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3 Color of Mn-bearing gahnite: A first example of electronic transitions in

4 heterovalent exchange coupled $^{IV}\text{Mn}^{2+}$ - $^{VI}\text{Mn}^{3+}$ pairs in minerals

5 ULF HÅLENIUS¹ AND FERDINANDO BOSI²

6

7 ¹Department of Geosciences, Swedish Museum of Natural History, Box 50007, SE-10405 Stockholm,

8

Sweden

9 ²Dipartimento di Scienze della Terra, Sapienza Università di Roma, Piazzale Aldo Moro, 5, I-00185

10

Rome, Italy

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ABSTRACT

13

14 A natural yellow and transparent crystal of Mn-bearing, Mg-rich gahnite (ZnAl_2O_4) from

15 Nordmark (Sweden) was studied by electron microprobe, single crystal X-ray diffraction and

16 optical absorption spectroscopy. The empirical structural formula of the crystal,

17 $^T(\text{Zn}_{0.52}\text{Mg}_{0.34}\text{Mn}^{2+}_{0.08}\text{Al}_{0.06})_{\Sigma 1.00}^M(\text{Al}_{1.88}\text{Mn}^{3+}_{0.06}\text{Fe}^{3+}_{0.01}\text{Mg}_{0.04}\text{Mn}^{2+}_{0.01})_{\Sigma 2.00}\text{O}_4$, shows that

18 Mn^{2+} and Mn^{3+} are almost completely ordered at the T and M sites, respectively. The

19 electronic absorption spectrum of the spinel shows, in addition to a strong UV-absorption

20 edge (O^{2-} - Mn^{3+} and O^{2-} - Fe^{3+} ligand-metal charge transfers) and two broad $^{VI}\text{Mn}^{3+}$ -bands, a set

21 of relatively narrow absorption bands at energies comparable to those caused by spin-

22 forbidden electron transitions in tetrahedrally coordinated Mn^{2+} in oxide spinel. However, the

23 set of bands in the present spectrum is shifted to lower energies and they are intensified by

24 approximately an order of magnitude compare to those recorded for spin-forbidden $^{IV}\text{Mn}^{2+}$ -

25 bands in spinel. These characteristic differences in combination with the determined cation

26 distribution in the present gahnite demonstrates that electronic transitions in exchange

27 coupled $^{IV}\text{Mn}^{2+}$ - $^{VI}\text{Mn}^{3+}$ pairs cause its color.

28

29 **Key words:** Gahnite, crystal structure, electronic spectra

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31

INTRODUCTION

32 Color is one of the fundamental physical properties that commonly is used for mineral
33 identification. Color in minerals has also an economic dimension. The value of gems and also
34 of several industrial minerals is strongly dependent on their colors. Consequently, knowledge
35 about the processes that create mineral colors is important. The majority of the common
36 ferromagnesian rock-forming minerals owe their characteristic colors and pleochroism to
37 electronic transitions in iron cations that are present in their structures. These electronic
38 processes occur at energies corresponding to those of visible light and they may either be
39 electron excitations that take place between energy levels within single ions of Fe^{2+} or Fe^{3+} or
40 more importantly, they may represent electron hopping between heterovalent neighboring
41 cations in the crystal, the so called Fe^{2+} - Fe^{3+} intervalence charge transfer processes (IVCT).

42

43 Numerous examples of Fe-related IVCT processes as the main cause for color and
44 pleochroism in the common silicates of the amphibole, pyroxene and mica mineral groups
45 have been observed (Burns 1993 and references therein). Color caused by electron transitions
46 in magnetically exchange coupled homonuclear cation pairs (ECP), trimers or larger clusters
47 have also been documented in minerals (e.g., Rossman 1975; Rossman 1976). Intensity
48 enhanced absorption bands caused by exchange coupled heterovalent Fe^{2+} - Fe^{3+} pairs have
49 been reported in spectra of tourmaline (Mattson and Rossman 1986) and spinel minerals
50 (Taran et al. 2005). However, there exist a limited number of documented examples of non-
51 iron IVCT or ECP processes that create color in minerals. Examples of heteronuclear

¹ E-mail: ulf.halenius@nrm.se

52 processes involving Fe are though known. The most common ones are $\text{Fe}^{2+}\text{-Ti}^{4+}$ IVCT
53 processes in silicates and oxide minerals (e.g., Mattson and Rossman 1988; Taran and Koch-
54 Müller 2011).

