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2	Visible and Short Wave Infrared Reflectance Spectroscopy of REE Fluorocarbonates
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12	А р стр а <i>с</i> т
15	ADSTRACT
14	The mineralogy of rare earth element deposits is critical in understanding their petrogenesis and has
15	significant implications for their economic viability. Lanthanide-bearing compounds are known to
16	produce sharp absorption features in the visible to short wave infrared region (VIS-SWIR), however, a
17	significant knowledge gap exists between the fields of hyperspectral reflectance spectroscopy and rare
18	earth element mineralogy. Reflectance spectra were collected from four bastnaesite samples, two parisite
19	samples and one synchysite sample from the visible into the shortwave infrared. These REE
20	fluorocarbonate mineral samples were characterized via scanning electron microscopy and electron
21	probe microanalysis. Sharp absorptions of REE-bearing minerals are mostly the result of 4 <i>f</i> -4 <i>f</i>
22	intraconfigurational electron transitions and for the light REE-enriched fluorocarbonates, the bulk of the
23	features can be ascribed to Nd ³⁺ , Pr ³⁺ , Sm ³⁺ and Eu ³⁺ . The lanthanide-related spectral responses of the
24	REE fluorocarbonates are consistent across the group, supporting the notion that the REE cation site is

Page 2 of 30

25	very similar in each of these minerals. Carbonate-related spectral responses differed between these
26	minerals, supporting the notion that the crystallographic sites for the carbonate radical differ between
27	bastnaesite, synchysite and parisite. Exploitable spectral differences include a distinct absorption band at
28	2243 nm that separates bastnaesite from synchysite and parisite. Similarly, for bastnaesite a dominantly
29	Pr^{3+} related absorption band located is at 1968 nm while in synchysite and parisite it occurs at 1961 nm.
30	
31	Keywords: Mineral spectroscopy, rare earth elements, hyperspectral, bastnaesite, parisite, synchysite
32	
33	INTRODUCTION
34	The mineralogy of rare earth element deposits is critical in understanding their petrogenesis and has
35	significant implications for their economic viability. Rapid determination of ore modal mineralogy for
36	these deposit types by hyperspectral reflectance spectroscopy would provide immediate feedback on the
37	strength, type and relevance of mineralization. Lanthanide-bearing compounds are known to produce
38	sharp absorption features in the visible to short wave infrared region (VIS-SWIR) and have been
39	conventionally viewed as unchanging features in the field of remote sensing (e.g., Clark 1999); however,
40	a significant knowledge gap exists between the fields of hyperspectral reflectance spectroscopy and rare

41 earth element mineralogy.

42

The lanthanides (*Ln*) are a series of 15 elements belonging to the *4f* block of the periodic table, and from lightest to heaviest are lanthanum (La), cerium (Ce), praseodymium (Pr), neodymium (Nd), promethium (Pm), samarium (Sm), europium (Eu), gadolinium (Gd), terbium (Tb), dysprosium (Dy), holmium (Ho), erbium (Er), thulium (Tm), ytterbium (Yb) and lutetium (Lu). In geological environments the lanthanides are commonly found together in their trivalent state, with the exception of divalent europium

Page 3 of 30

(Eu²⁺) and sometimes tetravalent cerium (Ce⁴⁺). The term "rare earth elements" (REE) comprises the
lanthanides, commonly yttrium (Y) and sometimes scandium (Sc) due to similar chemical
characteristics. Coordination states of the REE cations in the assorted mineral structures can be quite
variable (e.g., Miyawaki and Nakai 1996), ranging from a reasonably symmetrical 8-coordinated site in
xenotime, to britholite with two distinct sites with coordination numbers of 7 and 9, and the REE
fluorocarbonates with multiple bonding anions (oxygen and fluorine) with a coordination number of 9
(6O+3F).

55

56 The mineralogy of REE ore deposits is important because the metallurgy of only certain REE-bearing 57 phases is well established. Namely, most global REE production is derived from REE fluorocarbonate 58 minerals (e.g., bastnaesite, parisite and synchysite), REE phosphate minerals (e.g., monazite and 59 xenotime) and from ion-absorbed clays, however, recent advances have been made in the treatment of 60 REE silicate mineral ore (e.g., eudialyte and britholite) (e.g., Mariano and Mariano 2012). Furthermore, 61 REE deposits can show a complex mineral paragenesis with REE-mineral assemblages varying 62 significantly across a single geological system. Consequently, there exists significant potential for 63 reflectance spectroscopy to play an important role in the exploration and exploitation of REE ore 64 deposits.

65

The objectives of this study are to present mineralogical characterization and reflectance spectra in the visible to short wave infrared region of the main rare earth element fluorocarbonate minerals; bastnaesite, parisite and synchysite. Following mineralogical and spectroscopic background of REE bearing minerals we provide a band registry for these minerals as well as interpretations of spectral absorption features related to the lanthanides. The spectral features of the registry will be the focus of

Page 4 of 30

further study in the development of hyperspectral reflectance imaging spectroscopy to carry out REE
 mineral identification, REE mineral abundance estimates and rare earth element abundance estimates.

74

GENERAL SPECTROSCOPY OF THE LANTHANIDES

75 The outer radius of the 4f electron shells (~ 0.3 Å) for the lanthanides is much less than that of their filled 5s and 5p shells (~2 Å, ~1 Å). It can then be approximated that the local electronic environment of Ln^{3+} 76 77 cations interacts primarily with those outer shells, leaving the 4f electrons 'relatively sheltered' but not 78 completely non-participatory in bonding (e.g., Liu, 2005). Electrostatic repulsion of the base ion generates a first level splitting of spectroscopic states, ^{2S+1}L (e.g., ⁵I). Next, spin-orbit coupling splits 79 these into multiplets, or "J-levels" (e.g., ${}^{5}I_{8}$), and once placed into a crystal field the J-levels are then 80 81 split into "Stark Sublevels". Crystal field interactions for the Ln cation include variables such as ligand 82 type, coordination number and polyhedron asymmetry which all play a role in the location and intensity 83 of energy levels and the associated absorption (Görller-Walrand and Binnemans, 1998). Each of the 84 resulting sublevels provides the potential for promotion of a relaxed electron into an excited state, giving 85 rise to absorption of electromagnetic radiation (e.g., light) at a specific energy level (e.g., wavelength). 86

The spectroscopy of REE-bearing phases is well established in the fields of physics and chemistry,
however, the well understood principles and well studied doped compounds do not lend to direct
translation into mineralogy and hyperspectral remote sensing. For example, the Dieke Diagram (Dieke *et al.* 1968) details intraconfigurational *4f-4f* transitions for 'free ions' as deduced through studies of
largely mono-lanthanide synthetic compounds, however, the transitions shown do not include splitting of
energy levels due to a crystal field nor the complexities of naturally occurring REE minerals with
variable REE distributions and other elemental substitutions. Consequently, this diagram and other band

Page 5 of 30

- 94 registries can only act as proxies to help identify origins of absorption features in reflectance spectra95 from minerals.
- 96

97 In general, and excluding other physical parameters such as grain size, the strength of absorption 98 features by the lanthanides will be primarily a function of the concentration of the ion as well as the 99 specific absorptivity of that ion's intraconfigurational transitions within a given crystal structure. The 100 location of lanthanide-related absorption features will be primarily a function of the cation's specific 101 coordination and asymmetry in the host crystal structure.

