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**MSA Award Lecture: An accessory mineral and  
experimental perspective on the evolution of the early crust.**

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26 **Abstract.** As the only known mineral with confirmed ages >4 Ga, zircon is unmatched in the  
27 field of early Earth research. In the past two decades, researchers have continued to establish  
28 connections between zircon chemistry and the physical/chemical processes that shaped the early  
29 crust. This connection has benefited greatly from the application of high temperature/pressure  
30 laboratory experiments. This study presents: (i) new zircon U-Pb geochronology and strategies  
31 for characterizing and identifying ancient terrestrial material from the Inukjuak Domain in  
32 northern Québec, and the Jack Hills, Western Australia; and (ii) a blend of new laboratory  
33 experiments and measurements of isotope ratios and trace impurities of natural zircon. Research  
34 directions in need of future exploration, with emphasis on early Earth studies, are also  
35 explored. Topics include Hadean hydrous magmatism and the structural accommodation of  
36 “water” into the zircon lattice, Hadean subaerial crust and the identification of peraluminous or  
37 metaluminous source melts, methods to characterize the oxidation state of magmas and fluids,  
38 and the complementarity of the Si- and O- isotopic systems as proxies for crustal  
39 weathering. Finally, the implications of this work are discussed in the context of  
40 a possible transition from prebiotic to biotic chemistry on the early Earth.

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## Introduction

50 Zircon ( $\text{ZrSiO}_4$ ) is an excellent carrier of geochemical information through time; in many  
51 ways, it is unsurpassed in its physical and chemical durability, and it is readily dated because it  
52 incorporates U and Th which decay to isotopes of Pb. It is because of this robustness that  
53 zircons are often found in sediments. The mere presence of this mineral is not diagnostic or  
54 unique to any specific rock type, tectonic setting, or environment; zircons occur in the  
55 continental and oceanic crust, in kimberlites, and in some meteorites and lunar rocks (Ireland and  
56 Wlotzka, 1992; Valley et al., 1998; Watson et al., 2006; de Hoog et al. 2014; Barboni et al.,  
57 2017). It is because zircons do crystallize in diverse settings that researchers have turned to the  
58 investigation of trace element chemistry and key isotopic ratios as proxies for their formation  
59 environment.

60 The importance is amplified because detrital zircons yield ages as old as ~4.4 Ga, which  
61 exceeds the age of the oldest known, widely agreed upon rock by about 400 million years (Maas  
62 et al., 1992; Wilde et al., 2001; Mojzsis et al., 2001; Holden et al., 2009; Thern et al., 2012;  
63 Mojzsis et al., 2014). Over the past two decades, numerous geochemical investigations of  
64 Hadean zircon (>4.0 Ga) have been conducted (e.g., Wilde et al., 2001; Mojzsis et al., 2001;  
65 Cavosie et al., 2005, 2006; Trail et al., 2007; Harrison et al., 2008; Hopkins et al., 2008; Bell et  
66 al., 2011, 2017; Harrison et al., 2017). These studies, which are mainly based on measurements  
67 of: (i) oxygen and hafnium isotope ratios; (ii) rare earth element (REE) contents; and (iii) the  
68 composition of other minerals (inclusions) found within the zircons, have led researchers to  
69 conclude that the early Earth contained an evolved rock cycle including water-rock interaction,  
70 formation of granitic crust and probable sediment cycling. The above studies typically utilized

71 experimental zircon diffusion data (see Cherniak, 2010) to argue for primary retention of trace  
72 element contents and isotope ratios of ancient grains.

73 In a generally different strategy, researchers sought to develop zircon-based calibrations  
74 in the controlled setting of an experimental geochemistry laboratory. Such experiments have  
75 played a fundamental role in the quest to link chemical signatures preserved in ancient zircons  
76 with Hadean processes. For instance, the Ti content of zircon was calibrated as an indicator of a  
77 zircon's crystallization temperature and zircon Ce anomalies were investigated as a proxy for  
78 redox conditions of early Earth magmas (Watson and Harrison, 2005; Watson et al., 2006; Ferry  
79 and Watson, 2007; Trail et al., 2011a; Trail et al., 2012). This work led to new discoveries about  
80 the early Earth, including evidence for water-saturated or near water-saturated Hadean magmas,  
81 and suggestions that Hadean volcanic emanations that were broadly neutral (e.g., CO<sub>2</sub>) rather  
82 than uniformly reduced (e.g., CO).

83 This paper presents new data, some speculations, and highlights a handful of research  
84 directions that are presently of interest to the author. With little deviation, this contribution  
85 explores the chemistry of the terrestrial zircon age end-members. A progress report is presented  
86 for zircon U-Pb geochronology studies of the Inukjuak Domain and the Jack Hills aimed at  
87 identifying the oldest terrestrial zircons and new fragments of ancient crust. The chemistry of  
88 the youngest, newly-formed crystals from laboratory experiments are discussed in the context of  
89 the chemistry of the oldest detrital Eoarchean and Hadean zircons. Many of the topics explored  
90 are relevant to origin(s) of life questions and planetary habitability. Such discussions will remain  
91 important as evidence for an early terrestrial biosphere continue to mount.

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### **Eoarchean and Hadean material.**

94 **Inukjuak domain**

95           Rocks hosting Eoarchean zircons outcrop approximately 30 km south of Inukjuak (town),  
96 near to Porpoise Cove on the eastern shore of Hudson Bay, located at the western edge of the  
97 Northeast Superior Province in northern Quebec, Canada (David et al., 2009; Cates and Mojzsis,  
98 2009; Cates et al., 2013; Darling et al., 2013). Thus far, well studied rocks in the region include  
99 the Nuvvuagittuq Supracrustal Belt (NSB) with amphibolite supracrustal rocks, rare felsic  
100 schists, possible conglomeratic units, oxide-rich and quartz-rich iron formations, all of which is  
101 surrounded by Archean granitoid gneisses. The NSB surrounds tonalites which yield ages from  
102 3.4 to 3.66 Ga (David et al., 2009; O'Neil et al., 2013). The felsic schist and conglomeritic units  
103 within in the NSB contain detrital zircons with U-Pb or Pb-Pb depth profile ages of ~3.7-3.8 Ga  
104 (David et al., 2009; Cates and Mojzsis, 2009; Cates et al., 2013), constraining the emplacement  
105 age of the NSB to the Eoarchean (Cates et al., 2013). Others have suggested that the rocks  
106 hosting these zircons are not sedimentary, but igneous intrusions, and that the NSB is  $\geq 4.2$  Ga  
107 (e.g., Darling et al., 2013).

108           O'Neil et al. (2008) discovered NSB cummingtonite-rich amphibolites record deficits in  
109 the daughter product ( $^{142}\text{Nd}$ ) of the extinct radionuclide  $^{146}\text{Sm}$ . These deficits require isolation of  
110 an enriched low Sm/Nd crustal source of these rocks from the  $^{142}\text{Nd}$  isotopic evolution of the  
111 bulk silicate earth in the first few hundred million years of Earth history. More recently, Caro et  
112 al. (2017) conducted  $^{146}\text{Sm}$ - $^{142}\text{Nd}$  and  $^{147}\text{Sm}$ - $^{143}\text{Nd}$  measurements on samples from the Ukaliq  
113 Supracrustal Belt (USB) a few km northeast of the NSB. These authors report the discovery of  
114 amphibolites with  $^{142}\text{Nd}/^{144}\text{Nd}$  deficits. Caro et al. (2017) favor inheritance from an enriched  
115 mantle source as the explanation for the USB and NSB  $^{142}\text{Nd}/^{144}\text{Nd}$  deficits (cf. O'Neil et al.,  
116 2008, 2013). Even though there is not yet consensus about the meaning of  $^{142}\text{Nd}/^{144}\text{Nd}$

117 anomalies in this domain, their discovery is extremely important to early earth geology, and has  
118 led to a discussion about the possible preservation and volume of Hadean crust this remote area.

