

1 **Revision 2 of**
2 **Pyrite trace element and sulphur isotope geochemistry of Paleo-Mesoproterozoic**
3 **McArthur Basin: Proxy for Oxidative weathering**
4

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23 **Abstract:** Redox sensitive trace elements and sulphur isotope compositions obtained via in-
24 situ analyses of sedimentary pyrites from marine black shales, are used to track atmosphere-
25 ocean redox conditions between ~1730 and ~1360 Ma in the McArthur Basin, northern
26 Australia. Three black shale formations within the basin, (Wollogorang Formation 1730 ± 3
27 Ma, Barney Creek Formation 1640 ± 3 Ma and Upper Velkerri Formation 1361 ± 21 Ma)
28 display systematic stratigraphic variations in pyrite trace element compositions obtained
29 using LA-ICP-MS. The concentrations of several trace elements and their ratios, such as Se,
30 Zn, Se/Co, Ni/Co, Zn/Co, Mo/Co, Se/Bi, Zn/Bi, Ni/Bi, increase from the stratigraphically
31 lower Wollogorang Formation to the Upper Velkerri Formation. Cobalt, Bi, Mo, Cu and Tl,
32 show a consistent decrease in abundance while Ni, As and Pb show no obvious trends.

33
34 We interpret these trace element trends as a response to the gradual increase of oxygen in the
35 atmosphere-ocean system from ~1730 to 1360 Ma. Elements more mobile during erosion
36 under rising atmospheric oxygen show an increase up stratigraphy (e.g. Zn, Se) whereas

37 elements that are less mobile show a decrease (e.g. Co, Bi). We also propose the increase of
38 elemental ratios (Se/Co, Ni/Co, Zn/Co, Mo/Co, Ni/Bi and Zn/Bi) up stratigraphy are strong
39 indicators of atmospheric oxygenation.

40

41 Sulphur isotopic compositions of marine pyrite ($\delta^{34}\text{S}_{\text{pyrite}}$) from these formations, obtained
42 using SHRIMP-SI, are highly variable, with the Wollgorang Formation exhibiting less
43 variation ($\delta^{34}\text{S} = -29.4$ to $+9.5\%$; mean: -5.03%) in comparison to the Barney Creek ($\delta^{34}\text{S} =$
44 -13.8 to $+41.8\%$; mean: $+19.88\%$) and Velkerri Formations ($\delta^{34}\text{S} = -14.2$ to $+52.8\%$; mean:
45 $+26.9\%$). We propose that the shift in mean $\delta^{34}\text{S}$ to heavier values upsection corresponds to
46 increasing deep water oxygenation from ~ 1730 to 1360 Ma. Incursion of oxygenated waters
47 possibly caused a decrease in the areal extent of anoxic areas, at the same time, creating a
48 possibly efficient reducing system. A stronger reducing system caused the $\delta^{34}\text{S}$ of the
49 sedimentary pyrites to become progressively heavier. Interestingly, heavy $\delta^{34}\text{S}$ in pyrites
50 overlaps with the increase in the concentration of certain trace elements (and their ratios) in
51 sedimentary pyrites (Se, Zn, Se/Co, Ni/Co, Zn/Co, Mo/Co, Ni/Bi and Zn/Bi). This study
52 concludes that there was a gradual increase of atmospheric oxygen accompanied by ocean
53 oxygenation through the first ~ 400 million years of the Boring Billion (1800 - 1400 Ma) in the
54 McArthur Basin.

55

56 INTRODUCTION

57

58 Sedimentary rocks of the McArthur Basin in northern Australia have been the focus of
59 various studies on the application of bio-geochemical techniques to understand ocean and
60 atmospheric redox conditions during the Paleo-Mesoproterozoic. These sedimentary
61 packages are unmetamorphosed, relatively undeformed and believed to preserve imprints of

62 the biological and geochemical processes operative during their deposition (Brasier et al.,
63 1998; Anbar and Knoll, 2002; Shen et al., 2002; Arnold et al., 2004; Brocks et al., 2005;
64 Javaux et al., 2001, 2004; Scott et al., 2008; Johnston et al., 2008; Kendall et al., 2009;
65 Planavsky et al., 2011; Planavsky et al., 2014; Lyons et al., 2014; Large et al., 2014).
66 Particular interest in the McArthur Basin arises because it provides an insight into the
67 “Boring Billion” period (1800-800 Ma), which has been of great interest since being referred
68 to as “one of the dullest period in Earth’s history” (Buick, 1995) and as “a billion year of
69 environmental stasis” (Brasier and Lindsay, 1998). Several techniques have been used to
70 decipher the depositional conditions in order to interpret basin redox structure and its impact
71 on biological processes. Isotopic (C, Sr, S, Mo, Cr) and biomarker studies have been
72 undertaken to provide an overall perspective on the McArthur Basin redox conditions
73 (Brasier and Lindsay, 1998; Lindsay and Brasier, 2000; Shen et al., 2002; Chen et al., 2003;
74 Kah et al., 2004; Brocks et al., 2005; Johnston et al., 2008; Scott et al., 2008; Planavsky et al.,
75 2014; Luo et al., 2015). Bulk rock trace element concentrations in sedimentary rocks,
76 particularly black shales and ironstones, have also enhanced our understanding of Proterozoic
77 ocean biogeochemistry (Lyons et al., 2003; Tribovillard et al., 2006; Konhauser et al., 2009;
78 Sahoo et al., 2012; Reinhard et al., 2013; Partin et al., 2013; Planavsky et al., 2014).
79 Recently, studies by Large et al. (2014), Large et al. (2015 a, b), Gregory et al. (2014, 2015a,
80 2017), and Mukherjee and Large (2016, 2017) have shown that trace element concentrations
81 in marine sedimentary pyrite from black shales can also be used to infer ocean-atmosphere
82 redox relationships. Using the technique introduced by Large et al., (2014), potential
83 oxygenation during the Velkerri Formation depositional period (~1360 Ma) has been
84 proposed (Mukherjee and Large, 2016) on the basis of enrichment of certain redox sensitive
85 trace elements (Se, Mo, Ni, Zn) in pyrite in the black shales.

86

87 The aim of the present contribution is to further our understanding of marine conditions
88 (redox state of the atmosphere and ocean, nutrient trace element availability) in the McArthur
89 Basin by measuring trace element concentrations and S isotopic compositions of marine
90 pyrite in three black shale formations from the McArthur basin stratigraphy. These are the
91 Velkerri Formation (~1400 Ma), Barney Creek Formation (~1650 Ma) and Wollogorang
92 Formation (~1730 Ma) located in in the Roper, McArthur and Tawallah Groups respectively.

93

94 The McArthur Basin was chosen for this study for several reasons. First, the basin is a
95 multiphase tectonic system with five coherent packages (Wilton, Favenc, Glyde, Goyder and
96 Redbank), each with a characteristic age, lithostratigraphy, volcanism (style and
97 composition), basin architecture and micropaleontology (Rawlings, 1999). Of the five
98 packages, three comprise marine black shale formations, which makes the McArthur Basin
99 an ideal location to apply the technique introduced by Large et al. (2014; 2017), of using
100 trace element concentrations of sedimentary pyrite in marine black shales to infer redox
101 conditions of the ocean and atmosphere. The McArthur also provides an opportunity to
102 investigate possible factors such as atmosphere-ocean redox and basin tectonics that
103 controlled the pyrite trace element concentrations. Second, previous studies (Bull, 1998;
104 Jackson et al., 2000; Shen et al., 2002; Brocks et al., 2008; Johnston et al., 2008; Kendall et
105 al., 2009) have shown that the intracratonic McArthur Basin was connected to the global
106 ocean, possibly allowing us to infer global redox conditions between 1730 Ma and 1360 Ma.

107

108 **Importance of combining sedimentary pyrite trace element and sulphur isotope**
109 **compositions**

110

111 Certain redox-sensitive trace elements (Mo, U, Cr, V, Zn) in black shales have been used
112 previously as paleoredox indicators of the water column (Algeo et al., 2006; Tribovillard et
113 al., 2006; Algeo et al., 2009; Algeo et al., 2012; Meyer et al., 2008; Gordon et al., 2009;
114 Sahoo et al., 2012; Sahoo et al., 2016). Also, they have been used to track atmospheric
115 oxygenation through time (Scott et al., 2008; Partin et al., 2013; Lyons et al., 2014).
116 Recently, Large et al. (2014) proposed that trace element concentrations in sedimentary
117 pyrite formed in marine black shales could also be used as proxies for ocean trace element
118 chemistry and atmospheric oxygenation. The pyrite technique relies on the fact that most
119 redox-sensitive trace elements, in bottom waters and pore waters, are readily and efficiently
120 incorporated into sedimentary pyrites (Huerta-Diaz and Morse, 1992; Morse and Arakaki,
121 1993, Rickard et al., 2012; Gregory et al., 2014; Large et al., 2014; Gregory et al., 2015a;
122 Mukherjee and Large, 2016, 2017). The premise on which the technique is based is that
123 enhanced oxidative weathering on land causes an increase in the supply of redox-sensitive
124 trace elements in the riverine flux (dominant source) in the ocean (Bertine and Turekian,
125 1973; Taylor and McLennan, 1995; Scott et al., 2008; Sahoo et al., 2012; Crowe et al., 2013).
126 This is coupled with a decrease in areal extent of sulfidic bottom waters, a main sink for
127 several trace elements (Sahoo et al., 2016). On encountering a redox-boundary, these trace
128 elements become readily adsorbed by sedimentary pyrites forming in anoxic black shales,
129 either in the water column or in the upper most part of the sediments. Hence, an increase in
130 concentrations of redox sensitive trace elements in sedimentary pyrites can act as an indirect
131 proxy for an increase in atmospheric oxygenation (Gregory et al., 2014; Large et al., 2014;
132 Gregory et al., 2015a; Mukherjee and Large, 2016; 2017).

133

134 Bulk-rock sulphur isotope compositions of the McArthur Basin sediments have been studied
135 previously (Donnelly and Crick, 1988; Canfield, 1998; Shen et al., 2002; Shen et al., 2003;

136 Kah et al., 2004; Lyons et al., 2006; Johnston et al., 2008). Those studies produced $^{34}\text{S}/^{32}\text{S}$
137 compositions but additional information on depositional environment, microbial activity and
138 connectivity to the global ocean can be obtained using isotopes such as ^{33}S and ^{36}S
139 abundances (Johnston et al., 2008). Sulphur isotopic compositions of the McArthur Basin
140 sediments have also been reported previously to understand changes in sea water sulphate
141 concentration and the sulphate reservoir, biogeochemical sulphur isotope fractionation
142 patterns, global oxidative sulphur cycling and redox stratification of the Proterozoic oceans
143 (Shen et al., 2002; Kah et al., 2004; Johnston et al., 2008; Luo et al., 2015). In this study, we
144 carried out sulphur isotope analyses of sedimentary pyrites with SHRIMP-SI (Ireland, 2008)
145 and combined them with trace element concentrations to provide additional insight into
146 changes in putative biogeochemical processes operative during the deposition of the three
147 black shales formations.

148

149 There are several advantages of using a combined marine pyrite trace element (TE) and S-
150 isotope approach for understanding ocean chemistry. First, both techniques offer high spatial
151 resolution, so measurements can be performed within a single grain domain, avoiding the
152 problem of differing amounts of the detrital component that may adversely affect whole rock
153 analyses. Many TE, such as Mo, are partitioned between minerals in the shale (Tribovillard et
154 al., 2006), and thus variations in mineral composition can affect the bulk rock analyses
155 (Chappaz et al., 2014). Second, the effects of diagenesis, metamorphism and hydrothermal
156 activity, all of which affect trace element concentrations and sulphur isotopic compositions,
157 can be identified by textural studies of the pyrite (Large et al., 2014). For example, Large et
158 al. (2007; 2009) demonstrated that recrystallization of pyrite during metamorphism releases
159 most trace elements and results in subhedral to euhedral forms of pyrite with low trace
160 element abundances. Hydrothermal pyrites, on the other hand, may be enriched or depleted in

161 trace elements depending on the conditions of formation (temperature, salinity, proximity to
162 vents). These hydrothermal pyrite concentrations therefore do not reflect primary trace
163 element concentrations of the sea water (Large et al., 2014; Mukherjee and Large, 2017).

164

165 **GEOLOGICAL SETTING**

166

167 The McArthur Basin, northern Australia, preserves a Paleo-Mesoproterozoic volcano-
168 sedimentary history (Plumb, 1979a, b; Rawlings et al., 1999) (Fig. 1). The basin is divided
169 into northern and southern elements by the roughly east-west trending Urapunga Fault zone.
170 Both basin elements comprise a central N-S trending element (Walker and Batten Fault Zone
171 respectively) with flanking “shelves”. For this paper, we will refer to the stratigraphy of
172 southeastern part of the Batten Fault Zone, in which the basin fill is divided into four major
173 groups (Tawallah, McArthur, Nathan and Roper, Figure 1). With the exception of the Nathan
174 Group, each includes black shale formations i.e., Wollogorang Formation, Barney Creek
175 Formation, and Velkerri Formation in the Tawallah, McArthur, and Roper Groups
176 respectively (Fig. 2).

177

178 **Stratigraphy and depositional environments of the three black shale formations**

179

180 *Wollogorang Formation*

181

182 The Wollogorang Formation was deposited $1730 \text{ Ma} \pm 4 \text{ Ma}$ based on U-Pb zircon ages from
183 interbedded tuffs (Page et al., 2000). It is 100-150 m thick and is divided into upper and
184 lower units on the basis of lithology. The upper Wollogorang Formation comprises mudstone
185 and dolomitic sandstone whereas the lower part of the formation consists of black shales and

186 dolomitic siltstone (Jackson et al., 1985; Donnelly and Jackson, 1988). The Wollogorang
187 Formation black shales is interpreted to be deposited in a low sulphate, marine intracratonic
188 basin with oxic surface waters and euxinic deep water conditions prevalent during deposition
189 (Jackson et al., 2000; Page et al., 2000; Southgate et al., 2000; Shen et al., 2002).

190

191 *Barney Creek Formation*

192

193 The Barney Creek Formation has a depositional age of 1640 ± 3 Ma (Page and Sweet, 1998).
194 It is divided into the Cooley Dolostone member, W-Fold Shale member and HYC Pyritic
195 Shale member (~100 m). The HYC Pyritic Shale member consists of laminated pyritic,
196 dolomitic carbonaceous siltstone/mudstone and minor tuffaceous mudstone. The black shales
197 of the Barney Creek Formation were deposited in an anoxic environment, in a subsiding
198 basin in which morphological traps were forming due to local contemporaneous vertical
199 tectonism (Bull, 1998).

200

201 *Velkerri Formation*

202

203 The Velkerri Formation is 330 m thick and is composed dominantly of black carbonaceous
204 mudstones with minor grey siltstones and very fine-grained sandstones along with minor
205 calcite nodules, pyritic stringers and glauconitic siltstones (Abbott et al., 2000). The Velkerri
206 is essentially unmetamorphosed and is relatively undeformed. The depositional age of the
207 Velkerri Formation is 1361 ± 21 Ma for the Upper Velkerri Formation and 1417 ± 29 Ma for
208 the Lower Velkerri Formation based on according to Re-Os dating (Kendall et al., 2009). The
209 Velkerri Formation black shales were deposited in a low energy offshore environment
210 (Abbott et al., 2000) as a result of hemipelagic sedimentation under anoxic conditions with

211 cyanobacterial and algal biotic communities as the source of organic matter (Crick et al.,
212 1988; Javaux et al., 2001; Gortner and Grey, 2012).

213

214 **MATERIALS AND METHOD**

215

216 **Sampling Rationale**

217

218 The black shale units of these three formations were chosen for the present study after
219 considering their similarity in lithology (organic-matter rich shales; TOC% & their
220 depositional environments) and importantly, their stratigraphic position (black shales of three
221 different ages). The varying stratigraphic positions will allow the tracking of changes in
222 ocean chemistry. (See Appendix 1 (A1.1); Table A1). Mineralogically, they are similar
223 except that both the Wollogorang and Barney Creek Formations are associated with
224 carbonates while the Velkerri Formation is not (Ahmad et al., 2013).

