1 Revision 1, Nov. 3, 2013

Detection of Iron Substitution in Natroalunite-Natrojarosite Solid Solutions and Potential Implications for Mars

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19 **Abstract** – Natroalunite containing substantial amounts of Fe occurs as a prominent secondary 20 phase during acid-sulfate alteration of pyroclastic basalts in volcanic fumaroles in Nicaragua and 21 elsewhere, and has been observed in laboratory simulations of acid-sulfate alteration as well. 22 Reaction path models constrained by field and experimental observations predict that Fe-rich 23 natroalunite should also form as a major secondary phase during alteration of martian basalt 24 under similar circumstances. Here, we evaluate the potential to use spectroscopic methods to 25 identify minerals from the alunite group with compositions intermediate between natroalunite 26 and natrojarosite on the surface of Mars, and to remotely infer their Fe contents. X-ray 27 diffraction and spectroscopic measurements (Raman, visible/near infrared, mid-infrared, 28 Mössbauer) were obtained for a suite of synthetic solid solutions with a range of Fe contents ranging from natroalunite to natrojarosite. In the visible/near infrared, minerals with 29 30 intermediate compositions display several spectral features not evident in endmember spectra 31 that could be used to remotely identify these minerals and infer their composition. In addition, 32 Raman spectra, mid-infrared spectra, and X-ray diffraction peaks all show systematic variation 33 with changing Fe content, indicating that these methods could potentially be used to infer 34 mineral compositions as well. The results suggest that alunite group minerals with intermediate 35 Fe compositions may be able to account for some visible/near-infrared and Mössbauer spectral 36 features from Mars that had previously been unidentified or attributed to other phases. Overall, 37 our findings indicate that consideration of solid solutions may lead to new identifications of 38 alunite group minerals on the surface of Mars, and raise the possibility that minerals with 39 compositions intermediate between natroalunite and natrojarosite may be widely distributed on 40 the planet.

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Keywords: Mars, alunite group, visible/near infrared spectra, Raman spectra, hydrothermal, acidsulfate alteration.

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INTRODUCTION

47 Minerals in the alunite group have the ideal molecular formula $AB_3(SO_4)_2(OH)_6$, where the A site is most commonly occupied by the monovalent ions K^+ , Na^+ , and H_3O^+ (hydronium) and the 48 B site is occupied primarily by Al^{3+} and Fe^{3+} (for convenience, we use here the term "alunite 49 subgroup" to refer collectively to alunite and natroalunite, and "jarosite subgroup" to refer to 50 51 jarosite, natrojarosite, and hydroniumjarosite; Table 1) (Stoffregen et al. 2000; Mills et al. 2009). 52 In natural systems, minerals in the alunite and jarosite subgroups exhibit extensive solid solution 53 mixing in the A site. In contrast, reports of mixing between Al and Fe on the B site among these 54 minerals are more limited, with most published compositions having compositions close to either 55 the Al or Fe endmembers (Dutrizac and Jambor 2000; Stoffregen et al. 2000; Papike et al. 2006, 56 2007). Reports of minerals in the alunite subgroup containing more than a few mole percent Fe 57 in the B site have been particularly rare. In contrast, minerals with a full range of Fe-Al 58 compositions intermediate between endmembers from the alunite and jarosite subgroups can be 59 readily synthesized in the laboratory from mixtures of sulfate salts (Brophy et al. 1962; Härtig et al. 1984; McCollom et al. 2013a). More extensive solid solution mixing between Fe and Al in 60 61 the B site has been reported for other minerals in the broader alunite supergroup, but these 62 observations appear to be limited to those minerals where there is substantial (>55%) substitution of PO_4^{3-} or other trivalent anions in the sulfate site as well as >80% substitution of a divalent 63 64 cation in the A site (e.g., Scott, 1987).

In recent studies, natroalunites containing substantial amounts of Fe were observed in acidsulfate altered mineral deposits from fumaroles at Cerro Negro (CN) volcano and other volcanic sites in Nicaragua (Fig. 1) (McCollom et al. 2013b; Hynek et al. 2013). Precipitation of minerals

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68 from the alunite group with a broad range of compositions intermediate between natroalunite and 69 natrojarosite were also found to precipitate during incipient acid-sulfate alteration of basalt in 70 laboratory experiments (McCollom et al. 2013a). In both cases, the Fe-rich natroalunites occur 71 along with amorphous silica and Ca-sulfates (gypsum or anhydrite) as the predominant 72 secondary phases formed during the initial stages of acid-sulfate alteration, together with minor 73 hematite and other trace phases. The co-occurrence of Fe-bearing natroalunite with amorphous 74 silica \pm hematite during acid-sulfate alteration of basalt may be a common circumstance, since 75 apparently similar assemblages have been reported in altered basaltic tephra from Hawaii 76 (Morris et al. 2005; Bishop et al. 2007). In these cases, however, detailed chemical analyses of 77 the natroalunite were not provided.

The Fe-rich natroalunites found at Cerro Negro precipitate from Fe, Al, and Na released by dissolution of primary igneous phenocrysts in the basalt (including plagioclase, olivine, and augite) as they react with sulfuric acid (McCollom et al. 2013b). The sulfuric acid in the fumaroles is produced during cooling and condensation of SO₂-bearing volcanic vapors through reactions such as:

$$3\mathrm{SO}_2 + 2\mathrm{H}_2\mathrm{O} \rightarrow 2\mathrm{H}_2\mathrm{SO}_4 + \mathrm{S}^0. \tag{1}$$

Natroalunite is absent from the most extensively altered deposits at Cerro Negro, suggesting that this mineral may form as a transient phase during the early stages of alteration, and then disappears as Na, Al, and Fe are mobilized from the deposits during more extensive alteration. In both the natural samples and experimental products, the close spatial association of the Febearing natroalunite with amorphous silica and spheroidal hematite (Figs. 1, 2) suggests that these minerals precipitate concurrently. The apparent transient nature of the Fe-rich natroalunites in fumarole deposits together with the paucity of detailed chemical analyses of

alunite subgroup minerals found in such settings may account for the lack of previous literature
reports of natroalunites with elevated Fe contents.

93 Reaction-path models based on constraints provided by field and laboratory observations 94 predict that Fe-rich natroalunite should also occur as a major product during hydrothermal acid-95 sulfate alteration of pyroclastic basalts on Mars (McCollom et al. 2013a.b). The possibility that 96 Fe-bearing natroalunites might be present on Mars, where they could potentially serve as 97 indicators of incipient acid-sulfate alteration in fumarolic environments, led us to investigate 98 methods by which they could be detected remotely. We report here on the spectral 99 characteristics of solid solutions in this mineral group, focusing on the impact of Fe-Al solid 100 solution mixing in the B site. While there have been a number of previous studies on the spectral 101 properties of pure endmembers in the alunite and jarosite subgroups (e.g., Serna et al. 1986; 102 Sasaki et al. 1998; Bishop and Murad 2005; Frost et al. 2006a,b; Murphy et al. 2009; Maubec et 103 al. 2012), to our knowledge there are no previously published investigations of solid solutions 104 with compositions intermediate between the endmembers. Accordingly, spectral measurements 105 were made on synthetic solid solutions with compositions intermediate between natroalunite and 106 natrojarosite, and are compared with field and laboratory spectra of natural samples containing 107 Fe-bearing natroalunite.

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METHODS

109 Synthetic mineral solid solutions with compositions intermediate between natroalunite and 110 natrojarosite were produced from mixtures of Na-, Al-, and Fe-sulfate salts by heating the salts 111 with 0.1 M H_2SO_4 at 145 °C for 2-6 days in Teflon-lined reaction vessels, as described in 112 McCollom et al. (2013a). The chemical composition of the synthesized minerals was determined 113 by dissolving an aliquot in a combination of concentrated HNO₃ and HCl, and analyzing the 114 resulting fluids by inductively coupled plasma – atomic emission spectroscopy (ICP-AES). 115 Additional characterization of the morphology and chemical composition of synthetic and 116 natural minerals were performed using a combination of scanning electron microscopy (SEM) 117 equipped with an electron-dispersive X-Ray spectrometer (EDS), X-ray diffraction (XRD), and 118 electron microprobe analysis (EMPA). X-ray diffraction analyses were performed with a Terra 119 instrument (inXitu Incorporated, Campbell, CA) using $CuK\alpha$ radiation, which is functionally 120 equivalent to the ChemMin instrument deployed on the Curiosity rover on Mars (Blake et al. 121 2012). The SEM analyses were conducted on a JEOL 6480LV in backscattered electron (BSE) 122 mode with an accelerating voltage of 15 kV. The EDS analyses were performed using an Oxford 123 Instruments collector and processed using INCA software. Most analyses were performed on 124 samples mounted on aluminum stubs using double-sided carbon tape and analyzed uncoated. In 125 a few cases, samples were mounted in epoxy and polished in order to examine the minerals in 126 cross-section. Electron microprobe analyses were performed using a JEOL JXA 8600 equipped 127 with four wavelength-dispersive spectrometers. Samples were analyzed with an accelerating 128 voltage of 15 kV, a 5 µm beam, and a beam current of 10 nA. Anhydrite was used as the 129 standard for sulfur.

130 Laser Raman spectroscopic measurements of powdered synthetic natroalunite-natrojarosite solid solutions were made using a HoloLab5000-532 (Kaiser Optical System, Inc.). 131 The 132 instrument uses a 532 nm line from a frequency-doubled Nd:YVO₄ laser for excitation and a 133 20X, long-working-distance objective with numerical aperture (NA) of 0.4 for Raman signal 134 collection, which produced a well-focused laser beam of $\sim 6 \,\mu m$ diameter at the sampling spot. 135 The system is functionally equivalent to the Mars Microbeam Raman Spectrometer (MMRS) 136 under development for robotic planetary surface exploration (Wang et al. 2003). The

measurements cover a spectral range of 100-4000 cm^{-1} with a resolution of ~4-5 cm^{-1} . With a 137 standard wavelength calibration procedure, the wavelength accuracy is <0.5 cm⁻¹, and precision 138 is estimated to be <0.3 cm⁻¹. To check the chemical homogeneity of each sample, 5-10 Raman 139 140 spectra were obtained from each powdered sample. Raman spectra for the endmembers of this 141 series of synthetic solid solution were confirmed by comparison with standard spectra for 142 natroalunite and natrojarosite. Because of their small size, Raman spectra of natural samples 143 were recorded using a Horiba LabRAM HR confocal spectrometer with a 17 mW, 633 nm line of 144 a HeNe laser as excitation source. Individual spectra were recorded using a 40X long-distance objective. Measurements were made over a spectral range of 100-1200 cm⁻¹, and spectra 145 146 processed using LabSpec 5 software.

