

1 **Revision 1**

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3 **First-Principles Prediction of Pressure Enhanced Defect Segregation and Migration at**
4 **MgO Grain Boundaries**

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ABSTRACT

13 Understanding the ability of grain boundaries to accommodate point defects and enhance
14 diffusion rates in mantle materials represents an important but challenging problem. Extant
15 experimental studies and recent computational efforts are mainly limited to the ambient pressure.
16 Here, we investigate this problem for MgO at the atomistic level by performing density
17 functional theory-based simulations of the {310}/[001] tilt grain boundary in MgO at pressures
18 up to 100 GPa. Our results show that native defects and impurities (Ca, Al, and proton modeled
19 here) favorably segregate to the boundary, with the segregation considerably increasing with
20 pressure. They also imply that grain boundary diffusion is easier, and more anisotropic and
21 complex than bulk (lattice) diffusion: The calculated migration enthalpies for host ions and
22 impurities at the grain boundary are smaller than the bulk values, more so at higher pressures
23 with their values being as low as ~ 1.5 eV at 100 GPa compared to the bulk values of ~ 4 eV.
24 Thus demonstrated high defect activity of grain boundaries in MgO - a major phase of Earth's
25 lower mantle is expected to be relevant to our understanding of mantle rheology and
26 geochemical process.

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28 **Keywords:** Grain boundaries, defects and diffusion, high pressure, first-principles computation,
29 MgO, and mantle materials

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INTRODUCTION

32 Grain boundaries (GBs), which exist in most naturally occurring as well as synthesized
33 materials, are known to dramatically influence many physical properties of these materials. GBs
34 are considered to serve as effective sinks for point defects including impurities thereby
35 controlling chemical diffusion, electrical conduction, and creep process, and also as nucleation
36 sites for corrosion, precipitation, fracture, and plastic deformation (Duffy 1986; Sutton and
37 Balluffi 1995; Hiraga et al. 2004; Dohmen and Milke 2010). These effects can be considered to
38 arise mainly because the interfacial regions separating the component crystal grains contain
39 structural distortions and void spaces (e.g., Kingery 1974; Duffy 1986; McKenna and Shluger
40 2009; Wang et al. 2011). Such interfacial features are likely to be highly sensitive to pressure as
41 suggested by some calculations (Harris et al. 1999; Verma and Karki 2010). Much less is,
42 however, known about the way pressure controls the ability of GBs to accommodate native and
43 impurity defects, and influence diffusion rates.

44 Here, we investigate the effects of pressure on the interplay between point defects and GBs in
45 materials by considering MgO - a prototype metal-oxide and a technologically important
46 ceramic. Moreover, (Mg,Fe)O – thought to be the second most abundant mineral in the Earth's
47 lower mantle, which is essentially a polycrystalline rock subject to high pressure up to 136 GPa,
48 and to that end, the knowledge about grain boundary properties can have profound
49 geophysical/geochemical implications. In particular, the importance of grain boundaries for
50 understanding of chemical distribution in the mantle has been widely discussed (e.g., Hiraga et
51 al. 2004; Hayden and Watson 2007; Dohmen and Milke 2010). Extant experimental
52 measurements are limited to the ambient or low pressures. The tilt boundaries in MgO were
53 observed as atomic resolution images (e.g., Yan et al. 1998; Wang et al. 2011). Fast diffusion was
54 measured for cations and oxygen, and siderophile elements in polycrystalline MgO (Oishi and
55 Kingery 1960; Wuensch and Vasilos 1966; Hashimoto and Hama 1971; Chang et al. 1981;
56 Kizuka et al. 1998; Hayden and Watson 2007). To extend our understanding of the GBs to high-
57 pressure regime, we perform the density-functional theory-based computation, which is widely
58 accepted to make accurate prediction at the macroscopic (bulk) level as well as allow access to
59 atomistic information. Previous studies were mostly based on pair potentials, which allow fast

60 computation but have a disadvantage of being empirical (e.g., Harris et al. 1999; Browning et al.
61 1999; Karki and Kumar 2007; Adjaoud et al. 2012).

62 The first principles approach has so far been applied to tilt grain boundaries in materials with
63 focus on their atomic and electronic structures (Yan et al. 1998; McKenna and Shluger 2008;
64 2009; Ghosh and Karki 2014). We have previously simulated $\{n10\}/[001]$ tilt boundaries in
65 MgO to show that the boundary structures vary considerably with pressure and also show that
66 vacancy defects favorably segregate to the boundaries (Verma and Karki 2010). Here, we extend
67 our simulations to the incorporation and migration of impurities including Ca, Al, and proton at
68 MgO tilt boundaries as a function of pressure. These defects are likely to be present in MgO
69 under most environments including those of the Earth's interior. The issues we aim to address
70 here are whether these defects segregate to grain boundaries from the bulk, and if they do, how
71 these defects migrate through the boundary region over wide pressure regime.