119         The remoteness and challenges involved in sampling of this region perhaps warrant other  
120 exploration strategies. One possibility investigated here involves the collection of loose,  
121 unconsolidated sediments in strategic areas away from the present-day shore of the Hudson Bay.  
122 This tactic is based on the premise that such sediments are broadly regionally sourced, or at least  
123 are not dominated by loess or glacial till. In other words, unconsolidated sediments “sample”  
124 larger areas than possible with single hand samples, though in broadly restricted areas that may  
125 become the target of more intensive follow-up studies. This approach was tested in this domain  
126 for a few reasons. First, there is evidence for Eoarchean materials, and chemical signatures that  
127 have origins in the Hadean. Second, the rocks are reasonably well studied, in that the ages of the  
128 exposed surface area are broadly characterized. And finally, the younger Voizel suite contains  
129 inherited  $^{142}\text{Nd}/^{144}\text{Nd}$  deficits implying Hadean chemical remnants (O’Neil et al., 2008; Roth et  
130 al., 2013; Caro et al., 2017).

131         Unconsolidated sediment samples were collected during the July of 2016 field season.  
132 I226 and I247 were sampled in the NSB from small streams about 1 meter across and less than 1  
133 meter deep. I300 was collected from the shores of a small lake with a surface area of ~0.5 km,  
134 just to the east of the USB (**Figure 1**). Sediment samples were sieved and processed for zircon  
135 following standard heavy mineral separation procedures (e.g., Cates and Mojzsis, 2009; Trail et  
136 al., 2017). In this reconnaissance investigation, about 100 or more zircons from each of three  
137 samples were mounted on double-sided tape. Grains were cast in epoxy, polished to expose the  
138 cores, and U-Pb dated by LA-ICP-MS following the same procedure discussed in Trail et al.  
139 (2017). Briefly U-Pb data were collected with a Photon Machines 193 nm laser (25  $\mu\text{m}$  spot)

140 coupled to an Agilent 7900 quadrupole mass spectrometer, with ages standardized against AS-3  
141 zircon (Paces and Miller, 1993).

142 As expected, NSB stream samples yield U-Pb zircon ages dominated by the 2.7 Ga  
143 Boizard suite (**Figure 2**). Some ages are in broad agreement with Voizel suite and central  
144 tonalite ages (3.4 to 3.66 Ga). Note, however, that none of the ~3.7-3.8 Ga NSB zircon ages  
145 documented by others (e.g., David et al. 2009; Cates et al. 2013) have yet been discovered in  
146 these stream samples. Perhaps the most intriguing datum is from sediment sample I300,  
147 collected outside the NSB, which yielded a zircon age of 3740 Ma. The I300 sample was  
148 collected south and east of the area mapped by Caro et al. (2017), in which  $^{142}\text{Nd}/^{144}\text{Nd}$   
149 anomalies were identified. The importance of this find suggests that, like the NSB, this newly-  
150 documented supracrustal sequence may contain felsic rocks with Eoarchean zircon. On average,  
151 this small lake is likely to imply limited transport distance compared to the small stream samples  
152 collected (I226 and I247). There are two main outcomes from this reconnaissance sampling  
153 style in the Inukjuak domain. First, even low flow stream samples may prove too challenging to  
154 trace the origins rocks samples (I226 and I247), but this needs to be taken in context as only  
155 limited geochronology has been conducted. Second, the I300 sample predicts a hereto unreported  
156 felsic component in the USB that hosts >3.7 Ga zircons.

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## 158 **Jack Hills**

159 If zircon U-Pb geochronology is the criterion, then the Jack Hills classic locality  
160 metasedimentary outcrop, located in Western Australia, is the most well studied site in the world.  
161 Holden et al. (2009) presented a strategy and design for rapid Pb-Pb survey and U-Pb zircon age  
162 determination, using an automated ion microprobe. Thus far, approximately 200,000 zircons

163 been screened using this technique, with approximately 6000 grains yielding ages older than 4.0  
164 Ga (Harrison et al., 2017). Since the above Hadean zircon search commenced, advances in LA-  
165 ICP-MS (laser ablation inductively coupled plasma mass spectrometry) technology make it a  
166 viable geochronology alternative under certain circumstances.

167 Trail et al. (2017) presented U-Pb geochronology and other trace element data, for Jack  
168 Hills zircons by LA-ICP-MS. In this study of 275 zircons, ~12% of the grains are >4.0 Ga  
169 (**Figure 3a**). The volume of removed material is ~5000  $\mu\text{m}^3$ , leaving up to or greater than 90%  
170 of the sectioned grain for follow-up geochemical studies. Grains may be subsequently polished,  
171 leaving only the remaining outline of the ablation pits, if desired (**Figure 3b**). Consistent with  
172 previous results, a large fraction of the JH detrital zircon population is Paleoproterozoic, with an age  
173 peak centered at about 3.35 Ga. These Archean grains may be of value for certain studies (Bell  
174 et al., 2011), though the demand for such samples is diminished due to the presence of rocks of  
175 this age.

176 For those studies whose goal is to explore Hadean Earth, one possible Hadean zircon  
177 “concentration” strategy presently under exploration involves placing zircons mounted on  
178 double-sided tape directly into the same chamber of the LA instrument. Zircons remain  
179 unpolished and are not cast in epoxy. Approximately 1500 zircons from the same JH mineral  
180 separate as above were U-Pb dated at the University of Rochester (**Figure 3c,d**). Results reveal  
181 that 76/1514, or about 5% of the grains are older than 4.0 Ga. When compared to the previous  
182 data set, there is a shift in the peak of the old age population from 4.05 Ga (Figure 3a), to ~3.95  
183 Ga (Figure 3c). Moreover, a larger percentage of the apparent age population plot in the 3.8 to  
184 4.0 Ga interval. If these grains are included in the old age population, then the total increases



185 from 76/1514 to 173/1514 (or 11% of the population), which is comparable to the polished  
186 zircon U-Pb geochronology presented in Figure 3a.

187 Both zircon aliquots are from the heavy mineral separate and were picked from fresh  
188 separate. It is therefore reasonable to speculate that many of the unpolished grains with apparent  
189 ages between 3.8 to 4.0 Ga represent younger age domains or Pb loss mixed with older cores.  
190 This observation is consistent with Pb-Pb depth profiling and U-Pb spot mode geochronology  
191 conducted on Jack Hills zircons (e.g., Cavosie et al., 2004; Trail et al., 2007; Abbott et al., 2012).  
192 Since these zircons remain attached only to tape, the plan is to target grains with apparent ages  
193 older than 3.8 Ga. Samples will be mounted in epoxy, and prepared for follow-up U-Pb or stable  
194 isotope investigations. This strategy has clear advantages. First, a high percentage of ancient  
195 grains may be mounted in very close proximity to standards – a desired feature for ion  
196 microprobe work (e.g., Kita et al., 2009) – with relatively little effort. Second, in cases where a  
197 certain mass of Hadean material is required for a trace element or isotope measurement (e.g.,  
198 Amelin, 2004), ~90% of the original grain is retained and may be simply removed from the  
199 double-sided tape.

