

1 **Revision 1**  
2 **Slawsonite–celsian–hyalophane assemblage from a picrite sill (Prague Basin, Czech**  
3 **Republic)**

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10 **ABSTRACT**

11 The first European occurrence of slawsonite is reported from a picrite sill within  
12 Upper Ordovician strata of the Prague Basin near the village of Rovina, Czech Republic.  
13 The rare slawsonite forms an interstitial phase in association with abundant celsian and  
14 hyalophane, replacing the original calcic plagioclase (bytownite). A study of this curious  
15 natural slawsonite–celsian–hyalophane assemblage provides a valuable insight into  
16 feldspar stability and petrogenesis.

17 Whole-rock geochemical signatures of the picrite sill and underlying doleritic  
18 basalt intrusion show conspicuous enrichment in Sr and Ba superimposed on normal  
19 basaltic multielement patterns. These two elements were most likely introduced by  
20 intergranular fluids during diffusional seafloor metasomatism (rodingitization and  
21 serpentization) of the picrite. Strontian and barian feldspars precipitated directly from  
22 BaO–SrO–H<sub>2</sub>O bearing fluid, which caused decomposition of plagioclase to vuagnatite,  
23 aqueous SiO<sub>2</sub> and Al<sub>2</sub>O<sub>3</sub> at  $T \leq 350$  °C. Subsequently, vuagnatite decomposed to

24 hydrogrossular and excess  $\text{SiO}_2$  was consumed by serpentinization of olivine. At the  
25 expense of aqueous  $\text{Al}_2\text{O}_3$ , serpentine reacted to chlorite closing the picrite alteration at  
26 320–160 °C. Pressure did not exceed 0.5 GPa. The *in situ* EDS analyses indicate that the  
27 chemical composition of the slawsonite is  $\text{Sl}_{91}\text{Cn}_3\text{An}_3\text{Ab}_3$  (core) to  $\text{Sl}_{82}\text{Cn}_3\text{An}_4\text{Ab}_9\text{Or}_2$   
28 (rim), the celsians range from  $\text{Cn}_{96.9}\text{An}_{0.3}\text{Ab}_{0.2}\text{Or}_2\text{Sl}_{0.6}$  to  $\text{Cn}_{76.3}\text{An}_{4.7}\text{Ab}_3\text{Or}_{15.7}\text{Sl}_{0.3}$ , and  
29 the hyalophanes vary from  $\text{Cn}_{72.2}\text{An}_{1.4}\text{Ab}_{5.1}\text{Or}_{21.1}\text{Sl}_{0.2}$  to  $\text{Cn}_{57.3}\text{An}_{0.8}\text{Ab}_{3.1}\text{Or}_{38.5}\text{Sl}_{0.3}$ .

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32 **Keywords:** slawsonite, celsian, hyalophane, strontium and barium feldspars, vuagnatite,  
33 hydrogrossular, rodingitization, serpentinization, Prague Basin

34

35 Running title: Slawsonite–celsian–hyalophane assemblage from a picrite sill

36

## INTRODUCTION

37 Feldspar can be assigned as a barian once its chemical analysis contains more than  
38 2 wt% of BaO (Deer et al. 1992). Celsian ( $\text{BaAl}_2\text{Si}_2\text{O}_8$ ; Cn) is an end-member of an  
39 extensive solid solution with K-feldspar ( $\text{KAlSi}_3\text{O}_8$ ), whereby the term hyalophane (Hy)  
40 is used for intermediate compositions (10–80 mol% of Cn: Deer et al. 2001 or 15–75  
41 mol% of Cn: Essene et al. 2005), with the remainder being dominated by Or prevailing  
42 over Ab. Furthermore, hyalophanes with > 50 mol% Cn are referred to as barian, and  
43 those with < 25 mol% Cn as potassian. Natural celsians and hyalophanes usually have a  
44 monoclinic symmetry. However, even triclinic polymorphs of hyalophane (with less than  
45  $\text{Cn}_{30}$ ) were found in nature (Gay and Roy 1968). Celsian structure resembles monoclinic  
46 body-centered Si–Al unit cells of anorthite, in which an addition of relatively small  $\text{Ca}^{2+}$

47 leads to distortion of framework and origin of triclinic symmetry. Nevertheless, celsians  
48 do not show strict regularity in alternation of Si and Al atoms like anorthite but rather  
49 significant Al and Si disorder (Griffen and Ribbe 1976). An acceptance of large radius  
50 ion  $\text{Ba}^{2+}$  into celsian framework keep the monoclinic symmetry and explain why  
51 hyalophanes with  $< \text{Cn}_{30}$ , e.g., with prevalence of  $\text{Ca}^{2+}$ , have triclinic symmetry.  
52 Experimental works confirmed the solid solution between celsian and albite ( $\text{NaAlSi}_3\text{O}_8$ )  
53 but with a miscibility gap (Viswanathan and Harneit 1989). These results were applied to  
54 the ternary system  $\text{BaAl}_2\text{Si}_2\text{O}_8$ – $\text{NaAlSi}_3\text{O}_8$ – $\text{KAlSi}_3\text{O}_8$  (Lagache and Catel 1992;  
55 Viswanathan 1992). According to the most recent summary (Essene et al. 2005), 8 four-  
56 feldspar and 24 three-feldspar assemblages may be stable in the system  $\text{BaAl}_2\text{Si}_2\text{O}_8$ –  
57  $\text{CaAl}_2\text{Si}_2\text{O}_8$ – $\text{NaAlSi}_3\text{O}_8$ – $\text{KAlSi}_3\text{O}_8$ .

58         While the occurrence of Ba-bearing feldspars is relatively common, natural  
59 slawsonite ( $\text{SrAl}_2\text{Si}_2\text{O}_8$ ; Sl) is extremely rare. It is isostructural with paracelsian  
60 ( $\text{BaAl}_2\text{Si}_2\text{O}_8$ ) (Griffen et al. 1977; Matsubara 1985). Paracelsian is a metastable phase of  
61 the slawsonite–paracelsian series and reacts to celsian at 550 °C; however, its stability at  
62 a lower  $T$  remains a possibility (Lin and Foster 1968). Phase transition from the low- $T$   
63 triclinic to the moderate- $T$  monoclinic structure of natural slawsonite at ambient pressure  
64 has been observed at 320 °C (Tagai et al. 1995). Monoclinic slawsonite is stable up to  
65 500 °C (McCauley 2000) or 600 °C (Bambauer and Nager 1981), when it transforms into  
66 another monoclinic polymorph, Sr-celsian (melting  $T = 1650$  °C), although the reaction  
67 has not yet been reversed. The Sr-celsian is isostructural with celsian but has no natural  
68 counterpart (McCauley 2000). Partial replacement of Sr by Ca has been reported in  
69 slawsonite from metamorphosed limestones in Wallowa County, Oregon (Griffen et al.

