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15 Alburnite,  $\text{Ag}_8\text{GeTe}_2\text{S}_4$ , a new mineral species from the Roşia Montana Au-Ag epithermal  
16 deposit, Apuseni Mountains, Romania

17 REVISION 1

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33

34 Abstract:

35 Alburnite, ideally  $\text{Ag}_8\text{GeTe}_2\text{S}_4$ , was discovered in the Cărnicele vein from the Roşia Montana  
36 epithermal Au-Ag ore deposit, Apuseni Mountains, Romania. The new mineral is associated  
37 with tetrahedrite, galena, pyrite, sphalerite, chalcopyrite and tellurides (hessite, altaite and  
38 sylvanite). Associated gangue minerals are rhodochrosite, quartz, calcite and rhodonite.

39 Alburnite was observed only at microscopic scale as rounded to sub-rounded grains, veinlets  
40 or irregular inclusions hosted mainly by tetrahedrite, hessite and rhodochrosite. Due to the  
41 small size of alburnite grains observed so far it was not possible to determine some  
42 macroscopic properties; reported properties are based on microscopic observations. The  
43 mineral has a metallic luster and is opaque. It is non-fluorescent and has an estimated Mohs  
44 hardness of 4. The mineral shows no cleavage. Density could not be measured because of the  
45 small grain size, but calculated density based on the empirical formula is  $7.828 \text{ g/cm}^3$ . In  
46 plane-polarized light in air, alburnite is grey-blue with a bluish tint. It shows no pleochroism  
47 or bireflectance in air. Between crossed polars alburnite is isotropic and internal reflections  
48 have not been observed in air. In intense light the mineral decomposes. Reflectance minimum  
49 values in air (in %) are: 470 nm 29.70; 546 nm 28.00; 589 nm 27.35; 650 nm 26.95. The  
50 average chemical composition based on 18 electron-microprobe analyses from 9 different  
51 grains in one polished section is (in wt%): Ag 65.49, Ge 4.82, Te 20.16, S 9.66, total, 100.13.  
52 The ideal formula of alburnite,  $\text{Ag}_8\text{GeTe}_2\text{S}_4$ , based on 15 *apfu* requires Ag 65.43, Ge 5.50, Te  
53 19.35, S 9.72, total 100.00wt%. Features of the crystal structure of alburnite were determined  
54 based on electron backscattered diffraction and transmission electron microscopy. Alburnite is  
55 cubic, space group  $F\bar{4}3m$ , with unit cell parameters  $a = 10.4(1) \text{ \AA}$ ,  $V = 1125(30) \text{ \AA}^3$ ,  $Z = 4$ .  
56 The strongest eight calculated XRD lines [ $d$  in  $\text{\AA}(I)(hkl)$ ] are: 6.004(67)(111), 3.136(48)(113),  
57 3.002(100)(222), 2.600(26)(004), 2.123(33)(224), 2.002(61)(115), 1.838(76)(044) and  
58 1.644(12)(026). The name of the new mineral alburnite is derived from the Latin name of the  
59 locality. Roşia Montana Au-Ag deposit was known during the Roman period as *Alburnus*  
60 *Maior*. The mineral and the mineral name have been approved by the Commission on New  
61 Minerals, Nomenclature and Classification, IMA 2012-073.

62

63 Keywords: alburnite, new mineral, germanium-tellurium sulfosalts, Roşia Montana, Apuseni  
64 Mountains, Romania.

65

66

## 67 **Introduction**

68

69 The Roşia Montana deposit has a long history of mining that started at least in Roman times.

70 The deposit was known during the Roman period (106-273 AD) as *Alburnus Maior* (Pošepný

71 1870; Sîntimbrean 1989) and kilometers of Roman workings are still preserved (Benea and

72 Tămaş 2010; Cauuet and Tămaş 2012). Recent data (Tămaş et al. 2009; Baron et al. 2011)

73 indicate that several Au-Ag ore bodies exploited by the Roman miners in *Alburnus Maior* had

74 Te and Ge traces. The name of the mineral is thus derived from the Latin name of the locality.

75 The first mention of argyrodite from Roşia Montana based on SEM analysis was made by

76 Tămaş (2002), who studied an ore body exploited by the Roman miners in Cărnic massif,

77 Roşia Montana, while the first report of a Te and Ge-bearing mineral indicated by EPMA was

78 made simultaneously by Tămaş et al. (2004), Ciobanu et al. (2004) and Bailly et al. (2005),

79 who interpreted it as Te-bearing argyrodite.

80

81 The new mineral was observed so far only at the microscopic scale. The biggest alburnite

82 crystal is close to 50 micrometers in size, but usually grains are smaller and range between 10

83 and 30 micrometers. The small grain size is the main reason for some limitations of the

84 mineral description (i.e., single-crystal study, macroscopic physical properties etc.).

85

86 The presence of Ge-bearing minerals in Roşia Montana was postulated some time ago.

87 Andronescu (1962) studied the flotation concentrates from Roşia Montana and identified Ge

88 enrichment in pyrite-rich concentrates with low gold grades, but was not able to identify the  
89 Ge-bearing mineral. Since the first electron microscope evidence of a possibly new Te and  
90 Ge-bearing mineral, dating back to Autumn 2000, this mineral was observed in many samples  
91 from the Roşia Montana ore deposit, mostly from Ag-rich ore bodies (veins and breccias),  
92 some of them mined during Roman times and some of them mined only recently. The above  
93 observations indicate that the new mineral, in spite of its small grain size, occurs at the ore  
94 deposit scale. LA-ICPMS analyses carried out on sphalerite in alburnite-bearing veins from  
95 Roşia Montana (Cook et al. 2009), indicate an average of 73 ppm Ge (as well as 366 ppm Ga  
96 and more than 4wt% Mn) confirming the Ge-rich character of at least some parts of the Roşia  
97 Montana deposit.

