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Normal to inverse transition in martian spinel:

Understanding the interplay between chromium, vanadium,

and iron valence state partitioning through a crystal-chemical lens

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ABSTRACT

Spinel is a very important rock-forming mineral that is found in basalts from Earth, Mars,

the Earth's Moon, and basaltic meteorites. Spinel can be used as a sensitive indicator of

petrologic and geochemical processes that occur in its host rock. This paper highlights the role of

increasing fO_2 (from IW-1 to FMQ+2) in converting a >90% normal spinel to an ~25%

magnetite (inverse) spinel, the trajectory of $D_V^{spinel/melt}$ as it relates to the ratio of V^{3+}/V^{4+} in the

melt, and the crystal chemical attributes of the spinel that control the intrinsic compatibility of

both V^{3+} and V^{4+}. This work examines the nuances of the V partitioning and provides a crystal

chemical basis for understanding Fe^{3+}, Cr, and V substitution into the octahedral sites of spinel.

Understanding this interplay is critical for using spinels as both indicators of planetary parentage

and reconstructing the redox history of magmatic systems on the terrestrial planets. Three

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24   potential examples for this use are provided. In addition this work helps explain the ubiquitous

25   miscibility gap between spinels with changing ülvospinel contents.

26   **INTRODUCTION**

27   Spinels are a common accessory phase in many terrestrial and extraterrestrial materials

28   and their structure and composition are uniquely responsive to changing conditions (e.g.

29   Haggerty 1972, 1976; Lindsley 1976, 1991; Basaltic Volcanism Study Project 1981; McKay et

30   al. 2004a,b; Righter et al. 2006; Ishizawa et al. 2014). Specifically, the spinel structure and its

31   inherent ability to accommodate multi-valent cations such as V, Cr, and Fe is highly sensitive to

32   changing temperature and fO_2 conditions (e.g. Hill and Roeder 1974, McKay et al. 2004a,b;

33   Muan et al. 1972, Righter et al. 2006).

34   Papike et al. (2004) suggested that zoned, magmatic spinels could be used to determine

35   relative oxygen fugacities and planetary parentage. This study specifically noted that the

36   behavior of V is very different when comparing terrestrial, lunar, and martian basalts. In

37   terrestrial basaltic liquids, typically $V^{4+} > V^{3+}$, whereas in lunar basalts $V^{3+} > V^{4+}$, and in martian

38   basalts V^{3+} and V^{4+} are both significant. The core to rim trends for lunar spinels show a strong

39   positive correlation of V and Cr and negative correlation of V and Ti. For the Earth, the trends

40   are just the opposite, with a strong negative correlation of V and Cr and strong positive

41   correlation of V and Ti. Chromite in martian basalts shows intermediate trends. These

42   systematics suggest that at high oxygen fugacity (Earth), V^{4+} follows Ti^{4+}, whereas at low

43   oxygen fugacity (Moon), V^{3+} follows Cr^{3+}.

44   These qualitative interpretations of vanadium behavior were subsequently investigated

45   using a series of experiments to better understand the potential of V behavior in spinels as an

46   oxybarometer (e.g. Papike et al. 2004, 2005, 2013, 2014). The experiments illustrated the

47   relationships between Fe^{2+}-Fe^{3+} and V^{3+}-V^{4+}, in which the vanadium in spinel, V^{3+} is best

48   incorporated in normal spinel as $^{IV}Fe^{2+VI}V^{3+}_2O_4$ (coulsonite) and V^{4+} as $^{IV}Fe^{2+VI}(Fe^{2+}V^{4+})O_4$

49   (synthetic compound, inverse spinel) (Rodgers et al. 1963).

50   The intent of this study is to explore the relationships between V partitioning behavior

51   and spinel crystal chemistry at magmatic conditions over several orders of magnitude in fO_2. In

52   particular, we examine (1) the interplay among normal-inverse spinel structure, Fe^{2+}-Fe^{3+}, and V

53   partitioning and (2) the importance of using this interaction for interpreting petrologic conditions

54   for the origin of extraterrestrial materials.

55   **CRYSTAL CHEMISTRY OF SPINEL**

56   The potential interplay between vanadium and iron valance states in the context of the

57   spinel crystal chemistry is partially illustrated in Figure 1. This figure shows the crystal structure

58   of both normal and inverse spinel. The normal spinel structure (Fig. 1a) is represented by

59   magnesiocoulsonite, $^{IV}Mg^{VI}V^{3+}_2O_4$ (see Papike et al. 2005 and references therein). In naturally

60   occurring normal spinel, the sixteen octahedral sites per unit cell (Papike et al. 2005) are

61   occupied by trivalent cations Cr, Al, V^{3+}, etc., while the eight tetrahedral sites per unit cell are

62   occupied mainly by Mg and Fe^{2+}. The inverse spinel structure (Fig. 1b) is represented by

63   magnetite, $^{IV}Fe^{3+\,VI}(Fe^{2+}Fe^{3+})O_4$. In naturally occurring inverse spinel, the sixteen octahedral

64   sites per unit cell are occupied by tetravalent, trivalent, and divalent cations, Ti^{4+}, Fe^{3+}, Cr^{3+},

65   Al^{3+}, V^{3+}, Fe^{2+}, while the eight tetrahedral sites per unit cell are occupied mainly by cations such

66   as Fe^{3+}. In Figure 1c, synthetic inverse vanadium spinel is represented by $^{IV}Fe^{2+\,VI}(Fe^{2+}V^{4+})O_4$

67   (Rodgers et al. 1963). This spinel is referred to as V-inverse in this paper. This synthetic

68   compound has the same structure as ulvöspinel, with V^{4+} taking the place of Ti^{4+}. In a study by

69   Wechsler et al. (1984), the authors point out that (based on neutron diffraction), all Ti is confined

70   to the sixteen octahedral sites per unit cell, none enters the eight tetrahedral sites. We assume that

71   the same holds true in our own experiments. Thus even in spinel at reducing conditions such as

72   IW-1, which is largely chromite with the normal spinel structure ^{IV}(Mg, Fe^{2+})^{VI}Cr$^{3+}_2$O$_4$, small

73   domains of the inverse structure must exist in order to accommodate Ti.

