

1 **Revision 3: Crystal Structure of CaSiO₃ Perovskite at 28–62 GPa and 300 K**
2 **under Quasi-hydrostatic Stress Conditions**

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14

15 **Abstract**

16 In order to find the thermodynamically stable crystal structure of CaSiO₃
17 perovskite (CaPv) at high pressure and 300 K, we have conducted synchrotron X-ray
18 diffraction (XRD) on thermally stress-annealed samples in a Ne pressure medium in the
19 diamond-anvil cell at 28-62 GPa. Rietveld refinements of the diffraction patterns are
20 significantly improved in fitting the positions and intensities of the split lines of CaPv if
21 the starting model is a tetragonal perovskite-type structure with the SiO₆ octahedral
22 rotation around the tetragonal *c*-axis. The result is in contrast with other previous

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23 experiments, but is consistent with first-principles calculations, reconciling the
24 discrepancy between computations and experiments on the crystal structure of CaPv. We
25 attribute the observed difference to the formation of the thermodynamically more stable
26 phase under improved stress conditions in our experiments. Our fitting shows that the
27 bulk modulus of Ca-Pv is 223 ± 6 GPa when its pressure derivative fixed to 4, which is
28 also consistent with first-principles calculations. The previous observations of the
29 diffraction patterns of CaPv inconsistent with the first-principles studies could be due to
30 the formation of a metastable crystal structure of CaPv under elevated deviatoric stresses.

31

32 **Keywords:** CaSiO₃ perovskite; crystal structure; hydrostatic stress; X-ray diffraction;
33 equation of state.

34

35

Introduction

36 Perovskite-structured CaSiO₃ (hereafter CaPv) is the third most abundant mineral
37 (5–7 wt%) in the pyrolitic model composition of the Earth's mantle (Kesson et al. 1998;
38 Murakami et al. 2005; Ricolleau et al. 2009). CaPv is stable over a wide depth range from
39 the transition zone to the lowermost mantle (Shim et al. 2000; Ohta et al. 2008). In mid-
40 oceanic ridge basalt (MORB) in subducting slabs at lower-mantle pressures, CaPv
41 represents 23 wt% of the mineralogy, comparable to the amount of bridgmanite, that is 35
42 wt% (Hirose et al. 2005; Ricolleau et al. 2010, Grocholski et al. 2012).

43 Existing studies are in agreement in that CaPv has a phase transition from a non-
44 cubic to a cubic perovskite-type structure with heating ($T > 500$ K) (Komabayashi et al.
45 2007; Noguchi et al. 2013; Sun et al. 2016) at pressures (P) expected for the lower

46 mantle. However, there remain significant discrepancies between computations and
47 experiments on the crystal structure of the lower-temperature form of CaPv and its
48 compressibility. In the 1980s, it was believed that CaPv adopts the cubic perovskite
49 structure ($Pm\bar{3}m$) at 300 K and lower mantle pressures from energy-dispersive X-ray
50 diffraction (Liu and Ringwood 1975; Mao et al. 1989; Tamai and Yagi 1989; Wang et al.
51 1996). *Ab initio* calculations first predicted a lower symmetry structure for CaPv.
52 Stixrude et al. (1996) found the instability of the cubic structure with a SiO₆ octahedral
53 rotation of 7 degrees towards the tetragonal $I4/mcm$ subgroup. Caracas et al. (2005)
54 performed a more exhaustive search for the ground state of CaPv and tetragonal $I4/mcm$
55 was regarded as one of the most energetically competitive structures at low temperature
56 and high pressure.

57 Using high-resolution angle-dispersive diffraction technique, Shim et al. (2002)
58 found a non-cubic unit cell for CaPv at 20-46 GPa and 300 K. However, they found that
59 the positions and intensities of the split peaks (in particular the 200_{pc} peak, where the
60 subscript “pc” represents the Miller indices referred to the cubic aristotype unit cell of
61 perovskite) are more consistent with a longer *a*-axis than *c*-axis ($c_{pc}/a_{pc} < 1$). The ratio is
62 opposite to what is expected for the crystal structures predicted by computations, which
63 predicted the octahedral rotation ($c_{pc}/a_{pc} > 1$) as an origin of the non-cubic structure.
64 Instead, the experimental observation is more straightforward to understand if one of the
65 three axes of the SiO₆ octahedra is shorter than the other two (therefore, direct distortion
66 of the octahedra). Later XRD studies have shown a similar pattern of the 200_{pc} peak
67 splitting as Shim et al. (2002) in CaSiO₃ (Komabayashi et al. 2007; Sun et al. 2016),
68 peridotite (Ono et al. 2004) and MORB (Hirose et al. 2005) compositions at room

69 (ambient) T and high P . A multi-anvil press experiment identified an orthorhombic CaPv
70 at P - T up to 20 GPa and 1600 K (Uchida et al. 2009). However, such a low symmetry
71 phase in CaPv has not been identified at higher pressures. Therefore, the discrepancy
72 between theory and experiments remain unresolved for the stable crystal structure of
73 CaPv at 300 K and high pressure.

74 The rotations of the rigid octahedra are responsible for the low-symmetry crystal
75 structures of a wide range of perovskite-structured materials (Glazer 1975; Woodward
76 1997). For example, the rotation of the framework octahedra in a cubic perovskite
77 structure can result in a number of lower symmetry structures, including three tetragonal
78 subgroups, $I4/mcm$, $P4/mbm$, and $I4/mmm$. Rotation along the c -axis can lead to the
79 $I4/mcm$ and $P4/mbm$ structures with $c_{pc}/a_{pc} > 1$, as the rotation reduces the dimension in
80 the (110) plane. In the case of the $I4/mmm$ structure, rotation around the a axis can result
81 in $c_{pc}/a_{pc} < 1$. This symmetry also allows for octahedral distortion during the rotation.
82 Because of this relationship between the axial ratio and the crystal structure, Shim et al.
83 (2002) found that the observed $c_{pc}/a_{pc} < 1$ is difficult to explain with the octahedral
84 rotation alone and some degree of the octahedral distortion should be invoked. Therefore,
85 they use $P4/mmm$, which is purely based on the octahedral distortion along the c -axis, as
86 a proxy to explore the possibility for the distorted octahedra case. In $P4/mmm$, the Si-O
87 bonds along the c -axis are shorter than the bond along the a -axes in a tetragonal unit cell.
88 From the strong Si-O bonding, it is surprising that CaPv has a lower symmetry structure
89 through the direct distortion of the octahedra, which would require more energy than the
90 rigid octahedral rotation. Also, $P4/mmm$ structured perovskites are much less common
91 than many other subgroups (Comyn et al. 2004).

92 It is possible that the measured crystal structure of CaPv is sensitive to the deviatoric
93 stresses in diamond-anvil cell (DAC). Most experimental studies on CaPv have been
94 conducted in DACs where pressure medium play an important role for the stress
95 conditions of the sample. Earlier studies used Ar (Shim et al. 2002) or MgO (Kurashina
96 et al. 2004; Komabayashi et al. 2007) as pressure media, but some were conducted even
97 without a pressure medium (Ono et al. 2004, 2005). All these studies have observed non-
98 cubic CaPv (*P4/mmm*) upon temperature quench. Although laser annealing reduces the
99 deviatoric stress in a DAC, 40–80% of stress can be reintroduced upon temperature
100 quench in a MgO pressure medium (Kavner and Duffy 2001). Because computational
101 studies have shown that differences of free energy among competing crystal structures
102 are very small (Caracas et al. 2005), it is important to reduce the effects from non-
103 equilibrium conditions, such as deviatoric stresses, as much as possible. In this study, we
104 use Ne as a pressure medium combined with laser annealing to 1,700-2,300 K in order to
105 measure the stable crystal structure of CaPv under quasi-hydrostatic stress conditions at
106 300 K and high pressures.

107

108 **Experimental procedures**

109 We synthesized a CaSiO₃ glass using the laser levitation method (Tangeman et al.
110 2001) for a starting material. We measured the composition of the glass using wavelength
111 dispersive X-ray spectrometry (WDS) in an electron probe microanalyzer (EPMA; JXA-
112 8530F) at Arizona State University. The glass starting material has a Ca/Si molar ratio of
113 0.993±0.009 and is compositionally homogeneous.

114 The glass starting material was mixed with 10 wt% platinum (99.98% purity with
115 a grain size less than 3 μm ; Alfa Aesar). We used platinum as an internal pressure
116 standard (Ye et al. 2017) and a laser absorber. The sample+Pt powder mixture was
117 pressed into a foil and loaded into a chamber drilled in a Re gasket with the diamond
118 anvils mounted on symmetric-type DAC. We used 400 μm culet diamond anvils for a
119 peak pressure less than 35 GPa, and 200 μm culet diamond anvils for experiments above
120 this pressure. We loaded Ne using the gas-loading system at the GSECARS sector as a
121 pressure transmitting medium and thermal insulator (Rivers et al. 2008). Small CaSiO_3
122 glass grains (diameter < 10 μm and thickness < 2 μm) were placed below and above the
123 sample foil, in order to allow Ne to flow above and below the sample foil in the sample
124 chamber during the gas loading. This setup ensures that Ne forms layers between the
125 sample and the diamond anvils for thermal insulation. In one experiment, we loaded the
126 starting material without any medium, in order to examine the effects of deviatoric stress
127 and thermal gradients.