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METHODOLOGY

74 Computations were performed using the local density approximation and projector-
75 augmented-wave method as implemented in VASP software (Kresse and Furthmuller 1996). A
76 plane wave basis set with the cut-off of 450 eV was used to expand the valence electron wave
77 functions at Γ point. The Pulay correction of 1 to 2 GPa over the pressure regime (0 to 100 GPa)
78 studied was applied. The fully optimized configurations for the $\{n10\}/[001]$ tilt grain boundaries
79 ($n = 2, 3$ and 4) in MgO from our previous first-principles simulations (Verma and Karki 2010)
80 were used with the main focus on $\Sigma 5$ (310)/[001] (tilt angle of 36.8°) boundaries. The simulated
81 symmetric tilt GB stable at the zero pressure can be viewed as a series of dislocation pipes
82 forming a well-defined channel structure. Such open structure is less pronounced and becomes
83 denser in the asymmetric tilt GBs (in which the grains have shifted relative to each other along
84 the boundary plane) stable at 50 and 100 GPa (Verma and Karki 2010). Pressure thus suppresses
85 the excess volume, thereby affecting the type of coordination defects and the degree of bond
86 distortions at the interface. Further details on the input configuration, structural relaxation, and
87 point defects can be found in Verma and Karki (2010). For instance, we extract an Mg ion
88 leaving two valence electrons to create a cationic vacancy (V_{Mg}^{\bullet}) whereas we replace Mg ion
89 with a different cation of the same or different charge to create an impurity defect. Atomic

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90 positions were relaxed at fixed cell parameters so pressure was imposed by a constant volume
91 approach. The defect-defect coulomb interactions, arising due to periodic boundary conditions,
92 were treated as earlier (Karki and Khanduja 2006; Verma and Karki 2010). The supercells
93 contained 560 atoms for both types (symmetric and asymmetric) of grain boundaries studied.
94 Doubling the supercell did not affect the calculated results significantly.

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RESULTS AND DISCUSSION

97 Grain boundary segregation of defects

98 Unlike the bulk region, the interfacial region contains many nonequivalent sites and void
99 regions where defects can reside. Only Mg sites labeled *A*, *B*, *C* and *D* at the MgO {310}/[001]
100 tilt boundary region and interstitial sites which are at (or closest to) the boundary plane were
101 considered (Fig. 1). We have previously calculated the energetics and structures of Mg and O
102 vacancies at these boundary sites as a function of pressure and shown that both cation and anion
103 vacancies favorably segregate to the grain boundaries of MgO (Verma and Karki 2010). Here, we
104 characterize the tendency of impurities (Ca, Al and proton) to segregate to GBs by considering
105 substitution and interstitial mechanisms. For this purpose, we simulate a given defect one time in
106 the bulk region and the other time in the grain boundary region, and compare the formation
107 energies between the two cases. The segregation enthalpy of the concerned defect from the bulk
108 can be calculated as: $\Delta H_S = H_{GB} - H_{Bulk}$, where H_{GB} and H_{Bulk} are the enthalpies of the supercells
109 containing the defect in the boundary and bulk regions, respectively. A negative value of ΔH_S
110 means that the grain boundary site is energetically more favorable than the bulk site – the
111 corresponding defects from the bulk segregate to the grain boundary. Note that ΔH_S is
112 independent of both atomic and electronic chemical potentials. Also, for a charged defect, the
113 defect-defect corrections cancel out since the same correction (and band alignment) applies to
114 the both cases of the supercell: one with the defect at the bulk site and the other with the defect at
115 the boundary site (Karki and Khanduja 2006).

116 We find that all four Ca_{Mg}^{\times} configurations at the boundary have lower enthalpies than at the
117 bulk site with the site *B* being the most stable (by 0.9 eV) for the symmetric GB at zero pressure
118 (Table 1). The Ca substitution forms octahedral coordination at *B* unlike other three under-