74   Spinels with chromite contents between 0.00 and 0.48 mole fraction are characterized by

75   domains that are dominantly "normal" spinel, while those with ülvospinel contents greater than

76   0.88 mole fraction ülvospinel contain domains that are dominantly "inverse." See Haggerty

77   (1972) for documentation of this miscibility gap. The result is that the lowest free energy

78   assemblage is a two-phase region of a chromite-rich spinel and an ülvospinel-rich spinel

79   coexisting across a miscibility gap. This explains the ubiquitous miscibility gap between spinels

80   with ülvospinel contents less than 0.48 mole fraction ülvospinel/(ülvospinel + chromite) and

81   those with greater than 0.88 mole fraction ülvospinel.

EXPERIMENTAL AND ANALYTICAL APPROACHES

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Experiments

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84   All of the Yamato 980459 (Y98) composition partitioning experiments were conducted in

85   the gas mixing laboratory at NASA Johnson Space Center. The majority of the experiments were

86   performed using Re wire loops to prevent Fe-loss. Experiments run at oxygen fugacities at or

87   above the FMQ buffer employed Pt/Rh-alloy wire loops. Experimental fO$_2$ was controlled using

88   CO/CO$_2$ gas mixtures that were calibrated with a Y-stabilized zirconia electrochemical oxygen

89   sensor, specifically the SIRO2 from COF-Australia, housed in a dedicated reference furnace.

90   Experiments were conducted at discrete fO$_2$ intervals separated by approximately 1 log unit. The

91   experimental fO$_2$ values ranged from ΔIW-1 to ΔIW+5.4. Experiments were conducted with an

92   initial, eight hour super-liquidus dwell at 1450°C. At the end of the super-liquidus soak, the

93 temperature was dropped at 0.25˚C sec^{-1} until reaching the final target temperature of 1200˚C.

94 Experiments were terminated by drop quenching the charge into water. The synthetic Y98

95 composition used in the experiments was made from a mixture of synthetic oxide powders that

96 were fused to form a homogenous glass. This glass was milled to a fine-grained powder prior to

97 being used in the experiments. The composition of the glass was patterned after the reported

98 composition of the bulk Y98 meteorite. Additionally, starting glass was also doped with 0.20 wt.

99 % V_2O_3, EuO_2 and CeO_2.

Electron microprobe (EPMA)

101   Analyses were collected on the JEOL JXA 8200 electron microprobe at the Institute of

102 Meteoritics (IOM) and Department of Earth and Planetary Sciences (E&PS), at the University of

103 New Mexico (UNM). Backscattered electron imaging and analyses were collected at an

104 accelerating voltage of 15 kV. Oxide phases were analyzed using a probe current of 20 nA, and a

105 spot size of 0-2 μm. Glass analyses were collected at a beam current of 10 nA, and a spot size of

106 5 μm to limit beam volatilization. Elements were calibrated using C.M. Taylor Co. EPMA

107 standards, as well as additional standards developed in-house; a separate calibration was used for

108 the oxides and the glass. Vanadium measurements were made under typical analytical

109 conditions, i.e., 40 second peak counting and 20 seconds on both background positions. The Ti

110 interference on V was specifically determined for our spectrometer configuration (with the Ti Kβ

111 peak representing a known interference for V Kα), as was the Cr Kβ interference on Mn Kα.

112 Ferric iron in the spinel was calculated using the methodology of Droop (1987).

113   Uncertainties associated with the calculated partition coefficient values (e.g., for

114 $D_V^{spinel/glass}$) were propagated using the 1σ standard deviations and the mean of the EPMA

115 determined vanadium analyses. Typically, the 1σ standard deviation for a given population of

116   vanadium analyses was greater than the 2σ analytical uncertainty derived from x-ray counting

117   statistics; therefore, we feel the use of the 1σ standard deviation is a more accurate portrayal of

118   the actual uncertainties that arise from subtle compositional inhomogeneity.

119   **RESULTS**

120   Major and minor element chemistry of both spinel and glass are presented in Table 1.

121   Analyses were collected in samples spanning the fO_2 range of IW-1 to FMQ+2. Analytical

122   conditions were tailored to the phases being analyzed, and are discussed in detail in the

123   experimental and analytical approach section above.

124   The major element chemistry of the spinels is generally defined by an increase in

125   magnetite component at the expense of the chromite and spinel components with increasing fO_2.

126   This can be seen in changes in the concentration of the elements that define those components.

127   The average aluminum content of the spinels drops from 8.62 wt.% Al_2O_3 at IW-1 to 5.96 wt.%

128   at FMQ; at FMQ+1 and FMQ+2, Al_2O_3 jumps up slightly to 7.30 and 6.95 wt.%, respectively.

129   Chromium initially increases from 46.06 wt.% Cr_2O_3 at IW-1 to 54.87 wt.% at FMQ, but

130   increasing fO_2 results in a drop in Cr_2O_3 to 39.83 and 40.42 wt.% at FMQ+1 and FMQ+2,

131   respectively.   The calculated Fe_2O_3 concentration increases from 0.94 wt.% to 8.12 wt.% at

132   FMQ; increasing the fO_2 results in further increases in Fe_2O_3, where at FMQ+1 and FMQ+2, it

133   was calculated to be 17.27 and 17.53 wt.%, respectively (Fig. 2a). Figure 2b shows the modeled

134   Cr^{3+}/ total Cr in glass (discussed more below). Figure 2c illustrates measured vanadium in spinel,

135   and Figure 2d presents modeled ferric iron in the liquid. All of these are discussed below.

136   Vanadium concentration in spinel initially shows a decrease with increasing fO_2. The V_2O_3

137   concentration of the spinel drops from 9.16 wt.% to 1.58 wt.% in the IW-1 and FMQ

138   experiments, respectively (Fig. 2c). Increasing fO_2 further results in a slight bump in V, to 2.40

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139   and 2.22 wt.% V_2O_3, at FMQ+1 and FMQ+2, respectively. Remaining major and minor element

140   chemistry in the spinels is detailed in Table 1.