225

226 **Sample details**

227

228 Only drill core samples were used for the present study (Fig. 1, 3). Samples were obtained
229 from Mt Young 2, MBXDD001 and URAPUNGA-4 drill holes for Wollogorang, Barney
230 Creek and Velkerri Formations, respectively (Fig. 2). The Mount Young 2 (located in the
231 Batten fault zone) and Urapunga 4 (in the Roper Bar area – Urapanga fault zone according to
232 the map) drill cores were provided by Geoscience Australia. The MBXDD001 drill core was
233 provided by the Northern Territory Geological Survey (NTGS) drill core facility.
234 Considering that the Barney Creek Formation hosts the McArthur River SEDEX deposit, the
235 drillhole (MBXDD001) farthest away (~60 km south of the deposit) from the ore body was

236 selected in order to minimize possible effects of hydrothermal fluids (Mukherjee and Large,
237 2017). Black were selected from all three drill cores. Samples were collected every ~10 m
238 down-hole and rock chips from each interval were set in 2.5 cm diameter epoxy mounts and
239 polished with 1-micron diamond paste. Polished laser mounts were examined using reflected
240 light microscope.

241

242 **LA-ICP-MS pyrite trace element analyses**

243

244 The LA-ICP-MS analyses of trace elements in pyrite and the black shale matrix were
245 conducted at CODES, University of Tasmania. Analyses were carried out using a New Wave
246 Research UP-193ss laser microprobe coupled to an Agilent 7700s quadrupole ICP-MS for the
247 following elements and their respective isotopes, ^{13}C , ^{23}Na , ^{24}Mg , ^{27}Al , ^{29}Si , ^{34}S , ^{39}K , ^{43}Ca ,
248 ^{49}Ti , ^{51}V , ^{53}Cr , ^{55}Mn , ^{57}Fe , ^{59}Co , ^{60}Ni , ^{65}Cu , ^{66}Zn , ^{75}As , ^{77}Se , ^{85}Rb , ^{88}Sr , ^{90}Zr , ^{95}Mo , ^{107}Ag ,
249 ^{111}Cd , ^{118}Sn , ^{121}Sb , ^{125}Te , ^{137}Ba , ^{157}Gd , ^{178}Hf , ^{181}Ta , ^{182}W , ^{195}Pt , ^{197}Au , ^{202}Hg , ^{205}Tl , ^{206}Pb ,
250 ^{207}Pb , ^{208}Pb , ^{209}Bi , ^{232}Th and ^{238}U . STGL2b2 (in-house reference material for calibration of
251 relative element sensitivities; Danyushevsky et al., 2011), GSD-1G (USGS reference glass;
252 Jochum et al., 2005) and a stoichiometric pyrite PPP-1 crystal (Gilbert et al., 2014) were the
253 three primary reference materials used to quantify the abundances of chalcophile elements,
254 lithophile elements, and sulphur abundances, respectively. All three reference materials were
255 analysed twice, before and thereafter, 15 to 20 analyses (~every 1.5 hours) as well as at the
256 beginning and end of a run. Backgrounds were analysed for 30-seconds before the signal
257 from the ablated sample was acquired for 40-60 seconds. The laser was operated with ~3.5
258 J/cm^2 laser fluence and 5 Hz laser repetition rate. Samples were ablated in a He atmosphere
259 flowing at a rate of 0.8 l/min, followed by mixing of the ablation stream with the Ar carrier
260 gas (0.85 l/min) for introduction to the ICPMS.

261

262 The ICP-MS instrument was optimized to maximize sensitivity on mid- to high-mass
263 isotopes (>80 amu). Production of molecular oxide species (monitored as $^{232}\text{Th}^{16}\text{O}^+ / ^{232}\text{Th}^+$)
264 and doubly charged ion species (monitored as $^{140}\text{Ce}^{++} / ^{140}\text{Ce}^+$) were maintained at levels
265 below 0.2%. Dwell times on each mass varied between 5 and 30 msec, depending on the
266 count rates. Total sweep time (time required to measure all isotopes once) was 0.76 sec. A
267 total of ~10 spot analyses, with spot size ranging from 15-35 μm , of sedimentary pyrites were
268 performed on each sample. The pyrite (PPP-1) standard was analysed with same spot as
269 unknowns in order to account for Fe-S fractionation during ablation (Gilbert et al., 2014).
270 The glass standards were analysed with 51 μm spot size in order to alleviate heterogeneity of
271 the standards for key elements such as Au and Tl. Also, 5 spot analyses of the black shale
272 matrix (material surrounding the pyrites) were carried out on each sample to measure
273 concentrations of chalcophile and lithophile elements in siliciclastic matrix of the black
274 shales. Approximately, 10-12 analyses were performed on fine grained sedimentary pyrite
275 grains, containing abundant matrix inclusions, hence producing mixed analyses. Therefore, it
276 was necessary to deconvolute the analyses and extract the composition of clean pyrite. Both
277 matrix and pyrite analyses were used in the data reduction process in order to account for
278 mixing of pyrite and matrix components during LA-ICP-MS analyses of small pyrite grains.
279 The analyses of samples from the Velkerri and Wollongorang formations were processed by
280 an algorithm based on subtraction of the matrix component estimated from mass balance
281 (Large et al. 2014). The disadvantages of this method were subjectivity in selection of the
282 pyrite and matrix compositions and difficulty in determining uncertainties. The samples from
283 the Barney Creek Formation were reduced by the method of linear regression followed by
284 normalization of the total, which is described below. Uncertainties on analyses processed
285 using the regression method are presented in Appendix 2. The analysis of a number of

286 samples from the Barney Creek were processed by both methods and produced consistent
287 results within 10% (Appendix 3).

288

289 The LA-ICP-MS analyses of pyrite from the Barney Creek formation were processed by an
290 Excel based data reduction software developed in-house, which uses a linear regression based
291 algorithm for determining chalcophile and siderophile abundances relative to sulphur, for
292 calculation of sulfide composition. The conversion of raw data (counts per second) into
293 concentrations in ppm involved splitting the integration curve obtained from counts per
294 second vs analysis time, into five segments of equal duration. Concentrations represented by
295 each of these segments were reduced using time-equivalent portions of the signal obtained for
296 each of the calibration materials. The data in counts per second were then converted to
297 preliminary ppm values according to established methods (Longerich et al., 1996; Norman et
298 al. 1996), using Fe as the internal standard element. The method assumes a stoichiometric Fe
299 content of pyrite for calculating preliminary compositions, which are then normalized to a
300 100% total of all measured elements with the exception of C and Hg. To calculate the final
301 concentrations, a linear regression equation using S content was employed such that the sum
302 of the chalcophile elements is 100%. Regression fits for individual analyses were visually
303 inspected for all samples. The resulting compositions were close (Appendix 2) to major
304 element composition of stoichiometric pyrite. The analytical precision was estimated using
305 regression analysis for estimation of uncertainty on prediction, however ignoring
306 uncertainties on the measurements. For Se, As, Sb, Ni, Zn, Cu, Pb and Co at high
307 concentrations, the 90% relative uncertainty is <20% for the majority of the analyses.
308 However, uncertainty increases at lower concentrations. The relative analytical precision of
309 the replicate analyses of the standards was less than 5%. Considering that concentrations of

310 many elements vary by 2-3 orders of magnitude in different samples, we conclude the
311 observed trends are outside of analytical uncertainty.

312

313 Sedimentary pyrites from the three black shale units were analysed for Co, Ni, Cu, Zn, As,
314 Se, Mo, Tl, Pb, Bi concentrations using the LA-ICP-MS (Table A2 a, b, c in Appendix 2). A
315 total of 29 black shale samples were selected for the study (Wollogorang Formation: 5;
316 Barney Creek Formation: 15; Velkerri Formation: 9) and ~ten pyrite spot analyses were
317 performed on each sample.

318

319 The pyrite data from the two formations i.e., Velkerri and Barney Creek have been published
320 in Mukherjee and Large (2016) and Mukherjee and Large (2017) respectively. Pyrite data
321 from the Wollogorang formation have been recently published in Mukherjee et al (2018)
322 along with other global Proterozoic formations in a large database provided as background
323 information in the Supplementary Information section of that paper. Here we provide the data
324 for the three formations again as it is the subject of detailed statistical treatment, discussion
325 and interpretation in the context of the McArthur Basin specifically.

326

327 **SHRIMP sulphur isotopes**

328

329 Sulphur isotope compositions ($\delta^{34}\text{S}$ and $\Delta^{33}\text{S}$) of the pyrites were measured *in-situ* with the
330 SHRIMP-SI ion microprobe at the Research School of Earth Sciences, The Australian
331 National University (ANU). The same mounts used for laser ablation analysis were used for
332 SHRIMP analysis, in some cases the same grains Mounts were cleaned using ethanol, a dilute
333 alkaline cleaning solution, de-ionised water and dried in a vacuum oven at 60 °C for 24h.
334 Mounts were then coated with a film of ca. 40 nm of Au before pumping down in the sample

335 lock overnight. SHRIMP-SI measurements were performed with a Cs⁺ primary beam of ~ 2
336 nA focused to sputter an area of ~ 25 μm in diameter. The sample mount was held at -10 kV
337 resulting in a final impact energy at the target of 15keV. Analytical conditions of the analyses
338 are similar to those described in Ireland et al. (2014). Faraday cups were used for
339 simultaneous detection of ³²S⁻, ³³S⁻, and ³⁴S⁻. ³²S⁻ was collected on a 10¹¹ Ω resistor (50V
340 range), ³³S⁻ and ³⁴S⁻ on 10¹¹ Ω resistors (5V range). Under the operating conditions used for
341 these analyses, typical count rates on ³²S⁻ were ~ 0.2 – 1.2 × 10⁹ cps, on ³³S⁻ about 1.2 – 9.2
342 × 10⁶ cps, and on ³⁴S⁻ about 0.7 – 5.2 × 10⁷ cps. Data collection consisted of 1 or 2 sets of 6-
343 10 subsets, 20 s each, with each subset comprising ten 2 s integrations.
344 Analyses of unknown pyrites were bracketed by measurements of pyrite reference material
345 Balmat (δ³⁴S = +15.1 ± 0.2‰, Crowe and Vaughan, 1996). The ³³S/³²S and ³⁴S/³²S ratios,
346 corrected for instrumental mass fractionation, are expressed here in standard delta notation in
347 permil (‰) relative to V-CDT (VCDT: Vienna-Canyon Diablo Troilite) scale, with

$$\delta^{33}S_{V-CDT} = 1000 \times \left[\frac{(^{33}S/^{32}S)_{corrected}}{(^{33}S/^{32}S)_{reference}} - 1 \right] (\text{‰})$$

$$\delta^{34}S_{V-CDT} = 1000 \times \left[\frac{(^{34}S/^{32}S)_{corrected}}{(^{34}S/^{32}S)_{reference}} - 1 \right] (\text{‰})$$

348 Capital delta values (Δ³³S) have been calculated following Farquhar et al. (2003), where

$$\Delta^{33}S = \delta^{33}S - 1000 \times \left(\left(1 + \frac{\delta^{34}S}{1000} \right)^{0.515} - 1 \right)$$

349 Internal precision of single spot analyses of □³⁴S and Δ³³S are usually better than 0.08‰ and
350 0.3‰ (2SE; SE = standard error), respectively. External reproducibility of □³⁴S and Δ³³S
351 values, calculated as the 1sigma standard deviation (SD) of all analyses on the Balmat
352 reference material over the course of an analytical session, were between 0.2 and 0.3‰.

353

354 **RESULTS**

355

356 **Pyrite Textures**

357

358 Only pyrite with framboidal textures, aggregates of fine-grained pyrite, or layers of fine-
359 grained pyrite, identified using reflected light microscopy, were selected for LA-ICP-MS and
360 SHRIMP-SI analyses as these are interpreted as primary diagenetic pyrite (Fig A1 in
361 Appendix 1). Pyrites from all three formations exhibited similar textures i.e., fine-grained
362 pyrite aggregates (10-100 μm) with the size of individual crystals varying between 5-10 μm .
363 However, at two intervals in the Wollogorang Formation intersection (65.8 m and 82.8 m),
364 the pyrites were coarser grained and possibly of hydrothermal origin. These grains were
365 excluded from this study because recrystallized and hydrothermal pyrites are known to have
366 different trace element concentrations and therefore would not reflect contemporaneous
367 seawater trace element chemistry (Large et al., 2009; Large et al., 2014; Gregory et al.,
368 2015a; Mukherjee et al., 2017). For instance, both coarse grained and fine grained pyrites
369 analyses in the same laser mount in the Barney Creek Formation yielded different trace
370 element concentrations (Fig A2 in Appendix 1)

371

372 **LA-ICP-MS pyrite trace element and matrix analyses**

373

374 Geometric means of individual sample intervals from all three formations are presented in
375 Table 1. Geometric means were calculated because trace element analyses in pyrite tend to
376 approximate log-normal distributions. Multiplicative standard deviations (MSD) of trace
377 elements are presented in Table 2. For all the elements presented, MSD ranges between 1 and

378 3. Mean (geometric) pyrite trace element concentrations of the entire set of analyses in the
379 three black shale units are also compared in Table 3.

380

381 Copper, Co, Tl and Bi show marked decreases in concentrations upsection from the
382 Wollongorang Formation to the Upper Velkerri Formation (Fig 4). Mean Co concentrations
383 decrease from 457 ppm in the Wollongorang Formation to 48 ppm in the Upper Velkerri
384 Formation (~9-fold decrease), while Cu decreases from 377 ppm in the Wollongorang
385 Formation to ~81 ppm in the Upper Velkerri Formation (~5-fold decrease). Molybdenum
386 shows a gradual decrease in concentration with higher concentrations in the Wollongorang (77
387 ppm) and Barney Creek Formations (79 ppm) compared to the Upper Velkerri Formation (20
388 ppm). Thallium shows a gradual decrease up stratigraphy with mean values decreasing from
389 20 to 5 ppm. On the other hand, concentrations of Zn and Se increase up stratigraphy (Fig 4).
390 The Zn mean increases from 28 ppm to 96 ppm, almost 3 times, whereas the Se mean
391 increases from 9 ppm to 19 ppm, only 2 times. Nickel, As, Ag and Pb show no obvious
392 trends with stratigraphy (Fig 4).

393

394 Trace element ratios Se/Co, Se/Bi, Ni/Co, Ni/Bi, Zn/Co, Zn/Bi and Mo/Co for the three
395 formations all increase upsection from the Wollongorang Formation to the Velkerri Formation,
396 with Mo/Bi remaining roughly constant (varying little between 6-13) in all three formations
397 (Fig 4). Downhole plots of Se, Mo, Co, Se/Co and Mo/Co for all three formations (Fig 5) also
398 indicate that [Se], Se/Co and Mo/Co increase, and Co and Mo decrease, upsection. The down
399 hole trends are compared with global mean values obtained from analyses of ~5000 pyrites
400 by Large et al. (2014, 2015a, b), as denoted by the black dashed line in Fig 5.

401

402 Statistical T tests were performed on the dataset to confirm that the differences in the
403 geometric mean concentration of TEs of the three formations are significant (Table 4). Mean
404 Mo, Se, Zn, Co, Bi concentrations of the Velkerri Formation are statistically significantly
405 different from those of the Wollgorang and Barney Creek Formations. Mean concentrations
406 of Co and Bi differ significantly between the Barney Creek Formation and Wollgorang
407 Formation, with no significant differences in the means of Mo, Se and Zn concentrations. The
408 mean Ni concentrations of all three black shales do not differ significantly. Other statistical
409 tests such as Spearman Correlations and Analyses of Variance (AOVs) confirm the T-test
410 results (See A1.3 in Appendix 1 and AOVs in Appendix 5).

411

412 Black shale matrix analyses for Ti, Zr, Th and Cr for evaluation of provenance, are presented
413 in Table 5. The mean Ti/Zr of the Velkerri, Barney Creek, and Wollgorang Formations are
414 22, 37 and 25 respectively. The Th/Cr ratios of Velkerri and Barney Creek Formations are ~1
415 and are slightly lower in the Wollgorang Formation (~0.40). Statistical treatment of the data
416 on matrix elements confirm no significant differences between the means of Ti/Zr and Th/Cr
417 ratios in the three black shale formations (Refer Appendix 6).

418

419 **SHRIMP-SI pyrite sulphur isotopes analyses**

420

421 A total of 124 sedimentary pyrite spot analyses were obtained for the three black shale
422 formations (Wollgorang Formation: 31; Barney Creek Formation: 31; Velkerri Formation:
423 62) (Table 6). The $\delta^{34}\text{S}_{\text{VCDT}}$ values of the pyrites from the Wollgorang Formation range
424 between -29.4 to +9.5‰ with an average of $-5.03 \pm 13.62\%$ (n = 31; 1SD) (Table 6). Pyrites
425 from the Barney Creek Formation have a wider range of $\delta^{34}\text{S}_{\text{VCDT}}$ (-13.8 to +41.8‰) with an
426 average of $+19.88 \pm 17.7\%$ (n = 31; 1SD) (Table 6). The $\delta^{34}\text{S}_{\text{VCDT}}$ values of pyrites from the

427 Velkerri Formation range from -14.2 to +65.6‰ with an average of $+26.9 \pm 16.5$ ‰ (n = 62;
428 1SD), slightly higher than pyrites from the Barney Creek Fm. (Table 6). The $\Delta^{33}\text{S}$ values of
429 pyrites from all three formations are low, with values close to zero and an average of $0.36 \pm$
430 0.51 ‰ (n = 31; 1SD), 0.08 ± 0.28 ‰ (n = 31; 1SD) , and 0.01 ± 0.18 ‰ (n = 62; 1SD) for the
431 Wollogorang, Barney Creek, and Velkerri Formations, respectively (Table 6).

