

1 **Revision 3**

2 **Anisotropic growth of olivine during crystallization in basalts from**  
3 **Hawaii: Implications for olivine fabric development**

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## ABSTRACT

23 Olivine fabrics are crucial for deciphering anisotropy of many physical properties in the  
24 upper mantle, such as seismic velocity and thermal diffusivity. Under circumstances  
25 where crystallized mantle rocks don't suffer strong modification from later deformation,  
26 olivine fabrics can reflect the results of anisotropic crystal growth and viscous magmatic  
27 flow. Currently little is known about the crystallization habits of olivine in nature. Here  
28 we report crystallization habits of olivine phenocrysts in two Hawaii tholeiitic basalts.  
29 The low Mg numbers ( $< 88.1$ ) of olivine grains suggest that they are crystallized products  
30 from the host magma rather than captured ones from peridotite xenoliths. Many olivine  
31 grains have non-equidimensional euhedral crystal shapes due to anisotropic crystal  
32 growth rates. Electron backscatter diffraction (EBSD) analyses show that of the 115  
33 analyzed olivine grains, 84 grains are oriented with their long axes parallel to the  
34 [001]-axis, while 24 and 7 grains have their long axes parallel to the [100]-axis and the  
35 [010]-axis, respectively. This growth habit is different from that reported for olivine  
36 grown at subsolidus conditions without fluid/melt. Our results imply that olivine will  
37 most likely form the AG-type fabric during crystallization from magma chambers  
38 undergoing compaction, and the B-type fabric during pure or simple shear mode  
39 magmatic flow, provided that no or weak later plastic deformation superimposes on these  
40 magmatic fabrics. The AG-type and the B-type olivine fabrics in natural cumulates can be  
41 interpreted as results of anisotropic growth of olivine and viscous magmatic flow.  
42 Anisotropic olivine growth is expected to strongly affect the fabric development in the

43 upper mantle portion where there is extensive partial melting (e.g., the mantle wedge  
44 above subduction zones or the upper mantle beneath the middle ocean ridges) or reactive  
45 melt percolation leading to crystallization of olivine at the consumption of pyroxene.

46

47 **Keywords:** olivine fabric, crystallographic preferred orientation, anisotropic growth,

48 cumulate

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## INTRODUCTION

52 Crystallographic preferred orientation (CPO) of minerals is important for extrapolating  
53 deformation mechanism, shear sense, strain geometry, deformation temperature of rocks  
54 from deep crust and upper mantle of the Earth (e.g., [Law 1990](#)). It is generally believed  
55 that significant CPOs of minerals are produced by dislocation glide during dislocation  
56 creep of rocks (e.g., [Zhang and Karato 1995](#); [Tommasi et al. 2000](#); [Zhang et al. 2006](#);  
57 [Soustelle et al. 2010](#); [Michibayashi and Oohara 2013](#)). In addition, remarkable CPOs may  
58 also result from grain-size sensitive creep regimes, like diffusion creep or grain boundary  
59 sliding (GBS) (e.g., [Bons and den Brok 2000](#); [Hansen et al. 2012](#); [Miyazaki et al. 2013](#);  
60 [Getsinger and Hirth 2014](#); [Précigout and Hirth 2014](#)), although many studies have linked  
61 weak or random CPOs of minerals to deformation by diffusion creep or grain boundary  
62 sliding as well (e.g., [Storey and Prior 2005](#); [Warren and Hirth 2006](#); [Skemer and Karato](#)  
63 [2008](#); [Drury et al. 2011](#)). Apart from deformation-induced CPOs, distinct mineral CPOs  
64 are also frequently seen in undeformed rocks, like granites, anorthosites, gabbros, and  
65 peridotites of magmatic origin (e.g., [Jackson et al. 1975](#); [Benn and Allard 1989](#);  
66 [Mainprice and Nicolas 1989](#); [Müller et al. 2011](#); [Ji et al. 2014](#); [VanTongeren et al. 2015](#)).  
67 These magmatic CPOs are crucial for understanding the dynamical processes within  
68 magma chambers (e.g., [Benn and Allard 1989](#)) and to better constrain the CPO  
69 development and deformation mechanisms of deformed magmatic rocks, considering that  
70 CPOs arising from magmatic flow could be preserved during later deformation by

71 grain-size sensitive flow (e.g., [Cao et al. 2017](#)), but destroyed during strong deformation  
72 by dislocation creep. Thus, knowledge of CPO formation under supersolidus conditions is  
73 of significance.