200

## 201 **Experiments and applications**

202

### 203 **Hydrous magmatism.**

204 Several lines of evidence imply the presence of hydrous magmatism on the early Earth,  
205 including zircon crystallization temperatures, oxygen isotopes that are fractionated away from  
206 the canonical mantle values, and inclusions assemblages, such as muscovite (Wilde et al., 2001;  
207 Mojzsis et al., 2001; Watson and Harrison, 2005; Hopkins et al., 2008; Harrison, 2009). Yet, a

208 calibration that enables direct quantification of the water activity of magmas through the analysis  
209 of zircon chemistry is not available. Trail et al. (2011b) conducted experiments in the H<sub>2</sub>O-SiO<sub>2</sub>-  
210 ZrO<sub>2</sub> system, at 1650 and 1550 °C, and found that “water” – or OH – can be structurally  
211 accommodated by the zircon lattice. The OH measurements were performed by Fourier  
212 Transform Infrared (FTIR) spectroscopy and H<sub>2</sub>O concentrations were quantified using the  
213 following relationship:

$$214 \quad c(\text{wt}\%H_2O) = \frac{A_i(\text{cm}^{-1}) * 1.8015}{t(\text{cm}) \cdot D(\text{g} / \text{cm}^3) \cdot \varepsilon_i(\text{cm}^{-2} \text{ per mol } H_2O / L)}$$

215 where  $c$  is concentration,  $A_i$  is the total integrated area in the OH stretching region,  $t$  is the  
216 sample thickness,  $D$  is the density of zircon (4.65 g/cm<sup>3</sup>), and  $\varepsilon_i$  is the absorption coefficient. de  
217 Hoog et al. (2014) suggested that the absorption coefficient used in this study and reported by  
218 Bell et al. (2004) may be different than the true value by a factor of 2, and should be revisited.  
219 Thus, new data presented below are reported as total integrated area ( $A_i$ ; see Trail et al. 2011b for  
220 additional details), with rare references to concentration.

221 Trail et al. (2011b) did not rigorously establish a relationship between the solubility of  
222 water in zircon and temperature. Those results however, do suggest this relationship does exist,  
223 with measurable differences in the H<sub>2</sub>O concentration detected for 1650 and 1550 °C  
224 experiments. New data for zircon synthesized in the H<sub>2</sub>O-SiO<sub>2</sub>-ZrO<sub>2</sub> system are plotted as  
225  $\log(A_i, \text{ per mm of zircon thickness})$  vs.  $10^3/T(\text{K})$  and reveal an inverse relationship (**Figure 4**).  
226 Accepting for the moment that these experiments have water activities close to unity, then  
227 extrapolation of these data down to 700 °C predicts a solubility of 0.5 to 10 ppm H<sub>2</sub>O in zircon.  
228 This range of concentration includes propagation of fitting errors (see Figure 4), and assumes a  
229 factor of 2 uncertainty in  $\varepsilon_i$ . Silicate melts with water activities lower than one will lower the

230 apparent concentration further. While measuring such low concentrations will be challenging,  
231 this is unlikely the largest issue. Trail et al. (2011b) and de Hoog et al. (2014) demonstrated that  
232  $H^+$  charge couples with trivalent cations, such as the rare earth elements, Y, and Al. Such charge  
233 coupling can lead to  $H_2O$  concentrations 2-3 orders of magnitude higher than “pure” zircon.  
234 This would, in principle, need to be subtracted out to yield results that bear specifically on water  
235 activity of the crystallization environment.

236 This problem of quantifying water activity from ancient zircon with this strategy is  
237 further hindered because radiation damaged grains commonly contain secondary  $H_2O$  not bound  
238 to the lattice (e.g., Woodhead et al. 1991). To demonstrate some of these complexities, new  
239 polarized FTIR spectra for synthetic and Jack Hills (JH) zircons are presented (**Figure 5**).  
240 Zircons were double polished, and then aligned along principal crystallographic axes before  
241 collecting polarized spectra. The synthetic zircon contains absorption bands that change in  
242 intensity as a function of crystallographic orientation, strong evidence that  $H_2O$  is indeed  
243 associated with the zircon lattice. Several other examples representing lattice bound  $H_2O$  in  
244 zircon are provided in other studies (e.g., Trail et al. 2011b; de Hoog et al. 2014). These results  
245 contrast with the two polarized spectra collected //a and //c for the JH zircon. Subtle differences  
246 in the absorption features may indicate trace amounts of water present within lattice, but  
247 absorption intensity is dominated by isotropic (secondary) water.

248 However, some of these complexities may be irrelevant for certain types of zircon.  
249 Consider, for example, zircons found in kimberlites which are trace element poor and contain  
250 low actinide concentrations. In this case, the relationship between  $H_2O$  solubility and water  
251 activity is likely to be more straightforward. Crystals from African kimberlites Kimberley Pool,  
252 South Africa, and Orapa, Botswana (Haggerty et al. 1983) were characterized by FTIR from core

253 to rim (**Figure 6**). The Orapa zircon, with a radius of about half that of the other sample, exhibit  
254 no significant change in the H<sub>2</sub>O concentration from rim to core. On the other hand, the  
255 Kimberley Pool crystal exhibits a broadly systematic decrease in H<sub>2</sub>O contents from rim to core.  
256 It is at least possible that the results for the Kimberley Pool zircon represent a diffusion profile,  
257 in which case a systematic study to characterize H diffusion in zircon would be extremely  
258 valuable as a geospeedometer under certain circumstances. Qualitatively, the results imply that  
259 this zircon retains a record of different chemical potentials for H<sub>2</sub>O, perhaps related to  
260 differences of the H<sub>2</sub>O activity of the zircon source region and the mantle source region of the  
261 kimberlite. Quantitative constraints will require additional work. For instance, since H<sub>2</sub>O  
262 solubility in zircon is a function of temperature, the crystallization temperatures will need to be  
263 constrained. Estimates of Ti-in-zircon temperatures for mantle zircons (Page et al. 2007) may be  
264 complicated by cation site exchange of Ti from the Si site (Tailby et al. 2011; Ferry and Watson,  
265 2007) to the Zr site as a function of pressure (Ferriss et al. 2008). Thus, the solubility of H<sub>2</sub>O in  
266 zircon is at least a function of temperature, water activity, and trace element composition. Based  
267 on very limited experiments – 1.5 and 2.5 GPa only – no pressure-dependent solubility is  
268 observed (Trail et al., 2011b) though this needs to be addressed in more detail. Such constraints  
269 may help constrain the water activity of these zircon source regions.

270

## 271 **Oxidation state of magmas and fluids**

272 The oxygen fugacity ( $f_{O_2}$ ) of a magma or fluid influences mineral saturation and stability,  
273 the viscosity of magmas, and controls the speciation of volatiles in the C-O-H-S system  
274 (Carmichael and Ghiorso, 1990; Frost, 1991; Connolly and Cesare, 1993). Oxygen fugacity also  
275 defines the molecular speciation of gases exsolved during a volcanic eruption. Trail et al.

276 (2011a, 2012) showed that redox-sensitive Ce uptake in zircon was systematically sensitive to  $f_{O_2}$   
277 and temperature, and as such, a Ce zircon anomaly – defined as the abundance of Ce relative to  
278 bracketing REEs La and Pr – was proposed to form the basis of an  $f_{O_2}$  sensor for zircon-bearing  
279 igneous rocks.  $Ce^{4+}$  is more compatible in the zircon lattice (vs.  $Ce^{3+}$ ), which is due to: (i) the  
280 smaller ionic radius of  $Ce^{4+}$  ( $Ce^{4+} = 0.97 \text{ \AA}$  vs.  $Ce^{3+} = 1.143$ ; Shannon 1976), which means that  
281  $Ce^{4+}$  more readily substitutes for  $Zr^{4+}$  ( $0.84 \text{ \AA}$ ); and (ii) the absence of a charge-balancing cation  
282 required for  $Ce^{4+}$  in zircon. This work enabled a broad estimate of the oxidation state of Hadean  
283 magmas. Magmatic outgassing throughout Earth history is at least partially responsible for the  
284 chemical state of the surface environment (Kasting, 1993; Canil, 1997; Delano, 2001; Burgisser,  
285 and Scaillet, 2007; Trail et al., 2011a). The variable compatibility of Ce vs.  $f_{O_2}$  (Trail et al.,  
286 2011a, 2012; Burnham and Berry, 2012), led to the discovery that Hadean ( $\geq 4.0$  Ga) zircon  
287 source melts were not as reduced as lunar samples, but ranged up to values closer to the redox  
288 state of modern day magmas (Trail et al. 2011a).