Visible/near-infrared reflectance (VNIR) spectra were acquired in the laboratory at 147 148 wavelengths of 350 to 2500 nm using an Analytical Spectral Devices (ASD) FieldSpec3 fitted with a 5° foreoptic. The field of view was illuminated with a halogen light source at a 30° angle. 149 150 the spectrometer was calibrated using a spectralon reflectance standard, and spectra were 151 measured at an emission angle of approximately 30°. Spectra were collected from particulates for 152 synthetic samples, and from altered cinders and rock chips for natural samples. Several tens of 153 spectra were taken of each sample, rotated or turned between measurements, to assess 154 heterogeneity. Field VNIR spectra were measured using a TerraSpec4 high-resolution reflectance 155 spectrometer (Analytical Spectral Devices, Inc.) as described in Marcucci et al. (2013). Spectra were 156 measured for wavelengths from 350 to 2500 nm with a resolution of ~3-6 nm in the visible and ~12 157 nm in the infrared, over a spot size of 10 mm.

Mid-infrared reflectance (MIR) spectra were acquired at 4 cm⁻¹ sampling intervals on a Nicolet Magna 860 Fourier Transform Infrared (FTIR) spectrometer. Samples were first placed in a sample chamber purged of CO₂ and H₂O, and diffuse reflectance (DR) spectra were then

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measured between 5000 and 400 cm⁻¹ (2–25 μ m) using a KBr beamsplitter and mercury 161 162 cadmium telluride (MCT-B) detector. Each measurement included 1000 scans. A sample of 163 rough aluminum was used as the reflectance standard over the whole spectral range and was 164 assumed to have a reflectance of one. For particulate samples, additional measurements of 165 spectra were acquired using an attenuated total reflectance (ATR) accessory, the SensIR 166 Durascope. After acquiring a background of the crystal alone for calibration, each sample was pressed and held on a diamond crystal and spectra acquired over the wavenumber range 4000 to 167 400 cm⁻¹ (2.5–25 µm) using a KBr beamsplitter and uncooled deuterated triglycine sulfate 168 169 (DTGS) detector with a KBr window. For each sample, 50 scans per sample were made.

Several of the synthetic Fe-bearing natroalunites as well as several field and experimental samples were analyzed by Mössbauer (MB) spectroscopy. Measurements were made at room temperature using a conventional constant-acceleration spectrometer in transmission geometry with a 57 Co/Rh source. An α -Fe foil was used to calibrate isomer shifts and velocity scale at room temperature. The magnetic hyperfine parameters including isomer shift (IS) and quadrupole splitting (QS) were fit using the NORMOS program (Brand, 1987), which assumes a distribution of hyperfine parameters during spectral fitting.

177 Text tables listing the results of XRD, Raman, VNIR, MIR, and ATR measurements for the

178 synthetic solid solution series and natural samples are deposited as supplementary materials¹.

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RESULTS

180 Morphology and chemical composition of natural and synthetic minerals

¹ Deposit item AM-XX-XXXX, one Supplemental Figure and six Supplemental Tables. Deposit items are stored on the MSA web site and available via the *American Mineralogist* Table of Contents. Find the article in the table of contents (at GeoScienceWorld, ammin.geoscience-world.org or www.minsocam.org), and then click on the deposit link.

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The Fe-bearing natroalunite found in the acid-sulfate altered deposits from Nicaragua and produced in basalt alteration experiments have a fine-grained (mostly <10 µm), pseudocubic morphology, occurring as either single crystals or as dense monomineralic crusts, mostly

184 occurring in patches precipitated on the external surfaces of altered basalts or lining gas vesicles 185 (Fig. 1) (McCollom et al. 2013a.b). A summary of the relative amounts of Fe and Al in the B 186 site for several examples of natroalunite from natural systems and basalt alteration experiments, 187 expressed as Fe# $[=100 \times \text{Fe}/(\text{Al} + \text{Fe}), \text{ molar basis}]$, is provided in Table 2. The natural samples 188 exhibit considerable variability in the relative proportions of Fe and Al, both between samples 189 and for measurements of individual crystals within a single sample (the latter is reflected in the 190 relatively large standard deviations for some samples in Table 2). Additionally, individual 191 crystals from natural samples and from basalt alteration experiments often exhibit zonation in 192 Fe-Al contents (Fig. 1c,f). In all of these samples, EDS and EMPA indicate that the A site is 193 dominated by Na with little or no K present (McCollom et al. 2013a,b). The A site also appears to contain some amount of H_3O^+ , but uncertainties in the measurements of Na preclude 194 195 quantification of the amounts that may be present. For many of the samples listed in Table 2, the 196 sparse distribution and small grain size precluded analysis by electron microprobe, so only 197 compositional data from EDS are available. Although EDS is generally regarded as being semi-198 quantitative, comparison of measured values for Fe# determined by EDS with wet chemical and 199 EMPA results were found to agree within 10 % (McCollom et al. 2013a).

200 The synthetic solid solutions have a fine-grained (20 µm or less) pseudocubic morphology 201 very similar to that found in natural samples from volcanic environments (Fig. 3). The chemical 202 formulas of the synthetic minerals, as determined by wet chemical methods, are listed in Table 3 (complete results of the wet chemical analyses are provided as Supplemental Table 1¹). Based 203

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on the wet chemistry, the synthetic minerals contain significant amounts of H_3O^+ in the *A* site, as well as deficiencies in the *B* site relative to the occupancy expected for the ideal molecular formula (Fe + Al = 3 per formula unit). Examination of the synthetic minerals by EDS indicated that they were homogenous in composition, with a narrow range of compositions found for measurements of different crystals within each sample. Only trace amounts of other phases were detected in the samples.

210 X-ray diffraction

211 Analysis of the synthetic minerals by XRD shows regular shifts in the position of many 212 diffraction peaks with increasing Fe content (Fig. 4). For example, Fig. 4c shows a family of 213 diffraction peaks in the vicinity of $2\theta = 30^{\circ}$. With increasing Fe content, the single peak at 30.1° 214 for endmember natroalunite shifts to lower diffraction angles and splits into two peaks for Fe#> 215 \sim 60. Similarly, the prominent peak at \sim 18° shifts to lower values with increasing Fe, and minor 216 peaks at $\sim 15^{\circ}$ and $\sim 16^{\circ}$ become proportionally larger (Fig. 4b). For clarity, the diffractograms 217 shown in Fig. 4 have been adjusted to comparable vertical scales, but the overall amplitude of the 218 peaks also decreased substantially with increasing Fe content for the same integration time. In 219 all cases, only peaks that can be attributed to natroalunite-natrojarosite solid solution are present 220 in the diffactograms, confirming the purity of the samples. The peaks are somewhat broad, 221 however, which may be attributable to minor compositionally variability of the crystals owing to 222 minor variation in Fe-Al substitution or to the relative proportion of hydronium and water in the 223 A and B sites.

224 Laser Raman spectroscopy

225 The Raman spectra for the synthetic natroalunite-natroiarosite solid solution series exhibit 226 two trends with increasing substitution of Fe for Al (Fig. 5). First, there is a gradual and 227 systematic change in overall Raman spectral pattern from a typical spectrum for pure 228 endmembers of the alunite subgroup to one characteristic of endmembers in the jarosite 229 subgroup. For example, the natroalunite endmember (Natro6) has a doublet OH peak in the 3600-3400 cm⁻¹ spectral range (Fig. 5a) and a prominent Raman peak near 1026 cm⁻¹ (assigned to 230 231 v_1 of SO₄ tetrahedra, symmetric stretching vibration mode) that has the highest peak intensity in the spectral range between 1250 and 100 cm⁻¹ (Fig. 5b,c). With increasing Fe, there is a 232 transition to a triplet OH peak, the peak at $\sim 1010 \text{ cm}^{-1}$ becomes less prominent, and additional 233 peaks appear in the 1180-1100 cm⁻¹ range (Fig. 5a,b). The latter set of peaks has been observed 234 for all members of the jarosite subgroup regardless of A site occupancy (K, Na, H₃O, NH₄, Ag, 235 236 or Pb; Sasaki et al. 1998; Frost et al. 2006a; Chio et al. 2009), in which the intensities of a group of Raman peaks near 1110 cm⁻¹ (assigned to v_3 of SO₄ tetrahedra, asymmetric stretching 237 vibration mode) are more or less equal to the intensity of v_1 mode of SO₄ near 1010 cm⁻¹. This 238 239 pattern is very different from the Raman spectra of natroalunite (Fig. 5b) and other common sulfates, such as MgSO₄•*x*H₂O (x = 0-11), CaSO₄•*x*H₂O (x = 0-2), and Fe²⁺SO₄•*x*H₂O (x = 0-7), 240 where the v_1 peak has much higher peak intensity than that of the v_3 peak (Chio et al. 2006; 241

Wang et al. 2009; Ling and Wang 2010). One possible explanation for this change in spectral pattern could be a combined effect of the smaller size of $[Fe^{3+}O_6]$ octahedron in natrojarosite (in which the Fe^{3+} -O bond lengths are between 1.98 and 2.07Å) than $[MgO_6]$, $[CaO_6]$, and $[Fe^{2+}O_6]$, and the higher degree of covalence of Fe^{3+} -O bonding than Al-O that affect the vibrations of neighboring $[SO_4]$. Second, the positions of many peaks in the Raman spectra for the solid solutions systematically shift to lower wavenumbers with increasing Fe content. This shift is

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251 Visible/near infrared reflectance

252 Visible/near infrared reflectance spectra for the synthetic natroalunite-natrojarosite solid 253 solution series exhibit absorptions due to iron electronic transitions and *M*-OH vibrations in the 254 wavelength range 350-2500 nm. These spectral features vary systematically in intensity and 255 position with Fe content (Fig. 6). There are four measured absorptions related to Fe electronic absorptions: a broad absorption centered at 930 nm (${}^{6}A_{1g} \rightarrow {}^{4}T_{1g}$), a broad shoulder near 600 nm 256 $({}^{6}A_{1g} \rightarrow {}^{4}T_{2g})$, a strong, broad absorption near 500 nm that is the wing of the charge transfer band 257 that is centered in the UV, and a sharp, narrow feature near 433 nm (${}^{6}A_{1g} \rightarrow {}^{4}A_{1g}$, ${}^{4}E_{g}$) (Rossman 258 259 1976; Sherman and Waite 1985; Bishop & Murad, 2005). The pure natroalunite endmember 260 lacks the two broad absorptions, but all intermediate compositions exhibit these features, albeit 261 with decreasing intensity as Fe content decreases. The narrow 433 nm absorption feature has 262 been considered a sensitive indicator for the presence of jarosite (Hunt and Ashley, 1979), and is 263 absent in the endmember natroalunite sample. Interestingly, this diagnostic feature is strongest 264 in low Fe intermediates but weaker in higher Fe compositions, including the endmember 265 natrojarosite, likely because of masking by the presence of a stronger absorption with a band 266 center in the ultraviolet (UV) (Fig. 6b). As Fe content decreases, this UV band decreases in 267 strength, resulting in an apparent increase in intensity of the 433 nm absorption. Since all of the 268 synthetic samples had essentially the same grain size distribution, it is unlikely that grain size 269 plays a role in the decreasing depth of this feature for higher Fe contents.