74

75 Olivine is the most abundant and probably the weakest constituent mineral in the upper  
76 mantle ([Mackwell et al. 1990](#)). Numerous efforts have been made to elaborate olivine  
77 CPOs during deformation in nature and laboratory experiments (e.g., [Avé Lallemant 1975](#);  
78 [Zhang and Karato 1995](#); [Bystricky et al. 2000](#); [Jung and Karato 2001](#); [Tommasi et al.](#)  
79 [2008](#); [Ohuchi et al. 2011](#)). These studies demonstrated that olivine CPOs in response to  
80 dislocation creep vary as a function of temperature, stress, water content, pressure, and  
81 melt, in addition to deformation geometry. For instance, olivine develops fabrics  
82 characterized by the [001] axes parallel to the shear direction and the [010] axes or the  
83 [100] axes normal to the shear plane under water-rich conditions (i.e., the B- and C-type,  
84 respectively), and forms fabrics featured by the [100] axes parallel to the lineation and  
85 the [010] axes normal to the foliation under water-poor conditions (i.e., the A-type) ([Jung](#)  
86 [and Karato 2001](#)).

87

88 In comparison, considerably fewer studies had been dedicated to clarify how olivine CPO  
89 will develop under supersolidus (magmatic) conditions ([Benn and Allard 1989](#);  
90 [Zetterström 1996](#); [Jousselin et al. 2012](#)). Although olivine also show distinct CPOs under  
91 supersolidus conditions, such as the AG-type (also known as the [010]-fiber pattern; i.e.,

92 the [010] axes form a point maximum normal to the foliation plane, and the [100] and the  
93 [001] axes both form a great girdle in the foliation plane; see [Mainprice 2007](#); [Tommasi  
94 et al. 2008](#)) or the B-type, they cannot be linked to plastic deformation and explained by  
95 existing hypotheses, but are related to the crystal shape of olivine crystallized from  
96 magmas. Therefore, the crystallization habits of olivine during magma cooling can cast  
97 lights on olivine fabric development in olivine-bearing magmatic rocks. Many studies  
98 have explored in detail the textural evolution and compositions of olivine crystallized  
99 from magmas through experiments and natural observations (e.g., [Donaldson 1976](#); [Faure  
100 et al. 2003a, b](#); [Faure et al. 2007](#); [Welsch et al. 2013](#); [Welsch et al. 2014](#)). However, these  
101 studies mainly focused on the morphology and textures of olivine phenocrysts. Statistical  
102 analyses of growth habits of olivine are absent.

103

104 In this study, we carried out a detailed study on olivine phenocrysts from two Hawaii  
105 basalts. The mineral chemistry evidence is presented to demonstrate that the phenocrystic  
106 olivine was derived from fractional crystallization of the host basalts. Then the evidence  
107 of oriented growth of olivine phenocrysts preferentially along their [001]-axis over the  
108 [100]- and the [010]-axis is reported. Finally, the significance of anisotropic growth of  
109 olivine crystals for fabric development in cumulates is discussed.

110

111

## SAMPLES AND METHODS

112 Two vesicular Hawaii basalts, one from the O'ahu Island and the other from the Big

113 Island, were collected for this study. They are both tholeiitic lavas with MgO of  
114 46.99-47.38 wt% and Na<sub>2</sub>O + K<sub>2</sub>O of 1.49-2.61 wt% (see the supplementary **Table S1**).  
115 Nine thin sections with a final polishing of 0.05 μm by colloidal Al<sub>2</sub>O<sub>3</sub> had been prepared  
116 from these two samples. The major-element compositions of olivine phenocrysts were  
117 determined using a JEOL Superprobe JXA-8100 electron microprobe at the Institute of  
118 Geology and Geophysics, Chinese Academy of Sciences. The analyses were conducted at  
119 an accelerating voltage of 15 kV, a beam current of 20 nA, and a spot diameter of 5 μm.  
120 The counting time for SiO<sub>2</sub>, MgO, FeO was 20 s on peak and 10 s on lower and upper  
121 background positions. The counting time for other elements was 40 s on peak and 20 s on  
122 lower and upper background positions. All elements were detected by five analyzing  
123 crystals: two TAP crystals for Si, Mg, Al and Na, two LIFH crystals for Fe, Ni, Mn and  
124 Cr, and one PETJ crystal for K, Ca and Ti, respectively. SPI standard olivine-34 and  
125 diopside-21 were used as monitor standards. The detection limits for all elements are  
126 100-200 ppm. The precision (1SD) is 0.4-1.0% for major elements (Si, Mg and Fe), 3-9%  
127 for minor elements (Ca, Mn and Ni), and 19-90% for trace elements (Al, Ti, Cr, K and  
128 Na).

