

## 1 **Comparing Clays From Mars and Earth**

2 Highlights Article for American Mineralogist

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7 **Abstract:** A terrestrial analogue of the saponite identified by Mars Science Laboratory in the  
8 mudstone of Yellowknife Bay, Gale Crater has been identified. It is griffithite, from Miocene volcanic  
9 and sedimentary rocks in Griffith Park, Los Angeles. This is a ferrian, trioctahedral saponite with all  
10 the Fe<sup>3+</sup> in a distorted octahedral site.

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13 Since August 2012 the Mars Science Laboratory rover Curiosity has been studying ancient, clay-  
14 bearing sedimentary rocks of basaltic composition in Gale Crater. During sols (the Mars term for a  
15 day) 57-100, 180-292, Curiosity analysed soil and drilled fine grained sediment at the Rocknest sand  
16 shadow, then the John\_Klein and Cumberland outcrops of the 1.5 m thick Sheepbed unit.

17 The presence of CheMin – a novel transmission X-ray Diffraction instrument with a Co K $\alpha$  source  
18 [Blake et al. 2012] – in the body of the rover has allowed mineralogists to determine what the fine  
19 grained sediments of Yellowknife Bay are made of. The results of that work showed the presence of  
20 a saponite and magnetite-bearing assemblage in the mudstone and also amorphous material in both  
21 the scooped aeolian and drilled samples. In a remarkable step forward in our understanding of  
22 Mars, the clay-bearing assemblage was interpreted as diagenetic rather than detrital and forming in  
23 a circum-neutral, habitable, fluvio-lacustrine environment [McLennan et al. 2014; Grotzinger et al.  
24 2014].

25 However, as Treiman et al. [2014] point out in their paper in this issue, even Curiosity has its  
26 limitations, the 2 $\theta$  range of the CheMin detector 5-50 $^{\circ}$ , doesn't allow for the identification of all  
27 features normally used to help distinguish between di and trioctahedral structures. For instance, an  
28 060 peak, or 06l band which is present when there is rotational disorder between clay lattice layers,  
29 occurs at 61 $^{\circ}$  in saponite (Moore and Reynolds, 1997). Accurate estimations of the proportion of  
30 ferric iron and associated redox conditions are also challenging.

31 Studying Mars' mineralogy *in situ* sharpens the need for suitable terrestrial analogues that can  
32 provide detailed structural and compositional information to feed back into the martian studies. To  
33 this purpose, Treiman et al. have gathered such data on a saponite from Griffith Park, Los Angeles,  
34 found in basaltic amygdaloids of Miocene age and adjacent volcanoclastic sediments. This 'griffithite'  
35 [Larsen and Steiger, 1917] is a Mg-Fe<sup>3+</sup>-Fe<sup>2+</sup> trioctahedral smectite with all of the ferric iron (Fe<sup>3+</sup>/ $\Sigma$ Fe  
36 = 64 to 93%) considered on the basis of Mössbauer spectra, to be present in a distorted octahedral  
37 site. The griffithite has an 02l band at 4.59 Å, indicating some disorder resulting from rotations

38 between layers in the clay lattice, as does the saponite identified in the John\_Klein and Cumberland  
39 samples of the Sheepbed mudstone [Vaniman et al. 2014]. Vaniman et al. have already put this  
40 work to good use by showing that in one model of the Sheepbed saponite its composition is similar  
41 to the Griffith Park clay. By analogy between the two clays, this might mean that the Sheepbed  
42 saponite experienced oxidation after formation, in a late diagenetic event, to create its current  
43 ferrian composition.

44 Supporting evidence that such a saponite could be widespread in areas that have preserved  
45 alteration of the martian basaltic composition crust, comes from the Lafayette member of the  
46 nakhlite martian meteorite group. This meteorite contains hydrothermal veins including saponite of  
47 similar composition to the griffithite, though with greater iron content and some tetrahedral Fe<sup>3+</sup>  
48 site occupancy [Changela and Bridges, 2011; Hicks et al. 2014].

49 Treiman et al. also note that their infrared spectroscopy of the griffithite may help refine the  
50 identification of clay minerals by the OMEGA and CRISM spectrometers on the Mars Express and  
51 Mars Reconnaissance Orbiter spacecraft. The identification of clay from orbit was a key reason for  
52 the selection of Gale Crater as the Curiosity landing site [Milliken et al. 2010], and clays are high  
53 priority targets for the upcoming ExoMars 2018 and Mars2020 rover missions. Analogue studies like  
54 the one presented by Treiman et al. will be essential in the ongoing identification and  
55 characterisation of habitable environments on Mars.

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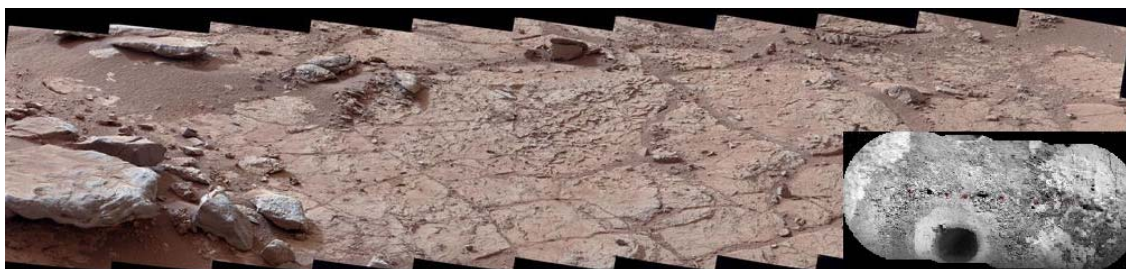
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107 Figure 1. The Sheepbed fluvio-lacustrine mudstone unit of Yellowknife Bay, Gale Crater showing  
108 characteristic flat bedding and raised ridges (Field of View 3.5 m, MastCam mosaic Sol 167,  
109 NASA/JPL-Caltech/MSSS). Inset shows a ChemCam Remote Micro-Imager view (Width of drillhole  
110 1.6 cm, Sol 183, NASA/JPL-Caltech/LANL/CNES/IRAP/LPGNantes/CNRS/IAS) of the John Klein drill  
111 hole, with laser spots visible, that was made within Sheepbed. The material from this drillhole, and  
112 the one in Cumberland, which was also in the Sheepbed unit, were analysed by the CheMin XRD  
113 instrument and found to contain a saponite which Treiman et al. 2014 argue is similar to terrestrial  
114 griffithite from Griffith Park, Los Angeles.