## 1 Revision 1

2	The origin of extensive Neoarchean high-silica batholiths and the nature of intrusive
3	complements to silicic ignimbrites: insights from the Wyoming batholith, U.S.A.
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9	
10	Abstract
11	Extensive intrusions composed entirely of biotite granite are common in Neoarchean cratons.
12	These granites, which have high silica and potassium contents, are not associated with intermediate and
13	mafic phases. One such Neoarchean granite batholith, herein named the Wyoming batholith, extends
14	more than 200 km across central Wyoming in the Granite and the Laramie Mountains. From field
15	characterization, petrology, geochemistry, and Nd isotopic data we establish that the magnesian
16	Wyoming batholith exhibits continental arc chemical and isotopic signatures. It is best interpreted as a
17	large, upper crustal silicic batholith that likely formed when the subducting oceanic plate steepened or
18	foundered, bringing mantle heat and mass to the base of the crust. Similar Cenozoic settings such as the
19	Altiplano-Puna plateau of the Andes and the volcanic provinces of the western United States, host large
20	volumes of silicic ignimbrite. The magma chambers supplying these eruptions are inferred to be silicic
21	but the structural, petrologic, and geochemical details are unknown because the batholiths are not
22	exposed. We suggest that the Wyoming batholith represents an analog for the plutonic complex
23	underlying these ignimbrite systems, and provides an opportunity to examine the shallow magma
24	chamber directly. Our work establishes that aside from more leucocratic margins, the sill-like magma
25	chamber is petrologically and chemically homogeneous, consistent with effective mixing by vertical

26	convection. Nd isotopic variations across the batholith indicate that horizontal homogenization is
27	incomplete, preserving information about the feeder system to the batholith and variations in magma
28	sources. The late Archean Earth may present optimal conditions for the formation of extensive granite
29	batholiths like the Wyoming batholith. By this time the majority of the planet's continental crust had
30	formed, providing the environment in which differentiation, distillation, and assimilation could occur.
31	Moreover, the Neoarchean Earth's relatively high radioactive heat production provided the power to
32	drive these processes.
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34	Keywords: granite, ignimbrite, continental arc, batholith, Archean
35	
36	Introduction
37	Continental arc batholiths form from magmas that are relatively oxidized, wet, and cool. These
38	rocks, dominated by granodiorite, are magnesian, calc-alkalic, and metaluminous (Christiansen, 2005;
39	Bachmann and Bergantz, 2008; Frost and Frost, 2014). Small volumes of true granite are components
40	of continental arc batholiths. These granite bodies tend to be younger than the granodiorites, consistent
41	with their origin by differentiation. If the granites are end-products of differentiation of a basaltic
42	parent, then they represent the less than 5% of the volume of original magma. Even if the starting
43	material is granodiorite and differentiation is accompanied by crustal assimilation, the volume of less
44	silicic cumulates will be 2-3 times greater than the volume of granite produced (Lee and Morton,
45	2015). As a result, it is not surprising that granites make up a relatively modest proportion of
46	continental arc batholiths.
47	In contrast, intrusions composed exclusively of granite are a notable feature of the Neoarchean
48	geologic record. Biotite and two-mica granites are a common component of every Archean craton, and
49	are the second most widespread lithology in Neoarchean terranes after the tonalite-trondhjemite-
50	granodiorite association (Laurent et al., 2014). These granites, which have silica contents of 70-76%

51 SiO<sub>2</sub>, are not associated with intermediate and mafic phases, they are calc-alkalic to alkali-calcic,

peraluminous, and high-K (Laurent et al., 2014). The granites are typically undeformed, extensive, and are widely distributed throughout the host craton. Most petrogenetic models for the formation of Neoarchean potassic granites call upon partial melting of older tonalite-trondhjemite-granodiorite crust or metasedimentary rocks (Moyen, 2011; Jaguin et al., 2012). Even if the magma sources lie dominantly within the crust, the production of such large volumes of granite across a broad area seems to require a large power input from the mantle.

58 This study documents a large batholith composed entirely of magnesian biotite granite in west-59 central Wyoming. It is petrologically indistinguishable from a late-stage granite pluton within a 60 contemporary Neoarchean continental arc composite batholith exposed in the neighboring Wind River 61 Range. This granite batholith, here named the Wyoming batholith, is exposed in Laramide basement-62 involved uplifts that extend 200 km from east to west. It is everywhere composed of undeformed 63 biotite granite, even though outcrops expose varying structural levels. We investigate its petrogenesis 64 and its potential similarity to the granitic batholiths that are inferred to underlie large silicic ignimbrite 65 provinces.

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#### 67 Geologic setting

68 The Neoarchean magmatic rocks of central Wyoming compose part of the Wyoming province, 69 an Archean craton that extends across most of Wyoming and parts of Montana, Idaho, Utah, and South 70 Dakota (Fig. 1). Today rocks of the Archean Wyoming province are exposed in the cores of basement-71 involved, Laramide uplifts. The craton is composed of several sub-provinces (Fig. 1). The Montana 72 metasedimentary province lies in the northwest part of the craton and is composed of Neoarchean 73 quartzite, pelite and carbonate rocks structurally intercalated with Paleoarchean quartzofeldspathic 74 gneiss (Mogk et al., 1988, 1992; Mueller et al., 1993, 2004; Mueller and Frost, 2006). The Beartooth-75 Bighorn magmatic zone (BBMZ) is dominated by ~2.8 Ga tonalites and granodiorites, although rocks

76	as old as 3.5 Ga and detrital zircons as old as 4.0 Ga have been identified (Mueller et al., 1996; 1998,
77	Frost and Fanning, 2006). The BBMZ is characterized by radiogenic whole rock and feldspar
78	<sup>207</sup> Pb/ <sup>204</sup> Pb ratios and Nd model ages up to 4.0 Ga (Mueller and Frost, 2006). The Wyoming province
79	grew southward (present-day coordinates) through accretion of terranes between 2.68 and 2.65 Ga.
80	Some of these southern accreted terranes are juvenile, and others contain pre-existing continental crust
81	(Mueller and Frost, 2006; Souders and Frost, 2006; Chamberlain et al., 2003).
82	Neoarchean granitic rocks dominate the Archean exposures of central Wyoming along the
83	boundary of the BBMZ with the southern accreted terranes (Fig. 1, 2). The 2.63 Ga Louis Lake
84	batholith occupies the southern half of the Wind River Range (Fig. 2; Stuckless et al., 1985; Frost et
85	al., 1998). This batholith, which is exposed over an area $>1,600$ km <sup>2</sup> , is composed of a suite of
86	undeformed, calc-alkalic diorites, granodiorites, and granites. It is intruded by a biotite granite body
87	known as the Bears Ears pluton, which varies from equigranular to porphyritic. This composite
88	batholith was derived from magmas composed of varying proportions of crustal and mantle sources,
89	and has been interpreted as a continental magmatic arc complex (Frost et al., 1998).
90	East of the Wind River Range, Precambrian exposures in the Granite Mountains, Shirley
91	Mountains, and Laramie Mountains consist exclusively of monotonous, undeformed, biotite granite
92	despite Laramide and younger basement-involved faulting that exposes varying structural depths. U-Pb
93	geochronology suggests that the granites are ~2.62-2.63 Ga (Ludwig and Stuckless, 1978; Wall, 2004;
94	Bagdonas, 2014). Precambrian exposures are extensive in the Laramie Mountains, but Tertiary
95	sedimentary rocks cover much of the Precambrian basement to the west in the Shirley Basin and in the
96	vicinity of the Granite Mountains. Nevertheless, structural and geophysical evidence is consistent with
97	the hypothesis that biotite granite extends at least 200 km east-west across the area (present
98	coordinates). There are no major faults between the Granite and Laramie Mountains, and the top of the
99	Precambrian basement between the two mountain ranges is relatively shallow (~2 km; Robbins and
100	Grow, 1992). Complete Bouger gravity anomaly and isostatic residual anomaly maps show similar

