1	Grain size measurement from two-dimensional micro-X-ray diffraction: Laboratory
2	application of a radial integration technique
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24	Abstract
25	Two-dimensional X-ray diffraction data contain information about not only the
26	type of mineral phases present in an assemblage, but also the textural or grain size
27	relationships between minerals in a sample. For minerals within a certain grain size
28	range, ~0.1 to 100 $\mu$ m, the appearance and characteristics of a Debye ring can reveal the
29	mean grain size of a sample. In this contribution, using mineral and rock samples of
30	known grain size ranges, we investigate the applicability of calculating the grain size of a
31	material using a two-dimensional X-ray diffraction crystallite size analysis method for
32	micron-sized materials. A radial integration technique was used to derive the number of
33	grains contributing to produce diffraction spots in the Debye ring. Monomineralic
34	pyroxene and magnetite samples of known grain size ranges were analyzed, and the
35	calculated grain size was observed to broadly correlate with the sample size except at the
36	upper and lower extremes. To evaluate the technique on broader geological materials,
37	polymineralic basalt samples with known grain size ranges were analyzed, and the
38	calculated grain sizes did not correlate with the size of the rock fragments, but did
39	correlate closely with the size of the individual mineral grains. Using a Bruker D8
40	Discover X-ray diffractometer with a 300 $\mu$ m nominal incident beam diameter, the
41	effectiveness of the applied method appeared limited to the grain size range of $\sim 15-63$
42	$\mu$ m for monomineralic samples. The method is further limited by several complicating
43	factors and assumptions, including the requirement for the crystallite size to correlate
44	with the sample grain size. The effective range of this method will vary with different
45	instrumental and experimental conditions. When applying this method to calculate the

46	grain size of geological materials, the calculated result should be interpreted as a
47	minimum estimate of the grain size.
48	
49	Keywords: micro-X-ray diffraction, two-dimensional X-ray diffraction, grain size,
50	crystallite size, chi-profile, gamma-profile.
51	
52	Introduction
53	
54	Throughout the century-long history of X-ray diffraction, methods have been
55	developed and applied to measure the size distribution of crystalline materials with two-
56	dimensional X-ray diffraction (2D XRD) images by studying the characteristics of
57	diffraction spots on the images and their relationship within a Debye ring. Deciphering
58	the grain size relationships with the progression of smooth Debye rings to 'spotty' rings,
59	and finally to large diffraction spots as the effective grain sizes of micron-sized minerals
60	increased was pursued in two manners: (1) qualitative description of the Debye ring
61	characteristics of minerals of known grain sizes with broad qualitative application to
62	other minerals; and (2) more quantitative attempts to measure parameters from 2D XRD
63	images and calculate a given grain size with some accuracy.
64	The qualitative method of grain size identification can be seen in the study of
65	Debye-Scherrer X-ray diffraction film characteristics by several authors (e.g. Azároff and
66	Buerger 1958; Klug and Alexander 1974; Cullity 1978) who presented observations of
67	the visual qualities of the Debye rings of samples with known grain size to which
68	samples with unknown grain size could then be compared. These observations can be

69	collectively summarized as follows: Below $\sim 0.1 \ \mu m$ , Debye rings will display line
70	broadening, and the lines will broaden with decreasing grain size until $\sim 0.01 \ \mu m$ where
71	the irradiated sample will begin a transition towards being X-ray amorphous. In the size
72	range of 0.1 to 10 $\mu$ m, a "perfect" powder X-ray diffraction pattern with thin, clearly
73	discernible rings will exist, although there is not complete agreement on the exact
74	transitions zones. Cullity (1978) stated that the transition from continuous rings without
75	spots to spotty diffraction rings occurs between 1 $\mu$ m and 10 $\mu$ m, whereas Azároff and
76	Buerger (1958) state that between 10 $\mu$ m to 40 $\mu$ m the sample has clearly discernable
77	diffraction rings consisting of very many spots that are closely spaced. Klug and
78	Alexander (1974) place continuous rings at $<5 \ \mu m$ and spotty rings at 15 to 50 $\mu m$ for
79	quartz. Beyond $\sim 50 \ \mu m$ Debye rings become progressively more discontinuous, and by
80	${\sim}200~\mu m$ or larger only a few diffraction spots are scattered on the film. Hörz and Quaide
81	(1973) give a summary of Debye ring characteristics pertaining to the grain block size in
82	several minerals.

83 In the finer grain size range where the Debye rings begin to broaden in the 84 transition towards becoming X-ray amorphous, well-established quantitative methods of 85 measuring the grain size of powdered materials exist (e.g. Klug and Alexander 1974; Rao and Houska 1986). These methods use data from one-dimensional diffractograms and 86 87 derive the grain size from equations such as the Scherrer equation. With decreasing grain 88 size, a nominally sharp diffractogram peak will begin to broaden at the base, and then 89 broaden uniformly throughout (Azároff and Buerger 1958). With extensive broadening 90 the peak height will decrease as well, and the area under the peak will remain constant (Azároff and Buerger 1958). Measuring these features and entering the results into the 91

92	Scherrer equation allows for the grain size to be inferred. The line profile method is
93	particularly applicable when the grain size is below approximately 0.1 $\mu$ m (He 2009).
94	Williamson and Hall (1953) combined equations for size and strain. The Rietveld
95	whole pattern crystal structure refinement method (Rietveld 1969) also includes grain
96	size and strain estimation (e.g. Balzar et al. 2004), and can refine on the crystal size of a
97	sample in a method akin to the line profile method. Therefore, calculating grain size via
98	the Rietveld method is applicable only for samples with grain size on the order of a few
99	microns or smaller. Methods of crystallite size-lattice strain estimation from powder
100	diffraction pattern line shapes were reviewed by Mittemeijer and Welzel (2008). One-
101	dimensional powder diffraction methods were reviewed by Lavina et al. (2014). In the
102	size range where the progression of changing Debye ring characteristics are seen, $\sim 0.1$ to
103	100 $\mu$ m, the above methods are not applicable and other methods need to be applied.
104	Quantitative methods of deriving grain size from 2D XRD involve measuring
105	parameters from 2D images and inputting these values into equations to calculate the
106	grain size of the sample. Interestingly, in addition to their descriptions of Debye ring
107	characteristics pertaining to grain size, Azároff and Buerger (1958) and Cullity (1978)
108	both outline how a quantitative method for measuring grain size from these images could
109	be attained. Their proposed methods were akin to the method applied in this study, but it
110	likely took the advent of precise computer analysis of digital micro-X-ray diffraction
111	$(\mu XRD)$ images for the method to become a feasible reality.
112	Early quantitative methods of grain size measurement from 2D XRD are
113	exemplified by Stephen and Barnes (1937). In applying a technique modified from
114	Shdanow (1935), they measured the grain size of materials by the comparison of the

115	number of diffraction spots on two photographs taken with differing exposure times, or
116	by counting the number of spots on one film taken under a standardized condition and
117	comparing the result to an empirical chart. The calculation method was applied to six
118	aluminum samples, and was shown to not be effective above 50 $\mu$ m, but the region from
119	50 to 10 $\mu$ m was shown to be useful, where results were found to agree within 10%.
120	Hirsch and Kellar (1952) applied a more generalized version of the method developed by
121	Stephen and Barnes (1937) for back-reflection patterns. Hirsch and Kellar (1952) stated
122	that the method used in the study was more appropriate than the line profile method for
123	materials which may have experienced strain, such as cold-rolled aluminum, because the
124	textural information which coincides with the 2D XRD images allows for the calculated
125	results to unequivocally be interpreted as grain size. The line profile method would
126	require the assumption that the line broadening was solely the result of the grain size, a
127	risky assumption when studying strained and deformed materials which can generate
128	their own broadening features.
120	He (2009) formulated a modern method of calculating crystallite size from 2D

