

1                   **Phase transitions in  $\epsilon$ -FeOOH at high pressure and ambient temperature**

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20 **Abstract**

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22 Constraining the accommodation, distribution, and circulation of hydrogen in the Earth's interior  
23 is vital to our broader understanding of the deep Earth due to the significant influence of  
24 hydrogen on the material and rheological properties of minerals. Recently, a great deal of  
25 attention has been paid to the high-pressure polymorphs of FeOOH (space groups  $P2_1nm$  and  
26  $Pnmm$ ). These structures potentially form a hydrogen-bearing solid solution with AlOOH and  
27 phase H ( $MgSiO_4H_2$ ) that may transport water ( $OH^-$ ) deep into the Earth's lower mantle.  
28 Additionally, the pyrite-type polymorph (space group  $Pa\bar{3}$  of FeOOH), and its potential  
29 dehydration have been linked to phenomena as diverse as the introduction of hydrogen into the  
30 outer core (Nishi *et al.* 2017), the formation of Ultra Low Velocity Zones (ULVZs) (Liu *et al.*  
31 2017), and the Great Oxidation Event (Hu *et al.* 2016). In this study, the high-pressure evolution  
32 of FeOOH was re-evaluated up to  $\sim 75$  GPa using a combination of synchrotron-based X-ray  
33 diffraction (XRD), Fourier transform infrared spectroscopy (FTIR), and optical absorption  
34 spectroscopy. Based on these measurements, we report three principal findings: (1) pressure-  
35 induced changes in hydrogen bonding (proton disordering or hydrogen bond symmetrization)  
36 occur at substantially lower pressures in  $\epsilon$ -FeOOH than previously reported and are unlikely to  
37 be linked to the high-spin to low-spin transition, (2)  $\epsilon$ -FeOOH undergoes a 10% volume collapse  
38 coincident with an isostructural  $Pnmm \rightarrow Pnmm$  structural transition at approximately 45 GPa,  
39 and (3) a pressure-induced band gap reduction is observed in FeOOH at pressures consistent with  
40 the previously reported spin transition (40 to 50 GPa).

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46 **Keywords:** Phase transitions, infrared-spectroscopy, optical absorption spectroscopy, X-ray  
47 diffraction, spin transition

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## 51 Introduction

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53 The accommodation, distribution, and circulation of hydrogen in the deep Earth is key to  
54 understanding the evolution of Earth's interior due to the significant influence of hydrogen on the  
55 material and rheological properties of high-pressure phases (e.g., Sarafian *et al.* 2017; Karato  
56 2010). Geophysical observations indicate that the Earth's upper mantle and transition zone, at  
57 least locally, host significant quantities of hydrogen (van der Meijde *et al.* 2003; Dixon *et al.*  
58 2004, Pearson *et al.* 2014; Palot *et al.* 2016; Tschauner *et al.* 2018). Additionally, tomographic  
59 evidence supports the idea that subducting lithospheric plates pierce the transition zone,  
60 potentially ushering water into the Earth's lower mantle (van der Hilst *et al.* 1997). Yet while  
61 hydrogen in the upper mantle and transition zone is hosted primarily in nominally anhydrous  
62 phases, these phases are not stable at the high pressure and temperature conditions of the lower  
63 mantle (Hirschmann 2006). If hydrogen introduced into the lower mantle remains in the lower  
64 mantle, it is almost certainly accommodated primarily in minor hydrous phases, because the  
65 dominant lower mantle minerals (bridgmanite, ferropericlase, calcium silicate perovskite) do not  
66 have the same capacity for water storage (Bolfan-Casanova *et al.*, 2002, 2003; Panero *et al.*,  
67 2015). A plausible lower mantle host is the isostructural FeOOH–AlOOH–MgSiO<sub>4</sub>H<sub>2</sub> system.

68 At moderate pressures (>6 GPa) goethite ( $\alpha$ -FeOOH, space group *Pbnm*), a widespread  
69 iron oxy-hydroxide, transforms into  $\epsilon$ -FeOOH, an orthorhombic (space group *P2<sub>1</sub>nm*, *Z* = 2)  
70 phase composed of edge-sharing FeO<sub>6</sub> units that are close-packed along the *c*-axis, with  
71 hydrogen atoms occupying the channels between these close-packed octahedra (Bendeliani *et al.*  
72 1972; Pernet *et al.* 1975; Bolotina *et al.* 2008) (Figure 1a). At increased pressure,  $\epsilon$ -FeOOH  
73 undergoes pressure-induced hydrogen bond symmetrization resulting in a second-order phase  
74 transition (space group *Pnmm*, *Z* = 2) (Figure 1c). However, the pressure at which this hydrogen  
75 bond symmetrization induced *P2<sub>1</sub>nm* → *Pnmm* transition occurs remains contested (Gleason *et*  
76 *al.* 2008; Xu *et al.* 2013; Thompson *et al.* 2017; Ikeda *et al.* 2019). At the pressures of the  
77 Earth's lower mantle,  $\epsilon$ -FeOOH forms a solid solution with  $\delta$ -AlOOH and phase H (MgSiH<sub>2</sub>O<sub>4</sub>),  
78 creating a viable hydrogen reservoir with *P-T* stability extending from lower mantle conditions  
79 to those of the core-mantle boundary (Sano *et al.* 2008; Nishi *et al.* 2015; Xu *et al.* 2019).

80 Hydrogen bond symmetrization is the process by which the hydrogen in an asymmetrical  
81 O—H···O bonding unit becomes centered with respect to the two oxygens with increased

82 pressure, as the longer hydrogen bridge bond ( $H\cdots O$ ) compresses more rapidly than the shorter  
83 but stiffer hydroxyl bond ( $O-H$ ) (Holzapfel 1972). Once the hydrogen bond is ‘symmetrized’,  
84 the two donor-acceptor distances are equivalent, and this shift in charge balance can lead to  
85 additional subtle structural changes (Figure 1c). In  $\epsilon$ -FeOOH and other isostructural MOOH  
86 phases, the  $MO_6$  octahedra shift from a slightly offset position to one centered along the two-fold  
87 axis, leading to the increase in symmetry from  $P2_1nm$  to  $Pnmm$  (e.g., Sano-Furukawa *et al.* 2009,  
88 2012). Alternatively, proton disorder in which both hydrogen and iron atoms are disordered over  
89 symmetric positions on the two-fold axis may also produce MOOH with  $Pnmm$  symmetry  
90 (Bolotina *et al.* 2008) with the resulting diffraction patterns nearly indistinguishable from  
91 symmetrized H-bond structures (e.g., Fujihara *et al.* 2002) (Figure 1b). Importantly, a recent  
92 neutron diffraction study by Sano-Furukawa *et al.* (2018), found that in  $\delta$ -AlOOH, which is  
93 isostructural to  $\epsilon$ -FeOOH, proton disordering is a precursor to hydrogen bond symmetrization.  
94 Accurately determining the pressure at which hydrogen bond symmetrization occurs is important  
95 because the phenomenon is linked to changes in compressibility (Vanpeteghem *et al.* 2003;  
96 Tsuchiya *et al.* 2005; Hushur *et al.* 2011) and due to isotopic effects may influence H/D  
97 fractionation in the deep Earth (e.g., Sano-Furukawa *et al.* 2009).

