1	Revision 1
2	Gasparite-(La), La(AsO4), a new mineral from Mn ores of the Ushkatyn-III deposit,
3	Central Kazakhstan and metamorphic rocks of the Wanni glacier, Switzerland
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Abstract

21 Gasparite-(La), La(AsO₄), is a new mineral (IMA 2018-079) from Mn ores of the Ushkatyn-22 III deposit, Central Kazakhstan (type locality) and from alpine fissures in metamorphic rocks 23 of the Wanni glacier, Binn Valley, Switzerland (co-type locality). Gasparite-(La) is named for its dominant lanthanide, according to current nomenclature of rare-earth minerals. 24 25 Occurrence and parageneses in both localities is distinct: minute isometric grains up to 15 µm 26 in size, associated with fridelite, jacobsite, pennantite, manganhumite series minerals 27 (alleghanyite, sonolite), sarkinite, tilasite and retzian-(La) are typically embedded into 28 calcite-rhodochrosite veinlets (Ushkatyn-III deposit), versus elongated crystals up to 2 mm in 29 size in classical alpine fissures in two-mica gneiss without indicative associated minerals 30 (Wanni glacier). Its chemical composition has been studied by EDX and WDX; crystal-31 chemical formulas of gaspatite-(La) from the Ushkatyn-III deposit (holotype specimen) and 32 Wanni glacier specimen) $(La_{0.65}Ce_{0.17}Nd_{0.07}Ca_{0.06}Mn_{0.05}Pr_{0.02})_{1.03}$ (cotype are 33 $((As_{0.70}V_{0.28}P_{0.02})_{1.01}O_4)$ and $(La_{0.59}Ce_{0.37}Nd_{0.02}Ca_{0.02}Th_{0.01})_{1.00}$ $((As_{0.81}P_{0.16}Si_{0.02}S_{0.02})_{1.00}O_4)$, 34 respectively. In polished sections, crystals are yellow and translucent with bright submetallic 35 luster. Selected reflectance values R_1/R_2 (λ , nm) for the holotype specimen in air are: 11.19/9.05 (400), 11.45/9.44 (500), 10.85/8.81 (600), 11.23/9.08 (700). Features of the 36 37 crystal structure of gasparite-(La) were studied by means of EBSD (holotype specimen), 38 XRD and SREF (cotype specimen). Gasparite-(La) has a monoclinic structure with the space 39 group $P2_1/n$. Our studies revealed that gasparite-(La) from the Ushkatyn-III deposit and 40 Wanni glacier have different origins. La/Ce and As/P/V ratios in gasparite-(La) could be used 41 as an indicator of formation conditions. 42 Keywords: gasparite-(La), new mineral, arsenate, REE, Mn ores, monazite-type structure,

- 43 Ushkatyn-III, Kazakhstan, Wanni glacier, Binn Valley, Switzerland
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45	Introduction
46	In the course of this study, we described a new rare-earth element (REE) arsenate-
47	gasparite-(La) (La(AsO ₄), IMA 2018-079) from the Ushkatyn-III deposit, Central Kazakhstan
48	and the Wanni glacier, Binn Valley, Switzerland. Gasparite-(La) is named for the dominant
49	lanthanide according to current nomenclature of REE minerals (Bayliss and Levinson 1988).
50	The type locality of gasparite-(La) is the Ushkatyn-III deposit, Central Kazakhstan.
51	Minute isometric grains of gasparite-(La) up to 15 µm in size were discovered in samples
52	from the Ushkatyn-III deposit collected during field work in 2017. The holotype specimen of
53	gasparite-(La) was deposited at the Mineralogical Museum of St. Petersburg State University,
54	St. Petersburg, Russia, catalogue number 19692.
55	The co-type locality of gasparite-(La) is the Wanni glacier, Binn Valley, Valais,
56	Switzerland. Elongated crystals of gasparite-(La) up to 2 mm in size were discovered in
57	autumn 2005 and visually classified as "monazite." Because of their unusual appearance, the
58	material was subjected to further analyses, and recognized as likely identical to the gasparite-
59	(La) later discovered in the Ushkatyn-III deposit. The cotype specimen from Wanni glacier is
60	preserved in Musée Cantonal de Géologie in Lausanne under catalogue number MGL
61	093518.
62	The use of crystals from different localities allowed us to describe the whole range of
63	physical and chemical properties of gasparite-(La), investigate its crystal chemistry and
64	identify some features characteristic of different genetic types of deposits.

65 *REE*-arsenates are among rare minerals: CNMNC IMA has approved only 14 mineral 66 species to date. Most of them representative of Ce-dominant species and only 3 La-dominant 67 minerals have been discovered (Dunn et al. 1984; Mills et al. 2010; Modresky 1983).

Despite the limited number of approved mineral species, REE-arsenates are widely
 distributed in distinct mineral assemblages on numerous localities. Mineralogical information

about arsenates (both discovered and crystal-chemical characteristics) provide a key to
understanding the occurrence and subsequent evolution of many localities (e.g., Campbell
and Nordstrom, 2014; Majzlan et al. 2014; Wu et al. 2018; Yang et al. 2018).

73 REE-arsenates have been reported from several postmagmatic and metasedimentary 74 rocks, whose mineral composition was strongly influenced by late hydrothermal fluids 75 (metasomatic replacement). In the Slovak rhyolites (Ondrejka et al. 2007), primary monazite-76 (Ce) and xenotime-(Y) were transformed into secondary gasparite-(Ce) and chernovite-(Y). 77 respectively. In case of the granite cupola at Zinnwald (Germany) or Cínovec (Czech 78 Republic), As-rich hydrothermal fluids dissolved and severely altered primary magmatic 79 REE-Y-Th-U-Zr mineralization and gave rise to the formation of REE-arsenates: arsenoflorencite-(Ce), chernovite-(Y), and hydrous xenotime(Y)-chernovite-(Y) solid 80 81 solutions (Förster et al. 2011). In the Hora Svaté Kateřiny granite (Czech Republic), reaction 82 with oxidizing As-bearing fluids caused the decomposition of xenotime-(Y), and led to the 83 precipitation of chernovite-(Y) and the incorporation of As into altered zircon and thorite 84 (Breiter et al. 2009). Migrating As-bearing solutions are also believed to have formed the 85 remarkable, classical assemblage of numerous arsenates and arsenites in the Wanni 86 glacier/Mt. Cervandone area at the frontier between the Binn Valley (Switzerland) and Alpe 87 Devero (Italy) (Graeser and Roggiani 1976, Guastoni et al. 2006; Hofmann and Knill 1996). 88 Besides indirect evidence of the presence of multiple REE-arsenates in close association, 89 several minerals show direct evidence of originating from such fluids: gasparite-(Ce) was 90 found as a reaction rim around synchisite-(Ce) (Graeser and Schwander 1987); deveroite-91 (Ce) was found as a dissolution product of cervandonite-(Ce) (Guastoni et al. 2013); agardite-92 (Y) (Gatta et al. 2018), rhabdophane-(La) and uranyl arsenates on cafarsite (Appiani et al. 93 2017).

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Manganese rocks of different genesis are often enriched in arsenic, reaching

concentrations several times higher than the mean values for both sedimentary rocks and for
the upper part of the continental crust as a whole (Li and Schoonmaker 2003; Maynard 2003).
More than half of all discovered Mn-arsenates were found in the famous Mn deposits in
Franklin/Sterling Hill, USA; Långban, Sweden and Moss Mine, Sweden.

Both Långban, Sweden and Franklin/Sterling Hill, New Jersey, USA are represented
by strongly metamorphosed Precambrian rocks of sedimentary origin (Frondel and Baum
1074; Holtstam and Langhof 1999; Lundström 1999). According to Frondel and Baum
(1974), the only primary As-bearing ore minerals in Franklin/Sterling Hill are löllingite,
arsenopyrite and the calcium arsenate svabite. In both cases (Långban and Franklin / Sterling
Hill), the greatest mineralogical diversity is found among the minerals in veins and fissures.

105 A limited number of small Fe-Mn-(Ba,V,As,Sb,Be,W,REE) deposits, containing both 106 arsenates and REE-arsenates, have been found and studied in the Swiss, Italian and Austrian Alps (Abrecht 1990; Brugger and Giere 1999; Brugger and Meisser 2006; Cabella et al. 107 108 1999). These occurrences are thought to represent syngenetic exhalative Fe-Mn 109 accumulations (Majzlan et al. 2014) metamorphosed during the Alpine orogeny (Abrecht 110 1990; Brugger and Meisser 2006). Brugger and Meisser (2006) argued that the chemical 111 composition of the rocks reflects the pre-metamorphic state. Cabella et al. (1999) reported 112 that the abundance of arsenates reduces sharply as a function of proximity to the Fe-Mn ores. 113 The same genetic conclusion was made for Mn-rich metamorphic rocks in the Hoskins 114 manganese mine, New South Wales, Australia (Ashley 1989).