432

433 **DISCUSSION**

434

435 **Factors controlling trace element concentrations in pyrite**

436

437 The trace element composition of primary marine pyrite is considered to be a reflection of the
438 coexisting seawater at the time the pyrite formed (Large et al., 2014; 2015). The trace
439 element composition of sea water, in turn, is a function of the coeval sources and sinks, and
440 the residence time of the element in seawater. The latter depends critically on the speciation
441 of the element in the ocean under ambient conditions at any given time. This study takes into
442 account the geological factors: oxidative weathering, tectonic setting and composition of
443 source rocks as determinants of the TE source flux. Whereas factors such as areal extent and
444 types of sinks (oxic, sub-oxic, anoxic) and residence time, likely determine the sink flux. Our
445 main focus in this paper is to evaluate whether trace element concentrations in diagenetic
446 pyrites were a function of changes in source fluxes or changes in areal extent of sinks, or
447 both, i.e., to evaluate possible causes for the trace element compositional variations measured
448 in sedimentary pyrites of the three formations.

449

450 *Oxidative weathering*

451

452 Oxidative weathering on the continents results in the release of redox sensitive trace elements
453 to solution. Trace elements are then transported by rivers to continental margins where they
454 are either sequestered into estuaries or continental margin sediments, or transported to
455 seawater etc. Once in the open ocean, they can be sequestered into pelagic sediments by
456 reduction upon encountering a redox boundary in the water column or sediment-water
457 interface where they may form complexes with organic acids and become incorporated into
458 authigenic sulphides.

459

460 Trace elements such as Mo, Zn, Se, Co, Ni, Cu, As are known to be redox-sensitive, (i.e., will
461 generally undergo a change in mobility at oxidizing or reducing geochemical barriers) and
462 are known to respond to changes in atmosphere-ocean redox conditions (Calvert and
463 Pedersen, 1993; Jones and Manning, 1994; Wignall, 1994; Crusius et al., 1996; Dean et al.,
464 1997, 1999; Yarincik et al., 2000; Morford et al., 2001; Pailler et al., 2002; Algeo and
465 Maynard, 2004; Algeo et al., 2009; 2012; Tribovillard et al., 2006; Nishri and Halicz, 2014;
466 Smith and Huyck, 1999) Table 7. While some elements are not redox-sensitive *sensu stricto*
467 (Bi, Pb), their speciation in the water column is controlled by redox conditions of the
468 atmosphere-ocean system (Refer Table 7). Not all redox sensitive elements in pyrite can be
469 used to infer atmospheric redox conditions because of partitioning into different phases
470 (organic, detrital) other than pyrite. For instance, Ni and Cu may be adsorbed onto organic
471 complexes and Fe-Mn oxides/hydroxides in the sedimentary process. Silver and Tl are not
472 redox sensitive, however Tl in pyrite can be used to screen out pyrites affected by
473 hydrothermal events (Large et al., 2009; Mukherjee and Large, 2016). Some redox sensitive
474 elements (Pb, As) in pyrites may not record any oxygenation trend solely due to their
475 abundant supply in the water column (high source flux). Arsenic is particularly problematic

476 due it being relatively more mobile under reducing conditions than other redox sensitive TE
477 (Smedley and Kinniburgh, 2002).

478

479 An increase in concentrations of redox sensitive elements Se and Zn from ~1730 Ma to 1360
480 Ma suggests an increased supply of these elements into the water column via oxidative
481 weathering on land. Decreases in concentrations of Co and Bi can also be used to support the
482 hypothesis that these elements may have been retained, relatively more than other elements,
483 by Fe-Mn hydroxides and oxides, an important product of oxidative weathering on land. The
484 premise on which the hypothesis is based is described below.

485

486 Experiments have shown that Fe, Al and Mn hydroxides and oxides, and particularly
487 manganese oxides (Gadde and Laitinen, 1974) are efficient scavengers of trace elements in
488 their cationic form (Jackson, 1998; Huang and Germida, 2002; Sparks, 2003; Violante et al.,
489 2008). It is also known that manganese oxide has a special preference for Co compared to
490 other elements in their cationic/anionic form (Smith, 1999). Further evidence lies in
491 experiments designed to understand interactions of certain elements with manganese
492 hydroxides/oxides (Murray et al., 1968, 1975) where the elements showed affinity in the
493 following order $Mg < Ca < Sr < Ba < Ni < Zn < Mn < Co$. Furthermore, manganese oxides
494 specifically, α - Mn_2O_3 , is known to adsorb cobalt much more than Ni, Zn or Cu, according to
495 experiments by McKenzie (1972). Therefore it is possible that the supply of cobalt to the
496 ocean is inhibited by oxidative processes on land where it is retained by oxides/hydroxides on
497 land. (Pickering, 1979; Jackson, 1998; Huang and Germida, 2002; Sparks, 2003; Violante et
498 al., 2008).

499

500 Another evidence supporting this hypothesis comes from three different lines of research on
501 Co concentrations through time. Thermodynamic modelling of the concentration of Co in the
502 oceans (Zerkle et al., 2005), Co concentrations in sedimentary pyrite (Large et al., 2014) and
503 Co/Ti in iron formations (Swanner et al., 2014) collectively confirm that Co concentrations in
504 the oceans have decreased through time coinciding with increases in atmospheric pO₂. Where
505 elements like Se and Mo have shown an increase through time in response to oxygenation,
506 Co shows a reverse trend (Large et al., 2014

507

508 Bismuth behaves similar to Co (despite not being strictly redox sensitive) in response to
509 increases in oxygenation. The element Bi enters the marine realm via two main sources;
510 atmospheric inputs i.e., eolian dust of volcanic origin (Lee et al., 1985; 1986) or river influx,
511 both being comparable source materials of Bi (Bertine et al., 1996). Bismuth, owing to its
512 extensive hydrolytic activity and strong particle reactivity, is also retained in oxyhydroxides,
513 particularly manganese phases (Barnes, 1967; Fowler et al., 2010), resulting in very low
514 concentrations of Bi in modern sea water (Bertine et al., 1996). Currently there is no
515 thermodynamically modelled trend for Bi in the oceans through time. Nevertheless, Large et
516 al. (2014) pointed out that Bi was one of the least abundant trace elements in the ocean today,
517 with a very short residence time (in the order of 20 years; Bertine et al., 1996) and, like Co,
518 possibly decreased in concentration in pyrite as it was retained by Mn-oxides forming in
519 response to oxidative weathering on land.

520

521 Surprisingly, our data shows that Mo concentrations in pyrite decreased from 1730 Ma to
522 1360 Ma. This is surprising because Mo concentration in shales has been used to track redox
523 changes in the ocean and atmosphere (Scott et al., 2008; Gordon et al., 2008; Kendall et al.,
524 2009; Sahoo et al., 2012) and, whereas most other trace element data indicates an increase in

525 oxygenation of the atmosphere, the Mo concentrations measured in the pyrites studied here
526 seem to suggest a decrease in atmospheric oxygen over this period. It is possible that Mo is
527 retained in the black shale matrix rather than pyrite, causing the values in pyrite to be
528 suppressed in the (~1360 Ma) Velkerri Formation. However, that is not the case for our data.
529 Whole rock values of Mo in the Velkerri Formation have previously been compared to their
530 concentration in pyrite by Mukherjee et al., (2016). Whole rock Mo concentrations range
531 between 2-37 ppm compared to 4-795 ppm in pyrite. The other reason for this surprising
532 behavior of Mo could be a source flux issue i.e., a possibly Mo-rich source for the Tawallah
533 Group. The possibility that Mo could be weathered and mobilized even under mildly oxic
534 weathering conditions (Anbar et al., 2007) and is very sensitive to sulphide oxidation
535 (William and Rickaby, 2012) could also potentially lead to Mo-rich pyrites of the
536 Wollogorang Formation.

537

538 This unusual decrease in Mo in pyrite suggests caution when using a single particular trace
539 element concentration or isotope ratio as a proxy for tracking changes in redox conditions.
540 Another way of tackling this issue is to use ratios of two elements that exhibit antithetic
541 behavior i.e., one element that might be expected to increase in concentration in pyrite with
542 atmosphere oxidation, and the other that might be expected to decrease in concentration in
543 pyrite. Thus, a more robust way to evaluate atmosphere oxygenation would be to couple
544 certain trace elements that show enrichments in pyrite under conditions of increasing
545 oxygenation (e.g., Se, Zn, Mo, Ni) with others that show depletions in pyrite under conditions
546 of increasing oxygenation (e.g., Co, Bi). For instance, even though pyrite from the
547 Wollogorang Formation shows increased Mo concentrations relative to pyrite from the
548 Velkerri shales, the higher Co and Bi contents of pyrite argue against the possibility of
549 oxygenation. Similarly, Ni in pyrite does not display a particular trend, but its concentration

550 increases with respect to Co at any given time through the 370-million-year period. All of the
551 Co/Bi ratios (Ni/Co, Se/Co, Mo/Co, Zn/Co and Zn/Bi ratios in Fig 6; Table 3) increase with
552 sedimentation age and provide robust evidence for increasing atmosphere oxygenation over
553 this interval of time (1730 to 1360 Ma).

554

555 *Source rock composition*

556

557 So far, our discussion of redox-sensitive trace element compositions in pyrite in the
558 Proterozoic marine sediments studied here has been linked to an increase in atmospheric
559 oxygen, which is assumed to be related directly to the intensity of oxidative weathering on
560 the continents. Nevertheless, source rock compositions (dominantly mafic or felsic) and basin
561 tectonics (active or less active) at the time of deposition of these black shales, are important
562 factors that need to be addressed in order to evaluate the nature of the trace element flux.
563 Below we summarise previous work on provenance analyses of the three black shale
564 formations studied here. Past studies of their sedimentology, including paleo current
565 measurements, sequence stratigraphy, and U-Pb zircon geochronology, have indicated that
566 the Wologorang and Barney Formation sediments were derived from weathering and erosion
567 of the basal Tawallah Group, i.e., local sources. Deposition of these formations was related to
568 stretching and thinning of the basin floor in a rift environment (Rogers, 1996; Giles et al.,
569 2002). On the other hand, the Roper Group was deposited in an intracratonic ramp setting
570 soon after the Isan Orogeny (Abbott and Sweet, 2000). Sediments of the Roper Group are
571 likely to have been derived from the weathering and erosion of the North Australian craton
572 basement rocks (Munsen et al., (2016). The question of whether the characteristic pyrite trace
573 element concentrations in the Velkerri Formation have been affected due to change in
574 provenance, needed to be further investigated.

575

576 We used Ti, Zr, Th, and Cr concentrations, augmented by literature data, to investigate
577 potential influences of source rock composition on trace element concentrations in pyrite.
578 Several studies have highlighted the importance of trace element ratios (Cr/Ni, V/Cr, V/Ni,
579 La/Yb, La/Sm, Ti/Zr, Y/Ni, Zr/Cr, etc.) in determining the type of source rocks from which
580 the sediments were derived (Bhatia, 1983; Condie and Wronkiewicz, 1990; McLennan et al.,
581 1990; McLennan and Taylor, 1991; Garver and Scott, 1995; Cingolani et al., 2003; Ali et al.,
582 2014). We have taken a similar approach and used Ti/Zr and Th/Cr ratios of the matrix to
583 infer differences in the types of source rocks (Table 5; black shale matrix analyses of Ti, Zr,
584 Cr, Th). This approach is based on the premise that Ti and Cr are more enriched in mafic
585 rocks (means 10,000 and 250 ppm respectively; Riemann and De Caritat, 1998) than felsic
586 rocks (means 3,000 and 10 ppm respectively), while Zr and Th are enriched in felsic rocks
587 (means of 200 and 15 ppm respectively) relative to basic rocks (means 120 and 2.2 ppm).
588 Figure 7 illustrates these compositions in a plot between Ti/Zr and Th/Cr with an assigned
589 point for mean basalt and rhyolite. It is evident from Fig 7 (Th/Cr and Ti/Zr ratios of
590 analyses) that all three black shale formations cluster together near the rhyolite (felsic) mean
591 rather than showing any consistent variable association. In all three cases, source rocks are
592 interpreted to be similar in terms of composition. The formation that most deviates from a
593 rhyolitic source towards a mafic source is the Velkerri Formation which is depleted in Co,
594 which would be expected to be enriched if the source of the trace elements was largely
595 detrital. The data suggests that changes in trace element concentrations in the pyrites are most
596 likely related to changes in the redox state of the atmosphere rather than changes in source
597 rock compositions available for erosion.

598

599 We have considered only the source flux of the trace elements in the above discussion.
600 However, changes in sink fluxes could be important in controlling trace element
601 concentrations in pyrite, as a result we have used sulphur isotope compositions of pyrite to
602 give us an insight into changes in sinks, described in the next section.

603

604 **Sulphur isotopes in the McArthur Basin**

605

606 Sulphur isotope compositions of sedimentary sulphides and sulphates of the McArthur Basin
607 sediments have been studied by a number of researchers (Donnelly and Jackson, 1988;
608 Canfield, 1998; Shen et al., 2002; Shen et al., 2003; Kah et al., 2004; Lyons et al., 2006;
609 Johnston et al., 2008). All three black shale units have been sampled for sulphur isotope
610 analysis in order to understand the amount and patterns of fractionation between sea water
611 sulphate and sulphide, including constraining sea water sulphate levels in the Middle
612 Proterozoic (Canfield, 1998). The previous studies (using measured values of $\delta^{34}\text{S}_{\text{CAS-PY}}$ and
613 $\partial\delta^{34}\text{S}_{\text{CAS}}/\partial t_{(\text{max})}$; CAS: Carbonate-associated sulphate and PY: pyrite) indicate that the Middle
614 Proterozoic was a period of very low sea water sulphate values (0.1-1.8mM) compared to
615 28mM today (Kah et al., 2004; Luo et al., 2015). Based on S-isotope patterns, it has also been
616 suggested that the Wollongorang and Velkerri Formations were deposited in a basin connected
617 to the global ocean (Shen et al., 2002; Shen et al., 2003; Johnston et al., 2008). The Barney
618 Creek Formation, on the other hand, has features that suggest it formed in a semi-restricted
619 basinal environment with progressive interaction with the global sulphate reservoir during its
620 deposition (Johnston et al., 2008).

621

622 As part of this study, sedimentary pyrite was analysed for ^{32}S - ^{33}S - ^{34}S , in order to understand
623 variations of $\delta^{34}\text{S}$ in pyrite between the three black shales. $\delta^{34}\text{S}_{\text{VCDT}}$ in the Velkerri Formation

624 displays the greatest range and the highest mean of all three formations (-14.2 to +52.8‰;
625 mean: +26‰) (Fig 8; Table 7). The Barney Creek Formation has a similar range of $\delta^{34}\text{S}_{\text{VCDT}}$
626 values but a lower mean (-13.8 to +41.8‰; mean: +19.5‰), (Fig 8). In contrast, the
627 Wollogorang Formation has a more restricted range of $\delta^{34}\text{S}_{\text{VCDT}}$, extending to the most
628 negative values observed in this study, and the lowest mean (-29.4 to +9.5‰; mean: -5.0‰)
629 (Fig 8).

630

631 The in-situ $\delta^{34}\text{S}$ values in pyrite obtained using SHRIMP-SI had a similar range and mean
632 when compared to the conventional techniques utilized in previous research (Johnston et al.,
633 2008; Shen et al., 2002). This is indicated in Figure 8 where the shaded regions represent the
634 range of data obtained by previous studies. However, in some instances our data show a
635 greater spread compared with other studies. This is to be expected when results of a micro-
636 beam technique are compared with those of a micro-drill technique. We have documented
637 extremely heavy pyrites (^{34}S enriched) in the Barney Creek and Velkerri Formations relative
638 to published data. Based on the $\delta^{34}\text{S}$ values presented in Tables 6, all three formations also
639 have negative $\delta^{34}\text{S}$ values, which could be indicative of microbial activity i.e., bacterial
640 sulphate reduction. Pyrite from the Wollogorang Formation records the most negative $\delta^{34}\text{S}$
641 values (-20 to -30‰) relative to the other two formations where the negative $\delta^{34}\text{S}$ values
642 range between 0 to -14‰ (see above quoted data). Similarly, pyrite in the Wollogorang
643 Formation has the least positive $\delta^{34}\text{S}$ values (maximum of +10‰), compared to >40‰ in the
644 Barney Creek Formation >50‰ in the Velkerri Formation. Overall, there is an increase in not
645 only the range of positive $\delta^{34}\text{S}$ values but also the $\delta^{34}\text{S}$ mean of each formation up
646 stratigraphy (Table 6).

