

1 **Revision 1:** Pressure-induced Pbc₂-P2₁/c phase transition of natural orthoenstatite: Compositional effect
2 and its geophysical implications

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10 Abstract

11 Raman spectroscopy has been employed to investigate possible compositional effects on the high-
12 pressure phase transition of Mg-rich orthoenstatite to a newly-discovered P2₁/c phase. Three natural
13 orthoenstatite samples were used in this study: near end-member Mg orthoenstatite (Zabargad Island,
14 Egypt), Al-free Fe-bearing orthoenstatite (Morogoro, Tanzania), and Al-rich Fe,Ca-bearing orthoenstatite
15 (Kilbourne Hole, New Mexico). Experiments were carried out at room temperature. For all samples, the
16 high-pressure phase transition is characterized by a splitting of the 660-680 cm⁻¹ doublet in the Raman
17 spectrum into a triplet, with a corresponding change of peak intensities. These spectral changes are caused
18 by the lowered symmetry of the high-pressure phase, as indicated by structural refinement from single-
19 crystal X-ray diffraction results. The high-pressure phase of all samples appears to have space group
20 P2₁/c. No evidence for a C2/c phase was observed. Our results indicate that upon compression, 10mol%
21 Fe decreases the onset pressure of formation of the high-pressure P2₁/c phase by about 1 GPa. Results for
22 the Kilbourne Hole OEN show that a combined enrichment of Al and Ca contents increases the onset
23 pressure of formation of HPCEN2 (upon compression) by over 3 GPa relative to Tanzania OEN. Upon

24 decompression, all samples revert to single-crystals of the orthoenstatite starting phase. Our
25 measurements suggest that orthoenstatite is the prevalent phase of Mg-rich pyroxene throughout the
26 uppermost mantle, whereas the newly discovered $P2_1/c$ phase might be present near the bottom of
27 uppermost mantle, slightly shallower than the top of the transition zone.
28 Keywords: orthoenstatite, high-pressure clinoenstatite, high-pressure phase transition, upper mantle,
29 Raman spectroscopy

30 Introduction

31 Mg-rich Fe-bearing pyroxene with approximate composition $(\text{Mg,Fe})\text{SiO}_3$, is one of the major
32 minerals in Earth's uppermost mantle. Four polymorphs of $(\text{Mg,Fe})\text{SiO}_3$ are potentially stable under
33 upper mantle conditions: orthoenstatite (OEN) with space group $Pbca$ (Morimoto and Koto 1969), low-
34 pressure clinoenstatite (LPCEN) with space group $P2_1/c$ (Morimoto et al. 1960), high-pressure
35 clinoenstatite (HPCEN) with space group $C2/c$ (Angel et al. 1992), and a newly discovered high pressure
36 monoclinic polymorph (HPCEN2) also with space group $P2_1/c$ (Zhang et al. 2012). Although the
37 equilibrium stability fields for these four polymorphs have not yet been firmly established, the discovery
38 of the new $P2_1/c$ structure has potentially important implications for the phase relations in the
39 $(\text{Mg,Fe})\text{SiO}_3$ system and for upper mantle mineralogy.

40 Experimentally determined pressures for the transition from OEN to the high-pressure polymorph
41 HPCEN2 span a wide range for different experiments using different samples with different techniques.
42 For example, Raman spectroscopy on synthetic flux-grown MgSiO_3 (containing minor amounts of Li,
43 Mo, V), compressed in a diamond anvil cell (DAC), indicate a transition pressure of between 6.1 GPa and
44 12.0 GPa in an argon pressure-transmitting medium (Chopelas 1999), and between 9.5 GPa and 10 GPa in
45 water (Lin 2003). Multi-anvil ultrasonic experiments constrained the transition pressure to be between 9.6
46 and 11.8 GPa (Kung et al. 2004). For natural orthoenstatite from San Carlos, containing 8 mol% Fe and
47 2.5 wt% Al, the transition does not occur until 14.26 GPa in a neon pressure medium, as determined from
48 single-crystal X-ray diffraction experiments (Zhang et al. 2012). For orthoenstatite from Kilbourne Hole,
49 containing 9 mol% Fe 5 wt% Al, there is no evidence for this phase transition up to 12.5 GPa by stimulated
50 light scattering in a DAC (Chai et al., 1997). These discrepancies on the transition pressure suggest that
51 chemical composition (both natural cation substitutions such as Al, Fe, Ca, as well as the incorporation of

52 artificial flux components such as Li, Mo, V), as well as the pressure medium used, could both have a
53 significant effect on the transition pressure and possibly the phase relations between OEN, HPCEN
54 LPCEN and HPCEN2. Thus, it is important to separate the effects of chemistry from effects of the
55 pressure medium in order to investigate the effect of composition on the stability of the $P2_1/c$ phase.
56 Moreover, orthoenstatite in Earth's mantle does not exist as $MgSiO_3$, but will almost certainly contain Fe,
57 Al, and Ca as minor elements, with the exact composition dependent on depth. In general, the Fe content
58 of OEN tends to be relatively constant with increasing depth, whereas Ca and Al content will likely
59 decrease as enstatite is progressively dissolved into Ca-rich clinopyroxene and garnet (Akaogi and
60 Akimoto 1977; Irifune and Ringwood 1987; Brey et al. 2008). Thus it is important to evaluate the phase
61 transition pressures of OEN for a variety of potential mantle compositions. We thus chose to study two
62 natural Fe-bearing samples representative of possible upper mantle compositions, and a composition near
63 the $MgSiO_3$ end-member for the sake of comparison with prior work and calibration.

64 In this paper we report the results of in-situ high-pressure room-temperature Raman experiments
65 with several chemically distinct samples loaded together in a sample chamber of a diamond-anvil cell
66 (DAC) to clarify the effects of compositional variations in natural orthoenstatite on the newly discovered
67 $Pbca \rightarrow P2_1/c$ phase transition.

68

69 Sample Description and Experimental Methods

70 Natural orthoenstatite samples from three different locations and with different chemical
71 compositions were used in this study: 1) near end-member $MgSiO_3$ orthoenstatite from Zabargad Island,
72 Egypt; 2) Al-free Fe-bearing orthoenstatite from the Morogoro Region, Tanzania; 3) Al-rich Fe,Ca-

73 bearing orthoenstatite from Kilbourne Hole, New Mexico. Samples of high optical quality were polished
74 into plate-like samples of ~20 μm thickness using Al_2O_3 abrasive film (down to 0.3 μm grain size). All
75 samples were scratch-free under optical examination, and were cut into pieces approximately 20 – 60 μm
76 wide for compositional analysis and DAC loading.

