

1 **Spatio-temporal constraints on magma storage and ascent conditions in a transtensional**  
2 **tectonic setting: The case of the Terceira Island (Azores)**

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4 Revision 2

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14 **Keywords**

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25 **Abstract**

26 The mafic magmatism of the last 50 ka on Terceira Island, Azores archipelago, occurred  
27 along three segments of the fissure zone that crosses the island. The two subaerial segments  
28 developed with different trends over pre-existing, quiescent or extinct, central volcanoes. The  
29 Serreta submarine ridge is the offshore segment of the fissure zone which erupted recently in AD  
30 1998-2001. The combined study of CO<sub>2</sub>+H<sub>2</sub>O fluid inclusions hosted in mafic minerals and rock  
31 geochemistry of the magmas, reveals different storage and ascent conditions among the fissure  
32 zone segments. The maximum pressure of fluid trapping for all the fissure systems occurred at  
33 the Moho Transition Zone, between 498 MPa and 575 MPa (20.3-21 km deep). At this depth  
34 interval all magmas stagnated for some time, before ascending towards the surface, experiencing  
35 fractional crystallization and degassing. Magmas of the southeastern and Serreta segments of the  
36 fissure zone ascended rapidly through the crust without further stops. Those of the central  
37 segment experienced a multi-step ascent, with fluid trapping at 406 MPa and 209 MPa (16.5-8.5  
38 km deep) and associated geochemical evolution towards trachybasalt.

39 The magma ascent below the different segments of the fissure zone varies from almost  
40 isochoric at the submarine segment, associated with minimum re-equilibration of the inclusions,  
41 to polybaric slow ascent at the central segment, associated to almost complete re-equilibration of  
42 the inclusions. Variable degrees of re-equilibration and multi-step ascent may be linked to both  
43 the presence of pre-existing intracrustal crystallized bodies of more evolved composition and the  
44 stress field acting on this area. The latter responds to the local and shallow conditions related to  
45 the presence of older central volcanoes and to the main regional spreading direction of the  
46 Terceira Rift, which at regional scale, is approximatively orthogonal to the fissure zone axis.

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## INTRODUCTION

50 The location and structure of magma storage systems, the amount of melt available for  
51 mobilization prior to eruption, and its path and rate of ascent are all fundamental information to  
52 achieve a better understanding of the behavior of volcanic systems. These elements are known in  
53 well-monitored and frequently erupting basaltic volcanoes, such as Etna, Stromboli, Kilauea and  
54 Piton de la Fournaise, and fairly understood in those with higher associated risk but more  
55 infrequent activity (e.g. Vesuvius, La Peleé, the Cascades volcanoes). However, aside from the  
56 abovementioned cases, such information is completely unknown for many oceanic islands, such  
57 as the Azores.

58 The Azores islands are volcanic systems predominantly formed by basaltic fissure zones,  
59 where magma, collected at the Moho Transition Zone, is mobilized in response to tectonic stress  
60 (Haase and Beier 2003; Zanon et al. 2013; Zanon and Frezzotti 2013). Typically, eruptive  
61 fissures open in the same area during a limited period (a few thousands of years) before  
62 migrating to another area nearby (Hildenbrand et al. 2008). Erupted lavas are poorly evolved  
63 basalts (e.g. Beier et al. 2007; 2008; 2012), which ascend from the mantle due to extensional  
64 tectonics. The interaction of extensional fissures with transtensional faults determines the  
65 formation of short-lived and shallow-level magma reservoirs, leading to the establishment of  
66 centralized feeding conduits or a series of closely-spaced feeder dykes, and, progressively, to the  
67 development of larger volcanic systems (Miranda et al. 1998).

68 On the Azorean island of Terceira eruptions occur at both central volcanoes and various  
69 segments of a fissure zone that crosses the island and extends offshore. Magmas from central  
70 volcanoes are generally highly differentiated (mugearites-trachytes and peralkaline rhyolites),  
71 while those from the fissure zone consist of basalt-trachybasalt associations (Self and Gunn

72 1976; Mungall and Martin 1995; Madureira et al. 2011). The mafic magmas erupted in the last  
73 50 ka were emitted from different segments of the fissure zone and are characterized by a  
74 variable degree of evolution. The subaerial fissure system is located between closely-spaced  
75 central volcanoes of different age.

76 Possibly, the presence at depth of partially crystallized bodies of evolved composition,  
77 related to the central volcanoes, influenced the magma ascent at the fissure zone segments of the  
78 island, promoting the formation of intracrustal storage areas and enhancing their chemical  
79 evolution through fractional crystallization. This hypothesis is here investigated through the  
80 microthermometric study of fluid inclusions trapped in mafic minerals found in the magmas  
81 from the different segments of the fissure zone. The data of samples from the two subaerial  
82 segments and the submarine segment are compared also in terms of geochemical characteristics  
83 of the magmas. This method is a rapid and reliable way to constrain the spatial and temporal  
84 evolution of the magma storage areas at crustal depths. In fact, the same methodology has been  
85 successfully applied to various volcanic systems, including the Aeolian (Zanon et al. 2003;  
86 Bonelli et al. 2004; Zanon and Nikogosian 2004; Di Martino et al. 2010) and Canary islands  
87 (Hansteen et al. 1991; Klügel et al. 1997; Hansteen et al. 1998; Klügel et al. 2005; Galipp et al.  
88 2006; Stroncik et al. 2009). More recently, it has been also applied to the islands of Pico and  
89 Faial, also in Azores (Zanon and Frezzotti 2013), where basalts are mobilized by extensional  
90 tectonics and ascend from the ponding area located at the Moho Transition Zone, without further  
91 stops in the crust.

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## GEOLOGICAL SETTING

95 Terceira Island is located in the central-north Atlantic Ocean, near the triple junction among  
96 the North American, Eurasian and Nubian plates. This area of the Atlantic is characterized by a  
97 complex geodynamic setting with the presence of the Mid-Atlantic Ridge to the west and the  
98 Gloria strike-slip fault to the east (Fig. 1a). A regional WNW-ESE trending system (the so called  
99 Terceira Rift), characterized by a slow and asymmetric spreading rate from northeast to the  
100 southwest (e.g. Marques et al. 2013), is located between the two abovementioned structures.

101 Terceira is formed by two extinct and two quiescent central volcanoes and a diffuse fissure  
102 zone (Fig. 1b). Ages of all eruption products show an overall east to west younging progression,  
103 although with contemporaneous periods of activity (Calvert et al. 2006 and references therein).  
104 Cinco Picos and Guilherme Moniz are the two oldest and extinct volcanoes (>388 ka and >270  
105 ka, respectively; Calvert et al., 2006), located in the eastern and central part of the island. The  
106 calderas are filled by more recent lavas erupted from the fissure zone. Pico Alto Volcano is older  
107 than 141 ka (Gertisser et al. 2010) and is located north of Guilherme Moniz, while the western  
108 third of the Terceira is formed by the well-preserved conical-shaped volcano of Santa Bárbara,  
109 the age of which is unknown.

110 The fissure zone crosses the island along a general WNW-ESE direction and extends  
111 offshore, forming submarine ridges. Its presence is marked by alignments of scoria and spatter  
112 cones and fissures which erupted basaltic (*sensu lato*) lavas. In comparison with fissure systems  
113 present on other islands of the archipelago, these alignments are made up of fewer cones,  
114 suggesting a lower number of eruptive events. A careful observation of the Digital Elevation  
115 Model (DEM) in Figure 1 reveals that the fissure zone is segmented. The overall WNW-ESE  
116 direction is interrupted in the area of Cinco Picos and Guilherme Moniz calderas, where the  
117 possible effect of NW-SE tectonics has exerted a clockwise rotation of the eruptive axis.

118 The age of the fissure zone is currently constrained between late Pleistocene (>43 ka; Calvert  
119 et al. 2006) and the present, although some activity may be older. The eruptive vents show a  
120 general progressive younger age from southeast to northwest, with historical eruptions inland  
121 (AD 1761) and offshore in the Serreta submarine ridge (AD 1867 and AD 1998-2001; Fig. 1b).

122 The tectonics of the island responds to the dynamics of the WNW-ESE trending Terceira  
123 Rift, as most structural features are sub-parallel to this regional direction (Fig. 1b; Self 1976;  
124 Tzanis and Makropoulos 1999; Quartau et al. 2014). The main feature is the NW-SE oriented  
125 Lajes graben in the northeast sector of the island. Other WNW-ESE to NW-SE trends including  
126 volcano-tectonic alignments of vents and faults provide evidence of a transtensive regime on the  
127 island.

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## METHODS

131 For this study we selected porphyritic lava samples emitted from the fissure zone segments  
132 across the island and offshore. Rocks containing olivine and clinopyroxene were chosen for the  
133 study of fluid inclusions, as this mineralogy is representative of the early stages of evolution of  
134 magmas.

