1 Revision 2

2	Interpretation of the infrared spectra of the lizardite-nepouite series in the near and middle
3	infrared range
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9	Abstract
10	A series of 1:1 silicate clays of the lizardite-nepoute series ($Si_2Mg_{3-x}Ni_xO_5(OH_4)$) with x =
11	0, 0.5, 1, 1.5, 2, 2.5, and 3) was synthesized at 220°C during 7 days from coprecipitated gels in
12	hydrothermal conditions. A clear relationship was evidenced between the d(06-33) and the Ni/Mg
13	ratio of the synthesized samples following a Vegard's law and suggested a rather random
14	distribution of octahedral cations. For the first time, infrared spectra of this series were given in
15	both near and mid infrared spectral regions (250 - 7500 cm ⁻¹). Notably, the bands due to the OH
16	stretching vibrations and those of their first overtones in the lizardite-nepouite series were
17	attributed. The combination bands observed in the near infrared region for both end-members
18	could be attributed thanks to combinations of two or three mid infrared features. Some of the
19	observed combination bands are clearly linked to combination of different vibrational groups.
20	Infrared spectroscopy is simple to use and is a powerful tool to study the crystal-chemistry
21	of garnierites. More broadly, the improvement of band attributions especially in near infrared
22	contributes to develop the infrared analyses in field geology and remote sensing.
23	Keywords: lizardite, nepouite, infrared spectroscopy, near infrared, middle infrared, synthesis,
24	nickel, clay minerals, serpentine, phyllosilicates, garnierite

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Introduction

Nickel ore deposits are found in lateritic or sulphide forms. The major part of produced Ni 26 comes from sulphide deposits (60 %) but Ni-laterite ores make up 60 to 70 % of the world's Ni 27 resources (Butt and Cluzel 2013). Ni concentration in laterites results from an intensive 28 weathering of ultramaphic rocks or their serpentinized equivalents under tropical or rainforest 29 conditions like in New Caledonia (Wells et al. 2009) or Brazil (Mano et al. 2014). These deposits 30 are divided in two ore types, an oxidized one and a hydrous silicate type currently named 31 "Garnierite" (Faust 1966; Trescases 1975, 1979). The garnierite group is a general name for an 32 intimate mixture of hydrous Ni-Mg silicates that commonly includes two or more of the 33 following minerals: serpentine, talc, smectite and chlorite (Faust 1966; Brindley and Hang 1973; 34 Springer 1974; Brindley et al. 1979; Gleeson et al. 2004). Ni-bearing minerals in the serpentine 35 group are mainly lizardites and chrysotiles, whose Ni-analogs are nepouite and pecoraite. 36 Nepouite was first described by Glassere (1907) as the Ni-lizardite analog, and Maksimovich 37 (1973) showed the occurrence of the lizardite-nepouite series. The phyllosilicates belonging to 38 the lizardite-nepouite series are generally mixed with other minerals in garnierite. The study of 39 the crystal chemistry of these natural phyllosilicates is then uneasy. Moreover, as far as the 40 authors are aware, spectral data concerning this series are not available in literature, notably in 41 the near infrared (NIR) range (4000 - 7000 cm⁻¹) although NIR field spectrometers could widely 42 be used by Ni-mining company for exploration. Field spectrometers are commonly used to 43 distinguish mineral compositional variations which may vary with the composition of altering 44 fluids, temperature and composition of host rocks in an alteration system. Mapping the 45 mineralogical variations at both geological scale with field/airborne/spaceborn infrared 46 spectrometer (Bowen et al. 2007; Di Tommaso and Rubinstein 2007; Chen et al. 2007; 47 Brandmeier et al. 2013) using NIR data and at the microscopic scale with infrared 48

49 microspectroscopy (IRMS) (Robin et al. 2013) using MIR data, contributes to improve the 50 understanding of the functioning of an alteration system and to determine the 51 mineralization/alteration relationships. In order to map mineralogical variations in lateritic Ni ore 52 deposit, spectral data in NIR range on the nepouite-lizardite series must be known. And 53 generally, even if spectral data exist for clay minerals, any band interpretations in NIR range are 54 made.

55 Therefore samples from the lizardite-nepouite series were synthesized to obtain pure 56 reference samples for a Fourier transform infrared (FTIR) study, in both NIR and mid infrared 57 (MIR) regions and band attributions were made.

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Experimental

60 Clay synthesis

Samples of the lizardite-nepouite series were synthesized using a coprecipitated gel as starting material as in Baron et al. (2016). Coprecipitated gels of $Si_2Mg_{3-x}Ni_xO_7$ theoretical composition with x = 0, 0.5, 1, 1.5, 2, 2.5, and 3 were obtained by mixing solutions of sodium silicate, Mg and/or Ni chloride and sodium hydroxide (reaction 1).