102 **REVIEW OF REE-RELATED REFLECTANCE SPECTROSCOPY STUDIES**

103 Few studies have been published addressing rare earth element bearing minerals and rocks in the field of 104 reflectance spectroscopy. Surveys of carbonatites and alkaline intrusive complexes in the literature can 105 be divided into coarse spectral resolution spaceborne studies (e.g., Oppenheimer 1998, Rowan and Mars 106 2003, Mars and Rowan 2011), fine spectral resolution airborne studies (e.g., Crowley et al. 1988, Rowan 107 et al. 1995, Bowers and Rowan 1996, Bedini 2009), and fine resolution ground and lab-based studies 108 (e.g., Rowan et al. 1986, McHugh et al. 2000). The early work by Rowan, Crowley and Mars recognized 109 the potential of exploiting REE absorption features with hyperspectral imaging, however, focus of their 110 future research shifted to other geological systems, use of ASTER multispectral satellite imagery, and 111 planetary remote sensing.

112

The studies with coarse resolution used ASTER imagery and were focused on lithological mapping via endmember extraction and band ratios. The general results are that lithologies of carbonatite complexes can be identified with moderate confidence using ASTER satellite data from areas with well-established geological maps (e.g., Khanneshin, Mountain Pass, Ol Doynio Lengai). Remote sensing using the finer

Page 6 of 30

117 resolution AVIRIS airborne system shows improved classification of rock types and even interpreted 118 detection of REE from Mountain Pass (e.g., Crowley et al., 1988). Bedini (2009) investigated the REE-119 mineralized Sarfatoq carbonatite complex of Greenland using the HyMap airborne hyperspectral imager 120 focusing on the classification of rock type and predictive mapping, however, there was no mention of 121 any specific REE phases being detected from the airborne platform. Similar results were achieved by 122 Bowers and Rowan (1996) at the Ice River Complex using AVIRIS data. Ground and laboratory based 123 studies by Rowan et al. (1986) of four carbonatite hosted rare metal deposits focused on rock type 124 classification and were able to discriminate between several carbonatite phases, as well as identify the sharp absorption features caused by Nd³⁺ in field samples. This earliest work is referenced in all of the 125 126 above literature and has served as a base for the spectral classification of rock types at carbonatite and 127 alkaline intrusive igneous complexes.

128

129 Early infrared spectroscopy investigations of carbonate minerals by Adler and Kerr (1963) included 130 bastnaesite and parisite because of their distinct carbonate arrangements, but did not investigate the 131 influence of the lanthanides and stopped short of the shortwave infrared region. The benchmark mineral 132 spectroscopy paper of Hunt (1977) lists monazite (misspelled as monzonite) in his tabulation of common minerals and respective spectral signatures, however, La^{2+} is listed as the origin of absorptions but La 133 134 would have 3+ valence in monazite and would therefore not have spectral features in the VIS-SWIR 135 (e.g., Liu 2005). Clark (1999) covers several REE oxides in his Spectroscopy of Rocks and Minerals 136 review and recognizes that the patterns seen in REE minerals are a combination of several lanthanides 137 but states that absorptions are independent of mineralogy. Reflectance spectra in the VNIR-SWIR of 138 bastnaesite and parisite with limited discussions have been presented by Kerr et al. (2011) and Morin-Ka

Page 7 of 30

(2012), with both publications stating that REE signatures are present in these REE minerals and thatmore detailed documentation is warranted.

141

142 CRYSTAL STRUCTURE REVIEWS OF BASTNAESITE, PARISITE AND SYNCHYSITE

143 Bastnaesite, parisite and synchysite are rare earth element fluorocarbonate minerals of economic

significance in many REE deposits and occurrences. Bastnaesite, CeCO₃F, is the most commonly

145 reported of these minerals, however, it is also the most familiar of the group. Synchysite, CaCe(CO₃)₂F,

146 is the most Ca-enriched of the REE fluorocarbonates, sometimes forms euhedral prismatic crystals and

147 is the second most reported. Parisite, CaCe₂(CO₃)₃F₂, can form distinct doubly terminated pyramidal

148 crystals and is the least reported of these rare earth fluorocarbonate minerals. Rontgenite,

149 Ca₂Ce₃(CO₃)₅F₃, is another rare REE fluorocarbonate mineral with a similar structure but was not

addressed in this study.

151

152 Total rare earth oxide (REE_2O_3) content increases from ~52 wt% in Ca-rich synchysite, to ~60 wt% in

parisite and up to ~75 wt% in Ca-absent bastnaesite. These minerals show preference for the light rare

earth elements (LREE) and have similar slopes and trends in Chondrite-normalized diagrams. These

three minerals are commonly found together, are structurally related, can show syntaxial intergrowth and

are also commonly found together in secondary mineral mixtures.

157

158 Structurally, each of the phases can be assembled using a common set of building blocks (Ni *et al.*

159 1993), which include layers of CeF, CO₃, and Ca. On an atom basis, bastnaesite is Ca-absent, parisite

160 shows REE:Ca=2:1 and synchysite shows REE:Ca=1:1. Accordingly, the stacking order of the blocks

161 gives rise to the different compositional proportions and mineral species (Fig. 1). Polytypism has been

Page 8 of 30

162	documented for these minerals (e.g., Meng et al. 2001). It is important to note that all of the REE will
163	occur in the Ce site but only Ce will be listed for brevity. Table 1 documents some basic crystallographic
164	data for the REE fluorocarbonates.
165	
166	In bastnaesite Ni et al. (1993) described the hexagonal crystal structure as being built by layers of REE-
167	F alternating with layers of CO_3 in (0001) arrangement. There is one REE site, which is coordinated
168	with 3 in-plane fluorine (F) atoms and two sets of 3 oxygen (O) atoms from bordering CO ₃ layers. Bond
169	lengths of REE with O are reported at 2.591, 2.542 and 2.579 Å while lengths with F are shorter and
170	between 2.403 and 2.416 Å. The resulting 9-coordinated tricapped trigonal prismatic REE site is
171	therefore asymmetrical with two distinct sets of bonding ligands and lengths.
172	
173	Ni et al. (2000) studied the crystal structure of parisite and determined that, unlike bastnaesite, it is
174	monoclinic. The structure was described as being built of two portions of bastnaesite layers connected
175	by a Ca layer, stacked along the c-axis (i.e., CeF-CO ₃ , CeF-CO ₃ , Ca-CO ₃). Coordination for REE is
176	similar to bastnaesite (3 x Ce-F, 6 x Ce-O) with Ce-O mean lengths between 2.55 and 2.57 Å and Ce-F
177	means between 2.38 and 2.41 Å. Minimum and maximum Ce-O bond lengths are 2.47 and 2.63 Å, while
178	for Ce-F they are 2.37 and 2.45 Å. The resulting 9-coordinated tricapped trigonal prismatic REE site is
179	therefore asymmetrical with two distinct sets of bonding ligands and lengths. The unit cell for parisite
180	contains three carbonate radical layers, two of which are influenced by both the REE-F and Ca layers
181	and one of which is influenced only by REE-F layers.
182	
183	The crystal structure of synchysite, CaCe(CO ₃) ₂ F, was investigated by Liben <i>et al.</i> (1994) and expanded
184	upon by Ni et al. (2000). Synchysite is also monoclinic and is built by stacking layers of CeF-CO ₃ and

Page 9 of 30

185 Ca-CO₃ in a one-to-one ratio along (001), making it the most Ca-rich member of the REE

186 fluorocarbonate group. Coordination for REE is similar to bastnaesite and parisite (3 x Ce-F, 6 x Ce-O)

187 with Ce-O mean lengths between 2.557 and 2.567 Å and Ce-F means between 2.394 and 2.407 Å.

