

1 **Revision 2**

2 **A micro-reflectance IR spectroscopy method for analyzing volatile species**
3 **in basaltic, andesitic, phonolitic and rhyolitic glasses**

4 Penelope L. King^{1, 2, 3} and Jessica F. Larsen⁴

5 1. Research School of Earth Sciences, Australian National University, Canberra ACT 0200
6 Australia.

7 2. Institute for Meteoritics, University of New Mexico, Albuquerque NM 87131 USA.

8 2. Dept. Earth Sciences, University Western Ontario, London ON N6A 5B7 Canada.

9 3. Geophysical Institute, University of Alaska, Fairbanks, AK 99775 USA.

10 * Present address.

11 **ABSTRACT**

12 Volatile contents of geologic glasses are used to model magma chamber and degassing
13 processes, thus, there is considerable interest in small-scale analytical techniques for analyzing
14 volatiles in glasses. Infrared (IR) spectroscopy has the advantage of determining volatile
15 speciation in glasses (e.g., OH⁻, molecular H₂O, molecular CO₂ and CO₃²⁻). However, sample
16 preparation for the most common IR method used, micro-transmission IR spectroscopy, is
17 complicated because glasses must be prepared as thin parallel sided-wafers. Raman analysis,
18 while valuable for Fe-poor samples, can be difficult to use for Fe-rich glasses.

19 We have calibrated a micro-reflectance infrared method for determining volatile species
20 using calculated Kramers-Kronig absorbance (KK-Abs.) spectra that requires that only one side
21 of a glass be polished. The method is easier to use than other reflectance methods where it is
22 difficult to determine the baseline for the IR bands. Total H₂O wt.% = $m * (3600 \text{ cm}^{-1} \text{ KK-Abs.})$,
23 where m , is the slope of the calibration line that is obtained from a fit to the data. The m value is

24 related to the calculated refractive index, n , for a range of aluminosilicate glass compositions
25 allowing the technique to be applied to samples with unknown calibration slopes. For calc-
26 alkaline andesite glasses we determined calibration slopes for micro-reflectance IR
27 measurements of molecular H₂O, molecular CO₂, and CO₃²⁻. The method has been calibrated for
28 glasses with up to 6.76 wt.% total H₂O (but is useful for glasses with more than 20 wt.% total
29 H₂O) and has been calibrated for glasses with up to 0.575 wt.% total CO₂.

30 This technique provides a means to analyze volatile abundances in samples that are not
31 possible to analyze or prepare for analysis with transmission micro-IR techniques. We have
32 determined volatile contents in fragile samples such as cracked, vesicular or crystal-bearing
33 glasses formed by volcanic or impact processes or in high pressure bubble nucleation
34 experiments and H diffusion experiments. We have monitored H uptake during weathering of
35 basaltic glasses that cannot be polished and determined volatiles in melt inclusions and pumice.

36

37 *Key words:* IR spectroscopy, glass properties, new technique, volatiles, H₂O, CO₂, CO₃²⁻

38

INTRODUCTION

39 Volatiles, such as H₂O and CO₂, are found in natural glasses and their abundances are
40 used to model the role of volatiles on magmatic processes and degassing (Wallace and Anderson
41 Jr., 2000; Larsen and Gardner, 2004; Métrich and Wallace, 2008); to determine volatile diffusion
42 rates in melts (e.g., Liu et al., 2004); and to measure volatile contents of impact glasses (Harris et
43 al., 2007). Thus, there is considerable interest in small-scale analytical techniques for analyzing
44 glasses. A variety of techniques are used, such as vacuum-line manometry, secondary ion mass
45 spectrometry, Karl-Fischer titration (for H₂O), Raman spectroscopy, and Fourier transform
46 infrared (FTIR) spectroscopy (see reviews in Ihinger et al., 1994; Deloule et al., 1995; Devine et

47 al., 1995; King et al., 2002; Behrens et al., 2006; Thomas et al., 2006; Aubaud et al., 2007). Of
48 these techniques, only Raman and FTIR spectroscopy provide information on volatile species
49 (e.g., OH⁻, molecular H₂O (H₂O_{mol}), molecular CO₂ (CO_{2 mol}) and CO₃²⁻) in small sample areas
50 (< 50 μm across). The abundances of these species are also used in volcanic 'geospeedometers'
51 (e.g., Zhang et al., 2000) and it is therefore advantageous to measure them accurately.

52 Micro-Raman spectroscopy has been developed to determine the total H₂O (H₂O_{tot}),
53 H₂O_{mol} and OH⁻ in silicic, Fe-poor glasses (Thomas, 2000; Chabiron et al., 2004; Behrens et al.,
54 2006; Thomas et al., 2006) and a few Fe-bearing glasses (Chabiron et al., 2004; Zajacz et al.,
55 2005; Behrens et al., 2006; di Muro et al., 2006; Thomas et al., 2008a; Mercier et al., 2009).
56 Raman methods use either internal calibrations (e.g., Behrens et al., 2006) or a comparator
57 method where samples are referenced to one calibration glass (Thomas et al., 2008b). However,
58 there are difficulties in applying the micro-Raman method to Fe-bearing glasses due to band
59 overlap in the 100–1100 cm⁻¹ region and absorption of the laser radiation with depth in Fe-
60 bearing glasses that decreases the OH band intensity (Behrens et al., 2006; Thomas et al., 2008a).
61 In addition, glasses with low melting temperatures, including water-rich samples, may lose
62 volatiles or melt due to strong light absorption and/or fluorescence radiation.

63 FTIR spectroscopy has been used to successfully analyze Fe-rich compositions like
64 basaltic glasses (e.g., Dixon et al., 1988; Dixon and Pan, 1995). However, the most common
65 method used, micro-transmission IR spectroscopy, is complicated because glasses are generally
66 prepared as thin, parallel-sided, stand-alone wafers (e.g., Paterson, 1982; Stolper, 1982a, b;
67 Newman et al., 1986; Dixon et al., 1988; Ihinger et al., 1994; Dixon and Pan, 1995).
68 Transmission IR techniques using ground glass powders pressed into a KBr disc have the
69 advantage that painstaking thickness measurements on doubly polished glass wafers are not

70 required; however KBr is hygroscopic and therefore it is challenging to ensure that water is not
71 adsorbed (e.g., Izawa et al., 2010). Furthermore, the KBr method is not appropriate for samples
72 in which spatially constrained, fine-scale spectra are needed such as melt inclusions and some
73 experimental glasses. A technique has been developed to investigate melt inclusions in doubly
74 polished olivine crystals (Nichols and Wysoczanski, 2007); however that technique is not
75 amenable to studying very small melt inclusions and some volatile species are not possible to
76 measure (e.g., $\text{H}_2\text{O}_{\text{mol}}$) due to band interferences from the host mineral. An alternative technique
77 is micro-reflectance IR analysis (e.g., Grzechnik et al., 1996; Moore et al., 2000; Hervig et al.,
78 2003), where only one sample face is polished.

79 The major advantage of the micro-reflectance method is that the sample can be prepared
80 as a single polished surface. Therefore, small samples, such as melt inclusions, do not need to be
81 removed from their matrix for mounting and fragile pumices and experimental charges may be
82 analyzed in textural context. Also, different depths within a single sample can be analyzed
83 sequentially by successively polishing the sample to deeper levels.

84 Previous work has shown that reflectance spectra for C-O-H volatiles have band positions
85 similar to that found in transmission IR spectra, but they have different shapes due to the
86 mathematical nature of the superposition of the volatile bands on the Si-O band. Example micro-
87 reflectance IR spectra are shown in Figure 1a and 1b.

88 For O-H species in glasses, the largest band occurs as a "negative band" in the 3500 cm^{-1}
89 region (Hervig et al., 2003). This band, referred to here as a "total H_2O " band, arises from a
90 combination of symmetric stretching of the H-O-H molecule ($\nu_{\text{ss}}\text{H-O-H}$), plus symmetric
91 stretching of Si-OH ($\nu_{\text{ss}}\text{Si-OH}$) in the aluminosilicate network, plus an overtone of the H-O-H
92 bending vibrations ($2\delta\text{H-O-H}$). The asymmetric stretching vibration (ν_{as}) of CO_2_{mol} gives rise to

93 a "sigmoidal" feature in the reflectivity spectrum (Moore et al., 2000). The dissolved CO_3^{2-}
94 concentrations (vC-O) in basanite and leucitite glasses have a "doublet" band in micro-
95 reflectance IR spectra that resembles the absorption bands obtained in transmission IR studies
96 (Grzechnik et al., 1996). In this study, we show that a "negative band" for molecular H_2O related
97 to H-O-H bending vibrations (δHOH) occurs in the $\sim 1620 \text{ cm}^{-1}$ region, in agreement with the
98 band observed in transmission IR spectra.

99 In micro-reflectance IR measurements, $\text{H}_2\text{O}_{\text{tot}}$ and total CO_2 ($\text{CO}_{2 \text{ tot}}$) concentrations have
100 been related directly to the height of the reflectance band for the species of interest (ΔR_i) relative
101 to the reflectance at a reference wavenumber position (R_{ref}) via an empirical equation of the
102 form:

$$103 \quad (\Delta R_i / R_{\text{ref}}) = k_1 * (\text{volatile species } \%) + k_2 \quad (1)$$

104 where the positions of the ΔR_i and R_{ref} and the constants k_1 and k_2 depend on the glass
105 composition (see King et al., 2004).

106 Despite, the ease of micro-reflectance IR measurements, the extraction of band heights
107 from raw R% spectra may be somewhat challenging (e.g., Fig. 1a, 1b). There are significant
108 uncertainties associated with the baseline position, especially for samples with low volatile
109 contents and for sigmoidal bands like $\text{CO}_{2 \text{ mol}}$ and $\text{H}_2\text{O}_{\text{mol}}$ (Fig. 1a, 1b). Also, assumptions are
110 needed regarding the location of R_{ref} for different composition glasses. Furthermore, calibrations
111 are not available for $\text{H}_2\text{O}_{\text{tot}}$ or $\text{H}_2\text{O}_{\text{mol}}$ in basalt, or C-O species or $\text{H}_2\text{O}_{\text{mol}}$ in andesite.

