

1 **Revision 1**

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3 **Thermodynamics of mixing in an isostructural solid solution:**  
4 **simulation methodologies and application to rutile-**  
5 **cassiterite system**

6 Xin Liu<sup>1,2</sup> Victor L. Vinograd<sup>3,\*</sup> Xiancai Lu<sup>1,\*</sup> Egor V. Leonenko<sup>4</sup> Nikolay N. Eremin<sup>4</sup>

7 Rucheng Wang<sup>1</sup> and Björn Winkler<sup>2</sup>

8 <sup>1</sup>State Key Lab for Mineral Deposits Research, School of Earth Sciences and  
9 Engineering, Nanjing University, 210023, Nanjing, China (xinliu.nju@gmail.com;  
10 xcljun@nju.edu.cn; rcwang@nju.edu.cn)

11 <sup>2</sup>Institute of Geosciences, Goethe Universität Frankfurt am Main, 60438, Frankfurt am  
12 Main, Germany (b.winkler@kristall.uni-frankfurt.de)

13 <sup>3</sup>Institute of Energy and Climate Research 6, Forschungszentrum Jülich, 52425, Jülich,  
14 Germany (v.vinograd@fz-juelich.de)

15 <sup>4</sup>Department of crystallography and crystal chemistry, Geological faculty, Moscow State  
16 University, 119991, Moscow, Russia (egorleo85@gmail.com; neremin@geol.msu.ru)

17

18 Corresponding Author

19 \*E-mail: v.vinograd@fz-juelich.de; xcljun@nju.edu.cn

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## ABSTRACT

27 The accuracies of two different approaches to model thermodynamic mixing properties of  
28 solid-solutions are explored using the rutile-cassiterite solid solution as an example. Both  
29 methods employ an expansion of the configurational enthalpy in terms of pairwise  
30 interactions energies. In the first method the partition function is directly computed from  
31 the excess energies of all Ti/Sn configurations within a  $2 \times 2 \times 4$  supercell. In the second  
32 method the free energy of mixing is calculated by a thermodynamic integration of the  
33 thermally averaged enthalpies computed with the Monte Carlo method using an  $8 \times 12 \times$   
34  $16$  supercell. The phase relations derived from Monte Carlo simulations agree well with  
35 the available experimental data, under the condition that the free energy is corrected for  
36 the effect of the excess vibrational entropy. The direct calculation of the partition  
37 function provides reasonable phase relations only when the configurational entropy is  
38 corrected to be consistent with the ideal mixing in the high-temperature limit. Advantages  
39 and drawbacks of the both approaches are discussed. The findings are generally  
40 applicable to models of isostructural solid solutions.

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42 **Keywords:** rutile-cassiterite, solid solution, first principles based calculations, pairwise  
43 interactions, Monte Carlo simulations, configurational statistics

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## 1. INTRODUCTION

47 Rutile-based solid solutions are important as geothermometers. The concentrations of  
48 certain trace elements, such as Zr, in rutile serve as indicators of the formation  
49 temperatures of its host rock (Cherniak et al., 2007; Watson et al., 2006; Zack et al.,  
50 2004a, 2004b). Rutile is also one of the main mineral phases of synroc, the synthetic  
51 ceramics proposed for nuclear waste storage (Ringwood et al., 1979a, 1979b; Xu and  
52 Wang, 2000; Zhang et al., 2001). Solid solutions with the rutile structure also attract  
53 attention as promising candidate materials for photocatalytic and photoelectrochemical  
54 energy conversion (Carp et al., 2004; Fujishima and Honda, 1972). The photocatalytic  
55 and semiconducting properties of TiO<sub>2</sub> can be modified by doping with various elements.  
56 Hence, it is important to understand the thermodynamic properties of this solid solution,  
57 and it is well suited as a benchmark to test different modelling approaches aimed at  
58 elucidating the properties of isostructural solid solutions.

59 In this study, a variety of state-of-the-art modelling approaches is applied to predict the  
60 extent of the miscibility gap in the rutile-cassitertite solid solution. The main focus is on  
61 methodical issues. Thermodynamic functions, which determine the stability of a solid  
62 solution, could be computed with methods of statistical mechanics if the energy spectrum  
63 of possible configurations of exchangeable atoms is known with sufficient detail. In  
64 essence, a model must provide the possibility to evaluate the energy and the Boltzmann  
65 weight of any configuration of exchangeable atoms within a supercell of a reasonably  
66 large size. However, practically, this detailed configuration-dependent information can be  
67 obtained only by computations based on first principles, while a direct sampling of all  
68 configurations within a large supercell (i.e. the computation of their energies *ab initio*) is

69 currently not feasible. In general, two main strategies are used to overcome this problem.  
70 One strategy is to directly compute the energies of all configurations, but within a  
71 supercell of manageable (smaller) size. The alternative way is to take a sufficiently large  
72 supercell, explicitly compute a small set of configurations and employ an interpolation  
73 method with which the energy of any configuration can be approximately estimated. For  
74 the second alternative there are two further options. Using an interpolation equation one  
75 can calculate the energies of all configurations and directly compute the partition function  
76 (Becker et al., 2000; Prieto et al., 2000), or apply a Monte Carlo (importance sampling)  
77 algorithm to compute approximately the equilibrium enthalpy of mixing and then  
78 evaluate the free energy indirectly by a thermodynamic integration (Bosenick et al., 2001;  
79 Dove, 2001; Myers, 1998; Reich and Becker, 2006; Warren et al., 2001). One of the aims  
80 of this study is to compare the advantages and disadvantages of the last two approaches.  
81 We adopt the following strategy. At first we explicitly compute the excess energies of 20  
82 specially chosen structures within a  $2 \times 2 \times 4$  supercell (32 cations) using density  
83 functional theory (DFT). Then we employ an interpolation method known as the J-  
84 formalism (Becker et al., 2000; Bosenick et al., 2000, 2001; Dove, 2001; Dove et al.,  
85 1996, 2000) to compute the pairwise interactions (i.e. the energies of the exchange  
86 reactions  $\text{SnSn} + \text{TiTi} = 2\text{SnTi}$ ) at all distances within a  $2 \times 2 \times 4$  supercell of rutile.  
87 These energies are then used to derive a generalized Ising-type Hamiltonian of this  
88 system, to compute the energies of all possible configurations of Sn and Ti in the  $2 \times 2 \times$   
89  $4$  supercell of rutile and to evaluate the partition function. A correction procedure after  
90 Becker et al. (2000) is applied to thermodynamics of mixing to make it consistent with  
91 the regular mixing behavior in the high-temperature limit. The same Hamiltonian is then

92 applied to compute the thermodynamic functions of the system with the aid of the Monte  
93 Carlo method by employing an  $8 \times 12 \times 16$  supercell. Finally, phase diagrams predicted  
94 with the both methods are compared to available experimental data (Garcia and Speidel,  
95 1972; Naidu and Virkar, 1998; Padurow, 1956; Park et al., 1975).

96

## 97 **2. STATE-OF-THE-ART MODELLING APPROACHES**

### 98 **2.1. Configuration dependent and configuration independent approaches.**

99 Although all configurations are needed for an adequate model, a variety of configuration-  
100 independent approaches exist. At an infinitely high temperature the Boltzmann weights of  
101 all configurations are equal and the distribution of exchangeable atoms over the lattice is  
102 perfectly random. The mixing behavior of such a solid solution can be assessed with the  
103 help of a quasi-random structure (Jiang, 2008; Wei et al., 1990; Zunger et al., 1990). A  
104 similar randomization occurs in a diluted solution, although for a different reason. The  
105 mixing properties in the dilute limit can be accurately estimated from the excess energy  
106 of a supercell structure of a pure (host) phase containing a single substitutional defect of a  
107 solute component (Sluiter and Kawazoe, 2003; Vinograd et al., 2013). The solid solution  
108 behavior both in the dilute and high-temperature limits can be adequately modeled with  
109 the regular or subregular model (Ganguly, 2001). The configuration independent methods  
110 provide very useful information on the limiting cases and can serve as good tests for the  
111 performance of more sophisticated models.

112 Configuration dependent approaches can be subdivided into two groups. The first group  
113 of methods attempts a direct evaluation of the partition function and the related  
114 thermodynamic quantities by sampling *all* configurations within a supercell (Grau-

115 Crespo et al., 2004, 2007; Liu et al., 2015; Todorov et al., 2004), while other methods  
116 compute the energies of only a limited set of configurations and, based on this set,  
117 evaluate the energies of the remaining majority of configurations via a parameterized  
118 equation (Becker et al., 2000; de Fontaine, 1992, 1994;; Finel, 1994; Laks et al., 1992;  
119 Sanchez and de Fontaine, 1978; Sanchez et al., 1984; van de Walle and Ceder, 2002a;  
120 Vinograd et al., 2007). Within the second group of methods there are variants, which  
121 depend on the approach chosen for the parameterization. One approach, which is now  
122 extensively used within materials science, is known as the cluster expansion (de Fontaine,  
123 1994; de Fontaine et al., 1992; Finel, 1994; Laks et al., 1992; Sanchez and de Fontaine,  
124 1978; Sanchez et al., 1984; van de Walle and Asta, 2002; van de Walle and Ceder, 2002a,  
125 2002b). A similar approach, known as the J-formalism became popular in geosciences  
126 (Becker et al., 2000; Bosenick et al., 2000; Bosenick et al., 2001; Dove, 2001; Dove et  
127 al., 1996, 2000; Ferriss et al., 2010; Jung et al., 2010; Kulik et al., 2010; Reich and  
128 Becker, 2006; Renock and Becker, 2011; Vinograd et al., 2007, 2009). Both approaches  
129 are briefly outlined below.