77 The chemical composition of each sample was analyzed using a JXA 8100 electron microprobe at
78 Nanjing University. All elemental analyses were performed using an accelerating potential of 15 kV, a
79 beam current of 20 nA, and a beam 1 μm in diameter. Element peaks and backgrounds were measured
80 with counting times of 20 and 10 s, respectively. A ZAF routine correction was used in the data reduction.
81 Hornblende was used as the standard for all the measured elements, yielding results as follows: Zabargad
82 Island (Zabg) $(\text{Mg}_{0.994}\text{Fe}_{0.002}\text{Al}_{0.004})_2(\text{Si}_{0.996}\text{Al}_{0.004})_2\text{O}_6$, or $\text{En}_{99}\text{Fs}_0\text{Di}_0\text{MgTs}_1$; Tanzania (Tan)
83 $(\text{Mg}_{0.897}\text{Fe}_{0.097}\text{Ca}_{0.004}\text{Al}_{0.002})_2(\text{Si}_{0.996}\text{Al}_{0.002})_2\text{O}_6$, or $\text{En}_{89}\text{Fs}_{10}\text{Di}_1\text{MgTs}_0$; and Kilbourne Hole (KBH)
84 $(\text{Mg}_{0.835}\text{Fe}_{0.090}\text{Ca}_{0.020}\text{Al}_{0.055})_2(\text{Si}_{0.945}\text{Al}_{0.055})_2\text{O}_6$, or $\text{En}_{76}\text{Fs}_9\text{Di}_4\text{MgTs}_{11}$, where En indicates the $\text{Mg}_2\text{Si}_2\text{O}_6$
85 component, Fs is $\text{Fe}_2\text{Si}_2\text{O}_6$, Di is $\text{CaMgSi}_2\text{O}_6$, and MgTs is $(\text{MgAl})(\text{SiAl})\text{O}_6$.

86 High-pressure experiments were performed with a membrane-type diamond anvil cell for precise
87 pressure control. Rhenium metal gaskets with an initial thickness of 250 μm were pre-indented to 0.070
88 mm using 500 μm culet, ultra-low fluorescence diamond anvils. A 250 μm diameter hole in the gasket
89 formed the sample chamber. Three different plate-like samples (maximum ~60 μm width) and several
90 ruby balls were loaded together in the diamond-anvil cell sample chamber (Figure 1). An alcohol mixture
91 (methanol:ethanol=4:1) was loaded as a pressure-transmitting medium. A heat gun was used to warm the
92 chamber (up to 60°C) to help relax differential stresses when needed. Pressure was determined from ruby
93 fluorescence (Mao et al. 1986) before and after the Raman experiment at each pressure; the maximum

94 pressure difference from the two readings was 0.14 GPa or less. Maximum stress differentials were
95 obtained during decompression runs.

96 In-situ high-pressure single-crystal Raman spectroscopy experiments were carried out at Ecole
97 Normale Supérieure de Lyon. Raman spectra were obtained using a multichannel Raman microprobe
98 (LabRam HR800 from DILOR) equipped with a confocal microscope configuration that enhances the
99 signal-to-noise ratio by eliminating most of the parasitic light from sample and diamond fluorescence.
100 Experiments were conducted in a backscattering geometry with a Mitutoyo objective that focused the
101 incident laser spot to less than 2 μm in diameter. The scattered Raman light is focused through a 100-
102 micron slit into a spectrograph equipped with a 1800 gr/mm grating and analyzed by a CCD detector,
103 giving a resolution of approximately 2.5 cm^{-1} (Auzende et al. 2004). The accumulation times for Raman
104 spectra were typically 60-120s over the spectral region from 100 cm^{-1} to 1250 cm^{-1} . Peak positions were
105 determined to within 2 cm^{-1} or better depending on sample quality. The 514.5 line of an argon ion laser
106 was used as an excitation source at an output power ranging from 500 to 1500 mW. Only 5 to 10% of this
107 power reaches the sample, due to absorption or reflection from the optical elements and diamonds in the
108 optical path.

109

110 Results

111 High pressure Raman data were collected at 47 different pressures between 0.30(1) GPa and
112 18.01(2) GPa in several compression and decompression cycles. The number of Raman-active modes in
113 orthoenstatite is predicted to be 120 (Ferraro 1975). However, the number of observed Raman bands is
114 smaller in any given run due to the weak intensity and/or overlap of many bands, and the strong
115 orientation dependence of intensity. Because the edge filter used to cut off Rayleigh scattering limits the

116 measured Raman shifts to above 100 cm^{-1} , we were not able to observe the lowest frequency bands of
117 OEN. Another limitation included the overlap of the strong methanol ethanol C-O stretching band at
118 $\sim 1030\text{ cm}^{-1}$ with OEN and HPCEN2 peaks near that frequency shift.

119 We consistently observed 34 bands in Zabg OEN, 40 bands in Zabg HPCEN2; 28 bands in Tan
120 OEN, 26 bands in Tan HPCEN2; 23 bands in KBH OEN, 24 bands in KBH HPCEN2. The results are
121 listed in Tables 1 and 2. Our observations are in good agreement with previous results (Lin 2003,
122 Chopelas 1999; Huang et al. 2000; Stalder et al. 2009; Reynard et al. 2008; Zucker and Shim 2009).
123 Small discrepancies were due to crystal orientation effects, overlap of peaks, and weak intensities of some
124 peaks.

125 Typical Raman spectra of OEN and HPCEN2 for all three samples are shown in Figure 2a. In the
126 low frequency range, the transition is characterized by the occurrence of HPCEN2 peaks (e.g. for Zabg
127 OEN at 12.5GPa, 199.5 cm^{-1} , 228.3 cm^{-1} , 341.9 cm^{-1}), with fading OEN peaks (206.6 cm^{-1} , 238.5 cm^{-1} ,
128 210.5 cm^{-1} , 218.1 cm^{-1} , 308.2 cm^{-1} , 312.7 cm^{-1}). The single most significant characteristic of this
129 transition is splitting of the strong $660\text{-}680\text{ cm}^{-1}$ doublet into a triplet (Figure 3), accompanied by a change
130 in the relative intensities of the peaks. The $660\text{-}680\text{ cm}^{-1}$ doublet is assigned to the Si-O-Si bending mode
131 of the tetrahedral chains. A change in the number of peaks in that region indicates a change in the number
132 of symmetrically distinct tetrahedral chains, which is diagnostic of the phase transition (Chopelas and
133 Boehler 1992; Ross and Reynard, 1999). The observation of an increase from two to three peaks is
134 consistent with X-ray structural refinements of the HPCEN2 phase, which indicate that the number of
135 distinct tetrahedral (T) chains doubles across the OEN-HPCEN2 transition (Figure 4). In detail, the A site
136 T-chain in the OEN structure splits into two T-chains with the O3-O3-O3 angle changing from
137 $160.71(16)^\circ$ to $148.44(79)^\circ$ and $216.61(77)^\circ$. The B site T-chain in OEN undergoes a change in O3-O3-

138 O3 angle from $135.44(15)^\circ$ to $131.94(72)^\circ$ and $133.13(49)^\circ$ (Zhang et al. 2012). The large change of angle
139 in the A chain is related to the split of the Raman doublet, whereas the change for the B chain is so small
140 that the associated Raman peak splitting cannot be resolved, thus yielding a spectral triplet instead of a
141 quadruplet. For similar reasons, the doublet of OEN transforms to a single peak in the same frequency
142 region upon transition to HPCEN with space group C2/c (Chopelas and Boehler 1992; Ross and Reynard,
143 1999). These clear spectral differences makes Raman spectroscopy an efficient tool to differentiate OEN
144 polymorphs from each other (Lin 2004).