101 anomalies in both mountain ranges, and a smooth gradient to values 15-20 mgals lower in the 102 intervening basin is interpreted to reflect low-density basin fill. These characteristics have been 103 interpreted to suggest that felsic basement extends from the Granite Mountains to the Laramie 104 Mountains (Robbins and Grow, 1992). If the entire area from the Granite Mountains to the Laramie Mountains is underlain by biotite granite, then it occupies an area of more than 6,900 km<sup>2</sup>. Because it 105 106 appears so extensive, we refer to this Neoarchean granite as the Wyoming batholith. 107 Few contacts of the Wyoming batholith with its country rocks are exposed. In the eastern 108 Granite Mountains the northern margin of the Wyoming batholith intrudes 3.30-3.45 Ga gneisses 109 (McLaughlin et al., 2013; Frost et al., 2015). Both the Neoarchean granite and Paleoproterozoic 110 gneisses are silicified and epidotized along the contact. In addition, the contacts are commonly 111 occupied by younger diabase dikes. The southern contact of the Wyoming batholith in the Granite 112 Mountains is fault-bounded. Pliocene motion on the southern Granite Mountains fault down-dropped 113 the Granite Mountains by a minimum of 600 meters (Love, 1970). No contacts with older rocks are 114 exposed in the Shirley Mountains. In the Laramie Mountains, the northern and southern contacts are 115 exposed but the eastern and western margins are covered by Cenozoic sedimentary rocks. The northern 116 and southeastern contacts of the granite are gradational and marked by development of biotite foliation 117 against the host gneisses. The southwestern contact is fault-bounded (Condie, 1969). 118 119 The Wyoming batholith 120 Previous studies have described portions of the Wyoming batholith. The Granite Mountains 121 granites were investigated by the U.S Geological Survey (e.g., Peterman and Hildreth, 1978; Stuckless 122 et al., 1977; Stuckless and Miesch, 1981) and were the subject of theses by Langstaff (1995), Wall 123 (2004), and Meredith (2005). The Bear Mountain granite, an exposure of the Wyoming batholith near 124 the southern Granite Mountains fault, was described by Bowers and Chamberlain (2006). Condie 125 (1969) reported on the mineralogy and petrology of the Laramie Mountains granites. For this study, the

lead author undertook extensive field observations across the entire Wyoming batholith, including the
first systematic study of the granites in the Shirley Mountains (McLaughlin et al., 2013; Bagdonas,
2014).

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# 130 Field relations and petrography

131 Very little petrographic diversity exists within the Wyoming batholith, although magmatic 132 fabrics are observed particularly adjacent to country rock. Two main lithologic units are identified 133 throughout the batholith on the basis of field observations and petrography: biotite granite and 134 leucocratic banded granite. Both of these units may be hydrothermally altered, as described by Ludwig 135 and Stuckless (1978). Where altered, the granites form blocky outcrops that are more resistant to 136 weathering than the unaltered rocks.

137 Leucocratic banded granite appears to be more common along batholith margins. It is present 138 along the northern contact with older country rocks in the northwestern Granite Mountains where 139 banding in the granite parallels the contact. Leucocratic banded granite may also form near the roof of 140 the batholith, as suggested by its location on the summit of Lankin Dome (Fig. 3a). Contacts with 141 biotite granite are abrupt to gradational, and are commonly scalloped to feathery (Fig. 3b,c). 142 Leucocratic banded granite tends to be coarse-grained along these contacts. In other places, leucocratic 143 banded granite is incorporated within biotite granite. On an inselberg near Sweetwater station in the 144 Granite Mountains leucocratic banded granite appears both on the summit and as folded enclaves and 145 lenses within biotite granite (Fig. 3d). We attribute these relationships to entrainment of solidified 146 banded granite from batholith margins into more melt-rich biotite granite crystal mush. 147 148 **Biotite granite.** Biotite granite is the dominant rock type of the Wyoming batholith, occupying 149 approximately 90% of outcrop. Most is medium- to coarse-grained, although fine-grained varieties

150 were observed in the Shirley and Laramie Mountains and slightly porphyritic biotite granite was found

151 in places in the Laramie Mountains. Biotite granite is composed of 30-40% microcline, 25-35% quartz, 152 25-35% plagioclase (An<sub>5</sub>-An<sub>30</sub>), and 1-7% biotite. Accessory minerals include epidote, titanite, 153 magnetite, zircon, and apatite (Fig. 4a-c). Trace amounts of muscovite are present in some localities. 154 Biotite granite generally is homogeneous, exhibits weak to no fabric, and lacks mafic enclaves. 155 Foliation in biotite granite is magmatic, and is marked by aligned plagioclase and biotite. No post-156 magmatic deformation is evident in these rocks. 157 Leucocratic banded granite. Leucocratic banded granite comprises approximately 10% of 158 outcrops and is present across the Wyoming batholith. It is characterized by millimeter- to decameter-159 scale compositional banding, typically defined by aligned biotite and most likely formed due to 160 magmatic flow. Leucocratic banded granite is composed of 35-40% quartz, 20-35% microcline, 20-161 30% plagioclase, 0-5% biotite and 0-5% magnetite. In some localities it contains sparse potassium 162 feldspar phenocrysts. Accessory minerals include epidote, titanite, zircon and apatite. Magnetite ranges 163 from mm- to cm-diameter and appears to form at the expense of biotite (Fig. 4d-f).

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#### 165 Geochronology

166 Past efforts to date Wyoming batholith granites have established that uranium concentrations of 167 Wyoming batholith zircon are high and alpha-recoil damage has led to considerable open-system 168 behavior (Ludwig and Stuckless, 1978). For this study, U-Pb isotopic data on zircon from three 169 Wyoming batholith samples were obtained using SHRIMP RG at the Australian National University. 170 Zircon grains mounted in epoxy together with chips of the Temora reference zircons (Black et al., 171 2003) and polished. Cathodoluminescence (CL) Scanning Electron Microscope (SEM) images were 172 taken for all zircons. The U-Th-Pb analyses were performed following procedures given in Williams 173 (1998, and references therein). Temora reference zircon grains were analyzed after every three 174 unknown analyses. The data were reduced using the SQUID Excel Macro of Ludwig (2001). The Pb/U 175 ratios were normalized relative to a value of 0.0668 for the Temora reference zircon, equivalent to an

176 age of 417 Ma (Black et al., 2003). Uncertainties given for individual analyses (ratios and ages) are at 177 the one sigma level with correction for common Pb made using the measured <sup>207</sup>Pb/<sup>206</sup>Pb ratios. Concordia plots, linear discordia regression fits and weighted mean  $^{207}$ Pb/ $^{206}$ Pb age calculations were 178 carried out using either ISOPLOT or ISOPLOT/EX (Ludwig, 2003). Weighted mean <sup>207</sup>Pb/<sup>206</sup>Pb ages 179 180 are calculated and uncertainties reported at the 95% confidence level. 53 analyses yielded uranium 181 contents of 279-4984 ppm and discordant U-Pb compositions (Supplementary Tables S1-3). Data 182 interpretation below focuses on those grains with the lowest uranium concentrations and most 183 concordant U-Pb isotopic compositions. The CL images for all samples show either simple igneous 184 internal structures, mostly oscillatory zoning, or are dominated by metamict areas, interpreted as 185 having replaced the primary igneous zoning. 186 187 Leucocratic banded granite (10LD2). Sample 10LD2 was collected from the summit of 188 Lankin Dome, a sample of leucocratic banded granite with very little feldspar and biotite alteration. A 189 single morphology of euhedral zircon grains yielded uranium concentrations from ~280-3115 ppm. 190 Data from three high uranium zircons that were >50% discordant were not considered. A 7-point 191 regression of the remaining analyses yielded an imprecise upper concordia intercept of  $2628 \pm 21$  Ma (MSWD = 4.7). The single concordant analysis gave a  ${}^{207}$ Pb/ ${}^{206}$ Pb age of 2627 ± 3 Ma (1 $\sigma$ ), which we 192 193 interpret as the best estimate of the magmatic age for this sample (Figure 5a).