98 In addition to the previously described  $P2_1nm \rightarrow Pnmm$  transition,  $\epsilon$ -FeOOH reportedly  
99 undergoes a high-spin to low-spin transition, determined on the basis of X-ray diffraction (XRD)  
100 and X-ray emission spectroscopy (XES) experiments (Gleason *et al.* 2013). The XES results  
101 from Gleason *et al.* (2013) are indicative of a high-spin to low-spin transition initiated at  $\sim 40$   
102 GPa and completed at  $\sim 60$  GPa. However, that study did not use a pressure-transmitting  
103 medium, which introduced significant deviatoric stress into the sample, inhibiting the detection  
104 of a sharp transition. A contemporary study by Xu *et al.* (2013), which used a pressure-  
105 transmitting medium, pinpointed this electronic transition more closely to  $\sim 45$  GPa using a  
106 combination of Mössbauer spectroscopy, single-crystal diffraction, and resistivity measurements.  
107 In both studies, the authors propose causality between the hydrogen bond symmetrization and  
108 spin transition in  $\epsilon$ -FeOOH, although Gleason *et al.* (2013) report that hydrogen bond  
109 symmetrization induces the spin transition, whereas Xu *et al.* (2013) conclude the spin transition  
110 induces hydrogen bond symmetrization. However, both studies agree that a large ( $\sim 11\%$ ) volume  
111 collapse occurs coincident with the high-spin to low-spin transition, attributable to an  
112 isostructural phase transition. Significant volume reductions have been previously reported due

113 to high-spin to low-spin transitions without invoking hydrogen bond symmetrization, although a  
114 volume reduction of ~11% is higher than the volume reductions reported for either ferroperricite  
115 ((Mg,Fe)O) (1-3%) or ferromagnesite ((Mg,Fe)CO<sub>3</sub>) (6-10%) (Lin *et al.* 2013), possibly due to  
116 the larger relative volume occupied by iron atoms in FeOOH compared to these phases.

117 In this study, we present experimental evidence of two distinct transitions in  $\epsilon$ -FeOOH in  
118 the explored pressure range (0-70 GPa): a low-pressure (~18 GPa) second-order transition  
119 connected to hydrogen bond symmetrization and an independent first-order transition at ~45  
120 GPa. Additionally, we report evidence of a rapid reduction in the bandgap of  $\epsilon$ -FeOOH at  
121 pressures consistent with the previously reported spin transition.

122

## 123 **Methods**

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125 Synthesis of  $\epsilon$ -FeOOH for this study was performed by Akio Suzuki and is described in  
126 Suzuki (2010). The high pressures necessary for this study were achieved by compressing  $\epsilon$ -  
127 FeOOH samples using symmetric-type diamond anvil cells (DACs) with 250  $\mu$ m culet type II  
128 diamond anvils. Seventy-micron diameter sample chambers were laser ablated into stainless steel  
129 or rhenium gaskets pre-indented to 22 or 28 GPa, respectively. Powdered  $\epsilon$ -FeOOH, kept in a  
130 desiccator prior to loading, was pressed into platelets ~3-5  $\mu$ m thick and ~30 microns in  
131 diameter. Platelets of  $\epsilon$ -FeOOH were loaded into the sample chamber, utilizing gas-loaded Ne or  
132 ~10  $\mu$ m thick platelets of KBr as a pressure medium to reduce pressure gradients within the  
133 sample. Neon was loaded as a pressurized gas at the Advanced Photon Source using the  
134 COMPRES/GSECARS gas loading system (Rivers *et al.* 2008) and acted as a secondary  
135 pressure standard (Fei *et al.* 2007), with reported uncertainties based on the standard error of the  
136 (111), (200), and (220) *d*-spacings. Pressures were independently determined via the *in situ*  
137 monitoring of the *R<sub>1</sub>* luminescence line (Dewaele *et al.* 2008) of 2-3 ruby grains placed in the  
138 sample chamber, with errors reflecting the larger of either the standard deviation or 3%.

139 Room-temperature X-ray diffraction experiments were performed at beamline 13-ID-D  
140 (GSECARS) at the Advanced Photon Source, Argonne National Laboratory. A monochromatic  
141 ( $\lambda=0.3344$ ) incident X-ray beam was used, measuring 3  $\mu$ m by 4  $\mu$ m at full width at half  
142 maximum of the focused spot. Sample-to-detector distances and tilt were calibrated using 1-bar  
143 diffraction of LaB<sub>6</sub>. Diffraction patterns were integrated to produce 2 $\theta$  plots using DIOPTAS

144 (Prescher and Prakapenka 2015), positions of individual diffraction peaks were determined using  
145 PeakFit (Systat Software), and lattice parameters were calculated from the fitted  $d$ -spacings  
146 using the author's own Mathematica script. Individual peaks were fit to single Gaussian curves,  
147 and unresolved overlapping peaks were not used in the calculation of lattice parameters. All  
148 samples prepared for X-ray diffraction experiments were loaded using Ne as a pressure medium.

149 Fourier transform infrared (FTIR) spectra in this study were collected using both  
150 synchrotron and globar sources. Synchrotron experiments were performed at the 1.4.3 beamline  
151 of the Advanced Light Source at Berkeley Laboratory using a Nicolet 760 FTIR spectrometer  
152 with a custom microscope and HgCdTe detectors. Offline experiments were performed at  
153 Brookhaven National Laboratory using a Bruker Vertex 80v spectrometer and Hyperion 2000  
154 microscope with a MCT detector. All FTIR spectra presented here reflect a spectral resolution of  
155  $4\text{ cm}^{-1}$ , measured wavelength ranges of 500 to  $8000\text{ cm}^{-1}$ , and were recorded as 1024 or 512  
156 scans. PeakFit was used for background subtraction and to obtain precise absorption peak  
157 positions using a least-squares refinement, and interference fringes in absorbance spectra were  
158 reduced mathematically in DatLab. Samples prepared for FTIR experiments used either Ne or  
159 KBr as pressure medium, and the three samples that went to the highest pressures used KBr.

160 Optical absorption measurements were collected at the University of Chicago. Spectra  
161 were recorded from 420 to 980 nm using a 0.3 meter focal length Princeton Instruments SP-  
162 2300i spectrograph. Transmitted light ( $I$ ), generated with a tungsten filament bulb, was measured  
163 through the pressure medium (KBr) only, while incident light ( $I_0$ ) was collected through the  
164 sample and pressure medium. Background measurements were collected both with the light  
165 source off and with the light source on but measured at the gasket, which is opaque, with  
166 comparable results. Interference fringes in the absorbance spectra were removed mathematically  
167 in DatLab. To account for variation in sample thickness ( $h$ ), reported absorption values [ $\alpha = h^{-1}$   
168  $\ln(I_0/I)$ ] have been scaled such that  $I_0/I$  is equal for all samples at 10 GPa. Samples prepared for  
169 optical absorption measurements used KBr as pressure medium.

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## 171 **Results and Discussion**

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### 173 *X-ray Diffraction*

174

175 High-pressure, room-temperature X-ray diffraction (XRD) measurements were used  
176 re-evaluate the phase diagram of FeOOH up to 75 GPa, as the existing literature exhibits a  
177 paucity of volume–pressure ( $V$ – $P$ ) data in the mid-pressure range (20–50 GPa) (Gleason *et al.*  
178 2008, 2013; Suzuki 2010, 2016; Ikeda *et al.* 2019). Additionally, using a gas-membrane  
179 diaphragm to remotely pressurize the DAC between diffraction measurements without manual  
180 action, data could be collected with greater pressure resolution than in previous studies. As X-ray  
181 diffraction inherently relies on electron density, the detection of hydrogen using powder XRD is  
182 virtually impossible. The  $P2_1nm \rightarrow Pnnm$  transition is primarily defined by the pressure-induced  
183 symmetrization of the hydrogen bonds, but associated with hydrogen bond disordering and  
184 symmetrization are subtle changes in the structure and strain accommodation, which enable the  
185 detection of these transitions indirectly (e.g., Sano-Furukawa *et al.* 2009, 2012; Kuribayashi *et*  
186 *al.* 2014).

