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5 Geochemistry and mineralogy of a saprolite developed on Columbia River Basalt: Secondary
6 clay formation, element leaching, and mass balance during weathering

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(Revision 3)

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23 Abstract

24 This study presents chemical and mineralogical data on weathering trends in a saprolite
25 that is preserved between flows of the Columbia River Basalt Group at Trinidad, WA. Bulk
26 chemistry, electron imaging, and X-ray mapping indicate early Fe and Mg depletion by
27 dissolution of ferromagnesian minerals, followed by depletion of alkalis, Al, Ti, and P,
28 corresponding to dissolution of feldspars, titanomagnetite, and apatite. Secondary coatings of
29 nontronite clay in the deep saprolite display intricate, sub-micron scale zoning in Fe and Mg
30 content. Distinct aluminous zones in these clays become more prominent at shallower depths.
31 The primary Fe-containing phase shifts from nontronite in deeper samples to hematite in shallow
32 samples; samples at the boundary contain the assemblage kaolinite + nontronite, which may
33 mark the transition from permeability-limited fluid flow to fully open-system behavior. This shift
34 is observed in rocks that have lost 30-40% of the total rock mass to leaching, and coincides with
35 the disappearance of feldspar, Fe-Ti oxides, and apatite. Rocks in the uppermost saprolite have
36 been converted to an assemblage of Al-smectite + hematite (+kaolinite). These results suggest
37 that the presence of nontronite in weathered samples may indicate weathering under conditions
38 of limited permeability; however, it does not necessarily indicate weathering in a chemically
39 closed system. These observations may be useful in interpreting the clay mineral assemblages
40 observed on Mars and what information they contain about near-surface conditions in the
41 planet's ancient past.

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Introduction

45 The weathering of fresh volcanic rocks, and particularly of fresh basaltic rock, is a major
46 factor controlling ocean chemistry, atmospheric CO₂ fixation, and ultimately, terrestrial climate
47 (Gaillardet et al. 1999; Kump et al. 2000; Benedetti et al. 2003; Dessert et al. 2003; Dupré et al.
48 2003; Gislason et al. 2009; Maher and Chamberlain 2014). These studies indicate that
49 weathering of basaltic lavas is responsible for approximately one third of global CO₂ fixation by
50 weathering of continental rocks, with much of this accounted for by weathering of young basalts
51 in warm, wet tropical localities. Subsurface weathering of volcanic rocks can be, quantitatively,
52 an even more significant ion source and carbon sink than surface weathering (Rad et al. 2007).
53 This study presents a detailed examination of chemical leaching and its relationship to secondary
54 clay mineralogy in a basalt saprolite.

55 The Columbia River Basalts (CRB) are continental flood basalt lavas located in the
56 interior northwestern United States including portions of Washington, Oregon, and Idaho (Tolan
57 et al. 1989; Reidel et al. 2013). The majority of these lavas were erupted during the mid-Miocene
58 Climatic Optimum, when the climate of the northwestern United States was considerably warmer
59 and wetter than in the present day (Takeuchi and Larson 2005). Lava flows exposed at the
60 surface weathered rapidly under these conditions to form soils. Those soils were later covered by
61 younger lava flows, preserving the buried saprolites and paleosols as records of past climatic
62 conditions (Sheldon 2003; Takeuchi and Larson 2005; Sheldon 2006; Hobbs and Parrish 2016).
63 Although some overprinting occurred where the capping lava flows reacted with the capped soils
64 (Baker 2017), the paleosols and their underlying saprolite zones also preserve a record of the
65 weathering reactions that converted basalt flows to soils. Weathering of CRB is considered to

66 have significantly contributed to both fixation of atmospheric CO₂ and the flux of dissolved ions
67 to the world's oceans via the Columbia River system (Taylor and Lasaga 1999).

68 The fixation of carbon and the flux of dissolved ions from a weathering basalt will
69 depend upon the individual weathering reactions occurring in the surface and subsurface. A
70 number of studies have examined the chemistry and mineralogy of basalt weathering (Colman
71 1982; Eggleton et al. 1987; Nesbitt and Wilson 1992). These studies show that weathering trends
72 of element depletion and preferential mineral dissolution vary from system to system. The early
73 weathering behavior of individual rocks is controlled by their primary mineralogy and by the
74 microstructural properties that control water movement through the rock (Meunier et al. 2007),
75 with element leaching being diffusion-limited until porosity reaches ~9% (Navarre-Sitchler et al.
76 2009; Navarre-Sitchler et al. 2013). Typically, olivine is the earliest-weathering phase if present,
77 but the order of subsequent dissolution of glass and of minerals such as pyroxenes, plagioclase,
78 and Fe-Ti oxides varies between different studies (Allen and Scheid 1946; Eggleton et al. 1987;
79 Nesbitt and Wilson 1992). Element leaching from basalts is also subject to redox control, with
80 Fe loss indicating anoxic or reducing conditions (Ohmoto 1996; Rye and Holland 1998; Rye and
81 Holland 2000; Babechuk et al. 2014). Thus, element release from weathering of freshly
82 emplaced basalts is likely to vary with time as different minerals undergo dissolution and as the
83 zone of leaching moves progressively downward from the flow surface.

84 Element leaching from weathering basalts is accompanied by precipitation of secondary
85 mineral phases. Fe-bearing smectites are typical secondary clays produced by early basalt
86 weathering, transitioning to kaolinite / halloysite and Fe (oxyhydr)oxides in more extensively
87 weathered samples (Glasmann and Simonson 1985; Eggleton et al. 1987; Vingiani et al. 2010).
88 The first low-temperature weathering product of CRB is nontronite clay, sometimes

89 accompanied by opaline silica and clinoptilolite (Benson and Teague 1982; Allen and Scheid
90 1946). Nontronite is a ferric smectite with nominal (Na-saturated) endmember composition
91 $\text{Na}_{0.3}\text{Fe}_2(\text{Si},\text{Al})_4\text{O}_{10}(\text{OH})_2 \cdot n\text{H}_2\text{O}$. Previous studies have suggested that nontronite forms via basalt
92 weathering only under conditions where water-rock interaction is limited (Ehlmann et al. 2011)
93 and that it disappears completely from rocks as they undergo more pervasive alteration (Allen
94 and Scheid 1946). However, Sherman et al. (1962) observed nontronite accompanying bauxite in
95 highly weathered basalt from Hawaii, and suggested it might persist longer than was thought
96 under extensive weathering. Coexisting nontronite and halloysite or kaolinite have since been
97 described by several other studies of basalt weathering. Vingiani et al. (2010) observed
98 nontronite + kaolinite/halloysite replacing feldspars in weathering basaltic pebbles in soil, and de
99 Oliveira et al. (1998) described it in a transitional zone of intermediate weathering between
100 saprock and soil. Cravero et al. (2014) described it in altered ignimbrites, indicating it may also
101 occur in rocks of silicic composition. These studies attributed the co-occurrence of ferric
102 smectites and kaolinite / halloysite to local variations in porosity and permeability that allowed
103 secondary mineral compositions to be controlled locally by dissolution. It is not clear whether
104 any of the clays described were Fe-bearing mixed-layer kaolinite-smectite phases, which may
105 also form via basalt weathering, and the compositions of which are also subject to local controls
106 on solution chemistry (Le Blond et al. 2015).

107 These observations suggest that both element leaching and secondary mineral formation
108 during basalt weathering are intimately tied to hydrologic conditions in the rock (Meunier et al.
109 2007; Navarre-Sitchler et al. 2011), not only during early weathering but well into intermediate-
110 stage weathering in which significant desilication is occurring. Therefore, secondary clay

111 mineralogy in saprolites may serve as an indicator of how hydrologic connectivity developed,
112 and how element leaching progressed, during the weathering process.

113 The purpose of this study was to examine weathering along a depth transect through a
114 basaltic saprolite, and to assess the primary minerals undergoing dissolution, major rock
115 components undergoing leaching, and clay and secondary minerals being formed. A previous
116 study (Baker 2017) presented whole-rock chemical data and bulk XRD data along this transect
117 that showed chemical patterns resulting from weathering (Table 1). This study presents
118 complementary element maps and synchrotron-based microanalyses examining the individual
119 reaction zones where alteration reactions were preferentially leaching certain elements and
120 minerals from the rock, and where secondary clay minerals were precipitating from the
121 weathering fluids. These data provide information on element release over time from weathering
122 Columbia River Basalts, and on the stability of nontronite during closed- and open-system
123 weathering.

124 The questions of element leaching and secondary mineral formation during weathering
125 are also of considerable interest on Mars, where the presence of extensive clay mineral deposits
126 in ancient rocks indicates that aqueous weathering may have taken place in the planet's ancient
127 past (Bibring et al. 2006; Mustard et al. 2008; Murchie et al. 2009; Bishop et al. 2013; Carter et
128 al. 2015). Terrestrial basalt weathering systems are useful possible analogs for rock alteration
129 and clay formation on Mars (Greenberger et al. 2012; Thomson et al. 2014; Greenberger et al.
130 2015). This system presents a likely geochemical parallel to terrestrial basalt weathering in terms
131 of the release of dissolved ions, carbonate fixation, and possibly climate feedbacks. A better
132 understanding of rock weathering and the formation of secondary clay minerals under different
133 climatic conditions will improve interpretation of the martian clay mineral deposits. If saprolite

134 secondary mineralogy preserves information about weathering conditions and element leaching,
135 then future close study of the secondary mineralogy of martian saprolites may provide detailed
136 information on crustal hydrology, element cycling, and weathering-controlled climate feedbacks
137 on early Mars.

138

139

Methods

140 Sample collection

141 The sample locality near Trinidad, WA was previously described by Baker (2017) and
142 further details of the regional geology are given in that study. At this locality, a saprolite overlain
143 by a thin (~50 cm) paleosol is developed on the Basalt of Gingko, and capped by a flow of the
144 Basalt of Sand Hollow, both units of the Frenchman Springs Member of the Wanapum Basalts of
145 the Columbia River Basalts (Tolan et al. 2009). The saprolite was sampled at depths of 10, 15,
146 20, 30, 60, and 80 cm below the contact between the capping basalt and the paleosol top. Fresh
147 Gingko basalt was not exposed at the transect site, so an unweathered sample was collected from
148 an exposure downslope.

149 In the laboratory, secondary mineral infillings were scraped or picked by hand from void
150 spaces in the 80 cm saprolite sample for separate spectroscopic analysis. These infillings were
151 sorted on the basis of color and included tan, brown, and black samples. Upon sampling it was
152 found that the so-called black clay actually consisted of a brown infill with a brittle, black
153 exterior coating; however, this sample was analyzed separately and will be referred to throughout
154 this study as the black clay sample. Multicolored secondary mineral infillings were also

155 observed in the 60 and 30 cm saprolite samples, but it was not possible to collect sufficient
156 amounts of these clays for separate analysis.

157

158 **Analytical methods**

159 Details of bulk chemical and mineralogical analysis, and the results of these analyses, are
160 given by Baker (2017). The bulk chemical data from that study are presented in Table 1, with
161 additional data on C and Zr concentrations and calculated weathering indices.

162 Bulk saprolite samples from each depth were crushed by hand using a mortar and pestle
163 and were analyzed by Fourier transform infrared spectroscopy (FTIR). Each sample was mixed
164 with optical-grade KBr at a clay:KBr ratio of 3:100 and analyzed on a Perkin-Elmer System
165 2000 spectrometer with diffuse reflectance accessory. FTIR spectra were processed using the
166 Kubelka-Munk algorithm provided in Perkin Elmer Spectrum 2.0 software.