119 coordinated sites (Fig. 1), and also each local structure is much more relaxed compared to bulk
120 site. The Ca segregation to particular atomic columns in the GB was experimentally observed at
121 the ambient pressure (Yan et al. 1998). Not all boundary sites are energetically favorable in
122 asymmetric GBs (Table 1); the sites *A* and *C* being the most and least stable, respectively. The Ca
123 segregation enthalpy increases with pressure by a factor of two reaching 2.1 eV at 100 GPa.
124 Similarly, the segregation enthalpy of $\text{Al}_{\text{Mg}}^{\bullet}$ (trivalent cation substituting Mg) was shown to
125 increase with pressure (0.2 to 0.5 eV between 0 and 100 GPa). The sites *A* and *D* are stable for
126 the symmetric boundary whereas only the site *C* is stable for the asymmetric boundaries at zero
127 and elevated pressures (Table 1). For proton defects, we show that the substitution is more
128 favorable at the boundary sites than at the bulk site with double proton preferred over single
129 proton. When two protons occupy a vacant Mg site, a $2\text{H}_{\text{Mg}}^{\times}$ defect is created. At zero pressure,
130 the boundary site *B* is preferred most with segregation enthalpy of -1.14 eV, which increases to -
131 2.63 eV at 100 GPa (Table 1).

132 We consider possible interstitials in the open space of the boundary versus the interstitial
133 bulk sites for $\text{Ca}_i^{\bullet\bullet}$ and $\text{Al}_i^{\bullet\bullet\bullet}$ as well as host cation ($\text{Mg}_i^{\bullet\bullet}$). The GB interstitials were found to be
134 distinctly preferred by each cation type over the bulk interstitials by several eVs (> 5 eV) at all
135 pressures, indicating strong segregation to the boundary. This tendency can be attributed to the
136 ion size mis-fit effect because the interstitial cations can easily fit in the excess free space of the
137 boundary and also the boundary atoms can undergo large relaxations. The impurities behave as
138 somewhat delocalized defects in the central part of the GB void space in both symmetric and
139 asymmetric boundaries though the sites are more constricted at the elevated pressures.

140 For interstitial proton (H_i^{\bullet} or p_i), we have located three inequivalent adsorption sites in the
141 symmetric GB open space at zero pressure, each involving binding of the proton to O (forming a
142 hydroxyl). These sites are energetically more favorable than the bulk site by up to 1.5 eV,
143 consistent with previous calculations (McKenna and Shluger, 2009). The boundary segregation
144 enthalpy of interstitial proton is -2.7 and -2.3 eV at 50 and 100 GPa, respectively (Table 1). At
145 the mid pressure, proton bridges two oxygen ions (B_L and B_R sites, Fig. 1) from two grains to
146 form a strongly bound state. An interstitial proton induces an unoccupied electronic state below
147 the conduction band and can trap an electron to form H atom, whose segregation enthalpy
148 increases from -1.2 eV at 0 GPa to -2.5 eV at 100 GPa.

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150 **Defects migration at grain boundary**

151 To study ionic self-diffusion at grain boundaries, we locate saddle points and possible
152 pathways for migration of host ions and impurities by trying a series of positions. For each
153 position, the migrating ion was allowed to have a perpendicular movement and the rest of the
154 super-cell was allowed to relax around it as in previous studies (Harris et al. 1997; Verma and
155 Karki 2009). Also, the migration energy for a charged defect is not influenced by the defect-
156 defect corrections (Karki and Khanduja 2006). The (310)/[001] tilt boundary involves two types
157 of diffusion pathways, one is *along* the boundary (the perpendicular direction to the tilt axis) and
158 the other is *down* the boundary (the direction of the tilt axis). These two directions are
159 represented by the *y*- and *z*-axes, respectively (Fig. 1 and 2), and note that two grains in the
160 simulated bi-crystal system are rotated about the *z*-axis (the tilt axis). Ionic diffusion along the
161 boundary involves migration between the open channels whereas the diffusion down the
162 boundary is confined within single GB core.

163 The directions of migration (via vacancy hopping) and associated activation enthalpies for
164 the host ions (Mg^{2+} and O^{2-}) in the symmetric (0 GPa) and asymmetric (100 GPa) tilt boundaries
165 are shown in Figure 2. The calculated migration enthalpy varies considerably for jumps between
166 consecutive sites, and the trend is similar between the cation and anion (Table 2). The vacancy
167 hopping along the boundary (*y*-direction) occurs from one dislocation pipe to an adjacent one so
168 migrating ion remains on the same grain. The activation energies for most steps are lower than
169 the bulk migration enthalpies, and the diffusion-rate determining step occurs between the sites *C*
170 and *D* with activation enthalpies of 2.3 eV (cation) and 2.2 eV (anion) for the {310} symmetric
171 boundary. For migration down the boundary (*z*-direction), the most favorable step occurs
172 between the opposite faces of the boundary via A_L and A_R sites with activation energies of 1.5 eV
173 (cation) and 1.3 eV (anion). Other energetically competitive path involves hopping between *B*
174 and *C* sites with migrating ion remaining on the same grain. The calculated migration enthalpies
175 for the {210} and {410} symmetric tilt grain boundaries at zero pressure are similar to the values
176 for the {310} boundary (Table 2).