270 The synthetic natroalunite-natrojarosite series has three sets of vibrational absorptions at 271 1400-1500 nm, 1700-1900 nm, and 2100-2300 nm, which are overtones and combinations of fundamental vibrations (near 1000 cm⁻¹), related to OH bending and stretching vibrations 272 273 (Bishop and Murad 2005). Natroalunite has a doublet with two absorptions of approximately 274 equal strength at 1440 and 1492 nm, shifting to 1448 and 1491 nm as Fe content increases to Fe# 275 = 25 (Fig. 6c). Natrojarosite has a doublet with a strong absorption at 1477 nm and a weaker 276 shoulder near 1547 nm. Intermediate compositions with Fe# = 25-92 display a single feature that 277 sharpens and increases in wavelength from 1452 to 1477 nm with increasing Fe content. The 278 shoulder at ~ 1547 nm is distinctly evident only for the samples Fe# = 84 and above. Near 1800 279 nm (Fig. 6d), natroalunite exhibits a minimum at 1760 nm, natrojarosite at 1850 nm, and all 280 samples of intermediate composition possess an absorption at 1803 nm that is strongest for 281 intermediate iron contents. Similarly, natroalunite has a prominent absorption at 2172 nm, 282 natrojarosite at 2268 nm, and all intermediate compositions have an absorption near 2218 nm 283 that is most prominent for compositions near the middle of the Fe-Al range.

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4 Mid-infrared attenuated total reflectance

Attenuated total reflectance spectra for the synthetic solid solutions are shown in Fig. 7. ATR spectra are similar to transmission spectra in that peaks are mostly dependent on the absorption coefficient, i.e., the imaginary component of the complex index of refraction, but with slightly shifted band centers and intensities. The spectra exhibit a strong v_{OH} band center located at 3452 cm⁻¹ for natroalunite and at 3351 cm⁻¹ for natrojarosite (Fig. 7a), with intermediate compositions displaying a relatively systematic shift to smaller wavenumbers for higher Fe contents (Fig. 7e). The peak broadens considerably at intermediate Fe contents.

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Several strong absorptions are present, centered between 1000 and 1250 cm⁻¹ (peak attributions from Bishop and Murad 2005 and Lane 2007). A doublet attributable to $v_3(SO_4^{2-})$ stretching vibrations is observed with one center shifting systematically from 1215 to 1174 cm⁻¹ as Fe content increases, while the other center occupies two discrete positions characteristic of each endmember (Fig. 7b,c). A doublet attributed to $\delta(OH)$ is also present, and its center shifts systematically from 1028 to 1007 cm⁻¹ (Fig. 7d) and from 1066 to 1060 cm⁻¹ with increasing Fe content, though the latter peak is not apparent in the endmember natrojarosite.

Strong absorptions are also present in the 400-650 cm⁻¹ range due to a number of vibrations 299 in the mineral structure. A doublet of $v_4(SO_4^{2-})$ vibrations is observed near 617 and 658 cm⁻¹ for 300 endmember natroalunite and 627 cm⁻¹ and 677 cm⁻¹ for endmember natrojarosite. Intermediate 301 302 compositions have variable center positions, though these do not appear to vary systematically (Fig. 7a). Natroalunite has an out-of-plane bend γ (OH⁻) at 595cm⁻¹, absent in natrojarosite and 303 more Fe-rich samples. A $v_2(SO_4^{2-})$ bending vibration is at 430 cm⁻¹ in natroalunite, 419 cm⁻¹ in 304 intermediate compositions, and 442 cm⁻¹ in natrojarosite. Lastly, M-O (Al, Fe) band centers are 305 at 486 and 511 cm⁻¹ in natroalunite, and 471 and 503 cm⁻¹ in natrojarosite. In most intermediate 306 samples, one band center shifts from 507 to 498 cm⁻¹ and the lower wavenumber center is not 307 308 discernible.

309 Mid-infrared reflectance

Unlike ATR spectra in the mid-infrared, MIR diffuse reflectance spectra depend strongly on both the real and imaginary components of the index of refraction, and are thus more complicated than ATR spectra (Fig. 8; Lane 2007). For wavelengths shortward of ~5-6 μ m, the real component of the complex index of refraction (*n*) is approximately constant while the imaginary component of the index of refraction (*k*), proportional to absorption, varies, causing

 d_k both vary and both reflectance minim

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315 reflectivity minima. At longer wavelengths, *n* and *k* both vary and both reflectance minima and 316 maxima are significant in mineral identification. Thermal emission spectra from orbiters (TES & 317 THEMIS) and rover-based instruments (MiniTES) utilize this wavelength range, and in emission 318 data, reflectivity maxima are emissivity minima.

319 MIR reflectance spectra for the synthetic solid solution series are shown in Figure 8. As in ATR spectra, some of the strongest features occur at wavenumbers near 1000 cm⁻¹ (Fig. 8b) 320 321 (Lane 2007). These features vary with changes in the mass of the metal cation and effects on the bonding environment (Sasaki et al. 1998). The $v_3(SO_4^{2-})$ stretching vibration has reflectance 322 maxima that decrease systematically from 1255 cm⁻¹ in natroalunite to 1205 cm⁻¹ in natrojarosite, 323 but the second v_3 mode varies only slightly and non-systematically between 1112 and 1118 cm⁻¹. 324 The $\delta(OH)$ vibration is strong in natrojarosite at 1028 cm⁻¹, but is not as prominent in 325 natroalunite or intermediates. The $v_4(SO_4^{2-})$ maxima in natroalunite are prominent with a doublet 326 at 696 and 628 cm⁻¹. The latter feature is shifted to 633 cm⁻¹ in Fe-rich samples with the second 327 328 maximum of the doublet absent in these samples. Finally, Al/Fe-O features are strong in natrojarosite at 478 and 526 cm⁻¹ but are weak to absent in natroalunite (possible feature near 329 520 cm⁻¹). The 598 cm⁻¹ feature in natroalunite is due to an out-of-plane OH bend (Lane 2007). 330

A number of additional features occur in the MIR spectra in the 1500-4000 cm⁻¹ range (Fig. 8a). The v(OH) shifts systematically from ~3500 cm⁻¹ for natroalunite ~3400 cm⁻¹ for natrojaosite. Multiple vibrational overtones of v₃ and δ (OH⁻) occur in the 1900-2400 cm⁻¹ wavenumber range (Bishop and Murad 2005). Samples with iron have a peak near 2500 cm⁻¹ and the position of features comprised of the bands decreases in wavenumber with increasing Fe content.

337 Mössbauer spectroscopy

338 Two examples of the Mössbauer spectra obtained for synthetic natroalunite-natrojarosite 339 solid solutions are shown in Fig. 9, together with a library spectrum for endmember jarosite. 340 The synthetic samples exhibit a doublet in the MB spectra whose peak positions closely 341 correspond to those of jarosite. Values of isomer shift (IS) and quadrapole splitting (QS) derived 342 from the MB spectra of these and several other synthetic solid solutions are listed in Table 4 and 343 displayed in Figure 10. For the most part, the synthetic solid solutions have MB parameters 344 within the range of reported values for natural and synthetic jarosites, although the most Fe-345 enriched samples have combined IS and QS values that fall slightly outside this range. Notably, 346 the natroalunites with relatively low Fe contents (Fe# = 10-25) have IS and QS values very 347 similar to those reported for many jarosites. In addition, the synthetic solid solutions display 348 decreasing values of QS with increasing Fe content, while IS shows no clear trend with Fe# (Fig. 349 11).

350 Analyses of natural deposits

351 In their natural setting, the Fe-bearing natroalunites found on Nicaraguan volcanoes occur 352 both as nearly monomineralic deposits and as relatively minor phases in mixtures with other 353 secondary minerals, most commonly amorphous silica, gypsum, and Fe-oxides/oxyhydroxides 354 including hematite (McCollom et al. 2013b; Hynek et al. 2013). In an initial effort to evaluate 355 the extent to which Fe-bearing natroalunite might be detectable in natural fumarolic settings, the 356 spectral characteristics of several samples of acid-sulfate altered deposits from Nicaraguan 357 volcanoes that were known to contain Fe-bearing natroalunite were measured and compared with 358 results from the synthetic solid solutions (see references above for more complete descriptions of 359 the samples and geological context). The mineral composition of the samples is summarized in 360 Table 5, and XRD analyses of the bulk samples are provided as a Supplemental Figure 1. The

natural samples included deposits composed predominantly (>90 %) of Fe-bearing natroalunite
(CN R36, Telica 7, and Masaya 3), and deposits where Fe-bearing natroalunite occurs as a
relatively minor component in combination with other secondary phases (CN R21, CN R28,
Telica 6). The relative proportions of Fe and Al in the natroalunites in the samples as determined
by EDS and EMPA are listed in Table 2. The natroalunites in the samples range from those
containing relatively small amounts of Fe (Telica 6, CN R28) to Fe-rich compositions (Telica 7,
Masaya 3).