55

56 In view of the large number of examples of Fe-related IVCT and ECP coloring processes in
57 minerals it is surprising that the number of reported IVCT and ECP cases in Mn-bearing
58 minerals are so scarce. Considering the terrestrial abundance of manganese, the natural
59 occurrence of differently charged Mn-cations, cation radii that proxy those of iron cations,
60 and the number of common Mn-bearing minerals, one would expect a comparable number of
61 documented examples of $\text{Mn}^{2+}\text{-Mn}^{3+}$ or $\text{Mn}^{2+}\text{-Me}^{n+}$ IVCT or ECP causing mineral coloration.
62 Rossman and Mattson (1986) suggested that $\text{Mn}^{2+}\text{-Ti}^{4+}$ IVCT occurred in Mn-rich yellow
63 tourmaline. In addition, exchange coupled pairs of octahedrally coordinated Mn^{2+} and
64 tetrahedrally coordinated Fe^{3+} in trioctahedral micas have been proposed as the cause for the
65 color and the reverse pleochroism in some phlogopite samples (Smith et al. 1982).

66 Furthermore, Mn-related ECP processes have been proposed for a small number of minerals:
67 Mn-bearing diopside (Hålenius and Skogby 1996), genthelvite-helvite solid solution minerals
68 (Hålenius 2011), tephroite and rhodochrosite (Hålenius et al. 2007).

69

70 It is notable that in contrast to the above very restricted number of mineral examples, the
71 coloring power of electronic transitions in exchange coupled pairs of Mn cations has long
72 been known and documented for a large number of inorganic compounds (e.g., Ferguson et al.
73 1965b and 1966b; Lohr and McClure 1968). These cases demonstrate that Mn-related ECP-
74 transitions may occur even in compounds diluted in manganese and also that ECP-transitions
75 in homonuclear as well as heteronuclear Mn-related pairs are common (e.g., Ferguson et al.
76 1965a). Ferguson et al. (1966b) summarized theoretical considerations on the nature of the

77 ECP-transitions and the character of the selection rule relaxations that allow enhanced
78 transition probabilities. Reviews on optical absorption spectroscopy of exchange coupled
79 cation pairs, providing several examples of spectroscopic active homonuclear and
80 heteronuclear *3d* transition element pairs in chemical compounds, have been published by,
81 e.g., Lohr (1972) and McCarthy (1988). Articles within different fields of physics and
82 chemistry demonstrate the broad and continued interest in exchange interactions in transition
83 metal pairs. Published studies on Ni²⁺-Mn²⁺ (Ferguson et al. 1966a), Cu²⁺-Mn²⁺ (Ferre and
84 Regis 1978; Cador et al. 2000), Mn²⁺-Mn³⁺ (Franz et al. 2003; Cox et al. 2011; Xie et al.
85 2012), Mn³⁺-Mn³⁺ (Pelletier 1995) and Mn³⁺-Mn⁴⁺ pairs (Birkelbach et al. 1998) are some
86 examples that reflect this interest.

87

88 With the aim to explore the potential occurrence of electronic transitions in exchange coupled
89 pairs of Mn in spinel minerals, we have studied a natural Mg-rich gahnite (ZnAl₂O₄) that
90 contains an appreciable Mn-content by means of electron microprobe, single-crystal X-ray
91 diffraction and single-crystal optical absorption technique.

92

93

EXPERIMENTAL

94 **Sample**

95 The studied gahnite single crystal was selected from a calcite-rich layer with hausmannite,
96 manganosite and feitknechtite at the manganese ore at Kittelgruvan, Nordmark, Värmland,
97 Sweden. The small (up 0.5 mm) gahnite single crystals occur as transparent, octahedra
98 displaying a yellow color similar to those observed in synthetic Mn-bearing spinels (Bosi et
99 al. 2007; Hålenius et al. 2011). The same selected crystal was used in all the performed
100 experiments.

101

102 **Single-crystal structure refinement**

103 X-ray diffraction measurements were performed at Earth Sciences Department, Sapienza
104 University of Rome, with a Bruker KAPPA APEX-II single-crystal diffractometer, equipped
105 with CCD area detector ($6.2 \times 6.2 \text{ cm}^2$ active detection area, 512×512 pixels) and a graphite
106 crystal monochromator, using MoK α radiation from a fine-focus sealed X-ray tube. The
107 sample-to-detector distance was 4 cm. A total of 4268 exposures (step = 0.2° , time/step = 10
108 s) covering a full reciprocal sphere were collected. The orientation of the crystal lattice was
109 determined from more than 500 strong reflections ($I > 100 \sigma_I$) evenly distributed in the
110 reciprocal space, and used for subsequent integration of all recorded intensities. Final unit-cell
111 parameters were refined by means of the Bruker AXS SAINT program from ca. 1328
112 recorded reflections with $I > 10 \sigma_I$ in the range $9^\circ < 2\theta < 90^\circ$. The intensity data were
113 processed and corrected for Lorentz, polarization and background effects with APEX2
114 software program of Bruker AXS. The data were corrected for absorption using multi-scan
115 method (SADABS). The absorption correction led to a significant improvement in R_{int} . No
116 violation of $Fd\bar{3}m$ symmetry was noted. Sporadic appearance of forbidden space-group
117 reflections was recognized as double reflections.

