1	Revision 3
2	Z-contrast imaging and <i>ab initio</i> study on "d" superstructure in sedimentary
3	dolomite
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23 ABSTRACT

24 Nano-precipitates with tripled periodicity along the *c*-axis are observed in a Ca-rich dolomite sample from Proterozoic carbonate rocks with "molar tooth" structure. This observation is 25 26 consistent with previous description of d reflections. High-angle annular dark-field STEM imaging (or Z-contrast imaging) that avoids dynamic diffraction as seen in electron diffraction 27 and high-resolution TEM imaging modes, confirms that d reflections correspond to nanoscale 28 29 precipitates aligned parallel to (001) of the host dolomite. The lamellae precipitates have a cation ordering sequence of Ca-Ca-Mg-Ca-Ca-Mg along the *c* direction resulting in a chemical 30 composition of $Ca_{0.67}Mg_{0.33}CO_3$. This superstructure is attributed to the extra or d reflections, 31 thus is referred to as the *d* superstructure in this study. The structure can be simply described as 32 interstratified calcite/dolomite. The crystal structure of the *d* superstructure calculated from 33 34 density functional theory (DFT) has a space group of P31c and has a and c unit cell parameters of 4.879 Å and 16.260 Å, respectively, values between those of dolomite and calcite. The 35 detailed structural characteristics and parameters obtained from *ab initio* calculations are also 36 37 reported in this paper. The method of combining Z-contrast imaging and *ab initio* calculation can be used for solving structures of other nano-precipitates and nano-phases. 38

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40 INTRODUCTION

The dolomite($R\bar{3}c$) structure has alternating Ca²⁺ and Mg²⁺ cation layers along the *c*-axis with 41 the triangular CO_3^{2-} anion layers lying between two cation layers. Since the size of Mg²⁺ ions is 42 smaller than that of Ca^{2+} , the CO_3^{2-} layers are closer to Mg^{2+} layers. The lack of c glides in the 43 dolomite structure due to the cation ordering causes the occurrence of extra reflections (b 44 reflections, $h\bar{h}0l$, l= odd) in dolomite diffraction patterns compared to that of calcite (a 45 46 reflections only) (Reeder, 1992). Two additional reflections (c and d reflections) have been observed in natural sedimentary dolomite. Previously reported c reflections were interpreted as 47 being a result of cation ordering within dolomite basal planes (Van Tendeloo et al., 1985). Our 48 recent STEM work confirmed that the "c"-reflections could result from multiple diffraction 49 50 between the host dolomite and twinned Mg-calcite nano-lamellae under TEM imaging and 51 diffraction modes (Shen et al., 2013). The d reflection was first observed in Devonian dolomite samples by Wenk and Zenger (1983) and have been also found in ankerite samples (Rekesten, 52 53 1990). The d reflections occur as satellites around a and b reflections with diffraction vector $\sim 1/3$ 54 $(0001)^*$ and are usually streaking along c^* direction (Wenk and Zenger, 1983; Wenk and Zhang, 55 1985; Van Tendeloo et al., 1985).

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57 Dynamical diffraction in transmission electron microscopy (TEM) mode has been a major 58 problem for structure determination as shown in previous work of analyzing "*c*" superstructures 59 (Shen et al., 2013). The scanning transmission electron microscopy (STEM) method uses the 60 high-angle annular dark-field (HAADF) detector to give the most highly localized 1s Bloch state 61 imaging, which eliminates most of the obvious effects of dynamical diffraction (Pennycook, This is a preprint, the final version is subject to change, of the American Mineralogist (MSA) Cite as Authors (Year) Title. American Mineralogist, in press. (DOI will not work until issue is live.) DOI: http://dx.doi.org/10.2138/am.2014.4647

62	2002). With the advantage of a spherical aberration corrector, the resolution of HAADF STEM
63	or Z-contrast images is only limited by the size of the 1s Bloch state that is \sim 0.6-0.8 Å
64	(Pennycook et al., 2000). The intensity of Z-contrast images is dependent on the atomic number
65	of atoms through the $\sim Z^2$ dependence of the Rutherford scattering cross-section, which thus
66	provides chemical information for the material (Kirkland, 1998; Pennycook et al., 2000). A study
67	of microstructures in natural dolomite samples using Z-contrast images may help find answers to
68	previous observations of superstructures and explore new microstructures in dolomite.