He (2009) formulated a modern method of calculating crystallite size from 2D 129 130 XRD using multi-wire detectors and computer software, but also featuring many parallels 131 with these earlier film-based techniques. Essentially, these methods calculate the volume 132 of material irradiated by the X-ray beam using constraints such as the diameter of the X-133 ray beam and the X-ray absorption characteristics of the material, and divide this volume by the number of irradiated crystallites that were calculated in some manner to arrive at a 134 calculated grain size. Here we present a laboratory investigation of the method of He 135 136 (2009), applied with a micro-X-ray diffractometer with reflection-mode geometry, with the aims of exploring the applicability of this method in calculating the grain size of 137

138	geological materials in an X-ray diffraction laboratory, and constraining the degree to
139	which differences in crystallite and grain size affect the method.
140	
141	Methods and Materials
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143	Grain size from 2D XRD after He (2009)
144	He (2009) formulated a method of crystallite size analysis that takes advantage of
145	2D XRD. This method, in its simplest form, relates the number of diffraction spots in a
146	spotty diffraction ring and the sample volume to the size of the crystallites. On the image
147	of the 2D XRD detector, a 2 $\theta$ by $\chi$ angular window is selected for a particular Debye ring
148	and is integrated along the $\chi$ direction. This produces an intensity versus $\chi$ plot which can
149	then be fitted with a polynomial or line of the average intensity. Half the number of times
150	the integrated profile crosses the plotted line is calculated as the number of crystallites
151	diffracting in the window. When this number is related to the sample volume, the
152	multiplicity of the diffracting plane, and the instrument parameters, the crystallite size
153	can be calculated. This $\chi$ -profile analysis method is useful when applied to samples with
154	crystallite sizes ranging from 0.1 $\mu$ m to 100 $\mu$ m, over the size range where Debye ring
155	characteristics rapidly evolve with changing crystallite size. Note that He (2009)
156	denominates this method as $\gamma$ -profile analysis, but here we have used the term $\chi$ to
157	correlate with the previous 2D XRD literature. The derivation of the crystallite size
158	measurement that follows is taken from He (2009), and the reader is referred there for a
159	thorough discussion.

The number of crystallites  $(N_s)$  contributing to a diffraction ring in a perfectly

160

161random powder can be given by162
$$N_s = p_{hkl} \cdot \frac{\nabla f_1}{v_l} \cdot \frac{\alpha}{4\pi}$$
 (1)163where the multiplicity of the diffraction ring is  $p_{hkl}$ ,  $V$  is the effective sampling volume,  $f_l$ 164is the volume fraction of the crystallites being measured,  $v_l$  is the volume of the165individual crystallites, and  $\Omega$  is the instrument angular window. Including the multiplicity166of the diffracting planes in the calculation allows for the mineral symmetry to be167considered. When counting the number of crystallites contributing to a given diffraction168ring, the multiplicity associated with the  $hkl$  index of the diffracting grains. The170instrumental angular window can be expressed as171 $\Omega = \beta_1 \beta_2 = 2\beta \arcsin[\cos \theta \sin(\Delta \chi/2)]$  (2)172where  $\beta_l$  and  $\beta_2$  are the instrument angular window in the 20 and  $\chi$  directions,173respectively, and is related to the integration area on the 2D detector. The diffraction174vector angular range corresponding to the azimuthal angular range can be expressed as175 $\Delta_X$ ,  $\beta_l$  can equal  $\beta$  for the window in the 20 direction when instrumental broadening of the176detector is neglected.177For the effective volume analyzed, several additional parameters need to be178introduced. When data are collected with a coupled scan where  $\theta_1$  equals  $\theta_2$  (see below),179the effective volume can be reduced to180 $V = \frac{4_0}{2\mu}$ , (3)181where  $\mu$  is the linear absorption coefficient (see below).

182 Substituting Equations 2 and 3 into Equation 1, the crystallite volume of the *i*th 183 phase can be expressed as 184  $v_i = \frac{p_{hkl}f_iA_0\beta \arcsin[\cos\theta\sin(\Delta\chi/2)]}{4\pi\mu N_s}$ . (4)

185 Using the diameter of the crystallites (*d*),

186 
$$v_i = \frac{\pi d_i^3}{6},$$
 (5)

187 to replace  $v_i$ , the crystallite size can then be expressed by

188 
$$d = \left\{ \frac{3p_{hkl}f_i A_0 \beta \arcsin[\cos\theta\sin(\Delta\chi/2)]}{2\pi^2 \mu N_s} \right\}^{1/3}.$$
 (6)

189 This formulation of the  $\chi$ -profile crystallite size analysis was applied in this study. He

190 (2009) also presents a version for transmission-mode geometry, as well as an alternate

191 form of each employing a scaling factor that removes all of the numeric constants,

192 simplifying the equation. This scaling factor can then be used as a calibration factor that

193 can be set using 2D XRD data from a known standard.

194 The laboratory investigation herein explores the applicability of this method for measuring the grain size of a variety of geological samples typically analyzed in an X-ray 195 196 diffraction laboratory. In this contribution, we apply the term grain size to identify the 197 mean of the distribution of crystallite sizes in a geological sample, i.e. a powder, polished 198 section, or hand sample. The use of the term grain size is somewhat varied in the 199 geological sciences and can be used to describe many physical properties such as the size 200 of individual mineral grains or crystals in coarse-grained igneous and metamorphic rocks 201 as well as the size of rock particles consolidated into a fine-gained sedimentary rock. For 202 the geological samples examined by  $\mu$ XRD in this study, we are applying a definition 203 akin to the former, meaning that the mean size of mineral crystals in a sample will

204	approximately correlate with the mean grain size of a sample. However, in fine-grained
205	rocks, the crystallite size, as measured by X-ray diffraction, will be smaller than the
206	particle size of the rock. Thus, the application of this crystallite size calculation method
207	should provide a minimum estimate of the grain size for all of these geological sample
208	types, when this definition and other assumptions discussed below, are taken into
209	consideration.
210	
211	Micro-X-ray diffraction
212	The Bruker D8 Discover at the University of Western Ontario was used for this
213	investigation, having theta-theta geometry, operating at 35 kV and 45 mA with a
214	radiation source of CoKa (1.79026 Å), and a Göbel mirror with a 300 $\mu$ m pinhole
215	collimator. A HI-STAR detector with General Area Detector Diffraction System
216	(GADDS; Bruker-AXS 2010) software was used. The sample-to-detector distance used
217	was 12 cm. Applications of micro-X-ray diffraction ( $\mu$ XRD) in the geological sciences
218	have shown it to be an effective technique for analyzing mineral textural information
219	(Flemming 2007; Izawa et al. 2011; Bramble et al. 2014). The 2D frame windows were
220	chosen so that the second frame was collected with the goniometer angles of both the
221	source and the detector from the sample plane were both approximately 45°. A window
222	in lower 2 $\theta$ was also chosen to allow for comparison with the approximately 45° window,
223	and because both magnetite and pyroxene have frequent and diagnostic XRD peaks in the
224	range of 30 to 60 °20. The specific parameters used in this experiment employed a two
225	frame coupled scan, where frame 1 was collected with $\theta_1 = \theta_2 = 20^\circ$ and frame 2 was
226	collected with $\theta_1 = \theta_2 = 29.5^\circ$ , and width = 19°. These parameters, when integrated,

generated an analysis range of 18–79 °20. The data collection time for each frame was
120 minutes. Unless specified otherwise, all data in this study were collected with the
same instrumental parameters.