Gasparite-(Ce) is, although generally rare, the most widely distributed REE-arsenate. The type locality for gasparite-(Ce), and several other REE-arsenates and arsenites, is Mt. Cervandone, a summit on the frontier of Italy (Cervandone, Val Devero) and Switzerland (Wanni glacier, Binntal), where it occurs in metasedimentary rocks (Graeser and Schwander 1987). Besides that, gasparite-(Ce) was found as an accessory mineral in the Black Range Tin District, New Mexico, USA (Foord et al. 1991); Tisovec-Rejkovo, Slovakia (Ondrejka et al.
2007); Beryllium Virgin Claim, New Mexico, USA (Anthony et al. 2000); Chudnoe and
Nesterovskoe occurrences, Maldynyrd Range, Prepolar Ural, Russia (Moralev et al. 2005);
Kesebol deposit, Sweden (Kolitsch and Holtstam 2004; Kolitsch et al. 2004); Grubependity
Lake cirque, Maldynyrd Range, Prepolar Ural, Russia (Mills et al. 2010); Artana, Carrara,
Apuane Alps, Italy (Mancini 2000); Tanatz Alp, Switzerland (Roth and Meisser 2013) and
Ponte dei Gonazzi, the Maritime Alps, Italy (Cabella et al. 1999).

In most cases, lanthanum is present in gasparite in subordinate amounts (< 15 wt% La₂O₃). However, La-dominant grains of gasparite were found in Mn-enriched metamorphosed rocks from the Ponte dei Gonazzi, the Maritime Alps, Italy (up to 45 wt% La₂O₃; Cabella et al. 1999) and in A-type rhyolite from Western Carpathians, Slovakia (up to 26 wt% La₂O₃; Ondrejka et al. 2007). Particle size was insufficient to allow for the investigation of the properties of La-dominant gasparite, and therefore these studies did not describe their findings as a new mineral phase.

134

Occurrence

135 Ushkatyn-III deposit, Central Kazakhstan

136 The Ushkatyn-III deposit (48°16'06"N, 70°10'43"E) is located in Central Kazakhstan 300 km southwest of the city of Karaganda and 20 km to the northeast of the village 137 138 Zhayrem. The deposit was discovered in 1962. Manganese ore mining started in 1982 and 139 continues to date. Beginning from 2015, barite-lead ores started to be mined. As of 2015, 140 manganese ore reserves amounted to 102 million tons, with an average Mn content of 24 wt% 141 and Fe 3.5 wt%, and barite-lead ore reserves of 42 million tons, with an average Pb-2.6 wt%, 142 BaSO₄-19 wt% (JSC "Zhayremsky ore mining and processing enterprise," 2015). 143 The geological structure of the deposit was considered in the works of Kayupova

144 (1974), Buzmakov et al. (1975), Mitryaeva (1979), Rozhnov (1982) and Skripchenko (1980,

145 1989). The Ushkatyn-III deposit is located in the western part of the Zhailinsky graben-146 syncline. This large riftogenic structure originated in the Late Devonian during the 147 destruction of the epi-Caledonian Central-Kazakhstan continental block. The clay-siliceous-148 carbonate rocks of the Famennian stage of the Upper Devonian are ore-bearing. In the eastern 149 part of the deposit, they are represented by reefogenic limestones containing a stratiform 150 barite-lead mineralization. In the western part of the deposit, these rocks are replaced by 151 detrital and nodular-layered siliceous limestones containing layers of manganese ores. 152 Altogether, there are fourteen ore layers, each of which has a well-marked, rhythmically-153 stratified structure with alternating layers of manganese ore and limestones. The thickness of individual rhythms range from 15 cm up to 1 m, and the total thickness of ore layers varies 154 155 from 5 to 25 m. A series of adjacent layers is grouped into a large pack, traced over a strike of 156 more than 1.5 km, a drop of 760 m and a thickness of 50-150 m. Volcanic rocks are present 157 on the deposit but in volumetric inferior amount (no more than 10% of the hole volume of the 158 ore-bearing strata).

159 Manganese ores are fine-grained rocks (average size of mineral grains 10-30 µm) 160 with layered and lenticular-banded textures. Altogether, more than 60 minerals have been identified in manganese ores of the Ushkatyn-III deposit by optical, electron microscopy, X-161 162 ray powder and microprobe analysis (Brusnitsyn et al. 2017, 2018; Kayupova 1974). The main minerals are braunite, hausmannite, quartz, calcite, rhodochrosite, tephroite, friedelite 163 164 and minerals of manganhumite series (sonolite, alleghanyite). The most characteristic 165 secondary minerals are hematite, jacobsite, rhodonite, caryopilite, pennantite, manganese 166 clinochlore, albite and barite. Among the most interesting accessory minerals are cinnabar, 167 pyrobelonite, cerianite-(Ce), fluorite and several REE and arsenate minerals: sarkinite, 168 svabite, tilasite, retzian group minerals.

169

Manganese ores could be divided into two types: 1) braunite: braunite + calcite +

170 quartz \pm albite, and 2) hausmannite: hausmannite + calcite + rhodochrosite \pm tephroite 171 $(sonolite, alleghanyite) \pm friedelite (caryopilite).$ These types of ores can form separate layers, 172 and can be combined within a single layer. In the latter case, the mineral composition of ores 173 change substitution of braunite for hausmannite/associated result of as а 174 silicates/rhodochrosite.

The first arsenates in the Ushkatyn-III deposit were discovered in the early 70s (Kayupova 1974). However, due to the lack of technical equipment, only relatively big grains of minerals (sarkinite, tilasite and brandtite) could be identified.

178 Gasparite-(La) was found in hausmannite ores, associated with other arsenates. The 179 mineral was found in microveins cutting layers of hausmannite, calcite, rhodochrosite. Gasparite-(La) is associated with fridelite, jacobsite, pennantite, manganhumite series 180 181 minerals (alleghanvite, sonolite), sarkinite, tilasite and retzian-(La). The microveins 182 containing gasparite-(La) range in size from microns to 1-5 mm in thickness and to 1-3 cm 183 in length. Gasparite-(La) forms grains of 2-25 µm in size, as well as aggregates with other 184 arsenates of irregular shape up to 50 µm and was found in association with retzian-(La) and 185 alleghanyite (Figure 1a).

186 Wanni glacier, Binn Valley, Valais, Switzerland

187 The Wanni glacier, located in the Binn Valley, Valais (Wallis), Switzerland, represents the Swiss side of the Scherbadung or Pizzo Cervandone, of which the Italian side 188 189 is located in Alpe Devero, Piemonte, Italy. Its mineral assemblage extends to both sides of 190 this mountain and is the type locality of seven REE-arsenates and REE-arsenites (Armbruster 191 et al. 1988; Demartin et al. 1994; Graeser 1966; Graeser and Schwander 1987; Graeser et al. 192 1994; Guastoni et al. 2006, 2013) and its geology has been summarized in Streckeisen et al. 193 (1974), Steck (1987), Klemm et al. (2004), Hettman et al. (2014) and Bergomi et al. (2017). 194 The REE-As mineralization is hosted in two-mica gneisses of the Monte Leone nappe and

extends multiple km westward to the Gischi glacier (Graeser and Roggiani 1967), Chummibort (Cuchet et al. 2005) and Mättital (Krzemnicki 1992, 1997) and eastwards to the Lercheltini area. According to Krezemnicki and Reusser (1998), several Pre-alpine ore concentrations within this nappe were locally re-mobilized during Alpine metamorphism, thus generating some unique hydrothermal mineralization.

The sample with gasparite-(La) was found in rocks of the Monte Leone nappe (46°19′20,N, 8°12′48E; Hettman et al. 2014; Klemm et al. 2004). The Monte Leone nappe includes fine-grained banded orthogneisses and minor coarse-grained augen gneiss interlayered with paragneisses, hornblende gneisses and amphibolites and shows a penetrative amphibolite-facies metamorphic overprint of Alpine age (Bergomi et al. 2017; Maxelon and Mancktelow 2005).

The specimen with gasparite-(La) was extracted from a small, classical Alpine fissure. The stratum containing the fissure with gasparite-(La) is a fine-layered, two-mica gneiss and is located outside the main Cu-As-F-mineralization which has cafarsite as the dominating Ascontaining mineral (Cuchet et al. 2014, 2016). Based on our observation, the As-enrichment is not very dense in this sublayer. Within the occurrence of *REE*-arsenates (chernovite-(Y), gasparite-(Ce)) is increased, whereas arsenites (cafarsite and asbecasite) are a diminished gasparite form of microcrystalline pseudomorphoses to synchysite-(Ce).

The specimen with gasparite-(La) (containing three elongated crystals) was located on one side of the cavity; the remainder was empty, apart from minerals that belong to the classical fissure parageneses: minor titanite, quartz, feldspar and albite. Crystals of gasparite-(La) are prismatic of a size up to 2 mm (Figure 1b). An interesting feature of the studied crystals is that their prism appears non-translucent (as if fractured) and yellow, whereas the summit faces are perfectly, gemmy translucent and more orange.

219

Elemental composition

Elemental compositions of rock-forming minerals were studied on the carbon-coated polished sections by means of a Hitachi S-3400N scanning electron microscope equipped with an Oxford X-Max 20 energy dispersive x-ray spectrometer (EDX). EDX spectra were obtained under the following conditions: 20 kV accelerating voltage and 2 nA beam current with an acquisition time of 30 s per spectrum.