70 1977; Matsubara 1985) and from slawsonite veinlets cutting pectolite veinlets in  
71 metamorphosed xenoliths (glaucophane to pumpellyite–actinolite schist of the Ino Fm.)  
72 in Rendai, Japan (Nakajima et al. 1978; Matsubara 1985). Barium substitution was  
73 detected in slawsonite veinlets from xenolith in an ultramafic rock exposed in weakly  
74 metamorphosed rocks of the Ino Fm. in Sarusaka, Japan (Matsubara 1985; Tagai et al.  
75 1995). Xenolith consisted of diopside, chlorite and altered sphene but was devoid of  
76 igneous texture and was not considered as primary magmatic product (Matsubara 1985).

77 In this paper, we report the mode of occurrence and chemical composition of a  
78 rare assemblage of strontian and barian feldspars from picritic sill intruding to sediments  
79 in Rovina (Prague Basin, Czech Republic). We discuss the temperatures of their  
80 formation and suggest the likely genesis. Observations and evaluation of this curious  
81 slawsonite–celsian–hyalophane assemblage provide valuable information on feldspars  
82 phase equilibria, particularly in cases when one coexisting phase – slawsonite – has been  
83 noticed only rarely in nature.

#### 84 **FIELD RELATIONS AND PETROLOGY OF THE HOST PICRITE**

85 The studied picrite sill, together with the underlying doleritic basalt intrusion, was  
86 emplaced within Ordovician (Hirnantian) quartzitic siltstones of the Kosov Fm. (N49°  
87 55.412', E14° 13.320'). The bottom section of the outcrop is partially covered by  
88 Quaternary sediments, but the lower contact of the Kosov Fm. with Silurian (Llandovery)  
89 black calcareous shales of the Motol Fm. is clearly tectonic. It is represented by the so-  
90 called Koda Fault – a synsedimentary structure that was reactivated during Variscan  
91 orogeny to thrust Ordovician over the Silurian strata (Kříž 1991). The present-day picrite

92 outcrop is preserved in the form of several 1 m thick blocks exposed within ca. 15 m  
93 wide zone on a vegetated slope.

94 The picrite shows a porphyric texture with idiomorphic olivine and clinopyroxene  
95 phenocrysts. Modal composition determined on the basis of point counting by Fiala  
96 (1947), and supplemented by our mineral analyses, is as follows: rare fresh olivine  
97 phenocrysts (15.2 vol% – Fo<sub>74-78</sub>) are intensely replaced by serpentine (28.7 vol%) with  
98 chlorite (25.6 vol% – Mg-rich chlorite) and usually bear spinel inclusions. Clinopyroxene  
99 phenocrysts (6.8 vol% – diopside, hedenbergite) carry tiny spinel (ulvöspinel, magnetite)  
100 and ilmenite inclusions; the phenocrysts form poikilitic intergrowths with plagioclase (9.8  
101 vol% – An<sub>70-84</sub>) and olivine. Ilmenite and spinel account for 1.9 vol% of the rock.  
102 Biotites (0.7 vol%) overgrow ilmenite and form transition zones to the interstitial matrix  
103 (11.3 vol%), which is dominated by chlorite. Partially decomposed plagioclases, chlorite  
104 pseudomorphs after calcic amphibole, apatite, ilmenite and spinel constitute the matrix.  
105 Recently, secondary mineral assemblage consisting of slawsonite, celsian, hyalophane,  
106 vuagnatite, hydrogrossular and pectolite has been identified in the interstitial phase by  
107 back-scattered electron (BSE) imaging.

## 108 **ANALYTICAL METHODS**

109 Representative picrite (20 kg) and basalt (40 kg) samples were collected in the  
110 field, crushed, homogenized and ground in an agate mill. Major-element analyses were  
111 carried out at the Central Laboratories of the Czech Geological Survey, Prague (wet  
112 silicate analyses: FAAS, ICP-OES, PMT, IR spectrometry). The trace-element data came  
113 from Activation Laboratories, Vancouver (ICP-MS package 4B2 Research, sample  
114 dissolution by a lithium metaborate/tetraborate fusion followed by rapid digestion in a

115 weak nitric acid solution). Whole-rock data were processed by the GCDkit package  
116 (Janoušek et al. 2006). Analyses and compositional maps of the picrite mineral  
117 assemblage were obtained using the scanning electron microscope TESCAN Vega  
118 equipped with EDS detector X-Max 50 (Oxford Instruments) at the Institute of Petrology  
119 and Structural Geology (Faculty of Science, Charles University in Prague). The raw data  
120 were processed with INCA software, employing the XPP matrix correction procedure  
121 (Pouchou and Pichoir 1988). Standards used with X-ray lines in parentheses were albite  
122 (Na  $K\alpha$ ), sanidine (K, Si  $K\alpha$ ), synthetic garnet (Al  $K\alpha$ ), wollastonite (Ca  $K\alpha$ ), celestine  
123 (Sr  $L\alpha$ ), barite (Ba  $L\alpha$ ), olivine (Mg  $K\alpha$ ), almandine (Fe  $K\alpha$ ), rhodonite (Mn  $K\alpha$ ),  
124 pentlandite (Ni  $K\alpha$ ), rutile (Ti  $K\alpha$ ), Cr<sub>2</sub>O<sub>3</sub> (Cr  $K\alpha$ ) and vanadium (V  $K\alpha$ ); the detection  
125 limits were typically 0.1 wt% of the respective element. All analyses were obtained at 15  
126 kV acceleration voltage and 1.5 nA beam current with a rastered 3×3 μm beam. Time  
127 tests showed no elemental mobilization under these conditions. The compositional  
128 variations of samples are represented by maps of X-ray intensity at defined energy  
129 intervals. The effect of interference of Sr and Si peaks was suppressed by manual  
130 adjustment of the considered energy range. However, only the picrite feldspars were  
131 suitable for EDS analyses, because the barian feldspars in the basalt intrusion were  
132 smaller than the beam current resolution.

133 In total, twelve analyses of a single slawsonite grain, representative analyses of  
134 one hundred and twenty-seven celsian and of nine hyalophane grains, and two analyses of  
135 discontinuously zoned barian feldspar obtained from seven thin sections of the picrite are  
136 the subject of this paper. Representative analyses of hydrogrossular, vuagnatite, pectolite,  
137 plagioclase, olivine, clinopyroxene, spinel, amphibole, chlorite, serpentine, talc and

138 ilmenite are provided in the Appendix. In case of hydrogrossular,  $\text{OH} = (3-\text{Si}) \times 4$   
139 substituting silica was determined from stoichiometry for  $8(\text{cations} + \text{OH}/4)$ . Results of  
140 chlorite thermometry calculations (Kranidiotis and MacLean 1987) are also listed in the  
141 Appendix.