65 (1)
$$2 \operatorname{SiO}_2\operatorname{Na}_2O + (3-x)\operatorname{MgCl}_2 + x\operatorname{NiCl}_2 + 2\operatorname{NaOH} \rightarrow \operatorname{Si}_2\operatorname{Mg}_{3-x}\operatorname{Ni}_xO_7 + 6\operatorname{NaCl} + \operatorname{H}_2O$$

After precipitation, the gel was collected by filtration and gently washed with deionized water to remove sodium chloride and then dried at 60°C for 48 hours before being crushed. 0.5 g of the powdered gel were placed into Teflon coated metallic bombs with 30 ml of deionized water and heated at 220°C during one week under equilibrium vapor pressure. The solid was separated from the solution by filtration and dried overnight at 60°C.

71	A synthetic Ni-Mg kerolite was also used in this study for comparison. The synthesis was
72	done using the same way than for lizardite-nepouite series from a coprecipitated gel with
73	$Si_4Mg_{1.5}Ni_{1.5}O_{11}$ theoretical composition obtained following the reaction (2):
74	(2) $4 \operatorname{SiO}_2\operatorname{Na}_2\operatorname{O} + 1.5\operatorname{MgCl}_2 + 1.5\operatorname{NiCl}_2 + 2\operatorname{HCl} \rightarrow \operatorname{Si}_4\operatorname{Mg}_{1.5}\operatorname{Ni}_{1.5}\operatorname{O}_{11} + 8\operatorname{NaCl} + \operatorname{H}_2\operatorname{O}$
75	Analysis
76	X-ray diffraction (XRD) powder patterns were performed with a Bruker D8 advance
77	diffractometer (CuK α radiation, 40 kV and 40 mA) over the range 2 - 65°2 θ , with 0.025°2 θ step
78	size and 0.6 seconds per step. The reflection (06-33) was recorded with a step of $0.025^{\circ}2\theta$ and a
79	counting time of 4 seconds over the range 57 - $63^{\circ}2\theta$. The position of the (06-33) reflection was
80	determined by fitting XRD patterns with a pseudo-Voigt component in the Fityk software
81	(Wojdyr 2010).
82	A Magna-IR 760 Nicolet spectrometer was used for mid infrared spectra over the range
83	400 - 4000 cm ⁻¹ , with an EverGlo source, a KBr beam splitter, a DTGS-KBr detector and a

400 - 4000 cm⁻¹, with an EverGlo source, a KBr beam splitter, a DTGS-KBr detector and a resolution of 4 cm⁻¹ with co-additions of 100 scans. Over the range 200 - 500 cm⁻¹, a Thermo Scientific Nicolet 6700 FT-IR spectrometer was used with an EverGlo source, a CsI beam splitter, a DTGS-CsI detector and a resolution of 4 cm⁻¹ with the co-additions of 100 scans. MIR spectra were obtained in transmission through KBr pellets, prepared with 1 mg of sample and 150 mg of KBr salt. The mixture was pressed during 5 minutes at 8 kbar and dried overnight in an oven at 110°C.

NIR spectra (3850 - 7500 cm⁻¹) were acquired with a Thermo Scientific Integrating
Sphere (diffuse reflectance) equipped with an internal InGaAs detector, set on a Thermo
Scientific Nicolet 6700 FT-IR spectrometer. A white light source was used, associated with a

 CaF_2 beam splitter. The resolution was set at 4 cm⁻¹ with the co-additions of 100 scans. Powder samples were analyzed directly on the integrating sphere without dilution.

The second derivative was carried out using the Savitzky–Golay filter (Savitzky and Golay 1964) in the Omnic[™] software with a set of 13 points and a polynomial order of 3. The second derivative was used to determine the position of NIR and MIR bands, and was given in inverted form in this study to facilitate the reading of spectra.

The deviation in band position possibly induced by the second derivative method was estimated from a mathematical model. The model generated theoretical spectra by varying the relative intensities of 2 pseudo-Voigt components, with a 4 cm⁻¹ resolution (as for experimental spectra). The second derivative of these theoretical spectra was carried out as described above for experimental spectra. The maximum deviation measured between the band positions introduced in the model and the band positions measured from the second derivative was +/- 3 cm⁻¹.

