

1 **Revision 2**

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3 **U-Pb LA-ICP-MS dating of apatite in mafic rocks: evidence for a major**
4 **magmatic event at the Devonian-Carboniferous boundary in the Armorican**
5 **Massif (France)**

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Abstract

30 Apatite is a ubiquitous accessory mineral found in most magmatic rocks and is often the only U-
31 bearing mineral available to date mafic rocks because primary zircon and/or baddeleyite are not
32 present. In this paper, U-Pb LA-ICP-MS dating of apatite was applied to 7 different dike and sill
33 samples of dolerite from the Variscan belt of Brittany (Armorican Massif, western France). These
34 dolerites, which are characterized by a within-plate tholeiite geochemical signature, are organized
35 in several dense swarms across the belt. Their geochemical compositions are rather homogeneous
36 although they intrude a large geographical area subdivided into several domains each
37 characterized by different tectono-metamorphic settings. Their emplacement ages were so far
38 poorly constrained due to the difficulty to date these mafic rocks using either the $^{40}\text{Ar}/^{39}\text{Ar}$ or the
39 U-Pb methods on classical minerals like mica, plagioclase or zircon. Although the closure
40 temperature of apatite is lower than the emplacement temperature of the magma, physical models
41 show that the time needed to solidify and cool these mafic dikes and sills below the apatite
42 closure temperature is basically of the order of 100 years or less. Consequently, the U-Pb dates
43 obtained on apatite can be interpreted as the emplacement ages for these mafic intrusions. Our
44 results demonstrate that, in all cases, the apatite grains do carry enough radiogenic Pb to be dated
45 by in-situ U-Pb analyses and yield a ^{207}Pb -corrected mean age of 363.4 ± 5.8 Ma. These results
46 reveal the existence of a major and short-lived magmatic event in the Variscan belt of Brittany
47 during the Devonian-Carboniferous transition, a feature further highlighted by field evidence.
48 Beyond the geological implications of these results, U-Pb LA-ICP-MS dating of apatite appears
49 to represent an ideal tool to date small size mafic intrusions.

50

51 **Keywords:** Apatite, LA-ICP-MS, U-Pb dating, mafic rocks, French Variscan belt.

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Introduction

54 Apatite is a common accessory mineral in metamorphic, sedimentary, mafic and felsic igneous
55 rocks (Watson 1979, 1980; Green and Watson 1982; Harrison and Watson 1984; Piccoli and
56 Candela 2002; Spear and Pyle 2002; Webster and Piccoli 2015). The structure of apatite is very
57 stable and can accommodate a number of substitutions including Pb (McConnell 1938) and U
58 which may substitute for Ca via a charge coupled substitution involving a Ca-deficiency
59 (Oosthuyzen and Burger 1973; Baumer et al. 1983; Hugues and Rakovan 2002). Furthermore, it
60 is the most ubiquitous U-bearing mineral in mafic rocks, unlike zircon or baddeleyite, and can
61 consequently represent a powerful tool for U-Pb geochronology. Empirical and experimental
62 closure temperatures for Pb diffusion in apatite are between ca. 375°C and 550°C (Cherniak et al.
63 1991; Chamberlain and Bowring 2000; Harrison et al. 2002; Schoene and Bowring 2007;
64 Cochrane et al. 2014), which is a range of temperatures that is lower than the U-Pb closure
65 temperatures for zircon, baddeleyite, and titanite but higher than the one for Ar-Ar in biotite
66 (Reiners et al. 2005 and references therein). Furthermore, in small-size mafic bodies (dikes, sills
67 and plugs), the K-Ar whole rock and $^{40}\text{Ar}/^{39}\text{Ar}$ on plagioclase methods are often hard to apply
68 either because of the excess Ar acquired from the host-rock or because of the loss or excess of K
69 resulting from chemical weathering and/or alteration (Ruffet et al. 1992, Kelley 2002, Faure and
70 Mensing 2005). The $^{40}\text{Ar}/^{39}\text{Ar}$ dating of biotite and amphibole can be used but these minerals are
71 rarely present in mafic dikes and sills. Because of its medium-range closure temperature, U-Pb
72 dating of apatite may not appear to be the best choice to date the emplacement of hot and large,
73 slowly cooling, mafic bodies. However, for mafic dikes or sills, objects that are limited in size
74 with a fast cooling rate, it may represent a very pertinent mineral to use provided these rocks
75 have not experienced a reheating event above the closure temperature of apatite.

76 Apatite was first dated with the U-Pb method by Aldrich et al. (1955) and Tilton et al.
77 (1955). However apatite accommodates significant amounts of common Pb in its crystal lattice
78 (Chamberlain and Bowring 2000; Hugues and Rakovan 2002; Cochrane et al. 2014) and has
79 typically low U and radiogenic Pb concentrations and, consequently, rather low radiogenic
80 Pb/common Pb ratios. It is therefore necessary to perform a so-called common Pb correction in
81 order to distinguish the radiogenic Pb from the common Pb. Although this problem can be easily
82 resolved using the isotope dilution thermal ionization mass spectrometry (ID-TIMS) technique, it
83 becomes more problematic for in-situ analyses (secondary ion mass spectrometry (SIMS) or laser
84 ablation inductively coupled plasma mass spectrometry (LA-ICP-MS)). Indeed, in-situ analysis
85 requires the use of a standard (most often an external standard, i.e. a matrix-matching mineral
86 with a known age) in order to correct the raw data for the analytical biases inherent to this
87 technique, namely the downhole fractionation and the instrumental drift. Although it is easy to
88 find high quality natural and concordant standards for minerals that carry no common Pb (such as
89 zircon, monazite, or baddeleyite), it is very difficult for common Pb bearing minerals such as
90 apatite. Therefore, these standards, while potentially usable to correct the analyses for the
91 instrumental biases, do often carry common Pb, which in turns means that they are not
92 concordant. Thus, a common Pb correction also needs to be applied to the standards themselves.
93 Recent developments using either MC-ICP-MS and/or new data treatment, as well as the
94 discovery of age homogeneous apatite standards (Storey et al. 2007; Carrapa et al. 2009;
95 Thomson et al. 2012; Chew et al. 2011, 2014) have considerably improved the applicability of U-
96 Pb dating of apatite by in-situ LA-ICP-MS analysis. In this study, we implemented a new
97 approach developed by Chew et al. (2014) using apatite standards carrying variable common Pb.

98 This paper focuses on the application of U-Pb dating of apatite by in-situ LA-ICP-MS in
99 order to obtain the emplacement ages of various dikes and sills of dolerite from the Armorican

100 Massif, which is located in the western part of French Variscan belt. These mafic rocks are
101 widespread in the region and form several swarms of dikes crosscutting a Neoproterozoic to
102 Paleozoic sedimentary series. Some dikes are spatially associated with Sb and Au mineralization,
103 especially in the central part of the Armorican Massif (Chauris et al. 1985; Pochon et al. 2016).
104 This magmatism, described as characteristic of within-plate tholéiite, is constrained by some
105 previous studies (Le Gall and Mary 1983; Mary and Le Gall 1985; Houlgatte et al. 1988; Lahaye
106 et al. 1995; Le Gall 1999; Aïfa et al. 1999) but the precise timing of the emplacement of these
107 mafic rocks remains poorly constrained, in particular because of the scarcity of minerals suitable
108 for dating. Apatite is present as an accessory mineral in all of these rocks and appears to represent
109 the best candidate to date them. Petrography, whole rock analyses, and apatite geochemical
110 analyses have been performed on samples from these rocks in order to check for their degrees of
111 alteration/weathering and to confirm a magmatic origin for the apatite grains before performing
112 the LA-ICP-MS analyses.

113

114 **Geological background**

115 **Main structural features**

116 The Armorican Massif is located in western France and is part of the West European Variscan
117 belt which resulted from the continental collision between the Gondwana and Laurussia plates
118 and the Armorica microplate (Ballèvre et al. 2009, 2014). This collision began during the late
119 Devonian and continued until the early Carboniferous (Matte 1986, Ballèvre et al. 2014). The
120 Armorican Massif comprises three main domains with contrasted structural styles and
121 deformation histories: the North Armorican, Central Armorican, and South Armorican domains.
122 These domains are bounded by two dextral crustal-scale shear zones, the North Armorican Shear

123 Zone and the South Armorican Shear Zone (NASZ and SASZ respectively, Fig. 1). The North
124 Armorican domain was essentially affected by Neoproterozoic deformations and constituted a
125 part of the upper brittle crust un-metamorphosed during the Variscan orogeny (Brun et al. 2001).
126 The Central Armorican domain is made of Late Neoproterozoic to upper Paleozoic sediments
127 affected by moderate deformations (Gumiaux et al. 2004a, 2004b) and a low-grade regional
128 metamorphism during the Variscan orogeny. The maximum temperature reached is estimated to
129 be close to 250-300°C, based on vitrinite reflectance (Donnot et al. 1973), chloritoid-bearing slate
130 occurrences (Le Corre 1969), the illite crystallinity (Le Corre 1975), and the chlorite
131 geothermometer (Gloaguen et al. 2007). The South Armorican domain belongs to the internal
132 metamorphic zones of the Variscan belt. It was affected by crustal thickening during late-
133 Devonian to Carboniferous times, with a high pressure-low temperature event at ca. 360 Ma
134 (Bosse et al. 2000; 2005), followed by late-Carboniferous extension (Gapais et al. 1993, 2015).

135

136 **Magmatism in the Armorican Massif**

137 Several magmatic events occurred during the Paleozoic in the Armorican Massif. These
138 are first represented by bimodal magmatism and volcanism during the late Cambrian-early
139 Ordovician, featuring calc-alkaline and rare peraluminous series (Ballèvre et al. 2014 and
140 references therein). The late Cambrian event is best developed in the North Armorican domain,
141 whereas the early Ordovician event is best developed in the Central and South Armorican
142 domains.

143 Numerous Variscan granitoids have been emplaced in the Armorican Massif and are
144 characterized by different geochemical associations (Capdevila 2010; Tartèse and Boulvais 2010;
145 Tartèse et al. 2011a, 2011b; Ballouard et al. 2015). According to Capdevila (2010), 5 main

146 associations are currently recognized: i) a ca. 370 Ma (U-Pb method on zircon; Cuney et al.
147 1993) calc-alkaline association containing orogenic biotite-hornblende granites resulting from the
148 partial melting of an enriched mantle; ii) a ca. 330 Ma Mg-potassic metaluminous association that
149 includes mafic to intermediate rocks and porphyroid biotite-hornblende monzogranites (e.g.
150 Plouaret granite, 329 ± 5 Ma - whole-rock Rb-Sr, Peucat et al. 1984); iii) a Mg-potassic
151 peraluminous association made up of monzodiorite with a mantle origin, monzogranite and
152 cordierite granite from partial melting of metasediments; and iv) a ca. 316-310 Ma two mica
153 peraluminous leucogranites formed by the partial melting of a similar metasedimentary sources
154 (Tartèse and Boulvais 2010), e.g. the Guérande leucogranite (309.7 ± 1.3 Ma by U-Pb method on
155 zircon and monazite; Ballouard et al. 2015) and the Lizio granite (316 ± 1.3 Ma by U-Pb on
156 zircons, Tartèse et al. 2011b); and v) ca. 300 Ma red monzogranites and porphyroid syenogranites
157 (e.g. the Ploumanac'h granite, 303 ± 15 Ma - whole-rock Rb-Sr, Vidal 1980).