98

99 The mineral and the mineral name have been approved by the Commission on New Minerals,  
100 Nomenclature and Classification (CNMNC), IMA 2012-073. The material used for the  
101 description of the new mineral was cut in several slices and nine polished sections. Alburnite  
102 was observed in six of these polished sections. Cotype material consists of three ore fragments  
103 (3230 – two fragments, and 3231), and five polished sections (3231B, 3231C, 1064a, 1064b  
104 and 1064c). Reflectance measurements were made on 1064a, but alburnite grains exist on the  
105 other sections as well. Microprobe data were gathered from the section 3231C, while  
106 crystallographic data (SEM and TEM) were obtained for both 3231C and 1064a sections.  
107 Alburnite grains were recognized optically in all polished sections selected as the cotypes  
108 (3231B, 3231C, 1064a, 1064b and 1064c). These sections were made from the three  
109 remaining ore fragments (3230 and 3231).

110

111 The three ore fragments (2x3230 and 3231) and four polished sections (3231B, 3231C, 1064a,  
112 and 1064c) are deposited in the systematic collection of the Mineralogical Museum,

113 Department of Geology, Faculty of Biology and Geology, Babeş-Bolyai University, 1 M.  
114 Kogălniceanu Street, Cluj-Napoca, Romania. Sample 3231 and the related polished sections  
115 are registered under inventory number 71a/1, and sample 3230 under inventory number 71a/2.  
116 The polished section 1064b is hosted by Zentrale Elektronen-Mikrosonde, Institute of  
117 Geology, Mineralogy and Geophysics, Ruhr University, Bochum, Germany.

118

### 119 **Occurrence**

120

121 The new mineral occurs in the Roşia Montana epithermal ore deposit, Apuseni Mountains,  
122 Romania (Fig. 1). Roşia Montana represents Europe's largest gold deposit (Manske et al.  
123 2006) with resources exceeding 400 Mt of ore at an average grade of 1.3 g/t Au. The deposit  
124 is located in the central-eastern part of Apuseni Mountains (Romania), within the so-called  
125 Golden Quadrilateral (Ghiţulescu and Socolescu 1941), an area hosting a large number of  
126 precious metals ore deposits. An overview of the Neogene tectonic setting, origin of the  
127 magmatic/volcanic activity and related porphyry copper and epithermal Au-Ag deposits from  
128 the Apuseni Mountains was made by Neubauer et al. (2005) and Roşu et al. (2005).

129

130 Roşia Montana is a low- to intermediate sulfidation Au-Ag deposit (Mârza et al. 1997; Tămaş  
131 and Bailly 1998, 1999; Leary et al. 2004; Tămaş et al. 2006) related to  $13.65 \pm 0.63$  Ma dacite  
132 intrusions (Roşu et al. 2004). The Cretaceous basement (shales, sandstones and micro-  
133 conglomerates) is partly overlain by a vent breccia interbedded with Neogene sedimentary  
134 rocks (clays, marls, sandstones, limestones, gypsum). Lava flows and volcanoclastics of  
135 andesitic composition, dated at  $9.3 \pm 0.47$  Ma (Roşu et al. 1997) cover the northern and eastern  
136 part of Roşia Montana area (Fig. 2). Mineralized and barren phreatomagmatic breccia  
137 structures as well as hydrothermal breccias are widespread at ore deposit scale (Tămaş 2010).

138

139 Alburnite was discovered in the Cârnicel vein, an intermediate sulfidation structure located in  
140 the southern part of Cârnic massif at Roşia Montana. The vein is accessible only underground  
141 where it crops out in the so-called Vercheş level or +853 mining level. The Cârnicel vein is  
142 hosted in dacitic vent breccias. The vein is oriented almost N-S and dips to the west ( $\sim 70^\circ$ ).  
143 The vein is exposed over 170 m and has an average width of about 25 cm. It has a banded  
144 texture with ore-rich layers mixed with carbonate-dominated gangue layers.

145

146 The ore mineral assemblage containing the new mineral is dominated by tetrahedrite  
147 associated with other sulfides (galena, pyrite, sphalerite and rarely chalcopyrite), and  
148 tellurides, mainly hessite but also altaite and sylvanite. Associated gangue minerals are  
149 rhodochrosite, quartz, calcite and rhodonite. Taking into account reflected light microscopic  
150 observations, the paragenetic sequence of common sulfides hosting Ag - Te - Ge minerals  
151 from Cârnicel vein is shown in Fig. 3.

152

### 153 **Appearance and physical properties**

154

155 To date, alburnite has been observed only under the microscope (Fig. 4). It occurs as rounded  
156 to sub-rounded grains (Fig. 4a,b), veinlets (Fig. 4c,d,l) or irregular inclusions (Fig. 4a,e,j)  
157 reaching a maximum size of 50 micrometers, hosted usually in tetrahedrite, as well as in  
158 hessite or located at the contact between these two minerals. It was also observed within  
159 galena or between galena and tetrahedrite (Fig. 4a). Large alburnite grains are hosted in late  
160 rhodochrosite veinlets that crosscut the tetrahedrite masses (Fig. 4f). Two-phase inclusions,  
161 e.g. hessite-alburnite and galena-alburnite, are more frequent than monophasic inclusions with  
162 alburnite alone (Fig. 4g). Alburnite-rhodochrosite assemblages occur within former vugs (Fig.

163 4i,j). Alburnite also forms intermingled assemblages with hessite (Fig. 4e), or hessite-  
164 rhodochrosite (Fig. 4h). The most complex mineral association with alburnite are assemblages  
165 with hessite, chalcopyrite, and galena with or without sphalerite hosted by tetrahedrite (Fig.  
166 4b,h). Hessite grains with micrometer-size alburnite inclusions may host electrum (Fig. 4k).  
167 Electrum was also observed as inclusions in alburnite and along the contact of alburnite with  
168 its tetrahedrite host (Fig. 4l).