130

## 131 **2.2. Cluster expansion method and the J-formalism method.**

132 The formalism of the cluster expansion method introduces the occupation variables  $\sigma_i$ ,  
133 which take values 1 or  $-1$  depending on whether a lattice point  $i$  is occupied by A or B  
134 atom. A cluster  $\alpha$  is defined as a set of points  $i, j, k, \dots$ , which has a certain fixed location  
135 within the lattice and forms a figure of a certain shape (point, pair, triangle, tetrahedron,  
136 etc.). The configuration  $\sigma_i, \sigma_j, \sigma_k, \dots$  of a cluster  $\alpha$  can be associated with the cluster  
137 function

138  $\varphi_\alpha(\sigma) = \sigma_i \sigma_j \sigma_k \dots$  (1)

139 It is shown that the cluster functions are orthonormal with respect to the scalar product  
140  $\langle \varphi_\alpha(\sigma) \varphi_\beta(\sigma) \rangle$  defined as the normalized sum of products  $\varphi_\alpha(\sigma) \varphi_\beta(\sigma)$  over all cluster  
141 configurations (de Fontaine, 1994; de Fontaine et al., 1992; Sanchez et al., 1984). From  
142 this orthonormality it follows that any function of configuration of a supercell, such as its  
143 enthalpy, can be expanded as

144  $H(\sigma) = \sum_\alpha J_\alpha \varphi_\alpha(\sigma)$  (2)

145 where

146  $J_\alpha = \langle \varphi_\alpha(\sigma) H(\sigma) \rangle$  (3)

147 The macroscopic ensemble averaged enthalpy of a model system (a phase) can be then  
148 written as

149  $\langle H \rangle = \sum_\alpha J_\alpha \xi_\alpha$  (4)

150 where  $\xi_\alpha$  is the ensemble averaged cluster function, which is called the correlation  
151 function of the cluster  $\alpha$ . The practical advantage of Eq. 4 consists in its rapid conversion  
152 with respect to an addition of terms corresponding to clusters of larger size. This  
153 mathematically rigorous expansion method includes a rigorous definition of the effective  
154 cluster interactions via Eq. 3. From this definition, the effective pair cluster interaction  
155 energy takes the form

156  $J_{pair(n)} = \frac{1}{4} \left( \langle H_{AA(n)} \rangle + \langle H_{BB(n)} \rangle - \langle H_{AB(n)} \rangle - \langle H_{BA(n)} \rangle \right)$  (5)

157 where  $H_{ij(n)}$  is the energy of a supercell with an  $ij$  pair in a fixed  $n$ -th location within a  
158 supercell at the fixed distance,  $d_n$ , between the atoms  $i$  and  $j$ , while the average is taken

159 over all possible configurations of the other atoms in the supercell. An important property  
160 of the cluster expansion is that all correlation functions of a completely disordered  
161 (A,B)R solid solution with the composition of  $x_A = 0.5$  are equal to zero. Thus, when all  
162 quantities are measured relative to a mechanical mixture of end-members, it appears  
163 necessary to introduce a constant,  $J_0$ , the so-called zero cluster interaction, and postulate  
164 that the correlation function of the zero cluster is equal to 1.  $J_0$  is equal to the enthalpy of  
165 mixing of a completely disordered (uncorrelated) solid solution with the composition of  
166  $x_A = 0.5$ . This completely disordered state ( $J_0 \neq 0$ ,  $\xi_0 = 1$ ,  $\xi_n = 0$ ,  $n > 0$ ) represents the  
167 central point of the cluster expansion of Sanchez et al. (1984) The terms of the cluster  
168 expansion with  $n > 0$  describe deviations of the enthalpy of a real solid solution from this  
169 reference state. In this sense the end-members and ordered intermediate compounds are  
170 treated on equal grounds as different examples of extremely correlated states.

171 The J-formalism (Bosenick et al., 2000, 2001; Dove, 1999, 2001; Dove et al., 1996, 2000;  
172 Palin et al., 2001; Warren et al., 2001; Will, 1998) is often understood as a restricted form  
173 of the cluster expansion, in which the cluster size is limited to a pair of points. However,  
174 this is not exactly true. Aside from being limited to pairs, the J-formalism assumes also a  
175 different reference state. The reference state of the J-expansion is not a disordered state  
176 with  $x_A = 0.5$ , but rather a mechanical mixture of the end-members AR and BR.  
177 Consequently, a constant,  $J_0$ , term is not required in the J-formalism. The basic idea of  
178 the J-formalism is that any deviation of the enthalpy of a solid solution from the enthalpy  
179 of a mechanical mixture is correlated with the production of pairs of AB type. Therefore,  
180 in the J-expansion the J terms are coupled not to correlation functions, but to probabilities  
181 of AB pairs. The pairwise interaction is thus defined as a contribution to the excess



182 enthalpy of a solid solution due to the formation of one AB pair (or one mole of AB  
183 pairs). As the production of two AB pairs is necessarily related to an annihilation of a  
184 couple of AA and BB pairs, the pairwise interaction is defined as

$$185 \quad J_{(n)} = 2\langle H_{AB(n)} \rangle - \langle H_{AA(n)} \rangle - \langle H_{BB(n)} \rangle \quad (6)$$

186 where the quantities in brackets have the same meaning as in Eq. 5. The  $J_{(n)}$  terms are  
187 thus proportional to the effective pair interactions in the cluster expansion (de Fontaine,  
188 1994; de Fontaine et al., 1992), but have a different sign. Similarly to the cluster  
189 expansion, the J-formalism can include clusters, which are larger than pairs. Indeed, the  
190 probabilities of AAB, ABB, AAAB, AABB, ABAB and ABBB are also constrained to be  
191 identically zero in the state of a mechanical mixture and thus the relevant ternary and  
192 quaternary terms could be easily included in the J-expansion. This is not done because of  
193 practical reasons. Ternary or higher-order interactions are usually much weaker than the  
194 pair interactions. The reason for the complete exclusion of the higher order interactions in  
195 applications to minerals comes from the observation that in oxides, silicates, carbonates  
196 or phosphates the exchangeable atoms (typically represented by cations) never appear in  
197 close contact with each other, as they do in alloys. In minerals the cations are usually  
198 separated by anion groups that remain inert to the mixing. Strong specific ternary or  
199 quaternary interactions, which could appear due to an association of several cations in a  
200 close proximity of each other, become less likely. The enthalpy of mixing for a  
201 configuration  $i$  is written as

$$202 \quad H_i = 1/2 \sum_{n=1}^N J_n Z_n P_{AB(n)} \quad (7)$$

203 where  $Z_n$  is the coordination number within a sublattice occupied by neighbors of the  
204 order  $n$ . The order of neighbors,  $n$ , usually corresponds to the distance between the atoms.

205

206 **2.3. Determination of the  $J_s$ .**

207 The methods for the evaluation of the  $J$  terms are similar in the cluster expansion and in  
208 the J-formalism. The cluster inversion (Connolly and Williams, 1983; de Fontaine, 1994;  
209 de Fontaine et al., 1992) procedure is based on the assumption that the quantities defined  
210 by Eq. 5 are relatively insensitive to the number and types of configurations included in  
211 the summation, such that they could be considered as configuration and composition  
212 independent constants. Thus, one is allowed to expand the energies of several compounds  
213 differing in symmetry and composition using the same set of the cluster terms by  
214 assuming that the effective interactions are the same in all compounds. In alloys these  
215 compounds are usually selected based on experimentally observed ordered phases. In this  
216 way one obtains a set of equations, which can be solved for the effective interactions. The  
217 J-formalism is typically applied to low-symmetry lattices, where the selection of ordered  
218 compounds is difficult. Therefore, the anchor structures/configurations are often selected  
219 at random within a supercell of a certain shape. The averaged normalized numbers of AB  
220 pairs in all these structures are counted at all distances and the system of equations of the  
221 type of Eq. 7 is solved for the  $J_n$  with the least-squares method (Becker et al., 2000;  
222 Ferriss et al., 2010; Reich and Becker, 2006; Renock and Becker, 2011; Vinograd et al.,  
223 2006, 2007).

224 An alternative way of evaluation of the  $J_s$  is to consider three supercells with the  
225 specified AA, BB and AB configurations and with the rest atoms in the same  
226 configurations and directly apply Eq. 5 or 6. Keeping the AA, BB and AB arrangement  
227 fixed, the rest atoms are permuted and the calculation is repeated. The  $J$  value can be

228 computed as the averaged energy obtained by a summation over a limited set of these  
229 configurations. It has been shown that this procedure can give converged values by  
230 averaging over about 20-50 randomly chosen configurations with fixed AA, BB and AB  
231 arrangements (de Fontaine et al., 1992; Wolverton et al., 1991). The double defect  
232 method (DDM) (Asato et al., 2001; Hoshino et al., 1993, 1996; Vinograd et al., 2009;  
233 Vinograd and Winkler, 2010) is a variant of this approach, where only two unique  
234 configurations are considered, namely, the configurations with “all A” and with “all B”.  
235 In the “all A” case the AB, AA and BB pairs are inserted in the supercell of the AR end-  
236 member. As in this case the A atoms of the pair cluster are indistinguishable from the rest  
237 A atoms of the host supercell, Eq. 5 or 6 effectively requires the computation of the  
238 energy difference between two supercells with a single defect B, a supercell of the pure  
239 end-member AR and a supercell with a double BB defect with the B atoms fixed at a  
240 certain distance from each other. When the energies of all structures are counted relative  
241 to the mechanical mixture of AR and BR, end-members, the recipe can be further  
242 simplified as follows

$$243 \quad J_{A(n)} = (2\Delta H_B - \Delta H_{BB})/D_n \quad (8)$$

244 where  $\Delta H_B$  and  $\Delta H_{BB}$  are the excess enthalpies of the supercells with a single, B, or a  
245 double, BB, defect, respectively, and  $D_n$  is the degeneracy factor, which counts the  
246 multiplicity of BB pairs due to the use of periodic boundary conditions in the model  
247 calculations. An analogous equation is written for the “all B” case

$$248 \quad J_{B(n)} = (2\Delta H_A - \Delta H_{AA})/D_n \quad (9)$$

249 As the DDM naturally provides two sets of the  $J_n$  values, it is convenient to model  
250 intermediate compositions by a linear mixing of  $J_{A(n)}$  and  $J_{B(n)}$

251  $J_n = x_A J_{A(n)} + x_B J_{B(n)} \quad (10)$

252 The composition dependence appears to be an important advantage of the DDM over the  
253 standard J-formalism. The method also allows to significantly reduce the number of  
254 supercell structures to be computed. Effectively, the determination of a single pairwise  
255 interaction requires the computation of a single supercell structure with a paired defect.