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Results and discussion

107 XRD data

XRD powder patterns of all samples exhibited an intense (001) reflection at 7.39 Å and a 108 (002) located at 3.67 Å for lizardite (x = 0) and at 3.64 Å for nepoute (x = 3), characterizing a 109 1:1 phyllosilicate structure. The (*hk*) reflections for lizardite (x = 0) and nepoute (x = 3) were 110 situated at 4.51 and 4.50 Å (020), 2.59 and 2.63 Å (200), 2.47 Å (201), and at 1.537 and 1.529 Å 111 (06-33), respectively (Fig. 1 and 2), revealing a trioctahedral structure. All (*hkl*) reflections were 112 in agreement with data reported in literature for natural lizardite and nepouite samples (Whittaker 113 and Zussman 1956; Mikheev 1957; Montoya and Baur 1963; Maksimovich 1973; Frondel and Ito 114 1975; Brindley and Brown 1980; Mellini 1982). XRD powder patterns revealed that no other 115 crystalline phases were present for all samples. 116

- 120 (Fig. 3) with the following regression line:
- 121 d(06-33) (Å) = -0.0026 x + 1.5374 (1)

This relation followed an apparent Vegard's law suggesting a random distribution of cations in the octahedral sheet of the synthesized samples. The progressive change of the d(06-33) along the lizardite-nepouite series is due to the atomic radii difference between Ni and Mg atoms in octahedral coordination (0.72 Å for Mg and 0.69 Å for Ni; Shannon 1976).

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127 MIR data

The OH-stretching zone. The spectrum of the synthetic lizardite (x = 0) was similar to the natural one (Serna et al. 1979; Fuchs et al. 1998; Balan et al. 2002), with a complex band at 3698 cm⁻¹ enlarged to 3688 cm⁻¹ (which will be discussed later), and a broad band at 3648 cm⁻¹ (Fig. 4). The former complex band at 3698 cm⁻¹ corresponded to the in-phase stretching of the inner-surface Mg₃-OH groups and the latter at 3648 cm⁻¹ to the two degenerate out-of-phase stretching modes of the inner-surface Mg₃-OH groups (Balan et al. 2002).

The increase of Ni contents in samples induced the progressive appearance of bands at 3686 cm⁻¹, 3668 cm⁻¹, and 3648 cm⁻¹. Finally, for nepouite (x = 3), only the band at 3648 cm⁻¹ was present with a broad band at around 3610 cm⁻¹ (Fig. 4). By analogy with the spectrum of the lizardite (x = 0) end-member, these two bands were assigned to the in-phase stretching of the inner-surface Ni₃-OH groups and to the two degenerate out-of-phase stretching modes of the inner-surface Ni₃-OH groups, respectively.

In trioctahedral phyllosilicates, each hydroxyl group is bonded to three octahedral cations 140 and the nature of these cations modifies the stretching vibration of hydroxyl groups. In the 141 142 lizardite-nepouite series, when Mg is partially replaced by Ni or conversely, four stretching OH (vOH) bands may be possibly distinguished due to the four possible cations combinations in the 143 three neighboring octahedral sites as it is well known for talc series: vMg_3-OH , $vMg_2R^{2+}-OH$, 144 $vMgR^{2+}$ 2-OH and vR^{2+} 3-OH where R^{2+} is the substituted cation (Wilkins and Ito 1967; Petit et al. 145 2004; Petit 2005, 2006; Madejová et al. 2011). Consequently, the two bands at 3686 and 3668 146 cm⁻¹ (Fig. 4) could be assigned to vMg₂Ni-OH and vMgNi₂-OH, respectively. 147

The 500-800 cm⁻¹ region. The spectrum of lizardite (x = 0) revealed three bands at 634, 148 607 and 562 cm⁻¹ (Fig. 4). These bands, characteristic of pure Mg-lizardite, were assigned, after 149 IR spectra calculation, to longitudinal (LO) and transversal (TO) modes of in-plane bending of 150 OH groups (Balan et al. 2002; Prencipe et al. 2009). Very often, natural lizardite samples contain 151 some Fe^{2+} or Ni^{2+} , and exhibit a broad band rather than these three distinguishable bands (Balan 152 et al. 2002). As a matter of fact, the three bands occurrence is a good criterion to easily determine 153 the Mg-purity of lizardite. For nepoute (x = 3), a broad and dissymmetric band at 674 cm⁻¹ with 154 a shoulder at 600 cm⁻¹ was observed. The intermediary samples (0.5 < x < 2.5) of the solid-155 156 solution exhibited a broad and complex multicomponent band (Fig. 4).

The lattice bands. Two main bands located at 984 and near 1080 cm⁻¹ (Fig. 4) were observed for all samples (Fig. 4). These bands were attributed to the two degenerate equatorial stretching of Si-O bonds and the symmetric stretch of apical Si-O bonds, respectively (Balan et al. 2002). The former band remained at 984 cm⁻¹ whatever the Ni content (x), whereas the latter band was at 1084 cm⁻¹ for lizardite (x = 0) and moves progressively to 1075 cm⁻¹ for nepoute (x = 3). A small shoulder at 1024 cm⁻¹ was observed for Mg-rich samples and disappeared with the increase of Ni content (x) (Fig. 4). This shoulder was attributed to a LO mode (Prencipe et al.