## 142 **RESULTS**

### 143 **Whole-rock geochemistry of the host basic volcanics**

144 On the basis of the total alkalis vs. silica diagram (TAS; Le Bas et al. 1986),  
145 Rovina picrite is characterized as a subalkaline picrobasalt and the underlying doleritic  
146 intrusion is a basalt, falling at the boundary between the alkaline and subalkaline domains  
147 (Fig. 1a).

148 Compared to NMORB (Fig. 1b; Sun and McDonough 1989), the basalt is rich in  
149 incompatible elements ( $\text{Zr}_N = 2.0$ ), while the picrite is depleted ( $\text{Zr}_N = 0.7$ ). The positive  
150 anomaly of Ba, the most prominent feature of both samples, corresponds to  
151 concentrations of 2,460 ppm ( $\text{Ba}_N = 390.5$ ) in the picrite and of 550 ppm ( $\text{Ba}_N = 87.3$ ) in  
152 the basalt. Also Sr concentrations are elevated (167 ppm in the picrite and 428 ppm in the  
153 basalt). Moreover, the NMORB-normalized patterns show positive anomalies of Ti,  
154 elevated Cs in the picrite (Cs in the basalt was below detection limit of 0.1 ppm) and a  
155 distinct trough for K in the basalt.

156 The chondrite-normalized (Boynnton 1984) REE patterns (Fig. 1c) are  
157 characterized by a moderate LREE over HREE enrichment ( $\text{La}_N/\text{Yb}_N = 3.5\text{--}4.0$ ) and near  
158 linear courses lacking any Eu anomaly ( $\text{Eu}/\text{Eu}^* = 1.0$ ).

159 **Chemical composition and textures of strontian and barian feldspars**

160       The BSE images reveal common barian feldspars and rare slawsonite in the  
161 interstitial phase between olivine and diopside phenocrysts and rare barian feldspars  
162 enclosed within diopsides (Fig. 2). Mode of occurrence of the barian feldspars is  
163 characterized in most cases by thin and rough mantling of plagioclase grains (Fig. 2a).  
164 These celsian-rim zones separate the plagioclases from hydrogrossular ( $\text{Ca}_3\text{Al}_2(\text{SiO}_4)_{3-x}(\text{OH})_{4x}$ ; Hgr) and vuagnatite ( $\text{CaAlSiO}_4(\text{OH})$ ; Vgt). Rarer are small plagioclase laths  
165 enclosed within the diopside phenocrysts, which are completely replaced by celsian, talc  
166 and possibly by Ca–Al silicate (Hgr or Vgt?) closely intergrown with Mg-rich mineral  
167 (serpentine or chlorite?; Fig. 2b). With the exception of talc, only mixture analyses were  
168 obtained from the original plagioclase due to small size of grains and their intergrowth.  
169 Described textures suggest resorption of original plagioclases followed by precipitation  
170 of Ba-rich feldspars, vuagnatite and hydrogrossular.  
171

172       Only one slawsonite grain has been detected in the seven thin sections of the  
173 Rovina picrite sample (Fig. 2a, c). The slawsonite replaces an original bytownite ( $\text{An}_{77}$ )  
174 lath (Fig. 2c), exhibits undulatory extinction, and is approximately 50  $\mu\text{m}$  across. The  
175 slawsonite composition varies from  $\text{Sl}_{91}\text{Cn}_3\text{An}_3\text{Ab}_3$  (core) to  $\text{Sl}_{82}\text{Cn}_3\text{An}_4\text{Ab}_9\text{Or}_2$  (rim).  
176 The BSE image (Fig. 2c) indicates the negligible content of Ba-rich inclusions. The  
177 slawsonite is rimmed by vuagnatite and celsian, which apparently infill fissures cutting  
178 the slawsonite grain. Chlorite and a serpentine–chlorite mixture were detected at the  
179 upper rim of the slawsonite grain.

180       The analyzed barian feldspars contain 27–41 wt% BaO, and the strontian  
181 feldspars 26–28 wt% SrO (Tab. 1). The nomenclature of the barian feldspars follows



182 Essene et al. (2005): celsian > 75 mol% Cn, hyalophane > 15 mol% and < 75 mol% Cn.  
183 An analysis of discontinuously zoned barian feldspar (Fig. 2d) revealed a hyalophane  
184 core  $\text{Cn}_{64.4}\text{An}_{1.1}\text{Ab}_{2.4}\text{Or}_{31.6}\text{Sl}_{0.5}$  and a celsian rim  $\text{Cn}_{91.8}\text{An}_{2.3}\text{Ab}_{0.5}\text{Or}_{4.7}\text{Sl}_{0.7}$ . The  
185 composition of interstitial celsians varies from  $\text{Cn}_{96.9}\text{An}_{0.3}\text{Ab}_{0.2}\text{Or}_2\text{Sl}_{0.6}$  to  
186  $\text{Cn}_{76.3}\text{An}_{4.7}\text{Ab}_3\text{Or}_{15.7}\text{Sl}_{0.3}$ , and the hyalophanes range from  $\text{Cn}_{72.2}\text{An}_{1.4}\text{Ab}_{5.1}\text{Or}_{21.1}\text{Sl}_{0.2}$  to  
187  $\text{Cn}_{57.3}\text{An}_{0.8}\text{Ab}_{3.1}\text{Or}_{38.5}\text{Sl}_{0.3}$  (Fig. 3).

188 Celsians and hyalophanes form rather thin rims, which cannot be readily  
189 distinguished from plagioclase under the optical microscope (Fig. 4a, b). Barium and  
190 strontium distribution maps (Fig. 4c, d) suggest the normal zoning of SrO and the reverse  
191 zoning of BaO in barian feldspars overgrowing the original plagioclases of  $\text{An}_{70-84}$ . The  
192 calcium distribution map (Fig. 3e) clearly distinguishes two other components enclosing  
193 plagioclases – vuagnatite from hydrogrossular. In addition, pectolite [ $\text{NaCa}_2\text{Si}_3\text{O}_8(\text{OH})$ ;  
194 Pct] was identified in the vuagnatite by the use of sodium distribution map (Fig. 3f).