112 In this paper, we present a method for analyzing glasses using micro-reflectance IR
113 spectroscopy that has been manipulated using a Kramers-Kronig (KK) transform. The technique
114 is advantageous because it is straightforward to prepare samples and easier to determine the
115 baseline position for IR bands than using the raw R% spectra. We provide new calibration curves

116 for basalt, calc-alkaline andesite, alkali andesite, phonolite and rhyolite. We also examine some
117 of the limitations of this technique.

118 **METHODS**

119 Volatile-bearing glasses were obtained (Perfit et al., 1983, 1999; Westrich, 1987; Moore
120 and Carmichael, 1998; King et al., 2002; Sinton et al., 2003; Larsen and Gardner, 2004; Lui,
121 2005; Mongrain et al., 2008) and their compositions were confirmed using the electron
122 microprobe if necessary (Table 1). The basaltic glasses were analyzed using transmission IR
123 spectroscopy and the H₂O_{tot} concentrations determined with the Beer-Lambert law (e.g., Stolper,
124 1982b):

$$125 \text{ molar concentration of H}_2\text{O}_{\text{tot}} = A / (d * \rho * \epsilon_{\text{H}_2\text{O}_{\text{tot}}}) \quad (2)$$

126 where, A is the linear absorbance (i.e. band height), d is the sample thickness which was
127 measured with a Mituyomo micrometer to $\pm 2 \mu\text{m}$ (<2% relative error for most samples); ρ is the
128 sample density that was calculated following methods in Silver (1988) and data from Church and
129 Johnson (1980) and Mandarino (1976); and $\epsilon_{\text{H}_2\text{O}_{\text{tot}}}$ is the extinction coefficient for H₂O_{tot} in
130 basalt (from Dixon et al., 1995). Calc-alkaline andesite glasses contained both C-O species as
131 well as H-O species. Table 1 shows that variations in volatile-free composition were negligible
132 for most glasses. In the case of the basalt and rhyolite 1 σ deviations are given because slightly
133 different compositions constitute the sample set. Errors in the volatile contents were taken from
134 the literature or for the basalts and phonolite conservatively estimated using curve fitting
135 uncertainties (10 to 14 % relative) (Table 1).

136 Micro-reflectance IR analyses of the glasses were collected using a Nicolet Nexus 670
137 spectrometer with a Global source, extended range KBr beamsplitter (XT-KBr), and a

138 Continuum microscope with a MCT-A detector (cooled with liquid nitrogen). The same
139 instrument was used in King's lab at both the Experimental Analysis Laboratory at the
140 University of Western Ontario and the New Mexico IR Analysis Laboratory at the University of
141 New Mexico. Analyses were made in triplicate if possible, but some glass samples were too
142 small for multiple analyses.

143 As with any IR analysis of C-O-H species, it is necessary to minimize the background
144 levels of atmospheric gases surrounding the sample. All analyses in this work were performed by
145 running a dry air purge into a volume around the sample that was $\sim 100 \text{ cm}^3$ (e.g., Fig. 47 in
146 King et al. 2004). We also monitored the total reflectance signal by examining the interferogram
147 peak-to-peak measurement before each analysis. We prefer to operate at conditions where the
148 peak-to-peak measurement is greater than 6.0 volts and we did not operate at conditions less than
149 4.0 volts.

150 Each spectrum was collected in a spectral range of $5300 - 650 \text{ cm}^{-1}$, over 300 scans, with
151 4 cm^{-1} resolution, using a $100 \times 100 \mu\text{m}$ sampling area. For each analysis, a background
152 spectrum was collected on a gold-coated glass slide because it has a reflection coefficient of
153 $\sim 100\%$ over the wavelength region measured, and so the absolute value of the sample reflectivity
154 is known without further calibration. The raw spectra were collected in units of percent
155 reflectance (%R). For our purposes, 300 scans appeared to provide sufficiently high signal-to-
156 noise ratio (SNR); however, if a small aperture IR beam is used (e.g., for small melt inclusions),
157 it is necessary to maximize the SNR by increasing the scan times to larger values (>1000 scans,
158 Grzechnik et al. 1996).

159 To convert the R% spectra to absorbance spectra we used the KK transform (McMillan
160 and Hofmeister, 1988), recently discussed in detail by Dufresne et al. (2009). The formula for the
161 KK transform is:

$$162 \quad \theta(\nu_i) = \frac{2\nu_i}{\pi} \int_0^{\infty} \frac{\ln r(\nu) - \ln r(\nu_i)}{\nu_i^2 - \nu^2} d\nu \quad (3)$$

163 where θ is the phase shift, ν is the real portion of the frequency, ν_i is imaginary portion of the
164 frequency and r is the real part of the reflectivity. In this study the reflectance spectra were
165 transformed into KK absorbance (KK-Abs.) units by first smoothing the data points as a function
166 of wavenumber over a $\sim 40.5 \text{ cm}^{-1}$ window (like Moore et al., 2000) using a “running average
167 type” algorithm (Savitsky-Golay). A smoothed spectrum minimizes error because the subsequent
168 KK transform requires extrapolating the wings of the reflectance spectrum (Efimov 1995). The
169 KK transform was performed using the OMNIC software (version 6.1) and to further ensure
170 uniform extrapolation, we sampled over a constant range of $5300 - 650 \text{ cm}^{-1}$.

171 RESULTS

172 Figures 1c and 1d show representative KK-Abs. spectra and illustrate the shapes of the
173 major bands that are observed in this type of spectra (discussed above). Each spectrum was
174 examined in the $700\text{-}1300 \text{ cm}^{-1}$ region to verify that only a broad glass band was present, ruling
175 out a significant volume of crystals (with sharp bands at specific wavenumbers) or issues related
176 to internal scattering of light. The volatile bands were then studied in detail. In contrast to the
177 subtle features observed in R% spectra (Fig. 1a, 1b), the KK-Abs. volatile species bands (Fig. 1c,
178 1d) form positive peaks and the CO_3^{2-} species forms a doublet, like transmission IR spectra. Each
179 band increases in KK-Abs. intensity with increasing volatile contents (Tables 2 and 3).

180 The baseline under the CO_2_{mol} band is relatively linear, whereas the baselines under the
181 CO_3^{2-} doublet and $\text{H}_2\text{O}_{\text{mol}}$ bands are curved due to superposition on the large Si-O band. We fit
182 the latter baselines using French curves following the techniques discussed in King et al. (2004).
183 That work evaluated the error associated with different baseline correction approaches (e.g.,
184 fitting with Gaussian curves) and found that the French curve method allows for a conservative
185 estimate of the fitting error (<10 % relative). We assessed reproducibility by fitting maximum
186 and minimum baselines to the spectra and the variations were all within this conservative error
187 (<10 % relative). The baseline under the 3600 cm^{-1} peak was assumed to be linear between 3800
188 and $\sim 2800\text{ cm}^{-1}$, and we assume that the line has a 10% error (e.g., King et al., 2002).

189 **Total H₂O contents**

190 The KK-Abs. for the $\text{H}_2\text{O}_{\text{tot}}$ band measured from a linear baseline has good
191 reproducibility, with: $1\sigma < 5\%$ for 3-4 replicates on most samples and always < 20% (e.g.,
192 Tables 2 and 3). For each composition, the $\text{H}_2\text{O}_{\text{tot}}$ KK-Abs. increases linearly as a function of
193 $\text{H}_2\text{O}_{\text{tot}}$ wt.% (Fig. 2), following the relation:

$$194 \quad \text{H}_2\text{O}_{\text{tot}} \text{ wt.}\% = m * (3550\text{ cm}^{-1} \text{ band KK-Abs.}) \quad (4)$$

195 where m , the slope of the calibration line, is a function of composition. The following are
196 compositions/ m/r^2 of the fit: calc-alkaline andesite/154/0.93; alkali andesite/156/0.99;
197 phonolite/172/0.90; and rhyolite/204/0.99. Our calibration slope for phonolite was presented in
198 Mongrain et al. (2008) and our approach was used by Pauly et al. (2011), but the calibration data
199 were not shown in either case. We chose to force equation 4 through zero because there should
200 be no IR signal when $\text{H}_2\text{O}_{\text{tot}}$ is absent.

201 The calibration slope for the basaltic glasses ($m = 125$, $r^2 = 0.97$) is essentially defined by
202 the sample with the high water content and so we suggest that this is a preliminary value. Also,

203 the basaltic glasses with low water contents are all natural samples and may show more scatter
204 because they have greater variation in composition and may have more intrinsic variation in
205 water contents. Nonetheless, we are encouraged that Pauly et al. (2011) found a similar slope for
206 palagonite ($m = 123$) when comparing KK-Abs. values to electron microprobe deficits.

207 **Molecular H₂O, molecular CO₂ and carbonate contents**

208 The H₂O_{mol} and CO_{2 mol} bands are relatively easy to identify in both the reflectance and
209 KK-Abs. spectra because both form strong, narrow peaks (low degeneracy in the molecular
210 vibration). In contrast, the CO₃²⁻ doublet at these concentrations is not possible to identify in
211 reflectance spectra (Fig. 1a; cf. Grzechnik et al., 1996; Moore et al., 2000) although the KK-Abs.
212 spectra have a discernable CO₃²⁻ doublet (Fig. 1c).