230

## 231 Omega scanning and sample oscillation

232 A key feature of  $\mu$ XRD is the versatility of the technique in the investigation of 233 crystalline matter *in situ*, without sample preparation and with modifications as seen in 234 the capabilities of omega scanning and sample oscillation. To constrain the effects of 235 these method modifications on the grain size equation, a select set of the pyroxene and 236 magnetite samples were analyzed by omega scan and sample oscillation in addition to the 237 coupled scan method and the grain sizes were calculated for comparison. Testing these 238 effects will aid in gauging the applicability of this grain size calculation method in 239 situations where the sample or the optics are in motion during data collection (e.g. Blake 240 et al. 2012).

241 Omega scanning is a feature of  $\mu$ XRD where the optics (source and detector) are 242 simultaneously rotated in the same direction (clockwise) by a certain angle omega ( $\omega$ ) 243 during data collection (see Flemming 2007). For each GADDS frame, the position of the 244 source starts at a low  $\theta_1$  angle relative to the sample and is rotated by the goniometer to 245 higher  $\theta_1$  angle relative to the sample (in degrees  $\omega$ ) while the detector, initially 246 positioned at high  $\theta_2$  angle relative to the sample, is rotated to lower  $\theta_2$  (in degrees  $\omega$ ). 247 The source and detector are rotated by the same omega angle at the same angular rate to 248 maintain a constant 2 $\theta$  position at the center of the detector (where  $\theta_1 + \theta_2 = 2\theta$ ). Rotation of the optics enables more lattice planes to enter the correct geometry to satisfy Bragg's 249

250	law, mimicking the rotation of a sample under fixed source-detector optics. The objective
251	of omega scanning is to increase the number of crystal lattices of a coarse-grained sample
252	that are irradiated by the X-ray beam, producing diffracted rays that reach the detector.
253	An omega scan will generate more diffracted spots on the detector for a particular
254	mineral phase than an equivalent coupled scan.
255	The $\chi$ -profile grain size calculation was derived for a stationary optical geometry.
256	This ensures that the window of area integrated on the detector directly matches the
257	irradiated sample volume for a specific instrument geometry. An omega scan should
258	increase the number of grains irradiated in a sample relative to the number measured in
259	an equivalent coupled scan, and the $N_s$ term should be similarly larger. This should result
260	in the grain size calculation underestimating the grain size, because more grains will be
261	counted and inserted into the equation than would be expected.
262	The micro-X-ray diffractometer used in this study also features a sample stage
263	capable of moving up to 10 cm in X, Y, and Z directions. This sample stage allows for
264	various materials to be placed on the stage and allows for spots of interest to be targeted
265	and focused in three dimensions. This stage allows for samples to be oscillated in one,
266	two, or three directions during data collection. Similar to the purpose of the omega scan,
267	sample oscillation is intended to increase the number of different crystal lattices passing
268	under the incident beam, which will then diffract X-rays towards the detector. Sample
269	oscillation aims to generate data that would be akin to analyzing a powdered sample.
270	Similar to the expected effects of omega scanning, sample oscillation should
271	increase the number of spots in a spotty ring for a given geometry than would be
272	expected without sample oscillation. The result should be an increase in the $N_s$ term for

- the grain size calculation and subsequently an underestimation of grain size.
- 274

### 275 Materials

276 The pyroxene investigated in this study has the formula

277  $[Mg_{1.753}Fe_{0.206}Ca_{0.025}Cr_{0.012}Mn_{0.004}Ni_{0.002}Co_{0.001}]_{\Sigma=2.003}(Si_{1.965}Al_{0.031})_{\Sigma=1.996}O_6 \text{ and was a}$ 

single large crystal that has been crushed and separated into 32 size fractions by dry and

279 wet sieving for previous investigation of reflectance properties (e.g. Cloutis et al. 2008).

280 The synthetic magnetite samples are from commercial sources and have been

investigated in a previous study of magnetic property variation with grain size (Yu et al.

282 2002), which included a grain size analysis by scanning electron microscopy (SEM).

283 Columbia River Basalt sample SA-51, used in this study, was taken from a Roza

284 Dike of Wallowa County, Oregon, of the United States of America. The Roza Member is

a geological unit of the Columbia River Basalt Group (Martin 1989, 1991; Thordarson

and Self 1998). This sample has been part of a geochemical analysis of the geological

unit (Atkinson 1990) and has been studied with reflectance spectroscopy (Cloutis et al.

288 2008). The sample was also separated into 32 size fractions by dry and wet sieving. The

289  $\gamma$ -profile grain size analysis of this sample was performed to test the application of this

290 method to multi-phase samples.

291

#### 292 Data processing procedure

The analysis of the collected GADDS 2D images involved their integration to
produce one-dimensional datasets. Initially, the entire frames were integrated normally to
produce intensity versus 20 plots. These were used to determine the exact location of

296	each Debye ring in $2\theta$ space in order to index these from a relevant International Centre
297	for Diffraction Data (ICDD) card, and determine the multiplicity of the diffracting plane.
298	Once the Debye rings and their multiplicities were identified, the frames were
299	integrated along the Debye rings ( $\chi$ dimension) to produce intensity versus $\chi$ plots. For
300	consistency and comparison, the same Debye rings (i.e. diffraction vectors) were
301	integrated for all samples of a given mineral. The integration windows were integrated by
302	delineating a selected $2\theta$ by $\chi$ angular window to include spotty rings (Fig. 1a); care was
303	taken to avoid splitting any particularly large or bright spots. A few fractions of one
304	degree of background $2\theta$ area were integrated on each side of the ring so that all intensity
305	across the full width of the ring was integrated.
306	The next step was to plot the $\chi$ -profiles along with a trend line (Fig. 1b). For the
307	initial pyroxene dataset, both a second-degree polynomial and a linear average intensity
308	line were used and the number of times the profile crossed the line $(N_s)$ was counted for
309	each. The polynomial often resulted in marginally higher numbers counted, but the linear
310	trend produced more uniform results. Therefore, while this pyroxene application used

each. The polynomial often resulted in marginally higher numbers counted, but the linear trend produced more uniform results. Therefore, while this pyroxene application used both methods, the magnetite and basalt applications used only the average intensity line. The  $N_s$  term was calculated visually for the pyroxene dataset as well as using a threshold crossing algorithm to check the accuracy of the algorithm, which was then applied to the magnetite and basalt data. While this profile method may bias the  $N_s$  term towards grains above a certain size, it should remove contributions from complicating factors such as very fine mineral dust coating the samples, as was observed below in select pyroxene 2D XRD images.

318

The divergence angle of the collimator along the primary beam ( $\beta$ ) was taken

319	from He (2009), where $\beta$ for a 300 µm single pinhole collimator was given as 0.225
320	degrees. This value was entered into the grain size equation as 0.003927 radians. All
321	calculations were performed with GNU Octave (Eaton et al. 2013).
322	The version of the $\chi$ -profile grain size measurement calculation applied in this
323	study was the one derived for reflection geometry $\mu XRD$ (He 2009). Therefore, the depth
324	of X-ray penetration must be known to derive the volume irradiated. The calculation uses
325	an area multiplied by height calculation where the area is the cross section of the beam
326	and the depth is given by the linear absorption coefficient ( $\mu$ ) and the instrumental
327	geometry. The linear absorption coefficients for all materials used in this study were
328	calculated in GNU Octave, and the calculation used was taken from Ladd and Palmer
329	(2003).