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Biotite granite sample 11SMG1. This sample also yielded euhedral zircon, with some grains coarser than ~200 $\mu$ m in length and others around 100 $\mu$ m; both smaller and larger grains yielded indistinguishable dates. Uranium contents range from ~450 to ~4985 ppm, and most areas analyzed have lost radiogenic Pb; up to ~50% discordant. A cluster of 5 analyses of zircon with U < 600 ppm and from  $\leq$ 4% discordant define an upper intercept age of 2626 +16/-9 Ma (MSWD = 0.60). These analyses yield a mean weighted <sup>207</sup>Pb/<sup>206</sup>Pb age of 2623 ± 4.3 Ma (MSWD = 1.02), which we interpret as the best estimate of the crystallization age (Fig 5b). A sixth low-U zircon gave a  $^{207}$ Pb/ $^{206}$ Pb age of ~2715 Ma, which we interpret as reflecting an inherited component.

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204Hydrothermally altered biotite granite sample 11SMG2. This sample of biotite granite205exhibits strong hydrothermal alteration fabrics, with complete replacement of feldspars by sericite,206biotite by chlorite and epidote, and significant quartz sub-grain development. Uranium contents of207zircon are high, up to 5615 ppm, and most analyses are strongly discordant. The two grains with the208lowest uranium contents yielded  $^{207}$ Pb/ $^{206}$ Pb ages of 2624 ± 4 Ma (-1% discordant) and 2614 ±7 Ma209(5% discordant; Fig 5c). We interpret the more concordant analysis as the best estimate of the210crystallization age.

These results suggest that the age of the Wyoming batholith is approximately 2625 Ma, similar to unpublished dates obtained by other workers (see Bagdonas, 2014 for a summary). The Wyoming batholith ages reported here are comparable to dates obtained for the Bears Ears pluton in the Wind River Range of  $2620 \pm 4$  Ma (Wall, 2004) and a granitic dike in the northern Wind River Range of  $2618.9 \pm 1.5$  Ma (Frost et al., 1998). The Wyoming batholith appears to be contemporaneous to slightly younger than the Louis Lake batholith of the Wind River Range, which has been dated at  $2629.2 \pm 2.8$ Ma and  $2629.5 \pm 1.5$  Ma (Frost et al., 1998).

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#### 219 Geochemistry

220 Twenty Wyoming batholith samples were analyzed by ICP-AES for major and minor element

concentrations and by ICP-MS for rare earth element abundances by ALS Minerals, Ltd. (Table 1).

222 These data, combined with analyses from the literature, indicate that the Wyoming batholith is

uniformly high in silica (70 to 77% SiO<sub>2</sub>; Table 1; Condie, 1969; Stuckless and Peterman, 1977;

224 Stuckless and Miesch, 1981; Wall, 2004; Meredith, 2005). The granites are mainly magnesian, calc-

alkalic to alkali-calcic, and peraluminous (Fig. 6). There is no variation in major element composition

226 related to location within the batholith, although the range in aluminum saturation index is greater 227 among samples from the Laramie Mountains than elsewhere in the batholith. The Bears Ears pluton 228 samples plot within the fields defined by Wyoming batholith samples. By contrast, samples of the 229 Louis Lake batholith exhibit a wide range in silica (52-76% SiO<sub>2</sub>). Like the Wyoming batholith, the 230 Louis Lake batholith is magnesian. However, Louis Lake batholith samples are mostly calc-alkalic and 231 metaluminous (Frost et al., 1998; Frost et al., 2000; Stuckless et al., 1985; Stuckless, 1989). 232 Variation diagrams of K<sub>2</sub>O, Rb, Sr, and Eu versus SiO<sub>2</sub> (Fig. 7) show elevated concentrations of 233 incompatible elements K<sub>2</sub>O and Rb in Wyoming batholith samples and Louis Lake batholith samples 234 with  $SiO_2 > 70\%$ , relative to the mafic to intermediate rocks of the Louis Lake batholith. These 235 characteristics are consistent with late crystallization of biotite and potassium feldspar. By contrast, 236 elements compatible in plagioclase, such as Sr and Eu, are lower in the high-silica rocks than in the 237 Louis Lake batholith. Elements compatible in hornblende, such as Nb and Y, are higher in the quartz 238 diorite and granodiorite Louis Lake batholith samples than in the Wyoming batholith and Bears Ears 239 granites. Both batholiths plot mainly in the volcanic arc granite field of Pearce et al. (1984; Fig. 8), 240 though Louis Lake batholith samples with cumulate hornblende, and accordingly higher Nb and Y, 241 extend beyond. 242 Rare earth element (REE) patterns for the Wyoming batholith exhibit some variability (Fig. 9). 243 Biotite granites from the Granite Mountains have uniform, LREE-enriched and HREE-depleted 244 patterns with negative Eu anomalies whereas banded granite has flatter and higher HREE contents (Fig. 245 9a). Some samples from the Shirley Mountains lack Eu anomalies, and one sample from the Laramie 246 Mountains has low LREE and a slight positive Eu anomaly (Fig. 9b, c). These variations are 247 interpreted to reflect different proportions of cumulate feldspar, zircon and other REE-bearing 248 accessory minerals in the samples. The REE patterns of hornblende-bearing Louis Lake batholith 249 samples are less HREE-depleted than Wyoming batholith samples (Fig. 9d). The Eu anomaly in Louis 250 Lake batholith samples varies from negative to slightly positive, consistent with differentiation by

fractionation (Stuckless et al., 1989; Frost et al., 1998). REE patterns of Bears Ears samples fall within
the field defined by Wyoming batholith samples, and variations likely reflect cumulate processes (Fig.
9d).

254

### 255 Nd isotopic compositions

256 Sm and Nd were isolated from homogenized powders of 17 Wyoming batholith samples using 257 the methods described in Frost et al. (2006) and were analyzed at the University of Wyoming on a Neptune MC-ICP-MS. Samples were normalized to  ${}^{146}$ Nd/ ${}^{144}$ Nd = 0.7219. La Jolla Nd analyzed after 258 every five unknowns gave  ${}^{143}$ Nd/ ${}^{144}$ Nd = 0.51185 ± 0.00001 (2 standard deviations). These data, 259 260 together with analyses reported by Frost et al. (2006), Fruchey (2002), Wall (2004), and Meredith 261 (2005), show that initial Nd isotopic compositions of the Wyoming batholith vary considerably from 262 +1.9 to -7.8. They mostly overlap data from the Louis Lake batholith and Bears Ears pluton but extend 263 to more negative initial  $\varepsilon_{Nd}$  (Fig. 10). Nd isotopic compositions correlate with location: in general, 264 samples from the Shirley and Laramie Mountains have less radiogenic initial ratios than samples from 265 the Granite Mountains and Bears Ears pluton (Fig. 11). This suggests that despite the uniformly high 266 SiO<sub>2</sub> of Wyoming batholith samples and their overall geochemical similarity, they do not share a single 267 magma source or common proportions of magma sources.