158 The North and Central Armorican domains are also marked by a dense swarm of dikes
159 and sills of dolerite that intrude the Late Proterozoic to Devonian sediments (Velde 1970; Ruffet
160 et al. 1992; Lahaye et al. 1995; Aïfa et al. 1999; Le Gall 1999). They are spatially distributed in
161 several groups (Fig. 1) defined by Le Gall (1999). The Mancellia and Saint-Malo group consists
162 of wide and dense swarms of dolerite dikes oriented N-S. The Laval Basin group comprises dikes
163 and sills of dolerite crosscutting Silurian and Devonian sediments and basaltic flows emplaced
164 within early Carboniferous series (Plaine 1976; Houlgatte et al. 1988; Pelhate 1994). The
165 Martigné-Ferchaud group consists of dike swarms striking N140° and spatially associated with
166 Sb mineralization (Chauris et al. 1985; Pochon et al. 2016). According to Le Gall (1999), the
167 dolerites of these groups present homogeneous geochemical compositions and appear to be
168 similar in composition to the thick mafic and felsic volcanosedimentary sequences of the
169 Châteaulin Basin in the western part of the Armorican Massif (Pelhate 1994; Caroff et al. 1996).

170 They are interpreted as within-plate tholeiites emplaced during a distensive/transtensive event or
171 during a transpressive phase prior to the late Carboniferous (Le Gall and Mary 1983; Mary and
172 Le Gall 1985; Houlgatte et al. 1988; Lahaye et al. 1995; Le Gall 1999; Rolet et al. 1994).
173 Although these dolerites are relatively common throughout the Armorican Massif (Fig. 1), their
174 ages are poorly constrained, due in particular to the difficulty to date this type of rock (Ruffet et
175 al. 1992). Only one doubtful whole-rock K-Ar age of 330 ± 10 Ma (Perroud et al. 1986) is so far
176 available. Because post-Devonian dikes are not observed in the field and because doleritic
177 fragments have been observed inside early Carboniferous volcanic breccias and conglomerate
178 (Plaine 1976), an age of around 360 Ma has been suggested for the emplacement of these mafic
179 intrusions (Le Gall, 1999).

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Description of the samples

182 Samples were collected within the un-metamorphosed North Armorican domain and the low-
183 grade Central Armorican domain (Fig. 1): two within the Saint Malo group (MED1 and Vi1);
184 three within the Laval Basin group (two in the southern part (PER1 and FOS9) and one in the
185 northern part (SDO1)); one within the Mancellia group (MGX1); and one from the Martigné-
186 Ferchaud group (MF1). Samples collected in the Saint Malo and Mancellia groups are very
187 similar and consist of dolerite dikes intrusive within the Neoproterozoic basement (Fig. 1 and 2a).
188 In the Laval Basin, two samples (PER1 and SDO1) consist of sills of dolerite intrusive within
189 Silurian black shales (Fig. 2b), and a third sample (FOS 9) consists of an altered sill of dolerite
190 found within Ordovician sandstones (Fig. 2c). In the Martigné-Ferchaud group, the sample
191 belongs to a dolerite dike crosscutting Ordovician slates and sandstones and was subsequently
192 partially hydrothermalized during the formation of Sb mineralization (Fig. 2d).

193 Despite their relatively wide geographical distribution, dolerites show similar textures and
194 mineralogy. Only the mafic sill FOS9 is different from the others because it contains cm-scale
195 phenocrysts of plagioclase that are strongly saussuritized and because it's magmatic texture is not
196 preserved. The textures of the other doleritic samples are mainly porphyritic and the principal
197 mineral assemblages consist of elongate plagioclase laths, clinopyroxene (mainly augite), and Fe-
198 Ti oxides such as ilmenite, rutile, and magnetite with exsolved ilmenite (Fig. 3a). Late stage
199 crystallization is represented by rare interstitial quartz (Fig. 3b, c and d) and/or quartz and alkali-
200 feldspar intergrowths (granophyric textures). The accessory mineral assemblage consists of
201 apatite, rare primary biotite and titanite, pyrite, chalcopyrite, pyrrhotite, chlorite, green
202 amphibole, and epidote (pistachite). Apatite is the most common of all the accessory minerals
203 and occurs in the groundmass of primary plagioclases but also in the mesostasis or within other
204 minerals such as quartz, amphibole, or clinopyroxene (Fig. 3a, b, c and d). Apatite appears as
205 euhedral and often acicular crystals up to 250 μm in length. Chemical weathering and
206 hydrothermal alteration of these Armorican dolerites are common and may be locally strong. It is
207 evidenced by the presence of sheets of silicates in interstitial glass, the destabilization of biotite
208 into chlorite, and the replacement of clinopyroxene and plagioclase by a typical propylitic
209 assemblage of chlorite, actinote, epidote, and carbonates.

210

211

Materials and methods

Major and trace elements analyses

212 Seven samples were collected mainly in quarries providing relatively fresh outcrop conditions.
213 Only the least altered rocks were finally selected for analyses. These selected samples were
214 crushed and powdered using agate mortars. Major and trace elements analyses, including REE,
215

216 were performed by the French analytical laboratory (SARM, CRPG-CNRS, Nancy) using an
217 inductively coupled plasma optical emission spectrometry (ICP-OES) and mass spectrometry
218 (ICP-MS), respectively, following the standard analytical procedures of Carignan et al (2001).

219 **Mineral compositions**

220 Apatite crystals were examined in polished thin-sections through a scanning electron microscope
221 (SEM equipped with EDS), and analyzed using an electron microprobe (Cameca SXFive –
222 Common analytical laboratory, BRGM-CNRS-University, Orléans, France). The major elements
223 Ca, P, and F as well as Cl, Fe, Mg, Si, Mn, Sr, Na, and S were analyzed in the apatite crystals. It
224 has been demonstrated that the measured X-ray flux from apatite is affected by crystal orientation
225 (Stormer et al. 1993), nevertheless, highly accurate analyses of F and Cl can be achieved using
226 low electron beam densities, large beams, and by analyzing grains with the electron beam
227 perpendicular to the apatite c axis (Goldoff et al. 2012). Analyses were performed using a 15 kV
228 accelerating voltage, a counting time of 10s, and a beam current of 6 nA for spot analyses, and
229 20nA for elemental mapping purposes, using the following mineral standards: Apatite (P $K\alpha$),
230 Topaz (F $K\alpha$), Hematite (Fe $K\alpha$), Andradite (Ca $K\alpha$, Mg $K\alpha$), $MnTiO_3$ (Ti $K\alpha$), Celestite (Sr $K\alpha$,
231 S $K\alpha$), Albite (Na $K\alpha$, Si $K\alpha$), and Vanadinite (Cl $K\alpha$). All the analyses were performed
232 unambiguously with the electron beam perpendicular to the apatite c axis because of the strongly
233 elongate shape of the grains. Several apatite grains were analysed per sample. The OH contents
234 of apatite have been estimated from EPMA measurements of F and Cl by charge balance,
235 assuming that the halogen sites are full (Piccoli and Candela 2002).

236 **U-Pb analyses**

237 Mineral separation procedures were applied to concentrate the apatite crystals using the facilities
238 available at Géosciences Rennes. Rocks were crushed and only the powder fraction with a

239 diameter < 250 μm has been kept. Heavy minerals were successively concentrated using the
240 Wilfley table and heavy liquids. Magnetic minerals were then removed with an isodynamic
241 Frantz separator. Apatite grains were carefully handpicked under a binocular microscope and
242 embedded in epoxy mounts. The grains were then grounded and polished on a lap wheel with a 6
243 μm and 1 μm diamond suspension successively. Apatite grains were imaged by
244 cathodoluminescence (CL) using a Reliotron CL system equipped with a digital color camera
245 available in Géosciences Rennes.

246 U-Pb geochronology was conducted by in-situ laser ablation inductively coupled plasma
247 mass spectrometry (LA-ICPMS) at Géosciences Rennes using a ESI NWR193UC excimer laser
248 coupled to a quadripole Agilent 7700x ICP-MS equipped with a dual pumping system to enhance
249 sensitivity (Paquette et al. 2014). The instrumental conditions are reported in Table 1.

250 The ablated material is carried by He, which is then mixed with N_2 and Ar, before injection into
251 the plasma source. The alignment of the instrument and mass calibration was performed before
252 each analytical session on the NIST SRM 612 reference glass, by inspecting the ^{238}U signal and
253 by minimizing the ThO^+/Th^+ ratio (<0.5%). During the course of an analysis, the signals of ^{43}Ca ,
254 $^{204}(\text{Pb}+\text{Hg})$, ^{206}Pb , ^{207}Pb , ^{208}Pb , and ^{238}U masses are acquired. The ^{235}U signal is calculated from
255 ^{238}U on the basis of the ratio $^{238}\text{U}/^{235}\text{U}=137.88$.

256 Single analyses consisted of 20 s background integration followed by 60s integration with
257 the laser firing and then a 10 s delay to wash out the previous sample. Ablation spot diameters of
258 50 μm with repetition rates of 5 Hz were used for most of the samples. In some samples,
259 however, the apatite grains were acicular and could not therefore accommodate a 50 μm spot.
260 The ESI NWR193UC laser is equipped with a rotational XY shutter allowing to perform
261 rectangular ablations while ensuring an even “dosage” of the laser energy to the sample during
262 the analyses. In these cases, we defined a rectangle that was suitable for most of the grains from a

263 given sample. This ablation rectangle, which can rotate freely around its center, was then used for
264 all the standards and the apatite crystals in order to use the same analytical conditions during the
265 analyses.

266 For each sample (one analytical session of 43 measurements), we used the following
267 standard bracketing procedure. Two analyses of the Madagascar apatite standard (ID-TIMS age
268 of 473.5 ± 0.7 Ma; Cochrane et al. 2014) used as the primary apatite reference material, one
269 analysis of the Durango apatite standard (31.44 ± 0.18 Ma; McDowell et al. 2005), one analysis
270 of the McClure apatite standard (523.51 ± 2.09 Ma; Schoene and Bowring 2006), followed by 6
271 analyses of the apatite grains. This sequence was then repeated 3 times with one analysis of the
272 Durango and two analyses of the Madagascar standards at the end of the session.

273 Data were corrected for U–Pb fractionation and for mass bias by the repeated
274 measurements of the Madagascar apatite standard. The Durango and McClure apatite standards
275 measurements were treated as unknowns and used to control the reproducibility and accuracy of
276 the corrections. During the course of the analyses, they provided ^{207}Pb corrected ages of $32.22 \pm$
277 0.52 Ma (MSWD = 0.69; probability = 0.93) and 526.4 ± 3.3 Ma (MSWD = 0.57; probability =
278 0.98) respectively (Fig. 4a and b respectively).

279 Data reduction was carried out using the data reduction scheme VizualAge_UcomPbine, a
280 set of Igor Pro procedures that work with Iolite (Chew et al. 2014). This data reduction scheme
281 performs a common Pb correction using the ^{207}Pb method for the initial Pb isotope composition
282 specified in the standard reference value file for the Madagascar apatite standard. Basically, a
283 model downhole $^{207}\text{Pb}/^{235}\text{U}$ fractionation curve is fitted with a ^{207}Pb -based common Pb correction
284 using an initial $^{207}\text{Pb}/^{206}\text{Pb}$ value of 0.8681 (Cochrane et al. 2014).