177 We find that pressure affects the boundary diffusion in a more complicated way than the bulk
178 ionic diffusion (Ita and Cohen 1997; Harris et al. 2001; Van Orman et al. 2003; Karki and

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179 Khanduja 2006). The predicted enthalpies for diffusion along the boundary strongly increase
180 with pressure and remain comparable with the corresponding bulk values (Table 2, Fig. 2).
181 However, their values for diffusion down the boundary at 50 and 100 GPa are slightly smaller
182 and larger, respectively, than those at the zero pressure (Table 2). The paths between B_L to B_R
183 constitute the most favorable pathway downward with the migration enthalpies (of 1.2 to 1.7 eV
184 at high pressures) being smaller by a factor of 2 to 3 compared to the bulk values (Fig. 2). It is
185 remarkable that the predicted grain boundary diffusion in MgO is highly anisotropic with much
186 faster migration down the boundary whereas the bulk (lattice) diffusion is isotropic.

187 We now discuss the migration of three cations (Mg_i^{2+} , Ca_i^{2+} and Al_i^{3+}) via interstitials in the
188 boundary region. At zero pressure, an impurity ion migrates down the boundary (z -direction) by
189 always remaining close to the center of the GB void space (Fig. 3, S1). The energy barriers are
190 small: 0.4 eV for Mg, 0.2 eV for Ca, and 0.7 eV for Al (Table 3) so these cations can easily
191 migrate down the boundary, and can cover large distances as shown by our direct first-principles
192 molecular dynamics simulations at 1000 K (S2 and S3). The impurity cation also shifts in the x - y
193 plane away from the center of void space at elevated pressures (Fig. 3). The diffusion of impurity
194 cation along the boundary also requires a cooperative process involving migration of host cation.
195 The migrating cation may eventually knock out Mg ion from one of two closest A_L and A_R sites
196 (Fig. 3, S1). The maximum barrier occurs just after the migrating cation crosses the B_L - B_R plane,
197 and the barrier height and location differ somewhat among three cationic types.

198 At high pressures, the migration has to cover dense (congested) regions separating two
199 successive void spaces and requires interstitial transitions involving more ions. At 100 GPa, the
200 cation passes through the B_L - B_R plane by pushing the nearest host cations away, and then passes
201 via one side (left or right) of site A (Fig. 3). It may eventually knock out Mg ion from at D to the
202 void space. This knocking out process occurs in two steps at 100 GPa (S4). The diffusion down
203 the boundary is essentially confined in the void region and may not require any such transition.
204 These differences make the cationic diffusion at the grain boundary highly anisotropic with the
205 activation enthalpies for migration along the boundary being larger than those for migration
206 down the boundary at all pressures (Table 3).

207 In the boundary region, a migrating proton tends to always remain close to O ion thereby
208 switching from one O-H bonding state to other (i.e., hopping between O ions). When the proton

209 migrates down the boundary (z -direction), it follows two energetically competitive paths at zero
210 pressure, with the saddle point lying at the boundary plane, as shown in Figure 4. One path is
211 bulged towards the B_L - B_R line (in the y -direction) with the saddle point lying at the intersection
212 of this line with the boundary plane. The other path is bulged in the opposite direction (i.e.,
213 towards the A_L - A_R line). In either case, the migrating proton switches side (from left to right or
214 vice versa). On the other hand, while migrating along the boundary (y -direction), the proton
215 remains bound to O ions on the same side. The saddle point lies at the A_L - A_R line. The calculated
216 activation enthalpies for proton migration *along* and *down* the boundary are comparable (1.0 to
217 1.2 eV) at zero pressure (Table 3). The effects of pressure on proton migration are non-
218 monotonic: The migration enthalpies for both pathways at 50 GPa are larger than those at zero
219 pressure and 100 GPa. This can be associated with more strongly bound proton site at the mid
220 pressure as pointed out earlier. It is important to note that the boundary saddle point energies at
221 all pressures are always lower than any bulk interstitial and saddle point energies. This means
222 that proton migration should remain confined in the grain boundary region.