289         Alternative approaches include direct detection of  $Ce^{4+}/Ce^{3+}$  by X-ray Absorption Near  
290 Edge Structure (XANES) in zircon (Trail et al. 2015). This approach holds some advantages  
291 over other techniques because it relies directly on the chemistry of zircon (cf. Smythe and  
292 Brenan, 2016), and Ce valence is expected to be independent of concentration, meaning that  
293 there is no need to "normalize" Ce contents against other La, Pr, or other REEs. Bishop Tuff  
294 zircons contain systematic core-to-rim zoning in Ce valence, where core regions range from ~40-  
295 60 %  $Ce^{4+}$ , and zircon rims range from ~70-100 %  $Ce^{4+}$ , with errors of ~5% of the absolute scale  
296 (Trail et al., 2015). However, Ce valence needs to be mapped onto  $f_{O_2}$  before robust quantitative  
297 constraints are possible.

298 Experiments reported here suggest it is possible to calibrate the relationship between  
299 oxygen fugacity and a direct Ce valence measurement in zircon. Zircon crystals were  
300 synthesized in the  $\text{ZrO}_2\text{-SiO}_2\text{-H}_2\text{O-CeO}_2$  system at 1125°C and 10 kbar in a piston cylinder,  
301 which implies the following substitution mechanism  $\text{Zr}^{4+} \rightarrow \text{Ce}^{3+} + \text{H}^+$ , after Trail et al. (2011b).  
302 The experiments were  $f_{\text{O}_2}$ -buffered after a design presented in Trail et al. (2012) and Trail  
303 (2018). Cerium L<sub>3</sub> edge XANES spectra were collected at Beamline 13 ID-E (GSECARS), at  
304 the 7 GeV Advanced Photon Source (APS), Argonne National Laboratory after the technique  
305 described in Trail et al. (2015). Briefly, the incident X-ray energy was selected using a Si (111)  
306 double crystal monochromator; each spectrum required ~15 minutes to collect. Randomly  
307 selected grains from each experiment were analyzed and valence calculated from the  $\text{Ce}^{3+}$  and  
308  $\text{Ce}^{4+}$  end-member compounds  $\text{Ce}^{3+}\text{PO}_4$  and  $\text{Ce}^{4+}\text{SiO}_4$ , respectively (**Figure 6**). At an  $f_{\text{O}_2}$  of  
309 approximately 6.8 log units above the fayalite-magnetite-quartz equilibrium (FMQ+6.8), >90%  
310 of the Ce measured in the zircon crystal is in the tetravalent state. At an  $f_{\text{O}_2}$  of FMQ+0.7,  
311 approximately 60% of Ce is present as  $\text{Ce}^{4+}$ , with the percentage dropping to ~10% at an  $f_{\text{O}_2}$   
312 close to the iron-wüstite equilibrium.

313 While promising, there are several areas that will be need to be explored to extract useful  
314 oxygen fugacity information from Ce valence measurements in natural zircons. For instance,  
315 measured Ce valence within a zircon at a given oxygen fugacity may depend on the substitution  
316 mechanism for  $\text{Ce}^{3+}$ , where the above experiments imply  $\text{Zr}^{4+} \rightarrow \text{Ce}^{3+} + \text{H}^+$ . Additional  
317 experiments that explore the following mechanisms of substitution  $\text{Zr}^{4+} \rightarrow \text{Ce}^{3+} + \text{Li}^+$  and  $\text{Si}^{4+} +$   
318  $\text{Zr}^{4+} \rightarrow \text{P}^{5+} + \text{Ce}^{3+}$  (Trail et al. 2016) are worth investigating. Also, there are no experiments that  
319 quantify the preservation potential of Ce valence in zircons as a function of temperature. Even if  
320 this method is discovered to have limited value for Hadean zircons, it may be useful for younger

321 samples. Multiple isotope dilution-TIMS ages obtained from carefully micro-fractured zircon  
322 growth domains are providing new insights into continuously changing histories of silicate  
323 magmas (Matzel et al., 2006; Rivera et al., 2013). Combining such high precision age  
324 measurements with micron-scale Ce valence quantification (Trail et al., 2015) may enable a  
325 direct assessment of changes in magma oxidation state that occur over timescales of zircon  
326 growth.

### 327 **Hadean subaerial crust**

328 When did a subaerial surfaces emerge and what evidence of it remains in the geologic  
329 record? The discovery of mature sedimentary rocks such as quartzites or conglomerates supports  
330 the existence of subaerial exposure, including the Eoarchean Inukjuak rocks (David et al., 2009;  
331 Cates et al 2013; cf. Oneil et al., 2013), though Eoarchean or Hadean siliciclastic sedimentary  
332 rocks older than 3.9 Ga have not yet been discovered.

333 However, it is understood that weathering of subaerial rocks results in a residuum  
334 enriched in Al, not Na, Ca, and K, which are water soluble. If this material is buried and  
335 assimilated into a magma, it is expected to produce melts enriched in Al relative to  $\text{Na}^+$ ,  $\text{Ca}^{2+}$ ,  
336  $\text{K}^+$ ; i.e.,  $\text{Al}_2\text{O}_3/(\text{CaO}+\text{Na}_2\text{O}+\text{K}_2\text{O}) > 1$ , or peraluminous rocks. Hadean zircon inclusion  
337 mineralogy studies point to the presence of muscovite and biotite derived from peraluminous  
338 parent rocks (Mojzsis et al., 2001; Hopkins et al., 2008; 2010; Bell et al., 2017). These  
339 observations support the idea of Hadean subaerial continental crust because ~1-2% of Hadean  
340 samples contain primary muscovite inclusions that are exposed on the surfaces of polished grains  
341 (Bell et al., 2015a). Complementary to inclusion studies, Trail et al. (2017) quantified zircon Al  
342 contents by LA-ICP-MS and ion microprobe from metaluminous and peraluminous rocks from  
343 18 different granitoids. The expectation is that zircons that crystallized in peraluminous melts

344 may have higher Al concentrations – as a trace impurity – than zircons derived from  
345 metaluminous silicate melts with  $\text{Al}_2\text{O}_3/(\text{CaO}+\text{Na}_2\text{O}+\text{K}_2\text{O}) < 1$ . Zircons from peraluminous  
346 rocks yield an average concentration of  $\sim 10$  ppm Al, which defines a different distribution –  
347 though with overlap – than crystals found in metaluminous rocks (average  $\approx 1.3$  ppm). Limited  
348 application of these observations to the Hadean Jack Hills detrital zircon record suggests a  
349 peraluminous origin for one pre-3.9 Ga zircon ( $n = 39$ ), while about 8% (out of 236 zircons  
350 analyzed) of the Archean zircons exhibit elevated Al concentrations consistent with peraluminous  
351 melts. The small number of pre-3.9 Ga data should be taken in context; these measurements need  
352 to be extended beyond the 40 zircons evaluated in this pilot study.

353       Recent experimental results involving zircon synthesis in a piston cylinder from granitic  
354 melts, with different ASI values but same temperature, pressure, and water content reinforce  
355 observations made in natural samples (Wang and Trail, 2017). For the metaluminous melts ( $T =$   
356  $1150$  °C) zircon Al concentrations were as low as  $\sim 30$  ppm, whereas for the most peraluminous  
357 melts, concentrations are on the order of  $\sim 200$  ppm. The notably higher Al concentration in  
358 experimental zircons compared with natural zircons (1 to 10 ppm), is likely due to the higher  
359 crystallization temperature, e.g.,  $> 1100$  °C, compared to  $< 800$  °C, for most natural zircons. While  
360 future work seeks to parameterize the solubility of Al in zircon as function of temperature and  
361 melt composition, an ulterior goal will involve the characterization Al solubilities in zircon with  
362 the activities of  $\text{Al}_2\text{SiO}_5$ ,  $\text{SiO}_2$ , and  $\text{H}_2\text{O}$  buffered. Trail et al. (2011b) showed that Al may  
363 substitute into the zircon lattice through the following reaction  $\text{Al}^{3+} + \text{H}^+ \rightarrow \text{Si}^{4+}$  and so  $\text{H}_2\text{O}$   
364 activity might also control Al uptake in zircon. Presently, there is no evidence that alkalis  
365 charge compensate for Al substitution in zircon. For instance, there is no correlation between Al  
366 and Li concentrations in zircon (Trail et al., 2017), which suggests that Li incorporation into



367 zircon (or other alkalies) are mostly controlled by the presence of rare earth elements (Trail et al.,  
368 2016).