256

257 **2.4. Calculation of thermodynamic quantities via direct sampling of all**  
258 **configurations.**

259 Here we assume that the total excess Gibbs free energy can be split into the  
260 configurational and vibrational parts, while the vibrational free energy is insensitive to  
261 configuration, but depends on the composition. This simplification can be justified when  
262 the excess vibrational free energy is small and its effect is expected to be significant only  
263 at high temperatures. This effect is modeled here with the aid of quasi-random structures.  
264 The configurational part is modeled with the J-formalism. First the  $J_n$  values are  
265 substituted into Eq. 7 to compute the energies of all possible configurations of a given  
266 supercell. When the energies of all configurations are known, the free energy per 1 mole  
267 of exchangeable atoms at a temperature  $T$  (K) can be computed as

268  $G = -(1/p)RT \ln Z \quad (11)$

269 where  $p$  is the number of exchangeable sites within the supercell,  $R=8.314$  J/K/mol is the  
270 gas constant and

271  $Z = \sum_i \exp(-H_i/(RT)) \quad (12)$

272 is the partition function. The summation is often taken only over symmetry non-  
273 equivalent configurations, which could be enumerated with available programs, for

274 example, SOD (Site Occupancy Disorder) (Grau-Crespo et al., 2004, 2007). Each  
275 exponential term is then counted with its multiplicity. The ensemble average enthalpy is  
276 computed as

$$277 \quad H = \frac{1}{pZ} \sum_i \exp(-H_i/(RT)) H_i \quad (13)$$

278 The configurational entropy is then calculated as

$$279 \quad S = \frac{H - G}{T} \quad (14)$$

280 It should be noted that the entropy computed with this approach would deviate from the  
281 ideal mixing entropy even in the high-temperature limit. Due to the limited size of the  
282 supercell, the maximum entropy is given by

$$283 \quad S_{\max} = (R/p) \ln \left( \frac{p}{k} \right) \quad (15)$$

284 where  $k$  is the number of atoms of type B. For a small cell this value could be  
285 significantly smaller than the entropy of ideal mixing

$$286 \quad S_{\text{ideal}} = -R(x_A \ln x_A + x_B \ln x_B) \quad (16)$$

287 For example, for a cell containing 24 sites, the maximum entropy computed with Eq. 16  
288 is about 11% smaller than the ideal mixing value. Becker et al. (2000) suggested an  
289 elegant procedure, which corrects for this drawback. In their procedure, the manifold of  
290 configurations is viewed as the probability vs. energy distribution scaled by the total  
291 number of configurations. The sum in Eq. 12 is then substituted with an integral over this  
292 distribution, while the number of configurations is taken out of the integral sign. The  
293 Gibbs free energy can be then written as

294 
$$G = -\frac{RT}{p} \ln\left(\frac{p}{k}\right) - \frac{RT}{p} \ln \frac{\int_0^{\infty} \exp(-H/(RT)) p(H) dH}{\int_0^{\infty} p(H) dH} \quad (17)$$

295 The correction consists in the replacement of the first term with the negative of the ideal  
296 entropy times  $T$

297 
$$G = RT(x_A \ln x_A + x_B \ln x_B) - \frac{RT}{p} \ln \frac{\int_0^{\infty} \exp(-H/(RT)) p(H) dH}{\int_0^{\infty} p(H) dH} \quad (18)$$

298 Eq. 18 correctly predicts the regular mixing behavior only in the high-temperature limit.  
299 The correction does not remove all drawbacks of a small-supercell model. An important  
300 problem of such a model is its inability to represent a proper relationship between the  
301 frequencies of ordered and disordered configurations. Indeed, in a small supercell the  
302 ratio of disordered to ordered configurations is inevitably underestimated. Consequently,  
303 in a small supercell short-range order has a weaker effect on the free energy of the  
304 disordered state. As the free energy of the disordered phase cannot be effectively  
305 decreased on cooling by SRO effects, it becomes less stable relative to an ordered phase  
306 (or to the mixture of end-members) at a rather high temperature. The order/disorder  
307 transition temperature is thus overestimated.

308

### 309 **2.5. Calculation of thermodynamic quantities via Monte Carlo simulations.**

310 The obvious remedy for the former method is to enlarge the supercell size. However, the  
311 alternative large-cell models have to deal with the problem of an intolerable increase in  
312 the number of configurations. There have been attempts to deal with the problem of  
313 “configurational explosion” by limiting the number of sampled configurations. A feasible  
314 approach is to select a subset of a few hundred of configurations by sampling them at

315 random (Allan et al., 2001, 2006; Purton et al., 2006; Todorov et al., 2004). This  
316 approach, however, often predicts free energies, which only marginally deviate from the  
317 regular mixing model. Configurations sampled in this way typically represent the largest  
318 pool of highly probable high-energy configurations. These configurations have similar  
319 (high) energies and similar Boltzmann weights. Therefore, configurational entropy  
320 computed from such a reduced set appears to be close to that of the ideal mixing model,  
321 while the configurational enthalpy closely follows the Margules model. A similar result  
322 can be more easily obtained with a quasi-random structure. Recently, D'Arco et al.  
323 (2013) offered a modified algorithm, in which the subset of states is obtained by random  
324 sampling over classes of symmetry independent configurations. The applicability of this  
325 algorithm to large supercells and thus for accurate computation of phase diagrams is still  
326 to be demonstrated.

327 The importance sampling via a Monte Carlo algorithm (Metropolis et al., 1953) can be  
328 conveniently applied to much larger supercells than those typically employed in direct  
329 methods. The average enthalpy of the sampled set energy is correctly shifted to a lower  
330 value, which reflects higher weights of low energy configurations at a finite temperature.  
331 The configurational enthalpy computed with a supercell containing few thousands of  
332 exchangeable sites is essentially correct. However, the entropy value computed from such  
333 a limited subset would still be close to that of the ideal mixing model as the energies of  
334 the converged set of configurations fluctuate only slightly from this average (although  
335 correct) value of the enthalpy. The evaluation of the correct entropy value requires  
336 computation of the partition function or the Gibbs free energy. While a direct evaluation  
337 of the former is difficult, the equilibrium Gibbs free energy can be computed relatively

338 easily using thermodynamic relations under the assumption that the average enthalpy  
339 computed by the Monte Carlo method is correct. The relationship between the free  
340 energy and the enthalpy is given by the integral (Dove, 2001; Myers, 1998; Warren et al.,  
341 2001):

$$342 \quad G = RT(x_A \ln x_A + x_B \ln x_B) + H_0 + \int_0^1 \langle H \rangle_\lambda d\lambda \quad (19)$$

343 where  $\lambda$ ,  $0 < \lambda < 1$ , is an artificial variable used to scale the magnitude of the effective  
344 pairwise interactions down, and  $\langle H \rangle_\lambda$  is the enthalpy averaged over an ensemble, in which  
345 the probability distribution of all configurations corresponds to the effective pairwise  
346 interactions scaled by  $\lambda$  value. In the application to the J-formalism, the evaluation of  
347  $\langle H \rangle_\lambda$  requires a Monte Carlo simulation of a system, in which the pairwise interactions are  
348 scaled by  $\lambda$ . It should be remembered the converged enthalpies resulting from the  
349 simulation should be rescaled back (i.e. divided by the same  $\lambda$  value). The first two terms  
350 in Eq. 19 correspond to the Gibbs free energy of the solid solution in the state of a  
351 complete disorder, which is known theoretically, while the integral describes the  
352 deviation of the real free energy at the given temperature from the reference state.

353 The power of the integration method is based on the ability of the Metropolis algorithm  
354 to arrive at the correct average (equilibrium) energy without the need of sampling all  
355 possible states. Consequently, the method can be applied to a sufficiently large cell, such  
356 that the averages are indistinguishable from the case of the thermodynamic limit. When  
357 the Markov chain of Monte Carlo stated converges to the Boltzmann distribution, each  
358 symmetry independent configuration occurs with the correct probability. Thus the correct  
359 average enthalpy can be computed by sampling a sufficient number of successive states.  
360 For example, the computation of an enthalpy value at a high temperature does not require



361 sampling of the ground state configuration. The probability of this configuration at a high  
362 temperature is vanishingly low and its contribution to the average energy is similarly  
363 small. This configuration, however, becomes very important at a low temperature, thus,  
364 as noted, for example, by d'Arco et al., 2013, there is a possibility that such a rare  
365 configuration could be missed in a simulation. The art of the simulation thus consists in  
366 gradually decreasing the temperature and in restarting from previously equilibrated  
367 configurations. Our experience suggests that the finding of the correct ground state is  
368 always possible in supercells containing less than about 4000 sites. There are different  
369 means that help to ensure that the correct state is found. One possibility is to check that  
370 the same final enthalpy is achieved in runs that start from different initial configurations.  
371 In the present case, where the ground state is represented by a miscibility gap, there  
372 appears a possibility to observe fine features within the isotherm shape, that are related to  
373 changes in the form of the interface between the two phases. Observing these fine  
374 features, in fact, means that true equilibration has been achieved.

