1 **REVISION 2**

Ab initio investigation of majorite and pyrope garnets: Lattice 2 dynamics and vibrational spectra 3 4 MARCO DE LA PIERRE^{1,*} AND DONATO BELMONTE² 5 6 ¹ Nanochemistry Research Institute, Curtin Institute for Computation, Department of Chemistry, 7 Curtin University, GPO Box U1987, Perth WA 6845, Australia 8 9 ² DISTAV, Università di Genova, Corso Europa 26, 16132 Genova, Italy 10 11 12 ABSTRACT 13 A detailed ab initio quantum-mechanical characterization is presented of the vibrational properties 14 of pyrope and majorite garnets, using the hybrid B3LYP functional and large all electron Gaussian 15 type basis sets. Discussed quantities include infrared (both TO and LO) and Raman frequencies, normal modes coordinates, spectroscopic intensities, mode Grüneisen parameters, isotopic 16 17 substitution, and infrared and Raman spectra. Comparison with data available in the literature 18 demonstrates the accuracy of the adopted method. Main spectral features of the two garnets are 19 interpreted in terms of either symmetry analysis or structural contributions to the vibrational modes. 20 Missing peaks in the experiments are discussed in light of the simulated spectra. The present high 21 quality vibrational data can be used to compute thermal expansivities at high pressure and high 22 temperature conditions. Calculated values for majorite at the bottom of the mantle transition zone $(\alpha_V = 2.2 \times 10^{-5} \text{ K}^{-1} \text{ at } \text{T} = 1500 \text{ K} \text{ and } \text{P} = 20 \text{ GPa})$ turn out to be sensibly greater (up to three times) 23 24 than those currently adopted in geophysical thermodynamic databases, thus calling for a careful 25 revision of the numerical models for thermo-chemical convection of the Earth's mantle. 26 27 28 Keywords Infrared spectrum, Raman spectrum, vibrational frequencies, mode Grüneisen 29 parameters, thermal expansivity, *ab initio* quantum mechanical calculations, CRYSTAL code

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33 INTRODUCTION

34 Thermodynamic properties of high pressure mineral phases in the system MgO-Al₂O₃-SiO₂ 35 are crucial to understand chemico-physical processes in the Earth's mantle and to constrain the 36 mineralogical constitution of its deeper parts. Among these phases, MgSiO₃ majorite garnet and its 37 solid solutions with pyrope $(Mg_3Al_2Si_3O_{12})$ are of relevant interest as they are major components of 38 the mantle transition zone, ranging in volume from 40% to 70% according to different petrological 39 models (Bass and Anderson 1984; Ringwood 1991; Vacher et al. 1998; Weidner and Wang 1998; 40 Akaogi et al. 2002). Furthermore, majorite-pyrope garnets undergo phase transitions at high 41 pressure and temperature conditions, which may be at the origin of multiple seismic discontinuities 42 between 520 km and 720 km depths (Hirose et al. 2001; Hirose 2002; Deuss et al. 2006; Saikia et 43 al. 2008).

44 Vibrational spectroscopy, in both its infrared and Raman variants, is an invaluable tool to 45 gather lattice dynamics data that can be used for thermodynamical modelling. Also, Raman 46 barometry techniques are emerging as a tool in the study of mineral inclusions in diamond (e.g. 47 Izraeli et al. 1999), of which majoritic garnets are common examples (cf. Harte 2010). In fact, these 48 methods have been shown to provide reliable estimates of the residual internal pressure exerted by 49 the host diamond on the inclusion, hence of their pressure of formation by the so called elastic 50 relaxation method (Nestola et al. 2011, 2012; Angel et al. 2014). These reasons make very desirable 51 the availability of accurate spectroscopic characterisations of the majorite-pyrope system, both at 52 ambient and high pressure conditions.

Experimental characterization of the vibrational properties of majorite-pyrope garnets is challenging, mainly due to problems affecting their high pressure synthesis in the laboratory. Pervasive merohedral or pseudomerohedral twinning is always present in the tetragonal endmember of the majorite-pyrope series (Mg₄Si₄O₁₂ or, for simplicity, MgSiO₃), as revealed by TEM studies (Angel et al. 1989; Hatch and Ghose 1989; Wang et al. 1993). The finding of single crystals for structural refinement is thus hindered, and IR and Raman spectra currently available for this phase could be obtained exclusively on polycrystalline samples (Kato and Kumazawa 1985; McMillan et al. 1989; Rauch et al. 1996; Manghnani et al. 1998; Chopelas 1999). Additional complications in the interpretation of the vibrational spectra arise from lowering of symmetry with respect to cubic garnets (*Ia-3d* to $I4_1/a$ space group) and from possible structural disorder.

63 Ab initio simulation has proved to be an excellent method to provide accurate descriptions of 64 the vibrational properties of silicate minerals, especially when exploiting hybrid density functionals 65 (Zicovich-Wilson et al. 2004, 2008; Prencipe et al. 2004; Pascale et al. 2005a, 2005b; Orlando et al. 66 2006; Valenzano et al. 2009, 2010; Ferrari et al. 2009; Dovesi et al. 2009, 2011; De La Pierre et al. 67 2011; Demichelis et al. 2012; Noël et al. 2012; Ulian et al. 2013; Maschio et al. 2013a, 2014; 68 Prencipe et al. 2014; Belmonte et al. 2014; Erba et al. 2015). Contrary to the experiment, it permits 69 to easily investigate "ideal" systems, free from the issues brought by polycrystallinity, defects, 70 impurities, disorder. Whereas several studies are available on the vibrational spectra of pyrope (e.g. 71 Zicovich et al. 2008; Dovesi et al. 2011; Maschio et al. 2013a; and references therein), only two 72 recent articles can be found, exploring the vibrational properties of majorite, by Yu et al. (2011) and 73 Hernández et al. (2015), using pure DFT (Density Functional Theory) functionals; both of them are 74 focused on the thermodynamic properties of the minerals, so that vibrational spectra (and related 75 accuracy issues) are discussed to a little extent.

76 In the present work, *ab initio* quantum mechanical calculations have been performed at an 77 hybrid Hartree-Fock/DFT level of theory, to get new insights into the vibrational properties and the 78 ensuing thermodynamics of majorite-pyrope end-member compositions.

The paper is organised as follows. The next Section describes the adopted computational framework. Section on Results and Discussion deals with the main aspects regarding structural properties, vibrational frequencies, infrared and Raman spectra, and additional issues like isotopic substitution and mode Grüneisen parameters. The main perspectives of this work are outlined in the Implications Section. Additional material is available on the web site in the form of Supplementary 84 Information.

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86 COMPUTATIONAL METHOD

Calculations were performed with the CRYSTAL14 periodic *ab initio* code (Dovesi et al.
2014a, 2014b), by using all electron Gaussian type basis sets and the hybrid B3LYP functional
(Becke 1993; Lee et al. 1988; Stephens et al. 1994).

Magnesium, aluminum, silicon and oxygen were described by (8s)-(511sp)-(1d), (8s)-(611sp)-(1d), (8s)-(831sp)-(1d) and (8s)-(411sp)-(1d) contractions, respectively. The exponents (in bohr⁻²) of the most diffuse functions are $\alpha_{sp} = 0.225$ and $\alpha_d = 0.50$ for Mg, $\alpha_{sp} = 0.380$ and $\alpha_d =$ 0.50 for Al, $\alpha_{sp} = 0.193$ and $\alpha_d = 0.61$ for Si and $\alpha_{sp} = 0.181$ and $\alpha_d = 0.60$ for O. These basis functions were already used in previous works on Mg-Al silicates and oxides (Ottonello et al. 2009; Ottonello et al. 2010a, 2010b; Dovesi et al. 2011; Demichelis et al. 2012; Belmonte et al. 2013, 2014; Erba et al. 2014).