135 Whole-rock analyses were performed by Activation Laboratories (Canada). Alkaline  
136 dissolution with lithium metaborate/tetraborate followed by nitric acid was used on 1 gram of  
137 rock powder before being fused in an induction furnace. The melt was then poured into a  
138 solution of 5% nitric acid containing cadmium as an internal standard and continuously stirred  
139 until complete dissolution was achieved (~30 minutes). The samples were contemporaneously  
140 analyzed by a Perkin Elmer 9000 inductively coupled plasma mass spectrometer (ICP-MS) and  
141 an Agilent 735 inductively coupled plasma atomic emission spectrometer (ICP-AES). Analytical

142 precision ( $2\sigma$ ) was better than 5% for most major elements and 10% for most minor and trace  
143 elements. Nine international rock standards were used to calibrate the two methods. Whole-rock  
144 and mineral chemistry are reported in the supplementary material.

145 A JEOL JXA 8200 Superprobe, equipped with five wavelength-dispersive spectrometers,  
146 energy-dispersive spectrometer and cathodoluminescence detectors (Dipartimento di Scienze  
147 della Terra, “Ardito Desio” University of Milan, Italy) was used to analyze mafic crystalline  
148 phases and glasses. A spot size of 1  $\mu\text{m}$  with a beam current of 15 nA was used for the mineral  
149 phases, whereas a spot size of 5-7  $\mu\text{m}$ , according to the available surface to be analyzed, and a  
150 beam current of 2 nA were applied to glasses. Count times were 30 s on the peak and 10 s on  
151 each background. Natural and synthetic minerals and glasses, used as standards, were used for  
152 calibration within 2% at  $2\sigma$  standard deviation. Raw data were corrected applying a Phi-Rho-Z  
153 quantitative analysis program. The typical detection limit for each element is 0.01%.

154 Geobarometric data were obtained by microthermometry of the fluid inclusions contained in  
155 mafic phenocrysts. About 250/300 olivines and 100/150 clinopyroxenes up to 500  $\mu\text{m}$  in size  
156 were handpicked from crushed rock, cleaned with deionized water, embedded in acetone-soluble  
157 epoxy and doubly polished up a final thickness of 120-80  $\mu\text{m}$ .

158 Microthermometry of fluid inclusions were carried out on a Linkam MDSG600 heating-  
159 cooling stage, calibrated using synthetic  $\text{H}_2\text{O}-\text{CO}_2$  inclusions (triple point for  $\text{CO}_2$  and  $\text{H}_2\text{O}$  at -  
160 56.6  $^\circ\text{C}$  and 0  $^\circ\text{C}$ , respectively). Melting and homogenization temperatures are reproducible to  
161  $\pm 0.1$   $^\circ\text{C}$ . For all the runs, the heating rate was in the range 0.2–0.5  $^\circ\text{C}/\text{min}$ . The equation of Span  
162 and Wagner (1996) was used to calculate density values of pure  $\text{CO}_2$  fluid. These values were  
163 then corrected for the probable presence of 10 mole% of  $\text{H}_2\text{O}$ , following the method suggested in  
164 Hansteen and Klügel (2008). Finally, isochores for  $\text{H}_2\text{O}-\text{CO}_2$  fluids were derived using the

165 model of Sterner and Bodnar (1991) by means of the computer program FLUIDS (Bakker,  
166 2003). Despite the fact that the model is experimentally calibrated up to 700 °C and 600 MPa, its  
167 application to higher temperatures provided a good fit. Final pressure calculations were obtained  
168 from the intersection of isochores with the maximum temperature of crystallization inferred from  
169 olivine thermometers (Putirka 2008; Beattie 1993; Shejwalkar and Coogan 2013).

170 Finally, direct density measurements were carried out on some samples with a resolution of 1  
171  $\text{kg}\cdot\text{m}^{-3}$ , using a MD 200s electronic densimeter and corrected for porosity.

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#### SAMPLE DESCRIPTION

175 The studied samples are lavas erupted from the different segments of the fissure zone between  
176 less than 50 ka and AD 2001 (Fig. 2). The main petrographic characteristics of these samples are  
177 summarized in Table 1. In contrast to lavas from Pico, Faial and São Miguel islands (Zanon et al.  
178 2013; Zanon and Frezzotti, 2013; Zanon 2014), the samples are almost aphyric (Fig. 3a) to  
179 poorly porphyritic, with a maximum phenocrysts content of ~22 vol% (Fig. 3b) and rare  
180 occurrences of megacrysts of olivines in only a few samples. The mineral assemblage consists of  
181 phenocrysts of olivine, clinopyroxene, plagioclase  $\pm$ oxides in variable proportions. Common  
182 textures are intersertal and intergranular, with a few glomeroporphyritic aggregates, while only  
183 in the lava balloons, from the AD 1998-2001 Serreta submarine eruption, textures range from  
184 vitrophyric with cryptocrystals to intersertal due to the extent of seawater quenching.

185 Olivine phenocrysts up to 2 mm in size are present in some samples in variable amounts ( $\leq 13$   
186 vol%). They are euhedral or subhedral, or rarely skeletal, and the core Mg number [ $\text{Mg}\# =$   
187  $100 \cdot \text{Mg}/(\text{Mg}+\text{Fe})$ ] ranges from 80 to 90%. Fluid inclusions are present in ~2-4% of these

188 crystals. This percentage increases to 8% only in the olivines from the lava balloons.  
189 Microphenocrysts <0.5 mm in size (Fig. 3b) are much more abundant and present in all samples,  
190 but are devoid of fluid inclusions.

191 Clinopyroxene phenocrysts are typically  $\leq 1$  vol% in most samples, with augite and augitic  
192 diopside composition ( $Wo_{33-38}$ ,  $En_{38-48}$ ,  $Fs_{3-11}$ ). They are <5 mm in size, euhedral and zoned, with  
193 rare evidences of disequilibrium (i.e. embayments and reaction rims). Their core Mg number  
194 [ $Mg\# = 100 \cdot Mg / (Mg + Fe^{+2} + Fe^{+3} + Mn)$ ] varies from 81-86% in the subaerial lavas and from 71-76  
195 in the lava balloons. Only rare crystals contain fluid inclusions. Microphenocrysts (~1.5 mm in  
196 size) are more abundant but devoid of fluid inclusions.

197 Plagioclase phenocrysts up to 5 mm in size are rare ( $\leq 4$  vol%) and are characterized by  
198 euhedral, either tabular or acicular morphology and oscillatory or reverse zoning. Their  
199 composition varies from andesine to labradorite ( $An_{63-72}$ ). Rare anhedral crystals with signs of  
200 disequilibrium (sieve texture) and bytownitic composition ( $An_{80}$ ) have also been found. No fluid  
201 inclusions have been found here. Plagioclase microphenocrysts (<0.5 mm) are abundant in many  
202 samples.

203 Finally, rare titanomagnetites are present as phenocrysts ( $\leq 0.2$  vol%) in few samples, while are  
204 common in all samples as microphenocrysts. Apatite and ilmenite are sometimes present in the  
205 groundmass.

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## ROCK GEOCHEMISTRY

209 The studied compositions are transitional to mildly alkaline basalts and trachybasalts (Fig.  
210 4a) with MgO content ranging from 4.5 to 11.6 wt%. The lowest MgO value characterizes a lava

211 of the central segment of the fissure zone (Cfz), while the highest derives from a more  
212 porphyritic rock of the southeastern segment (SEfz). With the exception of three samples which  
213 are mildly silica-undersaturated ( $ne \leq 2.14\%$ ), the others are silica-saturated, with up to 8.7%  
214 normative hypersthene (Table 2).

215 All major elements, except CaO, show a general increase with decreasing MgO, however  
216 data are scattered and not linearly correlated. Samples from the Cfz have lower MgO content  
217 ( $\leq 6.5$  wt%) and show, on average, higher Na<sub>2</sub>O, K<sub>2</sub>O, P<sub>2</sub>O<sub>5</sub> and TiO<sub>2</sub> contents than the  
218 corresponding rocks from the SEfz. The variations of compatible trace elements such as Ni, Sc,  
219 Cr, Cu and Co are positively correlated with MgO, whereas Large Ion Lithophile Elements  
220 (LILE: Rb, Ba, K, Cs) and High Field Strength Elements (HFSE: Ta, Nb, Zr, Hf, Th, U) have  
221 opposite trends.

222 All trace element patterns are quite similar, with marked K and Pb negative spikes and  
223 limited Th, U, Hf and Zr fractionation (Fig. 4b). The rocks from the Cfz, which are the most  
224 evolved, are consequently the most enriched in incompatible elements when compared to the  
225 composition of the primordial mantle (McDonough and Sun, 1995). All patterns of Rare Earth  
226 Elements (REE) are smooth and slightly enriched in lighter elements and without the Eu  
227 anomaly (Fig. 4c).