118 Structural refinements were carried out with the SHELXL program (Sheldrick 2008). Setting
119 the origin at $\bar{3}m$, initial atomic positions for oxygen atoms were taken from the structure of
120 spinel (Bosi et al. 2010). Variable parameters were overall scale factor, extinction coefficient,
121 atomic coordinates, site scattering values expressed as mean atomic number (m.a.n.), and
122 atomic displacement factors. No chemical constraint was applied during the refinement. To
123 obtain the best values of statistical indexes ($R1$ and $wR2$) the oxygen site was modeled with
124 neutral *versus* fully ionized oxygen scattering curves, while neutral curves were used for the
125 cation sites. In detail, the T site was modeled considering the presence of Zn and Mg
126 scattering factors, whereas the M site was modeled by Al scattering factor. Three full-matrix

127 refinement cycles with isotropic displacement parameters for all atoms were followed by
128 anisotropic cycles until convergence was attained, that is, when the shifts in all refined
129 parameters were less than their estimated standard deviation. No correlation over 0.7 between
130 parameters was observed at the end of refinement. Table 1 summarizes structural parameters
131 and refinement details.

132

133 **Chemical analyses**

134 Electron microprobe analyses were done by wavelength-dispersive spectroscopy with a
135 Cameca SX50 instrument at the “Istituto di Geologia Ambientale e Geoingegneria (Rome,
136 Italy), CNR”, operating at an accelerating potential of 15 kV, a sample current of 15
137 nA and a spot size of 1 μm . Minerals and synthetic compounds were used as standards:
138 magnetite (Fe), rutile (Ti), corundum (Al), periclase (Mg), sphalerite (Zn) and metallic Mn.
139 The PAP matrix correction procedure (Pouchou and Pichoir 1991) was applied to reduce the
140 raw data. The results are summarized in Table 2. Titanium was below its detection limits
141 (0.03 wt%).

142

143 **Optical absorption spectroscopy**

144 Unpolarised, room-temperature optical absorption spectra in the spectral range 270-1100 nm
145 ($37\,037\text{-}9091\text{ cm}^{-1}$) were recorded on a 46 μm thick double-sided polished section of the
146 single crystal at a spectral resolution of 1 nm using a AVASPEC-ULS2048X16 spectrometer
147 attached via a 400 μm UV fiber cable to a Zeiss Axiotron UV-microscope. A 75 W Xenon arc
148 was used as a light source and Zeiss Ultrafluar 10 \times lenses served as objective and condenser.
149 The size of the circular measure aperture was 64 μm in diameter. The wavelength scale of the
150 spectrometer was calibrated against Ho_2O_3 doped and $\text{Pr}_2\text{O}_3/\text{Nd}_2\text{O}_3$ doped standards (Hellma
151 glass filters 666F1 and 666F7) with accuracy better than 15 cm^{-1} in the wavelength range 300-

152 1100 nm. Recorded spectra were fitted using the Jandel PeakFit 4.12 software assuming
153 Gaussian peak shapes.

154

155 RESULTS AND DISCUSSION

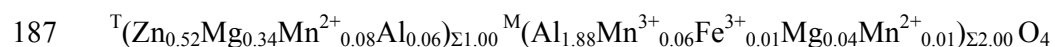
156 The results of the chemical analyses (Table 2), which represent mean values of 6 spot
157 analyses, were performed to obtain the average chemical compositions and estimates of
158 compositional homogeneity. The studied crystal shows some chemical variations in terms of
159 Mg, Mn and Zn. However, the Mn-content is substantial with a lowest concentration
160 corresponding to ~5.2 wt% MnO. In addition, the analyses demonstrate that the transition
161 element chemistry of the sample is quite simple comprising a major Mn component and very
162 low concentrations of Fe (less than 0.43 wt% FeO). The spinel formula, calculated on the
163 basis of 4 oxygen atoms, show that the present sample is dominated by Al (1.94 atoms per
164 formula unit, apfu) and Zn (0.52 apfu) along with Mg (0.38 apfu). Hence, the studied spinel
165 can be classified as Mg-rich gahnite (ZnAl_2O_4) containing significant amounts of both Mn^{2+}
166 and Mn^{3+} (0.10 apfu and 0.06 apfu, respectively) as confirmed by optical absorption spectra
167 (see below).

