2/26

1 2 3	Synthesis and characterization of the Mars-relevant phosphate minerals Fe- and Mg-whitlockite and merrillite and a possible mechanism that maintains charge balance during whitlockite to merrillite transformation
4 5 6 7 8	REVISION 1 Christopher T. Adcock ¹ *, Elisabeth M. Hausrath ¹ , Paul M. Forster ^{2,3} , Oliver Tschauner ^{1,3} and Kirellos J. Sefein ^{1†} Corresponding author email: <u>adcockc2@unlv.nevada.edu</u>
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17 18	Abstract
19	Memillite [Co NeWe(DO)] e como e e deminentenimente Co ale coloridade minentenimente
20	Merrillite [Ca ₉ NaMg(PO ₄) ₇] occurs as a dominant primary Ca-phosphate mineral in
21	martian meteorites and therefore presumably also on Mars. The mineral is an important phase in
22	exploring differences in geologic processes between Earth and Mars, and also has astrobiological
23	implications due to its potential role as a significant source of the bio-essential nutrient
24	phosphate. Merrillite does not occur terrestrially as a discrete mineral phase, making it difficult
25	to obtain for Mars-relevant studies. It can, however, be synthesized from a similar terrestrial
26	mineral, whitlockite (natural or synthetic), through dehydrogenation. Here we present methods
27	for synthesizing relatively large quantities (0.5 g or greater per batch) of coarse crystalline (75
28	μ m+) Mg-whitlockite, Fe-whitlockite, mixed Fe/Mg-whitlockites, and from these synthesized
29	minerals produce Mg-merrillite, ferrous and ferric Fe-merrillite and ferrous and ferric mixed
30	Fe/Mg-merrillite. Chemistry and atomic structures of synthesized Fe- and mixed Fe/Mg-
31	whitlockite and ferrous and ferric Fe- and mixed Fe/Mg- merrillite resulting from single-crystal
32	X-ray diffraction, infra-red spectroscopy, and electron microprobe analyses are presented. We
33	also present a mechanism for maintaining charge balance during the formation of merrillite from

34	whitlockite. Our results shed light on these mineral structures for future martian studies, and
35	provide methods for creating coarse crystalline merrillite for use in Mars-relevant
36	thermodynamic, kinetic, soil/dust simulant, crystallographic, astrobiological, and other studies.
37	Keywords: Whitlockite, merrillite, Mg, Fe, structures, synthesis, dehydrogenation,
38	phosphate, Mars, astrobiology.
39	
40	Introduction
41	
42	Merrillite is an anhydrous calcium phosphate mineral that occurs in lunar samples and
43	meteorites, including martian meteorites (Jolliff et al., 1993; McSween et al., 1996; Xie et al.,
44	2002; Terada et al., 2003). The IMA formula for merrillite is Ca ₉ NaMg(PO ₄) ₇ , though merrillite
45	can also be iron-bearing (Jolliff et al., 2006). Synthetic merrillite also lacks Na^+ and charge
46	balance is maintained by additional Ca^{2+} [i.e. $Ca_{9.5}Mg$ (PO ₄) ₇]. Merrillite is of particular interest
47	to martian studies because it is the dominant primary phosphate mineral in martian meteorites
48	(McSween et al., 1996; McSween and Treiman, 1998), and thus presumably also on Mars. This
49	is in contrast to Earth where fluorapatite $[Ca_5(PO_4)_3F]$ is the dominant terrestrial primary
50	phosphate mineral and the nearest analog to merrillite, the mineral whitlockite
51	[Ca ₉ Mg(PO ₃ OH)(PO ₄) ₆], occurs as a component of teeth and bones in vertebrates, but outside of
52	biological systems is a rare alteration product of phosphate pegmatites. Merrillite is also the
53	major carrier of Rare Earth Elements (REE) in martian meteorites (McSween et al., 1996;
54	McSween and Treiman, 1998; Shearer et al., 2011a) and likely one of the major sources of the
55	phosphorus observed in phosphorus rich martian soils and dust (Goetz et al., 2005; Greenwood
56	and Blake, 2006; Brückner et al., 2008). Phosphate minerals such as merrillite and apatite are

2/26

thought to react in late stage martian magmas (McCubbin and Nekvasil, 2008; Filiberto and 57 Treiman, 2009; Gross et al., 2013b), and their presence and intergrowth relationships with each 58 other within martian rocks (or meteorites) can be useful indicators of crystallization and melt 59 evolution, including the evolution of both water and halogen budgets over time (Patiño Douce 60 and Roden, 2006; Patiño Douce et al., 2011; Gross et al., 2013a; Gross et al., 2013b). Such use 61 of merrillite/apatite intergrowth relationships may also be applicable to other differentiated 62 bodies outside of Mars (Shearer et al., 2011b). Phosphate minerals have also been shown to be 63 altered in acid vapor environments, potentially recording important characteristics of the reacting 64 environment (Lane et al., 2008: Hausrath et al., 2013: Hausrath and Tschauner, 2013). In 65 addition, phosphate is a required component in fundamental biologic reactions as well as 66 considered critical to reactions that may have led to life on Earth (Gulick, 1955; Westheimer, 67 1987; Schwartz, 2006). The alteration of phosphate minerals such as merrillite has important 68 implications for phosphate availability and the potential for martian abiogenesis and 69 development of life on that planet (Adcock et al., 2013). Thus merrillite is an important phase in 70 exploring the differences in petrogenesis, mantle evolution, surface processes and other geologic 71 processes between Earth, Mars and possibly other bodies (McSween et al., 1996; Treiman, 2003; 72 Shearer et al., 2011a). 73

The importance of merrillite in martian processes makes a readily available source of the mineral desirable for laboratory study. However, single-phase natural merrillite only occurs in extraterrestrial materials and thus studies requiring merrillite must obtain the mineral directly from meteorites or lunar samples, or through synthesis. Here we present methods for producing fairly large quantities of crystalline Fe/Mg-bearing merrillite for use in Mars-relevant studies, the

chemistry and atomic arrangements of the synthetic minerals, and a proposed mechanism for
maintaining charge balance during the formation of merrillite from whitlockite.

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82 Background

Merrillite was first described in meteorites by Merrill (1915; 1917). Wherry (1917) was 83 the first to suggest applying "merrillite" as the name of the extraterrestrial mineral. Merrillite is 84 chemically and structurally similar to the terrestrially occurring mineral whitlockite; in fact, 85 86 though merrillite is not found terrestrially as a discrete mineral phase outside of biological systems, it does occur as a minor component within whitlockite, with which it forms a solid 87 solution (Hughes et al., 2008). In biological systems, merrillite forms in dental calculi and 88 urinary stones (Dorozhkin and Epple, 2002). The similarities of merrillite and whitlockite have 89 resulted in overlapping use of the mineral names within the literature, though they are not strictly 90 the same mineral. Whitlockite was first described by Frondel (1941) as a late stage calcium 91 phosphate mineral associated with pegmatites. Frondel and Prien (1946) later applied the term 92 "whitlockite" to synthetic tricalcium phosphate [β -Ca₃(PO₄)₂ or commonly β -TCP] because of 93 chemical and structural similarities of that phase. The similarity of merrillite, whitlockite, and β-94 TCP led Fuchs (1962) to argue, based primarily on powder X-ray diffraction data, that the three 95 phases were the same and that the term "merrillite" should be abandoned in favor of 96 "whitlockite". Subsequent research conclusively demonstrated that merrillite and β-TCP, while 97 isostructural and very similar to whitlockite, are distinct, and lack the required hydrogen of 98 99 whitlockite (Gopal and Calvo, 1972; Dowty, 1977; Rubin, 1997; Hughes et al., 2008). Dowty (1977) suggested "merrillite" as the more appropriate name for the hydrogen-free form found in 100 meteorites. However, because of the historical usage of "whitlockite" and the difficulty of 101

102	confirming the hydrous or anhydrous nature of the phases, the term "whitlockite" continues to be
103	occasionally used synonymously or interchangeably when describing merrillite (e.g. Ruszala and
104	Kostiner, 1980; Jolliff et al., 1993; McSween et al., 1996; Terada et al., 2003; Orlova et al.,
105	2009). This is especially the case in biomedical fields where synthetic merrillite compounds,
106	such as β -TCP, are important in bio-ceramics and are often referred to as whitlockite or β -
107	whitlockite (e.g. Jarcho et al., 1979; Lagier and Baud, 2003). In this paper, as in some other
108	studies that discuss both merrillite and whitlockite (e.g. Hughes et al., 2006; Jolliff et al., 2006)
109	the term "merrillite" is used to describe the anhydrous extraterrestrial (and synthetic) Ca-
110	phosphate mineral, and the term "whitlockite" to refer to the hydrogenated terrestrial mineral
111	form.
112	The most common method of synthesizing merrillite powder is solid state sintering,
113	which involves heating relevant powdered calcium phosphate chemistry to >1025 °C. The
114	method typically requires several cycles of regrinding the material during the synthesis process
115	(e.g. Lazoryak et al., 1996; Belik et al., 2002; Orlova et al., 2009), and therefore produces a very
116	fine powder. This general method is used to produce commercially available β -TCP powders
117	and ceramics, synthetic Ca-endmember merrillite materials. The same methods have been used
118	to synthesize other types of merrillite powders (Teterskii et al., 2007; Orlova et al., 2009),
119	including Mg-merrillite and ferric Fe-merrillite (Lazoryak et al., 1996). A drawback to this
120	general method is that it produces only very fine powders, which may not be preferred for some
121	studies (e.g. dissolution kinetics, crystallography, soil simulants). The commercial version is
122	also an endmember lacking in Fe or Mg and thus chemically dissimilar to the merrillite
123	commonly found in lunar samples or martian meteorites (Jolliff et al., 2006).