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286

Whole-rock geochemistry

287 To account for the degree of alkali-element mobility during hydrothermal alteration or
288 metasomatism, all the samples are plotted (Fig. 5a) in the $K_2O + Na_2O$ vs $K_2O/(K_2O + Na_2O)$
289 diagram of Hughes (1973). Most of the samples lie within the igneous spectrum field, with only
290 two samples (FOS9) and a sample from the Mancellia group analyzed by Le Gall (1999) plotting
291 outside of this field. This suggests that most of the mobile elements have not been significantly
292 affected after the magma emplacement. For the Laval Basin group, samples are slightly altered
293 with a loss of K_2O in some samples and almost a total loss of Na_2O in one sample (FOS9).
294 According to the Zr/TiO_2 vs Nb/Y plot (Fig. 5b), most of the samples display basaltic affinities
295 ($Zr/TiO_2 < 0.02$) with sub-alkalic ratios for the Saint-Malo group and the Martigné-Ferchaud
296 (MF1) dike ($Nb/Y < 0.8$) and with sub-alkali to alkali ratios for the Mancellia and the Laval
297 Basin groups ($0.4 > Nb/Y < 1.6$). Only one sample from the Mancellia group (MGX1) displays
298 an intermediate affinity. Most of the samples have a relatively low MgO content ($MgO < 6$ wt%)
299 except for PER1 (Table 2).

300 Data from this study, as well as the available values for different group of dolerites
301 (Lahaye et al. 1995; Le Gall 1999), have been plotted in a chondrite-normalized spider diagram
302 (Fig. 5c). Most of the dolerites display similar REE contents and sub-parallel REE patterns with a
303 slight to moderate fractionation of the light REE vs. the heavy REE [$(La/Lu)_N = 3.56-7.45$].
304 These patterns are similar to those reported for the Armorican dolerites by Le Gall (1999). Only
305 sample FOS9 yields a moderate negative Eu anomaly. Although this sample is strongly altered
306 (Fig. 5a), it shows a similar REE spectrum suggesting that the REE were not affected by
307 hydrothermal alteration and/or chemical weathering. The mantle-normalized spider diagram (Fig.
308 5d) displays similar incompatible element contents and ratios for all the samples, with the
309 exception of several elements such as Th, Hf-Zr, and Ti. Indeed, samples SDO1 and FOS9 show

310 slightly positive anomalies in Th that may be linked to the fact that the dolerites have
311 incorporated some country rocks during their emplacement. The Zr-Hf fractionations ($Zr/Hf =$
312 $35.65-46.07$) are very similar for all the samples. Their respective contents are relatively high in
313 sample MGX1 because of the presence of zircons inherited from the country rocks during
314 emplacement. Three samples display negative anomalies in Ti ($Ti/Ti^* = [Ti]_N/[Sm]_N+[Gd]_N =$
315 $0.11-0.30$) compared to the mean pattern for the Mancellia, Saint-Malo, and Laval Basin groups.
316 The largest negative anomaly is recorded in the dike MGX1 ($Ti = 1.87$ wt%). Finally, the
317 diagram also shows a light positive Y anomaly in sample PER1. According to mantle-normalized
318 trace element patterns proposed by Li et al. (2015), most of the samples are similar to average
319 compositions of alkaline continental flood basalts (CFB).

320

321

Apatite chemistry

322 We studied between 50 and 100 apatite grains for each of the 7 samples. In the MF1 and SDO1
323 samples, apatite grains appear as squat prisms up to 100 μm in length. In the PER1, MED1, and
324 Vi1 samples, apatite crystals have an elongate acicular shape up to 150 μm in length and a width
325 of ~ 25 μm . Finally, there are two types of apatite in samples MGX1 and FOS9: elongate crystals
326 with an acicular shape up to 250 μm in length and squat grains up to 150 μm .

327 SEM examinations, X-Ray maps (Fig. 6a, b, c and d), and electron microprobe data show that F
328 is always present in amounts above 0.62 apfu (Table 4), which indicates that all the apatite
329 crystals are fluorapatite. These fluorapatite crystals are unzoned or sometimes poorly zoned
330 around the rim. Oscillatory zoning has never been observed and, where present, the zoning is
331 mainly marked by differences in the Mg (0 to 0.04 apfu), Mn (0 to 0.07 apfu), and Fe (0 to 0.07
332 apfu) contents. X-Ray maps reveal that the edges are slightly enriched in Mg and Fe (Fig. 6b and
333 d) but the reverse has also been observed (eg sample SDO1). Average composition, standard

334 deviation, maximum and minimum (Table 4), demonstrate that these fluorapatite grains have a
335 narrow range of compositions and do not contain Sr despite the fact that Sr is available (between
336 48-455 ppm) in these rocks (Table 3). In the ternary OH-Cl-F diagram (Fig. 6e), most of the
337 grains match the compositions of fluorapatite from typically ultramafic/mafic complexes.
338 According to the data of Webster and Piccoli (2015), these compositions are the most frequently
339 encountered independently from the host-rock origin (mafic, felsic, and hydrothermal systems).

340

341 **Feasibility of U-Pb dating of apatite in mafic dike/sill**

342 In order to evaluate the feasibility to use apatite U-Pb to date the emplacement age of mafic
343 rocks, we performed some calculations and models dealing with the time necessary to crystallize
344 and cool these mafic dikes and sills. Firstly, we aim at estimating the time necessary to
345 completely crystallize a mafic dike or sill according to its thickness (up to 200 m) and the host-
346 rock temperatures (100, 200 and 300°C). We used the 1D analytical solution of the Stefan
347 problem transposed to dike and/or sill crystallization (see Turcotte and Schubert 1982 and Fig.
348 7a). For doleritic dikes and sills, we assumed that their rock parameters are close to basaltic ones.
349 Noteworthy, the solution implies that solidified doleritic magma and host rocks have the same
350 physical properties (i.e thermal diffusivity and specific heat capacity are similar for both
351 lithologies). The models show that the solidification can be considered as instantaneous (<20
352 years) on the geological time scale for a thickness less than 40 m whatever the host-rock
353 temperatures. When the mafic dikes or sills are thicker, the solidification times remain shorter
354 than 450 years. Consequently, the time necessary for the crystallization of the mafic dikes and
355 sills in this study is negligible.

356 Therefore, once solidified, we have to constrain the cooling temperature profile of a mafic
357 dike or sill in host rocks for different temperatures (100, 200, and 300 °C). At the end of the

358 solidification process, the temperature of the dolerite and its host rock has slightly changed when
359 compared to the solidus temperature (e.g. 1050 °C) and the initial host rock temperature.
360 However, in the following model, we consider that those changes are not significant considering
361 the thickness of the mafic dikes and sills sampled in this study. Therefore, we assume a
362 temperature of 1050 °C at the center of the dike or sill for $t=0$ (Fig. 7b). Models are built with a
363 2D finite element code using Comsol Multiphysics TM software in order to account for the
364 different physical properties between the dolerites and their host rocks, which is not possible with
365 the classical analytical 1D solution. We also calculated the apatite closure temperatures using the
366 Dodson (1973)'s equation with data from Cherniak et al. (1991). We used a radius ranging from
367 10 to 50 μm and the slowest cooling rate estimated for these dikes (7°C/year, Fig. 7b)
368 corresponding to a 60 m wide mafic sill or dike and a host rock temperature of 300°C. The
369 closure temperature obtained following these calculations range from 770 °C to 870 °C,
370 depending on the apatite radius. They are therefore well above the accepted closure range
371 between 375°C and 550°C because of the very fast cooling rate experienced by our samples.
372 Independent of the host rock temperature, the models show that time range for solidification of a
373 mafic dike or sill less than 60 m thick (maximum thickness of these intrusives) is very rapid (< 40
374 years). For a doleritic sill such as sample SDO1 (Fig. 1) with a thickness of ca. 60 m, the closure
375 temperature of apatite during post-solidification cooling is reached in less than 20 years.
376 Consequently, apatite appears to be a suitable mineral to date the emplacement age of such small-
377 size mafic bodies that have not been reheated above the closure temperature of apatite.

378 **Apatite dating**

379 We present the results obtained by U-Pb LA-ICP-MS dating on some of the apatite grains
380 described in the previous sections. For all the samples, the results are plotted in Tera-Wasserburg
381 diagrams using Isoplot/Ex (Ludwig 2012). All errors, given in Table S1 (supplementary data) and

382 the final results in Figures 8 and 9, are provided at 2σ level. In all the diagrams, the red dashed-
383 lines represent the unforced discordias, and the black lines represent the discordias calculated if
384 the initial common Pb value is forced to a $^{207}\text{Pb}/^{206}\text{Pb}$ value of 0.860 calculated following the Pb
385 evolution model of Stacey and Kramers (1975) for an age of 360 Ma.

386 In the Saint-Malo group, data obtained from samples MED1 (dolerite dike) and Vi1
387 (dolerite dike) are discordant with a relatively high proportion of common Pb ($^{207}\text{Pb}/^{206}\text{Pb}$ values
388 between 0.17–0.37 and 0.25–0.49, respectively). For sample MED1, the data define a lower
389 intercept date of 362 ± 13 Ma (MSWD = 1.4) with a $^{207}\text{Pb}/^{206}\text{Pb}$ initial value of 0.853 (Fig. 8a). If
390 the discordia is forced to a value of 0.860, we obtain a similar date of 363.0 ± 4.9 Ma (MSWD =
391 1.3). For sample Vi1, the unforced lower intercept date is 353 ± 17 Ma (MSWD = 1.8, initial
392 $^{207}\text{Pb}/^{206}\text{Pb}$ value = 0.81). If the discordia is forced to a $^{207}\text{Pb}/^{206}\text{Pb}$ value of 0.860, the resulting
393 lower intercept date is equivalent within error at 367.2 ± 5.6 Ma (MSWD = 2.1). The weighted
394 average ^{207}Pb -corrected dates (calculated using Stacey and Kramers (1975) terrestrial Pb
395 evolution model) are equivalent within error at 361.2 ± 6.3 Ma (MSWD = 0.62) for MED1 (Fig.
396 8c) and 363.9 ± 5.2 Ma (MSWD = 1.5) for Vi1 (Fig. 8d). The analyses of sample MF1 (dolerite
397 dike in Martigné-Ferchaud locality) are also discordant with $^{207}\text{Pb}/^{206}\text{Pb}$ values between 0.25 –
398 0.32 (Fig. 8e). The unforced discordia date is 346 ± 17 Ma (MSWD = 2.2, initial $^{207}\text{Pb}/^{206}\text{Pb}$
399 value of 0.774). If we calculate the lower intercept date by forcing the discordia to a $^{207}\text{Pb}/^{206}\text{Pb}$
400 initial value of 0.860, we end up with a more precise date, comparable within error, at 363.3 ± 3.1
401 Ma (MSWD = 2.6). The weighted average ^{207}Pb -corrected date is 364 ± 14 Ma (MSWD = 0.041;
402 Fig. 8f).