## 195 DISCUSSION

### 196 Comparison with published data

197 Among the reported natural slawsonites, the composition of the studied grain (82–  
198 91 mol% of Sl and 3–4 mol% of Cn) bears a striking resemblance to that from the  
199 Sarusaka ultramafic xenolith described by Matsubara (1985). This author assumed a low-  
200 *T* origin of the slawsonite–celsian assemblage (*ca.* 300–400 °C) and ruled out any  
201 significant substitution of Sr in celsian. The latter is also true for the Rovina celsians, in  
202 which the SrO content in barian feldspars does not exceed 0.75 wt% (2.6 mol% Sl).

203 Celsians (97–76 mol% of Cn) and hyalophanes (72–57 mol% of Cn) replace  
204 bytownites ( $\text{An}_{70-84}$ ) in the studied picrite. Element distribution maps document that the

205 celsians and hyalophanes attain maximal SrO concentrations next to the bytownites (Fig.  
206 4d) and that BaO concentrations increase rimwards (i.e., as the barian feldspar grew; Fig.  
207 2d, 4c). This could be caused by the incompatibility of Ba in the denser structure of the  
208 calcic plagioclase (Blundy and Wood 1991) during bytownite dissolution.

209       The entire dataset acquired spans the miscibility field of the ternary (Ba, K, Na)-  
210 feldspar system (Lagache and Catel 1992; Viswanathan 1992). Single discontinuously  
211 zoned feldspar (Fig. 2d), sufficiently large to obtain EDS analyses, with a hyalophane  
212 core (64 mol% Cn) rimmed by celsian (91 mol% Cn), could indicate the possibility of a  
213 solvus gap between celsian and hyalophane (Essene et al. 2005).

214       Albite content in barian feldspars correlates positively with  $T$  (Pan and Fleet  
215 1991; Chabu and Boulègue 1992; Morishita 2005). Essene et al. (2005) proposed a  
216 geothermometer on the basis of observations of barian feldspars with various Ab contents  
217 in rocks metamorphosed by various grades. The Ab contents of celsians and hyalophanes  
218 in the picrite vary from 0.1 to 6.7 mol%, which is, according to Essene et al. (2005),  
219 characteristic of diagenetic conditions ( $T = 100\text{--}250\text{ }^{\circ}\text{C}$ ). Nevertheless, the amount of Na  
220 in barian feldspar increases when buffered by albite. Hence, given the absence of albite in  
221 the Rovina picrite, the geothermometer is not applicable to our dataset.

## 222 Constrains on slawsonite–celsian–hyalophane genesis

223       Taking into account the partial dissolution of bytownites, replacement textures  
224 and the reverse zoning of BaO in Ba-feldspars, we suggest that the studied slawsonite,  
225 celsians and hyalophanes were not in equilibrium with the picrite magma. An  
226 introduction of the necessary Sr and Ba due to crustal contamination of the intruding  
227 magma is highly improbable, as the Sr- and Ba-rich phases replace the already

228 crystallized idiomorphic plagioclases. Most importantly, the picrite shows a pristine  
229 basaltic whole-rock geochemical signature with superimposed sharp Ba and Sr spikes  
230 (Fig. 1). In addition, apparent loss of K in the basalt is a characteristic feature of low  $T$   
231 (70–150 °C) basalt–seawater alteration (Bloch and Bischoff 1979; Seyfried 1979) and  
232 elevated concentration of Cs (0.9 ppm) in picrite relate to very low  $T$  stages of alteration  
233 and weathering (Hart et al. 1974). All these observations point to syn- or even post-  
234 depositional subsolidus open-system processes, most probably due to an interaction with  
235 fluids, rather than to an equilibration of crustal material with basaltic magma.

236         Textural relationships and the presence of vuagnatite, hydrogrossular and  
237 pectolite provide a valuable insight into genesis of the Sr- and Ba-rich phases.  
238 Vuagnatite, hydrogrossular and pectolite had been considered as products of  
239 rodingitization derived from replacement of plagioclase under the conditions of prehnite–  
240 pumpellyite facies (Sarp et al. 1976; Matsubara et al. 1977; Barriga and Fyfe 1983;  
241 Palandri and Reed 2004; Li et al. 2007; Bach et al. 2013). Rodingitization is a type of Ca-  
242 metasomatism leading to depletion of SiO<sub>2</sub> and Na<sub>2</sub>O, which ideally occurs at  $T < 350$  °C  
243 (Pirajno 2013). Rodingitization and serpentinization take place contemporaneously. The  
244 former is not caused by Ca-rich nature of serpentinization fluids but by the stability of  
245 Ca-silicates (i.e. hydrogrossular; Bach and Klein, 2009). Hence, the principal  
246 metasomatic mechanism of rodingitization is not a calcium addition but more importantly  
247 a silica loss during serpentinization process (Frost and Beard 2007; Frost et al. 2008;  
248 Bach et al. 2013).

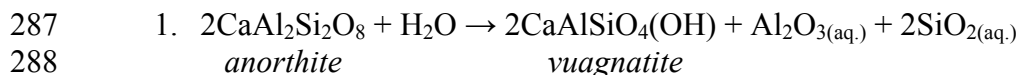
249         Taking into account constrains on genesis of the other reported Sr- and Ba-  
250 silicates, stronalsite and banalsite (Hori et al. 1987), which formed under silica-poor

251 conditions, we could infer that silica loss connected to rodingitization and  
252 serpentinization could represent a viable mechanism for formation of strontian and barian  
253 feldspars in our case. Moreover, Hori et al. (1987) suggested that at slightly Si-richer  
254 conditions, during decomposition of plagioclase to any Sr-excluding mineral such as  
255 hydrogrossular, formation of slawsonite instead of stronalsite would be favored. Besides  
256 that, Liferovich et al. (2006) proposed three major mechanisms for genesis of stronalsite  
257 and banalsite: crystallization from a silica-undersaturated sodic magma, formation during  
258 high-grade metamorphism and, the most common, deuteritic alteration of feldspathoid-  
259 bearing igneous rocks. None could reflect likely genesis of slawsonite, celsians and  
260 hyalophanes from the Rovina picrite – Ba and Sr concentrations are too high for primary  
261 magmatic composition, the picrite and any rocks within the Prague Basin are essentially  
262 unmetamorphosed (Chlupáč et al. 1998), and deuteritic alteration would require  
263 contamination by crustal rocks, which would not preserve the rest of primary magmatic  
264 signature of the picrite (Fig. 1).