330

## **331** Rietveld refinement of basalt sample SA-51

332 To quantify the modal mineralogy of the SA-51 basalt sample, a Rietveld 333 refinement was performed with data collected by a Rigaku Geigerflex D/MAX powder 334 X-ray diffraction system. The interest of the refinement was to acquire approximate 335 modal proportions of the major phases with an accuracy of a few weight percent. For 336 analysis, a portion of the  $<10 \mu m$  sieve fraction of SA-51 was ground for one hour with 337 an agate mortar and pestle to achieve a grain size less than approximately 5 µm. The ground sample was then placed in a sample holder with the aid of ethanol, which was 338 339 allowed to dry. The Rigaku X-ray diffractometer was operated at 40 kV and 35 mA. Data 340 were collected from 10–90 °2 $\theta$  at a step size of 0.02 °2 $\theta$  and a counting time of 42 s. The Rietveld refinement was performed using the TOPAS software (version 3, 341

342	Bruker-AXS 2005). Phases were identified using the Bruker-AXS DIFFRACplus
343	Evaluation software package in tandem with the Inorganic Crystal Structural Database
344	(ICSD). Rietveld refinement is a non-linear least squares calculation that fits a calculated
345	pattern to the observed diffraction data to determine crystal structural parameters for
346	powdered materials (Rietveld 1967, 1969; Young 1993; Pecharsky and Zavalij 2005).
347	The diffraction pattern peaks used the Thompson-Cox-Hastings pseudo-Voigt lineshape
348	for modeling. We refined on zero error, sample displacement, and surface roughness, as
349	well as scale and unit cell for each of the phases. Neither site occupancies nor the
350	chemical compositions of the phases were altered from the original input structures.
351	In addition to crystal structure refinement, the method produces modal
352	proportions of the phases in the refinement based on parameters such as the scaling
353	factors and unit cell volume. While these abundances are highly dependent on the initial
354	input parameters, they are accurate within a few percent, especially for fine-gained and
355	homogenous materials.
356	
357	Results
358	
359	Application to well-characterized pyroxene
360	Table 1 displays the grain sizes calculated for the pyroxene samples analyzed
361	using the polynomial (poly) and linear average intensity (lin) methods. A set of 2D
362	images are shown in Figs. 2 and 3 that display the progression of Debye ring
363	characteristics with increasing grain size, and images of the corresponding pyroxene
364	samples are shown in Fig. 4. All grain size calculations for Debye rings from a particular

sieve size are also averaged into a single value for the size fraction for both the
polynomial and average intensity values and are presented in their own column in Table
1, followed by their respective standard deviations. This process was followed for all of
the data presented below.

369 The calculated grain sizes appear to broadly correlate with the sieve-size bins. It 370 was difficult to draw conclusions about the "less than" sieve bins (e.g.  $<25 \mu m$ ) because 371 grain size did not have to fall within a specific limit. These sieves allowed for all material 372 below the given limit to pass. The calculated sizes for the <25 and <20 µm samples do 373 correlate with the sieve size, but the <10 and  $<5 \mu m$  samples are calculated above the 374 sieve mesh size. The discrepancies may result from equation parameters (such as using a 375 given beam divergence from He (2009)), or from the pyroxene physical properties that 376 may have affected the sieving results, such as a stubby prismatic habit. 377 For many of the pyroxene sieve size bins with upper or lower limits, the 378 calculated grain sizes correlate with the sieve size bins. The calculated grain sizes either 379 fell within the sieve size bins or just outside the bin by  $\pm -5 \mu m$ . This was not true for the 380 samples larger than 38  $\mu$ m, where the calculated size tends to become significantly 381 smaller than the bin size. These deviations in the larger sample sizes may be due to the above conditions, or may be a result of the instrumental setup. With the instrument 382 383 geometry used in this study, the close detector distance may have caused the equation to significantly underestimate grain size (by  $\geq 10 \mu$ m) as the sieve sizes increased above ~60 384 385  $\mu$ m. This may be the result of fewer diffraction spots reaching the detector area for a 386 given Debye ring. The standard deviations are similarly higher for the larger grain size samples and may reflect the uncertainty in measuring grain sizes above certain limits of 387

388 instrumental and physical conditions.

389	Table 1 shows the strong correlation between the polynomial and average
390	intensity methods for the majority of the $\chi$ -profiles analyzed. For the samples below 38
391	$\mu$ m, the difference in the calculated size between the polynomial and average intensity
392	methods differs by 1 $\mu m$ or less for the majority of the samples, and the sizes commonly
393	overlap within the calculated standard deviations. As the sieve sizes grow above 38 $\mu m$
394	the two methods begin to slightly diverge with the 38–45 and 75–90 $\mu m$ dry sieve
395	samples differing by about 5–6 $\mu$ m between the polynomial and average intensity
396	methods.
397	The visual analysis method of counting the number of times the $\chi$ -profile crossed
398	the trend line proved to generate more consistent results for the average intensity method
399	than the polynomial method. All of the following grain size calculations in this work use
400	the average intensity method.
401	
402	Application to well-characterized magnetite

403 The  $\gamma$ -profile grain size analysis was applied to a limited set of magnetite samples. 404 The samples analyzed were of the following grain sizes:  $0.065 \pm 0.036 \,\mu\text{m}$ ,  $0.24 \pm 0.07$ 405  $\mu$ m, 1.06 ± 0.71  $\mu$ m, and 18.2 ± 12.0  $\mu$ m (see Fig. 5), which were measured by Yu et al. 406 (2002) using SEM imaging. The  $\chi$ -profile method was proposed to be applicable to grain 407 sizes in the range of 0.1 to 100 µm (He 2009), but actual performance will vary within 408 this range, depending on the instrumental parameters and sample properties. These 409 samples will therefore be able to test the lower range of sizes calculable by this method. The results of  $\chi$ -profile analysis of these magnetite samples are shown in Table 2. 410

411	The measurement of the 18.2 $\pm$ 12.0 $\mu m$ magnetite sample produced a result of 31.14 $\pm$
412	7.43 $\mu$ m, and other magnetite samples appear to suggest that the method does not
413	calculate accurate grain sizes when the samples are at or below 1 $\mu$ m. He (2009)
414	suggested the applicability of the $\chi$ -profile method was in the range of 0.1 to 100 $\mu$ m, but
415	the 0.065, 0.24, and 1.06 $\mu m$ samples return approximately the same grain size of ${\sim}18$
416	$\mu m$ , suggesting that the problems arising with observation of the 0.065 and 0.24 $\mu m$
417	samples are also present with the 1.06 µm sample.
418	One possible explanation is that as the grain size decreases to 1 $\mu$ m, and continues
419	with decreasing grain size to 0.065 $\mu$ m and beyond, the spotty rings become more
420	continuous, and the number of grains providing spots in the detector window trends
421	asymptotically towards the pixel density of the detector. This is likely occurring in the
422	pyroxene samples as well, and may be why the grain size of the smaller samples appears
423	to level off at about 18 $\mu$ m as the grain size continues to decrease (Table 1). If this
424	observation is correct, than the lower limit of applicability of the $\chi$ -profile method with
425	the instrumental geometry used in this study, and with the physical properties of the
426	magnetite samples, is perhaps closer to 18 $\mu m$ rather than the proposed 0.1 $\mu m$ (He
427	2009).
428	In the 2D image of the 0.065 $\mu$ m sample, line broadening of the Debye rings is
429	observed. This visually occurs between the 0.24 $\mu m$ and the 0.065 $\mu m$ samples as shown
430	in Fig. 5. By analyzing the line broadening of these rings in $2\theta$ , the grain size could be
431	calculated via the line profile method.
432	The larger calculated grain size for the pyroxene, with respect to sieve size bin for

433 the smaller sieve sizes, and the larger calculated grain size for magnetite, with respect to