128 We measured XRD patterns at in-situ high P - T in the laser heated DAC at beamlines
129 13ID-D in the GSECARS sector (Prakapenka et al. 2008) and 16ID-B in the HPCAT
130 sector (Meng et al. 2006) at the Advanced Photon Source. Monochromatic X-ray beams
131 with beam sizes of 3×4 and 5×6 μm^2 at GSECARS and HPCAT, respectively, were
132 focused on the sample and the diffraction patterns were collected using a Mar-CCD
133 detector. We focused near-infrared laser beams (1- μm wavelength) on the sample with a
134 hot spot size of 20-25 μm in diameter and aligned them coaxially with the X-ray beam
135 (Meng et al. 2006; Prakapenka et al. 2008). We measured the temperature of the heated

136 spots by fitting Planck equation to the thermal radiation spectra from both sides of the
137 sample after subtracting the effects of the optical systems.

138 Before laser heating, we confirmed the amorphous state of CaSiO₃ glass through
139 XRD. In a typical experiment, we compressed the sample to target pressures at 300 K,
140 and then heated a spot on the sample foil for approximately 5 minutes. We increased the
141 laser power to the target temperature (1,700-2,300 K) within 1 minute to avoid the
142 possible synthesis of metastable phases at low temperatures. We measured several XRD
143 patterns during and after heating. CaPv forms within 1 minute upon heating above 1,700
144 K. The diffraction pattern of CaPv remains essentially the same during heating after the
145 formation. After temperature quench, we acquired diffraction patterns on the stress
146 annealed spot at high pressure.

147 It is challenging to detect the subtle peak splitting in a quasi-hydrostatic pressure
148 medium. Therefore, we have optimized X-ray energy and detector-sample distance to
149 achieve the maximum resolution of the systems we used. Our maximum resolution was
150 achieved at a detector distance of 249.92 mm for an X-ray energy of 25 keV. The use of a
151 low-energy X-ray beam limited the measurable *d*-spacing to >1.557 Å, but enabled us to
152 detect the peak splitting of the 200*pc* line from CaPv.

153 We used the software package Dioptas (Prescher and Prakapenka 2015) to integrate
154 the 2D diffraction images to 1D diffraction patterns. We identified phases and measured
155 full width at half maximum (FWHM) of XRD peaks by the software PeakPo (Shim
156 2017). The data obtained from the GSECARS and HPCAT beamlines agreed well with
157 each other for the peak positions and intensities of CaPv. We performed Rietveld
158 refinements of the XRD patterns using the general structure analysis system (GSAS)

159 software combined with EXPGUI (Toby 2001). We refined the phase fractions first, then
160 the atomic positions, the lattice parameters and the spherical harmonic terms for the
161 preferred orientation. After reaching a good visual fit, we refined all the parameters
162 together to minimize the fit residuals.

163 A few different possible crystal structures have been proposed for CaPv at low
164 temperature and high pressures (Stixrude et al. 1996, 2007; Akber-Knutson et al. 2002;
165 Shim et al. 2002; Caracas et al. 2005; Jung and Oganov 2005; Adams and Oganov 2006;
166 Caracas and Wentzcovitch 2006; Li et al. 2006; Uchida et al. 2009). The focus of our
167 experiments presented here is: should the direct distortion of the SiO₆ octahedra (for
168 example, *P4/mmm* in Shim et al. 2002) be invoked in order to explain the diffraction
169 patterns from the CaPv phase stable at high pressure and 300 K? Or is the octahedral
170 rotation (for example, *I4/mcm* in Stixrude et al. 2007) sufficient as has been shown in the
171 computation and seen from a wide range of perovskite-structured materials?

172 In order to demonstrate the differences in powder diffraction patterns, we
173 conducted calculations using the crystal structures proposed for CaPv (Figure 1) (Shim et
174 al. 2002; Caracas and Wentzcovitch 2006). The cubic 200 diffraction line of CaPv splits
175 the most in both tetragonal diffraction patterns (*P4/mmm* and *I4/mcm*). However, the
176 relative intensities of the two split lines from the 200_{pc} peaks differ depending on the
177 space groups. For example, the lower-angle diffraction line (004 in Figure 1) has a lower
178 intensity in *I4/mcm*, while it has a higher intensity in the *P4/mmm* structure (200 in
179 Figure 1). More importantly, the higher 2θ-angle split line of 200_{pc} existed at a similar
180 angle as the original cubic 200 line for the octahedral rotation cases. However, for the
181 octahedral distortion case, it is the lower angle split line that exists at a similar 2θ angle

182 as the cubic 200 line (*P4/mmm*). These peak position differences are related to the fact
183 that through the rotations of the SiO₆ octahedra in the (001) plane, the c_{pc} axis becomes
184 longer than the a_{pc} axis in the *I4/mcm* structure, resulting in $c_{pc}/a_{pc} > 1$. Although the
185 *I4/mcm* structure does not result in $c_{pc}/a_{pc} < 1$, shorter Si-O bonding distances in the
186 octahedra can make the c_{pc}/a_{pc} ratio is smaller than 1 in the *P4/mmm* space (Shim et al.
187 2002). Therefore, for $c_{pc}/a_{pc} < 1$, as observed by Shim et al. (2002), it is difficult to
188 explain the crystal structure of CaPv with the octahedral rotation alone and some degree
189 of octahedral distortion (shorter Si-O bonding along the *c* axis) needs to be invoked.
190 According to our calculation, the cubic 110, 211, and 220 diffraction peaks also show
191 splitting but with smaller degrees in both *I4/mcm* and *P4/mmm* structures. Therefore, the
192 splitting of these latter lines is more difficult to resolve.

193 While the peak splitting of the relatively intense lines are helpful to distinguish the
194 octahedral distortion (*P4/mmm*) versus octahedral rotation (*I4/mcm* and *P4/mbm*), the
195 presence and absence of some lines could further help to distinguish these two
196 possibilities. Also, those weak diffraction features would help to further distinguish a
197 range of different symmetries possible for the octahedral rotation case, such as *I4/mcm*,
198 *P4/mbm*, and *I4/mmm*. However, the weak diffraction lines are difficult to detect
199 convincingly through powder diffraction in DAC. Although our 2D diffraction images
200 show some weak spots in those angular ranges, it is difficult to unambiguously assign
201 them to the diffraction peaks of CaPv because they could result also from other materials
202 along the beam path, such as beam stoppers, X-ray optics, and X-ray guides. Such task
203 can be better addressed through multi-grain diffraction (Margulies et al. 2001; Nisr et al.
204 2012) or single crystal diffraction techniques. Therefore, for our powder diffraction study

205 presented here, we focus on peak splitting and their implications for the octahedral
206 rotation versus octahedral distortion.

207

208 **Results and Discussion**

209 During laser heating, we did not observe any splitting of the CaPv diffraction lines
210 at temperatures between 1,700 K and 2,300 K up to our highest resolution XRD setup (X-
211 ray wavelength = 0.4959 Å and sample-to-detector distance = 197.52 mm) (Figure 2).
212 The observations support the stability of the cubic structure at high temperatures as
213 proposed by previous studies (Ono et al. 2004; Stixrude et al. 2007; Kurashina et al.
214 2004).

215 After temperature quench, we observed the clear splitting of the CaPv 200_{pc} line
216 (Figure 2a) at pressures between 28 and 62 GPa. In the 2D diffraction images (Figure 2c),
217 we also observed diffraction spots split into two. This shows that the symmetry change
218 occurs during temperature quench.

219 Unlike previous experiments (Shim et al. 2002; Kurashina et al. 2004; Komabayashi
220 et al. 2007), we found that the higher-angle diffraction line of the split 200_{pc} line has a
221 greater intensity throughout our studied pressure range at 300 K in both 2D and 1D
222 diffraction data (Figures 2a and c). As shown in Figures 1b and 2a, our new results is
223 more consistent with the XRD patterns expected for the tetragonal structures induced by
224 octahedral rotation (such as, $I4/mcm$), which has been proposed by first-principles
225 calculations (e.g., Stixrude et al. 1996, 2007; Caracas et al. 2005), than octahedral
226 deformation ($P4/mmm$), which has been indicated by previous XRD experiments (e.g.,
227 Shim et al. 2002; Kurashina et al. 2004; Komabayashi et al. 2007).

228 Diffraction intensities can be strongly affected by preferred orientation in powder
229 samples in the DAC. Preferred orientation can be caused by such effects as plastic flow
230 and preferential twinning, in response to the stress field. These effects can be reduced
231 when deviatoric stress is minimized. We performed Rietveld refinements with fitting for
232 the preferred orientation. We refined the diffraction patterns measured under our best
233 resolution setup (Figure 3). These refinements were performed with two different crystal
234 structures for CaPv: *I4/mcm* (the octahedral rotation, Stixrude et al. 1996) and *P4/mmm*
235 (the octahedral deformation, Shim et al. 2002) (Table 1). In order to facilitate the
236 comparison, we used the same parameters for the background in the fitting with the
237 *I4/mcm* and *P4/mmm* starting structural models of CaPv.