213 For the calc-alkaline andesite glasses we found the following relations: H₂O_{mol} =
214 $435 \cdot (1620 \text{ cm}^{-1} \text{ KK-Abs.})$, $r^2 = 0.92$, CO_{2 mol} = $37.2 \cdot (2350 \text{ cm}^{-1} \text{ KK-Abs.})$, $r^2 = 0.91$; and CO₃²⁻
215 = $500 \cdot (1520 \text{ cm}^{-1} \text{ KK-Abs.})$, $r^2 = 0.74$. The data used to constrain these equations are shown in
216 Figures 3-5 and, again, we chose to force the fit through zero. The most difficult volatile to
217 quantify with our technique is CO₃²⁻ based on the large scatter about the calibration line (Fig. 5),
218 likely due to the nature of the CO₃²⁻ vibration and its resulting IR doublet band.

219 In sum, these results indicate that H₂O_{tot}, H₂O_{mol}, CO_{2 mol} and possibly CO₃²⁻ may be
220 quantified in aluminosilicate glasses using reflectance IR techniques. We anticipate that the
221 technique may be further developed to quantify C-O-H species in a range of aluminosilicate
222 compositions.

223

DISCUSSION

224 **Errors in regression analysis**

225 To examine the goodness of fit (precision) of the calibration line with an ordinary least
226 squares fit, we determined how the calibration line fit varied when using a robust regression,
227 employing the Isoplot program for Excel (Ludwig, 2003). The purpose of using the robust
228 regression is to examine whether outliers in the data might artificially influence the line fits and
229 to also estimate how the slope of the calibration lines may vary accounting for the scatter in the
230 data used for each composition. When forced through zero, the robust regression returns slopes
231 that are within 7 % of those determined using the ordinary least squares fit (Table 4). The errors
232 on the slope of the line determined using the robust regression method vary from 5 to 10 %
233 (Table 4). This is comparable with other methods, such as SIMS, where the data used for
234 calibrations has +/- 10% relative error. In contrast, errors on transmission FTIR and manometry
235 data can be significantly smaller in the case of high quality data where exceptional care is taken
236 with the analyses. However, the error on transmission FTIR analyses is commonly significant in
237 the case where thickness estimates are difficult, and thus the error on the analysis commonly
238 approaches 10% or more (King et al., 2004).

239 In order to examine the predictive nature (accuracy) of the linear regressions for
240 determining the compositions of unknown samples, we report the root mean square error
241 prediction (RMSEP). This value provides a way to estimate the accuracy error involved in using
242 the calibration line to predict the volatile content of an unknown sample. RMSEP values take
243 into account the measurement errors, random errors, prediction errors (residuals) and bias.
244 Values of RMSEP that are close to zero indicate smaller prediction errors for the linear
245 regression. The formula for RMSEP is:

246
$$\text{RMSEP} = \sqrt{\frac{1}{n} \sum_{i=1}^n (y_i - \hat{y}_i)^2}$$
 (5)

247 where, n is the number of measurements in the calibration set (including a sample with 0% of the
248 volatile), i refers to a particular observation, and \hat{y}_i is the predicted value derived from a linear
249 regression on the calibration set leaving out the measured value for that observation (i.e. a jack-
250 knife or segmented approach). For $\text{H}_2\text{O}_{\text{tot}}$, the RMSEP values given in Table 4 have a range of
251 0.10 wt.% (rhyolite) to 0.20 wt.% (alkali andesite) for the linear fits. For the calc-alkaline
252 glasses, very good prediction errors are observed for $\text{H}_2\text{O}_{\text{mol}}$ (RMSEP = 0.16 wt.%; Fig. 3) and
253 for CO_2_{mol} (RMSEP = 0.003 wt.%; Fig. 4) for linear fits. The CO_3^{2-} contents were so poorly
254 correlated with the 1520 cm^{-1} band height that we did consider it useful to calculate the RMSEP.
255 The RMSEP are larger for the robust linear regressions (Table 4), as would be expected since
256 those fits include errors due to outliers. The andesites, rhyolite and basalt glasses show a small
257 differences (up to 4 wt.%) between the RMSEP for the standard linear regression versus the
258 robust linear regression, which indicates that the accuracy is for these glasses not greatly affected
259 by outliers. The difference is greater for the phonolites (standard RMSEP = 0.19 wt.% versus
260 robust RMSEP = 0.38 wt.%), indicating that outliers do affect the phonolite calibration curves.
261 In general, the RMSEP indicate that this method has accuracy errors for samples with 1 wt.%
262 $\text{H}_2\text{O}_{\text{total}}$ that are less than 20% relative, with the relative accuracy improving at higher $\text{H}_2\text{O}_{\text{tot}}$
263 contents.

264 **A general model for micro-reflectance IR measurements of total H_2O**

265 A systematic increase in the calibration slope for $\text{H}_2\text{O}_{\text{tot}}$, m , with increasingly more mafic
266 compositions is observed in the linear least squares fit calibration lines (Fig. 2 and Table 4), as
267 well as in the robust regressions (Table 4). Such a systematic change suggests that there is a
268 compositional or optical control on the calibration slope, similar to that observed for extinction

269 coefficients determined by transmission IR spectroscopy (e.g., Dixon et al., 1995; Jakobsson,
270 1997; Mandeville et al., 2002). We observe a reasonable negative correlation between (Si + Al)
271 mole fraction and the calibration slope, m (Table 4). However, a better correlation is observed
272 between the calculated refractive index, n , and m (Table 4 and Fig. 6) for both types of fits. We
273 hypothesize that this correlation indicates that the optical properties of the glass are the major
274 control on the slopes of the calibration lines. This finding indicates significant promise in
275 developing a general reflectance IR calibration that covers a range of aluminosilicate glass
276 compositions.

277 **Analysis of glass wafers**

278 Figure 7a shows the type of spectrum observed when a calc-alkaline glass wafer (in this
279 case 90 μm) is placed on a Au-coated glass slide and analyzed in reflection mode (referred to
280 hereafter as "wafer reflectance"). Instead of a reflectance spectrum (Fig. 1a), the wafer
281 reflectance spectrum (Fig. 7a) has strong similarities to a transmission spectrum (Fig. 7b)
282 combined with a KK-Abs. spectrum (Fig. 1b). This effect is likely due to some of the light
283 transmitting through the sample, being reflected by the Au-coated glass slide substrate, and
284 returning to the detector in addition to some signal reflecting off the surface of the glass. Mixed
285 transmission-reflectance behavior is supported by the observation that both the transmission and
286 reflectance features are affected by thickness of the sample (Fig. 8). For example, in calc-
287 alkaline andesite samples that are $< \sim 100 \mu\text{m}$ both transmission and reflectance features are
288 observed whereas those that are thicker show reflectance features only (like Fig. 1a). However,
289 the thickness at which the behavior of the IR light through a wafer changes from mixed to
290 reflectance is expected to depend on refractive index (a function of composition). Previous
291 studies that only observed transmission features for rhyolites on a Au backing (Nowak and

292 Behrens, 1995) only examined the near-IR part of the spectrum where transmission features
293 dominate and therefore the reflectance features were not observed.

294 In the wafer reflectance spectrum, the KK-Abs.intensity is amplified: relative to the
295 transmission spectrum, the amount of absorbance amplification is a function of the wavenumber.
296 For example, absorbance for $\text{H}_2\text{O}_{\text{tot}}$ and CO_2_{mol} are amplified $\sim 2x$, whereas absorbance for CO_3^{2-}
297 is amplified $\sim 1.5x$ (Figs. 7a, 7b). We predict that the absorbance amplification is non-linear due
298 to variations in the optical constants of the glass at different frequencies (e.g. refractive index,
299 dielectric constant etc.). Therefore, glasses of different compositions likely have specific "wafer
300 reflectance calibration coefficients", $m_{\text{WR, species}}$, for each of the volatile species similar to
301 transmission IR spectroscopy. Until we have better constrained $m_{\text{WR, species}}$, it would be
302 advantageous to minimize the effect of transmission in a thin sample. For example, melt
303 inclusions are commonly $<100 \mu\text{m}$ and thin sections are $\sim 30 \mu\text{m}$; thus both types of samples
304 provide the impetus for developing this micro-reflectance IR technique for thinner glass wafers.

305 To this end, we experimented with roughening the back of andesite wafers that were less
306 than $100 \mu\text{m}$ thick. Figure 8a shows reflectance spectra for a calc-alkaline andesite glass wafer
307 (MHA31, initially a $60 \mu\text{m}$ thick wafer) roughened on the back to $1 \mu\text{m}$, $40 \mu\text{m}$ and $90 \mu\text{m}$. The
308 spectra are seen to become progressively less transmission-like and more reflectance-like with
309 increasing roughness. Unfortunately, the wafer was damaged when it was roughened with
310 $200 \mu\text{m}$ grit. However, Figure 8b shows a different andesitic glass wafer (MHA-44, initially a
311 $63 \mu\text{m}$ thick wafer) roughened with a $\sim 130 \mu\text{m}$ polishing grit and set in epoxy resulted in an
312 almost 'pure' reflectance spectrum. The reflectance spectrum from the roughened sample was
313 converted to KK-Abs. units, resulting in a spectrum that is very similar to a KK-Abs. spectrum
314 derived from a reflectance spectrum collected from a thick glass (Fig. 9). The total water content

315 calculated using these two KK-Abs. spectra was identical, with both spectra giving 4 wt.%
316 H₂O_{tot}. This value is within error of the measured H₂O_{tot} content of 3.34 wt.%.

317 In sum, glass wafers may be analyzed with reflectance techniques using samples where
318 the back has been roughened to ~130 μm and set in epoxy. However, glass wafer methods
319 deserve further development because spectra collected on such samples have the advantage of
320 higher absorbance signal (Fig. 7a) and would be useful for analysis of melt inclusion and thin
321 sections.

