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3 **Solving the iron quantification problem in low-kV EPMA: An essential step**

4 **toward improved analytical spatial resolution in electron probe**

5 **microanalysis—Fe-Sulfides**

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### **Abstract**

13 The use of the field emission gun in scanning electron microscopy permits the imaging of sub-  
14 micrometer-size features. However, achieving sub-micrometer analytical spatial resolution in  
15 electron probe microanalysis (EPMA) requires both reducing the electron beam size and reducing  
16 the accelerating voltage to achieve the desired sub-micrometer interaction volume. The resulting  
17 quantification of the first-row transition metals at low accelerating voltage, i.e., below 7-8 kV, is  
18 problematic as the main characteristic X-ray lines ( $K\alpha$ ) cannot be excited at these conditions.  
19 Furthermore, the use of the  $L\alpha$  and  $L\beta$  soft X-ray lines for quantification is complicated by bonding  
20 and self-absorption effects resulting in not-yet determined mass absorption coefficients and hence  
21 in the failure of the traditional matrix correction procedure. We propose two methods to circumvent  
22 these low kV analysis limitations: using the non-traditional Fe  $L\ell$  line, and using universal  
23 calibration curves for the more traditional Fe  $L\alpha$  and  $L\beta$  lines. These methods were successfully

24 applied to Fe-sulfide minerals showing accurate quantification results by EPMA at reduced kV,  
25 necessary for accurate quantification of sub-micrometer sulfide grains.

26 Keywords: EPMA, sulfides, low-kV, iron, microprobe, SXES, EMPA, X-ray

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## Introduction

29 Electron probe microanalysis (EPMA) is a characterization technique that is routinely used for  
30 qualitative and quantitative analysis of micrometer-sized volumes. The technique has seen many  
31 significant advances in the last decades (Llovet et al. 2020). One of them has been developing and  
32 integrating Schottky field emission guns (FEGs) in microprobe instruments. These types of  
33 electron sources produce a stable, bright and narrow electron beam, resulting in a smaller beam  
34 diameter compared to tungsten and LaB<sub>6</sub> sources and hence offer a significant increase of the  
35 spatial imaging resolution. However, to better use this resolution improvement for quantitative  
36 purposes, it is also essential that the electron beam energy is reduced, typically from the traditional  
37 15 or 20 kV to 7-8 kV or less. The combination of reduced beam size through the availability of  
38 FEGs and reduced electron interaction volume through using lower beam energy then permits the  
39 possible successful analysis and quantification of sub-micrometer features.

40 Figure 1 illustrates this for the Fe K $\alpha$  and La X-ray production volume diameters in FeS<sub>2</sub>, using  
41 the Monte Carlo code PENEPMA/PENELOPE (Llovet and Salvat 2017; Salvat 2019). By  
42 decreasing the accelerating voltage from 20 kV to 7 kV (assuming an electron beam diameter of  
43 80 nm), the diameter of the contour delimiting an area where 99% of the considered X-rays are  
44 produced, is reduced from 3  $\mu\text{m}$  to about 0.75  $\mu\text{m}$ , as shown on Figure 1. Hence, the analytical  
45 spatial resolution is increased by a factor of 4. Reducing the beam diameter will only reduce the  
46 X-ray production volume's diameter by the same amount, therefore not drastically improving the

47 analytical resolution. The critical parameter governing the analytical spatial resolution is the  
48 accelerating voltage. It is worth noting that at 10 and 20 kV, the diameter of the X-ray production  
49 volume for the Fe  $K\alpha$  and Fe  $L\alpha$  X-ray lines are similar, as summarized in Table 1 . However, at  
50 7 kV the maximum of the produced X-ray intensity for the Fe  $L\alpha$  X-ray line is about three times  
51 higher than the maximum of the produced X-ray intensity for Fe  $K\alpha$  X-ray line at 10 kV.

52 We are not, however, the first to investigate the use of low kV EPMA for sulfides. Desborough et  
53 al. (1971) ran a successful series of experiments, measuring S  $K\alpha$  at 6 kV on the USGS ARL  
54 EMX electron probe. The goal was not improving the analytical spatial resolution but rather that  
55 of effectively eliminating matrix effects. Using a suite of synthetic in-house sulfide reference  
56 materials (Czamanske and Ingamells, 1970), they concluded that “a linear calibration curve  
57 relating X-ray intensity to sulfur concentration eliminates the need for using complex computer  
58 corrections required at higher operating voltages” – recall the state of computing five decades  
59 ago.

#### 60 **Difficulties of EPMA at low accelerating voltage**

61 Two<sup>1</sup> methods can be distinguished to improve the analytical spatial resolution in EPMA (and as  
62 demonstrated by Desborough et al. (1971), reduce the magnitude of the matrix correction for some  
63 elements): low overvoltage, and low voltage. The low overvoltage approach employs lowering the  
64 accelerating voltage, or electron beam energy, to just above the ionization energy required to

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<sup>1</sup> There is a third possible method, if special sample preparation is utilized. Kubo et al. (2013) thinned a specimen using focused ion beam (FIB) technology and with a 30 kV electron beam achieved high spatial resolution EPMA results in a semi-conductor specimen.

65 produce the X-ray transition of interest (i.e., low overvoltage) such as 8-10 kV for the Fe  $K\alpha$  X-  
66 ray line. The beam electrons have enough energy to produce the ionization only in a very small  
67 volume near the landing point of the electron beam, resulting in high analytical spatial resolution,  
68 as shown on Figure 1c. One disadvantage of this method is that the X-ray count rates will also drop  
69 as the probability of producing the desired ionization is reduced when the electron energy gets  
70 closer to the ionization edge threshold. This is further compounded because less electrons will  
71 contribute to the production of the desired ionization and, due to other interaction mechanisms, will  
72 quickly bring the electron energy below the required ionization energy threshold.

73 The second analytical approach consists of reducing the accelerating voltage to a low absolute  
74 value, e.g., 7 kV instead of the conventional 15 kV or 20 kV. The primary electrons have fewer  
75 energy and thus travel, on average, shorter distances before losing too much energy and  
76 contributing to the production of the specific X-ray line. Therefore, the electron interaction volume  
77 is consequently reduced, constraining the production of the characteristic X-rays of interest within  
78 a much smaller volume. Going much lower in kV is counterproductive, as the electron beam tends  
79 to widen on some FEG instruments (Pinard and Richter 2014) and the count rates tend to decrease  
80 rapidly. Increasing the beam current can compensate for the lowered count rates but aggravates the  
81 increase in beam diameter.

82 However, three main issues arise when considering operating at low accelerating voltage: (1) X-  
83 ray lines conventionally used at high kV may not be excited at low kV. The  $K\alpha$  X-ray line of the  
84 first row transition elements cannot be excited for accelerating voltages below 5-8 kV. The same  
85 situation happens for some of the rare earth elements where the  $L\alpha$  X-ray line cannot be excited at  
86 such low accelerating voltages. For example, the element Fe whose critical excitation energy for  
87 the production of the Fe  $K\alpha$  X-ray line is 7.114 keV (Zschornack 2007) cannot be excited with

88 electron beam energies lower than this threshold. (2) The measurements will be more sensitive to  
89 the state of the sample surface because the electrons no longer penetrate deeply into the material.  
90 Hence, for both the unknown specimen and the standards, the quality of surface polish, the surface  
91 fidelity and cleanliness, and the conductive coating thickness will be critical to achieve accurate  
92 measurements. In particular, for maximum accuracy, the accurate determination of the coating  
93 thickness should be performed to account for thickness variations between the unknown and the  
94 standards. Coating unknowns and standards simultaneously can help minimize the difference in  
95 coating thicknesses. However, this may not be sufficient depending on the geometry of the coater.  
96 Rotating the samples during the coating process may be needed to homogenize the coating  
97 thickness between samples. At low accelerating voltages, these variations, resulting in differences  
98 in X-ray production and absorption, will be more pronounced. And (3), since the primary electrons  
99 no longer penetrate as deep or spread out radially as much, they deposit their energy in a very small  
100 volume and may lead to sample damage. These problems encountered at low kV are discussed in  
101 detail in several recent publications (Jercinovic et al. 2012; Gopon et al. 2013; Kearns et al. 2014;  
102 Saunders et al. 2014; Buse and Kearns 2015, 2018; Moy et al. 2019a, 2019b).

103 In the present work, Fe-sulfide minerals were studied because of their importance in geoscience  
104 and materials science in numerous ways. For example, the analyses of the Fe–S–As minerals in  
105 gold deposits can illuminate the deposition mechanism of Au in sulfides, to better understand and  
106 to improve the ore formation models (Zhang et al. 2014). The study of sulfides is also used to  
107 understand the formation and the evolution of our solar system. The composition and crystal  
108 structure of sulfide inclusions found in chondrites and as fine-grained rims, with the help of Fe-Ni-  
109 S isothermal phase diagrams, can be used to constrain oxygen fugacity and cooling history of such  
110 objects and illuminate the processes and conditions of chondrite petrogenesis (Harries and

111 Langenhorst 2013; Schrader et al. 2015; Schrader and Zega 2018). The study of sulfide inclusions  
112 is also relevant in the determination of origins of deep Earth diamonds (Deines and Harris 1995).  
113 Characterization of sulfide inclusions in steels is also of great importance because these inclusions  
114 can create numerous problems such as breakage of steel wires during drawing, fatigue failure or  
115 surface flaws.

116 We focus here on the study of the transition element iron in iron-sulfide minerals at low  
117 accelerating voltage. At low accelerating voltages (7 kV and below) the traditionally used  
118 characteristic K X-ray lines of iron will not be excited, and thus our study considers the use of the  
119  $L\alpha$  ( $L_3$ - $M_{4,5}$  electron transition) and  $L\beta_1$  ( $L_2$ - $M_4$  electron transition) X-ray lines to perform  
120 quantification by EPMA, using the Fe  $L\alpha+L\beta_1$  area k-ratio (i.e., the area of the  $L\alpha$  and  $L\beta_1$  X-ray  
121 lines measured on the unknown and divided by the area measured on a standard) associated to a  
122 calibration curve (Moy et al. 2019a; 2019b). Note that for the sake of simplicity the  $L\beta_1$  transition  
123 will be denoted  $L\beta$  in the rest of the text. We also investigate the possibility of using the Fe  $L\ell$  ( $L_3$ -  
124  $M_1$  electron transition) X-ray line for quantification.

125

### Materials and Methods

126 EPMA measurements were performed using a CAMECA SXFive-FE instrument located at the  
127 Eugene Cameron electron microscopy laboratory, Department of Geoscience, University of  
128 Wisconsin-Madison and using a JEOL JXA-8530FPlus instrument at the Department of Earth  
129 Sciences, University of Minnesota. Both electron microprobes are equipped with field emission  
130 electron guns and use the Probe for EPMA software to operate the instrument and to acquire the  
131 data (Donovan et al. 2021). Additionally, the resulting calibration curve was validated on a separate  
132 instrument, a CAMECA SX51 at UW Madison Department of Geoscience.

### 133 **Fe-sulfides Samples**

134 A total of 14 Fe-sulfides of known composition were used in this study: 11 Fe-sulfide standards  
135 were used for the measurements performed with the SXFive-FE instrument and 12 Fe-sulfides  
136 standards were analyzed with the JXA-8530FPlus. The iron elemental concentrations in these  
137 sulfides range from 5.79 wt% up to 63.53 wt%. On both instruments, a pure Fe-metal sample was  
138 used as a standard. The samples' Fe concentrations as well as their sources are given in Table 2.