168

169 **Site populations**

170 The single-crystal-structure refinement results show that the m.a.n. at the tetrahedrally
171 coordinated T site and octahedrally coordinated M site are 22.6 and 13.5, respectively. These
172 values indicate that the T site is mainly occupied by Zn ($Z = 30$) and Mg ($Z = 12$) and the M
173 site is not only occupied by Al ($Z = 13$) but also by heavier cations such as Mn ($Z = 25$).
174 These site population is also supported by the values of T-O and M-O bond distances. In
175 detail, the observed T-O = 1.954 Å reflects a T site mainly occupied by Zn (expected value for
176 $^{\text{T}}\text{Zn-O}$ equals 1.950 Å, where the M site is dominated by Al, Bosi et al. 2011) along with

177 significant amounts of ^TMg [ideal bond distance 1.966(1) Å, Lavina et al. 2002]. The
178 observed M-O = 1.921 Å reflects a M site dominated by Al [ideal distance 1.908(2) Å, Lavina
179 et al. 2002] along with relatively minor amounts of larger cations such $^M\text{Mn}^{3+}$ [ideal distance
180 2.03(2) Å, Bosi et al. 2002; Lavina et al. 2002]. In order to translate the crystal chemical and
181 structural refinement results into quantitative site populations at T and M, experimental and
182 calculated data from the atom distribution were optimized. By using a least-squares program
183 and the ideal bond distances of Lavina et al. (2002), except for $^T\text{Zn-O}$ (Bosi et al. 2011), the
184 residuals between calculated and observed data (i.e., m.a.n. and bond distances) were
185 minimized. Final residuals were within experimental uncertainties. The final structural
186 formula for the present spinel is:



188 This formula is characterized by the almost complete dominance of divalent cations (i.e., Zn,
189 Mg, Mn^{2+}) at the T site and trivalent cations (i.e., Al, Mn^{3+}) at the M site. A small degree of
190 inversion (about 0.05), expressed as the fraction of divalent cations at M, also occurs.

191

192 **Optical absorption spectrum**

193 The recorded optical absorption spectrum (Fig. 1) shows a set of relatively narrow absorption
194 bands centered at 26780, 25150, 23150, 21780 and 20300 cm^{-1} . In addition, there are two
195 relatively broad absorption features centered at 23600 and 21200 cm^{-1} . All these bands are
196 superimposed on a strong UV-edge absorption that rapidly falls off towards the visible
197 spectral region and becomes insignificant in the green region of the visible spectrum.

198

199 At a first glance, the spectrum seems, with the exception of the strong UV-absorption edge to
200 be almost identical to the spectrum of synthetic end-member galaxite, MnAl_2O_4 , (Fig. 2)
201 recorded by Hålenius et al. (2007, 2011). The strong UV-absorption edge in the present

202 spectrum is caused by O^{2-} - Mn^{3+} and O^{2-} - Fe^{3+} ligand-metal charge transfers (LMCT), which
203 cannot occur in the spectra of the exclusively Mn^{2+} -bearing galaxite. Detailed inspection of
204 the present spectrum discerns two additional major differences. Firstly, the narrow absorption
205 bands occur at distinctly lower wavenumbers than the absorption bands in the spectrum of the
206 end-member galaxite (Fig. 2). The absorption bands recorded in the galaxite spectrum
207 (Hålenius et al. 2007; 2011), which are caused by spin-forbidden transitions in tetrahedrally
208 coordinated Mn^{2+} , occur at 27780, 25970, 23390 [${}^6A_1(S) \rightarrow {}^4E_g^4A_{1g}(G)$ transition], 22250,
209 20245 cm^{-1} (Hålenius et al. 2007, 2011). The corresponding absorption bands in the present
210 Mn-bearing gahnite occur at energies that are up to 1000 cm^{-1} lower. Secondly, based on the
211 concentration of tetrahedrally coordinated Mn^{2+} , the molar absorption coefficients (ϵ) of the
212 bands in the present gahnite spectrum are in the range 3-25 L/mol·cm. This is one order of
213 magnitude higher than the values of 0.1-1.9 L/mol·cm recorded for the absorption bands
214 caused by spin-forbidden single ion transitions in tetrahedrally coordinated Mn^{2+} in spectra of
215 the spinel (*sensu stricto*)-galaxite series (Hålenius et al. 2007, 2011).

216

217 The two broad bands observed in the present gahnite spectrum at 23600 and 21200 cm^{-1} occur
218 at the energies recorded for spin-allowed *d-d* transitions in octahedrally coordinated Mn^{3+} in
219 spectra of synthetic Mn^{3+} -bearing spinels (Bosi et al. 2007). The widths of these two bands
220 are also comparable to those recorded in spectra of the synthetic samples. Based on the molar
221 absorption coefficients of these two bands in spectra of the synthetic samples, the content of
222 octahedrally coordinated Mn^{3+} in the present gahnite is calculated to be about 0.04 apfu. This
223 value is consistent with the content of 0.056 (± 0.020) apfu Mn^{3+} obtained from the electron
224 microprobe analyses.