The level of accuracy in evaluating the Coulomb and Hartree-Fock exchange series is controlled by five parameters (Dovesi et al. 2014a), for which the 8 8 8 9 18 values were used. The threshold on the SCF energy was set to 10^{-8} Hartree (E_h) for geometry optimisation and to 10^{-10} E_h for calculation of vibrational frequencies. The reciprocal space was sampled according to a sublattice with shrinking factor (Dovesi et al. 2014a) set to 3 (along the 3 lattice vectors), corresponding to 4 (pyrope) and 6 (majorite) independent \vec{k} vectors in the irreducible part of the Brillouin zone.

104 The DFT exchange-correlation contribution is evaluated by numerical integration over the 105 unit cell volume. In CRYSTAL, radial and angular points of the grid are generated through Gauss-106 Legendre radial quadrature and Lebedev two-dimensional angular point distributions. A (75,974)p 107 grid was used, corresponding to a pruned grid with 75 radial and 974 angular points (XLGRID 108 keyword in the CRYSTAL manual; Dovesi et al. 2014a). Accuracy in the integration can be 109 estimated by the error Δe on the integrated electronic charge density in the unit cell ($\Delta e = 2 \cdot 10^{-4} |e|$ and $3 \cdot 10^{-6}$ |e| for pyrope and majorite, respectively, on a total of 800 electrons).

Structure were optimised by using analytical energy gradients with respect to atomic coordinates and unit cell parameters (Doll 2001; Doll et al. 2001; Civalleri et al. 2001), within a quasi-Newton scheme combined with the BFGS algorithm for Hessian updating (Broyden 1970; Fletcher 1970; Goldfarb 1970; Shanno 1970). Convergence was checked on both gradient components and nuclear displacements, for which the default values (Dovesi et al. 2014a) were chosen.

117 The calculation of vibrational frequencies at the Γ point ($v_{0,n}$) was performed within the 118 harmonic approximation; the mass-weighted Hessian matrix *W* is constructed by numerical 119 differentiation of the analytical gradients with respect to the atomic Cartesian coordinates:

120
$$W_{\alpha i,\beta j}(\Gamma) = \frac{H_{\alpha i,\beta j}}{\sqrt{M_{\alpha}M_{\beta}}}$$
(1)

121 where $H_{\alpha i,\beta j}$ is the second derivative of energy (evaluated numerically starting from the analytical 122 gradients), M_{α} and M_{β} are the atomic masses; greek and latin indices refer to atoms and atomic 123 Cartesian coordinates, respectively. The calculated (optimised) equilibrium geometry is taken as 124 reference. More details on the calculation of vibrational frequencies can be found in Pascale et al. 125 (2004), Zicovich-Wilson et al. (2004) and Zicovich-Wilson et al. (2008).

Once the Hessian matrix *H* is calculated, frequency shifts due to isotopic substitutions can be obtained at zero computational cost, by changing the masses in Equation (1). This possibility was exploited to evaluate the participation of the various atoms to the vibrational modes in the different regions of the spectrum. In particular, the masses of Mg, Al, Si and O were augmented by 7.5% in separate runs.

131 Infrared absorption intensities II_n were computed for each *n*-th mode by means of the mass-132 weighted effective mode Born charge vector (Barrow 1962; Hess et al. 1986), evaluated through a 133 Berry phase approach (Dall'Olio et al. 1997; Baranek et al. 2001; Noël et al. 2002); see Zicovich-134 Wilson et al. (2008) for more details. Isotropic Raman intensities II_n were computed for each mode through a fully analytical, coupled perturbed Hartree-Fock/Kohn-Sham scheme (Maschio et al.2013b, 2013c).

137 A graphical representation of the infrared and Raman spectra S(v) was obtained as a 138 superposition of Lorentzian functions *F*, one for each mode:

139
$$S(\nu) = \sum_{n} F(\nu; \nu_{0,n}, II_n, \delta_n)$$
(2)

140
$$F(v; v_{0,n}, H_n, \delta_n) = \frac{H_n}{\pi} \left[\frac{\delta_n/2}{\left(v - v_{0,n}\right)^2 + \delta_n^2/4} \right]$$
(3)

141 where δ_n is the damping factor of the *n*-th mode, which is related to the phonon lifetime. Being 142 unable to compute this quantity, we used a constant value of 8 cm⁻¹. S(v) curves were evaluated in 143 the range 0-1100 cm⁻¹, in steps of 1 cm⁻¹.

144 Mode Grüneisen parameters (γ_n) define the volume dependence of the vibrational frequencies 145 in the framework of quasi-harmonic approximation (QHA) as follows:

146
$$\gamma_n = -\left(\frac{\partial \ln v_{0,n}}{\partial \ln V}\right)_0 = -\frac{V}{v_{0,n}} \left(\frac{\partial v_{0,n}}{\partial V}\right)_0$$
(4)

147 where V is the unit cell volume. Values for these quantities were obtained by least-square fitting of 148 the vibrational frequencies computed at four different volume conditions ($V/V_0 = 0.88, 0.99, 1.00$ 149 and 1.01 for majorite; $V/V_0 = 0.80$, 0.90, 0.95 and 1.00 for pyrope). For each mineral the three 150 frequency datasets close to the uncompressed volume condition (i.e. $V/V_0 = 1$) permit an accurate 151 mapping of the frequency behaviour close to the minimum of the potential energy surface. Besides, 152 the additional dataset at high compression accounts for the high pressure conditions. Second-order 153 polynomial fitting turns out to accurately describe the volume dependence of all the vibrational 154 frequencies.

155 Calculated and experimental frequencies v_i were compared through three global indices 156 defined as follows: This is a preprint, the final version is subject to change, of the American Mineralogist (MSA) Cite as Authors (Year) Title. American Mineralogist, in press. (DOI will not work until issue is live.) DOI: http://dx.doi.org/10.2138/am-2016-5382

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157
$$\overline{\left|\Delta\right|} = M^{-1} \sum_{i=1}^{M} \left| v_i^{calc} - v_i^{exp} \right|$$
(5)

158
$$\overline{\Delta} = M^{-1} \sum_{i=1}^{M} \left(v_i^{calc} - v_i^{exp} \right) \tag{6}$$

159
$$\left|\Delta\right|_{max} = max \left|v_i^{calc} - v_i^{exp}\right| \qquad i = 1,...,M$$
(7)

160 where *M* is the number of data in the set, $|\overline{\Delta}|$ is the mean absolute difference between the calculated 161 and the experimental values, $\overline{\Delta}$ is the mean difference, $|\Delta|_{max}$ is the maximum absolute difference.

Graphical animations of the normal modes were built using the Jmol 3D engine (jmol.sourceforge.net;www.theochem.unito.it/crystal_tuto/mssc2013_cd/tutorials/webvib/index.htm 1) and are available on the CRYSTAL Web site (www.crystal.unito.it/animations-of-vibrationalmodes.php); this tool is very helpful for a simple and intuitive interpretation of the "nature" of the modes (stretching, bending, libration, translation, etc). Graphs were realised with the Gnuplot utility (www.gnuplot.info).

168

169 **RESULTS AND DISCUSSION**

170 Structure

171 Structural parameters of pyrope (Prp) and majorite (Maj) as obtained from our calculations 172 are given in Table 1. Prp has a cubic conventional cell with *Ia-3d* space group symmetry and 4 173 irreducible atoms; the ordered, lowest energy Mg/Si configuration of Maj was considered in this 174 study, featuring a tetragonal conventional cell, the $I4_1/a$ space group and 13 irreducible atoms.

When compared to structural refinements by Pavese et al. (1995) (Prp) and Angel et al. (1989) (Maj), the computed lattice parameters (i.e. cell edges a and c) show a slight overestimation of 0.94% (Prp) and 0.96%, 0.54% (Maj), which are in line with the known overestimation due to the adopted functional (see e.g. Dovesi et al. 2011). Atomic fractional coordinates compared to the above mentioned experimental studies show deviations as small as 0.0002÷0.0009 for Prp, and 180 slightly larger in the case of Maj, where all but one lie in the range $0\div0.0066$, with the only larger 181 discrepancy (0.0104) occurring for the *z* coordinate of Mg in site Dod1. These can be considered as 182 quite accurate results, given that previous computational studies of Maj showed deviations in the 183 same range (Vinograd et al. 2006; Yu et al. 2011; Hernández et al. 2015; Pigott et al. 2015).