228 The variations of the ratio CaO/Al<sub>2</sub>O<sub>3</sub>, correlated to those of FeO<sub>T</sub>/MgO, and the positive  
229 correlation between the variations of compatible trace elements and MgO qualitatively would  
230 suggest that all these compositions are related to each other by a process of fractional  
231 crystallization, involving the removal of stable proportions of mafic phases from a common and  
232 more primitive parental magma (Fig. 5a). However, LILE and U and Th are incompatible within  
233 the structures of both clinopyroxene and olivine, therefore, during the fractional crystallization of



234 these phases ratios among these elements in the parental liquid cannot change. In our  
235 compositions these ratios are very scattered (Fig. 5b). Similar conclusions are derived from the  
236 analysis of the variation among light and heavy REE. These elements are generally incompatible  
237 in both olivine and clinopyroxene, although they have variable partition coefficients. Variations  
238 in the ratios of these elements are too large for a process of removal of mafic phases (Fig. 5c).  
239 Finally, the variable degree of silica saturation is not related to the Mg# of the rocks. All these  
240 elements indicate that the characteristic geochemical signature of each of these samples reflects  
241 variations of the degree of melting at the source or the extent of high pressure fractional  
242 crystallization from a pristine liquid. Therefore, each of these compositions does not result from  
243 the evolution of a common parental melt in a large magma reservoir at shallow-to-intermediate  
244 depth, but rather represents a different melt, which behaved as a single magma batch.

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#### **PETROGRAPHY AND MICROTHERMOMETRY OF FLUID INCLUSIONS**

248 In general, the occurrence of fluid inclusions in Terceira lavas is rare (Table 2); a small  
249 number of inclusions are present in ~2% of olivine phenocrysts and in ~0.5% of clinopyroxene  
250 of the lavas from the subaerial segments of the fissure zone. In the lava balloons from the Serreta  
251 submarine fissure zone (SERfz), a larger number of inclusions are present in ~10% of olivines  
252 and ~2% of clinopyroxenes. No inclusions have been found in the rare megacrysts and in  
253 microphenocrysts.

254 Textural characteristics of fluid inclusions allow us to distinguish two types of populations.  
255 Type-I inclusions are well rounded, either isolated or clustered, up to 25-35  $\mu\text{m}$  in size and  
256 randomly located at the core of phenocrysts. In some cases they are present together with small

257 crystalline phases and silicate melt inclusions between growth zones of the phenocrysts. The  
258 inclusions do not contain crystal phases. These textural characteristics indicate that these  
259 inclusions are trapped during the early stages of crystallization of the phenocrysts. Many  
260 inclusions show evidence of density partial re-equilibration i.e., micro-cracks departing from the  
261 main cavity and/or a halo of minute inclusions, around the main cavity, and/or a dark aspect (Viti  
262 and Frezzotti 2000; Viti and Frezzotti 2001).

263 Type-II fluid inclusions are more abundant in the mafic phases of the lava balloons, when  
264 compared to the subaerial samples. They commonly occur as trails of variable length and  
265 thickness, which lined completely healed fractures, limited by grain boundaries. Inclusions are  
266 typically  $<10\ \mu\text{m}$ , usually rounded and may coexist with both Type-I (Fig. 3c) and silicate melt  
267 inclusions with variable size and degree of crystallization. These inclusions formed during  
268 episodes of deformation/cracking of the host crystals after their growth.

269 Frozen inclusions melt instantaneously within a temperature interval between 57.3 and 56.4  
270 °C, with most of the data at 56.7 °C. These temperatures suggest that trapped fluids are  
271 composed of almost pure  $\text{CO}_2$  together with the presence of few moles (less than 1%; Van den  
272 Kerkhof 1990) of other volatile species (i.e.  $\text{N}_2$  and  $\text{SO}_2$ ). This small amount does not  
273 significantly affect the interpretation of entrapment conditions (Van den Kerkhof 1990; Frezzotti  
274 et al. 2002), and for this reason these inclusions can be treated as composed of pure  $\text{CO}_2$ .  
275 Clathrates are not detected but liquid  $\text{H}_2\text{O}$  was rarely optically observed in some large Type-I  
276 inclusions in olivines. According to Lamadrid et al. (2014),  $\text{H}_2\text{O}$  in small size inclusions is not  
277 detectable either optically or by Raman spectroscopy, as it forms a very thin film at the walls of  
278 the inclusion. Final homogenization occurs into the liquid ( $\text{Th}_L$ ). Only a single inclusion  
279 homogenizes into the vapor phase ( $\text{Th}_V$ ).

280 Type-I fluid inclusions in the olivines from the subaerial lavas have values of  $Th_L$  from 23.5  
281 °C to 31 °C (480-730 kg·m<sup>-3</sup>) and those in the clinopyroxenes from 30.6 °C to 31 °C (480-560  
282 kg·m<sup>-3</sup>). Inclusions inside the olivines of the lava balloons have values of  $Th_L$  from 18.8 °C to  
283 26.1 °C (690-790 kg·m<sup>-3</sup>).

284 Type-II fluid inclusions in the olivines from the subaerial lavas have values of  $Th_L$  from 24.3  
285 °C to 31 °C (480-720 kg·m<sup>-3</sup>). A single inclusion homogenized into the vapor phase at 31 °C  
286 (460 kg·m<sup>-3</sup>). Inclusions in the olivines from the lava balloons have values of  $Th_L$  from 23.7 °C  
287 to 30.2 °C (580-730 kg·m<sup>-3</sup>), while in clinopyroxenes from 27.2 °C to 31 °C (470-670 kg·m<sup>-3</sup>).  
288 The frequency-distribution histograms of  $Th_L$  of fluid inclusions in olivines and clinopyroxenes  
289 (Fig. 6a, b, c) is built by merging the data from lavas belonging to the same segment of the  
290 fissure zone. These plots show polymodal distribution of data. The same characteristics can be  
291 observed in histograms of density data distribution (Fig. 6d, e, f).

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## 294 **MAGMA PONDING AND ASCENT CONDITIONS**

### 295 **Barometric constraints from fluid inclusions**

296 After formation, a crystal may stretch or even crack, as a function of the strain rate, due to  
297 the variations of confining pressure and temperature. As result, trapped inclusions re-equilibrate  
298 to variable extents, changing their original density values. As an extreme case, new cracks form  
299 around the main inclusion cavity, allowing the fluid to redistribute into them, which results in an  
300 inclusion larger than the original. These processes are common in a magmatic assemblage during  
301 the ascent history of the carrier magma (Bodnar 2003). For this reason, all inclusions suffered at  
302 least a minimum degree of re-equilibration, although some features are too small to be resolved

303 by optical microscopy, and require the use of transmission electron microscopy (Viti and  
304 Frezzotti 2000; Viti and Frezzotti 2001). The smallest inclusions have the highest chance to have  
305 suffered minimum degrees of re-equilibration (Fig. 7) and we assume that those inclusions,  
306 without visible re-equilibration features, are the most adequate to provide information on the  
307 depth of the magma reservoir (Andersen and Neumann 2001). A latter event of fluid trapping  
308 will produce a second mode in a fluid density distribution histogram. Also in this case, the  
309 densest inclusion of this population reveals information useful to locate the magma reservoir.

310 The composition of trapped fluids is mostly CO<sub>2</sub> and few moles of other volatile species.  
311 However, H<sub>2</sub>O, as main component of exolved fluid, is expected to be originally present inside  
312 inclusions, and indeed it was directly observed at the microscope, in some rare inclusions. Either  
313 hydrogen diffused at high temperatures into the host olivine (Gaetani et al. 2012) or it probably  
314 reacted with the host mineral at low temperatures, forming hydrous carbonates (Andersen and  
315 Neumann 2001; Frezzotti et al. 2002). For these reasons, according to Dixon and Stolper (1995),  
316 Dixon (1997) and Hansteen and Klügel (2008) and here corroborated by direct observation of the  
317 presence of H<sub>2</sub>O in fluid inclusions, we hypothesized an original fluid composition at the time of  
318 trapping of  $x_{\text{H}_2\text{O}}=0.1$  and  $x_{\text{CO}_2}=0.9$ .

319 In order to obtain barometric information on the magma plumbing system, it is necessary to  
320 calculate the magmatic temperature at the moment of fluid trapping. We applied thermometers  
321 based upon the equilibrium of olivine-liquid to the groundmass glass to have a first order  
322 estimation of the eruptive temperature. Only two microlitic olivines from a lava balloon (Fo<sub>78-80</sub>)  
323 were found to be in equilibrium and provided a temperature of 1139 °C and 1146 °C (±30 °C)  
324 using the thermometer of Putirka (2008). The thermometer of Beattie (1993) yielded 1154 ±10  
325 °C. The calculation of the olivine crystallization temperature requires the compositions of melt

326 inclusions and of their hosts, but the composition published in Kueppers et al. (2012) for the AD  
327 1998-2001 submarine eruption did not equilibrate, due to post-entrapment crystallization. For  
328 this reason, we have applied a chemical thermometer based upon the Ca content in olivine  
329 (Shejwalkar and Coogan 2013), despite our olivine compositions are off the field of application.  
330 However, we hypothesized that the real conditions of crystallization of the olivines in this study  
331 were not too far from those required by this thermometer. The resulting temperatures are shown  
332 in Figure 8. Despite a minimum scattering of the data, many values peak at 1170 °C, which is  
333 only about 25-30°C higher than the eruptive temperature and 20 °C higher than the calculated for  
334 Pico and Faial islands (Zanon and Frezzotti 2013). However, the determination of a very precise  
335 temperature is not a critical element for the validation of the barometric data of fluid inclusions  
336 (Roedder 1965; 1983); in fact a difference in this temperature of  $\pm 20$  °C changes the resulting  
337 values of pressure only by  $\pm 8$  MPa (1.5%).