647

648 This increasing trend of $\delta^{34}\text{S}_{\text{pyrite}}$ values upsection could be interpreted in several ways.
649 Negative $\delta^{34}\text{S}$ values are generally interpreted as being associated with microbial sulphate
650 reduction because of the preferential uptake of ^{32}S relative to ^{34}S in sulphide, as the process is
651 chemically and kinetically more favourable in metabolic pathways. This causes low values of
652 $\delta^{34}\text{S}_{\text{pyrite}}$ and can be reflected in later diagenetic pyrite that forms from the ^{34}S -depleted pore
653 fluids. Thus, any increase in $\delta^{34}\text{S}_{\text{pyrite}}$, is difficult to attribute to microbial activity acting on an
654 effectively infinite reservoir of sulphate. However, diagenetic pyrite could either be more
655 depleted or more enriched in ^{34}S depending on whether it precipitates in an open or close
656 system, which regulates the supply of sulphate in the area of pyrite formation.

657

658 Generally, heavy pyrites (^{34}S enriched) are attributed to precipitation under a closed system
659 (with limited or no additional supply of sulphate) (Shen et al., 2002; Shen et al., 2003;
660 Johnston et al., 2008; Gregory et al., 2015b). Continuing sulphate reduction leads to
661 progressively ^{34}S -enriched pyrites as they begin to record the isotopic signature of the closed-
662 system sulphate. If that were the case, the data from the present study would indicate that the
663 increasing trend of $\delta^{34}\text{S}$ values upsection reflects a transition from an open system in the
664 Wologorang Formation to a closed system in the Velkerri Formation. However, such a
665 scenario is not compatible with sedimentological, paleontological and geochemical studies
666 that have confirmed an open marine origin for these rocks (Jackson et al., 1991; Abbott and
667 Sweet, 2000; Shen et al., 2003; Javaux et al., 2001; Kendall et al., 2009). Previous studies
668 have attributed the negative $\delta^{34}\text{S}$ values in the Velkerri Formation to reduction of sea water
669 sulphate in a basinal setting with non-limiting sulphate supply and vice versa. (Logan et al.,
670 1995; Shen et al., 2003). Johnston et al. (2008) used not only $\delta^{34}\text{S}_{\text{pyrite}}$ but ^{33}S and ^{36}S (larger
671 variability in $\Delta^{33}\text{S}$ inversely correlated to $\Delta^{36}\text{S}$) to show that pyrite deposition in the Velkerri
672 Formation black shales occurred in an open marine setting with connectivity to the global

673 ocean. In fact, Kah et al. (2004) and Luo et al. (2015) proposed an increase in sea water
674 sulphate levels based on their measured values of sulphur isotopes in the pyrites from the
675 Velkerri Formation.

676

677 The Barney Creek Formation is considered to have formed in a partly restricted basin, and
678 the positive $\delta^{34}\text{S}_{\text{pyrite}}$ values may be related to the closed system behavior. Indeed, previous
679 studies by Shen et al. (2002) and Johnston et al. (2008) have pointed out that the limited
680 supply of sulphate in a closed system may have rendered $\delta^{34}\text{S}$ values in pyrite close to
681 contemporaneous sea water (+25‰). However, we observe much heavier values than
682 reported in the literature so far. Closed system behavior could also be created locally within
683 sediments below the sediment-water interface. That could have led to a tendency toward
684 heavy values in the Barney Creek and Velkerri Formations, but interestingly, very heavy
685 values are absent from the Wologorang Formation. Therefore, it is not possible to link the
686 increasing trend of $\delta^{34}\text{S}$ values we observe in our pyrite to microbial sulphate reduction, lack
687 of connection to global ocean, or quantitative sulphate reduction in the sediments. The
688 increasing trend of $\delta^{34}\text{S}$ values upsection demands an alternative explanation, which we
689 explore below.

690

691 We suggest an alternative explanation for the increasing mean $\delta^{34}\text{S}$ trend with the help of a
692 simple schematic diagram (Fig 9). Two scenarios are proposed; scenario 1 represents
693 depositional conditions of the Wologorang Formation, and scenario 2 represents the
694 depositional conditions of the Velkerri Formation. The Barney Creek depositional condition
695 is assumed to be intermediate between these two scenarios. Both scenarios represent open
696 marine conditions with low global sulphate levels in the oceans, as is the case with the three
697 black shale formations. In scenario 1, there is less oxygen in the atmosphere impeding

698 oxidative weathering on land. This leads to a low supply of certain trace elements, sulphate
699 and other nutrients that, in turn, would cause low productivity and possibly growth of the
700 prevalent biologic community at 1730 Ma, and consequent lower organic matter production
701 and deposition. In this case, it is envisaged that black shales are forming in an anoxic
702 environment mainly due to low organic matter deposition and a lack of oxygen in the water
703 column related to low pO_2 in the atmosphere. Low productivity would cause a decrease in the
704 availability of reactive organic compounds for sulphate reduction. Scenario 1 reflects anoxic
705 conditions over a large area of the ocean floor, with lower sulphate reduction rates & lower
706 pyrite burial proportions, and consequently forming pyrites that lack heavy $\delta^{34}S_{\text{pyrite}}$ values.

707

708 Scenario 2 has higher oxygen contents in the atmosphere, causing greater oxidative
709 weathering on land. This leads to a higher input of trace elements, sulphate and nutrients in
710 the ocean, promoting an increase in organic matter production and organic carbon deposition.
711 This causes two important changes. First, increased productivity, leads to increased release of
712 oxygen into the ocean and atmosphere, and second, higher levels of organic carbon
713 production and deposition ensures ample availability of reactive organic compounds in the
714 sediments for sulphate reduction. Subsequently, the process leads to a strongly reducing
715 environment (anoxic to euxinic) with higher sulphate reduction rates and pyrite burial
716 compared to scenario 1. The pockets or sub-basins of anoxia are believed to be more focused,
717 possibly due to basin architecture and also the increase in oxygen of the overall ocean-
718 atmosphere system, causing the region of anoxia/euxinia to shrink (organic matter is
719 regionally oxidized and only locally preserved in sub-basins). However, because organic
720 matter deposition occurs over a smaller area, it causes sulphate reduction rates and pyrite
721 burial to be higher. In turn, this causes a relative increase in $\delta^{34}S$ values in pyrite compared

722 with scenario 1 (Fike and Grotzinger, 2008, Leavitt et al., 2013; Fike et al., 2015; Pasquier et
723 al., 2017; Liu et al., 2019).

724

725 The proposed oxygenation event is also consistent with an ocean oxygenation event at ~1400
726 proposed by Yang et al. (2017) using U isotopes and the previously proposed increase in sea
727 water sulphate concentrations (Kah et al., 2004; Luo et al., 2015) due to oxidative weathering
728 on land. The observed increased complexity of microorganisms in the Velkerri Formation
729 relative to Barney Creek and Wollogorang Formations (Javaux et al., 2001; 2004), may also
730 be attributed to these changes. Although this model is qualitative, speculative and requires
731 further study to evaluate, it is supported by marine S isotope fractionation at another critical
732 time in ocean evolution. The increase of $\delta^{34}\text{S}$ values in pyrite during the Neoproterozoic
733 Oxygenation Event (NOE or GOE 2) is apparent at the Precambrian-Cambrian transition
734 (Claypool et al., 1980; Strauss and Schieber, 1990; Strauss et al., 1992; Strauss, 1993) where
735 $\delta^{34}\text{S}_{\text{pyrite}}$ varies between +7.6 to +53‰ over the period 600 to 520 Ma, followed by the
736 Cambrian explosion of life. Although, a number of primary (Fike and Grotzinger, 2008; Ries
737 et al., 2009, Fike et al., 2015) and secondary (Cui et al., 2018) processes could explain the
738 phenomenon. Modern-day examples of isotopically heavy pyrites under the current high pO_2
739 are described in (Borowski et al., 2000; Borowski et al., 1996; Borowski et al., 2013;
740 Jørgensen et al., 2004; Lin et al., 2016). In addition, strongly reducing environments, such as
741 those found in the Black Sea and Cariaco Basin, do not impart heavy values of $\delta^{34}\text{S}$ in pyrite
742 compared to Proterozoic black shales. This could be due to the vast reservoir of sea water
743 sulphate (28mM) today compared to the Proterozoic (0.1-1.8mM) [Luo et al., 2015].

744

745 **IMPLICATIONS**

746

747 Trace element concentrations (and their ratios) of sedimentary pyrite in three black shale
748 formations of the McArthur Basin provide evidence for a gradual increase in atmospheric
749 oxygenation from 1730 to 1360 Ma. We also observe a marked change in pyrite sulphur
750 isotopic compositions in the three black shale formations i.e., a marked increase in mean
751 $\delta^{34}\text{S}_{\text{pyrite}}$ values from the Wollongorang Formation to the Velkerri Formation. This is possibly
752 indicative of expansion of oxygenated waters and decreasing areal extent of anoxia. Results
753 from both techniques have major implications on the atmospheric redox evolution in “Boring
754 Billion”, a period believed to witness non-fluctuating redox conditions.

755

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765 **References**

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Table 1 LA-ICP-MS pyrite analyses (geometric mean) for the Wologorang Formation (n=46), Barney Creek Formation (n=115) and Velkerri Formation (n=55) in ppm; m=met

Formation	Depth in meters (m)	Co	Ni	Cu	Zn	As	Se	Mo	Tl	Pb	Bi	Data source
Wologorang Formation	67.8	212	320	64	138	936	3	226	78	113	11	Mukherjee et al., 2018
Wologorang Formation	73.8	590	511	468	103	2958	14	201	32	666	13	Mukherjee et al., 2018
Wologorang Formation	74.3	148	314	110	75	2027	12	871	8	217	80	Mukherjee et al., 2018
Wologorang Formation	77	354	179	744	4	8472	4	9	62	169	1.29	Mukherjee et al., 2018
Wologorang Formation	85.4	2596	1102	964	1	6938	32	12	1	148	113	Mukherjee et al., 2018
Barney Creek Formation	47.5	183	475	109	11	965	8	357	21	970	3	Mukherjee and Large, 2017
Barney Creek Formation	54.2	76	218	59	323	262	7	59	3	236	14	Mukherjee and Large, 2017
Barney Creek Formation	60.8	94	242	102	229	393	6	79	10	222	15	Mukherjee and Large, 2017
Barney Creek Formation	68.5	53	141	92	88	375	4	106	5	114	11	Mukherjee and Large, 2017
Barney Creek Formation	70.2	775	1258	770	17	1587	24	71	23	1474	4	Mukherjee and Large, 2017
Barney Creek Formation	76.1	22	73	44	2	269	6	8	10	323	0.425	Mukherjee and Large, 2017
Barney Creek Formation	81.9	148	354	102	96	346	6	215	7	380	6	Mukherjee and Large, 2017
Barney Creek Formation	86.8	158	335	92	11	1192	8	169	15	462	9	Mukherjee and Large, 2017
Barney Creek Formation	93	160	479	293	45	1769	13	446	9	838	11	Mukherjee and Large, 2017
Barney Creek Formation	98.3	118	202	136	29	972	5	202	16	428	5	Mukherjee and Large, 2017
Barney Creek Formation	100.7	112	275	108	74	380	6	274	18	493	8	Mukherjee and Large, 2017
Barney Creek Formation	109.8	133	271	44	159	307	5	194	12	288	19	Mukherjee and Large, 2017
Barney Creek Formation	124.5	108	231	43	21	1473	5	67	51	274	18	Mukherjee and Large, 2017
Barney Creek Formation	128.4	106	170	121	9	1321	49	5	1	190	2	Mukherjee and Large, 2017
Barney Creek Formation	148.8	572	1246	240	5	7442	11	3	53	933	4	Mukherjee and Large, 2017
Velkerri Formation	115.7	2	85	48	22	20	5	5	0.412	5	1	Mukherjee and Large, 2016
Velkerri Formation	126.3	40	223	41	47	36	10	38	1	80	3	Mukherjee and Large, 2016
Velkerri Formation	140.1	40	685	100	107	7512	87	37	7	404	1	Mukherjee and Large, 2016
Velkerri Formation	151.2	34	444	64	1560	821	34	30	10	121	3	Mukherjee and Large, 2016
Velkerri Formation	161.8	187	1671	176	397	605	38	415	18	371	8	Mukherjee and Large, 2016
Velkerri Formation	172.4	22	144	39	118	1921	7	3	10	111	1	Mukherjee and Large, 2016
Velkerri Formation	183.4	105	504	104	8	4787	48	1	13	384	2	Mukherjee and Large, 2016
Velkerri Formation	193.9	85	260	112	165	173	8	14	4	97	2	Mukherjee and Large, 2016
Velkerri Formation	216.9	122	236	54	81	380	9	9	3	56	2	Mukherjee and Large, 2016

Table 2 Multiplicative standard deviation for the Wollogorang Formation (n=46), Barney Creek Formation (n=115) and Velkerri Formation (n=55) in ppm

Formation	Depth in meters (m)	Co	Ni	Cu	Zn	As	Se	Mo	Tl	Pb	Bi
Wollogorang Formation	67.8	1.27	1.09	1.11	1.84	1.07	1.15	1.06	1.19	1.28	1.04
Wollogorang Formation	73.8	1.09	1.12	1.32	1.33	1.06	1.11	1.14	1.31	1.10	1.08
Wollogorang Formation	74.3	1.07	1.08	1.21	1.04	1.28	1.18	1.24	3.11	1.52	2.12
Wollogorang Formation	77	1.37	1.34	1.49	2.46	1.37	1.39	2.12	1.19	1.37	1.51
Wollogorang Formation	85.4	1.09	1.28	1.30	3.08	1.19	1.12	1.13	1.88	1.19	1.20
Barney Creek Formation	47.5	1.65	1.57	1.26	2.68	1.49	1.18	1.18	1.38	1.68	2.74
Barney Creek Formation	54.2	1.16	1.14	1.31	1.14	1.13	1.49	1.22	1.80	1.20	1.11
Barney Creek Formation	60.8	1.11	1.11	1.31	1.29	1.14	1.33	1.32	1.48	1.10	1.15
Barney Creek Formation	68.5	1.22	1.11	1.12	1.56	1.21	1.38	1.21	1.36	1.65	2.32
Barney Creek Formation	70.2	2.46	2.04	1.55	3.07	1.33	1.41	2.15	1.60	1.39	2.42
Barney Creek Formation	76.1	1.23	1.17	1.03	1.25	1.15	1.03	1.17	1.07	1.02	1.14
Barney Creek Formation	81.9	1.08	1.10	1.06	1.42	1.14	1.15	1.10	1.10	1.06	1.05
Barney Creek Formation	86.8	1.02	1.02	1.12	1.52	1.07	1.04	1.19	1.06	1.04	1.15
Barney Creek Formation	93	1.19	1.28	1.49	1.59	1.53	1.39	1.17	1.13	1.17	1.07
Barney Creek Formation	98.3	1.01	1.02	1.05	1.11	1.04	1.07	1.05	1.04	1.03	1.05
Barney Creek Formation	100.7	1.11	1.14	1.15	1.18	1.17	1.13	1.18	1.19	1.07	1.20
Barney Creek Formation	109.8	1.11	1.05	1.08	1.37	1.08	1.76	1.31	2.14	1.08	1.09
Barney Creek Formation	124.5	1.09	1.08	1.23	1.25	1.16	1.31	1.16	1.09	1.11	1.03
Barney Creek Formation	128.4	1.26	1.35	1.52	2.35	1.56	1.29	3.09	1.97	1.25	1.51
Barney Creek Formation	148.8	1.19	1.19	1.25	1.08	1.46	1.48	2.34	1.38	1.24	1.31
Velkerri Formation	115.7	1.43	1.54	2.36	1.51	1.93	1.66	1.63	2.10	1.55	1.00
Velkerri Formation	126.3	1.56	1.42	1.36	1.14	1.33	1.51	1.24	1.27	1.23	1.25
Velkerri Formation	140.1	1.16	1.16	1.24	1.39	1.10	1.11	1.24	1.10	1.10	1.00
Velkerri Formation	151.2	2.09	1.83	1.15	1.41	1.26	1.19	1.08	1.94	1.83	1.55
Velkerri Formation	161.8	1.31	1.47	1.52	1.93	1.30	1.24	1.66	1.27	1.32	1.31
Velkerri Formation	172.4	1.67	1.72	1.41	2.60	2.86	2.06	2.50	1.25	1.67	1.14
Velkerri Formation	183.4	1.30	1.43	1.67	2.31	2.12	1.42	2.21	1.27	1.54	1.31
Velkerri Formation	193.9	1.68	1.68	2.40	1.53	2.19	1.64	2.46	2.13	1.63	1.41
Velkerri Formation	216.9	1.79	1.75	2.21	2.45	2.20	1.84	1.94	2.36	1.81	1.51

Table 3 LA-ICP-MS pyrite analyses (geometric mean) for all three black shale formations in ppm.