403 The analyses of MGX1 (dolerite dike from the Mancellia group) are discordant with
404 various amounts of common Pb ($^{207}\text{Pb}/^{206}\text{Pb}$ values between 0.25–0.55). They provide a lower
405 intercept date of 364.5 ± 9.7 Ma (MSWD = 0.39; $^{207}\text{Pb}/^{206}\text{Pb} = 0.837$; Fig. 8g). When the

406 discordia is calculated with an initial $^{207}\text{Pb}/^{206}\text{Pb}$ value of 0.860, the resulting lower intercept date
407 is similar within error at 361.9 ± 4.2 Ma (MSWD = 0.39). The weighted average ^{207}Pb -corrected
408 date is equivalent within error at 369.7 ± 6.3 Ma (MSWD = 0.101; Fig. 8h).

409 In the Laval Basin group, the analyses of FOS9 (intermediate volcanic rock), SDO1 (sill
410 of dolerite), and PER1 (sill of dolerite) are also discordant with variable amounts of common Pb.
411 For samples FOS9 and SDO1, the amount of common Pb is high ($^{207}\text{Pb}/^{206}\text{Pb}$ values between
412 0.50–0.77 and 0.48–0.57). The lower intercept dates are constrained at 376 ± 12 Ma (MSWD =
413 2.0; $^{207}\text{Pb}/^{206}\text{Pb}$ initial value of 0.877) and 344 ± 44 Ma (MSWD = 3.0; $^{207}\text{Pb}/^{206}\text{Pb}$ initial value of
414 0.83), respectively (Fig. 9a and b). If the discordias are forced to a $^{207}\text{Pb}/^{206}\text{Pb}$ initial value of
415 0.860, the resulting lower intercept dates are 362.0 ± 3.9 (MSWD = 2.5) and 367.1 ± 6.7 Ma
416 (MSWD = 3.1), respectively, while the weighted average ^{207}Pb -corrected dates are similar at
417 362.6 ± 8.5 Ma (MSWD = 0.21; Fig. 9c) for FOS9 and 363.1 ± 6.7 Ma (MSWD = 1.3; Fig. 9d)
418 for SDO1. LA-ICPMS analyses of apatite grains from sample PER1 show the highest amount of
419 radiogenic Pb ($^{207}\text{Pb}/^{206}\text{Pb} = 0.14\text{--}0.29$). The lower intercept date is at 354 ± 11 Ma (MSWD =
420 0.38; $^{207}\text{Pb}/^{206}\text{Pb}$ initial value of 0.81; Fig. 9e). If we calculate the lower intercept date by forcing
421 the discordia to a value of 0.860, we end up with a date comparable within error of 360.2 ± 3.8
422 Ma (MSWD = 0.44). The weighted average ^{207}Pb -corrected date is 359.8 ± 4.3 Ma (MSWD =
423 0.27; Fig. 9f).

424

425 Interpretation and discussion

426 In the Northern and Central domains of the Armorican Massif, especially in the eastern part, the
427 regional metamorphism is of very low-grade with maximum temperatures close to 250–300°C
428 (Donnot et al. 1973; Le Corre 1969, 1975; Gloaguen et al. 2007). Consequently, we can assume
429 that the isotopic systems of the dated apatite grains were not reset after the emplacement of the

430 dikes. In all the samples, the crystals are not hydrothermal but primary and magmatic in origin
431 because they are mainly found within magmatic plagioclase crystals.

432 In this study, the dolerite dikes are relatively thin with a maximum thickness of 10 m,
433 whereas the dolerite sill is up to ~ 60 m thick in the Laval Basin. Because of their small
434 thicknesses, numerical models show that the solidification and cooling of these mafic rocks was
435 very fast (Fig. 7a and b). Therefore, although the apatite closure temperature is far cooler than the
436 magma temperature, we can safely assume that the dates we obtain in this study can be
437 interpreted as the emplacement ages for the mafic intrusives.

438 Furthermore, the apatite U-Pb dates obtained for the 7 samples are similar within error. In
439 the case of the unforced discordias, the weighted average of the lower intercept dates is $360.3 \pm$
440 8.7 Ma (MSWD = 1.9), the ^{207}Pb -corrected mean date is 363.4 ± 5.8 Ma, and the weighted
441 average of the lower intercept dates forced to a $^{207}\text{Pb}/^{206}\text{Pb}$ initial value of 0.860 is 363.0 ± 1.7
442 Ma (MSWD = 1.05). Even the most altered samples collected in the Laval Basin (PER1 and
443 FOS9, Fig. 5a) yield a comparable age suggesting that the apatite U-Pb system is not necessarily
444 disturbed when the rocks suffered some degree of hydrothermal alteration and weathering.
445 Therefore, we conclude that all these samples were emplaced ca. 360 Ma ago. The $^{40}\text{Ar}/^{39}\text{Ar}$
446 dating of saussuritized plagioclase in a doleritic dike from the North Armorican domain give
447 anomalous dates (330 ± 10 Ma) because of the presence of excess argon (Ruffet et al. 1992).
448 Whereas apatite grains extracted from saussuritized plagioclase were not affected. This
449 demonstrates that apatite should be considered as the best mineral to date mafic dike and sill
450 emplacements even though they underwent propylitic alteration and very low-grade
451 metamorphism. Nevertheless, further studies are necessary to unravel precisely the action of
452 hydrothermal alteration and corrosive hot fluids on the apatite U-Pb geochronometer.

453 On the scale of our studied area (~160 x 200 km), within the Armorican Massif, we
454 demonstrate that a major mafic magmatic event took place ca. 360 Ma ago, at least in the eastern
455 part of the area. This age is confirmed by field evidence as doleritic fragments of the same nature
456 are found within early Carboniferous volcanic breccias and conglomerates (Plaine 1976), and
457 because dikes in post-Devonian rocks have never been observed (Le Gall 1999). This mafic
458 magmatic event, characterized by geochemically homogeneous compositions with no significant
459 crustal input, is interpreted as a within-plate type like alkaline continental flood basalts (Fig. 5d).
460 It occurred on a regional scale during a brief episode and appears to be equivalent to the mafic
461 and felsic volcanosedimentary sequences of the Châteaulin Basin in the western part of the
462 Armorican Massif (Pelhate 1994; Caroff et al. 1996). This similarity seems to indicate that this
463 major magmatic event can be generalized to the entire North and Central Armorican domains (ca.
464 40-45000 km²). This event appears coeval with a drastic change in the tectono-sedimentary
465 regime. Indeed, it coincides with the paroxysm of high pressure, low temperature metamorphism
466 recorded in the South Armorican domain (eclogite-facies in Champtoceaux complex, Bosse et al.
467 2000 and blueschist-facies in the Groix Island, Bosse et al. 2005) that marks the maximum depth
468 of the presently exposed rocks from the north Gondwanian margin. At the same time, the
469 distributed marine sedimentation becomes localized in the narrow Carboniferous basins (Fig. 1),
470 such as the Laval or Châteaulin Basins, where continental sedimentary deposits occur locally.
471 This period corresponds to the initiation of the collision *sensu stricto* of the Variscan orogeny in
472 the Armorican Massif (Ballèvre et al. 2014). The origin of this magmatism may result from a
473 melting of enriched lithospheric mantle (see the enrichment of LREE vs. HREE, Fig. 5c).
474 However, additional geochemical investigations (e.g. radiogenic isotopes) are necessary to
475 identify and characterize the geodynamical context and the deep sources responsible for this
476 major mafic magmatic event.

477

478

Implications

479 Mafic rocks are generally difficult to date due to a lack of suitable minerals. Consequently, the
480 dating of apatite by U-Pb LA-ICP-MS appears to be a very useful, quick, and pertinent method to
481 date mafic rocks, which were emplaced in a single injection as dikes or sills. Furthermore, it is
482 now evident that in-situ LA-ICP-MS dating, which is a fast and relatively inexpensive technique,
483 can be successfully applied. We can, however, consider that this dating approach may not be
484 suitable to date the emplacement age of large-scale layered mafic/ultramafic complexes or
485 intrusions characterized by repeated injection of magma. However, it might be convenient to
486 perform modelling of other examples of solidification and cooling in order to check the range of
487 sizes and emplacement contexts for mafic intrusions that could be successfully dated using the U-
488 Pb geochronometer on apatite. Prior to applying this dating technique, it is necessary to study the
489 regional geology, especially the metamorphic, hydrothermal alteration, and/or weathering events
490 that may have modified the chemistry and mineralogy of the studied rocks after their
491 emplacement. Because of the relatively good precision (ca. 5 Ma) obtained in the age
492 determination of the Variscan mafic rocks in this study, U-Pb LA-ICPMS geochronometry of
493 apatite appears to be a key mineral in unraveling the history of mafic magmatism.

494

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List of figure captions

728 **Figure 1:** Simplified geological map of the Armorican Massif modified from Chantraine et al.
729 (1996) and Le Gall (1999). P = Plouaret granite; G = Guérande granite; L = Lizio granite; Pl =
730 Ploumanac'h granite. Sample names and locations are reported in the WGS84 geographic
731 coordinate system: Vi1 (48°16'5.23"N, 1°45'20.03"W); MED1 (48°17'0.05"N, 1°40'1.67"W);
732 MGX1 (48°12'52.26"N, 0°41'55.74"W); PER1 (47°59'22.98"N, 0°44'30.32"W); SDO1
733 (48°0'34.05"N, 0°17'53.39"W); FOS9 (47°56'17.62"N, 0°42'20.79"W); MF1 (47°50'26.83"N,
734 1°18'0.71"W).

735

736 **Figure 2:** Representative field photographs showing: (a) A typical dolerite dike intruding
737 Neoproterozoic basement from the Saint-Malo group in the North Armorican domain. (b) A sill
738 of dolerite intruding Silurian black shales in the northern part of the Laval Basin. The white
739 dashed line represents the contact between the sill and the host rock. (c) A mafic intrusive facies
740 with phenocrysts of plagioclase within Ordovician sandstones in the southern part of the Laval
741 Basin. (d) A hydrothermalized dolerite dike associated with Sb mineralization in Ordovician
742 slates from the Martigné-Ferchaud group.

743

744 **Figure 3:** Selected thin section micro-photographs illustrating: (a) the typical texture of a dolerite
745 dike with laths of plagioclase intergrown with clinopyroxene and apatite crystals occurrence in
746 the groundmass of fresh and saussuritized plagioclase and acicular apatite in clinopyroxene. (b)
747 Primary amphibole and interstitial anhedral quartz (< 2 %) in dolerite dike. (c) Apatite growing in
748 primary amphibole and interstitial quartz. (d) Several apatite crystals in the groundmass of
749 plagioclase, interstitial quartz, and clinopyroxene. See Whitney and Evans (2010) for mineral
750 abbreviations.

751

752 **Figure 4:** Diagrams representing the weighted average ^{207}Pb -corrected measured ages for the
753 secondary apatite standards analyzed during this study: (a) Durango apatite and (b) McClure
754 apatite. Prob. = Probability.

755

756 **Figure 5:** (a) Whole-rock composition in the chemical diagram of Hughes (1973) for the dolerite
757 samples. (b) Zr/TiO₂ vs. Nb/Y diagram of Winchester and Floyd (1977) modified by Pearce
758 (1996) with alkaline, subalkaline, basic, and intermediate rocks subdivision. (c) Chondrite
759 normalized REE diagram and (d) mantle-normalized spider diagrams reporting the mean patterns
760 for the Mancellia, Laval Basin and Saint-Malo groups of dolerite. Chondrite and primitive mantle
761 normalizing values are taken from Sun and McDonough (1989).