139 The Fe-sulfide standards were initially precisely characterized by EPMA at 15 kV (20 kV for the  
140 arsenopyrite standard) and 20 nA with PHA in wide differential mode using the SXFive-FE  
141 instrument. At these conditions, traditional  $K\alpha$  X-ray lines were used for all the elements except  
142 for Sn where the  $L\alpha$  line was used instead (which is considered as a main characteristic X-ray line  
143 for this element). A minimum of 5 points were measured on different locations, and the PAP matrix  
144 correction algorithm (Pouchou and Pichoir 1991) with the MAC30 mass absorption coefficients  
145 (Heinrich 1987) was employed to determine the elemental composition of the different specimens.  
146 For each sample, the S  $K\alpha$  X-ray line position was re-determined by a peak scan to avoid any peak  
147 shift error. The measured compositions are given in Table 2. Some of the minerals used were  
148 synthesized by Czamanske at the USGS in the late 1960s (Czamanske and Ingamells 1970). The  
149 slightly low total of the synthetic Zn-Fe-S samples Czamanske37 and Czamanske35 can be  
150 partially attributed to the fact that these samples are composed of small grains (~5  $\mu\text{m}$  diameter);  
151 hence, there is a lack of characteristic X-rays produced by secondary fluorescence in the grains,  
152 leading to a deficit of the total produced X-rays, compared to a bulk sample of the same material  
153 ("size effect": Fournelle 2006). This deficit of characteristic X-rays can easily account for an  
154 underestimation of the Fe concentration of 1 wt% in these small grains, where the pure Fe standard

155 used was significantly larger in size. The measured iron compositions were used in the following  
156 work.

### 157 **Spectra acquisition**

158 Three analytical sessions using an accelerating voltage of 7 kV were performed with the SXFive-  
159 FE microprobe to acquire the Fe  $L\alpha$  and  $L\beta$  spectra, as well as the Fe  $L\ell$  spectrum, on the different  
160 samples. During the first session, spectra from the following samples were recorded: Fe metal,  
161 pyrite, stannite, pentlandite, pyrrhotite, chalcopyrite, bornite and cobaltite. During the second  
162 session of measurements, the following new samples were analyzed, as well as samples previously  
163 analyzed in order to verify the reproducibility of the measurements: Canyon Diablo troilite, CZ4  
164 FeS, Mnt8 FeS, pyrrhotite, pyrite and Fe metal. The third session consisted of the measurement of  
165 the arsenopyrite and Fe metal specimens. With the JXA-8530FPlus microprobe, two analytical  
166 sessions were also performed at 7 kV to record the Fe spectra. During the first session, the same  
167 samples as the ones analyzed during the first session on the SXFive-FE instrument were measured  
168 at 7 kV and 20 nA: Fe metal, pyrite, stannite, pentlandite, pyrrhotite, chalcopyrite, bornite and  
169 cobaltite, while for the second session, the following samples were measured at 7 kV and 50 nA:  
170 chalcopyrite, Czamanske37 ZnFeS, Czamanske35 ZnFeS, chalcocite, Canyon Diablo troilite and  
171 pure Fe.

172 Three wavelength-dispersive spectrometers (WDSs) were employed on the SXFive-FE instrument  
173 using LTAP ( $2d = 25.745 \text{ \AA}$ ), TAP ( $2d = 25.745 \text{ \AA}$ ) and PC0 ( $2d = 45.0 \text{ \AA}$ ) diffractor crystals with  
174 a takeoff angle of  $40^\circ$ . A Thermo Scientific UltraDry energy-dispersive spectrometer (EDS) was  
175 also used on that instrument to record the spectra, also at a  $40^\circ$  takeoff angle. With the JXA-  
176 8530FPlus instrument a different type of spectrometer was used to record the X-ray spectra: a Soft  
177 X-ray Spectrometer (SXES, JEOL Inc.) consisting of an extended range diffraction grating JS2000,



178 and a Peltier-cooled CCD camera (Terauchi et al. 2010). The SXES spectrometer records the whole  
179 X-ray spectrum between 240 and 2800 eV simultaneously (similar to a SiLi or SDD energy  
180 dispersive spectrometer but with a spectral resolution similar to WDS-TAP). This spectrometer  
181 also has a takeoff angle of 40°. The electron beam conditions were set to 7 kV and 90 nA for the  
182 SXFive-FE instrument and to 7 kV and 20 or 50 nA for the JXA-8530FPlus instrument.

183 On the SXFive-FE microprobe, three spectra of the Fe  $L\alpha$ - $L\beta$  X-ray lines were recorded on each  
184 sample. The spectra were measured using Probe for EPMA in step-by-step acquisition mode (in  
185 opposition to the continuous ROM mode), from 0.685 to 0.733 keV with 511 steps and a dwell  
186 time of 1.8 s/step for the LTAP and TAP crystals and from 0.583 to 0.893 keV with 1.5 s/step for  
187 the PC0 pseudocrystal, resulting in an acquisition time of 13 minutes per scan. The spectra recorded  
188 with the PC0 crystals were wide enough to encompass the Fe  $L\ell$  X-ray line. Other spectrometers  
189 did not record this line. The EDS spectrum was recorded from 0 to 20 keV using 2048 channels  
190 and a live time of 90 seconds. On the JEOL instrument, three SXES spectra were also recorded and  
191 aggregated together on each sample. The spectra were acquired from 0.236 to 2.830 keV with 4096  
192 channels and a total counting time of 15 min.

193 For each spectrum, the background near the studied X-ray lines was fitted and removed using a  
194 linear function for the TAP crystals, the SXES and the EDS, and using an exponential curve for  
195 the PC0 crystal. The corrected spectra were then integrated by numerical analysis using the  
196 trapezoidal rule to calculate the total recorded Fe  $L\alpha$ + $L\beta$  X-ray line intensity. To only calculate  
197 the Fe  $L\alpha$  or Fe  $L\beta$  X-ray intensity, as detailed below, the spectra were fitted with a set of pseudo-  
198 Voigt functions (weighted sum of a Gaussian and Lorentzian functions) using a non-linear fitting  
199 algorithm applying the Levenberg-Marquart (Levenberg 1944; Marquardt 1963) minimization

200 procedure and then integrated analytically. The Fe L $\ell$  X-ray spectrum was also fitted using pseudo-  
201 Voigt functions to resolve the X-ray line from the Fe L $\eta$  X-ray line.

## 202 **How does carbon contamination affect spectral shapes?**

203 Acquisitions of detailed spectra require prolonged counting time. The carbon contamination that  
204 builds up over time due to the electron beam interactions on the surface of the sample may affect  
205 the measurements (Buse and Kearns 2015; Gopon et al. 2015) and distort the recorded spectrum.  
206 Note that the use of a dry pumped system (without oil pumps) as well as a cold stage help reduce  
207 the formation of carbon contamination. Any significant amount of carbon contamination created  
208 during the acquisition may distort the spectra by decreasing the energy of the primary electrons  
209 reaching the Fe-sulfide material and by increasing the absorption of the emitted X-rays as they  
210 have to travel through the contamination layer before exiting the sample. These effects would be  
211 more marked at the end of the acquisition when the carbon contamination has accumulated  
212 significantly, and could produce different appearances between two spectra recorded in opposite  
213 movement directions on a same spectrometer. The UW-Madison SXFive-FE instrument is  
214 equipped with a custom cryo-chiller that cools the cold plate under the annular BSE detector to  
215 reduce the amount of carbon contamination on the analysis spot. This anticontamination system  
216 reduced the temperature of the cold plate in the microprobe chamber to values as low as about –  
217 70°C and improved the chamber vacuum to a pressure as low as  $5 \times 10^{-6}$  Pa.

218 To investigate the effects of carbon contamination on the spectral acquisitions we performed  
219 spectral scans at the same conditions as our analytical sessions with the SXFive-FE instrument by  
220 recording the spectrum using a TAP, a LTAP and a PC0 monochromator crystal, scanning from  
221 low to high X-ray energies and then from high to low X-ray energies on a fresh new spot. For each  
222 direction, three spectra were acquired and averaged together to reduce the fluctuations due to the

223 counting statistics. The spectra were recorded under similar conditions as previously described: in  
224 step-by-step mode, with 511 steps and for X-ray energies ranging from 0.685 to 0.733 keV for the  
225 LTAP crystal, from 0.680 to 0.728 keV for the TAP crystal, and from 0.583 to 0.892 keV for the  
226 PC0 pseudocrystal. At the time of these measurements, the temperature of the cryo-chiller was of  
227 -33°C, potentially exacerbating the effects of the carbon contamination, if any, compared to the  
228 temperature at which the previous spectra (reported in the “spectra acquisition” section) were  
229 recorded. The measurements were performed on three samples covering a broad range of Fe  
230 concentrations: pure Fe, pyrite and stannite.

231 As seen on Figure 2, the recorded spectra show no significant variations on either peak shape or  
232 intensity for both the background and the tails of the Fe L $\alpha$  and Fe L $\beta$  X-ray lines. The only  
233 noticeable difference is a small reduction of the Fe L $\alpha$  maximum X-ray intensity and a slight  
234 increase of the Fe L $\beta$  X-ray intensity when the spectra were recorded from high to low energies  
235 compared to spectra recorded in the opposite direction. This observation, which cannot be  
236 explained by counting statistics only, is in accordance with an increase of the carbon contamination  
237 during the course of the measurement. However, the majority of the disparities between the spectra  
238 are typically within the uncertainties of the measurements due to the counting statistics. These  
239 uncertainties are represented for every 3 points by the error bars on Figure 2. The variations of the  
240 X-ray line areas due to carbon contamination and counting statistics remain generally small, as  
241 seen by the difference spectra shown on Figure 2, with average relative standard deviations  
242 (calculated from areas of spectra measured in opposite directions) of 0.7% for the LTAP crystal,  
243 of 3.6% for the TAP crystal and of 1.6% for the PC0 crystal. The worst case observed amongst all  
244 the recorded spectra was measured on the pure Fe standard by the low detection efficiency TAP  
245 crystal and had a relative standard deviation of about 4%. The fact that the deviation on the LTAP

246 spectrometer is very low indicates that carbon contamination has limited effect on the  
247 measurements at these experimental conditions when a cryo-chiller is used. It also indicates that  
248 the higher deviations seen on the spectra measured with the TAP and PC0 spectrometers, which  
249 were measured at the same time that the spectra measured on the LTAP spectrometers, are mostly  
250 due to the counting statistics.

251 On the JEOL JXA-8530FPlus microprobe, no anti-contamination devices were utilized but the  
252 measured area k-ratios are in good agreement, within the uncertainties due to counting statistics,  
253 with the area k-ratios measured on the CAMECA SXFive-FE microprobe using the cryo-chiller  
254 anticontamination device. This can easily be explained by assuming that the contamination rate is  
255 the same between samples for a given instrument. Hence, each spectrum is affected by the same  
256 X-ray intensity reduction due to the carbon contamination which is then canceled out when  
257 calculating the area k-ratios. The area k-ratios obtained between instruments are then similar. This  
258 shows that the effect of the carbon contamination does not significantly affect the measurements,  
259 even when no cryo-chiller devices are used. Nevertheless, to be consistent, all the measurements  
260 used to derive the calibration curves were obtained from spectra acquired in the direction of  
261 decreasing X-ray energy (except for the SXES and EDS detectors where the spectrum is acquired  
262 simultaneously).

263

264

## Results

### 265 **Quantification of Fe using the traditional method with the Fe $L\alpha$ X-ray line**

266 An attempt was made to quantify the Fe-sulfides at 7 kV and 90 nA with the traditional  
267 quantification method using the background-corrected Fe  $L\alpha$  X-ray peak intensities (measured at

268 the top of the X-ray peak and off to the sides to remove the background) with the PAP model  
269 (Pouchou and Pichoir 1991) and the MAC30 mass absorption coefficients (MACs) (Heinrich 1987)  
270 and using pure Fe or pyrite as a standard. The measurements were performed on the SXFive-FE  
271 instrument using a LTAP monochromator crystal and at least three points were measured on each  
272 sample. Figure 3 shows the measured iron content in the different samples as a function of the Fe  
273 content (previously measured at 15 kV with the traditional Fe  $K\alpha$  X-ray line). The results present  
274 significant deviations from the known Fe concentration. The error in the quantification, increases  
275 with the Fe content and can account up to 20 wt% (in absolute value) in the worst case when using  
276 pure Fe as a standard and up to 30 wt% when using pyrite as a standard. The error bars of the Fe  
277 concentration obtained using the Fe  $L\alpha$  line are on average of 2% for both standards, with a  
278 maximum of 5.6% and 5.0% for the Czamanske35 ZnFeS sample when using pure Fe and pyrite  
279 as a standard, respectively.