169

170 Due to the small size of alburnite grains observed so far (less than 50 micrometers) it was not  
171 possible to determine the macroscopic properties of the new mineral. However, some  
172 properties are reported based on microscopic observations. The mineral has a metallic luster  
173 and it is opaque. It is non-fluorescent and has an estimated Mohs hardness of 4. The mineral  
174 shows no cleavage. Density could not be measured because of the small grain size, but  
175 calculated density based on the empirical formula is  $7.828 \text{ g/cm}^3$ .

176

### 177 **Optical properties**

178

179 In polished sections illuminated by plane-polarized light, alburnite is grey with a bluish tint  
180 (Fig. 4), having lower reflectance than tetrahedrite, the main host mineral. Te-bearing  
181 canfieldite -  $\text{Ag}_8\text{Sn}(\text{S},\text{Te})_6$ , first mentioned by Harris and Owens (1971), has the same bluish  
182 tint in plane polarized light in air (Soeda et al. 1984; Wimmers 1985). By comparison, Te-free  
183 argyrodite and canfieldite have a pinkish brown to violet tint in reflected light (Uytenbogaardt  
184 and Burke 1971). Alburnite shows neither pleochroism nor bireflectance (in air). Between  
185 crossed polars alburnite is isotropic. Internal reflections have not been observed in air. In  
186 intense light the mineral decomposes (Fig. 4l).

187

188 Reflectance measurements were carried out in air using a Hamamatsu C100838CA  
189 spectrometer attached to a Leitz Orthoplan microscope. The device was calibrated with a  
190 Zeiss SiC reflectance standard (No. 787). A 20x objective with an effective numerical  
191 aperture of 0.2 was used. The diameter of the measured area was about 10  $\mu\text{m}$ . The original  
192 spectrum was recorded at 712 different wavelengths. Reflectance minimum values and a  
193 selection of 16 values between 400 and 700 nm in steps of 20 nm together with those for  
194 argyrodite, canfieldite and putzite (Criddle and Stanley 1986, 1993; Paar et al. 2004) are given  
195 in Table 1.

196

197 The new mineral shows systematically greater reflectance than argyrodite and canfieldite (~  
198 3%). As compared with putzite, alburnite has a lower reflectance between 400 and 440 nm  
199 and is more reflective at wavelengths greater than 440 nm (Fig. 5).

200

#### 201 **Chemical data**

202

203 Chemical analyses (18 from 9 different grains in one polished section) were carried out using  
204 a JEOL JXA-8200 electron microprobe (WDS mode, 15 kV, 20 nA, < 1  $\mu\text{m}$  beam diameter,  
205 and a counting time of 30 seconds). Preliminary EDS analysis revealed the presence of only  
206 four elements, Ag, Ge, Te and S. Naumannite ( $\text{Ag}_2\text{Se}$ ), pyrite ( $\text{FeS}_2$ ), cassiterite ( $\text{SnO}_2$ ), Ge  
207 metal and Te metal were used as standards. Sn contents in all analyses were below detection  
208 limits (0.03 wt% Sn). We used the following lines for the analysis:  $SK\alpha$ ,  $AgL\alpha$ ,  $GeL\alpha$ ,  $SnL\alpha$   
209 and  $TeL\alpha$ . Detection limits (wt%) were 0.09 for Ag, 0.03 for Ge, 0.05 for S, 0.03 for Sn, and  
210 0.03 for Te. Analytical data are given in Table 2.

211

212 The empirical formula of alburnite based on 15 *apfu* is  $\text{Ag}_{8.04}\text{Ge}_{0.88}\text{Te}_{2.09}\text{S}_{3.99}$ . The simplified  
213 formula is  $\text{Ag}_8\text{GeTe}_2\text{S}_4$ , which requires Ag 65.43, Ge 5.50, Te 19.35, S 9.72, total  
214 100.00wt%.

215

## 216 **Crystallography**

217

218 The new mineral occurs only as microscopic grains up to 50 micrometers in size. All  
219 diffraction data were obtained by electron back scatter diffraction (EBSD) and electron  
220 diffraction (ED).

221

222 Argyrodite-type  $\text{M}_8\text{XC}_6$  structures (M = Ag, Cu; X = Si, Ge, Sn etc.; C = S, Te, Se) have a  
223 common cubic high temperature  $\gamma$  phase, space group  $F\bar{4}3m$ . At lower temperatures, several  
224 ordered structures exist, such as the  $\beta''$  structure with orthorhombic symmetry in which  
225 argyrodite crystallizes at room temperature. The phase with the lowest transformation  
226 temperature to the cubic structure is the synthetic compound  $\text{Ag}_8\text{GeTe}_6$  ( $-29^\circ\text{C}$ ,  $\beta \rightarrow \gamma$ ),  
227 whereas the cubic modification of argyrodite *sensu stricto* is only stable above  $234^\circ\text{C}$   
228 (Gorochov 1968; Katty et al. 1981). Tellurium on the C site seems to extend the stability field  
229 of the  $\gamma$  phase to lower temperatures. Putzite has the cubic  $\gamma$  structure at room temperature.  
230 The close synthetic analogue of putzite,  $\text{Cu}_8\text{GeS}_6$ , has also a low transformation temperature  
231 ( $55^\circ\text{C}$ ,  $\beta' \rightarrow \gamma$ ). Diffraction data for the new mineral can only be indexed as a cubic phase,  
232 which makes the  $\gamma$  phase structure a likely candidate.