164 2009) and could not be attributed to small amount of chrysotile because it would be associated to
165 an intense band at 960 cm⁻¹ (Farmer 1974; Anbalagan et al. 2010; Madejová et al. 2011).

The lizardite (x = 0) spectrum exhibited bands at 454, 436, 400, 383, and 303 cm⁻¹ while the nepouite (x = 3) spectrum exhibited bands at 455, 426, 377, 361, and 318 cm⁻¹ (Fig. 4). The lizardite (x = 0) experimental spectrum was in agreement with the calculated spectra (Balan et al. 2002; Prencipe et al. 2009). For the intermediary samples (0.5 < x < 2.5), the bands moved progressively and were at intermediate position between those of the lizardite (x = 0) and nepouite (x = 3) end-members.

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173 NIR data

The occurrence of a broad and intensive OH-water combination band around 5205 cm⁻¹ (Hunt and Salisbury 1970) in all spectra (Supplementary file a), revealed the relatively high water content in samples, as it is very common for synthetic phyllosilicates. The high hydration state of synthetic phyllosilicates is generally due to their small particle size (Dumas et al. 2013).

The first OH overtone region. By analogy with the MIR vOH region, the bands at 7226, 7198, 7164 and 7124 cm⁻¹ (Fig. 5) were assigned to $2vMg_3$ -OH, $2vMg_2Ni$ -OH, $2vMgNi_2$ -OH and $2vNi_3$ -OH. As for the MIR region, the presence of four 2vOH bands with relative intensities varying continuously all along the solid-solution series argued for a rather random distribution of Ni and Mg atoms in the octahedral sheet of the samples. The lack of a band at 7236 cm⁻¹ confirmed the absence of chrysotile (Madejová et al. 2011).

The anharmonicity constant between the vOH and 2vOH wavenumbers of the lizarditenepoute series is $X = -85.6 \text{ cm}^{-1}$, that is the same value measured by Petit et al. (2004) for talcs and others phyllosilicates (Fig. 6). This relation gives confidence for the above NIR assignments.

The NIR spectrum of lizardite (x = 0), exhibited a shoulder at 7203 cm⁻¹ which could not 187 be attributed to the 2vMg₂NiOH overtone (Fig. 5). This band may be attributed to the overtone of 188 the shoulder at 3688 cm⁻¹ observed in the MIR spectrum according to the anharmonicity constant 189 value of -86.5 cm⁻¹. These bands may be due to vibrations of structural Mg₃-OH groups in 190 lizardite (x = 0). In theory, only the transverse mode (TO) is IR active for infinite crystals. 191 However when crystals exhibit one (or more) dimension much smaller than the wavelength of the 192 IR radiations (0.78 - 1000 µm), both the LO and TO modes are observed (Farmer 1998; Balan et 193 al. 2002; Prencipe et al. 2009). Therefore the wavenumber of the observed IR bands are strongly 194 dependent on the lizardite particles size. Similar lizardites were synthesized and exhibited 195 essentially small rounded particles with 40 nm in diameter and few lathed particles with 0.2 µm 196 in length (Mondésir and Decarreau 1987). So, for lizardite (x = 0), the bands at 3698 and 7226 197 cm⁻¹ and the shoulders at 3688 and 7203 cm⁻¹ possibly correspond to the LO-TO splitting of the 198

199 Mg₃OH groups (Farmer 1998; Prencipe et al. 2009).

The OH combination bands. In the 3850 - 4500 cm⁻¹ region, the lizardite (x = 0) spectrum exhibited two main bands located at 4305 and 4275 cm⁻¹ while for nepouite (x = 3), only one main asymmetric broad band was observed at 4280 cm⁻¹ (Fig. 5). In the 4500 and 5000 cm⁻¹ region (Fig. 5) two bands at 4780 and 4709 cm⁻¹ were observed in the lizardite (x = 0) spectrum, which shifted with increasing Ni contents to 4722 and 4650 cm⁻¹, respectively.

Origins of combination bands for lizardite (x = 0) and nepouite (x = 3) end members. The combination bands due to water (5000 – 5500 cm⁻¹) are well known for clay minerals (Cariati et al. 1981, 1983a; b). But the observed combination bands in the 3850 – 5000 cm⁻¹ range are generally not assigned, except for the combination bands arising from combination of stretching and bending structural OH groups for some clay minerals (Frost et al. 2002; Gates 2008; Bishop et al. 2008; Andrieux and Petit 2010; Madejová et al. 2011; Petit et al. 2015).