Elemental analyses for gasparite-(La) were obtained using an Inca Wave 500 wavelength dispersive x-ray (WDX) spectrometer also equipped on the microscope mentioned above. WDX spectra collection conditions were: 20 kV, 10 nA, beam diameter 5 μm, 30 sec peak and 30 sec background collection per element, XPP matrix correction. Fe metal, Mn metal, V metal, InP, InAs, wollastonite, Th-, Y-, La-, Ce -, Nd- and Sm-bearing glass standard samples (MAC-standards) were used for spectrometer calibration.

Preliminary EDX analyses showed that gasparite-(La) from the Ushkatyn-III deposit had almost no chemical zoning or stable element ratios; all analyzed grains showed La>Ce>Nd and As>V>P. In the case of gasparite from the Ushkatyn-III deposit, Central Kazakhstan five analyses (WDX) from three different grains were performed in one carboncoated polished section (Table 1). The empirical formula of gasparite-(La) from the Ushkatyn-III deposit based on 4 oxygen is $(La_{0.65}Ce_{0.17}Nd_{0.07}Ca_{0.06}Mn_{0.05}Pr_{0.02})_{1.03}$ ((As_{0.70}V_{0.28}P_{0.02})_{1.01}O₄).

One elongated crystal of gasparite from Wanni glacier was studied by EDX and WDX analyses. EDX analysis showed that the crystal had chemical zoning: its La/Ce/Nd ratio varied significantly, whereas As/P ratio was stable (Figure 2). The optically more translucent summit of the crystal was La-dominant, while the prismatic part of the crystal was Cedominant; and thus, represents gasparite-(Ce). We performed five WDX analyses from the summit of the crystal and the overall empirical formula of gasparite-(La) from the Wanni glacier based on 4 oxygen is $(La_{0.59}Ce_{0.37}Nd_{0.02}Ca_{0.02}Th_{0.01})_{1.00}((As_{0.81}P_{0.16}Si_{0.02}S_{0.02})_{1.00}O_4)$. The simplified formula of gasparite-(La) from both the Ushkatyn-III deposit as the Wanni glacier was La(AsO₄).

Gasparite-(La) belongs to the monazite group, which contain seven monoclinic phosphate and arsenate minerals (Table 2). It is a La-dominant analog of gasparite-(Ce) (Graeser and Schwander 1987) and arsenate-dominant analog of monazite-(La). According to the Nickel-Strunz Classification, gasparite-(La) belongs to 8.AD (8: phosphates, arsenates, vanadates, A: phosphates, etc. without additional anions, without H₂O, D: with only large cations).

253

Physical properties and optical data

Gasparite-(La) crystals are yellow and translucent with bright submetallic luster. The Vickers Hardness Number (VHN) measured on gasparite-(La) from the Wanni glacier was 325 with a range 308–340 kg mm⁻² (load 20 g) by means of HMV-2T (Shimadzu). This data is in a good agreement with data on gasparite-(Ce) (VHN=327 kg mm⁻²; Graeser and Schwander 1987). Mohs hardness could not be determined because of the tiny size of the crystals. The Mohs hardness calculated from the VHN value was approximately $4\frac{1}{2}$.

260 As gasparite-(La) from the Ushkatyn-III deposit had no chemical zoning, one of its 261 grains was chosen for optical study. In polished sections, gasparite-(La) from the Ushkatyn-262 III deposit looked dark gray in reflected light. The mineral was slightly anisotropic with Δ $R_{589} = 2.04\%$. The reflectivity of gasparite-(La) in air (Table 3) was measured against a SiC 263 264 standard (Reflexions standard - 474251, No. 545) using MSF-21 spectrophotometer with a 265 monochromator slit of 0.4 mm and a 100 µm zone diameter. The measurement parameters 266 were as follows: lens magnification 21x, aperture 0.4 and $\Delta\lambda$ =10 nm, SiC. The reflectivity 267 spectrum is shown in Figure 3.

268

Raman spectra

269 Gasparite-(La) crystals from both localities were used for Raman studies. Raman 270 spectra (Figure 4) were recorded with a Horiba Jobin-Yvon LabRAM HR800 spectrometer

equipped with an Olympus microscope. The microscope comprised 50x and 100x objectives. Raman spectra were excited by an Ar ion laser at a wavelength of 514 nm and a maximum power of 50 mW. The spectra were obtained in the range of 100-4000 cm⁻¹ at a resolution of 2 cm^{-1} at room temperature. To improve the signal-to-noise ratio, the number of acquisitions was set to 20. The spectra were processed using licensed Labspec and Origin software . Band fitting was done using a Lorentz function with the minimum number of component bands used for the fitting process (Table 4).

278 Raman spectra of gasparite-(La) from both the Ushkatyn-III deposit and the Wanni 279 glacier were very close to synthetic La(AsO₄) (Fig. 4, Table 4). Bands in the region from 4000 to 1100 cm⁻¹ were not registered, which means that gasparite-(La) contained no (OH)⁻ 280 groups. Bands assigned to stretching vibrations v_1 and v_3 of arsenate ion are observed in the 281 region 900-800 cm⁻¹. There were bending vibrations v_2 and v_4 of arsenate ion in the region of 282 500–350 cm⁻¹. The lattice vibrations were located below 320 cm⁻¹. The main differences of 283 Raman spectra of the minerals in comparison with the pure synthetic phase were associated 284 with impurities of $(VO_4)^{3-}$ (Ushkatyn-III) and $(PO_4)^{3-}$ (Wanni glacier). Vibrations $v_1(PO_4)$ 285 was about 960 cm⁻¹, $v_1(VO_4)$ – near 840 cm⁻¹ (Song et al. 2018; Solecka et al. 2018). In the 286 region of the v_4 band the main differences connected with the overlapping of bands related to 287 $(AsO_4)^{3-}$, $(PO_4)^{3-}$ and $(VO_4)^{3-}$. Raman spectra pointed out predominance of La+Ce in the 288 mineral composition of $v_1(AsO_4)$ of 861–863 cm⁻¹. The Raman spectrum of Ho(AsO₄) 289 v_1 (AsO₄) was about 895 cm⁻¹ (Barros et al., 2009). Impurity tetrahedral cations (P,V) did not 290 291 influence the $v_1(AsO_4)$ band shift.

292

Crystallography

293 Because gasparite-(La) from the Ushkatyn-III deposit occurs as microscopic grains up 294 to 15 μm in size (Fig. 2) it was not possible to determine its crystal structure with a single 295 crystal X-Ray diffraction. All diffraction data were obtained by electron backscatter

diffraction (EBSD). In the case of gasparite-(La) from the Wanni glacier, we managed to
isolate the La-enriched zone of the elongated crystal and refine its crystal structure using
single crystal X-Ray diffraction. The same crystal fragment was used for powder X-ray
diffraction studies.

300 **Powder X-ray diffraction (XRD).**

301 The powder XRD pattern for gasparite-(La) from the Wanni glacier was recorded in Debye-302 Scherrer geometry by means of a Rigaku R-AXIS Rapid II diffractometer equipped with a 303 curved (cylindrical) imaging plate detector (r=127.4 mm). CoK α radiation (λ =1.79021 Å) 304 was generated by a rotating anode (40 kV, 15 μ A) with microfocus tube optics; exposure time 305 was set to 15 min. The data were processed using the osc2xrd program (Britvin et al. 2017) 306 and Stoe WinXPOW software (Stoe and Cie 2006). XRD data for gasparite-(La) from the 307 Wanni glacier is presented in Table 5 and was similar to synthetic $La(AsO_4)$ (Le Berre et al. 308 2007; JCPDS file 15-0756). Calculated data were obtained using Rietveld refinement of the 309 powder pattern [c 6.7087(3), b 7.1499(2), a 6.9429(2) Å, β 104.442(2) °].

310 Single-crystal X-ray diffraction and refinement (SREF)

REE arsenates, chromates, phosphates and vanadates of $Ln(XO_4)$ type can crystallize in monoclinic (monazite structure) or tetragonal (zircon-type structure) symmetry (Botto and Baran 1982; Clavier et al. 2011; Schwarz 1963). In the REE arsenate and phosphate series, the La, Ce, Pr and Nd end-members exhibit the monazite structure (Clavier et al. 2011). REE vanadates have more complex behavior: the majority of REE vanadates (Ce–Lu, Sc) exhibit the zircon-type structure, whereas La(VO₄) crystallize in both the zircon-type and monazitetype structures (Witzke et al. 2008).

The monazite-type structure was first reported by Mooney (1948) and then refined from this date by several authors for various *REE*-phosphates (Beall et al. 1981; Feigelson 1964; Ghouse 1968; Mullica et al. 1984; Mullica et al. 1985; Ni et al. 1995; Pepin and Vance

1981). The structural arrangement in monazite-type structures is based on the nine-fold coordination of the metallic cation and can be described as an equatorial pentagon interpenetrated by a tetrahedron (Clavier et al. 2011). The tetrahedron located out of the equatorial plane can then be described as a link between the REE O₉ polyhedra, leading to the formation of infinite chains along the c axis ([0 0 1] direction).