Although it is not possible to know the Nd isotopic composition of potential magma sources at depth in the Wyoming craton, the Nd isotopic composition has been established for various Archean country rocks that are exposed on the surface today. Country rocks to the Wyoming batholith in the Granite Mountains are composed of 3.3-3.4 Ga gneisses (Fruchey, 2002; Frost et al., 2015). At 2625 Ma, these gneisses had  $\varepsilon_{Nd}$  values of -9 to -14. Evolved quartzite and pelitic rocks deposited on this basement had slightly more radiogenic  $\varepsilon_{Nd}$  values at 2625 Ma of between -5 and -4 (Fruchey, 2002), and juvenile metasedimentary rocks accreted to the Wyoming province at approximately 2.67 Ga had

275  $\varepsilon_{Nd}$  values from 1 to +4 (Frost et al., 2006). The wide variation in proportions of juvenile crust or 276 mantle and crust incorporated into Wyoming batholith magmas indicated by the range in initial  $\varepsilon_{Nd}$  can 277 be quantified by the Neodymium Crustal Index (NCI; DePaolo et al., 1992). Assuming a juvenile end 278 member with  $\varepsilon_{Nd}$  of +4 and a Paleoarchean crustal source of -12, the proportion of Nd from crustal 279 sources varies from 13 to 73% with an average of 39% (n = 33). The crustal assimilant is likely to have 280 been felsic for two reasons. First, the older Archean crust, where exposed, is dominated by granite, 281 granodiorite, and trondhjemite (McLaughlin et al., 2013; Frost et al., 2015). Second, if the crustal 282 assimilant was of a strongly contrasting composition then a correlation should exist between amount of 283 assimilation and geochemical composition of the Wyoming batholith granite. 284 The 2630 Ma Louis Lake batholith, which ascended through the crust in the same area where 285 the Bears Ears was later emplaced, also exhibits a range of initial  $\varepsilon_{Nd}$  (-2.1 to +3.5; Fig. 10). This more 286 restricted range in initial  $\varepsilon_{Nd}$  compared to the Wyoming batholith is indicative of a lower proportion 287 older crustal sources in the Louis Lake batholith (NCI = 20-37%; n = 11) (Frost et al., 1998). The 288 initial  $\varepsilon_{Nd}$  of the Bears Ears and the western part of the Granite Mountains overlap those of the Louis 289 Lake batholith. Nd isotopic data permit that the Bears Ears pluton and Wyoming batholith in the 290 Granite Mountains formed either from Louis Lake residual magmas, or from similar proportions of the 291 same sources. In either case, Nd isotopic systematics require that the Wyoming batholith granite farther 292 to the east in the Shirley and Laramie Mountains must have incorporated a larger proportion of older 293 continental crust (Fig. 10, 11). 294 295 Discussion 296

#### 297 Extent of the Wyoming Batholith

298 This study has identified a large 2625 Ma biotite granite batholith that extends east-west at least 299 200 km across the Laramie, Shirley, and Granites Mountains of central Wyoming. In addition, we have 300 shown that the Bears Ears pluton of the Wind River Range is petrographically, temporally, and 301 chemically similar to the Wyoming batholith, and that together these biotite granites dominate a large 302 area of the exposed Archean crust (Fig. 2). It is possible that the Wyoming batholith intrusive event is 303 even more extensive than defined here. For example, it is unclear whether the Prospect Mountain 304 granite of the southwestern Wind River Range is part of the Louis Lake or the Wyoming batholith: 305 Sutherland and Luhr (2011) describe biotite granite that we have confirmed is petrographically similar 306 to the Wyoming batholith. Another possible part of the batholith lies in the Lewiston district at the 307 southeastern end of the Wind River Range, where the granite exposed there is unstudied. It has been 308 assumed to belong with the Louis Lake batholith and Bears Ears pluton, but may equally belong to the 309 Wyoming batholith. Smaller outcrops of ~2.6 Ga granite are also present within the Owl Creek 310 Mountains and on Casper Mountain at the northern end of the Laramie Mountains (Fig 2). In places 311 these granites are garnet-bearing, and their relationship to the Wyoming batholith is uncertain. Even if 312 none of these granites should be considered part of the Wyoming batholith, it remains a large body of 313 true granite, dwarfing the size of the neighboring, roughly contemporary Louis Lake continental arc 314 batholith to the west.

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#### 316 **Petrogenesis of the Wyoming batholith**

Silica-rich granite is found in many tectonic environments, from extensional regimes and orogenic margins to continental collision zones. Even though their mineralogies are similar, silica-rich granites retain in their major and trace element geochemical compositions evidence of their petrogenesis (Frost et al., 2016). Large volumes of magnesian granite commonly form by differentiation, usually accompanied by some degree of crustal assimilation (Bachmann and Bergantz, 2008). Two tectonic settings are most likely to the generate magnesian, calc-alkalic to alkali-calcic

323 granite of the Wyoming batholith. The first, a magmatic arc batholith setting, is suggested by the 324 spatial and temporal association of the Wyoming batholith with the Louis Lake continental arc 325 batholith. In this model, the biotite granite of the Wyoming batholith is part of an incrementally 326 assembled composite batholith that also is composed of large volumes of granodiorite and more mafic 327 rocks. The second, a large, shallow silicic system related to a continental arc, is suggested by the 328 abundance of granite and lack of more mafic components. In this model, changes in subduction 329 dynamics allow for a pulse of mantle heat to power melting and the formation of a large silicic 330 magmatic system in the upper crust. Each of these is evaluated in more detail below.

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332 A continental arc batholith? Though dominated by granodiorite, continental arc batholiths 333 also contain true granite (Bateman and Chappell, 1979; Lee et al., 2007). These granites are thought to 334 form by fractional crystallization, and commonly also involve crustal assimilation (e.g., Ague and 335 Brimhall, 1988; Putirka et al., 2014). Such a process may have formed the Bears Ears granite from 336 residual magmas of the Louis Lake continental arc batholith. Geochemically, the Bears Ears pluton 337 samples lie at the high-silica end of Louis Lake batholith differentiation trends. Bears Ears pluton 338 initial  $\varepsilon_{Nd}$  values are within the range of Louis Lake samples, permitting formation of the Bears Ears 339 granite from Louis Lake granodiorite by differentiation alone. In the field, contacts between the Bears 340 Ears pluton and Louis Lake granodiorites are complex. They are gradational in places, but in others the 341 Bears Ears has intruded the granodiorites. Taken together, all observations are consistent with the 342 crystallization of the Bears Ears granite at the top of the Louis Lake batholith in the latter stages of its 343 development.

The similarities in petrology, age, and geochemistry could suggest that the Wyoming batholith, like the Bears Ears pluton, is also related to a continental arc magmatic system. A difficulty with a continental arc model is the sheer quantity of granite exposed in the Wyoming batholith. To form granite magma by differentiation of a basalt parent requires approximately 95% of the original melt

348 volume to have crystallized (Lee and Morton, 2015). However, none of these voluminous mafic 349 components of such a magmatic system is evident in the Wyoming batholith, which is composed solely 350 of true granite. One possibility is that the structural level of exposure of the Wyoming batholith is 351 higher than the Louis Lake batholith in the Wind River Mountains, such that only the granitic top of 352 the continental arc batholith appears at the erosional surface. Current topographic relief is 353 approximately 1.3 km, and Laramide and Cenozoic faulting has exposed the Wyoming batholith to a 354 minimum of 0.75 km additional structural relief. Yet no other rock types are exposed, even in the 355 deepest sections. Although possible, it seems unlikely that faulting would have brought none of the 356 more voluminous, more mafic components of the continental arc batholith to the surface in the Granite, 357 Shirley, and Laramie Mountains as it did in the Wind River Range. 358 359 **Plutonic equivalent of silicic ignimbrites?** Alternatively, the Wyoming batholith may 360 represent the plutonic record of a large rhyolitic system associated with a continental arc. A possible 361 analog system is preserved in the western United States, where large volumes of silicic, high-K 362 ignimbrites formed during the middle Cenozoic. The 35-27 Ma Southern Rocky Mountain volcanic 363 field (SRMVF) of Colorado and northern New Mexico erupted voluminous silicic ignimbrites. The 364 field is composed of dominantly dacitic volcanic tuffs originating from a number of different calderas

along with several small, subvolcanic plutons (Lipman, 2007; Lipman and Bachmann, 2015). Farther

to the west, the 36-18 Ma southern Great Basin ignimbrite province in Nevada and Utah is dominated

367 by dacites that share a volcanic arc chemical signature with the continental arc rocks to the west (Best

368 et al., 2013a). These volcanic rocks were deposited on a high plateau in an area of orogenically

thickened crust known as the Great Basin altiplano (Best et al., 2009). Magmatism is thought to have

370 resulted from progressive steepening of the subducting oceanic Farallon plate (Best et al., 2013a). As

- 371 the slab rolled back, basaltic parental magmas to the ignimbrites formed by partial melting in the
- 372 lithospheric mantle and asthenosphere (Christiansen and Best, 2014). Throughout the western United