#### 280 **Quantification of Fe using the Fe $L\ell$ X-ray line and the traditional method**

281 Gopon et al. (2013) showed that the Fe  $L\ell$  X-ray line ( $L_3$ - $M_1$  transition) can successfully be used  
282 to quantify Fe in Fe-silicide samples. The spectra measured with the PC0 crystal were wide enough  
283 to record the Fe  $L\ell$  and Fe  $L\eta$  ( $L_2$ - $M_1$  transition) X-ray lines. The X-ray intensity, measured at the  
284 peak maximum, for the Fe  $L\ell$  line was obtained by fitting the experimental spectrum (Figure 4).  
285 The background in this energy range and for this monochromator crystal is best fitted by an  
286 exponential function rather than a linear function. The proximity of the more intense Fe  $L\alpha$ - $L\beta$   
287 lines prevents the acquisition of the background on the high energy side of the Fe  $L\eta$  line. This  
288 requires the acquisition of the X-rays far from the Fe  $L\ell$  line on the high energy side of the Fe  $L\alpha$ -  
289  $L\beta$  spectrum to be able to correctly evaluate the background. The Fe  $L\ell$  and  $L\eta$  lines were  
290 deconvoluted using pseudo-Voigt functions. As seen on Figure 4, the  $L\eta$  line does not interfere

291 strongly with the  $L\ell$  line: its intensity, at the X-ray energy of the maximum of the Fe  $L\ell$  line, is  
292 about ~1% to 1.5% of the maximum X-ray intensity of the Fe  $L\ell$  line. This value, however, could  
293 change with other monochromator crystals having different spectral resolutions (e.g., TAP vs PC0  
294 crystals). It should be noted that this interference is not a problem as long as the  $L\ell$  and  $L\eta$  X-ray  
295 intensity ratio remains constant between unknowns and standards, and can be seen as the classical  
296  $K\alpha_1$ ,  $K\alpha_2$  interference. The net X-ray intensities corresponding to the maximum of the Fe  $L\ell$  lines  
297 were extracted from the fits and used to calculate k-ratios relative to pure Fe. These k-ratios were  
298 processed using Probe for EPMA (Donovan et al. 2021) and BadgerFilm (Moy and Fournelle 2021)  
299 by specifying the other element abundances (previously measured at 15 kV). The PAP  $\phi(\rho z)$  model  
300 (Pouchou and Pichoir 1991) was used in the matrix correction procedure, as well as the FFAST  
301 MACs (Chantler et al. 2005) for Probe for EPMA and the MAC30 MACs for BadgerFilm (the  
302 MAC30 MACs cannot be used for the  $L\ell$  line in Probe for EPMA). The obtained Fe concentrations  
303 are displayed in Table 3.

304 Generally good quantification numbers were obtained at 7 kV using the Fe  $L\ell$  X-ray line compared  
305 to the results obtained at 15 kV using the Fe  $K\alpha$  line. The greatest deviation was obtained for the  
306 cobaltite sample with a relative deviation of -5.72%. One disadvantage of using the Fe  $L\ell$  X-ray  
307 line for quantification is that compared to the  $L\alpha$  X-ray line, the X-ray intensity of the  $L\ell$  line is  
308 weak, which can be problematic especially for low Fe concentration specimens. This leads to  
309 greater statistical fluctuations in the measured spectrum, to a less accurate fitting and hence to a  
310 less accurate measured net X-ray intensity making its use difficult to quantify low Fe concentration  
311 samples. Higher beam currents or longer counting times can be used to overcome this problem,  
312 with the downside of potentially damaging the sample and increasing carbon contamination effects.  
313 It should also be noted that the Fe  $L\ell$  line is interfered by the second order of diffraction of the Mg

314  $K\alpha$  X-ray line, making its use for quantification difficult in Mg-bearing samples (not the case here).  
315 For these reasons, the authors evaluated alternative Fe  $L\alpha$ - $L\beta$  approaches for low kV EPMA of  
316 silicates (Moy et al. 2019b).

317

### 318 **Fe $L\alpha$ + $L\beta$ area k-ratio vs. Fe content**

319 As shown above (Figure 3) and in previous works (Gopon et al. 2013; Llovet et al. 2016; Buse and  
320 Kearns 2018; Moy et al. 2019a, 2019b), the traditional quantification method is not suitable for  
321 quantification using the Fe  $L\alpha$  X-ray line. This is due to a multitude of reasons, such as large  
322 uncertainties in the knowledge of the MACs close to the absorption edges, which is the case for  
323 the soft X-ray lines such as Fe  $L\alpha$ . These MACs can vary strongly from one material to another  
324 due to chemical bonding effects. Attempts have been made to measure the  $L\alpha$  MACs, as well as  
325 other atomic parameters such as the fluorescence yield, of Ni in Ni-Si samples (Heikinheimo et al.  
326 2016; Llovet et al. 2016) and of Fe in olivines (Buse and Kearns 2018). These works showed, to  
327 some extent, an improvement of the experimental results. However, in addition to these problems,  
328 we have shown that the theory behind the matrix correction procedure is not valid close to the  
329 absorption edges, e.g., the Fe  $L\alpha$  X-ray line, and it is not possible to accurately measure the MACs  
330 close to absorption edges using traditional WDSs (see Moy et al., 2019b for a full explanation).

### 331 **$L\alpha$ + $L\beta$ area k-ratio calibration curve**

332 The recent studies by Moy et al. (2019 a, b) describe an alternative approach for quantifying Fe  
333 contents at low kV that overcomes the accuracy problem using the traditional Fe  $L\alpha$  EPMA  
334 approach and when Fe  $L\ell$  is not applicable. We evaluate below its applicability to Fe-sulfides. In  
335 this approach, an empirical curve is defined which uses “area k-ratios”, integrating the background-

336 subtracted areas under the Fe  $L\alpha+L\beta$  peaks and ratioing them to the same integrated area of an Fe-  
337 rich standard, here Fe metal. This approach actually is a re-tooled version of one of Castaing's  
338 original methods for quantification in his original electron probe dissertation (Castaing 1951),  
339 which became known as the alpha-factor correction method (Ziebold and Ogilvie 1963) and then  
340 the Bence-Albee correction method (Bence and Albee 1968).

341 The area k-ratios of the combined Fe  $L\alpha-L\beta$  X-ray lines, measured at 7 kV and a takeoff angle of  
342  $40^\circ$ , were obtained on each Fe sulfide sample relative to the same X-ray lines' area obtained on a  
343 pure Fe standard. Figure 5 shows the measured area k-ratios as a function of the known Fe  
344 concentration. The first significant observation is that the area k-ratio values are very close to each  
345 other for a given sample, irrespective of the spectrometer type (WDS, EDS or SXES) or instrument  
346 model used – and in accordance with the theoretical derivations given in the following section. The  
347 experimental area k-ratios follow a smooth increase with increasing Fe concentration up to about  
348 40 wt%. However, at higher Fe concentrations, i.e., for the pyrite, pyrrhotite and troilite materials,  
349 the area k-ratio increases abruptly, deviating from the expected correlation. Indeed, pyrite which  
350 has an Fe concentration of about 47 wt%, has an area k-ratio similar to the area k-ratio measured  
351 on pyrrhotite and troilite, which have an Fe concentration of about 60-63 wt%. These materials,  
352 which have a similar matrix composition, should see an increase of the Fe  $L\alpha$  X-ray intensity with  
353 increasing iron content. However, as shown in Figure 6, the opposite behavior occurs: pyrite has a  
354 much stronger Fe  $L\alpha$  X-ray line intensity than pyrrhotite and troilite despite its lower Fe  
355 concentration. This behavior is not seen for the Fe  $L\beta$  X-ray intensity resulting in a similar Fe  
356  $L\alpha+L\beta$  area k-ratios for these materials.

357 To circumvent this problem for modeling of those high-Fe sulfides, instead of calculating the Fe  
358  $L\alpha+L\beta$  area k-ratio, the Fe  $L\alpha$  and Fe  $L\beta$  area k-ratios are calculated independently and then



359 averaged together. To calculate the two areas separately, because of their overlap, each spectrum  
360 is fitted with a set of Pseudo-Voigt functions which are then integrated analytically. The  
361 background was previously removed by fitting the tails of the spectrum with a linear or exponential  
362 function depending on the monochromator crystal used. Using this method, the area k-ratios of the  
363 pyrite, pyrrhotite and troilite can be distinguished from each other. Hence, two procedures have  
364 been developed here: the combined Fe  $L\alpha+L\beta$  area k-ratios can be used to obtain a calibration  
365 curve and to quantify Fe-sulfides with an Fe concentration lower than about 45 wt%, and the  
366 averaged  $k_{av}=(\text{Fe } L\alpha \text{ area k-ratio} + \text{Fe } L\beta \text{ area k-ratio})/2$  quantities can be used for higher Fe  
367 concentrations. Note that this technique is harder to use with data acquired with an EDS or with a  
368 WDS equipped with a PC0 or PC1 ( $2d = 60.79 \text{ \AA}$ ) monochromator crystal because the low spectral  
369 resolution of these detectors does not allow the distinction between the  $L\alpha$  and the  $L\beta$  X-ray lines.  
370 This can be partially mitigated by fitting the low-resolution spectra using strong constraints such  
371 as the position of the  $L\beta$  line relative to the position of the  $L\alpha$  line. However, accuracy of such fits  
372 is difficult to assess and may lead to inaccuracies in the derivation of a calibration curve for high  
373 Fe concentration specimens.

374 Because the experimental k-ratios are almost identical from one spectrometer to another, in  
375 agreement with the theoretical derivations in the next section, data across instruments and  
376 spectrometers were averaged subsequently. The two sets of data (Fe concentrations below and  
377 above 45 wt%) follow a smooth increasing trend, as shown in Figure 7. However, the dataset of  
378 low Fe concentration samples has some “irregularities”: the stannite, cobaltite and arsenopyrite  
379 data are slightly off the curve defined by the other elements. The data were fitted by part, with  
380 polynomial functions with the conditions that the k-ratios are 0 and 1 for Fe concentrations of 0

381 and 100 wt%, respectively. The low Fe concentration data were fitted with and without the stannite,  
382 cobaltite and arsenopyrite samples. The obtained polynomial equations (displayed in Figure 7) are:

$$C = -797.44 \times k^4 + 599.35 \times k^3 - 61.966 \times k^2 + 72.372 \times k \quad \text{for } C \leq 45 \text{ wt\%} \quad (1)$$

$$C = 32719.17 \times k^5 - 36588.61 \times k^4 + 13959.12 \times k^3 - 2063.446 \times k^2 + 176.4047 \times k \quad \text{for } C \leq 45 \text{ wt\%} \quad (2)$$

$$C = 75.865 \times k_{av}^4 - 107.29 \times k_{av}^3 + 36.104 \times k_{av}^2 + 95.323 \times k_{av} \quad \text{for } C > 45 \text{ wt\%} \quad (3)$$

383

384 where C is the Fe concentration in wt%, k is the Fe L $\alpha$ +L $\beta$  area k-ratio (where pure Fe is the  
385 standard), and k<sub>av</sub> is the (Fe L $\alpha$  area k-ratio + Fe L $\beta$  area k-ratio)/2 quantity. Very good regression  
386 factors of 0.996 and 0.999 were found for the fitting polynomials. The best fit of the low Fe  
387 concentration dataset was obtained for Fe-sulfides with Cu and/or Ni whereas the addition of  
388 samples containing other elements in the sulfide structure yielded more complicated and slightly  
389 poorer fit. This can be attributed to changes in X-ray production and emission caused by bonding  
390 effects due to the presence of non-divalent ions like Sn and As. Consequently, the area k-ratios  
391 from these Fe-sulfide samples differ from those only containing divalent cations like Fe, Cu and  
392 Ni. Therefore, we recommend using the simpler calibration curve that excludes the stannite,  
393 cobaltite and arsenopyrite data points (Eq. 1) when analyzing sulfides with only 4th-period  
394 transition metals as cations. For sulfide compositions outside of this compositional range such as  
395 arsenopyrite the calibration curve that includes all data points (Eq. 2) should be favored. It should  
396 also be noted that for the low Fe concentration calibration curve, there is no experimental data for  
397 the relevant Fe concentrations above 34.32 wt%. Consequently, the fitting is unconstrained above  
398 this value. This may result in some inaccuracies in the quantification predictions, in particular close  
399 to the limit of 45 wt%.