368 Partial X-ray diffraction patterns for the natural samples in the $2\theta = 28-31^{\circ}$ range are shown 369 in Figure 12 (the full diffraction patterns are provided as Supplemental Figure 1). All of the 370 diffraction peaks in this range for the natural samples are shifted to lower 2θ values than that of 371 the endmember natroalunite, consistent with Fe-for-Al substitution in the mineral structure (note 372 that sample CN R21 has a peak for augite at $\sim 29.8^{\circ}$ that obscures the natroalunite peak in this 373 region). The peak positions, however, do not align precisely with the positions expected for the 374 measured Fe contents of the minerals, nor do they all vary systematically with measured Fe#. 375 Potential explanations for these discrepancies include inaccurate measurements of bulk 376 composition by EDS and EMPA (which were measured for only a limited number of crystals 377 that may not be representative of the bulk sample), or variations in the unit cell dimensions of the 378 natural samples relative to the synthetic solid solutions caused by additional substitutions in the 379 A and B sites. For instance, varying degrees of deficiency in occupancy of the B site by Fe and 380 Al can result in slight shifts in peak positions (e.g., Grey et al. 2013), and the natural minerals 381 may have a different extent of deficiency at this position than the synthetic minerals. In any 382 case, the XRD measurements exhibit shifts from the expected peak position for endmember 383 natroalunite, and are shifted in a direction that would be indicative of Fe enrichment. The

384 presence of singlet peaks for the natural samples also indicates that these are true solid solutions 385 rather than intimate, micron-scale mixtures of two separate phases like those described for some 386 occurrences of jarosite and natrojarosite (Desborough et al. 2009).

387 Raman spectra for natroalunite in two samples from Cerro Negro volcano are shown in 388 Figure 13 (CN R21, CN R36). In both cases, the spectra were obtained from analysis of single 389 crystals of fine-grained, pseudocubic natroalunite located with the aid of a visible-light 390 microscope; analyses of multiple crystals from the same sample yielded essentially identical 391 spectra. The Raman spectra for both of the Cerro Negro samples show peak shifts to lower 392 values consistent with Fe-bearing natroalunite with Fe# ≈ 10 (e.g., Fig. 13b,c). This value is 393 slightly lower than the average Fe contents for these samples determined by EDS/EMPA, but 394 within the range of measured values (Table 2). Also shown in the Figure 13 is a Raman 395 spectrum measured for natroalunite produced in a laboratory simulation of acid-sulfate alteration 396 of Cerro Negro basalt that has a substantially higher Fe content (ADSU6). The Raman peaks for 397 this sample are shifted to even lower values than for the natural samples, reflecting the higher Fe 398 content. In this case, the Raman peak positions are consistent with the measured amount of Fe in 399 the sample.

VNIR spectra for several natural samples are shown in Figure 14, which includes spectra obtained both in the laboratory and in the field. The top three spectra in Figure 14a are samples composed predominantly of Fe-bearing natroalunite. These spectra all show characteristics indicative of natroalunite with minor amounts of Fe, including overall spectral shape, a doublet at ~2172 and ~2218 nm with subequal absorption depth, a doublet in the 1448-1491 nm range, a broad absorption at ~930 nm, and decreasing reflectance for wavelengths <700 nm. However, only Masaya 3 exhibits a clear absorption band at 1803 nm, a feature observed in the 407 intermediate synthetic solid solutions. Curiously, Masaya 3 is also the only one of these samples 408 that displays the prominent absorption at ~433 nm seen in synthetic samples, although this 409 feature would have been expected for all three samples based on measured Fe content (Fig. 14b). 410 It is possible that this feature is suppressed in the other samples by absorption bands for other Fe-411 oxide/oxyhydroxide phases present in small amounts, which may also affect the overall shape of 412 the spectra in the <700 nm region. The relative depths of the doublets at $\sim 2172/2218$ nm for 413 these samples indicate Fe contents in the Fe# \approx 15-30 range (Fig. 6e), consistent with 414 compositions indicated by EDS and EMPA analyses.

415 The middle group of spectra in Figure 14 shows samples where the natroalunite is present in 416 combination with other phases, and illustrate some of the potential complications that may arise 417 in evaluating the composition of natroalunites that occur in mixtures. Because the spatial 418 distribution of minerals in these samples is highly heterogeneous, it was not possible to 419 quantitatively estimate the proportions of phases present in the field of view during spectroscopic 420 measurements. Natroalunite is a major component of sample Telica 6, but the deposit also 421 contains kaolinite as well as minor amounts of amorphous silica and gypsum (Hynek et al. 422 2013). The VNIR spectrum for this sample exhibits spectral features in the 350-1000 nm region 423 consistent with Fe-bearing natroalunite, including a small absorption feature at 433 nm, a broad 424 trough at ~ 900 nm, and the overall shape of the < 700 nm region. However, the natroalunite 425 doublet at ~2172/2218 nm is obscured by a strong absorption band from kaolinite centered at 426 \sim 2205 nm, and absorptions in other regions of the spectrum reflect substantial contributions from 427 amorphous silica and gypsum.

428 Natroalunite is a relatively minor phase in samples CN R21 and CN R28, where the 429 secondary mineralogy is dominated by amorphous silica and gypsum (McCollom et al. 2013b).

430 The VNIR spectra for these samples are strongly influenced by contributions from these other 431 hydrated phases. Nevertheless, sample CN R21 shows several features indicative of Fe-bearing 432 natroalunite, including a doublet at 2190 and 2220 nm with subequal absorption depth, doublet 433 peaks at ~1438 and ~1490 nm, a feature near 2320 nm, an absorption band at 433 nm, and a 434 broad absorption at ~ 900 nm. The overall shape of the spectra at wavelengths below 1000 nm. 435 however, appears to show some influence of the hematite that occurs in small amounts in this 436 sample (Fig. 2). Sample CN R28 is largely composed of micron-sized natroalunite crystals 437 distributed across the surface of large prismatic gypsum crystals. The spectra indicate a 438 substantial contribution from gypsum, but the shape of the spectra in the 2140-2290 nm and 439 <1000 nm region has some attributes of natroalunite-natrojarosite except for the absence of the 440 433 nm feature. However, the shape of these features is most similar to solid solutions with Fe# 441 \approx 50, which is much higher than the Fe contents indicated by EDS and XRD analysis (Table 2; 442 Fig. 12). It is likely that the spectral features in the region are influenced by the presence of 443 gypsum and other minerals in the sample, obscuring the contribution from natroalunite.

444 The MIR spectra of three natural samples are shown in Figure 15. The spectra exhibit 445 several SO₄ and OH vibration features indicative of the presence of sulfate and hydroxylated 446 minerals. However, features that are uniquely indicative of the presence of natroalunite or its Fe 447 contents are not obvious in all samples. The presence of small reflectance peaks at ~ 1255 cm⁻¹ and the absence of a peak at 1028 cm⁻¹ are both consistent with natroalunite with Fe# $<\sim$ 25. The 448 broad peak at ~2500 cm⁻¹ is consistent with the presence of an Fe-bearing member of the 449 450 natroalunite-natrojarosite series, but provides no specific information about the Fe content. Potentially diagnostic features in the 400-700 cm⁻¹ are only weakly present in sample CN R36, 451 452 and not evident in the other samples. Specific assignment of any of these features to natroalunite

would require linear unmixing models to evaluate potential contributions from other phases

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454 present in the samples.

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455 A Mössbauer spectrum for one of the natural samples (CN R21) is shown in Figure 9. 456 Deconvolution of this and other spectra for altered basalt cinders from Cerro Negro identified an 457 alunite group component in addition to several other primary and secondary Fe-bearing phases, 458 and similar results were found for examples of altered basalt cinders from laboratory 459 experiments (McCollom et al. 2013a,b). Mössbauer parameters derived for the Fe-bearing 460 natroalunite component of three natural samples from Cerro Negro as well as several products of 461 laboratory simulations of acid-sulfate alteration are shown in Figures 10 and 11. The natural and 462 experimental samples exhibit a considerable range of IS and QS values. Values of IS for the 463 natural and experimental samples are generally slightly higher than those determined for the 464 synthetic solid solution. The experimental samples and one of the Cerro Negro samples have QS 465 values similar to that of the synthetic solid solutions with relatively low Fe contents (1.15-1.25 466 mm/s), but two of the CN samples have substantially lower QS values (0.99-1.06 mm/s) more 467 similar to solid solutions with high Fe contents (1.05-1.11 mm/s). In one of these samples (CN R21), the average Fe content for natroalunite is Fe# = 17, while the other appears to have a 468 469 bimodal distribution of minerals in the alunite group ranging from Fe-rich natroalunite (Fe# \approx 470 20-30) to Al-rich natrojarosite (Fe# \approx 60-50) (see McCollom et al. 2013b). For both the natural 471 and experimental samples, values of IS and QS fall within the range of previously reported 472 values for natural and synthetic members of the jarosite subgroup (Fig. 10; Table 4).

473

DISCUSSION

474 Determination of relative Fe-Al contents of alunite-jarosite group minerals

475 All of the methods used to investigate natroalunite-natrojarosite solid solutions exhibited 476 systematic variations in spectral patterns and peak positions with Fe content that could 477 potentially be used to infer the composition of natural minerals in this group, either on Earth or 478 in other planetary contexts. Of course, development of a comprehensive calibration to infer the 479 compositions of solid solutions among these minerals using spectral or XRD methods will 480 require further investigation of the impact of additional compositional factors (e.g., Swayze et al. 481 2008). For instance, solid solution mixing of Na, K, and H_3O^+ on the A site is common in minerals from this group, and many natural minerals in the alunite and jarosite subgroups that 482 483 precipitate in low temperature environments also have deficiencies in occupancy of the B site 484 relative to the ideal molecular formula (Stoffregen et al. 2000; Grey et al. 2013). Inclusion of 485 additional H₂O into the crystal structure to compensate for these deficiencies can result in 486 lowering the symmetry of the minerals from rhombohedral to monoclinic (Scarlett et al. 2010; 487 Grev et al. 2011, 2013). Further work will be required to determine how these and other factors 488 impact the spectral and XRD characteristics of minerals from this group. Nevertheless, at least 489 for the case of laser Raman spectroscopy, previous studies (e.g., Murphy et al. 2009) have 490 indicated that substitutions in the A site result in only small shifts in peak position relative to 491 those observed here for Fe-Al substitutions in the B site, which appear to have a much larger 492 influence on the vibrations of [SO4] and [OH] groups. In addition, investigation of the variation 493 of peak positions on ambient environmental conditions such as temperature may be required to 494 employ these methods on Mars or other planetary surfaces (e.g., Mills et al. in review).