762

763 **Figure 6:** Chemistry of representative poorly-zoned fluorapatite crystals using X-ray maps: (a) P
764 K α , (b) Fe K α , (c) F K α , and (d) Mg K α . (e) plot of the halogen contents (in molar proportions)
765 with fields from various mafic layered intrusions modified after Boudreau (1995).

766

767 **Figure 7:** (a) Solidification time (time necessary for a complete crystallization of the magma) vs.
768 thickness of doleritic dikes and/or sills for three different host rock temperatures (100, 200, and
769 300 °C). Calculation was performed using a crystallization temperature of 1050 °C, a latent heat
770 of 400 kJkg⁻¹, a mean specific heat capacity of 1 kJkg⁻¹K⁻¹ and a mean thermal diffusivity of 7e⁻⁷
771 m²s⁻¹. (b) Cooling temperature of the center of a mafic dike/sill vs. time as a function of different
772 thicknesses and host rock temperatures. For a mafic rock density of 3000 kg m⁻³, the specific heat
773 capacity is 800 Jkg⁻¹K⁻¹ and the thermal conductivity is 3.5 Wm⁻¹K⁻¹. Host rocks properties
774 correspond to averages for silico-clastic metasediments with a density of 2800 kg m⁻³, specific

775 heat capacity of $750 \text{ Jkg}^{-1}\text{K}^{-1}$, and thermal conductivity of $2 \text{ Wm}^{-1}\text{K}^{-1}$. Colors used correspond to
776 a dike or a sill with a thickness of 1 m (orange), 10 m (blue), 20 m (black), and 60 m (red).

777

778 **Figure 8:** Tera-Wasserburg concordia diagram with the corresponding ^{207}Pb -corrected average
779 dates for the dolerites from the Saint-Malo group with samples MED1 and Vi1 (a, b, c, and d),
780 the Martigné-Ferchaud group with sample MF1 (e and f) and the Mancellia group with sample
781 MGX1 (g and h). The red dashed-lines represent the unforced discordias and the black lines
782 represent the discordias calculated if the initial common Pb value is forced to a $^{207}\text{Pb}/^{206}\text{Pb}$ value
783 of 0.860 calculated following the Pb evolution model of Stacey and Kramers (1975) for an age of
784 360 Ma. N corresponds to the number of apatite grains that have been analyzed per sample.
785 Ellipses and errors are reported at 2σ .

786

787 **Figure 9:** Tera-Wasserburg concordia diagram and the corresponding ^{207}Pb -corrected average
788 dates for dolerites from the Basin Laval for samples FOS9, SDO1 (a, b, c and d) and PER1 (e and
789 f). See legend in the Figure 8.

790

Table 1. Operating conditions for the LA-ICP-MS equipment

U-Pb apatite analyses	
Laboratory & Sample Preparation	
Laboratory name	Géosciences Rennes, UMR CNRS 6118, Rennes, France
Sample type/mineral	Magmatic apatite
Sample preparation	Conventional mineral separation, 1 inch resin mount, 1µm polish to finish
Imaging	CL: RELION CL instrument, Olympus Microscope BX51WI, Leica Color Camera DFC 420C
Laser ablation system	
Make, Model & type	ESI NWR193UC, Excimer
Ablation cell	ESI NWR TwoVol2
Laser wavelength	193 nm
Pulse width	< 5 ns
Fluence	6.5 J/cm ²
Repetition rate	5 Hz
Spot size	50 µm (round spot) or 70x30 µm (rotational XY shutter)
Sampling mode / pattern	Single spot
Carrier gas	100% He, Ar make-up gas and N ₂ (3 ml/min) combined using in-house smoothing device
Background collection	20 seconds
Ablation duration	60 seconds
Wash-out delay	15 seconds
Cell carrier gas flow (He)	0.75 l/min
ICP-MS Instrument	
Make, Model & type	Agilent 7700x, Q-ICP-MS
Sample introduction	Via conventional tubing
RF power	1350W
Sampler, skimmer cones	Ni
Extraction lenses	X type
Make-up gas flow (Ar)	0.87 l/min
Detection system	Single collector secondary electron multiplier
Data acquisition protocol	Time-resolved analysis
Scanning mode	Peak hopping, one point per peak
Detector mode	Pulse counting, dead time correction applied, and analog mode when signal intensity > ~ 10 ⁶ cps
Masses measured	⁴³ Ca, ²⁰⁴ (Hg + Pb), ²⁰⁶ Pb, ²⁰⁷ Pb, ²⁰⁸ Pb, ²³² Th, ²³⁸ U
Integration time per peak	10-30 ms
Sensitivity / Efficiency	28000 cps/ppm Pb (50µm, 10Hz)
Dwell time per isotope	5-70 ms depending on the masses
Data Processing	
Gas blank	20 seconds on-peak
Calibration strategy	Madagascar apatite used as primary reference material, Durango and McClure apatites used as secondary reference material (quality control)
Reference Material info	Madagascar (Thomson et al. 2012) Durango (McDowell et al. 2005) McClure (Schoene and Bowring 2006)
Data processing package used	Iolite (Paton et al. 2010), VizualAge_UcomPbine (Chew et al. 2014)
Quality control / Validation	Durango: Weighted average ²⁰⁷ Pb corrected age = 32.22 ± 0.52 Ma (MSWD = 0.69; probability=0.93) McClure: Weighted average ²⁰⁷ Pb corrected age = 526.4 ± 3.3 Ma (MSWD = 0.57; probability = 0.98)

Table 2. Major elements from studied rock samples.

Sample	MED1	Vi1	SDO1	PER1^a	FOS9	MF1	MGX1
Facies	Dolerite	Dolerite	Dolerite	Dolerite	Andesite	Dolerite	Dolerite
SiO₂	47.43	43.96	50.91	42.32	51.82	46.30	54.74
Al₂O₃	13.26	15.19	13.92	11.75	20.23	15.58	12.42
Fe₂O₃	14.413	15	12.707	12.86	12.40	14.24	14.05
MnO	0.20	0.20	0.15	0.11	0.18	0.19	0.26
MgO	5.51	5.95	3.43	11.15	3.53	5.99	2.00
CaO	9.22	7.53	6.63	6.04	0.64	7.58	5.31
Na₂O	2.84	2.99	2.67	1.45	0.04	2.46	3.61
K₂O	0.56	0.99	1.83	0.1	1.94	0.99	1.85
TiO₂	3.56	3.38	3.10	2.58	1.81	2.81	1.87
P₂O₅	0.46	0.43	0.74	0.38	0.47	0.28	0.90
PF	1.35	3.50	2.95	11.6	8.22	2.84	2.05
Total	98.79	99.11	99.03	100.34	101.26	99.27	99.05

Notes: ^a The chemical composition of PER1 sample were performed by Le Gall (1999)

Table 3. Trace elements and REE from studied rock samples.

Sample	MED1	VII	SDO1	PER1	FOS9	MF1	MGX1
As	bdl	2.369	bdl	-	36.39	bdl	bdl
Ba	123.4	254.6	479.1	47	442.9	457.7	413.4
Be	1.54	1.609	2.326	-	6.587	1.169	3.339
Bi	bdl	bdl	bdl	-	bdl	bdl	bdl
Cd	0.286	0.233	0.421	-	0.411	0.296	1.499
Co	41.34	44.03	28.92	69	27.01	47.04	15.36
Cr	118.2	69.54	19.45	384	55.5	50	bdl
Cs	1.572	3.426	0.676	-	3.294	8.876	4.24
Cu	48.65	26.85	18.53	57	30.12	31.3	22.12
Ga	23.74	22.34	27.48	-	29.19	23.2	31.66
Ge	1.766	1.634	1.534	-	1.309	1.512	2.08
Hf	6.43	5.673	11.31	-	8.844	4.644	32.32
In	0.139	0.13	0.153	-	bdl	0.14	0.228
Mo	1.364	1.543	2.142	-	1.917	bdl	3.85
Nb	26.27	22.29	34.69	0	22.56	13.73	58.18
Ni	60.46	54.86	23.19	345	52.99	65.33	bdl
Pb	2.2269	1.9958	9.0967	-	9.4729	2.4961	8.926
Rb	13	43.54	41.94	bdl	58.25	56.89	42.94
Sc	32.34	24.09	21.03	0	23.48	28.24	28.8
Sb	bdl	0.251	bdl	-	4.723	bdl	bdl
Sn	2.96	2.536	3.331	-	3.446	2.302	6.048
Sr	307.9	318.6	281.1	222	48.66	455.1	234.1
Ta	2.087	1.716	2.638	-	1.746	1.114	4.166
Th	2.759	1.701	8.165	bdl	7.839	1.673	6.246
U	0.708	0.475	1.812	-	3.808	0.406	2.084
V	343.1	215.1	176.7	226	134.5	220.9	38.43
W	0.6	0.474	0.746	-	16.67	bdl	0.993
Y	37.32	34.47	50.66	25.17	55.81	32.81	89.84
Zn	142.4	132.7	165.9	bdl	195.8	161.9	251.4
Zr	267.6	239.4	403.2	bdl	351.6	166	1489
La	24.42	20.41	43.2	14.17	33.86	14.97	55.23
Ce	53.73	45.1	97.35	35.05	71.49	32.7	123.8
Pr	7.519	6.354	12.71	-	9.641	4.796	18.26
Nd	33.25	28.6	55.22	21.71	40.88	21.94	82.27
Sm	8.377	7.289	13.21	5.84	9.448	5.973	20.66
Eu	2.73	2.611	3.788	1.99	2.571	2.036	6.989
Gd	8.138	7.158	12	5.75	9.412	6.256	19.55
Tb	1.235	1.113	1.801	-	1.451	1.017	2.965
Dy	7.506	6.875	10.51	4.87	9.139	6.252	17.7
Ho	1.495	1.355	2.031	-	1.999	1.292	3.518
Er	3.682	3.364	5.012	2.03	5.272	3.282	8.922
Tm	0.484	0.46	0.646	-	0.742	0.442	1.216
Yb	3.191	3.014	4.245	1.76	4.982	2.886	7.943
Lu	0.451	0.435	0.602	0.26	0.776	0.436	1.217

Notes : bdl = below detection limit

Table 4. Average electron microprobe analyses and corresponding structural formulas of apatite