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In addition to experimental studies in mineralogy, the application of *ab initio* calculations of 70 crystal structure, phase stability, and physical properties of minerals at given pressure and 71 72 temperature has increased in the past few years (Ogonov et al., 2006; Barnard and Xu, 2008; Chatterjee and Saha-Dasgupta, 2010; Stackhouse et al., 2010). Density functional theory (DFT) 73 74 uses the functional of electron density to solve the Schrödinger equation for a many electron 75 system to acquire the minimum energy of the system (Scholl and Steckel, 2009). This method can calculate the enthalpy of a system at 0K and the corresponding structure. Since STEM work 76 only provides topological information for the crystal structure, the use of the DFT method is 77 needed in this study to calculate the detailed crystal structure of the d phase and confirm our 78 model for the *d* superstructure by comparing the energetic stability of our model to another 79 model proposed in literature. 80

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The "molar tooth" structure refers to vertical and horizontal ribbons and blobs of fine-grained
calcite in a dolomite host (Bauerman, 1885). This structure has only been recognized in

Mesoproterozoic and Neoproterozoic marine carbonate rocks with a few exceptions (Frank and Lyons, 1998). The origin and the temporally limited occurrence of molar tooth carbonate have bewildered geologists for over a century (James et al., 1998; Pratt, 1998; Meng and Ge, 2002; Marshall and Anglin, 2004; Pollock et al., 2006; Long, 2007; Kuang et al., 2012). However, no previous work on the nano-scale mineralogy of carbonates with the "molar tooth structure" has been conducted. In this present study, STEM imaging and DFT calculations were combined to provide a complete crystallographic description of the *d* superstructure.

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92 SAMPLES

outcrops of the Helena Formation of the Mesoproterozoic Belt Group (Frank and Lyons, 1998). 94 The sample HHL-00H was chosen for STEM analysis. In the hand specimen of HHL-00H 95 96 (Figure 1), the thin sinuous vertical and horizontal ribbons with width of ~0.2-1.5cm intersect each other. The "molar tooth" is composed of clean and fine-grained calcite crystals with similar 97 size (~10 µm). The "molar tooth" host has three major phases: dolomite, calcite, and quartz 98 99 (Figure 2). K-feldspar, illite, chlorite, rutile, and apatite are also present in the "molar tooth" host rock. The XRD powder analysis of the "molar tooth" host rock shows that the dolomite is cation-100 ordered with the presence of sharp (105) and (009) peaks. The dolomite has a composition of 101 \sim Ca_{1.02}Mg_{0.98}(CO3)2 based on the d₁₀₄ value (2.890Å) and the relationship between d₁₀₄ values 102 and MgCO₃ content in ordered dolomite (Goldsmith and Graf, 1958; Zhang et al., 2010). 103

The "molar tooth" carbonate samples were collected near Hungry Horse Dam, Montana, from

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105 **METHODS**

106	Specimens for STEM measurements were prepared by ion milling. Ion milling was performed
107	with a Fischione 1010 ion mill operated at an accelerating voltage of 4 kV and an incident ion-
108	beam angle of 10°, followed by gentler milling at an accelerating voltage of 2.6 kV and an
109	incident angle of 10° in order to reduce surface amorphous material. The ion-milled samples
110	were lightly carbon coated. The microstructures in the dolomite crystals were examined by using
111	a spherical aberration-corrected FEG- STEM (Titan 80-200) operating at 200 kV at the
112	University of Wisconsin-Madison. This instrument can image single atoms with ~ 0.1 nm or
113	better spatial resolution in STEM mode. Probe current was set at 24.5 pA. Camera length for the
114	image acquisition was set at 160 mm. Collection angle of the HAADF detector for acquiring the
115	Z-contrast images ranges from 54 to 270 milliradians (mrads).
116	
117	The DFT calculations were performed by using the Vienna <i>ab initio</i> simulation package (VASP)
118	(Kresse et al., 1996). The general gradient approximation (GGA) with the Perdew, Burke, and
119	Ernzerhof (PBE) parameters was employed (Perdew et al., 1996). The projector-augmented wave
120	(PAW) method with an energy cutoff of 600 eV was used. A conventional hexagonal supercell
121	of calcite derived structures including 30 atoms or 6 chemical formula units was used. We tested
122	k-point convergence and a mesh of $3 \times 3 \times 1$ was found to be sufficient for the system. Two
123	starting structures for the d phase: one with calcite's unit cell parameters from experiments
124	(Graf, 1969) and another with dolomite's (Beran and Zemann, 1977). All the initial structures
125	were optimized using the static energy minimization scheme, where both the shape and volume
126	of the cell were allowed to relax. The structure with minimum energy calculated from the
127	previous step was further calculated by <i>ab initio</i> molecular dynamics simulations at 10K to better