141   Analyses collected in the glasses in the experiments indicate that Cr decreases in the melt

142   with increasing fO_2, while V increases (Table 1). The Cr_2O_3 concentration varies from 0.37

143   wt.% in the IW-1 experiment to 0.12 wt.% in the FMQ+2 experiment. The V_2O_3 concentration

144   increases from 0.20 wt.% in the IW-1 experiment, to a high of 0.33 wt.% in the FMQ+2

145   experiment. Partition coefficient values, calculated for V between spinel and melt ($D_V^{spinel/melt}$),

146   generally drop from the low fO_2 experiments through the more oxidized samples (Fig. 3, Table

147   1). The calculated values range from 45.2 at IW-1, down to 6.2 at FMQ. The experiments at

148   even higher fO_2 show some variability in the partition coefficient; $D_V^{spinel/melt}$ is calculated at 7.6

149   and 6.8 at FMQ+1 and FMQ+2, respectively.

150   **DISCUSSION**

151   **The Effect of fO_2 on Spinel Composition and the Compatibility of V^{3+} and V^{4+}**

152   Not only does fO_2 profoundly change the spinel-melt partitioning systematics of V, but it

153   also exerts a significant influence on the valence ratios of Fe^{3+} to Fe^{2+} and Cr^{3+} to Cr^{2+} in the

154   melt.   The availabilities of Cr^{3+}, Fe^{2+}, and Fe^{3+} in the melt should, in some way, also be

155   manifested in the equilibrium composition of the spinel.   Changes in the major element

156   composition of spinel have been indirectly observed to have important consequences for the

157   compatibility of V in the spinel structure (Canil 2002; Righter et al. 2006 and Mallmann and

158   O'Neill 2009). The underlying crystal chemical reasons for these effects are discussed in detail

159   in the following section. In this section we focus the discussion on how changes in Cr and Fe

160   valence influence both the composition and structure of the equilibrium spinels.

161   The spinel components have been plotted as a function of fO_2 in Figure 4. This figure

162   clearly illustrates two important points: (1) the major element composition of the spinel is

163   relatively constant as fO_2 is increased from IW-1 to IW+3.4 (i.e., FMQ), and (2) at fO_2 values

164   >IW+3.4, the composition of the spinel shows a significant increase in its magnetite component

165   accompanied by a reduction in the spinel and ferro-chromite components.

166   The availability of Cr^{3+} in the melt is an important parameter that plays a critical role in

167   defining the stability field of spinel with respect to both its position on the liquidus, as well as its

168   equilibrium composition, and therefore has important implications for V stability in spinel as

169   well. Several recent studies have demonstrated the Cr^{3+}/Cr^{total} (Fig. 2b) in basaltic melts changes

170   dramatically as fO_2 is increased from values near the IW buffer to values near the FMQ buffer

171   (Hanson and Jones 1998; Berry et al. 2006 and Bell et al. 2014). Curiously, the increase in

172   Cr^{3+}/Cr^{2+} of the melt does not appear to significantly increase the chromite content of the

173   experimental spinels. Importantly, this observation suggests that fO_2 has only a minor effect on

174   spinel composition in the fO_2 space where the concentration of ferric iron in the melt is minimal.

175   The increase in the Cr^{3+}/Cr^{2+} ratio that accompanies the oxidation of the melt appears to be

176   accommodated through the crystallization of additional Cr-rich spinel. This observation is

177   corroborated by the monotonic decrease in the Cr^{total} content of the melt that occurs with

178   increasing fO_2. In contrast, the effect of ferric iron on the composition of the spinel is relatively

179   pronounced. The increasing abundance of Fe^{3+} in the melt results in the spinel being steadily

180   enriched in the magnetite component and depleted in the $MgAl_2O_4$ component.

181   From a structural point of view, the observed change in spinel composition can be

182   understood as follows: at low oxygen fugacity, ~IW-1 to IW+3, the structure is that of normal

183   spinel, in which the sixteen octahedral sites per unit cell (Papike et al. 2005) are occupied by

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184   trivalent cations, Cr, Al, V^{3+} etc., while the eight tetrahedral sites per unit cell are occupied

185   mainly by Mg and Fe^{2+}. Spinel at IW-1 to IW+3 must be dominated by the normal spinel

186   structure, only containing isolated domains of the inverse structure to accommodate elements

187   such as Ti^{4+} and limited quantities of Fe^{3+}. However, in the oxygen fugacity range above IW+3,

188   the increasing Fe_2O_3 in the melt causes a significant shift in the spinel composition, where

189   domains of the inverse spinel structure (which have lower free energy than the normal spinel

190   structure) become increasingly prevalent.

191   The observed changes in the spinel composition have important implications for the

192   compatibility of V^{3+} and V^{4+} in the octahedral sites of the spinel. For example, V^{3+} is highly

193   compatible in Cr-rich normal spinel as it can readily substitute for Cr^{3+}. Normal spinels have

194   difficulty accommodating higher valence cations in the octahedral sites due to charge balance

195   issues, therefore small, highly charged cations such as Ti^{4+} and V^{4+} only partition weakly into

196   these sites. In contrast to this behavior, 4+ cations are much more compatible in the inverse

197   spinel structure where divalent cations occupy the tetrahedral sites in order to maintain charge

198   balance. The consequences of this change in spinel composition and structure will be explored

199   in detail in the next section of the work.

200   **The variation of $D_V^{spinel/glass}$ with Oxygen Fugacity**

201   The observed trend in the vanadium partition coefficients (Fig. 3a) must be interpreted

202   with respect to both the ratio of V^{3+} to V^{4+} in the melt, as well the changing major element

203   composition and structure of the spinel. It is represented schematically in Figure 5. In an attempt

204   to disentangle the crystal chemical effects on $D_V^{spinel/glass}$ from that of the changing V^{3+}/V^{4+} in the

205   melt, we have modeled the $D_V^{spinel/glass}$ as a function of fO_2 for the experimental spinels (Fig. 3b).