233

234 The  $\gamma$ -phase structure ( $Z = 4$ ) has three partially filled metal positions ( $M$ ; Wyckoff positions  
235  $24g$ ,  $48h^1$ ,  $48h^2$ ), one germanium position ( $4b$ ) and 3 positions for the chalcogenides, two of  
236 them completely filled ( $4d$ ,  $16e^1$ ) and one with an occupancy of 0.25 ( $16e^2$ ). The  $16e^1$  position

237 contains thus 2/3 of the total chalcogenides in the formula unit and the two other sites ( $4d$  and  
238  $16e^2$ ) are occupied by the remaining 1/3. The ratio of 0.5 between tellurium and sulfur in the  
239 new mineral may reflect the distribution of the chalcogenides among the three sites, i.e.,  
240 tellurium occupying the  $16e^2$  and  $4d$  sites and sulfur the  $16e^1$  site.

241

#### 242 **Electron backscattered diffraction**

243

244 Two ore samples were diamond polished and then chemically-mechanically lapped with a  
245 colloidal silica suspension (20 nm) in order to remove polishing damage. EBSD  
246 measurements, carried out on polished sections 3231C and 1064a, were performed using a  
247 FEI XL30 Sirion FEG SEM equipped with an EDAX-EBSD OIM system and a DigiView  
248 detector, and using Delphi software for data interpretation. The samples were tilted  $70^\circ$  from  
249 the horizontal SEM stage plane. The indexing quality is given by the confidence index (CI),  
250 which allows to discriminate between different solutions (values ranging between 0 and 1,  
251 solutions  $CI > 0.2$  are considered to be reliable, Mauler et al. 2000), and the fit factor, which  
252 describes the angular deviation between calculated and measured Kikuchi lines (good fit for  
253 deviations  $< 1.0^\circ$ ).

254

255 Patterns of tetrahedrite present in the matrix were collected to calibrate the EBSD system.  
256 Calibration settings were verified for sphalerite, galena and rhodochrosite. Afterwards, 25  
257 patterns from 7 different alburnite grains were collected (Fig. 6). The low temperature  
258 orthorhombic structure of argyrodite gave no match at all between calculated and measured  
259 patterns. Matching patterns were obtained for all 25 patterns using the  $\gamma$  phase structure.

260

#### 261 **Transmission electron microscopy**

262

263 Tetrahedrite grains, the main host of the new phase, were crushed in a mortar and the  
264 resulting powder deposited onto carbon-coated copper grids. The samples were analyzed with  
265 a Philips CM200 Transmission Electron Microscope operated at 200 kV, calibrated using a  
266 gold standard. Two grains of the new phase were found and analyzed by selected area  
267 electron diffraction. The diffractograms were indexed for a cubic phase with a lattice  
268 parameter of  $10.4 \pm 0.1 \text{ \AA}$  using JEMS software (Stadelmann 1987).

269

270 Due to the large thickness of the grains, we were only able to obtain good (001) SAED  
271 patterns (Fig. 7). They can be indexed based on a cubic cell and are compatible with an *F*-  
272 centered lattice. Simulation (JEMS software; Stadelmann 1987) of the (001) diffraction  
273 pattern using the  $\gamma$  phase structure with the chalcogenide distribution indicated above  
274 reproduced for several thicknesses the characteristic square pattern with weak (220)  
275 reflections. Other chalcogenide distribution were tested i.e. sulfur on  $16e^2$  or on  $4d$  and one of  
276 the  $16e$  sites. None of these distributions gave the characteristic weak intensity observed for  
277 the  $2h2k0$  reflections.

278

279 Powder XRD data could not be collected, but a calculated data set is given in Table 3. The  
280 powder pattern was calculated with the software “Crystaldiffract V5.2“ (Crystallmaker  
281 Software Ltd) for the  $\gamma$ -phase structure taking the atomic coordinates given by Paar et al.  
282 (2004) for putzite, the *a*-lattice parameter determined from the SAED pattern ( $10.4 \text{ \AA}$ ) with the  
283 chalcogenides distributed onto the following sites: Te on  $4d$  (full occupation) and  $16e^2$   
284 (occupation: 0.25) and sulfur on  $16e^1$  (full occupation). The peak positions were determined  
285 for monochromatic Cu- $K\alpha_1$  radiation. Peak intensities were determined using scattering

286 factors given by Doyle and Turner (1968) and taking in account Lorentz polarization and  
287 reflection multiplicity.

288

289

290 **Relation to other species**

291

292 Alburnite is considered to be part of the argyrodite-canfieldite series. The stoichiometry of the  
293 new mineral is the same as for the minerals of the argyrodite group (Strunz classification  
294 02.BA.35), which contains three members, one with orthorhombic symmetry, argyrodite  
295  $\text{Ag}_8\text{GeS}_6$ , one with cubic symmetry, putzite,  $(\text{Cu}_{4.7}\text{Ag}_{3.3})_{\Sigma 8}\text{GeS}_6$  (Paar et al. 2004), and one  
296 whose structure has not been determined but is usually given as orthorhombic, canfieldite,  
297  $\text{Ag}_8\text{SnS}_6$ . Recently a further Ge-bearing sulfosalt related to the argyrodite group has been  
298 described from Argentina, catamarcaite  $\text{Cu}_6\text{GeWS}_8$  (Putz et al. 2006). The chemical  
299 composition of the new mineral is closest to argyrodite, in which 2 of the 6 sulfur atoms are  
300 replaced by tellurium, though with cubic symmetry.

301

302 Tellurium as minor element was already reported for different species of the argyrodite-  
303 canfieldite group. Te-bearing canfieldite was first reported by Harris and Owens (1971) from  
304 Revelstoke, Canada, with the following composition (wt%): Ag 65.12(100), Sn 10.57(50), Ge  
305 -, Te 8.69(50), S 13.95(100), total 98.33. Wimmers (1985) reported Te-bearing canfieldite  
306 from Panasqueira, Portugal, with the following average composition (wt%): Ag 57.39, Sn 9.8,  
307 Te 19.095, S 9.86, Cu 2.44, Fe 1.95, Zn 0.08, total 100.62.