188 Minimum and maximum Ce-O bond lengths are 2.50 and 2.62 Å, while for Ce-F they are 2.38 and 2.41

189 Å. The resulting 9-coordinated tricapped trigonal prismatic REE site is therefore asymmetrical with two

190 distinct sets of bonding ligands and lengths. In the case of synchysite, the carbonate layer is now always

191 influenced by bonding with both Ca and REE-F layers.

192

193 Each of the REE fluorocarbonate minerals hosts REE in 9 fold tricapped trigonal prismatic coordination 194 with 6 oxygen atoms at the apices of the prism and 3 fluorine atoms in planar configuration through the 195 faces of the prism (Fig. 1). Minor differences in the bond lengths of REE-F and REE-O amongst the 196 REE fluorocarbonates are reported in the literature, however, a significant difference between each of 197 the minerals is the local environment surrounding the carbonate radical. In bastnaesite, the oxygen 198 apices of the CO_3 radical interact only with REE cations both below and above the CO_3 plane. In 199 parasite, two of the three repeating CO₃ layers interact with REE and Ca above and below the plane 200 while the third repeating CO_3 layer interacts only with REE. For synchysite, the apices of the CO_3 201 radical always interact with both REE and Ca on either side of the plane. Consequently, bastnaesite and 202 synchysite show only one configuration for cation bonding with the CO_3 radical whereas parasite shows 203 two configurations that are unevenly populated. Furthermore, the geometry of the CO_3 polyhedra is 204 much more variable for parisite than for synchysite or bastnaesite (Fig. 2, Table 1). 205 206 The cation sites for REE in these minerals are very similar and one would therefore expect that the

207 spectral features due to intraconfigurational electron transitions of the REE would also be very similar. It

Page 10 of 30

1/22

208 would then be expected that variations between spectra of the same mineral would be predominantly in 209 the relative strengths of REE-specific absorption features while variations of REE-related spectral 210 *features between the different minerals* might include slight shifts in the locations of REE-related 211 absorption features as well as the strength of some of these features. The coordination environment of 212 the carbonate radical is quite different for each of these minerals, relatively speaking, and one would 213 expect to see a difference in the carbonate related absorption features between the minerals. Other 214 variables that may influence spectra include the signal to noise ratio of the spectral data, crystallographic 215 orientation, variations in other trace elements, and possibly syntaxial intergrowths with other REE 216 fluorocarbonates.

217

218 The infrared spectra of bastnaesite and parisite were presented in Adler and Kerr (1963). Bastnaesite is 219 characterized by carbonate related absorptions common to what is seen in calcite and dolomite, however, the strongest features for bastnaesite are the $v_2(CO_3)$ and $v_3(CO_3)$ modes at 11.52 µm (868 cm⁻¹) and 220 6.93 μ m (1443 cm⁻¹), respectively, and are asymmetric. The v₂(CO₃) and v₃(CO₃) modes in parisite are 221 the most prominent and are located at 11.49 μ m (870 cm⁻¹) and 6.90 μ m (1449 cm⁻¹), respectively, and 222 223 are asymmetric. Both the $v_2(CO_3)$ and $v_3(CO_3)$ absorption features are broader for parisite than for 224 bastnaesite. For parisite Adler and Kerr (1963) also made note of the multiple non-equivalent carbonate radicals and observed doublets for the $v_1(CO_3)$ (9.19 and 9.27 µm, or 1088 and 1079 cm⁻¹) and $v_4(CO_3)$ 225 226 (13.40 and 13.62 μ m, or 746 and 734 cm⁻¹) vibrational modes while the v₁(CO₃) and v₄(CO₃) modes for bastnaesite occur at 9.21 µm and 13.74 µm (1086 and 728 cm⁻¹). Raman and infrared spectroscopy 227 228 studies by Frost and Dickfos (2007) and Yang et al. (2008) revealed that OH stretching bands are present 229 in most REE-fluorocarbonate samples studied, although the bands recorded are variable in number and 230 position.

Page 11 of 30

231

EXPERIMENTAL METHODS

232 Samples

233 Three bastnaesite samples were borrowed from the Canadian Museum of Nature's Mineral Collection 234 (CMNMC) and were labeled as originating from Burundi (#39382), the Karonga Mine (Congo, #56255) and Madagascar (#50588). These samples were all single crystal fragments, honey brown in colour, and 235 approximately 1 cm³. A fourth bastnaesite crystal sample from the Diao Lou Shan area in Sichuan 236 237 (China) was obtained from A. Mariano, was light brown in colour and measured 3 x 3 x 2 cm. Two 238 sample sets of parisite crystals from Snowbird (Montana, USA) and Muzo (Colombia) were obtained 239 from A. Mariano. The Snowbird set comprised four euhedral elongate tapering prismatic crystal 240 fragments, each approximately 1 x 1 x 2.5 cm and light grey-brown in colour. The Muzo set comprised 241 two euhedral tapering prismatic crystal fragments, each approximately 0.5 x 0.5 x 0.5 cm and golden 242 brown with high translucency. Scanning electron microscopy (SEM) investigations showed that all the 243 bastnaesite and parisite samples were not compositionally zoned, did not show syntaxy, and did not 244 show any other mineral inclusions. Three samples of synchysite were studied, originating from 245 Narsarsuk (South-west Greenland, #UBC-3376), the White Cloud Mine (Colorado, USA) and Morris 246 County (New Jersey, USA). Both the White Cloud Mine (from CMNMC, #37320) and Morris County 247 (from CMNMC, #37321) samples were fine grained and SEM investigations showed that the complex 248 mineralogy was not suitable for baseline spectroscopic characterization studies. The Narsarsuk hand 249 sample (from UBC Museum Collection) included a cluster of ~25 light brown elongate prismatic 250 euhedral synchysite grains up to ~ 1 mm long, hosted on a feldspar-biotite dominated matrix. SEM 251 examination and electron microprobe analyses (EMPA) data also revealed very minor amounts of 252 ancylite and an unknown REE-Ca-Sr-Ba phase. Reagent-grade lanthanide oxide powders and REE-

Page **12** of **30**

1/22

doped spectralon wavelength calibration samples were also investigated in order to aid in bandassignment.

255 Reflectance spectroscopy

256 Reflectance spectroscopy was primarily carried out using the sisuROCK instrumentation (manufactured 257 by SPECIM Spectral Imaging Ltd.) at the University of Alberta's CoreSensing Facility, and data was 258 handled using ENVI 4.4, a widely used and commercially available software package. Two imaging 259 spectrometers ("cameras") acquired reflectance spectra in the visible-near infrared (VNIR, 396 nm to 260 1003 nm over 784 channels for an average spectral resolution of 0.77 nm) and shortwave infrared 261 (SWIR, 928 nm to 2530 nm over 256 channels for an average spectral resolution of 6.26 nm) portions of 262 the electromagnetic spectrum in high spatial resolution mode. Spatial resolution of the cameras in this 263 mode was approximately 0.079 mm / pixel in the VNIR and 0.241 mm / pixel in the SWIR. Noise was 264 very prevalent in the shortest wavelength portion of the VNIR camera below ~550 nm and moderate 265 from 550 nm to \sim 650 nm. In the high spatial resolution mode, averaging \sim 16 pixels resulted in reliable 266 spectra in the noisier ranges that would be useable in spectral libraries. Spectra presented originate from 267 single crystals for parisite and bastnaesite. The synchysite spectrum is an aggregate of multiple single 268 crystals. The euhedral crystals of bastnaesite from Sichuan, parisite from Muzo, and parisite from 269 Snowbird were also large enough to permit imaging both parallel and perpendicular to the grains' *c*-axis. 270 Spectra documented here are nominally an average of 2004 pixels for the VNIR camera and 1029 pixels 271 for the SWIR camera. Samples were placed on a matte black surface that translates the samples under 272 the camera and has very low reflectance across the sampled wavelength range. Some samples were 273 propped up with foam blocks to ensure surfaces of interest faced the spectrometers. All samples were 274 also substantially thick enough to assume the reflectance spectra are representative of the mineral target. 275