129

130 To clarify the anisotropic growth behavior of olivine during crystallization from magmas,  
131 we selected those euhedral and anisometric olivine grains. Some rounded olivine grains  
132 can sometimes be observed in these samples. However, they are neglected because it is  
133 difficult to determine their long/short axes and aspect ratios although they may also

134 represent crystallization products from magma (Welsch et al. 2013). The crystal  
135 orientations of selected olivine grains were determined by indexation of electron  
136 backscatter diffraction (EBSD) patterns using a Quanta 450 field emission gun scanning  
137 electron microscope (FEG-SEM) at the State Key Laboratory of Geological Processes  
138 and Mineral Resources, China University of Geosciences (Wuhan). The 'match unit' for  
139 olivine in our HKL software is a forsterite with a Pbnm space group (lattice parameters of  
140 the unit cell are  $a = 4.76 \text{ \AA}$ ,  $b = 10.21 \text{ \AA}$ , and  $c = 5.98 \text{ \AA}$ ). The measurements were  
141 conducted with an accelerating voltage of 20 kV, a working distance of 24-25 mm, and a  
142 spot size of 6. Diffraction patterns were collected manually and indexed using the HKL  
143 Channel 5 + software. Only those measurements with mean angular deviation (MAD)  
144 values below 1.0 (between detected and simulated Kikuchi bands) were accepted and  
145 recorded. Moreover, the corresponding backscatter electron (BSE) image for each  
146 measured grain was also saved at the same time.

147

148 The crystal orientations of olivine phenocrysts are presented by stereographic projection  
149 of the [100], [010], and [001] axes measured by EBSD. To derive the true growth aspect  
150 ratios between two perpendicular crystallographic axes, only those stereographic  
151 projections with the third crystallographic axis projected approximately in the center  
152 ( $<10^\circ$  to the center) are chosen for comparison with the corresponding BSE images. In  
153 such cases, the aspect ratio can be easily determined.

154



155

## RESULTS AND DISCUSSION

156 Under a petrographic microscope, the studied olivine grains (0.2-3.6 mm in length) are all  
157 euhedral, tabular, and free of signs indicative of intracrystalline plastic deformation, like  
158 undulose extinction and subgrains (Fig. 1). Sometimes, inclusions of basaltic glass or  
159 plagioclase can be observed within olivine (Figs. 1c and 1d). These features clearly  
160 demonstrate that these olivine grains grew from the host basalts.

161

162 18 olivine phenocrysts from the studied Hawaii basalts have been analyzed for the  
163 major-element compositions (see the Supplementary Table S2). They show a considerable  
164 compositional variation with MgO contents ranging from 41.97 wt% to 47.60 wt% and  
165 FeO contents varying between 11.43 wt% and 18.67 wt% (Fig. 2). The Mg number (=  $\text{Mg}/[\text{Mg} + \text{Fe}_{\text{total}}] \times 100$  atomic ratios) values of olivine range from 80.0 to 88.1, with  
166 most of them falling below 86.0 (Supplementary Table S2). They are generally richer in  
167 FeO and more depleted in MgO compared with those from peridotite xenoliths entrained  
168 by the Hawaii basalts (e.g., Clague 1988; Sen 1988; Fig. 2), indicating that the olivine  
169 phenocrysts are crystallization products from the Hawaii basalts during cooling, rather  
170 than relics of captured peridotite xenoliths.

172

173 The crystallographic orientations of 546 randomly selected euhedral olivine phenocrysts  
174 were determined. Among them 115 olivine grains with one crystallographic axis  
175 projected approximately close to the center (i.e., normal to the thin section surface) are

176 chosen to derive the growth aspect ratios between two crystallographic axes. These 115  
177 grains can be divided into three groups: 1) 42 grains with the [100]-axis normal to the  
178 thin section surface (Group #1), 2) 46 grains with the [010]-axis normal to the thin  
179 section surface (Group #2), and 3) 27 grains with the [001]-axis normal to the thin section  
180 surface (Group #3) (Fig. 3 and Table 1). For Group #1 olivine grains, the long axes are  
181 always the [001]-axis and the short axes are always the [010]-axis with an average aspect  
182 ratio of  $1.7 \pm 0.6$  (error represents two times the standard deviation; Fig. 4a). For group  
183 #2 olivine grains, the majority (42 out of 46 grains) shows also the [001] axes being the  
184 long axes while the [100] axes being the short axes (with a corresponding average aspect  
185 ratio of  $1.6 \pm 0.7$ ; Fig. 4b). The long axes of rest 4 grains (8.7%; aspect ratios of 1.2-1.7)  
186 are the [100]-axis. For Group #3 olivine grains, the long axes of 20 out of 27 grains are  
187 the [100] axes (average aspect ratio =  $1.4 \pm 0.4$ ; Fig. 4c) while those of the rest (aspect  
188 ratios ranging from 1.3 to 2.3) are the [010] axes. The observed crystal faces for these  
189 olivine grains include {021}, and rarely, {101}, {120} and {110}, in addition to the  
190 low-index crystal faces {010}, {100}, and {001} (Fig. 3). These crystal faces were also  
191 observed for tabular olivines from peridotite xenoliths, which were explained as growth  
192 in the presence of grain boundary fluid (see Drury and van Roermund, 1989).