145

146 Discussion and conclusion

147 For all three OEN samples, peak splitting associated with the $Pbca \rightarrow P2_1/c$ phase transition was
148 observed. Evidence for a C2/c phase, which was claimed to be the stable high-pressure phase above ~ 7 -9
149 GPa (Angel et al. 1992), was not observed over the pressure range of the present experiments. We were
150 able to observe the formation of domains as the transition proceeded (e.g. Figure 2b and c). There are two
151 distinctive characteristics of these domains. Firstly, they developed upon initiation of the transition, and
152 disappeared after the transition was completed (e.g. Figure 2b). All samples were single crystals by
153 optical examination before initiation and after completion of the high-pressure transition. Secondly, the
154 formation of $P2_1/c$ phase domains started from edges of the sample and progressed toward the center (e.g.
155 Zabg OEN Figure 2c). Edges may have a high density of crystal defects that act as nucleation sites for the
156 new phase.

157 The onset pressure of the $Pbca \rightarrow P2_1/c$ transition (P_{Tr}) is different for the three samples. For near
158 Mg-end member Zabg OEN, P_{Tr} is between 13.13(11) - 13.41(8) GPa on compression, and between

159 11.03(1) - 11.13(11) GPa on decompression. For near-Al-free Fe-bearing Tan OEN, P_{Tr} was constrained
160 between 12.38(3) - 12.42(4) and 11.03(1) - 11.13(11) GPa in compression and decompression cycles,
161 respectively. For Al-rich Ca, Fe-bearing KBH OEN, P_{Tr} was constrained as between 15.73(12) - 16.25(14)
162 and 12.18(1) - 12.58(1) GPa during compression and decompression, respectively. A comparison between
163 Zabg OEN and Tan OEN indicates that 10mol% Fe content decreased the onset pressure of the initial
164 formation of HPCEN2 by ~1GPa; however, the comparison between KBH and Tan OEN is more complex,
165 including differences in both Al and Ca contents. KBH OEN contains 0.22 Al per formula unit (pfu,
166 based on 6 oxygens), which is 27.5 times more than in Tan OEN. The Ca content of KBH OEN is 0.040
167 pfu, which is 5 times more than in Tan OEN. Considering the fact that the Fe contents are very close to
168 each other, the presence of 0.212 pfu more Al and 0.032 pfu more Ca increased the onset pressure of the
169 initial formation of HPCEN2 by over 3 GPa. Notably, it seems that the transition pressure during
170 decompression is less sensitive to composition – ~10mol% Fe does not change decompressional transition
171 pressure while 0.212 pfu more Al and 0.032 pfu more Ca increased it only by about 1 GPa.

172 Increased repulsion between the M2 site cation and A-site Si was believed to be a key factor in
173 determining the topology of pyroxene structures at higher pressure. Fe prefers M2 sites over the M1 sites
174 as suggested from the structure refinement for both OEN and HPCEN2 (e.g., Zhang et al. 2012). The
175 substitution of Fe atoms into the structure results in a larger M2 site, thus giving the Si-O tetrahedral
176 chains more spatial flexibility, and making this $Pbca \rightarrow P2_1/c$ transition, which is characterized by
177 tetrahedral rotation, easier to occur at lower pressure. Similar qualitative arguments for the effects of Ca
178 and Al are less obvious, because the effects of these two elements cannot be decoupled in this experiment.
179 We do not rule out the possibility that a small amount of Ca might have a significant effect on the
180 transition pressure, as it has a resolvable effect on the OEN structure (Tribaudino and Nestola 2003).

181 Additional experiments would be needed to determine pure Ca or Al effects using synthetic samples.
182 However, the much greater abundance of Al over Ca in KBH OEN suggests to us that the enrichment of
183 Al is the primary cause of the increase in the onset of P_{Tr} . The case of Al substitution into the crystal
184 structure is complicated since it involves a coupled substitution mechanism, with half the Al atoms
185 occupying the tetrahedral A site substituting for Si, and half occupying the M1 sites substituting for Mg,
186 resulting in a Mg-Tschermaks component in the KBH OEN. In this case, two competing processes, due to
187 Al occupying two polyhedral sites, will be operative. Although the individual effects of these two
188 substitutions are not resolved by the present experiments, net effect is to increase the phase transition
189 pressure.

190 In the upper mantle, the Fe content in OEN does not change much with depth, whereas Ca and Al
191 contents decrease with increasing depth (Akaogi and Akimoto 1977; Irifune and Ringwood 1987; Brey et
192 al. 2008). Al and Ca will be enriched in Ca-rich clinopyroxene and garnet in comparison to OEN. Thus, in
193 the 300-450km depth range where this phase transition might occur, the composition of OEN is probably
194 between San Carlos OEN (Zhang et al. 2012) and Tan OEN, but likely closer to Tan OEN. Neglecting
195 thermal effects, our results suggest a transition depth of about 350-400km (if the OEN has not yet
196 completely dissolved into the garnet structure), just above the 410km discontinuity. Additional studies,
197 including high temperature measurements of the stability of HPCEN2, and single-crystal elasticity
198 measurements, will be required to obtain better estimates of the depth at which the OEN-HPCEN2
199 transition occurs in the mantle, and its seismic signature.

200

201

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271 Figure Captions

272 Figure 1 – Diamond anvil cell sample chamber

273 Figure 2 –

274 a. Selected Raman spectra of OEN and HPCEN2 for all three samples at high pressures. The 660-680
275 cm^{-1} doublet-to-triplet changes for each of the samples are indicated with red arrows.

276 b. Optical observation of the OEN→HPCEN2 phase transition in the mixed phase region for sample Tan.

277 c. Optical observation of the OEN→HPCEN2 phase transition in the mixed phase region for sample Zabg.

278 Figure 3 – Raman frequency shifts of the 660-680 cm^{-1} doublet/triplet as a function of pressure

279 a. Zabargad Island, Egypt

280 b. Morogoro Region, Tanzania

281 c. Kilbourne Hole, New Mexico

282 Figure 4 –Structural change associated with the OEN→HPCEN2 phase transition. O3-O3-O3 angle is
283 marked as θ in the figure. The largest change of θ occurs on the A site tetrahedral which is shown on the
284 top right.

