

34 INTRODUCTION

35 The transport, burial, and transformation of organic carbon along active margins is an
36 important but still poorly constrained component of the global biogeochemical carbon cycle.
37 Along active continental-oceanic tectonic boundaries associated with small mountainous rivers,
38 such as Cascadia, Alaska, Taiwan, and South America, exported kerogen (fossil carbon) often
39 dominates riverine sediments, and can constitute as much, and potentially more than 75% of the
40 total particulate organic load (Leithold et al. 2006; Drenzek et al. 2009; Blair et al. 2010; Hilton
41 et al. 2010). High sediment transport rates, coupled with close proximity to sediment sources,
42 promotes the efficient burial and preservation of organic carbon in the marine environment.
43 Kerogen forms by diagenetic to metasomatic processes imposed on organic matter, resulting in a
44 highly recalcitrant and structurally complex compound capable of persisting in the sedimentary
45 record on geologic time scales and likely exceeding the mass of carbon in living organic matter
46 by four orders of magnitude (Durand 1980). Understanding the high-pressure, high-temperature
47 properties of kerogen will lead to a better understanding of the behavior of kerogen during
48 subduction.

49 Raman spectroscopy has been widely used to characterize carbonaceous material (CM) in
50 Earth and planetary materials for some time. Detection of biosignatures in CM is a primary focus
51 in the study of early life on Earth (Pasteris and Wopenka 2002; Schopf et al. 2002; Schopf and
52 Kudryavtsev 2005; Kremer et al. 2012) with potential application to Mars (e.g. Ellery et al. 2004;
53 Dunn et al. 2007; Dartnell et al. 2012; Steele et al. 2012), where the ExoMars rover, expected to
54 launch in 2018, will deploy a Raman spectrometer to analyze potential organic compounds using
55 a 532 nm continuous excitation wavelength (Courrèges-Lacoste et al. 2007; Rull et al. 2013;
56 Hutchinson et al. 2014). Characterization of CM by Raman spectroscopy has also been used to

57 infer the thermal history of meteorites (Quirico et al. 2003; Hochleitner et al. 2004; Bonal et al.
58 2006; Busemann et al. 2007; Matsuda et al. 2010) and interplanetary dust particles (Wopenka
59 1988; Sandford et al. 2006), as well as to characterize riverine particulate organic carbon (Galy
60 et al. 2007, 2008; Bouchez et al. 2010). Interpretation of CM Raman spectra in Earth and
61 planetary materials requires caution because of its sensitivity to the laser (Kagi et al. 1994;
62 Morishita et al. 2011), ionizing radiation (Dartnell et al. 2012) and polishing in sample
63 preparation (Ammar et al. 2011).

64 In terrestrial meta-sedimentary rocks, the evolution of CM crystallinity measured by
65 Raman spectroscopy has been used to develop a geothermometer of peak metamorphic
66 temperature in the 300-700 °C range (Wopenka and Pasteris 1993; Beyssac et al. 2002a, 2002b;
67 Rahl et al. 2005). Application to lower grade metamorphism or metasomatism at 25-300 °C is
68 not as well developed, although there has been recent progress (Lahfid et al. 2010). In general,
69 thermometry studies rely on the influence of temperature on Raman spectra of the ordered G
70 (graphite) band at $\sim 1580 \text{ cm}^{-1}$ and disordered or defect bands, D1 (1350 cm^{-1}) and D2 (1620 cm^{-1}).
71 Because the two primary defects bands are absent in well-crystallized graphite, the evolution
72 of ratios $R1 = [D1/G]_{\text{intensity}}$ and $R2 = [D1/(G+D1+D2)]_{\text{area}}$ estimate the degree to which graphitic
73 ordering has occurred. In metamorphic belts, there is a strong correlation between the R2 ratio
74 and peak metamorphic temperature based on mineral assemblages (e.g. Beyssac et al. 2002b).

75 Whereas geothermometry of CM in metasediments examines CM that has been subjected
76 to simultaneous pressures and temperatures on geologic timescales, experiments can only
77 examine essentially instantaneous effects. The graphitization of CM has been studied in heating
78 experiments (e.g. Beny-Bassez and Rouzaud 1985; Rouzaud and Oberlin 1989). Because the
79 effect of pressure promotes the graphitization process, metamorphic samples display

80 graphitization features at significantly lower temperatures than can be reproduced in the
81 laboratory (Huang et al. 2010). Furthermore, laboratory heating experiments up to 600 °C on
82 kerogen not previously subjected to regional metamorphism showed complete reversibility in the
83 temperature dependence of the Raman shift of G, D1, and D2 bands to lower wavenumbers
84 (Huang et al. 2010), despite evidence for bond-breaking changes associated with changes in the
85 intensity ratio of D2/G bands above ~350 °C (Huang et al. 2010). Pressure shifts the first-order
86 bands of kerogen in the opposite sense, to higher wavenumbers, so it might be expected that
87 simultaneous pressure and temperature will influence the temperature effect differently than at
88 room pressure. Although some experiments have examined the effect of pressure and
89 temperature on kerogen Raman spectra post-heating ex situ (e.g. Beyssac et al. 2003a), the
90 evolution of kerogen Raman has not yet been explored at simultaneous high pressures and high
91 temperatures.

92 Characterizing the evolution of CM in mixtures of marine and terrestrial sediments in
93 near-shore sediments at subduction zones has the potential to constrain fluxes of organic carbon
94 removal from the biosphere (e.g. Mackenzie et al. 2004; Hayes and Waldbauer 2006). However,
95 primarily because CM is dark brown to black and highly absorbing of typical Raman-laser
96 excitation wavelengths, it is imperative to characterize the influence of laser-induced heating on
97 interpretation of the character of CM Raman spectra (Everall et al. 1991; Kagi et al. 1994;
98 Morishita et al. 2011).

99 In this study, we first examined the effect of laser-induced heating by the Raman
100 excitation source in the blue-green region (458 nm) on the evolution of Raman D and G bands of
101 kerogen. We examine isolated kerogen extracted from rocks of an accreted shale-rich mélange at
102 laser powers ranging from 0.05 to 6 mW, and compare the laser-heating effects to controlled

103 heating over the temperature range of 25-500 °C using a laser power below the damage
104 threshold, determined in the first part of the study. To study the effect of pressure on the high-
105 temperature stability and vibrational properties of kerogen, we measured first-order Raman
106 spectra at simultaneous *P-T* conditions up to 3.2 GPa and ~400 °C, within the range of colder
107 subduction geotherms where kerogen is most likely to persist to deeper depths. The spectra were
108 analyzed for the *P-T* evolution of primary vibrational band widths, intensities, and ratios in order
109 to shed light on the molecular evolution of carbon along *P-T* paths typical of subduction
110 environments.

111

112 **METHODS**

113 Kerogen was extracted by HF methods (Durand and Nicaise 1980) from fine-grained,
114 shale-rich mélangé (Franciscan Complex) of the Eel River watershed in northern California
115 (Blair et al. 2003). Although this extraction can potentially alter or dissolve some organic
116 compounds (Durand and Nicaise 1980), these effects were minimized by the maturity of sample
117 and room-temperature digestion in an inert atmosphere. Although part of the uplifted
118 accretionary wedge material associated with the coastal mountain region, the kerogen examined
119 in this study has experienced little if any metamorphic-grade alteration (Blair et al. 2004;
120 Leithold et al. 2006). The mature, isolated kerogen used in this study exhibits first-order D and
121 G-bands very similar to petrogenic carbon from riverine sediment loads (Galy et al. 2008;
122 Bouchez et al. 2010).

123 Raman spectra were obtained using 458 nm solid-state diode laser (Melles Griot BLS 85-
124 601) with ~250 mW output power at the source. A custom-built optical system was constructed
125 around an Olympus BX microscope with confocal apertures before and after the sample. Spectra

126 were obtained using a 0.3 m Andor Shamrock spectrograph with a 60 μm entrance slit and 1200
127 lines per mm diffraction grating. The Andor Newton DU970-CCD camera was
128 thermoelectrically cooled to $-90\text{ }^{\circ}\text{C}$. Spectra were collected using either a 100x Mitutoyo M Plan
129 Apo objective with 6 mm working distance and 0.7 numerical aperture (Figure 1 and laser
130 heating experiments) or an extra-long working distance 20x Mitutoyo M Plan Apo objective with
131 20.2 mm working distance and 0.42 numerical aperture (all other measurements). The laser
132 power at the sample was always measured using a hand-held power meter with $\pm 0.01\text{ mW}$
133 precision near the $\sim 2\text{ }\mu\text{m}$ diameter focal spot of the objective. For the spectra collected to study
134 the effects of laser heating, laser powers ranged from 0.05 to 20 mW, as measured at the sample
135 position. In the laser heating experiments, the focused laser was held steady on the sample for 5
136 minutes prior to data collections of 120 seconds, averaged over three accumulations.

137 High-temperature experiments at room pressure were performed using a resistive heater,
138 where the sample was placed on the flat culet of a diamond anvil mounted in a tungsten carbide
139 seat wound with a Mo-wire resistive heater. The sample was placed directly on the culet of the
140 diamond anvil with Ar gas flowing over the sample. A K-type thermocouple attached to the
141 diamond was used to monitor the temperature, which was increased at $25\text{ }^{\circ}\text{C}$ intervals and held
142 for 10 minutes at each temperature prior to data collection for three accumulations of 60 seconds.

143 Simultaneous high-pressure, high-temperature experiments were performed using a
144 resistively-heated sapphire-anvil cell. Due to overlap of the first-order longitudinal phonon of
145 diamond at 1332 cm^{-1} , sapphire anvils were required to obtain kerogen Raman spectra at high
146 pressures. We used sapphire anvils of 3.1 mm diameter and 1.7 mm height with 300 or 400 μm
147 culet sizes. The 400- μm culet anvils routinely cracked at high temperature above about 2 GPa, so

148 we used 300- μm culets to reach 3.2 GPa maximum pressure at high temperature without
149 breakage. The isolated kerogen samples were loaded with a KBr pressure medium into 0.20 mm
150 thick Inconel gaskets, pre-indented to 0.15 mm thickness with sample chambers drilled at about
151 half the diameter of the culet. Pressures were determined using the Sm:YAG scale (Trots et al.
152 2013). The nominal precision in all pressure measurements is ± 0.1 GPa. K-type thermocouples
153 attached to both upper and lower sapphire anvils were used to monitor the temperature, which
154 was increased at 25-50 $^{\circ}\text{C}$ intervals and held at each temperature for 15 minutes prior to data
155 collection lasting 60 seconds accumulated three times. For long-duration heating experiments,
156 we increased the pressure to ~ 3 GPa and heated the same sample to 450 $^{\circ}\text{C}$ for periods lasting 2-
157 8 hours, returning to room temperature in between each heating cycle. Laser power measured at
158 the objective using a hand held power meter was 0.09 mW. Attenuation through the sapphire
159 anvil resulted in a laser power of 0.05 mW at the sample, as measured using an open cell.