Page 13 of 30

276 A TerraSpec Pro point spectrometer manufactured by Analytical Spectral Devices Inc. (ASD) was used 277 in the earliest studies and records 2151 channels from 350 to 2500 nm, for a spectral resolution of ~ 1 278 nm. Spot size of the point spectrometer is approximately 1 cm in diameter, which significantly restricted 279 the use of this instrument. Methodology of acquiring spectra followed the manufacturer's 280 recommendation of ~30 second acquisition time with periodic darkfield and white reference spectralon 281 panel normalization and wavelength calibrations using Ho, Er and Dy doped spectralon samples 282 (manufactured by Labsphere). Though the ASD instrument provided limited spectra (e.g., for bastnaesite 283 from Mountain Pass), they were valuable for confidence in band assignment below 600 nm because of 284 lower noise levels than using the sisuROCK system. 285 286 Averaged spectra from the imaging spectrometer were preferred over spectra from the spot spectrometer 287 because some of the samples investigated had grain sizes smaller than were confidently resolvable by 288 the point spectrometer. This allowed for non-destructive testing of mineral specimens and the ability to 289 exclude spectral effects from other minerals that would have otherwise been in the field of view of the 290 spot spectrometer. For larger single crystal samples, a simple Region Of Interest (ROI) was used to 291 select the target mineral's pixels for averaging. For finer grained samples, a priori knowledge about the 292 sample allowed several baseline spectra to be isolated from single pixels. These spectra were then used 293 to run mixture tuned matched filtering within the ENVI software package on the entire scene, from 294 which a strict qualitative threshold allowed a discrete selection of pixels to be averaged (see Fig. 3). 295 Reflectance spectra did not have the continuum removed so as to present the data unmodified and to 296 facilitate comparison against other earlier publications.

Page 14 of 30

297 Scanning electron microscopy and electron microprobe analysis

298 The Philips XL30 scanning electron microscope (SEM) at the University of British Columbia, which is 299 equipped with an energy-dispersion X-ray spectrometer (EDS), was used for preliminary examination of 300 mineral mounts of selected minerals and rock fragments studied by reflectance spectroscopy. 301 302 Selected samples were then analyzed by electron microprobe at the Saskatchewan Research Council's 303 Advanced Microanalysis Centre using a Cameca SX-100 equipped with 5 tunable wavelength dispersive 304 spectrometers. Operating conditions were: 40° takeoff angle, beam energy of 15 keV, beam current of 20 305 nA, beam diameter of 5 µm. The MAN background intensity data was calibrated and continuum 306 absorption corrected. Elements were acquired using analyzing crystals **LLIF** for FeK α , TaL α , PrL α , 307 EuLa, DyLa, TmLa, MnKa, LaLa, NdLa, GdLa, HoLa, YbLa, BaLa, CeLa, SmLa, TbLa, ErLa, LuLa, 308 **PET** for CaKa, KKa, ClKa, TiKa, NbLa, YLa, SrLa, ZrLa, PKa, UMa, ThMa, and **LTAP** for MgKa, 309 FKα, NaKα, SiKα, AlKα. Counting times were 10 seconds for Zr and P and 15 seconds for all other 310 elements, with off peak count times of 10 seconds. The standards (with elements) were SPI-Barite (Ba), 311 SPI-Celestite (Sr), SPI-YAG (Y, Al), Smithsonian Cr-augite (Mg, Ca), Smithsonian Ilmenite (Fe, Ti), 312 Smithsonian Apatite (F, P), Smithsonian Microcline (K), Smithsonian Zircon (Zr), Harvard Albite (Si, 313 Na), Cameca Mn (Mn), SPI2-TlCl (Cl), SPI2-Nb (Nb), SPI2-La (La), SPI2-Ce (Ce), SPI2-Pr (Pr), SPI2-314 Nd (Nd), SPI2-Sm (Sm), SPI2-Eu (Eu), SPI2-Gd (Gd), SPI2-Tb (Tb), SPI2-Dy (Dy), SPI2-Ho (Ho), 315 SPI2-Er (Er), SPI2-Tm (Tm), SPI2-Yb (Yb), SPI2-Lu (Lu), SPI2-Ta (Ta), SPI2-Th (Th), and SPI2-U 316 (U). Amounts of CO_2 (as CO_3) and OH^2 (as H_2O) were determined by stoichiometry based on 4, 11 and 317 7 anions for bastnaesite, parisite and synchysite, and for full occupation of the F atomic site by F, Cl⁻ 318 and OH⁻.

Page **15** of **30**

319

RESULTS

320 Samples described here were selected from a larger set and were characterized by imaging reflectance 321 spectroscopy, scanning electron microscopy and microprobe analysis. Several samples have spectra 322 collected from a point spectrometer. Table 2 documents electron microprobe results for the REE 323 fluorocarbonates in this study and Figure 4 shows Chondrite-normalized patterns of the samples. Tables 324 3 and 4 document prominent absorption features for the REE fluorocarbonates and includes probable 325 origins of the features, as chosen through comparison with reflectance spectra from unpublished reagent 326 grade lanthanide oxide spectra, REE-doped calibration standard spectra, and other REE-bearing mineral 327 spectra for which compositional data exists, as well as REE spectroscopy literature. The position and 328 shape of these features were recorded using reflectance spectra (i.e., not continuum removed spectra). 329 Tables 5 and 6 compare significant chemical and spectral differences amongst the bastnaeiste and 330 parisite sample sets, respectively, and include the number of pixels averaged from the sisuROCK

imaging scenes by the VNIR and SWIR cameras.

332 Bastnaesite

Chemical variability amongst the four samples is moderate, with the samples from Burundi (#39382) and Karonga (#56255) showing greater amounts of heavier REE, such as Nd, Sm and Gd (Table 2 and Fig. 4). None of the crystals show unusual concentrations of other cations or notable compositional zoning. Published analyses of bastnaesite from other localities are comparable to the samples analyzed here (e.g., Zaitsev *et al.* 1998, Holtstam and Andersson 2007).

338

Comparing the spectra (Fig. 5) of the four bastnaesite samples in both the VNIR and SWIR ranges,

340 differences are primarily the relative strengths of absorption features generated by lanthanides. The

341 sample from Sichuan is taken as the baseline because it shows the cleanest spectrum, is derived from the

Page 16 of 30

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most number of pixels, and has supporting spectra from a point spectrometer. The sample from
Madgascar shows the largest differences in that, in general, its features are more subdued than its peers.
Spectra were collected both perpendicular and parallel to the *c*-axis for the Sichuan sample, however, no
differences were noted.

346

347 Parisite

348 Chemically, the Snowbird sample shows a greater proportion of heavy rare earth elements (Table 2 and

Fig. 4) and especially thorium than the sample from Muzo and neither shows notable compositional

350 zoning. Published analyses of parisite from other localities are comparable to the samples analyzed here

351 (e.g., Zaitsev et al. 1998, Ni et al. 2000).

352

353 Spectra of the two samples show strong similarities (Fig. 6) with variations predominantly in the

354 strength of absorption features and relative depths between specific features. Spectra from Muzo

355 samples also show narrower absorption features. Spectra were collected both perpendicular and parallel

to the *c*-axis for both parisite samples, however, no differences were noted.