184

185 Vibrational frequencies of pyrope

186 Symmetry analysis permits to classify the vibrational modes in terms of irreducible187 representations of the *m3m* point group:

188 $3 A_{1g} 5 A_{1u} 5 A_{2g} 5 A_{2u} 8 E_g 10 E_u 14 F_{1g} 17 F_{1u} 14 F_{2g} 16 F_{2u}$ (8) 189 There are 17 infrared (IR) active modes with F_{1u} symmetry, and 25 Raman active modes (A_{1g} , E_g 190 and F_{2g} symmetries); the remaining ones are silent.

The computational setup adopted in the present study (B3LYP functional and all electron Gaussian type basis sets) had already been used to analyse the IR (Zicovich et al. 2008; Dovesi et al. 2011) and Raman (Maschio et al. 2013a) properties of Prp. However, at variance with these previous studies, we have here widened the comparison with experiments by including a larger number of datasets, to better assess the accuracy that can be achieved in computational *versus* experimental vibrational spectroscopies.

197 Vibrational properties of Prp are reported in Tables S1-S5 in the Supplementary Information. 198 Computed IR active frequencies, both Transverse Optical (TO) and Longitudinal Optical (LO) are 199 reported in Table S1 and compared with one experimental study reported by McMillan et al. (1989), 200 three experimental studies carried out by Hofmeister and coworkers (Hofmeister and Chopelas 1991; Hofmeister et al. 1996, 2004) and a lattice dynamics simulation (Chaplin et al. 1998). 201 202 Statistics on the TO modes (Table S2), as compared with the three datasets by the Hofmeister group, reveal that our simulation has a mean absolute deviation $\overline{|\Delta|}$ in the range 5.2÷6.8 cm⁻¹, and a 203 maximum absolute deviation $\left|\Delta\right|_{max}$ of 25.7 cm⁻¹. Interestingly, cross-statistics among the three 204

datasets by Hofmeister and coworkers show very similar values: $\overline{|\Delta|}$ goes from 4.9 to 6.3 cm⁻¹, 205 while $\left|\Delta\right|_{max}$ is as large as 50.0 cm⁻¹. Thus accuracy of our computed IR frequencies is the same than 206 207 the typical precision of IR measurements carried out on Prp. 208 The only other theoretical study on Prp that is found is the literature was performed by 209 Chaplin et al. (1998) and made use of lattice dynamics. Compared to the experimental IR-TO datasets, $|\Delta|$ ranges from 19.1 to 23.6 cm⁻¹, with a maximum deviation of 44.5 cm⁻¹. These values 210 211 give evidence for the higher accuracy of *ab initio* B3LYP calculations with respect to classical force 212 field techniques. 213 Computed Raman frequencies (Table S3) are compared with data from two experimental 214 studies by Hofmeister and coworkers (Hofmeister and Chopelas 1991; Hofmeister et al. 2004), two 215 experimental studies by Kolesov and Geiger (1998, 2000) and once again the lattice dynamics simulation of Chaplin et al. (1998). When comparing our dataset with the experimental ones, $|\Delta|$ 216 217 values (Table S4) fall in the range $4.9 \div 7.1 \text{ cm}^{-1}$, showing essentially the same accuracy than the IR data. $\left|\Delta\right|_{\text{max}}$ can be as large as 46.0 cm⁻¹, but it lowers to 25.9 cm⁻¹ when considering only the most 218 recent sets from each experimental group. Experimental precision is higher than computational 219 accuracy in the case of Raman, as $\overline{|\Delta|}$ among experimental datasets is only 1.3÷3.9 cm⁻¹, with $|\Delta|$ 220 of 12.8 cm⁻¹. As for the case of IR-TO frequencies, lattice dynamics simulations (Chaplin et al. 221 222 1998) reveal large uncertainties in reproducing the Raman frequencies of Prp.

223

224 Vibrational frequencies of majorite

225 Classification of the modes according to the irreducible representations of the 4/m point group226 yields:

227 $25 A_g \quad 32 A_u \quad 27 B_g \quad 31 B_u \quad 28 E_g \quad 33 E_u$ (9)

228 The lower symmetry of Maj compared to Prp results in a larger number of distinct vibrational

modes: there are 65 IR active modes (A_u and E_u symmetries), 80 Raman active modes (A_g , B_g and E_g), and 31 silent modes (B_u). Data on the silent modes are reported in Table S6 in the Supplementary Information.

232 For each one of the 65 IR active modes, Table 2 shows computed TO and LO frequencies 233 together with the corresponding IR absorption intensity. Two experimental datasets are reported, 234 collected by Kato and Kumazawa (1985) and McMillan et al. (1989). Neither of the two experiments registered data below 350 cm⁻¹; our simulation provides information for 22 IR modes 235 236 in this spectral range. Of the remaining 43 modes above 350 cm⁻¹, the two experiments identify 15 and 22 of them, respectively. A closer look to the data reveals that 5 and 7 experimental peaks 237 238 (italic values in Table 2), respectively, can be assigned to pairs of computed modes, due to 239 frequency proximity of the latter, thus accounting for part of the modes that are missing in the 240 experiment. Five missing modes (bold values in Table 2) show a quite large intensity, but have not 241 been identified in the experimental curves probably due to superposition with very broad and intense bands; they lie around 480, 845, 915, 928 and 975 cm⁻¹. The remaining missing modes all 242 243 have relatively low intensities. Comparison between computed and experimental data is very good and of similar quality than for Prp: $\overline{|\Delta|}$ values are 5.7 and 4.6 cm⁻¹, respectively; in both cases only 244 one discrepancy is larger than 10 cm⁻¹ (it corresponds to the computed modes at 685 and 989 cm⁻¹, 245 246 respectively).

247 Table 3 reports computed Raman frequencies and intensities for the 80 Raman active modes; 248 the former are compared with four experimental datasets, and a DFT-LDA study by Yu et al. 249 (2011). Experiments identified 11 (Manghnani et al. 1998) to 34 (McMillan et al. 1989) peaks; at variance with the IR case, they include data from the whole spectral range from 130 to 1100 cm⁻¹, 250 251 so that missing peaks are distributed all along this span. Taking as a reference the most complete set 252 by McMillan et al. (1989), we note that as many as 12 experimental frequencies (italic values in 253 Table 3) can be associated to either pairs or triplets of neighbouring computed modes. All the remaining missing modes have relatively low intensity; a look at the spectra in the experimental 254

255 papers suggest that these could not be clearly identified due to the low signal-to-noise ratio. Note that three experimental datasets have a spurious frequency in the range 690÷750 cm⁻¹, which could 256 257 not be assigned to any computed mode. Statistics in Table 4 reveal a very good agreement of our simulated data with the experimental ones, with Δ ranging between 3.6 and 4.9 cm⁻¹, the best 258 259 agreement relating to the most complete dataset by McMillan et al. (1989). Only one computed mode shows discrepancies larger than 10 cm⁻¹, namely the low frequency one at 116.4 cm⁻¹ 260 (experiments at 136÷138 cm⁻¹); underestimation of modes in this range has been already reported in 261 262 the case of Prp (see this study, and Dovesi et al. 2011) and aragonite (De La Pierre et al. 2014).