238 Even when we fitted up to 14 parameters in a spherical harmonic function, we found
239 significant mismatch between measured and calculated intensities of 200_{pc} and 211_{pc} for
240 the octahedral distortion case (Figure 3b). We found significantly better fits for the
241 octahedral rotation case. The improvement is mainly because the octahedral rotation
242 requires the higher- 2θ split line to be closer in peak position to that of the cubic 200_{pc}
243 line. In contrast, the octahedral distortion requires the lower- 2θ split line to be closer in
244 peak position to that of the cubic 200_{pc} line. Therefore, if the observed diffraction is more
245 consistent with the octahedral rotation, the fit with the octahedral distortion would result
246 in large positive residue at the lower-angle side of the 200_{pc} split lines. The expected
247 positive residue can be indeed found in Figure 3b. In other words, the intensity residual
248 for the 200_{pc} line is significantly impacted by the relative position of the split lines (and
249 therefore the c/a ratio), which is not sensitive to the preferred orientation, as well as the
250 intensities of the split lines, which is sensitive to the preferred orientation.

251 From our Rietveld refinements, we obtained the c_{pc}/a_{pc} of CaPv (Figure 4). The
252 c_{pc}/a_{pc} remains constant within the pressure range we studied. Therefore, we averaged
253 c_{pc}/a_{pc} over our pressure range and the value is 1.0054 ± 0.0005 . Shim et al. (2002) showed
254 that the octahedral rotation will lead to a c_{pc}/a_{pc} higher than 1, while the ratio can be
255 either greater or smaller than 1 for the octahedral distortion depending on the shape of the
256 SiO_6 octahedra. Shim et al. (2002) reported 0.995, which cannot be explained by the
257 octahedral rotation alone but requires octahedral distortion. However, in our study, we
258 found that the c_{pc}/a_{pc} value is higher than 1, suggesting octahedral rotation as the main
259 source of symmetry lowering in CaPv. In other words, it is not necessary to invoke direct
260 deformation of the octahedra, which is less common among perovskite-structured
261 materials (Woodward et al. 1997).

262 Based on *ab initio* calculation, Stixrude et al. (2007) suggested an increase in c_{pc}/a_{pc}
263 from 1.013 to 1.020 with an increase in pressure from 1 bar to 140 GPa at 0 K from the
264 octahedral rotation. Because our pressure range is smaller, such a small increase per unit
265 of pressure would be difficult to resolve from our data. On the other hand, because the
266 value from the *ab initio* calculation is performed for 0 K, thermal effects could further
267 decrease c_{pc}/a_{pc} . Stixrude et al. (2007) estimated that the value would decrease to 1.0094
268 at 25.2 GPa and 300 K, which is more similar to our new results.

269 As indicated by Caracas et al. (2006), several competing structures can explain
270 c_{pc}/a_{pc} greater than 1 based on octahedral rotation, such as $I4/mcm$, $P4/mbm$ or $I4/mmm$,
271 all of which cause the elongation of the a -axes. Because of the subtle differences in their
272 expected diffraction patterns as discussed earlier and the limitation of power XRD, the
273 space group of the CaPv at 300 K and high pressure remains uncertain. The space group

274 *I4/mcm* is a reasonable option because it is known to occur as an intermediate distortion
275 between the aristotype $Pm\bar{3}m$ and the common *Pbnm* orthorhombic distortion, including
276 for example in the prototype compound CaTiO_3 (Yashima and Ali, 2009).

277 Using our Rietveld results with the *I4/mcm* structure, we fit the pressure–volume
278 data at 300 K to the Vinet equation (Vinet et al. 1989). Using the Pt scale of Ye et al.
279 (2017), we obtained an isothermal bulk modulus of 223(6) GPa (Table 2). In the fitting,
280 we fixed the pressure derivative of the isothermal bulk modulus (B_0') to 4.0, while we
281 also fit the volume at 1 bar and 300 K (V_0), because CaPv is unstable at these conditions
282 and therefore V_0 is unknown. From the fitting we obtained $V_0 = 46.3(1) \text{ \AA}^3$.

283 Former studies have used the V_0 value projected through the extrapolation of CaPv
284 equation of state (EOS) measured between 0.59 and 10.07 GPa by Wang et al. (1996).
285 We also conducted fitting for the bulk modulus with V_0 fixed to the value by Wang et al.
286 (1996). As shown in (Figure 5c), we found severe misfit for our data points below 35
287 GPa as well as the data points from Shim et al. (2002) at the similar pressure range. We
288 also attempted fitting B_0 and B_0' by fixing V_0 to the value by Wang et al. (1996).
289 However, we found unusually low B_0' (2.3) for $B_0 = 290$ GPa with much larger fit
290 residual.

291 Although Wang et al. (1996) measured most of the data points of CaPv outside of
292 the stable pressures, all of our data points are measured within the stability field of CaPv.
293 For the case of bridgmanite, Wang et al. (1994) discussed that the phase may show
294 abnormal volume behavior and such effects appear to be particularly severe near the
295 conditions where the phase is amorphized. We note that CaPv amorphizes as pressure
296 approaches 1 bar. It cannot be ruled out that CaPv may have different crystal structure

297 outside of its stability field because of the subtle free energy differences among candidate
298 structures and subtle differences in their expected XRD patterns. We note that Wang et
299 al. (1996) used a cubic perovskite structure to fit their data.

300 Our bulk modulus value is consistent with the computational studies (Table 2)
301 (Chizmeshya et al. 1996; Akber-Knutson et al. 2002; Magyari-Köpe et al. 2002; Jung and
302 Oganov 2005). If the thermal effect is taken into consideration, the bulk modulus of CaPv
303 at 300 K based on the 0 K value from some theoretical studies (such as, Caracas et al.
304 2005; Stixrude et al. 2007; Kawai and Tsuchiya 2014) would become more similar to our
305 value and therefore are also consistent with our results.

306 We compare our data points with those from Shim et al. (2002) in Figure 5. In this
307 comparison, we use the Pt and Au pressure scales by Ye et al. (2017), which are inter-
308 calibrated and consistent with each other within ± 1 GPa. Our new EOS agrees well the
309 data from Shim et al. (2002) at 20-60 GPa. Two experimental studies have reported the
310 zero-pressure bulk modulus of CaPv projected from its cubic stability field (>700 K): 249
311 GPa from Sun et al. (2016) and 225 GPa from Noguchi et al. (2013). The latter is more
312 consistent with our value based on measurements performed directly at 300 K. The
313 discrepancy between two high temperature studies may be because they need
314 extrapolation and the bulk modulus can be also changed by the crystal structure. In
315 addition, in EOS fitting, strong correlation among fit parameters can result in differences,
316 and thermal gradients can contribute to the uncertainties.

317 In this study, we improve the stress conditions of experiments for CaPv by
318 combining a soft pressure medium (Ne) and laser annealing. The improved stress
319 conditions in our new experiments may be important for explaining the different results

320 we obtained in this study. In order to test this hypothesis, we conducted a laser-heated
321 DAC experiment for the same starting material but without a pressure medium. In this
322 “no medium” experiment, we found that the CaPv 200_{pc} diffraction peak is much broader
323 and much more asymmetric than runs with Ne both during and after laser heating (Figure
324 6). We note that even at 1550 K 200_{pc} shows splitting at 31 GPa. The peak width of CaPv
325 200_{pc} without a medium is a factor of 2.3 greater than that in a Ne pressure medium at
326 similar P - T conditions (Figure 6). We also find that other lines have asymmetric peak
327 shapes, which were not observed in our experiments with a Ne medium. Because thermal
328 annealing does not relieve the deviatoric stress completely (Kavner and Duffy 2001),
329 some degree of deviatoric stresses still remain in our “no-medium” experiments likely to
330 much higher degree than our Ne medium experiments. Therefore, we interpret that the
331 non-hydrostatic stresses cause the broadening and splitting of the 200_{pc} peak in the “no
332 medium” experiments.

333 We estimate pressures from individual diffraction lines of Pt. For the experiments
334 with a Ne medium and laser annealing, we obtain -0.5~2.5 GPa differences in pressure
335 between the 111 and 200 lines of Pt. The value is comparable to that reported for laser-
336 annealed Pt in a Ne medium by Ye et al. (2017) where they estimated unit-axial stress
337 component (t) of -0.5~1.0 GPa. Instead, for the no medium data point, we found much a
338 higher value for the pressure difference, 7.5 GPa. The elevated deviatoric stresses can
339 affect the crystal structure of CaPv. They may enhance the existing distortions, or even
340 induce different distortions and consequently stabilize a metastable structure.

341 It is also possible that the thermal gradients during the heating without an insulation
342 medium and the pressure gradients remaining after heating would produce a range of

343 CaPv crystals with different degree of distortions or even with different crystal structures
344 along the X-ray path. Indeed, the diffraction patterns are difficult to fit with the candidate
345 structures we tried for the data measured in a Ne medium.

346 High shear strength pressure-transmitting media can support larger amounts of
347 deviatoric stress in the DAC. Magnesium oxide (MgO), which is used for CaPv
348 (Komabayashi et al. 2007), is a highly nonhydrostatic medium, with a 10 GPa shear
349 strength at pressures between 60 and 100 GPa (Duffy et al. 1995). The shear strength of
350 Ar, which was used in Shim et al. (2002), is about 0.8 GPa at 65 GPa when the stress is
351 not thermally annealed, while thermal annealing at 700 K reduces the stress to 0.56 GPa
352 at 65 GPa (Marquardt et al. 2013). The pressure gradient using a Ne medium is less than
353 0.5 GPa at a pressure of 50 GPa without annealing (Klotz et al. 2009).

354 In this study, we used cold compressed foils of the powdered starting material.
355 Under a microscope, we found that the foil contains small pore spaces and cracks, which
356 should allow significant amounts of Ne to penetrate the grain boundaries. We also found
357 that the foils disintegrated during decompression to 1 bar. Therefore, we believe that the
358 soft Ne medium should have controlled the stress field of individual CaPv grains in our
359 experiment.