187 In the lower pressure interval (0.5–43 GPa) the lattice parameters of  $\epsilon$ -FeOOH were  
188 determined using at a minimum the (110), (101), (011), (020), (210), (211), (121), (220), (002),  
189 and (301)  $hkl$  peaks. The resulting unit cell volumes and lattice parameters are plotted in Figures  
190 2a and 2b and are reported in tabulated form in Table S.1. The measured lattice parameters and  
191 unit cell volumes of  $\epsilon$ -FeOOH in this pressure interval are in excellent agreement with the results  
192 of Suzuki (2010, 2016) and Ikeda *et al.* (2019), albeit slightly lower than those reported by the  
193 Gleason *et al.* (2013). The diffraction data from Gleason *et al.* (2013) exhibit more scatter, likely  
194 due to the lack of pressure medium in that study, which likely contributed to the slight offset in  
195 measured  $P$ – $V$  values.

196 Based on the Bragg peaks indexed in this study, the  $P2_1nm$  and  $Pnnm$  structure cannot be  
197 distinguished from one another, as even at the lowest pressures of this study no peaks were  
198 observed that violated the additional systematic absences characteristic of the  $Pnnm$  symmetry ( $h$   
199  $+ l$  odd for  $h0l$ ,  $k + l$  odd for  $0kl$ ). Additionally, based on the findings of Sano-Furukawa *et al.*  
200 (2009), who evaluated the  $P2_1nm \rightarrow Pnnm$  transition in isostructural  $\delta$ -AlOOH, the volume  
201 reduction associated with this second-order transition is likely unresolvable, rendering  
202 determinations of this phase boundary on this basis alone dubious. Thus, employing the approach  
203 of Sano-Furukawa *et al.* (2009, 2012), we evaluated the ratios of lattice parameters as a means of  
204 elucidating the subtle  $P2_1nm \rightarrow Pnnm$  transition.

205 The evolution of the  $a/b$  and  $b/c$  lattice parameter ratios (Figure 2c) reveal a change in the  
206 axial compression of  $\epsilon$ -FeOOH with pressure, and the inflection points evident at  $18\pm 1$  GPa  
207 indicate a shift in the accommodation of strain in this system. Similar changes in axial  
208 compression have been used previously as an indicator of second-order phase changes in similar  
209 materials (e.g., Sano-Furukawa *et al.* 2009, 2012; Kuribayashi *et al.* 2014). On the basis of these  
210 lattice parameter ratios, we confirm that  $\epsilon$ -FeOOH undergoes a second-order phase transition  
211 from the  $P2_1nm$  to  $Pnmm$  structure at  $18\pm 1$  GPa, in good agreement with previous predictions  
212 based on density functional theory-based calculations (Thompson *et al.* 2017) and recent  
213 experiments (Ikeda *et al.* 2019). As any volume reduction across this presumed phase boundary  
214 is minimal (Figure 2a), the  $V$ - $P$  data from the entire pressure interval (0 to 45 GPa) was fit to a  
215 third-order Birch-Murnaghan equation of state (EoS) (Birch 1978):

$$216 \quad P(V, T) = 3K_0 f_E \left[ (1 + 2f_E)^{\frac{5}{2}} \left( 1 + \frac{3}{2} [K'_0 - 4] f_E \right) \right]$$

217 which relates pressure ( $P$ ), volume ( $V$ ), ambient pressure bulk modulus ( $K_0$ ), and its pressure  
218 derivative ( $K'_0$ ) in terms of finite Eulerian strain ( $f_E$ ):

$$219 \quad f_E = \frac{1}{2} \left[ \left( \frac{V_0}{V} \right)^{\frac{2}{3}} - 1 \right]$$

220 which is a measure of the volume compression of a solid relative to its initial volume ( $V_0$ ).  
221 Residuals from this fit are shown in Figure S1.

222 To obtain EoS parameters with the most comprehensive data set available, the  $V$ - $P$  data  
223 from this study and Suzuki (2010, 2016) were combined, and the resultant equation of state  
224 parameters are presented alongside parameters fit to the data from only this study. In addition to  
225 fitting  $V$ - $P$  data from the entire 0-45 GPa pressure range, we also report EoS parameters derived  
226 from fitting the  $V$ - $P$  data of the  $P2_1nm$  and  $Pnmm$  structures separately. The resultant EoS  
227 parameters, as well as previously published results, are reported in Table 1. We find that EoS  
228 parameters derived from fitting the combined data set across the entire pressure range (0-45 GPa)  
229 are in reasonable agreement with those reported by Gleason *et al.* (2008) and Suzuki (2010,  
230 2016). However, fitting  $V$ - $P$  data for the  $P2_1nm$  and  $Pnmm$  structures independently (using a  
231 fixed  $K'_0$  value of 4 because fewer fitted parameters are justified over the limited pressure  
232 ranges), reveals that the  $Pnmm$  structure has a bulk modulus  $\sim 20\%$  higher than that of the  $P2_1nm$   
233  $\epsilon$ -FeOOH, in good agreement with the  $\sim 19\%$  predicted from DFT calculations (Thompson *et al.*



234 2017). These findings are also in good agreement with a prior study by Xu *et al.* (2013) that  
235 identified a reduction in the compressibility of FeOOH at ~16 GPa but did not connect this  
236 change in material properties to the onset of pressure-induced hydrogen bond symmetrization or  
237 proton disordering. Similar increases in bulk modulus have been reported for other hydrous  
238 phases due to the onset of pressure-induced hydrogen bond symmetrization or proton disordering  
239 (Vanpeteghem *et al.* 2003; Tsuchiya *et al.* 2005; Hushur *et al.* 2011).

240 Additionally, XRD data from this study revealed a second, independent structural  
241 transition in  $\epsilon$ -FeOOH at ~45 GPa. The transition occurs over a small (<3 GPa) pressure interval  
242 and is evident in the appearance of multiple new Braggs peaks and a simultaneous reduction in  
243 the intensity, and eventual disappearance, of the Braggs peaks of the high-spin *Pnmm* structure  
244 (Figure 3). This high-pressure (> 45 GPa) structure of FeOOH is stable at ambient temperatures  
245 and pressures exceeding 70 GPa. While an earlier study identified a potential unit cell volume  
246 reduction in this pressure range (Gleason *et al.* 2013), the improved pressure resolution of the  
247 current study clarifies the pressure at which this transition occurs. The sharpness of this phase  
248 boundary and the significant volume reduction contrast with the gradual, displacive second-order  
249 transition produced by the gradual change in charge density due to hydrogen bond  
250 symmetrization. Conversely, the structural transition observed at ~45 GPa is within the pressure  
251 range in which the high-spin to low-spin transition has been observed (Gleason *et al.* 2013), and  
252 the contraction of the Fe<sup>3+</sup> ionic radius across the high-spin to low-spin transition has been  
253 reported to produce significant volume reductions in other iron-bearing minerals (e.g., Lin *et al.*  
254 2013). Although the <3 GPa interval in which the first-order transition is observed in this study  
255 is significantly narrower than the 20 GPa pressure interval across which the high-spin to low-  
256 spin transition was reported (Gleason *et al.* 2013), this discrepancy may be due to the lack of a  
257 pressure medium in the earlier study, an omission that can produce pressure gradients that could  
258 lead to significantly broader spin transitions.