193

194 In summary, the EBSD analyses show that olivine, like many other minerals from the  
195 crust and the upper mantle, also grows anisotropically during crystallization in magma  
196 chambers. Our measurements demonstrate unambiguously that olivine phenocrysts

197 grown from the Hawaii basalts have mostly a short prism shape with growth of the  
198 [001]-axis preferred relative to the [100]- and the [010]-axis and that of the [100]-axis  
199 preferred relative to the [010]-axis ([Table 1](#)). The preferential growth of the [001]-axis  
200 over the other two axes has also been reported by several previous studies ([Donaldson](#)  
201 [1976](#); [Miyazaki et al. 2013](#); [Welsch et al. 2013](#)).

202

203 Previous studies have shown that olivines with euhedral, tabular morphology are  
204 common in natural mantle peridotites (e.g., [Boullier and Nicolas 1973, 1975](#); [Drury and](#)  
205 [van Roermund 1989](#); [Vauchez et al. 2005](#); [Vonlanthen et al. 2006](#); [Tommasi et al. 2008](#)).

206 Generally, these olivines are considered to form as a result of grain growth under  
207 subsolidus conditions, assisted by the presence of fluid/melt (e.g., [Boullier and Nicolas](#)  
208 [1975](#); [Drury and van Roermund 1989](#); [Vauchez et al. 2005](#)). Existing observations  
209 suggest that, although differing in growth environments (for instance, free space,  
210 available constituents, etc.), olivine growth under subsolidus conditions is also  
211 anisotropic (e.g., [Drury and van Roermund 1989](#); [Tommasi et al. 2008](#); [Miyazaki et al.](#)  
212 [2013](#)). Nevertheless, the growth habits of olivine are still poorly constrained. For

213 example, [Drury and van Roermund \(1989\)](#) found that the olivine crystal faces formed at  
214 subsolidus conditions are similar to those formed from a fluid phase, but they didn't  
215 investigate the specific growth habits of olivine. [Tommasi et al. \(2008\)](#) suggested that the  
216 Tok tabular harzburgites experienced selective grain growth during static recrystallization  
217 in the presence of melt, forming tabular olivine crystals flattening normal to the

218 [010]-axis and a strong [010]-fiber CPO. These observations suggest clearly that the  
219 [010]-axis is the shortest axis during static recrystallization of olivine in the Tok  
220 harzburgites, but it is still not clear from their results the aspect ratios between the [100]-  
221 and the [010]-axes. One way to determine the growth habits of euhedral, tabular olivine  
222 in natural peridotites is via the CPOs of these olivines by assuming that the CPOs are  
223 results of anisotropic growth of olivine. Under such an assumption, the A-type olivine  
224 CPOs reported in [Drury and van Roermund \(1989\)](#) and [Vauchez et al. \(2005\)](#) for tabular  
225 olivines imply that olivine grows preferentially with  $a > c > b$ , while the [010]-fiber  
226 CPOs reported by [Tommasi et al. \(2008\)](#) imply that olivine grows preferentially with  $a >$   
227  $c > b$  or  $c > a > b$ . Therefore, no exclusive conclusions on the growth habits of olivine  
228 under subsolidus conditions can be reached based on the current dataset. Moreover, the  
229 driving forces (reduction of interface free energy, reduction of stored energy associated  
230 with defects, or chemically induced grain boundary migration; see [Evans et al. 2001](#)) for  
231 grain growth/recrystallization of olivine are generally not clearly stated by previous  
232 studies (e.g., [Vauchez et al. 2005](#); [Vonlanthen et al. 2006](#); [Tommasi et al. 2008](#)), which  
233 makes the discussion on the controlling factors for anisotropic growth of olivine very  
234 difficult. In summary, the general picture for the growth habits of olivine in the presence  
235 of fluid/melt is currently far from complete due to data limitation except that the  
236 [010]-axis seems to be generally the shortest axis (e.g., [Tommasi et al. 2008](#)). On the  
237 other hand, for olivine growth in the absence of fluid/melt, [Miyazaki et al. \(2013\)](#)  
238 concluded based on experimental observations that olivine grows anisotropically with  $a >$

239  $b > c$  at  $T < 1250$  °C and with  $a > c > b$  at  $T > 1250$  °C, pointing out a different growth  
240 habit of olivine under subsolidus conditions from what we observed in Hawaii basalts  
241 under supersolidus conditions.