322 **Interlaboratory comparison, comparison of IR detectors and caveats**

323 To determine if our reflectance IR technique depends on the instrument set-up we
324 undertook two separate studies. First, measurements were made using different detector types
325 (MCT-A* versus MCT-A in King's lab); and second, we measured the phonolite glasses at the
326 University of Alaska (Larsen's lab). The MCT-A detector has a lower detectability rating ($D^* =$
327 4.5×10^{10}) than the MCT-A* ($D^* = 6.5 \times 10^{10}$) according to the manufacturer (ThermoFisher).
328 Larsen's lab uses a Nicolet 6700 FTIR (compared to King's Nicolet Nexus 670), but both used a
329 Global source, XT-KBr beamsplitter, and a Continuum microscope with a liquid nitrogen cooled
330 MCT-A detector.

331 Figure 10a shows that the KK-Abs. value for H₂O_{tot} using the MCT-A* detector is
332 consistently lower than that determined using the MCT-A detector (Fig.10a). This result is
333 surprising because the MCT-A* detector is designed with higher detectability, thus we cannot
334 rule out detector alignment issues. Nonetheless, the data show that such variables (detector,
335 alignment etc.) are important for accurate calibration curves.

336 Figure 10b shows the results of phonolite analyses in Larsen's lab using the spectrometer
337 configuration noted above. The reflectance peak heights from Larsen's lab are systematically

338 higher than those obtained in King's lab using an older instrument, although some results are
339 within error (Fig. 10b).

340 Together, the results of analyses with different machine configurations indicate that
341 volatile calibration curves depend on a particular instrument's configuration (e.g., light
342 throughput and detector efficiency). Therefore, to use these techniques it is necessary to create
343 instrument-specific calibration curves and to check these periodically. This is similar to other
344 micro-analytical techniques currently used for volatiles that also depend on instrument set-up and
345 efficiency, such as secondary ion mass spectrometry (SIMS) and Raman spectroscopy. One of
346 the advantages of the IR method relative to SIMS is that "blanks" as well as calibration standards
347 do not need to be measured at each analytical session. Furthermore, we have shown that the IR
348 calibrations for H_2O_{tot} are related to calculated refractive indices, which means that once
349 calibrations have been done in a laboratory using certain glass compositions that it is possible to
350 estimate the slope of the calibration line, m , to determine H_2O_{tot} in other compositions.

351 As implied above, the KK transform results are sensitive to the input parameters,
352 especially the wavelength range. Smoothing to window spacings that are slightly different results
353 in negligible changes to the calculated KK-Abs. spectrum. However, we find that spectra that are
354 measured over different wavelength ranges do not necessarily give the same KK-Abs. as our
355 technique measured over $5300 - 650 \text{ cm}^{-1}$ due to variations in extrapolating the wings of the
356 reflectance spectrum prior to KK transformation. Despite the potential *accuracy* errors associated
357 with the wavelength extrapolation, this study shows that two laboratories or different detectors
358 produce results with similar *precision* errors. Furthermore, the technique produces slopes for the
359 calibration lines for a range of glass compositions that may be logically explained through a

360 physical property of the glass: the refractive index. We recommend using the same spectral range
361 as our measurements when using our approach.

362 **Application of the micro-reflectance IR techniques to challenging samples**

363 We have been able to determine volatile contents in fragile experiments on bubble
364 nucleation (Mongrain et al., 2008) and H_2O_{tot} diffusion from wet basaltic to dry rhyolitic melt
365 (glass) (Lui, 2005). We have monitored H_2O_{tot} uptake during weathering of basaltic glasses that
366 cannot be polished (our unpublished data) and others have measured H_2O_{tot} contents of fragile
367 palagonites (Pauly et al., 2011). Also, we have measured the H_2O_{tot} content of fragile impact
368 melt glasses (Harris et al., 2007). The impact glasses had in excess of 20 wt.% H_2O_{tot} determined
369 by alternate methods (Harris et al., 2007) and would have been exceedingly difficult to prepare
370 as very thin wafers for micro-transmission IR analyses; thus, the reflectance method is
371 advantageous for unusually H-rich samples. We recommend that the μ -R-IR technique is best for
372 samples with high volatile contents. In such samples, the method allows for rapid determination
373 of the H_2O_{tot} and H_2O_{mol} contents.

374 Our findings are consistent with technique development using attenuated total reflection
375 (ATR; Lowenstern and Pitcher, submitted) and synchrotron reflectance IR techniques. Combined
376 with these techniques, IR analysis with a mapping stage with overlapping spectral collection
377 (several micron stepsize) should improve the areal resolution of micro-reflectance IR analyses
378 and potentially allow smaller volumes to be sampled than with transmission IR methods which
379 are commonly limited by the thickness of the sample. Spectra can be obtained on glassy areas
380 and checked for the presence of minerals by verifying that the 700-1300 cm^{-1} region only
381 contains a broad glass band.

382

383

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543

544 **Figure Captions**

545 **FIGURE 1:** (a) Reflectance spectrum for calc-alkaline andesite, MHA-23, with insets showing the
546 total H₂O (H₂O_{tot}), molecular CO₂ (CO_{2 mol}) and CO₃²⁻ bands. Note that it is difficult to
547 evaluate where a baseline should be drawn in the case of the C-O species bands. (b)
548 Reflectance spectrum for calc-alkaline andesite MHA-41 showing the sigmoidal
549 molecular H₂O (H₂O_{mol}) band at ~1620 cm⁻¹. Note that it is difficult to evaluate where a
550 baseline should be drawn in the case of this sigmoidal band. (c) KK-Abs. spectrum for
551 MHA-23 with insets showing H₂O_{tot}, CO_{2 mol} and CO₃²⁻ bands. (d) KK-Abs. spectrum for
552 MHA-41 showing the H₂O_{mol} band at ~1620 cm⁻¹. The volatile contents for both of these
553 samples are given in Table 2.

554 **FIGURE 2:** KK-Abs. band height for the ~3600 cm⁻¹ band versus total H₂O (H₂O_{tot}) wt.%
555 determined via manometry or secondary ion mass spectrometry for a range of glass
556 compositions: basalt, calc-alkaline andesite, alkali andesite, phonolite and rhyolite.

557 **FIGURE 3:** KK-Abs. band height for the ~1620 cm⁻¹ band versus molecular H₂O (H₂O_{mol}) wt.%
558 determined via transmission IR spectroscopy for calc-alkaline andesitic glasses (King et
559 al., 2002).

560 **FIGURE 4:** KK-Abs. band height for the ~2350 cm⁻¹ band versus molecular CO₂ (CO_{2 mol}) wt.%
561 determined via transmission IR spectroscopy for calc-alkaline andesitic glasses (King et
562 al., 2002).

563 **FIGURE 5:** KK-Abs. band height for ~1520 cm⁻¹ band versus CO₃²⁻ wt.% determined via
564 transmission IR spectroscopy for calc-alkaline andesitic glasses (King et al., 2002).

565 **FIGURE 6:** Slope of the calibration line (*m*) for total H₂O (H₂O_{tot}) concentration versus calculated
566 refractive index, *n* for the range of compositions studied. Data are shown for both a linear

567 least squares regression approach (no error bars for m) and also a robust regression
568 approach (with error bars for m). Note that a range of basaltic compositions were used in
569 the calibration and only an average is reported.

570 **FIGURE 7:** (a) Reflectance spectrum for a wafer of representative calc-alkaline andesite glass
571 (MHA-23) placed on a gold-coated glass slide. The numbers refer to the KK-Abs. of the
572 band. (b) Transmission spectrum in absorbance units for a representative calc-alkaline
573 andesite glass (MHA-23). Note that the KK-Abs. values are greater for the wafer in
574 reflectance mode than the wafer in transmission mode.

575 **FIGURE 8:** (a) Reflectance spectra of an initially 60 μm thick calc-alkaline andesite MHA-31
576 wafer that has been polished or roughened on its back to 1 μm , 40 μm and 90 μm . (b)
577 Reflectance spectra of an initially 63 μm thick calc-alkaline andesite MHA-44 wafer that
578 has been polished to 1 μm or roughened to 130 μm on its back and set in epoxy. Note that
579 with increased roughening the intensity decreases in the downward pointing transmission
580 bands (T) related to C-O-H species and the intensity increases in the reflectance band (R)
581 at $\sim 1000\text{ cm}^{-1}$ related to O-Si-O asymmetric stretching.

582 **FIGURE 9:** KK-Abs. spectra for two samples of calc-alkaline andesite MHA-44 wafer: (i) solid
583 spectrum is from a wafer (initially 63 μm thick) roughened to 130 μm on its back and set
584 in epoxy; and (ii) dashed spectrum is from a thick glass that was singly polished.

585 **FIGURE 10:** (a) KK-Abs. band height for the $\sim 3600\text{ cm}^{-1}$ band versus total H_2O ($\text{H}_2\text{O}_{\text{tot}}$).wt.% in
586 calc-alkaline andesites measured using either a MCT-A or MCT-A* detector. (b) KK-
587 Abs. band height for the $\sim 3600\text{ cm}^{-1}$ band versus total H_2O ($\text{H}_2\text{O}_{\text{tot}}$).wt.% in phonolites
588 measured in King's lab and Larsen's lab.

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Table 1: Compositions of glasses used for calibrations in King's lab.

	Basalt	Calc-alk. andesite	Alkaline andesite	Phonolite	Rhyolite
wt.%, normalized to 100%					
SiO₂	53.69 (2.69)	60.62	63.68	57.31	75.58 (1.52)
TiO₂	1.17 (0.95)	0.89	0.64	0.24	0.17 (0.13)
Al₂O₃	14.29 (2.17)	18.57	17.60	23.12	13.57 (0.83)
FeO	10.71 (2.77)	5.56	3.89	1.63	1.31 (0.39)
MgO	6.51 (3.00)	2.71	2.70	0.10	0.18 (0.12)
CaO	10.82 (2.18)	6.05	5.74	0.66	0.88 (0.43)
Na₂O	2.25 (0.58)	4.67	4.12	11.22	4.36 (0.09)
K₂O	0.56 (0.47)	0.92	1.64	5.72	3.93 (0.30)
Reference*	1	2	3	4	5

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594 *1. Galapagos basalt # 1652-5 (Perfit et al., 1983, 1999), Manus Basin basalts #15-4, 2249 and
595 18-3 (Sinton et al., 2003) and experimental basalt DL414 (Lui, 2005). 2. Mount Hood andesite
596 (King et al., 2002). 3. Mascota andesite (Moore and Carmichael, 1998). 4. Laacher See Tephra
597 phonolite (Larsen and Gardner, 2004). 5. PCD and M3M rhyolites (Westrich, 1987).