167 Paleosol and saprolite samples were impregnated with Petropoxy under vacuum for thin-
168 section preparation. Many of these samples contained a large amount of pore space (primary
169 vesicle space, or voids expanded by dissolution), but were not high permeability (due to a lack of
170 interconnection between void spaces), which made complete epoxy impregnation challenging.
171 As a result of this, of the extreme friability of these samples, and of the need to preserve the
172 delicate spatial structures of secondary minerals infilling void spaces, multiple impregnation
173 steps were used to stabilize the samples before and between individual cutting and polishing
174 steps.

175 Carbon-coated thin sections were analyzed at the Washington State University
176 GeoAnalytical Laboratory on a JEOL 8500F field-emission electron microprobe. Backscattered

177 electron images and wavelength-dispersive X-ray maps, showing the relative distribution of
178 individual elements within the mapped area, were collected of secondary mineral infills and their
179 spatial relationships with the parent basalt. Because the secondary clay minerals in the samples
180 tended to pluck during polishing, resulting in a poorly polished surface, quantitative analyses of
181 clay composition were not performed.

182 Thin sections of all samples were taken to Stanford Synchrotron Research Laboratory
183 (SSRL) for synchrotron-based microanalysis on Beamline 2-3. Selected sample regions were
184 element mapped by μ -X-ray fluorescence (μ -XRF) with an incident beam energy of 7500 eV,
185 tuned with a Si(111) monochromator. Electron beam energy was 3 GeV and beam current was
186 350 mA. Beamline 2-3 uses Kirkpatrick-Baez Pt-coated focusing optics to achieve a beam spot
187 size as small as 2 by 2 μ m. Thin sections were attached to an XYZ stage oriented 45° to the
188 beam. The fluorescence yield was detected using a single channel Vortex Si detector. Element
189 maps were collected using a 2 μ m spot size and a dwell time per point of 250 milliseconds. The
190 μ -XRF data were analyzed and the maps were produced using the software package SMAK
191 version 1.1 (Webb et al. 2011). The regions mapped by μ -XRF were selected to overlap those
192 previously mapped by electron microprobe. Selected spots of interest, chosen due to their
193 relatively high Fe content, were analyzed by μ -X-ray absorption near edge spectroscopy (μ -
194 XANES) at the Fe K-edge. Three to four spectra from each spot were aligned and averaged using
195 the software package SixPack version 1.3 (Webb 2005).

196 Subsamples of three hand-picked clay infillings from the 80 cm saprolite sample were
197 analyzed by extended X-ray absorption fine structure spectroscopy (EXAFS) at SSRL Beamline
198 7-3. These samples were crushed and rubbed in an even layer on filter paper (Whatman #2). The
199 sample-impregnated filter paper was cut into strips, stacked three layers deep in a sample holder,

200 and held in place with Kapton tape. The monochromator for this beamline consisted of two
201 parallel Si(220) crystals with a 6-mm entrance slit. All samples were run in a liquid He-cooled
202 cryostat at a temperature of 10 K. Fluorescence data were collected using a passivated implanted
203 planar silicon (PIPS) detector. One to three EXAFS scans per sample were calibrated to a
204 standard Fe foil and averaged using SixPack (Webb 2005). Processing and shell fitting were
205 carried out using the programs Athena and Artemis (Ravel and Newville 2005). Shell fitting was
206 performed using the method described by Baker and Strawn (2012), to confirm that the brown
207 and tan clay infillings were typical of high-Fe nontronites with similar bond distances and
208 parameters.

209

210

Results

211 Hand samples and mineralogy

212 The Gingko parent basalt contained plagioclase phenocrysts, with abundant quench-
213 textured acicular plagioclase and skeletal Fe-Ti oxides within a finely textured groundmass. No
214 olivine or pyroxene phenocrysts were present in this flow. This is consistent with previous
215 descriptions of this basalt (Martin et al. 2013) and with results from bulk XRD analysis (Baker
216 2017).

217 Samples from 20-80 cm deep in the saprolite were medium to light gray in appearance,
218 progressively lighter-toned at shallower depths, and contained considerable void space. The bulk
219 mineralogy of these samples was dominated by primary plagioclase feldspar and minor
220 titanomagnetite, with a small admixture of secondary smectite and, at 20 cm, kaolinite (Baker

221 2017). Bulk FTIR analyses indicate that the clay mineral halloysite was also present in samples
222 from depths of 30-80 cm (Figure 1).

223 Open void spaces were observed in saprolite samples from 20-80 cm depth. Some void
224 spaces contained no obvious secondary minerals. Some spaces exhibited partial interior coatings
225 of tan, brown, green, or black minerals, with the different colors appearing in different void
226 spaces. Black terminal coatings were sometimes present on tan or brown clay infillings, but the
227 black material was never coated by clay minerals of other colors. In many cases, these partial
228 interior coatings had a mossy, vermicular appearance (Figure 2), with tendrils extending from the
229 void edge into its interior. These tendrils are observable in thin sections where they were
230 preserved by epoxy impregnation. Other void spaces were completely filled with clay of uniform
231 tan or brown color.

232 Samples from 10 and 15 cm below the contact were visually different from those below;
233 they exhibited a streaky white and pale green mottling. Open or obviously infilled void spaces
234 were not observable in these samples, although their epoxy consumption during thin sectioning
235 suggested they had relatively high porosity. Bulk XRD results indicated that plagioclase feldspar
236 was absent from these samples and that the dominant mineral present was a smectite (Baker
237 2017). Clay separates from these depths were smectites with d-spacings of 14.7 – 15 Å (Mg-
238 saturated), expanding slightly upon glycerol solvation (15-15.5 Å), and collapsing upon K-
239 saturation (13-13.2 Å) and subsequent heating (10 Å). The relatively small expansion upon
240 glycerol solvation suggests this smectite may have high layer charge. Alternatively, interlayer
241 hydroxide precipitates could be present in the sample, preventing complete glycerol penetration
242 (Barnhisel and Bertsch 1989; Harris and White 2008).

243

244 **Bulk chemical trends and weathering indices**

245 Bulk chemical analyses of the parent basalt and saprolite samples were presented by
246 Baker (2017) and are shown in Table 1. Also shown are the oxidizing and reducing mafic index
247 of alteration (MIA-O and MIA-R respectively) (Babechuk et al. 2014). Table 2 shows mobility
248 ratios (τ) (Brimhall et al. 1992; Sheldon 2003) calculated with respect to Zr from the data in
249 Table 1 corrected to an anhydrous basis. The mobility ratio compares abundance of an element in
250 a given sample to its abundance in the parent rock. The concentration of the element of interest is
251 ratioed to the concentration of a supposedly immobile element. Calculation of mobility ratios
252 with respect to Zr thus includes an assumption that Zr remains immobile throughout weathering;
253 this behavior has been observed for weathering of other Columbia River Basalts (Thomson et al.
254 2014) but, from the bulk chemical data in Table 1, is clearly not appropriate at Trinidad for
255 depths shallower than 20 cm, and ratios calculated for 0-15 cm should be treated with caution.
256 No elements analyzed were observed to remain immobile in the Trinidad saprolite. Elements
257 often treated as immobile in weathering systems include Zr, Ti, and Al. All of these elements
258 become mobile in the Trinidad saprolite above 20 cm (Tables 1 and 2) and calculations of τ for
259 the Trinidad samples with respect to each of these three elements yielded similar values.

260 These data show that at 80 cm depth, the majority of Mg and K had already been leached
261 from the rock along with some Ca, Si, and Fe. This leaching trend progressed up-section to 20
262 cm. Elements including Al, Ti, Na, and P as well as many trace elements were relatively
263 immobile at these depths, and passively accumulated as other elements were leached. This
264 observation is consistent with dissolution of any glass present and of ferromagnesian groundmass
265 minerals. The immobility of Al in combination with the detection of halloysite suggests that any
266 Al leached from aluminous phases such as plagioclase was immediately re-precipitated. At 10-15

267 cm depth, loss of P, Ti, and Zr indicates dissolution of refractory phases such as apatite and
268 titanomagnetite, but other elements appeared to have been enriched with respect to the deeper
269 saprolite layers. In the paleosol, elements including Fe, Si, and particularly Mn were enriched to
270 higher values than in the parent basalt. This enrichment likely arose due to chemical reaction
271 with the overriding basalt, mediated by hydrothermal circulation of soil water (Baker 2017).

272 **Electron microprobe element mapping**

273 The electron microprobe images and element maps show that primary minerals present
274 included plagioclase and titanomagnetite, in agreement with the previously published
275 observations from bulk chemistry and XRD. At 80 cm below the surface, large void spaces were
276 lined and small void spaces were filled with clay minerals. These clays exhibited a vermicular
277 texture (Figure 2). Compositionally zoned Fe-rich clays lined the contact between groundmass
278 and clay infill, and individual clay tendrils had Fe-rich, Mg-poor cores (Figure 3). These early
279 clays were coated with alternating bands of relatively Mg-rich and relatively Mg-poor clays
280 (Figure 4). In some void spaces, the clay coating was terminated by a layer of black Mn oxides
281 (Figure 3). Most void spaces were coated rather than completely filled by clays, but some clay-
282 filled embayments penetrated the walls of the void spaces, apparently having formed by
283 dissolution of more soluble groundmass components (Figure 5).

284 At 60 cm below the surface, textural evidence suggested continued preferential
285 dissolution of ferromagnesian components. Relict plagioclase crystals projected into open void
286 spaces and were coated with Fe-Mg clay minerals (Figure 6). Although plagioclase showed
287 textural evidence of some dissolution, the crystals were largely intact and maintained their
288 overall lath-like shape and sharp boundaries. The texture of intact plagioclase laths projecting
289 into open void spaces, and of open void spaces with square terminations against plagioclase

290 crystal edges, suggests the void spaces were formed around the crystals by preferential
291 dissolution and partial removal of the matrix around them, leaving the clay precipitates behind.
292 As in the 80 cm sample, these clays were zoned in Fe and Mg.

293 At 30 cm, evidence appeared for partial dissolution of plagioclase in addition to
294 ferromagnesian minerals. Plagioclase crystals no longer projected into void spaces, but
295 terminated against intricately zoned clay infillings. Void spaces were lined with layered Fe-Mg
296 clays as in deeper samples, but distinct Al-rich zones were also present in the clay linings (Figure
297 7). These Al-rich zones defined a boundary between the layered Fe-Mg clay infillings and more
298 morphologically complex Fe-Mg clays that appeared to be possibly pseudomorphic after parent
299 mineral(s). These may define the edges of crystals that were originally clay-coated but unaltered
300 (Figure 6), and they indicate that the clay-filled zones of alteration expanded outwards into the
301 rock mass from the boundaries of the originally clay-lined or clay-filled void spaces.

302 At 20 cm below the surface, the texture changed from rock with localized altered zones to
303 a pervasively recrystallized appearance. Relict plagioclase phenocrysts showed abundant
304 evidence of dissolution (Figure 8). Quench-textured Fe-Ti oxide crystals disappeared from the
305 groundmass and were replaced by a Mg-rich phase, possibly a smectite. Aluminous zones
306 became more prominent in clay void infills. The Fe-Mg-rich clays filling the older void spaces
307 were still present, but were now surrounded by aluminous secondary clay minerals. This sample
308 appeared to represent a transition state between leaching of ferromagnesian components and
309 removal of aluminosilicate and less mobile components, and from precipitation of Fe-rich clays
310 to precipitation of aluminous clays.