The wavenumber of observed combination bands was measured for both lizardite (x = 0)and nepouite (x = 3) end-members using the second derivative treatment of spectra (Fig. 7.a for lizardite (x = 0) and Fig. 7.b for nepouite (x = 3)). Generally in infrared spectroscopy, the signal treatment like the second derivative is more suitable to reveal the band position than the decomposition process, due to the relatively broad and complex bands (Gionis et al. 2006; Chryssikos et al. 2009). The 21 observed features (Fig. 7) were listed in Table 1 for lizardite (x = 0)and in Table 2 for nepoute (x = 3).

Absorption bands in the combination region result from the sharing of IR radiation energy 218 between two or more fundamental vibrations which are observed in MIR range (Ciurczak 2006). 219 Consequently, a second derivative treatment was realized on MIR spectra of each end-members 220 in order to reveal the position of MIR features (Fig. 8.a for lizardite (x = 0) and Fig. 8.b for 221 nepouite (x = 3)). First, all combinations between one vOH and one lattice or δ OH bands were 222 calculated (Table 3.a for lizardite (x = 0) and Table 3.b for nepoute (x = 3)). The wavenumber of 223 224 the combination band corresponds then more or less to the summation of the wavenumber of the implicated vibrations. The proposed attribution of bands was made using a difference limit (Δ) 225 arbitrary fixed at 10 cm⁻¹ between calculated and observed NIR band positions (Table 1 and 2). 226 227 This relatively high limit value is assumed to be due to the shift which could result from the actual position of a band and the position of the band determined by the second derivative of NIR 228 and MIR spectra. The difference limit is also assumed to include a possible contribution of 229 230 anharmonicity.

Most of the bands observed in the NIR spectra of lizardite (x = 0) (Table 1) and nepouite (x = 3) (Table 2) can be explained by the summation of two MIR bands (bands # 1, 3 to 12, 14, and 19 to 21 for lizardite (x = 0) and bands # 1 to 9, 11 to 13, and 19 to 21 for nepouite (x = 3)). Some of them could have several origins (bands # 6, 7, 9, and 19 for lizardite (x = 0) and bands #

235 2, 4 to 7, and 11 for nepouite (x = 3)). For nepouite (x = 3), the proximity between NIR bands (14 236 cm⁻¹ between bands # 8 and 9; and 15 cm⁻¹ between bands # 5 and 6), did not allow to 237 discriminate between the two possible combinations (e.g. bands # 8 and 9 and bands # 5 and 6, 238 respectively - Table 2).

A rather good agreement is observed between the experimental NIR spectra and the calculated NIR band positions for both lizardite (x = 0) and nepouite (x = 3). But some of the observed NIR bands remained unexplained (# 2, 13, 15 to 18 for lizardite (x = 0) and # 10, 14 to 18 for nepouite (x = 3)). The band # 2 in lizardite (x = 0) spectrum cannot be explained by any combination within the fixed difference limit (Δ). However, this band is at an intermediate position between NIR bands # 1 and 3 and then could result from an overlap of these two bands. Such an overlap may generate a broad band centered at the position of the observed band # 2.

Note that some of the calculated combinations (Table 3) were not observed on the NIR spectra of both end-members. These combinations may be forbidden and are not valid or maybe of too low intensity and hidden by another band.

When it was possible to identify the MIR features corresponding to a same vibrational group in both lizardite (x = 0) and nepouite (x = 3) end-members spectra, they were labelled with the same letter (a to i - Fig. 8). The attribution of NIR combinations was then compared between both end-members (Table 4). The combination bands concerning equivalent vibrational groups for lizardite (x = 0) and nepouite (x = 3) spectra give confidence on the attribution of the observed NIR combination bands.

In both end-member spectra, some bands between 4300 and 4600 cm⁻¹ could not be explained by combinations between two MIR features. Combinations with three MIR features were tested (Supplementary files b and c for lizardite (x = 0) and nepoute (x = 3), respectively) to tentatively attribute the observed NIR bands in the 4300 - 4600 cm⁻¹ range. Many

combinations between three MIR features could explain the observed NIR bands which were not attributed previously. These combinations between three MIR features could correspond to combinations between one MIR feature and one NIR feature (previously attributed; Table 1 and 2) localized in the low wavenumbers range of NIR spectra (3850 - 4300 cm⁻¹).