495 As demonstrated by the analysis of natural samples, a potential limitation on the use of these 496 methods to identify and evaluate the composition of minerals in the alunite and jarosite 497 subgroups are interferences from other phases that are likely to be present in many acid-sulfate

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498 deposits. These additional phases can inhibit both the detection of minerals from this group and 499 the capacity to infer their composition. For instance, kaolinite and other Al-phyllosilicates have 500 absorption bands in the ~ 2200 nm region of the VNIR spectra that overlap with the diagnostic 501 features of natroalunite/natrojarosite in this region (Fig. 14c), and gypsum may interfere with 502 detection of absorbance bands in the 1400-1500 nm and 1750-1850 nm regions. Similarly, 503 hematite and other Fe-oxides/oxyhydroxides have absorption bands in the VNIR spectral region 504 below 1000 nm that can obscure diagnostic features for Fe-bearing minerals in the alunite group 505 and their compositions (Fig. 14b).

506 Conversely, the overlap in spectral bands with other minerals presents the possibility for 507 misidentification of phases if spectra for solid solutions in the alunite group are not considered. 508 For example, the increasing band depth at ~ 2218 nm in VNIR spectrum with increasing Fe 509 content observed for natroalunite-natrojarosite solid solutions could lead to spectral properties 510 that could be mistaken for kaolinite or other Al-phyllosilicates. It is worth noting, however, that 511 there is a slight difference between the position of the absorption feature for intermediate 512 natroalunite-natrojarosite compositions at ~ 2218 nm and that of kaolinite at ~ 2205 nm. Telica 6, 513 which is the only natural sample shown in Figure 14 where kaolinite was identified during XRD 514 and SEM/EDS analysis, is the only sample with a feature centered at ~2205 nm, while the other 515 natural samples have features in this region shifted to slightly higher wavelengths consistent with 516 Fe-bearing natroalunite. Potentially, this shift could be used to differentiate between kaolinite 517 and Fe-bearing natroalunite during spectral analysis, provided that the position of the absorbance 518 feature at ~2218 nm does not shift significantly with other compositional variables (such as 519 substitution of H_3O^+ for Na in the A site). In addition, the sharp 433 nm absorption band of 520 VNIR spectra is often considered to be diagnostic for the presence of jarosite, but in the solid

521 solution spectra this band was observed in Fe-bearing natroalunites and decreased in magnitude 522 with increasing Fe content. Consequently, detection of this band could be consistent with any 523 Fe-bearing member of the alunite group, not just those in the jarosite subgroup. Together, these 524 observations indicate that it would be very worthwhile to include solid solutions from the alunite

525 mineral group in reference libraries for interpretation of spectral data from Mars and elsewhere.

526 Examination of the natural samples by VNIR indicates that identification of natroalunite-527 natrojarosite solid solutions and estimation of their Fe-content should be relatively 528 straightforward in samples where they dominate the mineralogy in the field of view. Significant 529 complications may arise, however, when they occur in mixtures with other minerals including 530 Fe-oxides/oxyhydroxides, phyllosilicates, other hydrated sulfate minerals, and hydrated forms of silica. Since these are precisely the types of minerals that would be expected to co-occur with 531 532 natroalunite/natrojarosite (and other minerals in the alunite/jarosite subgroups) in natural acid-533 sulfate environments, a key objective for future work involving solid solutions will be to 534 evaluate the extent to which they may be masked by accessory phases. In addition, a better 535 understanding of how the presence of other Fe-bearing phases leads to suppression of the 433 nm 536 feature would aid interpretation of spectra in the <1000 nm region.

Laser Raman spectroscopy appears to have a high potential to provide compositional information for solid solutions in the alunite and jarosite subgroups, particularly during in situ Raman measurements where the spectrum is obtained from a single mineral phase excited by a well-focused laser beam (10-20 µm diameter) during planetary surface exploration. Complications may arise in cases where mineral grain size is less than the spot size of the laser beam, allowing multiple phases to be present in the field of view. However, this problem may be less severe than in the cases of VNIR and MIR because of the sharpness of Raman spectral peaks and less extensive overlap for different phases, especially because of the very different peak positions (on the order of a few hundreds in wavenumber) for sulfates, silicates, and Feoxides/oxyhydroxides.

547 The Mössbauer results do not show any consistent variation in either QS or IS with 548 increasing Fe content. Values of these parameters are shown in Figure 11 for the samples 549 discussed here as well as a few additional samples from the literature where compositional data 550 While our synthetic solid solutions display decreasing values of OS with are available. 551 increasing Fe content, other samples do not. Morris et al. (2006) suggested that substitution of 552 Al for Fe in members of the jarosite family would increase quadrapole splitting, evidently 553 because two Al-bearing jarosites included in their study had QS values near the high end of the 554 range reported in the literature (Fig. 11). However, other samples with higher Al contents do not 555 have QS values higher than the Al-bearing jarosites.

556 As a consequence, it appears that Mössbauer spectroscopy may not be a particularly useful 557 method for determining the Fe content of minerals in the alunite-jarosite group. To the contrary, 558 the apparent insensitivity of MB spectroscopy to Fe contents suggests that it is not possible to 559 differentiate Fe-bearing natroalunite from members of the jarosite family using this method, even 560 when only a few mole% Fe is present in the B site. Indeed, a doublet in a MB spectrum with QS 561 \approx 1.05-1.25 and IS \approx 0.36-0.40 may be sufficient only to indicate the presence of an Fe-bearing 562 mineral in the alunite group, without providing more specific information about the identity or 563 composition of the phase.

564 Implications for sulfate deposits on Mars

565 Minerals from the alunite group have been inferred to be present at several locations across 566 Mars based on interpretation of data from orbital remote sensing and surface exploration.

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567 Members of the jarosite subgroup have been reported to be present in the layered sulfate deposits 568 at the Opportunity rover landing site in Meridiani Planum based on results from Mössbauer 569 spectroscopy (Klingelhöfer et al. 2004), and may be present at the Gusev Crater landing site as 570 well (Morris et al. 2008). The use of high spatial and spectral resolution instruments from Mars 571 orbit has led to additional reported detections of alunite (Swavze et al. 2008a: Wrav et al. 2011) 572 and jarosite (Milliken et al. 2008; Farrand et al. 2009; Wray et al. 2011; Ehlmann and Mustard, 573 2012; Thollot et al. 2012) from multiple additional locations on the Martian surface, in regionally 574 significant deposits. For example, Farrand et al. (2009) noted occurrences of jarosite near Al-575 phyllosilicates in Mawrth Vallis and this assemblage possibly reflects early hydrothermal 576 alteration of Mars' crust. Elsewhere, alunite and kaolinite have been reported in putative acid-577 saline paleolake deposits within craters in Terra Sirenum (Swayze et al. 2008b; Wray et al. 578 2011). All of these interpretations of orbital data, however, are based on endmember spectra, and 579 consideration of spectra for intermediate solid solutions may lead to additional identifications as 580 well as providing constraints on the chemical composition of phases in this mineral group where 581 they have already been identified.

582 Several of the spectroscopic methods investigated in this study have the potential to be 583 useful to identify intermediate Fe-Al solid solutions from the alunite group on Mars, and to infer 584 their chemical composition. The reflectance VNIR spectra obtained in this study are relevant to 585 OMEGA and CRISM onboard the Mars Express and MRO orbiters. The systematic changes in 586 XRD patterns with varying Fe-Al content should provide also useful information for the CheMin 587 instrument onboard the Curiosity rover if members of the alunite group are encountered in Gale 588 Crater. Flight models of a laser Raman spectrometer are under development (Wang et al. 2003), 589 and one of them was selected for the ExoMars rover (a mission run by the European Space Agency scheduled for launch in 2018). Future research should allow development of calibrationcurves for evaluation for solid solution compositions using all of these methods.

592 There are several outstanding problems in Mars exploration where consideration of 593 intermediate compositions may be relevant to regional studies. For example, multi-component 594 fits of thermal emission spectra (MiniTES) from the sulfate-bearing bedrock at Meridiani 595 Planum were not significantly improved by inclusion of endmember jarosite at the level inferred 596 from Mössabauer results (10%) (Glotch et al. 2006). However, MB spectra of solid solutions 597 indicate that the component interpreted as jarosite could be equally well represented by Fe-598 bearing natroalunite or other members of the natroalunite subgroup (Fig. 10), suggesting that 599 models of Meridiani mineralogy should be expanded to consider solid solutions encompassing 600 the entire alunite group. The synthetic minerals with compositions intermediate between 601 natroalunite and natrojarosite exhibit distinctive spectral characteristics at MiniTES wavelengths (1600-200 cm⁻¹) (Fig. 8), and including solid solutions in the reference library could potentially 602 603 improve fits to the observed MiniTES spectra.

604 As a second example, a "doublet material" with absorptions at 2205–2218 nm and 2265– 605 2278 nm has been identified in CRISM data from several areas but has not vet been associated 606 with a particular mineral (Roach et al. 2010; Weitz et al. 2012; Thollot et al. 2012). The position 607 of these features correspond closely to those of the synthetic natroalunite-natroajarosite solid 608 solutions containing relatively high Fe, suggesting that they could be consistent with the 609 presence of a member of the alunite group with an intermediate composition of Fe# ≈ 65 . The 610 "doublet material" has additional absorptions at 1400-1420 nm and 1910-1920 nm that are not 611 observed in spectra for the pure natroalunite-narojarosite series, so further work will be required 612 to determine whether mixtures with additional phases such as hydrated silica, gypsum, and

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613 hematite can reproduce the observed spectra. In this respect, it may be worth noting that in most 614 instances the natroalunite at Cerro Negro is found in close association with amorphous silica and 615 gypsum, and samples of these deposits exhibit features in the 1400 and 1910 nm regions (e.g., 616 CN R21 and CN R28; Fig. 14). In addition, the presence of Al-phyllosilicates has been inferred 617 in some locations based on absorptions at 2170 and 2210 nm that are characteristic of the 618 kaolinite doublet, but with the relative intensities of these features reversed from what is 619 expected for kaolinite (Fig. 14c). As an alternative, these spectral features could potentially be 620 explained by Fe-bearing natroalunite. Whether natroalunite-natrojarosite solid solutions can 621 account for these and other features will require detailed comparisons of martian spectra with 622 spectra obtained for synthetic minerals, but these possibilities raise the prospect that alunite 623 group minerals with intermediate Fe compositions could be widespread on the surface of Mars. 624 625 ACKNOWLEDGEMENTS 626 This research was supported by funds from NASA Mars Fundamental Research Program grants 627 NNX12AI02G (TMM) and NNX10AM89G (AW), NASA Exobiology Award #NNX08AQ11G

Rossman for assistance and use of his FTIR spectrometer and to Frieder Klein for help with the Raman analysis of natural samples. The authors are grateful for reviews by Gregg Swayze and Stuart Mills, whose thoughtful comments helped to improve the manuscript.