311 At 10-15 cm below the surface, open void spaces such as were present deeper in the
312 section were absent. Minor Fe clays were present in isolated pockets (Figure 9), but most Fe-rich

313 zones were Si-poor. An Fe-rich phase was present as lacy coatings or infills and was free of Si or
314 other elements, suggesting a hydroxide or oxide phase. Aluminous clay minerals were
315 pseudomorphic after plagioclase laths (Figure 9 upper right), or Al-Mg clays alternated in bands
316 with an Fe-rich, Si-poor phase (Figure 9 upper left).

317 **X-ray absorption spectroscopy**

318 Synchrotron-based X-ray absorption (XAS) data including bulk EXAFS and μ -XANES
319 spectra were used to help identify Fe-rich secondary minerals present in the clay infillings. Fe K-
320 edge XAS spectra are a sensitive indicator of Fe valence, speciation, and structural environment
321 in crystalline and non-crystalline materials (Waychunas et al. 1983; Wilke et al. 2001; O'Day et
322 al. 2004; Wilke et al. 2007). They are also useful for distinguishing phyllosilicates from oxides
323 and other major Fe-bearing phases (Baker et al. 2010), for distinguishing between dioctahedral
324 and trioctahedral phyllosilicates (Finck et al. 2015), and for analyzing cation content and
325 distribution in clay minerals (Manceau et al. 1988; Vantelon et al. 2003; Baker and Strawn
326 2012). The use of synchrotron microprobe XRF and XANES analysis in particular allows in-situ
327 identification of very small or low abundance phases that cannot be analyzed in bulk.

328 The tan, brown, and black clay infillings that were handpicked from the 80 cm sample
329 were available in sufficient quantity for bulk XAFS analysis (Figure 10). These bulk samples are
330 likely to be mixtures of the compositional range of zoned clays observed by electron microprobe
331 (Figures 2-4). The black clay corresponds to the infillings observed by electron microprobe to be
332 Fe-rich clays coated with black Mn oxides (Figure 3). The EXAFS spectra of all these samples
333 resembled those of nontronite standards and were typical of high-Fe clays; the spectrum of UI-
334 Garfield nontronite (Baker and Strawn 2012), a high-Fe ferric nontronite, is shown in Figure 10
335 for comparison. Spectra of the tan and brown samples were very similar to one another, with

336 only a few minor differences that may correspond to slight variations in composition. If other
337 secondary phases were present such as high-Al or high-Mg smectites, trioctahedral smectites
338 such as saponite, or other Fe-rich phases such as (oxyhydr)oxides, their presence would be
339 evident in the XANES and EXAFS spectra. No such phases were detected in the bulk samples
340 from 80 cm, only high-Fe nontronite. The presence of some phases such as Fe (oxyhydr)oxides
341 would be evident from unique features in the XANES spectra (Figures 9a and 9b), whereas the
342 presence of trioctahedral smectites or of compositionally heterogeneous dioctahedral smectites
343 would be apparent from shifts in the XAFS oscillations (Figure 10c) and from changes in peak
344 location and intensity in the Fourier transformed spectra (Figure 10d).

345 The Fe edge in the near-edge portion of the black clay spectrum contained a lower energy
346 shoulder at 7120 eV that was not present in the spectra of the tan and brown clays and the
347 nontronite standard; this inflection was more obvious in the first derivative of the spectrum
348 (Figure 10b, arrow). This feature suggests that some Fe(II) was present in this clay sample
349 (Waychunas et al. 1983). The XANES and EXAFS spectra were otherwise similar to those of
350 high-Fe nontronite samples, except that the black clay XAFS spectrum had a lower amplitude
351 than the tan, brown, and Garfield clay spectra (Figure 10c and 10d), suggesting this clay may
352 also be somewhat less ordered.

353 **μ -XRF and XANES**

354 The tan, brown, and black clays described above were collected from the same hand
355 samples that were thin-sectioned and probed using μ -XRF and XANES. Because they represent
356 the same population of secondary minerals, they were used in this study as spectroscopic
357 standards for interpretation of the microbeam data. This is particularly useful for the Fe(II)-

358 containing clays with black Mn-rich coatings, as similar clays have not, to our knowledge, been
359 previously described.

360 The point XANES spectra of clay coatings from the 60 and 80 cm samples (Figure 11)
361 were nontronite-like, and most resembled the bulk XANES spectrum of the black clay, including
362 the reduced Fe shoulder at 7120 eV as described in the previous section. This strongly suggests
363 that most of the clay coatings analyzed contained similar amounts of Fe(II) to the black clay, and
364 that the fully oxidized tan and brown clays were not typical of the clays analyzed by point
365 analysis. However, some beam-induced reduction of Fe may have also occurred in these
366 samples, as the microprobe XANES spectra were collected at room temperature whereas the bulk
367 XAFS spectra were collected at 10 K.

368 One XANES spectrum was deliberately collected on a primary Fe-Ti oxide from the 80
369 cm sample. Although this spectrum resembled a hematite standard spectrum in its overall shape
370 (Figure 12), the very high pre-edge peak and the position of the first main edge inflection (7120
371 eV) more closely matched that of the magnetite spectrum. As in the black clay, the inflection at
372 7120 eV indicated the presence of some Fe(II) in this sample. Hematite was not a primary phase
373 in the basaltic parent rock, whereas titanomagnetite was, and was detected in this sample by
374 XRD (Baker 2017). Therefore this spectrum will be taken as typical of primary titanomagnetite
375 in the parent basalt. All but one of the point XANES spectra collected from the 30 cm sample
376 (Figure 12) partly matched the XANES spectrum of this titanomagnetite. The remaining, non-
377 oxide-like spectrum from 30 cm resembled that of the brown clay. Two of the oxide-matching
378 points had very large pre-edge features (position indicated by arrow on Figure 12) and best
379 matched the titanomagnetite spectrum from the 80 cm sample. The other two had relatively small
380 pre-edge features similar to clay mineral pre-edges rather than the large pre-edge features typical

381 of Fe oxide and hydroxide minerals; however, their main edge features were similar to the
382 titanomagnetite main edge and not to clay spectra. Those two spectra did not perfectly match any
383 single standard.

384 All but two point XANES spectra collected from the 20 cm sample resembled the black
385 clay spectrum and were indistinguishable from Fe-rich clay XANES spectra collected on the
386 deeper samples (Figure 13). This indicated that nontronite was still present in this kaolinite-
387 bearing sample. The two spectra (shown lowest on Figure 13) from the 20 cm sample that did not
388 match clay spectra do not have distinctive edge features, suggesting they may be poorly
389 crystalline phases such as Fe (oxyhydr)oxides. Point XANES spectra collected from the 10 cm
390 sample best match that of the hematite standard (Figure 14).

391

392

Discussion

393 **Patterns of mineral dissolution, chemical depletion, and clay formation**

394 Between the parent basalt and the deepest 80 cm saprolite sample, Al, Ti, P, and Zr
395 appeared to behave in a conservative manner (Table 1). Most other major elements appeared to
396 have undergone some leaching from the bulk rock, whereas volatile components (H₂O and C;
397 Tables 1 and 2) were added. The most mobile elements at this depth were Mg and K; the 80 cm
398 sample is depleted of 80% of the K₂O and 77 % of the MgO that were present in the parent rock.
399 Less depleted but still mobile elements included Ca (43 % CaO leached), Fe (33% Fe₂O₃), Si (30
400 % SiO₂), and Na (20% Na₂O) (Table 2). Overall, at the 80 cm mark, 28% (calculated on a
401 volatile-free basis) of the rock's mass had been removed by leaching. This calculation assumes
402 Zr behaved in a completely conservative manner, but similar numbers are arrived at using Ti and

403 Al contents. The volatile content of this rock is dominated by H₂O (Table 1) and are likely
404 primarily due to its clay content. The nontronite clay present as linings and infillings in void
405 spaces represents elements (mostly Fe and Si, with minor Mg) that were leached from primary
406 minerals but were redeposited rather than being lost to groundwater. The depletion of these
407 elements in the bulk rock, and the presence of relatively thin secondary smectite coatings, shows
408 that formation of the brown and tan secondary nontronite coatings does not account for all of the
409 Fe and Si leached from the parent rock. By contrast, as noted above, the presence of halloysite
410 and the immobility of Al suggests that any Al liberated from aluminous phases such as
411 plagioclase was immediately re-precipitated in this clay mineral.

412 Between 30 and 80 cm depth, sample bulk chemistry indicated ongoing removal of
413 ferromagnesian minerals (loss of Mg, Fe, and Ca) and passive accumulation of Al, Ti, P, and
414 remaining alkalis. This is consistent with the observation by XRD and EPMA that groundmass
415 was leached but plagioclase feldspar was mostly unaltered at these depths. Some plagioclase
416 crystals immediately around void spaces did show evidence of dissolution in the sample from 30
417 cm depth (Figure 7). As minimal depletion of Al was observed in this sample (Table 2), it
418 appears that Al released by feldspar dissolution was immediately precipitated in the Al-rich, Fe-
419 poor clay layers that were also observed in this sample (Figure 7).

420 As discussed above, two point XANES spectra from the 30 cm sample had large main
421 edge features resembling the titanomagnetite standard, but small pre-edge features more similar
422 to clay minerals. The identity of these mineral phases is unclear. These spectra could indicate
423 primary minerals undergoing alteration and intimately mixed with a secondary phase, or they
424 could be mixtures produced by analyzing thin clay coatings over embedded Fe-Ti oxides. The
425 XANES points from the 30 cm sample were all located on what appeared to be clay coatings

426 similar to the coatings in the deeper samples. However, any Fe oxides embedded in the clay
427 coatings (such as the clay coated titanomagnetites visible in Figure 6, and the titanomagnetites
428 embedded in the clay near the clay-groundmass boundary in Figure 7) would have contained the
429 highest mapped Fe contents and would therefore have presented likely targets for point analysis.

430 The total volatile content (LOI; Table 1) of the samples from between 30 and 80 cm did
431 not vary with depth, suggesting that the total clay mineral content also did not vary up-gradient
432 through this depth interval. Weathering in this depth range therefore appears to have primarily
433 taken the form of leaching of chemical components and precipitation of secondary minerals
434 including halloysite and nontronite.

435 By the 20 cm depth interval, 42% of the rock (by mass) had been removed by progressive
436 leaching, calculated on a volatile-free basis assuming Zr was completely conservative.
437 Plagioclase dissolution was observed in electron microprobe images of this sample (Figure 8),
438 and aluminous zones became more prominent in smectite clay linings. Element maps showed
439 that Fe-Ti oxides had been replaced by a Mg-rich phase (Figure 8), possibly a clay such as
440 montmorillonite. The bulk rock Al and Ti contents peaked at this depth (Table 1), indicating that
441 although these elements' host minerals such as feldspar and titanomagnetite were being etched
442 (Figure 8), the elements themselves had not been leached from the bulk rock. The volatile
443 content of this sample was higher than those beneath it (Table 1), indicating a higher clay
444 mineral content, which may account for a portion of the Al that had been leached from
445 plagioclase but was not lost from the rock. A broad 7 Å inflection seen in XRD patterns suggests
446 kaolinite may be present at this depth in addition to smectite (Baker 2017). Isolated Ti-rich zones
447 (Figure 8) may indicate that Ti from oxides was transported only on the micron scale, and re-
448 precipitated as oxides.