160 Deconvolution of the spectra was performed using PeakFit (Systat Software Inc.). Raman
161 bands of kerogen are typically a combination of a Gaussian and Lorentzian fit, therefore a Voigt
162 Amplitude function is often used (Beysac et al. 2002b, 2003b; Rahl et al. 2005). Sadezky et al.
163 (2005) demonstrate that first-order bands, with the exception of D3, are Lorentzian. We fitted
164 spectra using both Voigt and Lorentzian fits, finding no significant difference in the results. All
165 first order-bands in this study are reported using Lorentzian Amplitude and Area-based fits with
166 standard errors reported from PeakFit.

167

168 **RESULTS**

169 **Effect of laser heating on kerogen Raman spectra**

170 Raman spectra of the kerogen in this study are characterized by first-order (1100 – 1800
171 cm^{-1} ; Figure 1a) and second-order (2200 – 3500 cm^{-1} ; Figure 1b) vibrational bands typical of
172 mature kerogen (e.g. Wopenka and Pasteris 1993). Within the first-order region, the G band
173 (1580 cm^{-1}) associated with C–C in-plane stretching of aromatic layers (e.g. Ferrari and
174 Robertson 2000) overlaps with the D2 (1620 cm^{-1}) disordered carbon band. Also present in the
175 first order region are the D1 (1350 cm^{-1}) and D3 (1500 cm^{-1}) bands (Beysac et al. 2002b). The D
176 bands, shown in Figure 1a, are attributed to a variety of graphitization disorder effects, including
177 reduced graphite lattice symmetry near edges, edge vibrations, heteroatoms, and defects in or
178 between the aromatic plane (Beny-Bassez and Rouzaud 1985; Wang et al. 1990; Beysac et al.
179 2002a).

180 The second order region shows bands at approximately 2700, 2900, and 3250 cm^{-1}
181 labeled S1, S2, and S3 in Figure 1b. These bands are attributed to overtones and combination of
182 the first-order region (Kelemen and Fang 2001). The S1 band is possibly an overtone of the D1
183 band ($2 \times 1350 \text{ cm}^{-1}$), S2 a result of combination scattering from the D1 and G bands ($1350 +$
184 1580 cm^{-1}), and S3 an overtone of D2 ($2 \times 1620 \text{ cm}^{-1}$) (Marshall et al. 2009). Another
185 explanation for the second order region is the presence of aliphatic or aromatic C–H stretching
186 associated with the S2 band, which has been observed in shale samples (Tsu et al. 1978; Czaja et
187 al. 2009; Marshall et al. 2009). In this study, we focus on the first-order Raman bands of
188 kerogen.

189 Band intensity or area ratios are commonly used parameters for the assessment of thermal
190 maturity or metamorphic grade (e.g. Wopenka and Pasteris 1993; Beysac et al. 2002b; Huang et
191 al. 2010). R1 is calculated as $[(D1/G)_{\text{intensity}}]$ and R2 is calculated by $[D1/(G + D1 + D2)_{\text{area}}]$
192 (Beysac et al. 2002b; Huang et al. 2010). Wopenka and Pasteris (1993) note that R1 and R2 are

193 less reliable for the interpretation of kerogen metamorphic grade than FWHM, due to the fact
194 that more mature kerogen will have much narrower first order peaks coupled with stronger
195 second order features. Decreases in R1 and R2 provide evidence for increased thermal maturity
196 and metamorphic grade (Beysac et al. 2002b). Another less commonly used parameter, R3, is
197 calculated by $[S2/(S1 + S2)_{area}]$, however with generally weaker bands in the second order region
198 for kerogen this parameter may not be sensitive to small changes (Beysac et al. 2002b).
199 Additional quantitative measurements to indicate increased thermal maturity/metamorphic grade
200 include a decrease in the intensity of the D1 band, decrease in the area of the D1 band, decrease
201 in the area of the D3 band, narrowing of the G band, decrease in the intensity of the D2 band,
202 decrease in the intensity of the S2 band, shift in the G band to a lower center position (Beysac et
203 al. 2002a).

204 Increasing the excitation laser power results in an overall decrease of normalized
205 intensity (Figure 2a), but can also clearly influence band intensity ratios, as illustrated in Figure
206 2b. Therefore, it is imperative to quantify the damage threshold of excitation lasers in Raman
207 studies of kerogen (e.g. Morishita et al. 2011). In order to characterize the heating potential of
208 the 458 nm excitation laser used in this study, we collected spectra under otherwise identical
209 collection conditions at various laser powers ranging from 0.05 to ~6 mW, as measured by a
210 Melles-Griot hand-held laser meter near the focal point of the objective lens. Spectra in Figure 2a
211 and 2b are shown only to 6 mW maximum power because at higher powers (up to 20 mW) the
212 sample instantaneously degraded. Figure 3 illustrates the effect of temperature (direct heating
213 under a stream of Ar gas) on the first-order Raman spectra of kerogen, measured using a laser
214 power of 0.05 mW.

215

216 **Effect of temperature on kerogen Raman spectra**

217 In order to compare the decrease of band intensities on heating by the laser and by direct
218 heating, Figure 4 illustrates the normalized intensities of G, D1, and D2 bands as a function of
219 laser power (normalized to intensity at 0.05 mW) and temperature (normalized to 22 °C, laser
220 power of 0.05 mW). With increased laser power (Figure 4a) and increased direct thermal heating
221 (Figure 4b) normalized intensities decrease, reflecting alteration and eventual degradation above
222 ~6 mW and 525°C. While greater initial intensities may be achieved by increased laser power, as
223 shown in Figure 4a between the 0.05 and 0.76mW step, prolonged exposure to increased laser
224 power will damage the sample.

225 Figure 5a illustrates the influence of laser power on R1 (intensity) and R2 (area) ratios.
226 Taking 0.05 mW as the reference, the $R1 = (D1/G)_{intensity}$ ratio decreases from approximately 0.9
227 to 0.65 with an increase in laser power to ~1.5 mW, but does not decrease significantly further
228 up to 6 mW, suggesting the most drastic changes occur in R1 below 2 mW power. The R2 area
229 ratio is somewhat invariant up to 2 mW but decreases by almost 50% at 3-6 mW laser power
230 (Figure 5a). For comparison to temperature, we plot in Figure 5b the R1 ratio as a function of
231 temperature up to 500 °C, where the drastic drop in R1 from ~0.9 to ~0.6 is observed above
232 approximately 300 °C, implying that 1-2 mW of laser power can cause an equivalent of 300 °C
233 heating. The drastic decrease in R1 upon raising the laser power from 0.05 mW to greater than 1
234 mW (Fig 5a) suggests that a practical laser-damage threshold for 458 nm is significantly below
235 1mW.

236 Figure 6 illustrates the effect of laser power on D1, D2, and G-band widths, expressed as
237 full-width at half maximum (FWHM, cm^{-1}). The G and D1 bands (Fig. 6a, b) both decrease in
238 FWHM with increasing laser power, while the disordered carbon band D2 (Fig. 6c) increases. As

239 a function of temperature (Figure 6 d-f) all bands show an increase in FWHM with increasing
240 temperature. The increase in FWHM is largest in the D2 band (Figure 6f), however the majority
241 of band width increase does not occur until temperatures greater than 400°C.

242 The most robust comparison of laser-heating to direct heating below the laser-damage
243 threshold can be seen in G, D1, and D2 peak positions. Figure 7a overlays the shift in band
244 positions with both laser power and temperature. Whereas the D1 peak position is essentially
245 invariant with laser power and temperature, the G-band position decreases significantly (by
246 almost 20 cm⁻¹) from 0.05 to 6 mW, an equivalent reduction in band position on heating to 500
247 °C. The D2 band position decreases slightly with laser power and temperature to approximately
248 300°C, but then shifts back towards higher wavenumber with both laser power and temperature.
249 Considering the general decrease in G-band position with laser power and temperature in Figure
250 7a, we conclude that for 458 nm laser excitation, there is roughly 100°C of equivalent heating per
251 mW of increased laser power between 0.05 and 6 mW. The slope of band position with
252 temperature and laser power is further confirmed by comparison to the high-temperature data of
253 Huang et al. (2010) (Figure 7b).

254 Although many of the observed spectral changes by laser-induced heating are comparable
255 to bulk-heating the sample at laser powers below the damage threshold, namely, reduction of
256 band intensities, band intensity ratios, and peak positions, some aspects of the spectral evolution
257 of kerogen are different. In particular, on heating to above ~350 °C we observe an abrupt drop in
258 the R1 ratio, which is not observed with laser powers up to 6 mW (Figure 5). Whereas the
259 FWHM of the D2 band increases with both laser power and heating temperature, the FWHM of
260 G and D1 increases before sharpening up above 2 mW while the FWHM of G and D1 increases
261 gradually with bulk sample heating (Figure 6). We do not have an obvious explanation for these

262 observed differences, requiring further study, but maintain a broad equivalency exists between
263 laser-induced heating and bulk heating based upon the band shifts shown in Figure 7.

264 To summarize the results of our tests on the influence of laser power on Raman spectra of
265 kerogen, blue-green excitation lasers should not to exceed 1 mW, and preferably spectra should
266 be collected at laser powers less than 0.1 mW. With a focal spot diameter of $\sim 2 \mu\text{m}$, 0.1 mW
267 corresponds to a total irradiance of about 3 kW/cm^2 . For the purpose of studying the evolution of
268 kerogen Raman spectra with simultaneous pressure and temperature, we used a laser power of
269 0.05 mW ($\sim 1.6 \text{ kW/cm}^2$), as measured focusing through the sapphire of an open cell, ensuring no
270 effects of laser heating.

271

272 **High-pressure high-temperature Raman spectra of kerogen**

273 The use of a resistively-heated sapphire cell permitted collection of first-order Raman
274 spectra of kerogen as a function of pressure and temperature up to 3.2 GPa and 380 °C. In the
275 current configuration, achieving pressures $>3.2 \text{ GPa}$ while heating to above 400 °C was not
276 possible, and several anvil breakages occurred. Future experiments with different gasket
277 materials and/or culets smaller than 300 μm should allow for higher P-T conditions to be
278 reached.

279 We measured first-order Raman spectra of kerogen at around 80 different pressure-
280 temperature points between 0-3.2 GPa and 25-380 °C. Representative spectra are shown in
281 Figure 8a as a function of pressure at room temperature, and in Figure 8b at simultaneous high
282 pressures and temperatures. We observe no significant change or breakdown of kerogen over the
283 experimental P-T range.

284 Kerogen spectra taken in the sapphire cell as a function of pressure up to 3.2 GPa at room
285 temperature are shown in Figure 8a, illustrating very little change of no more than 10 cm⁻¹ shift
286 in the position of the G+D2 band, but a small shift of ~20 cm⁻¹ to higher wavenumbers of the D1
287 band position. Figure 8b shows spectra recorded at simultaneous P-T conditions up to 3.2 GPa
288 and 380 °C, also showing no major changes in the bonding configuration of kerogen. To
289 systematically assess the influence of pressure and temperature on band positions, Figure 9
290 shows the evolution D1, D2, and G band positions from all P-T conditions visited during this
291 study. We separated each plot to show data at high temperatures but various pressures (Figure
292 9a,c,e), and separately as a function of pressure at various temperatures (Figure 9b,d,f). For the
293 purpose of analyzing the effect of pressure on the temperature evolution of the bands, we fitted
294 the high-temperature variability of band positions at all pressures above 0.8 GPa (Figure 9).