(TMM and BMH), and NASA Early Career Award #NNX12AF20G (BMH). Thanks to George

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633 References

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- Bishop, J.L., and Murad, E. (2005) The visible and infrared spectral properties of jarosite and
- alunite. American Mineralogist, 90, 1100-1107.
- Blake, D., Vaniman, D., Achilles, C., Anderson, R., Bish, D., Bristow, T., Chen, C., Chipera, C.,
- 639 Crisp, J., Des Marais, D., Downs, R.T., Farmer, J., Feldman, S., Fonda, M., Gailhanou, M.,
- Ma, H., Ming, D.W., Morris, R.V., Sarrazin, P., Stolper, E., Treiman, A., and Yen, A. (2012)
- 641 Characterization and calibration of the CheMin mineralogical instrument on Mars Science
- 642 Laboratory. Space Science Reviews, 170, 341-399.
- 643 Brand, R.A. (1987) Improving the validity of hyperfine field distributions from metallic alloys.
- Part I: Unpolarized source. Nuclear Instruments and Methods in Physics Research B, 28,
 398-405.
- 646 Brophy, G.P., Scott, E.S., and Snellgrove, R.A. (1962) Sulfate studies II: Solid solution between
- 647 alunite and jarosite. American Mineralogist, 47, 112-126.
- 648 Chio, C.H., Sharma, S.K., Ming, L., and Muenow, D.W. (2005) Micro-Raman studies of
- hydrous ferrous sulfates and jarosites. Spectrochimica Acta Part A, 61, 2428–2433.
- 650 Chio, C.H., Sharma, S.K., Ming, L., and Muenow, D.W. (2009) Raman spectroscopic
- 651 investigation on Jarosite-Yavapaiite stability. Spectrochimica Acta Part A, 75, 162-71.
- 652 Desborough, G.A., Smith, K.S., Lowers, H.A., Swayze, G.A., Hammarstrom, J.M., Diehl, S.F.,
- 653 Leinz, R.W., and Driscoll, R.L. (2010) Mineralogical and chemical characteristics of some
- natural jarosites. Geochimica et Cosmochimica Acta 74, 1041-1056.

- 655 Dutrizac, J.E., and Jambor, J.L. (2000) Jarosites and their application in hydrometallurgy. In C.
- 656 N. Alpers, J. L. Jambor, and D. K. Nordstrom, Eds. Sulfate Minerals: Crystallography,
- 657 Geochemistry, and Environmental Significance. Reviews in Mineralogy and Geochemistry,
- 658 Mineralogical Society Of America, Washington, D. C., pp. 405-452.
- Ehlmann, B.L., and Mustard, J.F. (2012) An in-situ record of major environmental transitions on
- 660 early Mars at Northeast Syrtis Major. Geophysical Research Letters, 39, L11202,
- 661 doi:10.1029/2012GL051594.
- 662 Elwood Madden, M.E., Madden, A.S., and Rimstidt, J.D. (2009) How long was Meridiani
- 663 Planum wet? Applying a jarosite stopwatch to determine the duration of aqueous diagenesis.
- 664 Geology, 37, 635-638.
- Elwood Madden, A.S., Rimstidt, J.D., Zahrai, S., Kendall, M.R., and Miller, M.A. (2012)
- Jarosite dissolution rates and nanoscale mineralogy. Geochimica Cosmochimica Acta, 91,306-312.
- 668 Farrand, W.H., Glotch, T.D., Rice, Jr., J.W., Hurowitz, J.A., and Swayze, G.A. (2009) Discovery
- of jarosite within the Mawrth Vallis region of Mars: Implications for the geologic history of
- 670 the region. Icarus, 204, 478–488.
- 671 Frost, R., Wills, R., Weier, M.L., Martens, W., and Mills, S. (2006a) A Raman spectroscopic
- 672 study of selected natural jarosites. Spectrochimica Acta Part A, 63, 1-8.
- 673 Frost, R., Wills, R., Weier, M.L., Martens, W., and Kloprogge, J.T. (2006b) A Raman
- 674 spectroscopic study of alunites. Journal of Molecular Structure, 785, 123-132.
- Glotch, T.D., Bandfield, J.L., Christensen, P.R., Calvin, W.M., McLennan, S.M., Clark, B.C.,
- 676 Rogers, A.D., and Squyres, S.W. (2006) Mineraology of the light-toned outcrop at
- 677 Meridiani Planum as seen by the Miniature Thermal Emission Spectrometer and

- 678 implications for its formation. Journal of Geophysical Research, 111, E12S03, doi:10.1029/
- 679 2005JE002672.
- 680 Golden, D.C., Ming, D.W., Morris, R.V., and Graff, T.G. (2008) Hydrothermal synthesis of
- 681 hematite spherules and jarosite: Implications for diagenesis and hematite spherule formation
- in sulfate outcrops at Meridiani Planum, Mars. American Mineralogist, 93, 1201–1214.
- 683 Grey, I.E., Scarlett, N.V.Y., Bordet, P., and Brand, H.E.A. (2011) Jarosite-butlerite intergrowths
- 684 in non-stoichiometric jarosites: crystal chemistry of monoclinic natrojarosite-
- hydroniumjarosite phases. Mineralogical Magazine, 75, 2775-2791.
- 686 Grey, I.E., Scarlett, N.V.Y., and Brand, H.E.A. (2013) Crystal chemistry and formation
- 687 mechanism of non-stoichiometric monoclinic K-jarosites. Mineralogical Magazine, 77, 249688 268.
- Härtig, C., Brand, P., and Bohmhammel, K. (1984) Fe-Al-Isomorphie und Strukturwasser in
- Kristallen vom Jarostie-Alunit-Typ, Zeitschrift fur Anorganische und Allgemeine Chemie,508, 159-164.
- Hrynkiewicz, A.Z., Kubisz, J., and Kulgawczuk, D.S. (1965) Quadrupole splitting of the 14.4
- gamma line of 57Fe in iron sulfates of the jarosite group. Journal of Inorganic and Nuclear
- 694 Chemistry, 27, 2513–2517.
- Hunt, G.R. and Ashley, R.P. (1979) Spectra of altered rocks in the visible and near infrared.
- 696 Economic Geology, 74, 1613–1629.
- 697 Hynek, B.M., McCollom, T.M., Marcucci, E.C., Brugman, K. and Rogers, K.L. (2013)
- 698 Assessment of environmental controls on acid-sulfate alteration at active volcanoes in
- 699 Nicaragua: Applications to relic hydrothermal systems on Mars. Journal of Geophysical
- 700 Research Planets, 118, doi:10.1002/jgre.20140.

- 701 Johnson, J.H. (1977) Jarosite and akaganeite from White Island volcano, New Zealand: An X-
- ray and Mossbauer study. Geochimica Cosmochimica Acta, 41, 539–544.
- 703 Klingelhöfer, G., Morris, R.V., Bernhardt, B., Schröder, C., Rodionov, D.S., de Souza, P.A.,
- 704 Yen, A., Gellert, R., Evlanov, E.N., Zubkov, B., Foh, J., Bonnes, U., Kankeleit, E.,
- Gu tlich, P., Ming, D.W., Renz, F., Wdowiak, T., Squyres, S.W., and Arvidson, R.E.
- 706 (2004) Jarosite and hematite at Meridiani Planum from Opportunity's Mössbauer
- 707 spectrometer. Science, 306, 1740–1745.
- Lane, M.D. (2007) Mid-infrared emission spectroscopy of sulfate and sulfate-bearing minerals.
- American Mineralogist, 92, 1–18.
- 710 Ling, Z.C., and Wang, A. (2010) A systematic spectroscopic study of eight hydrous ferric
- sulfates relevant to Mars. Icarus, 209, 422-433.
- 712 Marcucci, E.C., Hynek, B.M., Kierein-Young, K.S. and Rogers, K.L. (2013) Visible-near
- 713 infrared reflectance spectroscopy of volcanic acid-sulfate alteration in Nicaragua: Analogs
- for early Mars. Journal of Geophysical Research Planets, 118, doi:10.1002/jgre.20159.
- 715 Maubec, N., Lahfid, A., Lerouge, C., Willie, G., and Michel, K. (2012) Characterization of
- alunite supergroup minerals by Raman spectroscopy. Spectrochimica Acta A, 96, 925-939.
- 717 McCollom, T.M., Robbins, M., Moskowitz, B., Berquó, T.S., Jöns, N., and Hynek, B.M. (2013a)
- Experimental study of acid-sulfate alteration of basalt and implications for sulfate deposits
- on Mars. Journal of Geophysical Research, 118, 1-38.
- 720 McCollom, T.M., Hynek, B.M., Rogers, K., Moskowitz, B., and Berquó, T.S. (2013b) Chemical
- and mineralogical trends during acid-sulfate alteration of pyroclastic basalt at Cerro Negro
- volcano and implications for early Mars. Journal of Geophysical Research, 118, 1719-1751.

- 723 Milliken, R.E., Swayze, G.A., Arvidson, R.E., Bishop, J.L., Clark, R.N., Ehlmann, B.L., Green,
- R.O., Grotzinger, J.P., Morris, R.V., Murchie, S.L., Mustard, J.F., and Weitz, C. (2008)
- 725 Opaline silica in young deposits on Mars. Geology, 36, 847–850.
- 726 Mills, S. J., Hatert, F., Nickel, E. H., and Ferraris, G. (2009) The standardization of mineral
- group hierarchies. European Journal of Mineralogy, 21, 1073-1080.
- 728 Morris, R.V., Klingelhöfer, G., Schröder, C., Rodionov, D.S., Yen, A., Ming, D.W., de Souza,
- Jr., P.A., Wdowiak, T., Fleischer, I., Gellert, R., Bernhardt, B., Bonnes, U., Cohen, B.A.,
- 730 Evlanov, E.N., Foh, J., Gu tlich, P., Kankeleit, E., McCoy, T., Mittlefehldt, D.W., Renz,
- 731 F., Schmidt, M.E., Zubkov, B., Squyres, S.W., and Arvidson, R.E. (2006) Mossbauer
- 732 mineralogy of rock, soil, and dust at Meridiani Planum, Mars: Opportunity's journey across
- sulfate-rich outcrop, basaltic sand and dust, and hematite lag deposits. Journal of
- 734 Geophysical Research, 111, doi:10.1029/2006JE002791.
- 735 Morris, R.V., Klingelhöfer, G., Schröder, C., Fleischer, I., Ming, D.W., Yen, A.S., Gellert, R.,
- 736 Arvidson, R.E., Rodionov, D.S., Crumpler, L.S., Clark, B.C., Cohen, B.A., McCoy, T.J.,
- 737 Mittlefehldt, D.W., Schmidt, M.E., de Souza, Jr., P.A., and Squyres, S.W. (2008) Iron
- mineralogy and aqueous alteration from Husband Hill through Home Plate at Gusev Crater,
- 739 Mars: Results from the Mössbauer instrument on the Spirit Mars Exploration Rover. Journal

740 of Geophysical Research, 113, E12S42, doi:10.1029/2008JE003201.