434	SEM grain size for the smaller grain sizes, both appear to result from erroneous
435	circumstances with the equations. As the pyroxene sieve sizes became smaller, the
436	calculated grain sizes began to level off at about 15 $\mu$ m. No pyroxene grain size was
437	calculated below ~14 $\mu m.$ It is possible that the window that is integrated in $\chi$ becomes
438	saturated as the grains are approaching ${\sim}15~\mu\text{m},$ and as the grains pass below this size
439	(~15 $\mu$ m) the N <sub>s</sub> term trends asymptotically towards that of a smooth Debye ring. The
440	magnetite grains suggest that this hypothesis is correct, because for the 18.2 $\mu m$
441	magnetite sample, the observed grain size correlates with the calculated grain size within
442	given uncertainties, but for the three smaller samples (1.06 $\mu$ m and below),
443	approximately the same grain size was calculated (~18 $\mu m$ ). In this case we are simply
444	measuring the $\chi$ -profile of a smooth ring, which would not contain any grain size
445	information in the $\chi$ -dimension.
446	
447	Omega scanning and sample oscillation
448	A test of the effects of the omega scan method (rotating optics) was executed on
449	four samples. Three binned pyroxene samples of sieve sizes 10–15, 20–25, and 25–38

450  $\mu$ m, and one magnetite sample (18.2 ± 12.0  $\mu$ m) were analyzed. A two-frame omega scan

451 was employed, with  $\theta_1 = 15^\circ$ ,  $\theta_2 = 25^\circ$ , and  $\omega = 10^\circ$  for the first frame, and  $\theta_1 = 19^\circ$ ,  $\theta_2 =$ 452 40°, and  $\omega = 25^\circ$  for the second frame. To produce similar quality data as were collected 453 for the stationary pyroxene and magnetite above, the data were collected for 120 min per

454 frame.

The results of the χ-profile grain size calculations of the omega-scanned samples
are shown in Table 3 along with the corresponding grain sizes for the coupled-scanned

457	(stationary) samples. The sizes measured from the omega scan correlate strongly with the
458	sizes measured from the coupled scan, but they all underestimate the coupled scan sizes
459	as the above theory suggested. The omega scan of the 10–15 $\mu$ m pyroxene sample falls
460	within the sieve size bin, but the two larger-sized samples have their omega scan sizes
461	fall about 5 $\mu$ m below the lower limit of the bin size. The omega scan magnetite sample
462	was calculated to be about 6 $\mu$ m below the coupled scan (stationary) sample, but the
463	omega scan size is actually closer to the size measured by SEM (though both calculated
464	sizes fall within the uncertainty of the magnetite SEM size).
465	The effects of sample oscillation were investigated by analyzing the four samples
466	analyzed using the omega scan above. The coupled scan method was applied, and the
467	samples were oscillated in the Y direction by 3.5 mm during data collection.
468	The results of the $\chi$ -profile grain size calculations on the oscillated samples are
469	shown in Table 4 along with the corresponding grain sizes for the non-oscillated samples.
470	The observations of the oscillated data were similar to those of the omega scan data. As
471	the theory suggested, the data collected while oscillating underestimated the calculated
472	grain size as compared to that calculated for same samples when not oscillated. This was
473	true for all pyroxene samples analyzed as they underestimated their non-oscillated sizes
474	by up to ~5 $\mu m.$ The magnetite sample was also underestimated by ~7 $\mu m,$ but, as was
475	seen with the omega scan data for magnetite, the oscillated sample was closer to the
476	magnetite grain size as measured by SEM.
477	

# 478 Application to the Roza Member—Columbia River Basalt Group

479 Using the same parameters for data collection as the pyroxene and magnetite,

480	Roza Member basalt samples that were sieved into size fractions with an upper and lower
481	limit were analyzed with the Bruker D8 Discover $\mu$ XRD diffractometer. The 2D XRD
482	frames were analyzed in the same manner as the data presented above.
483	As the basalt samples are composed of multiple mineral phases, the $\chi$ -profile
484	grain size equation required an estimate of the modal fraction of the selected mineral in
485	the assemblage. The modal proportion is entered into the equation as the $fi$ term, which is
486	the modal fraction of the <i>i</i> th mineral phase.
487	The results of the Rietveld refinement for modal mineralogy and the initial
488	structures employed are given in Table 5. The $R_{wp}$ value of the calculation was 10.6. This
489	refinement was executed only to ascertain the approximate modal proportions of the
490	major phases present. For a thorough crystal structure refinement, a refinement with
491	stronger fit statistics and smaller step size would be required. Only the two most
492	abundant basalt phases, anorthite (60 %) and augite (23 %), were used for the grain size
493	calculation.
494	Multiple phases in the collected diffraction pattern displayed preferred
495	orientation, the evidence of which was observed to continue into the residual from the
496	refined pattern. The preferred orientation was likely a result of sample preparation.
497	Grinding the sample for one hour may have not created a homogeneous assortment of
498	crystallites, especially for the anorthite and augite which likely remained somewhat lath-
499	like. Preferred orientation was expected as the crushed basalt sample was suspended in
500	ethanol in the sample holder, which was then allowed to dry. As the interest in the
501	refinement was to acquire approximate modal proportions of the major phases with an
502	accuracy of a few weight percent, preferred orientation was not taken into consideration

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503 nor refined in this calculation.

000	
504	The calculated grain sizes for a set of sieved SA-51 basalt samples are shown in
505	Table 6. In contrast to our previous applications of the $\chi$ -profile method, the majority of
506	the SA-51 grain sizes for anorthite and augite calculated did not correlate with the sieve
507	size bins. The smaller sieve sizes up to the 15–20 $\mu$ m sieve size did correlate with the
508	calculated grain sizes, within their standard deviations. The overall characteristic of the $\chi$ -
509	profile grain sizes appears that they seem to converge at a value of about 13–14 $\mu$ m, even
510	in the 63–75 $\mu$ m sieve size fraction of the sample. This was especially apparent when a
511	'whole rock' grain size was calculated by averaging all the calculated grain sizes for
512	different minerals in a single sieve size sample. One standard deviation of this 'whole
513	rock' grain size for all sieve sizes analyzed is 1.41 µm.
514	Roza Member basalt samples appear to be composed of agglomerated minerals of
515	${\sim}13{-}14~\mu m$ in size that dominate the calculated grain size, regardless of the rock particle
516	size (sieve size), whether the minerals occur as separate grains or as part of a larger rock.
517	This is expected, as X-rays will only investigate the size of coherent scattering domains
518	of the mineral grains or crystallites within the rock fragment and not size of the larger
519	rock fragments. The mineral grain size of the rock fragments was measured qualitatively
520	by reflected light petrography and appeared to agree with the $\sim 14 \ \mu m$ grain size
521	calculated by the $\chi$ -profile method.
522	Difficulty arose in integrating particular Debye rings for these samples as it was

523 not straightforward which particular spots belonged to which mineral phase. This

524 occurred, in part, because basalt consists of multiple minerals with low symmetry. The

525 low symmetry causes the mineral to have a low multiplicity that results in the mineral

	having many closely-spaced Debye rings (cf. Figs. 2 and 3 with Fig. 5). A mixture of
527	minerals of relatively low symmetry made it difficult to avoid overlapping diffraction
528	spots. Additionally, augite and pigeonite have many diffraction peaks at similar $2\theta$
529	locations. The augite Debye rings used for the basalt grain size calculations were at 40.7
530	and 41.5 °20, which were also the approximate locations of similar intensity pigeonite
531	peaks. Some proportion of the augite diffraction spots used to calculate grain size in this
532	study may in fact have been due to pigeonite peaks, and therefore the uncertainty in these
533	calculations is increased.
534	Further difficulty arose from the size of spots on the 2D images of multi-phase
535	materials. Frequently, the XRD spots were so large in area that they spanned an entire
536	degree of $2\theta$ , making the identity of particular diffraction spots difficult for multi-phase
537	materials. This was particularly difficult for the larger sieve size fractions, where the
538	Debye rings were discontinuous and individual diffraction spots were large.
538 539	Debye rings were discontinuous and individual diffraction spots were large.
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539 540 541 542 543 544	Discussion Progression of Debye ring characteristics with decreasing grain sizes Our collected data on the suite of pyroxene samples allows us to observe the changing characteristics of Debye rings with changes in grain size (see Figs. 2 and 3).

see that the diffraction spots may be deviating somewhat from the Debye ring diffraction
vectors. With the additional difficulty of the low symmetry of pyroxene, it can be
difficult to index large diffraction spots in data for pyroxene grains that are 38 µm or
larger in size.