369

### 370 **Si and O isotopes in zircon**

371

372 About 75% of the crust is either Si or O by weight, and therefore the identity of  
373 weathered material strongly depends on lithospheric cycling of these two elements. Oxygen  
374 isotope measurements in silicate rocks are well established, and recent advances in analytical  
375 techniques are now at the precision required to resolve variations in igneous Si isotope  
376 compositions (e.g. Georg et al., 2006; Zambardi and Poitrasson, 2011; Savage et al., 2011,  
377 2012). These two isotopic systems have the potential to significantly advance our understanding  
378 of silicate weathering during the Hadean and Eoarchean and to constrain the identity Hadean  
379 weathered material.

380 Moreover, in the absence of a sedimentary/evolved component, Si isotopic composition  
381 varies linearly with Si content in igneous rocks, with the more differentiated samples yielding  
382 higher  $\delta^{30}\text{Si}$  values (Savage et al., 2011). In cumulate igneous rocks Si isotope compositions can  
383 be used to broadly predict the rock's normative mineralogy (Savage et al., 2013). Altogether,  
384 this information illustrates that the Si isotope composition of whole rocks record information of  
385 their past history, including weathering and magmatic differentiation.

386 To explore Si isotopes at the mineral scale, mineral-mineral Si isotope fractionation  
387 factors are needed. The magnitude and sense of Si isotope fractionation between zircon and  
388 quartz have been experimentally investigated based on the 3 isotope exchange method, in which  
389 zircon and quartz progressively approached their equilibrium isotopic composition. Direct

390 synthesis experiments similar to the zircon-quartz oxygen isotope fractionation study of Trail et  
391 al. (2009) have also been conducted. Preliminary data yield calculated fractionations that are  
392 broadly consistent with those measured between zircon-quartz pairs in natural samples (Trail and  
393 Savage 2015, 2017); that is,  $\delta^{30}\text{Si}(\text{quartz}) - \delta^{30}\text{Si}(\text{zircon}) \approx 0.4\text{‰}$ . With this information,  
394 simultaneous multicollecion of O and Si isotopes using the CAMECA *ims1290* ion microprobe  
395 at UCLA will add a new dimension to the nature of silicate weathering of the early crust (Figure  
396 3b). Specifically, the goal is to better define the composition of Hadean material involved in  
397 crust-atmosphere-hydrosphere interactions, which may include felsic/mafic sediments, chemical  
398 sediments, and altered basalt.

399

400

## Implications

401 Hadean zircon chemistry has transformed our view of the earliest Earth, even with full  
402 awareness that our understanding is likely biased by the restricted provenance of the material thus  
403 far explored. No single example highlights this more than the Jack Hills zircons, as the vast  
404 majority of Hadean zircons studied in detail come from a single outcrop (Harrison et al., 2017).  
405 These samples continue to serve an integral role in the quest to decipher early Earth  
406 environments, though a more accurate picture of Hadean/Eoarchean surface states and  
407 environments requires investigations that extend beyond single outcrops. There are 13 sites for  
408 which pre-4.0 Ga material has been identified (Harrison, et al., 2017), which will surely lead to  
409 information about the Hadean Earth if interrogated to the same level of detail as the Jack Hills.

410 As recorders of crustal evolution, perhaps the most intriguing aspect of Hadean zircon  
411 chemistry is that many of them appear to have formed in volatile rich magmas that captured and  
412 archived a fraction of Earth's ancient surface chemistry through re-melting of sediments. The

413 combination of Si- and O-isotopes together may lead to new insights that define the weathered  
414 product with greater detail (Abraham et al. 2011). Given the prospect of a Hadean biosphere  
415 (Bell et al., 2015b), reconstructions of the near surface environment may be essential to  
416 understand the transition from prebiotic to biotic chemistry.

417

418

419 For instance, the presence (and abundance) of subaerial continental exposure is important  
420 for some origin of life models. First, subaerial weathering of continental crust may have provided  
421 key components (e.g., clays?) needed for early pre-biotic chemistry (Ferris and Ertem, 1993); Al  
422 concentrations and Si-/O-isotope ratios in zircon will help define the identity of this material.  
423 Second, the presence of subaerial exposures help facilitate model prebiotic chemical reactions  
424 conducted in the laboratory. Consider, for example, that RNA may be a potentially important  
425 link between prebiotic chemistry and modern DNA biochemistry (Zaug and Cech, 1986).  
426 However, ribose, which is a major component of the RNA molecule, is unstable in many aqueous  
427 environments. One possible solution to this dilemma centers on a geological model in which  
428 RNA developed in an aqueous environment enriched in borate. Borate acts as a complexing  
429 agent that effectively stabilizes ribose against decomposition in aqueous solutions. Specifically,  
430 ribose is stabilized in solution in the presence of borate-buffering evaporate minerals such as  
431 colemanite,  $\text{CaB}_3\text{O}_4(\text{OH})_3(\text{H}_2\text{O})$ , ulexite,  $\text{NaCaB}_5\text{O}_6(\text{OH})_6(\text{H}_2\text{O})_5$ , and kernite,  
432  $\text{Na}_2\text{B}_4\text{O}_6(\text{OH})_2(\text{H}_2\text{O})_3$ , implying a desert-like subaerial environment (Ricardo et al., 2004;  
433 Benner et al., 2011). If the simple organic precursors existed, the presence of borate could have  
434 allowed the accumulation of ribose in prebiotic environments. In this regard, evidence for

435 subaerial exposure, whether in the form of inclusion mineralogy or trace elements in zircon (Al)  
436 is needed to support the above model.

437 Another question is whether complex molecules such as amino acids formed on the early  
438 Earth. One of the first breakthroughs is the now-famous Miller-Urey experiment in which an  
439 “atmosphere” comprised of the reducing gases methane (CH<sub>4</sub>), ammonia (NH<sub>3</sub>), and hydrogen  
440 (H<sub>2</sub>) were subjected to simulated lightning (Miller, 1953). In the mix of experimental products  
441 were many of the amino acids that are major components of cells. Later experiments showed  
442 that the products also contained high energy intermediates like cyanide and acetylene (Stribling  
443 and Miller, 1987). This was an important observation because cyanide and acetylene can  
444 assemble in subsequent steps to form nucleotides, which are the building blocks of genetic  
445 material. Knowledge of the early atmospheric composition would enable researchers to explore  
446 reaction pathways that could lead to molecules of increasing complexity such as amino acids, but  
447 without knowledge of the reactants, the products of such experiments are of limited value. As  
448 discussed, experiments help constrain volatiles emanating from the Hadean Earth, which  
449 probably dominated by neutral species. While a useful constraint, defining the relative  
450 abundance of possible reactants as inputs for Miller-Urey type experiments would benefit from a  
451 more precise oxygen fugacity sensor, coupled with zircon-melt partition coefficients for elements  
452 that comprise volatile species (vs. oxygen fugacity). While H in zircon has certain complications  
453 (section 3.1), Harrison et al. (2017) speculated about the possible incorporation of C into natural  
454 zircon.

455 The picture of the early Earth will likely become clearer as new geochemical tools and  
456 additional crustal remnants are discovered. Some of these problems may not be solvable in  
457 realistic timeframes, or may give sufficiently ambiguous results with limited value.