225

226 A third relatively strong and broad absorption band at 10800 cm^{-1} was recorded in the

227 spectrum of synthetic Mn^{3+} -bearing spinels and was assigned to a spin-allowed $d-d$ transition
228 in tetrahedrally coordinated Mn^{3+} (Bosi et al. 2007). This band is not observed in the present
229 spectrum of a Mn-bearing gahnite (Fig. 1 and 2). Consequently, it is concluded that Mn^{3+} is
230 ordered at the octahedrally coordinated site in the present sample.

231

232 The origin of the set of relatively narrow and intense absorption bands detailed earlier remains
233 to be discussed. We can initially conclude that the present spinel contains four different
234 transition element cations that may potentially give rise to narrow spectral bands through
235 spin-forbidden electronic single ion transitions: $^{\text{VI}}\text{Mn}^{2+}$, $^{\text{IV}}\text{Mn}^{2+}$ and $^{\text{VI}}\text{Mn}^{3+}$, $^{\text{VI}}\text{Fe}^{3+}$. The last
236 species can be excluded as the content of octahedrally coordinated ferric iron is simply too
237 low to account for the intensity of the observed bands. Furthermore, the energies of the most
238 intense spin-forbidden bands due to $^{\text{VI}}\text{Fe}^{3+}$ in spinel oxides are at 23000, 21600 [$^6\text{A}_1(\text{S}) \rightarrow$
239 $^4\text{E}_g$ $^4\text{A}_{1g}(\text{G})$ transition] and 15050 cm^{-1} (Taran et al. 2005). We do not observe these bands in
240 the present spectra. On the basis of band intensities and comparisons with spectra of synthetic
241 Mn^{3+} -bearing spinel (Bosi et al. 2007), we can exclude spin-forbidden transitions in $^{\text{VI}}\text{Mn}^{3+}$ as
242 the origin for these bands. We can also exclude single ion $^{\text{VI}}\text{Mn}^{2+}$ transitions on the basis of
243 the comparatively low energy of the narrow absorption bands in the present spectrum. For
244 instance, the characteristic, narrow and intense band due to the $^6\text{A}_1(\text{S}) \rightarrow ^4\text{E}_g$ $^4\text{A}_{1g}(\text{G})$
245 transition in octahedrally coordinated Mn^{2+} occurs at distinctly higher energies (24000-25000
246 cm^{-1}) in spectra of oxygen-based minerals (Burns 1993 and references therein, Hålenius et al.
247 2007). As mentioned previously, the set of narrow absorption bands in the present spectrum
248 shows a strong similarity with the absorption spectrum due to $^{\text{IV}}\text{Mn}^{2+}$ in spinels. However, the
249 bands are displaced towards lower energies their intensities are an order of magnitude higher.
250 We can therefore conclude that the recorded narrow absorption bands observed in the
251 spectrum of the present Mn-bearing gahnite are not caused by any spin-forbidden single ion

252 transitions.