206   It is clear from this figure that increasing fO_2 depresses the $D_V^{spinel/glass}$ value.   The model

207   assumes that: 1) V^{3+} and V^{4+} are the two valence states present in the melt and 2) the $D_{V3+}^{spinel/melt}$

208   and $D_{V4+}^{spinel/melt}$ values are static (i.e., they are independent of the major element composition of the

209   the spinel and comply with Henry's Law).   The total partition coefficient for vanadium can

210   therefore be expressed as:

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212    $D_V^{spinel/melt} = (xV^{3+}_{melt} D^{V^{3+}}) + (xV^{4+}_{melt} D^{V^{4+}})$

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214   In the case of our experiments, the effects of fO_2 on the major element compositions of the

215   experimental spinels are minimal at fO_2 values lower than IW+2. Therefore, the use of static D

216   values in this fO_2 range is justified. However, calculating the $D_V^{spinel/melt\ (effective)}$ requires

217   knowledge of the molar concentrations of V^{3+} and V^{4+} in the melt. Previous spectroscopic studies

218   of V valence in silicate melts have convincingly demonstrated that V^{3+} comprises >80% of the

219   total V in melts of basaltic compositions (Righter et al. 2006 and Sutton et al. 2005) at reduced

220   conditions (i.e., fO_2 <IW).   Through trial and error, we found that modeling the $D_V^{spinel/melt}$ curve

221   using a V^{3+}/V^{total} value of 0.90 for the IW-1 experiment and $D_{V3+}^{spinel/melt} \sim 48$ and $D_{V4+}^{spinel/melt} =$

222   3 yielded the best match for the observed trend of our partitioning data.

223   This modeling also helps clarify the role of Fe^{3+} on the inverse component of the spinel.

224   In the most oxidized experiments, the model curve shows a significant divergence from the

225   measured $D_V^{spinel/melt}$ curve. The divergence of the model curve and measured $D_V^{spinel/melt}$ values

226   can be explained by an increase in the compatibility of V^{4+} in the spinel structure (Papike et al.

227   2004, 2005), as it becomes more inverse. This change in structure is facilitated by the increasing

228   concentration of Fe^{3+} in both the spinel (as magnetite) and the melt. Furthermore, the modeling

229   clearly shows that in the spinels that contain little Fe^{3+} and have a constant Cr/Al, the $D_V^{spinel/melt}$

230   can be modeled as a simple function of the V^{3+}/V^{4+} present in the melt. This simplified approach,

231   however, cannot be used for spinels that have a substantial portion of inverse domains. In the

232   case of our experiments, the rise of the inverse structure coincides with rise of Fe^{3+} in the melt

233   with increasing fO_2. Although not explicitly addressed by our experiments, this effect is also

234   important for spinels enriched in the ülvospinel component.

Spinels as indicators of planetary parentage and redox histories

236   This manuscript illustrates the interplay among multivalent cations ($Fe^{2+,3+}$, $Cr^{2+,3+}$,

237   $V^{3+,4+}$), the spinel structure, and oxygen fugacity in determining the partitioning behavior of

238   vanadium. Understanding this interplay is critical for using spinels as both indicators of planetary

239   parentage and reconstructing the redox history of magmatic systems on the terrestrial planets.

240   Three examples for this use are given below.

241   Example 1. Papike et al. (2005) and Karner et al. (2003) illustrated the usefulness of

242   silicates (olivine, pyroxene, and plagioclase) for determining planetary parentage of

243   extraterrestrial materials. These approaches have been used in countless studies of meteorites

244   (e.g. Fagan et al. 2002; Agee et al. 2013; Elardo et al. 2013). Papike et al. (2004, 2013)

245   illustrated the qualitative usefulness in relating spinel chemistry to both a parent body source and

246   the fO_2 of a planetary body mantle and its associated basalt systems. As discussed in those

247   papers, differences in fO_2 among basaltic systems associated with the Earth's moon and

248   primitive, differentiated small bodies such as the eucrite parent body (~IW-1), the Earth

249   (FMQ±1), and Mars (IW+1 to FMQ) have profound implications for the valence state of multi-

250   valent cations and the spinel structure. The planetary environments that are reduced will be

251   dominated by normal spinels and the reduced species of multi-valent cations. More oxidizing

252   environments, such as the Earth (FMQ±1), and Mars (IW+1 to FMQ), will be dominated by the

253   transition between normal and inverse spinels and the more oxidized species of multi-valence

254   cations. These differences strongly influence partitioning behavior during mantle melting and

255   subsequent crystallization, and thereby result in differences in spinel compositions.

256   Example 2. A prominent geochemical feature of the shergottites is the large range in

257   initial Sr isotopic ratios (~0.702 – 0.724) and initial ε^{Nd} values (~ -10 to +50) (e.g., Nyquist et al.

258   2006; Borg et al. 2012; Brandon et al. 2012; Shearer et al. 2008; Symes et al. 2008). Within this

259   range of isotopic characteristics, the shergottites also exhibit a systematic variation in the oxygen

260   fugacity of crystallization (Wadhwa 2001; Herd 2006; Borg et al. 2003, 2005; Borg and Drake

261   2005; Shearer et al. 2008; Symes et al. 2008). To decipher the petrogenesis of the shergottites, a

262   fuller understanding of the petrologic linkages between these geochemical characteristics and

263   intensive parameters (fO_2) must be reached. One approach to deciphering these petrogenetic

264   linkages is to examine the crystallization history preserved in spinel compositions of a single

265   magma and compare the history preserved by spinel populations in different magmas.

266   Spinels in the shergottites exhibit substantial compositional differences among early

267   spinels from different shergottites and along crystallization trajectories within individual

268   shergottites (e.g. Goodrich et al. 2003; McKay et al. 2004a,b; Peslier et al. 2010; Gross et al.

269   2011). Early spinels in the isotopically depleted shergottites such as Y980459, NWA 5789, SaU

270   005, and Dhofar 019 have relatively low magnetite (1 to 2%) and inverse spinel components

271   (magnetite + ülvospinel ~ 4 to 5%). Therefore, fO_2 estimates from V partitioning in these

272   lithologies should be relatively unaffected by the presence of an inverse structural component.