Another method of synthesizing merrillite is through the heating of whitlockite to >1000 124 °C for at least 24 hours to drive off the "essential hydrogen" (Gopal and Calvo, 1972; Hughes et 125 al., 2008). This method has the advantage of being simple to perform and has been applied to 126 both synthetic Mg-whitlockite and natural whitlockite crystals to produce crystalline merrillite 127 rather than powders. However, natural terrestrial whitlockite is itself uncommon and generally 128 restricted to small quantities associated with pegmatites. Thus, in order to synthesize quantities 129 of high purity crystalline Mg- and Fe-merrillite for use in Mars-related kinetic, thermodynamic, 130 131 astrobiological, or other studies, the synthesis of Mg- and Fe-whitlockite in large quantity is required. 132

Whitlockite can be synthesized by the Rowles method (Rowles, 1968) which involves 133 hydrolysis of brushite in magnesium chloride solution to create Mg-whitlockite. The method can 134 create pure Mg-whitlockite crystalline material and large output masses (500+g). The 135 crystalline material generated by this method is very fine with specific surface areas (SSA) of 3-136 7 m²/g (geometrically equivalent to 0.3-0.6 μ m diameter average particle size)(Hamad and 137 138 Heughebaert, 1986) making it of limited usefulness in kinetic, simulant, crystallographic, or 139 other studies where coarser crystalline material is preferred. Hydrothermal methods, however, can create coarser crystalline material. Gopal et al. (1974) synthesized coarse, up to 500 µm per 140 side, crystalline Mg-whitlockite using such a method. The method required the pre-synthesis 141 and processing of a calcium phosphate gel and a large (6mm x 100mm) single-use gold tube to 142 contain the reagents during heating to 300 °C, making it less practical for synthesizing in 143 quantity. Hughes et al. (2008) inadvertently produced a small amount (i.e. "a few crystals") of 144 145 Mg-whitlockite by similar methods while hydrothermally recrystallizing hydroxyapatite $[Ca_5(PO_4)_3OH]$. Their method involved first synthesizing hydroxyapatite powder from a calcium 146

147	phosphate solution and suspending it in 10ml of distilled water with the pH adjusted to 2.79
148	using phosphoric acid. The resulting suspension was sealed in a reusable 23ml Teflon lined acid
149	digestion vessel and heated to 240 °C for 5 days. The output mass from the hydrothermal
150	treatment consisted mainly of elongate, recrystallized hydroxyapatite along with a small number
151	of large, well-formed, Mg-whitlockite crystals. The apparent cause of whitlockite synthesis was
152	Mg contamination from an undetermined source. Here we present methods based on Hughes et
153	al. (2008) to generate large quantities of single phase Mg-whitlockite and, from that whitlockite,
154	Mg-merrillite by intentionally supplying Mg^{2+} to the synthesis. We also developed methods to
155	generate large amounts of crystalline Fe-whitlockite, as well as mixed Fe/Mg-whitlockites, and
156	ferrous, ferric, and mixed Fe/Mg-merrillites. Atomic parameters are presented for ferrous, ferric,
157	and mixed merrillite and Fe- and mixed whitlockite, as well as a proposed mechanism for
158	maintaining charge balance during the formation of merrillite from whitlockite.
159	
160	Methods
161	Whitlockite synthesis
162	Based on preliminary experiments (deposit items Appendix A and Table A1 ¹), the
163	following synthesis methods were developed for Mg- and Fe-whitlockite production: a solution
164	containing 90ml of high purity (18.2 M Ω) water, laboratory grade hydroxyapatite (Spectrum,
165	reagent grade), and either magnesium nitrate hexahydrate for Mg-whitlockite (J.T. Baker, ACS
166	grade) or troilite for Fe-whitlockite (Alfa Aesar, Fe[II]S 99%) were combined in a 125ml Parr
167	acid digestion vessel (Parr #4748) with an acid washed Teflon liner. Specific volumes and
168	

¹ Deposit items are available two ways: For a paper copy contact the Business Office of the Mineralogical Society of America (see inside front cover of recent issue) for price information. For an electronic copy visit the MSA web site at http://www.minsocam.org, go to the American Mineralogist Contents, find the table of contents for the specific volume/issue wanted, and then click on the deposit link there.

concentrated phosphoric acid (Alfa Aesar, ACS grade). The vessel was then sealed and 169 incubated in an oven at 240 °C for 7 days. At the end of 7 days, the vessel was removed from 170 the oven and guenched in a water bath in an effort to prevent any further reaction. After cooling, 171 the vessel was opened and the solution decanted leaving the solids. Solid material was rinsed 172 from the vessel using ethanol, allowed to air dry for 24 hours, weighed, and inspected by optical 173 microscopy for preliminary phase identification. Impurities, first determined by powder XRD 174 and then optically by morphology thereafter, were typically hydroxyapatite, monetite (CaHPO₄). 175 In the case of Fe-bearing whitlockite, additional opaque materials presumed to be Fe-phases 176 occurred in quantities too minor for powder XRD analysis. These impurities were primarily 177 confined to the <75 µm fraction and output masses were brush sieved on 200 mesh screen to 178 remove that size fraction. Crystals were confirmed as Mg- or Fe-whitlockite by Scanning 179 180 Electron Microscopy (SEM) with Energy Dispersive Spectroscopy (EDS), Electron Microprobe Analysis (EMP), Powder X-ray Diffraction (XRD), and Single Crystal X-ray Diffraction (SC-181 182 XRD).

183 Based on results from endmember experiments described above, mixed Fe/Mgwhitlockites were also successfully synthesized by combining both troilite and magnesium 184 185 nitrate hexahydrate (J.T. Baker ACS grade, and Alfa Aesar Fe[II]S 99%, respectively) in the 186 solution used within a single synthesis experiment. Table 1 contains the specific masses and volumes used to produce the mixed Fe/Mg-whitlockite that was further characterized in this 187 study and it includes sodium (as NaCl) as part of the solution. Subsequent experiments indicated 188 189 that the addition of NaCl in this method has no observable effect and can be omitted (see Results and Discussion and deposit items Appendix A and Table A1), but it is included in Table 1 as this 190 was the material further characterized in this study. Outside of solution chemistry, all other 191

2/26

192	parameters used in mixed Fe/Mg-whitlockite synthesis were the same as for the endmember
193	whitlockite synthesis (e.g. temperature, acidity, incubation time). Impurities found in mixed
194	Fe/Mg synthesis batches (typically large polycrystalline aggregates, hydroxyapatite and opaque
195	phases as determined optically) were primarily confined to the <75 μm and >150 μm fractions
196	and output masses were brush sieved to isolate the 75-150 μm size fraction. Remaining
197	crystalline material was confirmed as Fe/Mg-whitlockite by EMP and SC-XRD.

198

199 Merrillite synthesis

Merrillite can be synthesized from whitlockite (natural or synthetic) by dehydrogenation through heating to >1000 °C for at least 24 hours (Gopal et al., 1974; Hughes et al., 2008). In this study, crystalline Mg-merrillite was synthesized from synthetic Mg-whitlockite by heating to 1055 °C (+/-5 °C) for a minimum of 24 hours in a loosely covered 10 ml Pt crucible nested within a larger (100ml) alumina crucible for easier handling. After heating, the crucibles were removed and allowed to air cool. After cooling, resulting materials, typically white to clear in color, were removed from the crucible using a synthetic brush.

Crystalline ferric Fe-merrillite was synthesized in an identical manner as the Mg-207 merrillite described above, in a Pt crucible in ambient atmosphere. The process produces an 208 orange/red colored material. Ferrous Fe-merrillite was synthesized in an evacuated (10⁻³ Torr) 209 sealed SiO₂ tube in place of the Pt crucible to inhibit oxidation of Fe from ferrous to ferric in the 210 Fe-whitlockite. Tubes were purged with argon three times to remove oxygen before final 211 evacuation and sealing. SiO₂ glass wool (average 4 µm diameter) was used to prevent the loss of 212 material from the tubes during purge and evacuation cycles. Whitlockite heat treated in this 213 manner ranged in color from white/translucent to black and nearly opaque. 214

Mixed Fe/Mg-merrillite was synthesized both in a 10 ml Pt crucible in ambient 215 atmosphere, and also in a triple argon purged evacuated (10^{-3} Torr) sealed SiO₂ tube to inhibit 216 oxidation of the Fe in the Mg-Fe whitlockite, both as described above. Resulting materials from 217 all merrillite syntheses were analyzed by optical microscopy, EMP and SC-XRD to confirm the 218 phases. Oxidation of Fe during merrillite synthesis can mask detection of dehydrogenation by 219 SC-XRD, so Fe-containing phases were additionally analyzed by Infrared Spectroscopy (IR). As 220 with the Fe-endmember whitlockite treated in the same manner, the output material ranged in 221 2.2.2 color from white/translucent to black and nearly opaque. 223 **Analytical methods** 224 Powder X-ray Diffraction (XRD) analyses were performed in the UNLV Geoscience 225 XXL facility. Samples of synthesized Mg- and Fe-whitlockite, as well as impurities from 226 synthesis, were powdered by hand in an agate mortar and pestle and subjected to powder XRD in 227 a Panalytical X'pert Pro diffractometer using Cu Ka radiation. Scans were run at 45Kv and 228 40mA from 5-75° 20 in continuous mode with step-size of 0.0084° and scan rate of 0.06° sec⁻¹. 229 Patterns were identified and confirmed using Panalytical's X'pert Pro High Score Plus software. 230 Optical microscopy was used to identify synthesized phases based on crystal 231 habit/morphology and previous XRD analyses. These identifications were carried out with a 232 Barska tri-ocular stereomicroscope with a magnification range of 7x to 45x. An adaptor 233 mounted Canon XTi camera was used for all optical photomicrographs. 234 Scanning Electron Microscopy (SEM) in secondary electron imaging mode (SEI) and 235 Energy Dispersive Spectrometry (EDS) was carried out on a JEOL 5600 SEM at the UNLV 236 Electron Microanalysis and Imaging Laboratory (EMiL). Operating conditions were 20 KeV 237