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Table 1: Comparison of Raman frequencies from previous studies

OEN, ambient condition							HPCEN2, normalized to 12.5GPa				
Stalder et al.	Huang et al.	Lin	Chopelas	This study			Lin	Chopelas	This study		
2009	2000	2003	1999	ZABG	TAN	KBH	2003	1999	ZABG	TAN	KBH
		83	83				94.7	94.9			
		106.8					135.9	139.1			
		115.4	115				166	<i>168.8</i>	168.3		
133.8		134.3	134	137.4			174.6	178.1	<i>179.7</i>		
		155.3	153		155.7	<i>151.4</i>	183				
		<i>161.5</i>				160	196.4	197.5	199.2	195.7	
		166.7	166	170.2	169.3	169.7	204				
				190.1	190.9		224.4	225.5	227.9	223.7	223.8
197.5		196.9	197	196.5	197.1	193.7	231.8				
				201.1			250.5	251.7	254.4	251.7	251.3
206.4		<i>206.5</i>		208.3			262.1		267.5		
			237	240.1			272	273.9	276.2		<i>269</i>
239	238	238.2	239	241.1	238.2	238.4			280.8		
		244.8	245	246.3	248.9		290.2	291.6	295.3	<i>288.7</i>	<i>289.7</i>
			261	260.8	<i>257.9</i>		304.1				<i>298.3</i>
				271.6			320.3	<i>314.5</i>	323.6	318.9	320
		278.2			276.6				329.4		
				286.5		284	337.2	338	341.3	338.4	335.7
302.5	298	301.5	302	303.2	300.7	<i>300.5</i>			<i>345.6</i>		
		<i>323.3</i>				320.1			<i>358.1</i>		
		<i>327.7</i>			328.9		373.4		<i>377.7</i>	368.5	
				<i>340.3</i>			381.5	<i>383.4</i>	<i>388.7</i>	377.6	<i>382.9</i>
				<i>342.9</i>			393.6	395.1	399.6	395.3	394.5
343.9	344	343.5	343	345.4	339.9	345.3	406.3	407	412.1	404.4	<i>404.7</i>

				374			428.9	431.2	430.2	427.8	
385.2		383.8		381	381.7	387		434.9	436.3		
402.8	407	402.1	402	402.6	399	407.6	445.8	447.3	452.9	444.4	443.6
422.7	422	421.7	422	424.2	418.5		462.7	466.3	469.3	<i>463.7</i>	
445.9	447	445.5	446	<i>445.4</i>	443.2	444.3	476.1	476.5	480.2	475.6	475.3
				448			493.8	496.8	499.4	<i>493.9</i>	488.5
		<i>458.8</i>	457	458.6	<i>459.9</i>		511.7			<i>507.6</i>	
		<i>472.9</i>		<i>478.7</i>	<i>476.4</i>		533.6		528.9	<i>525.4</i>	
			487	<i>488.3</i>			557.5	559.8	562.9	557.1	557.8
527	522	524.9	519	528.1	521	523.5	565.6		570.9	<i>569.8</i>	
						<i>533.7</i>	587.8	<i>578.7</i>	583.3		<i>581</i>
540.8	543		540	541.9	544.7	<i>538</i>			<i>600.9</i>		<i>598.5</i>
553.5	551	551.5	553	554.5		550.2			<i>607.2</i>		
581	582	580.3	580	582.1	578.5	574.9			620.4	616.1	<i>616</i>
		<i>594.5</i>			<i>593.9</i>		704.1	704.1	708.4	706.9	701.3
664.8	665	663.8	665	664.9	662.1	661.8	716.5	716.2	719.6	716.5	716
687.1	685	686.1	687	687.7	684.2	684.2	725.3				
						<i>736.1</i>	735.2	736.1	739.8	735.9	735.3
	751	750.7		<i>752.7</i>	<i>753.3</i>	753.9			<i>780.1</i>		
853.2	856	851.1		851	852.9	848.7			<i>814</i>		
			886				847	851.9	856.1		855.8
927.9	931	926.6	927	929.6	929.7	929.3	873.5		864.5	<i>861.7</i>	<i>865.4</i>
936.5		935.2	937		938.6	<i>940.9</i>	887.6	888.9	893.1	893.8	<i>893.2</i>
1013.2	1013	1011.3	1014				924.1	924.7			
1034.7	1035	1032.9	1034				963.6	955.3	969.5	<i>969.4</i>	<i>970</i>
							1030.2				
							1056.5	1055.5			
							1075	1073.3			
							1092.8	1093.3			

Note: **strong** peaks are in **bold** font;

weak peaks are in *inclined grey* font;

Intermediate peaks are in normal font.