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IMPLICATIONS

225 Many defects of both native and impurity types in multiple charge states other than those
226 considered here are possible. This is more so in MgO because of its existence in the Earth's
227 mantle, which is thought to contain many other elements including Fe, C, S, Cr, Na, K, Si, etc. in
228 varying amounts. To maintain overall charge neutrality, defects of opposite charges must be
229 present. For instance, two substituted Al ions require one charge-compensating Mg vacancy. If
230 Al impurity exists as GB interstitial, 3 vacancies are formed for two Al ions. Thus, boundary
231 interstitial cations can induce more vacancies, preferably, at boundary sites, perhaps forming
232 bound defect complexes. Al impurity and Mg vacancies can form stable $2Al_{Mg}^{\bullet} + V_{Mg}''$
233 configuration (Ammann et al. 2012) that we have also found to favor the boundary over the bulk.
234 It is important to note that impurity-controlled defects are of extrinsic origin (i.e., they are not
235 thermally produced). Our calculated migration enthalpies thus represent effective activation
236 enthalpies (E_A 's) for extrinsic diffusion, which can be used in the Arrhenius equation for the
237 diffusion coefficient: $D = D_0 \exp[-E_A / RT]$ (where D_0 , R and T are the pre-exponential
238 diffusion coefficient, gas constant and temperature, respectively) to make some quantitative

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239 comparison between the grain boundary and bulk diffusion contributions. Considering E_A 's of
240 ~ 1.5 eV for migration down the boundary (and assuming the same pre-exponential factors for the
241 boundary and bulk diffusion), we predict GB diffusion rates (D_{GB}) to be several orders (10^4 to
242 10^6) magnitudes higher than the bulk (lattice) diffusion rates (D_L). Our predicted D_{GB}/D_L ratios
243 are comparable with those estimated for dominant grain boundary diffusion of different ions for
244 grain size of 1 μm or smaller expected in mantle minerals (Dohman and Mike 2010).

245 Our first-principles calculations suggest that not only vacancy defects but also impurities
246 readily segregate to the boundary, and pressure increasingly stabilizes all types of defects at the
247 boundaries, i.e., enhances segregation effects. Moreover, pressure enhancement of the ion
248 mobility at the boundary region relative to the bulk (lattice) becomes more effective at higher
249 pressures. The predicted high defect activity can be considered to arise from the electrostatic and
250 elastic forces present at the boundary region increasing with pressure. Our results can have
251 profound implications for the lower mantle because (Mg,Fe)O is considered to be the second
252 most abundant material of this region. First, the predicted large segregation effects and low
253 activation migration enthalpies imply that the grain boundaries could be regions of primary
254 storage for point defects (impurities) in mantle rocks, consistent with suggestions based on low-
255 pressure experimental observations of large segregation of incompatible elements, in particular,
256 (Yan et al. 1998; Hiraga and Kohlstedt 2007). Also, the ionic transport in the deep mantle
257 mediated by grain boundary diffusion could cover long distances (several kilometers) over the
258 geological time scales thereby controlling geochemical processes via elemental redistribution
259 and fractionation (Hiraga et al. 2007; Hayden and Watson 2007). Second, enhanced ionic
260 diffusion could facilitate creep process, which is thought to drive mantle convection (Poirier
261 1985). To further assess these implications, we anticipate further study of more grain boundaries
262 and their effects on point defects in other key mantle minerals. Also important is to study the
263 effects of temperature on the boundary structures, and the boundary incorporation and migration
264 of defects, which we anticipate to simulate using the first-principles molecular dynamics method.

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352 **TABLE 1.** Segregation energies (in eV) for various impurities (Ca, Al, and proton) at different
 353 sites (*A*, *B*, *C*, and *D* shown in Fig. 1) of the {310} tilt symmetric (0 GPa) and asymmetric (0, 50
 354 and 100 GPa) boundaries. All energies are defined with respect to the bulk site of the boundary-
 355 bearing supercell. Mg-vacancy segregation energies are shown for comparison.

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	Site	V_{Mg}''	$\text{Ca}_{\text{Mg}}^{\times}$	$\text{Al}_{\text{Mg}}^{\times}$	$2\text{H}_{\text{Mg}}^{\times}$	p_i
{310} sym 0 GPa	<i>A</i>	0.05	-0.10	-0.19	-1.00	-1.5
	<i>B</i>	0.18	-0.93	0.02	-1.14	
	<i>C</i>	-0.21	-0.72	0.17	-0.77	
	<i>D</i>	0.73	-0.05	-0.13	-0.64	
{310} asym 0 GPa	<i>A</i>	0.13	-0.03	0.01	-0.79	-1.5
	<i>B</i>	-1.30	0.19	0.31	-0.46	
	<i>C</i>	0.24	0.47	-0.17	-0.85	
	<i>D</i>	-1.13	0.00	0.37	-1.06	
{310} asym 50 GPa	<i>A</i>	1.17	-0.42	0.02	-1.41	-2.7
	<i>B</i>	-1.82	-0.15	0.62	-2.27	
	<i>C</i>	1.16	1.52	-0.37	-0.90	
	<i>D</i>	-1.21	0.0	0.54	-2.15	
{310} asym 100 GPa	<i>A</i>	-1.01	-2.10	0.63	-1.42	-2.3
	<i>B</i>	-2.67	-1.73	0.77	-2.37	
	<i>C</i>	-0.90	0.36	-0.53	-0.98	
	<i>D</i>	-2.50	-0.54	0.16	-2.63	