259 High-pressure (>45 GPa) X-ray diffraction data from this study were indexed to a *Pnmm*  
260 ( $Z=2$ ) unit cell, using at a minimum the (110), (200), (011), (121), (310), (301), and (112) *d*-  
261 spacings. The resultant lattice parameters and unit cell volumes are consistent with a volume  
262 reduction of 10% without a loss of symmetry, in reasonable agreement with the isostructural,  
263 11% volume collapse proposed by Gleason *et al.* (2013) (Figure 2a). However, a recent *ab initio*  
264 study has predicted an intermediate *Pbca* ( $Z=8$ ) structure is stable between the *Pnmm* and the

265 pyrite ( $Pa\bar{3}$ ) phases of AlOOH (Verma *et al.* 2018). The high-pressure (>45 GPa) XRD data  
266 from this study was indexed to the proposed  $Pbca$  structure, but the reduction in volume across  
267 this potential  $Pnmm \rightarrow Pbca$  transition appears implausibly large (26%) and resulted in a reduced  
268 quality of fit. An attempt was made to index the high-pressure Bragg peaks to the  $\alpha$ -PbO<sub>2</sub>  
269 structure ( $Pbcn$ ), a structure adopted by CaCl<sub>2</sub>-type ( $Pnmm$ ) oxides at high pressures, but this  
270 failed to produce a reasonable fit to the data. Additionally, we can definitively state that this is  
271 not the appearance of the high-pressure pyrite structure of FeOOH, to which  $\epsilon$ -FeOOH is  
272 expected to transition at comparable pressures and elevated temperatures (Nishi *et al.* 2017; Hu  
273 *et al.* 2016). Therefore, we interpret the first-order transition at  $45 \pm 2$  GPa as an isostructural  
274 volume collapse (i.e.,  $Pnmm \rightarrow Pnmm$ ), accompanied by an increase in the  $a/b$  ratio in the higher-  
275 pressure structure (Figure 2d). The change in the  $a/b$  ratio may be due to a relaxation of the  
276 distortion of the FeO<sub>6</sub> units, in agreement with Xu *et al.* (2013) and linked to the previously  
277 reported high-spin to low-spin transition (Gleason *et al.* 2013; Xu *et al.* 2013). In the future,  
278 single crystal X-ray diffraction structure refinements might help elucidate the nature of this  
279 structural transition (displacive versus reconstructive) and refine the symmetry of the high  
280 pressure, low-spin structure.

281

### 282 *Infrared and Visible Absorption Spectroscopy*

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284 Room temperature infrared absorption spectra of  $\epsilon$ -FeOOH were collected at ambient  
285 pressure and at regularly increasing pressure intervals of  $\sim 2$  GPa up to  $\sim 50$  GPa. Qualitatively,  
286 the ambient pressure  $\epsilon$ -FeOOH spectrum bears similarity to that of  $P2_1nm$   $\beta$ -CrOOH, with which  
287  $\epsilon$ -FeOOH is isosymmetric (Jahn *et al.* 2012) (Figure 4). At ambient pressure, the O-H stretching  
288 region presents as a broadband feature ( $2600$ - $3100$  cm<sup>-1</sup>), consistent with previously reported  
289 spectra of low-pressure polymorphs of FeOOH (e.g., Williams and Garnero 1996; Kagi *et al.*  
290 2008), and with increased pressure, the broadband O-H stretching region further broadens and  
291 shifts to lower wavenumbers. Unfortunately, the increased breadth of the O-H stretching feature  
292 and its increased proximity to the diamond absorption region prevents tracking the pressure  
293 dependence of this feature above  $\sim 17$  GPa. Although the O-H stretching band could only be  
294 monitored over a limited pressure range, its negative frequency shift pressure dependence was  
295 clearly observable at low pressures (Figure S3). A linear fit of the pressure-frequency data of the

296 O-H stretching band results in an approximate pressure dependence of  $-14 \text{ cm}^{-1}\text{GPa}^{-1}$ ,  
297 demonstrative of a softening behavior consistent with eventual hydrogen bond symmetrization.  
298 However, it isn't feasible to predict the pressure of eventual hydrogen bond symmetrization, as  
299 the relationship between the 1-bar stretching frequency and O-H pressure dependence is poorly  
300 understood in  $\epsilon$ -FeOOH and similarly structured phases, and the pressure range of available  $\epsilon$ -  
301 FeOOH O-H stretching frequency data is limited.

302 The ambient pressure infrared spectrum of  $\epsilon$ -FeOOH also contains three distinct  
303 absorption bands at 965, 1145, 1175  $\text{cm}^{-1}$  that were identified as O-H bending vibrations based  
304 on comparison to similar structures (Williams and Garnero, 1996; Mashino *et al.* 2016; Pinney  
305 and Morgan 2013). Spectra at successive pressure steps were evaluated to determine the  
306 frequency shifts in the O-H bending as a function of pressure (Figure 5). As with the OH-  
307 stretching region, these features broaden with increased pressure, and the two higher frequency  
308 bands become indistinguishable above  $17.5 \pm 1$  GPa. The lowest-frequency bending band remains  
309 distinct to higher pressures, but the pressure dependence of this band changes dramatically above  
310  $17.5 \pm 1$  GPa. The sudden change in the pressure dependence of this band suggests a reorientation  
311 of the O-H bonds, suggestive of the onset of pressure-induced hydrogen bond symmetrization or  
312 proton disordering. A similar shift in the pressure dependence of the OH-absorption bands in  
313 isostructural  $\delta$ -AlOOH was observed by Kagi *et al.* (2010) at approximately 10 GPa, consistent  
314 with the onset of proton disordering in that phase (Sano-Furukawa *et al.* 2018).

315 During the course of these IR measurements, a reduction in the transmitted IR signal was  
316 observed in  $\epsilon$ -FeOOH samples compressed above  $\sim 40$  GPa (Figure S4). In addition to the  
317 changes in IR transmission properties, changes in the optical properties of FeOOH were  
318 observed upon compression. The samples were translucent orange at ambient and low pressures,  
319 but reddened with increased pressure and became opaque to transmitted visible light at 45 GPa,  
320 slightly preceding the loss of IR transmission (Figure S4). To quantify these observations, visible  
321 to near-infrared absorption measurements of  $\epsilon$ -FeOOH were collected up to pressures exceeding  
322 70 GPa, to track the pressure dependence of the absorption edge. Representative high-pressure  
323 optical absorption spectra of  $\epsilon$ -FeOOH are shown in Figure 6a and indicate a systematic  
324 pressure-induced increase in absorption concordant with qualitative observations. Normalized  
325 data for five samples are plotted in Figure 6b, in which there are two key features. First, each  
326 sample underwent a transient minimum in optical transmission between 15-20 GPa before

327 gradually recovering to previous values over the subsequent ~5 GPa. Second, all five samples  
328 became opaque in the light range probed (1.45 eV to 2.05 eV) above 50 GPa (Figure 6b), above  
329 which the ratio of transmitted light ( $I$ ) to incident light ( $I_0$ ) that can propagate through the sample  
330 decreases to zero. These vis/NIR absorption findings suggest a significant reduction in the  
331 bandgap in this pressure interval.