357 Synchysite

358 Three samples of synchysite were studied, however, only the sample from Narsarsuk (Greenland)

359 produced reliable and unambiguous spectra. Spectra indicative of REE-fluorocarbonates could be

- 360 resolved from the White Cloud Mine (CO, USA) and Morris County (NJ, USA) samples, however, SEM
- 361 investigations of both revealed complex REE mineralogy, including synchysite, and very fine grained
- 362 material. Consequently only the Narsarsuk sample is suitable for baseline characterization and

Page 17 of 30

discussed in detail. Microprobe results are in accordance with other published analyses (e.g., Forster
2001, Guastoni *et al.* 2009).

365

The VNIR and SWIR spectra included 1651 pixels and 175 pixels, respectively. The spectrum (Fig. 7) of the Narsarsuk sample is dominated by the high relative concentrations of spectrally active lanthanides and resembles spectra of bastnaesite and parisite. Synchysite shows carbonate related minima at 2337 and 2518 nm with a shoulder at 2499 nm.

370

371 DISCUSSION ON THE SPECTRAL VARIABILITY OF REE-FLUOROCARBONATE 372 MINERALS

The reflectance spectra from samples of bastnaesite, parisite and synchysite, all members of the REEfluorocarbonate mineral group, are consistent amongst their respective mineral species. The location of absorption features related to the lanthanides will be a function of the specific REE coordination and asymmetry of crystal structure; because all three minerals have very similar cation sites for the REE consistent patterns were observed. As alluded to in the review of these minerals' crystal structures, the differences in coordination of the CO₃ radical result in the greatest distinction between the REE fluorocarbonates.

380

In general, 15 'clusters' of spectral features are present between 500 nm and 2500 nm in bastnaesite, the REE-fluorocarbonate with the highest concentration of REE and simplest crystal structure. Spectra from parisite and synchysite show strong similarities to bastnaesite spectra, however, noise at shorter wavelengths reduces confidence in this region for these two minerals. Tables 3 and 4 document resolvable absorption features of the fluorocarbonates based primarily on spectra from the Sichuan

Page 18 of 30

(bastnaesite), Snowbird (parisite) and Narsarsuk (synchysite) localities. Figures 8 and 9 show
representative VNIR and SWIR spectra of bastnaesite, parisite and synchysite with bolded index lines
and approximate Full Width at Half Max regions for the dominant absorptions of each group. Probable
origins for absorptions are described in Tables 3 and 4.

390

391 Intraconfigurational 4f-4f electronic transitions

The REE-related absorptions in the VNIR-SWIR for these LREE-enriched minerals can be largely attributed to Pr^{3+} , Nd^{3+} and Sm^{3+} , with much lesser influence from Eu^{3+} . Trivalent Y, La^{3+} and Lu^{3+} do not have free electrons in the *f* orbitals to generate absorptions, and absorptions related to Ce^{3+} do not fall within the VNIR-SWIR range. Similarly, Gd^{3+} shows absorptions at too high of energy levels, and Dy^{3+} , Tb^{3+} , Ho^{3+} , Er^{3+} , Tm^{3+} and Yb^{3+} are either too low in concentration and/or overlapping with other stronger absorption bands to have a noticeable effect in the spectra of the studied minerals.

398

399 Most differences between the REE fluorocarbonate mineral spectra are small shifts in the location of the 400 absorption minima or apparent strengths of overlapping absorptions represented as shoulders or weak 401 local minima. Most of the differences in absorption minima locations are usually on the scale of one or 402 two spectral channels suggesting that in some cases it may be the result of noise or a function of 403 spectrum slope. Noise at shorter wavelengths in spectra from the sisuROCK Imaging Spectrometer 404 precludes delineation of some features below ~650 nm, however, data from the ASD point spectrometer 405 with greater signal to noise performance supports subtle features from 650 nm down to 450 nm. Better 406 resolution of local minima as opposed to overlapping absorptions or shoulders in spectra is primarily 407 observed for bastnaesite over synchysite and parisite, likely because of the greater concentration of REE 408 in the Ca-absent bastnaesite. For example, the local minima at 689 nm and 898 nm (Fig. 8) are generally

Page 19 of 30

1/22

409	better defined in the bastnaesite samples, and two resolved absorptions are prevalent with minima at
410	1547 and 1578 nm (Fig. 9). Parisite from Muzo and synchysite from Narsarsuk do, however, show more
411	resolved features than parisite from Snowbird near 676 nm, 889 nm and 1547 nm (Figs. 6 and 7).
412	Bastnaesite consistently shows a deep and relatively broad absorption centered at 1968 nm, while
413	parisite and synchysite both show the minimum of this absorption at 1961 nm. A water combination
414	band occurs in this region and likely plays a role, however, Pr ³⁺ also exhibits a strong absorption here.
415	This wavelength shift is consistent with Görller -Walrand and Binnemans (1996) listing this
416	intraconfigurational transition (excitation from ground to the ${}^{7}F_{2}$ multiplet) as being particularly
417	sensitive to the crystal field.
418	
419	Vibrational bands of the carbonate radical
420	The most important charge stice amongst the spectre of DEE flueroearchemotes is that a distinct
420	The most important observation amongst the spectra of REE fluorocarbonates is that a distinct
421	absorption at 2243 nm (4458 cm ⁻¹) distinguishes bastnaesite from parisite and synchysite, neither of
422	which have this feature (Fig. 8). Bastnaesite also shows a well resolved doublet with absorption bands at

423 2312 and 2324 nm (4325 and 4303 cm^{-1}), a shoulder at 2355 nm (4246 cm^{-1}) and an additional band at

424 2518 nm (3971 cm⁻¹). Notable is that the bastnaesite sample from Madagascar only shows a weak

425 shoulder at 2243 nm and at present the reason for this is uncertain (Fig. 5). It is possible that structurally

bound hydroxyl present in the bastnaesite samples may be the origin of the 2243 nm feature, however,

427 the same substitution of fluorine could be said for both parisite and synchysite and neither show this

428 absorption. Furthermore, EMPA composition of the Madagascar sample does not suggest it as being

429 OH-absent and F-rich. Moreover, our unpublished spectra of uncharacterized bastnaesite from other

430 localities do show the 2243 nm absorption feature.

431

Page 20 of 30

Both parisite and synchysite have broader absorption features that are likely related to the carbonate
radical. Parisite shows a weak absorption band at 2324 nm (4303 cm⁻¹) and another near 2499 nm (4002 cm⁻¹) (Fig. 9). Synchysite shows a more resolved absorption band at 2337 nm (4279 cm⁻¹) and another at
2518 nm (3971 cm⁻¹).

436

437 Spectral effects of the actinides

Most REE minerals will also host the actinide elements Th^{4+} and U^{4+} . The electronic structure of Th^{4+} is 438 similar to La^{3+} , and so has no spectral response related to *f*-orbital transitions. Tetravalent U has a similar 439 structure to Pr^{3+} and Binnemans *et al.* (1999) show that although similar, U⁴⁺ shows stronger but broader 440 absorption features than Pr^{3+} because the 5*f* electrons are less efficiently shielded in actinides than the 4*f* 441 electrons in lanthanides. Dominant absorptions of U^{4+} have been recorded near ~1111 nm (~9,000 cm⁻¹) 442 and ~666 nm (15,000 cm⁻¹). For the REE fluorocarbonate minerals studied here, Th and U contents were 443 low enough that no actinide-related spectral features were observed. Unpublished spectra of Th-U-444 445 bearing zircon do, however, show spectral features in these regions. 446

447

IMPLICATIONS

The REE are strategic elements in many high-tech industries due to their special properties in permanent magnets, metal alloys, catalysts and phosphors, among other applications. Recent volatility in supply and demand has increased exploration in North America for these critical elements but ore mineralogy can be significantly variable and can have significant implications for later beneficiation (e.g., Mariano and Mariano 2012).