400

## Discussion

## 401 **Theoretical derivation of the calibration curves**

402 The statement above that these calibration curves are valid for all electron probe WDS, SXES or  
403 EDS detectors, provided they are acquired at 7 kV and with a takeoff angle of 40°, can be supported  
404 by a rigorous examination of the physics of X-ray generation, of spectral characteristics, and of  
405 detector behavior.

406 The X-ray intensity of a given X-ray line, emitted from the sample before detection by the  
407 spectrometer, can be written as follows:

408

$$I_X(E) = C_A \frac{N_a}{A_r} n_{el} \sigma_X^A(E_0) \mathcal{L}(E) \int_0^\infty \varphi_i(\rho z) e^{-\frac{\mu(E)}{\rho} \frac{\rho z}{\sin \theta_d}} d\rho z \mathcal{F} \quad (4)$$

409

410 where  $I_X(E)$  is the number of characteristic X-rays originating from the inner shell transition of  
411 interest and emitted at the photon energy  $E$  per unit of time.  $C_A$  and  $A_r$  represent the concentration  
412 and the atomic mass of the element  $A$  of interest, respectively.  $N_a$  denotes the Avogadro number.

413 It is worth noting that the product  $C_A \frac{N_a}{A_r}$  by the material density  $\rho$  gives the number of atoms  $A$  per

414 unit volume.  $n_{el}$  is the number of primary electrons reaching the sample per unit of time. This  
415 quantity can be calculated by  $j/q$  where  $j$  is the current of the primary electron beam, in amperes,

416 and  $q$  is the electric charge of an electron, i.e.,  $\sim 1.602 \times 10^{-19}$  coulomb.  $\sigma_X^A(E_0)$  is the X-ray

417 production cross section, in  $\text{cm}^2$ , of the considered X-ray line by impact of an electron with an

418 energy  $E_0$ .  $\sigma_X^A(E_0)$  takes into account the ionization cross section by electron impact of the shell

419 (or subshell)  $i$  generating the considered X-ray line and also the ionization cross section of all the

420 inner-shells that can produce an electron vacancy in the shell  $i$  during the atomic relaxation process

421 – this includes radiative, non-radiative, Coster-Kronig and super-Coster-Kronig transitions (Moy

422 et al. 2013). For example, the X-ray production cross section of the  $L\alpha_1$  X-ray line by transition of  
 423 an electron from the  $M_5$  subshell to the  $L_3$  subshell can be written:

$$\sigma_{L\alpha_1}^A(E_0) = \omega_{L_3}^A \frac{\Gamma_{L_3-M_5}^A}{\Gamma_{L_3-total}^A} [Q_{L_3}^A(E_0) + Q_{L_2}^A(E_0) \times f_{23} + Q_{L_1}^A(E_0) \times (f_{13} + f_{12}f_{23} + f'_{13}) + Q_K^A(E_0) \times (\eta_{K-L_1}(f_{13} + f_{12}f_{23} + f'_{13}) + \eta_{K-L_2}f_{23} + \eta_{K-L_3})] \quad (5)$$

424

425 where  $\omega_{L_3}^A$  is the fluorescence yield of the element A for an electron vacancy in the subshell  $L_3$

426 (i.e., the probability to emit a photon during the relaxation of atom A with an initial vacancy in the

427 electronic subshell  $L_3$ ).  $\Gamma_{L_3-M_5}^A$  is the partial radiative width associated to the  $L_3-M_5$  transition and

428  $\Gamma_{L_3-total}^A$  is the total radiative width for all possible transitions to the  $L_3$  subshell. This ratio

429 corresponds to the probability that amongst all the possible electron transitions, the relaxation

430 occurs by the transition of an electron from the subshell  $M_5$  to the vacancy located in the subshell

431  $L_3$ .  $Q_i^A(E_0)$  is the ionization cross section of the shell or subshell  $i$  by impact of electrons of energy

432  $E_0$ .  $f_{ij}$  is the Coster-Kronig yield (i.e., the probability that a vacancy in the subshell  $L_i$  of a singly

433 ionized atom shifts to a higher subshell  $L_j$  of the same shell) and  $f'_{13}$  is the radiative transition yield

434 between the  $L_1$  and  $L_3$  subshells. The terms  $\eta_{K-L_i}$  are the radiative plus non-radiative yields for the

435 transitions of vacancies from the K shell to the  $L_i$  subshell. In Eq. (4),  $\mathcal{L}(E)$  represents the natural

436 shape of the X-ray line, whose area is normalized to unity. Due to the uncertainty principle of

437 quantum mechanics, the characteristic X-rays are emitted with a distribution following a

438 Lorentzian function of the form:

$$\mathcal{L}(E) = \frac{H_L}{1 + 4 \left( \frac{E - E_X}{\Gamma} \right)^2} \quad (6)$$

439

440 where  $H_L$  is the maximum of the function,  $\Gamma$  is the full width at half maximum and  $E_X$  is the energy

441 on which the function is centered. The  $\varphi_i(\rho z)$  term is the so-called phi-rho-z function corresponding

442 the ionization depth distribution of the electron shell (or subshell)  $i$  of element A at mass depth  $\rho z$   
443 inside the sample for an incident electron beam of energy  $E_0$ .

444 The next term in Eq. (4), represented by the exponential factor, accounts for the attenuation of the  
445 X-rays generated inside the sample before they reach the surface with an angle  $\theta_d$  (corresponding  
446 to the takeoff angle of the spectrometer).  $\frac{\mu}{\rho}(E)$  is the mass absorption coefficient, in  $\text{cm}^2/\text{g}$ , of the  
447 material composing the sample for a photon of energy  $E$ . It is worth noting that the term  $\frac{\mu}{\rho}(E)$   
448 varies rapidly as a function of the photon energy close to the ionization edges. Finally,  $\mathcal{F}$  represents  
449 the enhancement of X-rays by secondary fluorescence from other characteristic X-rays and from  
450 the bremsstrahlung ( $\mathcal{F} \geq 1$ ).

451 The emitted X-ray intensity  $I_X(E)$  is then recorded by the spectrometers at the X-ray energy  $E_i$  (not  
452 necessarily the line center  $E_X$ ):

$$I(E_i) = \int_{-\infty}^{+\infty} I_X(E) \frac{\Delta\Omega}{4\pi} \varepsilon(E) \mathcal{G}(E_i - E) dE \approx \int_{E_i - \Delta E/2}^{E_i + \Delta E/2} I_X(E) \frac{\Delta\Omega}{4\pi} \varepsilon(E) \mathcal{G}(E_i - E) dE \quad (7)$$

453  
454 where  $\varepsilon(E)$  and  $\frac{\Delta\Omega}{4\pi}$  are the intrinsic detection efficiency and the geometric detection efficiency of  
455 the spectrometer for photons of energy  $E$ , respectively. Because of the “imperfect” spectrometer  
456 response function, X-rays with energies in a small interval (or energy bin width) centered around  
457 the recorded X-ray energy  $E_i$  and with a width  $\Delta E$  will also be detected. The spectrometer response  
458 function can be better described by a gaussian function, here denoted  $\mathcal{G}(E_i - E)$  with a full width  
459 at half maximum of  $\Gamma_G$  (here  $3\Gamma_G \approx \Delta E$ ), a height  $H_G$  and centered at the photon energy  $E_i$ .

460 In the traditional case where  $\varepsilon(E)$  and  $\frac{\mu}{\rho}(E)$  are almost constant in the energy interval  $\Delta E$  (which is  
461 almost always the case for  $\varepsilon(E)$ , except, for example, close to an absorption edge caused by the

462 spectrometer separation window or by an absorption edge of the gas used in proportional counters),  
 463 the detected X-ray intensity becomes:

$$I(E_i) = C_A \frac{N_a}{A_r} n_{el} \sigma_X^A(E_0) \int_0^\infty \varphi_i(\rho z) e^{-\frac{\mu}{\rho} \frac{\rho z}{\sin \theta_d}} d\rho z \mathcal{F} \frac{\Delta\Omega}{4\pi} \varepsilon \int_{-\infty}^{+\infty} \mathcal{G}(E_i - E) \mathcal{L}(E) dE \quad (8)$$

464  
 465 Hence, in the traditional case, after cancelling out identical factors between the numerator and  
 466 denominator, the k-ratio is given by:

$$k = \frac{I_u(E_i)}{I_s(E_i)} = \frac{C_A^u \int_0^\infty \varphi_i^u(\rho z) e^{-\frac{\mu}{\rho} \frac{\rho z}{\sin \theta_d}} d\rho z \mathcal{F}_u}{C_A^s \int_0^\infty \varphi_i^s(\rho z) e^{-\frac{\mu}{\rho} \frac{\rho z}{\sin \theta_d}} d\rho z \mathcal{F}_s} \quad (9)$$

467 where the indices u and s stand for unknown and standard, respectively. It was assumed that the  
 468 same beam current was used, i.e., the same number of electrons (per unit of time) reached the  
 469 unknown and the standard. As seen in Eq. (9), this k-ratio is independent of the spectrometer  
 470 characteristics. However, when the MACs vary in the energy bin width  $\Delta E$ , the emitted X-ray  
 471 intensity  $I_X(E)$  cannot be extracted from the integral in Eq. (7) and thus, the k-ratio expression  
 472 cannot be simplified (the spectrometer contributions remain). This is the case close to the ionization  
 473 edge where the MACs strongly vary with respect to the X-ray energy. Also, the k-ratio expression  
 474 cannot be simplified when the natural line shape  $\mathcal{L}(E)$  of the X-ray line is different between the  
 475 unknown and the standard, which could be the case for the soft X-rays where bonding effects can  
 476 change the density of state of the outer electron shells and hence the shape of the associated  
 477 characteristic X-ray line, or when the X-ray production cross section differs from the standard and  
 478 the unknown, which can also be caused by bonding effects. In these situations (present near  
 479 absorption edges), the k-ratio remains:

480

$$k = \frac{I_u(E_i)}{I_s(E_i)} = \frac{C_A^u \sigma_X^{A,u}(E_0) \int_{-\infty}^{+\infty} \int_0^{\infty} \varphi_i^u(\rho z) e^{-\frac{\mu(E)}{\rho} \Big|_u \frac{\rho z}{\sin \theta_d}} d\rho z \mathcal{G}(E_i - E) \mathcal{L}(E) dE \mathcal{F}_u}{C_A^s \sigma_X^{A,s}(E_0) \int_{-\infty}^{+\infty} \int_0^{\infty} \varphi_i^s(\rho z) e^{-\frac{\mu(E)}{\rho} \Big|_s \frac{\rho z}{\sin \theta_d}} d\rho z \mathcal{G}(E_i - E) \mathcal{L}(E) dE \mathcal{F}_s} \quad (10)$$

481

482 and cannot be further simplified. These k-ratios depend on the spectrometer response function  $\mathcal{G}$

483 and then may be different from one spectrometer to another. The determination of  $C_A^u$  is also almost

484 impossible as most of the parameters used in Eq. (10) (especially  $\frac{\mu}{\rho}(E)$ ,  $\mathcal{G}(E_i - E)$  and  $\mathcal{L}(E)$ ) are

485 not well known, particularly for the soft X-rays. Fortunately, by calculating the area k-ratio, i.e.,