According to published data, both synthetic La(AsO₄) (Schmidt et al. 2005) and natural gasparite-(Ce) (Kolitsch et al. 2004) had a monoclinic structure with the space group $P2_1/n$.

329 The crystal structure of gasparite-(La) from the Wanni glacier (cotype specimen) was 330 solved by direct methods and refined to $R_1=0.014$ using a SHELX-2015 set of programs 331 (Scheldrick 2015) via Olex2 v.1.2.8 graphical user interface (Dolomanov et al. 2009) (Tables 332 6-8). Data collection and structure refinement details are given in Tables 6-7 and in the attached CIF file. The La site is coordinated by nine O atoms, with (La-O) distances of 2.586 333 334 (Table 8). The As site was found to be almost fully occupied by As (As_{0.84} $P_{0.16}$), in 335 agreement with the chemical data. The unit cell parameters (ucp) of gasparite-(La) are larger 336 than values reported for gasparite-(Ce) (Kolitsch et al. 2004), smaller than values reported for 337 synthetic La(AsO₄) (Schmidt et al. 2005) and in good agreement with powder XRD for the 338 same crystal.

339 Electron backscatter diffraction (EBSD)

EBSD measurements were performed on an Oxford HKLNordlys Nano EBSD detector equipped on a Hitachi S-3400N scanning electron microscope. Operating conditions are listed in Table 10. Both acquisition and analysis of Kikuchi-patterns were made by Oxford AZtecHKL software. Synthetic La(AsO₄) structural data (ICSD) were used as input (Schwartz et al. 2009). The sample was polished with progressively smaller polycrystalline diamond suspensions with the ending step of Ar ion etching for 10 min at the final stage

(Oxford IonFab 300) in order remove amorphized layers for EBSD analysis. The sample was

347	pre-tilted 70° along the normal to the EBSD detector. The fit factor (mean angular deviation
348	(MAD)), which describes the angular deviation between the calculated and measured Kikuchi
349	lines (good fit for deviations $<1.0^{\circ}$) was less than 0.3° for synthetic La(AsO ₄) and Ce(AsO ₄)
350	structures (Brahim et al. 2002; Kang and Schleid 2005; Schwartz et al. 2005).
351	Eighteen electron backscatter patterns from three different gasparite-(La) grains were
352	collected (Fig. 6). Good matches were obtained for all patterns using monoclinic La(AsO ₄)
353	structure with the space group $P2_1/n$ (Schmidt et al. 2005).

354

346

Discussion

355 Crystal chemistry and substitution mechanisms

The REE ratio in gasparite-(La) from the Ushkatyn-III deposit was stable and in all analyzed grains, La>Ce. The REE ratio in crystals of gasparite-(La) from the Wanni glacier varied significantly from Ce-dominant to La-dominant species in the same crystal (Fig. 2).

Our data revealed a very limited P-for-As substitution (P up to 0.02 apfu) and intense V-for-As substitution (V up to 0.28 apfu) in gasparite-(La) from the Ushkatyn-III deposit (Fig. 7). On the other hand, gasparite-(La) from the Wanni glacier contained no V and had intense P-for-As substitution (P up to 0.16 apfu).

Impurity of tetrahedral cations (P,V) did not appear to influence the (AsO₄) band shift as isolated tetrahedra ((AsO₄), (PO₄) and (VO₄)) in the gasparite structure surrounded by polyhedral ((LaO₉), (CeO₉), (NdO₉)). As a result, cations in the polyhedra appeared to have a major impact the position of the bands related to As-O vibrations in $(AsO_4)^{3-}$. Both P-for-As and V-for-As were reflected in the appearance of additional bands or broadening the main bands. The band around 390 cm⁻¹ in the case of P-for-As (Wanni glacier), shifted upwards, and in the case of V-for-As (Ushkatyn-III deposit), downwards (Fig. 4).

370

No natural vanadates of monazite structure-type have been reported yet. Only

371 wakefieldite group minerals with the zircon structure were reported (Deliens and Piret 1986; 372 Witzke et al. 2008). No experimental evidence for the existence of AsO_4 – PO_4 substitution 373 was reported. Cabella et al. (1999) described gasparite-(Ce) enriched with P and V and 374 reported a wide range of V-for-As substitutions (V up to 0.30 apfu) and P-for-As 375 substitutions (P up to 0.15 apfu). Kolitsch et al. (2004) reported that 10% of the As atoms 376 were substituted by P (P up to 0.12 apfu) in gasparite-(Ce).

The crystal chemistry of the MXO_4 monazite-type compounds in general (M=La, Ce; X=As, P, V) have been studied intensively (e.g., Clavier et al. 2011; Kolitsch et al. 2004). Kolitsch et al. (2004) assumed that the substitution of P for As les to the substantial decrease of the *c*-parameter, which can be explained by the stacked arrangement along the [0 0 1] direction the XO_4 groups in the structure.

Our data confirmed, that both *a* and *c* ucp are influenced mainly by the size of XO_4 tetrahedra (r=0.99, r=0.92, respectively; Fig. 8a), whereas the *b* parameter is mainly influenced by the size of the MO_9 polyhedra (r=0.92; Fig. 8b).

Comparison of $LaXO_4$ monazite-type compounds showed that the (La-O) distance in MXO₄ monazite-type compounds varied significantly: 2.579 (La(PO₄); Ni et al. 1995) to 2.599 Å (La(AsO₄); Schmidt et al. 2005). This could explain the distortion of LaO₉ polyhedra, which is well correlated to (X-O) distance (r=0.99; Table 9) and could be the reason for *b vs.* (*M*-O) correlation.

390 Origin of gasparite-(La)

As presented above, gasparite-(La) was found in occurrence with distinct geological situations. Gasparite-(La) from the Wanni glacier, Binn Valley was found in an alpine fissure with evidence of hydrothermal alteration (well-shaped crystals in an almost empty cavity), whereas gasparite-(La) from the Ushkatyn-III deposit was found in primary Mn ores (irregular-shaped grains in calcite-rhodochrosite-fridelite veins).

396 We believe that in the case of gasparite-(La) from the Wanni glacier, the source of 397 arsenic was an As-rich hydrothermal fluid, as in the case of gasparite-(Ce) (Graeser and 398 Schwander 1987) from the same area. More challenging to explain was the source of 399 lanthanum. The host mineral could be the main source of REE, as was shown in the case of 400 synchisite-(Ce) (Graeser and Schwander 1987), monazite-(Ce) (Ondrejka et al. 2007) and 401 cervandonite-(Ce) (Guastoni et al. 2013). High P content in gasparite-(La) from the Wanni 402 glacier could indicate that monazite could be the host mineral. However, additional process 403 of REE redistribution needed as (1) crystal of gasparite-(La) had La/Ce zoning and (2) (Ce)-

404 and (Y)-dominant species were the rule in this region, supported by Ce/Y >> La in whole 405 rock samples (Hofmann and Knill 1996). Another possible explanation to La-enrichment 406 were the partial oxidation of Ce^{3+} to insoluble Ce^{4+} during fluid transportation and subsequent 407 LREE (La, Nd, Sm) enrichment. This mechanism is well recorded in uranyl minerals (e.g., 408 Meisser et al. 2010).

Thus, gasparite-(La) from the Wanni glacier was formed as a metasomatic mineral in the process of alteration of primary REE mineralization by As-rich hydrothermal fluid and the ensuing La/Ce separation.

The Ushkatyn-III deposit is considered as an object of hydrothermal-sedimentary
genesis, transformed by processes of low-grade regional metamorphism (Brusnitsyn et al.
2018; Mitryaeva 1979; Rozhnov 1982; Skripchenko 1980).

Gasparite-(La) from the Ushkatyn-III deposit was found in microveins in hausmannite ores. It was found in association with fridelite, jacobsite, pennantite, manganhumite series minerals, sarkinite, tilasite and retzian-(La) and embedded into calcite-rhodochrosite veinlets. Similar veinlets in other rocks have a different composition. Calcite and quartz were found in microveins in the enclosing limestone. Calcite, kutnohorite, rhodonite, axinite-(Mn), friedelite, hematite, jacobsite and barite were found in microveins in braunite ores. A regular

421 change in the composition indicates the segregation mechanism of the genesis of such 422 microveins. They were formed by local redeposition of a substance from the rocks in which 423 they developed into thin cracks. Probably these processes occurred during the period of 424 tectonic deformation of the region during Late Paleozoic time.

We believe that As and REE were accumulated syngenetically with manganese sediments in the Ushkatyn-III deposit. Most likely, the initial manganese oxides absorbed these elements as it happens in modern oceans, there the absorption of REE to manganese oxides occurs much more intensively than many other precipitation minerals (Dubinin 2006).