386 **TABLE 2.** Activation enthalpies (in eV) for Mg²⁺ and O²⁻ ion migration *along* and *down* the
 387 {310}, {210} and {410} symmetric boundaries at 0 GPa and {310} asymmetric boundaries at 50
 388 and 100 GPa. The bulk values are from Karki and Khanduja (2006). Previous MD results for the
 389 {410} symmetric boundary are from Harris et al. (1997).

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	Along boundary		Down boundary		Bulk	
	Mg	O	Mg	O	Mg	O
{310}, 0 GPa	2.1	2.2	1.5	1.3	2.6	2.4
{210}, 0 GPa	2.3	2.4	1.2	1.4		
{410}, 0 GPa	2.3 (2.3)	2.2 (2.3)	1.7 (1.1)	1.5 (1.0)		
{310}, 50 GPa	3.1	3.9	1.2	1.2	3.2	3.5
{310}, 100 GPa	3.5	4.3	1.6	1.7	3.8	4.2

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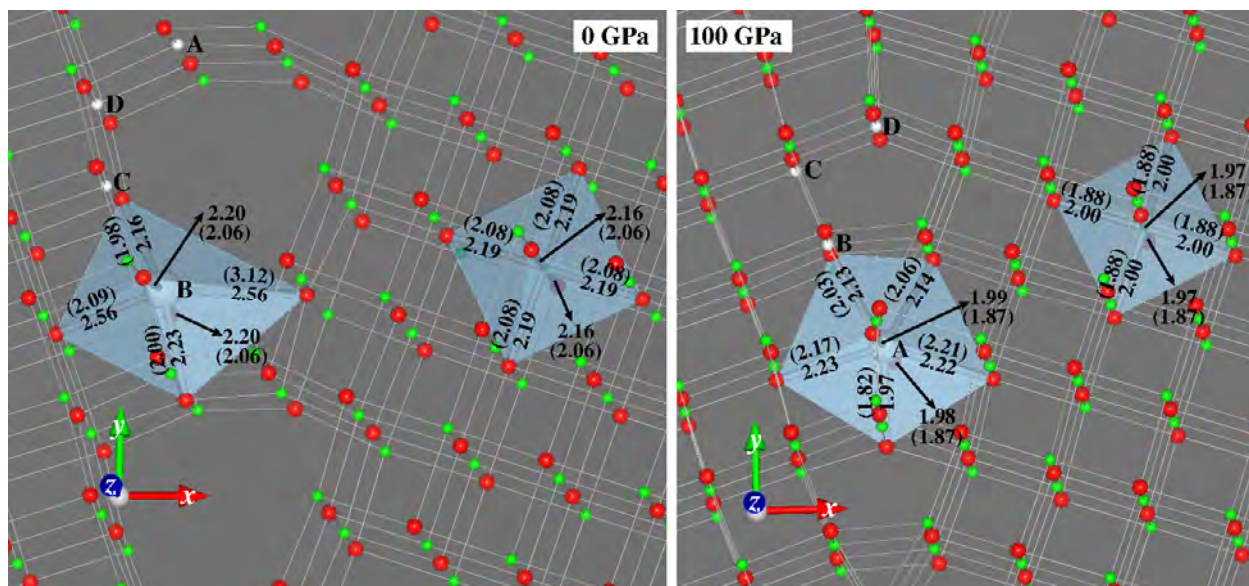
396 **TABLE 3.** Activation enthalpies (in eV) for interstitial migration of impurities (Ca^{2+} , Al^{3+} ,
397 and proton) and host cation (Mg^{2+}) *along* and *down* the {310} symmetric (0 GPa) and
398 asymmetric (50 and 100 GPa) boundaries.