Table 2: Results of KK-Abs. measurements on calc-alkaline andesite glasses in King's lab.

Sample	Method for $\text{H}_2\text{O}_{\text{tot}}$	Volatile concentrations (wt.%)				KK-Abs.values ($\pm 10\%$ rel.unless noted)				
		$\text{H}_2\text{O}_{\text{tot}}$ $\pm 10\%$ rel.	$\text{H}_2\text{O}_{\text{mol}}$	$\text{CO}_{2\text{mol}}$	CO_3^{2-}	3600 cm^{-1} ($\text{H}_2\text{O}_{\text{tot}}$)	1620 cm^{-1} ($\text{H}_2\text{O}_{\text{mol}}$)	2350 cm^{-1} ($\text{CO}_{2\text{mol}}$)	1520 cm^{-1} (CO_3^{2-})	3600 cm^{-1} ($\text{H}_2\text{O}_{\text{tot}}$) with MCT-A*
MHA23	SIMS	0.76	nd	0.077 (0.002)	0.337 (0.005)					
23-1						0.0035		0.0019	0.00074	
23-2						0.0046		0.0024	0.00100	
23-3						0.0041		0.0019	0.00076	
23y						0.0031		na		
MHA26	SIMS	2.43	0.75 (0.07)	0.033 (0.002)	0.373 (0.028)					
26-1						0.013	0.00130	0.0012	0.0005	0.0084 (0.0002)
26-2						0.014	0.00137	0.0013	0.0006	
26-3						0.013	0.00150	0.0011	0.0007	
26y						0.013	0.00160			
MHA27	SIMS	0.85	nd	0.077 (0.001)	0.283 (0.016)					
27-1						0.0038		0.00200	0.0007	0.0012
27-2						0.0032		0.00187	0.0008	0.0013 (0.0003)
27-3						0.0043		0.00204	0.0004	
MHA30	Man.	2.48	1.09 (0.09)	0.004	0.076 (0.011)					
30-1						0.0140	0.00213	nd	0.0004	
30-2						0.0134	0.00180	nd	0.0002	
30-3						0.0144	0.00218	nd		
MHA31	Man.	1.09	0.36 (0.03)	0.009 (0.001)	0.061 (0.032)					
31-1						0.0076	0.00047	nd	0.0001	
31-2						0.0079	0.00050	0.00024	0.0002	
31-3						0.0073	0.00044	0.00044	0.0001	
MHA41	Man.	2.69	1.34 (0.06)	0.014 (0.002)	0.323 (0.048)					
41-1	Man.					0.019	0.00271		0.0004	
41-2	Man.					0.020	0.00327		0.0005	
41-3						0.020	0.00300			
MHA44	Man.	3.34	2.02 (0.19)	0.014 (0.002)	0.483 (0.040)					
44-1	Man.					0.0240	0.0051		0.0012	0.0188 (0.0006)
44-2	Man.					0.0230	0.0047	0.00037	0.0013	0.0156 (0.0002)
44-3						0.0245	0.0051	0.00034	0.0012	0.0175 (0.0003)
J12	SIMS	3.39	2.20 (0.28)	0.009 (0.001)	0.575 (0.054)					
J12-1	SIMS					0.0251		na	0.0012	0.0165 (0.0017)
J12-2	SIMS									0.0169 (0.0017)
J12-3										0.0175 (0.0018)

Table 3: Results of KK-Abs. measurements on four glass compositions.

Sample	Analysis #	H ₂ O _{tot} method	H ₂ O _{tot} wt.% (± 10% rel. unless noted)	KK-Abs.values (± 10% rel. unless noted)	
				King's lab 3600 cm ⁻¹ (H ₂ O _{tot})	Larsen's lab 3600 cm ⁻¹ (H ₂ O _{tot})
Alkali andesites					
MA512	MA512-1	Manometry	0 (0.01)	0	
	MA512-2	Manometry	0 (0.01)	0	
M12-1	M12-1-1	Manometry	2.62 (0.01)	0.0163	
	M12-1-2	Manometry	2.62 (0.01)	0.0161	
	M12-1-3	Manometry	2.62 (0.01)	0.0173	
M12-2	M12-2-2	Manometry	5.03 (0.01)	0.0298	
	M12-2-3	Manometry	5.03 (0.01)	0.0312	
	M12-2-4	Manometry	5.03 (0.01)	0.0323	
M12-4	M12-4-1	Manometry	6.76 (0.01)	0.0450	
	M12-4-2	Manometry	6.76 (0.01)	0.0453	
Basalts					
ManB_2249	Man2249-1	Trans IR	1.19	0.0103	
	Man2249-2	Trans IR	1.19	0.0101	
ManB15-4	Man15-4-1	Trans IR	1.32	0.0134	
	Man15-4-2	Trans IR	1.32	0.0129	
Gal1652	Gal1652-5a	Trans IR	1.38	0.0123	
ManB-18-3	Man18-3-1	Trans IR	1.44	0.0127	
	Man18-3-2	Trans IR	1.44	0.0126	
	Man18-11-1	Trans IR	1.49	0.0154	
	Man18-11-2	Trans IR	1.49	0.0119	
DL0414	DL0414_r1	Trans IR	4.63	0.0368	
	DL0414_r2	Trans IR	4.63	0.0357	
	DL0414_r4	Trans IR	4.63	0.0354	
Phonolites					
LPG1	LPG1	Trans IR	4.68 (0.50)	0.0273	
	LPG1-2	Trans IR	4.68 (0.50)	0.0262	
	LPG1-3	Trans IR	4.68 (0.50)	0.0267	
LPG1a	LPG1a_1	Trans IR	4.13 (0.23)	0.0222	
	LPG1a_2	Trans IR	4.13 (0.23)	0.0223	
	LPG1a_3	Trans IR	4.13 (0.23)	0.0205	
	LPG1a_4	Trans IR	4.13 (0.23)	0.023	
LPG3	LPG3-1	Trans IR	4.55 (0.08)		
	LPG3-1	Trans IR	4.55 (0.08)		
	LPG3-2	Trans IR	4.55 (0.08)		
	LPG3-3	Trans IR	4.55 (0.08)		
LHAE2	LHAE2-1	Trans IR	2.93 (0.15)	0.0181 (0.0007)	
	LHAE2-2	Trans IR	2.93 (0.15)		
	LHAE2-3	Trans IR	2.93 (0.15)	0.0176 (0.0009)	
	LHAE2-4	Trans IR	2.93 (0.15)	0.0170 (0.0013)	
LHAE3	LHAE3-1	Trans IR	4.99 (0.72)	0.0279 (0.0001)	0.0330
	LHAE3-2	Trans IR	4.99 (0.72)	0.0276 (0.0002)	
	LHAE3-3	Trans IR	4.99 (0.72)	0.0282 (0.0003)	
LHAE4	LHAE4-1	Trans IR	4.25 (0.25)	0.0248 (0.0005)	0.0268
	LHAE4-2	Trans IR	4.25 (0.25)	0.0250 (0.0001)	
	LHAE4-3	Trans IR	4.25 (0.25)	0.0249 (0.0003)	
LHAE5	LHAE5-1	Trans IR		0.0290 (0.0003)	0.0312
	LHAE5-2	Trans IR		0.0287 (0.0001)	0.0329
	LHAE5-3	Trans IR		0.0282 (0.0001)	
LPG4	LPG4-1	Trans IR	3.77 (0.05)	0.0260 (0.0001)	0.0302
	LPG4-2	Trans IR	3.77 (0.05)	0.0277 (0.0004)	
	LPG4-3	Trans IR	3.77 (0.05)	0.0275 (0.0001)	
LPG6	LPG6-1	Trans IR	3.39 (0.10)	0.0177 (0.0005)	0.0251
	LPG6-2	Trans IR	3.39 (0.10)	0.0168 (0.0007)	0.0270
	LPG6-3	Trans IR	3.39 (0.10)	0.0194 (0.0006)	0.0266
	LPG6-4	Trans IR	3.39 (0.10)	0.0199 (0.0004)	
LSN1	LSN1-1	Trans IR		0.0294 (0.0005)	0.0430
	LSN1-2	Trans IR		0.0302 (0.0001)	0.0396
	LSN1-3	Trans IR		0.0299 (0.0006)	

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Table 3 (ctd.)