242

243

### IMPLICATIONS

244 Olivine crystallization under supersolidus conditions is common in the Earth's upper  
245 mantle where there is extensive partial melting, such as the mantle wedge above a  
246 subducting slab and the upper mantle beneath a middle ocean ridge, or a reactive melt  
247 percolation zone with the crystallization of subhedral to euhedral olivine at the expense  
248 of pyroxene (e.g., [Vauchez et al., 2005](#); [Tommasi et al., 2008](#)). Anisotropic olivine growth  
249 is expected to play an important role in fabric development under these circumstances.  
250 Therefore, the crystallization habit of olivine has an immediate influence on olivine  
251 fabric development in a magma chamber or channel. In a static magma chamber,  
252 gravitational sinking (similar to a compaction deformation) will cause the largest crystal  
253 face, i.e., the (010)-plane, of most crystallized olivine grains to align horizontally to form  
254 the magmatic foliation, whereas the [100] and the [001] axes will spread in the foliation  
255 plane, thus shaping an AG-type or a [010]-fiber olivine fabric ([Fig. 5a](#)). In a magma  
256 channel undergoing continuous pure shear viscous flow, it is reasonable to expect that the  
257 longest axis, i.e., the [001]-axis, will orient preferentially parallel to the flow direction,  
258 whereas the [010]-axis will be aligned normal to the flow plane ([Fig. 5b](#)). Crystallized  
259 olivine grains in a magma channel undergoing simple shear viscous flow will form a

260 similar olivine CPO with the [001] axes and the (010) planes oblique in an acute angle to  
261 the shear direction and the shear plane, respectively (Fig. 5c). This fabric pattern is  
262 analogous to the B-type olivine fabric that had been demonstrated experimentally to be  
263 formed by the (010)[001] slip during dislocation creep of olivine under hydrous or  
264 high-pressure conditions (e.g., Jung and Karato 2001; Jung et al. 2009; Ohuchi et al.  
265 2011). The B-type olivine fabric is crucial to the interpretation of upper mantle structure  
266 and dynamics, yet many mechanisms have been proposed to interpret its origin under  
267 different geological settings, including slip system transition induced by high water  
268 contents or high pressure (e.g., Jung and Karato, 2001; Ohuchi et al., 2011), strain  
269 partitioning between melt-rich and melt-deficient zones (Holtzman et al., 2003), diffusion  
270 creep controlled by interface reaction (Sundberg and Cooper, 2008), grain boundary  
271 sliding (Précigout and Hirth, 2014), static topotactic growth of olivine after  
272 high-temperature breakdown of foliated serpentinite (Nagaya et al., 2014), and  
273 low-temperature dissolution creep (Liu et al., 2018). Our results provide another  
274 alternative explanation for the formation of B-type olivine fabric in nature, i.e., oriented  
275 alignment of anisometric olivine grains resulting from anisotropic growth. A similar  
276 hypothesis has been recently proposed for the asymmetric B-type olivine CPO in the  
277 Songshugou dunite cumulates (Cao et al., 2017). Getsinger and Hirth (2014) showed that  
278 oriented grain growth and rigid grain rotation during diffusion creep led to significant  
279 SPO (shape preferred orientation) and CPO in amphibole growing from a basalt. Because  
280 anisotropic grain growth usually leads to a significant SPO, which helps to stabilize grain

281 rotation and prevent destruction of former CPOs (Wheeler, 2009, 2010), the resultant  
282 CPOs could be preserved during later diffusion-mediated creep (diffusion creep or  
283 diffusion-accommodated GBS) deformation. This mechanism can explain CPOs of  
284 olivine in natural peridotites deformed by the diffusion-mediated creep (e.g., Précigout  
285 and Hirth, 2014; Cao et al., 2017).

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- 437

438 **Figure Captions**

439 **Fig. 1.** Photomicrographs showing the typical euhedral and anisometric olivine (Ol)  
440 phenocrysts within the Hawaii basalts. Note the inclusions of plagioclase (Pl) (c) and  
441 basaltic glass (d) inside the olivine phenocrysts. Scale bars 1 mm.