308

309 A Te-bearing silver-germanium sulfosalt found in association with hessite, Ag-Au tellurides,  
310 altaite and base metal sulfides (tetrahedrite, galena, sphalerite, chalcopyrite, etc.) was first

311 mentioned by Tămaş et al. (2004) from Roşia Montana, based on EPMA data, but no  
312 crystallographic data were determined. Based on the sole chemical data, Ciobanu et al.  
313 (2004), Bailly et al. (2005), and Tămaş et al. (2006) interpreted this mineral as a Te-bearing  
314 argyrodite. Tămaş et al. (2006) suggested the existence of a possible solid solution between  
315  $\text{Ag}_8\text{GeS}_6$  and  $\text{Ag}_8\text{GeTe}_6$ , which is known only as a synthetic phase (Boucher et al. 1993). The  
316 XRD data obtained by Boucher et al. (1993) for synthetic  $\gamma\text{-Ag}_8\text{GeTe}_6$  (space group  $F\bar{4}3m$ ,  $a$   
317 = 11.5656 Å,  $Z = 4$ ) are given in Table 4 for comparison.

318

319 Gu et al. (2012) reported a new member of the argyrodite-canfieldite series with orthorhombic  
320 symmetry, tellurocanfieldite  $\text{Ag}_8\text{SnTe}_2\text{S}_4$ , from Bajiazi Pb-Zn ore deposit, Jianchang County,  
321 Liaoning Province, China. Although initially approved by IMA Commission on New  
322 Minerals, Nomenclature and Classification as the new mineral IMA 2012-013, this approval  
323 was subsequently withdrawn (Williams et al. 2012) following a single-crystal study on Te-  
324 rich canfieldite from Lengenbach quarry, Binntal, Switzerland (Bindi et al. 2012). This study  
325 indicated that this phase is cubic and not orthorhombic and that the distribution of substituted  
326 Te atoms is disordered over the three sulfur sites. The problem is that the structure of pure  
327 canfieldite has not been determined so far. It is usually given as orthorhombic. If this should  
328 turn out to be true, the structural relationship between the rejected “tellurocanfieldite” and  
329 canfieldite may be similar to the relationship between alburnite and argyrodite, where  
330 tellurium substituting for sulfur stabilizes the cubic structure to lower temperatures.

331

332

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341

342

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464

465

466 Figure captions

467

468 Figure 1. (a) General map of the Carpathians and the Apuseni Mountains; (b) location of  
469 Roşia Montana ore deposit in the Apuseni Mountains.

470

471 Figure 2. Simplified geology of Roşia Montana area (courtesy of Roşia Montana Gold  
472 Corporation, with changes).

473

474 Figure 3. Paragenetic sequence of Cârnicel vein, Roşia Montana, showing the main species  
475 present in alburnite-bearing mineralization. Abbreviations: Alb- alburnite; Ccp – chalcopryrite;  
476 El – electrum; Gn – galena; Hes – hessite; Qz - quartz; Sp – sphalerite; Ttr – tetrahedrite; Rds  
477 – rhodochrosite.

478

479 Figure 4. Optical microscope images (plane polarized light). (a) Rounded and irregular  
480 alburnite inclusions in galena hosted by tetrahedrite and at the contact between galena and  
481 tetrahedrite, sometimes associated with hessite (sample 1064C); (b) alburnite-hessite-  
482 chalcopryrite-galena inclusion related to sphalerite hosted in tetrahedrite (sample 3231C); (c)  
483 irregular alburnite grains and veinlets with hessite, galena, and chalcopryrite within

484 tetrahedrite (sample 1064B); (d) alburnite-hessite and chalcopyrite deposited in vugs  
485 (presently filled with rhodochrosite) and as veins in tetrahedrite (sample 1064A); (e) irregular  
486 inclusion of alburnite in hessite associated with chalcopyrite hosted in tetrahedrite (sample  
487 1064B); (f) alburnite attached to tetrahedrite in a rhodochrosite vein crosscutting tetrahedrite  
488 (sample 1064A); (g) galena-alburnite inclusion in tetrahedrite (sample 3231C); (h) mixed  
489 alburnite-hessite-rhodochrosite inclusions in tetrahedrite, and alburnite-chalcopyrite-  
490 sphalerite in a void filled with rhodochrosite in tetrahedrite; inclusions of hessite-alburnite ±  
491 chalcopyrite within tetrahedrite (sample 3231E); (i) alburnite-rhodochrosite associated with  
492 galena, hessite, sphalerite, chalcopyrite in tetrahedrite (sample 1064B); (j) detail of (i); (k)  
493 hessite-alburnite inclusion with associated electrum in a vug at the contact of tetrahedrite-  
494 sphalerite (sample 3231C); (l) an alburnite bleb hosting electrum; alburnite is associated with  
495 hessite and chalcopyrite and occurs in the same mineral association also as veinlets in  
496 tetrahedrite; the grey spots on alburnite formed within a few minutes in intense light (sample  
497 1064A). Abbreviations: Alb- alburnite; Ccp – chalcopyrite; El – electrum; Gn – galena; Hes –  
498 hessite; Sp – sphalerite; Ttr – tetrahedrite; Rds – rhodochrosite. The scale bar is 100  
499 micrometers for g, 50 micrometers for a,b,c,d,f,h,i,l, and 20 micrometers for e,j,k.

500

501 Figure 5. Reflectance spectra in air for alburnite, putzite (Paar et al. 2004), argyrodite (Criddle  
502 and Stanley, 1993), and canfieldite (Criddle and Stanley 1986).