To simplify the attribution of observed NIR bands for both end-members and to reduce 263 the error of position, combinations between one MIR feature with one observed NIR feature in 264 the low wavenumbers range of the NIR spectra were made (Table 5 for lizardite (x = 0) and 265 Table 6 for nepouite (x = 3)). The bands # 13, 15, 16, 17, and 18, in the lizardite (x = 0) spectrum 266 and bands # 14, 15, 16, 17, and 18, in the nepoute (x = 3) spectrum, could be explained by 267 several combinations with one MIR and one NIR features. Note that the bands observed in the 268 4300 - 4600 cm⁻¹ range could not be explained by combinations of two MIR features (Table 1 and 269 2 for lizardite (x = 0) and nepoute (x = 3), respectively) but they could be explained by 270 combinations with three MIR features. More precisely, all these bands could be explained by 271 272 combinations between observed NIR features in the low wavenumbers range of the NIR spectra and the MIR band noted c (at 401 cm⁻¹ for lizardite (x = 0) and at 379 cm⁻¹ for nepoute (x = 3), 273 Fig. 8). The origin of this MIR band noted c, was not found in the literature as far as the authors 274 are aware. The complex pattern of IR spectra in the 400 - 200 cm⁻¹ was assumed to involve mixed 275 vibrations involving the silicon-oxygen network, the octahedral cations and the hydroxyl group 276 (Farmer 1974). 277

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Implications

NIR spectroscopy is widely used for remote-sensing and field studies to identify clay minerals. However, NIR spectra are often complex and uninterpreted even for monomineral samples, especially in the combination region where the origin of bands is unknown for many 283 clay minerals. Substantial progress in the interpretation of the IR spectra can be realized from chemical series of monomineral samples with controlled distribution of atoms. Hydrothermal 284 syntheses allowed to obtain those chemical series. In this paper, samples from the lizardite-285 nepoute series with a random distribution of Ni and Mg were synthesized to obtain robust 286 reference spectra. Because the $2vR^{2+}$ -OH vibrations were attributed, these reference spectra 287 288 could be used in field geology to discriminate the Ni distribution in single phase or in intimate mixture between 1:1 (lizardite-nepoute) and 2:1 (kerolite-pimelite) phyllosilicates from lateritic 289 290 nickel ore deposits as in garnierite. Note that the 2vOH in NIR region provided more resolved features compared to the vOH in MIR region and may help to distinguish between the 1:1 and the 291 2:1 Mg-Ni phyllosilicates even if partial overlapping of bands (Fig. 9). 292

293 The bands observed in the combination region were attributed with confidence by 294 comparing both lizardite (x = 0) and nepoute (x = 3) spectra. The observed combination bands result from combinations of two or three MIR features. Those combined MIR features can result 295 296 from the same vibrational group (e.g. combination between δOH and vOH vibrations) or from different vibrational groups (e.g. combination between vOH and vSi-O vibrations). The work of 297 interpretation by analogy using chemical series is essential for aiding the interpretation of IR 298 299 spectra. Indeed calculations of theoretical spectra (e.g. Balan et al. 2002; Prencipe et al. 2009) cannot be performed up to now for minerals which contain transition metals (e.g. Ni, Fe...). 300

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447 **Figure captions:**

448

Figure 1. Powder XRD patterns of synthesized lizardite-nepoute series $(Si_2Mg_{3-x}Ni_xO_5(OH)_4)$.

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Figure 2. XRD patterns of (06-33) reflections of synthesized lizardite-nepouite series (Si₂Mg_{3-x}Ni_xO₅(OH)₄).

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Figure 3. d(06-33) values (Å) versus Ni content (x) in lizardite-nepoute synthetic series (Si₂Mg₃₋

455 $_xNi_xO_5(OH)_4$). The size of symbol denotes the error on the d(06-33) measurements.

456

Figure 4. Selected regions of the MIR spectra of synthesized lizardite-nepouite series (Si₂Mg₃₋ $_xNi_xO_5(OH)_4$).

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Figure 5. Selected regions of the NIR spectra of synthesized lizardite-nepouite series (Si₂Mg₃₋
 xNi_xO₅(OH)₄).

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Figure 6. Wavenumbers of the v OH bands versus wavenumbers of the 2v OH first overtones for the synthesized lizardite-nepouite samples - this study (•), for various OH groups in talc (\Box and Ni-Mg \blacksquare) (Petit et al. 2004) and for palygorskite (Δ) (Gionis et al. 2006). The size of symbols denotes the error on the wavenumber measurements.