360 Although noble gas media in general support smaller deviatoric stresses, CaPv
361 appears to be still sensitive to such smaller deviatoric stress in the media. The different
362 crystal structures of CaPv, which are metastable in the stress-free state, can appear easily
363 with small deviatoric stress, because of their similar free energies (Caracas et al. 2005).

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Implication

367 Experimental and theoretical mineral physics studies on the properties of minerals
368 play an important role in understanding the structures and processes in the Earth's
369 interior. In theoretical studies, in order to compute the physical properties accurately, it
370 is important to know the stable crystal structures of the mantle phases. In this study, we
371 resolve long-standing discrepancy on the crystal structure of CaPv at 300 K and mantle-
372 related high pressure between theoretical and experimental studies. For experimentalists,
373 our finding of extreme sensitivity of CaPv on deviatoric stress highlights the importance
374 of conducting measurements on crystal structures and physical properties under better
375 controlled/characterized stress conditions relevant to those of the Earth's interior.

376

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388 for EOS are available in Table 1. Other data for this paper are available by contacting

389 SHDShim@asu.edu.

390 Table 1: Structural parameters of CaPv determined by Rietveld refinements. The
391 numbers in the parentheses are 2σ uncertainties for the last digit. Pressure was
392 calculated from the Pt scale by (Ye et al. 2017).

393

Run#	<i>P</i> (GPa)	Structure	Tilting angle (°)	<i>a</i> (Å)	<i>c</i> (Å)
71013	28.824	<i>I4/mcm</i>	8.0(4)	4.8896(3)	6.9677(8)
		<i>P4/mmm</i>	0	3.4690(2)	3.4496(4)
81021	29.138	<i>I4/mcm</i>	8.6(7)	4.8941(1)	6.9546(4)
		<i>P4/mmm</i>	0	3.4685(3)	3.4582(5)
81030	32.307	<i>I4/mcm</i>	9.4(2)	4.8805(2)	6.9611(8)
		<i>P4/mmm</i>	0	3.4610(2)	3.4477(4)
71050	41.972	<i>I4/mcm</i>	10.2(1)	4.8351(1)	6.8683(3)
		<i>P4/mmm</i>	0	3.4297(1)	3.4134(2)
81066	47.878	<i>I4/mcm</i>	7.8(1)	4.8070(2)	6.8354(3)
		<i>P4/mmm</i>	0	3.4119(1)	3.3922(3)
81074	53.884	<i>I4/mcm</i>	9.2(3)	4.7870(2)	6.8065(4)
		<i>P4/mmm</i>	0	3.3978(2)	3.3784(4)
71088	62.477	<i>I4/mcm</i>	8(1)	4.7535(1)	6.7571(3)
		<i>P4/mmm</i>	0	3.3737(1)	3.3546(2)

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395

396 Table 2: Compressibility of CaPv at high pressure and 300 K.

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Refs	V_0 (Å ³)	B_0 (GPa)	B_0'	Symm. ^a	Notes ^c
Experiments					
This work	46.3(1)	223(6)	4 ^d	Tet.	300 K, V, Pt-Y, Ne
Wang et al. (1996)	45.58(4)	232(8)	4.8 ^d	Cub.	300 K, BM, NaCl, NaCl
Shim et al. (2002)	45.58 ^d	255(5)	4 ^d	Tet.	300 K, BM, Pt-H, Ar
Shim et al. (2002)	45.58 ^d	259(8)	4 ^d	Tet.	300 K, V, Pt-Y, Ar
Ricolleau et al. (2009) ^b	45.60 ^d	244(1)	4 ^d	Tet.	300 K, BM, Au-F, Ne
Noguchi et al. (2013) ^c	45.80	225	4 ^d	Cub.	300 K, BM, Pt-F, NaCl
Sun et al. (2016) ^c	45.4(1)	249(4)	4 ^d	Cub.	300 K, BM, Pt-F, KCl/NaCl
Computation					
Kawai and Tsuchiya (2014)	46.17	203.5	4.76	Cub.	1000 K, V, LDA
Magyari-Köpe et al. (2002)	45.69	216	4.82	Orth.	300 K, V, LDA
Chizmeshya et al. (1996)	45.62	227	4.29	Cub.	300 K, BM, LDA
Caracas et al. (2005)	44.537	249	4.09	Tet.	0 K, BM, LDA
Stixrude et al. (2007)	44.00	252	4.1	Tet.	0 K, BM, LDA
Jung and Oganov (2005)	46.89	219	4.08	Tet.	0 K, BM, GGA
Akber-Knutson et al. (2002)	45.90	228	4.3	Orth.	300 K, BM, VIB

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399 Notes: The numbers in the parentheses are 2σ uncertainties for the last digit.

400 ^a Cub., Tet., and Orth. refer to cubic, tetragonal, and orthorhombic CaPv structures
 401 respectively.

402 ^b They used pyrolitic composition as a starting material.

403 ^c The values were projected from high temperature to 300 K and 1 bar.

404 ^d The values were fixed during EOS fitting.

405 ^e Temperature, EOS, Pressure scale or Computational methods, Pressure medium. V and
 406 BM refer to the Vinet and third order Birch-Murnaghan equations, respectively. Pt-H, Pt-
 407 Y refer to the Pt scales Holmes et al. (1989) and (Ye et al. 2017) respectively. Au-F and
 408 Pt-F refer to the Au and Pt scales by Fei et al. (2007). LDA (Local Density
 409 Approximation); GGA (Generalized Gradient Approximation); VIB (Variationally
 410 Induced Breathing.)