373 States, Cenozoic episodes of increased magma production and large ignimbrite eruptions appear to 374 coincide with a regional transition from low-angle subduction to an extensional regime (Lipman, 375 2007). The Neogene Altiplano-Puna plateau represents a second example of a continental arc setting 376 that yielded voluminous silicic ignimbrites (de Silva and Gosnold, 2007). The thick crust of the 377 Altiplano-Puna plateau in the central Andes formed during a period of flat slab subduction. Steepening 378 of the slab in mid-Miocene time introduced mantle heat and partial melt at the base of the crust, leading 379 to a flare-up in ignimbrite activity (Kay et al., 1999). 380 In both the western United States and the Altiplano-Puna plateau, the ignimbrites are inferred to 381 be surface manifestations of the formation of large silicic batholiths at depth (Best et al., 2013b; de 382 Silva and Gosnold, 2007; Lipman, 2007), although the plutonic underpinnings of these ignimbrites are 383 not exposed in the Andes and Great Basin, and only small subcaldera intrusions are present in the 384 Southern Rocky Mountain field. However, in the Andes an extensive low-velocity zone has been 385 imaged and interpreted as the sub-volcanic plutonic complex. This shallow low-velocity zone is

approximately 200 km in diameter and 11 km thick. The uppermost parts of this zone are suggested to

387 be composed of shallow silicic magma bodies that feed the ignimbrite eruptions (Ward et al., 2014).

388 Similarly, geophysical anomalies in the Southern Rocky Mountain volcanic field may image

underlying shallow batholiths of similar dimension (Lipman and Bachmann, 2015).

390 One potential difficulty with the model of the Wyoming batholith as an intrusive complement to 391 silcic ignimbrites is compositional: most ignimbrites in these settings are crystal-rich dacites, less 392 siliceous than the Wyoming batholith granites. However, crystal-poor ignimbrites tend to be rhyolitic. 393 For example, the second-largest ash-flow sheet in the SRMVF, the Carpenter Ridge tuff, is composed 394 of a basal, crystal-poor densely welded rhyolite with 77-78%  $SiO_2$  (Bachmann et al., 2014). Although 395 the most voluminous SRMVF eruption, the Fish Canyon crystal-rich tuff, is dacitic at around 68% 396 SiO<sub>2</sub>, the glass in the tuff has much higher silica, from 76.5-78% SiO<sub>2</sub> (Bachmann et al., 2002). Farther 397 west in the Great Basin, glass within Cottonwood Wash tuff is rhyolitic (Ross et al., 2015) and the

398 crystal-poor, glass-rich portions of the Lund tuff extend into the low-silica rhyolite field (Maughan et 399 al., 2002). These observations have led to the hypothesis that crystal-rich monotonous intermediates 400 represent erupted mush, and that liquid segregated from this mush is the source of high-silica rhyolites 401 (Charlier et al., 2007). 402 Similar compositional relationships have been documented from the central Andes. For 403 example, the Pliocene La Pacana caldera system on the Altiplano-Puna plateau includes two, large-404 volume ignimbrites. The older Toconao unit is crystal-poor and rhyolitic (76-77% SiO<sub>2</sub>) whereas the 405 younger Atana ignimbrite is crystal rich and dacitic ( $SiO_2 - 66-70\%$ ) but contains glass that is rhyolitic 406 (Lindsay et al., 2001). Another example is provided by ignimbrites from the Cerro Galán caldera on the 407 eastern edge of the Puna plateau (Folkes et al., 2011). Pumice clasts separated from ignimbrite, which 408 are interpreted as erupted crystal mushes, are composed of matrix glass, plagioclase, biotite, and quartz, 409 with lesser amounts of hornblende, sanadine, oxides, and accessory phases. Pumice clasts are 410 rhyodacites, with  $SiO_2$  of most samples between 68.5 and 70.5%. Matrix glass has higher silica, up to 411 81%, indicating that interstitial liquid was highly evolved (Folkes et al., 2011). Al-in-hornblende 412 barometry suggests that crystals formed in magma chambers at both intermediate (14-18 km) and 413 shallow (6-10 km) depths. This finding is consistent with other workers who have described a model 414 with multiple levels of magma accumulation (e.g., de Silva et al., 2006; de Silva and Gosnold, 2007;

415 Kay et al., 2010). The lowest level is at the base of the crust where mantle melts pond, induce lower

416 crustal melting, and differentiate to produce intermediate magmas. These intermediate magmas rise to a

417 mid-crustal level and fractionate further. Finally, the evolved melts ascend to shallow levels to form

418 silicic sills that feed eruptions. A sill-like magma chamber geometry impedes sidewall crystallization

and allows buoyant rhyolitic melt to segregate at the top of the chamber (De Silva and Wolff, 1995;

420 Christiansen, 2005; Bachmann and Bergantz, 2008). The bulk of the magma chamber is composed of

421 silicic cumulates and trapped rhyolitic liquid (Lee and Morton, 2015).

422 In this model, the Wyoming batholith represents an exposed example of a solidified, shallow 423 silicic magma body that may have fed voluminous felsic ignimbrites like those that blanket the Great 424 Basin and Altiplano-Puna plateau. The  $\sim 200$  km extent of the Wyoming batholith supports this 425 interpretation: it is comparable to the inferred size of the plutonic complexes supplying large volume 426 silicic ignimbrites. The size of the batholith supplying the Altiplano-Puna ignimbrites is inferred from 427 the extent of a low velocity zone at depths of 4-25 km below sea level. Seismic images of the 428 Altiplano-Puna crust by Ward et al. (2014) determined that this low velocity zone, known as the 429 Altiplano-Puna magma body, exceeds 200 km in diameter. Lipman and Bachmann (2015) interpret two 430 large negative gravity anomalies along the axis of the SRMVF image subvolcanic batholiths, each 431 approximately 100 km in diameter. The size of the batholith supplying the Great Basin ignimbrites is 432 less well-known. If the locations of calderas are interpreted to encompass the minimum size of the 433 batholith beneath, and if corrections are made for subsequent Basin and Range extension (40-50% in 434 this area; Best et al., 2013b), then it appears that the batholith (or batholiths) beneath the Great Basin 435 altiplano were also approximately ~200 km in diameter. 436 Lacking samples of the batholiths that supplied the Cenozoic ignimbrites it is not possible to 437 ascertain that the compositions of those plutonic rocks are equivalent to the Wyoming batholith. 438 However, inferences can be made by comparing the composition of the Wyoming batholith to 439 ignimbrites erupted from systems inboard of continental magmatic arcs. Ignimbrites from both the 440 Altiplano-Puna and western USA ignimbrites fields are dominated by crystal-rich dacite and 441 rhyodacite. In both locations, the volcanic systems also erupted high-K rhyolite ignimbrites, though 442 these are volumetrically minor (De Silva and Gosnold, 2007; Best et al., 2013b). And in both areas, 443 glass compositions identify the presence of rhyolitic liquid at depth. The ignimbrite compositions 444 encompass those of the Wyoming batholith granites. To illustrate, Figure 12 compares Wyoming 445 batholith granites to the ignimbrites of Cerro Galán, which produced magmas with nearly identical

geochemistries for over 3.5 Ma (Folkes et al., 2011). The K<sub>2</sub>O and silica contents of the Cerro Galán