238	and a 20 mm working distance. Analyses by Electron Microprobe (EMP) wavelength dispersive
239	spectroscopy (WDS) were carried out in a JEOL JXA-8900 at the UNLV EMiL facility on
240	polished epoxy mounts. Analysis conditions were 20 KeV and 10 nA using a 10 μ m beam.
241	Analyses were standardized using Smithsonian mineral standards of olivine (Mg, Fe) and apatite
242	(Ca, P)(Jarosewich, 2002), and Harvard (Amelia) albite (Na)(McGuire et al., 1992).
243	Single Crystal X-ray Diffraction (SC-XRD) analyses were carried out using a Bruker
244	APEXII single crystal diffractometer. An appropriate crystal was selected under paratone oil on
245	a polarizing microscope and attached to a Kapton mount. A full sphere of data was collected
246	with the sample cooled to 100 K using an Oxford nitrogen cryostream. An absorption correction
247	utilized the Bruker software SADABS, structure solution was performed using a beta-version of
248	SHELXT, and refinement was carried out using a combination of SHELX (Sheldrick, 2007) and
249	OLEXII (Dolomanov et al., 2009). Bond valence calculations from SC-XRD data were
250	conducted using VALIST (Wills, 1999).
251	Infrared (IR) spectra were obtained with a Digilab FTS 7000 series FT-infrared
252	spectrometer on powdered samples using a photo-acoustic chamber attachment. Analyses were
253	performed under ambient conditions at 2.5 KHz and a resolution of 8 cm ⁻¹ for 64 total scans.
254	Data were collected over the range of 400 to 4000 cm ⁻¹ . Carbon black was used to obtain a
255	background versus subsequent spectra.
256	Synchrotron X-ray diffraction experiments were conducted on two samples (a red ferric
257	Fe/Mg-merrillite crystal and a dark colored ferrous Fe/Mg-merrillite crystal) at the
258	superconducting bending magnet beamline 12.2.2 at the Advanced Light Source (ALS),
259	Lawrence Berkeley National Laboratory, with the goal of identifying any trace fine-grained
260	phases within the merrillite. Primary beam energy was 20.000 keV and a MAR345 image plate

261	detector was utilized. Detector parameters were calibrated and corrected for geometric
262	distortions based on a LaB6 NIST powder diffraction standard using the Fit2D software
263	(Hammersley et al., 1996). The X-ray beam at beamline 12.2.2 of the ALS is focused by
264	Kirkpatrick-Baesz mirrors vertically and horizontally to 10 x 15 μ m ² . Crystal samples were
265	mounted on the tips of quartz glass fibers of 30 μ m diameter and centered on the phi rotation
266	axis. Exposures were taken over 160 s while the samples were oscillated by +/-40 degrees
267	around the phi axis. Recorded images were integrated using Fit2D (Hammersley et al., 1996).
268	We first used Fit2D to mask strong reflections from merrillite and any remaining minor
269	untransformed whitlockite in the diffraction image frames and integrated the remaining patterns.
270	Integrated patterns were still dominated by merrillite reflections and less intense whitlockite
271	reflections. The contributions from both these phases were fitted by a Le Bail refinement (Le
272	Bail, 2005). Remaining reflections were then identifiable as fine-grained inclusions. We used
273	the first strong diffraction peaks of these fine-grained inclusions for a search in the American
274	Mineralogist database (Downs and Hall-Wallace, 2003) and commercial data bases (ICSD) for
275	identifying possible inclusion phases (Bergerhoff and Brown, 1987).
276	
277	Results and Discussion
278	Optimization of Whitlockite Synthesis
279	Results of preliminary Mg-whitlockite experiments indicated that the concentration of
280	solids in solution (g/L) was a strong control on Mg-whitlockite formation (Table A1 and
281	Appendix A, on deposit ¹). Concentrations of greater than approximately 22 g/L produced
282	primarily monetite (CaHPO ₄) and hydroxyapatite (HAP) crystalline material rather than
283	whitlockite. Concentrations of less than 22 g/L, including the 14.5 g/L used in the finalized

284	method, produced yields consisting primarily of well formed crystals of Mg-whitlockite
285	confirmed by morphology, SEM EDS, powder XRD, and SC-XRD (Figures 1A and 2A).
286	Solution concentrations of significantly less than 14.5 g/L (e.g. 11.1 g/L) produced whitlockite,
287	but in smaller amounts (Table A1, on deposit ¹). Variations in incubation time (5-14 days)
288	indicated that most crystal growth occurred in the first 7 to 8 days. Longer incubation times (up
289	to 14 days) did not produce significantly larger crystals. Variations in pH from 2.1 to 2.8 did not
290	appear to have any significant effect. Scaling up from 23ml vessels to 125ml vessels had no
291	effect as long as solution concentrations were maintained. Temperature (i.e. 240 °C) was not
292	varied and represented the highest safe temperature rated for the vessels used. The ratio of HAP
293	to magnesium nitrate hexahydrate (3.33: 1) was not varied significantly, but minor changes
294	(2.93:1 to 3.52:1) had no noticeable effect on syntheses. Overall yields from multiple batches of
295	Mg-whitlockite synthesis using the final method had an average yield of >700 mg (Figs. 1A and
296	2A, deposit item Table A1). The $<75\mu$ m fraction was removed to minimize impurities (i.e.
297	monetite and hydroxyapatite); this fraction averaged $<2\%$ of the total output mass.
298	Imperfections within some crystals, mainly in the larger size fraction, were apparent optically,
299	but centralized clusters of inclusions similar to those seen by Hughes at al. (2008), while present,
300	were uncommon. Most crystals (>80%) were over 150 μm and the largest crystals were 2mm.
301	Polycrystalline aggregates of whitlockite occurred in the >75 μ m fraction but the majority of the
302	material were single crystals. EMP analyses of synthesized crystals (Table 2) confirmed Mg-
303	whitlockite stoichiometry as Ca _{9.0} Mg _{0.9} (PO ₃ OH)(PO ₄) ₆ .
304	Attempts to synthesize Fe-whitlockite using FeCl ₂ as the source of Fe failed to produce
305	the mineral in amounts detectable by optical microscopy, even at total solid in solution
306	concentrations as low as 12.0 g/L. Instead, crystal morphologies consistent with monetite and

307	recrystallized HAP were produced as well as minor opaques (presumably Fe phases). The
308	subsequent use of Fe(II)S as the source of Fe did produce Fe-whitlockite. Aggregate and
309	hydroxyapatite production were minimized by decreasing the concentration of solids in solution
310	from an initial 16.1 g/L to $<$ 12.0 g/L. To minimize iron sulfide phases, the ratio of
311	hydroxyapatite to Fe(II)S in solution was increased from ~9:1 (near stoichiometric ideal) to
312	between 16:1 and 20:1 in an effort to "starve" the system of Fe. Yields from Fe-whitlockite
313	syntheses using the final method described in the methods section averaged >500 mg of largely
314	well formed whitlockite crystals (Figs. 1B and 2B, deposit item Table A1). Impurities (i.e. iron
315	sulfides, hydroxyapatite, and monetite) mainly occurred in the <75 μ m fraction which could be
316	removed through sieving. The <75 μm fraction accounted for 33% of the total output mass on
317	average. Inter-crystalline Fe-films were apparent optically in some aggregates (Fig. 1C). Some
318	Fe-whitlockite crystals themselves contained minor intra-crystalline Fe-film deposits
319	(determined optically) on earlier growth surfaces (Fig. 1B), with the crystals separable by
320	picking. Minor amounts of Mg (0.14 wt. % as MgO) were also detected in Fe-whitlockites by
321	EMP (Table 2). Like Hughes et al. (2008), the source of Mg is unknown, but is most likely from
322	trace impurities in the reagents used. EMP analyses of synthesized crystals (Table 2) confirmed
323	synthetic Fe-whitlockite stoichiometry as $Ca_{9.0}Fe_{1.0}(P_{1.0}O_3OH)(PO_4)_6$.
324	While Mg- and Fe-endmember whitlockite (and the merrillite synthesized from them)
325	represent closer analogs to naturally occurring minerals than a phase like β -Ca ₃ (PO ₄) ₂ , mixed
326	Fe/Mg phases are closer analogs still. Two early experiments to produce mixed Fe/Mg-
327	whitlockite failed to produce any significant whitlockite (Table A1 on deposit ¹). With
328	understanding gained from further endmember experiments, another set of experimental

329 conditions was tested with a lower solution concentration (13.0 g/L as opposed to >17 g/L for the

330	earlier attempts). This experiment included a source of Na ⁺ as part of the solution (included in
331	the solution concentration noted above). Though Na ⁺ does not generally occur in whitlockite, the
332	ion does occur in a specific Ca^{2+} site in natural merrillite, including merrillite found in martian
333	meteorites where Na^+ content can be as high as 2 wt.% (Jolliff et al., 2006). The batch produced
334	Fe/Mg mixed whitlockite (Figs. 1D and 2C) along with impurities of hydroxyapatite and
335	monetite, and large (>150 μm) polycrystalline aggregates. Most impurities were in the >150 μm
336	and <75 μ m size fractions, leaving the 75-150 μ m size fraction the purest. This fraction
337	comprised 39% of the total output mass. EMP analyses showed no incorporation of Na^+ into any
338	of the synthesized material and an average stoichiometry of $Ca_{9.0}Fe_{0.7}Mg_{0.3}(PO_3OH)(PO_4)_6$
339	(Table 2). The whitlockite seems to preferentially incorporate Mg, as the whitlockite from this
340	batch had a Fe to Mg molar ratio of \sim 7:3 but the molar ratio of Fe to Mg in the reactants added
341	was ~9:3. EMP analyses of results from additional mixed Fe/Mg-whitlockite synthesis
342	experiments support this observation (Table A2, on deposit). It is also of note that the mixed
343	Fe/Mg-whitlockite synthesis experiments had a higher failure rate than those of the endmembers,
344	even after finding a generally successful set of conditions. Further investigation might yield the
345	source of the failures or better refine the method. The lack of Na^+ incorporated within the
346	whitlockite synthesized in these experiments suggests it may not be possible to incorporate the
347	ion into the structure under hydrothermal conditions. Thus, mixed Fe/Mg-merrillite
348	subsequently synthesized from the mixed Fe/Mg-whitlockite would also lack the sodium
349	generally found in martian and lunar merrillite. Further investigation to determine conditions
350	that favor Na ⁺ substitution in merrillite are warranted as they could potentially reveal important
351	information about the specific martian environments in which merrillite forms. Nevertheless,
352	sodium-free mixed Fe/Mg-bearing merrillite represents a better analog to martian merrillite than