486 the area of the X-ray lines calculated over the X-ray energy for both the unknown and the standard,

487 and by using the property of the convolution product that states that the area under a convolution

488 of factors is the product of areas under these factors, the area k-ratio can be written as:

$$k_{Area} = \frac{\int_{-\infty}^{+\infty} I_u(E_i) dE_i}{\int_{-\infty}^{+\infty} I_s(E_i) dE_i} = \frac{C_A^u \sigma_X^{A,u}(E_0) \mathcal{F}_u \int_{-\infty}^{+\infty} \int_0^{\infty} \varphi_i^u(\rho z) e^{-\frac{\mu(E)}{\rho} \Big|_u \frac{\rho z}{\sin \theta_d}} d\rho z \mathcal{L}(E) dE \int_{-\infty}^{+\infty} \mathcal{G}(E_i - E) dE}{C_A^s \sigma_X^{A,s}(E_0) \mathcal{F}_s \int_{-\infty}^{+\infty} \int_0^{\infty} \varphi_i^s(\rho z) e^{-\frac{\mu(E)}{\rho} \Big|_s \frac{\rho z}{\sin \theta_d}} d\rho z \mathcal{L}(E) dE \int_{-\infty}^{+\infty} \mathcal{G}(E_i - E) dE} \quad (11)$$

489

490 The spectrometer broadenings  $\int_{-\infty}^{+\infty} \mathcal{G}(E_i - E) dE$  cancel out in the numerator and denominator and

491 the remaining quantities are independent of the spectrometer used. Although the area k-ratio cannot

492 be calculated analytically because of our lack of knowledge of the different parameters, a

493 calibration curve can be obtained by measuring the area k-ratios on samples of known

494 compositions. This calibration curve is independent of the spectrometer used and can thus be

495 employed on any instrument with any spectrometer, as long as the electron beam energy  $E_0$  and the

496 spectrometer takeoff angle  $\theta_d$  used to acquire the calibration curve are the same as used here.

497 The derivation of a calibration curve using the traditional k-ratio obtained by measuring the X-ray  
498 intensity at the maximum of the peak and on each side to remove the background, on the unknown  
499 and on a standard, is also possible, but that curve will not be spectrometer independent and must  
500 be redetermined on each instrument and each spectrometer, as shown in Eq. (10). The derivation  
501 was done for one characteristic X-ray line. In the case of several overlapping X-ray lines, as it is  
502 generally the case for the Fe  $L\alpha$  and  $L\beta$  X-ray lines using PC0 diffracting crystals, the area k-ratio  
503 can be obtained similarly by integrating over the extent of the overlapping lines. This theoretical  
504 derivation explains why the area k-ratio values displayed on Figure 5 are similar between the  
505 different instruments and the different spectrometers.

506

#### 507 **Using the calibration curve on a different instrument**

508 For an initial test of the validity of our calibration curves, the same calibration samples were  
509 examined using a different microprobe instrument, a W-filament CAMECA SX51 microprobe  
510 located at the Eugene Cameron electron microscopy laboratory, Department of Geoscience,  
511 University of Wisconsin-Madison. The Fe  $L\alpha$  and  $L\beta$  spectra were recorded at 7 kV and 50 nA  
512 using two TAP crystals ( $2d = 25.745 \text{ \AA}$ ) and a PC0 ( $2d = 45.0 \text{ \AA}$ ) crystal. Three spectra were  
513 recorded for each sample and aggregated together to reduce the counting fluctuations and to smooth  
514 the spectra. The spectra were recorded, with 300 steps and a dwell-time of 2 s/step, from 688 to  
515 733 eV for the TAP crystals and from 588 to 878 eV for the PC0 crystal. The takeoff angle of the  
516 spectrometers used was  $40^\circ$ . To remove the background, the spectra were fitted with a linear  
517 function for the TAP crystals and with an exponential function for the PC0 crystal. The Fe  $L\alpha$  and  
518  $L\beta$  X-ray lines were fitted with pseudo-Voigt functions to calculate area k-ratios relative to pure



519 Fe. The calibration curves deduced from the previous results (Eqs. 1–3) and shown in Figure 7  
520 were used to quantify Fe in the samples. As shown on Figure 8, and Table 4, generally good  
521 quantification results were obtained, especially when compared to the traditional matrix correction  
522 quantification method (Figure 3). The quantification of the sample Czamanske37, with nominal Fe  
523 concentration of 23.16 wt%, is systematically underestimated by the calibration curve method with  
524 an obtained average Fe concentration of 20.08 wt%. Despite this outlier sample, the maximum  
525 quantification error was of 3.2 wt% obtained on the pyrrhotite specimen with a TAP crystal. The  
526 quantification error, in average, was of 1.3 wt% for the TAP crystals and of 0.9 wt% for the PC0  
527 crystal. The observed errors (deviation from the diagonal) are consistent with counting statistics.

528 The results are in generally good agreement with the expected Fe concentrations, demonstrating  
529 the robustness of the calibration curve approach, irrespective of instrument, spectrometer, and  
530 crystal choices. It should also be noted that the area calibration curve is not affected by spectrometer  
531 drifts as the entire spectrum is recorded and integrated.

532 To show the improvements obtained on the quantification of Fe by using the calibration curve with  
533 regard to the traditional quantification method using the Fe  $L\alpha$  X-ray line, a direct comparison of  
534 the results obtained with the two methods is displayed in Table 5. The quantification results  
535 obtained with the calibration curve on the SX51 instrument, using the spectrometer #1 (TAP  
536 crystal) were compared to the results obtained on the SXFive-FE instrument with the spectrometer  
537 #1 (LTAP crystal) using the Fe  $L\alpha$  line and the traditional method. Both methods were using pure  
538 Fe as standard. The results were also compared to the quantification results obtained at 15 kV using  
539 the Fe  $K\alpha$  X-ray lines.

540 Better quantification results were obtained with the calibration curve method except for the stannite  
541 specimen where the Fe concentration is underestimated (2.18 wt% below the nominal value).  
542 Systematic substantial improvements were obtained for the high Fe concentration samples.

#### 543 **Using the calibration curve to quantify unknown Fe-sulfide specimens**

544 Unknown Fe-sulfide samples were used to test the predictions of the area k-ratio calibration curves  
545 as well as quantification results obtained using the Fe L $\ell$  X-ray lines. The three samples used were  
546 pyrrhotite from Galax, VA, chalcopyrite from Durango, Mexico and bornite from the Magma mine,  
547 AZ. These samples were first analyzed at 15 kV and 20 nA using the traditional K $\alpha$  lines on the  
548 SXFive-FE instrument to determine their exact compositions. PET crystals were used to measure  
549 the S K $\alpha$  lines and LLiF crystals were used to measure the Fe K $\alpha$  and Cu K $\alpha$  X-ray lines. The  
550 quantification results are displayed in Table 6.

551 Spectra of the Fe L $\alpha$  and L $\beta$  X-ray lines as well as the spectrum of the Fe L $\ell$  X-ray line were  
552 recorded at 7 kV and 70 nA using a LTAP and a PC0 crystals on the SXFive-FE instrument. PHAs  
553 were set to wide differential mode. Three spectra were recorded and averaged together for each  
554 sample. The averaged spectra were corrected from the continuum and fitted with pseudo-Voigt  
555 functions to calculate the area k-ratios, relative to pure Fe. The obtained Fe L $\alpha$ +L $\beta$  area k-ratios  
556 for the chalcopyrite and bornite specimens and the (Fe L $\alpha$  k-ratio + Fe L $\beta$  k-ratio)/2 quantity for  
557 the pyrite specimen were used with the calibration curves to determine the Fe content for each  
558 sample. The Fe L $\ell$  X-ray intensities were processed with CalcZAF using the FFAST MAC  
559 database, as well as with BadgerFilm using the MAC30 MACs. Both programs were using the PAP  
560 matrix correction algorithm. Quantification results are displayed in Table 7.

561 Using the calibration curves or the Fe L $\ell$  X-ray line intensity leads to significant quantification  
562 improvements compared to the traditional quantification method using the Fe L $\alpha$  X-ray line. For  
563 the low Fe concentration sample, the calibration curve method underestimates the concentration by  
564 4 wt% in average while the Fe L $\ell$  method gives very good results with an overestimation of only  
565 0.22 wt% in average.

566

567

### Implications

568 Sulfides are the source for most of the world supplies of non-ferrous metals and their synthetic  
569 analogues are of interest to physicists and material scientists for their electrical, magnetic, and  
570 optical properties. However, grain sizes can be in the sub-micrometer range or occur as narrow  
571 overgrowths or exsolutions below the spatial limit of traditional EPMA analysis at 15 kV or higher.  
572 Quantifying iron contents in sulfides at low accelerating voltage successfully is a necessary  
573 procedure for measuring these sub-micrometer sulfide grains by EPMA. We have shown that when  
574 accelerating voltages much lower than 15-20 kV must be used, two new approaches can avoid the  
575 problems encountered when utilizing Fe L $\alpha$  X-ray peak intensities combined with traditional  
576 matrix correction methods and provide generally good results at that reduced voltage. The first one  
577 utilizes the low intensity Fe L $\ell$  X-ray line but may be limited in its use to samples (1) with Fe-  
578 concentration larger than trace level because of the low Fe L $\ell$  X-ray line intensity, and (2) to Mg-  
579 poor phases like sulfides because of an interference by the second order of diffraction of the Mg  
580 K $\alpha$  line. The second approach uses a calibration curve method and exploits the more intense Fe  
581 L $\alpha$ -L $\beta$  X-ray peak area intensities to calculate area k-ratios relative to a standard. This approach  
582 yields generally good Fe quantification results even at low Fe abundances. The best quantification

583 results were obtained for Fe-sulfides with Cu and/or Ni whereas the presence of other elements in  
584 the sulfide structure made the quantification more difficult. This probably reflects changes in  
585 production and emission of X-rays due to changes in bonding environment in the presence of ions  
586 like Sn and As. This can result in area k-ratios differing from sulfides only containing divalent  
587 cations like Fe, Cu and Ni, and would explain why these specimens are more difficult to fit with a  
588 general calibration curve. Another advantage of the calibration curve method is that labs can readily  
589 apply the curve determined in this study to data acquired on their instrument and spectrometers as  
590 long as they also utilize an accelerating voltage of 7 kV and the spectrometer has a takeoff angle  
591 of 40°. Our universal calibration curve utilizes pure Fe metal, a widely available standard material,  
592 and thereby offers an easy and fast method to quantify Fe in sulfides using the Fe L $\alpha$ -L $\beta$  and/or Fe  
593 L $\ell$  X-ray lines at low kV. Alternatively, a new calibration curve can be determined, provided that  
594 a similar suite of reference materials is available, for either a different accelerating voltage and/or  
595 an instrument with a different takeoff angle (e.g., SEM-WDS). In addition, the method is well  
596 suited to high-spectral resolution, parallel channel X-ray detectors such as the SXES (JEOL Inc.),  
597 and to a lesser extent EDS, as these detector types acquire the entire Fe L X-ray spectrum  
598 simultaneously. The time-advantage presented by this method may be further compounded if other  
599 transition metals like Cr, Ni, Cu, Zn are present, and it can be established that they also can be  
600 quantified by calibration curves (von der Handt, Moy and Fournelle, in progress). Although the  
601 presence of the SXES detectors in microanalysis labs is currently still scarce, they could certainly  
602 play an essential role in the future, especially in the geosciences.

603 The successful application of the calibration curve method to low kV quantitative analysis of iron  
604 in sulfides extends our findings on Fe-silicides (Moy et al. 2019a) and olivines (Moy et al. 2019b)

605 and provides a roadmap to future investigation of more complex iron-bearing phases and a wider  
606 range of accelerating voltages.

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612 (AVDH).