449 The 20 cm sample thus appears to represent a turning point in the weathering evolution of
450 this basalt, from leaching of ferromagnesian groundmass components to leaching of all primary
451 phases including feldspar, titanomagnetite, and apatite. It contained both nontronite and
452 kaolinite, suggesting that clay mineralogy was still controlled by local porosity at this stage, but
453 that open-system leaching was initiating in portions of the weathering rock. This sample also
454 represents the point above which none of the elements analyzed behaved in a conservative
455 manner, so that mass losses above this depth can no longer be evaluated quantitatively. However,
456 the extensive leaching of all primary phases, and the fact that 30-40% of the rock mass had
457 already been previously lost from the rock, suggest that the majority of the initial basalt mass
458 (and volume) had likely been leached from this sample. It is therefore not surprising that the
459 transition to a very different mineral assemblage occurs within a few cm.

460 No relict primary basaltic mineral phases were observed in the sample from 10 cm. Lacy
461 Fe oxides were observed in this sample (Figure 9) and Fe-rich clays were not, suggesting it is
462 likely that hematite is the predominant Fe-bearing phase in this sample. An increase in LOI
463 indicated increased accumulation of clay minerals, and bulk XRD analysis indicated that a
464 smectite was abundant in this sample (Baker 2017). However, none of the XANES point spectra
465 resemble the spectrum of Fe in a clay mineral, suggesting that smectites in this sample are all
466 low-Fe in composition. Blocky concentrations of lacy Fe oxides and of aluminosilicates,
467 probably clays, appeared to be pseudomorphic after feldspar laths (Figure 9). The XANES
468 spectra from this sample indicated that high-Fe spots contained hematite rather than a clay
469 mineral. However, the shallow saprolite and paleosol samples have been geochemically altered
470 by reaction with the capping basalt flow and Si-rich sediments, and contain excess Fe, Si, Mg,

471 and Mn (Baker 2017). The mineral phases containing these elements are likely to partly reflect
472 formation in a contact metamorphic environment rather than a pedogenic environment.

473 **Open system vs closed system alteration**

474 It has been proposed that nontronite clay forms by alteration of mafic parent rocks under
475 closed-system conditions, whereas open-system alteration is dominated by leaching of mafic
476 components such as olivine (Ehlmann et al. 2011) . The results presented here suggest that there
477 is not a sharp distinction between closed nontronite-forming systems and open leaching-
478 dominated systems, because the two processes may be contemporaneous where water-rock ratio
479 is limited.

480 A completely closed system is not likely to be a common occurrence in nature. In the
481 case of aqueous weathering of basalt, unless alteration solely involves magmatic fluids (Meunier
482 et al. 2012), the system must at least be sufficiently open to permit the altering fluids to enter.
483 However, where fluid transport is limited by permeability, fluid chemistry will be controlled by
484 equilibrium with the mineral phases present (Navarre-Sitchler et al. 2011) and overall water-rock
485 ratio will be low. These are probably the conditions implied by “poor drainage” – the description
486 used by Allen and Scheid (1946). Nontronite formation is promoted by anoxic or reducing
487 conditions and relatively high pH (Harder 1976), conditions that suggest that fluid chemistry is
488 controlled by basalt dissolution. The behavior of Fe in solution is strongly affected by oxidation
489 state, because the solubility of Fe(II) is much greater than that of Fe(III). As a result, Fe will
490 leach from the weathering rock under anoxic conditions, whereas under oxidizing conditions it
491 will be retained as ferric oxides or hydroxides. To capture this difference, Babechuk et al. (2014)
492 developed the Mafic Index of Alteration (MIA), a weathering index that includes mafic minerals
493 containing Fe and Mg. The MIA can be separately calculated for oxidizing conditions (MIA-O),

494 when Fe is treated as an immobile element, or for reducing conditions (MIA-R), when Fe is
495 treated as a mobile element. As with other weathering indices, MIA values increase with
496 increasing alteration of the rock; a specific increase in MIA-R values as opposed to MIA-O is
497 indicative of Fe leaching.

498 At Trinidad, MIA-R values increased upwards in the transect from 80 to 20 cm (Table 1),
499 whereas MIA-O values remained relatively steady. This suggests Fe was leached from the rock
500 at depths greater than 20 cm, and therefore that conditions in the weathering fluid were anoxic or
501 reducing below this depth.

502 The Trinidad samples from between 30 and 80 cm depth have lost 30-40% of the initial
503 parent rock mass (assuming that Al, Ti, and Zr behave in a conservative manner) including Fe,
504 Mg, and alkalis. Given this, the Trinidad weathering system cannot reasonably be described as
505 closed; however, transport of fluid and dissolved ions out of the weathering system must have
506 been sufficiently slow to permit intermittent nontronite saturation. The physical form of the
507 nontronite as coatings and tendrils in open void spaces suggests that the clay may have
508 precipitated from thin films of water coating rock surfaces. A reasonable physical model for the
509 nontronite forming system at Trinidad would be periodic water saturation followed by slow
510 drainage of porewater to the groundwater system. Precipitation of clay minerals from successive
511 pulses of porewater would be consistent with the oscillatory zoning observed in the Mg and Fe
512 content of the clay linings, which suggest that the fluid composition was not continuously
513 buffered by equilibrium with dissolving primary phases.

514 Although the Trinidad nontronite-forming system was not closed to leaching of major
515 components, it appears to have coexisted with the primary basaltic mineral assemblage as long as
516 that assemblage remained. The appearance of more aluminous smectites and the disappearance

517 of nontronite coincided spatially with the onset of dissolution of more resistant phases including
518 feldspar, apatite, and titanomagnetite. In the Trinidad saprolite transect, this transition took place
519 in a relatively narrow zone between 10 and 20 cm depth. The stable neoformed Fe phase above
520 20 cm depth appeared to be hematite (Figure 9), indicating that weathering conditions changed
521 from anoxic to oxic across this transition zone. Extensive leaching of most analyzed elements
522 from the 10-15 cm samples (Table 1) suggest that the onset of oxidizing weathering conditions
523 corresponded with an increase in water-rock ratio. Future work examining stable isotope ratios
524 may illuminate this further.

525 Initially, the primary rock porosity consisted of vesicles and a network of cooling cracks
526 that effectively controlled rock permeability (Meunier et al. 2007; Mazurier et al. 2016). The
527 cracks permitted limited transport of water through the rock and into vesicles, and limited
528 transport of dissolved ions out of the weathering system. Over time, dissolution of the primary
529 rock matrix and precipitation of clay infillings altered this primary permeability network. The
530 nontronite-forming conditions would have existed for as long as drainage of porewater was
531 limited by permeability. The transition from nontronite formation to formation of aluminous
532 smectites, and from anoxic to oxic conditions, was therefore most likely controlled by the
533 development of permeability in the weathering basalt and the transition to a completely open
534 system.

535

536 **Mass contributions to surface water over time**

537 If the Trinidad saprolite is typical of weathering of CRB in the central Columbia Plateau,
538 this suggests that emplacement of fresh basalt flows would have been followed first by an initial

539 pulse of Mg and K to the groundwater system, then by slower release of Fe, Si, Na, and Ca.
540 Although nontronite + halloysite was the stable clay mineral assemblage present during early
541 weathering, relatively little leached Fe, Mg, or Si were retained as clay coatings within the rock,
542 whereas dissolved Al was apparently precipitated immediately as halloysite. Precipitation of
543 clays in cracks may have temporarily offset increases in subsurface permeability caused by
544 leaching, as anoxic or reducing conditions persisted below the surface during leaching of
545 approximately one third of the rock mass.

546 The switch to high water-rock ratio and to oxidizing conditions during subsurface
547 weathering coincided, likely due to increased permeability as a result of continuing dissolution of
548 the rock matrix. This later phase of weathering corresponded with release of Al, Ti, P, and
549 associated alkalis by dissolution of more resistant phases including plagioclase, apatite, and
550 titanomagnetite. Leaching of Fe ceased from the oxidized portion of the weathering column.
551 Release of Al and Ti to groundwater would have been controlled by their limited solubility, but
552 bulk chemistry suggests they were transported on at least a meter scale. This weathering
553 transition is also marked by the appearance of aluminous smectites, minor kaolinite, and
554 hematite as stable secondary phases. The sample from 20 cm depth appears to capture this
555 transition point.

556 The transport and fate of dissolved ions released to groundwater through basalt
557 weathering would depend on local hydrologic conditions. The present-day Columbia River is a
558 gaining stream through much of the Columbia Plateau (Whitehead 1994), although local
559 hydrologic connections between surface and groundwater can be complex due to the presence of
560 many low-permeability layers (Burt et al. 2009; Lindsey et al. 2009; Lite 2013). In the late
561 Miocene, much of this subsurface structure was in existence, but the surface hydrology of the

562 Columbia differed considerably from the present day. The emplacement of extensive CRB lava
563 flows had repeatedly disrupted established drainages, damming and redirecting streams including
564 the Columbia (Reidel and Tolan 2013). Redevelopment of drainages was affected by ongoing
565 tectonism as well as the architecture of the newly emplaced basalt. During late Wanapum time
566 (circa 14.5 Ma), the time period of weathering at Trinidad, the Columbia's course through the
567 western Columbia Plateau apparently consisted of a series of interconnected shallow lakes rather
568 than a well-developed river channel (Reidel and Tolan 2013). Interbedded sedimentary deposits
569 indicate wet surface environments in the immediate vicinity of the Trinidad paleosol (Ebbinghaus
570 et al. 2014). Thus, communication between groundwater and surface water was likely complex
571 during Wanapum time, and dissolved ions may have followed tortuous paths before discharging
572 to the Pacific Ocean (Taylor and Lasaga 1999).

573

574 **Potential applications to Mars**

575 Some clay-bearing outcrops on Mars contain up to 55% nontronite as estimated from
576 remote sensing (Poulet et al. 2008), much more of this clay than do any portions of the Trinidad
577 saprolite. As discussed above, although nontronite is present in the Trinidad saprolite, the
578 majority of Fe and Si dissolved from the parent basalt was lost to leaching rather than being
579 redeposited in secondary phases. If the martian nontronite-forming system was more poorly
580 drained than the Trinidad saprolite, then dissolved Mg, Fe, and Si may have re-precipitated in
581 clay minerals, resulting in the high observed nontronite abundances. Although Fe-rich martian
582 clays are often referred to in the literature as “nontronite”, some spectra suggest these clays may
583 contain both Fe(II) and Mg (Bishop et al. 2008), unlike typical terrestrial nontronites which
584 dominantly contain Fe(III). At Trinidad, bulk compositions suggest very early leaching of Mg

585 from the parent rock, and the Fe-Mg compositional banding observed in clay linings may
586 preserve information about this phase of early weathering. Ongoing studies of very early-stage
587 basalt weathering may help illuminate the relative timing of leaching and clay formation during
588 this phase.

589 The clay-bearing outcrops on Mars may have formed by surface weathering in a
590 pedogenic environment (Carter et al. 2015), although it has also been proposed that they may
591 have formed under other conditions (Murchie et al. 2009; Ehlmann et al. 2011). If the aluminous
592 clay outcrops are pedogenic in origin, then the relationships between leaching, element release,
593 and clay formation at Trinidad may be useful in interpreting the sequence of events on Mars. At
594 both high and lower abundance nontronite occurrences, the clay is accompanied by primary
595 minerals such as plagioclase (Poulet et al. 2008); this is consistent with the observed sapolite
596 mineralogy at Trinidad. Specific dissolution and leaching patterns of particular elements would
597 have depended on the mineralogy of the parent rock undergoing weathering. Furthermore, as
598 observed in ancient terrestrial soils formed before Earth's atmosphere became oxidized, soluble
599 Fe(II) would be leached in a similar manner to Mg (Babechuk et al. 2014). However, the
600 transition from Fe-Mg clays to aluminous smectites would still be expected to mark the transition
601 to fully open-system weathering with loss of all primary mineral phases and extensive leaching
602 of most elements.