446

447	rhyodacite pumice clasts are slightly lower than the Wyoming batholith, and analyses of rhyolite glass
448	are overlapping to higher (Fig. 12). We conclude that if it had erupted, the Wyoming batholith could
449	conceivably have produced a similar spectrum of ignimbrite compositions by eruption of segregated
450	rhyolitic liquid and of dacitic crystal mush.
451	
452	Implications
453	We suggest that the Wyoming batholith is best interpreted as an upper-crustal silicic plutonic
454	system generated in a continental arc setting (Fig. 13). The plutonic roots of the continental arc are
455	represented by the 2.63 Ga Louis Lake batholith, a magnesian, calc-alkalic composite arc exposed in
456	the Wind River Range. The Louis Lake batholith intruded older Archean crust and juvenile graywackes
457	interpreted as an accretionary prism (Frost et al., 1998; Frost et al., 2006). The Louis Lake batholith
458	crystallized at between 3 and 6 kb (10 to 20 km; Frost et al., 2000). It is a composite batholith
459	composed of a wide range of rock types from gabbro to granite, and the youngest intrusion, the Bears
460	Ears pluton, is geochemically indistinguishable from the Wyoming batholith.
461	The Wyoming batholith is a large body of true granite. Except for slightly more leucocratic
462	margins, it exhibits no apparent zoning or heterogeneity despite being exposed at different structural
463	levels. These characteristics are consistent with a sill-shaped magma chamber, where the rise of
464	buoyant evolved liquids and sinking of dense cumulates will promote convective stirring and mixing of
465	liquids and crystals (Christiansen, 2005). This magma chamber is envisioned as the upper-crustal
466	product of a distillation process involving differentiation at various depths (Fig. 13; Bachman and
467	Bergantz, 2008; Lee and Morton, 2015).
468	Although the Wyoming batholith is chemically homogeneous, its initial Nd isotopic
469	composition is not uniform. In the western end of the batholith, the initial Nd isotopic composition
470	matches the Bears Ears pluton, suggesting that similar magma sources supplied the Louis Lake
471	batholith and western Wyoming batholith. The less radiogenic $\epsilon_{Nd}$ of the batholith exposed in the

472 Shirley and Laramie Mountains requires a higher proportion of Archean crustal sources or involvement 473 of a less radiogenic crustal source. This isotopic variability implies that although vertical stirring was 474 effective, the batholith is less well-homogenized horizontally. Moreover, it suggests that the batholith 475 was supplied by multiple conduits that tapped intermediate magma chambers of varying Nd isotopic 476 composition (Fig. 13). 477 The characteristics of the large silicic batholiths associated with continental arcs have been 478 inferred from the compositional and temporal record of their voluminous ignimbrite products 479 (Christiansen, 2005; de Silva and Gosnold, 2007), and in the case of the Altiplano-Puna volcanic field, 480 from the seismic evidence of the magma system beneath (Ward et al., 2014). We interpret the 481 Wyoming batholith as an exposed example of one of these batholiths, and as such it provides an 482 important opportunity for direct study of the processes producing large volumes of siliceous magma. 483 As noted above, potassic granites are a common component of every Archean craton, and are 484 the second most widespread lithology in Neoarchean terranes after the tonalite-trondhjemite-485 granodiorite association (Laurent et al., 2014). Like the Wyoming batholith, these Neoarchean granites, 486 which have high silica contents (typically 70-76% SiO<sub>2</sub>), are not associated with intermediate and 487 mafic phases (Laurent et al., 2014). Extensive areas composed exclusively of granite are notable 488 feature of Neoarchean terrains. For example, the Turfloop and Lekkersmaak biotite granites dominate 489 the southern Pietersburg block of the Kaapvaal province (Laurent et al., 2014). Individual granite suites 490 in the Slave Province occupy fully 60% of a 130 km x 110 km map sheet (Davis et al., 1994) and 491 together they extend across the entire craton (Davis and Bleeker, 1999). Neoarchean potassic granite 492 occupies 20% of the area of the Yilgarn craton, and is distributed throughout the entire craton (Cassidy 493 et al., 2006). The formation of these granites seems to require a large power input from the mantle 494 applied across a broad area. Slab breakoff, retreat, and/or lithospheric delamination have been proposed 495 to expose the base of the crust to heat from the mantle that ultimately is responsible for extensive 496 granite magmatism (Laurent et al., 2014; Davis and Bleeker, 1999; Davis et al., 1994). The late

497	Archean Earth may present optimal conditions for the formation of large volumes of silicic magma. By
498	this time the majority of the planet's continental crust had formed (e.g. Hawkesworth et al., 2010),
499	providing the environment in which differentiation, distillation, and assimilation could occur. In
500	addition, the Earth's internal radioactive heat production was still relatively high, providing the power
501	to drive these processes.
502	
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508	
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#### 716 Figure Captions

35.

- Figure 1. Location of the Wyoming province and subprovinces, and the project area of this
- 718 investigation (blue dashed line).
- Figure 2. Map of central Wyoming showing extent of Archean outcrop, including 2.62-2.63 Ga granitic
- rocks of the Wyoming batholith and Louis Lake batholith. The largest of the other ~2.62 Ga

721 granites is the Bears Ears pluton, which intrudes the Louis Lake batholith. Laramide basement-722 involved thrust faults brought the Archean rocks to the present-day erosional surface. 723 Subsequent Cenozoic normal faulting down-dropped the Granite and Shirley Mountains a 724 minimum of 700 meters (Love, 1970). 725 Figure 3. Field photographs of Wyoming batholith rocks. (a) Lankin Dome, central Granite Mountains, 726 is composed of biotite granite at the base and is capped by leucocratic banded granite. The 727 contact is sharp along the west end (red arrow) and feathered in the center and east end (black 728 arrow) of Lankin Dome. The summit is at 2354m, 420 m above the valley floor. (b) View of the 729 southern face of Lankin Dome taken from the base of the outcrop. The boundary between 730 leucocratic banded granite (top) and the biotite granite (bottom) is interfingered (note people 731 below the shadowed horizontal ledge in central-left of image for scale). (c) Contact between 732 leucocratic banded and biotite granites as viewed from near the contact on the southern face of 733 Lankin Dome. The foreground shows the coarse texture of the leucocratic banded granite. The 734 black arrow indicates the scalloped nature of the boundary with the biotite granite. (d) Unnamed 735 peak in the central Granite Mountains near Sweetwater Station, showing incorporation of 736 leucocratic banded granite within biotite granite. The large, folded enclave of leucocratic 737 banded granite is interpreted to have detached from the batholith roof and been deformed 738 during magmatic flow along with other smaller and wispy inclusions of the leucocratic banded 739 granite. 740 Figure 4. Photomicrographs of biotite granite samples 11SMG2, 13SM05, and 12GM04 are shown in 741 images a, b, and c. (a) Quartz has undergone deformation and grain-boundary migration. 742 Plagioclase is weakly sericitized. Epidote and chlorite are secondary after biotite. (b) 743 Microcline is perthitic. Biotite commonly contains inclusions of zircon and magnetite. 744 Although not visible in the photomicrograph, rutile needles are common as inclusions within 745 quartz. (c) Finer-grained biotite granite sample showing serated grain boundaries formed by

746	subsolidus grain boundary migration. Potassium feldspar may also have undergone high
747	temperature grain-size reduction. Photomicrographs of leucocratic banded granite samples
748	12SMG9, 13LR02, and 12SMG3 are shown in images d, e, and f. (d) Graphic granite texture is
749	fairly common, as shown in this photomicrograph. (e) Plagioclase is sericitized and biotite has
750	undergone deuteric alteration. Magnetite is associated with biotite. (f) Grain-size of quartz and
751	feldspars varies widely in leucocratic banded granite samples; this photomicrograph is
752	dominated by a large quartz crystal. A small crystal of muscovite is present in the upper part of
753	the photomicrograph.
754	Figure 5. (a) Concordia diagram for zircon from leucocratic banded granite sample 10LD2. The
755	$^{207}$ Pb/ $^{206}$ Pb age of the concordant analysis, 2627 ± 3 Ma (1 $\sigma$ ), is interpreted as the best estimate
756	for the age of this sample. (b) Concordia diagram for zircon from biotite granite sample
757	11SMG1. 5 analyses of zircon with U < 600 ppm define a mean weighted age of $2623 \pm 4.3$
758	Ma (1 $\sigma$ ), which is interpreted as the best estimate of the crystallization age of this sample. (c)
759	Concordia diagram for zircon from hydrothermally altered biotite granite sample 11SMG2. The
760	two grains with the lowest uranium contents are shown in red. The most concordant analysis (-
761	1% discordant) yielded a $^{207}$ Pb/ $^{206}$ Pb age of 2624 ± 4 Ma (1 $\sigma$ ), which is interpreted as the best
762	estimate for the age of this sample.
763	Figure 6. Plots of (a) Fe index [FeO <sup>tot</sup> /(FeO <sup>tot</sup> + MgO)], (b) modified alkali-lime index (MALI; Na <sub>2</sub> O +
764	K <sub>2</sub> O – CaO), and (c) aluminum saturation index (ASI; molecular ratio Al/Ca – 1.67P + Na + K)
765	for the Wyoming and Louis Lake batholiths. The Wyoming batholith has a restricted range of
766	silica, and is mainly magnesian, calc-alkalic to alkali-calcic, and peraluminous. No geochemical
767	variations by geographic location are apparent. By contrast, the Louis Lake batholith includes
768	intermediate to mafic components and is calc-alkalic and metaluminous. Data from Table 1;
769	Condie, 1969; Stuckless and Peterman, 1977; Stuckless and Miesch, 1981; Frost et al., 1998;