253

254 In parallel with the assignment of intensity enhanced narrow absorption bands in spectra of
255 gahnite to electronic transitions in exchange coupled ${}^{\text{VI}}\text{Fe}^{3+}$ - ${}^{\text{IV}}\text{Fe}^{2+}$ pairs (Taran et al. 2005),
256 we must seek the origin of the observed narrow absorption bands (Fig. 1) in transitions taking
257 place in manganese cation pairs. The chemistry and cation distribution of the present gahnite
258 sample allow for a number of possible cation pair candidates: ${}^{\text{VI}}\text{Mn}^{2+}$ - ${}^{\text{VI}}\text{Mn}^{2+}$, ${}^{\text{IV}}\text{Mn}^{2+}$ -
259 ${}^{\text{VI}}\text{Mn}^{2+}$, ${}^{\text{IV}}\text{Mn}^{2+}$ - ${}^{\text{IV}}\text{Mn}^{2+}$, ${}^{\text{VI}}\text{Mn}^{2+}$ - ${}^{\text{VI}}\text{Mn}^{3+}$ and ${}^{\text{IV}}\text{Mn}^{2+}$ - ${}^{\text{VI}}\text{Mn}^{3+}$. Most of these potential cation
260 pairs can be excluded by comparisons with the recorded spectra of crystals belonging to the
261 spinel (*sensu stricto*)-galaxite solid solution (Hålenius et al. 2011). The narrow absorption
262 bands in the spectra of their synthetic samples showed constant molar absorption coefficients
263 (ϵ) with respect to $[{}^{\text{IV}}\text{Mn}^{2+}]$ -concentration and almost constant band energies within the entire
264 solid solution series although pair concentrations of their samples increased with increasing
265 total Mn-content (Hålenius et al. 2011). This suggests that the bands recorded in their spectra
266 of spinel-galaxite crystals were caused by spin-forbidden electronic transitions in single-ion
267 Mn^{2+} cations. No trivalent manganese was present in their synthetic samples that contained
268 appreciable contents of divalent manganese at the octahedral M-site (up to 0.16 apfu) and the
269 tetrahedral T-site (up to 0.84 apfu). Hence, their samples contained ${}^{\text{VI}}\text{Mn}^{2+}$ - ${}^{\text{VI}}\text{Mn}^{2+}$, ${}^{\text{IV}}\text{Mn}^{2+}$ -
270 ${}^{\text{VI}}\text{Mn}^{2+}$ and ${}^{\text{IV}}\text{Mn}^{2+}$ - ${}^{\text{IV}}\text{Mn}^{2+}$ cation pairs at concentration levels considerable higher than those
271 of the present gahnite sample (${}^{\text{IV}}\text{Mn}^{2+} = 0.01$ apfu and ${}^{\text{VI}}\text{Mn}^{2+} = 0.08$ apfu). Still, their spectra
272 did not show any intense and narrow absorption bands at the band energies recorded in the
273 present spectra. Consequently, the three pair configurations ${}^{\text{VI}}\text{Mn}^{2+}$ - ${}^{\text{VI}}\text{Mn}^{2+}$, ${}^{\text{IV}}\text{Mn}^{2+}$ -
274 ${}^{\text{VI}}\text{Mn}^{2+}$ and ${}^{\text{IV}}\text{Mn}^{2+}$ - ${}^{\text{IV}}\text{Mn}^{2+}$ may be excluded as causes for the bands observed in the present
275 spectrum. Of the two remaining pair configurations, ${}^{\text{IV}}\text{Mn}^{2+}$ - ${}^{\text{VI}}\text{Mn}^{3+}$ is the more likely one as
276 it necessitates smaller shifts in terms of energy and molar absorption coefficients with respect

277 to the corresponding formal single ion transitions in Mn^{2+} . For the proposed pair transition
278 related to the formal single ion ${}^6\text{A}_1(\text{S}) \rightarrow {}^4\text{E}_g {}^4\text{A}_{1g}(\text{G})$ transition in Mn^{2+} , the energy shift for
279 the corresponding ${}^{\text{IV}}\text{Mn}^{2+} - {}^{\text{VI}}\text{Mn}^{3+}$ pair transition is -240 cm^{-1} [23390 cm^{-1} in the spectrum of
280 end-member galaxite (Hålenius et al. 2011) compared to 23150 cm^{-1} in the present spectrum]
281 and its molar absorption coefficient increases ~ 9 times [$\epsilon \sim 1.90 \text{ L/mol}\cdot\text{cm}$ in end-member
282 galaxite (Hålenius et al. 2007) compared to $\epsilon \sim 17 \text{ L/mol}\cdot\text{cm}$ in the present sample]. As single
283 ion ${}^{\text{VI}}\text{Mn}^{2+}$ bands occur at distinctly higher energies and display molar absorption coefficients
284 approximately one order of magnitude smaller than single ion ${}^{\text{IV}}\text{Mn}^{2+}$ bands (Burns 1993 and
285 references therein), the resulting band energy shifts and band intensity enhancements would
286 be considerably larger for a ${}^{\text{VI}}\text{Mn}^{2+} - {}^{\text{VI}}\text{Mn}^{3+}$ pair transition assignment of the bands recorded in
287 the present spectra. Based on published energy data for spin-forbidden single ion ${}^{\text{VI}}\text{Mn}^{2+}$
288 bands (Burns 1993; Hålenius et al. 2007) the band energy of the present ${}^6\text{A}_1(\text{S}) \rightarrow {}^4\text{E}_g {}^4\text{A}_{1g}(\text{G})$
289 related transition is ~ 850 to $\sim 1850 \text{ cm}^{-1}$ lower. The energy shift of the ${}^6\text{A}_1(\text{S}) \rightarrow {}^4\text{E}_g {}^4\text{A}_{1g}(\text{G})$
290 related transition in $\text{Mn}^{2+} - \text{Me}^{n+}$ pairs is generally very small with respect to the energy of the
291 corresponding transition in the unperturbed Mn^{2+} cation, although shifts of up to -300 cm^{-1}
292 have been reported (Ferguson et al. 1966a). Consequently, we prefer the assignments of the
293 observed absorption bands to electron transitions in ${}^{\text{IV}}\text{Mn}^{2+} - {}^{\text{VI}}\text{Mn}^{3+}$ pairs.