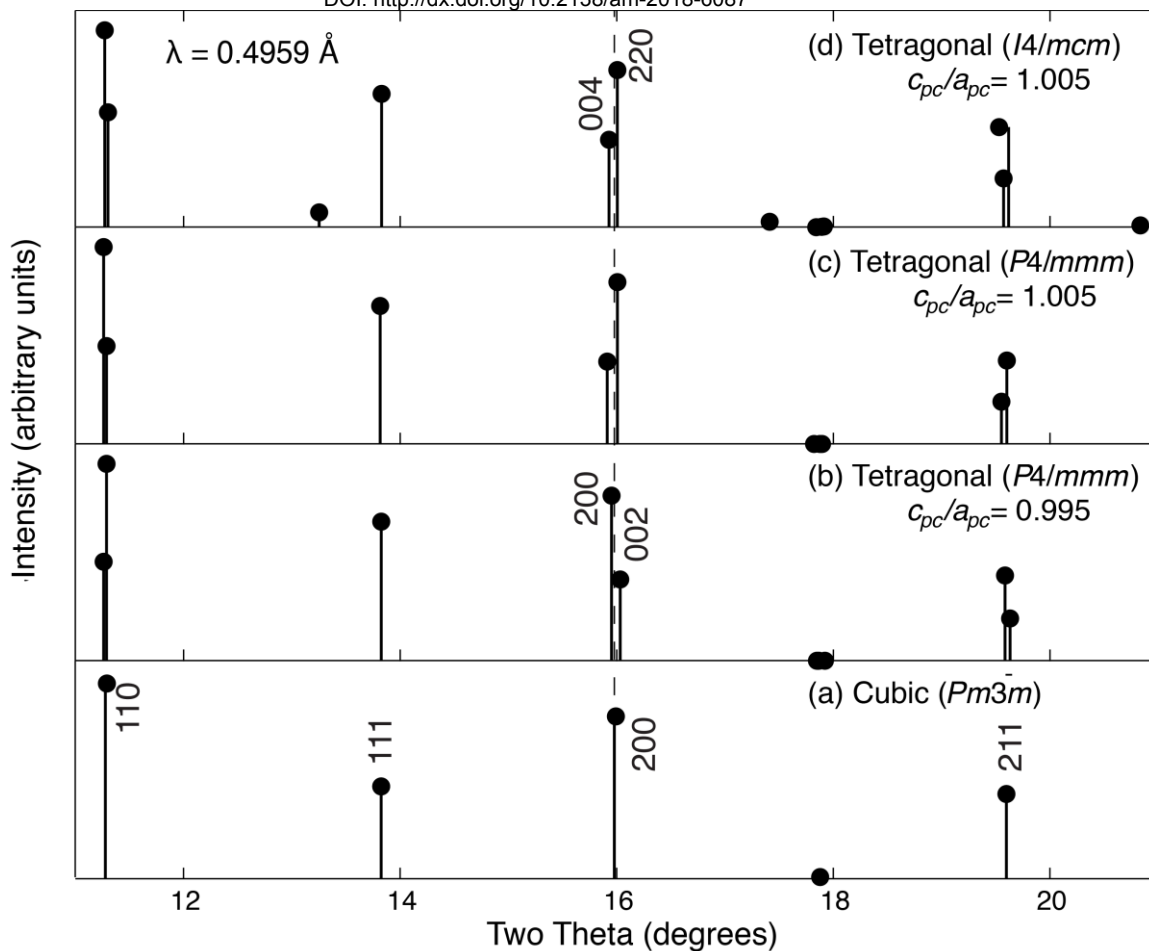
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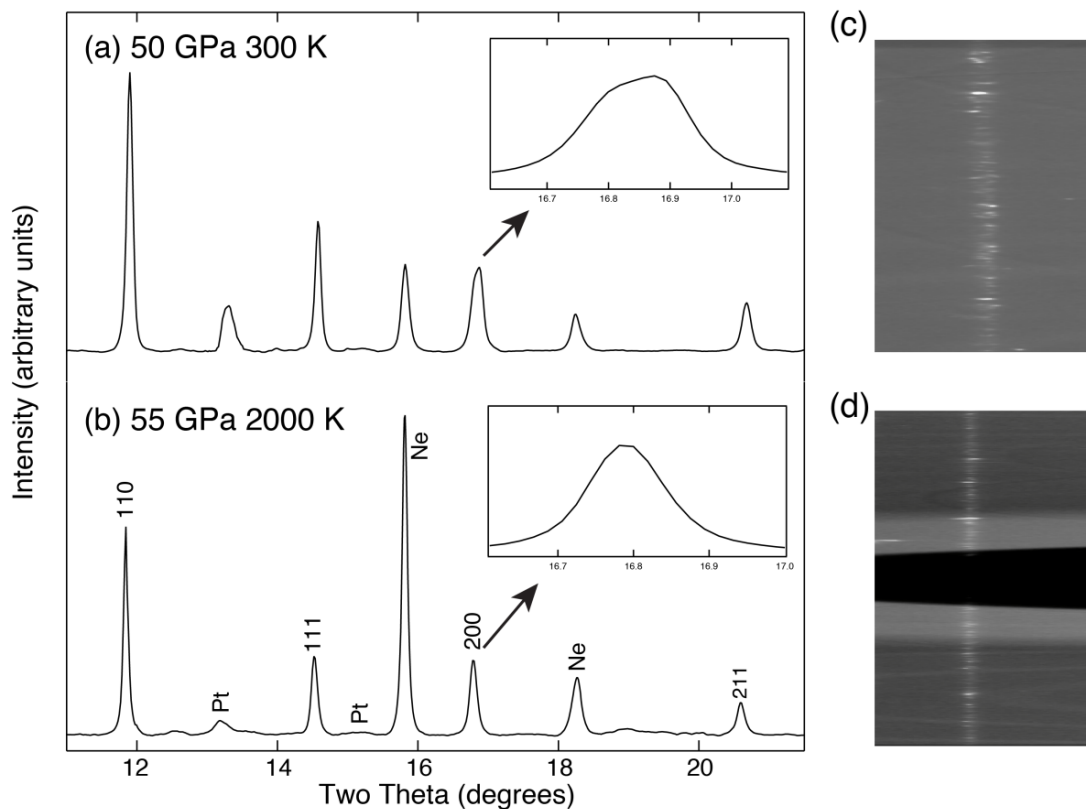
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 417 Figure 1: Calculated XRD patterns of CaPv in different crystal structures at 1 bar and 300
 418 K ($V_0 = 45.58 \text{ \AA}^3/Z$, X-ray wavelength = 0.4959 \AA): (a) cubic $Pm\bar{3}m$ (Caracas and
 419 Wentzcovitch 2006) (b) tetragonal $P4/mmm$ ($a = 3.5781 \text{ \AA}$ and $c = 3.5601 \text{ \AA}$) (Shim et
 420 al. 2002) (c) tetragonal $P4/mmm$ ($a = 3.5662 \text{ \AA}$ and $c = 3.5840 \text{ \AA}$) and (d) tetragonal
 421 $I4/mcm$ ($a = 5.0426 \text{ \AA}$ and $c = 7.1658 \text{ \AA}$) (Caracas and Wentzcovitch 2006). In (a), the
 422 diffraction peaks are indexed with the cubic $Pm\bar{3}m$ structure. For the tetragonal cases in
 423 (b-d), we assumed the same magnitude of deviation in the c/a ratio from 1: $c_{pc}/a_{pc} = 0.995$
 424 for (b) and $c_{pc}/a_{pc} = 1.005$ for (c) and (d). The thin vertical dashed line extended from the
 425 200_{pc} diffraction peak highlights the different angle shifts expected for the split lines in
 426 $I4/mcm$ and $P4/mmm$.
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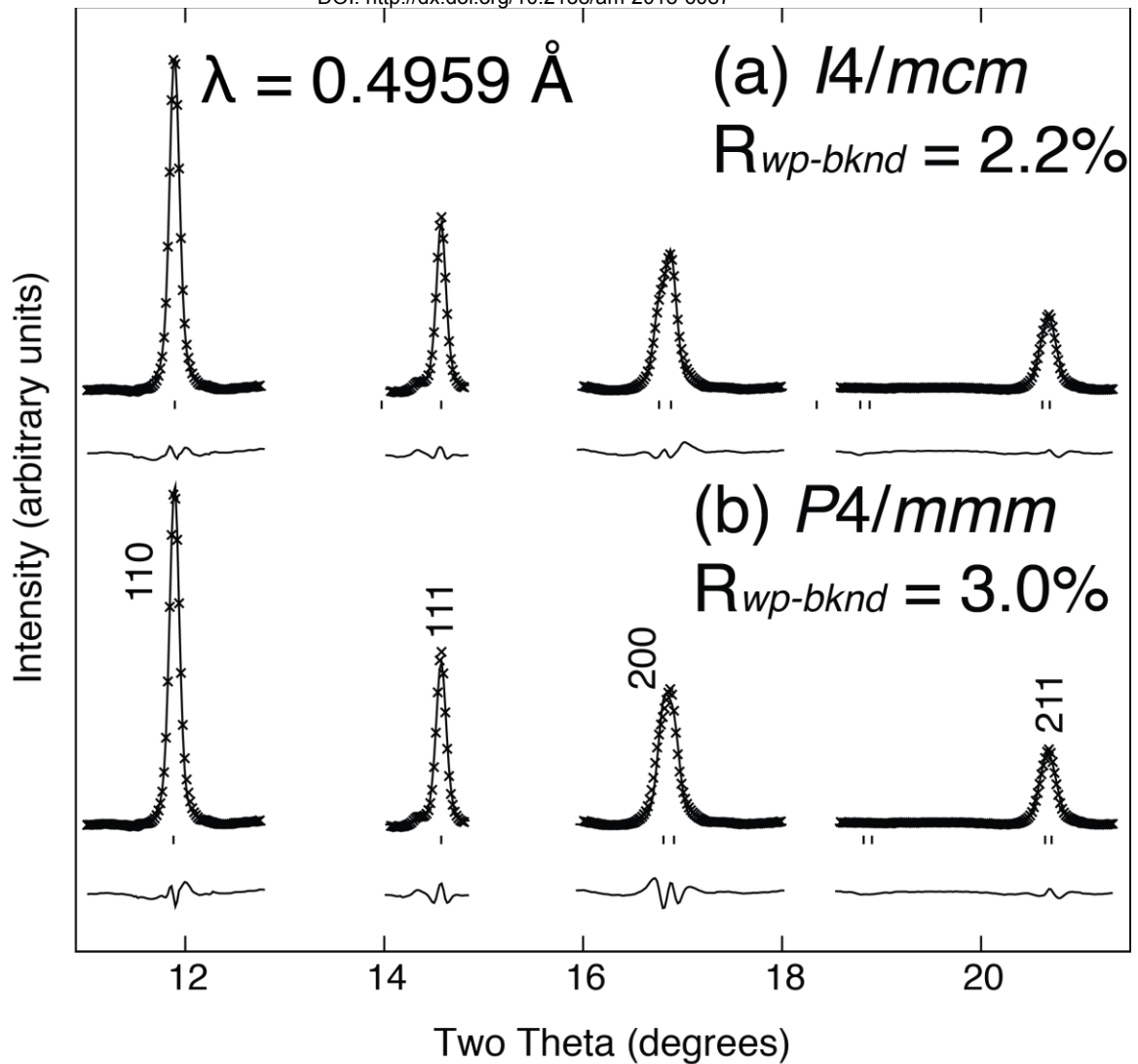
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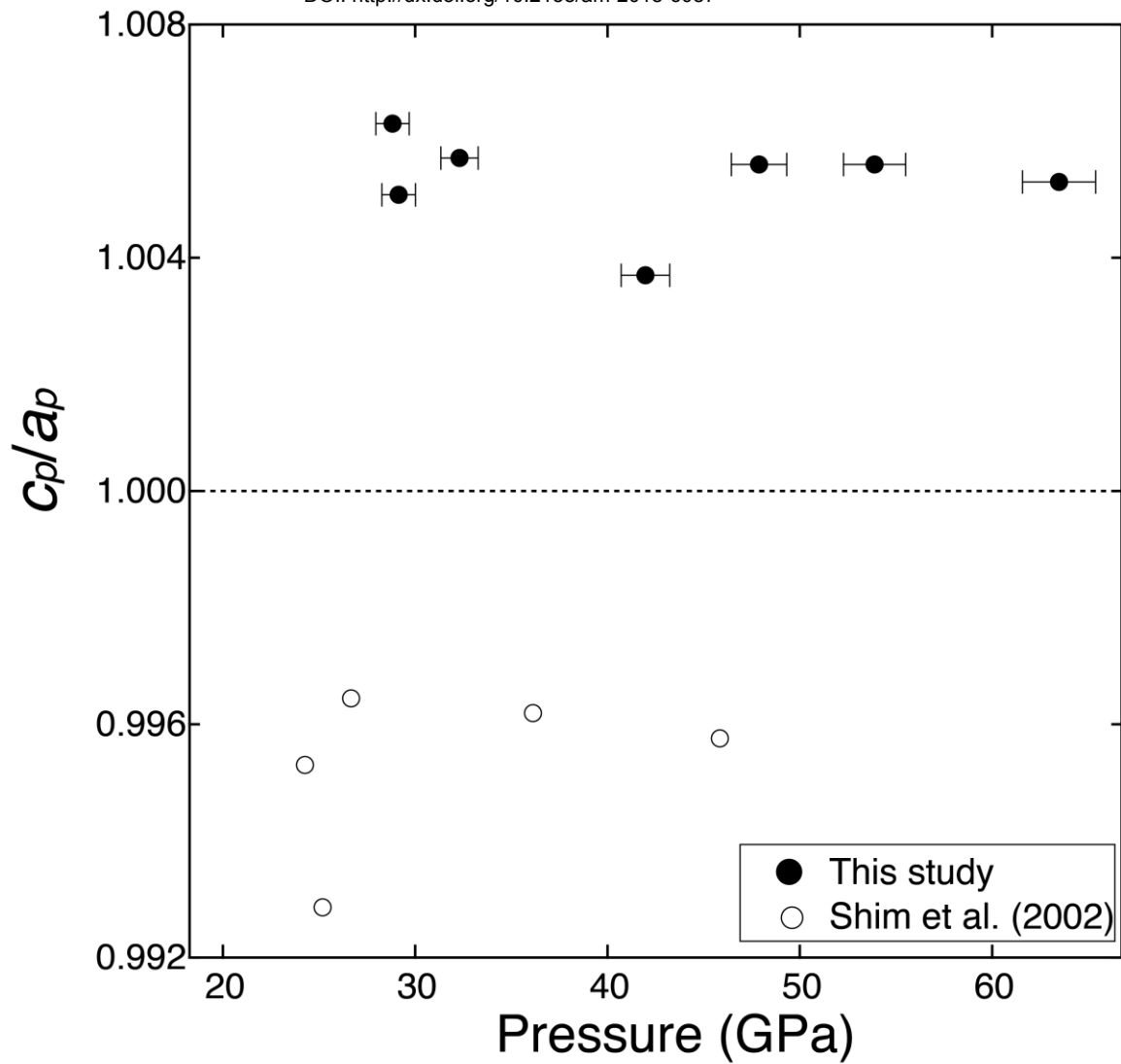
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431 Figure 2: X-ray diffraction patterns of CaPv at 50-55 GPa and 300 K (a) and 2000 K (b).
432 The samples include CaSiO_3 perovskite (CaPv), platinum (Pt, laser coupler) and neon
433 (Ne, pressure medium). The peaks are indexed with the cubic $Pm\bar{3}m$ space group. The
434 inset images show the (a) splitting and (b) no-splitting of the 200_{pc} diffraction peak at 300
435 K and 2000 K, respectively. The wavelength of the X-ray beam is 0.4959 \AA . The unrolled
436 projections of the 2D raw diffraction areas are shown for temperatures at (c) 300 K, and
437 (d) 2000 K for the same 2θ range (1.2°). In (d), the black area in the middle is from the
438 laser mirrors for heating.