375

376

### 3. COMPUTATIONAL DETAILS

377 The pairwise interactions were computed from the excess energies of  $2 \times 2 \times 4$  supercells  
378 of rutile structure ( $\text{TiO}_2$  or  $\text{SnO}_2$ ) containing substitutional defects of Ti and Ti-Ti in the  
379 case  $\text{SnO}_2$  being the host or Sn and Sn-Sn defects in the case of  $\text{TiO}_2$ . The paired defects  
380 were placed at 10 different distances in the range of 2-9 Å. The size of the supercell was  
381 chosen such that the computed interaction between the pair of defects placed at the  
382 longest distance practically vanishes. The total energies of the single- and double-defect  
383 structures were calculated with density functional theory in the Wu-Cohen (WC GGA)

384 approximation (Wu and Cohen, 2006) using the “on-the-fly” ultrasoft pseudopotentials  
385 integrated in the CASTEP distribution, version 7.03 (Clark et al., 2005; Segall et al.,  
386 2002). The plane wave expansion cutoff was 900 eV. The k-space was sampled using a 3  
387  $\times 3 \times 3$  Monkhorst-Pack grid.

388 Although Eqs. 8 and 9 could be directly applied for computing the pairwise interactions,  
389 the resulting values of the  $J_s$  often lack self-consistency (Vinograd et al., 2009; Vinograd  
390 and Winkler, 2010). The requirement of self-consistency is based on the observation that  
391 although the  $J_s$  are directly determined by the excess energies of the defect structures via  
392 Eqs. 8 and 9, the excess energies themselves are functions of the  $J_s$  via Eq. 7. This  
393 implies that the excess energies of the single- and double-defect structures should be  
394 exactly reproduced via Eq. 7. This requirement is often violated due to slight  
395 inaccuracies in the computed excess energies of the single- and double-defect structures  
396 or due to the presence of higher-order interactions. The self-consistency can be restored  
397 by a slight variation in the excess energies of the single defect structures, which occur in  
398 Eqs. 8 and 9. Practically, small parameters  $\delta$  and  $\delta$  are added to  $\Delta H_A$  and  $\Delta H_B$   
399 respectively, and their values are varied until the excess energies of the double-defect  
400 structures are reproduced exactly via Eqs. 7-10.

401 Monte Carlo simulations were performed with an  $8 \times 12 \times 16$  supercell containing 3072  
402 exchangeable atoms. The size of the supercell was chosen based on previous experience.  
403 It can be shown that the critical temperature in the well-studied two-dimensional square  
404 lattice with nearest-neighbor interactions [Onsager, 1944] can be reproduced with a  
405 relative accuracy of better than 5% using supercell containing about 2500 lattice sites.  
406 However, to our experience, a significant increase over this number could cause

407 difficulties in finding correct ground states. Our previous experience (Jung et al., 2010;  
408 Vinograd et al., 2007) with various three-dimensional lattices suggest that the optimal  
409 supercell should contain about 3000 or 4000 sites. The average enthalpy at a given  
410 composition and temperature was calculated assuming a canonical ensemble using the  
411 Metropolis algorithm (Metropolis et al., 1953). At each simulation step a pair of different  
412 atoms was chosen at random and the swap was attempted. The swap was either accepted  
413 or rejected depending on the enthalpy change due to the swap,  $\Delta H$ . A new configuration  
414 was accepted for  $\Delta H \leq 0$  or with the probability of  $\exp(-\Delta H/k_B T)$  for  $\Delta H > 0$ . The  
415 distribution was assumed to converge to the Boltzmann probability distribution after  
416  $2 \times 10^7$  steps. The next  $2 \times 10^7$  steps were used to calculate the averages. The simulations  
417 were performed with the  $J$  values computed with Eq. 10, the temperature was varied  
418 between 600 K and 2500 K with an interval of 100 K and the composition was varied  
419 with a step of 0.03125. For each temperature and composition the calculations were  
420 repeated 26 times at different  $\lambda$  values, where  $\lambda$  was varied from 0 to 1 with a step of  
421 0.04.

422 The excess vibrational entropy of mixing was calculated with an empirical force-field  
423 model. The interatomic potentials were fitted to the structure data of rutile (Swope et al.,  
424 1995), brookite (Baur, 1961), anatase (Horn et al., 1972) and cassiterite (Baur, 1956), and  
425 to the elastic stiffness tensors of rutile and cassiterite (Ahrens, 1995). The parameters are  
426 listed in Table 1. The unit-cell parameters and the elastic stiffness constants of rutile and  
427 cassiterite calculated with these potentials are given in Table 2. The vibrational free  
428 energies of the end-member structures and of the quasi-random structures (QRS) were  
429 computed with the GULP program (Gale, 1997, 2005; Gale and Rohl, 2003). using the

430 zero static internal stress approximation (ZSISA) (Allan et al., 1996). The force-field  
431 model was tested against the *ab initio* data (Sikora, 2005) for pure rutile. The  
432 thermodynamic properties (entropy and heat capacity at constant volume) of rutile  
433 deviate over the temperature interval of 300-2500 K by less than 3% from the DFT  
434 results. The effective excess vibrational entropy of the QRS was computed as the  
435 negative of the excess vibrational free energy divided by  $T$ . The entropies were averaged  
436 in the temperature range of 1473-1773 K and fitted to the equation

$$437 \quad S_{\text{vib}} = x_{\text{A}}x_{\text{B}}(x_{\text{A}}W_{12}^s + x_{\text{B}}W_{21}^s) \quad (20)$$

438 where  $x_{\text{A}}$  and  $x_{\text{B}}$  are mole fractions of rutile and cassiterite in the solid solution,  
439 respectively. The contribution of the excess vibrational free energy to the free energy of  
440 mixing was computed as  $-TS_{\text{vib}}$ .

441 The quasi-random structures ( $x_{\text{Sn}}=0.25$ ,  $x_{\text{Sn}}=0.5$  and  $x_{\text{Sn}}=0.75$ ) were found by a direct  
442 search among all possible configurations within a  $2 \times 2 \times 4$  supercell. During the search  
443 the frequencies of Ti-Sn pairs at 10 different distances were compared to their theoretical  
444 probabilities. The misfit function was defined as the total squared difference between the  
445 actual and the theoretical frequencies. Structures (configurations) corresponding to the  
446 minimum misfit were selected.

447

## 448 **4. RESULTS**

### 449 **4.1. Quasi-random and single-defect structures.**

450 The excess enthalpies of the structures with single defects (0.3899 eV per supercell for Ti  
451 in  $\text{SnO}_2$  and 0.4056 eV per supercell for Sn in  $\text{TiO}_2$ ) provide constraints for the  
452 subregular model of the enthalpy of mixing

453  $H_{\infty} = x_A x_B (x_A W_{12}^h + x_B W_{21}^h)$  (21)

454 The Margules parameters  $W_{12}^h = 39.13$  kJ/mol and  $W_{21}^h = 37.62$  kJ/mol are  
455 straightforwardly computed by applying the conversion factor of 96.485 kJ/mol/eV to the  
456 excess enthalpies of the single-defect structures (Vinograd et al., 2013). The slightly  
457 larger value of  $W_{12}^h$  reflects the increased difficulty of inserting a larger cation (Sn) into a  
458 structure with a smaller mole volume (TiO<sub>2</sub>). The excess energies of the quasi-random  
459 structures are similar to that obtained from Eq. 21 (Fig. 1), because the structures with  
460 single defects comply with the criteria of QRS. This consistency between the excess  
461 enthalpies confirms the good quality of the QRSs. These structures can hence be used for  
462 modelling the properties of a completely disordered solution. The same conclusion  
463 applies to the single-defect structures as well.

464 The regular model of the enthalpy of mixing,  $(W_{12}^h + W_{21}^h)/2 = 38.4$  kJ/mol,  
465 straightforwardly predicts the critical temperature of the phase separation at  $x_{\text{Sn}} = 0.5$  as  
466  $T_c = W/2R$ , i.e. 2309 K (Glynn, 2000; Park et al., 1975; Urusov et al., 1996). This value  
467 significantly overestimates the experimental temperatures of 1703±5 K (Park et al., 1975)  
468 and 1689 K (Naidu and Virkar, 1998), showing that the regular (or subregular) model is  
469 not adequate for this system. Within the regular model description this temperature could  
470 be brought into correspondence with the experiment only under the assumption that the  
471 excess vibrational entropy constitutes about 25% of the configurational entropy. As we  
472 will show below, the vibrational contribution appears to be much smaller. A more  
473 adequate model requires consideration of short-range order (SRO) effects. Indeed, SRO  
474 effects allow a significant stabilization of the disordered phase, which leads to a decrease  
475 in the transition temperature. SRO is primarily driven by pairwise interaction energies.

476

477 **4.2. Pairwise interactions.**

478 Fluctuations in excess energies of the paired defects Ti-Ti in rutile and Sn-Sn in  
479 cassiterite reflect variations in the defect-defect interaction at different distances (Table  
480 3). As these interactions should decrease with the increase of the distance between the  
481 defects, the excess energies of the defect-defect structures are expected to converge to  
482 twice the excess energy of the single-defect structure. Simultaneously, the effective  
483 interactions should converge to zero value (Fig. 2). This behavior is indeed observed.

484 The small final values of the  $\delta_1$  and  $\delta_2$  parameters, -0.000081 and 0.000011 eV,  
485 respectively, show that the computed excess energies of the single- and double-defect  
486 structures are very close to internal consistency. The small deviations could be caused by  
487 the presence of interactions of higher order although these interactions must be very  
488 weak. The  $J$  values determined from this fit are listed in Table 3. Fast convergence of the  
489  $J$  values as a function of interatomic separation (Fig. 2) implies the applicability of these  
490 parameters for the calculation of the excess enthalpies using much larger supercell.