353	pure endmember phases. EMP analyses revealed standard deviations of Fe and Mg contents
354	were slightly higher than those of endmember synthesized material (Table 2). However, the
355	minor increase in inhomogeneity should not limit the usefulness of the material for most studies.
356	

357 Merrillite Synthesis

Merrillite and whitlockite structures are too similar to be easily distinguished by powder 358 XRD, thus heat-treated materials were analyzed by SC-XRD and EMP to determine if the 359 whitlockite to merrillite transformation had occurred. In the case of Mg-merrillite, SC-XRD 360 (discussed in more detail below under Crystallography) confirmed synthesis had occurred and 361 EMP analyses of the resulting Mg-merrillite (Table 2) indicated a stoichiometry of 362 $Ca_{9.4}Mg_{1.1}(PO_4)_{7.0}$, compared to $Ca_{9.0}Mg_{0.9}(PO_3OH)(PO_4)_6$ of the starting Mg-whitlockite. The 363 higher stoichiometric value for Ca²⁺ results from Ca substitution for charge balance due to 364 dehydrogenation. Dehydrogenation is also the likely cause of higher observed totals for EMP 365 analyses of Mg-merrillite versus the Mg-whitlockite used to synthesize it (99.35% versus 366 98.02% respectively). The synthesized Mg-merrillite maintained the crystallinity and 367 morphology of the starting whitlockite material (Fig. 1E). 368

Fe-merrillite synthesis proved more complex than Mg-merrillite synthesis due to the apparent oxidation of ferrous Fe²⁺ to ferric Fe³⁺ when heated in open air. Fe-whitlockite samples heat-treated in open air (i.e. in the Pt crucible) changed color from white to orange or yellow. Like the Mg-merrillite, EMP analyses of the heat treated material showed overall totals rise from 98.15% for the original whitlockite to 98.80% for the heat treated material (with Fe calculated as FeO, see Table 2 and associated table notes), likely due to the loss of H⁺ during heating. However, EMP analyses also indicated a stoichiometry that did not significantly change from the Fe-whitlockite of $Ca_{9.0}Fe_{1.0}(PO_4)_{7.0}$. This EMP result, and the color change, suggested oxidation of Fe^{2+} to Fe^{3+} to produce a ferric Fe-merrillite phase balancing the change in charge due to the loss of H⁺ as shown below:

379
$$Ca_9Fe^{2+}(PO_3OH)(PO_4)_6 \rightarrow Ca_9Fe^{3+}(PO_4)_7.$$
 (1)

Bond valence calculations (Skowron and Brown, 1990) and Fe – O bond lengths (Table 3) based 380 on SC-XRD analyses of the heat treated material were also consistent with Fe^{3+} . To confirm 381 dehydrogenation, IR analyses were performed on both the unreacted Fe-whitlockite and the heat-382 treated Fe-merrillite to determine the presence or absence of O-H bonds (Fig. 3A). The IR 383 spectra show a large reduction in absorption between the starting material and the heat-treated 384 material over the range of 3330 to 2570 cm⁻¹ where a broad feature is associated with an O–H 385 stretch (White, 1990; Belik et al., 2002). There is also an O–H absorption peak at 2380 cm⁻¹ 386 (Belik et al., 2002) which disappears between the starting whitlockite material and the samples 387 heated treated in the Pt crucible open to the atmosphere (Fig. 3A), further supporting the 388 conclusion that the material is ferric (Fe^{3+}) Fe-merrillite. Previous research using solid state 389 methods has also produced a ferric Fe-merrillite (e.g. Lazoryak et al., 1996; Belik et al., 2002), 390 although the material in those experiments was a fine powder. 391

In order to prevent Fe oxidation and promote the formation of ferrous Fe-merrillite, Fewhitlockite was heated to 1055 °C in triple argon purged evacuated fused SiO₂ tubes instead of the open air Pt crucible. After heat treatment, EMP analyses indicated a stoichiometry of $Ca_{9,2}Fe_{1,0}Mg_{0,1}(P_{1,0}O_4)_7$ and overall totals rose compared to the starting material (Table 2). Bond valence calculations performed using VALIST (Wills, 1999) and Fe–O bond lengths (Table 3) from SC-XRD data are consistent with an Fe oxidation state of 2+, indicating that the phase is ferrous. However, IR results showed a remaining peak at 2380 cm⁻¹ indicating that full

399	dehydrogenation did not occur during the heat treatment (Fig. 3A). Thus the resulting phase
400	consists of ferrous Fe-merrillite containing some whitlockite. Extending the incubation time at
401	1055 °C to three days had no effect on the IR results. The result of heat treating the Fe-
402	whitlockite in a sealed SiO ₂ appears to be ~80-90% transformation to merrillite based on SC-
403	XRD and is discussed in more detail in the Crystallography section. The incomplete
404	dehydrogenation of the ferrous merrillite may be the result of changing partial pressure due to $\boldsymbol{H}^{\!+}$
405	loss in the sealed system and warrants further exploration. Many of the heat treated crystals
406	were darker in color suggesting the formation of other, possibly Fe-containing, phases, although
407	BSE and EMP analyses were unable to detect a separate phase within crystals or a color
408	dependent difference in chemistry.
409	Formation of mixed Fe/Mg-merrillite followed similar trends to that of endmember Fe-
410	merrillite synthesis, with samples heated in open air oxidizing to Fe ³⁺ /Mg-merrillite and samples
411	heated in an evacuated SiO_2 tube forming Fe^{2+}/Mg -merrillite. Fe/Mg-whitlockite heat treated in
412	a Pt crucible in open air showed a rise in EMP totals over the starting material beyond that
413	attributable only to total calculations based on Fe ₂ O ₃ rather than FeO (Table 2 and associated
414	table notes). IR analyses, as described above, (Fig. 3B) indicated dehydrogenation had occurred.
415	Bond valence calculations based on SC-XRD data showed that Fe had oxidized to Fe ³⁺ , and
416	EMP analyses of the reacted material showed a stoichiometry of $Ca_{9.2}Fe_{0.7}Mg_{0.4}(P_{1.0}O_4)_7$
417	compared to the unreacted whitlockite of $Ca_{9.0}Fe_{0.7}Mg_{0.3}(P_{1.0}O_3OH)(PO_4)_6$ indicating Ca
418	mobility to balance the change in charge due to the loss of H^+ had occurred. Heat treatment also
419	caused a color change in the material from white/translucent to red/orange.
420	Dehydrogenation of mixed Fe/Mg-whitlockite by heating in an evacuated fused SiO ₂ tube
421	resulted in crystalline material that varied in color from clear translucent to nearly opaque dark

422	black. SC-XRD analysis performed on a clear crystal indicated it to be Fe ²⁺ /Mg-merrillite that
423	contained some whitlockite. Fe - O distances determined by SC-XRD indicate that the Fe had
424	not significantly oxidized. A smaller, sharp O-H absorption peak at 2380 cm ⁻¹ remained (Fig.
425	3B) indicating that the phase contained some whitlockite. Similar to the ferrous endmember Fe-
426	merrillite extending the incubation time to three days had no effect on the IR results. Estimates
427	of the merrillite/whitlockite content are discussed in more detail in the Crystallography section,
428	but indicate that most of the mixed Fe/Mg-reacted material is merrillite (~70%). The color
429	change observed in post heat treatment crystals, as with the Fe-merrillite treated in a SiO ₂ tube,
430	suggested the possible formation of another phase. However, as with the similarly treated
431	endmember Fe-whitlockite, BSE and EMP analyses were unable to detect a separate phase
432	within the crystals or a difference in chemistry between lighter or darker colored material.
433	Surface deposits or films on crystals of synthesized endmember and mixed merrillite
434	crystals are apparent under SEM imaging (Figs. 2D to 2F) and by optical microscopy (e.g. Fig.
435	1E). Large (> 150 μ m) Mg-merrillite crystals were completely covered with the film, which is
436	generally sub-micron in thickness. The amount of coverage appeared to decrease with
437	decreasing grain size (i.e. smaller crystals exhibited less complete or thinner coverage). In
438	contrast, deposits on Fe ²⁺ -bearing crystals rarely covered the entire crystal surface and in the case
439	of ferric Fe-merrillite, most crystals showed no coating (Fig. 1F). SC-XRD analyses did not
440	detect a second phase, although SC-XRD analysis would be unlikely to detect such small
441	amounts of a second phase. The surface textures also lack any morphology that would suggest
442	crystallinity and thus may be amorphous. The coatings were also not detectable in EMP
443	prepared samples due to their sub-micron thickness and similar chemistry. EDS in SEM of the
444	reacted surfaces indicated that the surface deposits were mainly Ca and P with no evidence of

Mg or Fe, and therefore suggest that they may be the result of some form of diffusion to the surface during heating.