603 Oxidized Fe compounds are present on the martian surface. Aqueous alteration observed
604 in martian meteorites appears to have sometimes produced reduced phases such as siderite and
605 sulfide minerals, suggesting alteration under moderately reducing conditions; however, other
606 alteration assemblages contain ferric (oxyhydr)oxides, indicating more oxidizing conditions
607 (Bridges et al. 2001). Therefore, iron may have been leached during weathering of martian rocks

608 under some conditions. The chemical components leached from the parent basalt would have
609 been released to the martian paleo-hydrologic system, perhaps ultimately contributing to the salt
610 and oxide content of the present-day surface.

611

612 **Implications**

613 Patterns of element leaching and clay mineral deposition in a saprolite at Trinidad, WA
614 show that precipitation of nontronite clay does not necessarily indicate weathering occurred in a
615 chemically closed system; nontronite is stable even in rocks that have undergone significant
616 leaching of major elements. Instead, formation of nontronite may indicate weathering under
617 conditions of limited permeability. As permeability increases, aluminous clays such as kaolinite
618 and/or montmorillonite become the stable secondary phase(s) present, but nontronite may persist
619 in pockets of limited permeability. These observations suggest that the mineralogy and
620 distribution of secondary clays formed by basalt weathering are likely to be strongly influenced
621 by the initial physical characteristics of the basalt, such as the presence of cracks, joints, pillows,
622 or other structures – many of which are in turn influenced by the conditions of lava
623 emplacement. Release of dissolved ions and fixation of CO₂ may thus also depend on these
624 factors.

625 The observed mineralogy of clay-bearing outcrops on Mars is consistent with similar
626 patterns of mineral dissolution to those observed at Trinidad, although more detailed information
627 on the martian outcrops is needed to determine whether saprolites are indeed a good analog for
628 them. The high clay abundances observed in some martian outcrops may indicate more
629 secondary mineral deposition and less element leaching (e.g., a more closed system) than

630 observed in the Trinidad saprolite, suggesting that martian clay outcrops may preserve valuable
631 information about martian hydrologic conditions in the ancient past.

632

633

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855

Figure Captions

856 Figure 1: FTIR spectra of bulk saprolite samples, showing bulk clay mineralogy through the
857 transect. In samples from depths of 30, 60, and 80 cm below the contact, halloysite is present as
858 identified by peaks at 3695, 3620, 1100, 1035, and 915 cm^{-1} . Although smectites were identified
859 in clay separates from these samples, smectite absorption bands are not identifiable in the bulk
860 samples, suggesting halloysite is the more abundant phase. In samples from 10, 15, and 20 cm
861 depth, characteristic absorption bands of nontronite (3565 and 815 cm^{-1}) and/or montmorillonite
862 (3620, 875 cm^{-1}) are present.

863 Figure 2: Photo of vermicular tendrils in clay infilling in 80 cm sample (image 1 mm wide).

864 Figure 3: Clay tendril cross-section in 80 cm sample with relatively Fe-rich, Mg-poor center and
865 Mn oxide coating.

866 Figure 4: Complex Mg-Fe zonation in clay infill in 80 cm sample. Images are 120 μm across.

867 Figure 5: Clay-filled embayment in vesicle edge in 80 cm sample. Brighter rings in Fe map
868 (right) are cross-sections of vermicular tendrils with Fe-rich cores. Groundmass texture is also
869 evident in this image, with skeletal Fe-Ti oxides.

870 Figure 6: BSE image of clay-lined vesicle in 60 cm sample, with clay-coated relict plagioclase
871 laths and clay tendril cross sections (rounded / ovoid islands) and blocky embayments suggesting
872 this void space was expanded by dissolution. Bright skeletal crystals are titanomagnetite.

873 Figure 7: Two examples of zoned clay coating in 30 cm sample, with complex Fe, Al, and Mg
874 zonation.

875 Figure 8: Plagioclase dissolution and pervasive alteration of groundmass in 20 cm sample.
876 Element maps do not overlap context map, but show a zoomed area of a nearby contact between
877 original Fe-rich zoned clay infilling and altered groundmass. Patchy Ti distribution indicates
878 dissolution of primary Fe-Ti oxides, and skeletal Mg pseudomorphs suggest replacement of Fe-
879 Ti oxides by a low-Fe clay mineral.

880 Figure 9: Al-rich pseudomorph and feathery Fe oxides surrounded by Mg silicates in 10 cm
881 sample.

882 Figure 10: XANES (a), first derivative of XANES (b), EXAFS (c), and Fourier transformed
883 spectra (d) of picked clay infillings from 80 cm sample. Also shown is the UI-Garfield high-Fe
884 nontronite (Baker and Strawn 2012) for reference. Inset in (a) shows pre-edge spectrum in
885 greater detail. Arrow in (b) indicates reduced Fe shoulder. The Fourier transformed XAFS
886 spectrum (d) shows peaks corresponding to atomic shells around an octahedrally coordinated Fe
887 atom: six oxygen atoms at 2 Å, three first-neighbor Fe atoms at 3 Å and four Si neighbors at 3.25
888 Å, and additional backscattering O, Si, and Fe atoms at greater distances (Baker and Strawn
889 2012).

890 Figure 11: Synchrotron microprobe XANES (a) and first derivative of XANES (b) for high-Fe
891 clays in 60 and 80 cm samples. Bulk spectra for brown and black clays are shown for
892 comparison above and below point XANES spectra. Dotted line in (b) indicates reduced Fe
893 shoulder observed in the black clay. These data show that the analyzed clays are all high-Fe
894 nontronites. No Fe is present in an Al-rich or Mg-rich clay, a trioctahedral smectite, or in Fe
895 (oxyhydr)oxides.

896 Figure 12: Synchrotron microprobe XANES (a) and first derivative of XANES (b) for
897 titanomagnetite from 80 cm sample, hematite and magnetite standards, and high-Fe points in 30
898 cm sample (unlabeled curves). These results indicate that the analyzed high-Fe points in this
899 sample likely contain primary titanomagnetite.

900 Figure 13: Synchrotron microprobe XANES (a) and first derivative of XANES (b) for high-Fe
901 points in 20 cm sample and black clay separate. These results indicate that the high-Fe points
902 analyzed in this sample contain a high-Fe nontronite with a reduced component.

903 Figure 14: Synchrotron microprobe XANES (a) and first derivative of XANES (b) for high-Fe
904 points in 10 cm sample and hematite standard. These results indicate that the high-Fe points
905 analyzed in this sample all contain hematite; no Fe-rich clays were detected.

906

907 Table 1. Bulk compositions in weight percent of basalt, saprolite and paleosol samples (Baker 2017), and calculated mafic index of
 908 alteration for both reducing (MIA-R) and oxidizing (MIA-O) conditions (Babechuk et al. 2014). All values are in wt % unless noted.
 909 LOI = loss on ignition. Bas = parent basalt. P = paleosol.

Sample	SiO ₂	Al ₂ O ₃	Fe ₂ O ₃	MgO	CaO	Na ₂ O	K ₂ O	TiO ₂	P ₂ O ₅	MnO	LOI	C	Sum	Zr (ppm)	MIA -O	MIA- R
P	59.22	5.19	18.96	0.95	1.11	0.02	0.03	0.34	0.05	0.75	13.2	0.42	99.95	34	79	24
10 cm	46.60	13.58	15.47	1.94	2.92	0.71	0.08	0.53	0.03	0.02	18.0	0.4	99.95	68	67	39
15 cm	45.75	13.76	16.43	1.77	3.03	0.84	0.08	0.49	0.05	0.03	17.7	0.5	99.95	64	68	39
20 cm	48.13	21.67	3.33	1.46	5.70	2.62	0.31	5.02	1.15	0.06	10.2	0.17	99.82	319	56	51
30 cm	46.54	18.80	9.31	0.92	6.39	3.14	0.35	4.39	1.02	0.17	8.7	0.16	99.87	278	56	43
60 cm	46.66	19.32	8.71	0.92	6.29	3.08	0.34	4.51	1.05	0.13	8.7	0.06	99.87	282	56	44
80 cm	46.21	17.54	11.20	1.01	6.40	3.00	0.33	4.09	0.96	0.18	8.8	0.08	99.89	250	56	40
Bas	50.93	13.64	12.93	3.38	8.60	2.92	1.31	3.12	0.66	0.18	2.0	-	99.85	193	42	26

910

911

912 Table 2. Mobility ratios calculated on a volatile-free basis with respect to Zr from the data in Table 1.

	SiO ₂	Al ₂ O ₃	Fe ₂ O ₃	MgO	CaO	Na ₂ O	K ₂ O	TiO ₂	P ₂ O ₅	MnO
T 0 P	5.600	1.160	7.324	0.595	-0.267	-0.961	-0.870	-0.381	-0.570	22.652
T 10 cm	1.597	1.826	2.396	0.629	-0.036	-0.310	-0.827	-0.518	-0.871	-0.685
T 15 cm	1.709	2.042	2.832	0.579	0.062	-0.132	-0.816	-0.526	-0.772	-0.497
T 20 cm	-0.428	-0.039	-0.844	-0.739	-0.599	-0.457	-0.857	-0.027	0.054	-0.798
T 30 cm	-0.366	-0.043	-0.500	-0.811	-0.484	-0.253	-0.815	-0.023	0.073	-0.344
T 60 cm	-0.373	-0.031	-0.539	-0.814	-0.499	-0.278	-0.822	-0.011	0.089	-0.506
T 80 cm	-0.300	-0.007	-0.331	-0.769	-0.425	-0.207	-0.806	0.012	0.123	-0.228