453

Page 21 of 30

The reflectance spectra of bastnaesite, parisite and synchysite presented and discussed here form the first systematic spectral characterization and comparison of these three important REE minerals. Their spectra are feature-rich, distinct, and comprise many narrow overlapping absorption bands that can be mostly attributed to 4f-4f intraconfigurational electron transitions of Nd³⁺, Pr³⁺, and Sm³⁺ and possibly Eu³⁺.

459

460 Consistent lanthanide-related spectral responses of the REE fluorocarbonates support the notion that 461 their REE cation site is very similar. Variations in these consistent lanthanide-related responses suggest 462 REE concentration information could be extrapolated from reflectance spectra. Variable carbonate-463 related spectral responses support the notion that the crystallographic sites for the carbonate radical 464 differ between bastnaesite, synchysite and parisite.

465

Exploitable differences for mineral identification using hyperspectral imaging include the carbonaterelated absorption at 2243 nm for bastnaesite, as well as a resolved doublet with minima at 2312 nm and 2324 nm. Synchysite shows a carbonate-related minimum at 2337 nm and parisite shows a less resolved carbonate-related minimum near 2324 nm. A Pr³⁺ and water related absorption is exhibited at 1968 nm for bastnaesite while in synchysite and parisite it occurs at 1961 nm. Collectively, the spectral patterns of this mineral grouping are distinct from other REE minerals phases, such as monazite, xenotime, and eudialyte, among others.

473

This work starts to address the knowledge gap that exists between reflectance spectroscopy and rare
earth element mineralogy and future research will include continued characterization and comparison of
other important REE-bearing minerals. Accordingly, although hyperspectral remote sensing is becoming

Page 22 of 30

477 increasingly prevalent in the mining and mineral exploration industries, its adoption into REE-focused
478 programs could be of particularly great benefit, such as in defining and assessing ore mineralogy or
479 rapidly logging diamond drill core. Incorporating these spectra into larger spectral databases harnessed
480 by a variety of users will also facilitate the identification of REE fluorocarbonate minerals that may have
481 otherwise gone unnoticed.
482

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571	
572	FIGURE CAPTIONS
573	
574	Figure 1. (a) Coordination polyhedron for the Ce1 site in parisite (Ni et al. 2000); also applicable to
575	bastnaesite and synchysite. Ce1 (green) atom is coordinated with F1, F2 and F3 (lavender) atoms,
576	whose plane is roughly perpendicular to c-axis, and 6 oxygen (red) atoms O11, O23, O32, O42,
577	O53 and O61. Overall coordination number of 9 in a distorted tricapped trigonal prismatic
578	arrangement. (b) Parisite crystal structure from Ni et al. (2000), (c) bastnaesite crystal structure
579	from Ni et al. (1993), and (d) synchysite crystal structure from Wang et al. (1994). Atom
580	colouring: red=oxygen, green=REE, lavender=F, brown=C. Polyhedra colouring: green=REEO ₆ F ₃ ,
581	dark blue= CaO_8 , brown= CO_3 .
582	
583	Figure 2. Bond characteristics for the CO ₃ radical polyhedra in REE fluorocarbonate. Data from Table 1
584	and references therein, Min=minimum, Mid=middle, Max=maximum.
585	
586	Figure 3. (a) Hyperspectral imagery of the synchysite-bearing sample from Narsarsuk in the short wave
587	infrared (SWIR). (b) Mixture tuned matched filtering (MTMF) results for synchysite in SWIR
588	scene and (c) regions of interest (ROI) derived from thresholded MTMF results for generating an
589	average spectrum. (d) Annotated digital photograph of the same sample from nominally the same
590	perspective. Arrows in (c) and (d) point to the same patches of synchysite. Monomineralic patches
591	of synchysite are on the order of 2 mm by 2 mm, large enough for successful imaging with a
592	spatial resolution of 0.241 mm by 0.241 mm, to enable isolation of pure monomineralic pixels
593	with confidence for averaging of their spectra.

Page 28 of 30

1/22

594	
595	Figure 4. Chondrite-normalized REE plot from microprobe results. Normalization values from
596	McDonough and Sun (1995). Erbium (Er), Tm, Yb and Lu are below detection for all samples.
597	
598	Figure 5. Stacked spectra of bastnaesite in VNIR (left, 500 to 1000 nm) and SWIR (right, 975 to 2530
599	nm) from the two sisuROCK instrument cameras. Spectra from top to bottom: Sichuan (S),
600	Burundi (B), Madagascar (M), Karonga (K).
601	
602	Figure 6. Stacked spectra of parisite in VNIR (left, 500 to 1000 nm) and SWIR (right, 975 to 2530 nm)
603	from the two sisuROCK instrument cameras. Spectra from top to bottom: Muzo (M), Snowbird
604	(SB).
605	
606	Figure 7. Spectra of synchysite from Narsarsuk in VNIR (left, 500 to 1000 nm) and SWIR (right, 975 to
607	2530 nm) from the two sisuROCK instrument cameras.
608	
609	Figure 8. VNIR (500 to 1000 nm) spectra of bastnaesite (B, top), parisite (P, middle) and synchysite (S,
610	bottom). Italic numbers denote groups with probable origin described in Table 3. Lines denote
611	prominent absorption features with wavelength position, shaded boxes represent the approximate
612	Full Width at Half Max for each absorption or absorption cluster, borderless box indicates narrow
613	feature. Stacked spectra from sisuROCK instrument.
614	
615	Figure 9. SWIR (975 to 2530 nm) spectra of bastnaesite (B, top), parisite (P, middle) and synchysite (S,
616	bottom). Italic numbers denote groups with probable origin described in Table 4. Lines denote

Page 29 of 30

- 617 prominent absorption features with wavelength position, shaded boxes represent the approximate
- 618 Full Width at Half Max for each absorption or absorption cluster, borderless box indicates narrow
- 619 feature. Stacked spectra from sisuROCK instrument.