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712 **Figure 1.** X-ray production volumes of the Fe  $K\alpha$  and Fe  $L\alpha$  lines in pyrite ( $\text{FeS}_2$ ) at 7, 10 and 20  
713 kV, obtained by Monte Carlo simulations using the PENEPMA/PENELOPE code (Llovet and  
714 Salvat 2017; Salvat 2019). An electron beam diameter of 80 nm was used. The outer green contour  
715 shows the limit within which 99% of the considered X-rays are produced.

716  
717  
718 **Figure 2.** Fe  $L\alpha$  and Fe  $L\beta$  X-ray lines measured on stannite, pyrite and Fe metal at 7 kV and 90  
719 nA using WDS detectors to study the impact of carbon contamination. Spectra recorded by moving  
720 the monochromator crystal from low-to-high or high-to-low energies are very similar, indicating  
721 that the carbon contamination has limited impact on the measurements. The continuous grey line  
722 at the bottom of each plot represents the intensity difference (in c/s/nA) between spectra measured  
723 in opposite directions.

724  
725 **Figure 3.** Fe concentration determined by the traditional quantification method, with either pure  
726 Fe or pyrite as a standard, at 7 kV using the Fe  $L\alpha$  X-ray line versus the actual Fe concentration  
727 (in wt%). Error bars are of the size of the symbols or smaller. The large deviations from the 1:1  
728 diagonal indicate that the Fe  $L\alpha$  X-ray lines are not suitable for use with the traditional  
729 quantification method.

730  
731 **Figure 4.** Fe L spectrum measured on a pyrite standard at 7 kV and 90 nA. Panel (a) shows the L  
732 lines of Fe. Panel (b) is a close-up of the spectrum on the Fe  $L\ell$  and  $L\eta$  X-ray lines. Symbols are  
733 the experimental data, continuous lines are the pseudo-Voigt functions fitting the different X-ray  
734 lines, dot dashed lines represent the background, and the dashed line is the sum of the fitting  
735 functions.

736

737 **Figure 5.** Experimental Fe  $L\alpha+L\beta$  area k-ratios plotted as a function of the Fe concentration.

738 Panel (b) is a close-up of the lower Fe concentration Fe-sulfides. The area k-ratios for pyrite,  
739 pyrrhotite and troilite are similar even though their Fe concentrations are different.

740

741 **Figure 6.** Spectra of the Fe  $L\alpha$  and Fe  $L\beta$  X-ray lines measured on pyrite, pyrrhotite and troilite  
742 at 7 kV and 90 nA using a LTAP crystal. Despite containing less iron, pyrite has a more intense  
743 Fe  $L\alpha$  X-ray line intensity than pyrrhotite and troilite.

744

745 **Figure 7.** Calibration curves obtained for (a) low Fe concentrations using the Fe  $L\alpha+L\beta$  area k-  
746 ratio and (b) high Fe concentration using an average of the Fe  $L\alpha$  and  $L\beta$  area k-ratios (see text  
747 for details). The area k-ratios were averaged between the different spectrometers and fitted with  
748 polynomial functions.

749

750 **Figure 8.** Quantification of Fe in Fe-sulfide samples at 7 kV using the calibration curves. Three  
751 different spectrometers were used to obtain area k-ratios. The X axis represents the measured Fe  
752 concentration while the Y axis represents the known concentration. Results fall closely to the 1:1  
753 line.

754

755 **Table 1.** Diameter of the X-ray interaction volume where 99% of the considered characteristic X-  
756 rays are produced as a function of the accelerating voltage for the Fe K $\alpha$  and Fe L $\alpha$  X-ray lines.

X-ray line	7 kV	10 kV	20 kV
Fe K $\alpha$	–	1.10 $\mu\text{m}$	2.80 $\mu\text{m}$
Fe L $\alpha$	0.56 $\mu\text{m}$	0.99 $\mu\text{m}$	3.15 $\mu\text{m}$

757

758

759 **Table 2.** Compositions of the iron-sulfide standards used in this study (in wt%). The Fe  
 760 compositions were measured at 15 kV (and 20 kV for the arsenopyrite standard) using the  
 761 traditional K $\alpha$  line. (AMNH: American Museum of Natural History; MAC: Micro-Analysis  
 762 Consultants Ltd, England; NMNH: National Museum of Natural History)

Sample name	Fe	Zn	S	Sn	Co	Ni	Cu	As	Total	Source
Fe	100	–	–	–	–	–	–	–	100	
FeS (Czamanske4)	63.41±0.24	–	35.84±0.2 0	bdl	bdl	bdl	–	bdl	99.25±0.31	Czamanske collection
Canyon Diablo troilite	62.98±0.33	–	36.02±0.1 2	bdl	bdl	bdl	–	bdl	99.00±0.35	NMNH
FeS (Mnt 8)	62.97±0.20	–	36.04±0.1 1	bdl	bdl	bdl	–	bdl	99.01±0.22	AMNH
Pyrrhotite	60.41±0.17	bdl	39.51±0.1 1	bdl	–	–	–	–	99.92±0.21	MAC
Pyrite	46.73±0.23	bdl	53.48±0.1 7	bdl	–	–	–	–	100.21±0.2 8	MAC
Arsenopyrite	34.32±0.14	–	18.02±0.2 7	–	bdl	bdl	–	48.33±0.2 1	100.67±0.3 7	MAC
Pentlandite	30.62±0.22	0.16±0.0 4	33.34±0.0 9	bdl	0.83±0.11	35.30±0.1 5	–	–	100.25±0.3 0	MAC
Chalcopyrite	30.29±0.16	bdl	34.72±0.1 2	bdl	bdl	–	34.28±0.2 4	–	99.29±0.31	MAC
ZnFeS (Czamanske37)	23.16±0.21	42.00±0. 46	33.64±0.3 0	–	–	–	–	–	98.80±0.59	Czamanske collection
Stannite	11.68±0.14	1.46±0.1 7	29.33±0.0 5	18.87±0.2 2	bdl	–	38.86±0.1 5	–	100.20±0.3 5	MAC
Bornite	11.37±0.24	bdl	25.76±0.1 4	bdl	bdl	–	62.85±0.6 5	–	99.98±0.70	MAC
Cobaltite	9.40±0.21	–	19.95±0.1 3	bdl	17.15±1.0 5	9.87±0.69	–	42.95±1.0 7	99.32±1.67	MAC
ZnFeS (Czamanske35)	5.62±0.08	60.74±0. 60	32.59±0.1 2	–	–	–	–	–	98.95±0.62	Czamanske collection
Chalcocite	0.44±0.08	–	20.16±0.2	–	–	–	79.4±0.8	–	100.0±0.83	MAC

763 bdl: below detection limit

764 –: not measured

765

766 **Table 3.** Quantification of the Fe-sulfide samples using the Fe L $\ell$  X-ray line with two different  
 767 analysis software: Probe for EPMA and BadgerFilm (see text for details). Expected concentration  
 768 as well as the relative error are also given.

Sample name	Probe for EPMA			BadgerFilm	
	Fe wt% (K $\alpha$ line, 15 kV)	Fe wt% (L $\ell$ line, 7 kV)	Rel. Err. (%)	Fe wt% (L $\ell$ line, 7 kV)	Rel. Err. (%)
FeS (Czamanske4)	63.41±0.24	62.40	-1.60	64.12	1.12
Canyon Diablo troilite	62.98±0.33	62.76	-0.35	63.95	1.54
FeS (Mnt 8)	62.97±0.20	62.96	-0.02	63.94	1.54
Pyrrhotite	60.41±0.17	59.68	-1.20	60.47	0.10
Pyrite	46.73±0.23	45.87	-1.84	46.50	-0.49
Arsenopyrite	34.32±0.14	32.79	-4.45	33.01	-3.82
Pentlandite	30.62±0.22	30.85	0.76	30.42	-0.65
Chalcopyrite	30.29±0.16	30.01	-0.93	30.90	2.01
Stannite	11.68±0.14	11.36	-2.77	11.30	-3.25
Bornite	11.37±0.24	11.23	-1.25	11.30	-0.62
Cobaltite	9.40±0.21	8.86	-5.72	9.41	0.11

769

770 **Table 4.** Fe concentration (wt%) obtained using the area k-ratio calibration curve at 7 kV with  
 771 the CAMECA SX51 instrument and with three different WDSs.

Sample name	Nominal Fe wt%	SX51, Sp1 TAP	SX51, Sp2 TAP	SX51, Sp4 PC0
FeS (Czamanske4)	63.41	65.28	-	63.75
Canyon Diablo troilite	62.98	60.89	65.02	62.28
FeS (Mnt 8)	62.94	61.28	62.92	62.27
Pyrrhotite	60.41	62.83	63.60	60.37
Pyrite	46.73	48.17	47.76	49.05
Pentlandite	30.62	31.63	30.14	28.84
Chalcopyrite	30.29	29.90	29.92	30.08
ZnFeS (Czamanske37)	23.16	20.87	19.34	20.05
Stannite	11.68	9.50	9.82	10.56
Bornite	11.37	11.82	11.97	11.74
Cobaltite	9.4	8.16	9.06	9.16
ZnFeS (Czamanske35)	5.62	6.56	-	5.42

772

773



774 **Table 5.** Comparison of the Fe quantification results (in wt%) obtained using the traditional method  
 775 with the Fe L $\alpha$  X-ray intensity and using the area k-ratio calibration curve method on the CAMECA  
 776 SX51 instrument. The relative error, relative to the nominal Fe concentration, is also indicated.

Sample name	Nominal Fe wt%	Traditional method	Rel. Err. (%)	Calibration curve	Rel. Err. (%)
FeS (Czamanske4)	63.41	47.74	-24.7	65.28	3.0
Canyon Diablo troilite	62.98	46.88	-25.6	60.89	-3.3
FeS (Mnt 8)	62.94	47.33	-24.8	61.28	-2.6
Pyrrhotite	60.41	47.53	-21.3	62.83	4.0
Pyrite	46.73	65.48	40.1	48.17	3.1
Pentlandite	30.62	33.52	9.5	31.63	3.3
Chalcopyrite	30.29	27.26	-10.0	29.90	-1.3
ZnFeS (Czamanske37)	23.16	20.76	-10.3	20.87	-9.9
Stannite	11.68	12.66	8.3	9.50	-18.7
Bornite	11.37	12.43	9.3	11.82	4.0
Cobaltite	9.4	14.12	50.2	8.16	-13.2
ZnFeS (Czamanske35)	5.62	6.73	19.8	6.56	17.1

777

778

779 **Table 6.** Nominal composition, in wt%, of the unknown samples determined at 15 kV using the  
780 traditional  $K\alpha$  X-ray lines.

Sample name	S	Fe	Cu	Total
Pyrrhotite	38.85±0.15	60.88±0.19	bdl	99.73±0.24
Chalcopyrite	34.96±0.14	30.20±0.09	34.35±0.12	99.51±0.21
Bornite	25.77±0.09	10.95±0.05	62.69±0.18	99.41±0.21

781 bdl: below detection limit.