913

Figure 1

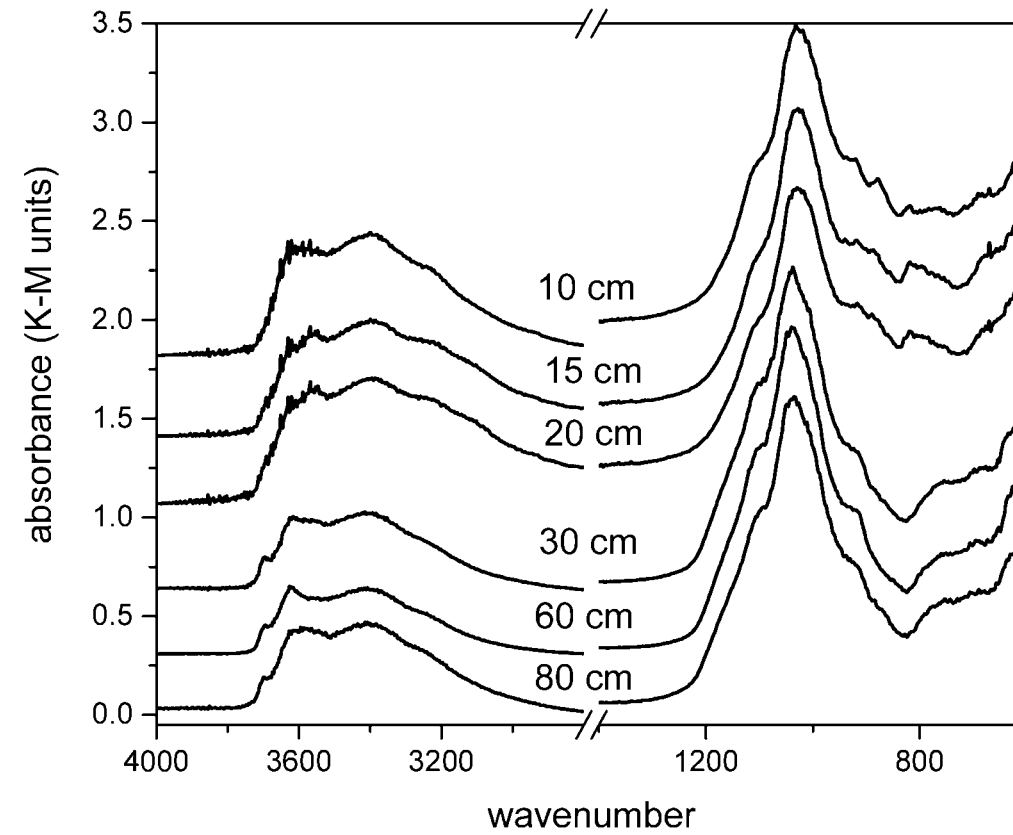


Figure 2

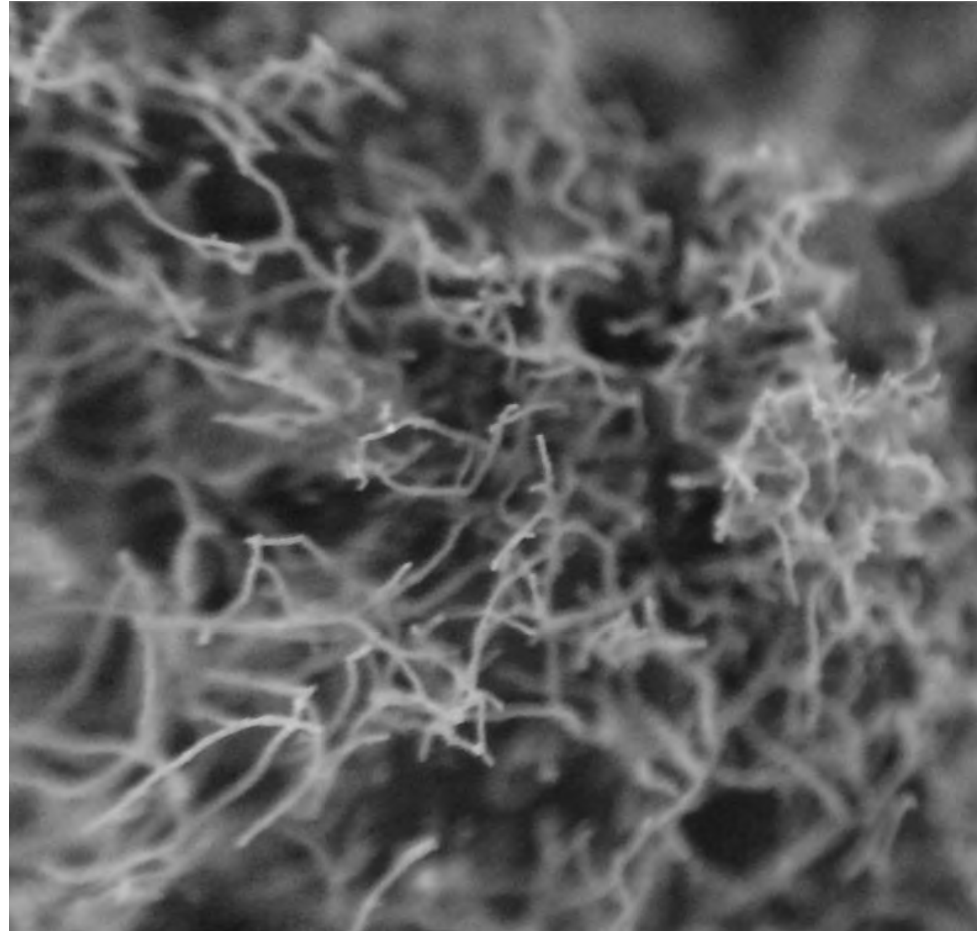


Figure 3

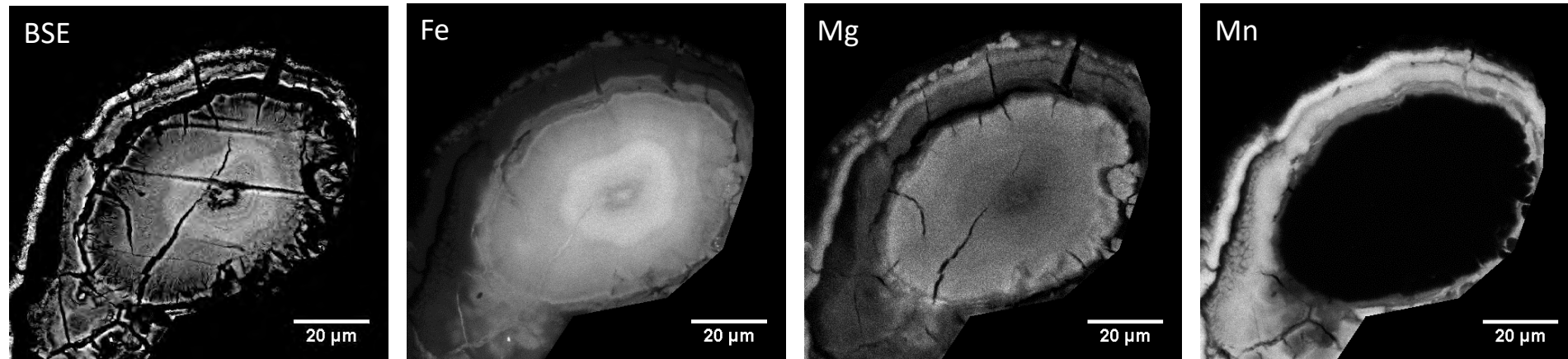


Figure 4

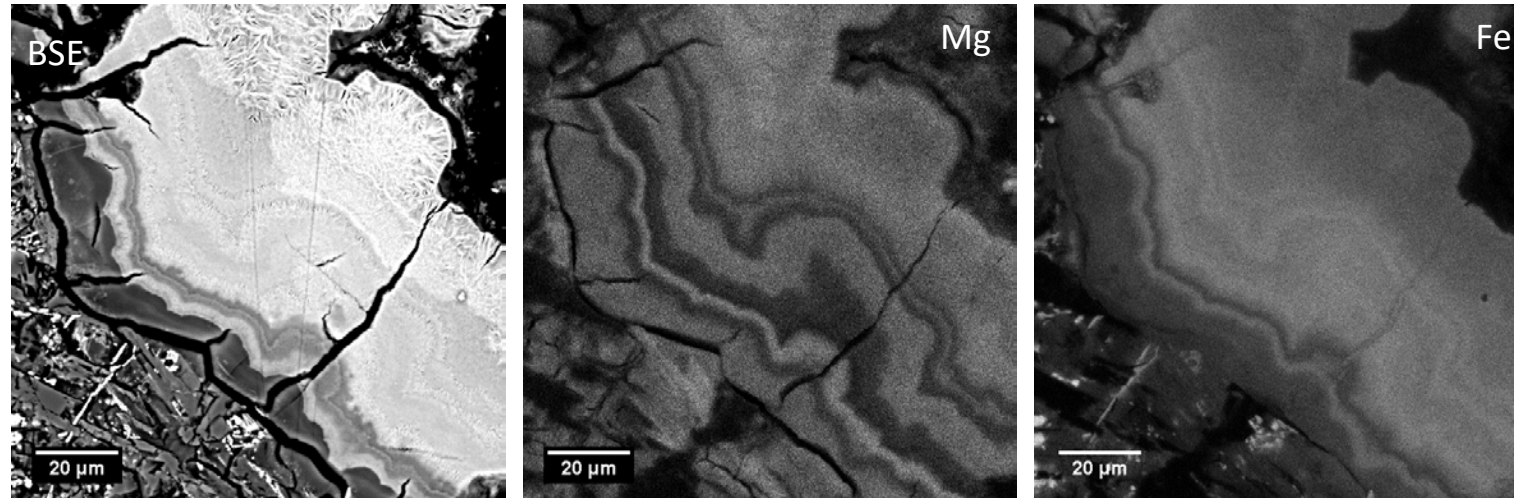


Figure 5