338 The pressure conditions for magma storage were obtained from isochores distribution at 1170  
339 °C (Fig. 9). The error associated to this calculation is  $\pm 1$  MPa. The possibility that the trapped  
340 fluids preserve their original density depends on the strength of the crystal to external stresses, on  
341 the P-T path during magma ascent. The effect of re-equilibration on the fluids trapped in the  
342 mafic phases from lava compositions of the three segments of the fissure zone is shown in the  
343 pressure scheme of Figure 10. The highest lithostatic pressure conditions are recorded by fluids  
344 in the olivines of the lava balloons from the SERfz (575 MPa). Differently, inclusions hosted in  
345 clinopyroxenes are partly re-equilibrated. In both cases this information is provided by Type-II  
346 inclusions, which indicates that the host minerals were already crystallized. The ascent of the  
347 carrier magma was continuous and relatively fast, as indicated by the lack of late stage trapping  
348 events and by the partial degree of re-equilibration of the inclusions.

349 The magmas erupted from the Cfz are more differentiated and the distribution of inclusions  
350 population is polymodal (Fig. 6). The densest inclusion in an olivine brings to  $754 \text{ kg}\cdot\text{m}^{-3}$  which  
351 corresponds to a pressure of 484 MPa. Further trapping of Type-I inclusions with a maximum  
352 density of  $692 \text{ kg}\cdot\text{m}^{-3}$  and  $497 \text{ kg}\cdot\text{m}^{-3}$  occurred respectively at 406 MPa and 209 MPa.  
353 Accordingly, the degree of re-equilibration of these inclusions is rather high, and many of them  
354 did not survived to the multiple ponding.

355 Finally, the magmas erupted from the SEfz show a similar evolutionary degree to the SERfz,  
356 but experienced a non-isochoric ascent, which caused partial re-equilibration in most of the  
357 samples. The densest Type-I inclusion ( $765 \text{ kg}\cdot\text{m}^{-3}$ ) was trapped at 498 MPa, which is quite  
358 similar to the value recorded by inclusions from the Cfz. Only a single olivine crystal had a  
359 second Type-I fluid trapping event ( $497 \text{ kg}\cdot\text{m}^{-3}$ ) at 219 MPa.

360

### 361 **Magma ponding depths**

362 The assumption that magmas ascend through the lithosphere, stop and accumulate at the  
363 Moho Transition Zone (i.e. Stolper and Walker 1980; Bureau et al 1999; Fodor and Galar 1997;  
364 Klügel et al. 2005, 2007; Stroncik et al 2009; Hansteen et al 1998; Geist et al 1998; Zanon and  
365 Frezzotti 2013) is essential for our calculation of the ponding depth, represented in Figure 11. A  
366 country rock density of  $2800 \text{ kg}\cdot\text{m}^{-3}$  was used for the crustal rocks beneath the Serreta submarine  
367 ridge, resulting from the direct measurement of our sample, and considered as representative for  
368 the crust. This value is in agreement with the measurements performed on lavas from fissure  
369 zones in Pico, Faial and São Miguel islands (Zanon and Frezzotti 2013; Zanon 2014). A value of  
370  $2500 \text{ kg}\cdot\text{m}^{-3}$  was used for the crust including the rocks of the island, according to literature  
371 (Montesinos et al. 2003; Self 1976) and to the average of three direct density measurements. The

372 deepest level of magma ponding is common to all fissure zone segments and is located between  
373 20.3 and 21 km deep, similarly to the depths reported for Pico and Faial fissure zones (Zanon  
374 and Frezzotti, 2013) and is interpreted to represent the Moho Transition Zone.

375 The magmas of the Cfz show a multi-step ascent with a first crustal ponding depth of 16.5  
376 km, common to all magmas, followed by at least one local ponding 8.5 km deep. At the first  
377 ponding site, magmas evolve towards trachybasalt and re-equilibrate the previously trapped fluid  
378 inclusions. A new generation of Mg-poor olivines, hosting Type-I fluid inclusions, records this  
379 process. The shallower local ponding level is associated to a final step of evolution, characterized  
380 by higher degree of fractional crystallization, as revealed by the geochemical characteristics of  
381 sample TRS07.

382 Finally, most magmas of the SEfz ascended without ponding into the crust. Only the magma  
383 of a single eruption ponded at 8.9 km deep, however, this stop must have been short, as the  
384 evolution degree of the magma is perfectly comparable with that of the other magmas of this  
385 segment.

386

387

#### 388 **STRESS FIELD OR/AND OCCURRENCE OF DENSITY FILTERS**

389 The petrological characteristics and the microthermometry of fluid inclusions of magmas  
390 erupted from the three fissure zone segments reveal slightly different ponding and ascent  
391 conditions through the crust. According to our interpretation, these differences were caused by  
392 two factors possibly related to one another: the local variation of the stress field and the presence  
393 of density barriers that hamper magma ascent.

394 The Moho Transition Zone is the most efficient barrier, which filters all primitive magmas,  
395 promoting fractional crystallization of mantle olivine and pyroxene in rising sub-lithospheric  
396 mafic magmas and underplating (e.g. Caress et al. 1995; Charvis et al. 2005; Klügel et al. 2005;  
397 Schwarz et al. 2004). At oceanic islands, such as the Azores, it generally corresponds to the  
398 transition from ultramafic cumulitic dunites, harzburgites and clinopyroxenites (density  $\cong 3200$   
399  $\text{kg}\cdot\text{m}^{-3}$ ) to partly serpentinized basaltic rocks and/or gabbros with density  $< 2800 \text{ kg}\cdot\text{m}^{-3}$  (e.g.  
400 Neumann et al. 1999; Clague and Bohrson 1991; Gaffney 2002).

401 The orientation of the SEfz is NW-SE, almost orthogonal to the WSW-ENE spreading  
402 direction of the Terceira Rift (Marques et al. 2013), i.e.  $\sigma_2$  and  $\sigma_3 \gg \sigma_1$ , enabling magma ascent.  
403 However, this segment is also influenced by older feeding system of the extinct Cinco Picos  
404 central volcano, which hampered the ascent of magma, causing the general re-equilibration of  
405 the inclusions.

406 The Cfz of the island, active in historical times, seems to be dominated by a stress field with a  
407 WNW-ESE direction. Due to its position and orientation, the Cfz represents an area of stress  
408 transfer from Santa Barbara central volcano, interested by a WNW-ESE tectonic system, to the  
409 central volcanoes of Pico Alto and Guilherme Moniz. The ascent of magmas is here both  
410 hampered by the direction of the spreading and the presence of intracrustal crystallized bodies of  
411 more evolved compositions from the central volcanoes. We interpreted that the latter acted as a  
412 density barrier and their presence beneath silicic volcanoes even influenced the local stress field,  
413 hampering the propagation of the fissure zone. The presence of density filters beneath Terceira  
414 has been previously discussed as a factor that may contribute to the marked bimodal volcanism  
415 of the island (e.g. Self and Gunn 1976).



416 Finally, the SERfz shows a WNW-ESE trend, but the crust here is thought to be free from the  
417 presence of pre-existing storage areas of evolved composition. For this reason magmas can  
418 ascend rapidly without significant obstacles.

419

420

421

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606

607

## 608 **Figure Captions**

609 **FIGURE 1.** a) Geographic sketch map of the Azores archipelago showing main structural features  
610 as dashed lines. MAR: Mid-Atlantic Ridge; TR: Terceira Rift; EAFZ: East Azores Fracture  
611 Zone; GF: Gloria Fault. Blue circle marks location of Terceira Island. b) Digital elevation model  
612 of Terceira showing axes of fissure zone segments, main volcanic and tectonic features and vents  
613 of historical eruptions (basaltic magmas). Stars locate sample sites and grey arrows indicate main  
614 paths of lava flows interpreted from aerophotographs. Cfz: Central fissure zone segment; SEfz:  
615 Southeastern fissure zone segment; SERfz: Serreta submarine fissure zone segment. UTM  
616 coordinates in km.

617

618 **FIGURE 2.** Temporal sequence of eruption of the lavas from different segments of the fissure  
619 zone. Sample numbers indicated in red. Coordinates correspond, whenever possible, to the  
620 source vent of the samples, measured as easting (UTM coordinates in km). Age correlations from  
621 Self (1976), updated with Calvert et al. (2006). The thick dashed line represents the Lajes-Angra  
622 Ignimbrite formation temporal marker (~21 ka). Fissure zone acronyms as in Fig. 1.

623

624 **FIGURE 3.** Photomicrographs of selected basaltic samples from Terceira. a) Poorly porphyritic  
625 lavas of the Cfz (TRS01) showing few microphenocrysts of plagioclase and rare mafic phases  
626 dispersed in a crystalline groundmass of feldspars, olivines, clinopyroxenes and oxides (not  
627 visible). b) Porphyritic basalt of the SEfz (TRS06) showing the occurrence of phenocrysts (0.8-  
628 0.5 mm) and microphenocrysts (~0.3 mm) of olivine and a basal section of a clinopyroxene. c)  
629 Coexistence of scattered Type-I with short trails of Type-II fluid inclusions and silicate melt  
630 inclusions (MIs).