explore the local minimum. The powder and electronic diffraction patterns of calculated

structures were generated by CrystalDiffract[®] and SingleCrystalTM respectively.

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131 HIGH-RESOLUTION (S)TEM OBSERVATIONS

In TEM images, the modulated microstructures with strong strain contrast are prevalent through 132 the "molar tooth" host dolomite. Calcite inclusions that are a common phenomenon in 133 sedimentary dolomites can be easily recognized because of being free of modulations (Figure 3). 134 Diffuse streaks along c^* occur in the electron diffraction pattern of the host dolomite (Figure 4). 135 Some maxima of the streaks are about one third of d_{006}^* (Fig. 4, and also see diffraction patterns 136 of Wenk and Zenger, 1983; Wenk and Zhang, 1985). According to the Fast Fourier Transform 137 (FFT) patterns of different areas in the dolomite images, the streaks in the diffraction pattern 138 139 come from the precipitates with linear features that are parallel to (001) (Figure 5). However, the precipitates themselves are arranged in such a way that they are roughly parallel to $(1\overline{1}4)$. The 140 141 observations above match the features of d reflections that were first described by Wenk and 142 Zenger (1983). It was proposed that the streaks parallel to c^* could be from stacking disorder 143 (Van Tendeloo et al., 1985).

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The dark areas of the precipitates in the bright field (BF) image under STEM mode become bright areas in the HAADF image (Figure 6), which suggests higher Ca contents in the precipitates than the host dolomite. The occurrence of *d* reflections in the FFT pattern of the HAADF image excludes the possibility that they are caused by multiple diffraction. In the HAADF of high magnification or Z-contrast image (Figure 7), the alternating bright and dark

150	layers along c axis represent the alternating Ca and Mg layers in the normal dolomite structure.
151	However, the <i>d</i> dolomite precipitate has the cation sequence of Ca-Ca-Mg-Ca-Ca-Mg- (Figure
152	7). The FFT pattern of this domain shows that the d reflections are attributed to this
153	superstructure. This observation is different from the previous model for the d superstructure in
154	which every third Mg layer is replaced by a Ca layer producing a sequence of Ca-Mg-Ca-Ca-Ca-
155	Mg- (Wenk and Zhang, 1985). But both sequences produce the same chemical stoichiometry of
156	$Ca_{0.67}Mg_{0.33}CO_3$. However, the repetition along the <i>c</i> -axis is doubled in the previous model with
157	respect to the observed Ca-Mg ordering in the Z-contrast image (Figure 7).
158	
159	Mg-bearing calcite precipitates are also observed in the dolomite (Figure 8). Along the $(\overline{1}02)$
160	trace, the calcite exsolution region has six consecutive Ca layers as opposed to the dolomite host
161	that has alternating Ca and Mg layers. In the line profile of the calcite region, some Ca columns
162	have slightly lower intensities than the pure Ca columns in the dolomite host, which suggests the
163	existence of a small amount of Mg in this calcite precipitate (Figure 8). This calcite exsolution is

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166 **DFT CALCULATIONS**

167 The structural parameters for the optimized *d* superstructure from DFT calculations are listed in 168 Table 1. Tables 2 and 3 compare the calculated lattice parameters of dolomite and calcite 169 structures with the reported experimental data. The calculated equilibrium volumes for both 170 dolomite and calcite are slightly smaller than the reported data, the underestimation of the *c* 171 parameter being the major contribution, but are still within the range of previous theoretical

similar to Mg-calcite precipitates in Ca-rich dolomite (Shen et al., 2013).