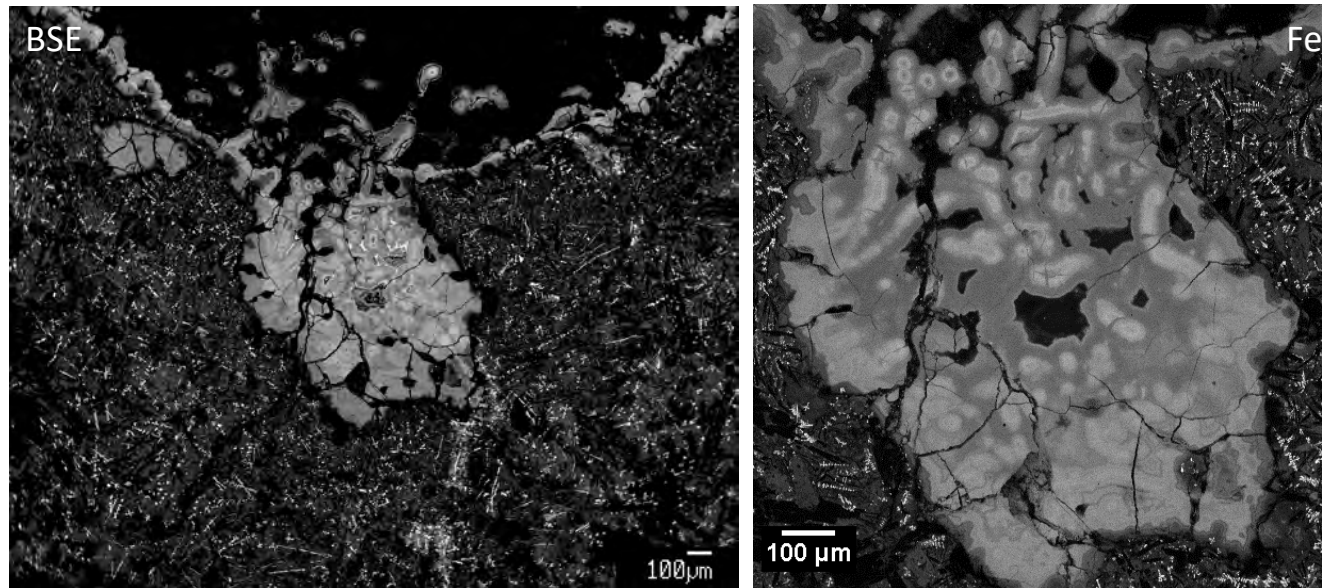


Figure 6



Figure 7

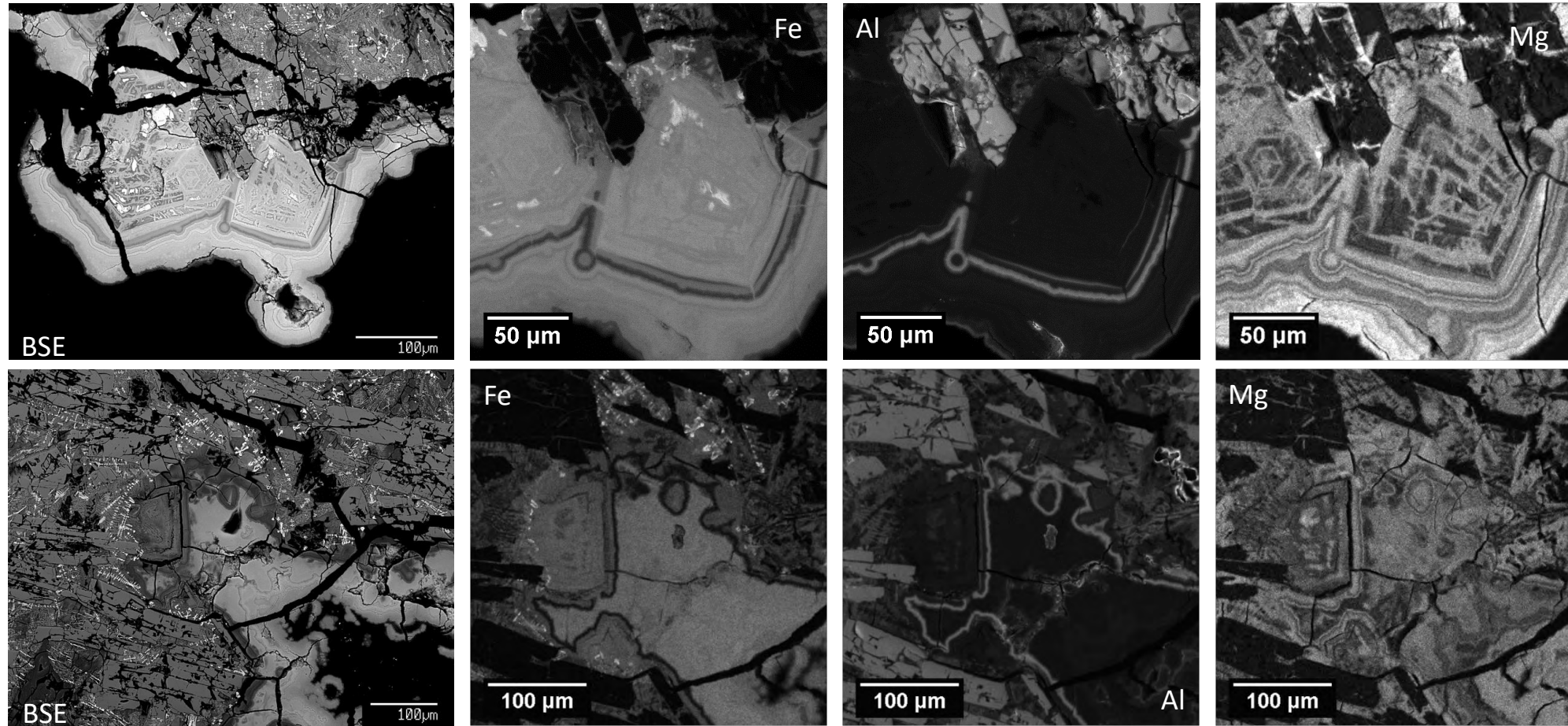


Figure 8

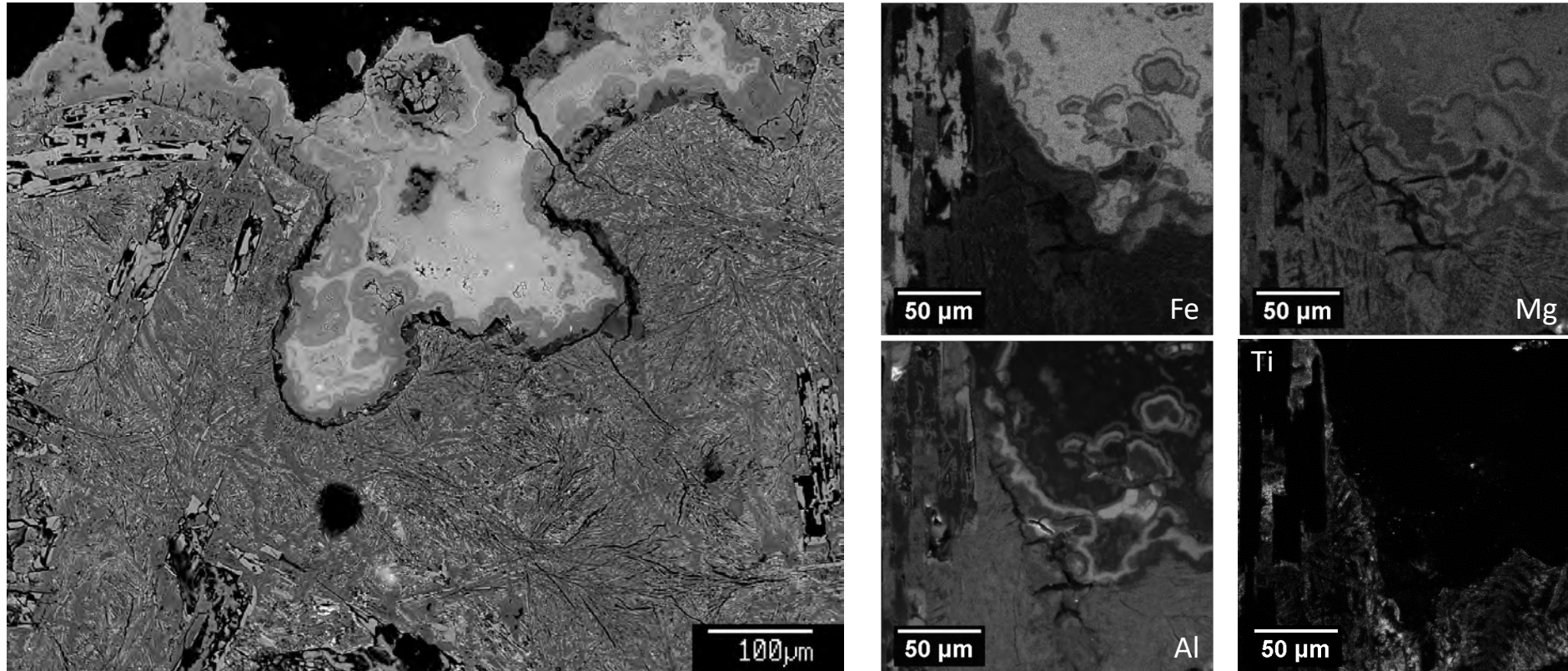


Figure 9

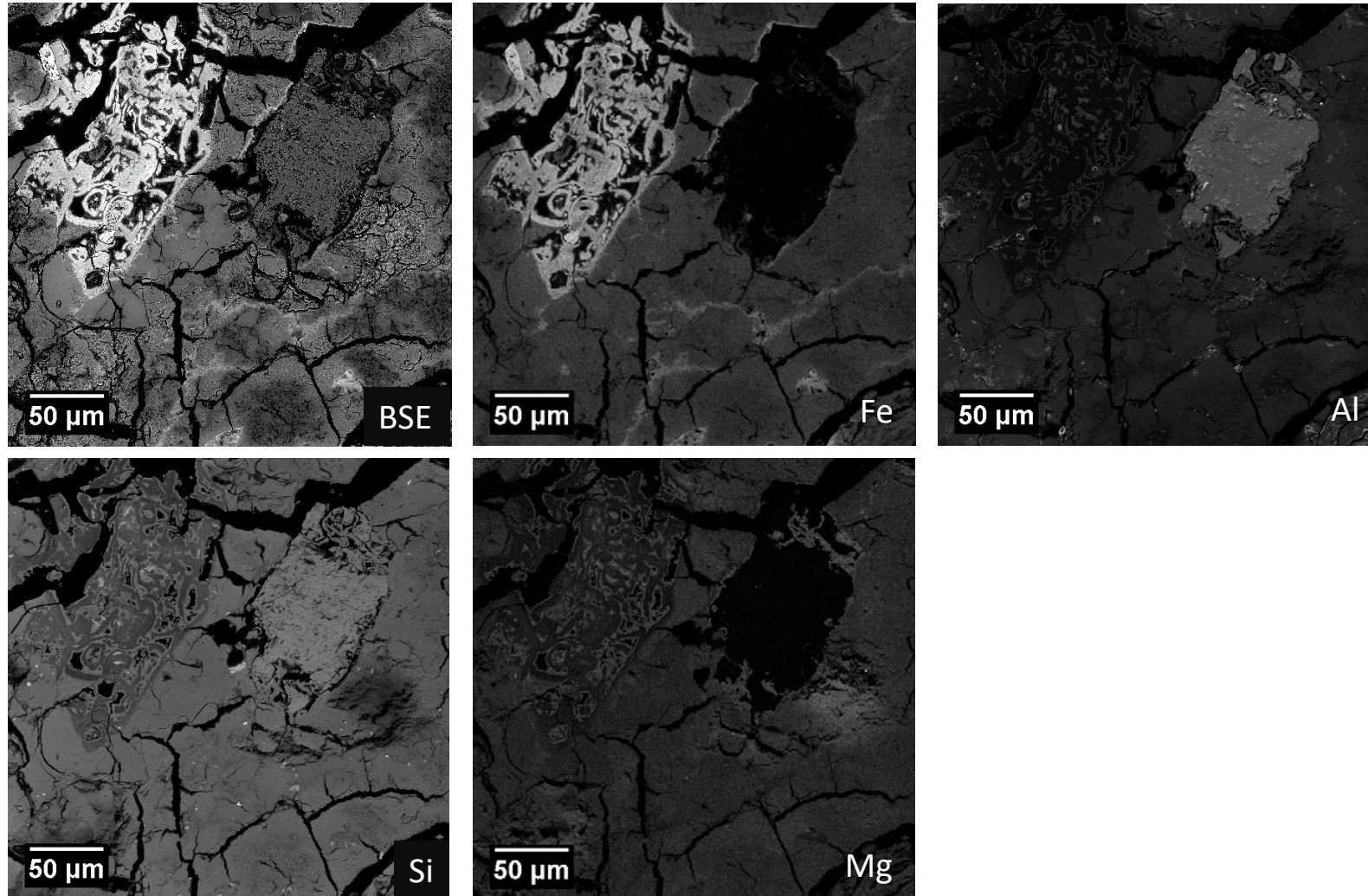


Figure 10