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	Along boundary				Down boundary			
	Mg	Ca	Al	H	Mg	Ca	Al	H
0 GPa	2.6	3.0	3.2	1.2	0.4	0.2	0.7	1.0
50 GPa	2.3	2.6	1.9	2.1	0.8	1.5	1.3	3.0
100 GPa	2.4	2.7	2.1	1.2	1.3	1.3	1.2	1.9

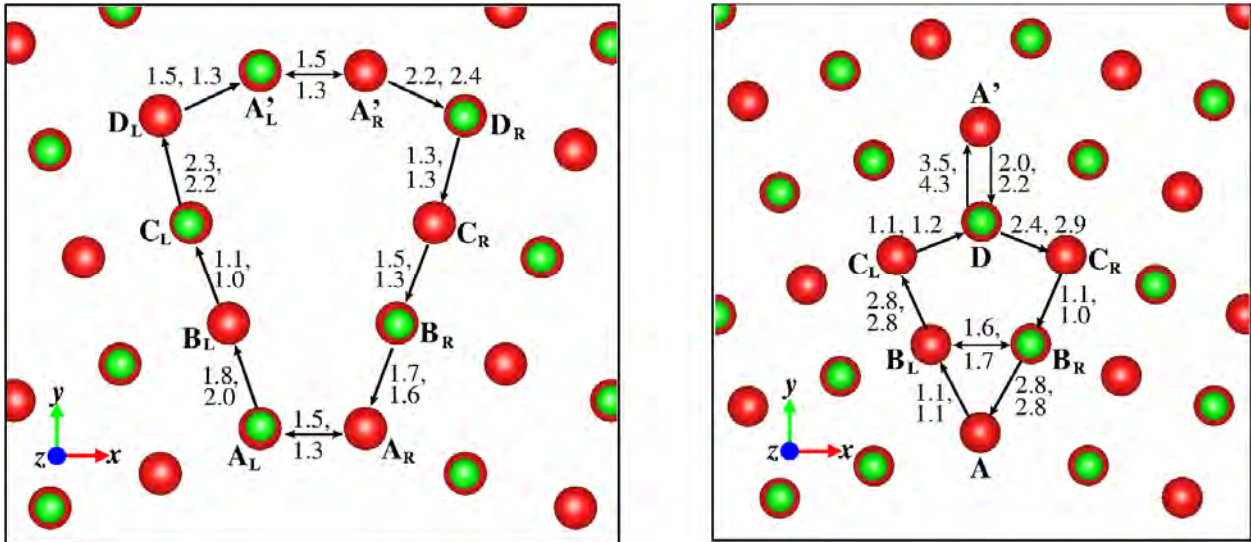
400 **FIGURE 1.** Local structures of Ca impurity substituted the boundary and bulk sites at zero
401 pressure (left) and 100 GPa (right) shown as polyhedra. Ca-O bond lengths are compared with
402 Mg-O bond lengths (parenthesis). All boundary sites (A, B, C and D) are also marked. The green
403 (small) and red (large) spheres represent Mg and O atoms, respectively.

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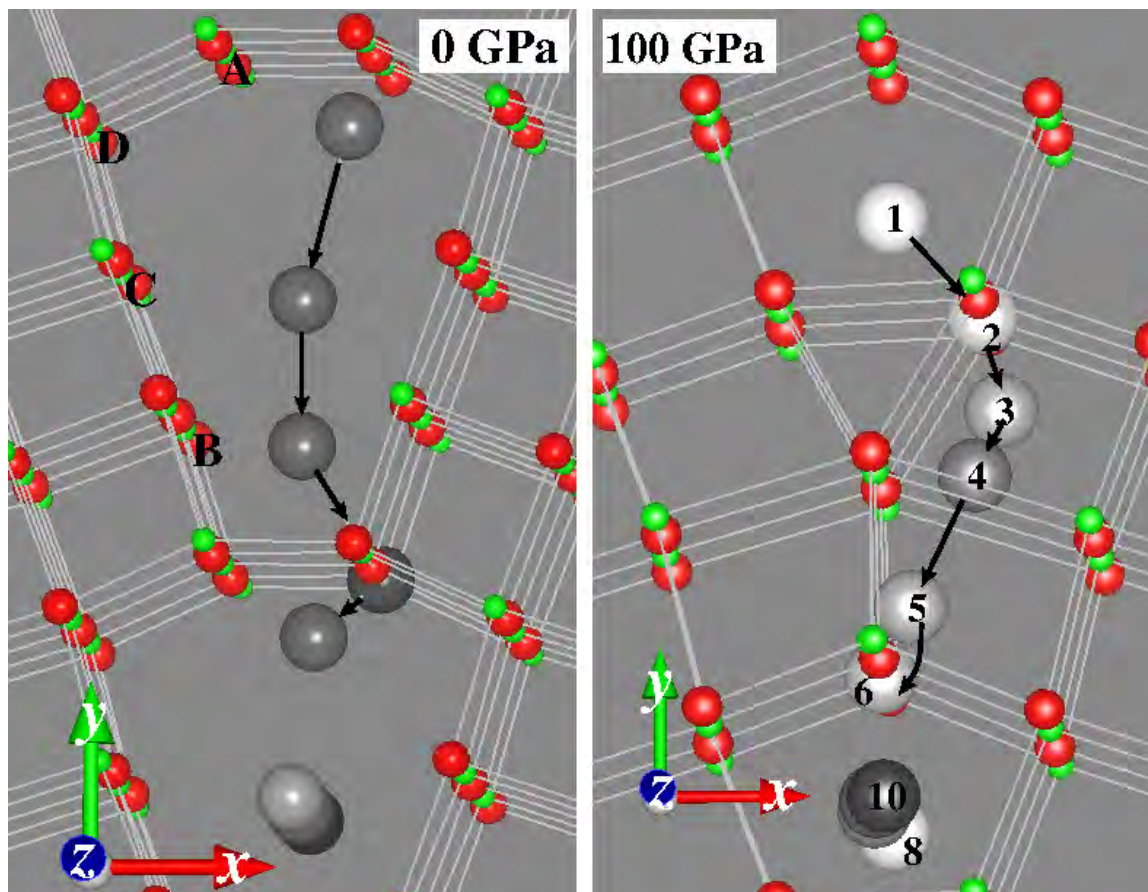
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406 **FIGURE 2.** Local migration directions and energies (in eV) for Mg²⁺ ion (first number) and O²⁻
407 ion (second number) in the {310} symmetric boundary at 0 GPa (left) and asymmetric boundary
408 at 100 GPa (right). The green (small) and red (large) spheres represent Mg and O atoms,
409 respectively. The sites are labeled with the subscript L for the left grain and R for the right grain.
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423 **FIGURE. 3.** Paths for cation (Mg/Ca/Al) migration at 0 and 100 GPa. The successive positions
424 of the migrating ion are displayed by large sphere with gray color representing the z -coordinate.
425 The green (small) and red (large) spheres represent Mg and O, respectively.