172 calculation data (see Table 2 and 3). The calculated values are for the structures at 0K. The reported experimental values were measured at ambient environment. Temperature could be a 173 174 factor for the small discrepancy between calculated and measured the unit cell volumes. A small 175 discrepancy between calculated and experimental values of lattice parameters is not uncommon for DFT calculations and may result from the use of an approximate exchange-correlation 176 potential (Hossain et al., 2011). In spite of this discrepancy, the trend found in experimental data 177 178 is maintained in our DFT calculations. The calculated d superstructure has an a parameter close to that of dolomite but has a *c* parameter closer to stoichiometric calcite. This finding is 179 180 consistent with the observation from the diffraction patterns of the d superstructure that the difference between *a* parameters of dolomite and the *d* superstructure is smaller than that 181 between *c* parameters, even though the DFT calculations are unconstrained bulk structure 182 183 calculations. This trend is reasonable because superstructure precipitates share the (001) interface 184 with the host dolomite. Smaller differences in the *a* dimensions between the host dolomite and d 185 superstructure would cause less strain at the interface.

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The C-O bonds in CO_3^{2-} groups are rigid and the C-O distances in both experimental and 187 188 calculation data are constant no matter what the actual composition of the carbonate mineral is. It is interesting to note that the C-O distances in the *d* superstructure are divided into two categories 189 (Table 1): 1.294Å when a CO_3^{2-} layer is between Ca^{2+} and Mg^{2+} layers along the c axis; and 190 1.297Å when a CO_3^{2-} layer is in between two Ca^{2+} layers. DFT calculations predict slightly 191 shorter Ca-O and Mg-O bonds in dolomite and calcite than experimental data (Table 2 and 3). 192 The Mg-O distance in the *d* superstructure increases from 2.069Å in calculated dolomite to 2.085 193 194 Å. The Ca-O distance in the d superstructure differs depending on the oxygen positions; the Ca-

195	O distances (2.374 Å) are larger when the oxygen ions are from CO_3^{2-} group sitting in between
196	Ca^{2+} and Mg^{2+} layers than those (2.355 Å) from CO_3^{2-} group between two Ca^{2+} layers. This is
197	because that the CO_3^{2-} layers are closer to Mg^{2+} layers than Ca^{2+} layers due to the smaller Mg^{2+}
198	radius. The inversion center is missing in the d superstructure while the c glide is retained. The
199	existence of (100) and (200) reflections in the diffraction pattern proves that it is not a
200	rhombohedral unit cell, but a primitive hexagonal (Figure 9C). Therefore, the space group is
201	determined to be P31c (No. 159). The atom coordinates and symmetry equivalent positions of
202	the <i>d</i> superstructure are reported in Table 1.

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204 The superstructure with cation sequence of Ca-Mg-Ca-Ca-Ca-Mg-Ca that was proposed to 205 explain the d reflections was referred to as δ structure (Wenk and Zhang, 1985; and Wenk et al., 206 1991) (see Figure 9A for details). The calculated δ structure has slightly larger unit cell parameters (a = 4.883Å and c = 16.281Å) than the *d* superstructure indicated here. The enthalpy 207 208 of a unit cell of δ structure is slightly higher than that of the *d* superstructure by 0.54 kJ/mol per MCO₃ unit. The previously proposed δ structure is unstable with respect to the current d 209 210 superstructure. The proposed δ structure model was based on the assumption that overlapped 211 diffraction patterns from host dolomite and d superstructure were from the precipitates only 212 (Wenk and Zhang, 1985). Careful examination of their diffraction pattern (Fig. 1A of Wenk and 213 Zhang, 1985) and our FFT pattern (Figure 9D), show that the position of 003 is not half way between 002 and 004 of the d superstructure, and the position of 009 is not midway between 008 214 and 0010 of the d superstructure. The d superstructure does not have reflections with odd l due to 215 216 its c-glide. A powder X-ray diffraction pattern with d_{104} value of 2.930Å for the d superstructure 217 is also calculated (Fig. 10).