491

492 **4.3. Direct method with a  $2 \times 2 \times 4$  supercell.**

493 All configurations were permuted with a self-written C program and their excess energies  
494 were evaluated via Eq. 7. Figure 3a, 4a and 5a show the results of the excess enthalpy,  
495 the free energy and the configurational entropy computed via Eqs. 13, 11 and 14,  
496 respectively.

497 The wavy shapes of the low temperature isotherms reflect ordering, which would have  
498 occurred in the system if the phase separation were forbidden (Fig. 3a, 5a,b). The minima

499 correspond to structures with alternating layers fully filled with Sn or Ti atoms, which are  
500 parallel to (100) or (010). Figure 4a shows that the phase separation at intermediate  
501 compositions occurs at about 2000 K. Thus the ordered states are never stable with  
502 respect to a mixture of Ti- and Sn-rich phases. Figure 5a shows also that at high  
503 temperatures the configurational entropy converges to values, which are significantly  
504 smaller relative to the ideal mixing model. This drawback of the  $2 \times 2 \times 4$  model leads to  
505 a significant overestimation of the free energy of the disordered phase at high  
506 temperatures and, consequently, to an overestimation of the solvus temperature.

507

#### 508 **4.4. Direct method with a $2 \times 2 \times 4$ supercell and with the entropy correction.**

509 In the method of Becker et al. (2000) the free energy is evaluated via Eq. 18. This leads  
510 to a significantly more negative free energy of mixing at high temperatures and,  
511 consequently, to a lower temperature of the solvus closure (Fig. 4b). The corrected  
512 configurational entropy becomes consistent with the high-temperature limit of the ideal  
513 mixing model (Fig. 5b). However, the low-temperature entropy isotherms shift to larger  
514 values.

515

#### 516 **4.5. Monte Carlo simulation using an $8 \times 12 \times 16$ supercell.**

517 The enthalpy of mixing, the free energy and the configurational entropy resulting of these  
518 simulations are given in Figure 3b, 4c and 5c, respectively. The shapes of low-  
519 temperature isotherms differ significantly from those computed with the direct method.  
520 This is the consequence of the phase separation (Fig. 6), which occurred during the  
521 Monte Carlo simulation runs. In contrast, in the direct method the phase separation does

522 not occur within the small supercells and thus does not affect the shape of the isotherms.  
523 The direct method includes configurations, which formally could be interpreted to  
524 represent phase separation. However, due to the relatively large interface energy, the  
525 energies of these configurations appear to be higher than the corresponding ordered  
526 structures and thus do not contribute much to the shapes of low-temperature isotherms.

527

#### 528 **4.6. Vibrational entropy correction and the phase diagram.**

529 Figures 3-5 correspond to the model, which is based on static energies only. However,  
530 each configuration is also characterized by its own vibrational entropy and free energy.  
531 Our calculations based on quasi-random structures suggest that the vibrational entropy  
532 adds about 10% to the total entropy of mixing at high temperatures and intermediate  
533 compositions (Fig. 7). The curve of the vibrational entropy is asymmetric ( $W_{12}^s=1.66$ ,  
534  $W_{21}^s=3.02$  J/K/mol). The sense of asymmetry is similar to that of the enthalpy of mixing.  
535 The phase diagram of rutile-cassiterite solid solution can be computed from the free  
536 energy isotherms via common tangent analysis. The phase relations derived from the  
537 different methods are shown in Figure 8 together with the experimental data (Naidu and  
538 Virkar, 1998; Park et al., 1975). It is shown that the addition of the  $-TS_{\text{vib}}$  term  
539 significantly stabilizes the disordered phase, causing a shift of the solvus line to lower  
540 temperatures by about 250 K. The solvus line calculated based on Monte Carlo  
541 simulations shows a critical temperature of 1800 K, which agrees well with the  
542 experimental data,  $1703\pm 5$  K reported by Park et al. (1975) and 1689 K reported by  
543 Naidu and Virkar (1998).

544



545

## 5. DISCUSSION

546 The results obtained with the Monte Carlo method differ noticeably from those obtained  
547 via a direct calculation of the partition function. First of all, the comparison of Figures 3a  
548 and 3b shows that the enthalpy of the disordered phase in the Monte Carlo model  
549 decreases more rapidly. This is the result of SRO, which appears to be more important in  
550 the simulations involving a realistically large supercell. The stabilizing effect of SRO  
551 seems to be the main cause of the smaller value of the critical temperature predicted with  
552 the Monte Carlo simulations. In the disordered limit the enthalpies of mixing calculated  
553 with the both methods are similar. Secondly, the low-temperature isotherms in Figures 5b  
554 and 5c are strikingly different. This difference can be easily explained. It is important to  
555 note that direct calculations apply to a much smaller supercell, in which phase separation  
556 typically cannot occur. Indeed, the phase separation implies the contribution of an  
557 interface energy. As the formation of such an interface in a small cell is much more  
558 unfavorable than in a large cell, the energy of such a cell appears to be much higher than  
559 that of a mixture of two phases. Consequently, configurations with the occurrence of such  
560 interfaces contribute very little to the total free energy and the entropy. On the other  
561 hand, in Monte Carlo simulations, where a much larger supercell is employed, the phase  
562 separation occurs. In supercells with approximately intermediate composition the  
563 interface is parallel to (100) or (010) (Fig. 6b). These are the same planes across which  
564 the compositional fluctuation occurs in the ordered layered structures. This shows that the  
565 same interatomic interactions are responsible for both the ordering and the phase  
566 separation effects. A phase separated structure can be viewed as an ordered structure  
567 composed of layers of infinitely large thickness.

568 The free energy of a system with a composition falling within a miscibility gap is  
569 necessarily a combination of the free energies of the two phases. Ideally, this  
570 combination should be represented by a straight line connecting the properties of the two  
571 phases. Such straight lines are indeed observed in the Figures 3b, 4c, 5c, however, they  
572 occur within a rather narrow intervals at intermediate compositions. The common tangent  
573 analysis predicts the phase separation over much wider composition ranges. Small  
574 inflections in the isotherms within the miscibility gap are related to changes in the shape  
575 of the interface boundary. At intermediate compositions the interface (i.e. the surface  
576 which separates phases (slabs) with different compositions) is represented by two parallel  
577 planes (Fig. 6b), while at more diluted compositions the interface is a closed curved  
578 surface. A similar observations were made by Vinograd and Winkler (2010) when  
579 modelling the system NaCl-KCl. Indeed, when the amount of the solute phase is small,  
580 there are not enough atoms to make a continuous slab throughout the supercell. As the  
581 amount of the second phase increases, the interface area grows until a slab is formed, and  
582 further becomes constant. Transitions between these regimes are marked with breaks in  
583 the isotherms. It is important to note that in an ideal case of an infinitely large supercell,  
584 the interface energy would not have any influence on the thermodynamic functions. This  
585 implies that although a reasonably large supercell was employed in our Monte Carlo  
586 simulations, our results still correspond to a constrained equilibrium, because the  
587 interface energy still matters. Due to the presence of the interface, the enthalpy of the  
588 Monte Carlo simulated two-phase assemblage is about 1 kJ/mol higher than this expected  
589 for a mechanical mixture.

590 The metastability that is caused by a formation of an interface or by an inability of a cell  
591 to phase separate can be removed by a post-simulation common tangent analysis of the  
592 shape of the predicted free energy surface. The phase diagram shows that the Monte  
593 Carlo simulations agree quite well with the experimental data, while the direct  
594 calculations based on a  $2 \times 2 \times 4$  supercell considerably overestimate the size of the  
595 miscibility gap. The latter is certainly related to the underestimation of the  
596 configurational entropy of the disordered phase. The method of Becker et al. (2000)  
597 improves the result significantly by decreasing the solvus temperature by 100-200 K.  
598 However, the solvus predicted with the direct method still occurs about 100-200 K higher  
599 relative to the Monte Carlo result. This effect is likely caused by an underestimation of  
600 the relative importance of high-energy configurations within a small cell. Consequently,  
601 SRO effects develop to a lesser extent than in a larger cell leading to an overestimation of  
602 the free energy of the disordered phase.

603

604

## 6. IMPLICATIONS

605 Currently, two groups of methods offer feasible strategies for the calculation of  
606 thermodynamic mixing properties of solid solutions, the direct methods, and the cluster  
607 expansion based Monte Carlo methods. The first group of methods is attractive due to the  
608 very straightforward computational algorithm, i.e. the direct evaluation of the partition  
609 function, and due to a similarly straightforward possibility of evaluation of functions  
610 other than the enthalpy and the Gibbs free energy. It appears, however, that direct  
611 computation could be performed for rather small supercells only. It also appears that a  $2$   
612  $\times 2 \times 4$  supercell (32 exchangeable atoms) is still too small to model correctly the

613 configurational statistics. The correction procedure proposed by Becker et al. (2000) can  
614 significantly improve the description of the disordered phase. Monte Carlo simulations  
615 provide an alternative indirect approach to the evaluation of the Gibbs free energy, and  
616 since a much bigger supercell can be involved in the calculation, the configurational  
617 statistics can be treated more accurately. The present study compared the free energies in  
618 the rutile-cassiterite system computed with the direct sampling and with the Monte Carlo  
619 methods. The same the cluster expansion model was used in the both approaches. This  
620 exercise suggested that the Monte Carlo simulations provide a superior description of the  
621 free energy of mixing. Thus, in cases when the enthalpy of mixing cannot be computed  
622 directly, and where an expansion (parameterization) of the enthalpy should be employed,  
623 Monte Carlo simulations seem to offer the best strategy. We note, however, that even an  
624  $8 \times 12 \times 16$  supercell does not allow achieving a fully adequate description of the phase  
625 separation. The interface energy still matters and affects the shape of free energy  
626 isotherms. A small metastability caused by the presence of the interface can be removed  
627 by a post-simulation analysis of the curvature of the free energy surface. Our study also  
628 shows that including the effect of excess vibrational entropy is important for bringing the  
629 results of simulation in an agreement with experiments. In conclusion, we wish to note  
630 that the currently available methods of atomistic simulations already provide means for  
631 accurate computation of phase diagrams in oxide systems. The present ability to compute  
632 an equilibrium content of a solute component in a mineral at a given temperature and  
633 pressure offers the possibility of developing geothermometers and geobarometers fully ab  
634 initio.  
635

636

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869 **Figure 1.** Enthalpy of mixing in the rutile-cassitertite system at an infinitely high  
870 temperature. The solid line is a subregular (two-parameter) fit to the excess energies of  
871 the structures with single defects.