442

443 **Fig. 2.** Co-variation diagram of MgO versus FeO for olivine (Ol). Data source: Ol  
444 phenocrysts in the Hawaii basalts (this study); Ol grains in peridotite xenoliths entrained  
445 by the Hawaii basalts ([Claque 1988](#); [Sen 1988](#)).

446

447 **Fig. 3.** Photomicrographs showing the euhedral olivine phenocrysts and their  
448 corresponding crystal orientations determined by EBSD. The crystal faces for these  
449 olivines are also indicated. The olivines in a) and b) belong to the Group #1, while those  
450 in c) and d) belong to the Group #2 and Group #3, respectively.

451

452 **Fig. 4.** Histograms showing the long/short axis ratios for each group of olivine  
453 phenocrysts.

454

455 **Fig. 5.** Schematic olivine fabrics. a. The AG-type olivine fabric that can form by  
456 compaction or gravitational sinking of euhedral olivine in a magma chamber. b. The  
457 B-type olivine fabric that can form during pure shear magmatic flow. c. The B-type  
458 olivine CPO that can form in response to viscous magmatic flow in simple shear. The

459 foliation plane (S) is normal to the paper along the dashed line, while the lineation (L) is  
460 horizontal and trends along the dashed line.  
461



462

463

464 **Table 1**

465 Summary of EBSD analyses.

466

Planes parallel to the thin section surface	(100) (Group #1)		(010) (Group #2)		(001) (Group #3)	
	[010]	[001]	[100]	[001]	[100]	[010]
Grains	0	42	4	42	20	7