295

296 **Pressure cycling**

297 To test the reversibility in the shift of G, D1, and D2 bands with pressure over the current
298 experimental pressure range, we performed a pressure cycling experiment where the initial
299 spectrum was taken at room pressure in the sapphire cell, loaded to 3.1 GPa, and allowed to rest
300 for 72 hours, during which time the pressure increased marginally to 3.3 GPa. The cell was then
301 brought back to room pressure for re-measurement. Figure 10a demonstrates complete
302 reversibility in the kerogen spectra up to 3.3 GPa at room temperature. This is in contrast to
303 pressurization to 9 GPa by Huang et al. (2010), who observed reversibility in the G, D1, and D2
304 band positions, but the sample recovered from 9 GPa had considerably less intensity and broader
305 peak widths.

306

307 **Heating-cooling cycles**

308 One of the surprising results of Huang et al. (2010) was the observed reversibility in shift
309 with temperature of D1, D2, and G band positions; being fully recoverable up to 600 °C. We
310 measured the D1, D2, and G band positions along a cycling heating path up to the instantaneous
311 breakdown observed at temperatures around 500 °C. Figure 11 shows the D1, D2, and G band
312 position on heating to 200, 300, 400, and 450 °C, as well as band positions at room temperature
313 between each subsequent thermal maximum. In the D1 band (Figure 11a), spectra taken while at
314 high temperature show a shift toward lower peak position wavenumbers with increasing
315 temperature, as also shown by (Huang et al. 2010). Room temperature D1 band spectra taken
316 between each high temperature measurement have peak positions shifted higher than the original
317 room temperature measurement, and shift higher following each successive heating. The D1
318 band represents the A_{1g} breathing vibration mode of aromatic rings (Ferrari and Robertson 2000;
319 Beyssac et al. 2003b) and the shift to lower frequencies of D1 at high temperature is the result of
320 softening of the C-C bond strength (Raravikar et al. 2002). The shift toward higher D1 band
321 position at room temperature following each high temperature measurement is the result of
322 catagenic thermal degradation of non-aromatic structures (Vandenbroucke and Largeau 2007)
323 and the prevalence of smaller aromatic clusters, which are known to shift the D1 band higher
324 (Ferrari and Robertson 2000). The D2 and G bands both represent the E_{2g} -derived stretching
325 mode (e.g Marshall et al. 2009). Similar trends with sequential heating and cooling are seen in
326 the D2 and G band position through 400 °C, after which the bands shift towards each other as the
327 sample moves toward graphitization, potentially signifying destruction of bonds at sp^2 sites
328 (Ferrari and Robertson 2000).

329 Temperature cycling also shows irrecoverable changes in the R1 and R2 ratios. Figure 12
330 shows R1 and R2 as a function of temperature before and after heating to 200, 300, 400, and 450
331 °C, while holding temperature for 10 minutes. The irrecoverable changes in R1 and R2 show the
332 loss of signal intensity with thermal degradation, which more degradation attributable to the D1
333 band in R1 and both defect bands in R2. At temperatures above 450 °C under Ar-gas flow, open-
334 cell heating of kerogen results in major degradation of the sample, similar to that seen by Tissot
335 et al. (1974). This is indicated in Raman spectra by a loss of signal intensity and the sample is
336 visually altered, becoming lighter in color before degradation. The reversibility in D1, D2, and G
337 band frequency following heating to 600 °C in Huang et al. (2010) is within 10 cm^{-1} , which is
338 comparable to the total change observed in our temperature-cycling experiments (Figure 11). In
339 both studies, band intensities do not recover from heating experiments. In addition, there are
340 likely differences owing to the fact that the sample from Huang et al. (2010) may have a different
341 depositional history and subsequent geologic thermal maturation pathway.

342 To examine the reversibility of peak intensity and position after longer heating durations
343 at high pressure, we performed several continuous heating experiments on the same sample, each
344 lasting between 2-8 hours. The same sample was heated to 450 °C at 2.7-3.0 GPa and held for 4,
345 8, 5, 5, and 2 hours, with recovery to room temperature in between each heating experiment. As
346 an example, a spectrum collected at 450 °C and ~2.7 GPa is shown in Figure 10b before and after
347 heating for five hours. Whereas we observed instantaneous and irreversible changes in the
348 Raman spectra collected above ~350 °C at room pressure (Figures 3, 5), the spectra recovered
349 from 450 °C and ~3 GPa from long-duration heating experiments is completely reversible
350 (Figure 10b), indicating that pressure effectively stabilizes the isolated kerogen to higher
351 temperatures.

352 **DISCUSSION**

353 The use of laser power in the blue-green region greater than 0.1 mW (3 kW/cm²) has the
354 potential to alter the integrity of kerogen samples by simulating natural thermal
355 maturity/metamorphic grade. Evidence for this comes from irreversible loss of intensity in all
356 peaks with increasing laser power, particularly the loss of intensity in the D1, D2, and G bands,
357 considered indicative of natural thermal maturity/metamorphic grade (Figure 4a). Direct thermal
358 alteration by heating reveals decreases in the intensity of these bands as well, and suggests that
359 laser heating is able to simulate natural thermal maturity.

360 The laser degradation of CM analogues to natural thermal maturity/metamorphic grade is
361 also indicated by several other diagnostic band parameters. Decreases in area of the D1 and D3
362 band are considered indicative of natural thermal maturity, however similar results are seen with
363 laser induced heating of the sample (Figure 13). Decreased intensity of the D1 and D2 peaks is
364 also considered indicative of natural thermal maturity, and laser induced heating resulted in a
365 similar decrease (Figure 4). Shifts to lower wavenumber of the G band are also used to indicated
366 natural thermal maturity, however laser induced heating of the sample was able to produce this
367 effect as well (Figure 7).

368 Direct heating experiments in this study were compared with results from other kerogen
369 heating experiments. The peak position of the D1, D2, and G bands shows good correlation
370 between data of this experiment and that reported by Huang et al. (2010) (Figure 7). Similar
371 trends of each band, with slightly offset values for each band are likely the result of differences
372 between provenance of sample as well as the laser wavelength. Comparison of laser induced
373 band shifts with direct heating and heating results of Huang et al. (2010) suggests that laser
374 induced heating can cause shifts in band position similar to large natural thermal maturity ranges

375 (Figure. 7). Laser power as low as 1 mW can allow the interpretation that the sample has
376 experienced thermal alteration to 100 °C. At laser powers of 6 mW, sample peak position is
377 comparable to temperatures greater than 500 °C.

378 Previous studies interpreting kerogen Raman bands for assessment of thermal maturity
379 have used laser power at the sample of greater than 1 mW, and as high as 15 mW (e.g. Pasteris
380 and Wopenka 1991; Wopenka and Pasteris 1993; Beyssac et al. 2002a, 2002b; Huang et al.
381 2010), although the wavelength of the excitation laser needs to be taken into consideration of
382 damage thresholds. The absorption coefficient of typical type-II kerogen, for example, is about
383 20% less at 532 nm compared with 458 nm (Ferralis et al. 2015). We found at 458 nm, the use of
384 laser power greater than even 0.1 mW has the potential to induce thermal alteration to kerogen
385 samples, thus complicating the interpretation of bands for thermal maturity.

386 Huang et al. (2010) measured first-order Raman spectra of kerogen between 25-600 °C at
387 room pressure, and up to 9 GPa at room temperature, finding that in general, the stretching
388 frequency (ν , in cm^{-1}) of major bands (D1, D2, and G) decreases with increasing temperature and
389 increases with increasing pressure. These structural changes signify softening of the mode
390 frequencies with temperature and stiffening of the mode frequencies with pressure, as expected.
391 Huang et al. (2010) reported that the D1, D2, and G band positions decrease in frequency shift
392 with temperature ($d\nu/dT$) by -0.020(5), -0.029(6), and -0.025(6) $\text{cm}^{-1}/^\circ\text{C}$, respectively. On
393 compression at room temperature, D1, D2, and G band positions increase in frequency shift with
394 pressure ($d\nu/dP$) by 5.15, 3.34, and 3.66 $\text{cm}^{-1}/\text{GPa}$, respectively. Here we examine the net
395 influence of pressure and temperature on mode-frequency shifts.

396 Figure 9 shows the D1, D2, and G-band frequencies (in Raman shift, cm^{-1}) as a function
397 of simultaneous pressure and temperature. To display these trends, we plotted data as a function
398 of temperature at various pressures between 0-3.5 GPa in the upper panels of Figure 9, and data
399 as a function of pressure at various temperatures between 20-400 °C in the lower panels. We
400 performed error-weighted, linear fits to the band positions as a function of temperature at all
401 pressures above 0.8 GPa, finding $-0.017(4)$, $-0.023(3)$, and $-0.027(4)$ $\text{cm}^{-1}/^\circ\text{C}$ for the D1, D2, and
402 G bands, respectively. These shifts are comparable within error to those found at room pressure
403 by Huang et al. (2010), with $dv/dT = -0.020(5)$, $-0.029(6)$, and $-0.025(6)$ $\text{cm}^{-1}/^\circ\text{C}$ for D1, D2, and
404 G, respectively. However, the slight reduction in shift of vibrational frequencies of the
405 disordered carbon modes with temperature (dv/dT) at pressures >1 GPa further indicates that
406 pressure acts to increase the thermal stability of kerogen. We find no dependence of R2 on
407 temperature at room pressure or high pressure up to 450 °C (Figures 5 and 14). However, at
408 room pressure, the R1 value drops sharply (from about 0.8 to 0.3) above ~ 350 °C (Figure 5b),
409 indicating a significant change in the kerogen structure. However, at high pressures (>0.8 GPa),
410 the R1 value is invariant up to nearly 400 °C (Figure 14a), indicating that pressure acts to
411 stabilize disordered kerogen at high temperatures. In contrast to heating at room pressure, where
412 the R1 and R2 ratios decrease with increasing temperature (Figure 9), at high pressures (>1 GPa)
413 there is no discernable variation in R1 and R2 as a function of temperature up to 380 °C.

414

415 **IMPLICATIONS**

416 Kerogen, broadly defined as insoluble macromolecular organic matter (OM), is
417 considered the most abundant form of OM on Earth (Vandenbroucke and Largeau 2007). In
418 addition to being the source material of oil and natural gas, fossil kerogen passing through the oil

419 window in subduction zones along active margins becomes part of the global carbon cycle.
420 Understanding the evolution of kerogen structure and stability at high pressures and temperatures
421 will therefore illuminate the interface between surface and mantle carbon (diamond). Our
422 experiments suggest that pressures of just a few GPa raises the temperature at which bonds begin
423 to cleave in kerogen to over 450 °C.