332

### 333 **Implications**

334

335 Close evaluation of the high-pressure behavior of  $\epsilon$ -FeOOH at ambient temperature  
336 indicates two sequential, structural phase transitions in the 0-75 GPa pressure range. A second-  
337 order transition at  $18 \pm 1$  GPa was determined on the basis of three observations: (1) a change in  
338 axial compressibility evident in the lattice parameter ratios measured using powder XRD, (2) a  
339 change in the OH-bending modes observed with FTIR spectroscopy, and (3) a transient decrease  
340 in optical absorption. In accordance with the previous work of Sano-Furukawa *et al.* (2009,  
341 2012, 2018) that evaluated isostructural MOOH phases, this  $18 \pm 1$  GPa transition is the  $P2_1nm \rightarrow$   
342  $Pnmm$  transition produced by either pressure-induced hydrogen bond symmetrization or proton  
343 disordering, with future neutron diffraction work needed to differentiate between these  
344 mechanisms. Furthermore, this transition is observed to lead to a ~20% increase in bulk modulus  
345 and could be linked to the increase in compressional velocity ( $V_P$ ) reported by Ikeda *et al.*  
346 (2019).

347 At  $45 \pm 2$  GPa,  $Pnmm$   $\epsilon$ -FeOOH undergoes a first-order structural phase transition,  
348 identifiable due to the appearance of new Bragg peaks coincident with the disappearance of the  
349 high-spin  $Pnmm$  peaks. This high-pressure structure of FeOOH can also be indexed to  $Pnmm$  with  
350 a ~10% volume reduction, in good agreement with previous studies (Gleason *et al.* 2013; Xu *et*  
351 *al.* 2013). Therefore, we propose that the first-order transition at  $45 \pm 2$  GPa is a nearly  
352 isostructural volume collapse (i.e.,  $Pnmm \rightarrow Pnmm$ ), accompanied by a small (~1%) increase in  
353 the  $a/b$  ratio in the higher-pressure structure (Figure 2d). As such, low-spin  $Pnmm$  FeOOH is  
354 stable across a wide pressure range (more than 30 GPa). However, the phase diagram of FeOOH  
355 at high pressures and moderate temperatures should be assessed to determine the relevance of  
356 low-spin  $Pnmm$  FeOOH to the Earth's interior, as the formation of the pyrite ( $Pa\bar{3}$ ) structure is  
357 reported at comparable pressures and high temperatures (Nishi *et al.* 2017; Hu *et al.* 2016). In the

358 future, single crystal X-ray diffraction structure refinements may be beneficial in verifying the  
359 symmetry of this higher pressure, low-spin structure.

360 In addition to the first-order structural transition identified in FeOOH at ~45 GPa, a color  
361 change (translucent orange to black) was observed at high pressure, and the optical opacity of  
362 *Pnmm*  $\epsilon$ -FeOOH was quantified using visible/NIR transmission and infrared absorption. Visible  
363 to near-infrared absorption measurements of  $\epsilon$ -FeOOH also identified a systematic pressure-  
364 induced increase in absorption consistent with band gap reduction in the high-spin *Pnmm* phase,  
365 although complete band gap collapse (i.e., metallization) of this phase cannot be established on  
366 the basis of these experiments alone. FeOOH becomes opaque in the 1.45-2.05 eV range and  
367 also in the mid-IR at 45 GPa, indicating a change in optical properties along with the high-spin to  
368 low-spin transition in this material, as both phenomena are tied to changes in the d orbitals of the  
369 iron atoms. In fact, a similar increase in optical opacity has been previously linked to the high-  
370 spin to low-spin transition in siderite (Lobanov *et al.* 2015) and ferropericlasite (Keppler *et al.*  
371 2007). The apparent width of this transition, when compared to the sharpness of the structural  
372 transition observed using diffraction, may be attributable to the difference in the hydrostaticity of  
373 the pressure medium (KBr and Ne, respectively).

374 These results demonstrate that bandgap reduction, the previously observed high-spin to  
375 low-spin spin transition (Gleason *et al.* 2013; Xu *et al.* 2013), and the first-order phase transition  
376 in  $\epsilon$ -FeOOH all occur at a pressure ( $45 \pm 2$  GPa) greatly exceeding the  $P2_1nm \rightarrow Pnmm$  symmetry  
377 transition associated with a change in hydrogen bonding ( $18 \pm 1$  GPa). Therefore, hydrogen-bond  
378 symmetrization is not the driver of spin transition in  $\epsilon$ -FeOOH or vice versa. Rather, FeOOH  
379 undergoes a second-order  $P2_1nm \rightarrow Pnmm$  symmetry change caused by a shift in hydrogen  
380 bonding at 18 GPa, and a subsequent, unrelated first-order transition at 45 GPa that involves a  
381 significant volume change, spin transition, and optical opacity. Furthermore, although  
382 endmember  $\epsilon$ -FeOOH is expected to transition to the pyrite ( $Pa\bar{3}$ ) structure at the concurrent  
383 pressures and high temperatures of the Earth's mantle (Nishi *et al.* 2017; Hu *et al.* 2016),  
384 aluminum substitution is expected to increase the thermodynamic stability of (Fe,Al)OOH such  
385 that the low-spin *Pnmm* structure may be stable at the conditions of the deep Earth (e.g., Nishi *et al.*  
386 *et al.* 2015; Xu *et al.* 2019). We anticipate that the high-spin to low-spin spin transition and  
387 associated volume collapse in intermediate compositions in the FeOOH–AlOOH–MgSiH<sub>2</sub>O<sub>4</sub>  
388 system are likely to occur at lower pressures than in the Fe-endmember, as has been observed in

389 ferropericlasite (Fei *et al.* 2007). Additional experiments are needed to assess the influence of  
390 cation substitution on the phase boundaries of the FeOOH–AlOOH–MgSiO<sub>4</sub>H<sub>2</sub> system, as this  
391 may play a pivotal role in the transport of water (OH<sup>-</sup>) deep into the Earth’s lower mantle (Ohira  
392 *et al.* 2014, Xu *et al.* 2019) and may contribute to geophysical heterogeneities observed in the  
393 deep Earth (Thompson *et al.* 2017; Ohira *et al.* 2019).

394

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396

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412

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594 **Tables**

595

596 **Table 1.** Birch-Murnaghan equation of state (EoS) parameters of  $\epsilon$ -FeOOH, including ambient-  
 597 pressure unit volume ( $V_0$ ), bulk modulus ( $K_0$ ), and the pressure derivative of the bulk modulus  
 598 ( $K_0'$ ). Also included is the pressure range of the included room-temperature XRD data and the  
 599 number of  $V$ - $P$  data points included in each EoS fitting. The first and second lines are  
 600 parameters obtained from fitting the same data to second- and third-order BM EOSs,  
 601 respectively. Values in parentheses are uncertainties on the last digit. Residuals from the fit to a  
 602 third-order Birch-Murnaghan equation of state are shown in Figure S1.