467

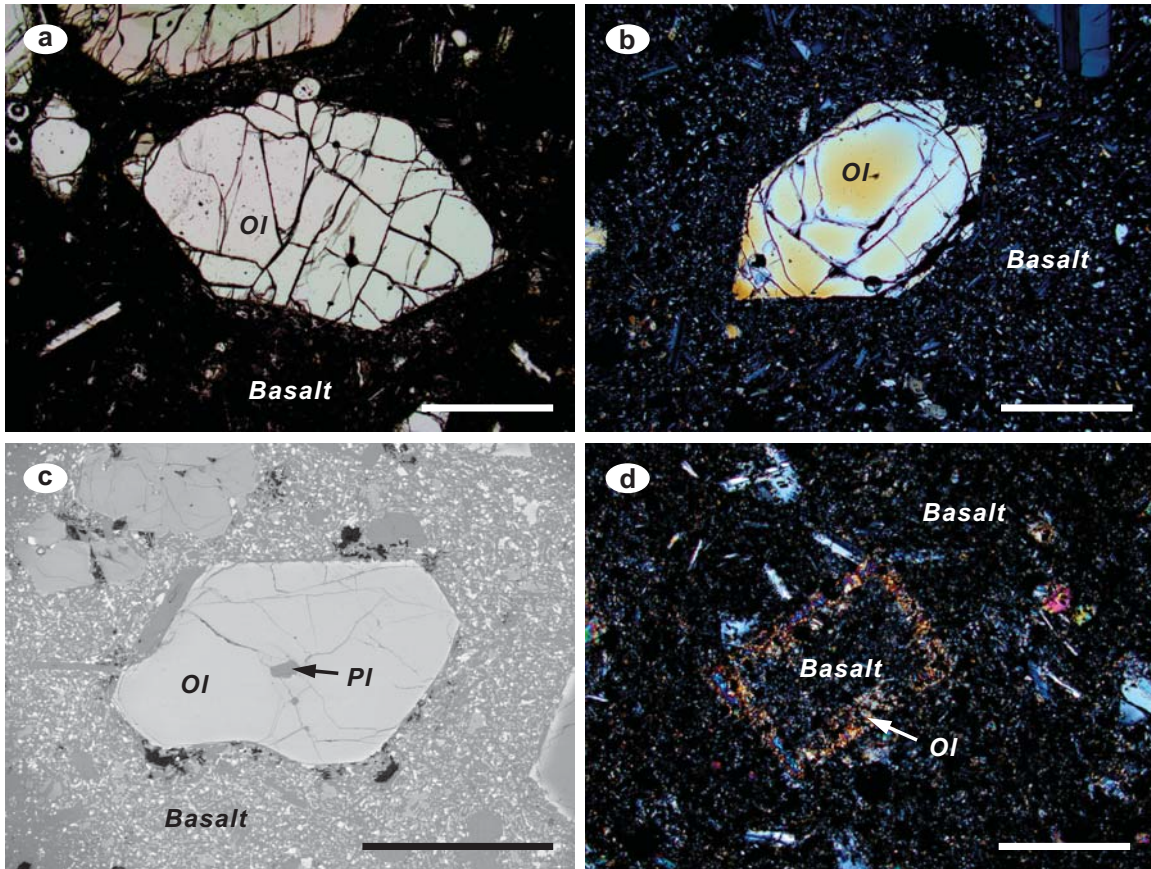


Fig. 1

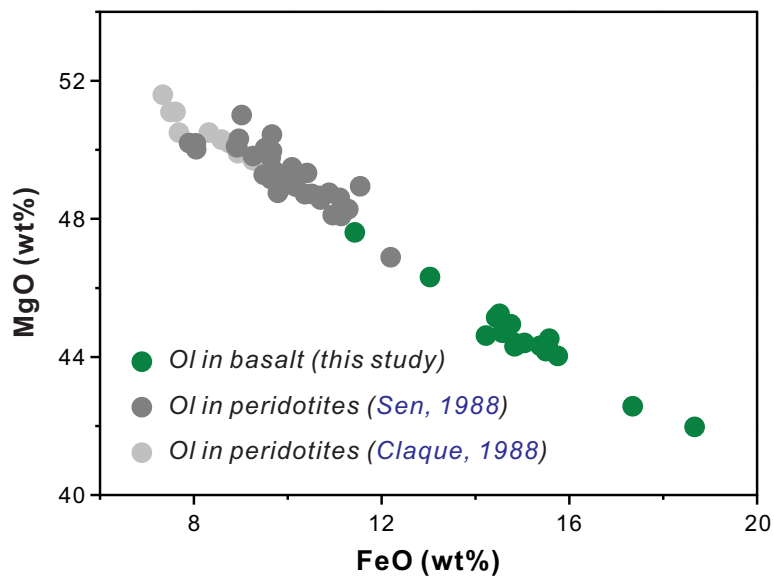


Fig. 2

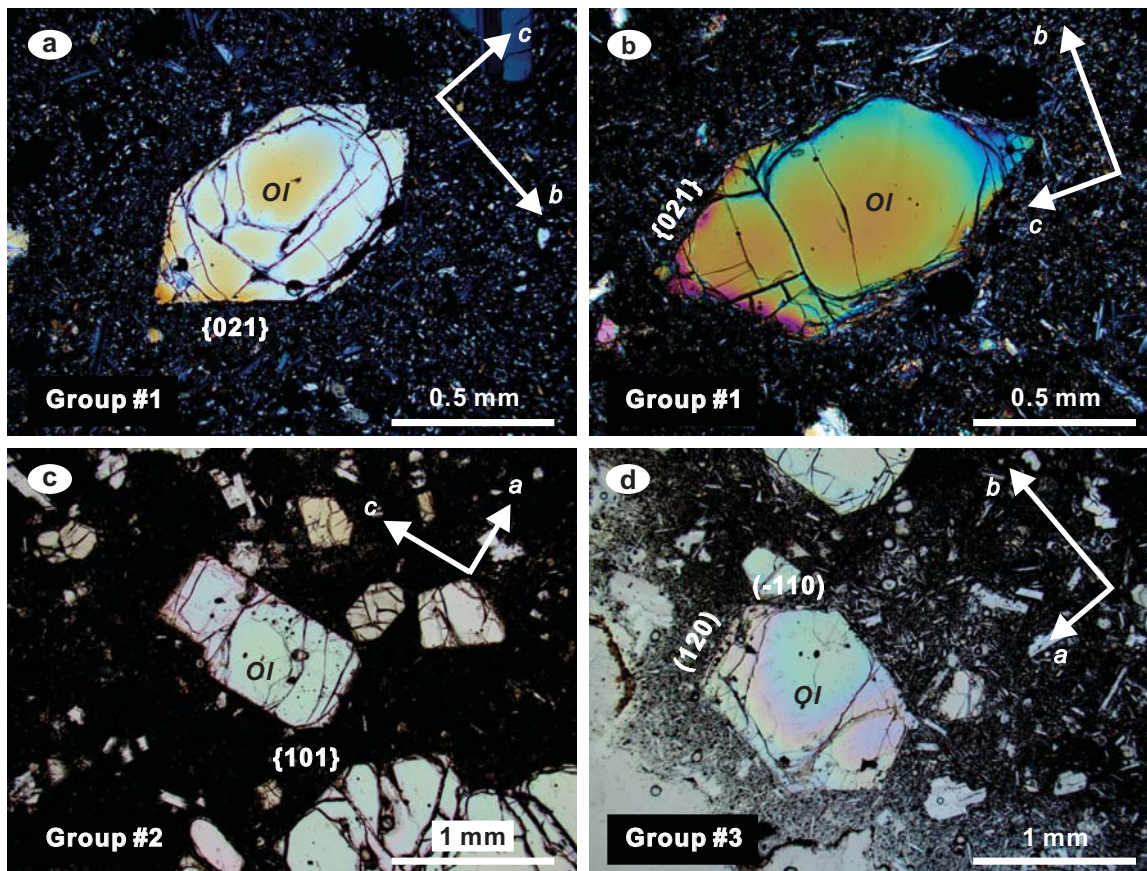