Sample	Analysis #	H ₂ O _{tot} method	H ₂ O _{tot} wt.%	KK-Abs.values (± 10% rel. unless noted)	
				King's lab 3600 cm ⁻¹ (H ₂ O _{tot})	Larsen's lab 3600 cm ⁻¹ (H ₂ O _{tot})
Phonolites (ctd.)					
LST3-40-1	LST3-40-1-1	Trans IR	0 (0.10)	0.0035 (0.0004)	0.0045
	LST3-40-1-2	Trans IR	0 (0.10)	0.0030 (0.0001)	
	LST3-40-1-3	Trans IR	0 (0.10)	0.0030 (0.0002)	
Rhyolites					
UTR2	UTR2	Manometry	0.09 (0.04)	0.0012	
PCD	PCDA	Manometry	0.13 (0.04)	0.0013	
	PCDB	Manometry	0.17 (0.04)	0.0015	
2N	2N	Manometry	2.07 (0.29)	0.0097	
	2N_2	Manometry	2.07 (0.29)	0.0103	
	2N_3	Manometry	2.07 (0.29)	0.0089	
M3N	M3N	Manometry	2.94 (0.16)	0.0142	
	M3N_3	Manometry	2.94 (0.16)	0.0145	
	M3N_4	Manometry	2.94 (0.16)	0.0141	
M6N	M6N	Manometry	5.11 (0.38)	0.0239	
	M6N_3	Manometry	5.11 (0.38)	0.0257	
	M6N_4	Manometry	5.11 (0.38)	0.0259	

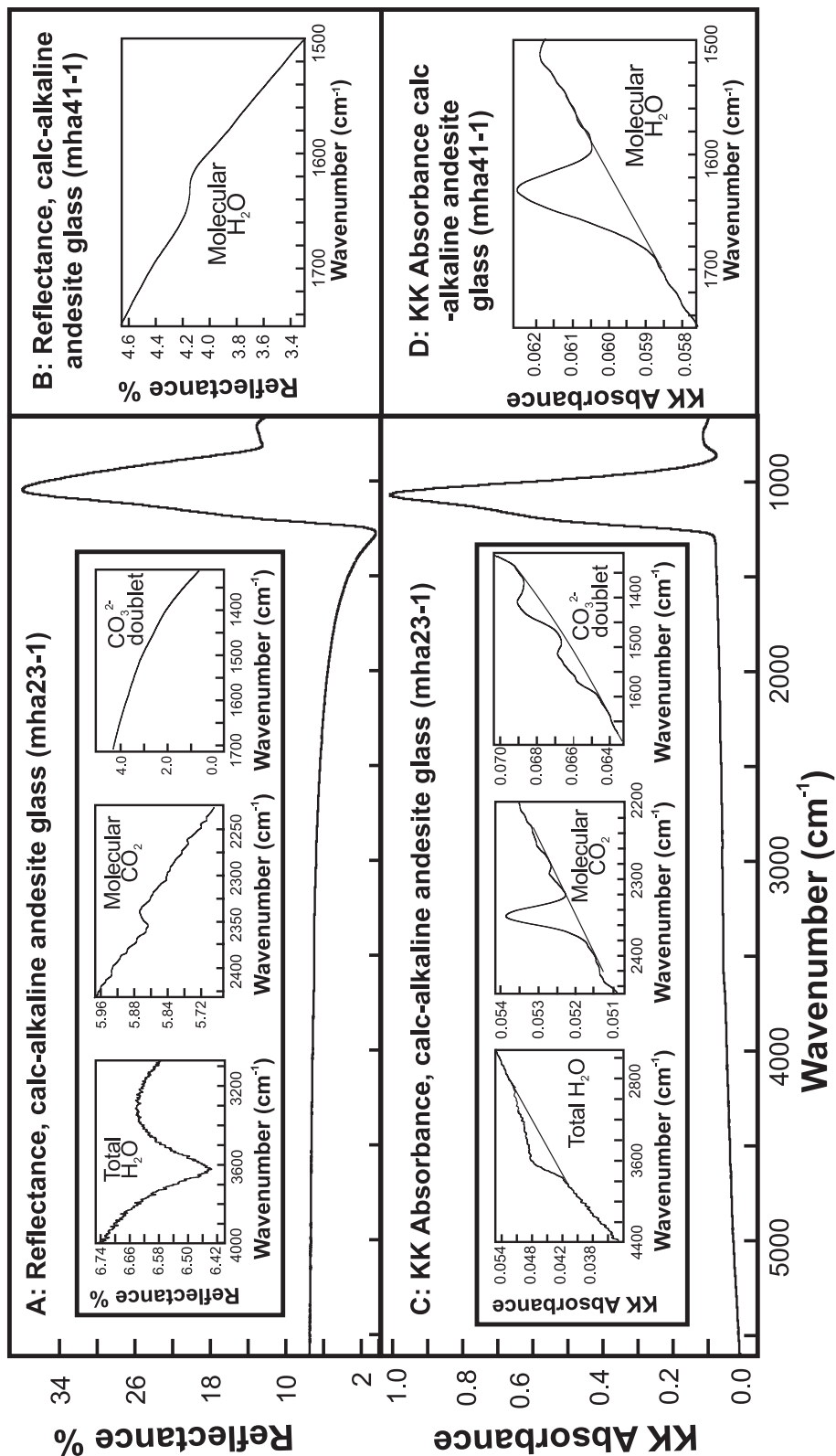
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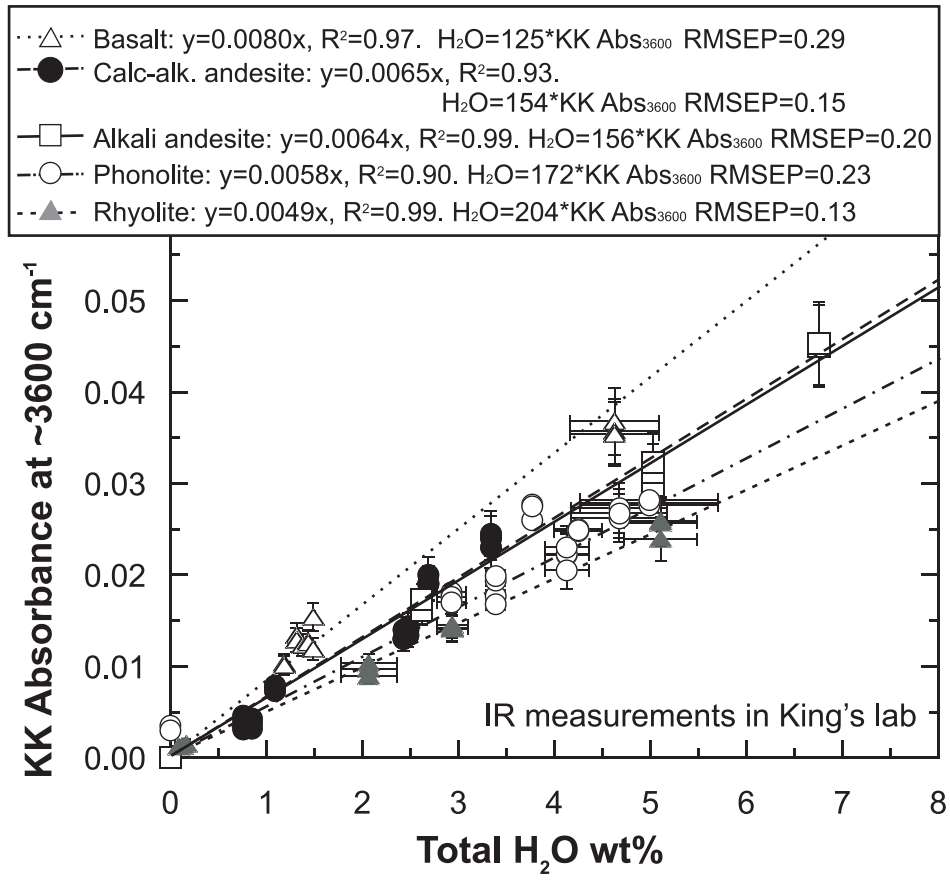
603 **Table 4: Molar Si+Al, calculated refractive indices, fit parameters and root mean square**
604 **error of prediction for H₂O_{tot} calibrations for five different glass compositions measured in**
605 **King's lab.**

	Basalt	Calc-alk. andesite	Alkaline andesite	Phonolite	Rhyolite
Molar Si+Al	0.66 (0.04)	0.76	0.78	0.73	0.85 (0.00)
Refr. index	1.590 (0.013)	1.544	1.533	1.506	1.491 (0.005)
Least squares fit results for H₂O_{tot}					
Slope	0.0080	0.0065	0.0064	0.0058	0.0049
<i>m</i>	125	154	156	172	204
RMSEP	0.12	0.14	0.20	0.19	0.10
Robust regression fit results for H₂O_{tot}					
Slope	0.00795 (+0.00080 -0.00071)	0.00697 (+0.00034 -0.00150)	0.00627 (+0.00044 -0.00034)	0.00522 (+0.00043 -0.00033)	0.00480 (+0.00018 -0.00024)
<i>m</i>	126 (+12/-12)	143 (+39/-7)	159 (+9/-10)	192 (+13/-15)	208 (+11/-8)
RMSEP	0.18	0.18	0.23	0.38	0.13