424 Figure 15 illustrates the pressure-temperature points visited in this study in comparison to
425 several slab geotherm models. The survival of kerogen to pressures above 3 GPa and 450 °C
426 indicates that a significant fraction of kerogen in marine sediments may survive deeper into the
427 subduction zone and uppermost mantle. If modern marine sediment subduction along all
428 convergent boundaries is equal to 0.2 Tmol OC yr⁻¹ (Hayes and Waldbauer 2006), a total mass of
429 60 Tg OC yr⁻¹ could be recycled into the mantle. The total terrestrial organic carbon burial flux
430 to the oceans of 58±17 Tg yr⁻¹ (Burdige 2005) is disproportionately contributed by convergent
431 margins due to high burial efficiencies (Blair and Aller 2012). In these convergent margin
432 systems 10 – 75% of the buried terrestrial organic carbon is kerogen, resulting in ~2.5 – 45 Tg
433 yr⁻¹ of kerogen burial. Therefore ~4 to 75% of organic carbon subducted could be kerogen. The
434 presence of carbonate melts deeper in the mantle eventually leads to diamond formation, and
435 some diamonds indicate a significant fraction of organic source carbon, with light δ¹³C values of
436 less than -10‰, and ranging as low as -24‰. (Walter et al. 2011) Kerogen isotope values
437 typically range from approximately -20 to -30‰ (Galimov 1980; Whiticar 1996).

438 Interpretation of kerogen Raman spectra requires care because of the very high sensitivity
439 of kerogen to laser heating by blue-green excitation wavelengths. We observed irreversible
440 changes in the band positions and ratios at laser powers for 458 nm above about 1 mW, and
441 recommend laser power, as measured at the sample, to be below 0.1 mW for tightly-focused,

442 confocal measurements of kerogen in the blue-green region. The Raman Laser Spectrometer
443 (RLS) for the ExoMars 2018 rover mission produces a spot size of 50 μm diameter, and the
444 output intensity has been fixed with an upper limit to produce a total irradiance of 0.8-1.2
445 kW/cm^2 (Rull et al. 2013). The upper limit of the ExoMars-RLS was set specifically to remain
446 below the threshold of thermal damage to powder grains of oxides, hydroxides, and potential
447 organics. Although the ExoMars RLS uses a slightly longer wavelength (532 nm versus 458 nm
448 in this study), a rough comparison can be made. We conservatively put the damage threshold for
449 458 nm at 0.05-1.0 mW, and conducted most of our experiments at 0.05 mW. For our
450 approximate beam focus diameter of 2 microns (nearly diffraction limited), using a power of
451 0.05 mW corresponds to a total irradiance of about 1.6 kW/cm^2 , which is just above the current
452 irradiance limit set on the ExoMars RLS at 1.2 kW/cm^2 (Rull et al. 2013). Because the
453 absorbance of kerogen decreases roughly 20% between 458 and 532 nm (Ferralis et al. 2015),
454 the damage threshold we found for 458 nm (above $\sim 3 \text{ kW}/\text{cm}^2$) could be conservatively reduced
455 to about 2.4 kW/cm^2 for a 532 nm laser. We therefore confirm the ExoMars-RLS laser output
456 has been set to an appropriate limit for the investigation of potential carbonaceous material on
457 Mars.

458 **ACKNOWLEDGMENTS**

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463 fabrication of custom sapphire-anvils and W.A. Bassett for help with construction of the high-
464 temperature sapphire cell.

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645 **Figure Captions**

646 Figure 1. First-order (a) and second-order (b) Raman spectra of isolated kerogen from the Eel
647 River, CA, USA. Sum of the deconvolution is shown by the solid black line, with misfit shown
648 below each spectrum. These spectra were collected using 458 nm excitation with 0.05 mW laser
649 power at the sample, collected through a 100x objective (0.7 numerical aperture) for 300
650 seconds, averaged over three accumulations.

651

652 Figure 2. (a) First-order and second-order Raman spectra of kerogen used in this study at various
653 laser power settings measured at the sample position. (b) First-order region detail measured using
654 various laser power settings and plotted with intensity normalized to the G+D2 band.

655

656 Figure 3. First-order Raman spectra of kerogen on heating under Ar-gas stream from 22 to 500
657 °C. Dashed grey lines at 1360 cm^{-1} and 1600 cm^{-1} are drawn for reference. Spectra at different
658 temperatures have been shifted for clarity.

659

660 Figure 4. Fitted intensity of G, D1, and D2 bands as a function of laser power (a) and
661 temperature (b). In (a), the intensity was normalized to peak intensities measured using 0.05 mW
662 laser power. In (b) all spectra were collected using 0.05 mW laser power and the peak intensities
663 were normalized to the values at 22°C.

664

665 Figure 5. (a) Variation of the R1 and R2 ratios (see text) as a function of laser power. (b)

666 Variation of the R1 and R2 ratios as a function of temperature, measured using 0.05 mW laser

667 power. The abrupt decrease in $R1 = (D1/G)_{Intensity}$ at ~ 350 °C indicates an increase in the amount
668 of ordered (graphitic) structure.

669

670 Figure 6. Variation of the G, D1, and D2 band widths (FWHM) as a function of laser power (a-c)
671 at room temperature, and as a function of temperature as measured using a laser power of 0.05
672 mW (d-f).

673

674 Figure 7. (a) Variation of the G, D1, and D2 band positions as a function of laser power (solid
675 circles) and as a function of temperature using 0.05 mW laser power (open symbols) from the
676 current study. (b) G, D1, and D2 band positions as a function of laser power (solid symbols)
677 from the current study compared with the effect of temperature from the study of Huang et al.
678 (2010), shown by open symbols.

679

680 Figure 8. First-order Raman spectra of kerogen in a high P-T sapphire cell (a) as a function of
681 pressure up to 3.1 GPa and (b) at various pressures and temperatures up to 3 GPa and 450 °C.
682 Spectra at 450 °C were collected after 24 hours at 3 GPa and 450 °C. Spectra intensity was
683 normalized to a common signal-to-noise ratio. Dashed grey lines for reference are drawn at 1360
684 cm^{-1} and 1620 cm^{-1} .

685

686 Figure 9. Mode frequencies (in Raman shift, cm^{-1}) of the D1 (a,b), D2 (c,d), and G bands (e,f) of
687 kerogen at high pressure and high temperature. In all six panels, data collected at high
688 temperature are shown by pink shaded circles, and data collected at room temperature are shown

689 by blue shaded squares. Fits to the data as a function of temperature at high pressure (>1.0 GPa)
690 are shown by dashed lines with fits given in the text.

691

692 Figure 10. The effect of pressure (a) and high-pressure high-temperature (b) on first-order
693 kerogen Raman spectra are recoverable below 3.3 GPa and 450 °C. (a) Shown in black, spectrum
694 measured before compression and in red the spectrum immediately after direct compression to
695 3.1 GPa. After holding for 72 hours, the pressure increased to 3.3 GPa and was measured again
696 (green). On decompression back to room pressure, the spectrum recovered completely (grey) to
697 its original state. (b) Shown in black, spectrum measured before compression and heating. In red,
698 spectrum taken at 2.7 GPa and 450 °C. Several heating durations of 2-8 hours each were made,
699 with one example shown in green, taken following five continuous hours of heating at 450 °C
700 and 2.7 GPa. Following 24 hours of cumulative heating, the spectrum in grey was obtained at 1.4
701 GPa and room temperature.

702

703 Figure 11. Temperature cycling experiments at room pressure show the positions of D1, D2, and
704 G bands at room temperature (blue shaded squares) prior to and following heating (pink shaded
705 circles) to the temperatures indicated along a heating-cooling thermal sequence.

706

707 Figure 12. Temperature cycling experiments at room pressure show irreversible effects of
708 heating on the R1 and R2 ratios at room temperature (blue shaded squares) prior to and following
709 heating (pink shaded circles) to the temperatures indicated along a heating-cooling thermal
710 sequence.

711

712 Figure 13. Area of disordered D1 and D3 bands as a function of laser power (a) and temperature
713 (b), measured at 0.05 mW laser power at the sample.

714

715 Figure 14. R1 (a,b) and R2 (c,d) ratios of kerogen spectra at high pressure and high temperatures.
716 Data measured at high temperature are shown by pink shaded circles, and data measured at room
717 temperature are shown blue shaded squares. In contrast to heating at room pressure, where the
718 R1 and R2 ratios decrease with increasing temperature (see Figure 9), at high pressures (>1 GPa)
719 there is no discernable variation in R1 and R2 as a function of temperature up to 450 °C.

720

721 Figure 15. Pressure-temperature points visited in the current experiments in comparison to
722 several calculated geotherms in subducted slabs (Syracuse et al. 2010) along oceanic-continental
723 crust. The graphite-diamond phase boundary is shown by the solid black line (Kennedy and
724 Kennedy 1976). The structural integrity of kerogen to the highest P-T condition indicates that
725 instantaneous breakdown of kerogen will not occur in the upper part of slabs, potentially leading
726 to deeper subduction of carbon back into the mantle.