Fig. 3

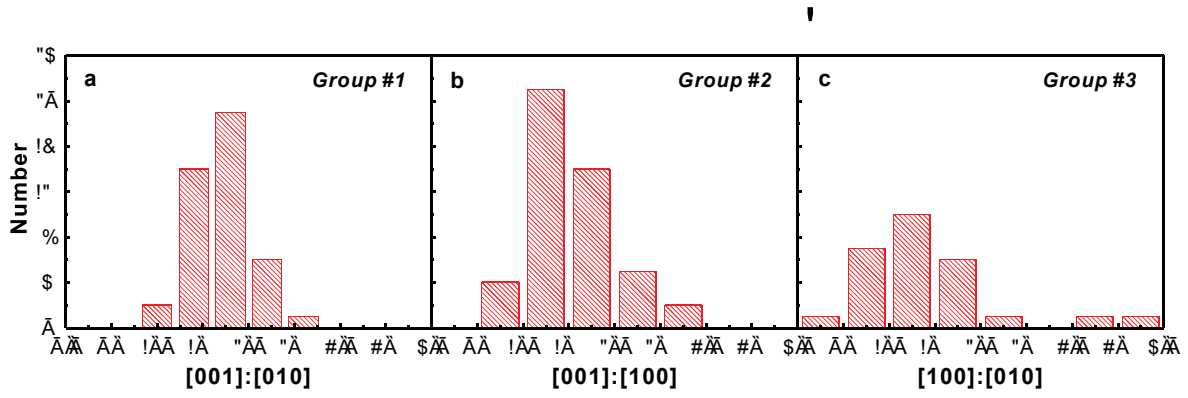


Fig. 4

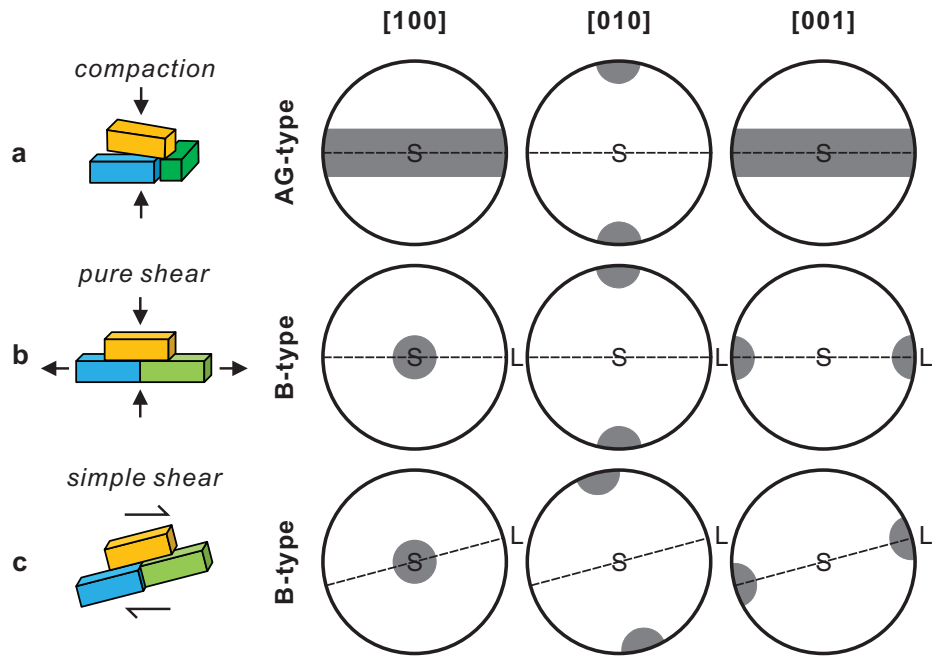


Fig. 5