in soils in glacial deposits of
423 the northern Rocky Mountains (Wyoming-Montana, USA): Influence of climate and
424 characteristics of the parent material. *Chemical Geology*, 105(1-3), 17-29.

425 Hall, R.D. and Michaud, D. (1988) The use of hornblende etching, clast weathering, and soils
426 to date alpine glacial and periglacial deposits: a study from southwestern Montana. *Geological*
427 *Society of America Bulletin*, 100(3), 458-467.

428 Hecht, M.H., Kounaves, S.P., Quinn, R.C., West, S.J., Young, S.M.M., Ming, D.W., Catling,
429 D.C., Clark, B.C., Boynton, W.V., Hoffman, J. and DeFlores, L.P. (2009) Detection of
430 perchlorate and the soluble chemistry of martian soil at the Phoenix lander site. *Science*,
431 325(5936), 64-67.

432 Hoch, A.R., Reddy, M.M., and Drever, J.I. (1996) The effect of iron content and dissolved
433 O₂ on dissolution rates of clinopyroxene at pH 5.8 and 25°C: preliminary results. *Chemical*
434 *Geology*, 132, 151–156.

435 Hochella, M.F. and Banfield, J.F. (1995) Chemical weathering of silicates in nature; a
436 microscopic perspective with theoretical considerations. *Reviews in Mineralogy and*
437 *Geochemistry*, 31(1), 353-406.

438 Hodson, M.E. (2003) The influence of Fe-rich coatings on the dissolution of anorthite at pH

439 2.6. *Geochimica et Cosmochimica Acta*, 67, 3355-3363.

440 Kounaves, S.P., Carrier, B.L., O'Neil, G.D., Stroble, S.T. and Claire, M.W. (2014) Evidence
441 of martian perchlorate, chlorate, and nitrate in Mars meteorite EETA79001: Implications for
442 oxidants and organics. *Icarus*, 229, 206-213.

443 Lee, M.R., Brown, D.J., Hodson, M.E., MacKenzie, M., and Smith, C.L. (2008) Weathering
444 microenvironments on feldspar surfaces: implications for understanding fluid-mineral reactions
445 in soils. *Mineralogical Magazine*, 72, 1319–1328.

446 Lee, M.R., Tomkinson, T., Mark, D.F., Stuart, F.M. and Smith, C.L. (2013) Evidence for
447 silicate dissolution on Mars from the Nakhla meteorite. *Meteoritics and Planetary Science*, 48(2),
448 224-240.

449 Legett, C., Pritchett, B.N., Elwood Madden, A.S. and Elwood Madden, M.E. (2014)
450 Measuring Mineral Dissolution Rates in Perchlorate Brines: Method Development and
451 Applications. *Lunar and Planetary Sciences Conference*, p.2492.

452 Ling, Z. and Wang, A. (2015) Spatial distributions of secondary minerals in the Martian
453 meteorite MIL 03346, 168 determined by Raman spectroscopic imaging. *Journal of Geophysical
454 Research: Planets*, 120(6), 1141-1159.

455 MacInnis, I.N. and Brantley, S.L. (1993) Development of etch pit size distributions on
456 dissolving minerals. *Chemical Geology*, 105(1-3), 31-49.

457 McSween, H.Y. (1994) What we have learned about Mars from SNC meteorites. *Meteoritics*,
458 29(6), 757-779.

459 Mikesell, L.R., Schaetzl, R.J. and Velbel, M.A. (2004) Hornblende etching and
460 quartz/feldspar ratios as weathering and soil development indicators in some Michigan
461 soils. *Quaternary Research*, 62(2), 162-171.

- 462 Miller, J.L., Madden, A.E., Phillips-Lander, C.M., Pritchett, B.N. and Madden, M.E. (2016)
463 Alunite dissolution rates: Dissolution mechanisms and implications for Mars. *Geochimica et*
464 *Cosmochimica Acta*, 172, 93-106.
- 465 Mustard, J. F., Poulet, F., Gendrin, A., Bibring, J.-P., Langevin, Y., Gondet, B., Mangold, N.,
466 Bellucci, G., and Altieri, F. (2005) Olivine and Pyroxene Diversity in the Crust of Mars. *Science*,
467 307, 1594–1597.
- 468 Navarro-González, R., Vargas, E., de La Rosa, J., Raga, A.C. and McKay, C.P. (2010)
469 Reanalysis of the Viking results suggests perchlorate and organics at midlatitudes on Mars. *JGR:*
470 *Planets*, 115(E12).
- 471 Olsen, A.A. and Rimstidt, J.D. (2008) Oxalate-promoted forsterite dissolution at low
472 pH. *Geochimica et Cosmochimica Acta*, 72(7), 1758-1766.
- 473 Ojha, L., Wilhelm, M.B., Murchie, S.L., McEwen, A.S., Wray, J.J., Hanley, J., Massé, M.
474 and Chojnacki, M. (2015) Spectral evidence for hydrated salts in recurring slope lineae on Mars.
475 *Nature Geoscience*, 8(11), 829-832.
- 476 Parnell, S.P., Phillips-Lander, C.M., McGraw, L.E., and Elwood Madden, M.E. (2016)
477 Carbonate dissolution rates in high salinity brines, *Lunar and Planetary Sciences*, p. 1460.
- 478 Phillips-Lander, C.M., Legett IV, C. Elwood Madden, A.S., and Elwood Madden, M.E.
479 (2017) Can we use pyroxene weathering textures to interpret aqueous alteration conditions? Yes
480 and No. *Lunar and Planetary Sciences*, p.1109.
- 481 Phillips-Lander, C.M., Legett IV, C. Elwood Madden, A.S., and Elwood Madden, M.E.
482 (2016) Pyroxene dissolution rates in high salinity brines: Implications for post-Noachian aqueous
483 alteration on Mars. *Lunar and Planetary Sciences*, p.1313.
- 484 Phillips-Lander, C.M., Fowle, D.A., Taunton, A., Hernandez, W., Mora, M., Moore, D.,