In an attempt to determine what additional phases might be present in the merrillite 447 samples, two samples were analyzed by synchrotron XRD; a translucent red-colored crystal of 448 the Fe/Mg-merrillite heat treated in open air (ferric Fe/Mg-merrillite), and a crystal containing a 449 hazy grey to black core within a clear crystal matrix of Fe/Mg-merrillite heat treated in a SiO₂ 450 tube (ferrous Fe/Mg merrillite). Specimens were 100 to 150 μ m in diameter. In both cases, 451 diffraction patterns exhibited three distinct sets of features: reflections that belong to multiple 452 domains of merrillite, reflections that belong to domains from whitlockite, and Debye fringes 453 from very fined grained (grainsize $< 1 \mu m$) inclusions. Thus the samples were not true single 454 crystals. 455

Modeled and observed patterns of the red-colored ferric Fe/Mg-merrillite sample along 456 with the residual of the fit appear in Figure 4A. In addition to the expected merrillite and 457 whitlockite, monetite and a minor amount of holtedahlite [Mg₂(PO₄)(OH)] were also identified. 458 Holtedahlite occurs as a fine grained phase which generated a smooth powder diffraction pattern. 459 With the given experimental settings and based on the observed peak profiles the average grain 460 size of holtdahlite is less than 1 µm and more than 100 nm. Since the diffraction signals from 461 whitlockite, merrillite, and monetite were not powderous, a Reitveld analysis for phase 462 quantification could not be conducted. However, the strength of the diffraction signal of 463 holtedahlite indicates that it is significantly less than merrillite, whitlockite, and monetite. 464 The ferrous Fe/Mg-merrillite specimen with the grey-black core region was composed of 465 a solid solution whitlockite/merrillite matrix (>3 micron grain size) containing small amounts of 466

467 monetite and an additional unidentified fine grained ($< 1\mu m$) phase. Limited diffraction data are

consistent with a Na-bearing Ca-ferrate, although a confident identification is not possible and
oxidized Fe would be surprising in the ferrous Fe/Mg-merrillite. The observed pattern with the
whitlockite/merrillite contributions masked and modeled patterns of monetite, a Na-bearing CaFerrate, and the residual of the fit appear in Figure 4B.

Presumably, some of these additional phases may also occur in the endmember merrillite materials, though we were not able to confirm that as part of the current study. Synchrotron diffraction patterns of monetite might explain the calcium and phosphorus phase observed as films on crystals. SC-XRD analyses did not detect monetite, holtedahlite or another unidentified phase, but the technique and sample selection criteria are designed to refine single structures from single crystals and are not applicable to detecting minor phases that exhibit reflections from multiple domains or powder fringe patterns.

The mechanism by which calcium balances the charge imbalance due to the loss of H^+ 479 from whitlockite and increases the stoichiometric Ca in non-ferric merrillite is unclear. Gopal 480 and Calvo (1972) and Gopal et al. (1974), who synthesized merrillite from both natural and 481 synthetic Mg-whitlockite, present no mechanism. Hughes et al. (2008), who synthesized 482 merrillite from synthetic Mg-whitlockite, suggest Ca-rich reagents entrained in the synthetic 483 whitlockite crystals yielded Ca sufficient to counter the charge imbalance due to H^+ loss, with 484 the Ca presumably diffusing through the crystal. However, this mechanism fails to explain 485 enhanced Ca concentrations in the whitlockite to merrillite dehydrogenation experiments of 486 Gopal and Calvo (1972) who used natural whitlockite, which would contain no entrained 487 reagents. 488

The apparent enrichment of Ca relative to phosphate during the whitlockite to merrillite transformation could take place through the formation of additional phases that either diffuse to

the surface or are dispersed and difficult to detect within the bulk crystals. This would be 491 consistent with the synchrotron diffraction data that indicate additional phases were present 492 within Fe/Mg-merrillites. An additional phase or phases, like the holtedahlite or Fe-bearing 493 phase indicated by synchrotron diffraction, would be required in order to take up extra Mg or Fe 494 made available during the transformation from whitlockite to merrillite. Synchrotron diffraction 495 data from the ferric Fe/Mg-merrillite heat treated open to atmosphere indicated holtedahlite as 496 the only additional crystalline phase beyond monetite, which would take up remaining Mg. It is 497 plausible that if the Fe is oxidizing during transformation as per Equation 1, an additional Fe-498 bearing phase may not necessarily form. However, in the case of the ferrous Fe/Mg-merrillite 499 treated in the SiO₂ sealed tube, both Fe- and Mg-bearing additional phases would be expected 500 and synchrotron diffraction data only indicate a possible Fe-bearing crystalline phase. It is 501 possible that an undetected amorphous Mg-phase is taking up the excess Mg, and that a similar 502 Fe-bearing phase may also form to take up any excess Fe in the Ferric merrillite that is not 503 accounted for by oxidation (Equation 1). Candidates for such a phase include Fe- and Mg-504 pyrophosphates $[X_2(P_2O_7)]$ where X is Mg or Fe], which can be produced in high temperature 505 metal/phosphate systems and can be amorphous (Bensalem et al., 1997; Lee et al., 2012). These 506 Fe- and Mg-pyrophosphate phases may also act as transitional phases – for example, a 507 hypothetical Mg-pyrophosphate may be a transitional phase to the holtedahlite observed by 508 synchrotron XRD (Equation. 2), which would explain why holtedahlite was detected in one 509 sample, but no Mg-containing phase in the other: 510

511
$$Mg_2(P_2O_7) + 2 H_2O \rightarrow Mg_2(PO_4)(OH) + H_3PO_4$$
 (2)

A general reaction reflecting the transformation from whitlockite to merrrillite may therefore include either holtedahlite or an amorphous metal pyrophosphate, and be represented by:

514

Whitlockite + Heat \rightarrow Merrillite + Monetite + Additional Phase(s) + H₃PO₄ + H₂O,

where non-merrillite phases take up any remaining chemistry resulting from the whitlockite-to-515 merrillite transformation. A balanced example that includes both monetite and pyrophosphate as 516 the metal phosphate is: 517 $24 \text{ Ca}_9 \text{X}(\text{PO}_3\text{OH})(\text{PO}_4)_6 \rightarrow 22 \text{ Ca}_{9.5} \text{X}(\text{PO}_4)_7 + 7 \text{ Ca}(\text{HPO}_4) + \text{X}_2(\text{P}_2\text{O}_7) + 5 \text{ H}_3\text{PO}_4 + \text{H}_2\text{O}$ (3). 518 519 Equations similar to 1, 2, or 3, or combinations thereof, are consistent with the observations of 520 521 this study and allow for a mechanism of charge balance that does not require entrained synthesis reagents. Such a mechanism would better explain the transformation of natural whitlockite to 522 merrillite confirmed by Gopal and Calvo (1972), where no entrained reagents from synthesis 523 were present. 524

525

526 Crystallography

The atomic parameters of Mg-whitlockite and Mg-merrillite synthesized by conceptually 527 similar methods to those used here have been previously measured by SC-XRD and discussed 528 (Gopal et al., 1974; Hughes et al., 2008), and therefore we only focus on differences between the 529 Fe- and Fe/Mg-whitlockite and merrillite structures determined in this work. The structures of 530 Fe-whitlockite and ferric Fe-merrillite have been previously refined using powder XRD 531 diffraction (Corlett and Keppler, 1966; Keppler, 1968; Lazoryak et al., 1996; Belik et al., 2002). 532 Belik et al. (2002) utilized neutron powder XRD and located two deuterium sites for Fe-533 whitlockite; hydrogen cannot be located by SC-XRD methods. Since significantly more 534 535 accuracy is possible with single crystal XRD, we report the structures of Fe-whitlockite and ferric Fe-merrillite here (Table 4). Atomic parameters of synthetic ferrous Fe-merrillite, and 536 mixed Fe/Mg-merrillite do not appear to be in the literature, and we present atomic coordinates 537 and full listing of bond distances as Tables A2 and A3 in deposit materials¹. A .cif file 538 containing all six crystal structures has also been deposited¹. 539

Broadly speaking, the six crystal structures determined (Fe- and Fe/Mg-whitlockite, as 540 well as ferrous and ferric Fe- and Fe/Mg-merrillite)(Table 4) are similar with differences related 541 mainly to the chemistry. The position of the proton has previously been established as 542 protonating the phosphate group located on the 3-fold axis (labeled P3 in this work). Protonated 543 phosphates point in the opposite direction compared to non-protonated phosphates. Modeling of 544 this disorder for the synthetic Fe-whitlockite found 89.0% of P3 on the inverted site; 100% 545 should be inverted if the structure were fully protonated. It is possible that some protons occupy 546 other sites in the structure, some Fe^{3+} is present, minor merrillite is present, or a combination of 547 the three. Small amounts of merrillite within natural whitlockite and synthetic Mg-whitlockite 548 have been previously noted (Hughes et al., 2008). 549 Heat treating the Fe-whitlockite in a Pt crucible open to the atmosphere yields an Fe-550 merrillite sample where 97.1% of the P3 are not on the inverted site, in agreement with 551 essentially full conversion to merrillite. This does not require 2.9% protonation – it is possible 552 that a fraction of the P3 phosphate remains inverted after deprotonation. The observed Fe - O553 distances (Table 3) are highly sensitive to the oxidation state of the iron. We find distances in 554 the Fe-whitlockite sample (Fe-O bond lengths of ~ 2.03 Å), and in the ferric merrillite sample 555 (Fe-O bond length of ~2.10 Å) that match expected metal – oxygen distances for Fe^{2+} and Fe^{3+} , 556 respectively. Note that the approximate difference of 0.1 Å between Fe for Fe^{2+} and Fe^{3+} is 557 larger by roughly a factor of 50 than standard deviations on these bond distances (Table 3). 558 Bond valence calculations are normally only an approximate method of identifying oxidation 559 state as bond distances can be influenced by a variety of structural features in a specific 560 compound. In this particular instance, bond valence calculations are more conclusive as the 561 crystal structures and compositions are very similar to each other. While the observed bond 562