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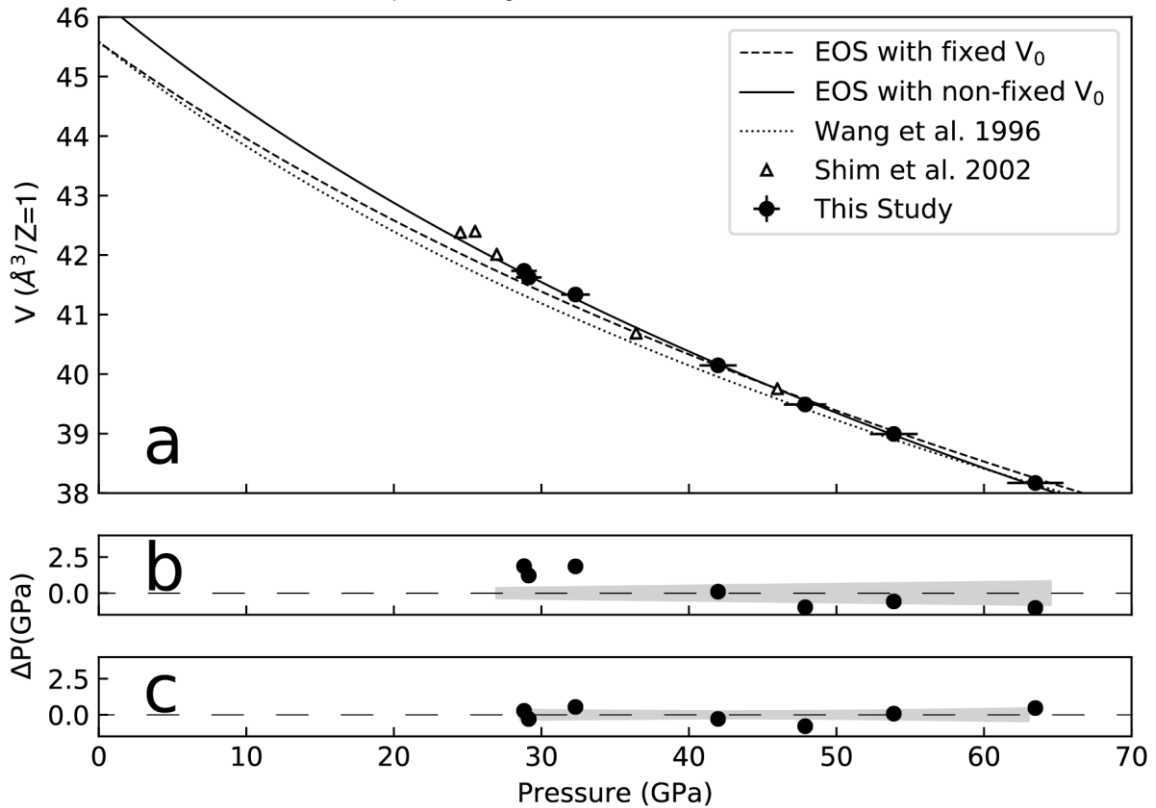


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444 Figure 3: Rietveld refinements of the XRD pattern of CaPv at 53.8 GPa and 300 K with
445 (a) the $I4/mcm$ and (b) $P4/mmm$ space groups. The crosses are observed intensities and
446 the line behind crosses is the calculated diffraction patterns. Differences between
447 observed and calculated intensities are plotted below. The small bars represent calculated
448 CaPv peak positions. We did not include the angular ranges for the Pt and Ne peaks in the
449 Rietveld refinements, in order to make $R_{wp-bknd}$ (residual factor after background
450 subtraction) only sensitive to CaPv.
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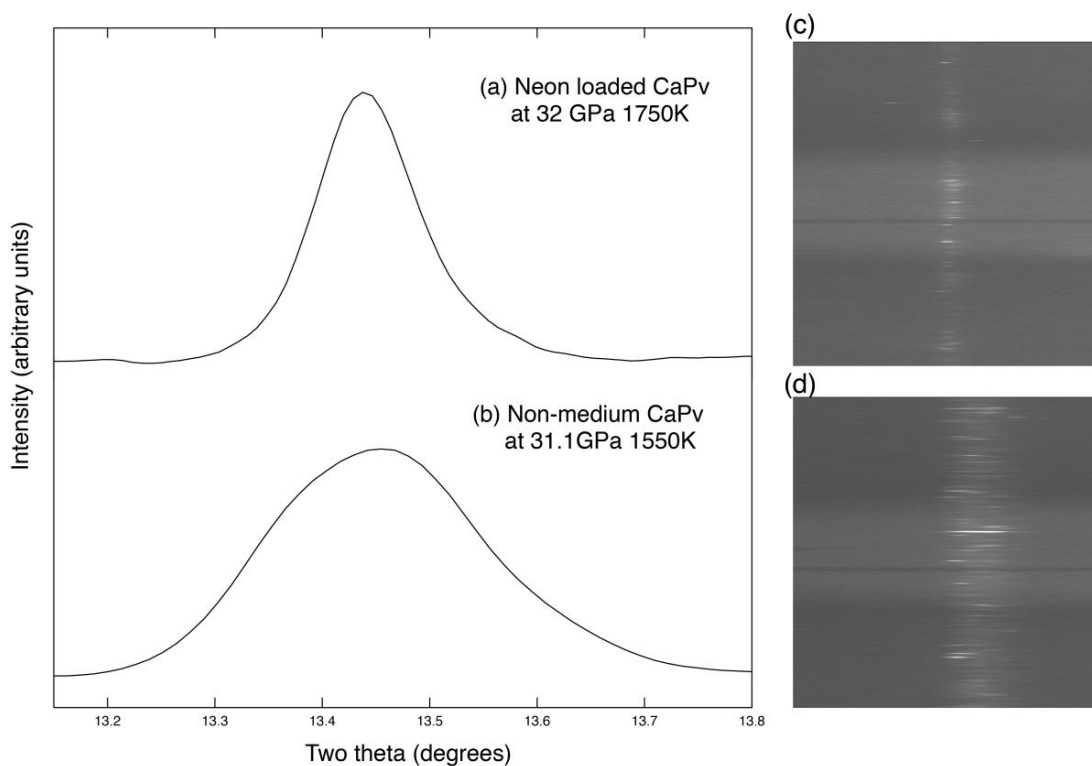
454 Figure 4: The axial ratio, c_{pc}/a_{pc} , of the tetragonal CaPv unit cell from Shim et al. (2002)
455 (open circles) and this study (close circles). Error bars represent the estimated 1σ
456 uncertainties.
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459 Figure 5: The pseudo-cubic unit cell volume (calculated with the same number of atoms
460 in a unit cell as the $Pm3m$ structure) of CaPv measured at high pressure (circles). The
461 solid line shows fitting our data to the Vinet equation (Table 2). Pressure was calculated
462 from Ye et al. (2017). We also include the compressional curves from Shim et al. (2002)
463 and Wang et al. (1996) for comparison. Error bars represent estimated 1σ uncertainties.
464 Z is the number of formula unit per unit cell and for pseudo-cubic cell Z is 1. ΔP is the
465 fit residues in pressure with fixed V_0 in (b) or fitted V_0 in (c).