782

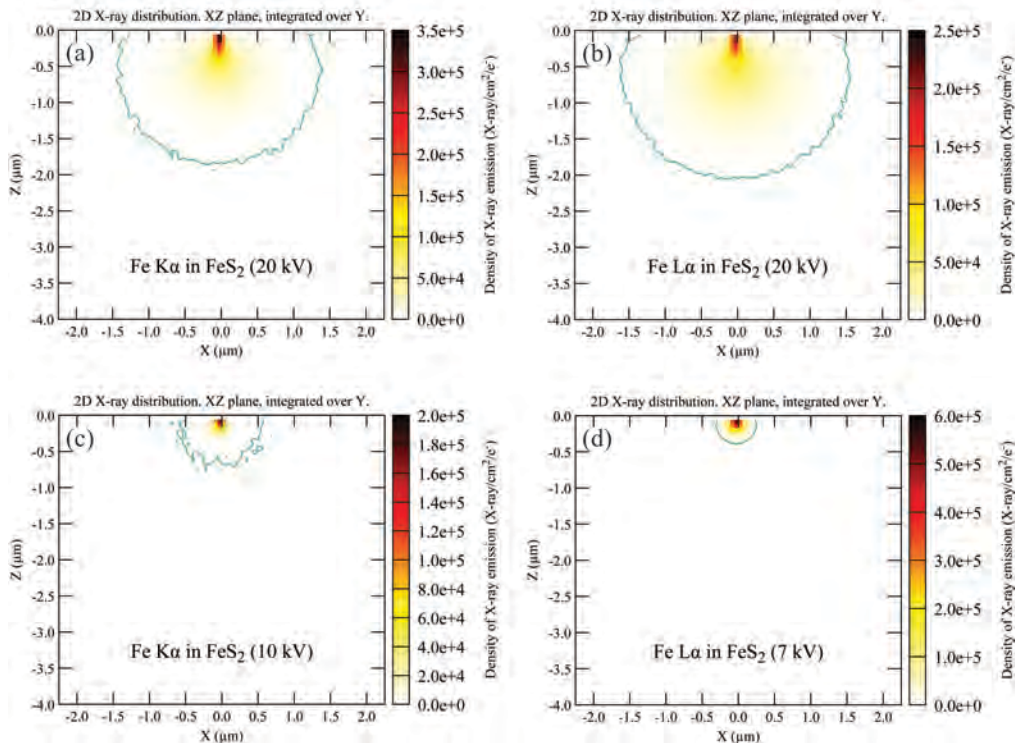
783 **Table 7.** Quantification of Fe at 7 kV using the traditional method with the Fe L $\alpha$  X-ray line, using  
 784 the area k-ratio calibration curves and using the traditional method with the Fe L $\ell$  line. Pure Fe  
 785 was used as a standard for all the methods. Concentrations are given in wt%.

786

Sample name	Nominal	Traditional method with the L $\alpha$ X-ray line		Area k-ratio calibration curves		Traditional method with the L $\ell$ X-ray line CalcZAF		Traditional method with the L $\ell$ X-ray line BadgerFilm	
		LTAP	PC0	LTAP	PC0	LTAP	PC0	LTAP	PC0
		Pyrrhotite	60.88	46.94	52.39	63.61	58.06	61.19	62.51
Chalcopyrite	30.20	28.07	32.45	30.93	30.85	31.19	31.42	31.24	31.47
Bornite	10.95	7.94	7.72	7.51	6.67	10.74	10.8	10.71	10.78

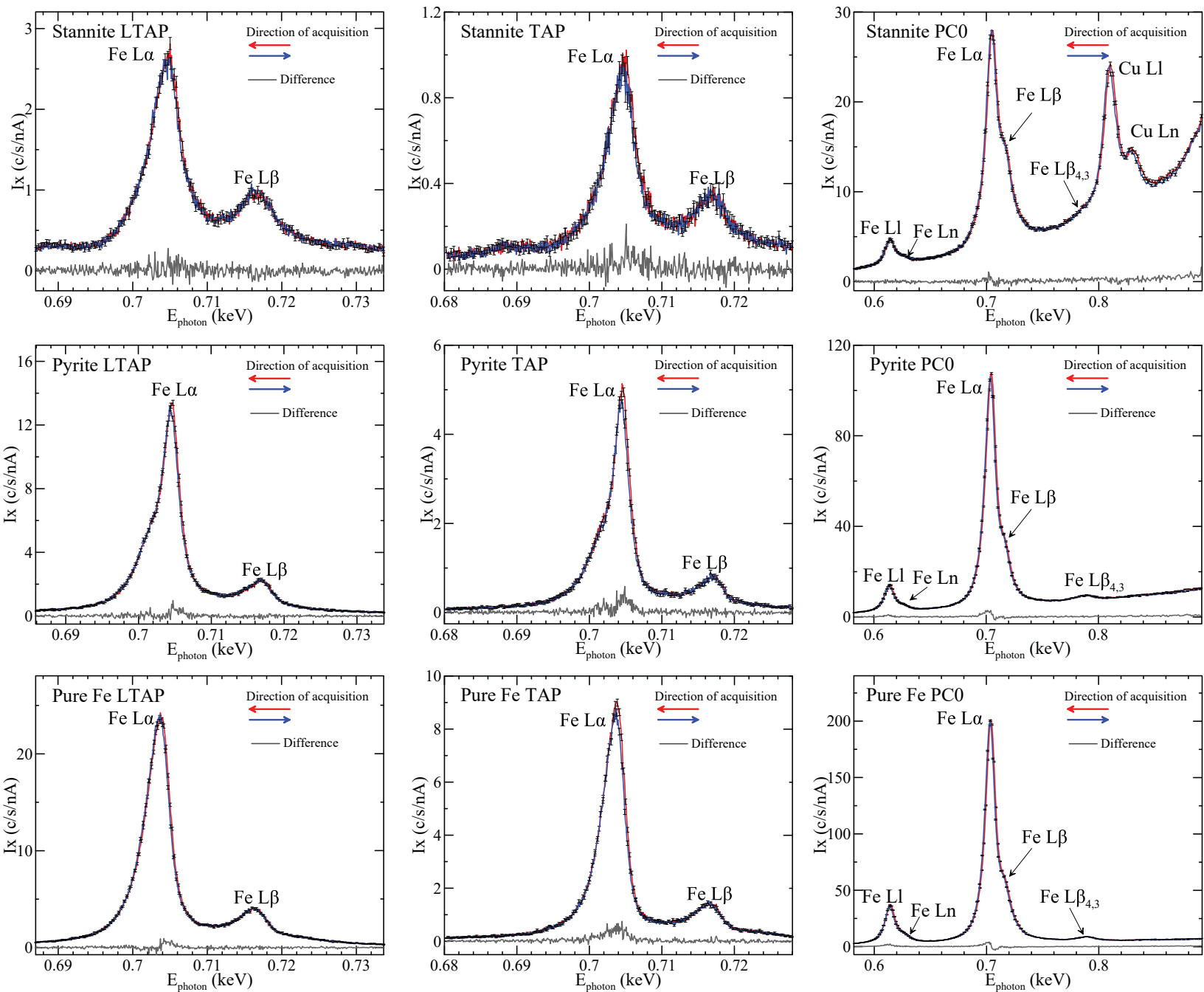
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Figure 1



Always consult and cite the final, published document. See <http://www.minsocam.org> or GeoscienceWorld