- 741 Murphy, P.J., Smith, A.M.L., Hudson-Edwards, K.A., Dubbin, W.E., and Wright, K. (2009)
- Raman and IR spectroscopic studies of alunite-supergroup compounds containing Al, Cr^{3+} ,
- Fe³⁺ and V^{3+} at the *B* site. Canadian Mineralogist, 47, 663–681.

- 744 Papike, J. J., Karner, J. M., and Shearer, C. K. (2006) Comparative planetary mineralogy:
- 745 Implications of martian and terrestrial jarosite. A crystal chemical perspective. Geochimica
- 746 Cosmochimica Acta, 70, 1309-1321.
- 747 Papike, J.J., Burger, P.V., Karner, J.M., Shearer, C.K., and Lueth, V.W. (2007) Terrestrial
- analogs of martian sulfates: Major, minor element systematic and Na-K zoning in selected
- samples. American Mineralogist, 92, 444-447.
- 750 Papike, J.J., Karner, J.M., Shearer, C.K., and Burger, P.V. (2009) Silicate mineralogy of martian
- 751 meteorites. Geochimica Cosmochimica Acta, 73, 7443-7485.
- 752 Ripmeester, J.A., Ratcliffe, C.I., Dutrizac, J.E., and Jambor, J.L. (1986) Hydronium ion in the
- alunite-jarosite group. Canadian Mineralogist, 24, 435-447.
- Roach, L.H., Mustard, J.F., Swayze, G., Milliken, R.E., Bishop, J.L., Murchie, S.L., and
- 755 Lichtenberg, K. (2010) Hydrated mineral stratigraphy of Ius Chasma, Valles Marineris.
- 756 Icarus, 206, 253–268.
- 757 Rossman, G.R. (1976) Spectroscopic and magnetic studies of ferric iron hydroxyl sulfates: the
- series $Fe(OH)SO_4 \cdot nH_2O$ and the jarosites. American Mineralogist, 61, 398–404.
- 759 Sasaki, K., Tanaike, O., and Konno, K. (1998) Distinction of jarosite-group compounds by
- raman spectroscopy. Canadian Mineralogist, 36, 1225-1235.
- 761 Scarlett, N.V.Y., Grey, I.E., and Brand, H.E.A. (2010) Ordering of iron vacancies in monoclinic
- 762 jarosites. Amercian Mineralogist, 95, 1590-1593.
- 763 Scott, K.M. (1987) Solid solution in, and classification of, gossan-derived members of the
- alunite-jarosite family, northwest Queensland, Australia. American Mineralogist, 72, 178-
- 765 187.
- Serna, C.J., Parada Cortina, C., and Garcia Ramos, J.V. (1986) Infrared and Raman study of

767	alunite	iarosite com	pounds S	pectrochimi	ca Acta I	Part A 42	729-734
101	aranne		poundo. D				

- 768 Sherman, D.M. and Waite, T.D. (1985) Electronic spectra of Fe3+ oxides and oxide hydroxides
- in the near IR to near UV. American Mineralogist, 70, 1262–1269.
- 770 Stoffregen, R.E., and C.N. Alpers (1992) Observations on the unit-cell dimensions, H₂O
- contents, and δD values of natural and synthetic alunite. American Mineralogist, 77, 1092-
- 772 1098.
- 573 Stoffregen, R.E., Alpers, C.N., and Jambor, J. L. (2000) Alunite-jarosite crystallography,
- thermodynamics, and geochronology. In C.N. Alpers, J.L. Jambor, and D.K. Nordstrom,
- Eds. Sulfate Minerals: Crystallography, Geochemistry, and Environmental Significance.
- Reviews in Mineralogy and Geochemistry, Mineralogical Society Of America, Washington,
- 777 D. C., pp. 453-479.
- 578 Swayze, G.A., Desborough, G.A., Smith, K.S., Lowers, H.A., Hammarstrom, J.M., Diehl, S.F.,
- Leinz, R.W., and Driscoll, R.L. (2008a) Understanding jarosite—From mine waste to Mars.
- 780 In P.L. Verplanck, Ed. Understanding Contaminants Associated With Mineral Deposits,
- 781 United States Geological Survey Circular, 1328, 8–13.
- 782 Swayze, G.A., Ehlmann, B.L., Milliken, R.E., Poulet, F., Wray, J.J., Rye, R.O., Clark, R.N.,
- 783 Desborough, G.A., Crowley, J.K., Gondet, B., Mustard, J.F., Seelos, K.D., and Murchie,
- 784 S.L. (2008b), Discovery of the acid-sulfate mineral alunite in Terra Sirenum, Mars, using
- 785 MRO CRISM: Possible evidence for acid-saline lacustrine deposits? Eos Transactions of the
- American Geophysical Union, 89(53), Fall Meeting Supplement, Abstract P44A–04.
- 787 Thollot, P., Mangold, N., Ansan, V., Le Mouélic, S., Milliken, R.E., Bishop, J.L., Weitz, C.M.,
- Roach, L.H., Mustard, J.F., and Murchie, S. L. (2012) Most Mars minerals in a nutshell:

12/11

- 789 Various alteration phases formed in a single environment in Noctis Labyrinthus. Journal of
- 790 Geophysical Research, 117, E00J06, doi:10.1029/2011JE004028.
- Wang, A. (2012) In situ laser Raman Spectroscopy for Mars Sample Return Mission. 43rd Lunar
 and Planetary Science Conference, abstract #2149.
- 793 Wang, A., Haskin, L.A., Lane, A.L., Wdowiak, T.J., Squyres, S.W., Wilson, R.J., Hovland, L.E.,
- Manatt, K.S., Raouf, N., and Smith, C.D. (2003) Development of the Mars Microbeam
- Raman Spectrometer (MMRS). Journal of Geophysical Research, 108, 5005,
- 796 doi:10.1029/2002JE001902.
- Wang, A., Freeman, J. J., and Jolliff, B. L. (2009) Phase transition pathways of the hydrates of
- magnesium sulfate in the temperature range 50°C to 5°C: Implication for Sulfates on Mars.
- Journal of Geophysical Research, 114, doi:10.1029/2008JE003266.
- 800 Weitz, C.M., Noe Dobrea, E.Z., Lane, M.D., and Knudson, A.T. (2012) Geologic relationship
- 801 between gray hematite, sulfates, and clays in Capri Chasma. Journal of Geophysical
- 802 Research, 117, E00J09, doi:10.1029/2012JE004092.
- 803 Wray, J.J., Milliken, R.E., Dundas, C.M., Swayze, G.A., Andrews-Hanna, J.C., Baldridge, A.M.,
- 804 Chojnacki, M., Bishop, J.L., Ehlmann, B.L., Murchie, S.L., Clark, R.N., Seelos, F.P.,
- 805 Tornabene, L.L., and Squyres, S.W. (2011) Columbus crater and other possible groundwater-
- fed paleolakes of Terra Sirenum, Mars. Journal of Geophysical Research, 116, E01001,
- doi:10.1029/2010JE003694.

End-member	Formula
Alunite subgroup (Al	> <i>Fe</i>)
Alunite	$KAl_3(SO_4)_2(OH)_6$
Natroalunite	$NaAl_3(SO_4)_2(OH)_6$
Jarosite subgroup (Fe	e > Al)
Jarosite	$KFe_3(SO_4)_2(OH)_6$
Natrojarosite	$NaFe_3(SO_4)_2(OH)_6$
•	$(\mathbf{U} \mathbf{O}^{+})\mathbf{E}_{2}$ (SO) (OU)

809 Table 1. Ideal end-member molecular formulas810 for common minerals from the alunite group.

814
 Table 2. Relative Fe-Al contents of Fe-bearing
 815 natroalunites (expressed as Fe#) in natural samples from Cerro Negro, Telica, and Masaya 816 volcanoes, Nicaragua, and from a laboratory 817 simulation of acid-sulfate alteration of basalt 818 819 (ADSU6).

Sample	Method	п	Average Fe# (std. dev.)
CN R21	EDS EMPA	24	17 (12)
CN R36	EMPA EDS	10	15(12)
	$\mathrm{EMPA}\ \mathrm{XRF}^{\#}$	14 2	16 (2) 16.7 (0.1)
CN R28	EDS	9	6(1)
Telica 6	EDS	14	4 (2)
Telica 7	EDS	38	26 (16)
Masaya 3	EDS	14	39 (14)
ADSU6	EMPA	26	38 (5)

820 $Fe# = 100 \times [Fe/(Fe + Al)]$, molar basis. n = number of

821 822 analyses. [#]Bulk rock analysis of sample CN R36, which

is composed predominantly (>90%) of natroalunite

823 (McCollom et al. 2013b).

825
826 Table 3. Calculated molecular formulas of synthetic natroalunite827 natrojarosite solid solutions based on wet chemical analyses (see
828 McCollom et al. 2013a).