872

873 **Figure 2.** Pairwise interactions computed from the excess enthalpies of the structures  
874 with single and paired defects.

875

876 **Figure 3. (a)** Isotherms of enthalpy of mixing in the rutile-cassitertite solid solution  
877 computed with the direct method (DM) using a  $2 \times 2 \times 4$  supercell. The enthalpies of the  
878 individual configurations are computed with Eq. 7. **(b)** Isotherms of the enthalpy of  
879 mixing computed with Monte Carlo simulations using an  $8 \times 12 \times 16$  supercell.  
880 Isotherms are 100 K apart.

881

882 **Figure 4.** Isotherms of Gibbs free energy of mixing in the rutile-cassitertite solid  
883 solution, **(a)** computed with the direct method using a  $2 \times 2 \times 4$  supercell, **(b)** computed  
884 with the direct method with the correction after Becker et al. (2000) using a  $2 \times 2 \times 4$   
885 supercell, **(c)** computed with Monte Carlo simulations using an  $8 \times 12 \times 16$  supercell.  
886 Isotherms are 100 K apart.

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889 **Figure 5.** Isotherms of the configurational entropy of mixing in the rutile-cassitertite  
890 solid solution, **(a)** computed with the direct method (DM) using a  $2 \times 2 \times 4$  supercell, **(b)**  
891 computed with the direct method with the correction after Becker et al. (2000) using a  $2 \times$



892  $2 \times 4$  supercell, (c) computed with Monte Carlo simulations using an  $8 \times 12 \times 16$   
893 supercell. Isotherms are 100 K apart.

894

895 **Figure 6.** Phase separation occurred during the Monte Carlo simulation runs for different  
896 compositions simulated at 600 K, (a)  $x_{\text{Sn}}=3/32$  and (b)  $x_{\text{Sn}}=16/32$ . Blue and brown balls  
897 represent Ti and Sn atoms, respectively. O atoms are not shown.

898

899 **Figure 7.** Excess vibrational entropy computed from the force-field model. The vertical  
900 bars reflect the variation of the values within the temperature interval of 1473-1773 K.

901

902 **Figure 8.** The subsolidus phase diagram of the rutile-cassiterite system. Dash lines and  
903 solid lines are the results of the simulations before and after vibrational entropy  
904 correction, respectively. Squares and the black curve denote the experimental data.

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912 **Table 1.** Parameters of the interatomic potentials for rutile and cassiterite.

Buckingham potential	$A$ , eV	$\rho$ , Å	$C$ , eV * Å <sup>6</sup>	cutoff, Å
Ti core - O shell	84045.732	0.159447	0.000	12.0
Sn core - O shell	376724.00	0.148615	0.000	12.0
O shell - O shell	1294.9157	0.317190	27.961	12.0
Charge and Spring potential	core	shell	$K_2$ , eV/Å <sup>2</sup>	
Ti	2.239195			
Sn	2.239195			
O	0.517015	-1.636611	25.830181	

913

914 **Table 2.** Structural and elastic properties of TiO<sub>2</sub> and SnO<sub>2</sub>: experimental data and results  
 915 of the force-field calculation.

	TiO <sub>2</sub>		SnO <sub>2</sub>	
	experiment <sup>a</sup>	calculation	experiment <sup>b</sup>	calculation
$a$ , Å	4.592	4.569	4.737	4.755
$c$ , Å	2.957	2.994	3.185	3.160
$V$ , Å <sup>3</sup>	62.366	62.497	71.468	71.450
$C_{11}$ , GPa	269.0	306.8	261.7	285.9
$C_{12}$ , GPa	177.0	175.2	177.2	199.8
$C_{13}$ , GPa	146.0	147.4	155.5	151.7
$C_{33}$ , GPa	480.0	462.5	449.6	427.7
$C_{44}$ , GPa	124.0	125.1	103.1	103.4
$C_{66}$ , GPa	192.0	179.1	207.4	198.0

916  $a$ : Structure properties are from Swope et al. (1995) and elastic properties are from  
 917 Ahrens (1995).

918  $b$ : Structure properties are from Baur (1956) and elastic properties are from Ahrens  
 919 (1995)..

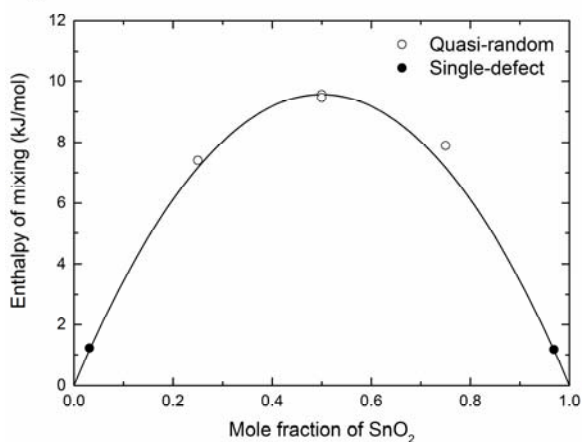
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921 **Table 3.** Excess enthalpies of double defects and the  $J$  values.

$n$	Distance, Å	$D_n$	$\Delta H$ , eV per supercell		$J$ values, kJ/mol	
			Ti-Ti defects in SnO <sub>2</sub>	Sn-Sn defects in TiO <sub>2</sub>	$J_{Ti}$	$J_{Sn}$
1	2.959	1	0.6816	0.7286	7.4613	9.5523
2	3.569	1	0.7616	0.7855	1.9713	1.8335
3	4.593	2	0.6964	0.7122	4.5218	4.0621
4	5.464	2	0.7428	0.7496	2.7175	1.8237
5	5.500	1	0.7681	0.7854	1.9809	1.2063
6	5.918	2	0.8064	0.9465	-6.7814	-1.2445
7	6.495	4	0.8003	0.8377	-0.7663	-0.4751
8	7.138	4	0.7801	0.8247	-0.4527	0.0121
9	7.491	4	0.7545	0.7508	1.3298	0.6296
10	8.787	8	0.7790	0.7938	0.1463	0.0193

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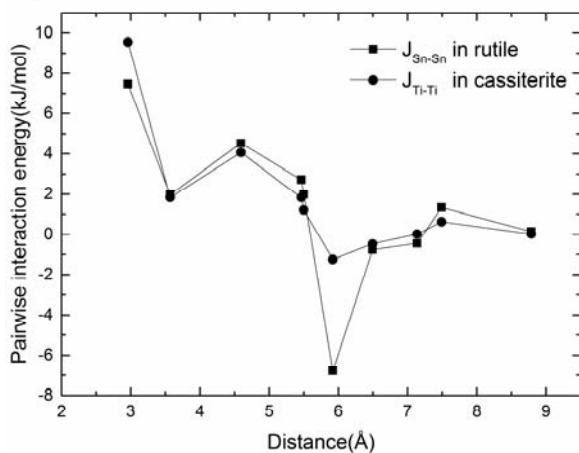
Figure 1



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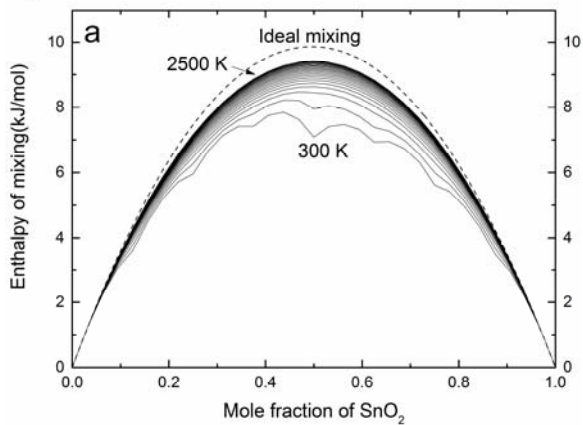
Figure 2



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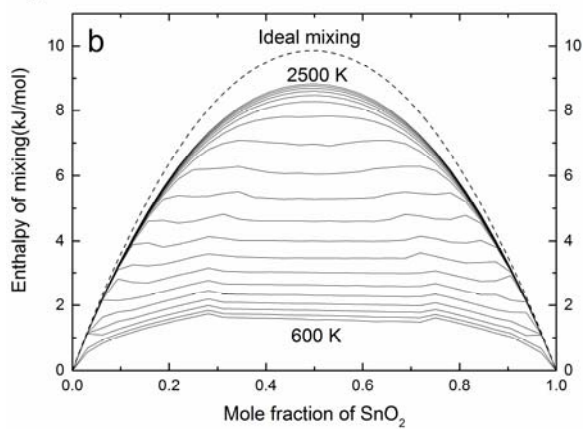
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Figure 3a



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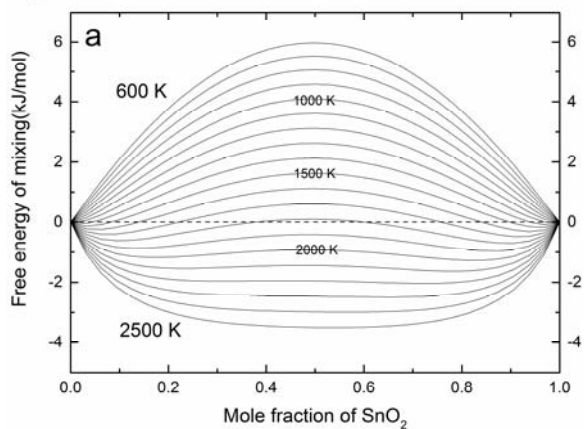
Figure 3b