552 Beginning at approximately 38 µm sieve fraction, the shape of the Debye rings 553 becomes more discernible with decreasing grain size. The diffraction spots continue to 554 have significant spacing between them, and this continues with decreasing grain size to 555 about the 20 µm sieve fraction. Beginning at about 20 µm, the Debye rings become more 556 continuous, but retain a spotty, often discontinuous appearance. A low intensity, 557 somewhat-continuous ring was seen with very high intensity smaller diffraction spots 558 embedded in the ring. Depending on the intensity of these high intensity spots, the lower 559 intensity spots may not be significant enough to count as a grain in the  $\chi$ -profile analysis. 560 The  $<5 \mu m$  sample, which was twice ground for one hour in this study, displays 561 the most continuous rings of all the pyroxene samples. Note that these rings are still highly discontinuous and at times highly spotty. The  $<5 \mu m$  sample continues to display 562 563 sporadic diffraction spots of relatively high intensity embedded in the rings. This is likely 564 the result of a distribution of grain sizes contained in each sieved sample.

Selected results from our findings here are in contrast to selected literature
observations of spotty Debye rings. In the introduction, several qualitative observations
relating Debye ring characteristics to grain size were presented. For example, Cullity
(1978) stated that for grain sizes from 10 to 1 µm there was a transition from spotty
diffraction rings to continuous rings without spots. Similarly, Hörz and Quaide (1973)

570	stated that samples in the range of 0.2 to 10 $\mu$ m should display "smooth rings, perfect
571	powder pattern". While we did observe this behaviour in the two magnetite samples that
572	bracket this range, it was not observed in the pyroxene samples. Our ${<}5~\mu m$ pyroxene
573	data cannot be described as continuous rings without spots, nor as a perfect powder
574	pattern.

Hörz and Quaide (1973) also stated that their samples in the range of 10 to 40 μm
displayed "clearly discernible diffraction lines". This Debye ring behaviour was not
clearly observed in our pyroxene data. In the range of 10 to 40 μm discernible rings begin
to appear but they remain strongly discontinuous.

579 These differing observations of Debye ring characteristics on 2D XRD images as 580 a function of grain size suggest that the qualitative method of constraining the grain size 581 of a sample by comparison of ring characteristics with literature observations should be done with caution. A potential source of the discrepancies between our observations and 582 583 the literature may be the result of the difference between a Debye-Scherrer film camera 584 and a HI-STAR multi-wire detector. Our detector has a diameter of 11.5 cm and a 585 resolution up to 1024 pixels along this diameter. A film camera has a width of ~2.5 cm. 586 Other sources of variance between the Debye-Scherrer observations and those of this study may be the use of different sample preparation methods, or the difference in X-ray 587 588 beam diameters applied in each system, which would result in different irradiated sample 589 volumes. As a result of the scaling involved between the different sample-to-detector 590 distances, the beam diameters employed, and imaging areas involved in the two methods, 591 it is possible that some of our discontinuous Debye rings would become smoother in film 592 cameras, and result in our discrepancies from the film-based grain size literature.

593	Similarly, the degree to which the grain size of a sample correlates with its crystallite size
594	will also affect the relationship of Debye ring characteristics as a function of grain size. If
595	the sample is not well constrained in this manner and the correlation is not strong then
596	there could be great variability at which grain size changes in the progression of ring
597	characteristics occurs.

598

599 Complications of sample heterogeneity

Samples that display a variation in grain size or mineralogy complicate the  $\chi$ profile grain size analysis method; the primary reason being that the method analyzes only a small volume of sample. A sample can fill a large sample holder but the method will only analyze a volume equal to the area of the beam times either the depth of penetration or the thickness of the transmission sample holder.

605 Voluminous or varied samples create a circumstance where the volume irradiated may not be representative of the sample. For example, if the sieve size is  $20-25 \mu m$  with 606 607 an uncertainty of about 10  $\mu$ m above or below the size bin, the small irradiated volume 608 may give a grain size of 10 µm or 35 µm, neither of which might be representative of the 609 whole. To limit the error due to sample inhomogeneity, we recommend collecting data at 610 multiple locations on larger samples, and averaging the grain size over these irradiated 611 volumes to provide a more representative calculated grain size. Analyzing data collected 612 at multiple targets of a single sample for comparison was employed in the magnetite 613 sample analysis in this study.

614

Variation in sample composition can also affect the certainty of the grain sizes

615	calculated due to the presence of the linear absorption coefficient in the calculation of
616	irradiated sample volume. Composition variation was shown to have a minor effect when
617	the variation was small. For example, when cation substitutions are on the order of a few
618	tens of percent the variation in calculated grain size can be on the order of a few microns.
619	This was confirmed by replacing the end-member enstatite (MgSiO <sub>3</sub> ) composition by a
620	pyroxene with the composition measured by microprobe (see Materials section above) in
621	the calculation. Little change was seen between the two calculated datasets. The change
622	in crystal chemistry resulted in a decrease in the calculated grain sizes of about one
623	micron or less for many of the samples, with the exception of few where the pyroxene
624	composition caused a drop of a few microns. The addition of iron into the equation
625	caused the linear absorption coefficient to marginally drop, corresponding to the X-ray
626	beam irradiating a smaller volume of material. Therefore, when the calculation was
627	executed assuming the same number of crystallites contributing to the Debye ring, the
628	calculated grain size was smaller.

For samples with larger chemical variation, or uncertain chemical nature, the volume of material analyzed may be greatly affected by the absorption characteristics of all the materials in the irradiated volume. If the linear absorption coefficient used in the calculation does not reflect the absorption characteristics of the minerals analyzed, the calculated volume irradiated, and the corresponding grain size, will be incorrect.

634

### 635 Complications of multi-phase materials

636 The  $\chi$ -profile grain size analysis method becomes more complicated when applied

637	to multi-phase materials with multiple minerals and varying modal mineralogies. With
638	single-phase materials, the method assumes that the diffraction spots correspond to the
639	entire irradiated volume divided into a corresponding number of grains per spots with
640	factors taken into consideration such as multiplicity. Multi-phase materials will subdivide
641	the calculated irradiated volume further to represent the volume fraction of the phase of
642	interest in the sample. This can be done by including the modal proportion of the mineral
643	in the grain size calculation.

644 Several concerns and complications arise from the inclusion of modal mineralogy. 645 To acquire modal mineralogy of a sample, time intensive and perhaps expensive 646 techniques are needed. A common method is Rietveld refinement, but for certain rock 647 types other methods such as normative mineralogy could be used. Rietveld refinement 648 requires high quality data that are not easily acquired using the Bruker D8 Discover 649 (Ning and Flemming 2005), and may require grinding the sample to provide the small 650 grain sizes needed for Rietveld-quality data as well as analyzing the sample on another 651 instrument. Both situations were used for the basalt sample analyzed in this study.