485 Shinogle, H. and Roberts, J.A. (2014) Silicate dissolution in Las Pailas thermal field:
486 Implications for microbial weathering in acidic volcanic hydrothermal spring
487 systems. *Geomicrobiology Journal*, 31(1), 23-41.

488 Pritchett, B.N., Madden, M.E. and Madden, A.S. (2012) Jarosite dissolution rates and
489 maximum lifetimes in high salinity brines: Implications for Earth and Mars. *Earth and Planetary
490 Science Letters*, 357, 327-336.

491 Sanemasa, I. and Katsura, T. (1973) The dissolution of $\text{CaMg}(\text{SiO}_3)_2$ in acid
492 solutions. *Bulletin of the Chemical Society of Japan*, 46(11), 3416-3422.

493 Schaetzl, R.J., Mikesell, L.R. and Velbel, M.A. (2006) Soil characteristics related to
494 weathering and pedogenesis across a geomorphic surface of uniform age in Michigan. *Physical
495 Geography*, 27(2), 170-188.

496 Schneider, C.A., Rasband, W.S. and Eliceiri, K.W. (2012) NIH Image to ImageJ: 25 years of
497 image analysis. *Nature Methods*, 9, 671-675.

498 Schott, J., and Berner, R. (1983) X-ray photoelectron studies of the mechanism of iron
499 silicate dissolution during weathering. *Geochimica et Cosmochimica Acta*, 47, 2233–2240.

500 Sidhu, P.S., Gilkes, R.J., Cornell, R.M., Posner, A.M., and Quirk, J.P. (1981) Dissolution of
501 iron oxides and oxyhydroxides in hydrochloric and perchloric acids. *Clays and Clay Minerals*,
502 29, 269–276.

503 Siever, R. and Woodford, N. (1979) Dissolution kinetics and the weathering of mafic
504 minerals. *Geochimica et Cosmochimica Acta*, 43, 717–724.

505 Stieglitz, R.D. and Rothwell, B. (1978) Surface microtextures of freshwater heavy mineral
506 grains. *Geoscience Wisconsin*, 3, 21-34.

507 Steiner, M.H., Hausrath, E.M., Madden, M.E., Tschauer, O., Ehlmann, B.L., Olsen, A.A.,

508 Gainey, S.R. and Smith, J.S. (2016) Dissolution of nontronite in chloride brines and implications
509 for the aqueous history of Mars. *Geochimica et Cosmochimica Acta*, 195, 259-276.

510 Thomas-Keptra K.L., Clemett S.J., McKay D.S., Gibson E.K., and Wentworth S.J. (2009)
511 Origins of magnetite nanocrystals in Martian meteorite ALH84001. *Geochimica et*
512 *Cosmochimica Acta* 73, 6631–6677.

513 Treiman, A.H. (2005) The nakhlite meteorites: Augite-rich igneous rocks from Mars. *Chemie*
514 *der Erde-Geochemistry*, 65(3), 203-270.

515 Treiman, A.H., Barrett, R.A. and Gooding, J.L. (1993) Preterrestrial aqueous alteration of the
516 Lafayette (SNC) meteorite. *Meteoritics*, 28(1), 86-97.

517 Velbel, M.A. (2007) Surface textures and dissolution processes of heavy minerals in the
518 sedimentary cycle: examples from pyroxenes and amphiboles. *Developments in Sedimentology*,
519 58, 113-150.

520 Velbel, M.A. (2011) Microdenticles on naturally weathered hornblende. *Applied*
521 *Geochemistry*, 26(8), 1594-1596.