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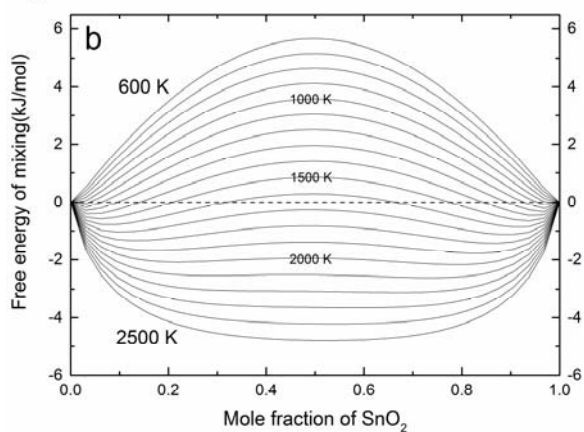
Figure 4a



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Figure 4b

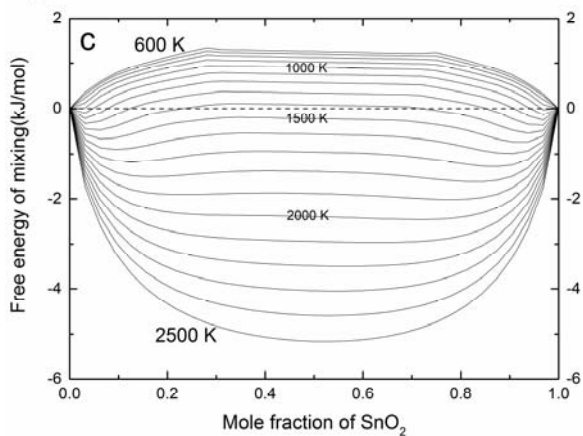


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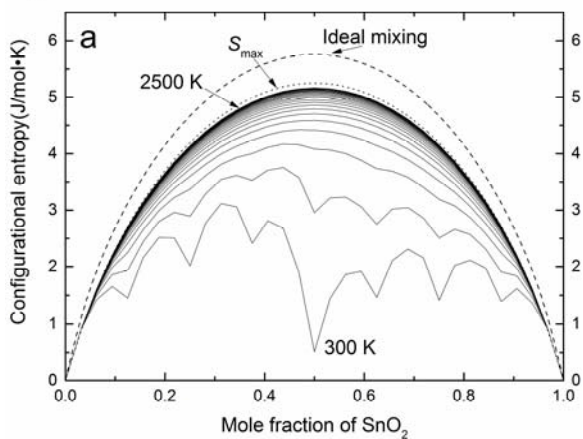
Figure 4c



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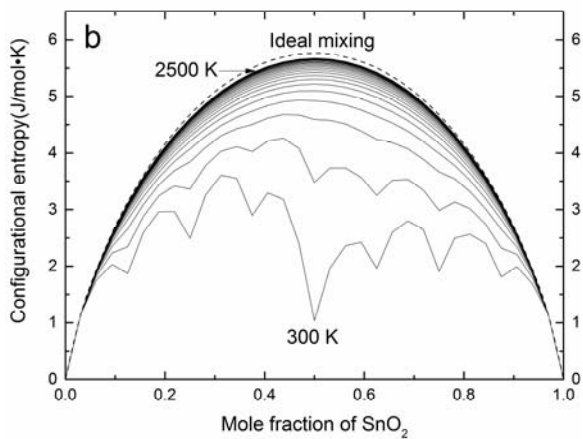
Figure 5a



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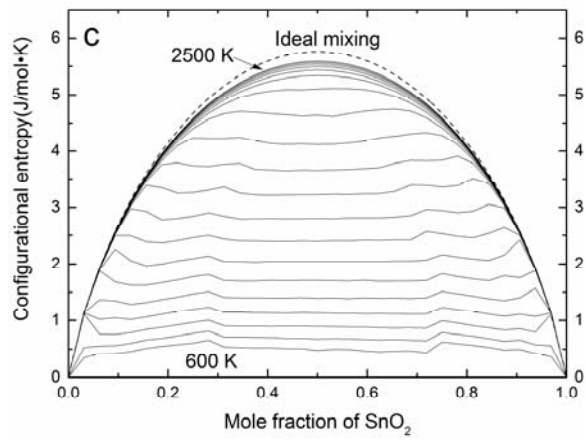
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Figure 5b



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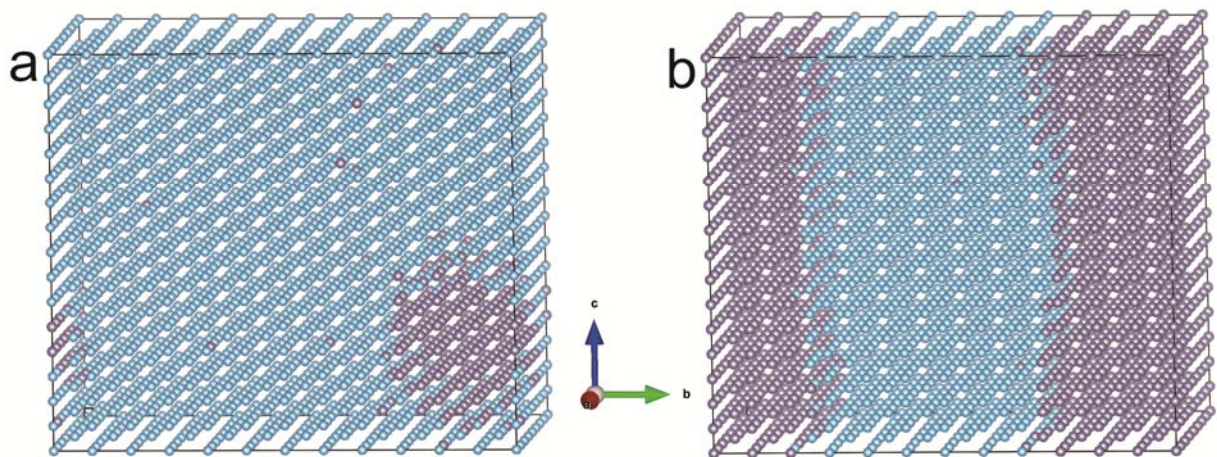
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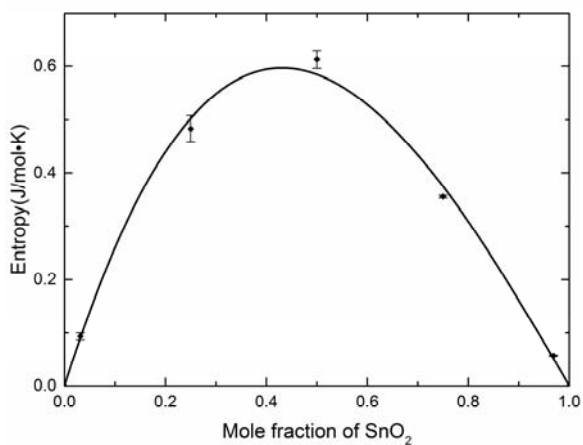
Figure 6



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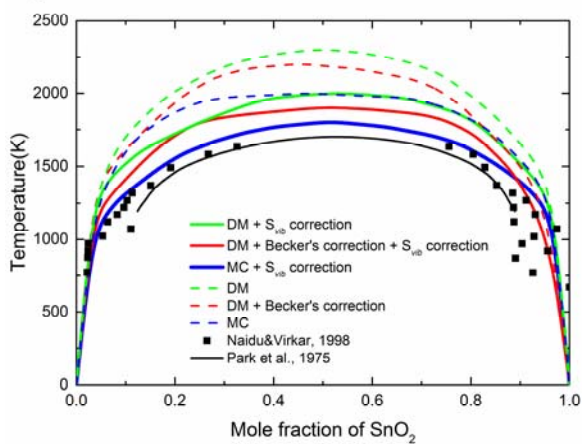
Figure 7