Table 2. Pressure dependence of the Raman frequencies for all three samples

OEN				$\nu = \nu_0 + a \times P + b \times P \times 10^{-2}$				HPCEN				$\nu = \nu_0 + a \times P$											
ZABG	near pure MgSiO ₃			TAN	Fe-bearing Al-free OEN			KBH	Fe,Al-bearing OEN			ZABG	near pure MgSiO ₃			TAN	Fe-bearing Al-free			KBH	Fe,Al-bearing		
ν_0	a	b	R^2	ν_0	a	b	R^2	ν_0	a	b	R^2	ν_0	a	R^2	ν_0	a	R^2	ν_0	a	R^2	ν_0	a	R^2
137.4(10)	5.2(5)	-38.8(41)	0.954	155.7(5)	0.8(2)	4.0(12)	0.989	151.4(10)	1.3(3)	-0.3(17)	0.969	157.8(7)	0.8(1)	0.926	181.5(15)	1.1(1)	0.818	202.6(22)	1.7(2)	0.919			
170.2(3)	2.1(1)	-5.4(7)	0.993	169.3(6)	1.6(2)	-1.6	0.985	160.0(9)	0.3(2)	3.5(11)	0.925	169.4(7)	0.8(1)	0.919	211.0(7)	1.0(1)	0.938	235.0(11)	1.3(1)	0.963			
190.1(9)	2.1(2)	-6.0(11)	0.981	197.1(5)	0.4(2)	0.9(11)	0.909	169.7(5)	2.0(1)	-5.3(5)	0.98	191.6(7)	0.6(1)	0.828	237.9(8)	1.1(1)	0.932	248.4(45)	1.6(3)	0.805			
196.5(10)	1.6(3)	-4.0(16)	0.951	190.9(36)	6.0(8)	-21.2(42)	0.972	193.7(7)	1.4(2)	-1.9(9)	0.958	215.3(9)	1.0(1)	0.895	268.3(31)	1.6(2)	0.724	257.1(41)	2.6(3)	0.902			
201.1(19)	3.8(7)	-19.8(50)	0.847	238.2(5)	1.7(2)	-3.7(11)	0.986	238.4(4)	1.7(1)	-3.8(5)	0.986	241.8(7)	1.0(1)	0.945	300.4(10)	1.5(1)	0.938	191.2(219)	8.6(15)	0.938			
208.3(9)	3.0(2)	-4.5(14)	0.992	248.9(13)	0.0(4)	10.2(28)	0.936	284.0(7)	2.0(1)	-2.9(7)	0.985	249.9(19)	1.4(1)	0.805	315.4(16)	1.8(1)	0.905	298.5(25)	1.7(2)	0.909			
240.1(6)	1.6(2)	-2.7(12)	0.976	257.9	3.4(2)	-3.6(17)	0.997	300.5(34)	1.0(8)	8.0(50)	0.869	256.7(13)	1.6(1)	0.915	338.1(20)	2.4(1)	0.923	309.5(24)	2.1(2)	0.936			
241.1(8)	2.9(2)	-7.3(15)	0.989	276.6(9)	3.5(2)	-10.4(14)	0.984	320.1(9)	4.5(2)	-5.7(10)	0.993	255.5(19)	2.0(1)	0.902	343.9(16)	2.7(1)	0.951	360.8(66)	1.8(5)	0.725			
246.3(10)	3.9(3)	-10.1(17)	0.989	300.7(6)	2.0(2)	-2.7(15)	0.99	345.3(6)	6.2(1)	-10.7(7)	0.998	273.4(17)	1.8(1)	0.891	371.2(11)	1.9(1)	0.957	370.6(13)	1.9(1)	0.977			
260.8(7)	3.9(2)	-6.7(13)	0.995	328.9(11)	2.8(3)	2.5(18)	0.994	387.0(68)	1.4(33)	14.4(226)	0.922	305.3(10)	1.5(1)	0.935	364.8(17)	3.2(1)	0.96	365.9(98)	3.1(6)	0.849			
271.6(61)	5.3(13)	-19.2(67)	0.853	339.9(12)	6.8(4)	-12.6(27)	0.995	407.6(6)	3.9(1)	-7.8(6)	0.995	310.2(16)	1.5(1)	0.876	396.6(26)	2.5(2)	0.868	397.7(50)	3.7(3)	0.913			
286.5(7)	2.8(2)	-5.5(11)	0.993	381.7(8)	3.7(3)	-3.0(19)	0.994	444.3(11)	2.5(2)	-0.4(13)	0.976	319.8(13)	1.7(1)	0.93	398.6(29)	3.7(2)	0.923	450.7(11)	2.0(1)	0.983			
303.2(5)	2.6(1)	-4.6(9)	0.994	399.0(10)	5.4(3)	-14.4(24)	0.992	523.5(7)	2.7(2)	-0.4(9)	0.991	312.8(56)	2.6(4)	0.788	433.4(23)	2.4(2)	0.948	438.7(52)	4.0(4)	0.933			
342.9(119)	3.6(21)	-7.2(93)	0.947	418.5(18)	2.3(7)	3.5(53)	0.972	538.0(9)	1.1(5)	8.7(29)	0.999	339.8(35)	1.5(2)	0.793	451.9(11)	1.9(1)	0.958	516.1(23)	3.3(2)	0.977			
340.3(81)	6.4(15)	-23.6(69)	0.893	443.2(7)	2.8(2)	-3.5(18)	0.991	550.2(8)	2.7(2)	-0.7(10)	0.989	342.6(22)	2.8(2)	0.957	457.6(18)	2.9(1)	0.956	551.7(37)	2.3(2)	0.949			
345.4(11)	6.9(3)	-12.2(22)	0.996	459.9(16)	2.0(7)	1.1(52)	0.961	574.9(11)	2.5(2)	-2.8(14)	0.971	351.6(30)	3.0(2)	0.936	473.8(42)	2.7(3)	0.913	571.1(45)	2.2(3)	0.855			
374.0(24)	6.1(6)	-13.7(39)	0.979	476.4(14)	5.5(4)	-11.4(26)	0.988	661.8(4)	3.8(1)	-3.3(6)	0.998	371.9(19)	2.2(1)	0.905	478.5(43)	3.8(3)	0.927	598.5(27)	1.4(2)	0.847			
381.0(32)	6.2(10)	-13.1(64)	0.947	521.0(6)	3.1(2)	-1.0(14)	0.996	684.2(5)	3.9(1)	-5.5(6)	0.997	374.3(36)	3.0(3)	0.856	524.3(22)	2.6(2)	0.916	658.1(57)	3.5(4)	0.926			
402.6(9)	5.3(3)	-12.8(17)	0.994	533.7(28)	2.2(10)	4.5(71)	0.925	736.1(42)	3.4(13)	-12.2(72)	0.802	401.2(17)	2.3(1)	0.929	545.8(47)	1.9(3)	0.625	674.1(26)	3.3(2)	0.969			
424.2(43)	2.9(15)	-1.7(106)	0.8	544.7(20)	4.0(6)	-6.8(42)	0.962	753.9(13)	2.6(3)	3.4(16)	0.985	407.6(27)	2.3(2)	0.843	592.1(17)	1.9(1)	0.908	690.6(22)	3.6(1)	0.981			
445.4(68)	1.1(14)	8.0(68)	0.978	578.5(4)	2.0(1)	-0.8(9)	0.995	848.7(9)	1.9(2)	5.7(11)	0.99	423.9(46)	2.3(3)	0.644	679.0(25)	2.2(2)	0.875	810.2(47)	3.7(3)	0.922			
448.0(9)	2.4(3)	0.4(19)	0.986	593.9(9)	3.5(4)	-8.4(29)	0.991	929.3(8)	4.0(2)	-4.3(10)	0.993	441.7(20)	2.2(1)	0.903	680.6(15)	2.9(1)	0.964	841.1(23)	1.9(2)	0.982			
458.6(17)	3.1(5)	-1.0(34)	0.965	662.1(9)	4.0(3)	-3.3(21)	0.994	940.9(38)	5.7(10)	-6.3(64)	0.974	456.8(15)	1.9(1)	0.916	697.8(16)	3.1(1)	0.962	844.7(24)	3.9(2)	0.988			
478.7(22)	2.3(5)	1.8(29)	0.983	684.2(5)	3.8(2)	-4.1(12)	0.997					463.9(22)	2.8(2)	0.923	840.8(29)	1.7(2)	0.706	947.7(21)	1.8(1)	0.935			
488.3(12)	3.4(4)	1.3(24)	0.99	753.3(20)	2.6(8)	4.8(72)	0.967					489.7(35)	3.1(3)	0.869	850.4(54)	3.5(4)	0.911						

528.1(11)	2.1(3)	3.5(23)	0.982	852.9(16)	1.1(5)	9.1(37)	0.963
541.9(5)	1.1(2)	10.1(13)	0.997	929.7(19)	3.5(5)	-0.2(29)	0.993
554.5(18)	2.3(5)	2.9(36)	0.96	938.6(22)	5.7(5)	-5.8(32)	0.995
582.1(6)	2.0(2)	-0.3(11)	0.992				
664.9(8)	4.0(3)	-3.8(17)	0.994				
687.7(7)	4.0(2)	-5.6(13)	0.996				
752.7(12)	3.2(4)	1.7(25)	0.989				
851.0(11)	1.2(3)	8.4(22)	0.979				
929.6(8)	3.7(2)	-2.1(15)	0.995				

Note: **strong** peaks are in **bold** format;

weak peaks are in *inclined grey* format;

Intermediate peaks are in normal format.