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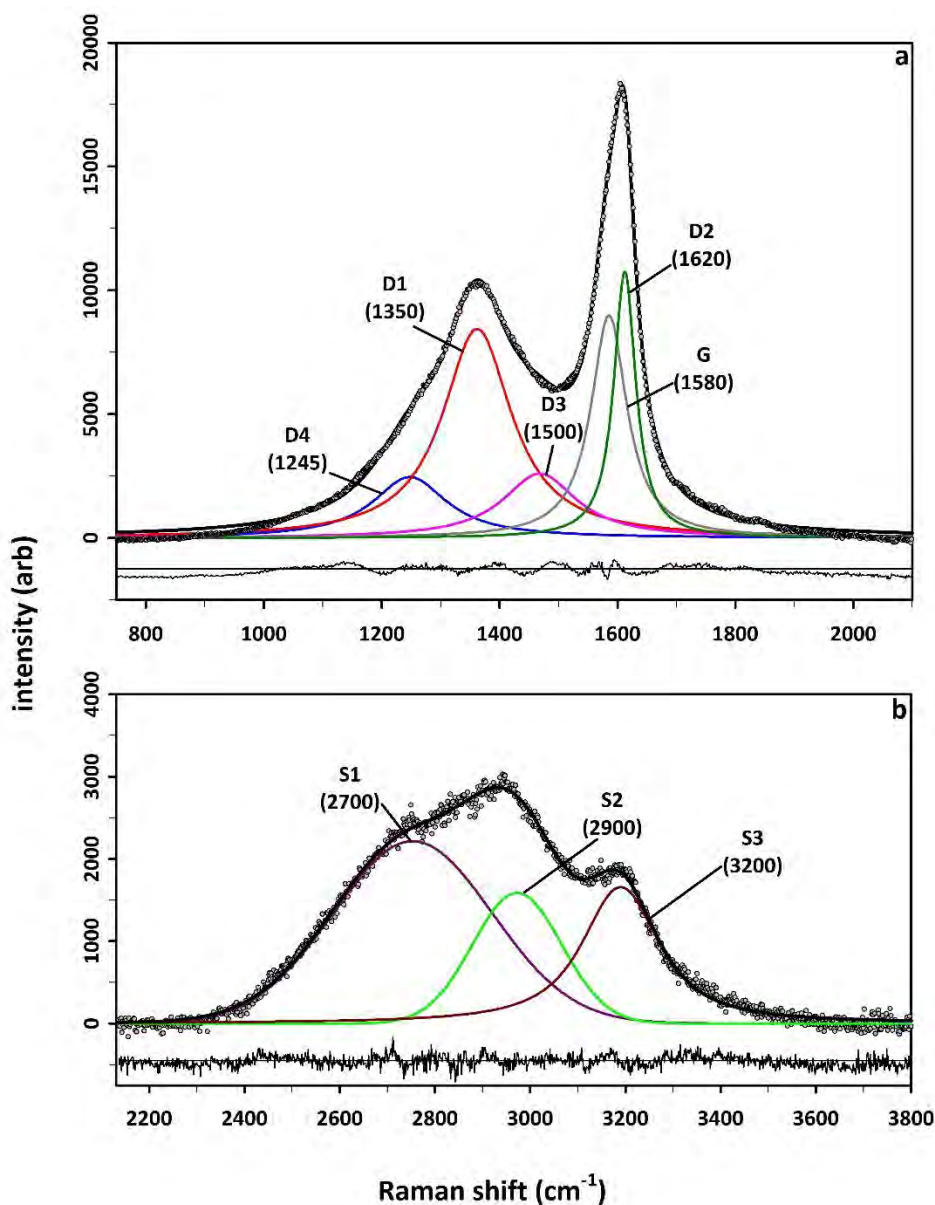
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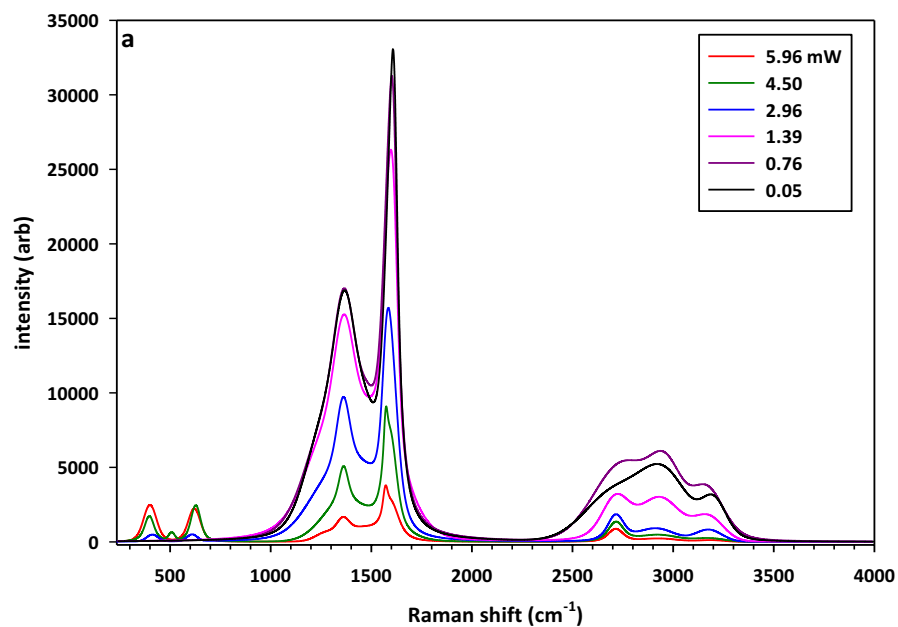
734 **Figure 1**



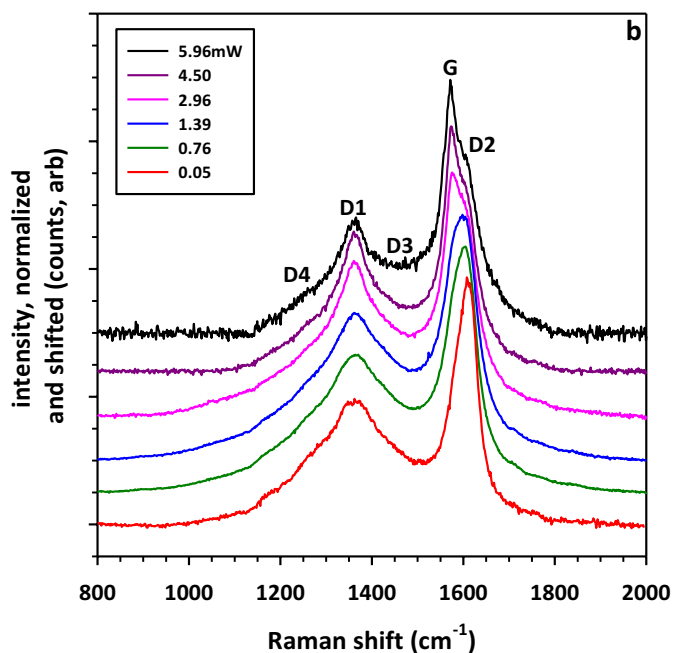
735

736 Figure 1. First-order (a) and second-order (b) Raman spectra of isolated kerogen from the Eel
737 River, CA, USA. Sum of the deconvolution is shown by the solid black line, with misfit shown
738 below each spectrum. These spectra were collected using 458 nm excitation with 0.05 mW laser
739 power at the sample, collected through a 100x objective (0.7 numerical aperture) for 300
740 seconds, averaged over three accumulations.

741 **Figure 2**



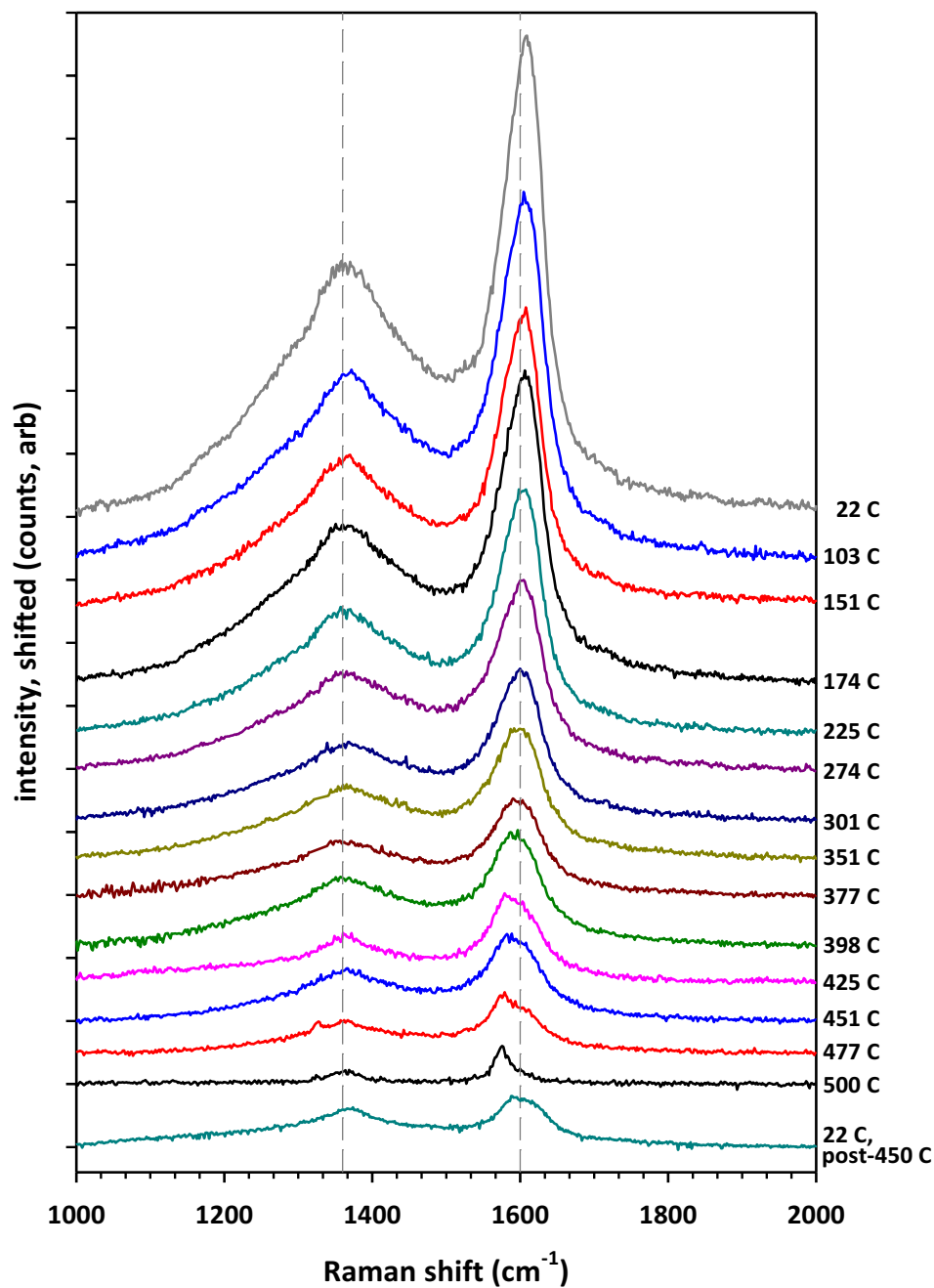
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744 Figure 2. (a) First-order and second-order Raman spectra of kerogen used in this study at various
745 laser power settings measured at the sample position. (b) First-order region detail measured using
746 various laser power settings and plotted with intensity normalized to the G+D2 band.

747 **Figure 3**



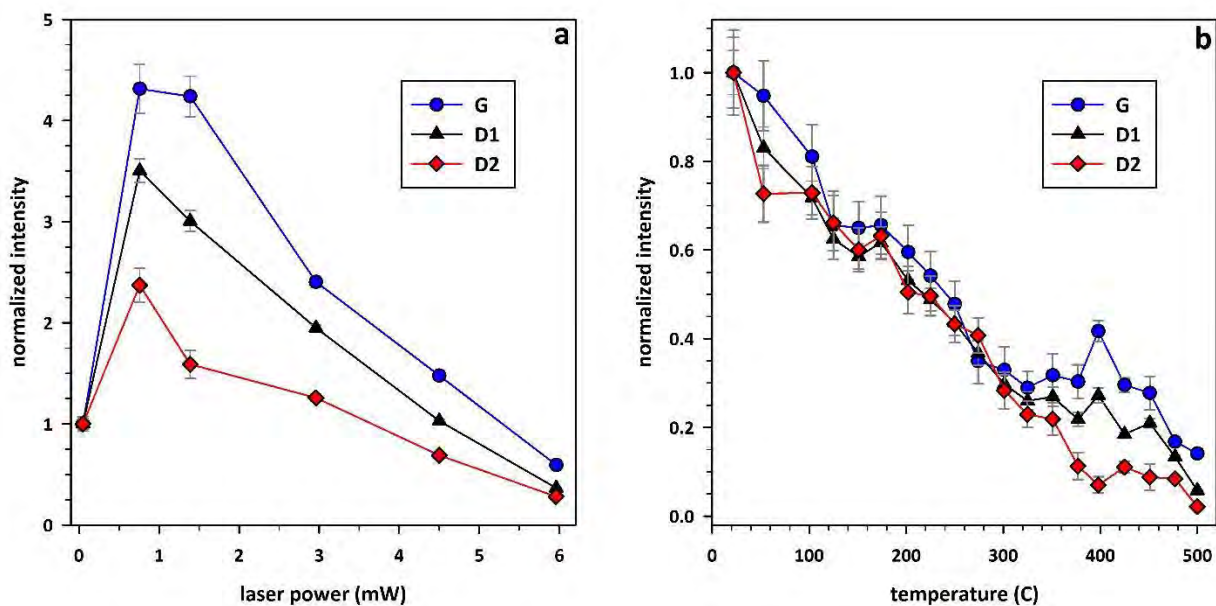
748

749 Figure 3. First-order Raman spectra of kerogen on heating under Ar-gas stream from 22 to 500
750 °C. Dashed grey lines at 1360 cm^{-1} and 1600 cm^{-1} are drawn for reference. Spectra at different
751 temperatures have been shifted for clarity.