263 The DFT study of Yu et al. (2011) reports frequencies for 22 out of 80 Raman active modes. 264 Statistical comparison with the experiments, carried out over the subset of matching frequencies, indicates a slighter worse match than our simulations, with $\overline{|\Delta|}$ between 4.9 and 5.4 cm⁻¹. A look at 265 266 the dataset in Table 3 shows that 8 out of 22 modes by that study cannot be put in relation with any 267 experimental value, once the latter have been classified by symmetry according to our analysis. The 268 latter has been performed based on symmetry classification of the computed modes (via an 269 algebraic decomposition of the reducible representation built on the basis of the Cartesian atomic 270 coordinates) and on the comparison of computed Raman intensities (a mode with higher computed 271 intensity is more likely to be observed experimentally). This suggests that symmetry labeling by Yu 272 et al. (2011) should be taken with some care. Note that a second set of DFT data on Maj has been 273 published very recently by Hernández et al. (2015), in a study focused on modeling a variety of 274 phases in the MgO-SiO₂ binary system, using the Perdew-Wang '91 GGA functional (Wang and 275 Perdew 1991). Despite their detailed and quite accurate results regarding structural, high pressure, 276 and thermodynamic properties, absolute values for the vibrational frequencies show a large underestimation, around 20 cm⁻¹ on average, compared to both our data and the experiments; this is 277 278 in line with the known trends for the level of theory they adopted.

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Overall, the accurate datasets presented here constitute a complete and comprehensive

investigation of the lattice dynamics of end-member Maj, representing a reliable reference not only
for spectroscopic and analytical purposes, but also for further thermodynamic investigations.

282

283 Isotopic substitution

Simulated isotopic substitution has been applied as a tool to quantitatively characterise the contribution of each chemical species to the vibrational modes of Prp and Maj. To this end, an artificial mass shift of +7.5% has been applied in turn to all atoms of a given species (Mg, Al, Si, or O), that results in negative frequency shifts along the spectral range.

Percent isotopic shifts for Prp are represented in Figure S1 in the Supplementary Information. Let us arbitrarily take a -1% shift as the threshold to identify modes contributed by a given species; then, the $100\div350$ cm⁻¹ and the $300\div600$ cm⁻¹ ranges are dominated by Mg and Al, respectively. Of the modes with frequency shift larger than -1% for Si, three lie around $250\div400$ cm⁻¹, one lies around 600 cm⁻¹, and the remaining ones are in the range $800\div1100$ cm⁻¹ (the latter correspond to Si-O stretching modes). Oxygen contributes to the vibrational modes along the whole spectral range.

295 Simulated isotopic substitution applied to Maj results in the frequency shifts represented in 296 Figure 1. Taking again a threshold of -1%, Maj shows features similar to Prp when dodecahedral 297 Mg, tetrahedral Si and O are considered; in particular, dodecahedral Mg can be associated to modes in the 100÷400 cm⁻¹ range, and tetrahedral Si to two modes around 300÷400 cm⁻¹, to one mode at 298 about 600 cm⁻¹, and to the range 800÷1100 cm⁻¹. Two separate substitutions were carried out for 299 300 Mg and Si occupying the octahedral sites. Octahedral Mg can be strictly related to three modes in the range $300 \div 450 \text{ cm}^{-1}$; if a slightly lower threshold is considered, around -0.75%, then the larger 301 302 250÷500 cm⁻¹ range can be assigned, moved to higher frequencies with respect to the dodecahedral Mg range. Octahedral Si relates to the $450 \div 750 \text{ cm}^{-1}$ range, at lower frequencies than the principal 303 304 tetrahedral Si range. Note how the two species occupying octahedral sites can be assigned to rather 305 distinct spectral ranges.

306

307 IR spectra

308 Computed IR spectra of Prp and Maj are represented in Figure 2, as blue and red curves, 309 respectively.

310 All modes above 750 cm⁻¹ are pure internal Si-O stretchings of the SiO₄ tetrahedra; there are 3 311 and 12 IR active modes in this range in Prp and Maj, respectively (see Table S1 and Table 2). These 312 modes are both limited in number and perturbed by a little extent by the structural differences 313 between Prp and Maj; thus it is possible to establish relations between the two garnets, by 314 exploiting correlation rules of irreducible representations (see e.g. Ferraro and Ziomek 1975): A_{1u} 315 and A_{2u} modes in Prp relate to A_u and B_u modes in Maj, respectively; E_u modes split into pairs of A_u 316 and B_u modes; F_{1u} and F_{2u} modes turn into pairs of $A_u + E_u$ and $B_u + E_u$ modes, respectively. 317 Correspondences are summed up in Table 5. Out of the 10 peaks visible in the IR spectrum of Maj, 318 6 of them are 3 pairs of A_u and E_u modes that can be related to the 3 F_{1u} peaks of Prp. The other 4 visible peaks relate to silent modes in Prp: 3 E_u modes at 876, 915 and 954 cm⁻¹ (resulting from F_{2u} 319 modes of Prp) and 1 A_u mode at 928 cm⁻¹, related to the E_u mode at 913 cm⁻¹ of Prp. 320

The frequency range below 750 cm⁻¹ features a large number of modes involving motions of 321 322 cations in the dodecahedral and octahedral sites, so that correspondences between the two spectra 323 are loosened and of little significance. Given the variety of structural motions involved in this 324 spectral range, it is interesting to characterise the most relevant peaks in the two spectra in terms of 325 the structural subunits dominating the corresponding normal modes. Both simulated isotopic 326 substitution (discussed above) and graphical inspection of the modes (see CRYSTAL Web site, 327 www.crystal.unito.it/animations-of-vibrational-modes.php) can be used to this end. Results are 328 summarised in Table S7 in the Supplementary Information. Of the 5 intense modes in the spectrum of Prp, the three low frequency ones at 119, 187 and 333 cm⁻¹ relate to Mg cations motions, with 329 330 the latter also involving translations of the SiO_4 tetrahedra; the remaining high frequency ones at 380 and 458 cm⁻¹ are motions of the Al cations, coupled with rocking motions of the SiO₄ 331

tetrahedra. In the case of Maj, starting from the lower frequencies, the four intense modes in the 332 range 340÷370 cm⁻¹ (labelled as 21, 22, 23 and 25 in Table 2) are related to Mg motions in both 333 334 dodecahedral and octahedral sites, coupled with SiO₄ librations and either translations or rockings. The three intense modes in the range 380÷400 cm⁻¹ (28, 30 and 31 in Table 2) relate to Mg motions 335 336 only in the octahedral sites, coupled with SiO₄ librations and rockings. The very intense mode at 337 441 cm⁻¹ involves motions of all the structural subunits. Finally, the two intense modes at 510 and 550 cm⁻¹ relate to motions of the Si atoms in the octahedral sites, coupled with SiO₄ librations and 338 339 rockings. Note the variety of structural motions in the case of Maj, arising from the octahedral sites 340 being occupied by both Mg and Si atoms.

341

342 Raman spectra

343 Computed Raman spectra are shown in Figure 3. Starting from the high frequency range above 750 cm⁻¹, Prp and Maj have 6 and 18 Raman active modes, respectively, 4 and 11 of which 344 345 are visible in the spectrum (see Table 5 for symmetry correspondences). In particular, 8 visible modes of Maj are in relation with active modes of Prp: the highly intense A_{1g} peak of Prp at 927 346 cm⁻¹ relates to the A_g peak of Maj at 926 cm⁻¹; the relatively intense peak of Prp at 864 cm⁻¹ is the 347 superposition of a F_{2g} and a Eg mode and results in four low intensity peaks in Maj (1 Ag, 2 Bg and 1 348 E_g), spread over the range 810÷880 cm⁻¹; the small F_{2g} peaks of Prp at 899 and 1072 cm⁻¹ relate to 349 the small E_g peaks of Maj at 910 and 1070 cm⁻¹; the A_g peak of Maj at 960 cm⁻¹ comes from a little 350 Eg feature of Prp at 939 cm⁻¹ (hidden by the very intense mode at 927 cm⁻¹). The remaining 3 351 352 visible modes of Maj result from silent Prp modes: the Ag mode representing the largest contribution to the peak found at 886 cm⁻¹ relates to the F_{1g} silent mode of Prp at 887 cm⁻¹; the two 353 small peaks at 796 cm⁻¹ (A_g) and 803 cm⁻¹ (E_g) both relate to the F_{1g} silent mode of Prp at 848 cm⁻¹. 354

As regards the spectral range below 750 cm⁻¹, isotopic substitution and graphical animation of the modes can be used again to inspect the nature of the corresponding vibrational modes. Findings are reported in Table S8 in the Supplementary Information. Of the 6 intense peaks of Prp the first

two, at 208 and 338 cm⁻¹, correspond to Mg motions coupled with SiO₄ librations; the two intense 358 peaks at 357 and 566 cm⁻¹ are pure SiO₄ librations; the two small peaks at 515 and 657 cm⁻¹ are 359 360 rocking motions of the SiO_4 tetrahedra. In the case of Maj, of the 11 modes which contribute to the most intense features in the spectrum, the 4 ones in the range $150 \div 230$ cm⁻¹ (labelled as 3, 4, 7 and 361 13 in Table 3) and the one at 369 cm⁻¹ are Mg motions coupled with SiO_4 librations. The small 362 peaks at 317 and 402 cm⁻¹ as well as the very intense mode at 602 cm⁻¹ are pure SiO₄ librations. 363 Finally, the three small peaks at 458, 541 and 653 cm^{-1} are SiO₄ rocking motions. Note how, despite 364 365 differences in peak positions and intensities, the Raman spectrum in the low frequency domain 366 presents similar structural contributions in Prp and Maj.