Sample	APATITE													
	MF1 (n=20)	σ	FOS9 (n=43)	σ	SDO1 (n=11)	σ	MED1 (n=23)	σ	Vi1 (n=10)	σ	MGX1 (n=17)	σ	PER1 (n=12)	σ
CaO	54.35	0.24	54.26	0.28	54.82	0.33	54.34	0.36	54.20	0.33	54.63	0.97	54.33	0.36
SrO	0.00	0.00	0.00	0.02	0.00	0.00	0.00	0.00	0.00	0.00	0.04	0.02	0.09	0.04
MgO	0.21	0.04	0.19	0.07	0.01	0.03	0.11	0.10	0.07	0.08	0.28	0.12	0.24	0.04
MnO	0.01	0.06	0.01	0.05	0.00	0.00	0.00	0.00	0.00	0.00	0.09	0.03	0.05	0.04
FeO	0.43	0.20	0.55	0.26	0.00	0.00	0.20	0.25	0.12	0.20	0.75	0.66	0.35	0.12
Na ₂ O	0.00	0.00	0.03	0.07	0.06	0.09	0.06	0.10	0.15	0.08	0.05	0.03	0.02	0.02
P ₂ O ₅	40.96	0.25	41.62	0.24	41.32	0.27	41.37	0.37	41.34	0.14	41.63	0.79	40.86	0.44
SiO ₂	0.63	0.12	0.14	0.07	0.23	0.05	0.47	0.21	0.33	0.11	0.31	0.39	0.58	0.14
SO ₃	0.00	0.00	0.00	0.01	0.00	0.00	0.00	0.02	0.00	0.00	0.00	0.00	0.00	0.00
F	2.48	0.09	2.94	0.17	3.34	0.14	2.71	0.21	3.12	0.41	2.84	0.29	2.91	0.15
Cl	0.18	0.04	0.16	0.05	0.37	0.04	0.18	0.05	0.09	0.03	0.17	0.03	0.25	0.04
Total	99.25	0.28	99.91	0.38	100.15	0.50	99.45	0.41	99.42	0.39	100.78	0.81	99.68	0.62
O=F	1.04	0.04	1.24	0.07	1.40	0.06	1.14	0.09	1.32	0.17	1.20	0.12	1.23	0.06
O=Cl	0.04	0.01	0.04	0.01	0.08	0.01	0.04	0.01	0.02	0.01	0.04	0.01	0.06	0.01
Total*	98.16	0.29	98.64	0.36	98.67	0.49	98.26	0.39	98.09	0.29	99.54	0.78	98.40	0.59
Structural formula on the basis of a 12.5 oxygen equivalent														
Ca	4.95	0.02	4.93	0.03	5.00	0.02	4.95	0.04	4.95	0.03	4.92	0.05	4.96	0.03
Sr	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
Mg	0.03	0.01	0.02	0.01	0.00	0.00	0.01	0.01	0.01	0.01	0.04	0.02	0.03	0.00
Mn	0.00	0.02	0.04	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.01	0.01	0.01
Fe	0.03	0.01	0.04	0.02	0.00	0.00	0.01	0.02	0.01	0.01	0.05	0.05	0.02	0.01
Na	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.02	0.03	0.01	0.01	0.01	0.00	0.00
P	2.95	0.01	2.99	0.01	2.98	0.01	2.98	0.02	2.98	0.01	2.96	0.03	2.95	0.01
Si	0.05	0.01	0.01	0.01	0.02	0.00	0.04	0.02	0.03	0.01	0.03	0.03	0.05	0.01
S	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
F	0.67	0.02	0.79	0.05	0.90	0.04	0.73	0.06	0.84	0.11	0.76	0.08	0.78	0.04
Cl	0.03	0.01	0.02	0.01	0.05	0.01	0.03	0.01	0.01	0.00	0.02	0.00	0.04	0.01
OH ^a	0.30	0.03	0.19	0.05	0.05	0.04	0.24	0.05	0.15	0.11	0.22	0.08	0.18	0.04

Notes: Oxide contents in wt.%, cationic contents in apfu and σ : standard deviation

Figure 1

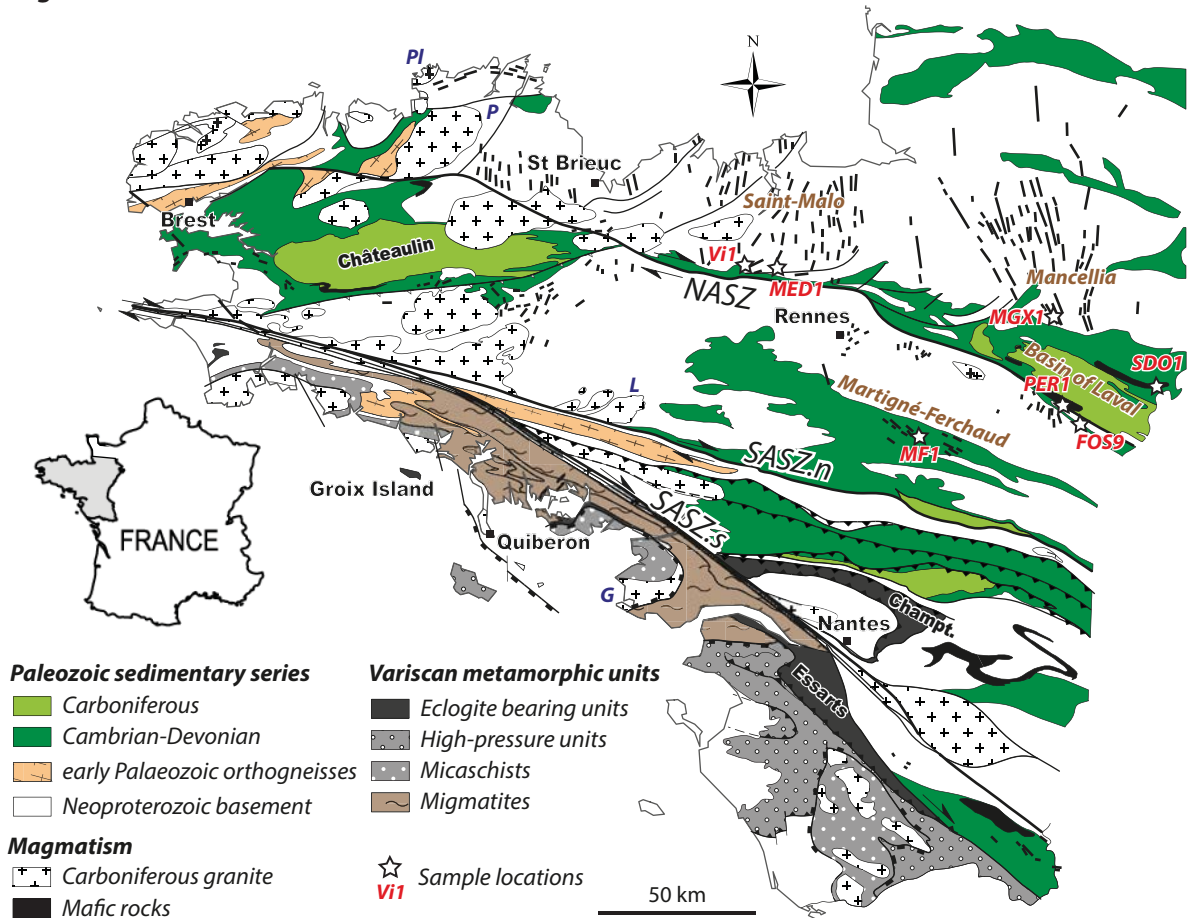
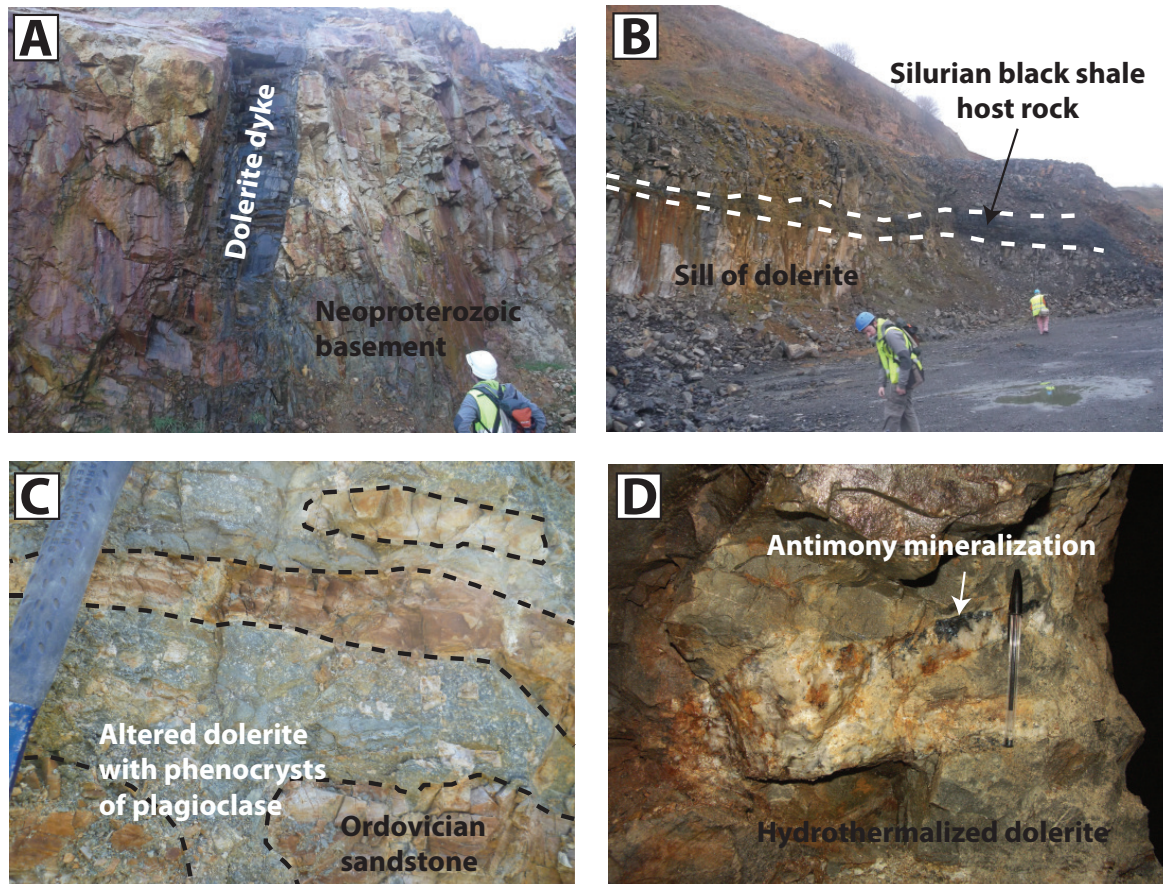
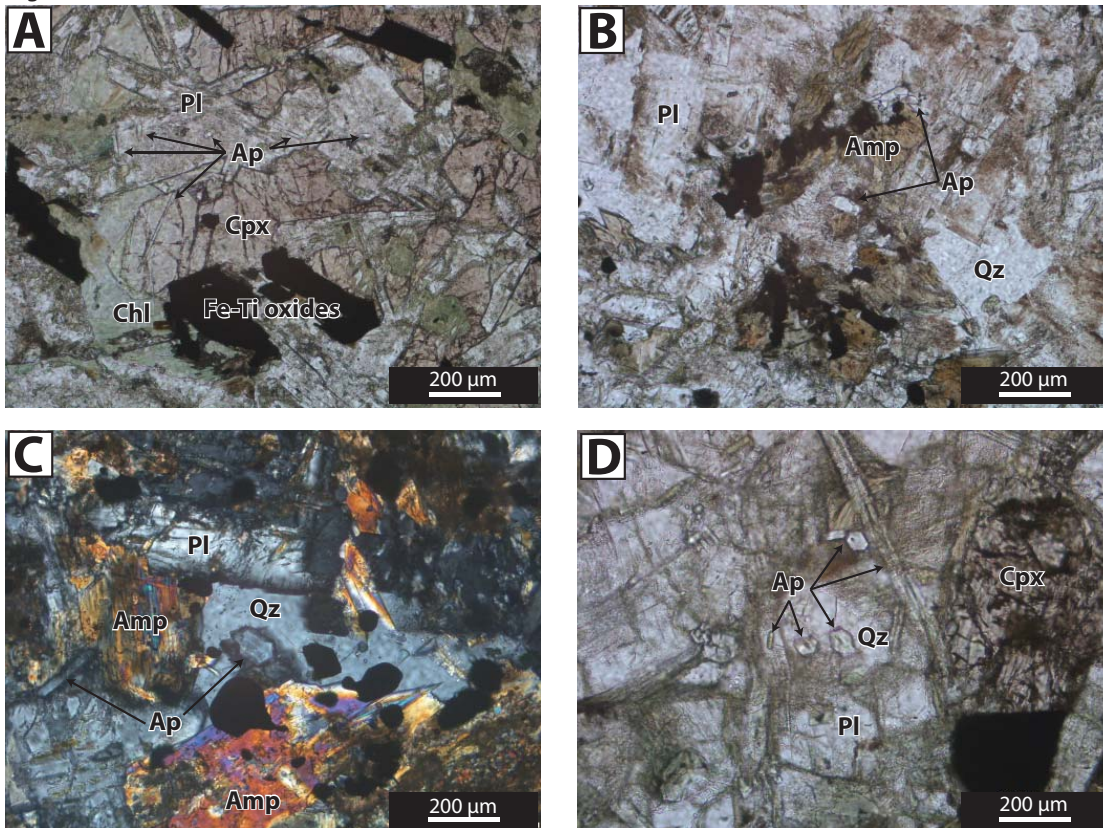