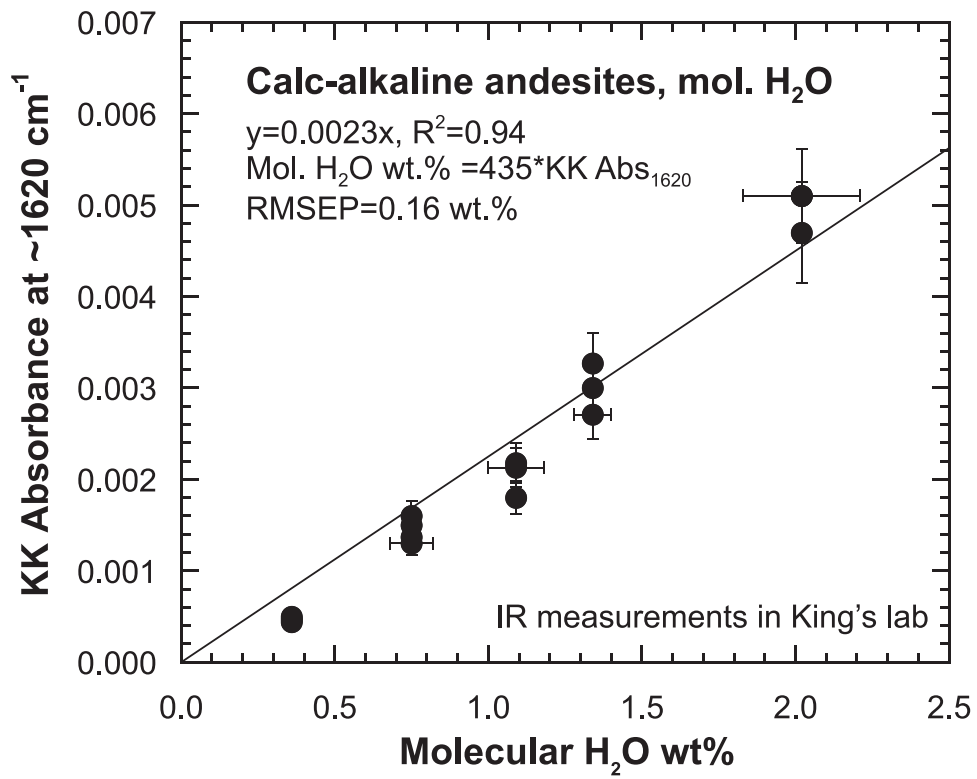
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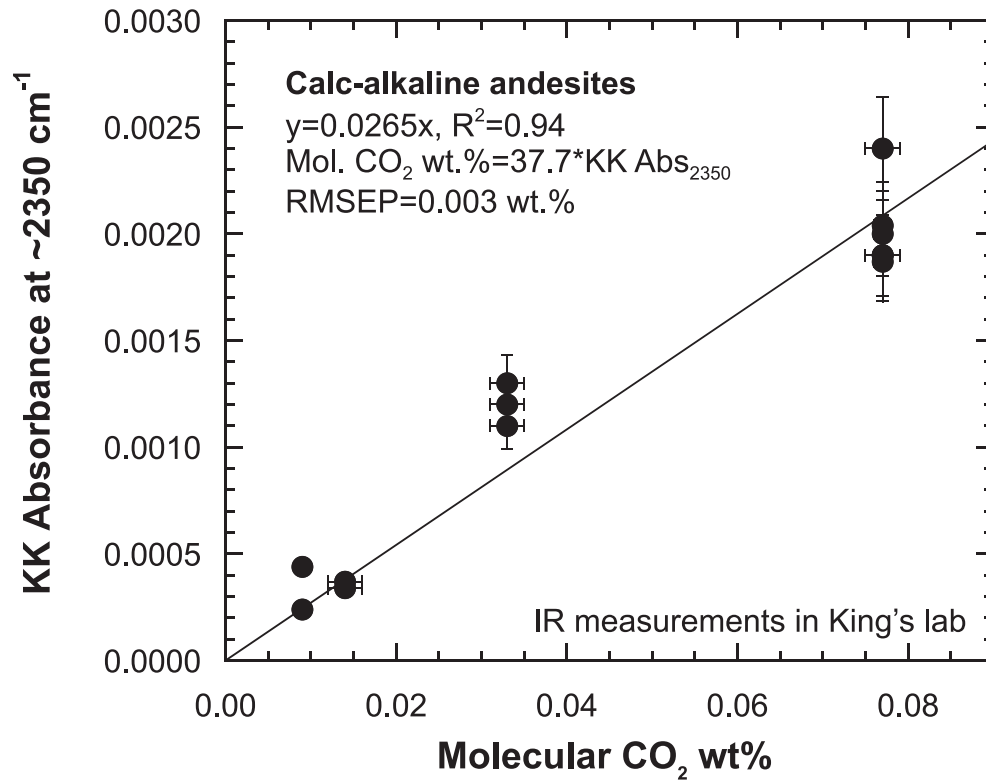
King & Larsen - FIGURE 1