367

368 High pressure vibrational properties and mode Grüneisen parameters

369 Mode Grüneisen parameters allow us to test the accuracy of our simulations at high pressure 370 conditions, through direct comparison with the experimental data available for both Prp (Chopelas 371 2006) and Maj (Rauch et al. 1996; Manghnani et al. 1998; Chopelas 1999). Data for the Raman 372 modes can be found in Table S9 in the Supplementary Information and Table 6 in the case of Prp 373 and Maj, respectively; data for the remaining modes are reported in Tables S5, S6, S10, S11 in the 374 Supplementary Information. Accurate mode Grüneisen parameters are crucial to determine 375 fundamental thermodynamic properties, such as thermal expansion (Ottonello et al. 2010a, 2013; 376 Prencipe et al. 2011; Erba 2014; Belmonte et al. 2014; Ulian and Valdrè 2015).

The mode Grüneisen parameters calculated for Prp and Maj up to pressures of 25 GPa compare well with available experimental data, except for some low frequency modes displaying relatively large values, up to 3.7 for Prp and 3.5 for Maj in the frequency ranges 105 - 119 cm⁻¹ and 101 - 116 cm⁻¹, respectively; the corresponding experimental data are no larger than 1.9 and 1.8, respectively. Note that in the case of Maj the relatively large computed γ values are in excellent agreement with the very recent GGA results by Hernández et al. (2015) (see their Supplementary content): compare our values of 3.51 and 3.36 for the two lowest frequency modes, namely the IR

active E_u mode at 101 cm⁻¹ and the Raman active E_g mode at 116 cm⁻¹, respectively, with the 384 385 literature values of 3.50 and 3.59. Furthermore, the Raman spectra measured for Maj by Manghnani 386 et al. (1998) and Chopelas (1999) reveal that the lowest frequency experimental modes, observed in the range 136 - 220 cm⁻¹, have γ values twice as large than the medium-to-high frequency modes 387 (and the averaged γ_i , as well). Although mode Grüneisen parameters sensibly greater than 1.0 are 388 389 possible (see for instance Lu et al. 2012), the apparently anomalous values associated with the 390 vibrational modes with lowest frequencies may have a lattice dynamical justification. In particular, the mismatch between the relatively small Mg^{2+} cation and the size of the dodecahedral sites in the 391 392 garnet structure is the most cited cause for the strongly anisotropic character of the thermal motions 393 of these modes (Armbruster et al. 1992; Pavese et al. 1995; Kolesov and Geiger 2000); anharmonic 394 effects could become relevant in this context (Winkler et al. 2000). The fact that both in Prp and 395 Maj these modes are dominated by motions of Mg atoms in the dodecahedral sites, as confirmed by 396 the analysis in the Isotopic Substitution Section above, clearly supports this interpretation.

397 Apart from the case of the just discussed low frequency modes (which represent only the 2% 398 of all the modes), our calculated high pressure properties are in very good agreement with 399 experimental results. In particular, Figure 4 shows the calculated pressure dependence for all the 400 Raman active modes of Maj (Ag, Bg and Eg symmetries) up to 25 GPa, as compared with the 401 experimental measurements by Rauch et al. (1996), Manghnani et al. (1998) and Chopelas (1999). 402 It turns out that all the vibrational frequencies increase linearly with pressure (and so all the mode 403 Grüneisen parameters are positive). By comparing the numerical values of both γ and dv/dP for the 404 Raman modes of Maj, as reported in Table 6, we may conclude that there is a general agreement 405 between experiment and simulation, within the limits of experimental and/or computational 406 uncertainty; the former is related to possible impurities in the samples (e.g. hydroxyl groups), cation 407 disorder, instrumental setup, whereas the latter mainly to the choice of the exchange-correlation 408 functional. A relevant feature of the high pressure Raman spectra of Maj is the large value of the 409 dv/dP slopes characterising the high frequency bands in three different experimental studies (Rauch 410 et al. 1996, Manghnani et al. 1998, Chopelas 1999). These slopes are between 3.7 and 4.3 cm⁻¹/GPa 411 in the frequency range above 750 cm⁻¹, whereas they never exceed 2.55 cm⁻¹/GPa in the range 412 below 750 cm⁻¹. Our calculated dv/dP slopes confirm this point: they lie within 3 and 5 cm⁻¹/GPa in 413 the high frequency range, and they rarely exceed 3 cm⁻¹/GPa in the low-to-medium frequency 414 range.

415 Finally, using the mode Grüneisen parameters we have computed the cell volumes of Prp and Maj 416 as corrected for vibrational effects at 300 K and 1 bar; the corresponding volume expansions 417 amount to 0.21% and 0.24%, respectively. The average value for the mode Grüneisen parameters of 418 both minerals is about 1; this value can be used (see Aliatis et al. 2015) to estimate that the change 419 in frequencies related to these volume expansions is about -0.2%. Thus frequency corrections due to vibrational pressure lie between 0 and -2 cm⁻¹ when spanning the spectral range from the lowest 420 frequency to the highest. We have mentioned above that a few modes in the range 100 - 120 cm⁻¹ 421 422 have peculiarly large mode Grüneisen parameters, up to 3.7; the corresponding maximum frequency correction for these modes is about -1 cm^{-1} (-0.8%). 423

424

425 **IMPLICATIONS**

In this work the vibrational properties of majorite-pyrope garnet end-members have been thoroughly investigated both at ambient and high pressure conditions by means of *ab initio* quantum mechanical simulations. An accurate and complete dataset of vibrational spectroscopic properties for majorite has been presented here for the first time, alongside the corresponding dataset for pyrope.

The present detailed analysis is particularly relevant in the case of the majorite end-member, for whose vibrational spectrum no complete experimental characterisation is available. Thermal expansivity (α_V) is among the thermodynamic properties stemming from vibrational frequencies and mode Grüneisen parameters of crystalline phases (see Anderson 1995 for details). Together with density, this quantity is of fundamental importance for the physical and compositional models 436 of the Earth's interior, since it controls the buoyancy forces that drive thermal convection in the437 mantle (Christensen 1995).

438 Despite the fact that majorite is a major constituent of the mantle transition zone, thermal 439 expansivity values for this mineral are poorly constrained by experiments due to several technical 440 problems, so that their extrapolation at high pressure and temperature conditions is affected by large 441 uncertainties. A survey of the literature values clearly shows a high degree of scattering, with α_V ranging from 1.6×10^{-5} K⁻¹ (Jacobs and de Jong 2007) to 2.2×10^{-5} K⁻¹ (Fabrichnava et al. 2004) at 442 443 298.15 K and 1 bar, according to different thermodynamic assessments. This is the reason why, in 444 most numerical simulations of mantle convection, either the thermal expansivity of majorite is 445 assumed constant or only its pressure dependence is taken into account, by means of simplified 446 parameterizations relying on extrapolations of low pressure data (Schmeling et al. 2003; Tosi et al. 447 2013).