294

295 **Implications**

296 The present study of a Mn-bearing gahnite demonstrates that electronic transitions in
297 exchange coupled pairs (ECP) of non-iron transition metal cations may have significant
298 effects on the physical properties of minerals. We have based our interpretation on arguments
299 related to energies and molar absorption coefficients (ϵ) of bands recorded in optical
300 absorption spectra. However, a large set of additional spectroscopies, ideally carried out at
301 different pressures and temperatures, and magnetic susceptibility measurements are available

302 for detection of exchange interactions in transition metal cation pairs. We believe that detailed
303 examinations of Fe-free or Fe-poor minerals rich in other transition element cations using
304 combinations of these methods will show that non-iron related IVCT and ECP transitions are
305 much more common as the cause for mineral color than so far realized. In particular, spectra
306 of close-packed mineral structures, as, e.g., oxide spinel minerals will most likely show a
307 large range of ECP and IVCT transitions, as, e.g., the recently recorded $\text{Cu}^+ \text{-Cu}^{2+}$ IVCT in
308 Cu-bearing spinels (Fregola et al. 2012).

309

310

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457 TABLES

458

459 TABLE 1. Selected X-ray diffraction data of the Mn-bearing gahnite from Nordmark (Sweden)

460 TABLE 2. Electron microprobe analyses of Mn-bearing gahnite from Nordmark (Sweden)

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466 FIGURE CAPTIONS

467

468 FIGURE 1. Optical absorption spectrum of Mn-bearing gahnite. The recorded spectrum is
469 shown by the black dotted line. Fitted UV-edge absorption, absorption bands and the total
470 fitted spectrum are shown by thin solid lines.

471

472 FIGURE 2. Optical absorption spectrum of Mn-bearing gahnite (black solid line) compared
473 with the spectrum of a synthetic end-member galaxite (dark grey solid line; from Hålenius et
474 al. 2007). The $^{IV}\text{Mn}^{2+}$ content in the galaxite sample is 0.84 apfu (Hålenius et al. 2011), but is
475 only 0.08 apfu in the present gahnite. Corresponding absorption bands in the two spectra are
476 connected by grey straight lines.

TABLE 1. Selected X-ray diffraction data for the Mn-bearing gahnite from Nordmark (Sweden)

Crystal size (mm)	0.20 × 0.20 × 0.22	Reciprocal range <i>hkl</i>	$-15 \leq h \leq 12$; $-10 \leq k \leq 14$; $-15 \leq l \leq 14$
<i>a</i> (Å)	8.1123(3)	Set of read reflections	1527
<i>u</i>	0.26404(4)	Unique reflections	120
T-O (Å)	1.9536(6)	Redundancy	11
M-O (Å)	1.9210(3)	Absorption correction method	SADABS
T-m.a.n.	22.57(8)	Refinement method	Full matrix least squares on F^2
M-m.a.n.	13.46(4)	EXTI	0.0024(2)
T- U^{11} (Å ²)	0.00452(7)	<i>R</i> int. (%)	0.51
M- U^{11} (Å ²)	0.00388(9)	<i>R</i> 1 (%) all reflections	0.90
M- U^{12} (Å ²)	-0.00035(6)	<i>wR</i> 2 (%)	1.79
O- U^{11} (Å ²)	0.00415(11)	Goof	1.123
O- U^{12} (Å ²)	-0.00030(9)	Diff. Peaks ($\pm e/\text{Å}^3$)	-0.20; 0.21

Notes: *a* = unit-cell parameter; *u* = oxygen fractional coordinate; T-O and M-O = tetrahedral and octahedral bond lengths, respectively; T- and M-m.a.n. = T- and M-mean atomic number; U^{11} = atomic displacement parameter; $U^{11} = U^{22} = U^{33}$ and $U^{12} = U^{13} = U^{23}$ (= 0 for T-site due to symmetry reasons); EXTI = extinction parameter; *R* int. = merging residual value; *R*1 = discrepancy index, calculated from *F*-data; *wR*2 = weighted discrepancy index, calculated from F^2 -data; Goof = goodness of fit; Diff. Peaks = maximum and minimum residual electron density. Radiation, Mo- $K\alpha$ = 0.71073 Å. Data collection temperature = 293 K. Total number of frames = 1539. Range for data collection $8^\circ < 2\theta < 86^\circ$. Origin fixed at $\bar{3}m$. Space group $Fd\bar{3}m$. *Z* = 8 formula units. Spinel structure has cations at Wyckoff positions 8a \equiv T (1/8, 1/8, 1/8) and 16d \equiv M (1/2, 1/2, 1/2), and oxygen anions at 32e (*u*, *u*, *u*).