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Figure 7. NIR spectra in the combination region and their second derivative of lizardite (x = 0) (Si₂Mg₃O₅(OH)₄) (a), and nepouite (x = 3) (Si₂Ni₃O₅(OH)₄) (b). Numbers 1 to 21 correspond to

470	combination bands listed in table 1 for lizardite $(x = 0)$ and in table 2 for nepoute $(x = 3)$. In
471	grey, combination band positions from the second derivative NIR spectra.
472	
473	Figure 8. Second derivative of MIR spectra of lizardite (x = 0) ($Si_2Mg_3O_5(OH)_4$) (a), and
474	nepouite (x = 3) (Si ₂ Ni ₃ O ₅ (OH) ₄) (b). Letters correspond to the same vibrational group for each
475	end-member.
476	
477	Figure 9. First OH-overtones (2vOH) region of synthesized Ni-lizardite (Si ₂ Mg _{1.5} Ni _{1.5} O ₅ (OH) ₄)
478	and Ni-kerolite $(Si_4Mg_{1.5}Ni_{1.5}O_{10}(OH)_2.nH_20)$.
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494 Tables:

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496	Table 1. Observed NIR combination band positions of lizardite $(x = 0)$ and calculated NIR
497	combination band positions from combination with two MIR features. Δ corresponds to the
498	difference between the calculated and the observed combination band positions. † corresponds to
499	small identified features (see Fig. 7a).

Band #	Observed NIR bands (cm ⁻¹)	Combi	nation feature	with two M es (cm ⁻¹)	Calculated NIR bands (cm ⁻¹)	∆ (cm⁻¹)	
1	3958	302	+	3648	=	3950	8
2	3985						
3	4001	302	+	3695	=	3997	4
4	4032	380	+	3648	=	4028	4
5†	4049†	401	+	3648	=	4049	0
c	4077	380	+	3695	=	4075	2
O	4077	434	+	3648	=	4082	5
7	4121	434	+	3695	=	4129	8
/	4121	481	+	3648	=	4129	8
8	4163	458	+	3695	=	4153	10
0	4100	504	+	3695	=	4199	0
9	4199	554	+	3648	=	4202	3
10†	4240†	554	+	3695	=	4249	9
11	4274	635	+	3648	=	4283	9
12	4307	607	+	3695	=	4302	5
13	4363						
14†	4396†	693	+	3695	=	4388	8
15	4409						
16	4442						
17†	4487†						
18†	4531†						
10	1675	974	+	3695	=	4669	6
19	4075	1024	+	3648	=	4672	3
20	4709	1024	+	3695	=	4719	10
21	4781	1086	+	3695	=	4781	0

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Table 2. Observed NIR combination band positions of nepouite (x = 3) and calculated NIR combination band positions from combination with two MIR features. Δ corresponds to the difference between the calculated and the observed combination band positions. † corresponds to small identified features (see Fig. 7b).

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Band #	Observed NIR bands (cm ⁻¹)	Combir f	natio eatui	n with two res (cm ⁻¹)	Calculated NIR bands (cm ⁻¹)	∆ (cm⁻¹)	
1	3934	318	+	3608	=	3926	8
n	2071	318	+	3649	=	3967	4
Ζ	3971	357	+	3608	=	3965	6
3	4000	357	+	3649	=	4006	6
4	4027*	379	+	3649	=	4028	9
4†	4037	427	+	3608	=	4035	2
E	4056	404	+	3649	=	4053	3
5	4056	456	+	3608	=	4064	8
C+	4071*	427	+	3649	=	4076	5
١٥	4071	456	+	3608	=	4064	7
7	4007	456	+	3649	=	4105	8
/	4097	493	+	3608	=	4101	4
8	4136	493	+	3649	=	4142	6
9	4150	493	+	3649	=	4142	8
10†	4182†						
11	4077	628	+	3649	=	4277	0
11	4277	676	+	3608	=	4284	7
12	4297	652	+	3649	=	4301	4
13	4321	676	+	3649	=	4325	4
14	4348						
15	4383						
16†	4426†						
17	4447						
18†	4522†						
19	4652	1049	+	3608	=	4657	5
20	4698	1049	+	3649	=	4698	0
21	4722	1079	+	3649	=	4728	6

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509 Table 3. Calculated combinations (in cm⁻¹) between vOH and lattice or δ OH bands; a: for

а	lizardite (x = 0)		b	nepouite (x = 3)				
Lattice or	Lattice or vOH bands		Lattice or	vOH bands				
δOH bands	3648	3695	δ OH bands	3608	3649			
1086	4734	4781	1079	4687	4728			
1024	4672	4719	1049	4657	4698			
974	4622	4669	1015	4623	4664			
693	4341	4388	978	4586	4627			
635	4283	4330	676	4284	4325			
607	4255	4302	652	4260	4301			
554	4202	4249	628	4236	4277			
504	4152	4199	614	4222	4263			
481	4129	4176	493	4101	4142			
458	4106	4153	456	4064	4105			
434	4082	4129	427	4035	4076			
401	4049	4096	404	4012	4053			
380	4028	4075	379	3987	4028			
302	3950	3997	357	3965	4006			
		318	3926	3967				

510 lizardite (x = 0) and b: for nepoute (x = 3).