503

504 Figure 6. Example of EBSD pattern indexation of the 25 patterns indexed using the  $\gamma$ -phase  
505 structure. The confidence indices and the fit factors are indicated; Euler angles  $\phi_1$  298° PHI  
506 154°  $\phi_2$  203°.

507

508 Figure 7. Experimental (001) SAED pattern (left) and simulated (001) diffraction pattern  
509 (right); thickness of slice 125nm. The simulated diffraction pattern was obtained using the  
510 calibrations (18,-4,0), LZ 0, center of Laue circle (0,0,0), zone axis: [0,0,1], tilt angle 0.00°,  
511 AV 200 kV, CL 1395 mm, ZA [0,0,1], FN [0,0,1]. The scale bar is 20 nm for each image.

512

513

514 Table 1. Reflectance data for alburnite compared with putzite (Paar et al. 2004), argyrodite  
515 (Criddle and Stanley 1993), and canfieldite (Criddle and Stanley 1986).

| $\lambda$ (nm) | R (%)<br>alburnite | R (%)<br>putzite | R (%)<br>argyrodite | R (%)<br>canfieldite |
|----------------|--------------------|------------------|---------------------|----------------------|
| 400            | 32.00              | 33.50            | 29.30               | 28.50                |
| 420            | 31.20              | 31.60            | 27.60               | 27.75                |
| 440            | 30.60              | 30.60            | 26.60               | 27.00                |
| 460            | 30.00              | 29.40            | 25.90               | 26.30                |
| 470            | 29.70              | 28.90            | 25.60               | 25.90                |
| 480            | 29.30              | 28.60            | 25.20               | 25.70                |
| 500            | 29.00              | 27.80            | 24.80               | 25.20                |
| 520            | 28.50              | 27.20            | 24.50               | 24.90                |
| 540            | 28.10              | 26.70            | 24.30               | 24.70                |
| 546            | 28.00              | 26.50            | 24.20               | 24.60                |
| 560            | 27.75              | 26.30            | 24.20               | 24.50                |
| 580            | 27.50              | 26.00            | 24.20               | 24.50                |
| 589            | 27.35              | 25.80            | 24.10               | 24.50                |
| 600            | 27.20              | 25.70            | 24.00               | 24.40                |
| 620            | 27.10              | 25.40            | 24.00               | 24.40                |
| 640            | 27.00              | 25.30            | 23.90               | 24.50                |
| 650            | 26.95              | 25.30            | 24.00               | 24.60                |
| 660            | 26.90              | 25.20            | 24.00               | 24.60                |
| 680            | 26.85              | 25.20            | 24.00               | 24.70                |
| 700            | 26.80              | 25.20            | 24.10               | 24.90                |

516

517

518 Table 2. Analytical data (wt%) for alburnite, Cârnicel vein, Roşia Montana, Romania.

| No. | Ag    | Ge   | Te    | S    | Total  |
|-----|-------|------|-------|------|--------|
| 1   | 65.32 | 4.91 | 20.68 | 9.35 | 100.26 |
| 2   | 65.65 | 5.06 | 19.80 | 9.68 | 100.19 |
| 3   | 65.36 | 5.09 | 20.22 | 9.59 | 100.26 |
| 4   | 66.03 | 5.02 | 20.03 | 9.57 | 100.65 |

|         |       |      |       |      |        |
|---------|-------|------|-------|------|--------|
| 5       | 64.74 | 4.74 | 19.75 | 9.51 | 98.74  |
| 6       | 65.05 | 4.77 | 19.99 | 9.53 | 99.34  |
| 7       | 65.48 | 4.65 | 20.69 | 9.58 | 100.40 |
| 8       | 65.35 | 4.97 | 19.99 | 9.75 | 100.06 |
| 9       | 66.02 | 4.86 | 20.20 | 9.68 | 100.76 |
| 10      | 65.64 | 4.80 | 20.32 | 9.70 | 100.46 |
| 11      | 65.48 | 4.82 | 19.73 | 9.81 | 99.84  |
| 12      | 65.20 | 4.83 | 19.38 | 9.95 | 99.36  |
| 13      | 65.14 | 4.70 | 20.17 | 9.73 | 99.74  |
| 14      | 65.92 | 4.81 | 20.27 | 9.74 | 100.74 |
| 15      | 65.73 | 5.00 | 19.96 | 9.65 | 100.34 |
| 16      | 65.79 | 4.64 | 20.72 | 9.68 | 100.83 |
| 17      | 66.49 | 4.13 | 20.53 | 9.43 | 100.58 |
| 18      | 64.40 | 4.87 | 20.38 | 9.96 | 99.61  |
| average | 65.49 | 4.82 | 20.16 | 9.66 | 100.13 |

