

1 **Diamonds from the lower mantle?**

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4 Natural diamonds and the mineral inclusions they trapped during growth provide a unique
5 opportunity to directly study material exhumed from great depths in the Earth. While the majority
6 of diamonds likely originate from depths of 140-220 km in cratonic mantle, a small subset appears
7 to have been exhumed from depths extending to > 800 km. These “superdeep” or “ultradeep”
8 diamonds are distinguished by mineral inclusions, or reverted inclusions, of high-pressure phases
9 (Walter et al. 2011; Pearson et al. 2014) and provide an unparalleled opportunity to study deep
10 mantle processes. In order to understand the information these samples provide it is critically
11 important that the depth of their formation is accurately constrained. Inclusions of magnesiowüstite
12 are among the mostly commonly described in sub-lithospheric diamonds and observation of these
13 inclusions are often assumed to indicate diamond provenance in the lower mantle. This is despite
14 the stability field of [Mg,Fe]O extending to ambient pressure conditions and experimental evidence
15 of magnesiowüstite stability in equilibrium with diamond throughout the upper mantle (Brey et al.
16 2004; Thomson et al. 2016). A new study by Uenver-Thiele et al. (2017) in *American Mineralogist*
17 places important new constraints on the formation and uplift history of inclusions containing
18 magnesioferrite.

19 Detailed studies of magnesiowüstite inclusions in diamonds from the Juina region of Brazil report
20 they often contain nanometre-sized crystals of magnesioferrite ([Mg,Fe²⁺]Fe³⁺₂O₄). These
21 precipitates occur at the interface between the diamond and inclusion, or as evenly distributed
22 dislocation “necklaces” within the inclusion interior (Harte et al. 1999; Wirth et al. 2014; Palot et al.
23 2016). Wirth et al. (2014) describe chains of globular [Mg_{0.5}Fe_{0.5}]Fe₂O₄ crystals, ~ 75 nm in size,
24 making up 6-11 vol.% of the entire [Mg₂₇Fe₇₁]O inclusion. This suggests an original Fe³⁺/ΣFe for
25 the inclusion of 11-14 %, compared with 7 ± 2 % in the recovered magnesiowüstite (McCammon
26 1997). Magnesioferrite is accompanied by small, ~ 10-30 nm, cubic voids, Al-bearing spinel and
27 Ni-Fe metal blebs. Palot et al. (2016) describe isolated 10-20 nm octahedra of
28 Mg[Fe_{0.75}Cr_{0.17}Al_{0.08}]₂O₄ throughout a [Mg₈₄Fe₁₆]O host with a recovered Fe³⁺/ΣFe content of 1-2
29 % that also contains ~ 30 ppm H₂O in brucite precipitates. Using the reported bulk inclusion
30 composition (~ [Mg₇₂Fe₂₈]O ignoring minor elements) implies the original magnesiowüstite must
31 have had an Fe³⁺/ΣFe of approximately 10-12 %. In both studies the magnesioferrite lamellae have
32 a topotaxial relationship with the [Mg,Fe]O host, confirming they must have formed during
33 exsolution from a homogenous magnesiowüstite grain. Using different arguments both studies
34 concluded that the magnesioferrite lamellae are indicative of the lower mantle provenance of these

35 diamonds. However, both their high ferric iron contents and new evidence from the phase relations
36 of magnesioferrite (Uenver-Thiele et al. 2017) might instead point to a shallower origin.

37 At low pressures (< 5 GPa) it is well understood that magnesiowüstite can incorporate significant
38 ferric iron, up to $\text{Fe}^{3+}/\Sigma\text{Fe}$ of 70 %, mainly charge balanced by negative cation vacancies (e.g.
39 Hazen and Jeanloz 1984; Dobson et al. 1998). With increasing pressure and decreasing oxygen
40 fugacity the ferric iron capacity of magnesiowüstite decreases, due to a high-pressure phase
41 transition of Fe_3O_4 (Huang and Bassett 1986; McCammon et al. 1998). Thus, at conditions just
42 within the lower mantle the maximum $\text{Fe}^{3+}/\Sigma\text{Fe}$ in $[\text{Mg}_{70}\text{Fe}_{30}]\text{O}$, similar to that observed by Palot et
43 al. (2016), is < 2% at the nickel-nickel oxide buffer (NNO) and < 0.5 % at the iron-wüstite buffer
44 (IW) (Otsuka et al. 2013). Similarly, a magnesiowüstite with $[\text{Mg}_{20}\text{Fe}_{80}]\text{O}$, similar to that observed
45 by Wirth et al. (2014), would have a $\text{Fe}^{3+}/\Sigma\text{Fe}$ capacity of ~ 7 - 14 % at the IW and NNO buffers
46 respectively. Further increasing pressure or decreasing temperature lowers the ferric iron solubility.
47 Thus, the magnesioferrite exsolution observed in diamond-hosted inclusions, presumably driven by
48 high ferric iron, appears inconsistent with the compositions of magnesiowüstite expected in the
49 lower mantle.