528.5(28)	2.8(2)	0.878	<i>937.0(47)</i>	<i>2.6(3)</i>	<i>0.844</i>
534.4(23)	2.9(2)	0.92			
550(23)	2.7(2)	0.912			
<i>582.5(33)</i>	<i>1.5(2)</i>	<i>0.791</i>			
<i>574.6(36)</i>	<i>2.6(3)</i>	<i>0.896</i>			
595.1(18)	2.0(1)	0.901			
673.4(23)	2.8(2)	0.913			
685.3(15)	2.7(1)	0.963			
701.9(17)	3.0(1)	0.957			
<i>738.8(24)</i>	<i>3.3(2)</i>	<i>0.962</i>			
<i>782.9(36)</i>	<i>2.5(3)</i>	<i>0.86</i>			
824.9(13)	2.5(1)	0.963			
846.4(23)	1.4(2)	0.764			
853.8(22)	3.1(2)	0.937			
939.6(27)	2.4(2)	0.858			

P=12.42GPa

ruby

Zabargad Island
MgSiO₃ OEN

ruby

Tanzania

(Mg_{0.90}Fe_{0.10})SiO₃

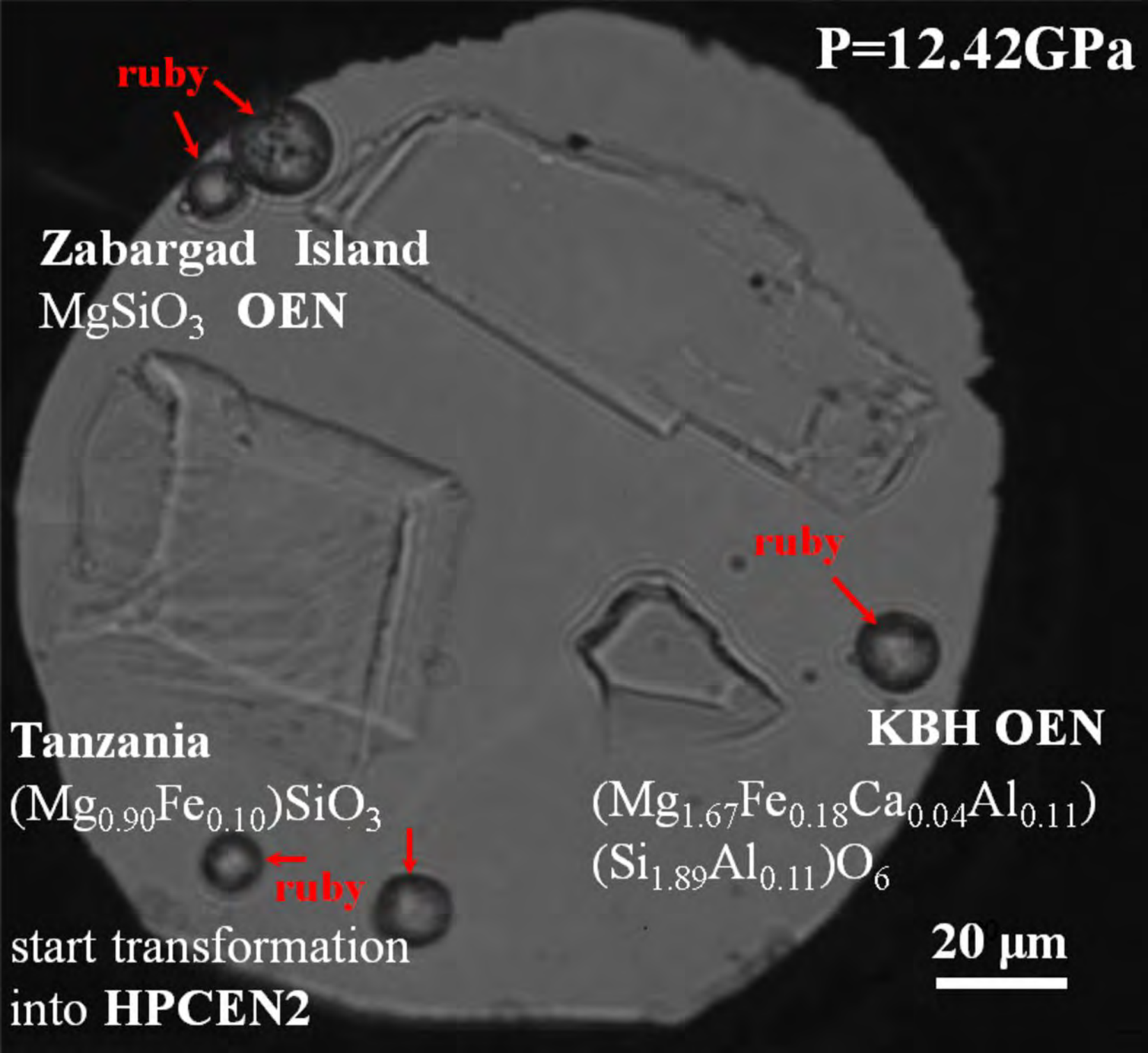
KBH OEN

(Mg_{1.67}Fe_{0.18}Ca_{0.04}Al_{0.11})
(Si_{1.89}Al_{0.11})O₆

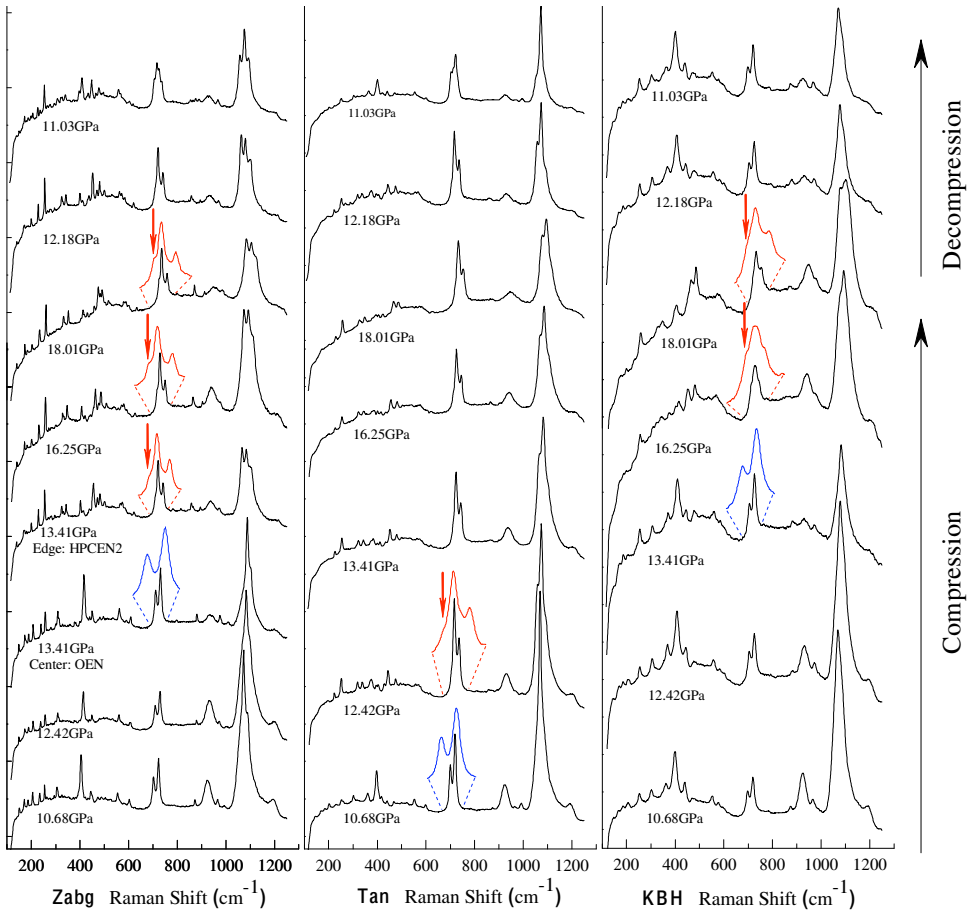
ruby