652 For mixtures in reflection-mode geometry, the effectiveness of measuring the 653 irradiated volume by incorporating the linear absorption coefficient must be questioned. 654 The separate phases in a sample will each have distinct absorption characteristics, and, 655 depending on the variations in atoms, the individual minerals may have greatly differing 656 linear absorption coefficients. The  $\chi$ -profile method assumes that, when measuring the 657 grain size of a particular mineral using its modal mineralogy, only the linear absorption 658 of this particular mineral must be employed. This is questionable for the X-rays will be 659 interacting with, and altered by, the absorption characteristics of all of the phases

660	encountered, especially if the sample is not completely homogenous. The situation is
661	more complex than the $\chi$ -profile method assumes. It changes the effective volume to suit
662	a particular circumstance of absorption and modal mineralogy when measuring one
663	particular mineral, and then allows the volume and absorption to change again when
664	measuring a different Debye ring for a different mineral in the same 2D image.
665	Multi-phase materials pose a significant homogeneity problem for samples of
666	larger grain sizes for the mixture of minerals being analyzed by the beam may not

667 necessarily correlate with a modal mineralogy measured on a finely ground sample.

668 When the grain size approaches  $\sim$ 50  $\mu$ m, the microscope context camera on the Bruker

D8 Discover shows that only the upper minerals will be targeted by the beam (see Fig. 4g

and 4h). The small size of the beam (nominally 300  $\mu$ m) suggests that the small volume

of material analyzed will not correlate with the bulk modal mineralogy, and thus the

671

672 calculated grain sizes will be inaccurate. As the grain size of the material analyzed

approaches a significant proportion of the X-ray beam footprint, the heterogeneity of

674 multi-phase materials may induce significant errors in the grain sizes calculated.

Similarly, He (2009) states that smaller X-ray beam cross sections should be
employed for accurate measurements. For the above reasons, multi-phase materials will
induce errors with smaller beam footprints for they will potentially subsample fewer
phases. An optimal crossover point may exist between the accuracy derived by smaller
cross sections and the need to sample a representative portion of multi-phase materials.
Further investigation of multi-phase materials is needed to constrain the above effects.

681 We reiterate that the  $\chi$ -profile method cannot be used to calculate the particle size

682	of rock fragments, because the diffracted X-rays relate to the individual irradiated
683	mineral grains in a sample. The surface of a rock particle will be composed of multiple
684	lattice planes corresponding to the constituent mineral grains. The X-rays will be
685	diffracted by these various planes and generate a 2D image that reflects the sum of
686	individual mineral grains in the rock sample, and not the particle size of the bulk sample.
687	As a rock is ground more finely, eventually the mineral grain size is reached (~13 $\mu m$ in
688	the case of Columbia River basalt), and the calculated grain size will correspond to the
689	particle size of the heterogeneous material analyzed (i.e. the sieve size). As the rock is
690	ground even finer, the particle size and crystallite size will decrease simultaneously, and
691	the calculated grain size by XRD will correlate with the sieve size, as was moderately
692	observed in the basalt dataset. More study is needed to aid in refining the application of
693	this method to multi-phase materials.

694

## 695 Grain versus crystallite size assumptions

696 For the above application of the  $\chi$ -profile technique, we used the collected data 697 from this crystallite size analysis technique to make inferences about the grain size of the 698 samples. This assumption that the grain size of a sample and the coherent scattering 699 domain size are the same may be the greatest limitation of applying this method to 700 geological materials. In some cases, these two physical properties may coincide, but the 701 provenance of and preparation techniques applied to many geological materials may exacerbate the disagreement between the coherent scattering domains of a mineral and 702 703 the grain size calculated.

704

The correlation between the coherent scattering domains and observable grain

705	and subgrains via other analytical measurement techniques, such as transmission electron
706	microscopy, suggests that there may not be a direct correlation between these phenomena
707	(e.g. Ungár et al. 2005). The presence of subgrains or small coherent scattering domains
708	within a mineral grain may bias a grain size measurement towards reporting sizes that are
709	smaller than the seemingly apparent whole grain. From this perspective, our observations
710	may be influenced by the subgrains present in the mineral samples. When applying this
711	method to calculate the grain size of the many types of geological materials, if taken from
712	a range where the method is known to be effective, the calculated result should be
713	interpreted as a minimum estimate of the grain size.
714	Certain observations could be made to constrain the strength of the assumption
715	that the grain size and coherent X-ray scattering domains correlate via 2D XRD images.
716	For example, a crystal, when experiencing strain or shock processes, will undergo a
717	transition from a discrete spot on a 2D XRD detector to a spot that will streak out until a
718	point where the strained lattice resolves into a series of discrete subgrains (Flemming
719	2007; Vinet et al. 2011; Izawa et al. 2011). The end result of clustered subgrains displays
720	an asterism pattern on the 2D XRD image (Flemming 2007; Vinet et al. 2011). Therefore,
721	it is possible to infer some degree of subgrain formation via 2D XRD data. We have not
722	observed spatially clustered diffraction spots akin to asterism in the data presented herein,
723	supporting that the mineral grains have not been exceedingly disrupted into small
724	subgrains. The crushing and sieving of the pyroxene samples could reasonably induce
725	defects into the crystals leading to poor correlation between the sieve size and the actual
726	grain size of the individual particles in the powders. The magnetite samples were
727	characterized by SEM, but despite the more-quantitative SEM characterization of

728	magnetite than sieving of pyroxene, the calculated grain sizes of each agreed with the
729	observed grain size at approximately the same correlation level.
730	
731	Implications
732	
733	Effective range of $\chi$ -profile grain size analysis
734	He (2009) states that the $\chi$ -profile crystallite size calculation method should be
735	effective for materials in the range of 0.1–100 $\mu$ m. The findings from our investigation
736	suggest that, when applied to calculate the grain size of varied geological materials, this
737	range should be constrained in both its upper and lower limit for reflection-mode
738	geometry µXRD.
739	The upper limit should be lowered to about 63 $\mu$ m, especially for closer sample-
740	to-detector distance as was used in this study. We found that the technique began to
741	significantly under-report grain sizes for samples larger than 63 $\mu$ m. The 45–63 $\mu$ m
742	pyroxene sample was the largest sample to provide grain size calculation that correlated
743	with sieve size. The grain size calculated for this sample was $56.77 \pm 24.68 \ \mu\text{m}$ . The 75–
744	90 $\mu$ m sample, for example, gave a calculated grain size of 42.12 ± 22.63 $\mu$ m. This may
745	be the result of not having enough grains diffracting at orientations within the limited
746	circumference of the Debye ring on the area detector to provide a statistically significant
747	$\chi$ -profile thus not providing a representative number of diffraction spots. The
748	measurement of only the partial Debye rings that fall within the 2D detector is a
749	potentially limiting factor of this method.