620

Page **30** of **30**





(a) SWIR False Colour

(b) MTMF Value Colour Ramp





(c) SWIR with ROI from MTMF (d) Digital Photo



















Mineral	bastnaesite	parisite	synchysite
Reference	Ni et al. (1993)	Ni <i>et al.</i> (2000)	Wang et al. (1994)
Formula	REECO ₃ F	$CaREE_2(CO_3)_3F_2$	CaREE(CO ₃) ₂ F
Symmetry	Hexagonal	Monoclinic	Monoclinic
Space group	$P\overline{6}2c$	Сс	C2/c
a (Å)	7.1175	12.305	12.329
b (Å)		7.1053	7.110
c (Å)	9.7619	28.25	18.741
B (°)		98.257	102.68
Structure			
REE coordination	9: 3*F, 6*0	9: 3*F, 6*O	9: 3*F, 6*0
REE site shape	tricapped trigonal	tricapped trigonal	tricapped trigonal
Number of unique REE			
cation sites	1	6	2
REE-O (Min, Å)	2.542	2.466	2.497
REE-O (Mean, Å)	2.571	2.564	2.558
REE-O (Max, Å)	2.591	2.631	2.619
REE-F (Min, Å)	2.403	2.370	2.376
REE-F (Mean, Å)	2.407	2.397	2.399
REE-F (Max, Å)	2.416	2.445	2.417
	Each CO ₃ layer	3 distinct CO ₃ layers, one	Each CO ₃ layer
	influenced by REE	influenced by REE only, two	influenced by Ca
CO ₃ influences	only	influenced by REE and Ca	and REE
Number of unique CO ₃			
polyhedra	1	9	3
Average C-O bond			
lengths (A)	1 207	1 2 4 7 0	1 2017
	1.287	1.24/8	1.281/
		1.2728	1.2854
CO ₃ #3		1.2/18	1.2752
<u> </u>		1.2636 a	
CO ₃ #5		1.2646 a	
CO ₃ #6		1.2736 a	
<u>CO3 #7</u>		1.3004	
CO ₃ #8		1.2887	
CO ₃ #9		1.2945	
Average CO_3 polyhedra	0.0017	0.0222	0.0146
Volume (A ³)	0.0017	0.0223	0.0146
Average CO ₃ polyheura	0.0004	0.0292	0.0150
Ceneral chemistry	0.0094	0.0282	0.0139
RFF ₂ O ₂ wrt ⁰ / ₂	73	59	51
$\Gamma_{2}O_{3}$ wt γ_{0}	0	10.5	17
F wt%	5	45	26
Γ_{0} wt%	20	24	2.0
003 111/0	20	<u> </u>	20

Table 1. Some basic properties of the REE fluorocarbonates

^a These CO₃ polyhedra in parisite are coordinated only to REE

Mineral	<u>bastnä</u>	site	bastnä	site	bastnä	site	bastnä	<u>site</u>	paris	site	paris	site	synch	ysite	
Sample ID	Sichu	<u>an</u>	<u>3938</u>	<u>32</u>	<u>5058</u>	<u>38</u>	<u>5625</u>	<u>55</u>	Mu	<u>zo</u>	Snow	bird	UBC3	376	
Locality	Diao I	Lou	Burur	ndi	Madaga	ascar	Karon	ga,	Color	nbia	Mont	ana,	Narsai	suk,	Detection
	Shan, C	China					Cong	go			US	A	Green	land	Limits
n	10	2σ	5	2σ	5	2σ	5	2σ	6	2σ	5	2σ	3	2σ	(average)
Nb_2O_5	0.01	0.02	0.00	0.00	0.01	0.03	0.00	0.00	0.00	0.02	0.00	0.01	0.00	0.00	0.04
P_2O_5	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02	0.00	0.00	0.00	0.00	0.05	0.17	0.02
SiO ₂	0.05	0.01	0.11	0.02	0.06	0.00	0.11	0.01	0.06	0.01	0.08	0.08	0.38	0.52	0.01
ZrO_2	0.00	0.00	0.01	0.03	0.00	0.00	0.01	0.03	0.00	0.00	0.00	0.00	0.01	0.03	0.04
UO_2	0.53	0.10	0.00	0.00	0.00	0.00	0.00	0.00	0.39	0.13	0.34	0.19	0.00	0.00	0.07
ThO ₂	0.14	0.13	0.00	0.00	0.37	0.48	0.00	0.00	0.83	0.64	1.62	0.26	0.54	0.12	0.06
Al_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.04	0.15	0.01
La_2O_3	27.77	1.88	19.44	1.00	25.22	0.64	20.76	0.75	14.03	0.53	12.87	0.37	12.17	1.33	0.03
Ce_2O_3	31.14	0.67	32.58	0.31	32.95	0.11	32.47	0.48	26.11	0.90	24.77	0.77	22.29	0.99	0.02
Pr ₂ O ₃	2.50	0.32	3.07	0.33	2.38	0.19	2.98	0.15	2.90	0.19	2.71	0.19	2.17	0.29	0.02
Nd_2O_3	7.25	1.18	11.05	1.07	6.62	0.15	10.28	0.20	11.90	0.59	12.14	0.17	8.25	0.57	0.02
Sm_2O_3	0.47	0.18	1.18	0.14	0.29	0.07	1.02	0.01	1.85	0.14	2.64	0.23	1.42	0.05	0.02
Eu_2O_3	0.00	0.02	0.16	0.08	0.02	0.05	0.11	0.05	0.03	0.07	0.10	0.10	0.17	0.01	0.02
Gd_2O_3	0.20	0.18	0.53	0.14	0.08	0.10	0.40	0.10	1.06	0.19	1.70	0.25	1.03	0.08	0.02
Tb_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.03	0.08	0.04	0.02	0.02
Dy_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.35	0.16	0.38	0.08	0.02
Ho_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.001	0.00	0.02
Er_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02
Tm_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02
Yb_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.03
Lu_2O_3	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.02
Y_2O_3	0.01	0.03	0.17	0.12	0.01	0.01	0.11	0.06	0.74	0.32	1.69	0.13	1.47	0.53	0.04
SrO	0.01	0.04	0.02	0.06	0.08	0.13	0.02	0.05	0.00	0.00	0.00	0.00	0.22	0.36	0.03
MgO	0.00	0.00	0.02	0.02	0.01	0.02	0.01	0.02	0.00	0.00	0.00	0.00	0.00	0.00	0.01
FeO	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.17	0.43	0.01
CaO	0.11	0.10	0.00	0.00	0.05	0.02	0.00	0.00	10.48	1.06	10.23	0.16	16.49	2.69	0.01
Na_2O	0.03	0.03	0.00	0.00	0.00	0.00	0.00	0.00	0.01	0.02	0.01	0.01	0.10	0.04	0.01
Cl	0.01	0.02	0.08	0.02	0.06	0.02	0.08	0.04	0.02	0.01	0.02	0.01	0.12	0.05	0.01
F	5.22	0.42	4.42	0.22	4.56	0.15	4.27	0.23	4.48	0.20	4.12	0.57	2.87	1.04	0.02
CO_2^*	18.86		18.36		18.32		18.35		24.2		24.46		27.26		

Table 2. Average EMPA compositions of REE fluorocarbonate samples in wt%

H ₂ O**	1.38	1.64	1.57	1.17	1.17	1.38	1.4	
-O≡Cl	0.00	-0.02	-0.01	-0.02	0.00	0.00	-0.03	
-O≡F	-2.20	-1.86	-1.92	-1.80	-1.89	-1.73	-1.21	
Total	93.50	90.96	90.73	90.32	98.37	99.51	97.80	
$[REE+Y]_2O_3$	69.35	68.19	67.57	68.12	58.61	58.99	49.38	

Ta, Ti, Mn, K and Ba were sought but not detected *Determined by stoichiometry, on the basis of 4, 11 and 7 anions, respectively **H2O calculated assuming ideal occupation of "F" site (F, Cl, OH)