218	The observed intermediate phases are analogous to those locally ordered domains in
219	mixed-layer clay minerals, such as interstratified chlorite/serpentine, chlorite/biotite, and
220	chlorite/pyrophyllite minerals (Banfield and Bailey, 1996; Xu and Veblen, 1996; Xu et al., 1996;
221	Wang and Xu, 2006). The observed intermediate phase between calcite and dolomite may be
222	described as interstratified calcite/dolomite, instead of δ -dolomite or <i>d</i> -dolomite.

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224 IMPLICATIONS

225	The calculated energies of calcite (Ca(CO3)), dolomite (CaMg(CO ₃) ₂), and d superstructure
226	$(Ca_2Mg(CO_3)_3)$ are 3620.64, 7132.37 and 10740.22 kJ/mol, respectively. The energy for the <i>d</i>
227	superstructure (Ca ₂ Mg(CO ₃) ₃) is higher (~4.2 kJ/mol per MCO ₃ unit) than the sum of the
228	calculated energies from dolomite (CaMg(CO ₃) ₂) and calcite (CaCO ₃) end members. The d
229	superstructures serve as a metastable phase with respect to the stoichiometric dolomite and
230	calcite. The <i>d</i> superstructures have a maximum of 4 repeats of Ca-Ca-Mg- along the c axis (<
231	35Å) in our sample and are only stable within dolomite host in the form of nano-precipitates. It
232	is very difficult for extra Ca ions in the dolomite structure to diffuse through carbonate layers
233	and to congregate to form lamellae parallel to the basal plane at low temperature (Wenk et al.,
234	1991; Shen et al., 2013). This explains all the reported occurrences of nano-precipitates of <i>d</i> -
235	reflections in natural samples. Wenk et al. (1991) summarized the various proposed
236	superstructures of Ca-Mg carbonates. By using Z-contrast imaging, we can image the cation
237	ordering directly and propose a more accurate structure model. By using the DFT method, we
238	can calculate the detailed structures and explore the energetics of these metastable nano-phases.

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240	or artifacts by using other methods.
239	The methods may be applied to understand many other nano-phases, where there are challenges

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Figure 1. (A)The outcrop shows the brittle deformation inside the carbonate layers with "molar tooth" structure. (B) Sinuous dark blue riboons of fine crystallined calcite ("molar tooth") exist in the dolomite host that is weathered into buff color. (C) Weathered surface of HHL-00H specimen . (D) "Molar tooth" becomes light color in fresh cut surface of HHL-00H specimen.



Figure 2. Powder XRD pattern of "molar tooth" host and "molar tooth" in sample HHL-00H.



Figure 3. TEM image shows that calcite inclusion that is free of modulations or strain contrast in Ca-rich dolomite host.



Figure 4. The diffuse streaks along c^* occur in the diffraction pattern of Ca-rich dolomite from Helena Formation, Montana. Some diffraction maxima occur at around 1/3 of d^*_{006} .



Figure 5. FFT patterns from TEM image show that the streaking and "*d*" reflections are from the precipitates (compare the FFT patterns from outlined areas of a and b).



Figure 6. The dark areas of d superstructure in the bright-filed image (left) under STEM mode become bright in HAADF image (right), which means higher Ca content in the precipitates of d superstructure than the host dolomite.



Figure 7. High magnification bright-field (A) and dark-filed (B) STEM images of d domains. Zcontrast (dark-filed) image shows that the Ca-rich precipitates have a cation sequence (bottomright) of Ca-Ca-Mg-Ca-Ca-Mg- along c axis as opposed to dolomite cation sequence as shown in the middle-right corner (B). The occurrence of streaking and splitting along c^* (i.e., dreflections) in the FFT patterns (bottom-left corner). 003 and 009 spots are from host dolomite, not from the d superstructure, because the position of 003 half way between 002 and 004 of the d superstructure, and the position of 009 is half way between 008 and 0010 of the d superstructure. Noise-filtered Z-contrast image (C) shows enhanced cation sequences in the d superstructure and dolomite host. Line profiles from line 1 (dolomite host) and 2 (d superstructure) also illustrate the ordering of Ca and Mg atoms.