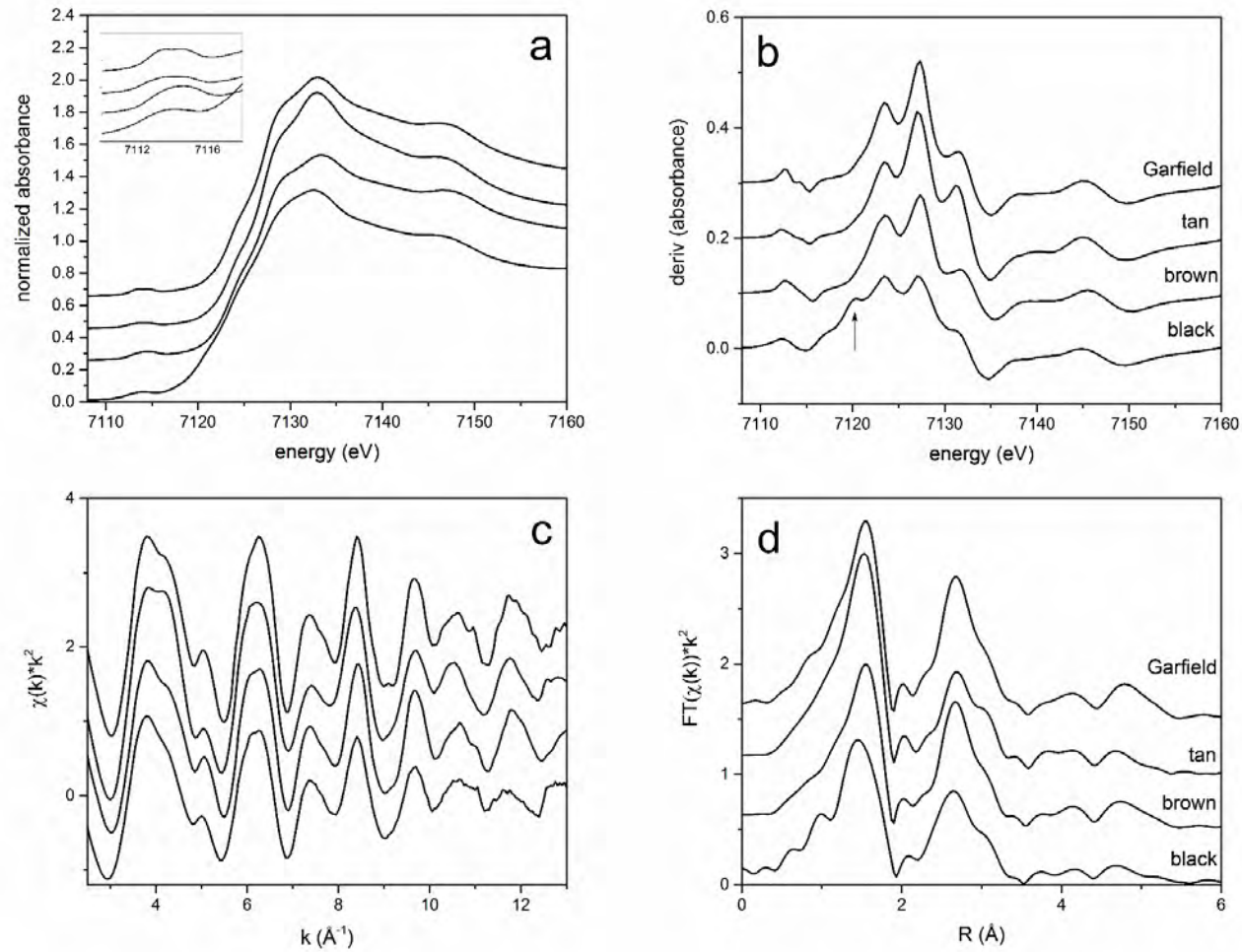


Figure 11

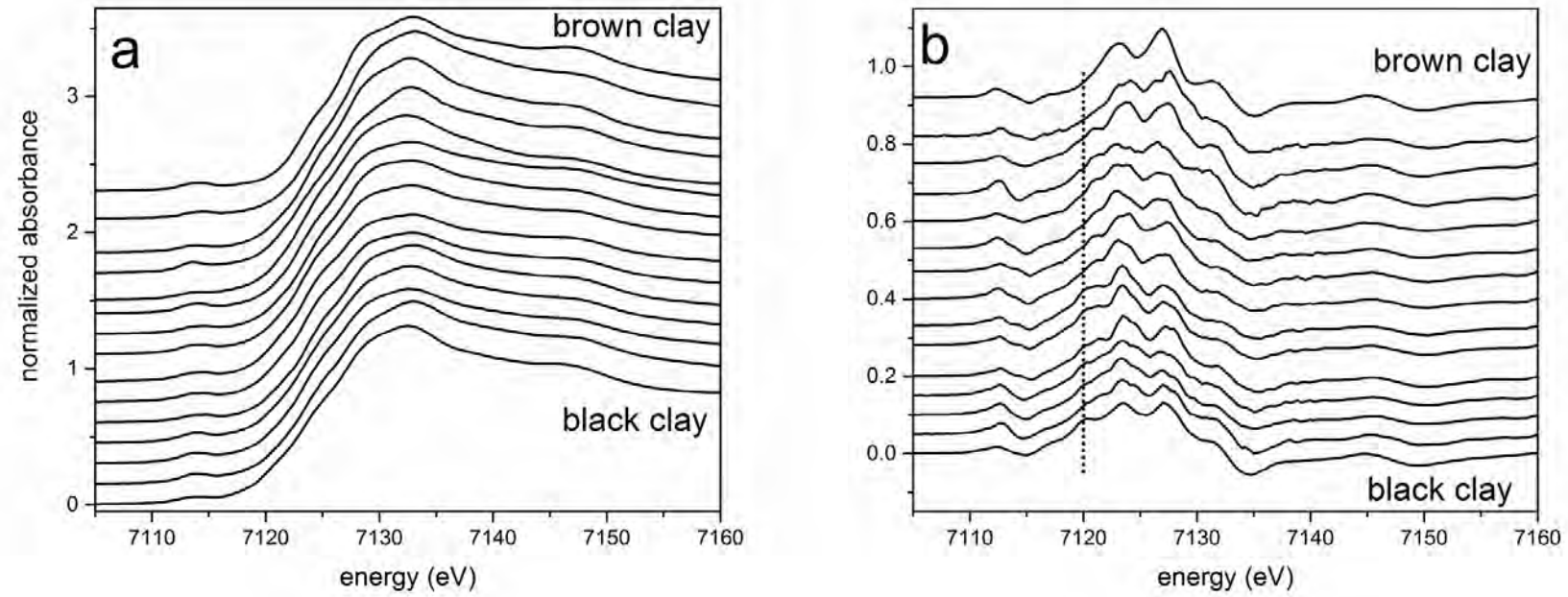


Figure 12

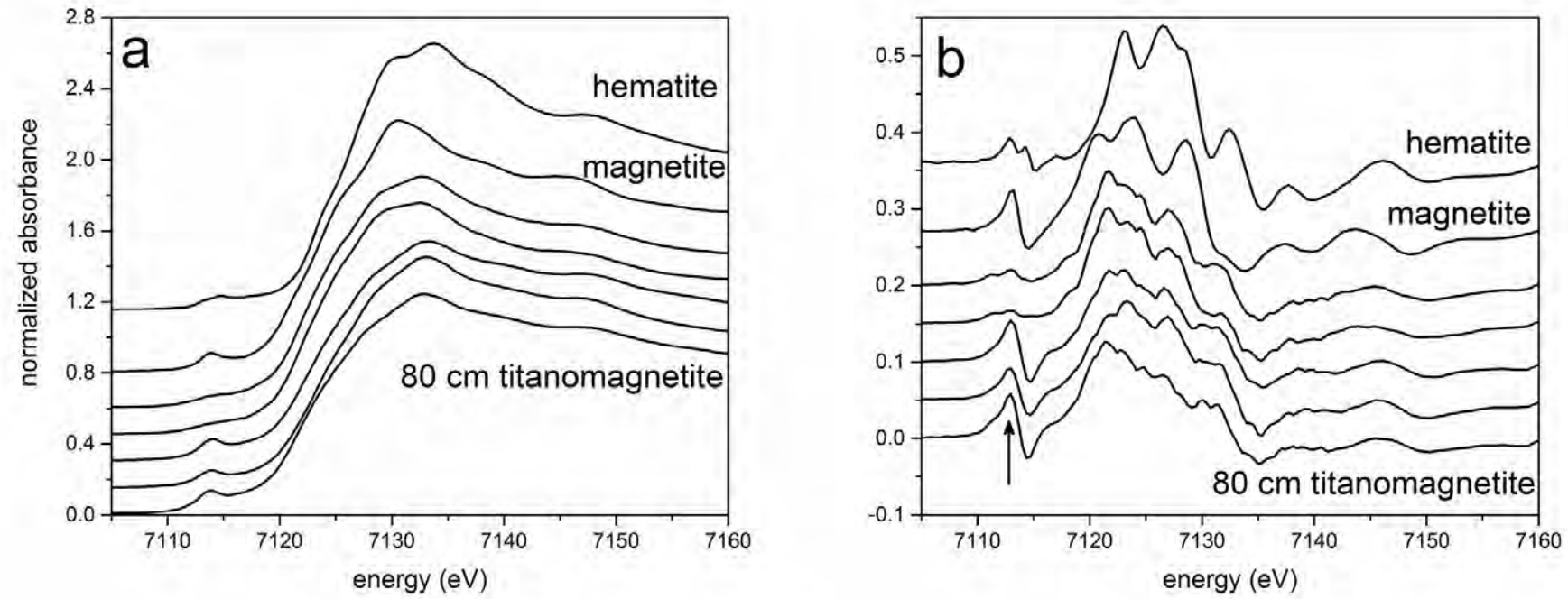


Figure 13

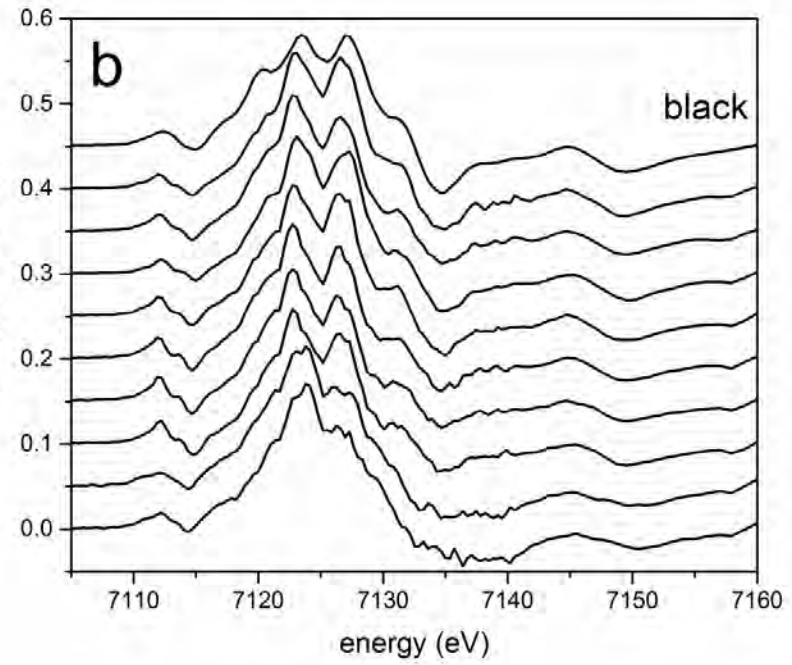
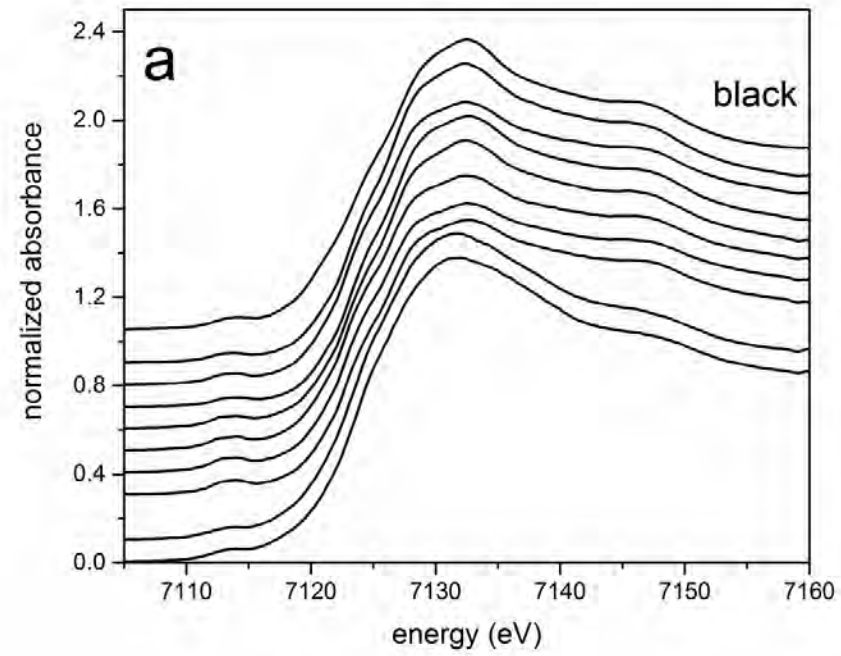


Figure 14