	Pressures (GPa)	# Data points	$V_0$ ( $\text{\AA}^3$ )	$K_0$ (GPa)	$K_0'$
This study	0 to 45	59	65.87(3)	152(4)	4.9(2)
This study	0 to 45	59	65.57(6)	170(1)	4 (fixed)
This study, Suzuki 2010, 2016	0 to 45	80	66.15(5)	141(3)	5.4(2)
This study, Suzuki 2010, 2016	0 to 45	80	65.88(4)	163(1)	4 (fixed)
This study, Suzuki 2010, 2016 ( $P2_1nm$ )	0 to 16	29	66.14(4)	146(2)	4 (fixed)
This study ( $Pnm$ )	18 to 45	50	65.3(1)	176(3)	4 (fixed)
Gleason <i>et al.</i> 2008	8 to 17	9	66.3(5)	158(5)	4 (fixed)
Suzuki 2010	0 to 9	8	66.20(3)	126(3)	10(1)
Suzuki 2016	0 to 11	13	66.278(6)	135(3)	6.1(9)

603

604 **Figure 1.** Low and moderate pressure structures of  $\epsilon$ -FeOOH including the (a)  $P2_1nm$  structure  
 605 with ordered hydrogen occupation and asymmetric hydrogen bonds, (b)  $Pnm$  structure with  
 606 disordered hydrogen (hydrogen positions are 50% occupied), and (c)  $Pnm$  structure with  
 607 symmetric hydrogen bonds. Dashed lines indicate unit cells, oxygen atoms are red spheres,  
 608 hydrogen atoms are grey spheres, and  $\text{FeO}_6$  units are tan polyhedra, which are stacked in edge-  
 609 sharing chains parallel to the  $c$ -axis. Image generated in VESTA [Momma and Izumi, 2008]

610

611 **Figure 2.** (a) Unit cell volume, (b) lattice parameters, (c)  $a/b$  and (d)  $b/c$  lattice parameter ratios  
612 of FeOOH as a function of pressure, including this study (solid symbols), Suzuki 2010 (black  
613 crosses), Suzuki 2016 (black X's), Gleason *et al.*, 2008 (open diamonds), Gleason *et al.* 2013  
614 (open triangles), and Ikeda *et al.* 2019 (open circles). Different solid symbols (circles, diamonds,  
615 squares) are used to show replicate samples in this study. Green, purple, and yellow symbols  
616 indicate that Bragg peaks have been indexed to the  $P2_1nm$  (green), high-spin  $Pnmm$  (purple),  
617 and low-spin  $Pnmm$  (yellow) structures. Light green and light purple lines in (a) represent the  
618 equations of state using the parameters from lines 3 and 4 of Table 1, respectively. Error bars  
619 reflect uncertainties in the lattice parameters from this study, and when not visible reflect that  
620 errors are smaller than the symbols.

621  
622 **Figure 3.** Integrated X-ray diffraction patterns of FeOOH at five increasing pressure steps from  
623 bottom to top: 44.3 GPa (red), 44.6 GPa, 45.0 GPa, 45.9 GPa, and 46.6 GPa (black), illustrating  
624 the structural transition in this pressure interval. Miller indices of  $\epsilon$ -FeOOH are shown in red for  
625 the lower-pressure  $Pnmm$  structure, and in black for the higher-pressure, presumably low-spin,  
626 structure that is approximately 10% smaller. The Miller indices for the pressure medium (Ne)  
627 and gasket material (Re) are shown in grey.

628  
629 **Figure 4.** Ambient pressure infrared spectra of  $P2_1nm$   $\epsilon$ -FeOOH in the (a) lattice and OH-  
630 bending vibrations region and (b) OH-stretching region.

631  
632 **Figure 5.** Evolution of the O-H bending frequencies of FeOOH as a function of pressure. The  
633 pressure of the  $P2_1nm \rightarrow Pnmm$  transition is indicated by a black dashed line. Different colors are  
634 used to distinguish the five replicate samples from this study using either Ne as a pressure  
635 medium (red, orange) or KBr as a pressure medium (green, blue, purple).

636  
637 **Figure 6.** Optical absorption data, including (a) representative optical transmission spectra,  
638 indicating the pressure dependent increase in the absorption of  $\epsilon$ -FeOOH, and (b) data from five  
639 samples of  $\epsilon$ -FeOOH, each indicated with a different symbol (diamonds, squares, circles,  
640 triangles, and horizontal dashes), showing the change in the ratio of transmitted light ( $I$ ) to

641 incident light ( $I_0$ ) as a function of pressure (X-axis) and the wavelength of the incident light,  
642 which is indicated by color.

643

#### 644 **Supplemental Information**

645

646 **Table S1.** Lattice parameters and unit cell volumes of  $\epsilon$ -FeOOH based on X-ray diffraction  
647 measurements. Values in parentheses reflect uncertainties. (ATTACHED CSV FILE)

648

649 **Figure S1.** Residuals from the third-order Birch-Murnaghan equation of state fit to  $\epsilon$ -FeOOH  
650 volume-pressure data from 0 to 45 GPa (Table 1, line 1). Closed black symbols show data  
651 collected during compression and open symbols show data collected during decompression.

652

653 **Figure S2.** Eulerian strain ( $f$ ) versus the normalized pressure ( $F$ ) of  $\epsilon$ -FeOOH, using a  $V_0$  of  
654  $65.87 \text{ \AA}^3$  (Table 1, line 1). Closed black symbols show data collected during compression and  
655 open symbols show data collected during decompression.

656

657 **Figure S3.** Stacked high-pressure IR spectra of  $\epsilon$ -FeOOH including (a) lattice and OH-bending  
658 vibrations region and (b) OH-stretching region. Spectra are color coded according to pressure,  
659 which are indicated (in GPa) in the legend.

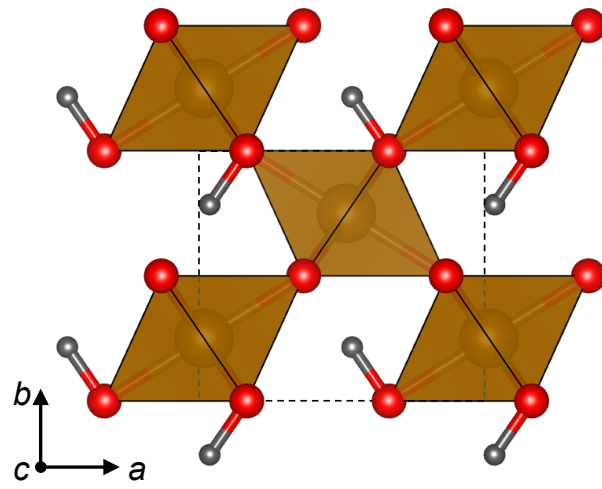
660

661 **Figure S4.** The % maximum FTIR signal intensity collected through the FeOOH sample as a  
662 function of pressure. Filled and open symbols are use to indicate duplicate samples used in this  
663 study. Dashed red line indicates the predicted pressure of the HS  $\rightarrow$  LS transition from Gleason  
664 *et al.* 2013.