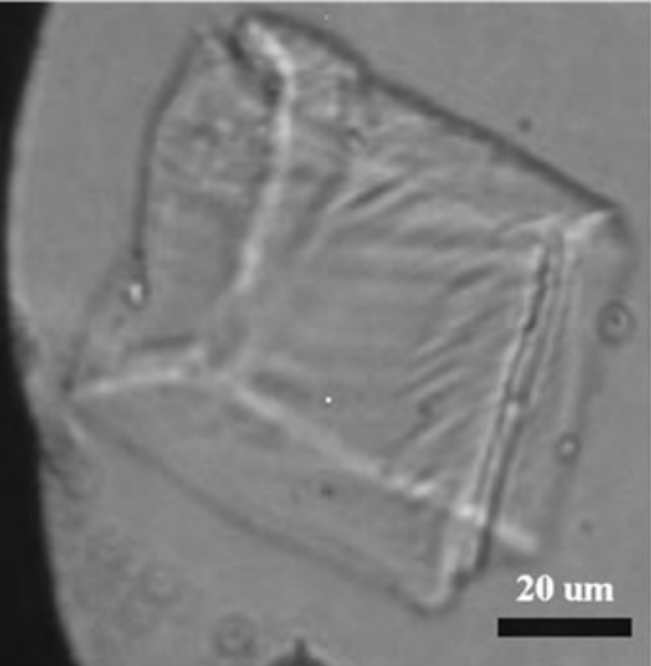
start transformation
into **HPCEN2**

20 μm

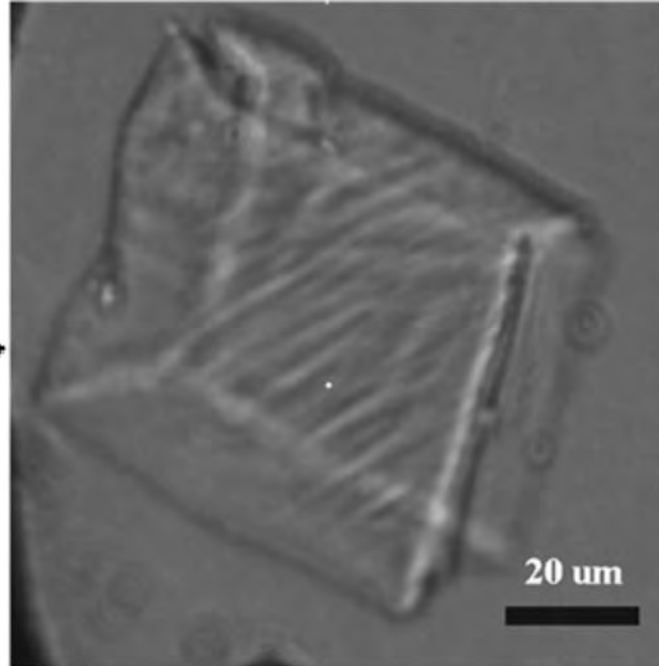


Intensity

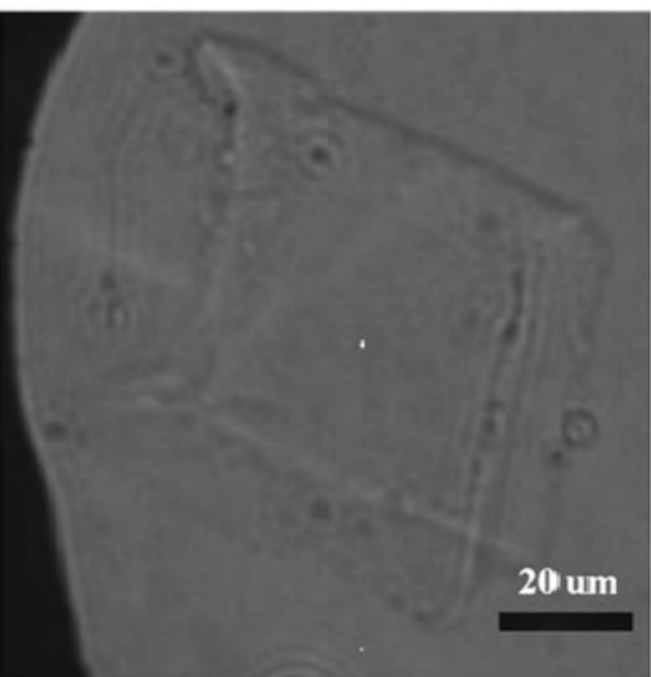




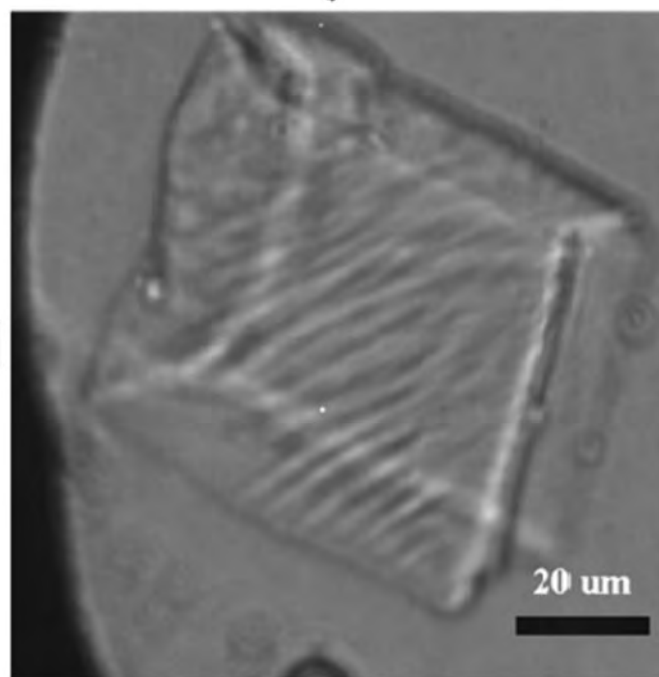
13.13 GPa



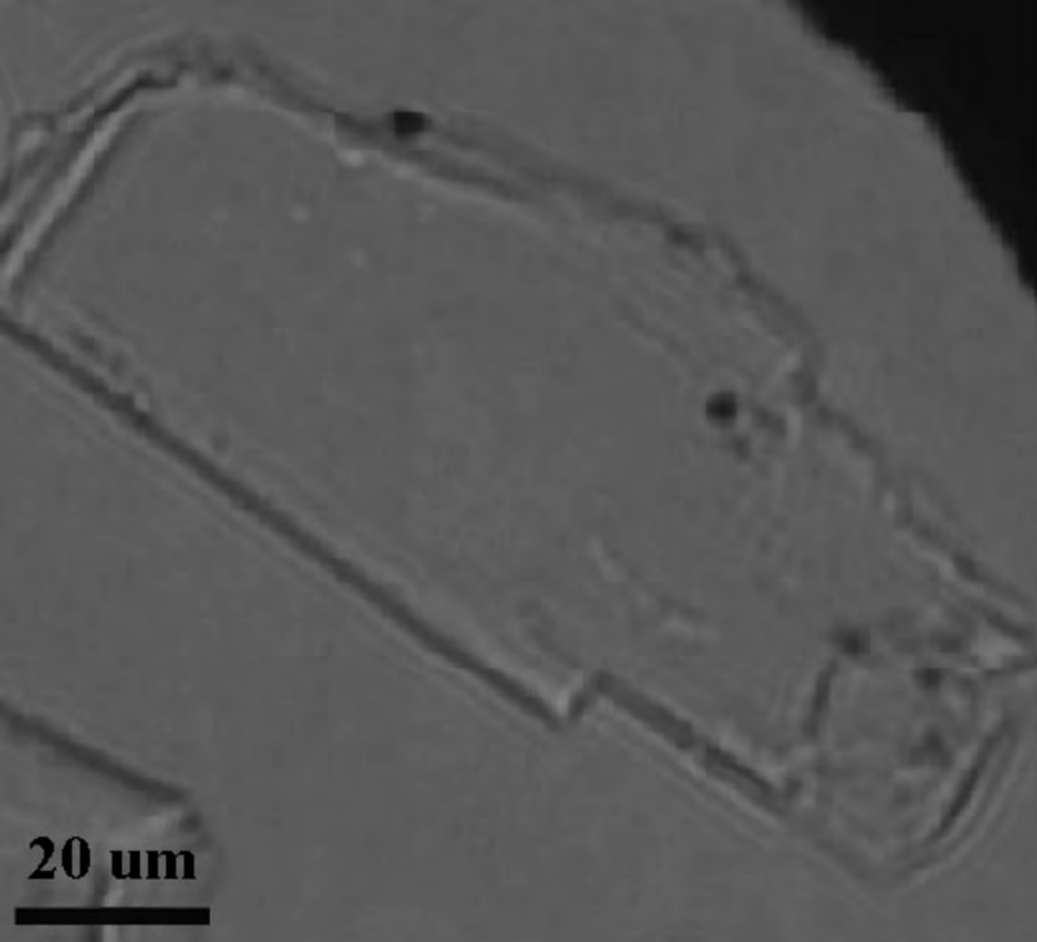
13.41 GPa



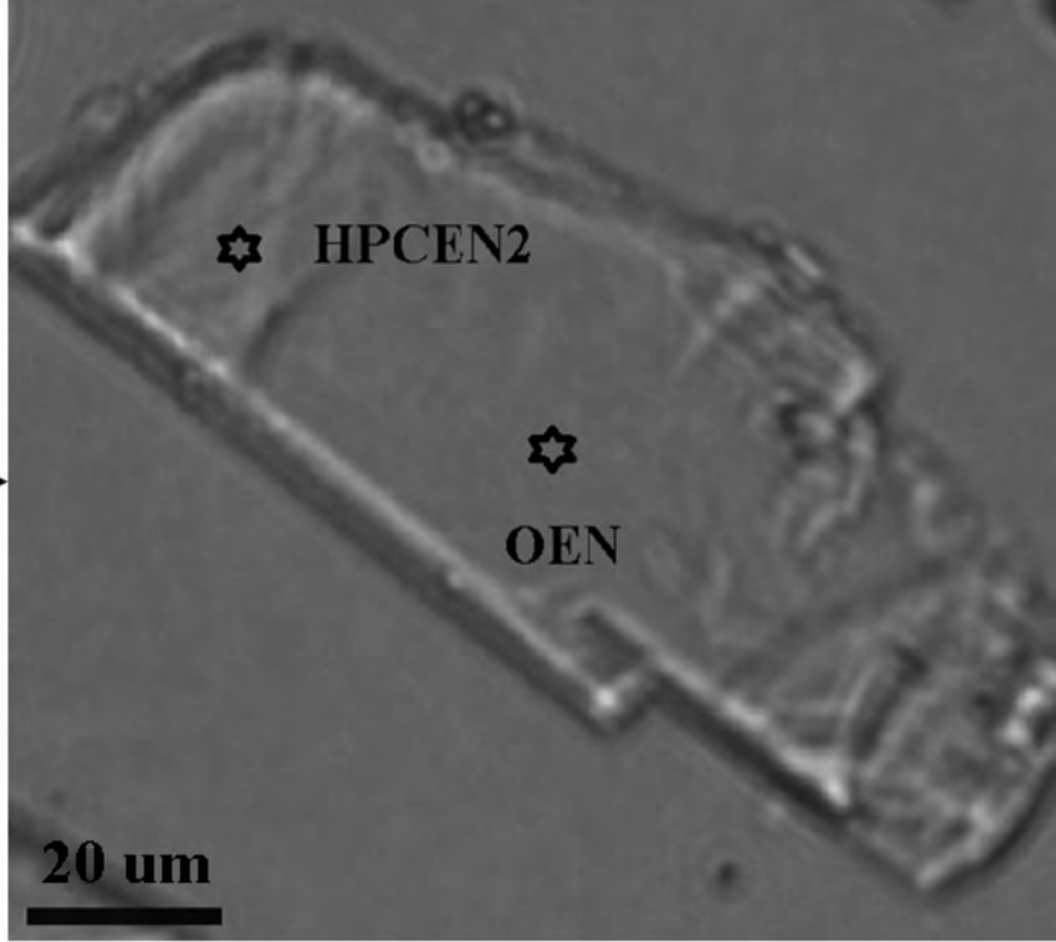
14.66 GPa



13.87 GPa



12.42GPa



13.87GPa