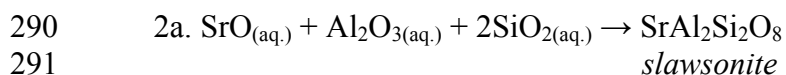
265         A possible clue to strontian and barian feldspar genesis in the studied picrite could  
266 be based on thermodynamic modeling of micro-rodingitization of olivine-rich troctolites  
267 (Frost et al. 2008). Thermodynamic reactions of plagioclase alteration to prehnite and  
268 grossular/hydrogrossular accompanied by serpentinization and chloritization of olivine  
269 were calculated by Frost et al. (2008) for maximum  $T = 350$  °C (based on a lower  
270 stability limit of clinozoisite) and  $P = 0.1$  GPa. Since vuagnatite shows the same mode of  
271 occurrence as prehnite, the former merely requiring more silica-deficient conditions  
272 (Matsubara et al. 1977), the micro-rodingitization model could be adopted for the Rovina  
273 picrite.

274 Hence, desilication of plagioclase to vuagnatite/hydrogrossular and  
275 serpentinization of olivine could have taken place simultaneously in the Rovina picrite. If  
276 true, the alteration reactions would have involved surplus H<sub>2</sub>O, SrO and BaO (Fig. 5). We  
277 suggest an addition of SrO and BaO by a fluid as celsians even replace plagioclase  
278 inclusions in magmatic diopside. The possibility that Ba and Sr were in situ products of  
279 original plagioclase decomposition requires a late-magmatic contamination of the picrite.  
280 Such case could be expressed by an accommodation of Ba and Sr in outer zones of  
281 interstitial plagioclases that were subsequently dissolved. On the contrary, Ba and Sr  
282 would not have been accommodated in plagioclases, which were fully crystallized before  
283 the diopside growth had been completed. At first, decomposition of plagioclase to  
284 vuagnatite occurred, releasing silica and alumina for the reaction of slawsonite or celsian  
285 with BaO and SrO supplied by the fluid:

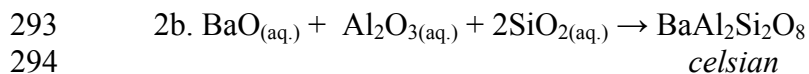
286



289



292 or

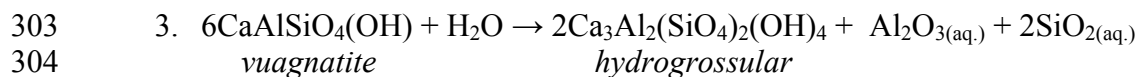


295

296 Genesis of slawsonite and celsian was most likely triggered by decomposition of  
297 plagioclase and these minerals precipitated directly from the fluid. Slawsonite seems the  
298 older phase because it is cut by veintlets filled by celsian. As the plagioclase is not a pure

299 anorthite but bytownite, the excess of Na could have been either compensated by a  
300 formation of pectolite or lost from the system. Furthermore, vuagnatite was transformed  
301 to hydrogrossular:

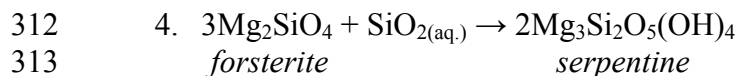
302



305

306 Hydrogrossular could also originate by the reaction:  $3\text{CaAl}_2\text{Si}_2\text{O}_8 + 2\text{H}_2\text{O} \rightarrow$   
307  $\text{Ca}_3\text{Al}_2(\text{SiO}_4)_2(\text{OH})_4 + \text{Al}_2\text{O}_{3(\text{aq.})} + 4\text{SiO}_{2(\text{aq.})}$ . However, hydrogrossulars enclosed within  
308 vuagnatites (Fig. 3e) indicate that vuagnatite decomposition was their most probable  
309 genesis. Excess silica provided by plagioclase dissolution was consumed by  
310 contemporaneous serpentinization of olivine:

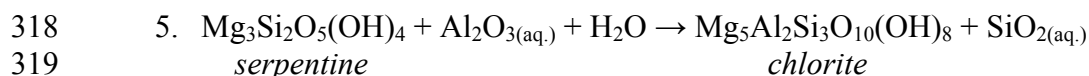
311



314

315 Alumina released by desilication led to the formation of chlorite and rarely talc from  
316 serpentine and closed the rodingitization of the Rovina picrite:

317



320

321 In summary, strontian and barian feldspars genesis was most probably connected to  
322 diffusional metasomatism – rodingitization and serpentinization reactions, which were  
323 caused by SrO–BaO–H<sub>2</sub>O intergranular fluid infiltration. Such reactions could operate at

324  $T \leq 350$  °C, which represents lower stability limit of clinozoisite. The lower limit of  $T$   
325 range suitable for proposed reactions could be constrained by the use of a chlorite  
326 thermometer, as chlorites represent the final stage of the Rovina picrite alteration.  
327 Discussion on chlorite thermometry (de Caritat et al. 1993) provides variable approaches  
328 on  $T$  calculations. For chlorites in Al-saturated environment, e.g., in the presence of other  
329 aluminous minerals, the Kranidiotis and MacLean (1987) thermometer is the most  
330 convenient one. Its calculation (Appendix) gives  $T$  range ca. 320–160 °C. Pressure  
331 constraint provides a transition of celsian to cymrite, which occurs at  $P \sim 0.5$  GPa and  
332 350 °C (Graham et al. 1992).

333 The source of Ba and Sr, enriched in the picrite and basalt, is a matter for  
334 discussion. Morishita (2005) suggested that Ba and Sr are important constituents in fluids  
335 related to the formation of metasomatized rocks in serpentinite–matrix mélanges formed  
336 during subduction, which dehydrated Ba-rich sediments – a possible source for Ba-  
337 bearing fluids. The Prague Basin, on the contrary, represented a different tectonic setting  
338 – rifted basin – during Ordovician (Havlíček 1982). Moreover, the process of  
339 rodingitization is characteristic of seafloor metasomatism associated with serpentinization  
340 (Pirajno 2013). Thus, range of formation  $T$  (350–160 °C) of the Ba- and Sr- feldspars  
341 points to their post-magmatic origin due to seafloor metasomatism.

## 342 **CONCLUDING REMARKS AND IMPLICATIONS**

343 The whole-rock geochemical signature of the Late Ordovician picrite sill and the  
344 underlying doleritic basalt intrusion are characterized by eye-catching enrichment in Ba  
345 and Sr. The accommodation of such significant amounts of Ba and Sr in rock samples  
346 resulted in the precipitation of a common barian (celsian and hyalophane) and a rare

347 strontian (slawsonite) feldspar assemblage, filling the interstitial space between the  
348 olivine and diopside phenocrysts. This represents the first known occurrence of  
349 slawsonite in Europe. From the study of this feldspar assemblage in the picrite, the  
350 following implications emerged:

351 1) Ba- and Sr-feldspars formed due to desilication of original plagioclases during  
352 interaction with SrO–BaO–H<sub>2</sub>O fluid.