Occurrence of As-minerals in the Ushkatyn-III deposit is connected to the manganese ores only and has not been found in the host rocks or tectonic deformation zones. Manganese rocks of different genesis are often enriched in arsenic, and more than half of all discovered Mn-arsenates were found in the famous Mn deposits (Abrecht 1990; Brugger and Giere 1999; Brugger and Meisser 2006; Cabella et al. 1999; Frondel and Baum 1974; Holtstam and Langhof 1999; Lundström 1999). Therefore, As and REE infiltration into ore layers from any external source is unlikely in this case.

Apparently, REE and As accumulated in the initial ore-bearing sediments as a component of Mn minerals. In the process of metamorphism and tectonic deformation, these elements were mobilized by pore solutions and re-deposited into secant ore veins. A similar mechanism for the genesis of REE-bearing arsenates is also assumed for metamorphosed Fe-Mn deposits in other regions (Cabella et al. 1999; Kolitschi et al. 2004).

Gasparite-(La) from the Ushkatyn-III deposit is characterized by relatively small amount of other REE (Ce+Nd+Pr <0.3 apfu) compared to other gasparite-(Ce) (Fig. 7) and have stable La/Ce ratio. Its occurrence with retzian-(La) (Fig. 2) may indicate specific conditions of mineral formation in which rocks were depleted by cerium.

445 Lanthanum and cerium have very close chemical properties, but the average content

of lanthanum in the Earth's crust is almost two times lower than cerium: 30 and 58 ppm, respectively (Li and Schoonmaker 2003). Gasparite-(La) and retzian-(La) formation require separating lanthanum from cerium. This could be done in two ways: (1) during the accumulation of manganese sediment or (2) later during lithification.

450 In favor of the first option is the fact that accumulations of manganese oxides of 451 hydrothermal genesis are characterized by cerium deficiency relative to the remaining REE. 452 In the REE spectra of such rocks, a negative cerium anomaly is usually well expressed 453 (Dubinin 2006; Bau et al. 2014). In other words, La/Ce is higher in them than in "normal" 454 marine sediments, which determines the possibility of the formation of lanthanum minerals. 455 If this assumption is correct, then lanthanum minerals, including gasparite-(La) and retzian-456 (La), should be considered as indicators of the hydrothermal-sedimentary genesis of 457 manganese ores. However, this issue requires further study.

458 According to the second option, the separation of lanthanum and cerium occurred at 459 the post-sedimentation stage of the development of the deposit. The very low Ce content of 460 both gasparite-(La) and retzian-(La) (less than 12 and 9 wt%, respectively) could be explained by formation from a strongly Ce-depleted source due to oxidation of Ce³⁺ and 461 462 subsequent formation of insoluble cerianite-(Ce) as shown in the case of wakefieldite-(La) 463 (Witzke et al. 2008). Cerianite-(Ce) was observed in several cases in the same samples from 464 the Ushkatyn-III deposit where La arsenates were found. According to experimental data 465 (Ohta, Kawabe, 2001), the oxidation of cerium with manganese oxides proceeds according to 466 the reaction:

$$2Ce^{3+} + MnO_2 + 2H_2O = 2CeO_2 + Mn^{2+} + 4H^+$$

468 Reaction will shift to the right by a weak alkaline mineral formation medium, typical 469 for carbonate associations. As a result, cerianite-(Ce) can coexist with Mn^{2+} minerals 470 (rhodochrosite, tephroite, friedelite, etc.). However, it is possible only in the absence of 471 organic matter. Otherwise, there will be a dissolution of cerianite-(Ce) with the restoration of

472 cerium by the reaction (Dubinin 2006):

473
$$CH_2O + 5H_2O + 11CO_2 + 4CeO_2 = 4Ce^{3+} + 12HCO_3^{-}$$

474 Accordingly, the formation of cerianite-(Ce) and, as a consequence, formation of La-rich 475 (Ce-depleted) minerals is controlled by local distribution of Mn, REE, and organic matter.

Both scenarios considered (cerium deficiency in initial sediments and cerium concentration in cerianite-(Ce) at post-sedimentary stages) do not contradict each other. Most likely, each of them contributed to the formation of gasparite-(La) and other La-rich minerals

479 in manganese ores of the Ushkatyn-III deposit.

Thus, gasparite-(La) from the Ushkatyn-III deposit was formed in primary ores and was not influenced by metasomatic processes. We believe that it is precisely the features of the chemical composition of the initial Mn-Fe ores that predetermined the possibility of the formation of lanthanum minerals in the Ushkatyn-III deposit.

484

Implications

Gasparite-(La) from Mn ores of the Ushkatyn-III deposit and metamorphic rocks of the Wanni glacier have different geological settings and different formation conditions. Thus, an occurrence of gasparite-(La) in rocks and its chemical composition could be used as a tool for geological reconstruction of their host rock formation.

In the case of gasparite-(La) from the Ushkatyn-III deposit, both REE and As were sourced from host Mn ores. In all analyzed grains, we have stable As/V/P and La/Ce ratios and As>V>P and La>Ce. Constant lanthanum predominance in analyzed gasparite grains indicate specific conditions of Mn ore formation: Ce depletion or La enrichment and no metasomatic process. Besides that, gasparite from Mn ores is characterized by low P and high V content (our data; Cabella et al. 1999). Modern metalliferous sediment is mainly composed of Fe- and Mn-oxy / hydroxides and smectite minerals (e.g., Vereshchagin et al. 2019), which

are carriers of V and *REE* (Emerson and Huested 1991; Gurvich, 2006). Ferromanganese ores
are sources of several *REE*-dominant vanadates (Moriyama et al. 2011; Witzke et al. 2008).
Thus, V content originated from primary Fe-, Mn- sediments and its value could be used as
an indicator of gasparite origin.
Gasparite-(La) from the Wanni glacier has different chemical features. It also has
stable As/V/P ratio, but in this case, As>P>V. High P content is a typical feature of gasparite
from metasomatic rocks (Graeser and Schwander 1987; Ondrejka et al. 2007). It also has

503 variable La/Ce ratio, which is typical for gasparite from metasomatic rocks, probably due to

- 504 La/Ce separation during recrystallization.
- 505

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777 Figure captions

- Figure 1. Morphology of gasparite-(La): (a) irregular-shaped grains and aggregates in calcite
- veins from the Ushkatyn-III deposit (BSE image), (b) elongated crystal (yellow) with albite
- 780 crystals (white) from the Wanni glacier (optical microscopy). Note: gasparite-(La)-Gas,
- 781 jacobsite-Jac, retzian-(La)-Ret, alleghanyite-All, calcite-Cal.
- 782 Figure 2. Chemical zoning of gasparite crystal from the Wanni glacier
- Figure 3. Reflectance spectra in air for gasparite-(La) from the Ushkatyn-III deposit.
- Figure 4. Raman spectra in air for gasparite-(La) from the Ushkatyn-III deposit, gasparite-
- 785 (La) from the Wanni glacier and synthetic La(AsO₄) (Botto and Baran, 1982)
- Figure 5. (a) EBSB pattern of the gasparite-(La) crystal taken (30 kV accelerating voltage,
- 0.3 nA beam current, 5 seconds exposure per frame, averaging of 20 frames, 1344x1024
- pixels image size), and (b) the pattern indexed with the $P2_1/n$ structure (MAD 0.15, 80 bands
- are represented on an image)
- 790 Figure 6. Chemical composition variations in gasparite group minerals: (a) La-Ce-Nd ratio,
- 791 (b) As-V-P ratio
- Figure 7. As vs. V+P+S substitution
- Figure 8. Unit cell parameters of $M(XO_4)$ compounds (M=La, Ce; X=As, P, V): (a) c vs. (X-
- 794 O), (b) *b* vs. (*M*-O)
- 795

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Table 1. Chemical composition of gasparite-(La)

Constituent	Ushkaty	n-III deposit	Wanni glacier		
Constituent	Wt%	Range	Wt%	Range	
Fe ₂ O ₃	0.05	0.00-0.16	0.00	0.00	
MnO	1.30	0.91-1.96	0.00	0.00	
CaO	1.33	0.97-1.54	0.34	0.24-0.48	
ThO ₂	0.00	0.00	0.58	0.37-0.72	
Y ₂ O ₃	0.01	0.00-0.05	0.01	0.00-0.03	
La ₂ O ₃	40.21	37.83-41.17	35.59	34.57-36.74	
Ce_2O_3	10.69	9.76-11.61	22.55	21.26-23.67	
Pr ₂ O ₃	1.46	0.00-1.99	0.29	0.00-0.88	
Nd_2O_3	4.24	3.53-4.69	1.04	0.50-1.30	
Sm_2O_3	0.09	0.00-0.33	0.07	0.00-0.35	
V_2O_5	9.77	8.58-11.21	0.00	0.00	
P_2O_5	0.64	0.22-0.99	4.29	4.18-4.45	
As ₂ O ₅ 30.32		29.46-31.41	34.48	33.27-35.22	
Total	100.11		100.13		