Formation	Age	Co	Ni	Cu	Zn	As	Se	Mo	Tl	Pb	Bi	Se/Co	Ni/Co	Zn/Co	Mo/Co	Se/Bi	Ni/Bi	Zn/Bi	Mo/Bi
Upper Velkerri (n=55)	~1360 Ma	48	387	81	96	502	19	20	5	137	2	0.40	8.00	1.98	0.42	8.29	166	41	9
Barney Creek (n=115)	~1640 Ma	138	318	120	33	772	8	79	11	419	6	0.06	2.30	0.24	0.58	1.36	51	5	13
Wollogorang (n=46)	~1730 Ma	457	377	377	28	3527	9	77	20	254	13	0.02	0.83	0.06	0.17	0.71	30.03	2.24	6.14

Table 4 Statistical t-test for equality of means in the three black shale formations; VF: Velkerri Formation, BCF: Barney Creek Formation, WF: Wologorang Formation (See supplement)

Elements	Formations	t-value	t-critical	p-value	Hypotheses of equal means ($\mu_1=\mu_2$)	Reason*	Difference in means
Mo	VF-BCF	-3.88	1.98	0.00	Reject	$-(t\text{-critical}) > t\text{-value}$	significant
	VF-WF	-3.19	1.98	0.00	Reject	$-(t\text{-critical}) > t\text{-value}$	significant
	BCF-WF	0.09	1.98	0.74	Cannot reject	$t\text{-value} < (t\text{-critical})$	insignificant
Se	VF-BCF	4.19	1.98	0.00	Reject	$t\text{-value} > (t\text{-critical})$	significant
	VF-WF	3.46	1.98	0.00	Reject	$t\text{-value} > (t\text{-critical})$	significant
	BCF-WF	-0.33	1.98	0.58	Cannot reject	$-(t\text{-critical}) < t\text{-value}$	insignificant
Ni	VF-BCF	0.99	1.98	0.32	Cannot reject	$t\text{-value} < (t\text{-critical})$	insignificant
	VF-WF	0.12	1.98	0.90	Cannot reject	$t\text{-value} < (t\text{-critical})$	insignificant
	BCF-WF	-1.27	1.98	0.20	Cannot reject	$-(t\text{-critical}) < t\text{-value}$	insignificant
Zn	VF-BCF	3.46	1.98	0.00	Reject	$t\text{-value} > (t\text{-critical})$	significant
	VF-WF	2.92	1.98	0.00	Reject	$t\text{-value} > (t\text{-critical})$	significant
	BCF-WF	0.44	1.98	0.65	Cannot reject	$t\text{-value} < (t\text{-critical})$	insignificant
Co	VF-BCF	-4.38	1.98	0.00	Reject	$-(t\text{-critical}) > t\text{-value}$	significant
	VF-WF	-8.82	1.98	0.00	Reject	$-(t\text{-critical}) > t\text{-value}$	significant
	BCF-WF	-7.31	1.98	0.00	Reject	$-(t\text{-critical}) > t\text{-value}$	significant
Bi	VF-BCF	-5.66	1.98	0.00	Reject	$-(t\text{-critical}) > t\text{-value}$	significant
	VF-WF	-5.56	1.98	0.00	Reject	$-(t\text{-critical}) > t\text{-value}$	significant
	BCF-WF	-2.33	1.98	0.02	Reject	$-(t\text{-critical}) > t\text{-value}$	significant

*Reject the hypotheses of equal means if $-(t\text{-critical}) > t\text{-value} > t\text{-critical}$ and $p < 0.10$ (α)

Table 5 Black shale matrix analyses of Ti, Zr, Cr, Th in ppm and their ratios

Formation	Depth in meters	Ti	Cr	Zr	Th	Ti/Zr	Th/Cr	Data source
Velkerri Formation	115.7	254.43	1.00	67.87	5.66	3.75	5.66	This study
Velkerri Formation	126.3	1156.76	55.71	74.27	7.72	15.57	0.14	This study
Velkerri Formation	140.1	8746.43	95.92	186.07	36.59	47.01	0.38	This study
Velkerri Formation	151.2	1609.06	88.71	109.70	3.01	14.67	0.03	This study
Velkerri Formation	151.2	1407.51	92.83	44.21	1.52	31.83	0.02	This study
Velkerri Formation	151.2	111.99	17.36	7.18	5.20	15.60	0.30	This study
Velkerri Formation	151.2	1042.86	66.30	53.70	3.25	19.42	0.05	This study
Velkerri Formation	161.8	2705.62	45.95	202.31	12.76	13.37	0.28	This study
Velkerri Formation	172.4	713.61	1879.43	72.14	15.72	9.89	0.01	This study
Velkerri Formation	183.4	890.25	1.00	43.49	3.59	20.47	3.59	This study
Velkerri Formation	193.9	4580.01	419.61	194.32	8.65	23.57	0.02	This study
Velkerri Formation	216.9	3947.04	91.32	73.88	7.96	53.43	0.09	This study
					Mean	22.38	0.88	
Barney Creek Formation	47.50	4587.66	5.60	206.66	61.56	22.20	10.99	This study
Barney Creek Formation	54.20	2149.17	57.77	229.04	16.80	9.38	0.29	This study
Barney Creek Formation	60.80	4409.26	42.65	93.34	10.84	47.24	0.25	This study
Barney Creek Formation	68.50	1038.77	20.66	38.61	6.96	26.90	0.34	This study
Barney Creek Formation	70.20	77793.00	81.04	424.32	24.97	183.33	0.31	This study
Barney Creek Formation	76.10	19067.05	129.94	649.72	72.99	29.35	0.56	This study
Barney Creek Formation	81.90	3001.76	64.46	139.79	14.89	21.47	0.23	This study
Barney Creek Formation	86.80	4728.15	62.08	107.23	26.64	44.09	0.43	This study
Barney Creek Formation	93.00	3864.95	87.86	150.62	18.39	25.66	0.21	This study
Barney Creek Formation	98.30	4351.58	58.97	269.40	29.19	16.15	0.50	This study
Barney Creek Formation	100.70	3430.80	70.19	133.39	19.57	25.72	0.28	This study
Barney Creek Formation	109.80	2294.94	44.83	100.62	20.59	22.81	0.46	This study
Barney Creek Formation	124.50	1106.57	24.27	30.34	4.97	36.47	0.20	This study
Barney Creek Formation	128.40	1751.79	80.22	160.18	26.67	10.94	0.33	This study
Barney Creek Formation	136.50	20401.09	90.28	293.52	32.43	69.50	0.36	This study
Barney Creek Formation	148.80	4484.17	213.02	698.77	196.19	6.42	0.92	This study
					Mean	37.35	1.04	
Wollogorang Formation	67.8	7108.46	71.76	340.36	28.81	20.88	0.40	This study
Wollogorang Formation	73.8	818.41	42.32	49.26	7.32	16.61	0.17	This study
Wollogorang Formation	74.3	527.13	30.96	61.57	3.01	8.56	0.10	This study
Wollogorang Formation	77	468.06	1.00	7.85	1.15	59.64	1.15	This study
Wollogorang Formation	85.4	660.50	27.28	38.30	6.41	17.25	0.23	This study
					Mean	24.59	0.41	

Table 6a SHRIMPI-SI pyrite analyses for all three black shale formations in ‰ (Data source: this study)

SPOT #	$\delta^{33}\text{S}_{\text{VCDT}}$	\pm Internal error (2 σ)	$\delta^{34}\text{S}_{\text{VCDT}}$	\pm Internal error (2 σ)	$\Delta^{33}\text{S}$	\pm Internal error (2 σ)	^{32}S cps (median)	^{33}S cps (median)	^{34}S cps (median)	Sets, scans
Velkerri Formation										
172.4-08-2.1	16.03	0.09	31.74	0.03	-0.20	0.10	9.5E+08	7.5E+06	4.3E+07	2, 10
172.4-08-1.1	8.60	0.07	17.45	0.04	-0.35	0.08	5.5E+08	4.4E+06	2.5E+07	2, 10
172.4-08-3.1	18.01	0.08	35.55	0.03	-0.14	0.08	1.1E+09	8.8E+06	5.0E+07	2, 10
161.8-03-1	7.92	0.15	15.74	0.04	-0.15	0.15	8.7E+08	6.8E+06	3.8E+07	2, 10
161.8-03-2	9.10	0.09	17.77	0.03	-0.01	0.09	1.0E+09	8.0E+06	4.5E+07	2, 10
161.8-03-5	7.95	0.13	16.21	0.03	-0.36	0.13	9.0E+08	7.1E+06	4.0E+07	2, 10
161.8-03-6	8.35	0.10	16.31	0.03	-0.01	0.11	9.1E+08	7.2E+06	4.0E+07	2, 10
144.3-01	12.49	0.10	24.43	0.03	-0.02	0.11	9.5E+08	7.5E+06	4.2E+07	2, 10
144.3-02	15.07	0.08	29.46	0.03	0.01	0.09	9.5E+08	7.6E+06	4.3E+07	2, 10
115.2_1	23.41	0.24	45.58	0.04	0.19	0.26	5.4E+08	4.3E+06	2.5E+07	1, 10
115.2_2	24.07	0.30	45.77	0.04	0.76	0.33	5.5E+08	4.4E+06	2.5E+07	1, 10
115.2_3	23.46	0.25	45.78	0.04	0.14	0.29	5.5E+08	4.4E+06	2.5E+07	1, 10
115.2_4	15.40	0.32	30.34	0.05	-0.11	0.33	5.6E+08	4.5E+06	2.5E+07	1, 10
115.2_5	22.25	0.16	43.95	0.04	-0.15	0.17	5.6E+08	4.5E+06	2.6E+07	1, 10
115.2_6	23.29	0.33	46.05	0.03	-0.16	0.34	5.4E+08	4.4E+06	2.5E+07	1, 10
140.1_1	11.70	0.21	22.40	0.04	0.22	0.22	5.9E+08	4.7E+06	2.6E+07	1, 10
140.1_2	11.35	0.20	22.07	0.04	0.04	0.21	5.8E+08	4.6E+06	2.6E+07	1, 10
140.1_3	11.07	0.41	22.16	0.04	-0.28	0.41	5.9E+08	4.7E+06	2.6E+07	1, 10
140.1_4	11.04	0.15	21.98	0.04	-0.22	0.16	5.9E+08	4.7E+06	2.6E+07	1, 10
140.1_5	11.46	0.22	22.50	0.04	-0.07	0.23	5.5E+08	4.4E+06	2.5E+07	1, 10
140.1_6	11.78	0.29	23.00	0.04	0.00	0.32	5.9E+08	4.7E+06	2.6E+07	1, 10
193.9-01	12.95	0.14	24.90	0.03	0.21	0.14	9.1E+08	7.3E+06	4.1E+07	2, 10
193.9-02	10.01	0.11	19.15	0.03	0.19	0.11	9.7E+08	7.7E+06	4.3E+07	2, 10
183.4-01	21.73	0.06	42.30	0.03	0.17	0.06	1.1E+09	8.7E+06	5.0E+07	2, 10
183.4-02	19.49	0.13	37.72	0.03	0.24	0.14	1.1E+09	8.8E+06	5.0E+07	2, 10
183.4-03	19.78	0.08	38.50	0.03	0.14	0.08	1.0E+09	8.0E+06	4.5E+07	2, 10
183.4-04	20.03	0.10	39.18	0.03	0.05	0.10	1.1E+09	8.7E+06	5.0E+07	2, 10
183.4-05	23.00	0.16	44.94	0.03	0.11	0.17	9.9E+08	7.9E+06	4.5E+07	2, 10
183.4-06	19.55	0.09	37.93	0.03	0.19	0.10	1.0E+09	8.2E+06	4.7E+07	2, 10
183.4-07	18.19	0.19	35.37	0.03	0.13	0.20	7.0E+08	5.6E+06	3.2E+07	2, 10
126.3-01	-3.06	0.14	-5.92	0.03	0.00	0.14	1.0E+09	8.1E+06	4.5E+07	2, 10
126.3-02	12.08	0.07	23.46	0.03	0.07	0.07	1.2E+09	9.2E+06	5.2E+07	2, 10
126.3-03	12.79	0.11	25.08	0.02	-0.05	0.12	1.2E+09	9.2E+06	5.2E+07	2, 10
126.3-04	2.62	0.10	4.95	0.02	0.07	0.10	9.7E+08	7.7E+06	4.3E+07	2, 10
126.3-05	-4.33	0.12	-8.23	0.03	-0.08	0.13	9.7E+08	7.6E+06	4.2E+07	2, 10
126.3-06	12.50	0.11	24.35	0.03	0.03	0.11	9.2E+08	7.3E+06	4.1E+07	2, 10
126.3-07	-3.26	0.14	-6.64	0.03	0.17	0.16	9.8E+08	7.7E+06	4.3E+07	2, 10
126.3-07	7.87	0.08	15.26	0.03	0.04	0.09	9.7E+08	7.6E+06	4.3E+07	2, 10
374.4-01	22.50	0.08	43.92	0.02	0.11	0.09	9.3E+08	7.4E+06	4.2E+07	2, 10
374.4-2	24.35	0.11	47.75	0.03	0.04	0.12	9.7E+08	7.8E+06	4.4E+07	2, 10
374.4-3	33.29	0.12	65.59	0.02	0.03	0.13	9.9E+08	8.1E+06	4.6E+07	2, 10
374.4-4	24.91	0.10	48.56	0.03	0.19	0.11	1.1E+09	8.6E+06	4.9E+07	2, 10
374.4-5	22.75	0.13	44.26	0.03	0.19	0.15	9.9E+08	8.0E+06	4.5E+07	2, 10
374.4-6	22.72	0.12	44.10	0.03	0.24	0.13	1.0E+09	8.2E+06	4.7E+07	2, 10
374.4-7	24.48	0.12	48.02	0.03	0.03	0.13	9.8E+08	7.9E+06	4.5E+07	2, 10
374.4-8	24.79	0.06	48.37	0.03	0.16	0.06	1.0E+09	8.3E+06	4.7E+07	2, 10
374.4-9	-7.36	0.07	-14.19	0.03	-0.02	0.07	1.1E+09	8.5E+06	4.7E+07	2, 10
374.4-10	26.93	0.09	52.79	0.03	0.08	0.10	1.0E+09	8.1E+06	4.6E+07	2, 10
183.4-01	12.28	0.09	24.14	0.03	-0.08	0.09	1.1E+09	8.4E+06	4.8E+07	2, 10
183.4-02	11.35	0.10	22.32	0.02	-0.08	0.11	1.4E+09	1.1E+07	6.4E+07	2, 10
151.2-01	11.83	0.24	22.89	0.03	0.11	0.25	5.7E+08	4.5E+06	2.5E+07	2, 10
151.2-02	13.54	0.08	26.68	0.03	-0.12	0.07	1.0E+09	8.0E+06	4.5E+07	2, 10
151.2-03	11.11	0.11	21.72	0.03	-0.01	0.12	7.2E+08	5.7E+06	3.2E+07	2, 10
151.2-04	12.15	0.10	23.80	0.03	-0.04	0.10	7.7E+08	6.1E+06	3.5E+07	2, 10
151.2-05	13.55	0.19	26.35	0.03	0.06	0.19	7.1E+08	5.7E+06	3.2E+07	2, 10
328.1-01	2.05	0.09	4.03	0.03	-0.03	0.09	9.3E+08	7.3E+06	4.1E+07	2, 10
328.1-02	3.57	0.13	7.29	0.02	-0.18	0.13	1.1E+09	8.7E+06	4.9E+07	2, 10
328.1-03	3.18	0.07	6.54	0.04	-0.19	0.07	8.1E+08	6.4E+06	3.6E+07	2, 10
328.1-04	2.02	0.11	4.03	0.04	-0.05	0.11	7.7E+08	6.0E+06	3.4E+07	2, 10
316.8_3	7.23	0.16	14.81	0.03	-0.37	0.16	6.3E+08	4.9E+06	2.8E+07	1, 10
316.8_1	7.58	0.15	14.98	0.04	-0.10	0.17	6.4E+08	5.1E+06	2.8E+07	1, 10
316.8_2	14.52	0.15	28.65	0.04	-0.14	0.15	5.0E+08	4.0E+06	2.3E+07	1, 10
			Mean	26.87		0.01				
			Std. Dev	16.48		0.18				
			Std. Error	2.09		0.02				
			Max	65.59		0.76				
			Min	-14.19		-0.37				