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Table 4. Comparison of some combination bands for both lizardite (x = 0) and nepoute (x = 3)

513 end-members. Letters correspond to MIR features which correspond to the equivalent vibrational

514 groups for each end-member. † corresponds to small identified features (see Fig. 7).

lizardi	te (x = 0)	Combinations from	nepouite (x = 3)		
Band #	Observed NIR bands (cm ⁻¹)	equivalent vibrational groups	Observed NIR bands (cm ⁻¹)	Band #	
1	3958	a + h	3934	1	
3	4001	a+i	2071	r	
4	4032	b + h	5971	Z	
c	4077	b + i	4000	3	
D	4077	d + h	4037†	4†	
7	4121	d + i	4071†	6†	
8	4163	e+i	4097	7	
21	4781	g + i	4722	21	

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Table 5. Calculated NIR combination band positions $(4300 - 4800 \text{ cm}^{-1})$ for lizardite (x = 0) from combination with one MIR feature and one NIR feature. Δ corresponds to the difference between the calculated and the observed combination band positions. *in letter in MIR spectra (Fig. 8) and in integer in NIR spectrum (Fig. 7). † corresponds to small identified features (see Fig. 7a).

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Band #	Observed NIR bands (cm ⁻¹)	Combination with one MIR band and one NIR band (cm ⁻¹)				Calculated NIR bands (cm ⁻¹)	∆ (cm⁻¹)	Observed features involved*
12	1262	380	+	3985	=	4365	2	b ; 2
15	4303	401	+	3958	=	4359	4	c;1
		380	+	4032	=	4412	3	b ; 4
1 5	4400	401	+	4001	=	4402	7	c ; 3
15	4409	434	+	3985	=	4419	10	d ; 2
		458	+	3958	=	4416	7	e ; 1
	4442	401	+	4032	=	4433	9	c ; 4
		401	+	4049	=	4450	8	c;5
16		434	+	4001	=	4435	7	d ; 3
		458	+	3985	=	4443	1	e ; 2
		481	+	3958	=	4439	3	-
		401	+	4077	=	4478	9	c;6
		434	+	4049	=	4483	4	d ; 5
17†	4487†	458	+	4032	=	4490	3	e ; 4
		481	+	4001	=	4482	5	-
		504	+	3985	=	4489	2	-
		401	+	4121	=	4522	9	c;7
		458	+	4077	=	4535	4	e ; 6
18†	4531†	481	+	4049	=	4530	1	-
		504	+	4032	=	4536	5	-
		554	+	3985	=	4539	8	-

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Table 6. Calculated NIR combination band positions (4300 - 4800 cm⁻¹) for nepouite (x = 3) from combination with one MIR feature and one NIR feature. Δ corresponds to the difference between the calculated and the observed combination band positions. *in letter in MIR spectra (Fig. 8) and

529 in integer in NIR spectrum (Fig. 7). † corresponds to small identified features (see Fig. 7b).

Band #	Observed NIR bands (cm ⁻¹)	Combina and	ation wi one NIR	th one MIR k R band (cm ⁻¹)	band	Calculated NIR bands (cm ⁻¹)	∆ (cm⁻¹)	Observed features involved*
		318	+	4037	=	4355	7	a ; 4
	42.40	357	+	4000	=	4357	9	b ; 3
14	4348	379	+	3971	=	4350	2	c ; 2
		404	+	3934	=	4338	10	d ; 1
		318	+	4056	=	4374	9	a ; 5
		318	+	4071	=	4389	6	a ; 6
15	4383	379	+	4000	=	4379	4	c;3
		404	+	3971	=	4375	8	-
		456	+	3934	=	4390	7	e ; 1
		357	+	4071	=	4428	2	b;6
		379	+	4037	=	4416	10	c ; 4
16*	4476+	379	+	4056	=	4435	9	c;5
101	4420	427	+	4000	=	4427	1	d ; 3
		456	+	3971	=	4427	1	e ; 2
		493	+	3934	=	4427	1	-
		318	+	4136	=	4454	7	a;8
		357	+	4097	=	4454	7	b ; 7
17	4447	379	+	4071	=	4450	3	c;6
		404	+	4037	=	4441	6	-
		456	+	4000	=	4456	9	e ; 3
		379	+	4136	=	4515	7	c;8
		379	+	4150	=	4529	7	c;9
10+	4500÷	427	+	4097	=	4524	2	d ; 7
TOI	43221	456	+	4056	=	4512	10	e ; 5
		456	+	4071	=	4527	5	e ; 6
		493	+	4037	=	4530	8	-

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