353 2) The surplus Sr and Ba was probably introduced by intergranular fluid during  
354 diffusional seafloor metasomatism (i.e. rodingitization and serpentinization).

355 2) Ba- and Sr-feldspars precipitated directly from fluid phase at 350 –160 °C and  
356 less than 0.5 GPa.

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364 the paper.

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#### 495 **FIGURE CAPTIONS**

- 496 Figure 1. Whole-rock geochemical analyses of the Rovina picrite and basalt. (a) TAS  
497 diagram (Le Bas et al. 1986; dividing line between the subalkaline and alkaline domains  
498 after Irvine and Baragar 1971). (b) Spiderplot normalized by average NMORB  
499 composition with model courses of the OIB and EMORB reservoirs (from Sun and  
500 McDonough 1989). (c) Chondrite-normalized (Boynton 1984) REE patterns.
- 501
- 502 Figure 2. Back-scattered electron images with positions of EDS analyses. (a) Interstitial  
503 phase – slawsonite and celsians. Rectangle indicates position of Fig. 4. (b) Original

504 plagioclase enclosed within diopside. Ca+Al+Si+Mg stand for mixture analysis of Ca–Al  
505 silicate (hydrogrossular or vuagnatite) and Mg mineral (serpentine or chlorite). (c)  
506 Slawsonite. (Dark stain below position of Sl<sub>4</sub> analysis represents surface external  
507 impurity.) (d) Discontinuously zoned barian feldspar. An = anorthite; Chl = chlorite; Cn =  
508 celsian, Di = diopside; Fo = forsterite; Hgr = hydrogrossular; Hy = hyalophane; Sl =  
509 slawsonite; Sp = spinel; Srp = serpentine; Tlc = talc; Vgt = vuagnatite. Chlorite *T* after  
510 Kranidiotis and MacLean (1987).

511

512 Figure 3. Analyses of barian and strontian feldspars with An content < 6 mol% in the  
513 ternary plots NaAl<sub>2</sub>Si<sub>2</sub>O<sub>8</sub>–KAlSi<sub>3</sub>O<sub>8</sub>–(SrAl<sub>2</sub>Si<sub>2</sub>O<sub>8</sub> + BaAl<sub>2</sub>Si<sub>2</sub>O<sub>8</sub>) (a) and BaAl<sub>2</sub>Si<sub>2</sub>O<sub>8</sub>–  
514 KAlSi<sub>3</sub>O<sub>8</sub>–SrAl<sub>2</sub>Si<sub>2</sub>O<sub>8</sub> (b).

515

516 Figure 4. Microphotograph of interstitial mineral assemblage: Plane polarized light (a)  
517 and crossed polars (b) Element distribution maps of Ba (c), Sr (d), Ca (e), and Na (f). An  
518 = anorthite; Chl = chlorite; Cn = celsian; Hgr = hydrogrossular; Pct = pectolite; Sl =  
519 slawsonite; Vgt = vuagnatite.

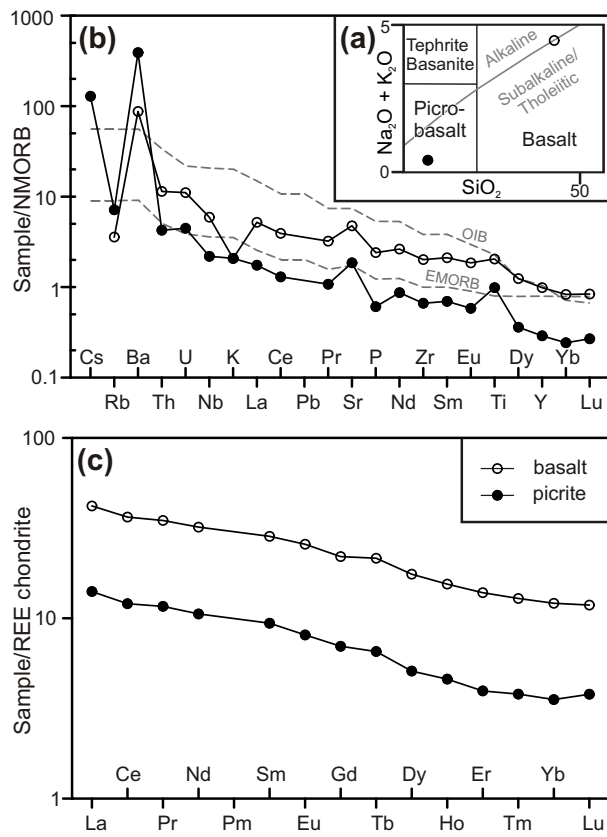
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521 Figure 5. Schematic sketch of textural relationships and reactions that operated during the  
522 alteration of plagioclase and olivine in the picrite.