631

632 **FIGURE 4.** a) Total Alkali Silica (TAS) classification diagram of the samples shows limited  
633 variability of composition. Porphyritic samples of the SEfz are basalt, while those crystal-poor of  
634 the Cfz are trachybasalt and basalt. b) Primordial-mantle-normalized patterns of incompatible  
635 elements (McDonough and Sun 1995) are all very similar, despite the degree of evolution of the  
636 samples. The lava balloons sample of the AD 1998-2001 eruption is among the least evolved,  
637 while the lavas of the Cfz are the most evolved. c) The chondrite-normalized patterns of Rare  
638 Earth Elements (Sun and McDonough 1989) confirm this observation. No negative Eu anomaly  
639 is observed in the studied samples.

640

641 **FIGURE 5.** a) Possible compositional evolution path by fractional crystallization of mafic phases  
642 and plagioclase. Each vector qualitatively indicates the direction of the compositional array in  
643 the case of single phase removal. b) Ratios among LILE (e.g. Rb/Ba) and U/Th are variable and  
644 beyond the analytical error, indicating that this chemical heterogeneity is inherited by the mantle  
645 source. A suite of rocks produced by the removal of mafic phases from the same parental liquid  
646 should preserve these ratios. d) Variations of the ratios among LREE/HREE in basaltic  
647 compositions are also too large to be explained by clinopyroxene and olivine fractionation.

648

649 **FIGURE 6.** a) Histograms of  $Th_L$  of fluid inclusions in olivines and clinopyroxenes. The data are  
650 polymodal and representative of the three fissure zone segments. Data population from the lava  
651 balloons sample is large, while that from the poorly porphyritic samples of the Cfz is small. The  
652 few data of  $Th_V$  are not reported here for convenience. b) Histograms of the total density values  
653 of the trapped fluid, divided in the same classes as for the temperatures of the homogenization.  
654 These histograms also include the few density data of the inclusions which homogenized into the  
655 vapor.

656

657 **FIGURE 7.** Relation between the size of the inclusions and their temperature of homogenization.  
658 Four small trails of inclusions in a single olivine show that the lowest  $Th_L$  characterizes the  
659 smallest inclusion in each trail. The increase in  $Th_L$  is directly related to the size, due to re-  
660 equilibration processes.

661

662 **FIGURE 8.** Diagram of the crystallization temperatures calculated from the application of Ca-in  
663 olivine geothermometer (Shejwalkar and Coogan 2013) to the composition of phenocrysts and  
664 microphenocrysts from three samples. The vertical bar represents the maximum error observed

665 for this algorithm ( $\pm 20$  °C). The data also include compositions of olivine phenocrysts from the  
666 lava balloons of the AD 1998-2001 submarine eruption (Kueppers et al. 2012).

667

668 **FIGURE 9.** Isochores for selected samples of fissure zone segments. The lines related to pure CO<sub>2</sub>  
669 fluid derived from the application of the equation of Span and Wagner (1996) and are  
670 represented by grey dashed lines. Those for H<sub>2</sub>O+CO<sub>2</sub> (molar ratio 1:9) fluids, represented by  
671 light-blue dashed lines, are obtained from the equation of Sterner and Bodnar (1991). The yellow  
672 vertical line is the intercepts of isochores with calculated temperature of trapping of the fluids.

673

674 **FIGURE 10.** Pressures scheme of trapping events and re-equilibration for the studied samples.  
675 Green and red colors discriminate information from inclusions trapped in clinopyroxenes and  
676 olivines, respectively. Arrows indicate re-equilibrated populations of inclusions. The black  
677 dashed line marks the limit between the pure fissure zone domain of the Serreta submarine ridge  
678 from the subaerial fissure zone segments passing through pre-existing storage systems.

679

680 **FIGURE 11.** Scheme of the distribution of magma storage systems of fissure zone segments  
681 across an ideal NW-SE profile of Terceira Island. The areas of fluid trapping are indicated by the  
682 pattern and by the thin blue dashed lines. The two thick blue dashed lines represent the possible  
683 location of the Moho Transition Zone. For further explications, please refer to the text.

684

685

686 **Table list**

687 **TABLE 1.** Brief description of the samples, representative of the different segments of the fissure  
688 zone. Unit names and ages from Self (1976) updated with Calvert et al. (2006). Phenocrysts and  
689 microphenocrysts are counted together in modal data.

690

691 **TABLE 2.** Summary of fluid inclusions data. Data report the range measured. Density values are  
692 re-calculated in consideration of a composition of the fluid of H<sub>2</sub>O+CO<sub>2</sub> with a molar ratio of  
693 1:9.

694

695

#### 696 **Supplementary material**

697 Supplementary datatable1 - Chemical database of whole-rock compositions.

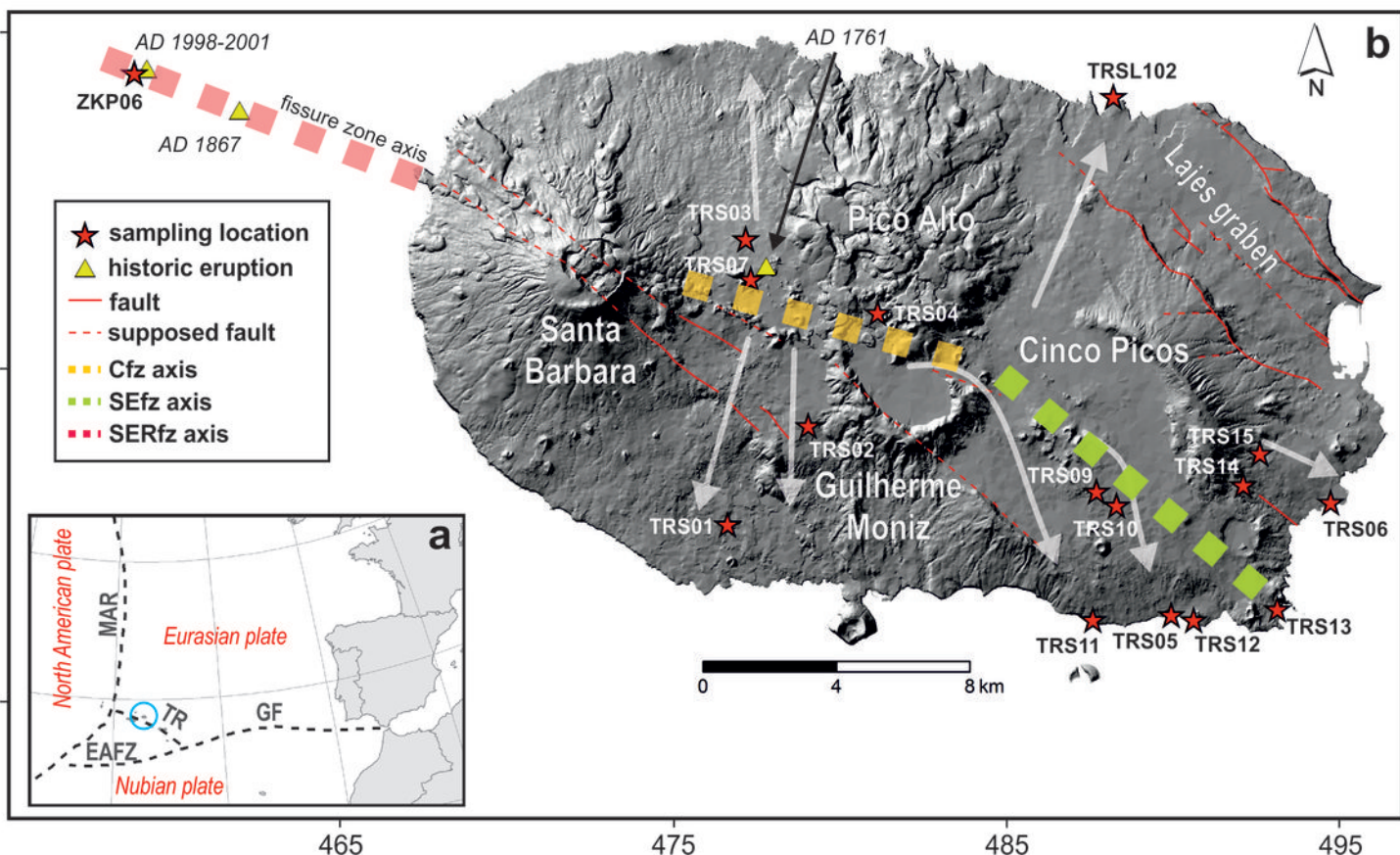
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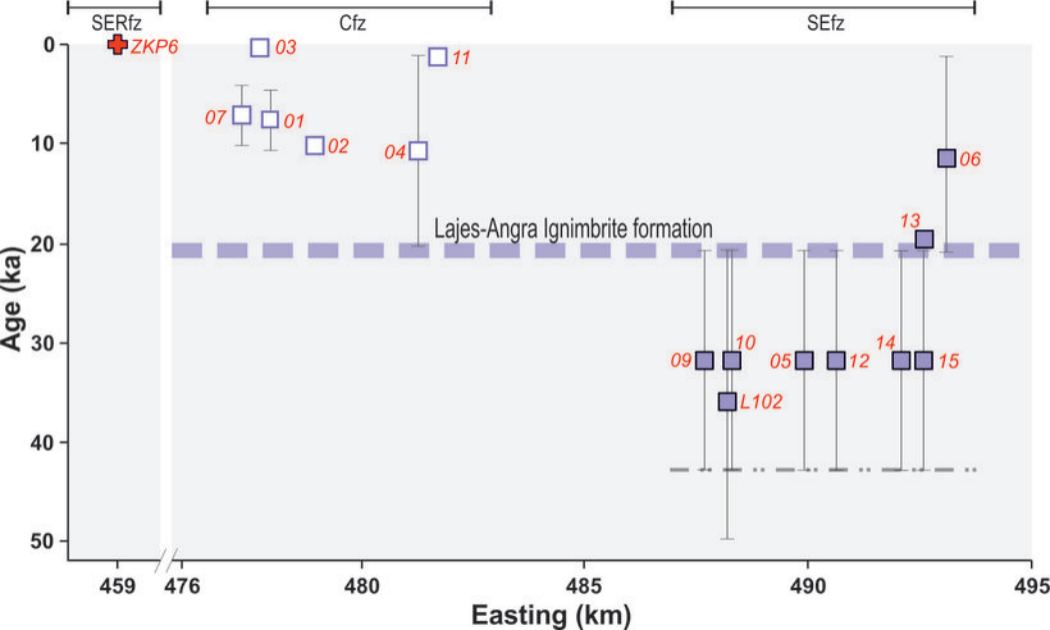
699 Supplementary datatable2 - Microprobe analyses of the mineral assemblage of selected samples  
700 and the groundmass glass. Mineral chemistry comes from the core of the crystals, using a spot  
701 beam. Groundmass glasses were analyzed with a defocalized beam. Most of the data are from  
702 single spot, except the trimmed mean (15% of extreme data) of 14 analyses. In this latter case,  
703 the standard deviation is reported in parentheses.