TABLE 2. Electron microprobe analyses of Mn-bearing gahnite from Nordmark (Sweden)

	Mean of 6 spots	Standard deviation
Al ₂ O ₃ wt%	58.42	0.35
MgO	9.10	0.77
*MnO	6.37	1.41
*FeO	0.32	0.07
ZnO	25.15	2.17
Total	99.37	

Cations per formula unit based on four oxygen atoms

Al	1.937	0.032
Fe ³⁺	0.008	0.002
Mn ³⁺	0.056	0.020
Mg	0.382	0.030
Mn ²⁺	0.096	0.026
Zn	0.522	0.040
Total	3.000	
e ⁻ _{EMPA} (epfu)	49.42	
e ⁻ _{SREF}	49.49	

*Notes: Fe³⁺ and Mn³⁺ calculated from stoichiometry. Standard deviation on cations calculated by error-propagation theory. e⁻_{EMPA} and e⁻_{SREF} = number of electrons per formula unit (epfu) derived from electron microprobe and structural refinement (respectively).

Figure 1

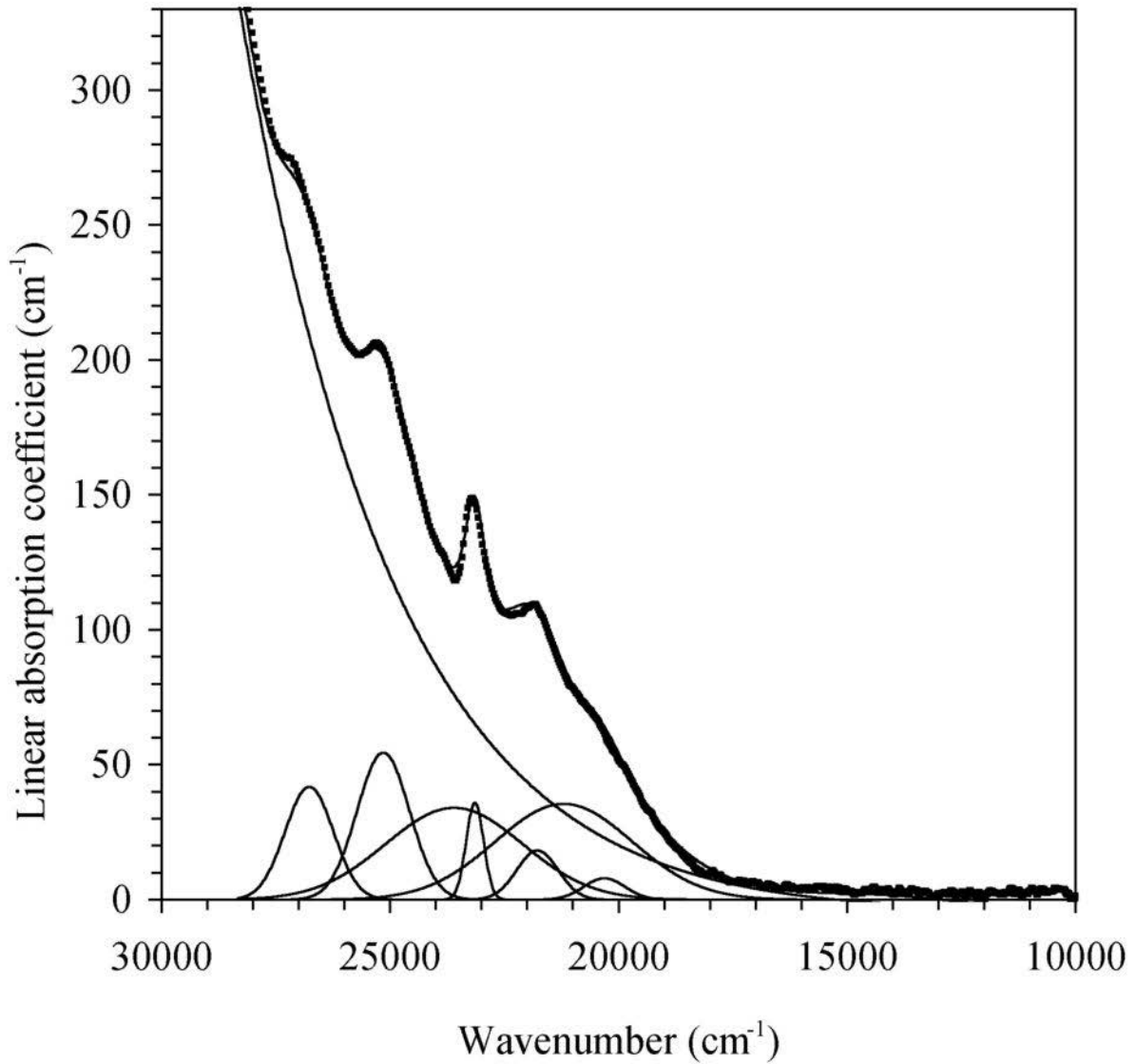


Figure 2