458 Nevertheless, models of the earliest Earth environmental conditions based on geochemical data  
459 from this Eon should be held in preference.

460

461

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471

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## Figure Captions

684 **Figure 1** Sketch map, identifying the major geological units (David et al. 2009; O’Neil et al.  
685 2013; Greer, 2013) and locations of sampled loose sediment. I226 and I247 were collected from  
686 small streams and I300 collected from the shore of a small lake. The Boizard suite is comprised  
687 of a heterogenous gneissic tonalite with ultra/mafic enclaves; zircon U-Pb geochronology yield  
688 ages of 2722 to 2750 Ma, and associated pegmatite dikes yield U-Pb monazite ages of 2688 Ma  
689 (David et al., 2009). The Voizel suite consists of tonalitic gneisses. The Nuvvuagittuq  
690 Supracrustal Belt (NSB), is dominated by volcano-sedimentary sequences (David et al., 2009).  
691 The Ukaliq Supracrustal Belt (USB) is located a few km north and east of the NSB; a detailed  
692 field map of the boxed in region is presented by Caro et al. (2017).

693

694 **Figure 2.** U-Pb geochronology of zircons separated from unconsolidated sediment, with ~85%  
695 of the ages (i.e., >2.65 Ga) reflecting the local geology (Figure 1). I226 (n = 207) and I247 (n =  
696 84) were collected from small streams passing through the NSB. I300 (n = 165) was collected  
697 near the Ukaliq Supracrustal Belt and contains a single grain with an age of 3.74 Ga, implying  
698 this belt may contain a sequence with Eoarchean zircons.

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700 **Figure 3. (a)** U-Pb LA-ICP-MS geochronology of Jack Hills zircons collected from the classic  
701 locality (e.g., Weiss et al., 2015). Of the 275 zircons, 34 are older than 4.0 Ga or about 12 %. **(b)**  
702 Secondary Electron and cathodoluminescence images of a 4.06 Ga grain, polished after U-Pb  
703 geochronology. The remaining outline of the LA-ICP-MS pit is visible, and to the left are ion  
704 microprobe spot analyses of Si and O isotopes. **(c)** U-Pb LA-ICP-MS geochronology of whole,  
705 unpolished grains reveal a decrease in identified grains older than 4.0 Ga, possibly due to the

706 intersection of younger overgrowths. **(d)** Stereoscope images of whole grains with U-Pb ages and  
707 the location of LA-ICP-MS pits identified.

708

709 **Figure 4.**  $\log(A_i)$ , per mm of zircon thickness) showing that absorption, which is proportional  
710 to concentration, is inversely related to temperature. All zircons were synthesized in the simple  
711  $\text{H}_2\text{O-SiO}_2\text{-ZrO}_2$  system. The total absorption,  $A_i$ , is calculated after the procedure discussed in  
712 Trail et al. (2011b). The  $A_i$  is plotted here, rather than concentration, because of the need to re-  
713 quantify the absorption coefficient for zircon (de Hoog et al., 2014).

714

715 **Figure 5.** Polarized spectra for doubly-polished crystals. Both zircons were treated in cold HF.  
716 When the electric vector is parallel to c ( $E//c$ ) vs.  $E//a$ , the synthetic zircon spectra exhibit clear  
717 changes in absorption bands. Absorption spectra for a Jack Hills zircon contains broadly similar  
718 features and intensity independent of whether  $E//c$  vs.  $E//a$ , suggesting that most of the “water” is  
719 isotopic and therefore likely to be secondary. Similar features were observed for four other Jack  
720 Hills zircons.

721

722 **Figure 6.** The total absorption,  $A_i$ , plotted vs. distance from the rim showing a broadly  
723 monotonic decrease in concentration for the Kimberly Pool zircon, though the same correlation  
724 for the smaller Orapa Botswana zircon is absent. Samples were double polished and  
725 measurements were taken  $//a$  from rim to core.

726

727 **Figure 7 (a)** Example Ce  $L_3$  edge XANES absorption spectra of synthetic zircons and standards.  
728 The bottom two spectra are end-member standards used to model valence of the unknowns. The



729 reduced→oxidized series represent spectra collected from synthetic zircons buffered at FMQ-  
730 3.4, FMQ+0.7, and FMQ+6.8 respectively. Spectra show increased abundances of Ce<sup>4+</sup>, as  
731 judged by comparisons with standard spectra, with increasing oxygen fugacity. All spectra are  
732 normalized to remove concentration effects for direct comparison of valence. **(b)** Backscattered  
733 electron image of synthesized grains and **(c)** APS beamline reflected light image during data  
734 collection. The scale bar applies to both images.

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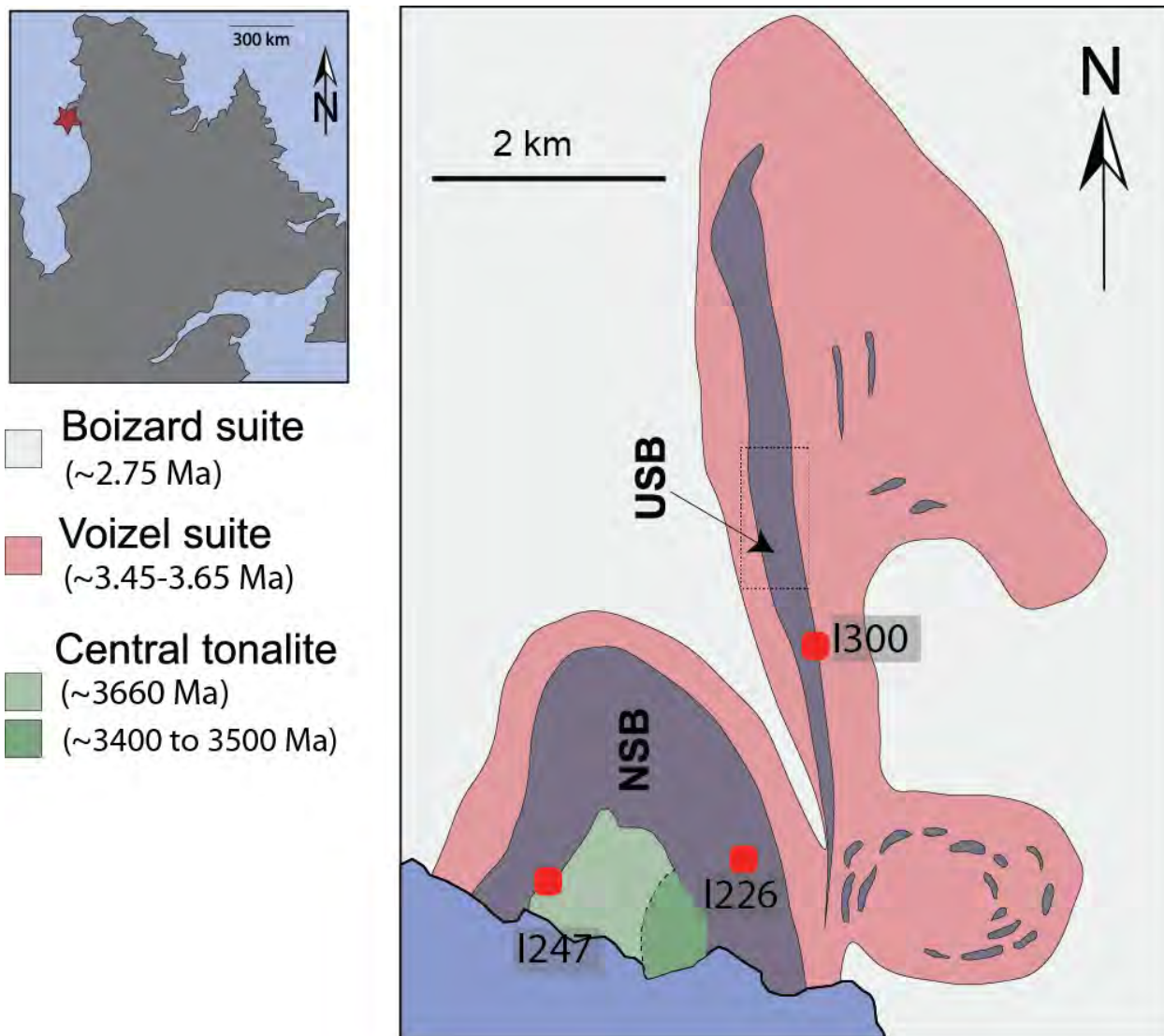
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745 **Figure 1.**