Figure 8. Z-contrast image shows calcite exsolution lamellae // (001) in dolomite host. The line profile 1 has six consecutive Ca columns along ($\overline{1}02$) trace. The line profile 2 of the dolomite region shows the normal dolomite cation sequence of alternating Ca and Mg columns. The line profile 3 shows one repeat of *d* superstructure. The lines with arrows show the boundaries between calcite and dolomite and between dolomite and *d* superstructure. The atomic models for dolomite/calcite and dolomite/*d* phase interfaces are shown at the bottom. Carbonate group are not shown proportionally in order to highlight the cations.



Figure 9. Previously proposed structure model for " δ dolomite" (A), and calculated model for *d* superstructure (B), and corresponding calculated electron diffraction pattern (C), compared to the FFT pattern (D) from an area dominated by the *d* superstructure. Very weak 003 and 009 reflections also occur in FFT pattern due to contributions from the dolomite host. However, the position of 003 is not half way between 002 and 004 of the d superstructure, and the position of 009 is not half way between 008 and 0010 of the d superstructure. The *d* superstructure does not have (*h0l*) reflections with odd *l* due to its *c*-glide.



Figure 10: Calculated powder diffraction patterns of the *d* superstructure and stoichiometric dolomite (copper K-alpha radiation). The peak of (105) in dolomite does not appear in the *d* superstructure phase. However, the peak of (106) in the *d* superstructure phase does not appear in dolomite.



Table 1. Fractional coordinates of atoms, lattice parameters, and bond distances for the d superstructure calculated using DFT method.

Atom	X	У	Z	Atom	X	У	Z	
Cal	0.00000	0.00000	0.00000	Ca2	0.33333	0.66667	0.18457	
Mg	0.66667	0.33333	0.34229	C1	0.66667	0.33333	0.09229	
C2	0.00000	0.00000	0.26849	C3	0.33333	0.66667	0.41608	
01	0.66667	0.06803	0.09229	O2	0.28361	0.04018	0.26739	
O3	0.04972	0.62648	0.41719					
<i>a</i> : 4.879Å <i>c</i> : 16.260Å V: 335.226Å ³ Space group: <i>P31c</i> (no. 159)								
C-O bond (Å) : 1.294, 1.297 Ca-O bond (Å) 2.355, 2.374 Mg-O bond (Å) 2.085								

Table 2. Calculated structural parameters of dolomite as compared to experimental and previously calculated data.

Lattice parameters	This work	Experimental	Theoretical
a (Å)	4.810	$4.808^{a}, 4.812^{b}$	4.787 ^c , 4.877 ^d , 4.858 ^e
c (Å)	15.704	$16.010^{a}, 16.020^{b}$	15.55 ^c , 16.285 ^d , 16.109 ^e
$V(Å^3)$	314.611	320.504 ^a ,321.251 ^b	308.623°,335.409 ^d ,329.248 ^e
C-O bond (Å)	1.294	1.233 ^a , 1.286 ^b	1.286 ^c , 1.299 ^d
Ca-O bond (Å)	2.358	$2.405^{a}, 2.382^{b}$	2.328 ^c , 2.401 ^d
Mg-O bond (Å)	2.069	$2.114^{a}, 2.087^{b}$	2.071 ^c , 2.314 ^d
O-Mg-O bond angles			89.335, 90.645, 180 ^c
(°)	89.03, 90.97, 180	89.17, 90.83, 180 ^a	88.546,91.454,180 ^d

^a Graf; ^b Beran and Zemann; ^c Hossain et al., LDA functional was used; ^d Hossian et al., GGA functional was used; ^e Bakri and Zaoui, fitted to Birch-Murnaghan equation of state.

Table 3. Calculated structural parameters of calcite as compared to experimental and previously calculated data.

Lattice	This		
parameters	work	Experimental	Theoretical
a (Å)	4.995	4.990 ^a	5.061 ^b , 4.981 ^c
c (Å)	16.685	17.062 ^a	17.097 ^b ,15.902 ^c
$V(Å^3)$	360.435	367.916 ^a	379.279 ^b , 341.676 ^c
C-O bond (Å)	1.295	1.286^{a}	
Ca-O bond (Å)	2.336	2.357 ^a	

^a Graf; ^b Ayoub, Zaoui, and Berghout; ^c Aydinol et al.