Mineral	Gasparite- (La)	Gasparite- (Ce)	Rooseveltite	Monazite- (La)	Monazite- (Ce)	Monazite- (Nd)	Monazite- (Sm)	Cheralite	
Chemical formula	La(AsO ₄)	Ce(AsO ₄)	Bi(AsO ₄)	La(PO ₄)	Ce(PO ₄)	Nd(PO ₄)	Sm(PO ₄)	CaTh(PO ₄) ₂	
Crystal system		•	·	Mo	noclinic				
Space group		$P2_{1}/n$							
<i>a</i> (Å)	6.9576(4)	6.929(3)	6.879(1)	6.8313(10)	6.7880(10)	6.7352(10)	6.6818(12)	6.7085(8)	
<i>b</i> (Å)	7.1668(4)	7.129(3)	7.159(1)	7.0705(9)	7.0163(9)	6.9500(9)	6.8877(9)	6.4152(6)	
<i>c</i> (Å)	6.7155(4)	6.697(3)	6.732(1)	6.5034(9)	6.4650(7)	6.4049(8)	6.3653(9)	6.4152(6)	
β (°)	104.414(1)	104.46(3)	104.84(1)	103.27(1)	103.43(1)	103.68(1)	103.86(1)	103.71(1)	
Z	4								
Reference	Our data	Kolitsch et al. 2004	Bedlivy and Mereiter 1982		Ni et a	al. 1995		Raison et al. 2008	

Table 2. Comparative crystallographic data for monazite group minerals

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Table 3. Reflectance of gasparite-(La) from the Ushkatyn-III deposit measured in air

λ(nm)	R ₁	R ₂	λ (nm)	R ₁	R ₂
400	11.19	9.05	560	10.92	8.90
420	12.04	9.82	580	10.88	8.84
440	12.35	10.15	589	10.87	8.83
460	12.35	10.13	600	10.85	8.81
470	12.16	9.95	620	10.85	8.77
480	11.97	9.77	640	10.89	8.74
500	11.45	9.44	650	10.92	8.77
520	11.15	9.17	660	10.94	8.79
540	11.00	9.00	680	11.09	8.86
546	10.98	8.98	700	11.23	9.08

803 804

Table 4. Raman spectral signatures of gasparite-(La)

Raman shift, cm ⁻¹ / Relative Intensity				Ram Rela			
Ushkatyn- III	Wanni glacier	Botto and Baran 1982	Assignment	Ushkatyn- III	Wanni glacier	Botto and Baran 1982	Assignment
92 / vw 104 / vw	94 / vw 107 / w		Lattice vibrations	364 / vw	367 / w	350 / vw 367 / w	v ₄ (AsO ₄)
126 / vw	125 / w 139 / vw			379 / st	395/ st	392 / m	v ₂ (AsO ₄)
153 / vw 1 190 / vw 1 203 / m 1	155 / w 190 / w 203 / vw			422 / w 452 / w	421 / m 462 / w	418 / w 440 / sh 461 / w	v4(AsO4)
264 / vw 2 320 / w 334 / vw 2	267 vw 337 / w	322 / w 336 / w		822 / m 843 / sh	812 / sh 826 / sh 848 / sh	798 / vw 827 / m 845 / st 872 / sh	v ₃ (AsO ₄)
				860 / vs	864 / vs	861 / vs	$v_1(AsO_4)$

Note: Vs - very strong; st - strong; m - medium; w - weak; vw - very weak; sh -shoulder.

808

Table 5. Powder X-ray diffraction data for gasparite-(La) from the Wanni glacier

	-						_	-	_	-			
I _{obs}	d _{obs}	I _{calc}	d _{calc}	h	k	l	I _{obs}	d _{obs}	Icalc	d _{calc}	h	k	l
9	5.39	5	5.39	1	0	-1	19	1.898	13	1.900	0	2	3
13	4.89	10	4.90	0	1	1	6	1.851	3	1.853	3	2	0
12	4.80	9	4.81	1	1	0	19	1.814	19	1.816	2	-2	-3
12	4.30	10	4.31	1	-1	-1			2	1.794	1	3	2
8	4.18	3	4.18	1	0	1			5	1.789	0	4	0
		5	3.61	1	1	1	25	1.780	22	1.780	2	3	1
20	3.574	16	3.577	0	2	0	11	1.727	10	1.728	0	4	1
64	3.361	58	3.363	0	0	2			3	1.681	0	0	4
6	3.249	5	3.249	2	0	0	9	1.673	8	1.673	2	0	-4
100	3.156	100	3.158	0	2	1			7	1.637	0	1	4
22	3.042	18	3.043	0	1	2	17	1.633	5	1.633	0	3	3
		26	2.968	2	-1	-1			2	1.633	4	-1	-1
77	2.956	69	2.958	2	1	0			2	1.631	4	0	-2
24	2.693	22	2.695	2	0	-2			5	1.624	4	0	0
19	2.519	18	2.524	2	-1	-2	8	1.619	3	1.618	2	1	3
19	2.511	12	2.506	2	1	1	8	1.587	6	1.590	4	-1	-2
8	2.446	5	2.450	0	2	2			2	1.584	4	1	0
6	2.407	6	2.410	2	-2	-1	12	1.579	5	1.579	2	-3	-3
5	2.301	3	2.304	1	0	-3			3	1.579	0	4	2
16	2.233	8	2.239	1	3	0	7	1.517	5	1.519	4	-2	-1
		8	2.230	3	0	-1			2	1.494	3	2	2
12	2.189	8	2.193	1	-1	-3	6	1.491	2	1.491	2	-4	-2
11	2.163	6	2.166	1	2	2	4	1.473	2	1.474	4	-1	-3
8	2.138	3	2.139	0	1	3	6	1.397	3	1.398	0	4	3
		2	2.021	2	-1	-3	10	1.379	8	1.380	4	2	1
33	2.004	31	2.006	2	1	2	8	1.374	6	1.372	4	-3	-1
9	1.941	4	1.942	1	-3	-2	15	1.362	3	1.364	2	-4	-3
30	1.923	28	1.925	2	-3	-1							

809

- 811 Table 6. Crystal data, collection and structure refinement details for gasparite-(La) from the
- 812

Wanni glacier

5								
Crystal Data								
Chemical formula	$(La_{0.60} Ce_{0.40}) ((As_{0.84} P_{0.16}) O_4)$							
$M_{ m r}$	271.39							
Crystal system, space group	Monoclinic, $P2_1/n$							
$a b a (\mathbf{\hat{\lambda}})$	6.7155 (4),							
u, v, c (A)	7.1668 (4), 6.9576 (4)							
β (°)	104.414 (1)							
$V(\text{\AA}^3)$	324.32 (3)							
Ζ	4							
$D_{\rm x} ({\rm g/cm^3})$	5.558							
Crystal size (mm)	$0.04 \times 0.04 \times 0.03$							
Data co	ollection							
Diffractometer	Bruker APEX-II CCD							
Radiation type	MoKα (0.71073 Å)							
$\mu (mm^{-1})$	21.95							
Absorption correction	Multi-scan							
No. of measured, independent								
and	3839, 944, 869							
observed $[I > 2s(I)]$ reflections								
2Θ range for data collection (°)	7.55 to 60.00							
Index ranges	$-9 \le h \le 9, -9 \le k \le 10, -9 \le l \le 9$							
Refin	ement							
$R_{\rm int}, R_{\sigma}$	0.0207, 0.0175							
$R_1[F^2 > 2\sigma(F^2)], wR_2(F^2), S$	0.014, 0.029, 1.12							
No. of reflections	944							
No. of parameters	57							
$\Delta \rho_{\text{max}}, \Delta \rho_{\text{min}} (e \text{ Å}^{-3})$	0.76, -0.55							

813

814 Table 7. Fractional atomic coordinates and isotropic displacement parameters (U_{iso} , $Å^2$) for

815

gasparite-(La) from the Wanni glacier

Site	x	у	z	$U_{ m iso}$	Occupancy (<1)
M(4e)	0.40018 (2)	0.34494 (2)	0.21899 (2)	0.00827 (6)	$(La_{0.6} Ce_{0.4}) *$
X(4e)	0.38653 (4)	0.16300 (4)	0.69644 (4)	0.00626 (11)	$(As_{0.84} P_{0.16})$
O1 (4 <i>e</i>)	0.2803 (3)	0.2142 (3)	0.8797 (3)	0.0131 (4)	
O2 (4 <i>e</i>)	0.5664 (3)	-0.0012 (3)	0.7481 (3)	0.0133 (4)	
O3 (4 <i>e</i>)	0.5031 (3)	0.3390 (3)	0.6113 (3)	0.0122 (4)	
O4(4e)	0.1793(3)	0.1075(3)	0.5181 (3)	0.0130(4)	

816 817 D4 (4e) | 0.1793 (3) | 0.1075 (3) | 0.5181 (3) | 0.0130 (4) |Note:* Occupancy of *M* site was fixed according to electron microprobe data.