519

520

521 Table 3. Calculated powder XRD data for alburnite.

| <i>h</i> | <i>k</i> | <i>l</i> | <i>d</i>     | <i>I<sub>rel</sub></i> |
|----------|----------|----------|--------------|------------------------|
| <b>1</b> | <b>1</b> | <b>1</b> | <b>6.004</b> | <b>67</b>              |
| 0        | 0        | 2        | 5.200        | 3                      |
| 0        | 2        | 2        | 3.677        | 11                     |
| <b>1</b> | <b>1</b> | <b>3</b> | <b>3.136</b> | <b>48</b>              |
| <b>2</b> | <b>2</b> | <b>2</b> | <b>3.002</b> | <b>100</b>             |
| <b>0</b> | <b>0</b> | <b>4</b> | <b>2.600</b> | <b>26</b>              |
| 1        | 3        | 3        | 2.386        | 2                      |
| 0        | 2        | 4        | 2.326        | 11                     |
| <b>2</b> | <b>2</b> | <b>4</b> | <b>2.123</b> | <b>33</b>              |
| <b>1</b> | <b>1</b> | <b>5</b> | <b>2.002</b> | <b>61</b>              |
| <b>0</b> | <b>4</b> | <b>4</b> | <b>1.838</b> | <b>76</b>              |
| 1        | 3        | 5        | 1.758        | 9                      |
| 0        | 0        | 6        | 1.733        | 3                      |
| <b>0</b> | <b>2</b> | <b>6</b> | <b>1.644</b> | <b>12</b>              |
| 3        | 3        | 5        | 1.586        | 8                      |
| 2        | 2        | 6        | 1.568        | 8                      |
| 4        | 4        | 4        | 1.501        | 12                     |
| 1        | 1        | 7        | 1.456        | 6                      |
| 0        | 4        | 6        | 1.442        | 1                      |
| 2        | 4        | 6        | 1.390        | 1                      |
| 1        | 3        | 7        | 1.354        | 6                      |
| 0        | 0        | 8        | 1.300        | 6                      |
| 3        | 3        | 7        | 1.271        | 3                      |
| 0        | 2        | 8        | 1.261        | 3                      |
| 0        | 6        | 6        | 1.226        | 7                      |
| 1        | 5        | 7        | 1.201        | 6                      |
| 2        | 6        | 6        | 1.193        | 12                     |

|   |   |   |       |   |
|---|---|---|-------|---|
| 0 | 4 | 8 | 1.163 | 9 |
| 1 | 1 | 9 | 1.142 | 7 |
| 2 | 4 | 8 | 1.135 | 1 |
| 4 | 6 | 6 | 1.109 | 5 |