563	distances as well as the bond valence calculations clearly suggest primarily Fe ²⁺ in the Fe-
564	whitlockite and ferrous merrillite, and primarily Fe ³⁺ in ferric merrillite, they only indicate which
565	oxidation state is predominant in each structure, not the relative amounts of Fe^{2+} and Fe^{3+} .
566	Heating the Fe-whitlockite samples in evacuated SiO ₂ tubes apparently prevented some
567	oxidation of Fe^{2+} to Fe^{3+} , as evidenced by little change in the measured Fe – O bond distances
568	(Table 3). Charge balance in the merrillite is maintained by the migration of Ca^{2+} cation sites to
569	around where protons are found in the whitlockite form. The Ca is distributed over two sites,
570	Ca4A and Ca4B, with refined occupancies of 14.8 and 29.6%, respectively. As it is not possible
571	from crystallography to assign elements to partially occupied sites such as these, we infer that
572	they are likely Ca based on its increase in microprobe data (Table 2) and the bonding
573	environments for both sites. Table 3 includes bond distances for these sites, all of which are
574	longer than expected for Fe^{2+} . In addition, the environments are quite irregular, which would be
575	unfavorable for d^6 Fe but not unusual for Ca^{2+} . As 50% occupancy of a divalent cation site
576	would provide sufficient charge balance to replace that lost by full dehydrogenation, these
577	occupancies are consistent with 88.8% conversion to merrillite. Of the P3 phosphates, 23.9% are
578	inverted, which is consistent with a modestly higher degree of protonation than suggested by the
579	Ca ²⁺ occupancies, though some of the P3 phosphate may have remained inverted after
580	deprotonation. The Ca ²⁺ occupancies of the Fe-whitlockite heat treated in a sealed SiO2 tube,
581	together with P3 observations, indicate ~80-90% conversion to merrillite. There was no
582	evidence of significant electron density on or around these sites in the Fe-whitlockite and Fe-
583	merrillite heat treated in a Pt crucible.
584	The mixed Fe/Mg-whitlockite sample refined to a similar overall structure as the pure Fe-

whitlockite. Refinement of the Fe/Mg ratio yielded a 63.7% Fe occupancy, with the balance Mg,

586	in general agreement with the microprobe results (Table 2). The Fe/Mg ratios refined for the two
587	heated samples refined to 75.4% (Pt crucible) and 64.6% (SiO ₂ tube). Inter-crystal standard
588	deviations on this ratio determined by microprobe were high, so it is quite possible that this value
589	varies some from crystal to crystal. P3 phosphates are inverted 84.0% of the time, suggesting the
590	whitlockite is 16% merrillite already, and the Fe – O distances are consistent with Fe^{2+} . The
591	sample heated in air ended up with 6.5% of the P3 site inverted, suggesting near complete
592	conversion to merrillite structure, and the Fe – O distances are consistent with Fe^{3+} .
593	Results of heating mixed Fe/Mg samples in sealed SiO ₂ tubes were also quite similar to
594	the pure Fe case. Fe – O distances are consistent with Fe^{2+} , and 30.9% of the P3 phosphate
595	group are inverted, indicating that the sample is largely merrillite. Two Ca sites with similar
596	environments and occupancies (9.8% and 26.4%) were also found.
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597 598	Implications for Martian Studies
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597598599600601	Implications for Martian Studies Merrillite is a potentially important indicator of martian geologic processes, including the evolution of halogen and water content over time in martian magmas. A deeper understanding of
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609	various chemistries and an understanding of the conditions under which these phases can be
610	formed, as presented here, will likely aid in facilitating a number of these studies as well as
611	further investigations into more fundamental thermodynamic and kinetic parameters of the
612	minerals. Complete dehydrogenation of ferrous merrillite in this study may have been hindered
613	by the sealed tube used. A flow through apparatus may eventually prove more successful, and
614	while Na ⁺ incorporation into merrillite may not be possible under hydrothermal conditions, a
615	high temperature approach may yield success. Future efforts in producing completely
616	dehydrogenated ferrous merrillite as well as incorporating Na ⁺ into synthetic merrillite are
617	warranted as the conditions under which these are achieved may reveal further insights into the
618	natural environments, martian or otherwise, in which these minerals form.
619	
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2/26

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636	
637	References
638	Adcock, C., Hausrath, E., and Forster, P. (2013) Readily available phosphate from minerals in early
639	aqueous environments on Mars. Nature Geoscience, 6(10), 824-827.
640	Belik, A., Izumi, F., Stefanovich, S.Y., Lazoryak, B., and Oikawa, K. (2002) Chemical and Structural
641	Properties of a Whitlockite-like Phosphate, Ca9FeD(PO4)7. Chemistry of Materials, 14(9), 3937-
642	3945.
643	Bensalem, A., Ahluwalia, M., Vijayaraghavan, T.V., and Ko, Y.H. (1997) Synthesis of amorphous
644	MgHPO4 \cdot x(R) [R = Ethanol; Ethylene glycol] in anhydrous media. Materials Research Bulletin,
645	32(11), 1473-1483.
646	Bergerhoff, G., and Brown, I. (1987) Crystallographic databases. International Union of Crystallography,
647	Chester, 77-95.
648	Brückner, J., Dreibus, G., Gellert, R., Squyres, S.W., Wänke, H., Yen, A., and Zipfel, J. (2008) Mars
649	Exploration Rovers: chemical composition by the APXS. In J. Bell, Ed. The Martian Surface -
650	Composition, Mineralogy, and Physical Properties, p. 58-100. Cambridge University Press,
651	Cambridge.
652	Corlett, M., and Keppler, U. (1966) Chemische Zusammensetzung des Whitlockits. Naturwissenschaften,
653	53(4), 105-105.
654	Dolomanov, O.V., Bourhis, L.J., Gildea, R.J., Howard, J.A., and Puschmann, H. (2009) OLEX2: a
655	complete structure solution, refinement and analysis program. Journal of Applied
656	Crystallography, 42(2), 339-341.
657	Dorozhkin, S.V., and Epple, M. (2002) Biological and medical significance of calcium phosphates.
658	Angewandte Chemie International Edition, 41(17), 3130-3146.
659	Downs, R.T., and Hall-Wallace, M. (2003) The American Mineralogist crystal structure database.
660	American Mineralogist, 88(1), 247-250.

2/26

- Dowty, E. (1977) Phosphate in Angra dos Reis: Structure and composition of the Ca3(PO4)2 minerals.
- Earth and Planetary Science Letters, 35(2), 347-351.
- Filiberto, J., and Treiman, A.H. (2009) The effect of chlorine on the liquidus of basalt: First results and
 implications for basalt genesis on Mars and Earth. Chemical Geology, 263(1-4), 60-68.
- ⁶⁶⁵ Frondel, C. (1941) Whitlockite: a new calcium phosphate Ca3 (PO4) 2. American Mineralogist, 26, 145-
- 666 152.
- 667 Frondel, C., and Prien, E.L. (1946) Deposition of Calcium Phosphates Accompanying Senile

668 Degeneration and Disease. Science, 103(2672), 326.

- ⁶⁶⁹ Fuchs, L.H. (1962) Occurrence of Whitlockite in Chondritic Meteorites. Science, 137(3528), 425-426.
- 670 Goetz, W., Bertelsen, P., Binau, C.S., Gunnlaugsson, H.P., Hviid, S.F., Kinch, K.M., Madsen, D.E.,
- Madsen, M.B., Olsen, M., and Gellert, R. (2005) Indication of drier periods on Mars from the
 chemistry and mineralogy of atmospheric dust. Nature, 436(7047), 62-65.
- Gopal, R., and Calvo, C. (1972) Structural Relationship of Whitlockite and βCa3(PO4)2. Nature, 237(71),
 30-32.
- 675 Gopal, R., Calvo, C., Ito, J., and Sabine, W.K. (1974) Crystal-Structure of Synthetic Mg-Whitlockite,
- 676 Ca18Mg2H2(PO4)14. Canadian Journal of Chemistry-Revue Canadienne De Chimie, 52(7),
 677 1155-1164.
- Greenwood, J.P., and Blake, R.E. (2006) Evidence for an acidic ocean on Mars from phosphorus
 geochemistry of Martian soils and rocks. Geology, 34(11), 953-956.
- 680 Gross, J., Filiberto, J., and Bell, A.S. (2013a) Water in the martian interior: Evidence for terrestrial
- MORB mantle-like volatile contents from hydroxyl-rich apatite in olivine–phyric shergottite
 NWA 6234. Earth and Planetary Science Letters.
- 683 Gross, J., Filiberto, J., Herd, C.D., Daswani, M.M., Schwenzer, S.P., and Treiman, A.H. (2013b)
- 684 Petrography, mineral chemistry, and crystallization history of olivine-phyric shergottite NWA
- 685 6234: A new melt composition. Meteoritics & Planetary Science.
- Gulick, A. (1955) Phosphorus as a Factor in the Origin of Life. American Scientist, 43(3), 479-489.

- Hamad, M., and Heughebaert, J.-C. (1986) The growth of whitlockite. Journal of Crystal Growth, 79(1),
 192-197.
- Hammersley, A., Svensson, S., Hanfland, M., Fitch, A., and Hausermann, D. (1996) Two-dimensional
 detector software: from real detector to idealised image or two-theta scan. International Journal of
- 691 High Pressure Research, 14(4-6), 235-248.
- Hausrath, E., Golden, D., Morris, R., Agresti, D., and Ming, D. (2013) Acid sulfate alteration of
- 693 fluorapatite, basaltic glass and olivine by hydrothermal vapors and fluids: Implications for
- 694 fumarolic activity and secondary phosphate phases in sulfate-rich Paso Robles soil at Gusev
- 695 Crater, Mars. Journal of Geophysical Research: Planets.
- Hausrath, E.M., and Tschauner, O. (2013) Natural Fumarolic Alteration of Fluorapatite, Olivine, and
- Basaltic Glass, and Implications for Habitable Environments on Mars. Astrobiology, 13(11),
 1049-1064.
- Hughes, J.M., Jolliff, B.L., and Gunter, M.E. (2006) The atomic arrangement of merrillite from the Fra
 Mauro Formation, Apollo 14 lunar mission: The first structure of merrillite from the Moon.
 American Mineralogist, 91(10), 1547-1552.
- Hughes, J.M., Jolliff, B.L., and Rakovan, J. (2008) The crystal chemistry of whitlockite and merrillite and
 the dehydrogenation of whitlockite to merrillite. American Mineralogist, 93(8-9), 1300-1305.
- Jarcho, M., Salsbury, R., Thomas, M., and Doremus, R. (1979) Synthesis and fabrication of β-tricalcium
- phosphate (whitlockite) ceramics for potential prosthetic applications. Journal of Materials
 Science, 14(1), 142-150.
- Jarosewich, E. (2002) Smithsonian microbeam standards. JOURNAL OF RESEARCH-NATIONAL

- Jolliff, B.L., Haskin, L.A., Colson, R.O., and Wadhwa, M. (1993) Partitioning in REE-saturating
- 710 minerals: Theory, experiment, and modelling of whitlockite, apatite, and evolution of lunar
- residual magmas. Geochimica et Cosmochimica Acta, 57(16), 4069-4094.