448 Due to the lack of experimental data, computational modelling is the best way to obtain 449 reliable and physically consistent values for the thermal expansivity of majorite at high pressure and 450 high temperature conditions. Employing the dataset of vibrational properties presented in this paper, a value of 2.1×10^{-5} K⁻¹ can be calculated for α_V at room pressure and temperature conditions. This 451 452 value lies within the range of the current thermodynamic databases for geophysical modelling and is in good agreement with the lattice dynamics calculation by Chopelas et al. (2000) (i.e. 2.24×10^{-5} 453 K^{-1}) and the first principles GGA calculations by Hernandez et al. (2015) (i.e. $2.15 \times 10^{-5} K^{-1}$). In 454 contrast, DFT simulations by Yu et al. (2011) using both LDA and GGA give sensibly 455 overestimated values, namely 2.5×10^{-5} K⁻¹ and 2.9×10^{-5} K⁻¹, respectively. Most importantly, if high 456 457 pressure and high temperature conditions are considered, $\alpha_{\rm V}$ values currently adopted for majorite 458 in thermodynamic databases (cf. Fabrichnaya et al. 2004; Jacobs and de Jong 2007) seem to be 459 drastically underestimated. By way of example, let us consider the pressure and temperature conditions relevant for the bottom part of the mantle transition zone. At 1500 K and 20 GPa, the 460 thermodynamic tabulations by Fabrichnava et al. (2004) give $\alpha_V \simeq 0.7 \times 10^{-5} \text{ K}^{-1}$. Using our first 461

462 principle results, the value obtained at the same conditions is 2.2×10^{-5} K⁻¹, i.e. more than three 463 times as large. The high quality data presented here therefore represent the ideal starting point for 464 an extensive revision of the thermodynamic and thermoelastic properties of mantle minerals at deep 465 Earth's conditions, aimed at improving the numerical modelling of thermo-chemical convection in 466 the mantle.

467

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476 TABLES

		Prp					Ma	j	
		а		V			а	с	V
Cell		11.546		1539.19	Cell		11.611	11.543	1556.19
Atom	Site	x	у	z	Atom	Site	x	у	z
Mg	Dod1	0.125	0	0.25	Mg	Dod1	0.6293	0.0147	0.2691
					Mg	Dod2	0.5	0.25	0.6232
Al	Oct1	0	0	0	Mg	Oct1	0.5	0	0.5
					Si	Oct2	0.5	0	0
Si	Tet1	0.375	0	0.25	Si	Tet1	0.5	0.25	0.375
					Si	Tet2	0.5	0.25	0.875
					Si	Tet3	0.6249	0.0107	0.7560
0	O(1)	0.0323	0.0498	0.6535	0	O(1)	0.5251	0.0574	0.6680
					0	O(2)	0.5437	-0.0450	0.8605
					0	O(3)	0.7238	0.1060	0.8045
					0	O(4)	0.7133	-0.0847	0.7025
					0	O(5)	0.4377	0.1629	0.4688
					0	O(6)	0.3979	0.2146	0.7823

Table 1 Calculated structural properties of pyrope and majorite. Cell parameters and cell volume

479 are in Å and Å³, respectively; atomic positions are expressed as fractional coordinates.

_			(DC)I will r	ot work unt	il issue	is live.) C)0I: ł	nttp://dx	.doi.or	rg/10.2	138/am-201	6-5382
				ν	'TO,exp							V	'TO,exp	
j	# Sym	V _{TO,calc}	II _{calc}	Kato '85	McMillan '89	$v_{LO,calc}$		# 5	Sym	v _{TO,calc}	II _{calc}	Kato '85	McMillan '89	$v_{LO,calc}$
Ī	1 E _u	101.3	1030			114.6	3	6	Eu	457.6	1111	450	450	454.3
ź	2 E _u	130.9	175			134.1	3	7	Au	464.0	186	438	439	466.5
	3 A _u	142.4	730			148.5	3	8	Au	479.8	544			486.3
4	4 A _u	156.4	673			178.4	3	9	Eu	491.1	780		490	482.3
:	5 E _u	172.3	577			177.7	4	0	Eu	504.3	248	501	501	500.6
	6 A _u	191.8	44			193.6	4	1	A _u	510.3	2172	501	501	555.0
1	7 E _u	197.7	170			198.7	4	2	E_{u}	522.5	1278	520	521	509.5
	8 E _u	206.8	679			213.0	4	3	E_{u}	550.1	2112	549	549	534.8
9	9 A _u	213.4	94			215.5	4	4	A_u	563.2	55			565.7
10	0 E _u	222.5	140			224.1	4	5	A_u	579.7	278	576	500	585.2
1	1 E _u	240.2	307			242.4	4	6	Eu	583.2	608	370	200	576.1
12	$2 A_u$	247.1	388			253.3	4	7	A _u	599.6	832		604	637.8
12	3 E _u	257.3	829			264.5	4	8	Eu	636.8	510	627	632	612.3
14	4 A _u	258.8	1			258.8	4	9	A_u	656.7	4			656.9
1:	5 E _u	277.0	62			277.6	5	0	Eu	684.8	200	670	675	687.4
10	5 A _u	289.8	386			294.2	5	1	A _u	701.6	82	693	693	703.4
1′	7 E _u	300.3	336			302.3	5	2	Eu	706.8	31			707.1
1	8 A _u	312.1	78			312.8	5	3	Au	730.4	32			730.9
1	9 A _u	322.6	57			323.1	5	4	Eu	826.3	8316	831	827	860.3
20	0 E _u	325.8	403			327.4	5	5	Au	845.1	469			847.3
2	1 E _u	341.4	1119			346.1	5	6	Eu	876.1	1256	076	070	885.1
22	$2 A_u$	345.1	1234			350.5	5	7	A _u	886.4	6185	870	879	1071.5
2	3 A _u	360.5	1596		254	373.7	5	8	Eu	901.4	2437	903	906	935.0
24	4 E _u	361.7	213		354	362.2	5	9	Eu	915.1	478			912.7
2:	5 E _u	368.8	1052		362	372.0	6	0	A _u	928.4	1011			920.2
20	5 A _u	372.4	22			371.7	6	1	A_{u}	950.4	48			949.5
2'	7 A _u	379.6	579		202	388.3	6	2	Eu	954.2	793	950	954	961.8
2	8 E _u	384.9	2453		303	391.4	6	3	A _u	974.5	1357			958.4
29	9 A _u	392.1	35			392.5	6	4	E_{u}	989.1	3867	998	1002	1067.2
30	0 E _u	396.5	1174	205	200	404.6	6	5	A_u	1055.5	37			1051.5
3	1 A _u	399.0	980	393	398	440.7	М					15	22	
32	2 E _u	417.1	1079		416	421.7	$\overline{ \Delta }$					5.7	4.6	
3.	3 Au	425.5	317		433	418.7	$\overline{\Delta}$					3.5	1.1	
34	4 Eu	440.5	4968	442	446	648.7	$\left \Delta\right _{ma}$	x				14.8	12.9	
3	5 Au	452.7	21			453.1			ļ					

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Table 2 Calculated and experimental IR properties of majorite. v and II are the frequencies [cm⁻¹]
and absorption integrated intensities [km/mol], respectively. Frequencies of both TO and LO modes
are given. Statistical indices (defined in the Computational Method Section) are reported on bottom.
Experimental frequencies in italic are assigned to pairs of contiguous computed modes (the most

- 496 intense one has been chosen to compute statistics); calculated frequencies in bold correspond to
- 497 intense peaks not identified in the experiments.