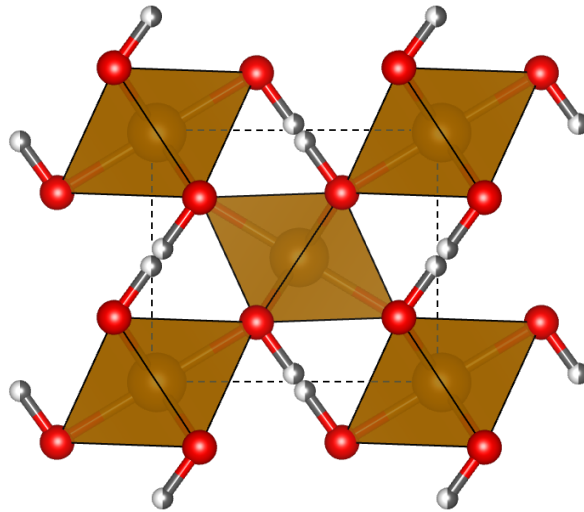
665

Figure 1

a)  $P2_1nm$ , ordered hydrogen



b)  $Pnmm$ , disordered hydrogen



c)  $Pnmm$ , symmetric hydrogen

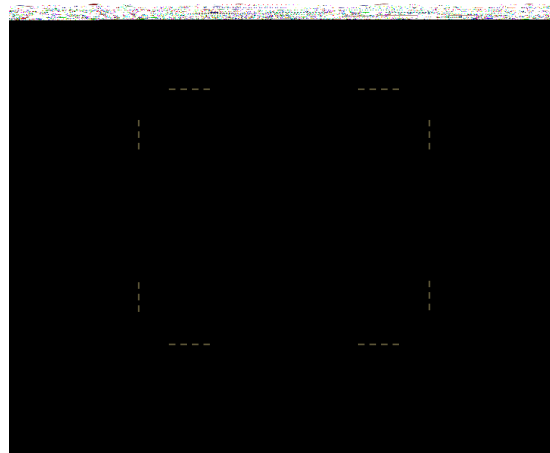




Figure 2