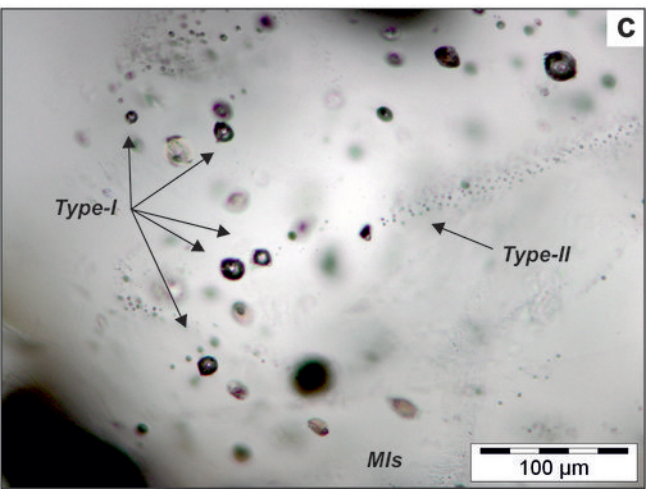
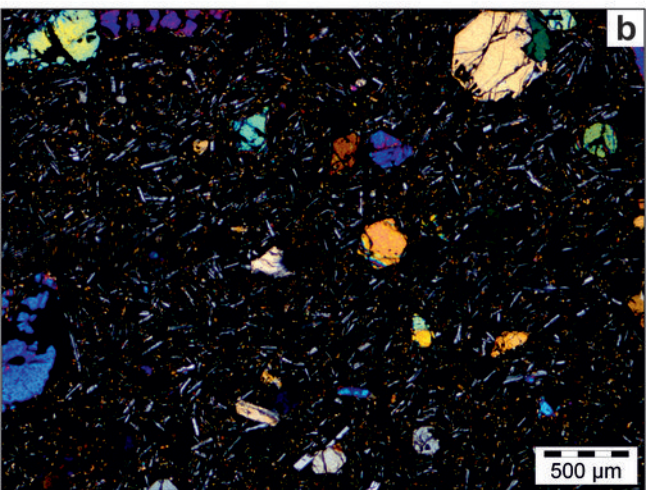
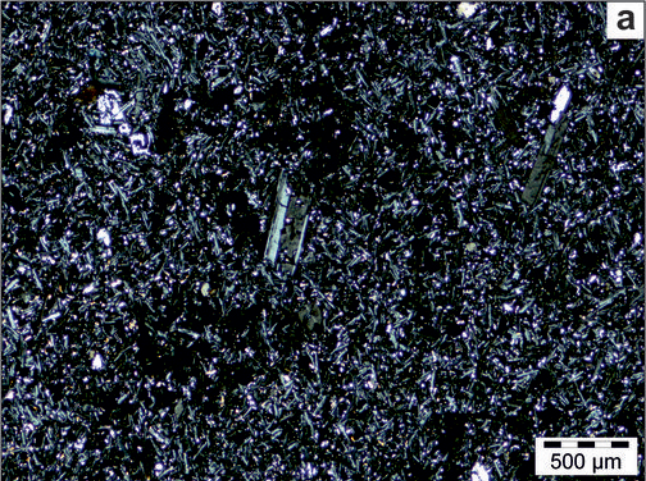
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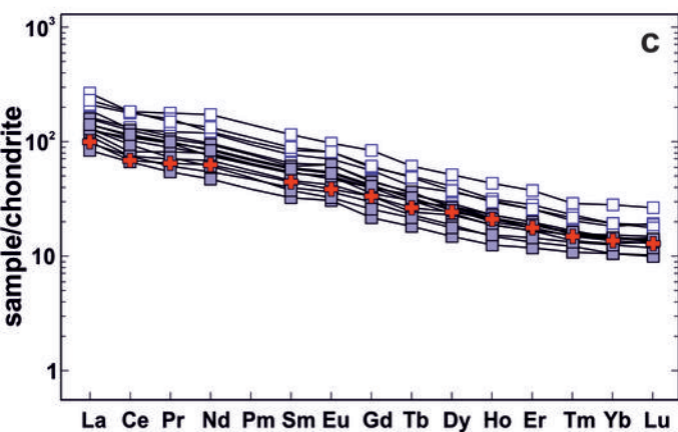
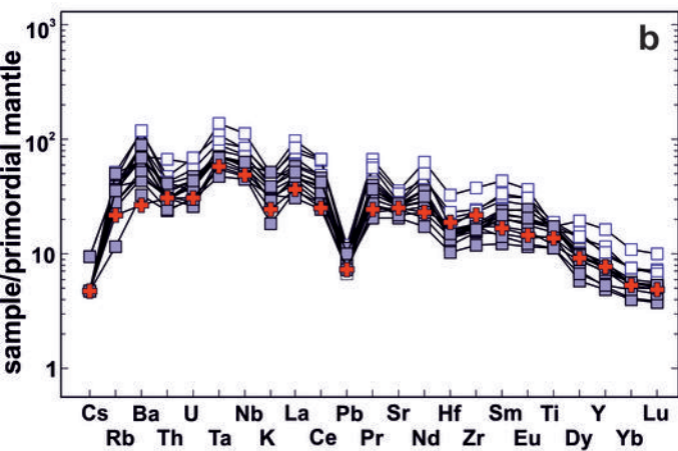
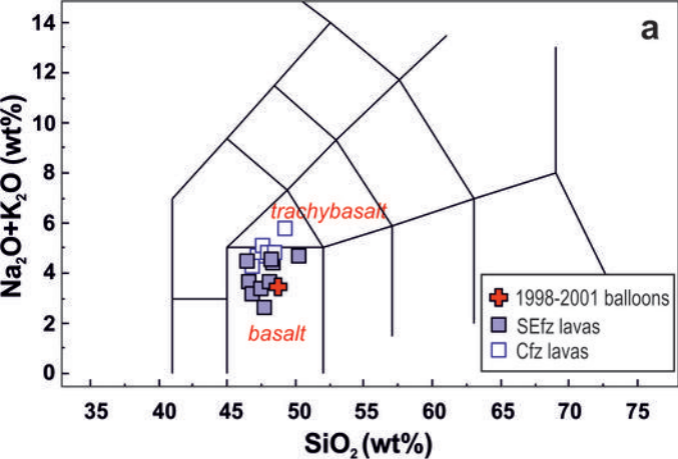
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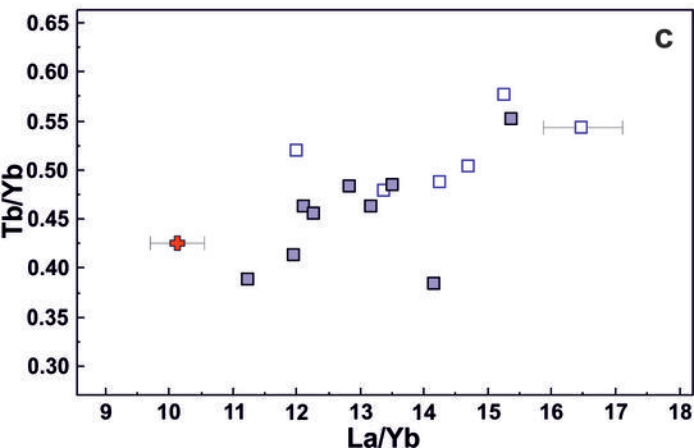
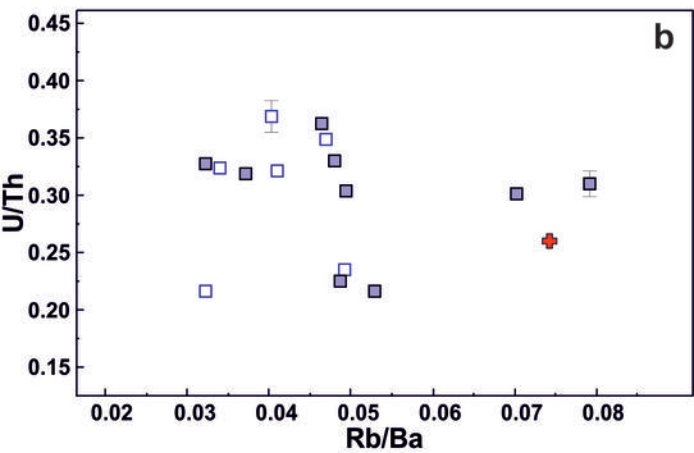
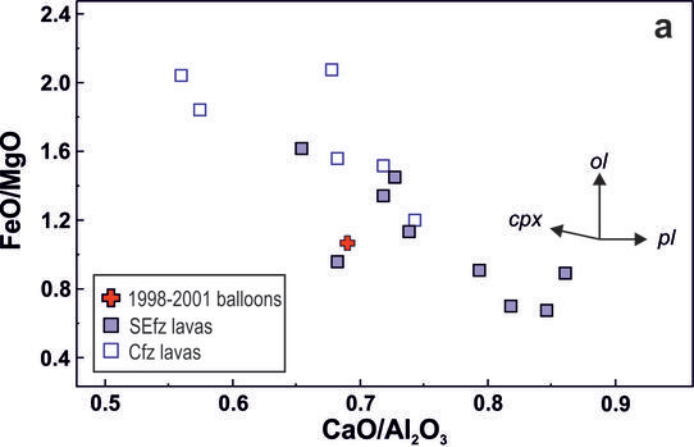