498

						Vexp		v_{calc}
#	Sym	ν.	п.	McMillan	Rauch	Manghnani	Chopelas	Yu
#	Sym	Vcalc	IIcalc	'89	'96	'98	'99	'11
1	E_{g}	116.4	87	138			136	130.34
2	E_{g}	138.0	9					
3	B_{g}	154.2	147	159				
4	$A_{g} \\$	154.5	165	107				
5	E_{g}	162.9	2					
6	$B_{g} \\$	164.6	6					
7	$A_{g} \\$	177.5	265	181				
8	E_{g}	189.6	96	107	196.1	107	105	198.90
9	$B_{g} \\$	200.1	45	197	200.0	197	195	194.46
10	$A_{g} \\$	210.3	30	205				
11	E_{g}	211.8	6					214.84
12	\mathbf{B}_{g}	212.1	2					
13	$A_{g} \\$	229.7	369	226	222.0	222	226	
14	E_{g}	237.2	105	238	222.0			
15	$\mathbf{B}_{\mathbf{g}}$	242.4	79	230	233.9			237.82
16	A_{g}	256.9	49	261	255.2			252.27
17	E_{g}	257.4	5					
18	A_{g}	276.4	28	275	272.2		277	281.34
19	$\mathbf{B}_{\mathbf{g}}$	281.6	24	275	273.2		277	
20	E_{g}	298.0	11	202				
21	$\mathbf{B}_{\mathbf{g}}$	299.0	39	293				
22	Eg	308.3	24					
23	$\mathbf{B}_{\mathbf{g}}$	315.7	19					320.71
24	$\mathbf{A}_{\mathbf{g}}$	315.9	0					
25	A_{g}	317.1	146	311	307.9		311	313.30
26	Eg	320.5	9					
27	A_{g}	338.8	33	336	332.4		334	342.43
28	Eg	344.1	16					
29	\mathbf{B}_{g}	346.7	12					
30	$\mathbf{B}_{\mathbf{g}}$	352.8	60	251	254.0	250		351.55
31	Eg	357.3	29	534	554.9	229		
32	Eg	363.9	11					372.98
33	$\mathbf{B}_{\mathbf{g}}$	367.4	51	267	265.0		267	
34	A_{g}	368.9	229	307	303.9		30/	
35	\mathbf{B}_{g}	382.5	3					
36	Eg	393.1	24					398.57
37	$\mathbf{A}_{\mathbf{g}}$	396.9	17					

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38	A _a	402.3	113	398	396.6		398	.001.019/10.2	1.
39	Ba	407.1	25				• • •		
40	Ea	413.4	1						
41	A	429.8	53	429	427.8			428.33	
42	A _o	458.0	116	458	456.4	456	458		
43	E,	462.2	20						
44	B	466.5	3					459.95	
45	Eg	481.2	8						
46	Bg	482.9	5	481					
47	Bg	500.4	59	498	493.6				
48	Ag	506.0	33	507	515.8				
49	Eg	509.7	1						
50	A_g	532.3	11						
51	E_{g}	540.6	110	535	533.1				
52	\mathbf{B}_{g}	559.4	60	550	5540				
53	A_{g}	567.7	59	559	554.0				
54	$\mathbf{B}_{\mathbf{g}}$	585.7	18						
55	E_{g}	593.0	6						
56	$A_{g} \\$	602.0	1667	602	599.0	597	602		
57	E_{g}	611.3	21						
58	\mathbf{B}_{g}	622.2	9					607.77	
59	$A_{g} \\$	636.5	2			641			
60	\mathbf{B}_{g}	639.1	16			041			
61	E_{g}	653.2	197	648	646.7		648		
62	\mathbf{B}_{g}	670.7	66					649.49	
	?			724		743	690		
63	A_{g}	795.9	79					804.06	
64	E_{g}	802.5	139	802	800.0	797	802		
65	\mathbf{B}_{g}	812.9	72						
66	E_{g}	850.0	97	852	852.8				
67	A_{g}	867.2	181	873					
68	\mathbf{B}_{g}	878.6	89						
69	Ag	886.2	399						
70	B_g	886.9	65	889	884.7	885	889	893.12	
71	Eg	888.4	45						
72	Eg	909.8	110						
73	Ag	926.1	1938	931	929.6	928	931	921.72	
74	Bg	957.0	28						
75	A _g	960.2	222	964	964.0			966.43	
76	E _g	984.5	7	989					
77	A _g	986.4	9	1024					
78	В _g	1029.6	14	1034					
79	E _g	1069.6	214	1065	1064.7	1058			
80	B_g	1073.0	57					1067.28	

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499 Table 3 Calculated and experimental Raman properties of majorite. v and II are the frequencies

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- 500 [cm⁻¹] and isotropic integrated intensities (in arbitrary units), respectively. Experimental
- 501 frequencies in italic are assigned to pairs or triplets of contiguous computed modes (the most
- 502 intense one has been chosen to compute statistics).
- 503

		McMillan '89	Rauch '96	Manghnani '98	Chopelas '99	Yu '11
This work	М	33	25	10	14	22
	$\overline{ \Delta }$	3.6	4.4	4.9	4.3	6.8
	$\overline{\Delta}$	-0.3	2.1	2.4	0.3	-0.1
	$\left \Delta\right _{max}$	21.6	9.8	11.6	19.6	21.2
McMillan '89	М		24	9	14	10
	$\overline{ \Delta }$		2.7	3.9	0.6	4.9
	$\overline{\Delta}$		1.6	2.8	0.3	1.4
	$\left \Delta\right _{max}$		8.8	7.0	2.0	9.3
Rauch '96	М			9	13	10
	$\overline{ \Delta }$			2.3	2.6	4.9
	$\overline{\Delta}$			1.4	-1.8	-1.0
	$\left \Delta\right _{max}$			6.7	5.0	10.0
Manghnani '98	М				7	3
	$ \Delta $				3.6	5.4
	$\overline{\Delta}$				-3.0	5.4
	$\left \Delta\right _{max}$				5.0	7.4
Chopelas '99	М					6
	$ \Delta $					5.1
	$\overline{\Delta}$					0.1
	$\left \Delta\right _{max}$					9.3



505 Table 3). Statistical indices are defined in the Computational Method Section.

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		IF	ł		Raman						
	Prp				Prp		Maj				
ν_{calc}	Sym	II _{calc}	ν_{calc}	Sym	II _{calc}	ν_{calc}	Sym	II _{calc}	ν_{calc}	Sym	II _{calc}
967 0	Б	12050	826.3	Eu	8316	0177	Б		795.9	$\mathbf{A}_{\mathbf{g}}$	79
007.9	F 1u	13939	845.1	$\mathbf{A}_{\mathbf{u}}$	469	047.7	Γ _{1g}		802.5	$\mathbf{E}_{\mathbf{g}}$	139
960.9	Б		821.2 B _u	962.0	Б	510	812.9	$\mathbf{B}_{\mathbf{g}}$	72		
809.8	Γ _{2u}		876.1	$\mathbf{E}_{\mathbf{u}}$	1256	803.9	Г _{2g}	510	850.0	$\mathbf{E}_{\mathbf{g}}$	97
909 C	Б	5702	886.4	$\mathbf{A}_{\mathbf{u}}$	6185	965 0	Б	10	867.2	$\mathbf{A}_{\mathbf{g}}$	181
090.0	r _{1u}	5795	901.4	$\mathbf{E}_{\mathbf{u}}$	2437	805.0	Еg	40	878.6	$\mathbf{B}_{\mathbf{g}}$	89
907.2	A_{2u}		878.5	\mathbf{B}_{u}		0071	Б		886.2	$\mathbf{A}_{\mathbf{g}}$	399
012.8	Б		925.5	B_{u}		00/.1	Γ _{1g}		888.4	E_{g}	45
912.0	Ľu		928.4	$\mathbf{A}_{\mathbf{u}}$	1011	800.2	F.	47	886.9	\mathbf{B}_{g}	65
015.0	Б		915.1	$\mathbf{E}_{\mathbf{u}}$	478	099.2	Г _{2g}	4/	909.8	$\mathbf{E}_{\mathbf{g}}$	110
915.0	r _{2u}		937.7	\mathbf{B}_{u}		926.6	$\mathbf{A}_{1\mathbf{g}}$	3072	926.1	$\mathbf{A}_{\mathbf{g}}$	1938
027 1	Б		950.4	A_{u}	48	020.2	Б	10	957.0	$\mathbf{B}_{\mathbf{g}}$	28
937.1	Ľu		951.9	\mathbf{B}_{u}		939.2	Ľg	19	960.2	$\mathbf{A}_{\mathbf{g}}$	222
056.1	Б		954.2	$\mathbf{E}_{\mathbf{u}}$	793	075 5	Б		984.5	E_{g}	7
930.1	Γ _{2u}		980.4	\mathbf{B}_{u}		913.3	г _{1g}		986.4	$\mathbf{A}_{\mathbf{g}}$	9
072 0	Б	5020	974.5	$\mathbf{A}_{\mathbf{u}}$	1357	1032.9	A_{2g}		1029.6	$\mathbf{B}_{\mathbf{g}}$	14
912.8	r _{1u}	3920	989.1	$\mathbf{E}_{\mathbf{u}}$	3867	1073 3	Б	205	1069.6	$\mathbf{E}_{\mathbf{g}}$	214
1052.5	A_{1u}		1055.5	A_u	37	1072.3	I ^r 2g	295	1073.0	$\mathbf{B}_{\mathbf{g}}$	57