770	Frost et al., 2000; Wall, 2004; and Meredith, 2005. Boundaries of fields are from Frost et al.
771	(2001) and Frost and Frost (2008).
772	Figure 7. Variation diagrams of SiO <sub>2</sub> versus (a) K <sub>2</sub> O, (b) Rb, (c) Sr, and (d) Eu. Incompatible elements
773	K and Rb increase modestly with silica in the Louis Lake batholith. Wyoming batholith and
774	high silica granites of the Louis Lake batholith are high-K. They exhibit a four- to five-fold
775	range in K <sub>2</sub> O and Rb indicating that potassium feldspar and biotite are late-crystallizing phases.
776	Compatible elements Sr and Eu show the opposite trends, and very low contents in the
777	Wyoming batholith and high-silica components of the Louis Lake batholith. Symbols and data
778	sources as in Fig. 7.
779	Figure 8. Diagrams of (a) Nb versus Y, and (b) Nb + Y versus Rb for samples of Wyoming and Louis
780	Lake batholith rocks, after Pearce et al. (1984). Most samples are within the volcanic arc granite
781	(VAG) field. Cumulate hornblende likely contributes to higher Nb and Y in Louis Lake
782	batholith samples. COLG = collisional granites, ORG = ocean ridge granites, WPG = within
783	plate granites. Symbols as in Fig. 7. Data from Table 1; Frost et al., 1998; Frost et al., 2000; and
784	Wall, 2004.
785	Figure 9. Chondrite-normalized rare earth element diagrams for Wyoming batholith samples from the
786	Granite Mountains (a), Shirley Mountains (b), and Laramie Mountains (c). Rare earth patterns
787	for the Louis Lake batholith and Bears Ears pluton of the Wind River Range are shown in (d).
788	Data sources: Table 1; Frost et al., 1998; and Stuckless, 1989.
789	Figure 10. Initial $\epsilon_{Nd}$ of Wyoming batholith samples compared to Louis Lake batholith and Bears Ears
790	pluton samples. All samples were intruded at ~2.62-2.63 Ga; the Louis Lake and Bears Ears are
791	displaced to the left of their intrusive age for clarity. Also shown are potential magma sources,
792	including depleted mantle, ~3.3-3.4 Ga crust of the Sacawee block immediately north of the
793	Wyoming batholith in the Granite Mountains, evolved metasedimentary rocks of the Granite
794	Mountains, and juvenile metasediments of the southern accreted terranes in the Wind River

795	Range and Granite Mountains. The Wyoming batholith in the Shirley and Laramie Mountains
796	contains less radiogenic Nd isotopic ratios, suggesting that these magmas incorporated a greater
797	proportion of old crust than the granites of the Granite Mountains and Wind River Range. Data
798	sources: Table 2, Frost et al., 2006; Fruchey, 2002; Wall, 2004; and Meredith, 2005.
799	Figure 11. Nd isotopic compositions and locations within the Wyoming batholith. Nd isotopic
800	compositions are grouped into three categories: red circles = samples with initial $_{Nd}$ < -3.4;
801	pink circles = samples with intermediate values of $N_{d}$ between -3.4 and -1.0; yellow circles =
802	samples with initial $_{Nd} > -1.0$ . Data from Table 2 and Wall, 2004.
803	Figure 12. Plot of K <sub>2</sub> O versus silica comparing the Wyoming batholith with eruptive products of Cerro
804	Galán caldera, northwest Argentina (Folkes et al., 2011). The small black dots represent
805	analyses of ignimbrite glass and the field encompasses compositions of pumice clasts. Data
806	sources as in Fig. 7 and Folkes et al., 2011.
807	Figure 13. Schematic east-west cross-section through the central Wyoming province in the
808	Neoarchean, depicting relative depth of emplacement of the Wyoming batholith and the Louis
809	Lake/Bears Ears composite batholith. The Louis Lake batholith, which was emplaced at 3-6 kb
810	(Frost et al., 2000), intrudes juvenile accreted terranes along its southern boundary. Nd isotopic
811	evidence suggests that similar magma sources supplied the Louis Lake batholith and western
812	Wyoming batholith but a larger proportion of evolved Archean crustal sources contributed to
813	the eastern part of the Wyoming batholith. BEP = Bears Ears pluton.