273   Comparison with our data suggests an fO_2 of crystallization for Y980459 of IW+1.

274   On the other hand, the more isotopically enriched shergottites such as NWA 1110 and

275   EETA 79001 have early spinel with higher magnetite (3 to 6%) and inverse spinel components

276   (magnetite + ülvospinel ~ 5 to 7%). Based on comparison with our data, the fO_2 of crystallization

277   for the enriched shergottites is more oxidizing than IW+2. All of the early spinels have similar

278   chromite-ülvospinel-spinel components and overlapping V_2O_3. This suggests similar bulk

279    $D_V^{spinel/melt}$ and slightly different bulk V in the magma. Examining the spinel crystallization

280   trajectories, the magnetite and inverse spinel components increase only slightly, with the

281   magnetite component increasing to 2-5% and the inverse spinel component increasing (magnetite

282   + ülvospinel ~ 8 to 15%). Also, there is an overall increase in the V component in these spinels.

283   Based on our results, this suggests only a subtle increase in fO_2 during crystallization.

284   In contrast, several of the depleted olivine-phyric shergottites exhibit a substantial

285   increase in both the magnetite (13-15%) and the inverse spinel components (magnetite +

286   ülvospinel ~ 18 to 20%). These data suggest that the fO_2 becomes increasingly oxidized during

287   crystallization. Very little V data exists for these spinels, but based our analysis would predict an

288   overall decrease in bulk $D_V^{spinel/melt}$ and therefore a decrease in the V content of the late spinels.

289   These initial observations are only the beginning. Combining the experimental data and

290   interpretations with a much more detailed analytical dataset of spinels from the shergottites

291   would yield a much more profound understanding of the fO_2 record preserved in martian spinels.

292   Example 3. The differentiation of asteroids is one of the most fundamental processes to

293   have occurred in the early Solar System. There are numerous families of achondrite meteorites

294   that represent partial melting of these small parent bodies (e.g. eucrites, diogenites, acapulcoites,

295   lodranites, brachinites, and unclassified Na-rich achondrites such as GRA 06129/06128). Several

296   of these achondrites represent magmatism on parent bodies with a well-defined fO_2 (e.g. eucrites

297   and diogenites ~ IW-1). Other achondrites are derived from parent bodies with fO_2 that are not

298   well-defined (e.g. brachinites, unclassified Na-rich achondrites such as GRA 06129/06128). As

299   has been demonstrated, differences in the fO_2 (IW-1 to IW+2) during partial melting on these

300   parent bodies will produce considerably different melt compositions (e.g. Jurewicz et al. 1995;

301   Garder-Vandy et al. 2013; Usui et al. 2014). Several studies have linked brachinites to unique

302   sodium feldspar-rich achondrite GRA 06129/06128 (e.g. Shearer et al. 2010; Usui et al. 2014).

303   Usui et al. (2014) illustrated that partial melting of a chondritic precursor would yield a

304   brachinite residuum and a GRA 06129 melt, under specific fO_2 conditions. As both lithologies

305   contain Cr-spinels, examining the systematic relationship between their chemistries and fO_2

306   would allow for an analysis of conditions of melting and crystallization. The spinels in

307   brachinites have a Cr/(Cr+Al) of 0.73 to 0.83. The extent of the magnetite component is debated,

308   but is thought to be approximately 1 to 3%. Complicating the interpretation of this data is the

309   subsolidus reequilibration experienced by both of these primitive lithologies.

310   This work demonstrates the nuances of the V partitioning and provides a crystal chemical

311   basis for understanding Fe^{3+}, Cr, and V substitution into the octahedral sites of spinel. The above

312   examples illustrate the importance of understanding this interplay for using spinels as both

313   indicators of planetary parentage and reconstructing the redox history of magmatic systems on

314   the terrestrial planets.

315   **IMPLICATIONS**

316   We seek to understand the origin and evolution of the solar system. There are many ways

317   to do this: theory, experimental approaches and space missions with orbital and landed

318   experiments. Arguably, the most powerful approach is to have a sample from the specific body

319   we want to understand. In some cases this is not currently possible given the costs or the lack of

320   critical technology to enable such missions (e.g., obtaining samples from Europa or the center of

321   Jupiter). The book "Basaltic volcanism in the terrestrial planets" (Basaltic Volcanism Study

322 Project, 1981) was an early attempt to address the issue of understanding processes in the solar

323 system as recorded by basalts. A team led by the first author (JJP) participated in this exciting

324 study by pulling together information on the silicates from planetary basalts. In the present study

325 we follow giants in the field (Haggerty 1976; Lindsley 1976, 1991; Basaltic Volcanism Study

326 Project, 1981) including S.E. Haggerty (Haggerty 1976; his contribution to Basaltic Volcanism

327 Study Project, 1981) who was a leader in the study of spinels. We hope our work, both preceding

328 and specifically in this paper, will make the use of spinel chemistry, crystal chemistry, and

329 valence state partitioning even more powerful.

330   The use of a single zoned spinel from (normal spinel, chromite) to (inverse spinel

331 ülvospinel-magnetite) can, with use of the ratios of Fe^{2+}, Fe^{3+}, Cr^{2+}, Cr^{3+}, $V^{2+,3+,4+,5+}$ compared to

332 the other elements in the core to rim zoning trend of a single spinel crystal, identify the planetary

333 parentage of the grain and (if the melt composition is known) the quantitative oxygen fugacity

334 trajectory during crystallization of the host basalt.

335

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524   **FIGURE CAPTIONS**

525   Figure 1. Spinel structural diagrams. a. The normal spinel structure represents

526   magnesiocoulsonite, $^{IV}Mg^{VI}V^{3+}_2O_4$ (see Papike et al. 2005 and references therein). b. The inverse

527   spinel structure (magnetite $^{IV}Fe^{3+} {}^{VI}(Fe^{2+}Fe^{3+})O_4$). c. The V-inverse synthetic spinel structure

528   ($^{IV}Fe^{2+} {}^{VI}(Fe^{2+}V^{4+})O_4$). The structures are color coded as follows: Light blue represents oxygen

529   ligands in all structures. Magnesiocoulsonite has yellow representing both Mg^{2+}-containing

530   tetrahedra and spheres. Red represents V^{3+} and V^{4+} octahedra and spheres. Magnetite has purple

531   representing Fe^{3+} tetrahedra, octahedra, and spheres, and green representing Fe^{2+}. V-inverse has

532   green representing Fe (both ferrous and ferric) and red representing V^{4+}. As the structure

533   evolves, from (a) normal spinel/ magnesiocoulsonite to (b) magnetite and from (a) to (c), with

534   increasing oxygen fugacity, V^{3+} in normal spinel is replaced by V^{4+} in inverse spinel and by Fe^{2+}

535   and Fe^{3+} in magnetite (an inverse spinel).