818 Table 8. Anisotropic displacement parameters (Å²) for gasparite-(La) from the Wanni glacier

Site	U^{11}	U^{22}	U^{33}	U^{12}	U^{13}	U^{23}
M	0.00902 (8)	0.00690 (9)	0.00836 (8)	0.00092 (6)	0.00117 (5)	0.00025 (6)
X	0.00599 (16)	0.00562 (17)	0.00700 (16)	-0.00001 (11)	0.00131 (11)	0.00013 (11)
01	0.0119 (10)	0.0165 (11)	0.0105 (9)	-0.0006 (8)	0.0024 (8)	-0.0027 (8)
02	0.0108 (10)	0.0086 (10)	0.0199 (10)	-0.0009 (8)	0.0029 (8)	-0.0012 (8)
03	0.0162 (10)	0.0097 (10)	0.0117 (9)	-0.0033 (8)	0.0054 (8)	0.0009 (8)
04	0.0119 (10)	0.0131 (11)	0.0120 (9)	0.0002 (8)	-0.0011 (8)	-0.0048 (8)

819

820 Table 9. Selected bond lengths (Å) and distortion of *M*-polyhedra for gasparite-(La)

821	from the Wanni glacier and	structurally related $M(XO_4)$	compounds (M=La, Ce; X=As, P, V).
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	Length									
Bond	Gasparite- (La)	Gasparite- (Ce)	La(AsO ₄)	Ce(AsO ₄)	La(VO ₄)	La(PO ₄)				
M — 01^{iii}	2.554 (2)	2.486(5)	2.571(3)	2.939(8)	2.521(4)	2.479(3)				
$M - 01^{\text{iv}}$	2.481 (2)	2.527(5)	2.489(3)	2.636(9)	2.497(3)	2.554(3)				
M — 02^{v}	2.479 (2)	2.902(5)	2.498(4)	2.460(7)	2.656(3)	2.783(3)				
$M - 02^{ii}$	2.559 (2)	2.544(4)	2.560(4)	2.619(8)	2.568(3)	2.589(3)				
$M - 03^{ii}$	2.901 (2)	2.632(5)	2.655(3)	2.472(8)	2.528(4)	2.615(3)				
$M - 03^{vi}$	2.562 (2)	2.620(6)	2.577(3)	2.551(7)	2.677(4)	2.503(3)				
М—03	2.6440 (19)	2.457(5)	2.912(4)	2.562(7)	2.887(4)	2.672(3)				
M—O4 ⁱⁱⁱ	2.624 (2)	2.488(5)	2.644(4)	2.476(7)	2.533(4)	2.466(3)				
M—O4 ^{vii}	2.472 (2)	2.555(4)	2.486(3)	2.543(7)	2.502(4)	2.548(3)				
(<i>M</i> -O)	2.586	2.579	2.599	2.584	2.597	2.579				
ΔM	0.035	0.036	0.035	0.038	0.037	0.030				
X-01	1.650 (2)	1.647(5)	1.670(4)	1.682(8)	1.724(4)	1.523(3)				
<i>X</i> —02	1.660 (2)	1.664(5)	1.680(4)	1.699(7)	1.720(4)	1.553(3)				
X03	1.669 (2)	1.659(5)	1.695(4)	1.692(7)	1.699(4)	1.541(3)				
<i>X</i> —04	1.6656 (19)	1.615(5)	1.691(3)	1.686(8)	1.693(4)	1.537(3)				
(X-O)	1.661	1.646	1.684	1.690	1.709	1.538				
Reference	Our data	Kolitsch et al. 2004	Schmidt et al. 2005	Brahim et al. 2002	Rice and Robinson 1976	Ni et al. 1995				

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Table 10. Parameters of Electron Backscatter Diffraction of gasparite-(La) 825

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from the Ushkatyn-III deposit

Electron microscope model	Hitachi S-3400N
EBSD detector	Oxford HKL Nordlys Nano
Acceleration voltage, kV	30
Beam current, nA	0.3
Incidence angle, degrees	70
Sample Tilt	70
Exposure time, seconds per frame	5
Averaging, frames	20
ICSD number of reference	155017 (Sabmidt at al. 2005)
structure	155917 (Schillidt et al. 2005)
Number of indexed bands	12 (max)
Mean angular deviation (MAD)	0.15

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Constituent	Ushkaty	n-III deposit	Wanni glacier		
Constituent	Wt% Range		Wt%	Range	
Fe ₂ O ₃	0.05	0.00-0.16	0.00	0.00	
MnO	1.30	0.91-1.96	0.00	0.00	
CaO	1.33	0.97-1.54	0.34	0.24-0.48	
ThO ₂	0.00	0.00	0.58	0.37-0.72	
Y ₂ O ₃	0.01	0.00-0.05	0.01	0.00-0.03	
La ₂ O ₃	40.21	37.83-41.17	35.59	34.57-36.74	
Ce_2O_3	10.69	9.76-11.61	22.55	21.26-23.67	
Pr ₂ O ₃	1.46	0.00-1.99	0.29	0.00-0.88	
Nd ₂ O ₃	4.24	3.53-4.69	1.04	0.50-1.30	
Sm_2O_3	0.09	0.00-0.33	0.07	0.00-0.35	
V_2O_5	9.77	8.58-11.21	0.00	0.00	
P_2O_5	0.64	0.22-0.99	4.29	4.18-4.45	
As ₂ O ₅	30.32	29.46-31.41	34.48	33.27-35.22	
Total	100.11		100.13		

Table 1. Chemical composition of gasparite-(La)

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1		Table 2. Comparative crystallographic data for monazite group minerals									
	Mineral	Gasparite- (La)	Gasparite- (Ce)	Rooseveltite	Monazite- (La)	Monazite- (Ce)	Monazite- (Nd)	Monazite- (Sm)	Cheralite		
	Chemical formula	La(AsO ₄)	Ce(AsO ₄)	Bi(AsO ₄)	La(PO ₄)	Ce(PO ₄)	Nd(PO ₄)	Sm(PO ₄)	CaTh(PO ₄) ₂		
	Crystal system Monoclinic										
Space group $P2_1/n$											
	<i>a</i> (Å)	6.9576(4)	6.929(3)	6.879(1)	6.8313(10)	6.7880(10)	6.7352(10)	6.6818(12)	6.7085(8)		
	<i>b</i> (Å)	7.1668(4)	7.129(3)	7.159(1)	7.0705(9)	7.0163(9)	6.9500(9)	6.8877(9)	6.4152(6)		
	<i>c</i> (Å)	6.7155(4)	6.697(3)	6.732(1)	6.5034(9)	6.4650(7)	6.4049(8)	6.3653(9)	6.4152(6)		
	β (°)	104.414(1)	104.46(3)	104.84(1)	103.27(1)	103.43(1)	103.68(1)	103.86(1)	103.71(1)		
	Z 4										
	Reference	Our data	Kolitsch et al. 2004	Bedlivy and Mereiter 1982		Ni et a	ıl. 1995		Raison et al. 2008		

2 Note: for comparison u.c.p parameters given as a > c

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Table 3. Reflectance of gasparite-(La) from the Ushkatyn-III deposit measured in air

λ(nm)	R ₁	\mathbf{R}_2	λ(nm)	R ₁	\mathbf{R}_2
400	11.19	9.05	560	10.92	8.90
420	12.04	9.82	580	10.88	8.84
440	12.35	10.15	589	10.87	8.83
460	12.35	10.13	600	10.85	8.81
470	12.16	9.95	620	10.85	8.77
480	11.97	9.77	640	10.89	8.74
500	11.45	9.44	650	10.92	8.77
520	11.15	9.17	660	10.94	8.79
540	11.00	9.00	680	11.09	8.86
546	10.98	8.98	700	11.23	9.08

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Table 4. Raman spectral signatures of gasparite-(La)

	Relative Intetsity (a.u.)				Relat			
Raman			Assignment	Raman	Intetsity	(a.u.)	Assignment	
shift (cm ⁻¹)	Ushkatyn-	Wanni	Assignment	shift, (cm ⁻¹)	Ushkatyn-	Wanni	Assignment	
	III	glacier			III	glacier		
92	0.6			364	3		$v_4(AsO_4)$	
94		3		367		7	$v_4(AsO_4)$	
104	0.8			379	73		$v_2(AsO_4)$	
107		6		395		100	$v_2(AsO_4)$	
125		6		421		28	$v_4(AsO_4)$	
126	0.3			422	9		$v_4(AsO_4)$	
138	2			452	10		$v_4(AsO_4)$	
139		4	Lattice	462		16	$v_4(AsO_4)$	
153	3		vibrations	812		4	$v_3(AsO_4)$	
155		5		822	21		$v_3(AsO_4)$	
190	2	7		826		9	$v_3(AsO_4)$	
203	29	4		843	8		$v_3(AsO_4)$	
264	0.8			848		74	$v_3(AsO_4)$	
267 320		1		860	100		$v_1(AsO_4)$	
	8			864		86	$v_1(AsO_4)$	
334	4							
337		12						

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Table 5. Powder X-ray diffraction data for gasparite-(La) from Wanni glacier