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753 **Figure 4**
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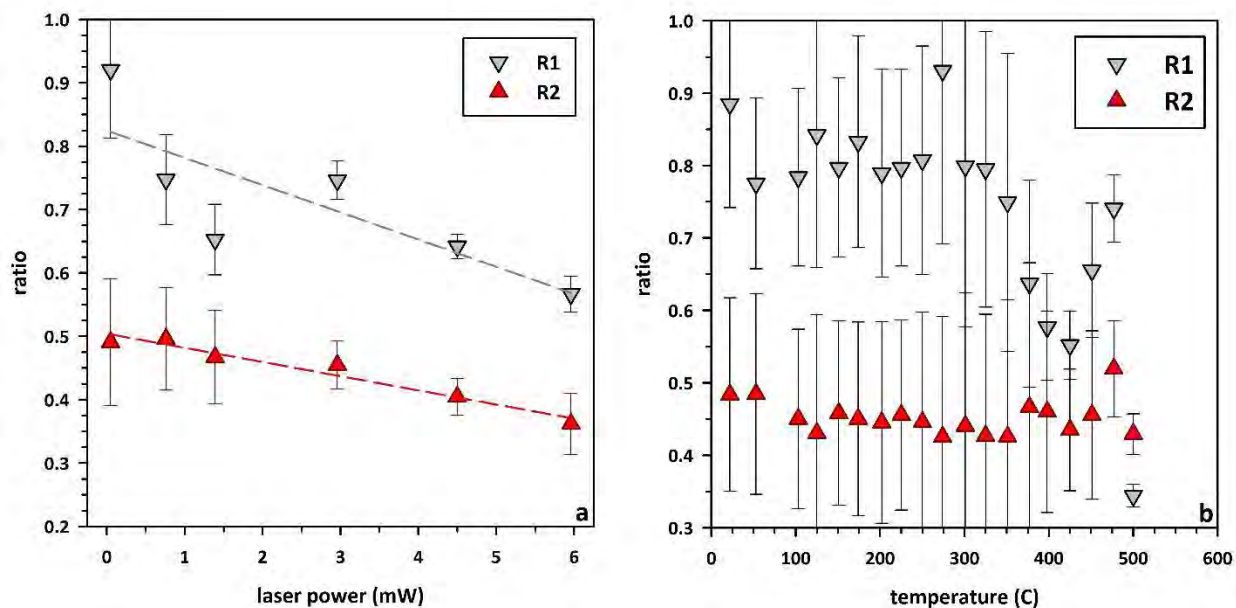
756

757 Figure 4. Fitted intensity of G, D1, and D2 bands as a function of laser power (a) and
758 temperature (b). In (a), the intensity was normalized to peak intensities measured using 0.05 mW
759 laser power. In (b) all spectra were collected using 0.05 mW laser power and the peak intensities
760 were normalized to the values at 22°C.

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762 **Figure 5**

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765 Figure 5. (a) Variation of the R1 and R2 ratios (see text) as a function of laser power. (b)

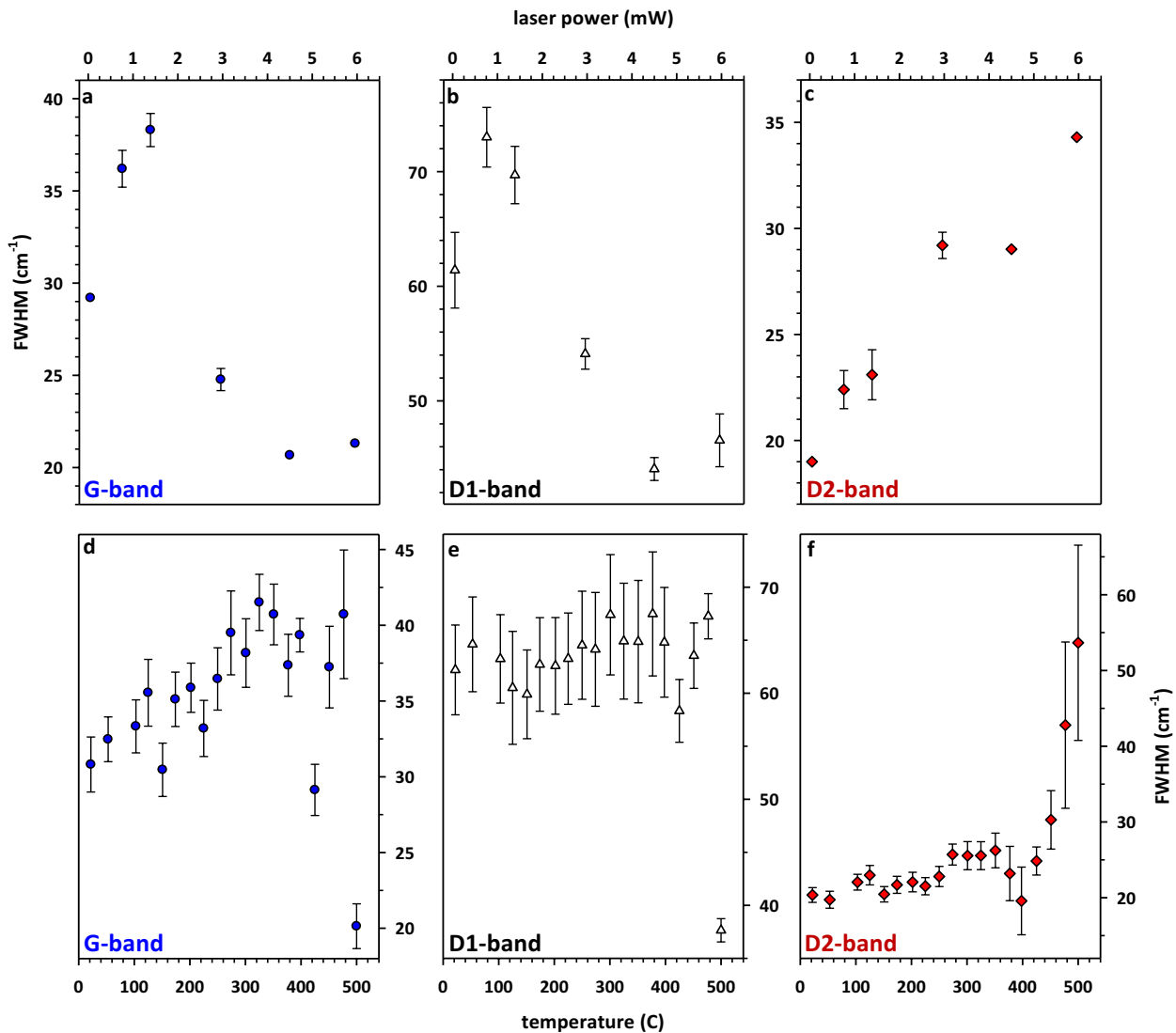
766 Variation of the R1 and R2 ratios as a function of temperature, measured using 0.05 mW laser

767 power. The abrupt decrease in R1 = $(D1/G)_{Intensity}$ at ~350 °C indicates an increase in the amount

768 of ordered (graphitic) structure.

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770 **Figure 6**



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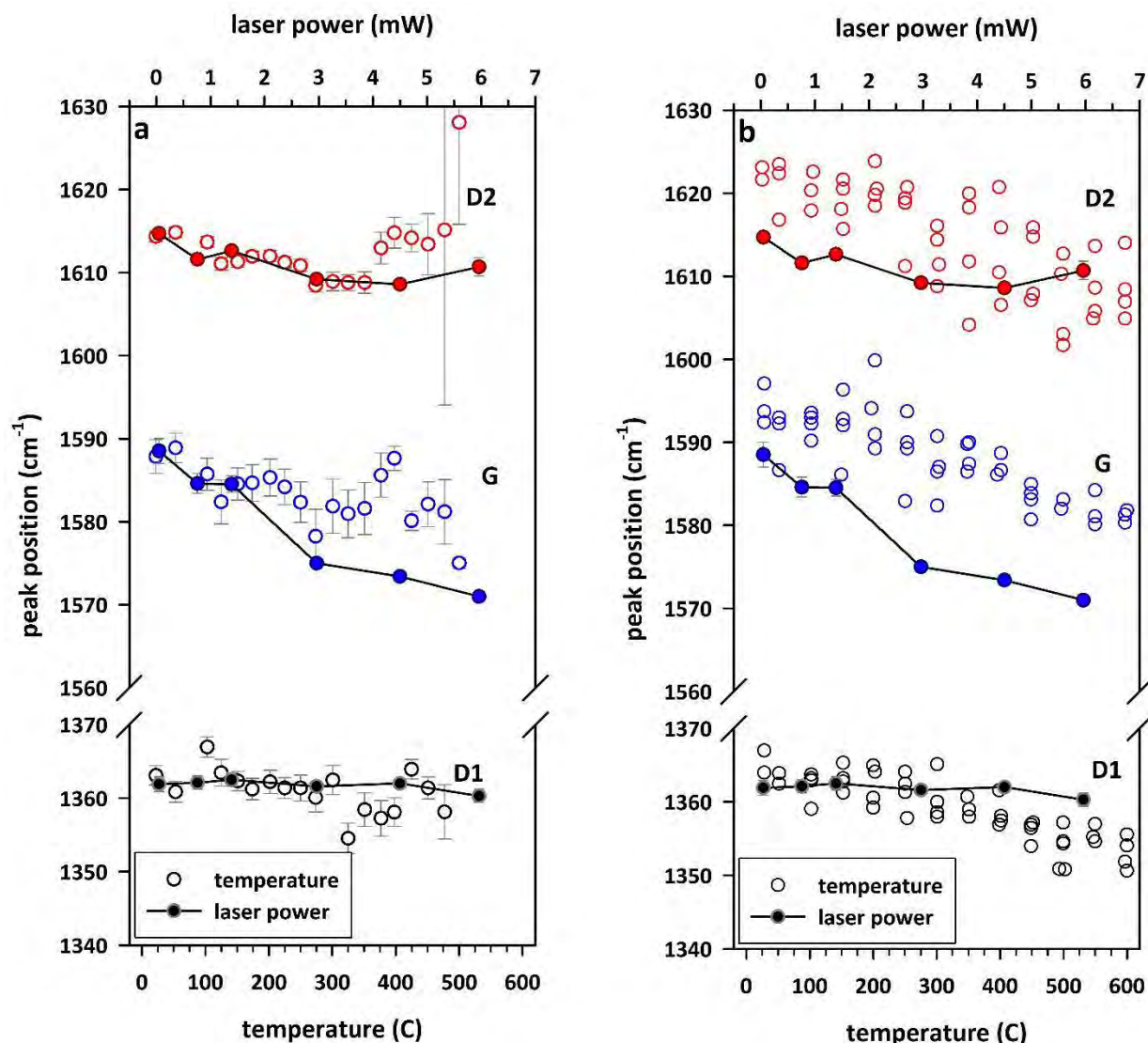
772 Figure 6. Variation of the G, D1, and D2 band widths (FWHM) as a function of laser power (a-c)

773 at room temperature, and as a function of temperature as measured using a laser power of 0.05

774 mW (d-f).