Figure 2



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Figure 3



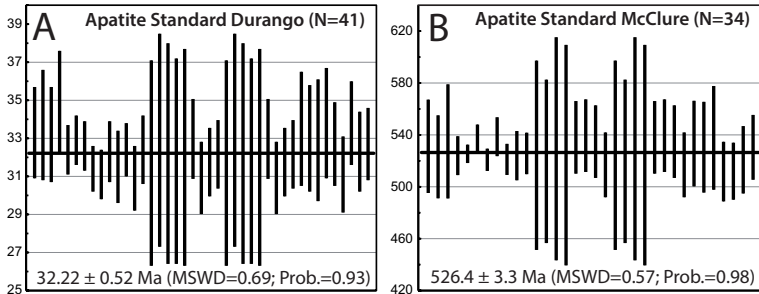
Always consult and cite the final, published document. See <http://www.minsocam.org> or GeoscienceWorld

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Cite as Authors (Year) Title. American Mineralogist, in press.

(DOI will not work until issue is live.) DOI: <http://dx.doi.org/10.2138/am-2016-5736>

Figure 4



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Figure 5

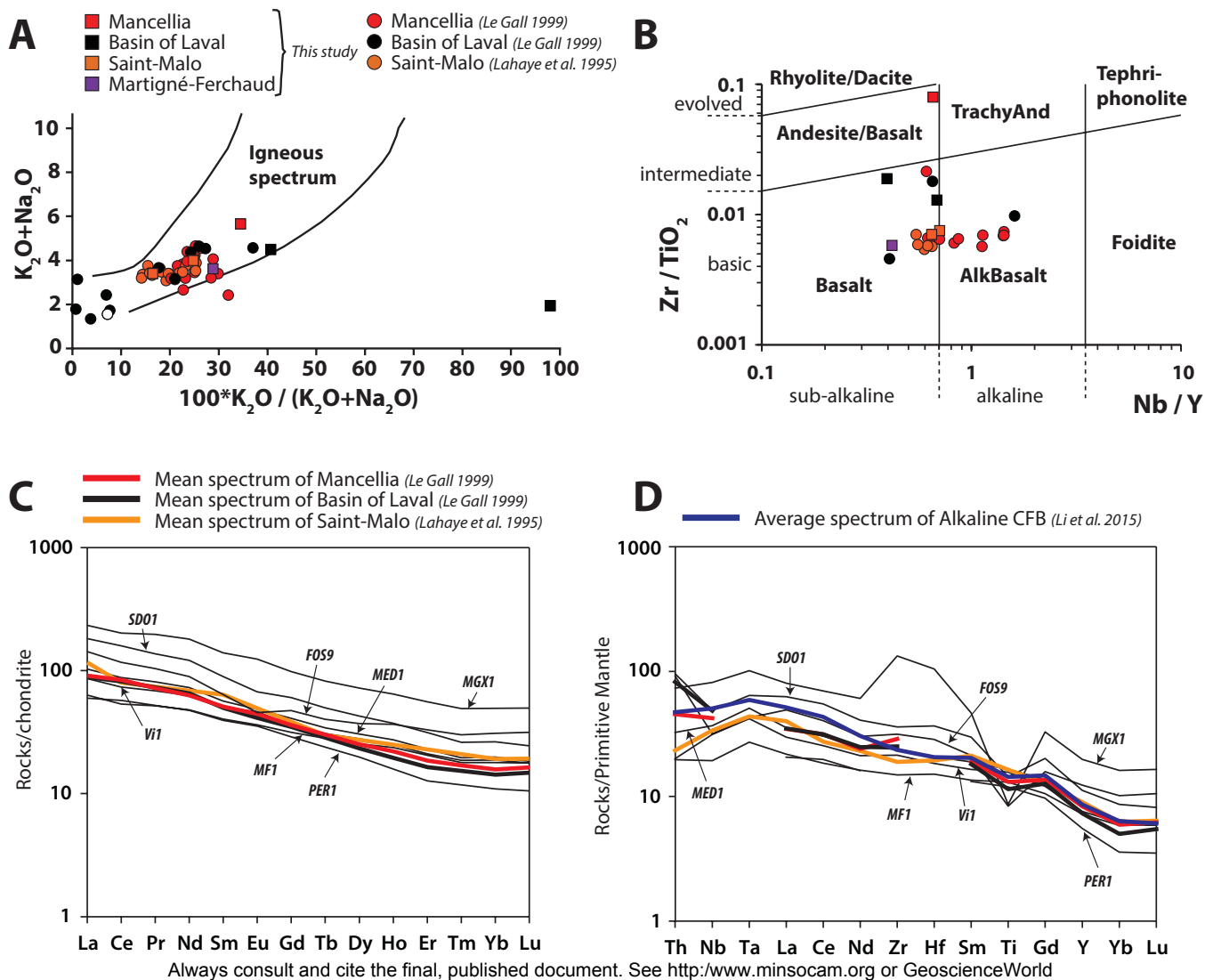


Figure 6

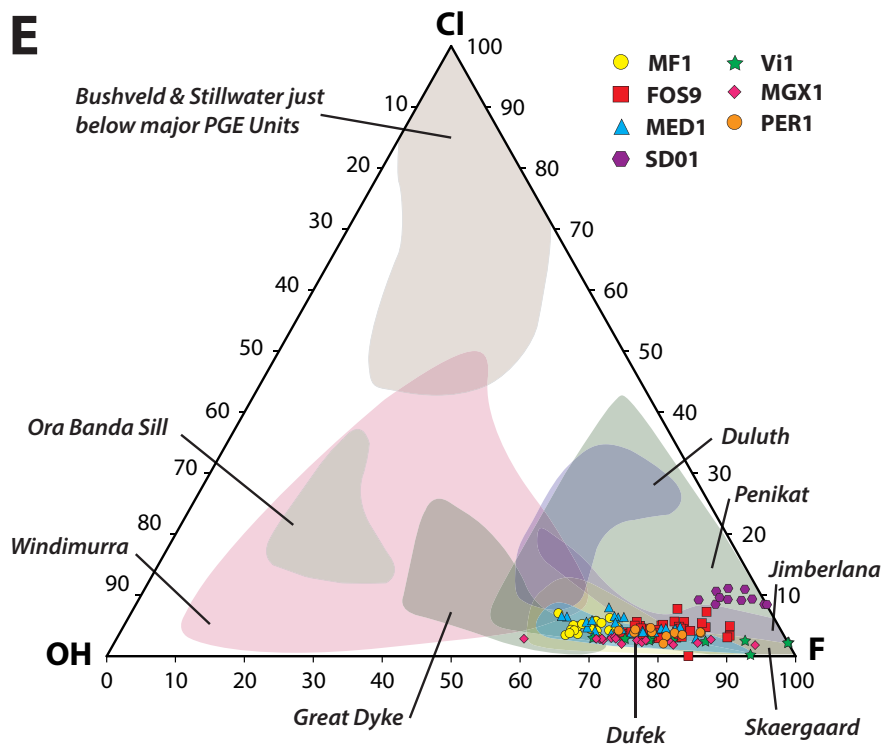
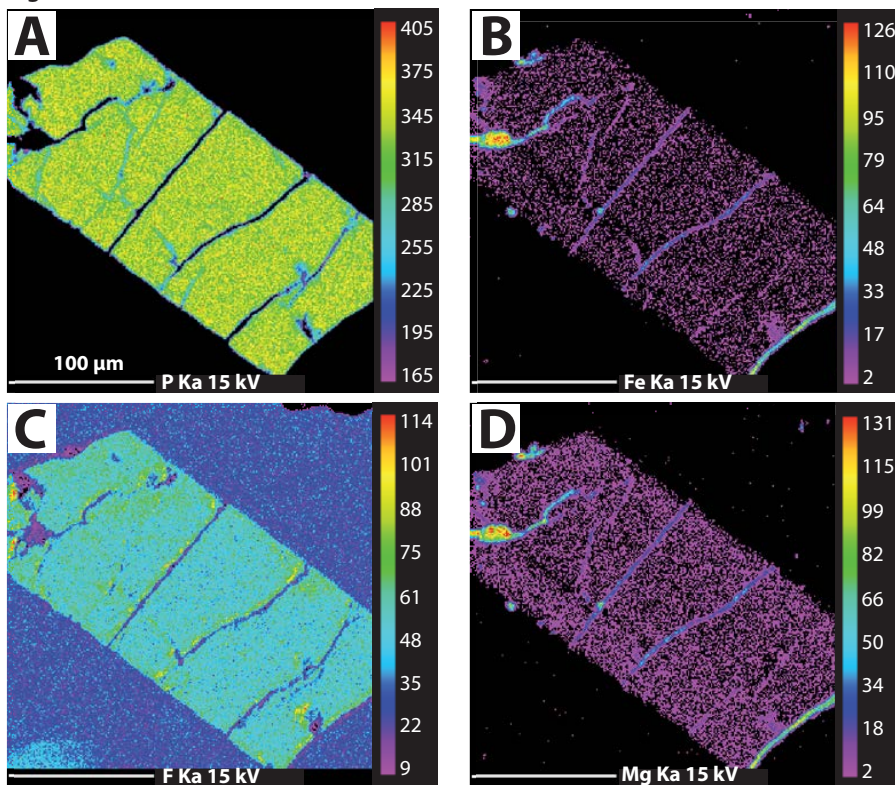
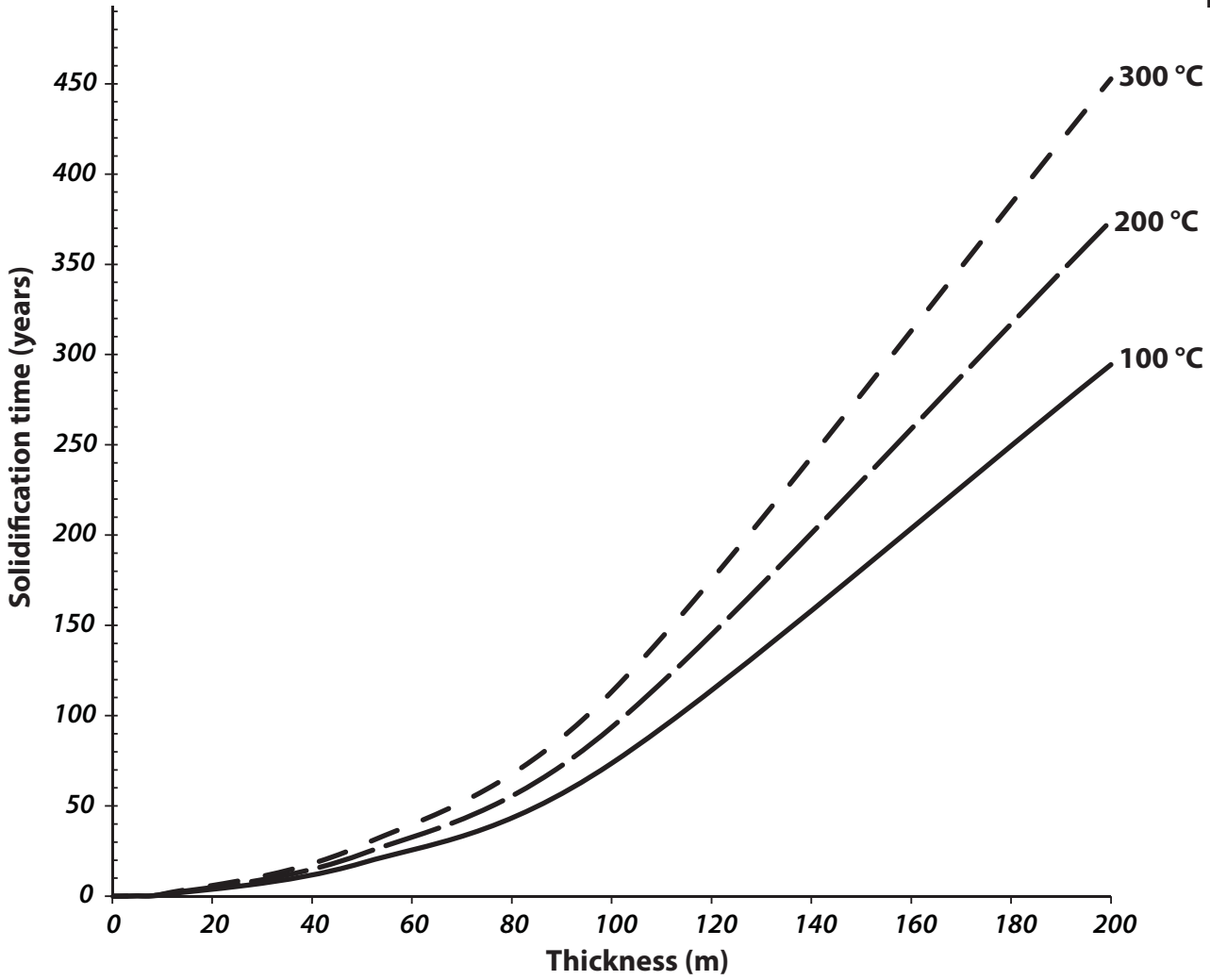