50 In this issue, Uenver-Thiele et al. (2017), have experimentally determined the high-pressure phase
51 relations of magnesioferrite (MgFe_2O_4) using the multi anvil apparatus. Prior to this study it was
52 believed that MgFe_2O_4 had a relatively simple phase diagram, with the ambient cubic spinel
53 structure (*Fd-3m*) stable until an isochemical phase transition to orthorhombic CaMn_2O_4 structure
54 (*Pbcm*), HP- MgFe_2O_4 , at ~ 17 GPa and temperatures above 1700 °C, or breakdown to $\text{MgO} +$
55 Fe_2O_3 at lower temperatures (Levy et al. 2004). Instead, the experiments of Uenver-Thiele et al.
56 (2017) have revealed a very different phase diagram, observing the spinel-structured MgFe_2O_4
57 decomposes at ~ 10 GPa. It forms a phase assemblage of $\text{MgO} + \text{Fe}_2\text{O}_3$ at temperatures below 1200
58 °C or $\text{Fe}_2\text{O}_3 +$ an unrecoverable phase of $\text{Mg}_5\text{Fe}_2\text{O}_8$ - $\text{Mg}_4\text{Fe}_2\text{O}_7$ stoichiometry at higher
59 temperatures. At pressures beyond ~ 13 GPa the unrecoverable phase(s) are replaced by
60 orthorhombic, CaFe_3O_5 structured (*Cmcm*), $\text{Mg}_2\text{Fe}_2\text{O}_5$ (Boffa Ballaran et al. 2015). HP- MgFe_2O_4
61 was not observed at any conditions up to 18 GPa and 1300 °C in this study. Further high-pressure
62 experiments are required in order to determine the structure(s) of the unrecoverable phase(s) using
63 *in-situ* methods, the full extent of the $\text{Mg}_2\text{Fe}_2\text{O}_5$ stability field and whether HP- MgFe_2O_4 becomes
64 stable at higher pressures as suggested by previous studies (Andrault and Bolfan-Casanova 2001;
65 Levy et al. 2004).

66 The phase relations determined by Uenver-Thiele et al. (2017), coupled with the low ferric iron
67 capacity of magnesiowüstite in the lower mantle, have very significant consequences for the
68 interpretation diamond formation pressure. Firstly, magnesioferrite is not stable at lower mantle

69 conditions where the diamond inclusions (Wirth et al. 2014; Palot et al. 2016) are believed to have
70 formed. Secondly, if the magnesioferrite did exsolve from (Mg,Fe)O as HP-MgFe₂O₄ in the lower
71 mantle, it could not have directly inverted to the spinel structure, due to the large stability field of
72 Mg₂Fe₂O₅ + Fe₂O₃ as suggested by Palot et al. (2016). The presence of an additional minor phase
73 between the magnesioferrite platelets (Wirth et al. 2014) does suggest inversion from exsolution
74 lamellae of alternative stoichiometry, and the conversion of Mg₂Fe₂O₅ + Fe₂O₃ into magnesioferrite
75 at ~ 300 km depth would appear to be consistent with these observations. The conditions of original
76 inclusion entrapment remain uncertain, however, high ferric iron contents and magnesioferrite
77 phase relations are more consistent with formation in the upper mantle or transition zone. The study
78 of Uenver-Thiele et al. (2017) highlights the potentially rich and unexplored chemography and
79 importance of post-spinel phase relations for understanding the Earth's fundamental geochemical
80 and geodynamic cycles.

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