536

537   Figure 2. a. Shows the variation of Fe_2O_3 in spinel (wt.%). b. Shows modeled Cr^{3+}/total Cr in

538   melt. Error bars represent the 1σ standard deviation. c. Shows V_2O_3 in spinel (wt. %). d. Shows

539   modeled Fe_2O_3 in the liquid (wt. %). The modeled Cr valence in the melt was produced with the

540   data and methods described in Bell et al. (2014). The ferric/ferrous ratios of the melts were

541   modeled using the Kress and Carmichael (1991) equation and the EPMA determined melt

542   composition.

543

544   Figure 3. a. $D_V^{spinel\,/melt}$ (partition coefficient). Error bars represent the 1σ propagated error. b.

545   Measured $D_V^{spinel\,/melt}$, as in Figure 3a., with modeled partitioning curve overlain.

546

547   Figure 4. Spinel components chromite, magnetite, and other normal spinel, as a function of fO_2.

548

549   Figure 5. A cartoon that illustrates the combined effects of vanadium valence in the melt and the

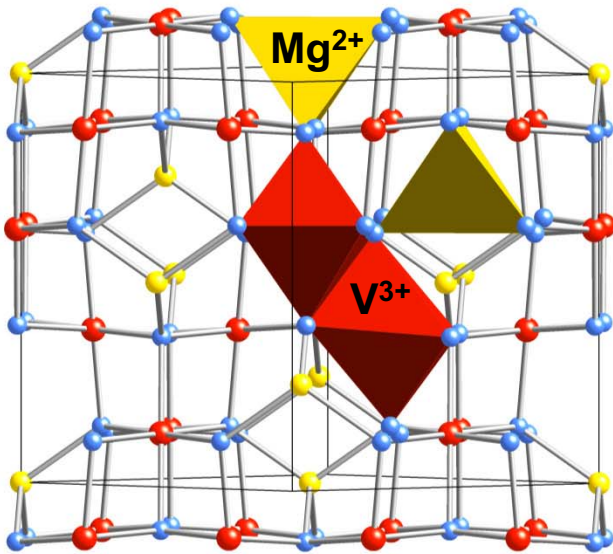550   increase in the magnetite component of the spinel.

Table 1. Experimental average values. Standard deviation (1σ) shown in parantheses.

fO_2:	IW-1		IW		IW+1		IW+2.2		FMQ		FMQ+1		FMQ+2	
Sample ID:	Y98C-21		Y98C-15		Y98C-10		Y98C-40		Y98C-19		Y98C-29		Y98C-31	
Oxide (Wt.%)														
SiO_2	0.26	(0.03)	0.18	(0.02)	0.26	(0.03)	0.16	(0.03)	0.15	(0.02)	0.17	(0.03)	0.17	(0.03)
Al_2O_3	8.62	(1.94)	8.40	(1.12)	8.14	(0.78)	6.74	(0.21)	5.96	(0.12)	7.30	(0.05)	6.95	(0.11)
$Fe_2O_3{}^*$	0.94	(0.25)	2.35	(0.24)	3.59	(0.41)	6.30	(0.33)	8.12	(0.25)	17.27	(0.29)	17.53	(0.39)
TiO_2	2.79	(0.21)	2.37	(0.25)	2.04	(0.05)	1.42	(0.14)	0.90	(0.04)	1.89	(0.02)	1.72	(0.07)
Cr_2O_3	46.06	(1.94)	48.43	(1.74)	49.08	(1.81)	51.68	(0.63)	54.87	(0.42)	39.83	(0.11)	40.42	(0.76)
MgO	6.76	(0.24)	6.67	(0.21)	6.97	(0.18)	6.66	(0.11)	9.44	(0.08)	7.45	(0.04)	7.39	(0.07)
FeO	25.64	(0.31)	25.26	(0.25)	24.50	(0.23)	23.82	(0.12)	19.17	(0.13)	22.79	(0.06)	22.65	(0.19)
MnO	0.60	(0.02)	0.57	(0.01)	0.58	(0.01)	0.60	(0.02)	0.53	(0.02)	0.60	(0.02)	0.62	(0.02)
V_2O_3	9.16	(0.44)	6.56	(0.40)	5.58	(0.23)	2.80	(0.53)	1.58	(0.05)	2.40	(0.01)	2.22	(0.07)
Total	100.83		100.78		100.74		100.19		100.72		99.69		99.67	
A.F.U.														
Si	0.01	(0.00)	0.01	(0.00)	0.01	(0.00)	0.01	(0.00)	0.01	(0.00)	0.01	(0.00)	0.01	(0.00)
Al	0.34	(0.07)	0.33	(0.04)	0.32	(0.03)	0.27	(0.01)	0.24	(0.00)	0.30	(0.00)	0.28	(0.01)
Fe^{3+*}	0.02	(0.01)	0.06	(0.01)	0.09	(0.01)	0.16	(0.01)	0.21	(0.01)	0.45	(0.01)	0.45	(0.01)
Ti	0.07	(0.01)	0.06	(0.01)	0.05	(0.00)	0.04	(0.00)	0.02	(0.00)	0.05	(0.00)	0.04	(0.00)
Cr	1.23	(0.06)	1.29	(0.05)	1.31	(0.05)	1.40	(0.02)	1.46	(0.01)	1.08	(0.00)	1.10	(0.02)
Mg	0.34	(0.01)	0.34	(0.01)	0.35	(0.01)	0.34	(0.01)	0.47	(0.00)	0.38	(0.00)	0.38	(0.00)
Fe^{2+}	0.72	(0.01)	0.71	(0.01)	0.69	(0.01)	0.68	(0.00)	0.54	(0.00)	0.66	(0.00)	0.65	(0.00)
Mn	0.02	(0.00)	0.02	(0.00)	0.02	(0.00)	0.02	(0.00)	0.02	(0.00)	0.02	(0.00)	0.02	(0.00)
V	0.25	(0.01)	0.18	(0.01)	0.15	(0.01)	0.08	(0.01)	0.04	(0.00)	0.07	(0.00)	0.06	(0.00)
Total	3.00		3.00		3.00		3.00		3.00		3.00		3.00	
Components (Mol.%)														
Chromite	0.77		0.77		0.76		0.76		0.77		0.59		0.60	
Spinel	0.21		0.20		0.19		0.15		0.12		0.16		0.15	
Magnetite	0.01		0.04		0.05		0.09		0.11		0.24		0.25	
Glass (Wt.%)														
Cr_2O_3	0.37	(0.02)	0.22	(0.03)	0.20	(0.02)	0.13	(0.02)	0.13	(0.04)	0.12	(0.02)	0.12	(0.02)
V_2O_3	0.20	(0.02)	0.18	(0.02)	0.21	(0.01)	0.23	(0.02)	0.26	(0.03)	0.32	(0.02)	0.33	(0.02)
Partition Coefficient														
$D_V^{spinel/melt}$	45.19		36.75		26.52		12.37		6.19		7.55		6.77	