775

776 **Figure 7**

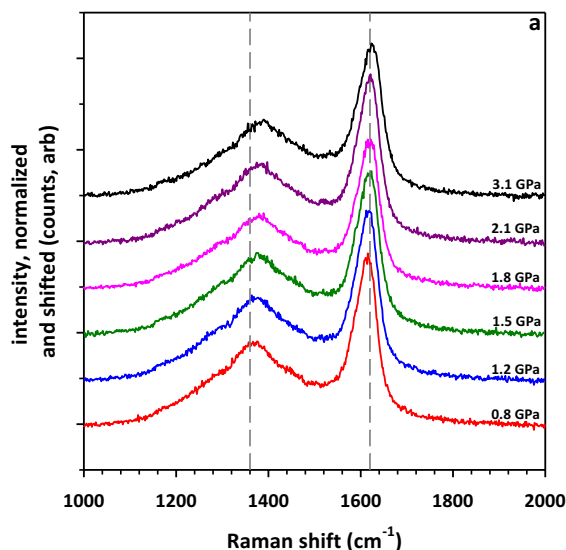


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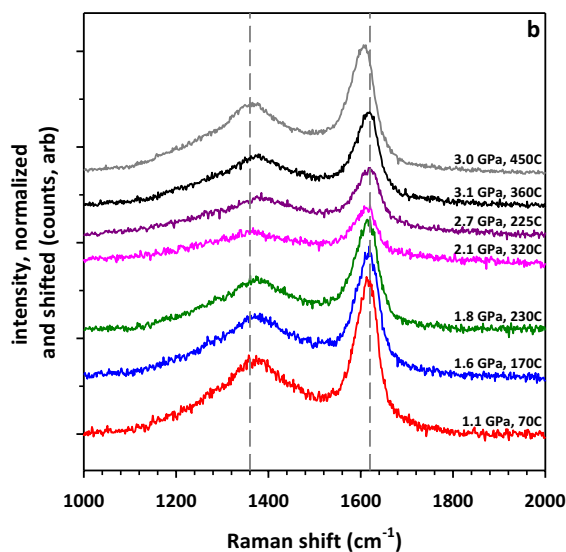
778 Figure 7. (a) Variation of the G, D1, and D2 band positions as a function of laser power (solid
779 circles) and as a function of temperature using 0.05 mW laser power (open symbols) from the
780 current study. (b) G, D1, and D2 band positions as a function of laser power (solid symbols)
781 from the current study compared with the effect of temperature from the study of Huang et al.
782 (2010), shown by open symbols.

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784 **Figure 8**



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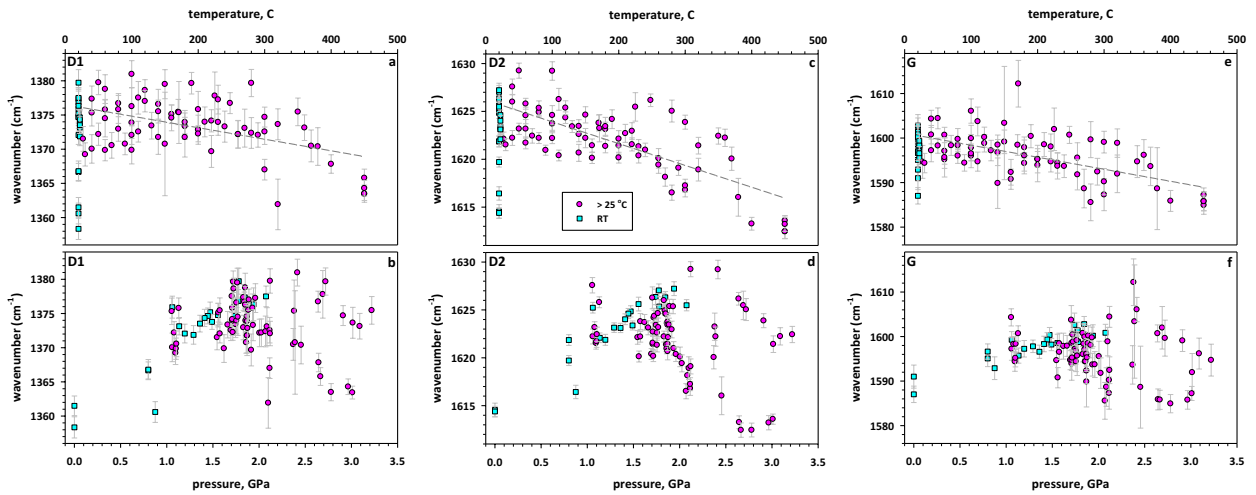


786

787 Figure 8. First-order Raman spectra of kerogen in a high P-T sapphire cell (a) as a function of
788 pressure up to 3.1 GPa and (b) at various pressures and temperatures up to 3 GPa and 450 °C.

789 Spectra at 450 °C were collected after 24 hours at 3 GPa and 450 °C. Spectra intensity was
790 normalized to a common signal-to-noise ratio. Dashed grey lines for reference are drawn at 1360
791 cm⁻¹ and 1620 cm⁻¹.

792 **Figure 9**



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794

795 Figure 9. Mode frequencies (in Raman shift, cm⁻¹) of the D1 (a,b), D2 (c,d), and G bands (e,f) of
796 kerogen at high pressure and high temperature. In all six panels, data collected at high
797 temperature are shown by pink shaded circles, and data collected at room temperature are shown
798 by blue shaded squares. Fits to the data as a function of temperature at high pressure (>1.0 GPa)
799 are shown by dashed lines with fits given in the text.

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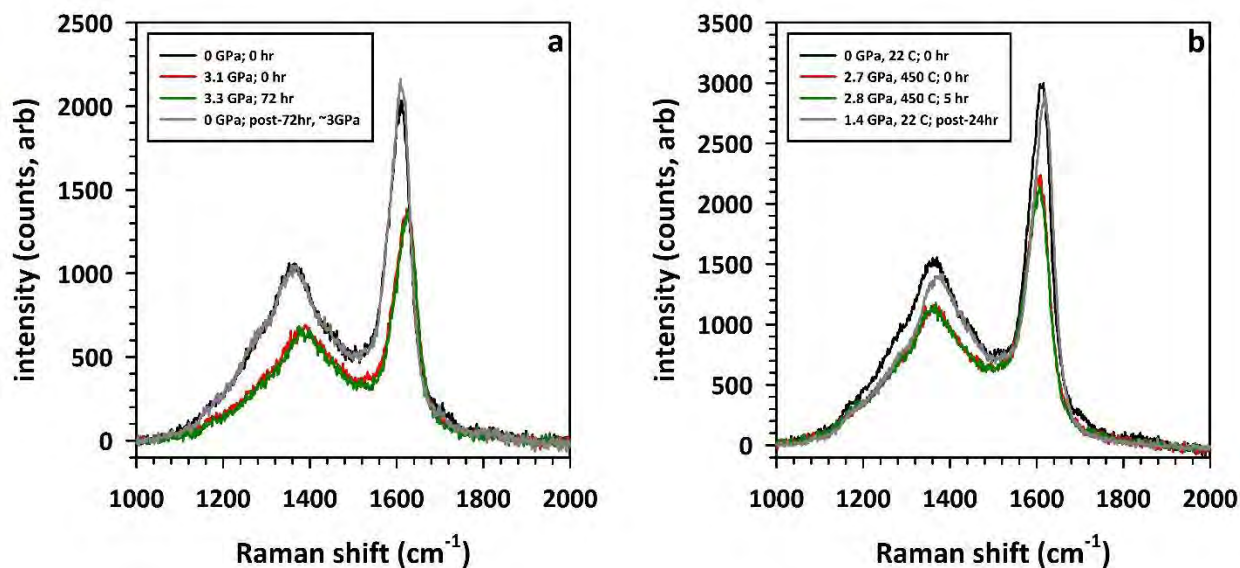
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809 **Figure 10**

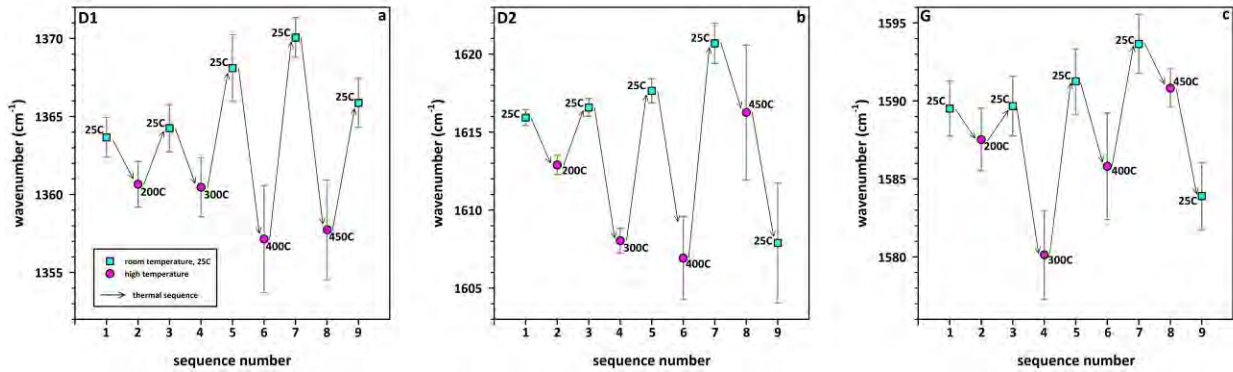


810

811 Figure 10. The effect of pressure (a) and high-pressure high-temperature (b) on first-order
812 kerogen Raman spectra are recoverable below 3.3 GPa and 450 °C. (a) Shown in black, spectrum
813 measured before compression and in red the spectrum immediately after direct compression to
814 3.1 GPa. After holding for 72 hours, the pressure increased to 3.3 GPa and was measured again
815 (green). On decompression back to room pressure, the spectrum recovered completely (grey) to
816 its original state. (b) Shown in black, spectrum measured before compression and heating. In red,
817 spectrum taken at 2.7 GPa and 450 °C. Several heating durations of 2-8 hours each were made,
818 with one example shown in green, taken following five continuous hours of heating at 450 °C
819 and 2.7 GPa. Following 24 hours of cumulative heating, the spectrum in grey was obtained at 1.4
820 GPa and room temperature.

821

822 **Figure 11**



823

824 Figure 11. Temperature cycling experiments at room pressure show the positions of D1, D2, and

825 G bands at room temperature (blue shaded squares) prior to and following heating (pink shaded

826 circles) to the temperatures indicated along a heating-cooling thermal sequence.