Table 6c SHRIMPI-SI pyrite analyses for all three black shale formations in ‰ (Data source: this study)

SPOT #	$\delta^{33}\text{S}_{\text{VCDT}}$	\pm Internal error (2σ)	$\delta^{34}\text{S}_{\text{VCDT}}$	\pm Internal error (2σ)	$\Delta^{33}\text{S}$	\pm Internal error (2σ)	^{32}S cps (median)	^{33}S cps (median)	^{34}S cps (median)	Sets, scans
Wologorang Formation										
MTY_67.8_1	3.84	0.10	2.47	0.04	2.57	0.09	5.8E+08	4.6E+06	2.6E+07	1, 10
MTY_70.4_1	2.10	0.10	4.19	0.04	-0.05	0.10	5.7E+08	4.5E+06	2.5E+07	1, 10
MTY_70.4_B2	-1.62	0.24	-3.32	0.04	0.09	0.25	5.4E+08	4.2E+06	2.3E+07	1, 10
MTY_70.4_B3	-1.38	0.24	-2.21	0.03	-0.24	0.26	5.7E+08	4.4E+06	2.5E+07	1, 10
MTY_70.4_B4	-0.16	0.27	-1.10	0.04	0.41	0.28	5.4E+08	4.2E+06	2.4E+07	1, 10
MTY_70.4_B5	-0.61	0.14	-1.04	0.04	-0.08	0.15	5.7E+08	4.4E+06	2.5E+07	1, 10
MTY_70.4_B6	-2.47	0.25	-4.63	0.04	-0.08	0.29	5.6E+08	4.4E+06	2.4E+07	1, 10
MTY_77_1	-13.79	0.18	-27.92	0.04	0.69	0.18	5.3E+08	4.1E+06	2.3E+07	1, 10
MTY_77_10	-12.65	0.10	-25.57	0.05	0.60	0.10	4.3E+08	3.3E+06	1.8E+07	1, 10
MTY_77_11	1.45	0.10	1.24	0.04	0.81	0.11	4.7E+08	3.7E+06	2.1E+07	1, 10
MTY_77_12	-13.98	0.26	-28.24	0.04	0.67	0.28	5.8E+08	4.5E+06	2.5E+07	1, 10
MTY_77_2	-12.70	0.15	-25.59	0.04	0.56	0.15	5.8E+08	4.5E+06	2.5E+07	1, 10
MTY_77_3	-8.27	0.22	-16.63	0.04	0.33	0.23	5.5E+08	4.3E+06	2.4E+07	1, 10
MTY_77_4	2.99	0.10	5.26	0.04	0.28	0.10	5.5E+08	4.3E+06	2.4E+07	1, 10
MTY_77_5	2.16	0.15	3.85	0.04	0.18	0.16	5.1E+08	4.0E+06	2.2E+07	1, 10
MTY_77_6	-13.74	0.13	-27.07	0.04	0.29	0.14	5.1E+08	3.9E+06	2.2E+07	1, 10
MTY_77_7	-9.30	0.16	-19.25	0.04	0.67	0.17	5.2E+08	4.0E+06	2.2E+07	1, 10
MTY_77_8	-14.70	0.22	-29.44	0.05	0.57	0.23	4.8E+08	3.7E+06	2.1E+07	1, 10
MTY_77_9	-13.15	0.22	-26.78	0.04	0.74	0.24	4.7E+08	3.6E+06	2.0E+07	1, 10
MTY_89_B1	1.89	0.30	3.50	0.03	0.09	0.31	8.4E+08	6.6E+06	3.7E+07	1, 10
MTY_89_B2	3.44	0.19	5.66	0.03	0.53	0.19	8.9E+08	7.0E+06	3.9E+07	1, 10
MTY_89_B3	2.57	0.34	4.45	0.02	0.27	0.37	1.0E+09	8.1E+06	4.5E+07	1, 10
MTY_89_B4	3.22	0.29	5.25	0.03	0.52	0.33	9.7E+08	7.6E+06	4.2E+07	1, 10
MTY_89_B5	2.89	0.24	5.25	0.02	0.19	0.29	1.0E+09	7.9E+06	4.4E+07	1, 10
MTY_89_B6	1.23	0.12	2.69	0.03	-0.16	0.11	8.3E+08	6.5E+06	3.6E+07	1, 10
MT_YOUNG_4	4.82	0.35	9.49	0.03	-0.02	0.32	8.9E+08	7.0E+06	4.0E+07	1, 10
MT_YOUNG_1	3.73	0.29	6.65	0.02	0.34	0.23	9.3E+08	7.3E+06	4.1E+07	1, 10
MT_YOUNG_5	3.16	0.28	5.63	0.03	0.30	0.24	9.2E+08	7.2E+06	4.1E+07	1, 10
MT_YOUNG_6	3.62	0.16	7.10	0.02	0.01	0.09	9.4E+08	7.4E+06	4.1E+07	1, 10
MT_YOUNG_3	2.37	0.33	5.25	0.03	-0.30	0.27	8.9E+08	7.0E+06	3.9E+07	1, 10
MT_YOUNG_2	2.79	0.21	4.90	0.04	0.30	0.13	9.2E+08	7.2E+06	4.1E+07	1, 10
		Mean	-5.03		0.36					
		Std. Dev	13.62		0.51					
		Std. Error	2.45		0.09					
		Max	9.49		2.57					
		Min	-29.44		-0.30					

Table 6b SHRIMPI-SI pyrite analyses for all three black shale formations in ‰ (Data source: this study)

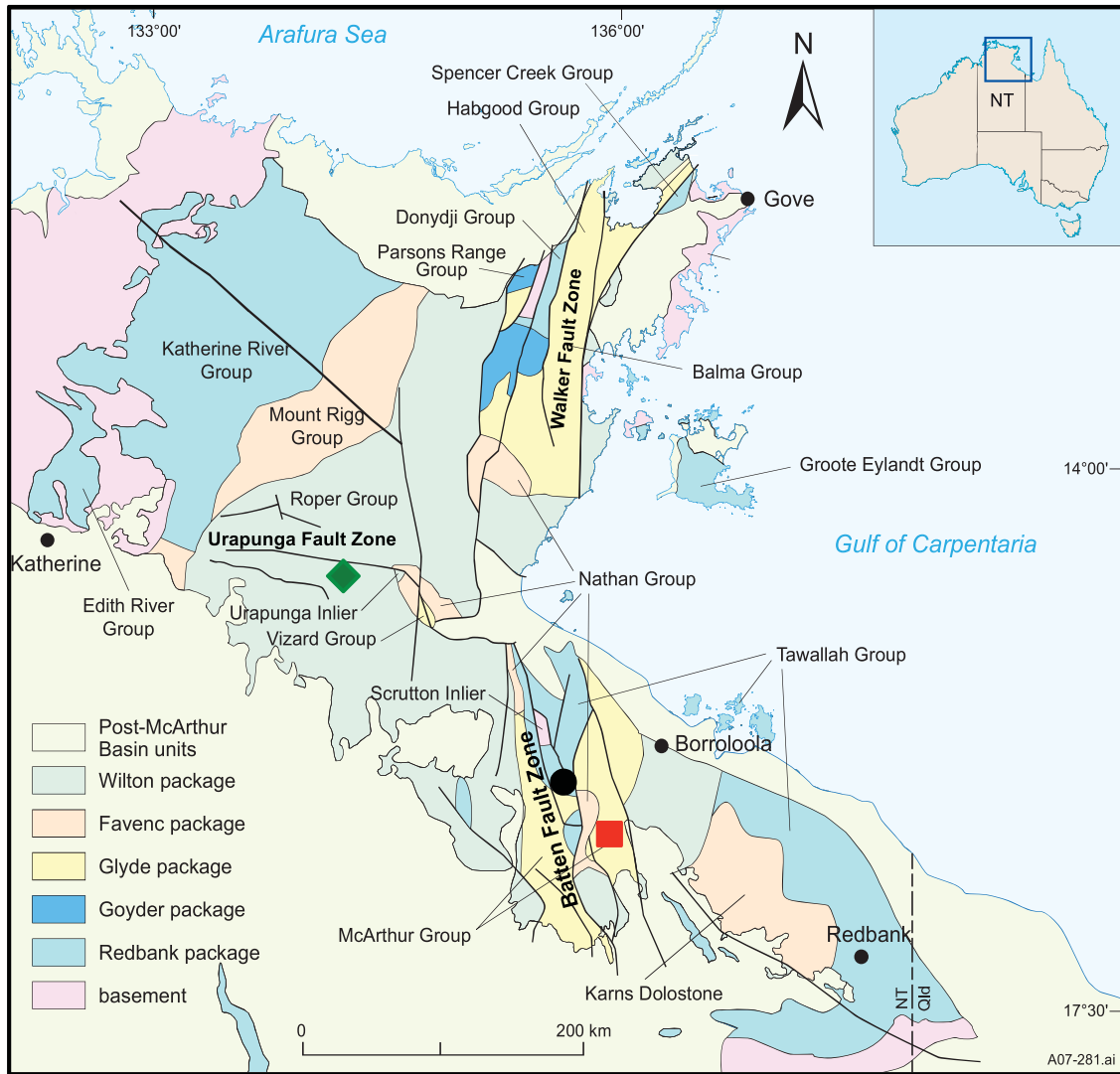
SPOT #	$\delta^{33}\text{S}_{\text{VCDT}}$	\pm Internal error (2σ)	$\delta^{34}\text{S}_{\text{VCDT}}$	\pm Internal error (2σ)	$\Delta^{33}\text{S}$	\pm Internal error (2σ)	^{32}S cps (median)	^{33}S cps (median)	^{34}S cps (median)	Sets, scans
Barney Creek Formation										
MBX_47.50_B1	-2.13	0.15	-4.76	0.05	0.33	0.15	6.2E+08	4.8E+06	2.7E+07	1, 10
MBX_47.50_B2	-3.30	0.17	-6.83	0.04	0.23	0.18	7.0E+08	5.5E+06	3.0E+07	1, 10
MBX_47.50_B3	-2.82	0.31	-5.28	0.04	-0.09	0.33	6.1E+08	4.8E+06	2.7E+07	1, 10
MBX_47.50_B4	4.39	0.32	8.83	0.04	-0.15	0.34	5.8E+08	4.5E+06	2.5E+07	1, 10
MBX_47.50_B5	7.41	0.11	14.71	0.04	-0.14	0.11	5.7E+08	4.5E+06	2.5E+07	1, 10
MBX_47.50_B6	-7.17	0.32	-13.79	0.05	-0.05	0.34	5.7E+08	4.5E+06	2.5E+07	1, 10
MBX124.5_1	17.92	0.25	34.68	0.04	0.21	0.25	4.6E+08	3.6E+06	2.1E+07	1, 10
MBX22.7_2	4.18	0.55	7.79	0.12	0.17	0.55	1.1E+08	8.7E+05	4.9E+06	1, 10
MBX22.7_4	5.93	0.19	12.01	0.04	-0.23	0.20	4.6E+08	3.6E+06	2.0E+07	1, 10
MBX30.10_7B	13.76	0.18	26.50	0.03	0.20	0.20	6.0E+08	4.8E+06	2.7E+07	1, 10
MBX30.10_1B	8.97	0.17	17.11	0.03	0.19	0.19	6.2E+08	4.9E+06	2.7E+07	1, 10
MBX30.10_2B	9.76	0.31	18.30	0.03	0.38	0.34	6.2E+08	4.9E+06	2.8E+07	1, 10
MBX30.10_3B	10.06	0.17	19.84	0.03	-0.11	0.18	6.5E+08	5.2E+06	2.9E+07	1, 10
MBX30.10_4B	-0.10	0.12	0.17	0.04	-0.19	0.14	6.2E+08	4.9E+06	2.7E+07	1, 10
MBX30.10_5B	-6.72	0.45	-12.88	0.04	-0.06	0.49	6.4E+08	5.0E+06	2.8E+07	1, 10
MBX30.10_6B	-4.80	0.29	-9.35	0.03	0.03	0.31	6.3E+08	4.9E+06	2.7E+07	1, 10
MBX-76.10_1	17.73	0.15	33.55	0.06	0.59	0.17	3.8E+08	3.0E+06	1.7E+07	1, 6
MBX-76.10_2	17.87	0.31	34.59	0.08	0.20	0.36	3.5E+08	2.8E+06	1.6E+07	1, 6
MBX-76.10_3	19.17	0.44	36.35	0.07	0.61	0.53	3.6E+08	2.9E+06	1.6E+07	1, 6
MBX-76.10_4	18.47	0.21	35.76	0.07	0.21	0.20	3.7E+08	2.9E+06	1.7E+07	1, 6
MBX-76.10_5	18.08	0.05	35.45	0.07	-0.02	0.06	3.6E+08	2.9E+06	1.6E+07	1, 6
MBX-76.10_6	16.69	0.12	32.33	0.06	0.17	0.12	3.7E+08	2.9E+06	1.7E+07	1, 6
MBX-76.10_7	17.33	0.41	33.69	0.08	0.11	0.47	3.6E+08	2.9E+06	1.6E+07	1, 6
MBX-76.10_8	16.47	0.06	32.12	0.08	0.06	0.06	3.7E+08	2.9E+06	1.7E+07	1, 6
MBX-76.10_9	17.39	0.14	33.06	0.08	0.50	0.14	3.6E+08	2.9E+06	1.6E+07	1, 6
MBX98.30_1	19.43	0.36	37.21	0.08	0.44	0.41	2.5E+08	2.0E+06	1.1E+07	1, 6
MBX98.30_2	11.69	0.37	22.50	0.10	0.17	0.38	2.7E+08	2.1E+06	1.2E+07	1, 6
MBX98.30_2B	10.52	0.21	21.85	0.07	-0.67	0.24	2.7E+08	2.1E+06	1.2E+07	1, 6
MBX98.30_4	20.24	0.29	39.89	0.06	-0.10	0.36	2.9E+08	2.3E+06	1.3E+07	1, 6
MBX98.30_5	19.94	0.35	39.23	0.10	-0.07	0.43	2.0E+08	1.6E+06	9.1E+06	1, 6
MBX98.30_6	20.99	0.98	41.80	0.13	-0.32	1.15	1.6E+08	1.2E+06	7.1E+06	1, 6
		Mean	19.88		0.08					
		Std. Dev	17.74		0.28					
		Std. Error	3.19		0.05					
		Max	41.80		0.61					
		Min	-13.79		-0.67					



Table 7 General characteristics of some chemical elements in simple surface or near-surface aqueous systems (modified after Smith and Huyck, 1999); n

ELEMENT	IONIC FORM	Species		REDOX-SENSITIVE; References	Type
		Oxic	Reduced		
Arsenic (As)	anionic	$H_2AsO_4^{4-}$ or $HAsO_4^{2-}$	$H_3AsO_3^0$	(Smith and Huyck, 1999; Hinkle and Polette, 1999; Hem et al., 1985)	n
Molybdenum (Mo)	anionic; cationic species exist but are rare in aqueous systems.	MoO_4^{2-} , MoO_2^{+} ;	$MoO_xS_{4-x}^{2-}$ MoS_4^{2-}	(Smith and Huyck, 1999; Algeo and Maynard, 2006)	c, n
Selenium (Se)	anionic	SeO_4^{2-} , SeO_3^{2-}	Se^0	(Smith and Huyck, 1999; Morita et al., 2007)	n
Nickel (Ni)	cationic	Ni^{2+} , Ni^{1+} , $3+$, $4+$, $NiCl^+$	NiS	(Khan et al., 2014; Young 1995; Clayton and Clayton 1994; Coogan et al. 1989; Barciela-Alonso et al., 2003)	n
Zinc (Zn)	cationic	Zn^{2+} , $ZnCl^+$	ZnS	(Algeo and Maynard, 2006)	n
Bismuth (Bi)	cationic	Bi^{3+} , BiO^+ BiO_2^{2+} , $BiOH^2$	Bi_2S_3	(Brookins., 1988)	s
Cobalt (Co)	cationic	Co^+	CoS	(Algeo and Maynard, 2006)	s
Copper (Cu)	cationic	$CuCl^+$, Cu^{2+} , Cu^{1+}	Cu_2S , CuS	(Algeo and Maynard, 2006; Smith and Huyck, 1999)	n, s
Lead (Pb)	cationic	Pb^+ , Pb^{2+} , $PbCl^+$	PbS	(redox sensitive only under extreme conditions; Smith and Huyck, 1999)	s