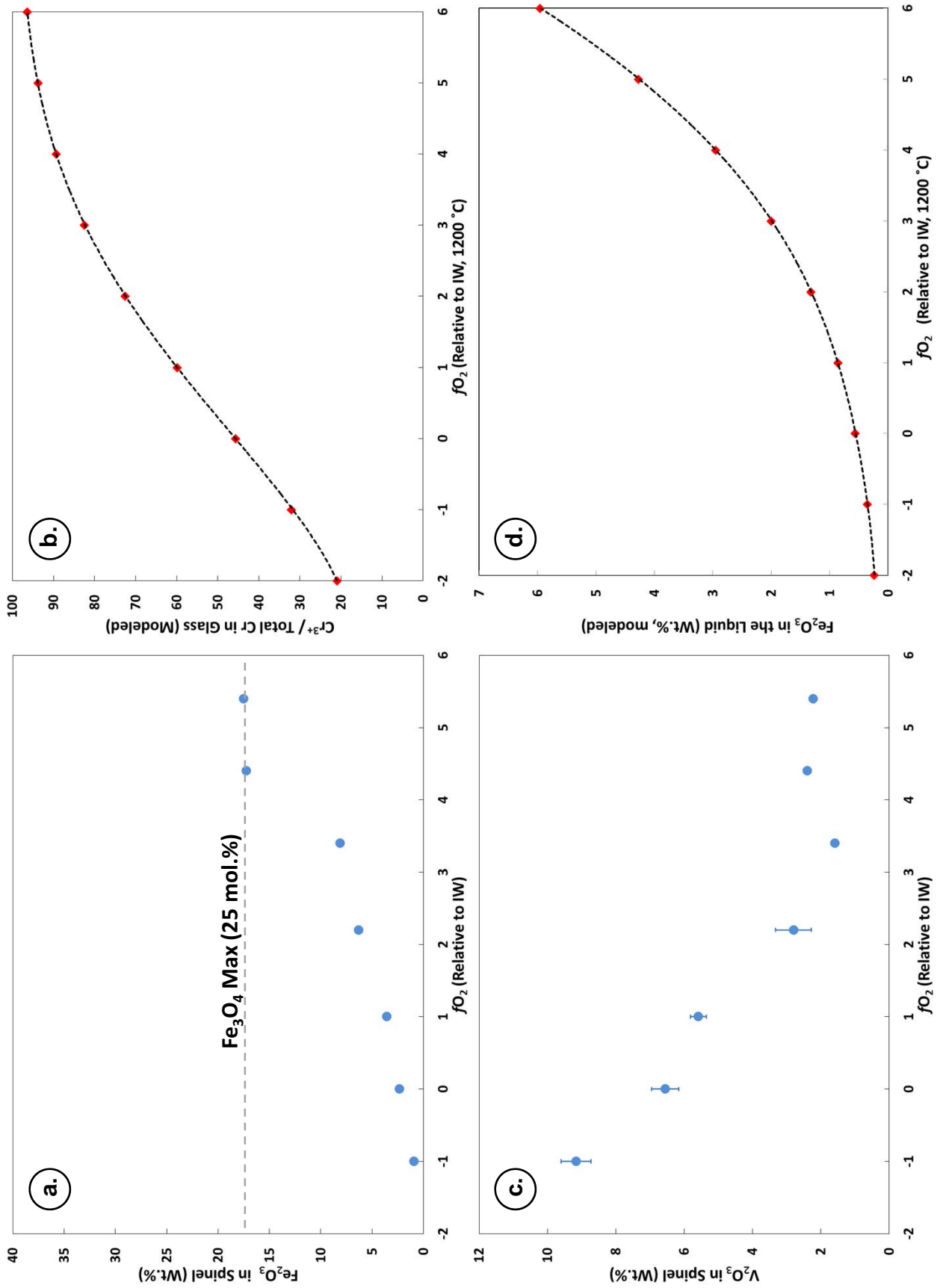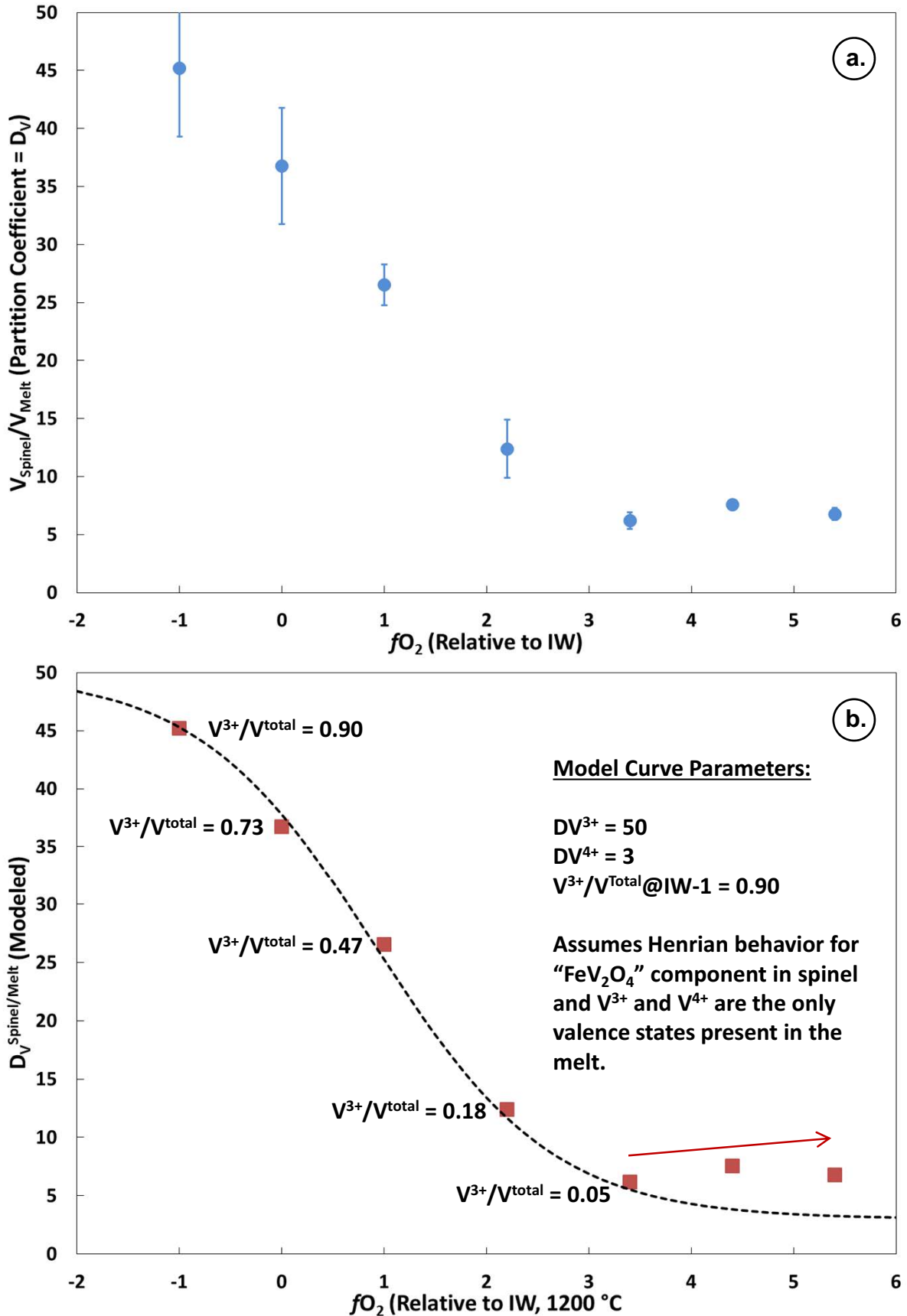*Estimated using the methodology of Droop (1987).