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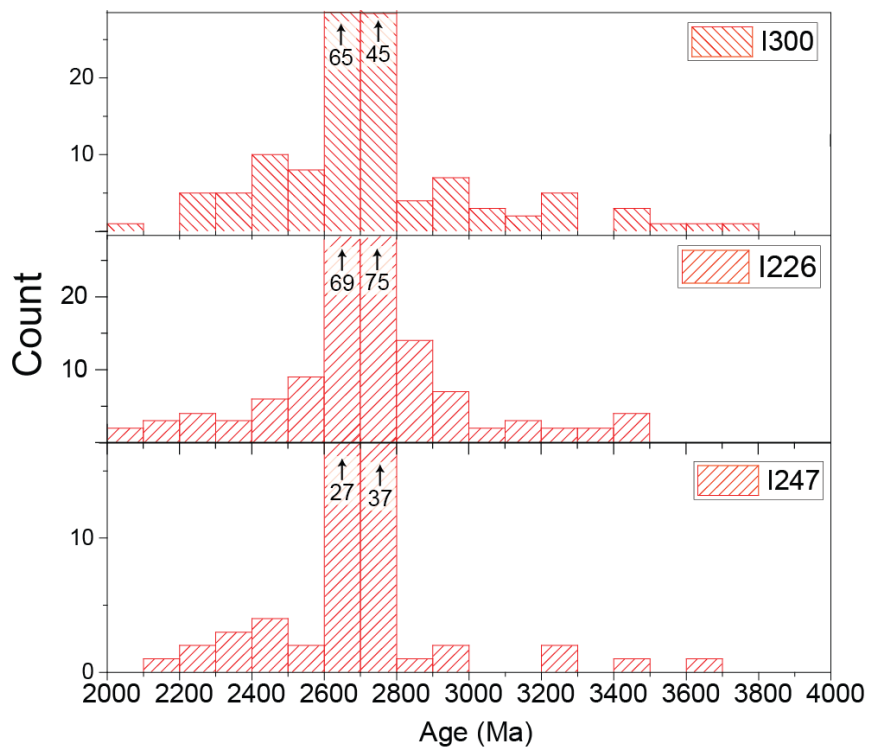
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754 **Figure 2**



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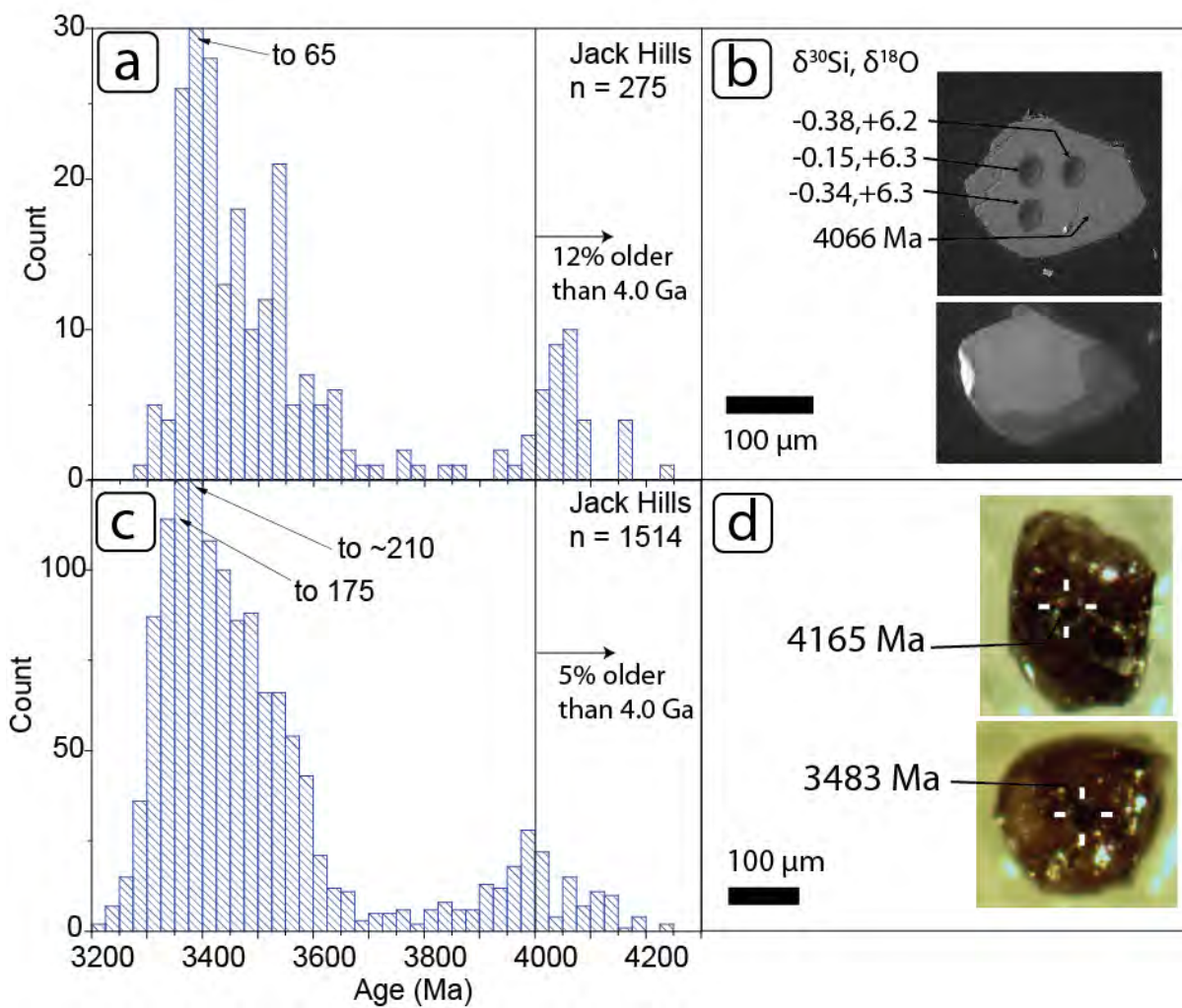
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768 **Figure 3.**