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471 Figure 6: Cubic 200 diffraction peaks of CaPv measured with (a) Ne and (b) without
472 pressure mediums at in situ high P and T . The diffraction patterns were measured with
473 an X-ray wavelength of 0.3344 \AA and a detector distance of 196.56 mm. Unrolled
474 projection of 2D diffraction area of CaPv 200_{pc} peak (c) with a Ne pressure medium and
475 (d) without a pressure medium with same range (1.4°).

476 **References:**

- 477 Adams, D.J., and Oganov, A.R. (2006) Ab initio molecular dynamics study of CaSiO₃
478 perovskite at P-T conditions of Earth's lower mantle. *Physical Review B*, 73,
479 184106.
- 480 Akber-Knutson, S., Bukowinski, M.S.T., and Matas, J. (2002) On the structure and
481 compressibility of CaSiO₃ perovskite. *Geophysical Research Letters*, 29, 4–1.
- 482 Caracas, R., and Wentzcovitch, R.M. (2006) Theoretical determination of the structures
483 of CaSiO₃ perovskites. *Acta Crystallographica Section B: Structural Science*, 62,
484 1025–1030.
- 485 Caracas, R., Wentzcovitch, R., Price, G.D., and Brodholt, J. (2005) CaSiO₃ perovskite at
486 lower mantle pressures. *Geophysical Research Letters*, 32, L06306.
- 487 Chizmeshya, A.V.G., Wolf, G.H., and McMillan, P.F. (1996) First-principles calculation
488 of the equation-of-state, stability, and polar optic modes of CaSiO₃ perovskite.
489 *Geophysical Research Letters*, 23, 2725–2728.
- 490 Comyn, T.P., McBride, S.P., and Bell, A.J. (2004) Processing and electrical properties of
491 BiFeO₃–PbTiO₃ ceramics. *Materials Letters*, 58, 3844–3846.
- 492 Duffy, T.S., Hemley, R.J., and Mao, H. (1995) Equation of State and Shear Strength at
493 Multimegabar Pressures: Magnesium Oxide to 227 GPa. *Physical Review Letters*,
494 74, 1371–1374.
- 495 Fei, Y., Ricolleau, A., Frank, M., Mibe, K., Shen, G., and Prakapenka, V. (2007) Toward
496 an internally consistent pressure scale. *Proceedings of the National Academy of*
497 *Sciences*, 104, 9182–9186.
- 498 Glazer, A.M. (1975) Simple ways of determining perovskite structures. *Acta*
499 *Crystallographica Section A*, 31, 756–762.
- 500 Grocholski, B., Catalli, K., Shim, S.-H., and Prakapenka, V. (2012) Mineralogical effects
501 on the detectability of the postperovskite boundary. *Proceedings of the National*
502 *Academy of Sciences*, 109, 2275–2279.
- 503 Hirose, K., Takafuji, N., Sata, N., and Ohishi, Y. (2005) Phase transition and density of
504 subducted MORB crust in the lower mantle. *Earth and Planetary Science Letters*,
505 237, 239–251.
- 506 Holmes, N.C., Moriarty, J.A., Gathers, G.R., and Nellis, W.J. (1989) The equation of
507 state of platinum to 660 GPa (6.6 Mbar). *Journal of Applied Physics*, 66, 2962–
508 2967.
- 509 Jung, D.Y., and Oganov, A.R. (2005) Ab initio study of the high-pressure behavior of
510 CaSiO₃ perovskite. *Physics and Chemistry of Minerals*, 32, 146–153.

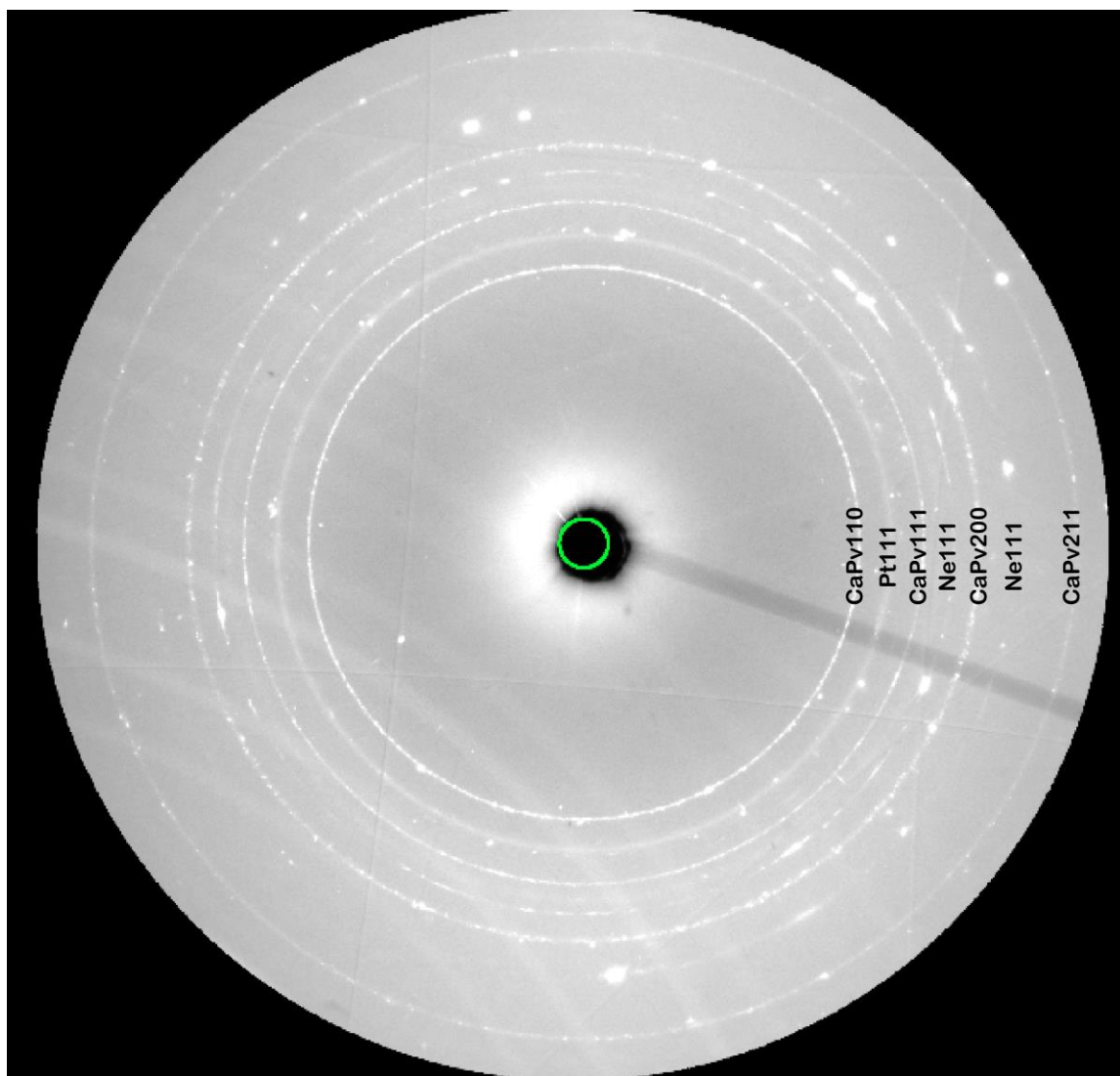
- 511 Kavner, A., and Duffy, T.S. (2001) Pressure–volume–temperature paths in the laser-
512 heated diamond anvil cell. *Journal of Applied Physics*, 89, 1907–1914.
- 513 Kawai, K., and Tsuchiya, T. (2014) P-V-T equation of state of cubic CaSiO₃ perovskite
514 from first-principles computation. *Journal of Geophysical Research: Solid Earth*,
515 119, 2013JB010905.
- 516 Kesson, S.E., Fitz Gerald, J.D., and Shelley, J.M. (1998) Mineralogy and dynamics of a
517 pyrolite lower mantle. *Nature*, 393, 252–255.
- 518 Klotz, S., Chervin, J.-C., Munsch, P., and Marchand, G.L. (2009) Hydrostatic limits of 11
519 pressure transmitting media. *Journal of Physics D: Applied Physics*, 42, 075413.
- 520 Komabayashi, T., Hirose, K., Sata, N., Ohishi, Y., and Dubrovinsky, L.S. (2007) Phase
521 transition in CaSiO₃ perovskite. *Earth and Planetary Science Letters*, 260, 564–
522 569.
- 523 Kurashina, T., Hirose, K., Ono, S., Sata, N., and Ohishi, Y. (2004) Phase transition in Al-
524 bearing CaSiO₃ perovskite: implications for seismic discontinuities in the lower
525 mantle. *Physics of the Earth and Planetary Interiors*, 145, 67–74.
- 526 Li, L., Weidner, D.J., Brodholt, J., Alfè, D., David Price, G., Caracas, R., and
527 Wentzcovitch, R. (2006) Elasticity of CaSiO₃ perovskite at high pressure and high
528 temperature. *Physics of the Earth and Planetary Interiors*, 155, 249–259.
- 529 Liu, L.-G., and Ringwood, A.E. (1975) Synthesis of a perovskite-type polymorph of
530 CaSiO₃, 28, 209–211.
- 531 Magyari-Köpe, B., Vitos, L., Grimvall, G., Johansson, B., and Kollár, J. (2002) Low-
532 temperature crystal structure of CaSiO₃ perovskite: An ab initio total energy
533 study. *Physical Review B*, 65, 193107.
- 534 Mao, H.K., Chen, L.C., Hemley, R.J., Jephcoat, A.P., Wu, Y., and Bassett, W.A. (1989)
535 Stability and equation of state of CaSiO₃-Perovskite to 134 GPa. *Journal of*
536 *Geophysical Research: Solid Earth*, 94, 17889–17894.
- 537 Margulies, L., Winther, G., and Poulsen, H.F. (2001) In Situ Measurement of Grain
538 Rotation During Deformation of Polycrystals. *Science*, 291, 2392–2394.
- 539 Marquardt, H., Speziale, S., Gleason, A., Sinogeikin, S., Kantor, I., and Prakapenka, V.B.
540 (2013) Brillouin scattering and x-ray diffraction of solid argon to 65 GPa and
541 700 K: Shear strength of argon at HP/HT. *Journal of Applied Physics*, 114,
542 093517.
- 543 Meng, Y., Shen, G., and Mao, H.K. (2006) Double-sided laser heating system at HPCAT
544 for in situ x-ray diffraction at high pressures and high temperatures. *Journal of*
545 *Physics: Condensed Matter*, 18, S1097.