712	Jolliff, B.L.	, Hughes, J.M.	Freeman, J.J.	, and Zeigler, R.A	. (2006) Cr	vstal chemistr	y of lunar merrillite
	, , , , , , , , , , , , , , , , , , , ,	, , , ,	, , , ,	, , ,			J

- and comparison to other meteoritic and planetary suites of whitlockite and merrillite. American
 Mineralogist, 91(10), 1583-1595.
- Keppler, U. (1968) Structural investigation of calciumphosphate and isotypic structures. Bulletin de la
 Société Chimique de France.
- Lagier, R., and Baud, C.-A. (2003) Magnesium whitlockite, a calcium phosphate crystal of special
- interest in pathology. Pathology-Research and Practice, 199(5), 329-335.
- Lane, M.D., Bishop, J.L., Darby Dyar, M., King, P.L., Parente, M., and Hyde, B.C. (2008) Mineralogy of
 the Paso Robles soils on Mars. American Mineralogist, 93(5-6), 728-739.
- 721 Lazoryak, B., Morozov, V., Belik, A., Khasanov, S., and Shekhtman, V.S. (1996) Crystal Structures and
- Characterization of Ca 9 Fe (PO 4) 7 and Ca 9 FeH 0.9 (PO 4) 7. Journal of Solid State
- 723 Chemistry, 122(1), 15-21.
- Le Bail, A. (2005) Whole powder pattern decomposition methods and applications: a retrospection.
 Powder Diffraction, 20(4), 316-326.
- Lee, G.-H., Seo, S.-D., Shim, H.-W., Park, K.-S., and Kim, D.-W. (2012) Synthesis and Li electroactivity
- of Fe2P2O7 microspheres composed of self-assembled nanorods. Ceramics International, 38(8),
 6927-6930.
- 729 McCubbin, F.M., and Nekvasil, H. (2008) Maskelynite-hosted apatite in the Chassigny meteorite:
- Insights into late-stage magmatic volatile evolution in martian magmas. American Mineralogist,
 93(4), 676-684.
- McGuire, A.V., Francis, C.A., and Dyar, M.D. (1992) Mineral standards for electron microprobe analysis
 of oxygen. American Mineralogist, 77, 1087-1087.
- McSween, H., and Treiman, A.H. (1998) Martian Meteorites. In J. Papike, Ed. Planetary Materials, p. F1-
- F53. Mineral Society of America, Washington, DC.

736	McSween, H.Y., Eisenhour, D.D., Taylor, L.A., Wadhwa, M., and Crozaz, G. (1996) QUE94201
737	shergottite: Crystallization of a Martian basaltic magma. Geochimica et Cosmochimica Acta,
738	60(22), 4563-4569.
739	Merrill, G.P. (1915) On the Monticellite-Like Mineral in Meteorites, and on Oldhamite as a Meteoric
740	Constituent. Proceedings of the National Academy of Sciences, 1(5), 302-308.
741	(1917) On the calcium phosphate in meteoric stones. American Journal of Science(256), 322-324.
742	Orlova, A.I., Khaĭnakov, S., Loginova, E., Oleneva, T., Granda, S.G., and Kurazhkovskaya, V.S. (2009)
743	Calcium thorium phosphate (Whitlockite-type mineral). Synthesis and structure refinement.
744	Crystallography Reports, 54(4), 591-597.
745	Patiño Douce, A.E., and Roden, M. (2006) Apatite as a probe of halogen and water fugacities in the
746	terrestrial planets. Geochimica et Cosmochimica Acta, 70(12), 3173-3196.
747	Patiño Douce, A.E., Roden, M.F., Chaumba, J., Fleisher, C., and Yogodzinski, G. (2011) Compositional
748	variability of terrestrial mantle apatites, thermodynamic modeling of apatite volatile contents, and
749	the halogen and water budgets of planetary mantles. Chemical Geology, 288(1), 14-31.
750	Rowles, S. (1968) The precipitation of whitlockite from aqueous solutions. Bull Soc Chim Fr, 1968, 802.
751	Rubin, A.E. (1997) Mineralogy of meteorite groups: An update. Meteoritics & Planetary Science, 32(5),
752	733-734.
753	Ruszala, F., and Kostiner, E. (1980) The hydrothermal synthesis and crystal growth of various
754	whitlockites and a manganese containing graftonite. Journal of Crystal Growth, 48(3), 473-474.
755	Schwartz, A.W. (2006) Phosphorus in Prebiotic Chemistry. Philosophical Transactions: Biological
756	Sciences, 361(1474), 1743-1749.
757	Shearer, C., Papike, J., Burger, P., Sutton, S., McCubbin, F., and Newville, M. (2011a) Direct
758	determination of europium valence state by XANES in extraterrestrial merrillite: Implications for
759	REE crystal chemistry and martian magmatism. American Mineralogist, 96(8-9), 1418-1421.

- Shearer, C.K., Burger, P.V., Papike, J.J., Sharp, Z.D., and McKeegan, K.D. (2011b) Fluids on
- differentiated asteroids: Evidence from phosphates in differentiated meteorites GRA 06128 and
 GRA 06129. Meteoritics & Planetary Science, 46(9), 1345-1362.
- Sheldrick, G.M. (2007) A short history of SHELX. Acta Crystallographica Section A: Foundations of
 Crystallography, 64(1), 112-122.
- Skowron, A., and Brown, I. (1990) Refinement of the structure of robinsonite, Pb4Sb6S13. Acta
- 766 Crystallographica Section C: Crystal Structure Communications, 46(4), 527-531.
- Terada, K., Monde, T., and Sano, Y. (2003) Ion microprobe U-Th-Pb dating of phosphates in martian
- meteorite ALH 84001. Meteoritics & Planetary Science, 38(11), 1697-1703.
- 769 Teterskii, A., Stefanovich, S.Y., Lazoryak, B., and Rusakov, D. (2007) Whitlockite solid solutions Ca9-
- x M x R (PO4) 7 (x= 1, 1.5; M= Mg, Zn, Cd; R= Ln, Y) with antiferroelectric properties. Russian
- Journal of Inorganic Chemistry, 52(3), 308-314.
- Treiman, A.H. (2003) Chemical compositions of martian basalts (shergottites): Some inferences on b;
- formation, mantle metasomatism, and differentiation in Mars. Meteoritics & Planetary Science,
 38(12), 1849-1864.
- Westheimer, F.H. (1987) Why Nature Chose Phosphates. Science, 235(4793), 1173-1178.
- Wherry, E. (1917) Merrillite, meteoritic calcium phosphate. American Mineralogist, 2, 119.
- White, R. (1990) Chromatography/Fourier transform infrared spectroscopy and its applications. CRC
 Press.
- Wills, A.S. (1999) VaList-bond valence calculation and listing Program, www.ccp14.ac.uk
- Xie, X., Minitti, M.E., Chen, M., Mao, H.-k., Wang, D., Shu, J., and Fei, Y. (2002) Natural high-pressure
- 781 polymorph of merrillite in the shock veins of the Suizhou meteorite. Geochimica et
- 782 Cosmochimica Acta, 66(13), 2439-2444.

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Figure Captions

Figure 1. Optical images of synthesized phases. A) Mg-whitlockite single crystals with a 785 786 polycrystalline aggregate in the upper left corner. **B**) Fe-whitlockite single crystal. Dark color inside crystal suggests some intra-crystalline iron film on growth faces. C) Fe-whitlockite polycrystalline 787 788 aggregates. Darker colors in some aggregates from inter-crystalline iron films. **D**) Mixed Fe/Mgwhitlockite single crystals. E) Mg-merrillite showing the hazy surface texture due to a thin film deposit 789 790 post heat treatment. The extent of coverage appeared to decrease with decreasing grain size (i.e. smaller 791 crystals exhibited less complete or thinner coverage). F) Ferric Fe-merrillite crystals showing the 792 incomplete hazy surface texture post heat treatment on some crystals (arrows). Deposits on the minerals rarely covered the entire crystalline surface. "White crystals" in this image are polycrystalline aggregates. 793 794 Figure 2. Scanning Electron Microscope images of synthesized materials. A) SEM SEI image 795 of synthesized Mg-whitlockite. B) SEM SEI image of synthetic twinned Fe-whitlockite. C) SEM SEI image of synthesized Fe/Mg-whitlockite. D) SEM SEI with inset (taken at 3000x) showing coating 796 797 developed on Mg-merrillite crystal during synthesis. Most Mg-merrillite crystals were completely 798 covered, although smaller grains exhibited less complete or thinner coverage. E) SEM SEI image of Ferric Fe-merrillite showing coating developed during synthesis. In contrast to Mg-merrillite, the deposit 799 800 on the mineral rarely covered the entire crystalline surface and in some cases crystal showed no coating, consistent with discussion of our proposed mechanisms. F) SEM SEI image of ferrous Fe-merrillite 801 showing coating developed during synthesis. 802

Figure 3. IR Spectra of A) Fe-whitlockite, ferrous merrillite treated in a SiO₂ tube, and ferric merrillite treated open to atmosphere. B) Fe/Mg-whitlockite, ferrous Fe/Mg-merrillite heat treated in a SiO₂ tube, and ferric Fe/Mg-merrillite heat treated open to atmosphere. Whitlockite data show an apparent absorption from 3330 to 2570 cm⁻¹ where a broad O-H stretch feature exists. The absorbance is mainly absent in material treated in a Pt crucible exposed to air (ferric Fe- and Fe/Mg-merrillite), with some absorption in the O-H stretching band present in material treated in a sealed SiO₂ tube (ferrous Fe- and Fe/Mg-merrillite) including a peak at 2380 cm⁻¹. Spectra offset for clarity.