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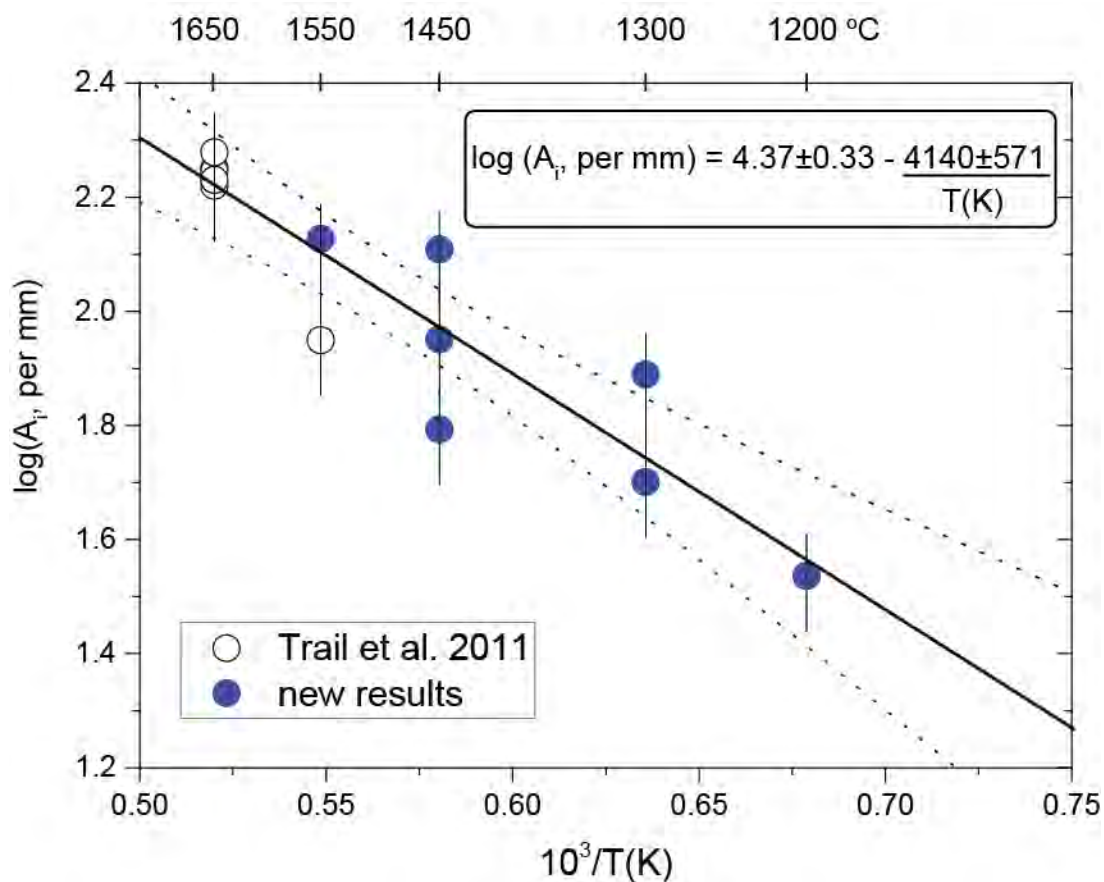
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782 **Figure 4**



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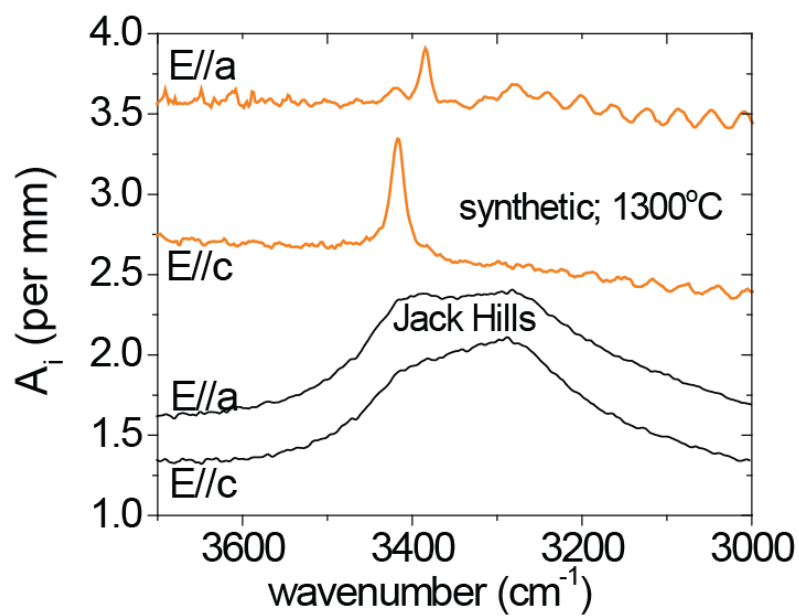
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794 **Figure 5.**



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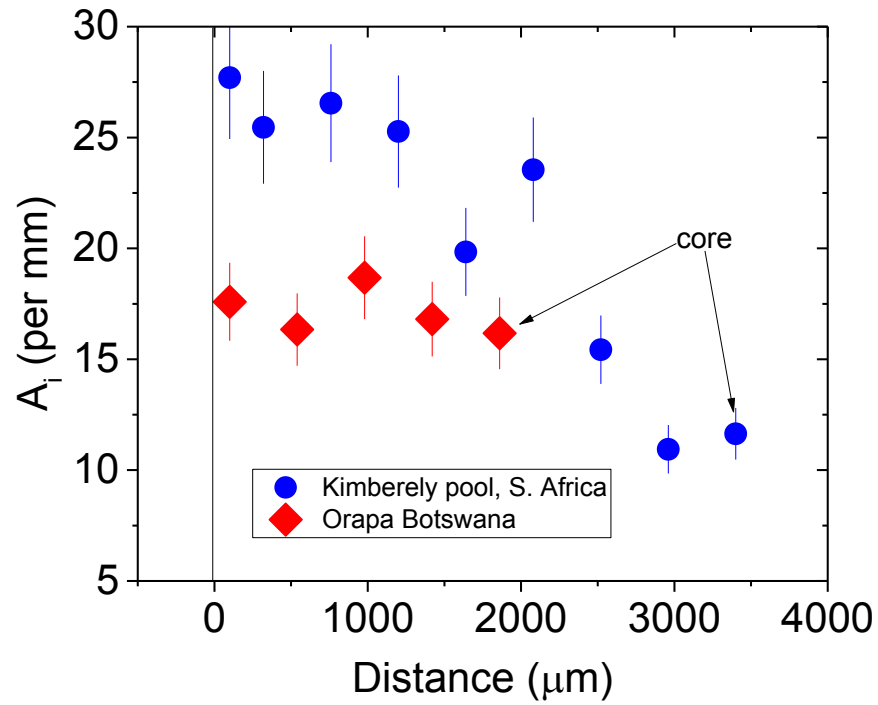
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810 **Figure 6**



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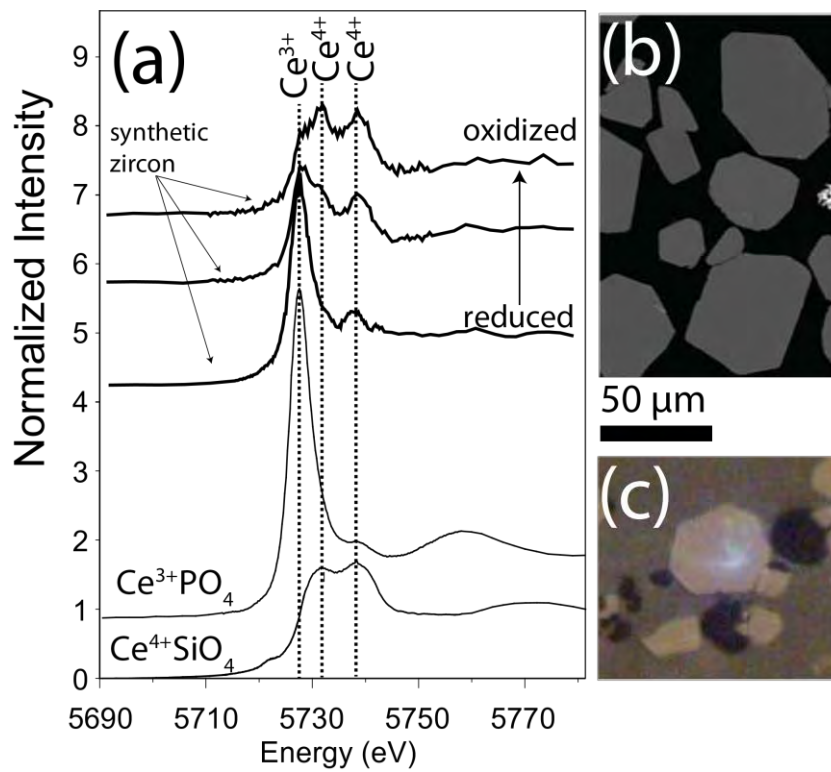
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823 **Figure 7**



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