$I_{\rm obs}$	d _{obs}	Icalc	d _{calc}	h	k	l	Iobs	d _{obs}	Icalc	d _{calc}	h	k	l
9	5.39	5	5.39	1	0	-1	19	1.898	13	1.900	0	2	3
13	4.89	10	4.90	0	1	1	6	1.851	3	1.853	3	2	0
12	4.80	9	4.81	1	1	0	19	1.814	19	1.816	2	-2	-3
12	4.30	10	4.31	1	-1	-1			2	1.794	1	3	2
8	4.18	3	4.18	1	0	1			5	1.789	0	4	0
		5	3.61	1	1	1	25	1.780	22	1.780	2	3	1
20	3.574	16	3.577	0	2	0	11	1.727	10	1.728	0	4	1
64	3.361	58	3.363	0	0	2			3	1.681	0	0	4
6	3.249	5	3.249	2	0	0	9	1.673	8	1.673	2	0	-4
100	3.156	100	3.158	0	2	1			7	1.637	0	1	4
22	3.042	18	3.043	0	1	2	17	1.633	5	1.633	0	3	3
		26	2.968	2	-1	-1			2	1.633	4	-1	-1
77	2.956	69	2.958	2	1	0			2	1.631	4	0	-2
24	2.693	22	2.695	2	0	-2			5	1.624	4	0	0
19	2.519	18	2.524	2	-1	-2	8	1.619	3	1.618	2	1	3
19	2.511	12	2.506	2	1	1	8	1.587	6	1.590	4	-1	-2
8	2.446	5	2.450	0	2	2			2	1.584	4	1	0
6	2.407	6	2.410	2	-2	-1	12	1.579	5	1.579	2	-3	-3
5	2.301	3	2.304	1	0	-3			3	1.579	0	4	2
16	2.233	8	2.239	1	3	0	7	1.517	5	1.519	4	-2	-1
		8	2.230	3	0	-1			2	1.494	3	2	2
12	2.189	8	2.193	1	-1	-3	6	1.491	2	1.491	2	-4	-2
11	2.163	6	2.166	1	2	2	4	1.473	2	1.474	4	-1	-3
8	2.138	3	2.139	0	1	3	6	1.397	3	1.398	0	4	3
		2	2.021	2	-1	-3	10	1.379	8	1.380	4	2	1
33	2.004	31	2.006	2	1	2	8	1.374	6	1.372	4	-3	-1
9	1.941	4	1.942	1	-3	-2	15	1.362	3	1.364	2	-4	-3
30	1.923	28	1.925	2	-3	-1							

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13 14 Table 6. Crystal data, collection and structure refinement detailsfor gasparite-(La) from Wanni glacier

for gasparite-(La) from wallin glacter							
Crystal Data							
Chemical formula	$(La_{0.60} Ce_{0.40}) (As_{0.84} P_{0.16}) O_4$						
$M_{ m r}$	271.39						
Crystal system, space group	Monoclinic, $P2_1/n$						
$a h a (\hat{\lambda})$	6.7155 (4),						
a, b, c (A)	7.1668 (4), 6.9576 (4)						
β (°)	104.414 (1)						
$V(\text{\AA}^3)$	324.32 (3)						
Z	4						
$D_{\rm x} ({\rm g/cm}^3)$	5.558						
Crystal size (mm)	0.04 imes 0.04 imes 0.03						
Data co	ollection						
Diffractometer	Bruker APEX-II CCD						
Radiation type	ΜοΚα (0.71073 Å)						
$\mu (mm^{-1})$	21.95						
Absorption correction	Multi-scan						
No. of measured, independent							
and	3839, 944, 869						
observed $[I > 2s(I)]$ reflections							
2Θ range for data collection (°)	7.55 to 60.00						
Index ranges	$-9 \le h \le 9, -9 \le k \le 10, -9 \le l \le 9$						
Refinement							
$R_{\rm int}, R_{\sigma}$	0.0207, 0.0175						
$R_1[F^2 > 2\sigma(F^2)], wR_2(F^2), S$	0.014, 0.029, 1.12						
No. of reflections	944						
No. of parameters	57						
$\Delta \rho_{\rm max}, \Delta \rho_{\rm min} ({ m e}{ m \AA}^{-3})$	0.76, -0.55						

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Table 7. Fractional atomic coordinates and isotropic displacement parameters (U_{iso}, Å²) for gasparite-(La) from Wanni glacier

Site	x	у	z	$U_{ m iso}$	Occupancy (<1)		
M(4e)	0.40018 (2)	0.34494 (2)	0.21899 (2)	0.00827 (6)	$(La_{0.6} Ce_{0.4}) *$		
X(4e)	0.38653 (4)	0.16300 (4)	0.69644 (4)	0.00626 (11)	$(As_{0.84} P_{0.16})$		
O1 (4 <i>e</i>)	0.2803 (3)	0.2142 (3)	0.8797 (3)	0.0131 (4)			
O2 (4 <i>e</i>)	0.5664 (3)	-0.0012 (3)	0.7481 (3)	0.0133 (4)			
O3 (4 <i>e</i>)	0.5031 (3)	0.3390 (3)	0.6113 (3)	0.0122 (4)			
O4 (4 <i>e</i>)	0.1793 (3)	0.1075 (3)	0.5181 (3)	0.0130 (4)			
Note:* Occupancy of <i>M</i> site was fixed according to electron microprobe data.							

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21 Table 8. Anisotropic displacement parameters $(Å^2)$ for gasparite-(La) from Wanni glacier

-				()	8 (·		8
	Site	U^{11}	U^{22}	U^{33}	U^{12}	U^{13}	U^{23}
	М	0.00902 (8)	0.00690 (9)	0.00836 (8)	0.00092 (6)	0.00117 (5)	0.00025 (6)
	X	0.00599 (16)	0.00562 (17)	0.00700 (16)	-0.00001 (11)	0.00131 (11)	0.00013 (11)
	01	0.0119 (10)	0.0165 (11)	0.0105 (9)	-0.0006 (8)	0.0024 (8)	-0.0027 (8)
	O2	0.0108 (10)	0.0086 (10)	0.0199 (10)	-0.0009 (8)	0.0029 (8)	-0.0012 (8)
	03	0.0162 (10)	0.0097 (10)	0.0117 (9)	-0.0033 (8)	0.0054 (8)	0.0009 (8)
	04	0.0119 (10)	0.0131 (11)	0.0120 (9)	0.0002 (8)	-0.0011 (8)	-0.0048 (8)

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Table 9. Selected bond lengths (Å) and distortion of *M*-polyhedra for gasparite-(La) from Wanni glacier

and structurally related MXO4 compounds (M=La, Ce; X=As, P, V).									
		th							
Bond	Gasparite- (La)	Gasparite- (Ce)	La(AsO ₄)	Ce(AsO ₄)	La(VO ₄)	La(PO ₄)			
M — 01^{iii}	2.554 (2)	2.486(5)	2.571(3)	2.939(8)	2.521(4)	2.479(3)			
$M - 01^{iv}$	2.481 (2)	2.527(5)	2.489(3)	2.636(9)	2.497(3)	2.554(3)			
M — 02^{v}	2.479 (2)	2.902(5)	2.498(4)	2.460(7)	2.656(3)	2.783(3)			
M — $O2^{ii}$	2.559 (2)	2.544(4)	2.560(4)	2.619(8)	2.568(3)	2.589(3)			
M — $O3^{ii}$	2.901 (2)	2.632(5)	2.655(3)	2.472(8)	2.528(4)	2.615(3)			
$M - 03^{vi}$	2.562 (2)	2.620(6)	2.577(3)	2.551(7)	2.677(4)	2.503(3)			
М—ОЗ	2.6440 (19)	2.457(5)	2.912(4)	2.562(7)	2.887(4)	2.672(3)			
M—O4 ⁱⁱⁱ	2.624 (2)	2.488(5)	2.644(4)	2.476(7)	2.533(4)	2.466(3)			
M—O4 ^{vii}	2.472 (2)	2.555(4)	2.486(3)	2.543(7)	2.502(4)	2.548(3)			
(<i>M</i> -O)	2.586	2.579	2.599	2.584	2.597	2.579			
ΔM	0.035	0.036	0.035	0.038	0.037	0.030			
<i>X</i> —01	1.650 (2)	1.647(5)	1.670(4)	1.682(8)	1.724(4)	1.523(3)			
<i>X</i> —O2	1.660 (2)	1.664(5)	1.680(4)	1.699(7)	1.720(4)	1.553(3)			
Х—ОЗ	1.669 (2)	1.659(5)	1.695(4)	1.692(7)	1.699(4)	1.541(3)			
<i>X</i> —O4	1.6656 (19)	1.615(5)	1.691(3)	1.686(8)	1.693(4)	1.537(3)			
(X-O)	1.661	1.646	1.684	1.690	1.709	1.538			
Reference	Our data	Kolitsch et al. 2004	Schmidt et al. 2005	Brahim et al. 2002	Rice and Robinson, 1976	Ni et al. 1995			

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Electron microscope model	Hitachi S-3400N					
EBSD detector	Oxford HKL Nordlys Nano					
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Beam current, nA	1					
Incidence angle, degrees	70					
Sample Tilt	0					
Exposure time, seconds per frame	0.5					
Averaging, frames	20					
ICSD number of reference	155017 (Sahmidt at al. 2005)					
structure	155917 (Schilldt et al. 2005)					
Number of indexed bands	12 (max)					
Mean angular deviation (MAD)	0.1053					

