Figure 1.
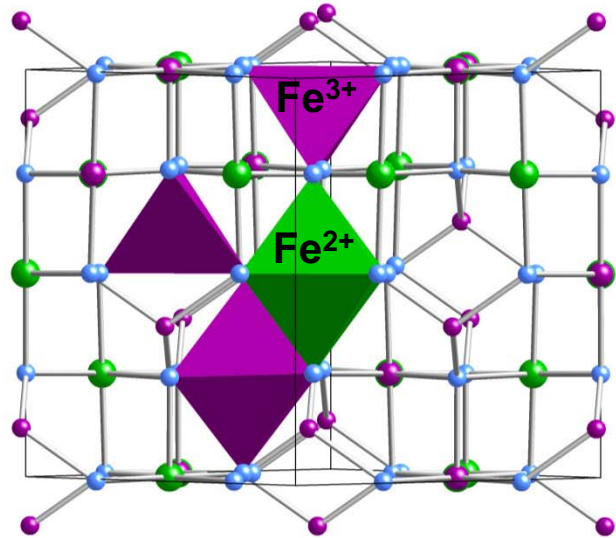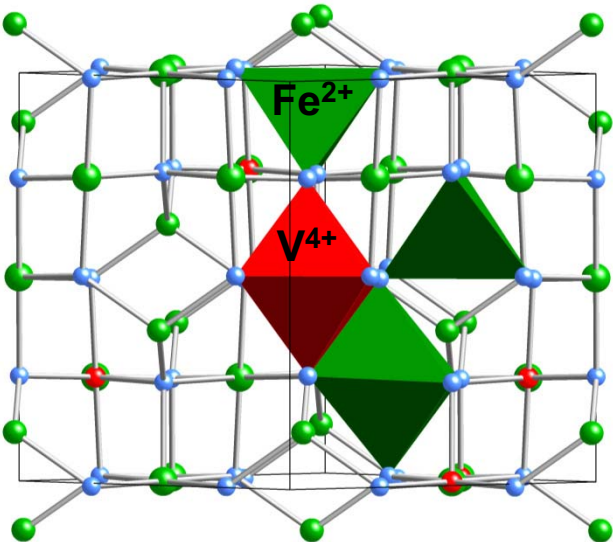
Figure 2.

Figure 3.

Figure 4.

Figure 5.

Increasing V^{4+} in melt (decreasing V^{3+}/V^{4+})

$D_V^{eff} = (xV^{3+} \text{ melt } DV^{3+}) + (xV^{4+} \text{ melt } DV^{4+})$

"Spinel" Contains Ferric Iron ~
D_V^{4+} increases & D_V^{3+} decreases with
Fe_3O_4 content of spinel
(i.e. the D values, themselves, are a
function the system fO_2)

Fe-Free "Spinels"~
D_V^{4+} & D_V^{3+} are constant
and $D_V^{4+} \ll D_V^{3+}$

log fO_2

D_V^{eff} (spinel-melt)