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Figure 8



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Figure 1

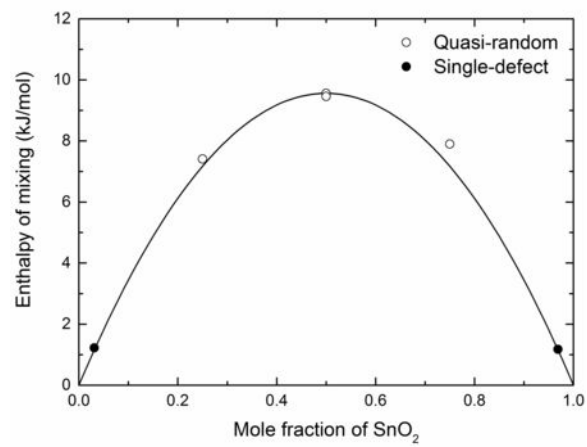




Figure 2

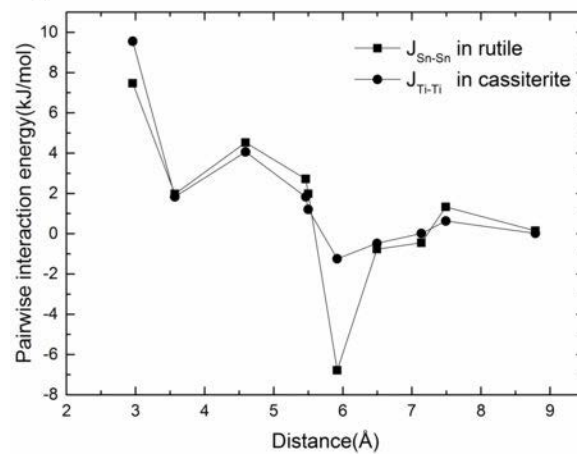


Figure 3a

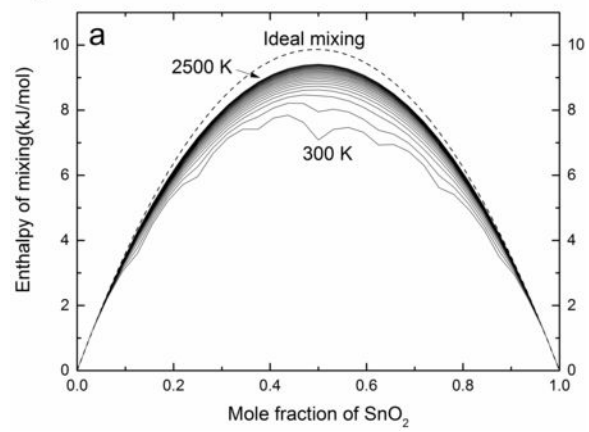


Figure 3b

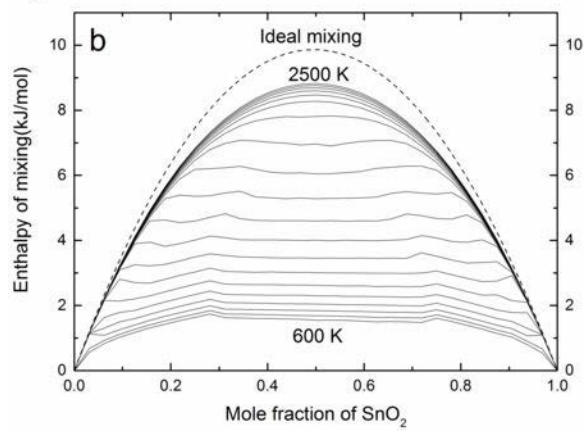


Figure 4a

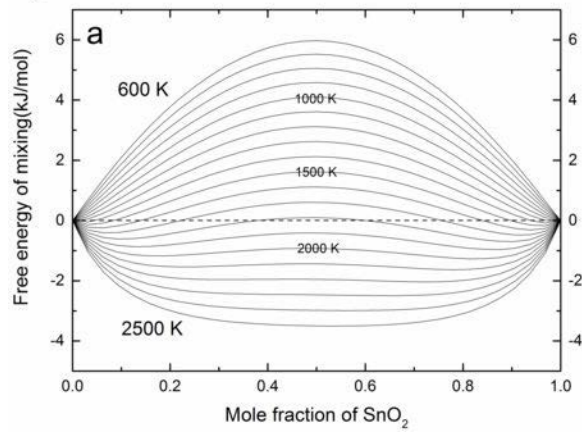


Figure 4b

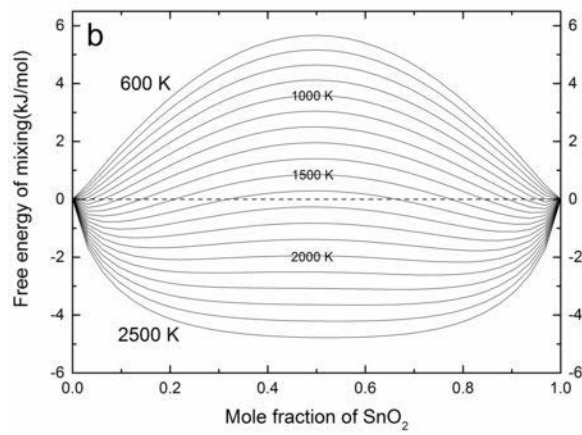


Figure 4c

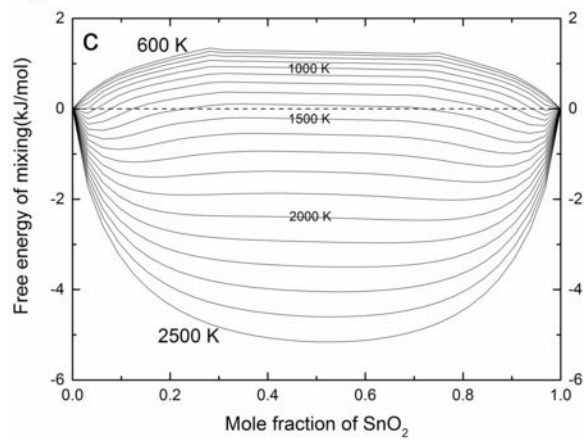


Figure 5a

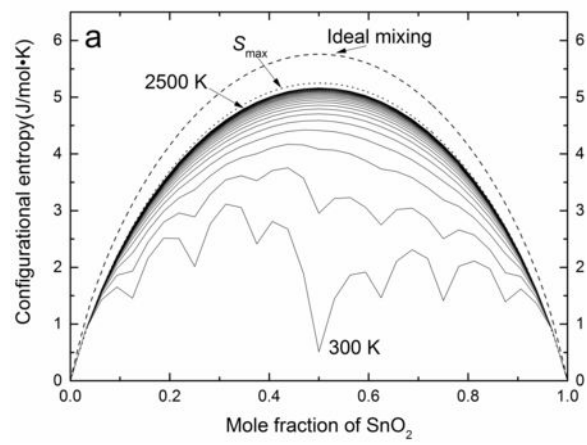


Figure 5b

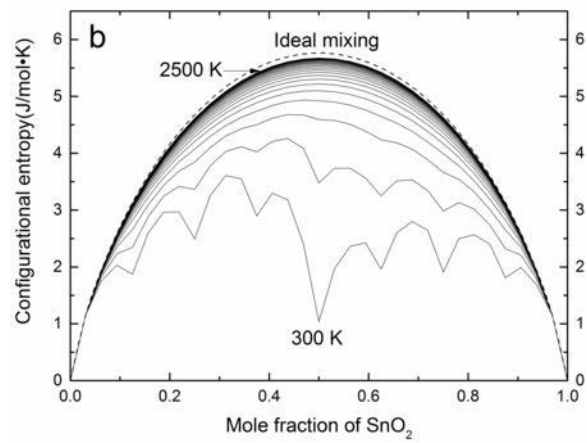




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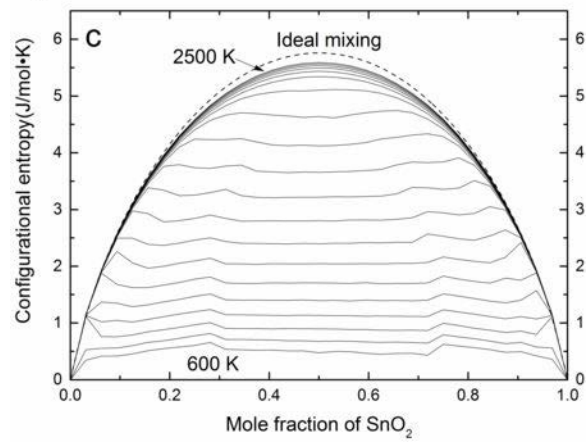
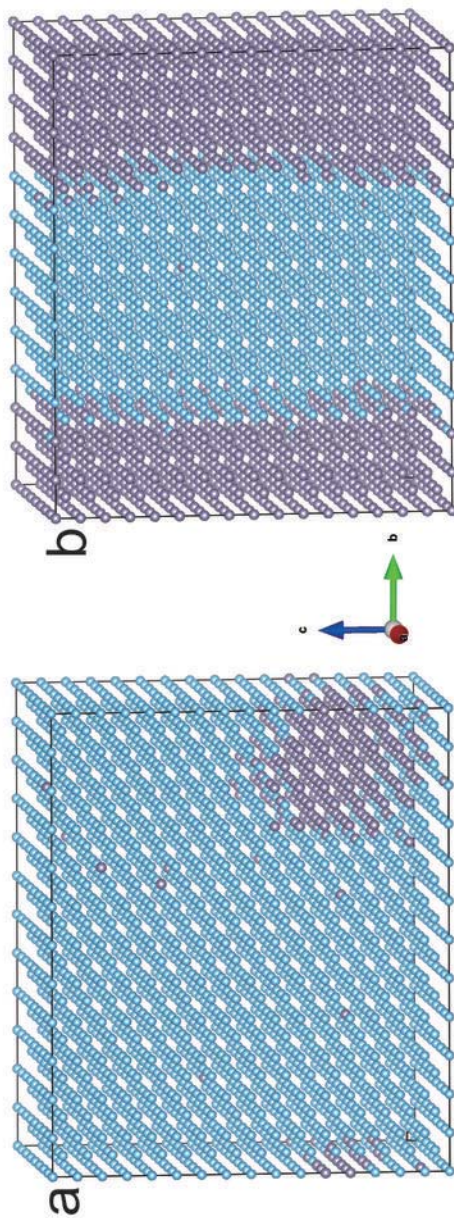


Figure 6



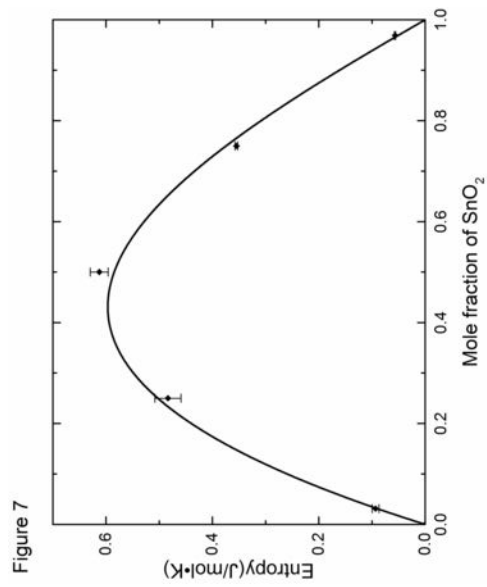


Figure 8