Group	Formation	Age (Ga)	Drill hole	Symbol
Roper	Velkerri	~1.4	Urap-4	◆
McArthur	Barney Creek	~1.64	MBXDD001	■
Tawallah	Wollogorang	~1.73	Mt Young	●

Fig 1 Drill hole locations for the three black shale formations (Figure adapted from Ahmad et al., 2013)

SOUTHEASTERN McARTHUR BASIN

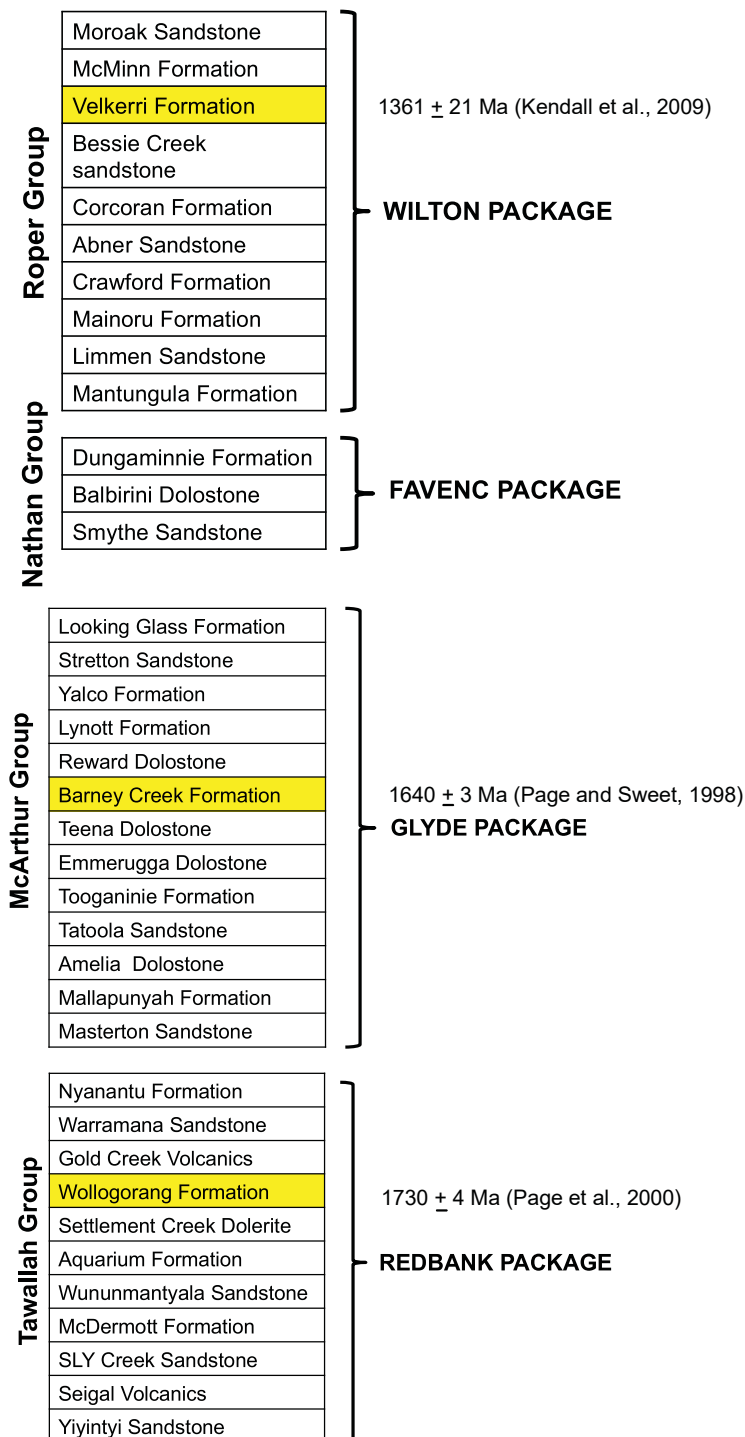


Fig 2 Generalized stratigraphy of the southern McArthur Basin, Australia adapted from Rawlings (1999); highlighted in yellow are the three black shale formations examined in this study.

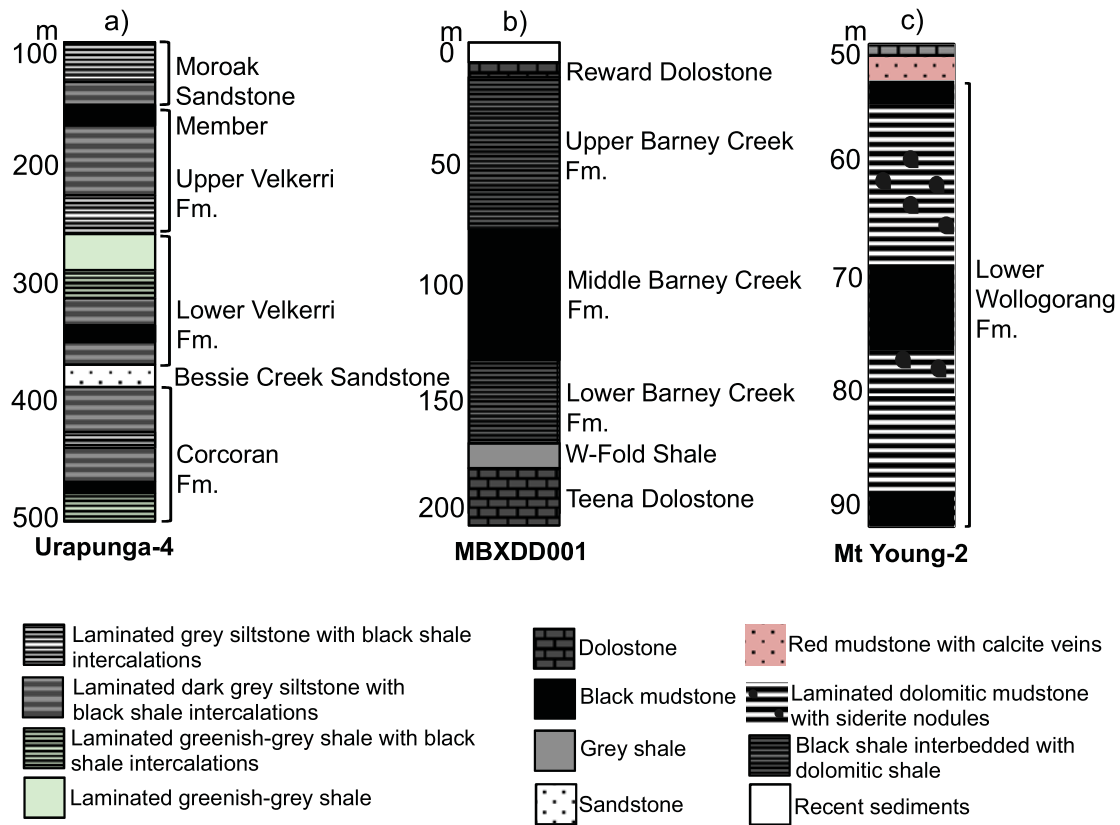
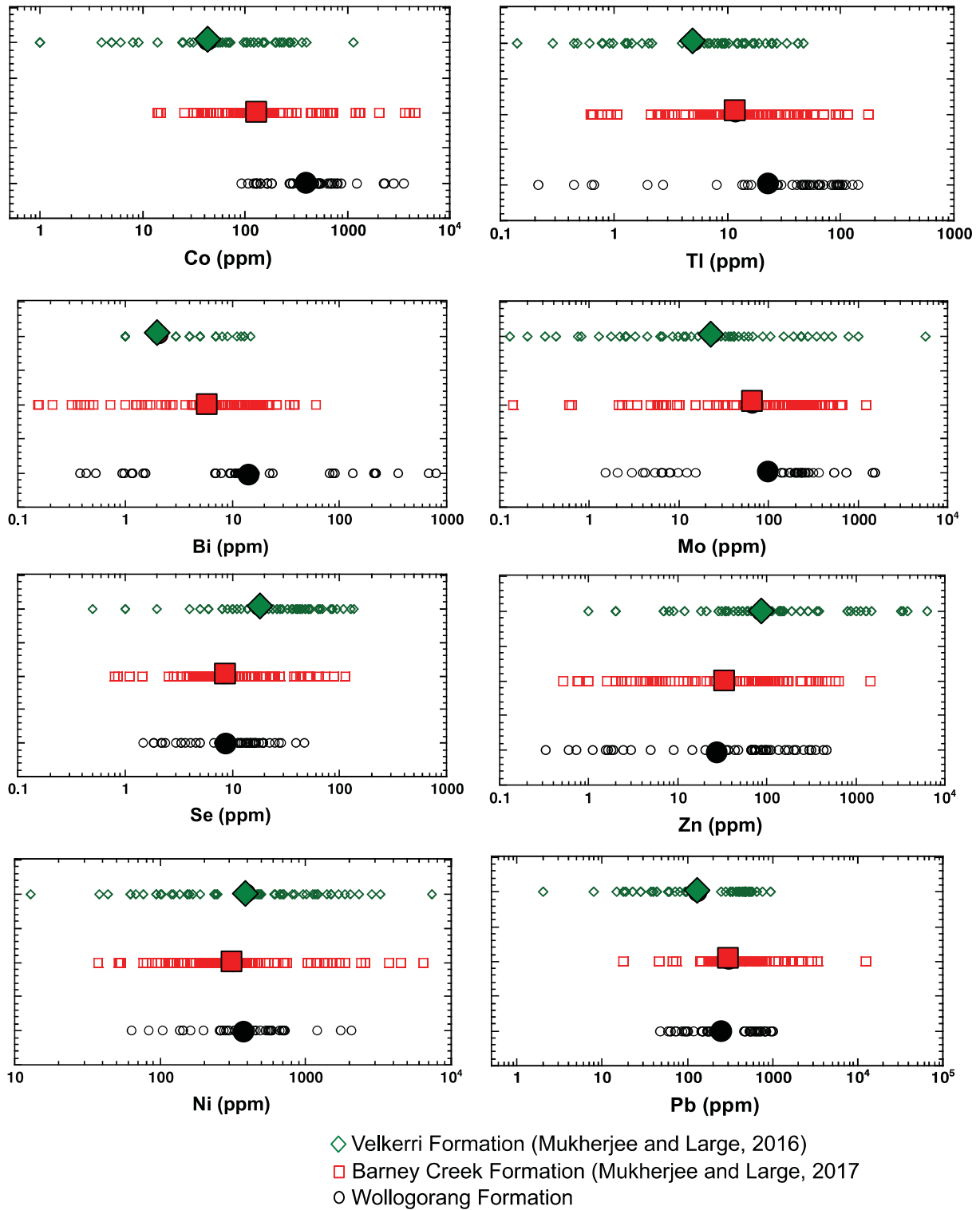


Fig 3a, b, c Simplified logs of the three drill holes used for the study.

a)



b)

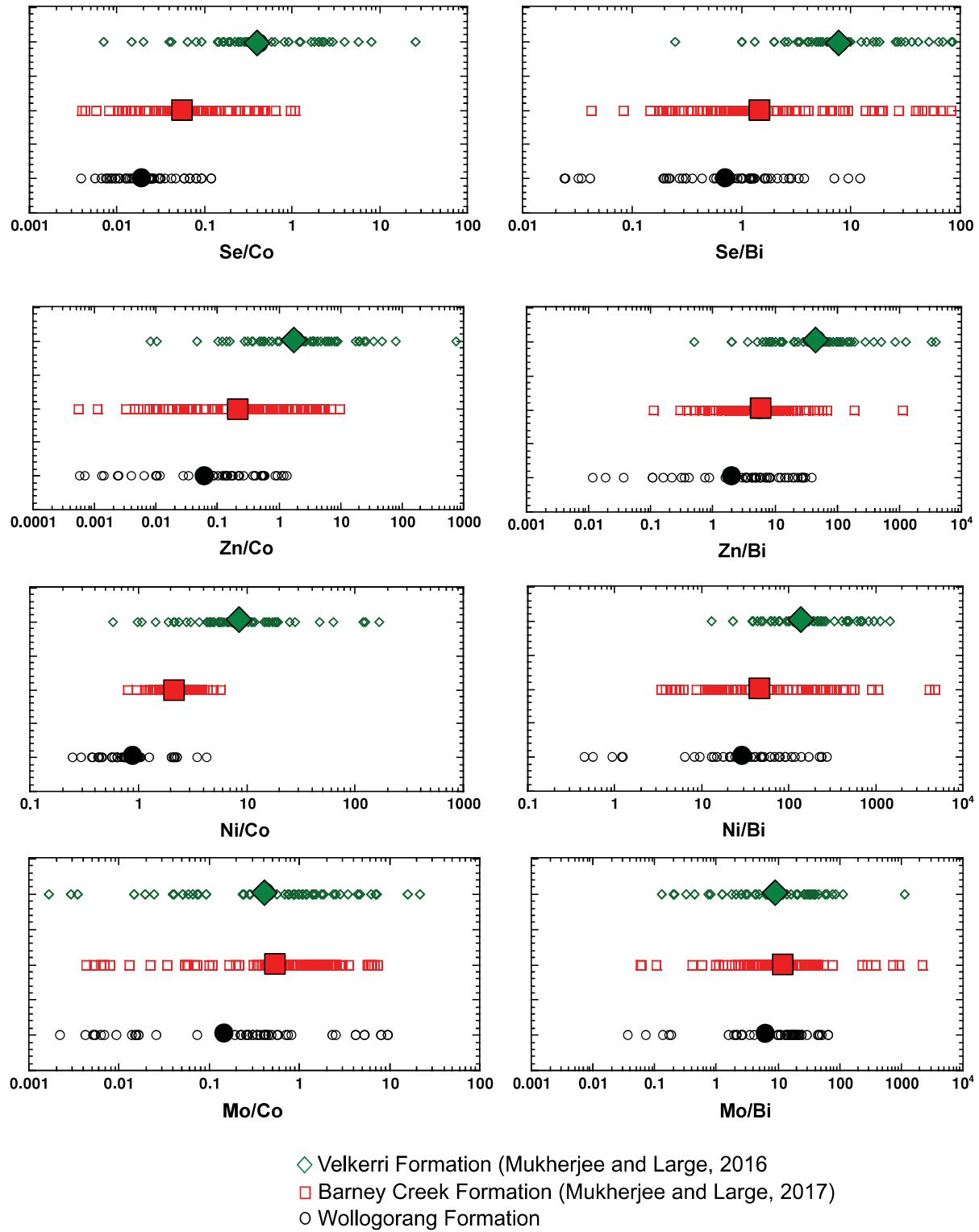


Fig 4a, b Comparison of trace elements in pyrite and their ratios in the three black shale formations