Sample	Calculated Formula	Fe#
Natro6	$(Na_{0.82}H_{3}O_{0.18}^{+})Al_{2.69}(SO_{4})_{2}(OH)_{5.07}(H_{2}O)_{0.93}$	0
Natro8	$(Na_{0.89}H_{3}O_{0.11}^{+})(Al_{2.59}Fe_{0.30})(SO_{4})_{2}(OH)_{5.67}(H_{2}O)_{0.33}$	10
Natro4	$(Na_{0.96}H_{3}O_{0.04}^{+})(Al_{2.36}Fe_{0.60})(SO_{4})_{2}(OH)_{5.88}(H_{2}O)_{0.12}$	20
Natro2	$(Na_{0.87}H_3O_{0.13}^+)(Al_{2.10}Fe_{0.70})(SO_4)_2(OH)_{5.40}(H_2O)_{0.60}$	25
Natro7	$(Na_{0.88}H_{3}O_{0.12}^{+})(Al_{1.57}Fe_{1.48})(SO_{4})_{2}(OH)_{6.15}$	49
Natro1	$(Na_{0.91}H_3O_{0.09}^+)(Al_{1.15}Fe_{1.79})(SO_4)_2(OH)_{5.82}(H_2O)_{0.18}$	61
Natro3	$(Na_{0.90}H_3O_{0.10}^+)(Al_{1.00}Fe_{1.89})(SO_4)_2(OH)_{5.67}(H_2O)_{0.33}$	65
Natro14	$(Na_{0.84}H_3O_{0.16}^+)(Al_{0.45}Fe_{2.40})(SO_4)_2(OH)_{5.55}(H_2O)_{0.45}$	84
Natro10	$(Na_{0.86}H_3O_{0.14}^+)(Al_{0.22}Fe_{2.51})(SO_4)_2(OH)_{5.19}(H_2O)_{0.81}$	92
Natro9	$(Na_{0.93}H_{3}O^{+}_{0.10})Fe_{2.96}(SO_{4})_{2}(OH)_{5.88}(H_{2}O)_{0.12}$	100

Table 4. Room temperature magnetic hyperfineMössbauer parameters for synthetic Fe-bearing natroalunite.

	QS (mm/s)	IS (mm/s)	%
Natro8 (Fe#=10)	$1.24(1) \\ 1.23(1) \\ 1.19(1) \\ 1.12(1) \\ 1.06(1) \\ 1.11(1) $	0.36(1)	100
Natro4 (Fe#=20)		0.38(1)	100
Natro2 (Fe#=25)		0.36(1)	100
Natro1 (Fe#=61)		0.36(1)	100
Natro14 (Fe#=84)		0.35(1)	100
Natro10 (Fe#=92)		0.36(1)	100

QS = quadrupole splitting and IS = isomer shift. Errors are quoted in parenthesis. Mössbauer parameters were fit with single components in all cases.

836 837

843 **Table 5.** Mineral composition of acid-sulfate fumarole deposits from Nicaraguan volcanoes 844 containing Fe-bearing natroalunites (M = major, m = minor, t = trace), as determined by a 845 combination of XRD and SEM/EDS analysis.

Sample	Phases present
CN R21	Glass (<i>M</i>), amorphous silica (<i>M</i>), plagioclase (<i>m</i>), augite (<i>m</i>), natroalunite (<i>m</i>), gypsum (<i>m</i>), magnetite (<i>t</i>), FeOx (<i>t</i>)
CN R36	Natroalunite (<i>M</i>), amorphous silica (<i>t</i>)
CN R28	Gypsum (M) , natroalunite (M) , amorphous silica (m) , FeOx (t)
Telica 6	Natroalunite (<i>M</i>), kaolinite (<i>m</i>), amorphous silica (<i>m</i>)
Telica 7	Natroalunite (M) , amorphous silica (m)
Masaya 3	Natroalunite (<i>M</i>), amorphous silica (<i>m</i>)
ADSU6	Glass (<i>M</i>), amorphous silica (<i>M</i>), plagioclase (<i>m</i>), augite (<i>m</i>), anhydrite (<i>m</i>), natroalunite (<i>m</i>), hematite (<i>t</i>), FeOx (<i>t</i>)

Relative abundances estimated from XRD and SEM analysis based on: $M \approx 10-90\%$, $m \approx 2-10\%$, t < -2%. FeOx

847 = unidentified micron- to submicron-sized iron oxide/oxyhydroxides.



Figure 1. Backscattered electron images of representative pseudocubic Fe-bearing natroalunite from Cerro Negro volcano (a-c) and laboratory experiments (d-f). The natroalunite often precipitates in close association with amorphous silica (si). Larger prismatic crystals in (d) are anhydrite, and "dg" is dried amorphous silica gel. Bright spots are spheroidal hematite. Images shown in (c) and (f) are from polished cross-sections of altered basalt cinders embedded in epoxy, while all others are from uncoated grains mounted on aluminum stubs. Arrows in (c), (d), and (f) point to psuedocubic crystals. In (c), the natroalunite is growing on amorphous silica, a common habitat in natural samples. Panel (f) shows cross-sections of larger psuedocubic crystals precipitated in gas vesicles from an altered basalt cinder. Crystals shown in panels (c) and (f) exhibit compositional zoning from darker, more Al-rich interiors to lighter, more Fe-rich exteriors. Cerro Negro images are from samples CN R20 and CN R21 (McCollom et al. 2013b) and laboratory images are from experiments ADSU5 and ADSU6 (McCollom et al. 2013a).

- Figure 2. Examples of spheroidal hematite associated
- with Fe-bearing natroalunite in acid-sulfate altered rocks from (a) Cerro Negro and (b) basalt alteration
- experiments.
- 879

885	Figure 3.	Backscattered	electron	images	of
886	synthetic nati	roalunite-natroja	rosite soli	id solutio	ns.
887	(a) Natro3 (Fe#= 65). (b)	Natro 4	(Fe#= 2	0).

Figure 5. Raman analysis of synthetic natroalunite-natrojarosite solid solutions. (a,b,c) Raman 908 peaks in the (a) OH/H₂O spectral range, (b) spectral range of SO_4^{2-} fundamental vibrations (v₁ & 909 v₃), and (c) spectral range of fundamental vibrations (v₂ & v₄) and lattice modes. (d,e,f) 910 Variations in Raman peak position with Fe content for selected peaks.

Figure 6. (a) V1stble/near-infrared (VNIR) spectra of synthetic natroalunite-natrojarosite solid solutions. (b-e) Continuum removed data show systematic changes in absorption with Fe content for (b) a sharp electronic transition $(6A_{1g} \rightarrow {}^{4}T_{1g}, 4E_{g})$ that is most prominent in intermediate Fe samples and for (c, d, e) combinations and overtones of OH-related bending and stretching vibrations which vary in position and intensity according to iron content, with the sample of higher Fe contents generally having absorption centers at longer wavelength.

990 Figure 7. (a) Attenuated total reflectance (ATR) Mid-IR spectra of synthetic natroalunite-991 natrojarosite solid solutions. (b-e) Plots of band center versus iron content of the sample (Fe#) indicate, depending on the band in question, systematic or gradual changes in the positions of the 992 strongest absorptions due to (b,c) v_3 (SO₄²⁻), (d) δ (OH) and (e) v_3 (OH), that correlate with 993 changes in iron content. In general, band centers shift to lower wavenumbers (longer 994 995 wavelengths) with increasing Fe# though some band centers (e.g. panels c, e) appear to instead 996 have discrete positions depending on whether the natroalunite structure is relatively Al- or Fe-997 enriched. 998

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1047 Figure 9. Mössbauer spectra for synthetic natroalunite with Fe# = 10 (Natro8) and Fe# = 25 (Natro2), a 1048 library spectrum for jarosite (RRUFF database), and an example of acid-sulfate altered basalt from Cerro 1049 Negro (CN R21). Black circles are measured values. Synthetic natroalunite samples are fit with a single 1050 component (blue line). For the CN sample, the black line represents the total fit to the data using the 1051 components shown by colored lines, including minerals from the alunite group (blue), augite (red, aqua), 1052 olivine (green), Fe-oxide/oxyhydroxide (brown) and titanomagnetite (purple, pink).

Synthetic naroalunite-natrojarosite s.s.

 Burns Formation, Mars
 Natroalunite, Cerro Negro
 Natural jarosite subgroup

 Natroalunite, basalt alt. expt.

Figure 10. Mössbauer parameters (isomer shift, IS, versus quadrapole splitting, QS) for synthetic natroalunite-natrojarosite solid solutions, for the alunite group component in acid-sulfate altered basalt cinders at Cerro Negro, and for Fe-bearing natroalunite produced during experimental simulation of basalt acid-sulfate alteration. Included for comparison are some published parameters for natural and synthetic members of the jarosite subgroup and for the Fe3D4 phase from Meridiani Planum on Mars interpreted to be jarosite (data from Morris et al. 2006, and references therein). Data for Cerro Negro and basalt alteration experiments are from McCollom et al. (2013a,b).

Synthetic natroalunite-natrojarosites.s.

- A Natural jarosites
- Natroalunite, Cerro Negro

• Natroalunite, basalt alter. expt.

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- 1073

Figure 11. Variation in Mössbauer parameters quadrapole splitting (QS) and isomer shift (IS) as a function of Fe content of alunite group minerals, expressed as Fe#. Green bars define range of values for natural minerals from the jarosite subgroup compiled by Morris et al. (2006). Data for Al-bearing natrojarosites from Hrynkiewicz et al. (1965) and Johnson (1977), and for Cerro Negro and basalt alteration experiments from McCollom et al. (2013a,b).

- 1079
- 1080

Figure 12. Partial X-ray diffractograms for acid-sulfate altered deposits from Cerro Negro, Telica, and Masaya volcanoes. The figure shows peaks between $2\Theta = 28$ and 31° , encompassing the range where primary diffraction peaks for natroalunite-natrojarosite solid solutions occur.

Figure 13. (a) Raman spectra of individual crystals in natural samples from Cerro Negro (CN R21, CN R36) and from a laboratory experiment (ADSU6). Shown for comparison are several spectra from the synthetic natroalunites. (b,c) Expanded views of spectra at Raman shifts in the ranges 680-610 cm⁻¹ (b) and 540-420 cm⁻¹ (c).

Figure 14. VNIR spectra of natural samples from Cerro Negro, Telica, and Masaya volcanoes. Samples Telica 7, Telica 6, and Masaya 3 measured in the field (Marcucci et al. 2013), other spectra measured in the laboratory. Mont. = montmorillonite.

Figure 15. Mid-infrared diffuse reflectance spectra of natural samples from Cerro Negro, with several spectra from synthetic minerals shown for reference.

Synthetic naroalunite - natrojarosite s.s.

 Burns Formation, Mars
 Natroalunite, Cerro Negro
 Natural jarosite subgroup
 Natroalunite, basalt alt. expt.
 Synthetic jarosite subgroup

Synthetic natroalunite-natrojarosite s. s.