Figure 7

A



B

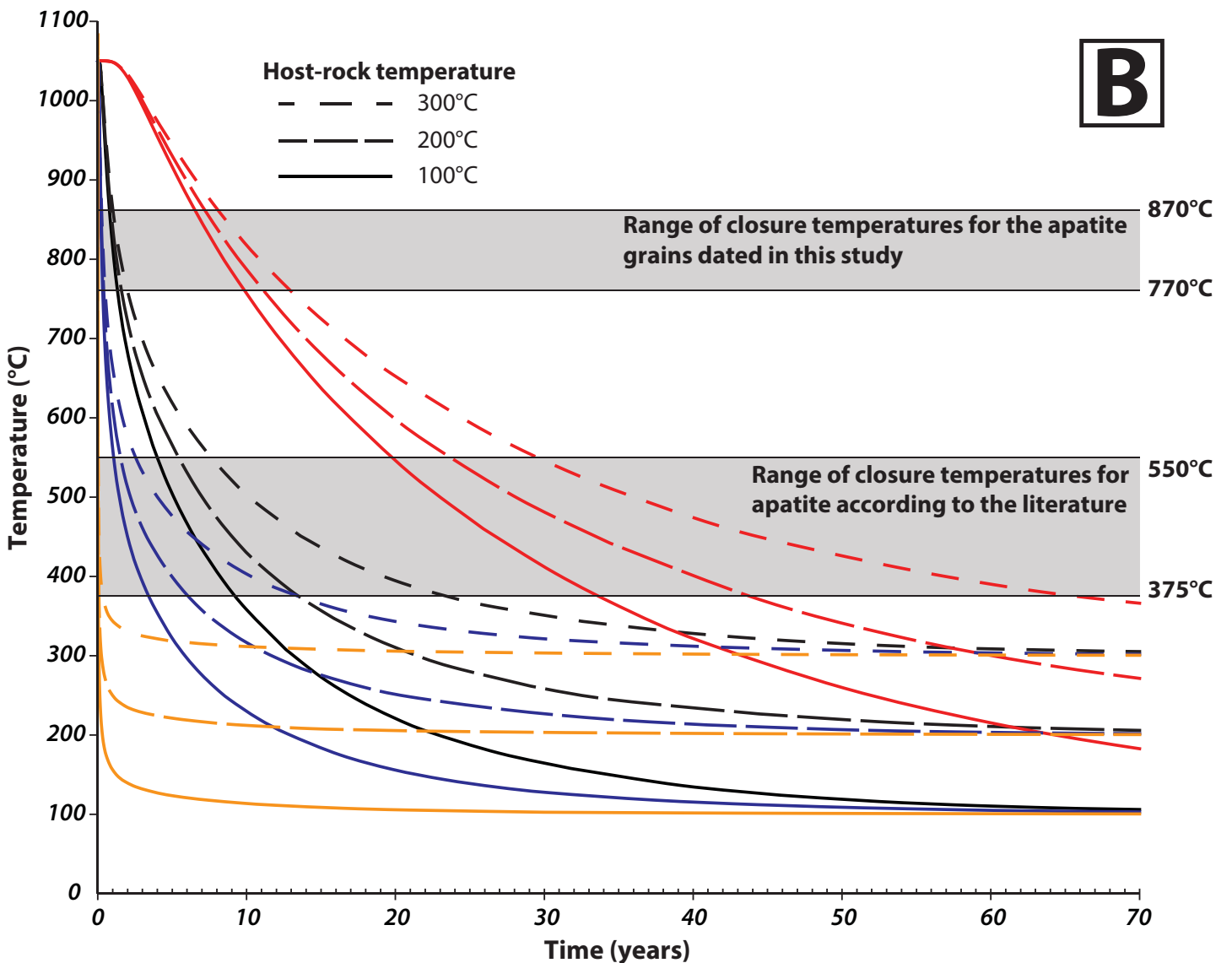


Figure 8

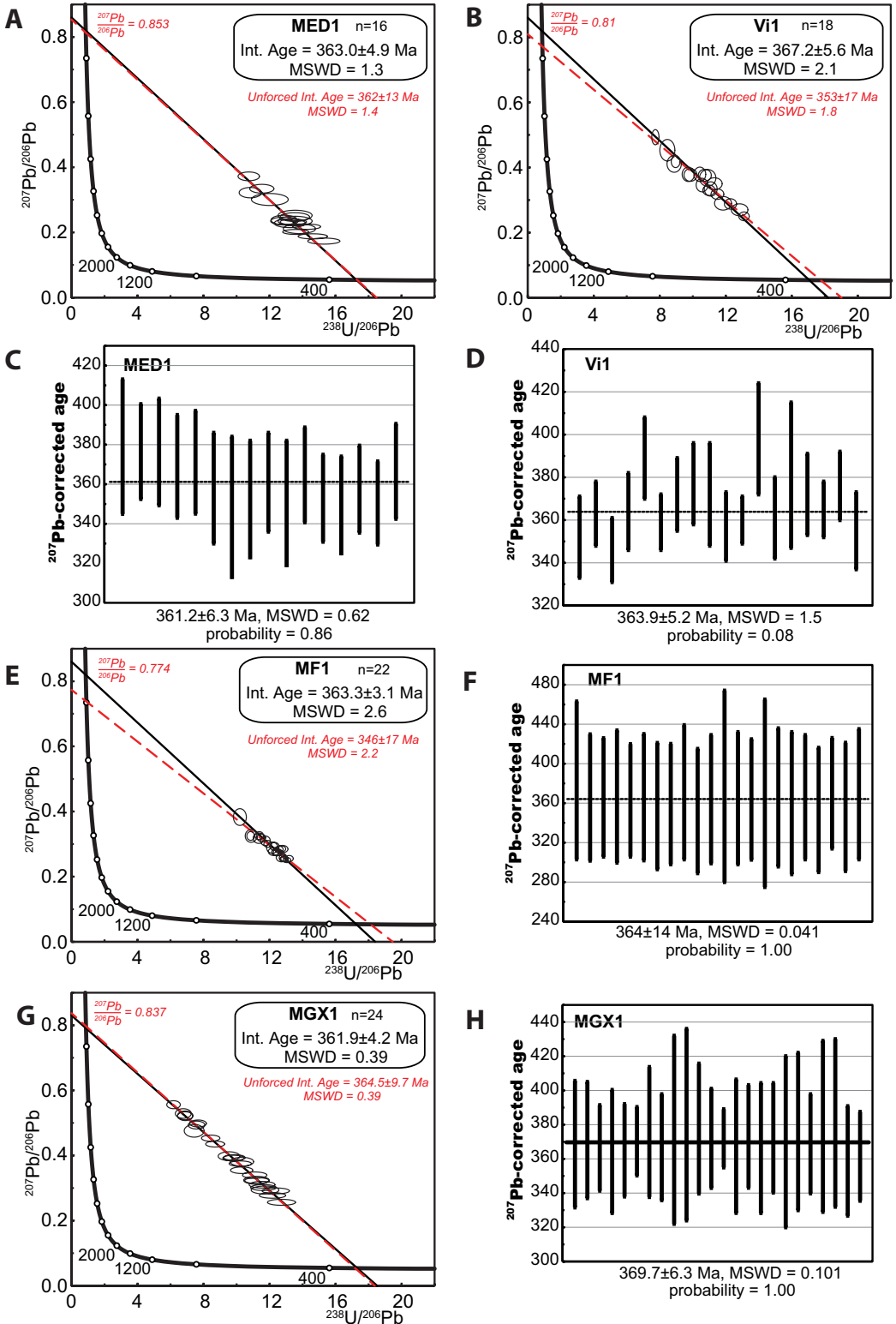


Figure 9