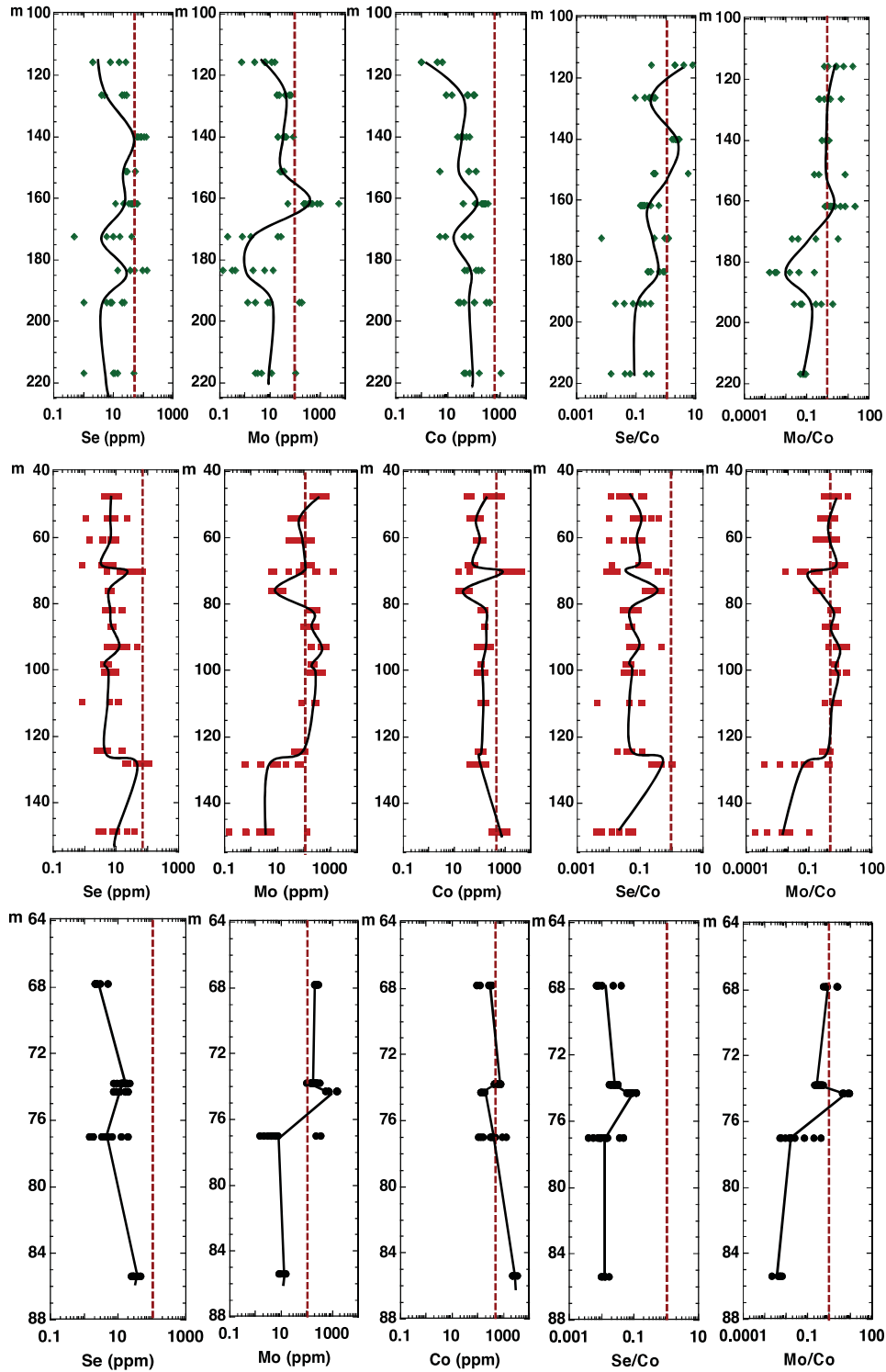
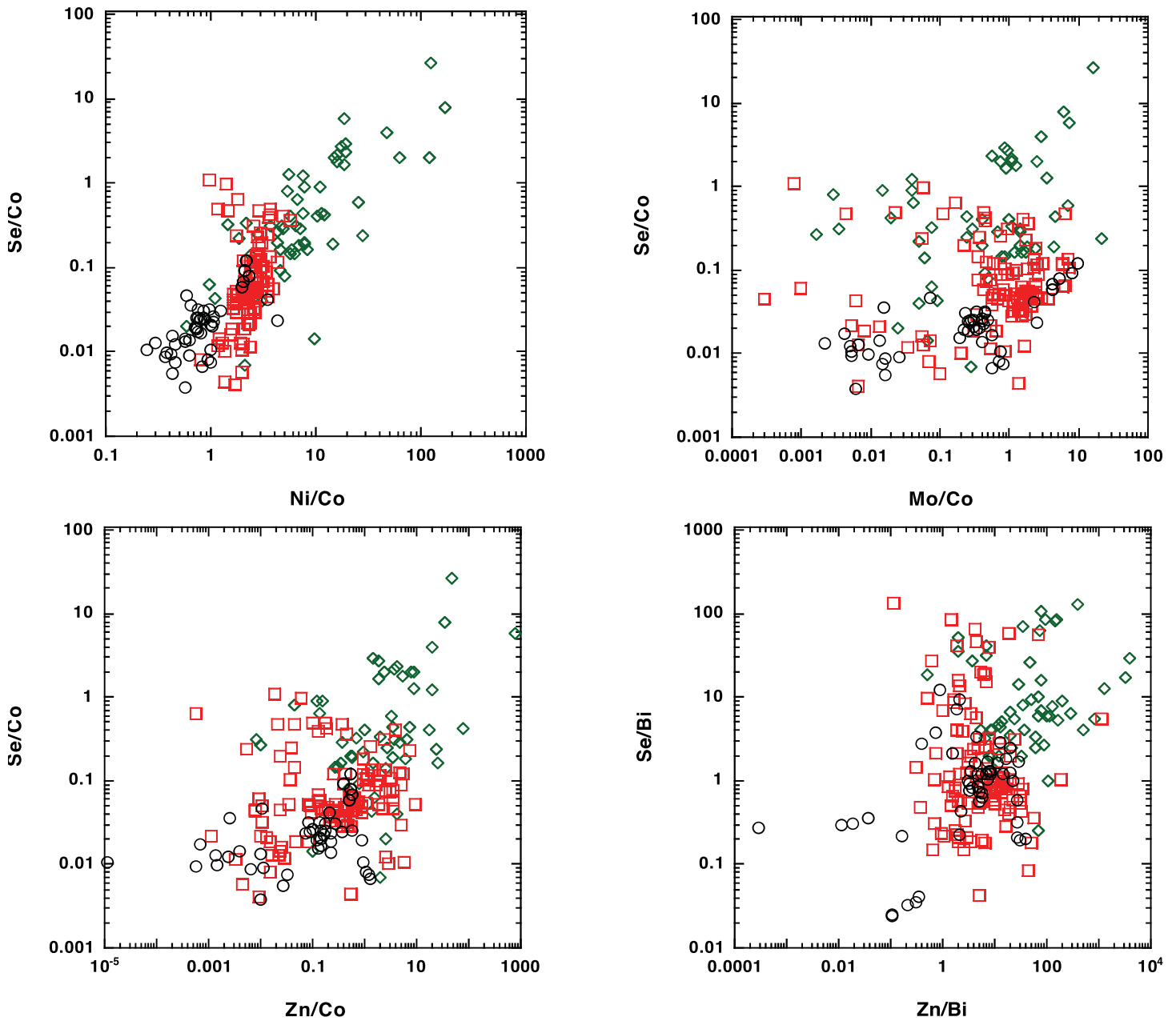


Fig 5 Downhole variation of TEs (Se, Mo, Co (in ppm) & Se/Co, Mo/Co) for the Velkerri, (top), Barney Creek (middle) and Wollogorang (bottom) formations; Red dashed line represents global mean value of ~5000 pyrites for Se, Mo, Co (from Large et al., 2014); and unity for Se/Co & Mo/Co (Data source for Velkerri Formation: Mukherjee and Large, 2016 and Barney Creek Formation: Mukherjee and Large, 2017)



- ◇ Velkerri Formation (Mukherjee and Large, 2016)
- Barney Creek Formation (Mukherjee and Large, 2017)
- Wollgorang Formation

Fig 6 Plot of Se/Co: Ni/Co; Se/Co: Mo/Co; Se/Co: Zn/Co; Se/Bi; Zn/Bi for the three black shale formations (black circles: Wollgorang; red squares: Barney Creek; green diamonds: Velkerri)

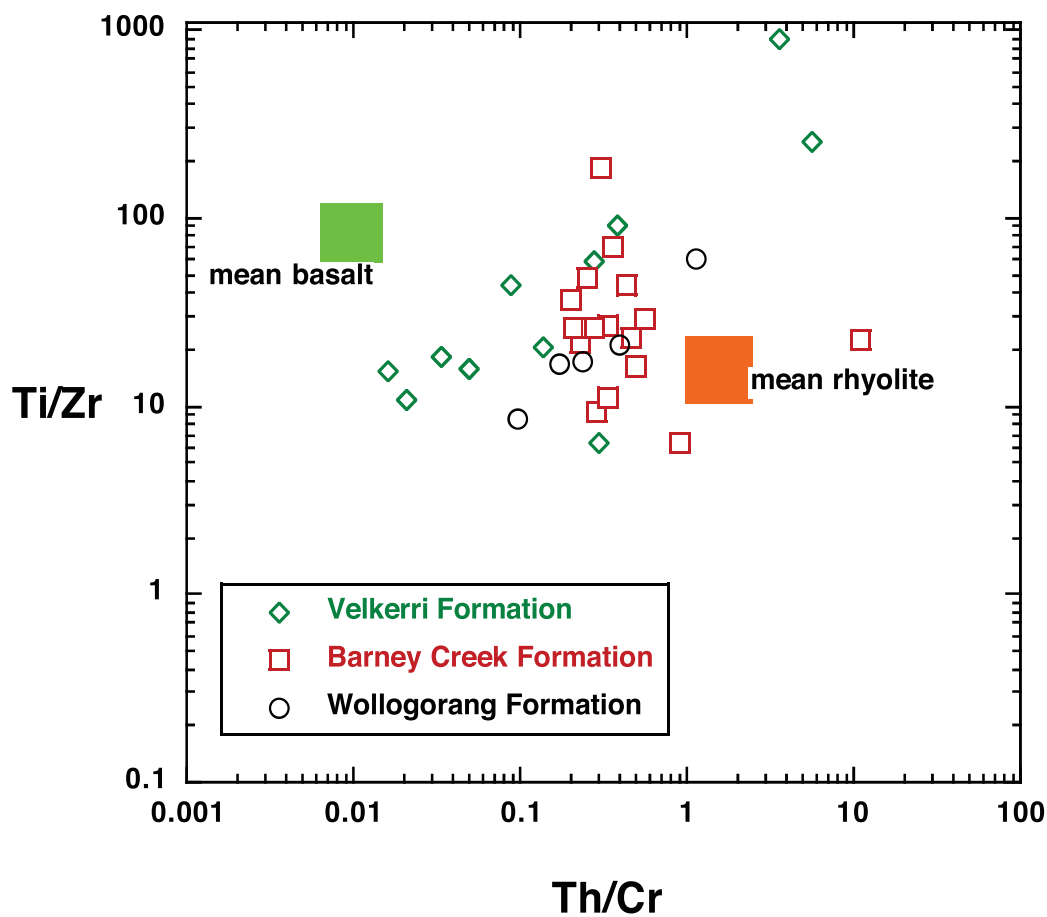


Fig 7 Plot of Th/Cr and Ti/Zr ratios for silicate matrix compositions of the black shales to identify source rock composition (felsic or mafic); Ti/Zr and Th/Cr ratios for basalt and rhyolite from Riemann and De Caritat (1998).

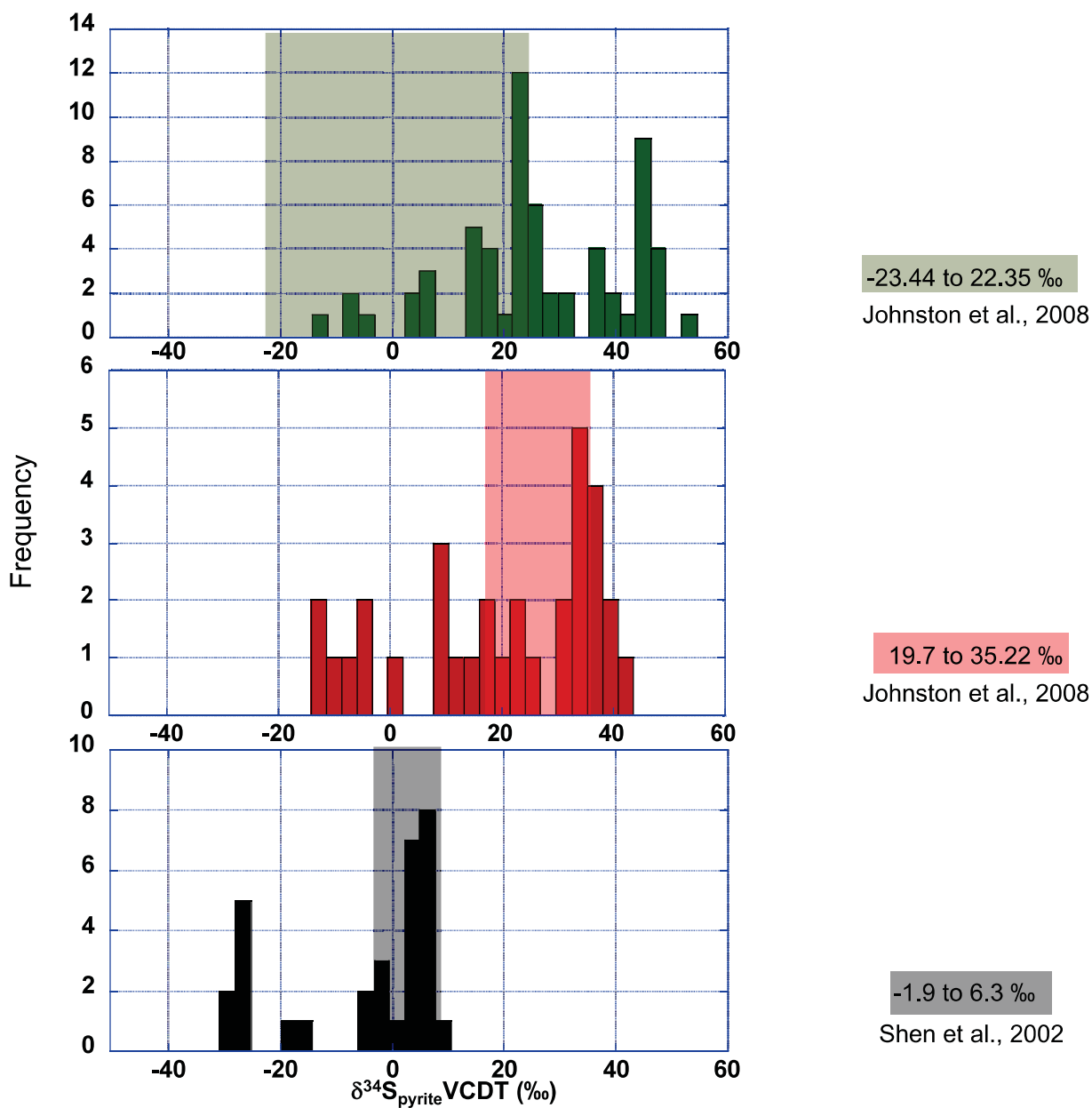
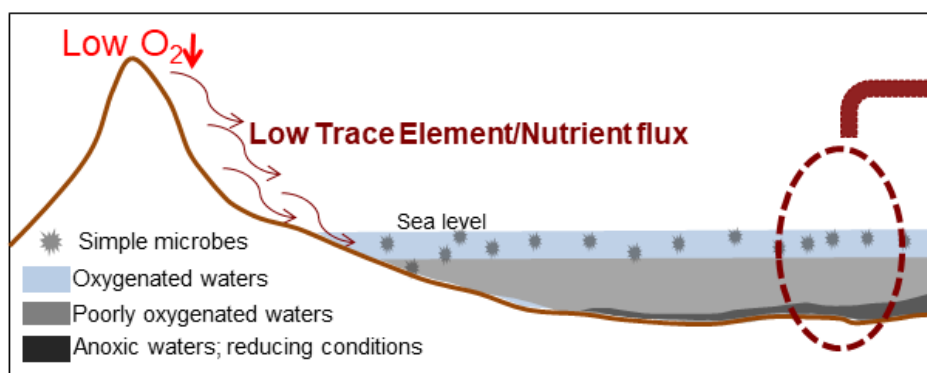


Fig 8 Pyrite sulphur isotopic compositions $\delta^{34}\text{S}_{\text{VCDT}}$ for the three black shale formations (black/bottom: Wollogorang; red/middle: Barney Creek; green/top: Velkerri); shaded regions represent range of sulphur isotopic compositions $\delta^{34}\text{S}_{\text{VCDT}}$ from Johnston et al. (2008) and Shen et al. (2002) in these formations.

Scenario 1

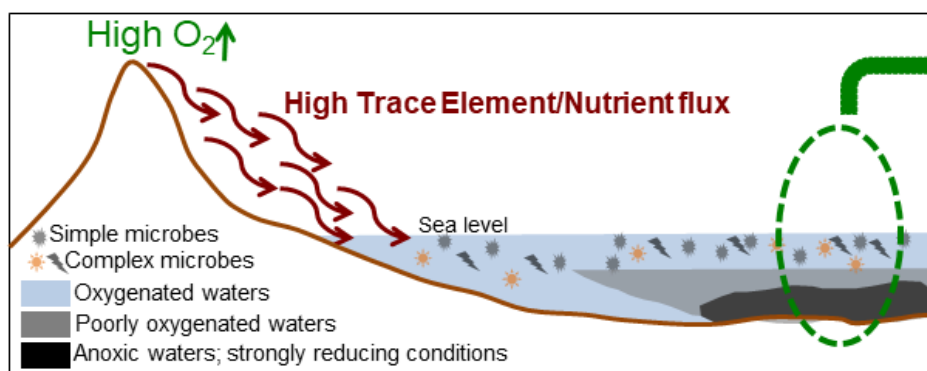


Less organic matter production and availability of reactive organic compounds

Lower sulphate reduction rates and proportion of pyrite burial

Low mean $\delta^{34}\text{S}_{\text{pyrite}}$

Scenario 2



More organic matter production and availability of reactive organic compounds

Higher sulphate reduction rates and proportion of pyrite burial

High mean $\delta^{34}\text{S}_{\text{pyrite}}$