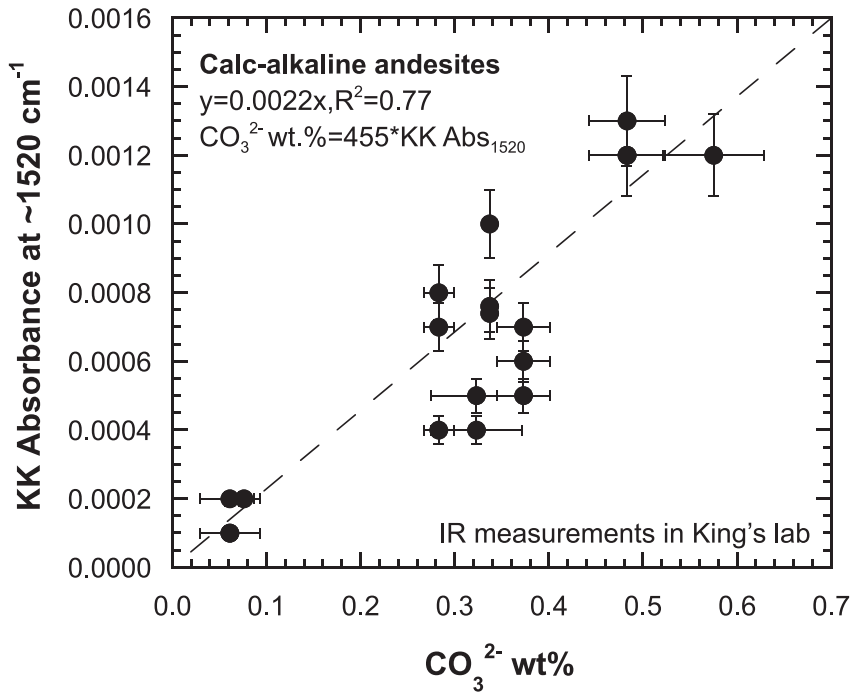
King & Larsen - FIGURE 2



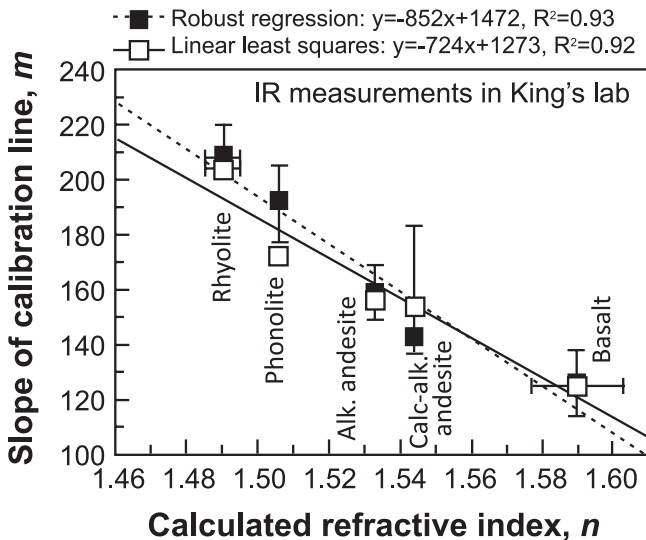
King & Larsen - FIGURE 3



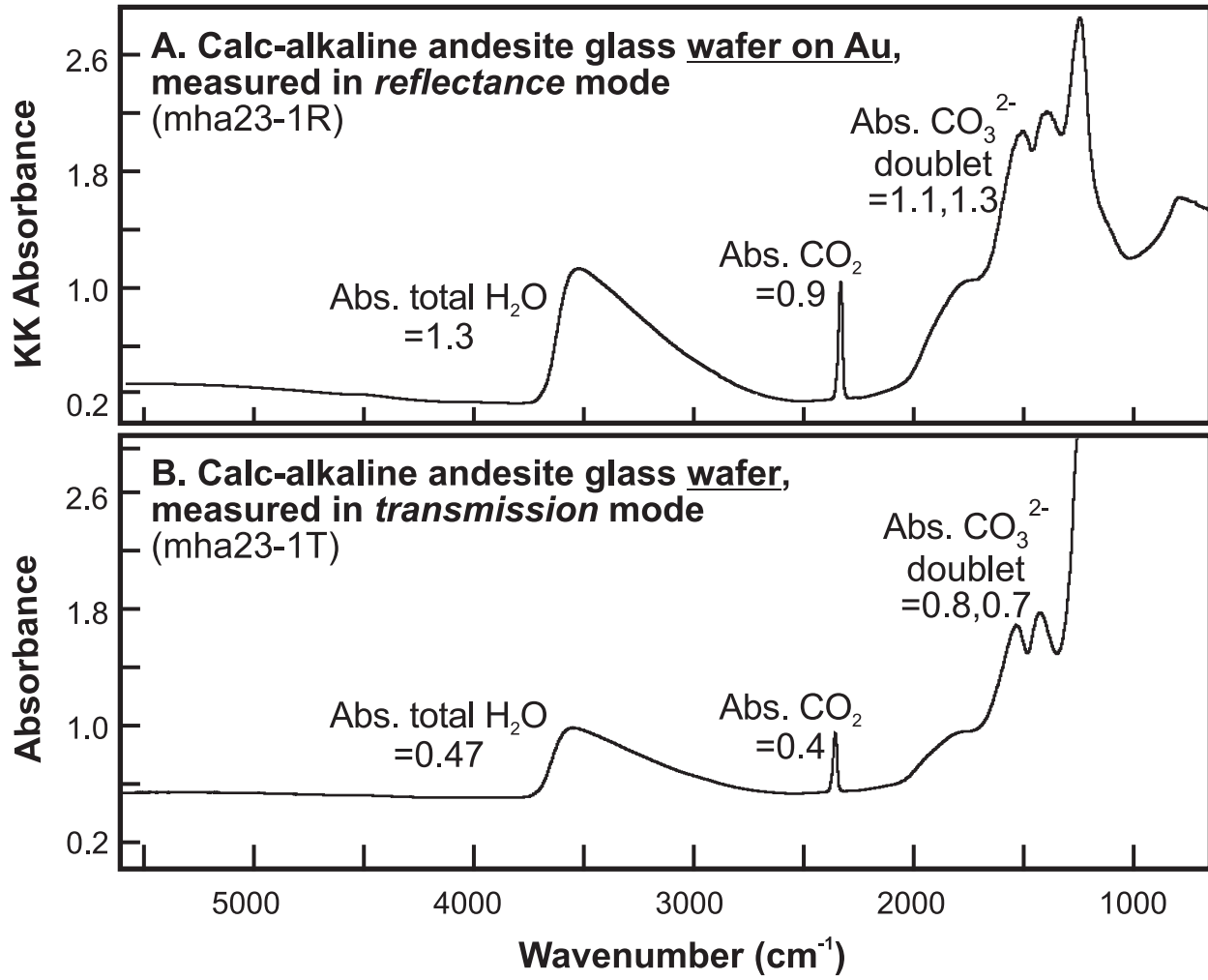
King & Larsen - FIGURE 4



King & Larsen - FIGURE 5

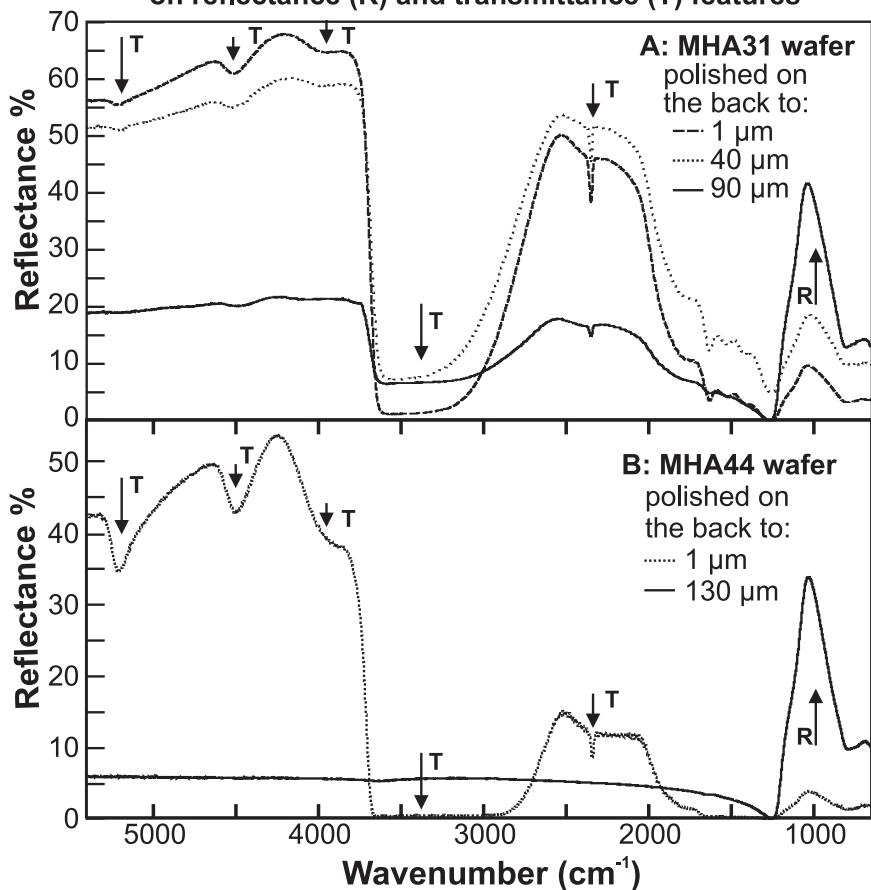


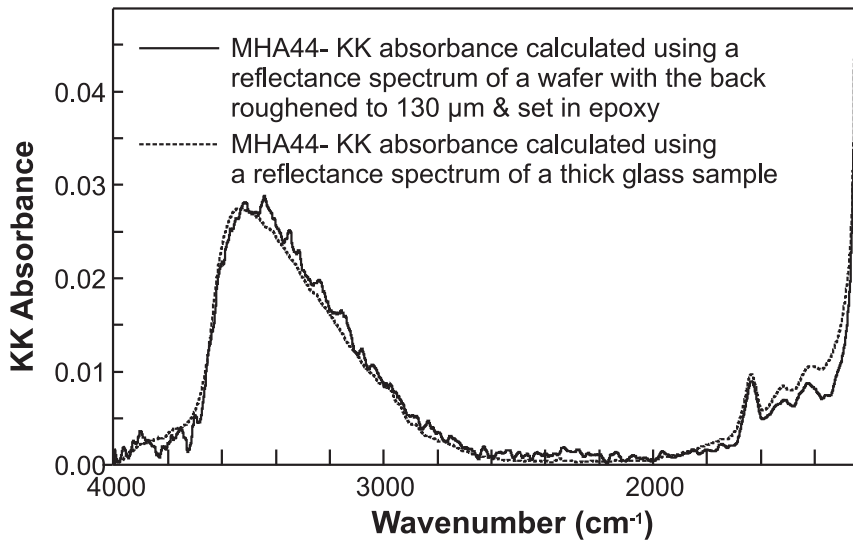
King & Larsen - FIGURE 6



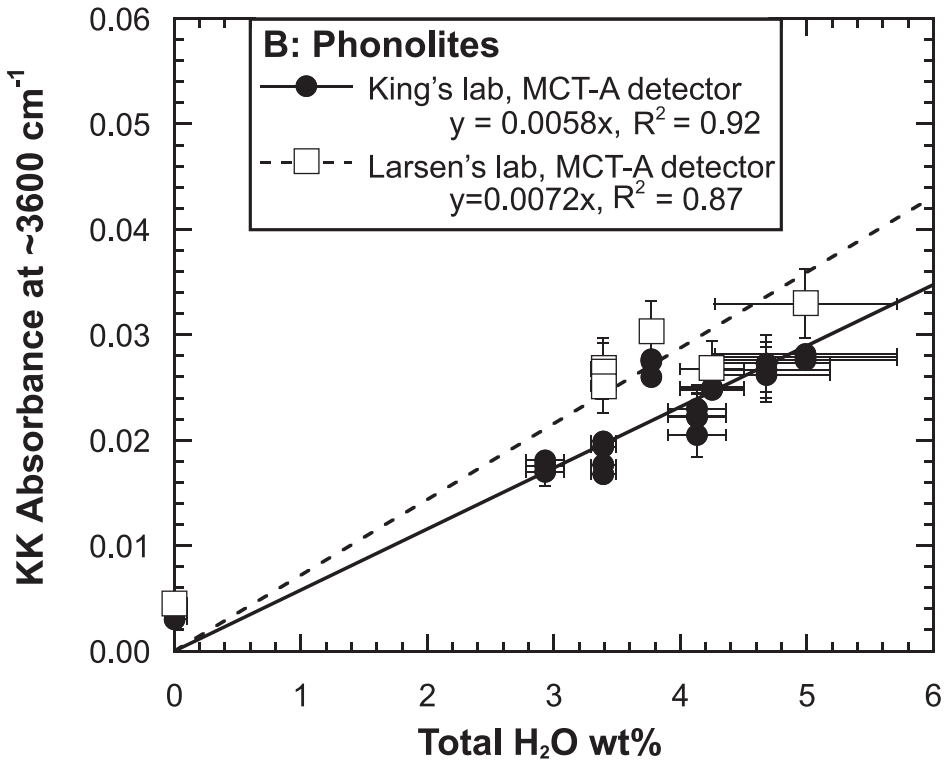
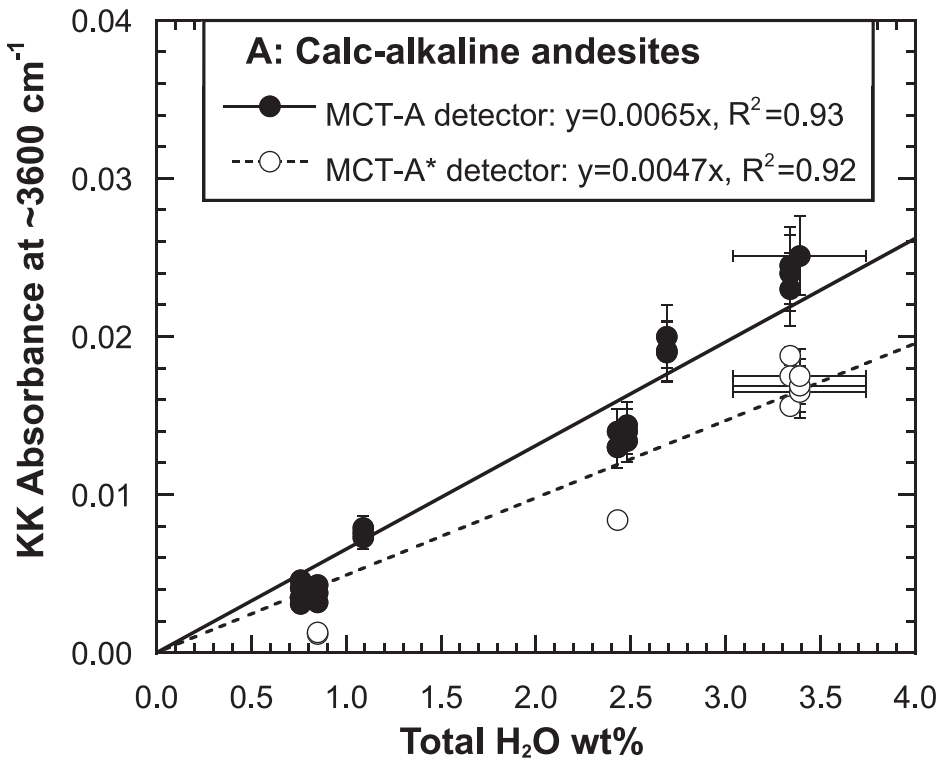
King & Larsen - FIGURE 7

Effect of back polishing with different grits on reflectance (R) and transmittance (T) features





King & Larsen - FIGURE 9



King & Larsen - FIGURE 10