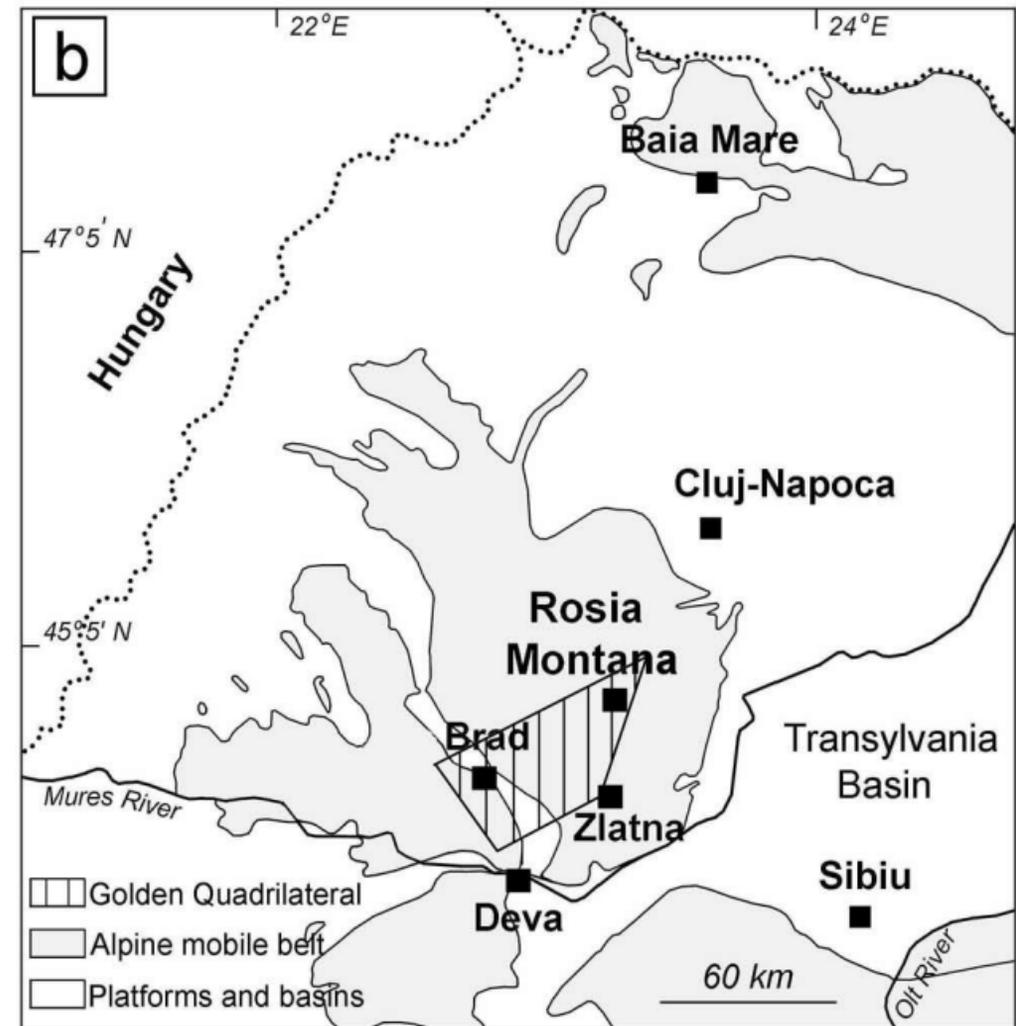
522

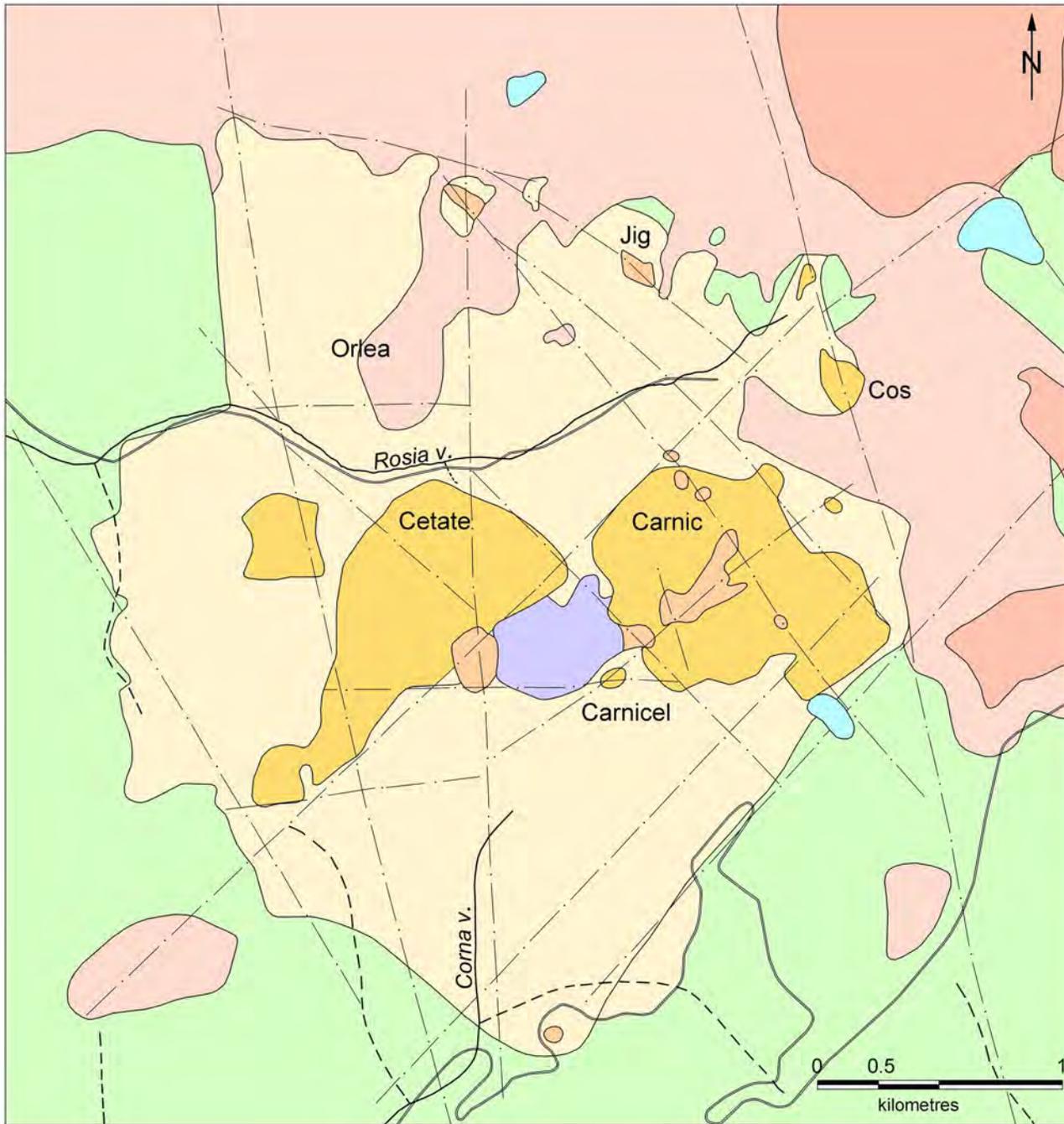
523

524 Table 4. XRD data for synthetic  $\gamma$ -Ag<sub>8</sub>GeTe<sub>6</sub> at room temperature (Boucher et al. 1993).

| <i>h</i> | <i>k</i> | <i>l</i> | <i>d</i> | <i>I</i> <sub>rel</sub> |
|----------|----------|----------|----------|-------------------------|
| 1        | 1        | 1        | 6.675    | 4                       |
| 2        | 2        | 0        | 4.090    | 4                       |
| 3        | 1        | 1        | 3.488    | 39                      |
| 2        | 2        | 2        | 3.340    | 100                     |
| 3        | 3        | 1        | 2.654    | 14                      |
| 4        | 2        | 2        | 2.361    | 37                      |
| 3        | 3        | 3        | 2.226    | 93                      |
| 4        | 4        | 0        | 2.044    | 35                      |
| 5        | 3        | 1        | 1.955    | 18                      |
| 4        | 4        | 2        | 1.927    | 24                      |
| 6        | 2        | 0        | 1.829    | 6                       |
| 5        | 3        | 3        | 1.764    | 14                      |
| 6        | 2        | 2        | 1.744    | 13                      |
| 5        | 5        | 1        | 1.620    | 8                       |
| 6        | 4        | 2        | 1.545    | 2                       |
| 5        | 5        | 3        | 1.506    | 11                      |
| 8        | 0        | 0        | 1.446    | 17                      |
| 7        | 3        | 3        | 1.413    | 1                       |
| 8        | 2        | 2        | 1.363    | 4                       |
| 5        | 5        | 5        | 1.336    | 9                       |
| 6        | 6        | 2        | 1.327    | 11                      |
| 8        | 4        | 0        | 1.293    | 2                       |
| 7        | 5        | 3        | 1.270    | 10                      |
| 6        | 6        | 4        | 1.233    | 3                       |
| 9        | 3        | 1        | 1.213    | 5                       |
| 8        | 4        | 4        | 1.180    | 10                      |
| 7        | 5        | 5        | 1.162    | 2                       |
| 8        | 6        | 2        | 1.134    | 2                       |
| 7        | 7        | 3        | 1.118    | 4                       |
| 6        | 6        | 6        | 1.113    | 7                       |
| 9        | 5        | 3        | 1.079    | 1                       |
| 7        | 7        | 5        | 1.043    | 2                       |
| 9        | 5        | 5        | 1.010    | 4                       |
| 10       | 6        | 2        | 0.977    | 5                       |

525





**LEGEND:**

-  Cretaceous basement
-  Vent breccia and synchronous Neogene sedimentary deposits
-  Dacite (Cetate type)
-  Breccia structures (mineralized)
-  Andesite pyroclastics
-  Andesite lava flows (Rotunda type)
-  Faults
-  Lake
-  Road