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Table 5 Correspondences between modes of pyrope and majorite in the high frequency range
(above 750 cm⁻¹) of the IR and Raman spectra. Modes in bold are the ones distinguishable in the
spectra (see Figures 2 and 3).

	(D	OI will not	work un	itil issue is live	e.) DOI: http	://dx.doi.org	g/10.213	38/am-2016-53	82	
				$d\nu/dP_{exp}$	<i>a</i> 1 1	dv/dP_{calc}		γ_{exp}		Yeale
# Sym	ν_{calc}	$d\nu/dP_{calc}$	Rauch '96	Manghnani '98	Chopelas '99	Yu '11	γ_{calc}	Manghnani '98	Chopelas '99	Yu '11
1 E _g	116.4	2.51			1.61	1.51	3.36		1.77	1.86
2 E _g	138.0	1.72					1.94			
3 B _g	154.2	0.22					0.22			
4 A _g	154.5	2.06					2.08			
5 E _g	162.9	1.49					1.43			
6 B _g	164.6	2.68					2.54			
7 A _g	177.5	1.57					1.38			
8 E _g	189.6	1.87	0.92			1.27	1.54			1.03
9 B _g	200.1	1.21	0.72	0.99	1.20	1.47	0.94	0.84	0.90	1.22
$10 \ A_g$	210.3	0.49					0.36			
11 E _g	211.8	2.27				1.97	1.67			1.47
12 B _g	212.1	2.08					1.53			
13 A _g	229.7	2.02	1.98	2.11	1.97		1.37	1.56	1.32	
$14 E_g$	237.2	1.93	1.25				1.27			
$15 B_g$	242.4	1.96				1.80	1.26			1.22
$16 A_g$	256.9	1.48	1.9			1.30	0.90			0.83
$17 E_g$	257.4	2.98					1.81			
18 Ag	276.4	1.52	1.69		1.70	1.67	0.86		0.97	0.95
$19 \ B_g$	281.6	1.46					0.81			
$20 \ E_g$	298.0	1.75					0.92			
$21 \ B_g$	299.0	1.88					0.98			
$22 \ E_g$	308.3	2.47					1.25			
$23 \ B_g$	315.7	3.21				2.21	1.59			1.11
$24 A_g$	315.9	3.92					1.93			
$25 \ A_g$	317.1	1.83	1.73		1.90	1.53	0.90		0.92	0.79
$26 \ E_g$	320.5	1.46					0.71			
$27\ A_g$	338.8	2.33	2.17		2.53	2.20	1.07		1.14	1.03
$28 \ E_g$	344.1	2.32					1.05			
$29\ B_g$	346.7	4.61					2.08			
$30 \ B_g$	352.8	2.34	2.05	1.65		1.91	1.04	0.74		0.87
$31 \ E_g$	357.3	2.50					1.09			
32 E _g	363.9	2.31				2.08	0.99			0.90
$33 \ B_g$	367.4	3.25					1.38			
34 A _g	368.9	2.73	2.6		2.32		1.15		0.95	
$35 \ B_g$	382.5	2.84					1.16			
$36 E_g$	393.1	2.77				2.00	1.10			0.81
37 A _g	396.9	2.63					1.04			
38 A _g	402.3	1.64	2.7		2.34		0.64		0.54	
$39 \ B_g$	407.1	2.59					0.99			
$40 \ E_g$	413.4	1.66					0.63			
41 A _g	429.8	2.37	2.1			2.77	0.86			1.04
42 A _g	458.0	1.31	1.62	1.91	1.83		0.45	0.69	0.60	

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43 E _g	462.2	2.14					0.72			
44 B _g	466.5	1.83				1.52	0.61			0.53
45 E _g	481.2	1.55					0.50			
46 B _g	482.9	2.26					0.73			
47 B _g	500.4	1.57	1.44				0.49			
48 Ag	506.0	2.32	1.61				0.72			
49 E _g	509.7	2.37					0.73			
50 A _g	532.3	2.50					0.73			
51 E _g	540.6	2.29	2.11				0.66			
52 B _g	559.4	2.27	2.5				0.63			
53 A _g	567.7	2.60					0.72			
54 Bg	585.7	1.61					0.43			
55 E _g	593.0	2.62					0.69			
$56 A_g$	602.0	3.18	2.49	2.54	2.51		0.83	0.70	0.63	
57 E _g	611.3	2.86					0.73			
$58 B_g$	622.2	2.67				2.38	0.67			0.63
$59 A_g$	636.5	2.04		2.43			0.50	0.63		
$60 \ B_g$	639.1	1.51					0.37			
61 E _g	653.2	2.25	2.31		2.34		0.54		0.54	
$62 \ B_g$	670.7	2.35				2.14	0.55			0.53
$63 A_g$	795.9	4.21				3.88	0.83			0.78
64 E _g	802.5	4.20	3.8	4.29	3.39		0.82	0.89	0.63	
$65 \ B_g$	812.9	3.61					0.69			
66 E _g	850.0	5.95	3.9				1.09			
$67 A_g$	867.2	3.27					0.59			
$68 \ B_g$	878.6	5.78					1.03			
$69\ A_g$	886.2	3.79	3.8	3.91	3.93		0.67	0.72	0.66	
$70 \ B_g$	886.9	4.75				4.00	0.84			0.72
71 E _g	888.4	4.34					0.76			
72 E _g	909.8	4.86					0.83			
73 A _g	926.1	4.59	3.7	3.67	4.20	3.56	0.77	0.66	0.67	0.62
$74 \ B_g$	957.0	4.86					0.79			
75 A _g	960.2	5.29	3.9			3.81	0.86			0.63
76 E _g	984.5	5.45					0.86			
77 A _g	986.4	5.56					0.88			
78 B _g	1029.6	5.09					0.77			
79 E _g	1069.6	5.13	4.1	3.72			0.75	0.59		
80 B _g	1073.0	5.07				3.91	0.74			0.59
$\overline{\gamma}$							1.02	0.81	0.87	0.92

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Table 6 Calculated and experimental mode Grüneisen parameters γ (dimensionless) and pressure dependences dv/dP [cm⁻¹/GPa] for the Raman modes of majorite. $\overline{\gamma}$ is the mean value computed over each set of γ data.

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527 FIGURES



534 Figure 1 Percentage isotopic shifts Δv % on the vibrational frequencies of majorite. Masses of each





536

537

Figure 2 Computed IR spectra of pyrope and majorite. The peak integrated intensity of each *n*-th mode is II_n [km/mol] (see Tables S1 and 2).



541 **Figure 3** Computed Raman spectra of pyrope and majorite. The peak integrated intensity of each *n*-

542 th mode is II_n (in arbitrary units, see Tables S3 and 3).





544 Figure 4 Calculated pressure dependence of Raman frequencies of majorite, as compared with 545 experimental data.

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