	Bastnaesite		Parisite		Syn	chysite	Probable Origin
Group	λ (nm)	Shape	λ (nm)	Shape	λ (nm)	Shape	-
· · ·	500	MIN	501	MIN-n	503	MIN-n	Nd ³⁺
1	511	MIN	510	MIN-n	512	MIN-n	Nd ³⁺
	523	MIN	523	MIN	525	MIN-n	Nd ³⁺
	533	MIN-w	533	MIN-n	534	MIN-n	Nd ³⁺ >Eu ³⁺
	545	SH	545	SH-n			Nd ³⁺
	557	MIN-w					Sm ³⁺
2	575	SH	574	SH-n	576	SH-n	Nd ³⁺
	580	MIN-st	580	MIN-st	580	MIN-st	Nd ³⁺
	595	SH	595	SH	595	SH	Nd ³⁺ >Pr ³⁺
3	625	MIN	625	MIN-w	625	MIN-w, n	Nd ³⁺
	636	MIN-w	637	MIN-w, n			Nd ³⁺
	642	SH-w					Nd ³⁺
4	676	MIN	676	MIN	678	MIN	Nd ³⁺
	681	SH	681	MIN-w, n	680	SH	Nd ³⁺
	689	MIN-w	689	MIN-w	691	SH	Nd ³⁺
	693	SH					Nd ³⁺
5	734	SH	732	SH	732	SH	Nd ³⁺
	738	SH	738	SH	738	SH	Nd ³⁺
	741	MIN-st	741	MIN-st	740	MIN-st	Nd ³⁺
	749	MIN-w	749	SH	749	SH	Nd ³⁺
	755	SH	755	SH	756	SH	Nd ³⁺
	761	SH	762	SH	762	SH	Nd ³⁺
	768	SH	770	SH	770	SH	Nd ³⁺
6	781	SH	781	SH	780	SH	Nd ³⁺
	784	SH-w	784	SH-w	784	SH-w	Nd ³⁺
	792	MIN-db	792	MIN-db	792	MIN-db	Nd ³⁺
	797	MIN-db	798	MIN-db	799	MIN-db	Nd ³⁺
	803	SH	804	SH	804	SH	Nd ³⁺
	813	SH	813	SH	813	SH	Nd ³⁺
	821	SH	822	SH	822	SH	Nd ³⁺
7	845	SH	845	SH			Nd ³⁺
	864	MIN-st	864	MIN-st	864	MIN-st	Nd ³⁺
	871	SH	871	SH	871	SH	Nd ³⁺
	889	MIN	889	MIN	889	MIN	Nd ³⁺
	898	MIN-w	899	SH	899	SH	Nd ³⁺
8	953	MIN	953	MIN	953	MIN-w	Sm ³⁺
	978	MIN-w	978	SH-w			Nd ³⁺ >Pr ³⁺

Table 3. Prominent absorption features of the REE fluorocarbonates in the VNIR range.

Bolded values are indexed on Fig 8. Shape and modifier abbreviations: MIN=local minimum, SH=shoulder, -n=noisy, -w=weak, -st=strong, -db=doublet.

U	Bastnaesite		Parisite		Sync	hysite	Probable Origin
Group	λ (nm)	Shape	λ (nm)	Shape	λ (nm)	Shape	
9	1004	MIN-w	1004	SH-w	998	SH-w	Pr ³⁺
	1023	MIN	1023	MIN-w	1023	SH	Pr ³⁺
	1093	MIN-st	1093	MIN-st	1093	MIN-st	Sm ³⁺
10	1225	SH-w	1225	SH-w	1225	SH-w	Sm ³⁺
	1251	MIN-st	1251	MIN-st	1251	MIN-st	Sm ³⁺
	1314	SH-w	1314	SH-w	1301	SH-w	Pr ³⁺ ?
11	1408	SH	1408	MIN-w	1408	MIN-w	H ₂ O
	1465	MIN	1465	MIN	1459	MIN	Pr ³⁺
	1496	MIN-w	1496	MIN-w	1496	MIN-w	Pr ³⁺ >Sm ³⁺
	1547	MIN-st	1547	MIN	1547	MIN-st	Pr ³⁺ >Sm ³⁺
	1578	MIN	1578	SH	1578	SH	Pr ³⁺ >Sm ³⁺ >Nd ³⁺
	1666	SH-w	1666	SH-w	1666	SH-w	Nd ³⁺
	1710	MIN	1710	MIN	1710	MIN	Nd ³⁺
	1742	SH-w	1742	SH-w	1742	SH-w	Nd ³⁺
	1761	SH-w	1767	SH-w	1767	SH-w	Nd ³⁺
	1817	SH-w	1817	SH-w	1817	SH-w	Nd ³⁺
	1886	SH-w	1880	SH-w			Pr ³⁺ ?
	1943	SH-w					Pr ³⁺
12	1968	MIN-st	1961	MIN-st	1961	MIN-st	Pr³⁺>Sm³⁺ & H₂O
	2037	SH	2037	SH-w	2037	SH-w	Pr ³⁺ >Eu ³⁺
	2112	SH-w	2118	SH-w	2105	SH-w	CO3?
	2143	MIN-w	2143	SH-w			Pr ³⁺
13	2243	MIN-st		ABSENT		ABSENT	CO ₃
14	2312	MIN-db	2312	SH	2312	SH-w	CO ₃ >Pr ³⁺ >Nd ³⁺ >Sm ³⁺
	2324	MIN-db	2324	MIN / SH	2337	MIN	CO₃>Pr ^{³+} >Nd ^{³+} >Sm ^{³+}
	2355	SH	2355	SH-w	2368	SH-w	CO ₃ >Pr ³⁺ >Nd ³⁺ >Sm ³⁺
			2393	SH-w	2393	MIN-w	CO ₃
15	2437	SH	2443	SH-w			$CO_{3}>Pr^{3+}>Nd^{3+}>Sm^{3+}$
	2480	SH-w	2474	SH	2474	SH-w	CO ₃
	2499	SH-w	2499	MIN	2499	SH	CO ₃
	2518	MIN	2518	SH-w	2518	MIN	CO ₃

Table 4. Prominent absorption features of the REE fluorocarbonates in the SWIR range.

Bolded values are indexed on Fig 9. Shape and modifier abbreviations: MIN=local minimum, SH=shoulder, -n=noisy, -w=weak, -st=strong, -db=doublet.

Sample	VNIR	SWIR	Notes
	pixels	pixels	
Diao Loa Shan	5715	4379	Chemically, shows intermediate values for REE. Spectrally, shows
– Sichuan,			the most resolved features in the VNIR.
China			
Burundi	2312	173	Chemically, shows the highest values for heavier REE, such as Gd. Spectrally, is very similar to the Sichuan sample. The 1093 nm and 1251 nm absorptions (Sm ³⁺ related) are particularly
			pronounced in the suite, which is correlative with this sample
			1400 nm feature is represented as a local minimum rather than a
			shoulder
Madagascar	613	387	Chemically, shows higher La and Ce but lower middle REE within
			the bastnaesite suite. Spectrally, it generally shows subdued REE
			absorptions as compared to other samples. However, a
			significant difference is that the carbonate related 2243 nm and
			2312/2324 nm doublet features are very subdued.
Karonga Mine,	1341	135	Chemically, shows intermediate values for REE with a greater
Congo			abundance of middle REE than the Sichuan sample. Spectrally, it
			is similar to the Burundi sample with respect to the 1093, 1251
			and 1408 nm Sm ³⁺ features, however, it displays less resolved
			Nd ³⁺ related features at 741, 792/797 and 864 nm.

 Table 5. Significant chemical and spectral differences between bastnaesite samples.

 Sample
 VNIP
 SWIP

Sample	VNIR	SWIR	Notes
	pixels	pixels	
Muzo,	265	162	Chemically, more enriched in La, Ce and Pr. Spectrally, tends to
Colombia			show better resolved and narrower features than for the
			Snowbird sample.
Snowbird,	2130	1790	Chemically, shows greater values for heavier REE, from Nd to Lu,
USA			but is still LREE-enriched. Spectrally, it shows broadened features
			and finer structure is generally subdued (e.g., clusters near 580
			nm and 1500 nm). The carbonate-realted features near 2324 and
			2474 nm are also less resolved. The sample's greater amount of
			Th will have subjected these crystals to greater radioactivity,
			which might contribute to broadening of spectral features as the
			crystal loses crystallinity.

Table 6. Significant chemical and spectral differences between parisite samples.