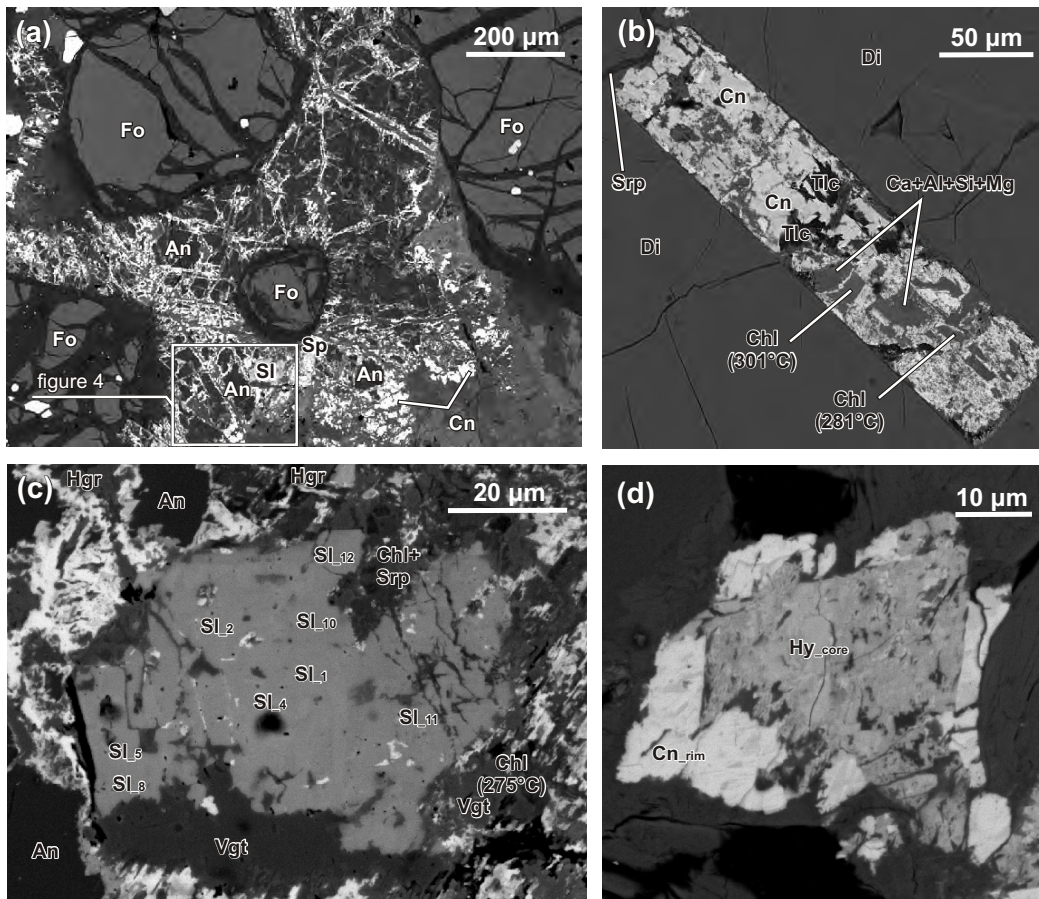
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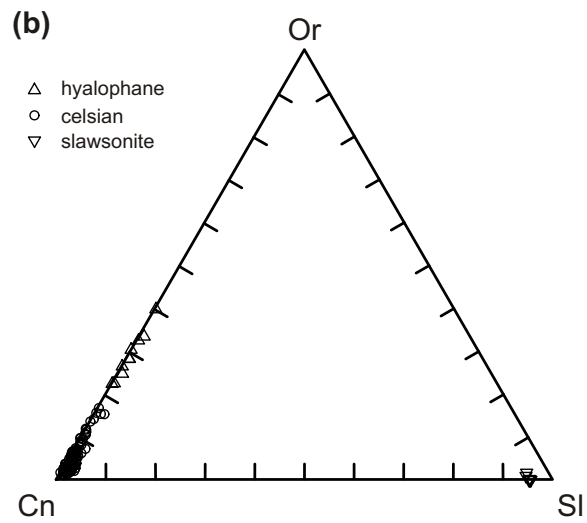
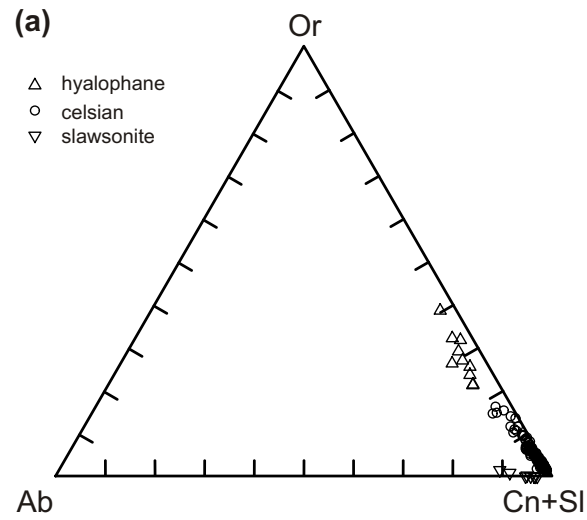
## TABLES

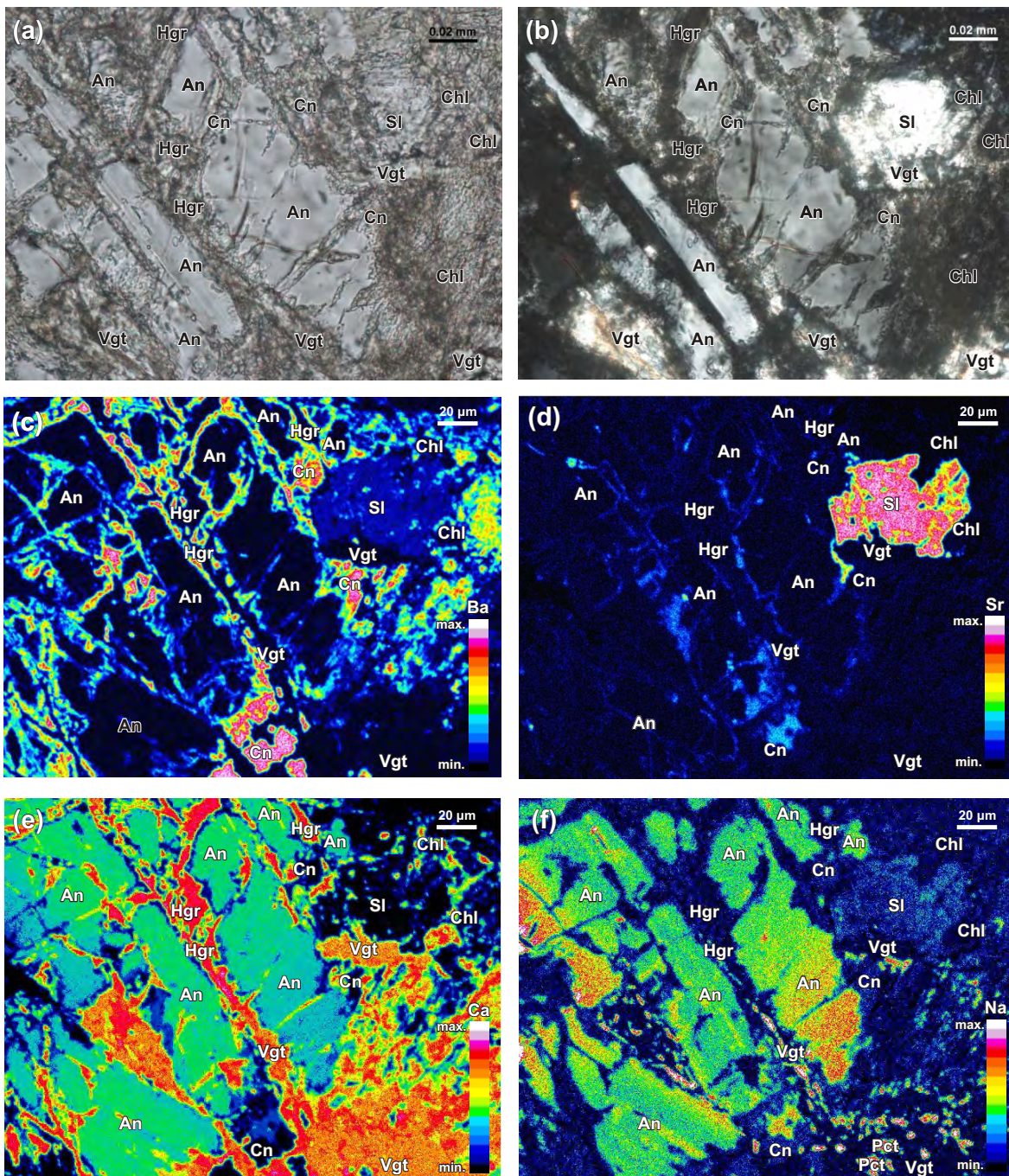
524 Table 1. Selected strontium and barium feldspar analyses (wt%) recalculated to structural  
525 formulae (apfu) and end members (mol%).