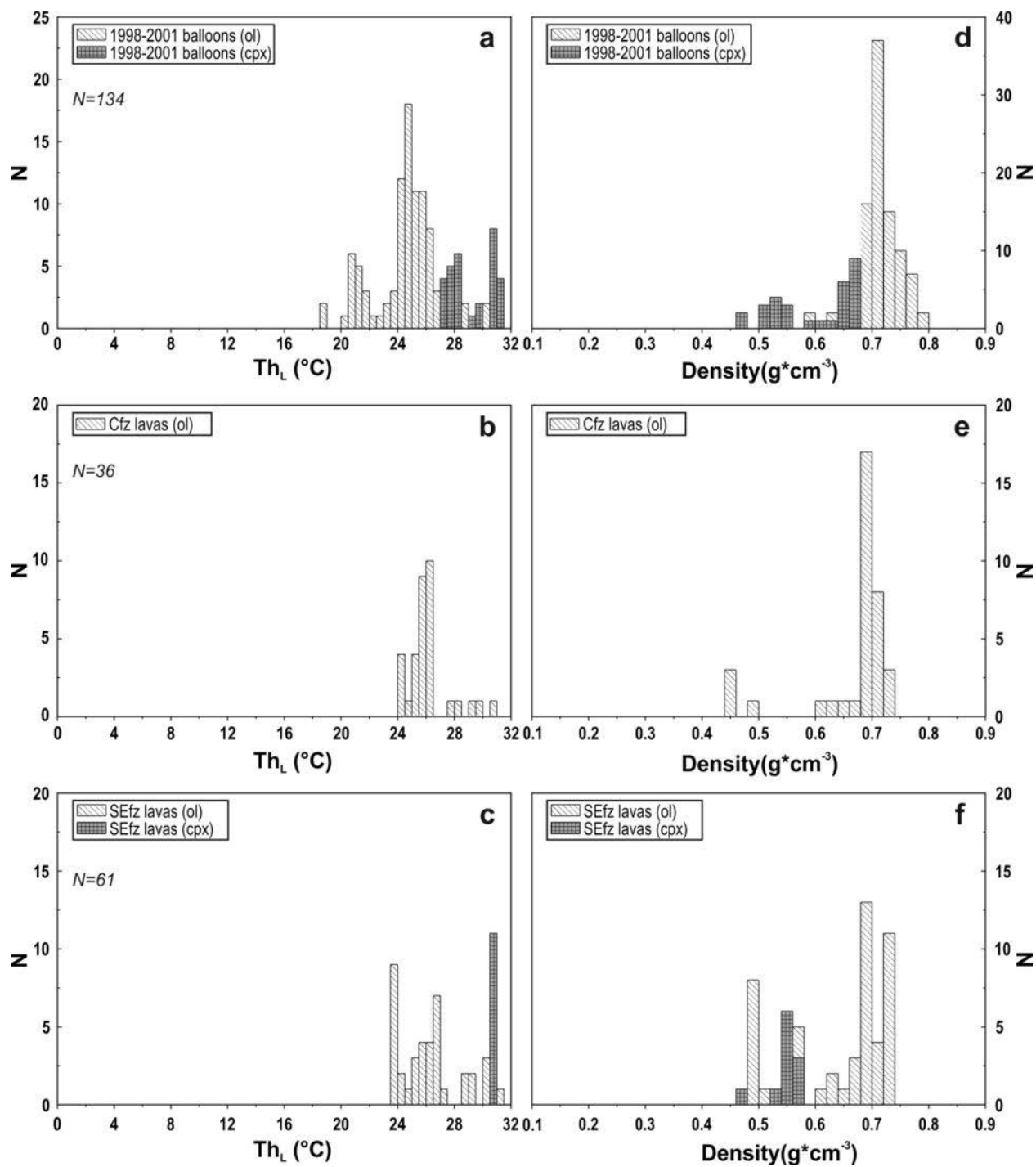


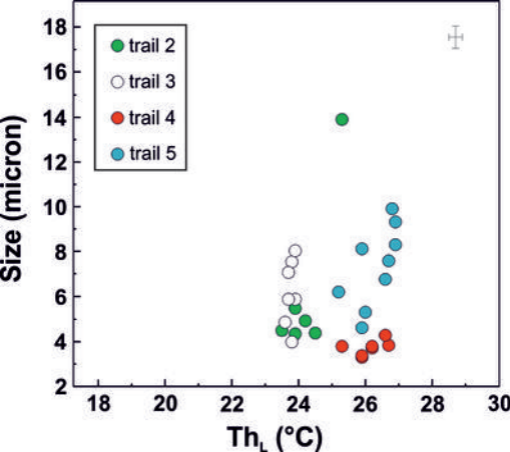


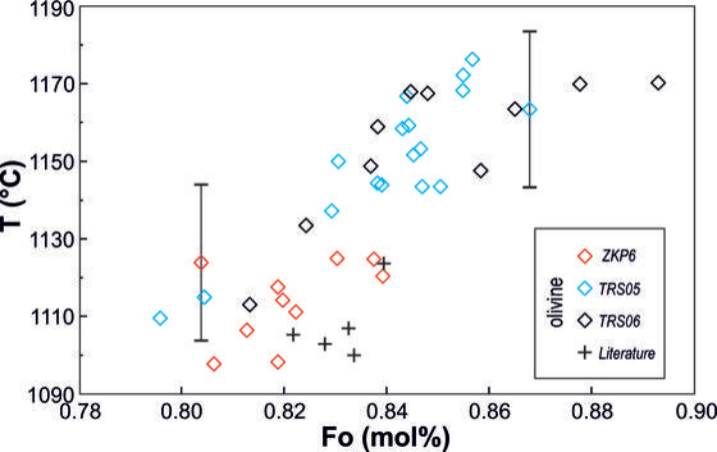


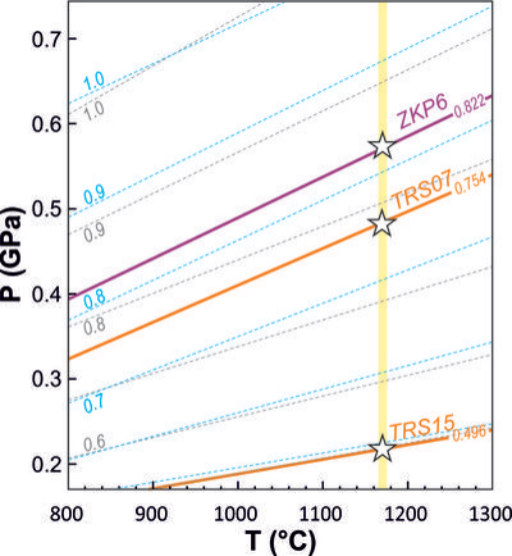


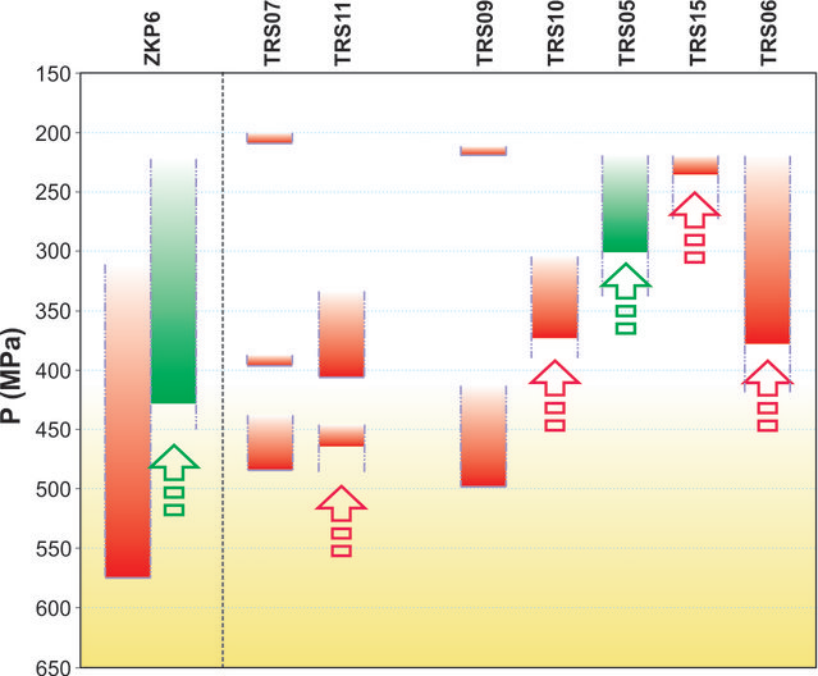


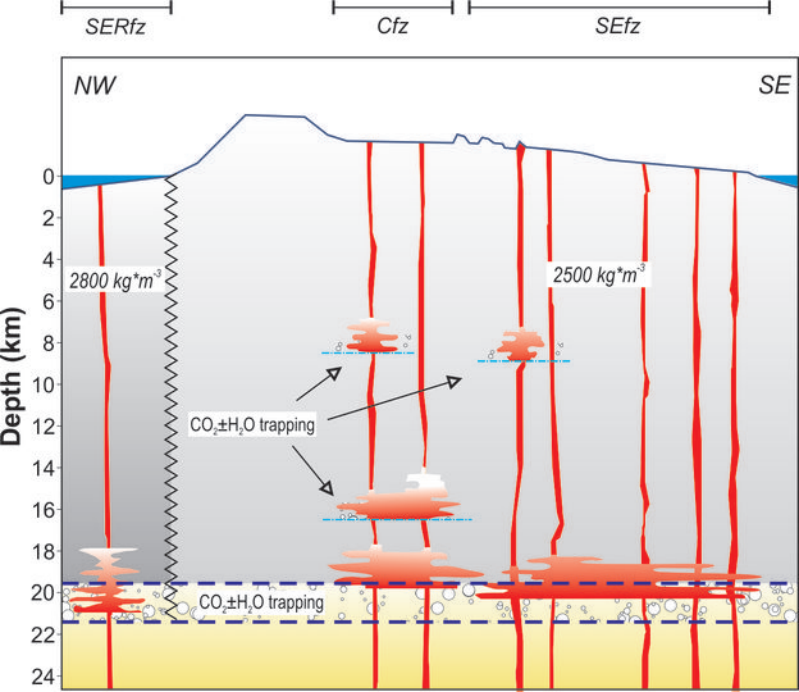














**TABLE 1.** Brief description of the samples, representative of different segments of the fissure zone. Unit names and ages from Self (1976) updated with Calvert et al. (2006).

Sample	Easting Northing	Unit/Location	Volcanic System	Age (years BP)	Texture	Modal analysis	Occurrence of fluid inclusions
TRS01	476611 E 4280357 N	Bagacina flow	Cfz	<10,090±50 and >4480±40	intergranular to intersertal	pl <sub>16</sub> +ol <sub>5</sub> +cpx <sub>2</sub> +ox <sub>6</sub> +gms <sub>71</sub>	not found
TRS02	479036 E 4283318 N	Vareiras flow	Cfz	10,090±50	intersertal	pl <sub>18</sub> +ol <sub>10</sub> +cpx <sub>5</sub> +ox <sub>10</sub> +gms <sub>57</sub>	not found
TRS03	477159 E 4288907 N	historic flow	Cfz	AD 1761	intersertal with glomeroporphyritic aggregates	pl <sub>26</sub> +ol <sub>3</sub> +cpx <sub>3</sub> +ox <sub>14</sub> +gms <sub>54</sub>	not found
TRS04	481117 E 4286700 N	unnamed cone	Cfz	>1910±35	intersertal	ol <sub>7</sub> +pl <sub>5</sub> +cpx <sub>2</sub> +ox <sub>7</sub> +gms <sub>79</sub>	not found
TRS05	489915 E 4277647 N	Porto Judeu	SEfz	>21,000	-	-	cpx
TRS06	494696 E 4281018 N	Porto S. Fernando	SEfz	<21,000	intersertal	pl <sub>18</sub> +ol <sub>16</sub> +cpx <sub>6</sub> +ox <sub>6</sub> +gms <sub>54</sub>	ol
TRS07	477312 E 4287701 N	Pico 599	Cfz	>4480±40 and <10,090±50	-	-	ol
TRS09	487671 E 4281350 N	base of Cinco Picos caldera	SEfz	>21,000	intersertal with glomeroporphyritic aggregates	ol <sub>17</sub> +pl <sub>15</sub> +cpx <sub>6</sub> +ox <sub>3</sub> +gms <sub>59</sub>	ol
TRS10	488280 E 4280926 N	base of Cinco Picos caldera	SEfz	>21,000	intersertal to seriate	pl <sub>17</sub> +ol <sub>7</sub> +cpx <sub>1</sub> +ox <sub>10</sub> +gms <sub>65</sub>	ol