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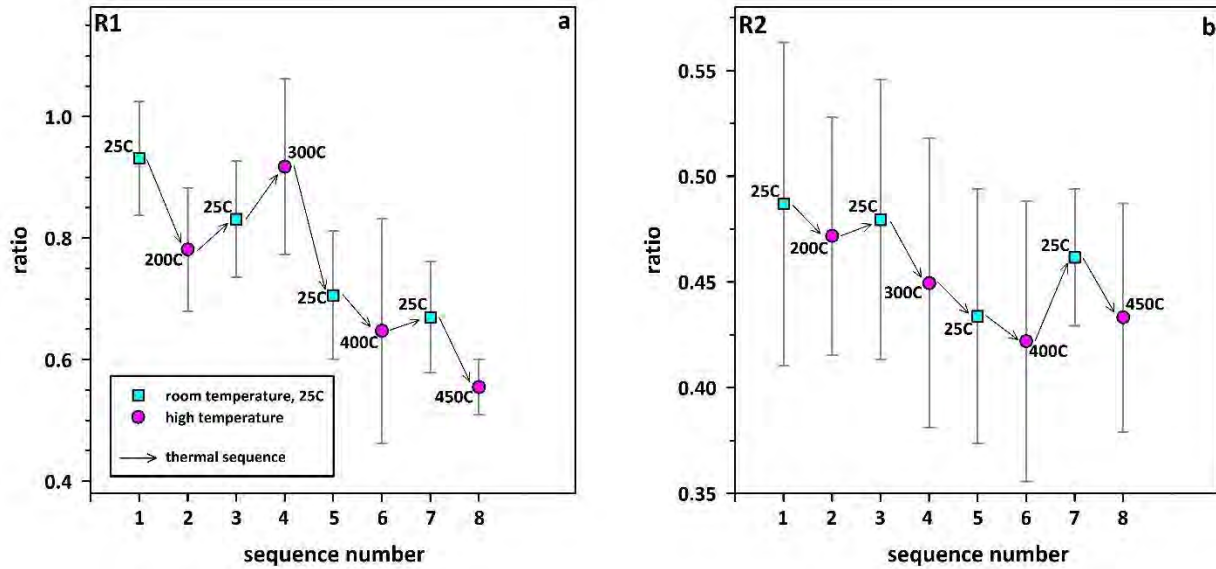
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842 **Figure 12**



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845 Figure 12. Temperature cycling experiments at room pressure show irreversible effects of
846 heating on the R1 and R2 ratios at room temperature (blue shaded squares) prior to and following
847 heating (pink shaded circles) to the temperatures indicated along a heating-cooling thermal

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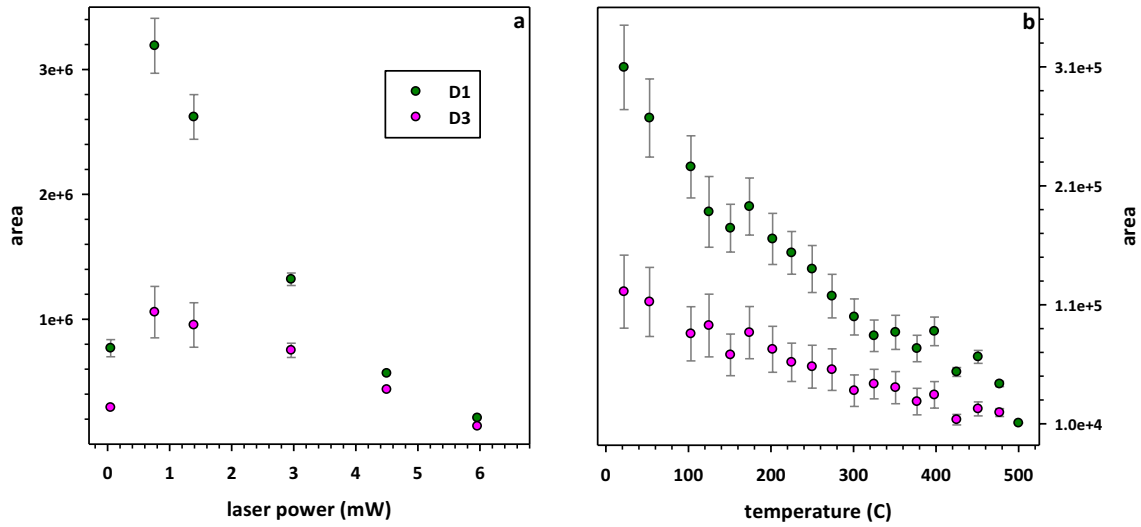
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860 **Figure 13**



861

862 Figure 13. Area of disordered D1 and D3 bands as a function of laser power (a) and temperature

863 (b), measured at 0.05 mW laser power at the sample.

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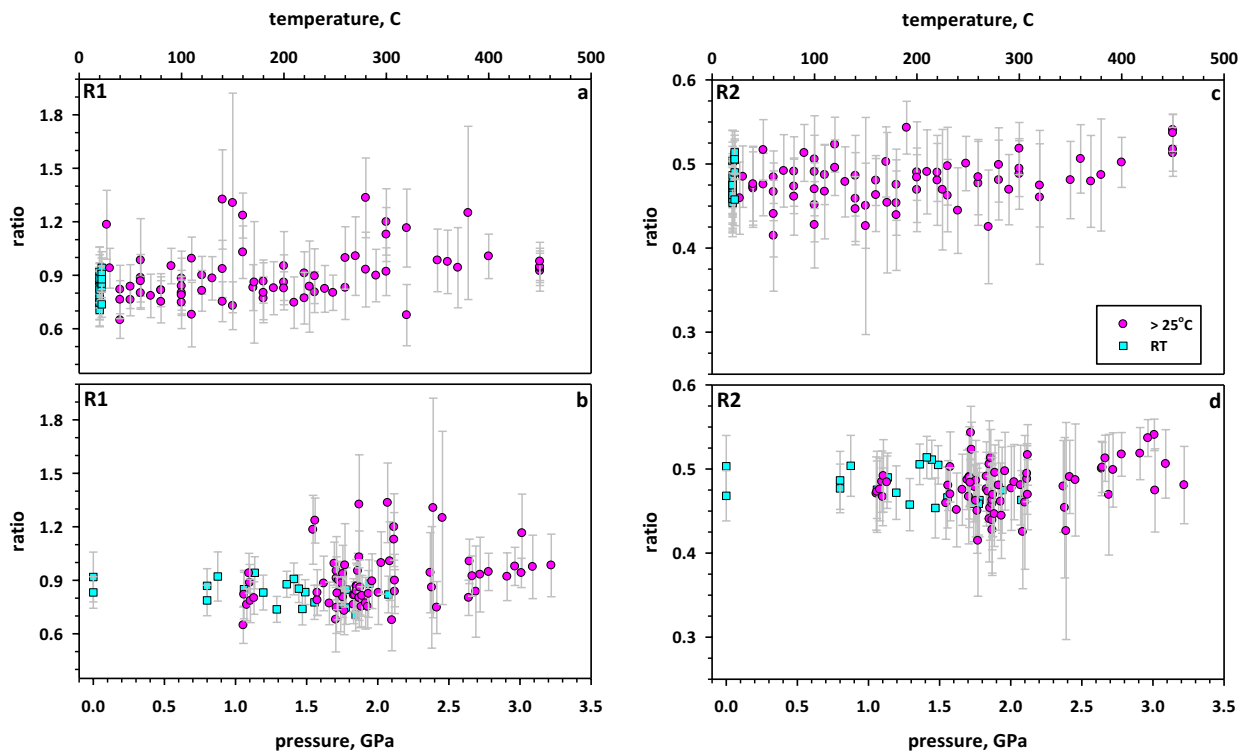
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878 **Figure 14**



879

880 Figure 14. R1 (a,b) and R2 (c,d) ratios of kerogen spectra at high pressure and high temperatures.

881 Data measured at high temperature are shown by pink shaded circles, and data measured at room

882 temperature are shown blue shaded squares. In contrast to heating at room pressure, where the

883 R1 and R2 ratios decrease with increasing temperature (see Figure 9), at high pressures (>1 GPa)

884 there is no discernable variation in R1 and R2 as a function of temperature up to 450 °C.

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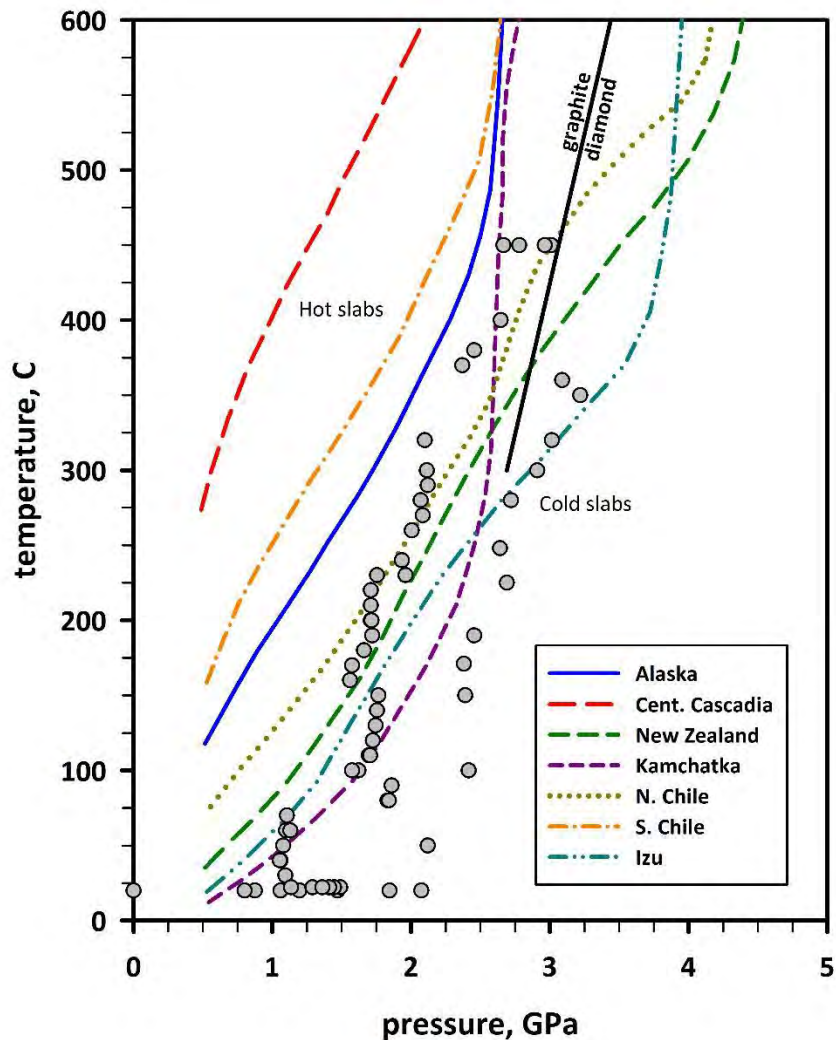
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892 **Figure 15**



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894 Figure 15. Pressure-temperature points visited in the current experiments in comparison to
895 several calculated geotherms in subducted slabs (Syracuse et al. 2010) along oceanic-continental
896 crust. The graphite-diamond phase boundary is shown by the solid black line (Kennedy and
897 Kennedy 1976). The structural integrity of kerogen to the highest P-T condition indicates that
898 instantaneous breakdown of kerogen will not occur in the upper part of slabs, potentially leading
899 to deeper subduction of carbon back into the mantle.