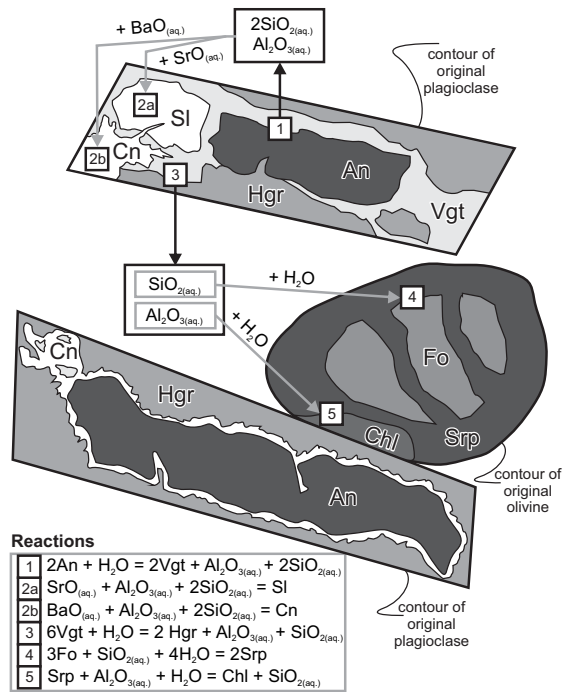












	SL_1	SL_2	SL_4	SL_5	SL_8	SL_10	SL_11	SL_12	Cn_1	Cn_2	Cn_117	Cn_128	Hy_1	Hy_10	Cn_rim	Hy_core
<b>Na<sub>2</sub>O</b>	0.19	0.28	0.25	0.23	0.44	0.46	0.71	0.88	0.02	0.02	0.17	0.27	0.46	0.29	0.04	0.23
<b>Al<sub>2</sub>O<sub>3</sub></b>	31.45	31.36	31.26	31.06	31.00	31.16	30.35	30.55	25.66	26.31	25.82	25.38	24.68	23.97	25.73	24.31
<b>SiO<sub>2</sub></b>	37.30	37.07	37.10	36.64	37.48	37.85	38.39	39.21	30.69	31.18	34.13	35.69	37.48	42.69	32.04	40.69
<b>K<sub>2</sub>O</b>			0.02		0.06	0.03	0.14	0.24	0.26	0.37	1.48	2.18	2.92	5.58	0.62	4.52
<b>CaO</b>	0.54	0.55	0.51	0.67	0.62	0.82	0.71	0.60	0.05		0.13	0.78	0.22	0.13	0.37	0.19
<b>FeO</b>	0.15	0.14	0.13	0.07	0.19	0.20	0.17	0.36	0.16	0.27	0.43	0.29	0.29	0.10	0.58	0.41
<b>SrO</b>	28.24	28.29	27.86	28.04	27.78	26.98	26.26	26.03	0.17		0.17	0.10	0.06	0.11	0.22	0.15
<b>BaO</b>	1.55	1.23	1.69	1.75	1.67	1.68	1.75	1.53	41.32	39.97	37.11	34.62	32.43	27.06	39.93	29.99
<b>Total (wt%)</b>	99.42	98.93	98.83	98.46	99.21	99.18	98.48	99.39	98.33	98.12	99.43	99.32	98.55	99.94	99.54	100.47
<b>Formulae (cations based on 8 oxygens)</b>																
<b>Na</b>	0.02	0.03	0.03	0.02	0.05	0.05	0.07	0.09	0.00	0.00	0.02	0.03	0.05	0.03	0.01	0.03
<b>Al</b>	2.00	2.01	2.00	2.00	1.98	1.98	1.93	1.92	1.96	1.99	1.87	1.81	1.74	1.59	1.92	1.64
<b>Si</b>	2.02	2.01	2.02	2.01	2.03	2.04	2.08	2.09	1.99	2.00	2.10	2.15	2.24	2.40	2.03	2.33
<b>K</b>	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.02	0.02	0.03	0.12	0.17	0.22	0.40	0.05	0.33
<b>Ca</b>	0.03	0.03	0.03	0.04	0.04	0.05	0.04	0.03	0.00	0.00	0.01	0.05	0.01	0.01	0.02	0.01
<b>Fe</b>	0.00	0.00	0.00	0.00	0.00	0.01	0.00	0.01	0.01	0.01	0.01	0.01	0.01	0.00	0.02	0.01
<b>Sr</b>	0.88	0.89	0.88	0.89	0.87	0.84	0.82	0.81	0.01	0.00	0.01	0.00	0.00	0.00	0.01	0.00
<b>Ba</b>	0.03	0.03	0.04	0.04	0.04	0.04	0.04	0.03	1.05	1.00	0.89	0.82	0.76	0.60	0.99	0.67
<b>Total (apfu)</b>	4.99	5.00	5.00	5.00	5.01	5.00	5.00	5.00	5.04	5.03	5.03	5.04	5.03	5.03	5.04	5.03
<b>End members</b>																
<b>An</b>	3.23	3.25	3.05	3.97	3.61	4.87	4.18	3.53	0.30	0.00	0.83	4.68	1.35	0.78	2.29	1.10
<b>Ab</b>	2.07	3.01	2.74	2.48	4.62	4.92	7.60	9.28	0.21	0.22	1.90	2.97	5.12	3.06	0.50	2.40
<b>Or</b>	0.00	0.00	0.17	0.00	0.39	0.19	1.00	1.63	1.98	2.92	11.08	15.67	21.13	38.48	4.65	31.59
<b>Cn</b>	3.39	2.68	3.70	3.79	3.56	3.63	3.75	3.26	96.91	96.85	85.61	76.36	72.20	57.34	91.80	64.44
<b>SI</b>	91.31	91.05	90.33	89.76	87.82	86.39	83.48	82.29	0.60	0.00	0.57	0.32	0.20	0.34	0.75	0.47