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811	Figure 4. Synchrotron X-ray diffraction. A) Observed and modeled patterns of the red-colored
812	ferric Fe/Mg-merrillite sample along with the residual of the fit. Beyond the expected
813	whitlockite/merrillite observed in the crystal, monetite and holtedahlite also appear to be present. B)
814	Similar data for a dark-colored ferrous merrillite crystal with whitlockite/merrillite pattern data masked
815	out. This crystal also contained monetite and an additional unconfirmed phase with structure similar to a
816	ferrate.

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Table 1. Masses and volumes used in finalized whitlockite synthesis methods

Synthetic Mineral	MgNO₃ (g)	FeS (g)	NaCl (g) ^a	$HAP\left(g ight)^{b}$	Total solid Mass (g)	H₂O (ml)	Solids in Solution (g/L)
Mg-whitlockite	0.300 (±0.003)	-	-	1.000 (±0.003)	1.300	90	14.4
Fe-whitlockite	-	0.055 (±0.005)		1.000 (±0.003)	1.055	90	11.7
Fe/Mg-whitlockite	0.050 (±0.003)	0.050 (±0.005)	0.067 (±0.005)	1.000 (±0.003)	1.167	90	13.0
^a Later experiments showed NaCl to have no effect on synthesis. ^b HAP = hydroxyapatite.							

Tables

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Table 2. Chemical analyses of whitlockite and merrillite by electron microprobe in wt%.

	Mg-v	vhitlockite	Mg-merrillite				
CaO	47.61	(0.20)	48.89	(0.21)			
P_2O_5	46.85	(0.27)	46.38	(0.32)			
MgO	3.54	(0.14)	4.06	(0.12)			
FeO	0.02	(0.03)	0.02	(0.03)			
Na₂O	B.D.		B.D.				
^ª H₂O	0.86		-				
Total	98.02	(0.34)	99.35	(0.36)			
Ν		17		16			
Stoic.	Ca _{9.0} Mg _{0.9}	9(PO ₃ OH)(PO ₄) ₆	Ca _{9.4} M	g _{1.1} (PO ₄) _{7.0}			
			F	erric	Fe	rrous	
	Fe-w	hitlockite	Fe-m	nerrillite	Fe-merrillite		
			(Pt cruc.)		(SiO ₂ tube)		
CaO	46.00	(0.45)	46.37	(0.27)	47.00	(0.19)	
P_2O_5	45.26	(0.47)	46.02	(0.50)	45.42	(0.46)	
MgO	0.14	(0.03)	0.34	(0.07)	0.39	(0.10)	
^b Fe ₂ O ₃	-		6.74	-	-		
FeO	6.75	(0.27)	-		6.85	(0.18)	
Na₂O	B.D		B.D.		B.D.		
^ª H₂O	0.86		-		-		
Total	98.15	(0.63)	99.47	(0.45)	99.65	(0.53)	
N		18		17	20		
Stoic.	Ca _{9.0} Mg _{0.9}	9(PO ₃ OH)(PO ₄) ₆	Ca _{9.0} Fe _{0.9}	Mg _{0.1} (PO ₄) _{7.0}	Ca _{9.2} Fe _{1.0}	Mg _{0.1} (PO ₄) _{7.0}	
				Ferric		Ferrous	
	Fe/Mg-whitlockite		Fe/Mg-merrillite (Pt cruc)		Fe/Mg-merrillite (SiOs tube)		
CaO	46.54	(0.59)	46.79	(0.33)	47.09	(0.22)	
P ₂ O ₅	45.83	(0.36)	45.96	(0.39)	45.40	(0.38)	
MgO	1.00	(0.20)	1.29	(0.09)	1.20	(0.16)	
b						• •	

^b Fe₂O₃	-		4.98	-	-		
FeO	4.62	(0.71)	-		5.68	(0.23)	
Na₂O	B.D.		B.D.		B.D.		
^ª H₂O	0.86		-		-		
Total	97.99	(0.49)	99.02	(0.60)	99.36	(0.48)	
n	15		13		19		
Stoic.	Ca _{9.0} Fe _{0.7} Mg _{0.3} (PO ₃ OH)(PO ₄) ₆		Ca _{9.1} Fe _{0.7} I	Mg _{0.3} (PO ₄) _{7.0}	Ca _{9.2} Fe _{0.9} Mg _{0.3} (PO ₄) _{7.0}		

Parenthetical values are 1 standard deviation. n = number of analyses averaged chemistry is based on. B.D. = below detection.

 a H₂O is based on ideal whitlockite (Hughes et al., 2008) and is not included in the EMP totals.

 b^{b} Fe was measured assuming Fe^{2+} and recalculated to Fe^{3+} with total and resulting stoichiometry adjusted

830 accordingly. For ferric Fe-merrillite, the original FeO wt. % = 6.07 with a standard deviation of 0.13, analysis total

831 of 98.80, and calculated stoichiometry of $Ca_{9.1}Fe_{0.9}Mg_{0.1}(PO_4)_{7.0}$. For Ferric Fe/Mg-merrillite the original FeO wt. % =

4.48 with a standard deviation of 0.08, analysis total of 98.53, and calculated stoichiometry of

833 $Ca_{9.2}Fe_{0.7}Mg_{0.4}(PO_4)_{7.0}$. NOTE: It is not possible by microprobe alone to actually determine Fe^{2+}/Fe^{3+} content and 834 these values are estimates.

Table 3. Select bond lengths discussed in text for synthetic minerals Å.

		Fe- whitlockite	Fe/Mg- whitlockite	Ferric Fe- merrillite (Pt cruc.)	Ferrous Fe- merrillite (SiO ₂ tube)	Ferric Fe/Mg- merrillite (Pt cruc.)	Ferrous Fe/Mg- merrillite (SiO ₂ tube)
Fe(1)-	O(4)	2.1131(16)	2.1067(13)	2.046(2)	2.122(3)	2.048(2)	2.104(2)
	O(4)#1	2.1130(16)	2.1068(13)	2.046(2)	2.122(3)	2.048(2)	2.104(2)
	O(4)#2	2.1130(16)	2.1068(13)	2.046(2)	2.122(3)	2.048(2)	2.104(2)
	O(5)	2.0893(16)	2.0806(12)	2.011(2)	2.108(3)	2.023(2)	2.092(2)
	O(5)#1	2.0893(16)	2.0806(12)	2.011(2)	2.108(3)	2.023(2)	2.092(2)
	O(5)#2	2.0893(16)	2.0806(12)	2.011(2)	2.108(3)	2.023(2)	2.092(2)
Ca(4A)-	O(6)#18	NA	NA	NA	2.736(5)	NA	2.758(6)
	O(6)#19	NA	NA	NA	2.736(5)	NA	2.758(6)
	O(7)#18	NA	NA	NA	2.541(4)	NA	2.541(4)
	O(7)#19	NA	NA	NA	2.541(4)	NA	2.541(4)
Ca(4B)-	O(6)#18	NA	NA	NA	2.516(3)	NA	2.518(2)
	O(6)#19	NA	NA	NA	2.516(3)	NA	2.518(2)
	O(7)#18	NA	NA	NA	2.846(5)	NA	2.851(4)
	O(7)#19	NA	NA	NA	2.846(5)	NA	2.851(4)

837 Parenthetical values are standard deviations. "Pt cruc." denotes merrillite produced in a platinum crucible in open

air. "SiO₂ Tube" denotes merrillite produced using a triple argon purged and seal glass tube. Only Fe-O and Ca-O

839 bonds discussed in text appear in table. A full table of bond lengths appears in deposit item Table A4.

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Mineral	Fe- whitlockite	Fe/Mg-whitlockite	Ferric Fe- merrillite (Pt cruc.)	Ferrous Fe- merrillite (SiO ₂ tube)	Ferric Fe/Mg- merrillite (Pt cruc.)	Ferrous Fe/Mg-merrillite (SiO2 tube)
Emp. formula	Ca ₉ FeHO ₂₈ P ₇	Ca ₉ Fe _{0.64} HMg _{0.36} O ₂₈ P ₇	Ca ₉ FeH _{0.02} O ₂₈ P ₇	Ca _{9 44} FeH _{0 12} O ₂₈ P ₇	Ca ₉ Fe _{0.75} Mg _{0.25} O ₂₈ P _{.7}	Ca _{9.36} Fe _{0.65} H _{0.28} Mg _{0.35} O ₂₈ P ₇
Formula wt.	1082.37	1070.93	1081.38	1099.12	1073.59	1084.98
a (Å)	10.3510(7)	10.3499(4)	10.3278(11)	10.3453(6)	10.3301(13)	10.3392(7)
c (Å)	37.059(2)	37.0715(16)	37.050(4)	37.118(2)	37.062(5)	37.081(2)
Volume (Å ³)	3438.6(4)	3439.1(2)	3422.4(8)	3440.3(5)	3425.1(10)	3432.8(5)
Density (calc.) (Mg/m ³)	3.136	3.103	3.148	3.183	3.123	3.149
Reflections collected	17516	17609	17589	11475	17501	17513
Independent reflections	2352	2355	2334	2238	2338	2342
R(int)	0.0221	0.0202	0.0248	0.0236	0.0249	0.0213
parameters	140	142	140	146	143	153
Goodness-of-fit on F2	1.148	1.085	1.095	1.078	1.053	1.085
Final R indices [I>2sigma(I)]						
R1	0.0185	0.0154	0.0172	0.0204	0.0171	0.0157
wR2 R indices (all	0.0552	0.041	0.0454	0.0517	0.0483	0.0411
data)						
R1	0.0185	0.0155	0.0175	0.0211	0.0178	0.016
wR2	0.0553	0.0411	0.0455	0.0521	0.0489	0.0414
Largest diff. e Å ⁻³						
peak	0.361	0.315	0.372	0.590	0.787	0.472
hole	-1 913	-0.956	-1 092	-1 032	-0 366	-0 425

841 **Table 4.** Single crystal X-ray diffraction refinement statistics

842 Note: 5.1° to 61.0° 2 Θ range, temperature 100 K, space group R3c. Parenthetical values are standard deviations.



- 845 846
 - Figure 1.

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852853 Figure 4.