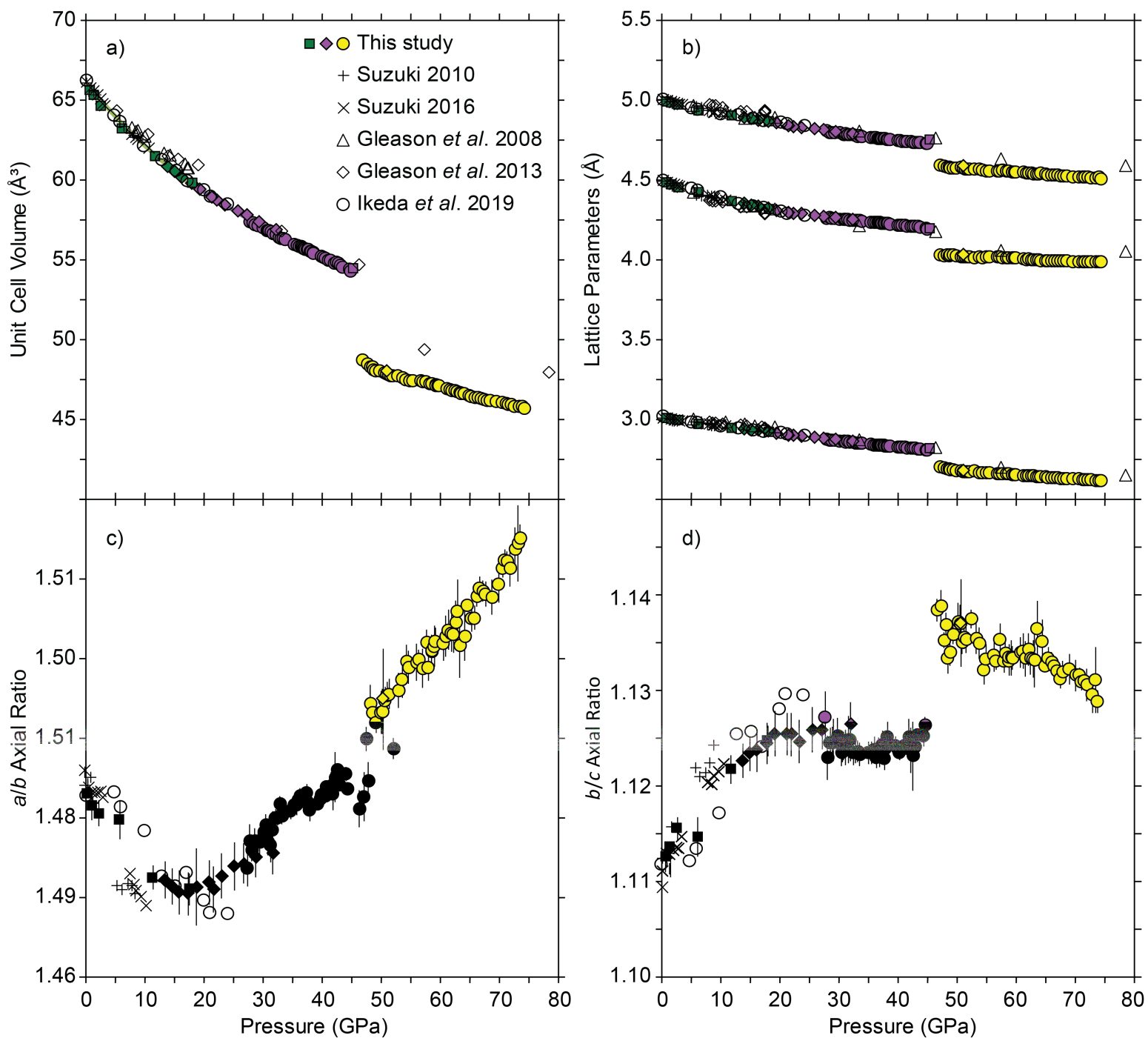
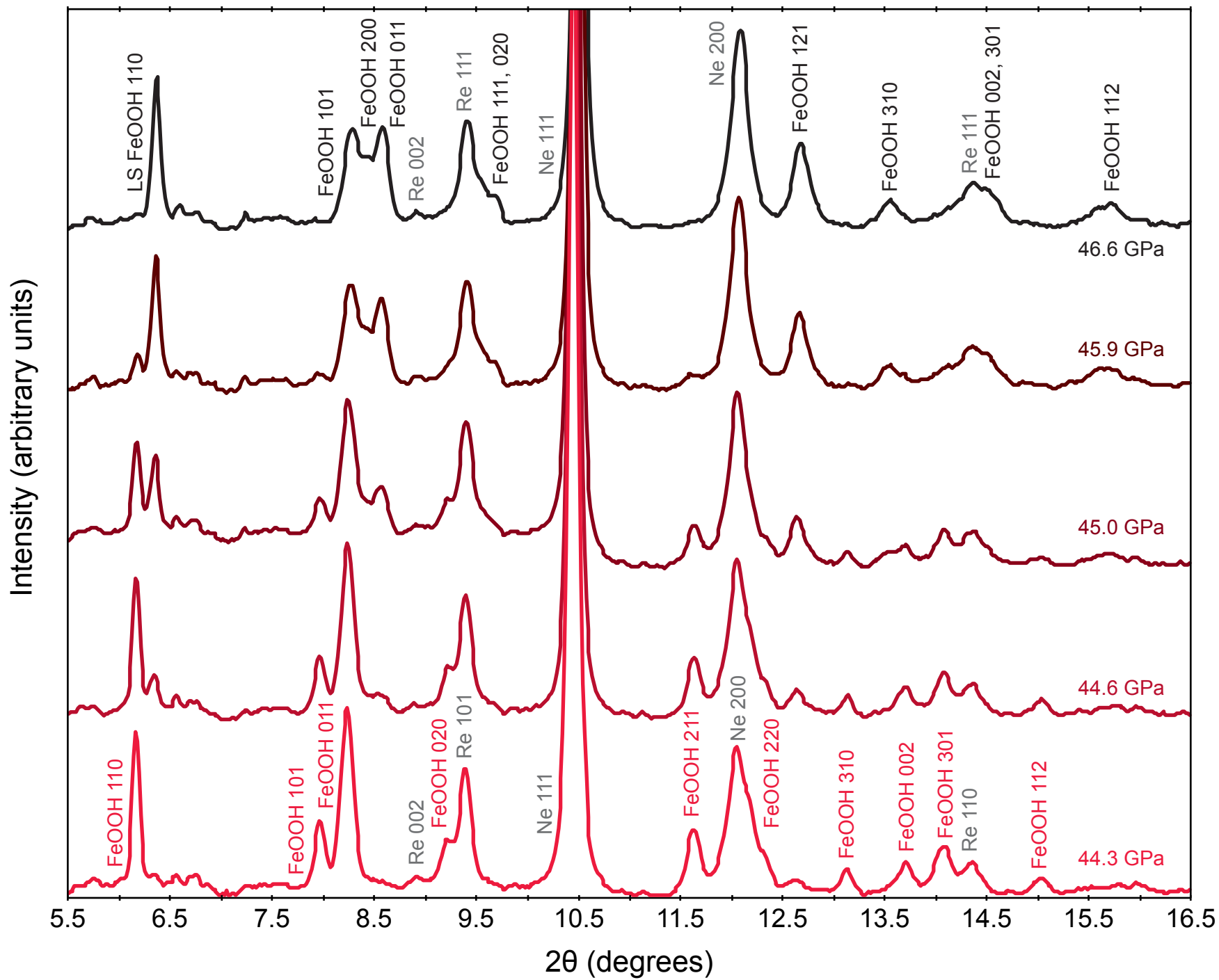
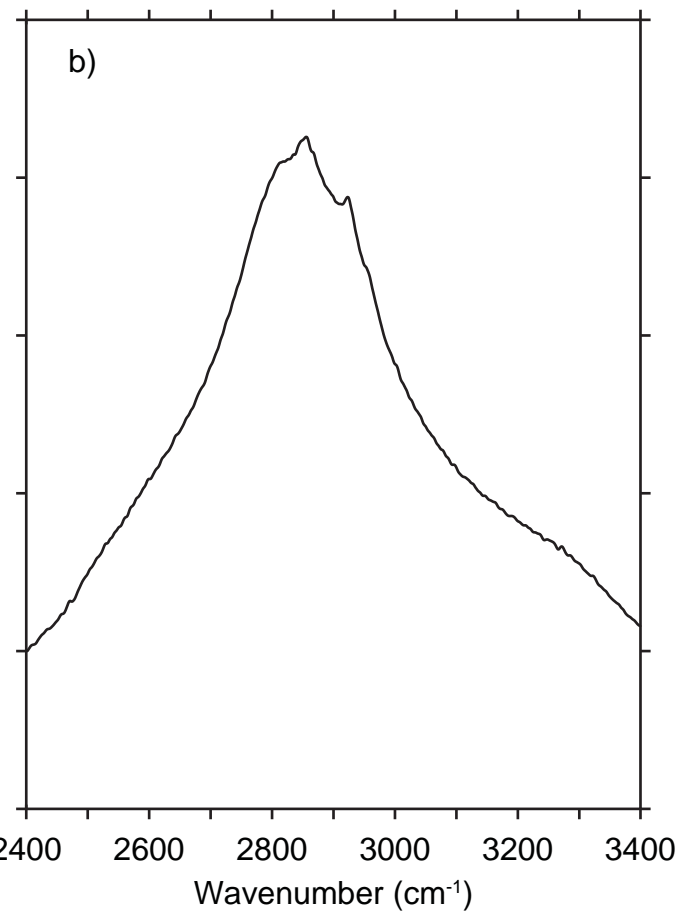
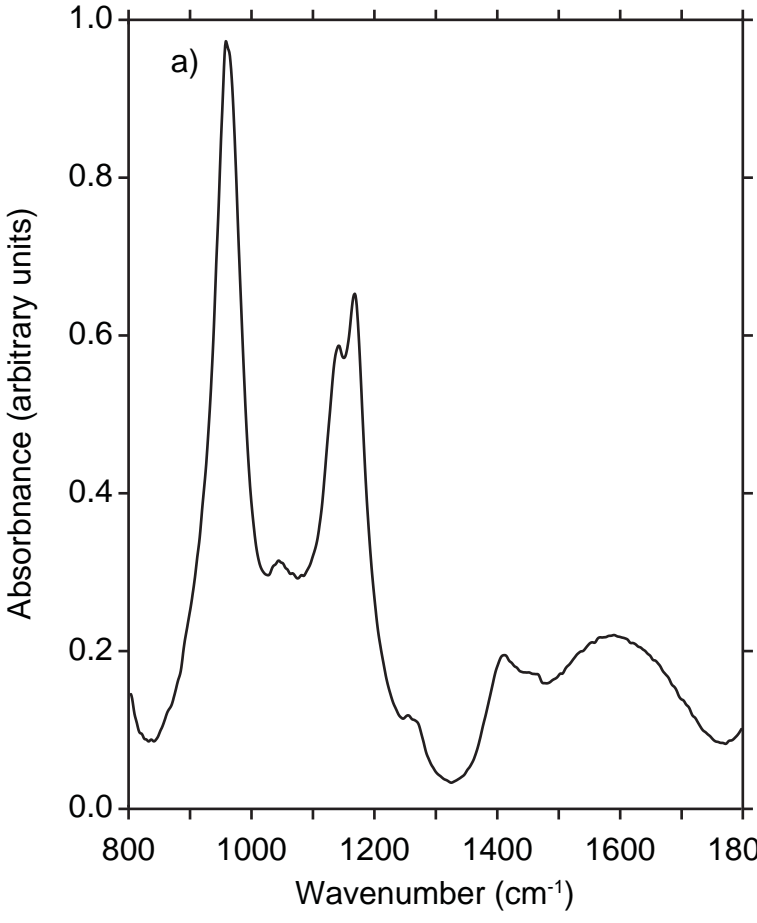
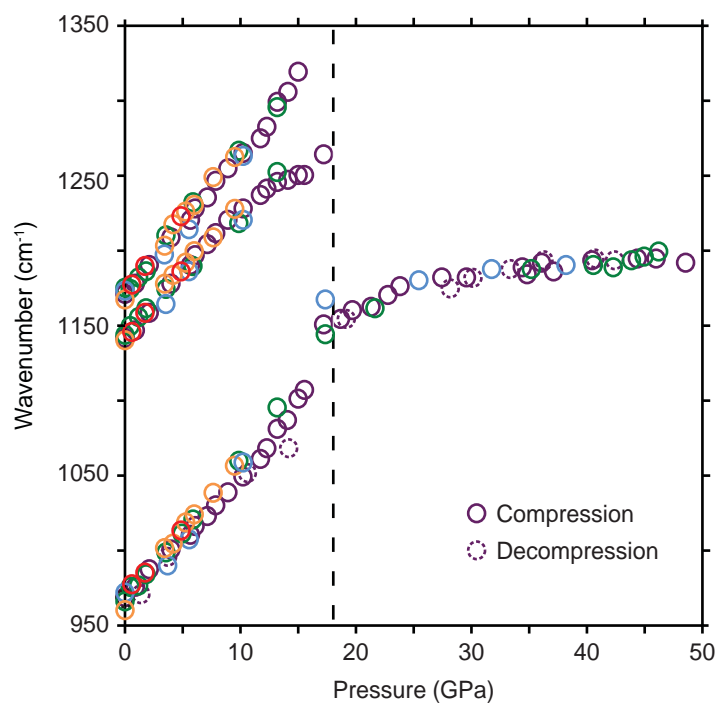


Figure 3







# Figure 6