Figure 2



# Figure 3

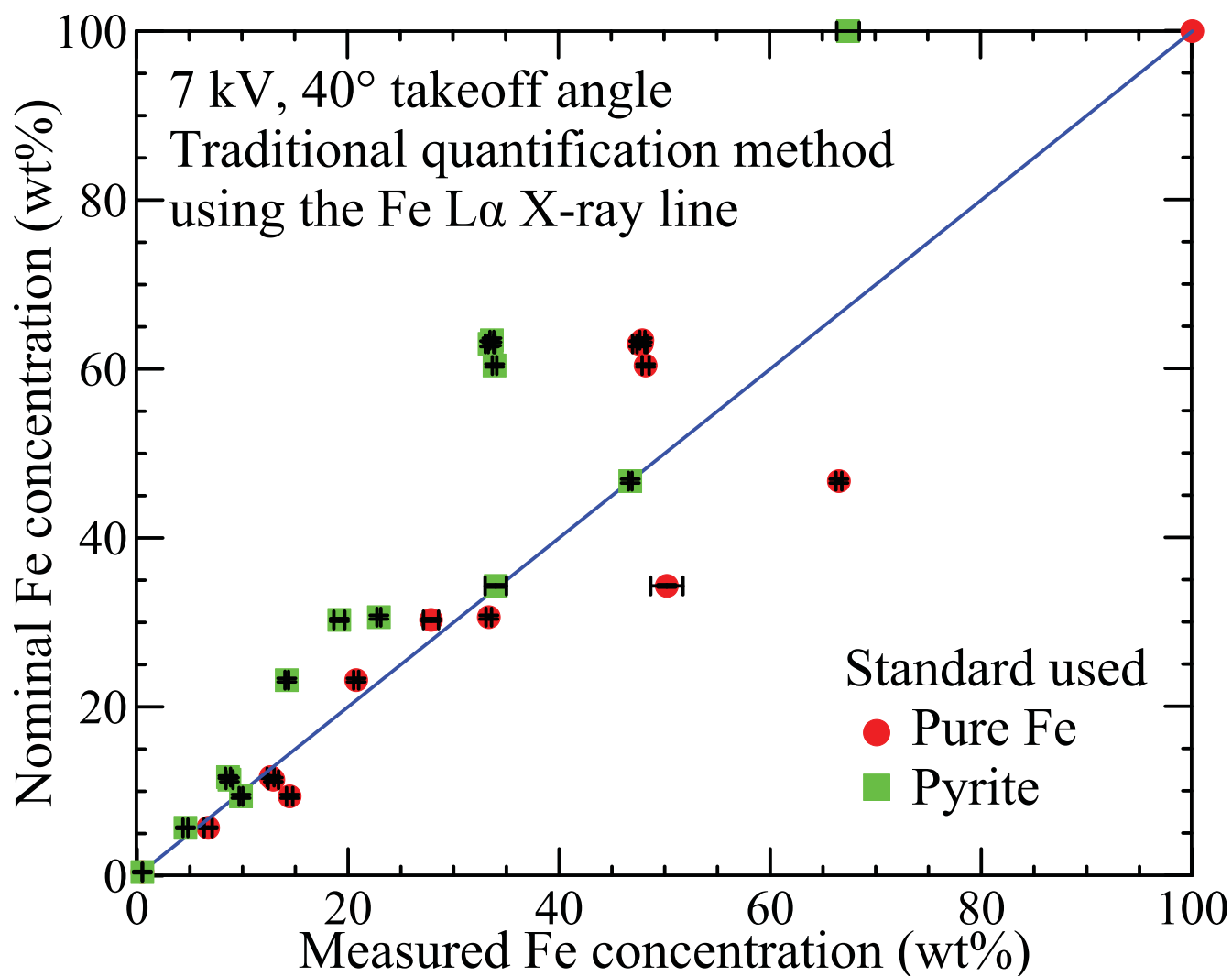


Figure 4

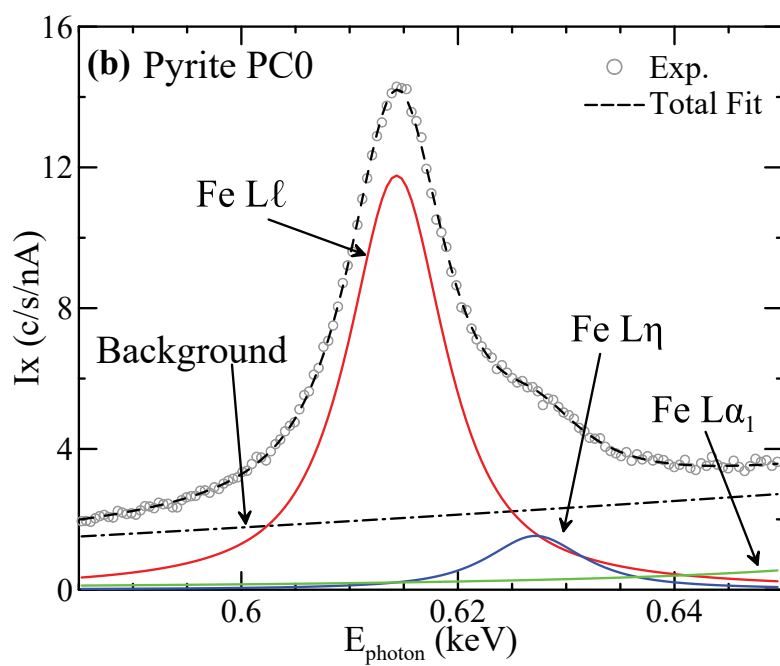
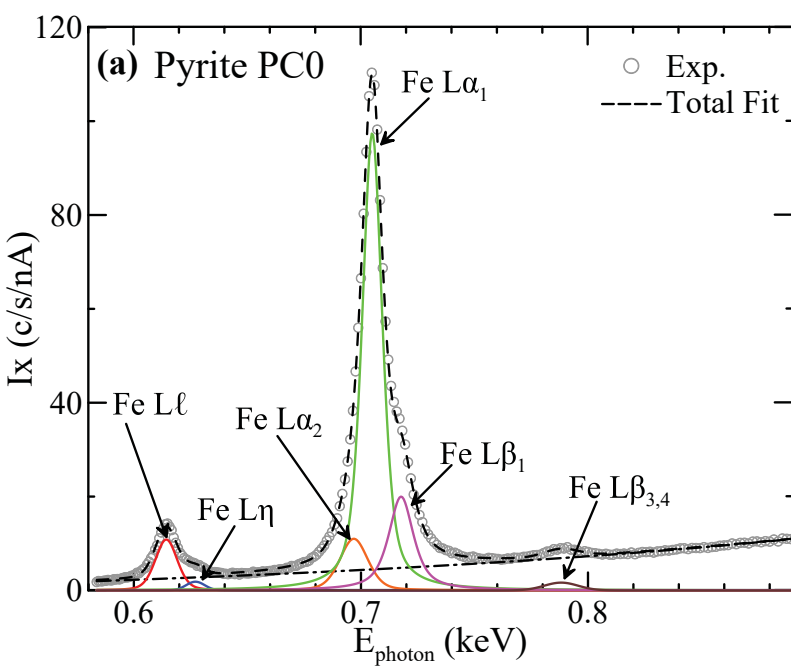
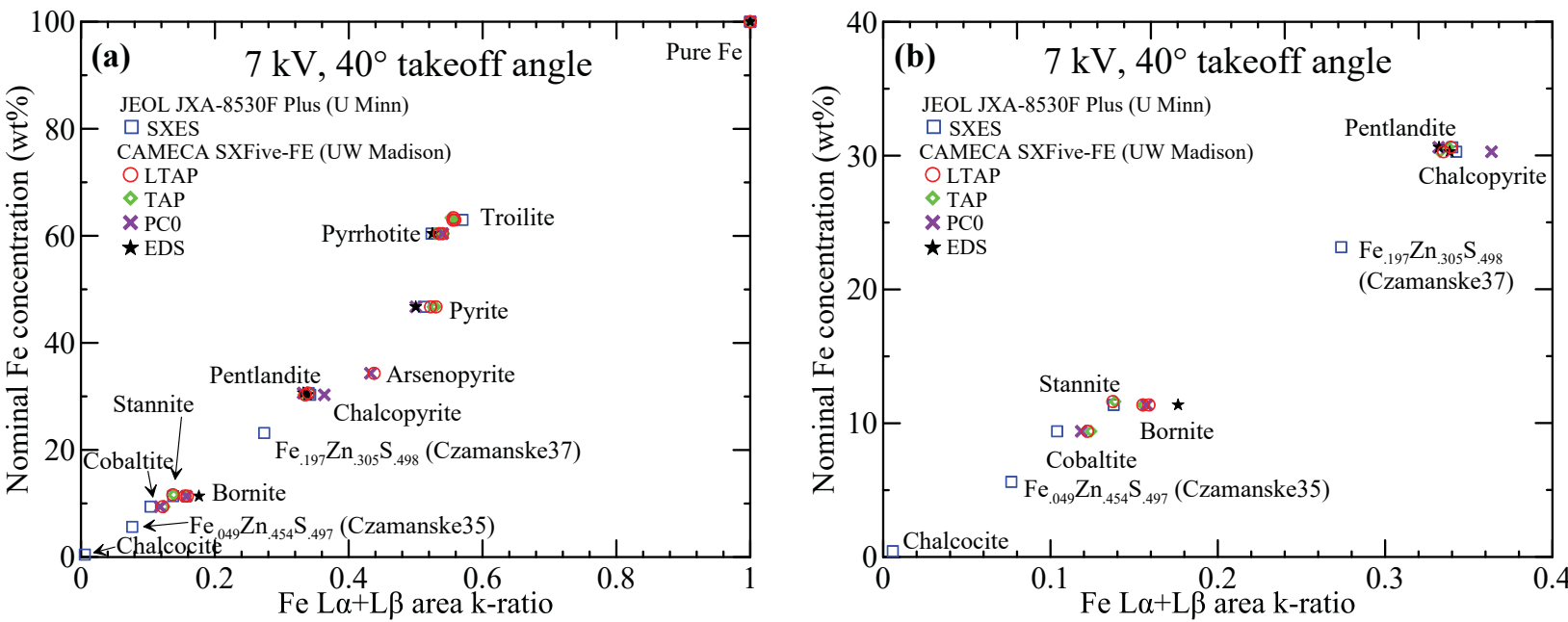


Figure 5





# Figure 6

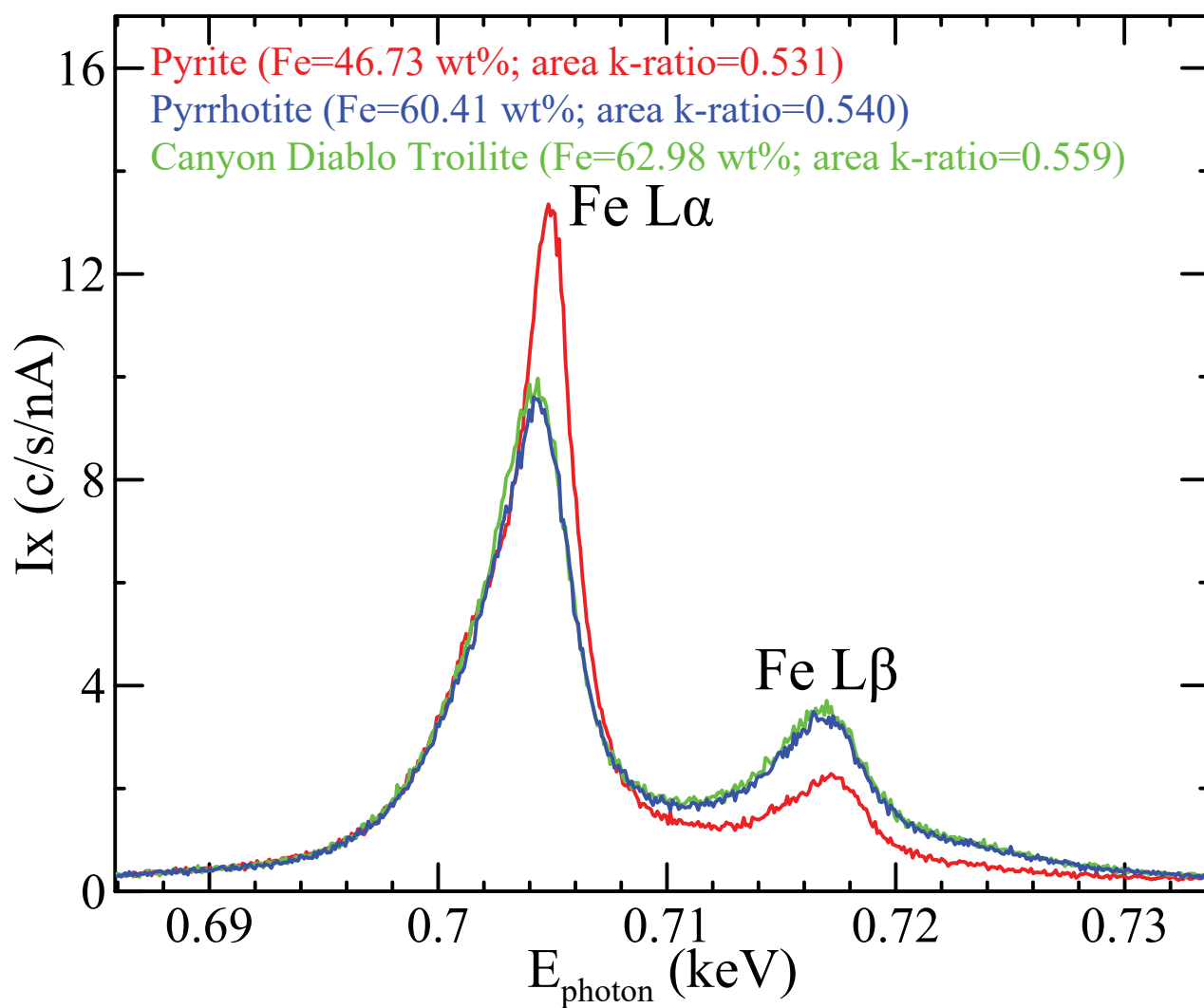
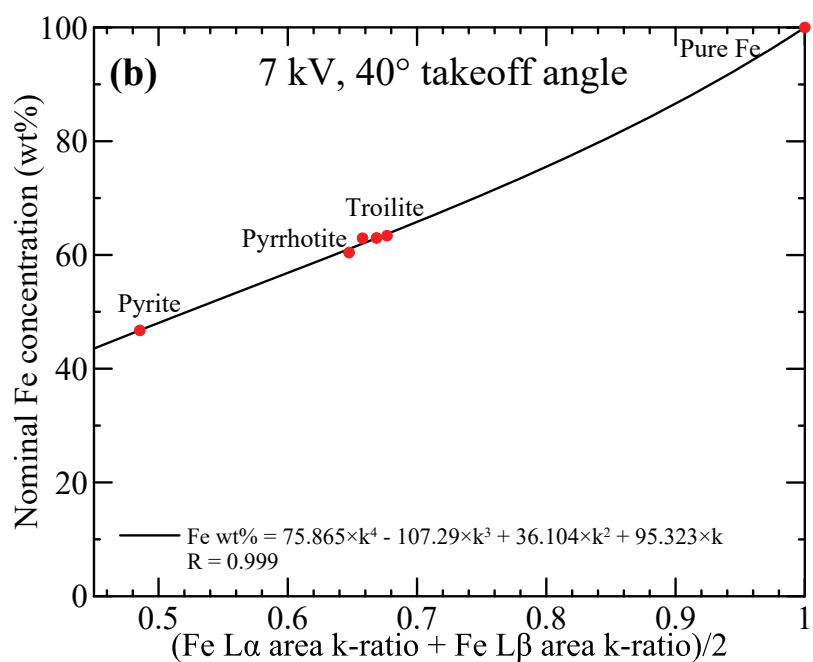
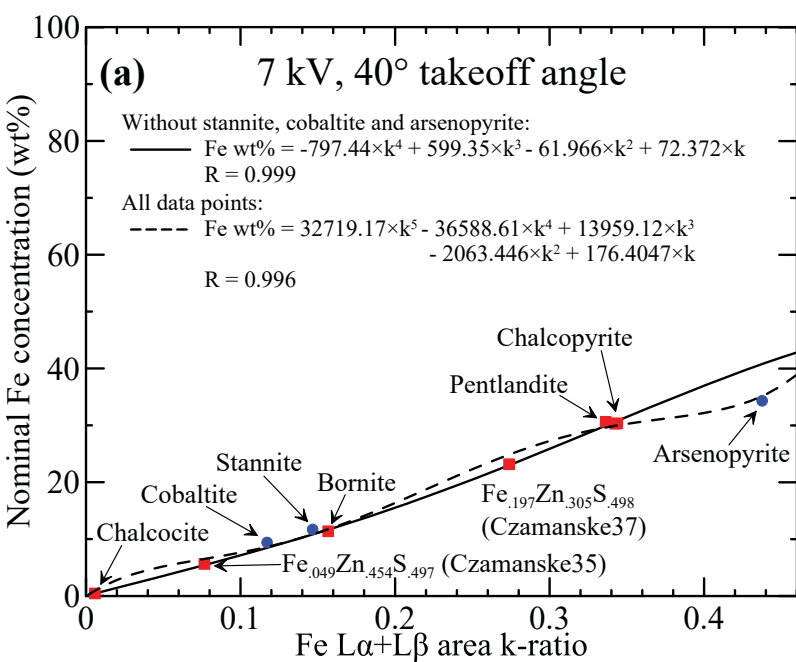


Figure 7



# Figure 8