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Figure 7



# Figure 8





Figure 9





Figure 11



0/



Rocks within the 2.62 to 2.63 Ga magmatic domain



Initial  $\mathcal{E}_{Nd}$  ranges • < -3.4 • -3.4 > -1.0 • > -1.0









				Gra	anite Mounta	ains						Shi	rley Mounta	ins				Laramie N	lountains	
	10LD-2	11SMG1	11SMG2	12SMG7	12GM01	12GM02 1	L2GM03	12SMG9	04TC17*	12GM04	13SM02	13SM05	13SM06	13SM07	13SM03B	13SM03A	13LR01	13LR03	13LR02B	13LR06
SiO <sub>2</sub>	75.02	75.07	7 72.99	71.30	76.30	71.91	75.40	74.10	75.50	75.30	73.60	73.90	74.90	72.40	74.00	74.30	77.70	74.70	73.50	73.50
TiO <sub>2</sub>	0.06	0.26	5 0.24	0.09	0.16	0.37	0.08	0.10	0.10	0.12	0.32	0.18	0.11	0.17	0.14	0.14	0.03	0.15	0.22	0.18
$Al_2O_3$	14.58	14.42	l 13.67	14.55	13.45	14.40	12.90	14.45	13.75	14.05	13.60	14.45	13.45	14.20	13.70	14.55	12.95	13.80	13.40	13.75
$Fe_2O_3^{total}$	0.51	0.62	l 1.21	1.40	0.81	3.06	0.71	1.16	1.26	1.16	2.98	1.54	0.96	1.76	1.60	1.62	0.25	1.66	2.24	0.85
MnO	0.00	0.02	1 0.02	0.01	0.02	0.02	0.01	0.01	0.02	0.01	0.03	0.02	0.01	0.02	0.03	0.02	0.01	0.03	0.02	0.01
MgO	0.23	0.33	3 0.26	0.25	0.28	0.66	0.21	0.30	0.26	0.26	0.74	0.31	0.17	0.43	0.52	0.44	0.09	0.36	0.45	0.19
CaO	0.59	2.17	7 0.98	3 1.10	0.81	1.95	0.73	1.36	0.82	1.46	1.95	1.29	0.87	1.13	0.33	0.37	0.13	1.04	0.22	2.50
Na <sub>2</sub> O	3.88	4.27	7 3.20	3.45	3.62	3.75	3.48	3.33	3.80	3.68	3.10	3.60	3.06	3.42	2.92	4.52	2.77	3.40	2.74	4.55
K <sub>2</sub> O	5.60	2.54	1 5.37	5.59	4.99	3.73	4.74	5.62	4.48	4.65	3.35	5.18	5.71	5.11	6.12	4.01	6.91	4.88	5.93	2.27
$P_2O_5$	0.04	0.07	7 0.07	0.05	0.02	0.15	0.04	0.10	0.07	0.03	0.12	0.07	0.02	0.05	0.04	0.04	0.02	0.04	0.08	0.07
Total	100.52	99.73	3 98.01	. 97.79	100.46	100.00	98.30	100.53	100.06	100.72	99.79	100.54	99.26	98.69	99.40	100.01	100.61	98.40	96.56	97.02
Fe-index	0.66	0.63	0.81	0.83	0.72	0.81	0 75	0 78	0.81	0.80	0.78	0.82	0.84	0 79	0 73	0.77	0 7142857	0 8058252	0 8175182	0 8010471
ΜΔΠ	0.00 8 89	4.6	5 7.60	. 0.03 ) 7.94	7.80	5 53	7 49	7 59	7.46	6.87	4 50	7 49	7 90	7.40	8 71	8 16	9.4709178	7 2476825	8 5720662	<i>A A</i> 178555
ASI	1.05	1.0 1.0	5 1.00	, ,, 5 1.06	1.00	1.06	1.06	1.04	1 10	1.03	1 12	1.45	1.05	1.40	1 14	1 17	1 0575032	1 0853025	1 193487	0 9548466
	1.00	2100		1.00	1.05	1.00	2.00	1.01	1.10	1.00		1.00	1.05	1.00		1.17	1.0070002	1.0033023	11100 107	0.00 10 100
Rb ppm	266.4	64	1 157.5	5 126.5	174	129.5	160	128	144.5	136	132	191.5	171	187.5	243	183	219	177	223	75.6
Sr ppm	65.9	240.2	2 137.9	192.5	99.4	246	130.5	233	89.1	191	142.5	166.5	193.5	124	106	115.5	107.5	142.5	87.1	335
Zr ppm	64.8	143.4	1 141	. 85	93	306	112	105	89	108	194	156	58	175	118	124	107.5	142.5	87.1	335
Y ppm	25.5	9.7	7 9.6	6 8.3	11.1	11.3	8.1	10.4	11.6	11.9	8.3	12.4	3.4	17.2	16.6	6.3	12.3	10.5	16.4	19.9
Nb ppm	6.8	3.6	5 4.9	6.4	10.3	10.6	5	6.7	7.8	9	9.6	9.8	4	8.7	14.9	4.4	3.1	11	20.8	15.4
La ppm	21.6	i		46.1	38.6	110.5	52.4	23.2	35.2	50.0	33.6	57.3	23.5	95.2	26.7	7.5	2.0	26.9	87.8	73.8
Ce ppm	43.2			93.2	72.7	204.0	100.5	42.7	63.9	89.3	83.2	116.0	46.2	185.5	61.8	16.5	4.6	94.1	184.5	147.5
Pr ppm	4.5			10.5	7.6	18.9	10.1	5.2	6.4	8.9	6.3	11.9	4.3	16.9	4.7	1.3	0.4	6.4	17.7	14.7
Nd ppm	15./			37.3	25.3	62.7	32.3	18.1	21.2	29.5	21.9	40.1	13.8	54.8	15./	4.3	2.0	22.6	56.5	47.0
Sm ppm	3.4			/.3	4.2	9.1	4.9	3.6	4.0	5.4	3.8	7.2	2.0	9.4	3.1	1.0	0.7	4.5	9.5	7.9
Eu ppm	0.3			0.7	0.6	0.8	0.7	0.7	0.4	0.7	1.0	0.7	0.6	0.6	0.3	0.3	0.5	0.6	0.7	0.9
Ga ppm	2.8			4.7	2.6	5.7	2.9	2.7	3.0	3.6	2.7	4.1	1.2	6.1	2.9	0.8	1.2	3.3	6.1	5.9
i b ppm	0.5			0.5	0.3	0.6	0.3	0.4	0.4	0.5	0.4	0.5	0.2	0.8	0.4	0.2	0.3	0.4	0.8	0.8
Dy ppm	3.8			2.2	1.9	2.6	1.7	1.9	2.3	2.4	1.9	2.2	0.8	3.2	2.7	0.9	2.0	2.0	3.7	4.3
Ho ppm	0.8			0.3	0.4	0.4	0.3	0.4	0.4	0.4	0.4	0.4	0.1	0.6	0.6	0.2	0.5	0.4	0.6	0.8
Er ppm	2.7			0.7	1.1	0.9	0.8	1.0	1.1	1.2	0.7	1.2	0.4	1.5	1.7	0.7	1.4	1.1	1.6	1.9
im ppm	0.4			0.1	0.2	0.1	0.1	0.1	0.2	0.1	0.1	0.2	0.1	0.2	0.3	0.1	0.2	0.1	0.2	0.3
YD ppm	2.6			0.4	1.1	0.7	0.6	0.8	0.8	1.0	0.6	1.1	0.3	1.2	1.5	0.7	1.3	1.0	1.2	1.6
Lu ppm	0.4	•		0.1	0.2	0.1	0.1	0.1	0.1	0.1	0.1	0.2	0.1	0.2	0.2	0.1	0.2	0.2	0.2	0.3

Table 1. Geochemistry of Neoarchean granite of the Wyoming batholith

\*major element and Rb, Sr, Zr, Y and Nb data from Meredith (2005)

Major and minor element data obtained by ICP-AES and REE data by ICP-MS at ALS Minerals.

Sample locations are shown on Fig. 1 and UTM provided in Bagdonas (2014).

Fe-index, MALI, and ASI defined in Frost et al. (2001)

Table 2. Sm-Nd isotopic data for the Wyoming batholith.

Sample	Uplift	Sm ppm	Nd ppm	<sup>147</sup> Sm/ <sup>144</sup> Nd	<sup>143</sup> Nd/ <sup>144</sup> Nd	<sup>143</sup> Nd/ <sup>144</sup> Nd at 2625 Ma	$\epsilon_{\text{Nd}}$ at 2625 Ma
04TC17	Granite Mountains	3.6643	19.8219	0.11175	0.51107	0.509135	-1.90
10LD2	Granite Mountains	2.9180	13.7863	0.12796	0.51148	0.509269	0.72
12GM01	Granite Mountains	3.9440	24.4931	0.09734	0.51088	0.509197	-0.70
12GM02	Granite Mountains	9.0222	60.7883	0.08971	0.51062	0.509064	-3.31
12GM03	Granite Mountains	4.2724	29.1711	0.08853	0.51070	0.509165	-1.32
12SMG7	Granite Mountains	7.0978	35.4693	0.12098	0.51139	0.509294	1.21
12SMG9	Granite Mountains	3.5159	17.4107	0.12208	0.51124	0.509122	-2.16
12GM04	Shirley Mountains	4.3180	24.4961	0.10655	0.51079	0.508944	-5.65
13SM02	Shirley Mountains	3.6425	20.9785	0.10495	0.51072	0.508900	-6.51
13SM03A	Shirley Mountains	0.8204	4.1118	0.12061	0.51114	0.509056	-3.45
13SM03B	Shirley Mountains	2.7920	14.8652	0.11354	0.51098	0.509011	-4.34
13SM05	Shirley Mountains	6.5575	39.3692	0.10068	0.51082	0.509073	-3.13
13SM06	Shirley Mountains	2.1609	14.9477	0.08738	0.51069	0.509179	-1.04
13SM07	Shirley Mountains	5.1969	31.4543	0.09987	0.51080	0.509074	-3.10
13LR02B	Laramie Mountains	9.5015	57.1816	0.10044	0.51083	0.509094	-2.71
13LR03	Laramie Mountains	4.5403	22.6236	0.12131	0.51094	0.508837	-7.75
14LR1/LR94-7	Laramie Mountains	7.71975	54.31734	0.085904355	0.5105583	0.509071	-3.17

Sm and Nd were isolated from whole rock powders using methods described in Frost et al. (2006). Isotopic ratios were obtained on a Neptune MC-ICP-MS at the University of Wyoming. Samples were normalized to  $^{146}$ Nd/ $^{144}$ Nd = 0.7219. La Jolla Nd analyzed after every five unknowns gave  $^{143}$ Nd/ $^{144}$ Nd = 0.51185 ± 0.000012 (2 s.d.)