TRS11	487566 E 4277502 N	Algar do Carvão flow	Cfz	1910±35	intergranular to intersertal	ol <sub>7</sub> +pl <sub>2</sub> +cpx <sub>18</sub> +ox <sub>11</sub> +gms <sub>63</sub>	ol
TRS12	490587 E 4277495 N	Pico do Refugio	SEfz	>21,000	intersertal	ol <sub>16</sub> +pl <sub>3</sub> +cpx <sub>1</sub> +ox <sub>4</sub> +gms <sub>76</sub>	not found
TRS13	493093 E 4277817 N	near Pico dos Comos	SEfz	19,120±50?	intersertal	pl <sub>18</sub> +ol <sub>7</sub> +cpx <sub>3</sub> +ox <sub>7</sub> +gms <sub>65</sub>	not found
TRS14	492075 E 4281540 N	Fonte do Bastardo	SEfz	>21,000	intersertal	ol <sub>4</sub> +cpx <sub>3</sub> +pl <sub>1</sub> +ox <sub>3</sub> +gms <sub>89</sub>	not found
TRS15	492578 E 4282484 N	Fonte do Bastardo	SEfz	>21,000	intersertal	pl <sub>14</sub> +ol <sub>6</sub> +cpx <sub>4</sub> +ox <sub>4</sub> +gms <sub>76</sub>	ol
TRSL102	488179 E 4293179 N	Ponta das Escaleiras	SEfz	>21,000 and <50,000±10,000	intersertal	pl <sub>19</sub> +ol <sub>11</sub> +cpx <sub>2</sub> +ox <sub>9</sub> +gms <sub>59</sub>	not found
ZKP06	458828 E 4293879 N	~10 km offshore Terceira	SERfz	AD 1998-2001	vitrophyric with cryptocrystals to intersertal	pl <sub>24</sub> +ol <sub>12</sub> +cpx+gms <sub>64</sub>	ol+cpx

Mineral modal counting (%) is indicated in subscripts. Acronyms: ol – olivine, pl – plagioclase, cpx – clinopyroxene, ox – oxides (unspecified), gms – groundmass; Cfz - Central fissure zone segment; SEfz - Southeastern fissure zone segment; SERfz - Serreta submarine fissure zone segment.

**TABLE 2:** Fluid inclusions data

<b>Sample</b>	<b>N° measures</b>	<b>Host</b>	<b>Th<sub>L</sub> (°C)</b>	<b>Th<sub>V</sub> (°C)</b>	<b>Density (kg·m<sup>-3</sup>)</b>	<b>P (MPa)</b>	<b>Depth (km)</b>
TRS05	11	cpx	30.5-31		496-590	222-301	9.1-12.3
TRS06	11	ol	28.6-31		497-669	219-378	8.9-15.4
TRS07	19	ol	24.3-28.0	31	482-754	209-484	8.5-19.8
TRS09	36	ol	23.5-31		497-765	219-498	8.9-20.3
TRS10	2	ol	28.8-30.4		598-664	303-373	12.4-15.2
TRS11	17	ol	25.2-29.8		629-739	334-464	13.6-18.9
TRS15	1	ol	31		497	219	8.9
ZKP06	30	cpx	27.2-31.0		487-703	222-428	8.1-15.6
ZKP06	104	ol	18.8-30.2		610-822	315-575	11.5-21.0