750	The lower limit results from saturating the detector, as discussed briefly above. As
751	the pyroxene and magnetite grain sizes grew smaller, the calculated grain size appeared
752	to trend asymptotically toward a value of about 15 $\mu$ m. Therefore, as grain sizes approach
753	this value, which may vary depending on the material being analyzed, caution should be
754	employed. Certain instrumental conditions, such as a larger sample-to-detector distance
755	or different beam footprint, may allow for smaller grain sizes to be calculated, but in the
756	instrumental setup used in this study, the lower limit appears to have been about 15 $\mu$ m.
757	Further experiments in measuring grain sizes in the range of 0.1 to 15 $\mu$ m would greatly
758	benefit the efficacy of the $\chi$ -profile grain size measuring method.
759	In the grain size range of 15–63 $\mu$ m, the $\chi$ -profile method of measuring grain size
760	via reflection geometry $\mu XRD$ appears to be moderately successful at providing an
761	estimate of the sample grain sizes. The reflection-mode version of this equation does
762	appear to be hindered by the many inherent free parameters and numerical constants
763	required for the calculation and the physics and geometry involved with multi-phase
764	materials. The method can provide a minimum estimate of the calculated grain size that is
765	accurate within several microns of the mean grain size under specific instrumental
766	conditions, but the requirement of modal mineralogy for multi-phase materials may limit
767	the effectiveness of this technique.
768	In calculating the grain size of geological materials, this method may be limited to

measuring monomineralic materials and further limited by heterogeneities of bulk

samples. The scope of this method may be limited to highly characterized and uniform

- samples where the variation in constituent grains is small. The scope may be further
- 1772 limited in its laboratory application of grain size measurement due to the assumption that

3/11

773	the coherent scattering domains correlate with the overall grain size of the sample, due to
774	instrumental factors that may be influencing the results, and likely other factors requiring
775	further investigation. If a particular sample were heavily influenced by subgrain
776	formation, then the grain sizes measured by the above method would correlate with the
777	coherent scattering domains of the crystallites and not closely related to the grain size of
778	the sample as measured by geological methods such as sieving.
779	In the geological and planetary sciences, 2D XRD has been increasing in use as a
780	method of quantifying the strain and shock levels of materials via the progression of 2D
781	XRD characteristics of minerals as they are shocked or strained by impacts or other
782	planetary processes. To further advance these and other analyses, understanding the grain
783	size of the samples being analyzed is of great importance, and being able to analyze the
784	grain size of a sample <i>in situ</i> via 2D $\mu$ XRD will be valuable. Similarly, the study of
785	planetary materials will benefit from the further development of a method of constraining
786	the grain size of a sample via non-destructive and sample-preparation-free methods, such
787	as µXRD.
788	
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790	
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Figure 3: Two-dimensional X-ray diffraction images of pyroxene samples of increasing
sieve size. Frames are shown for pyroxenes of dry sieve sizes (a) 25–38, (b) 38–45, (c)
45–63 μm, and (d) 90–125 μm. Note: Broad, diffuse Debye rings seen in the images are
likely due to fine-grained dust-sized particles of pyroxene generated during the original
crushing of the sample, as the sample was not washed after crushing.



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## 977 Table 1: Grain sizes of sieved pyroxene samples as calculated by $\chi$ -profile grain size

#### 978 analysis.

Sieve size (µm)	Avg. calc. poly (µm)	SD (µm)	Avg. calc. lin. (µm)	SD (µm)
<5 wet, ground	15.76	1.34	15.79	1.24
<10 wet	17.97	1.32	18.44	2.60
<20 dry	17.77	0.38	17.34	0.99
<25 dry	22.73	2.52	24.35	3.91
10 - 15 wet	20.22	2.14	20.85	2.57
15 - 20 wet	18.56	3.70	19.26	4.99
20 - 25 wet	22.32	0.48	23.15	1.40
25 - 38 wet	21.63	7.05	22.35	6.27
25 - 38 dry	25.88	4.24	28.27	5.77
38 - 45 dry	22.70	6.07	28.01	10.68
75 - 90 dry	36.16	20.67	42.12	22.63

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Notes: Samples are identified by the sieve size used and whether they were wet or dry
sieved. A value averaging the grain size from multiple Debye rings for the sample is
given as well as the standard deviation (SD). This process was executed for both the

983 polynomial dataset ('poly') and the linear average intensity line dataset ("lin.").

984

985 Table 2: Magnetite grain sizes as measured by scanning electron microscopy (SEM; Yu

986 et al. 2002) and calculated by  $\chi$ -profile grain size analysis.

SEM size (µm) Calc. (µm) SD (µ	ш)
$0.065 \pm 0.036$ 18.36 4	.28
$0.24 \pm 0.07$ 18.53 4	.37
$1.06 \pm 0.71$ 19.58 4	.66
$18.2 \pm 12.0$ $31.14$ 7	.43

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988 Note: The sample averaged calculated size (Calc.) is given along with the standard

989 deviation (SD).

# Table 3: Grain sizes of non-omega-scanned and omega scanned pyroxene and magnetite

samples. The samples are identified by either their wet sieve size (in  $\mu$ m) for pyroxene, or

992 scanning electron microscopy size for magnetite. The  $\chi$ -profile calculated grain sizes are

given along with the standard deviations (SD).

Original	Calc. (µm)	SD (µm)
10-15	20.85	2.57
20-25	23.15	1.40
25-38	22.35	6.27
$18.2\pm12.0$	31.14	7.43
Omega scan	Calc. (µm)	SD (µm)
Omega scan 10-15	Calc. (μm) 16.77	<b>SD (μm)</b> 0.63
0	<b>u</b> /	
10-15	16.77	0.63

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Table 4: Grain sizes of non-oscillated and oscillated pyroxene and magnetite samples.

997 The samples are identified by either their wet sieve size (in  $\mu$ m) for pyroxene, or

998 scanning electron microscopy size for magnetite (also in  $\mu$ m). The  $\chi$ -profile calculated

grain sizes are given along with the standard deviations (SD).

Original	Calc. (µm)	SD (µm)
10-15	20.85	2.57
20-25	23.15	1.4
25-38	22.35	6.27
$18.2 \pm 12.0$	31.14	7.43
Oscillated	Calc. (µm)	SD (µm)
Oscillated	<b>Calc. (μm)</b> 16.86	<b>SD (μm)</b> 0.87
10-15	16.86	0.87

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- 1002 Table 5: The results of the Rietveld refinement for modal mineralogy of the SA-51 basalt
- 1003 sample. The mineral phases employed in the refinement, their initial structure references,
- 1004 and the weight percent proportions are given.

Phase	Reference	wt.%
Anorthite	Facchinelli et al. 1979	60
Augite	Clark et al. 1969	23
Apatite	Fleet and Pan 1997	3
Magnetite	Fjellvag et al. 1996	3
Olivine	Merli et al. 2001	3
Pigeonite	Camara et al. 2003	3
Hematite	Blake et al. 1966	2
Orthopyroxene	Sueno et al. 1976	2
Ulvospinel	Bosi et al. 2009	1

1005

- 1007 Table 6: Grain sizes of sieved basalt samples as calculated by  $\chi$ -profile grain size
- 1008 analysis. Samples are identified by their sieve size, and whether they were wet or dry
- 1009 sieved. For each sieve size, the calculated grain size for each mineral is given, as well as
- 1010 an averaged value for the entire sample ("Whole rock"). Uncertainties are given as
- 1011 calculated by the standard deviation (SD). The N/A denotes a grain size consisting of a
- 1012 single Debye ring measurement.

Sieve Size (µm)	Anorthite (µm)	SD (µm)	Augite (µm)	SD (µm)	Whole rock (µm)	SD (µm)
less than 10	13.09	N/A	10.45	1.01	11.11	1.56
10 to 15 wet	16.34	9.98	11.91	2.03	13.69	5.73
15 to 20 wet	21.52	N/A	11.55	0.41	14.04	5.00
20 to 25 wet	9.88	1.57	15.97	6.33	12.93	5.30
25 to 38 wet	13.59	9.57	13.33	5.82	13.46	7.09
38 to 45 dry	9.51	2.43	12.60	6.13	11.06	4.50
45 to 63 dry	15.67	10.93	14.45	8.58	14.94	8.19
63 to 75 dry	14.52	8.85	13.80	9.09	14.09	7.81
1013						