522 Velbel, M.A. (2012) Aqueous alteration in Martian meteorites: Comparing mineral relations
523 in igneous-rock weathering of Martian meteorites and in the sedimentary cycle of
524 Mars. *Sedimentary Geology of Mars*, SEPM. Society for Sedimentary Geology Special
525 Publication, 102, 97-117.

526 Velbel, M.A. and Barker, W.W. (2008) Pyroxene weathering to smectite: conventional and
527 cryo-field emission scanning electron microscopy, Koua Bocca ultramafic complex, Ivory Coast.
528 *Clays and Clay Minerals*, 56(1), 112-127.

529 Velbel, M.A. and Losiak, A.I. (2010) Denticles on chain silicate grain surfaces and their
530 utility as indicators of weathering conditions on Earth and Mars. *Journal of Sedimentary*

531 Research, 80(9), 771-780.

532 Wentworth, S.J. and Gooding, J.L. (1994) Carbonates and sulfates in the Chassigny
533 meteorite: Further evidence for aqueous chemistry on the SNC parent planet. *Meteoritics*, 29(6),
534 860-863.

535 Werner, A.J., Hochella, M.F., Guthrie, G.D., Hardy, J.A. and Aust, A.E. (1995) Asbestiform
536 riebeckite (crocidolite) dissolution in the presence of Fe chelators: implications for mineral-
537 induced disease. *American Mineralogist*, 80(11-12), 1093-1103.

538 White, A.F. and Brantley, S.L. (2003) The effect of time on the weathering of silicate
539 minerals: why do weathering rates differ in the laboratory and field? *Chemical Geology*, 202(3),
540 479-506.

541 White, A., Yee, A., and Flexser, S. (1985) Surface oxidation-reduction kinetics associated
542 with experimental basalt-water reaction at 25°C. *Chemical Geology*, 49, 73–86.

543

544 **Figure Captions**

545 Figure 1: Diopside dissolution textures observed with SEM indicate dendrites in all images
546 are $\leq 1 \mu\text{m}$. Dendrite lengths and apparent apical angles were measured using ImageJ as shown in
547 (A). Systematic changes in dendrite lengths or apparent apical angles were not observable with
548 increasing salinity between (B) 0.35 M NaCl and (E) 3 M NaCl or (C) 0.35 M Na₂SO₄ and (F)
549 1.7 M Na₂SO₄ or decreasing activity of water. Microdendrites are observed on pyroxene in
550 almost all images, except (A) UPW, (B) 0.35 M NaCl, (F) 1.7 M Na₂SO₄, and (G) 4.5 M CaCl₂.
551 These samples represent a very dilute solution ($a_{\text{H}_2\text{O}}=0.99$) and a concentrated brine
552 ($a_{\text{H}_2\text{O}}=0.35$).

553 Figure 2: Box and whisker plots of diopside (A) dendrite lengths and (B) apparent apical

554 angles display no clear trends with respect to changes in activity of water. (A) Denticle lengths
555 are significantly longer in 2 M NaClO₄. (B) With the exception of 3 M NaCl, apparent apical
556 angles are within range of each other. However, if outlier points are considered, apical angles in
557 all samples are similar.
558

559

Tables

Table 1: Measured Diopside Denticle Lengths (μm) and Apical Angles ($^\circ$)

Solution	UPW	0.35 M NaCl	0.35 M Na₂SO₄	1.7 M Na₂SO₄	2 M NaClO₄	3 M NaCl	4.5 M CaCl₂
<i>a</i>H₂O	1.0	0.99	0.98	0.95	0.93	0.75	0.35
Number of Measurements	77	118	230	112	85	50	21
Denticle Length (μm)							
Maximum	2.9	1.6	0.79	2.6	5.9	1.7	2.1
Upper Quartile	1.6	0.70	0.27	0.94	2.8	0.47	1.3
Median	0.76	0.50	0.14	0.67	1.1	0.36	0.59
Lower Quartile	0.50	0.30	0.075	0.49	0.60	0.22	0.40
Minimum	0.14	0.069	0.028	0.17	0.18	0.10	0.20
Mean	1.1	0.54	0.19	0.76	1.9	0.41	0.88
Std. Dev.	0.79	0.32	0.15	0.43	1.5	0.28	0.59
Apparent Apical Angle ($^\circ$)							
Maximum	78.3	61.7	63.5	55.0	45.0	79.8	57.5
Upper Quartile	23.4	13.6	27.7	16.3	22.5	55.2	38.7
Median	18.6	9.52	23.6	11.1	18.1	41.6	21.4
Lower Quartile	14.8	6.81	19.8	8.30	14.4	33.6	15.3
Minimum	6.71	3.21	5.84	3.40	8.40	11.5	11.8
Mean	21.2	12.2	24.3	14.4	19.3	43.1	26.5
Std. Dev.	12.1	8.84	7.55	10.1	7.10	16.1	13.2

560