- 546 Murakami, M., Hirose, K., Sata, N., and Ohishi, Y. (2005) Post-perovskite phase
547 transition and mineral chemistry in the pyrolitic lowermost mantle. Geophysical
548 Research Letters, 32, L03304.
- 549 Nisr, C., Ribárik, G., Ungár, T., Vaughan, G.B.M., Cordier, P., and Merkel, S. (2012)
550 High resolution three-dimensional X-ray diffraction study of dislocations in
551 grains of MgGeO₃ post-perovskite at 90 GPa. Journal of Geophysical Research:
552 Solid Earth, 117, B03201.
- 553 Noguchi, M., Komabayashi, T., Hirose, K., and Ohishi, Y. (2013) High-temperature
554 compression experiments of CaSiO₃ perovskite to lowermost mantle conditions
555 and its thermal equation of state. Physics and Chemistry of Minerals, 40, 81–91.
- 556 Ohta, K., Hirose, K., Lay, T., Sata, N., and Ohishi, Y. (2008) Phase transitions in pyrolite
557 and MORB at lowermost mantle conditions: Implications for a MORB-rich pile
558 above the core–mantle boundary. Earth and Planetary Science Letters, 267, 107–
559 117.
- 560 Ono, S., Ohishi, Y., and Mibe, K. (2004) Phase transition of Ca-perovskite and stability
561 of Al-bearing Mg-perovskite in the lower mantle. American Mineralogist, 89,
562 1480–1485.
- 563 Ono, S., Ohishi, Y., Isshiki, M., and Watanuki, T. (2005) In situ X-ray observations of
564 phase assemblages in peridotite and basalt compositions at lower mantle
565 conditions: Implications for density of subducted oceanic plate. Journal of
566 Geophysical Research: Solid Earth, 110, B02208.
- 567 Prakapenka, V.B., Kubo, A., Kuznetsov, A., Laskin, A., Shkurikhin, O., Dera, P., Rivers,
568 M.L., and Sutton, S.R. (2008) Advanced flat top laser heating system for high
569 pressure research at GSECARS: application to the melting behavior of
570 germanium. High Pressure Research, 28, 225–235.
- 571 Prescher, C., and Prakapenka, V.B. (2015) DIOPTAS: a program for reduction of two-
572 dimensional X-ray diffraction data and data exploration. High Pressure Research,
573 35, 223–230.
- 574 Ricolleau, A., Fei, Y., Cottrell, E., Watson, H., Deng, L., Zhang, L., Fiquet, G., Auzende,
575 A.-L., Roskosz, M., Morard, G., and others (2009) Density profile of pyrolite
576 under the lower mantle conditions. Geophysical Research Letters, 36, L06302.
- 577 Ricolleau, A., Perrillat, J.-P., Fiquet, G., Daniel, I., Matas, J., Addad, A., Menguy, N.,
578 Cardon, H., Mezouar, M., and Guignot, N. (2010) Phase relations and equation of
579 state of a natural MORB: Implications for the density profile of subducted oceanic
580 crust in the Earth's lower mantle. Journal of Geophysical Research: Solid Earth,
581 115, B08202.

- 582 Rivers, M., Prakapenka, V.B., Kubo, A., Pullins, C., Holl, C.M., and Jacobsen, S.D.
583 (2008) The COMPRES/GSECARS gas-loading system for diamond anvil cells at
584 the Advanced Photon Source. *High Pressure Research*, 28, 273–292.
- 585 Shim, S.-H., Duffy, T.S., and Shen, G. (2000) The stability and P–V–T equation of state
586 of CaSiO₃ perovskite in the Earth’s lower mantle. *Journal of Geophysical*
587 *Research: Solid Earth*, 105, 25955–25968.
- 588 Shim, S.-H., Jeanloz, R., and Duffy, T.S. (2002) Tetragonal structure of CaSiO₃
589 perovskite above 20 GPa. *Geophysical Research Letters*, 29, 2166.
- 590 S.-H. Shim (2017) PeakPo - A python software for X-ray diffraction analysis at high
591 pressure and high temperature.
592 Zenodo. <http://doi.org/10.5281/zenodo.810199>
- 593 Stixrude, L., Cohen, R.E., Yu, R.C., and Krakauer, H. (1996) Prediction of phase
594 transition in CaSiO₃ perovskite and implications for lower mantle structure.
595 *American Mineralogist*, 81, 1293–1296.
- 596 Stixrude, L., Lithgow-Bertelloni, C., Kiefer, B., and Fumagalli, P. (2007) Phase stability
597 and shear softening in CaSiO₃ perovskite at high pressure. *Physical Review B*, 75,
598 024108.
- 599 Sun, N., Mao, Z., Yan, S., Wu, X., Prakapenka, V.B., and Lin, J.-F. (2016) Confirming a
600 pyrolitic lower mantle using self-consistent pressure scales and new constraints on
601 CaSiO₃ perovskite. *Journal of Geophysical Research: Solid Earth*, 121,
602 2016JB013062.
- 603 Tamai, H., and Yagi, T. (1989) High-pressure and high-temperature phase relations in
604 CaSiO₃ and CaMgSi₂O₆ and elasticity of perovskite-type CaSiO₃. *Physics of the*
605 *Earth and Planetary Interiors*, 54, 370–377.
- 606 Tangeman, J.A., Phillips, B.L., Navrotsky, A., Weber, J.K.R., Hixson, A.D., and Key,
607 T.S. (2001) Vitreous forsterite (Mg₂SiO₄): Synthesis, structure, and
608 thermochemistry. *Geophysical Research Letters*, 28, 2517–2520.
- 609 Toby, B.H. (2001) EXPGUI, a graphical user interface for GSAS. *Journal of Applied*
610 *Crystallography*, 34, 210–213.
- 611 Uchida, T., Wang, Y., Nishiyama, N., Funakoshi, K., Kaneko, H., Nozawa, A., Von
612 Dreele, R.B., Rivers, M.L., Sutton, S.R., Yamada, A., and others (2009) Non-
613 cubic crystal symmetry of CaSiO₃ perovskite up to 18 GPa and 1600 K. *Earth*
614 *and Planetary Science Letters*, 282, 268–274.
- 615 Vinet, P., Rose, J.H., Ferrante, J., and Smith, J.R. (1989) Universal features of the
616 equation of state of solids. *Journal of Physics: Condensed Matter*, 1, 1941.

- 617 Wang, Y., Weidner, D.J., and Guyot, F. (1996) Thermal equation of state of CaSiO_3
618 perovskite. *Journal of Geophysical Research: Solid Earth*, 101, 661–672.
- 619 Woodward, P.M. (1997) Octahedral Tilting in Perovskites. I. Geometrical
620 Considerations. *Acta Crystallographica Section B: Structural Science*, 53, 32–43.
- 621 Yashima, M.; Ali, R. (2009) Structural phase transition and octahedral tilting in the
622 calcium titanate perovskite CaTiO_3 . *Solid State Ionics* 180, 120-126.
623
- 624 Ye, Y., Prakapenka, V., Meng, Y., and Shim, S.-H. (2017) Inter-comparison of the Gold,
625 Platinum, and MgO Pressure Scales up to 140 GPa and 2,500 K. *Journal of*
626 *Geophysical Research: Solid Earth*, 2016JB013811.

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631 Figure S1, 2D diffraction pattern of CaPv, Pt and Neon. We show different diffraction
632 peaks with their Miller indices. We used cubic indices for CaPv. The diffraction pattern
633 was collected at 50 GPa and 300 K.