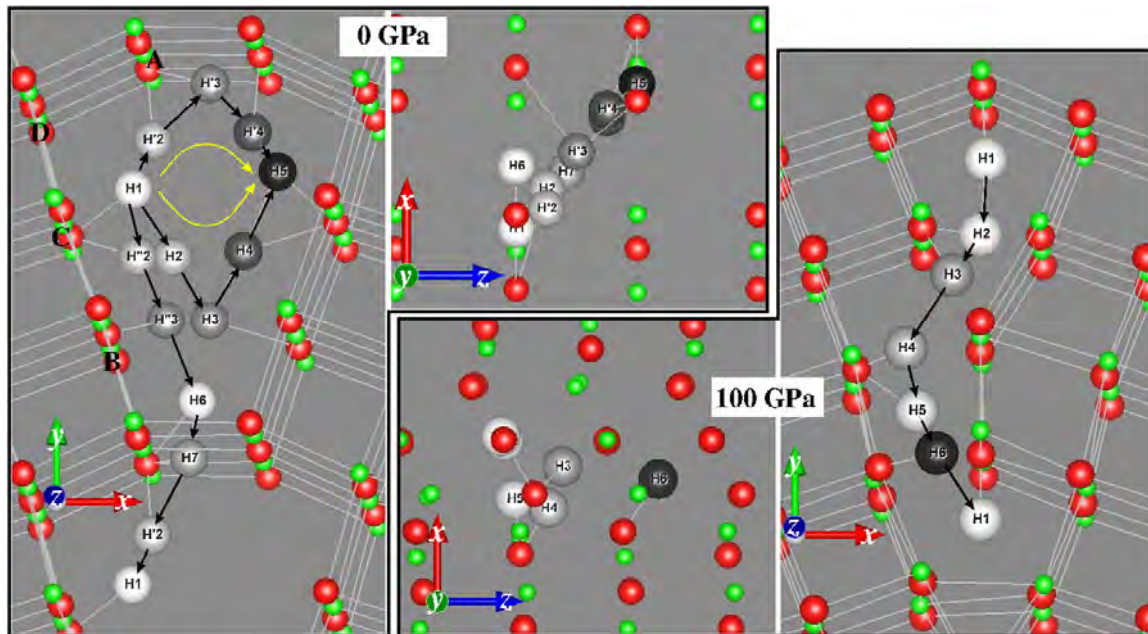
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437 **FIGURE 4.** Predicted migration paths for interstitial proton at 0 and 100 GPa. Proton at its
438 successive positions is shown by large sphere with grey-scale representing the z -coordinate.
439 Some Mg-O bonds at GB are removed for the clarity. The green (small) and red (large) spheres
440 represent Mg and O, respectively.

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